Evolution and applications of the RD51 VMM3a/SRS gaseous beam telescope

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Gas Detectors Development Group

Outline

1. Overview

- RD51 test beam campaigns and triple-GEM beam telescope
- RD51 Scalable Readout System (SRS)

2. Telescope's performance

- Detector characterisation in energy, space and time simultaneously
- Rate-capability

3. Detectors under test

Micro-resistive WELL (µRWELL)

4. Operation at NA61/SHINE

- Combining self-triggered and externally triggered readout
- Offline event-matching







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Test beam campaigns of RD51

- RD51: CERN-based R&D collaboration on the development of Micro-Pattern Gaseous Detectors (MPGDs)
- Test beam campaigns for detector tests: H4 beam line @ CERN's SPS (PPE 134)
- Infrastructure contains (amongst other things)
 two beam telescopes
 - \rightarrow Gas Electron Multiplier (GEM)-based
 - \rightarrow MicroMegas-based
 - \rightarrow Strip or pad readout anode
 - \rightarrow Read out with RD51 Scalable Readout System (SRS)



Test beam campaigns of RD51: Triple-GEM telescope



COMPASS-like triple-GEM detectors [1] for tracking \rightarrow filled with Ar/CO₂ (70/30 %) \rightarrow strips with 400 µm pitch

10 x 10 cm² active area (facing the beam)

RD51 SRS electronics with
 VMM3a front-end ASIC [2]
 More than 2k channels for DUTs

Readout electronics: RD51 Scalable Readout System (SRS)

 Common RD51 readout system for small R&D setups up to mid-sized experiments [1,2]

 \rightarrow Ideal for readout of beam telescope with multiple detectors

- Front-end ASICs available for beam telescopes
 - \rightarrow Since 2009: **APV25**
 - \rightarrow Since 2021: VMM3a
- Front-end ASIC plugged onto the detector via hybrid-PCB
- ASIC-specific adapter card with multiple hybrids
- Common Front-End Concentrator (FEC)
- Multiple FEC/adapter cards for larger systems



Readout electronics: ATLAS/BNL VMM3a front-end ASIC

- Specifically developed by BNL for multi-channel readout of gaseous detectors (ATLAS New Small Wheel) [1]
- Self-triggered continuous readout in SRS implementation
- High rate capability: **9 Mhits/s per VMM** in SRS implementation
- Provides only peak amplitude (10-bit charge ADC)
- Provides time of the peak with O(ns) time resolution (12+8-bit timing)
- Good for R & D applications
 - Adjustable peaking times
 - Adjustable electronics gains
 - Wide range of input capacitances (< 200 pF up to 2 nF)
- Full detector characterisation with charge, space and time information with same front-end electronics



Signal peak amplitude / ADC Values

Beam telescope's performance: Detector-based studies



Time residuals / ns

[1] <u>vmm-sdat</u>

Cathode

GFM1

GFM2

-HV

3 mm Drift

2 mm Transfer 1

2 mm Transfer 2

Beam telescope's performance: Track-based studies

- Position determination: Centre-of-gravity (COG), as well as alternative approaches [1,2]
- Event-building based on cluster time
- Tracking with Kalman filter via anamicom [3]



Efficiency not @ 100%

due to geometrical effects and

Beam telescope's performance: Rate-capability



- 80 GeV/c pion beam: particle flux from ~7 x 10⁴ particles per spill (~5 s) to 10⁷ particles per spill
- Bandwidth saturation with ~5 x 10⁶ particles per spill and more
- Limiting factors: SRS FEC's Gigabit Ethernet and 9 Mhits/s readout limit of VMM3a
 - \rightarrow Loss in number of recorded interactions
 - \rightarrow Decrease of quality of acquired data, as described in [1]

Detectors under test



Studies on the micro-resistive WELL (µRWELL)

- Attractive because of
 - \rightarrow low material budget
 - \rightarrow simple manufacturing process
- GEM-foil glued on readout PCB
- Third RD51 beam telescope to be built, based on µRWELL
- 10 x 10 cm²
- 256+256 x-y-strips, 400 µm pitch
- 3 mm drift gap
- Filled with Ar/CO₂ (70/30 %)
- Resistive layer with 40 M Ω/\Box



Not drawn to scale

[Courtesy of Djunes Janssens]

Studies on the micro-resistive WELL (µRWELL)

- µRWELL from Yi Zhou and Xu Wang from USTC [1]
- Optimised for equal charge sharing between top and bottom strips (X and Y strips)



• Simultaneous detector characterisation in energy, space and time

[1] Yi Zhou et al. (RD51 Coll. Meet.)

The NA61/SHINE experiment

- SPS Heavy Ion and Neutrino Experiment (SHINE)
 - Strong interactions (heavy ion collisions):
 - \rightarrow Phases of strongly interacting matter
 - \rightarrow Onset of deconfinement
 - Interactions of cosmic rays in the interstellar medium (light ion collisions): Pierre Auger Observatory, AMS, ...
 - Study of **target interactions** for neutrino experiments: J-PARC (**T2K**), Fermilab (**DUNE @ LBNF**), ...
- Here: T2K run (T2K replica target: ~1 m carbon rod)
 → Reduce uncertainties in neutrino oscillation experiments



[Courtesy of Marek Gazdzicki]



Self-triggered + externally triggered

- Ad-hoc interim solution for beam tracking in front of interaction target
- Five weeks non-stop operation without failure of detectors or VMM3a/SRS electronics!
- 31 GeV/c protons @ up to 50 x 10³ particles per SPS spill
- Challenge:
 - Beam telescope: self-triggered
 - NA61: externally triggered @ ~ 1 kHz
 - \rightarrow matching NA61 events with VMM3a/SRS tracks
- **Solution:** inject event-ID from NA61 trigger into the VMM3a/SRS data stream
- Split event-ID-bits on VMM readout channels
- Match tracks and events in the offline analysis







X-position / cm

Event matching: Position and alignment

- Positions needed in NA61/SHINE coordinate system
- Alignment run with external reference
- No target + **Delay Wire** Chambers (DWC) of beam instrumentation



[Courtesy of Brant Rumberger]



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Counts

150

Summary and outlook

- The GEM-based beam telescope of RD51 with VMM3a/SRS readout allows studying time resolution, position resolution and energy behaviour simultaneously
- Particle beams with up to 1 MHz interaction rate can be recorded
- Various detector technologies (not limited to MPGDs) can be read out
- **µRWELL** studied in view of **building a third telescope**
- Successfully operated as part of the NA61/SHINE experiment
- Currently **full self-triggered:** implement **externally triggered mode** to achieve lower thresholds
- Started efforts on distributed system: increase telescope's lever arm from 1 m to 40 m
- Integration into other experiments started (e.g. with colleagues from NA64 @ CERN or P2 @ Saclay and Mainz)

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Back-up slides

Improving the spatial resolution

- Position determination: Centre-of-gravity (COG) ۲
- With X-rays [1]: improvement of position • reconstruction for imaging applications by
 - Modification of COG (**Q² weighting**) •

- Review of COG systematics and modifications $x = \frac{\sum_{i} Q_{i}^{n} x_{i}}{\sum_{i} Q_{i}^{n}} \quad n = 2$ (MPGD2022): Igor Smirnov: Algebraic methods for reconstruction of coordinates in strip detectors
- VMM3a: neighbouring-logic to recover charge • below THL

Effective gain / 10³

Scanned: THL range from 1.5 fC to 5.5 fC

- **Q²** improves spatial resolution all the time
- NL only at low signal-to-threshold ratio

Modulation of readout pattern

Modulation of readout pattern

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Improving spatial resolution

Triple-GEM detector (256+256 x-y-strips, 400 µm pitch)

μRWELL (256+256 x-y-strips, 400 μm pitch)

