

PROJECT 8



Massachusetts
Institute of
Technology

FUTURE EXPERIMENTS FOR NEUTRINOS MASS DETERMINATION

ν Phys 2023, London

Wouter Van De Pontseele

December 20, 2023

wvdp@mit.edu

Massachusetts Institute of Technology

THE PREDICTION OF THE NEUTRINO

[PAULI, 1930]

Physikalisches Institut
der Eidg. Technischen Hochschule
Zürich

Zürich, 4. Des. 1930
Gloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich huldvollst anzuhören bitte, Ihnen des näheren auseinandersetzen wird, bin ich angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie des kontinuierlichen beta-Spektrums auf einen verzweifelten Ausweg verfallen um dem "Wechselsatz" (1) der Statistik und dem Energiesatz zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale Teilchen, die ich Neutronen nennen will, in den Kernen existieren, welche den Spin 1/2 haben und das Ausschließungsprinzip befolgen und sich von Lichtquanten außerdem noch dadurch unterscheiden, dass sie nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen müsste von derselben Grössenordnung wie die Elektronenmasse sein und jedenfalls nicht grösser als 0,01 Protonenmasse... Das kontinuierliche beta-Spektrum wäre dann verständlich unter der Annahme, dass beim beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert wird, derart, dass die Summe der Energien von Neutron und Elektron konstant ist.

- β -spectrum seems to violate energy conservation.
- Propose a new particle: the neutrino
- Must be neutral and have a super tiny mass ... or no mass at all?

Pauli & Fermi speak at Solvay 1933, and a year later, Fermi writes down a theory about β -decay

- Rejected by Nature
- Neutrino mass will affect the shape of the electron energy spectrum near the endpoint!
- Independently suggested by Perrin in 1933

TENTATIVO DI UNA TEORIA DEI RAGGI β

Nota (1) di ENRICO FERMI

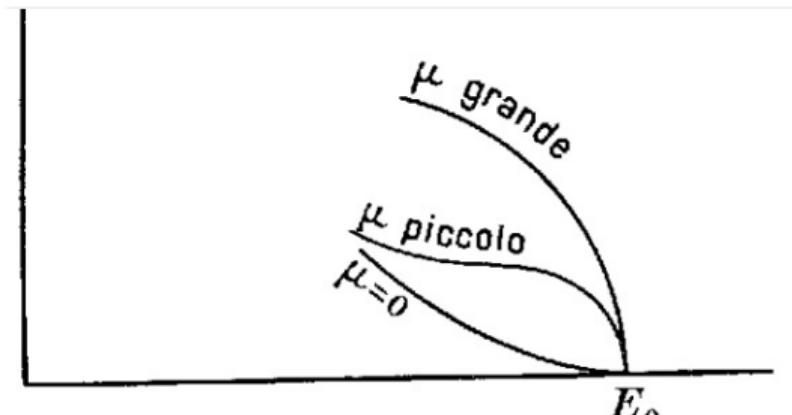


Fig. 1

Pauli & Fermi speak at Solvay 1933, and a year later, Fermi writes down a theory about β -decay

- Rejected by Nature
 - Neutrino mass will affect the shape of the electron energy spectrum near the endpoint!
 - Independently suggested by Perrin in 1933
- ⇒ Does not require detecting neutrinos!

TENTATIVO DI UNA TEORIA DEI RAGGI β

Nota (1) di ENRICO FERMI

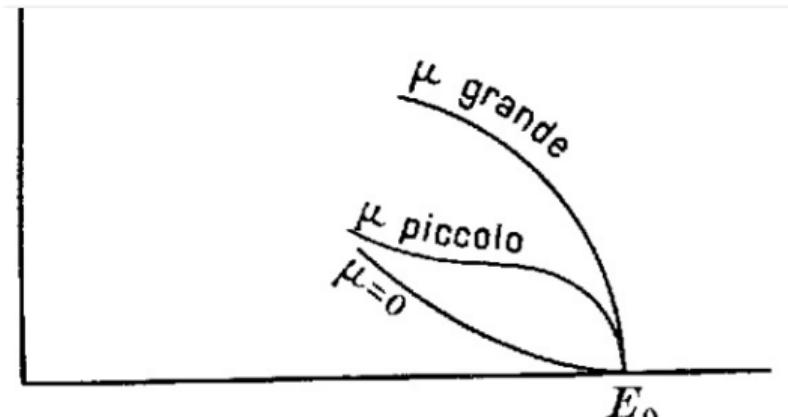
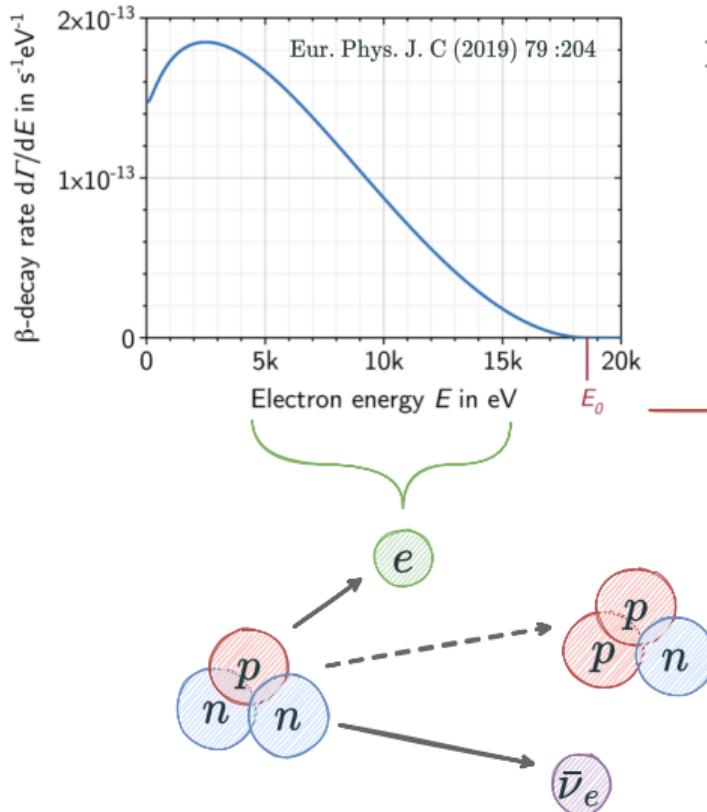


Fig. 1

NEUTRINO MASS MEASUREMENTS: TRITIUM β -DECAY

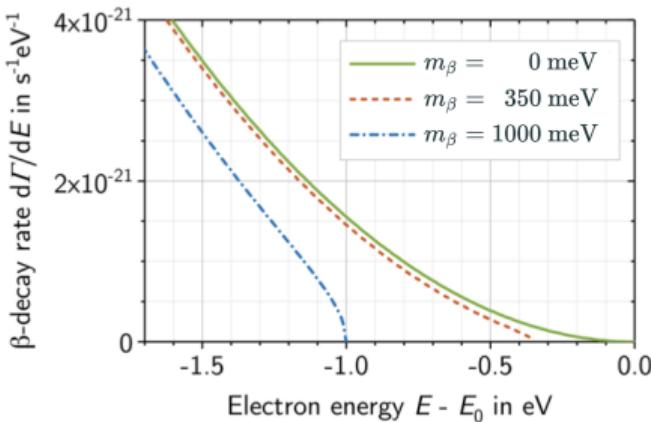


Electron energy distribution from tritium β -decay:

$$\frac{dN}{dE} \sim (E_0 - E) \sqrt{(E_0 - E)^2 - m_\beta^2}$$

$$m_\beta = \sqrt{\sum_{i=1}^3 |U_{ei}^2| m_i^2}$$

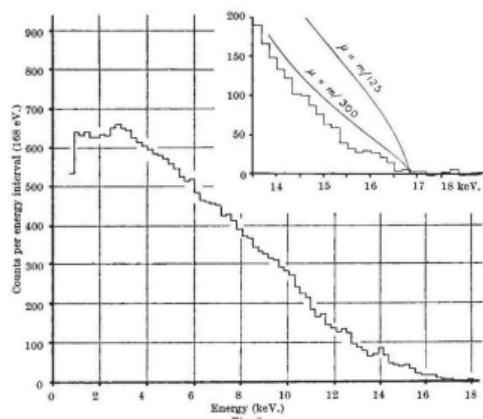
zoom in around endpoint E_0



Beta Spectrum of Tritium

S. C. CURRAN, J. ANGUS & A. L. COCKCROFT

Nature 162, 302–303 (1948)



© 1948 Nature Publishing Group

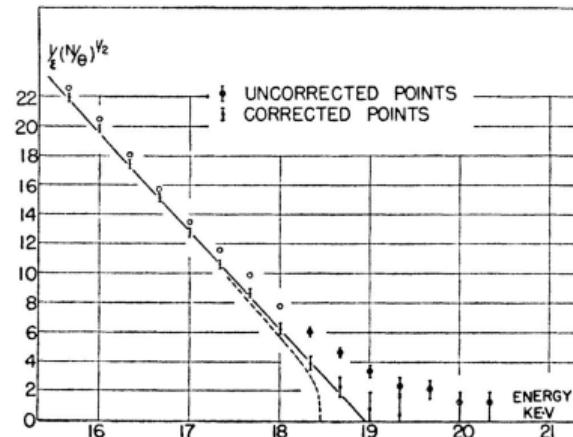
$$m_\beta \lesssim 1700 \text{ eV}$$

The β -Spectrum of H^3

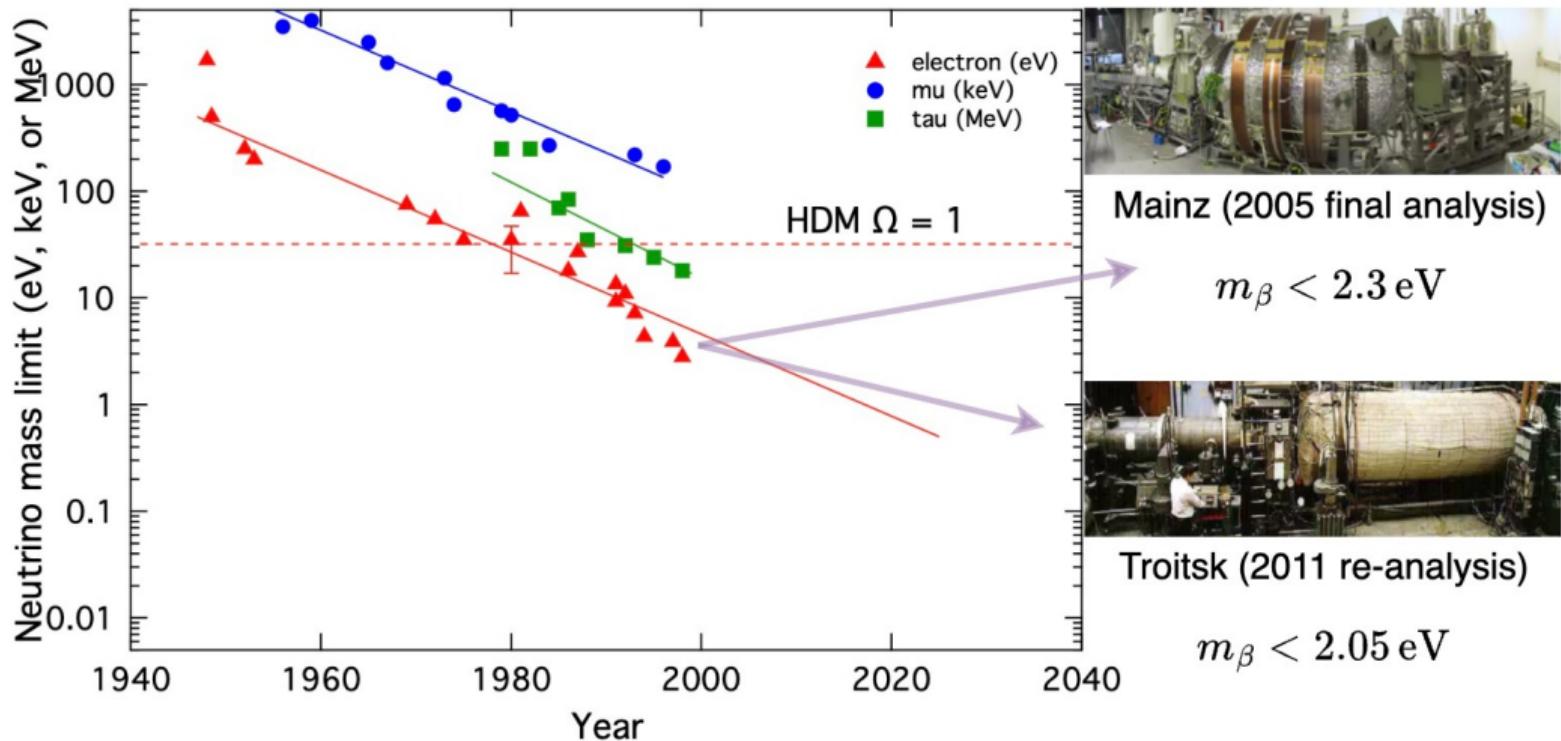
G. C. HANNA AND B. PONTECORVO

*Chalk River Laboratory, National Research Council of Canada,
Chalk River, Ontario, Canada*

January 28, 1949

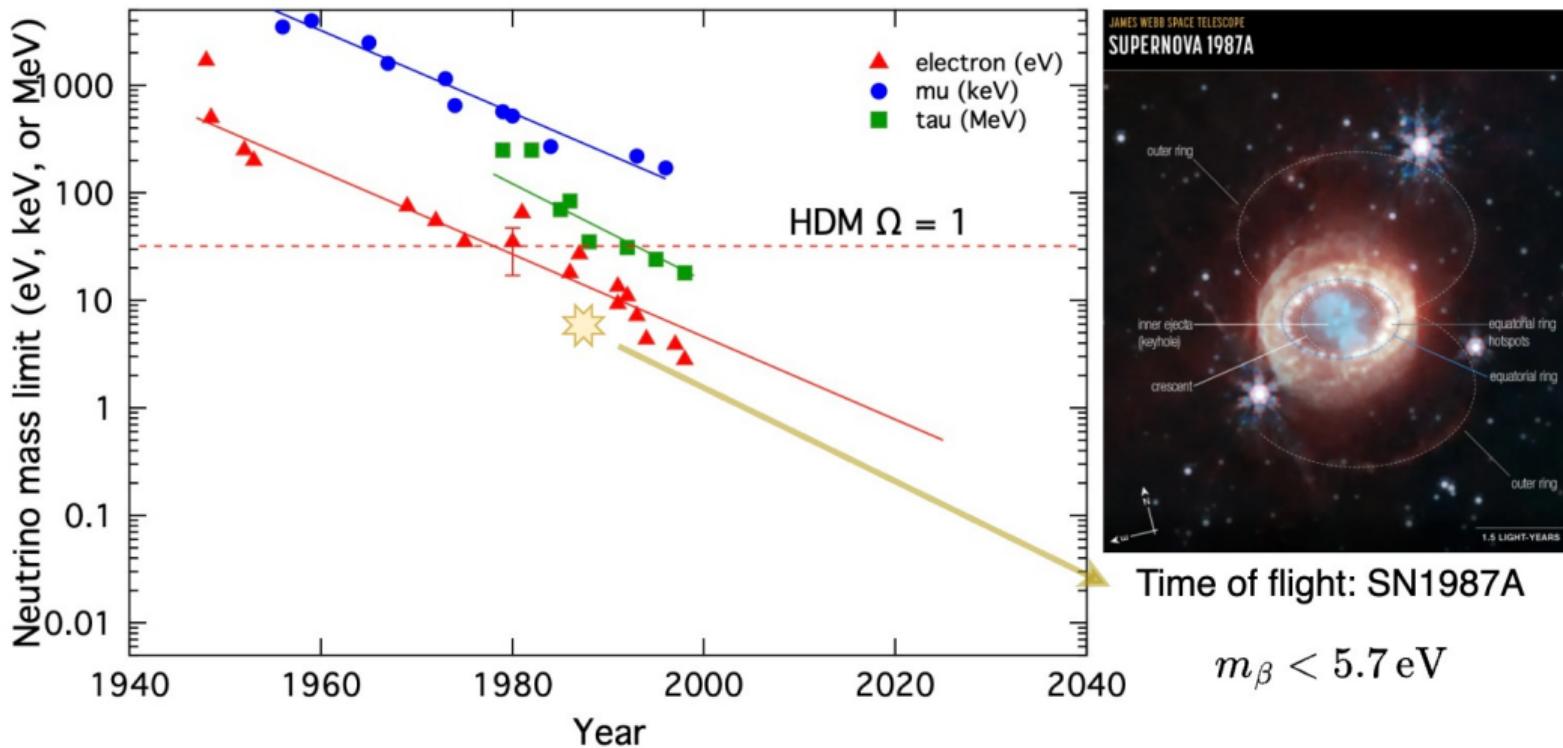


$$m_\beta \lesssim 500 \text{ eV}$$

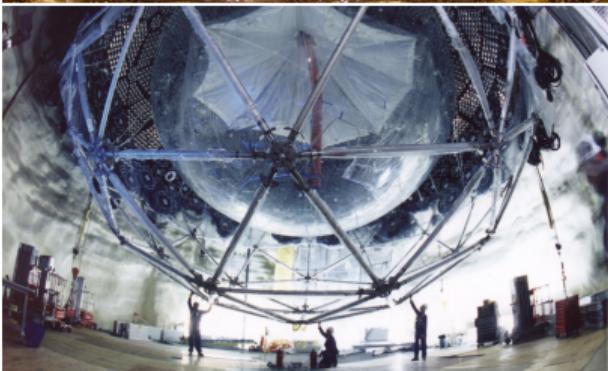
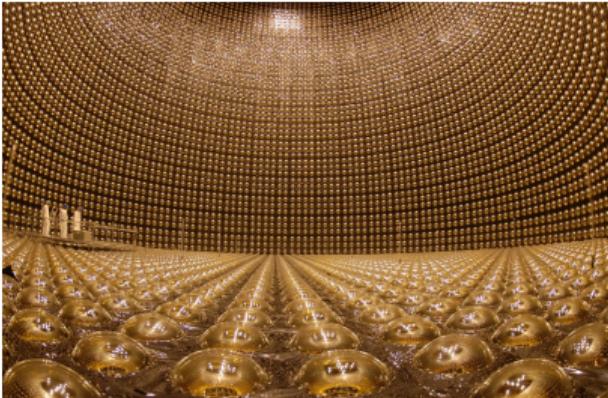


MOORE'S LAW OF NEUTRINO MASS MEASUREMENTS

THE EARLY DAYS



NEUTRINO OSCILLATIONS! NEUTRINOS MUST HAVE MASS



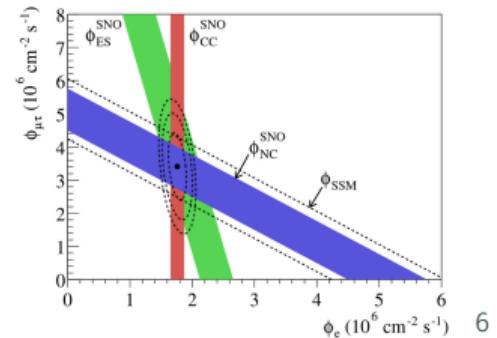
Wouter Van De Pontseele

1. Minimal SM prediction
 $m_\beta = 0$ must be wrong
2. We only need to measure m_β to fix the three mass eigenstates
3. Oscillation measurements cannot measure the absolute scale

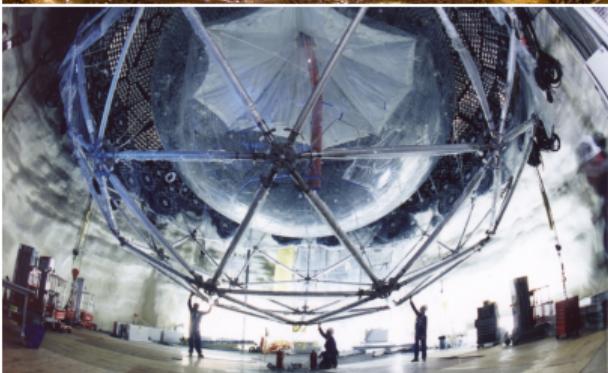
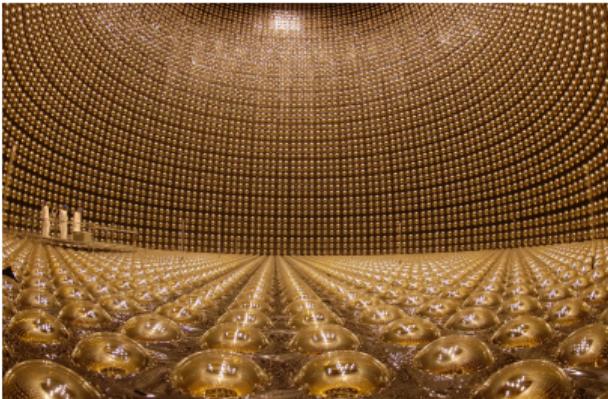
2015 NOBEL PRIZE
in Physics



NEUTRINO OSCILLATIONS
The discovery of these oscillations shows that neutrinos have mass.



NEUTRINO OSCILLATIONS! NEUTRINOS MUST HAVE MASS



Wouter Van De Pontseele

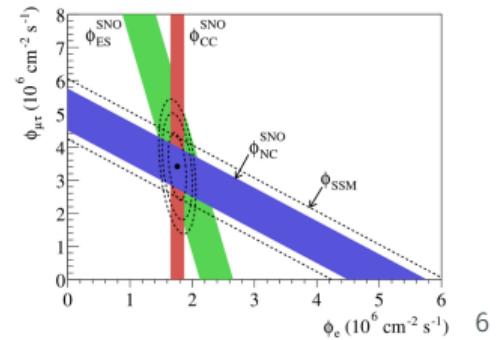
1. Minimal SM prediction
 $m_\beta = 0$ must be wrong
2. We only need to measure m_β to fix the three mass eigenstates
3. Oscillation measurements cannot measure the absolute scale

but set lower limit:
 $m_\beta \gtrsim 9 \text{ meV}$

2015 NOBEL PRIZE
in Physics

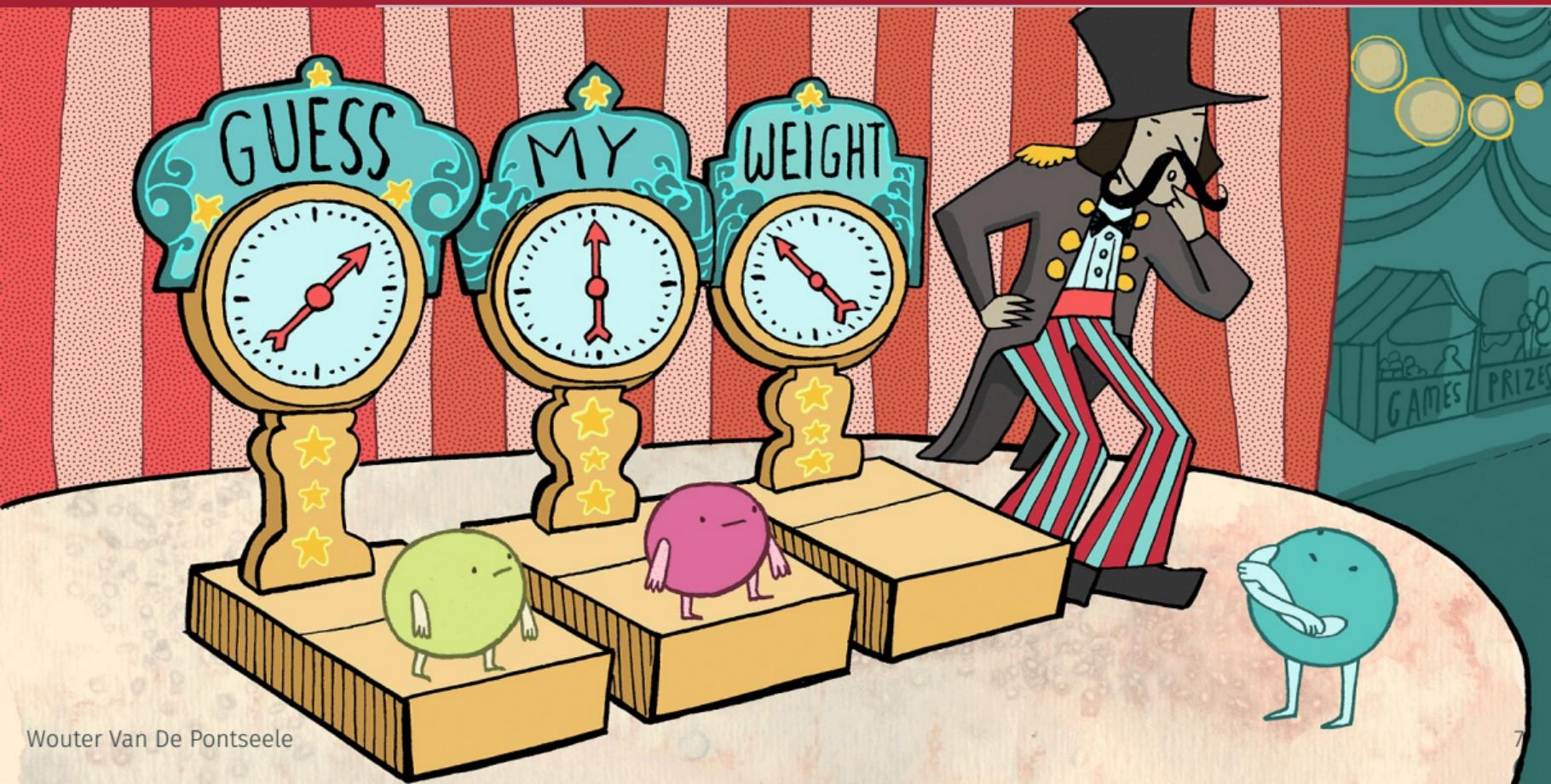


NEUTRINO OSCILLATIONS
The discovery of these oscillations shows that neutrinos have mass.

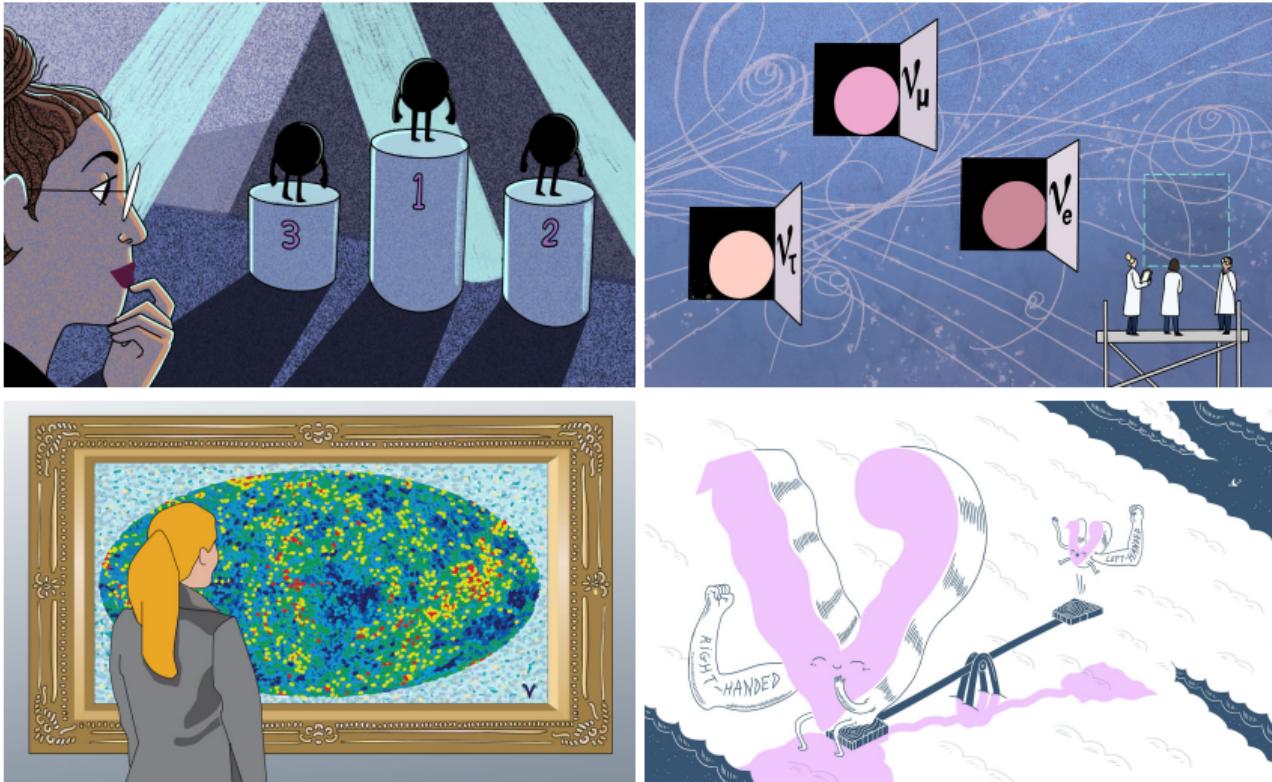


RECENT PROGRESS: WHY DO WE CARE & WHAT TOOLS CAN WE USE?

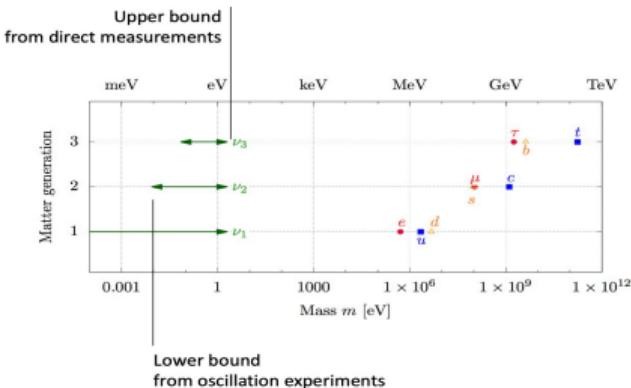
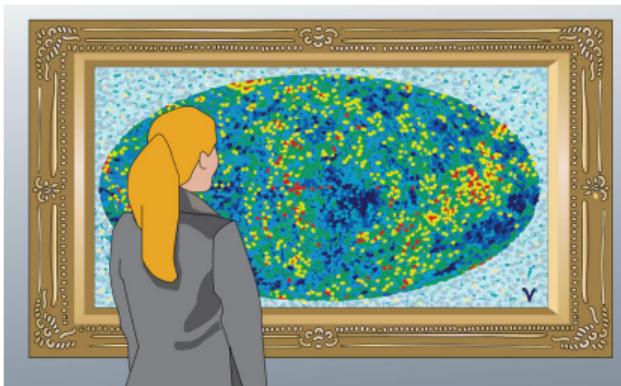
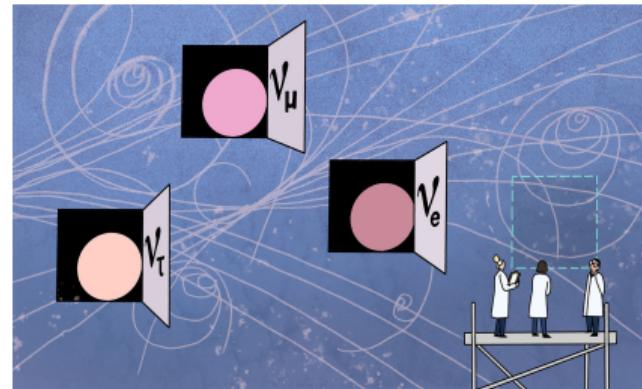
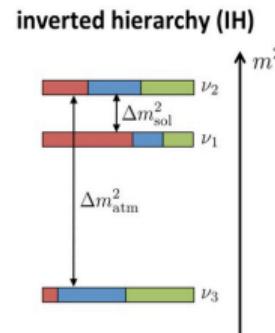
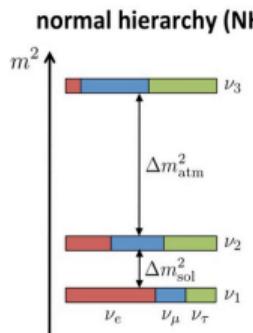
SYMMETRY MAGAZINE



- Neutrino Mass ordering
- Sensitive to sterile ν 's
- Ingredient in cosmological models
- Majorana or Dirac?



- Neutrino Mass ordering
- Sensitive to sterile ν 's
- Ingredient in cosmological models
- Majorana or Dirac?



PROCESSES THAT SHED LIGHT ON THE NEUTRINO MASS SCALE

Cosmology

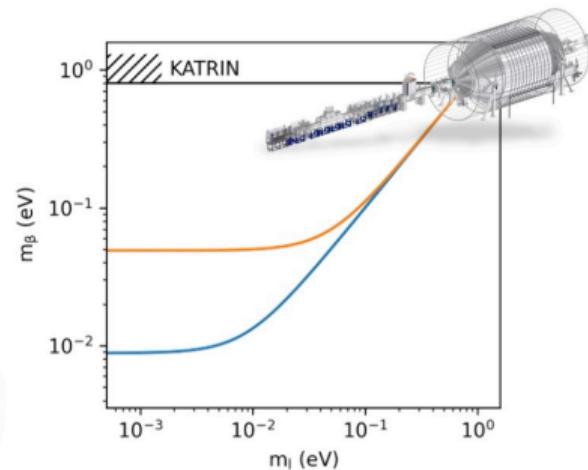
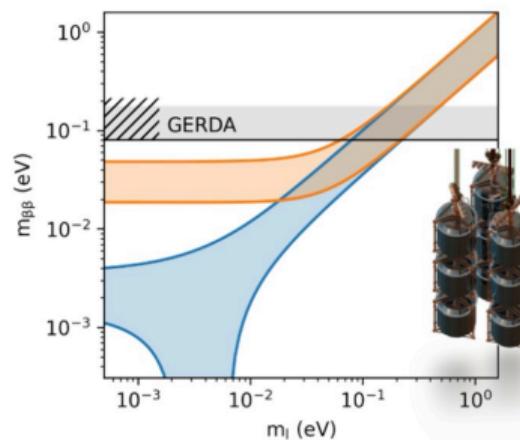
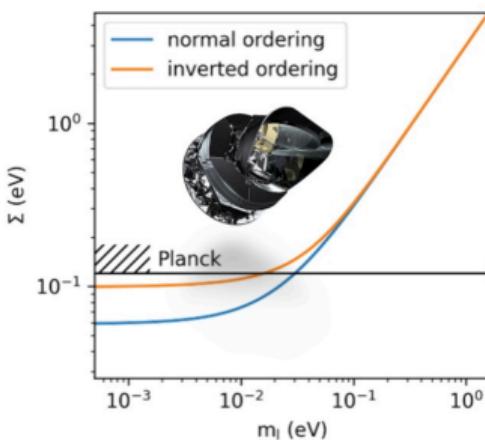
$$\Sigma = \sum_i m_i$$

$0\nu\beta\beta$ -decay

$$m_{\beta\beta} = \sum_i (U_{ei})^2 m_i$$

Kinematics

$$m_\beta = \sqrt{\sum_i |U_{ei}|^2 m_i^2}$$



PROCESSES THAT SHED LIGHT ON THE NEUTRINO MASS SCALE

Cosmology

$$\Sigma = \sum_i m_i$$

- Model dependent

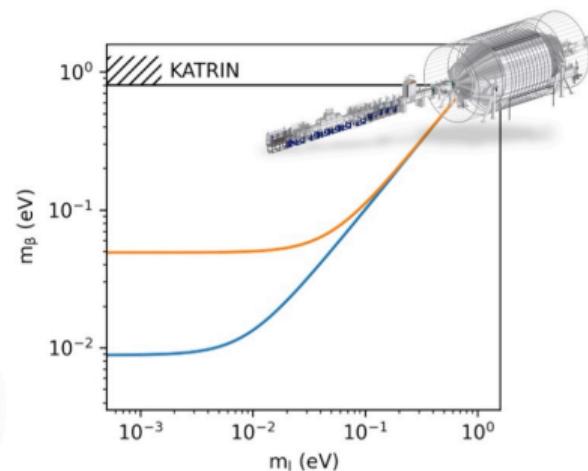
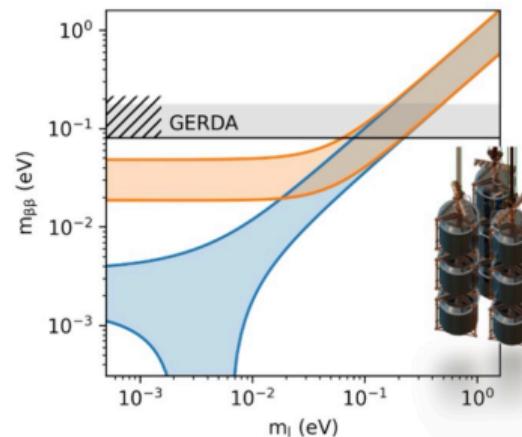
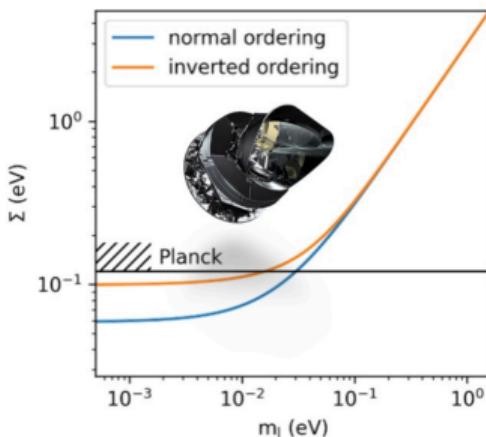
$0\nu\beta\beta$ -decay

$$m_{\beta\beta} = \sum_i (U_{ei})^2 m_i$$

- Only if Majorana

Kinematics

$$m_\beta = \sqrt{\sum_i |U_{ei}|^2 m_i^2}$$



WHAT SOURCES CAN WE USE?

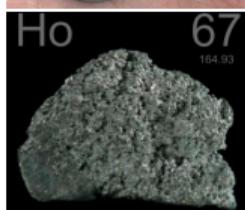


| Type | Element | Isotope | Half-life (years) | Mass diff (Q) |
|------------------|---------|-------------------|--------------------|---------------|
| β -decay | Tritium | $^3\text{H}_2$ | 12 | 18 keV |
| | Indium | ^{115}In | 4×10^{14} | 0.15 keV |
| | Cesium | ^{135}Cs | 1×10^6 | 0.44 keV |
| Electron capture | Rhenium | ^{187}Re | 4×10^{10} | 2.5 keV |
| | Holmium | ^{163}Ho | 4750 | 2.8 keV |

1. Low Mass difference: Number of events in last eV goes as $1/Q^3$
2. Low half-life time: increases statistics.

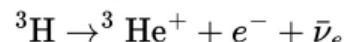
Ultra-low Q sources are under study,
Currently Tritium and Holmium are the favorites

WHAT SOURCES CAN WE USE?

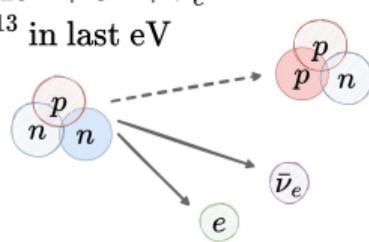


| Type | Element | Isotope | Half-life (years) | Mass diff (Q) |
|------------------|---------|-------------------|--------------------|---------------|
| β -decay | Tritium | $^3\text{H}_2$ | 12 | 18 keV |
| | Indium | ^{115}In | 4×10^{14} | 0.15 keV |
| | Cesium | ^{135}Cs | 1×10^6 | 0.44 keV |
| Electron capture | Rhenium | ^{187}Re | 4×10^{10} | 2.5 keV |
| | Holmium | ^{163}Ho | 4750 | 2.8 keV |

β -decay



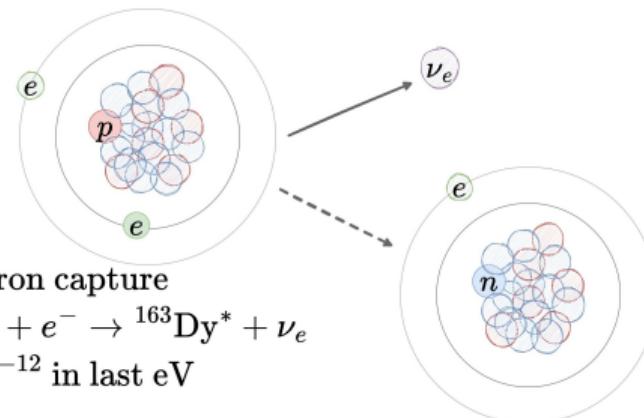
3×10^{-13} in last eV



Electron capture

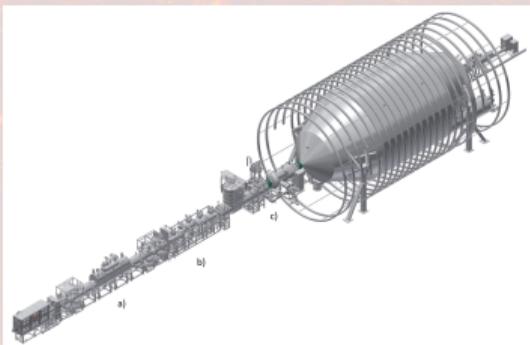


$\approx 10^{-12}$ in last eV



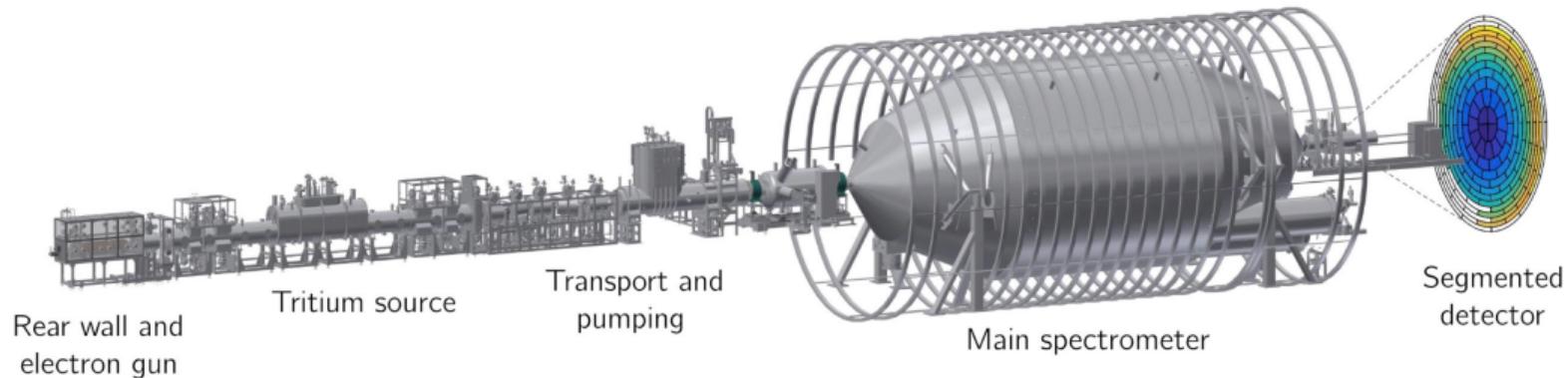
WHAT EXPERIMENTAL TECHNIQUES ARE ON THE TABLE?

| Method | Variable affected by m_ν | Resulting measurements | Isotope |
|----------------------|------------------------------|------------------------|---------|
| MAC-E filter | Electron energy | Counts above threshold | Tritium |
| Calorimetry | Deposited energy | Heat/Phonons | Holmium |
| CRES | Electron energy | Frequency spectrum | Tritium |
| Levitating particles | Recoil spectrum | Scattered light | TBD |

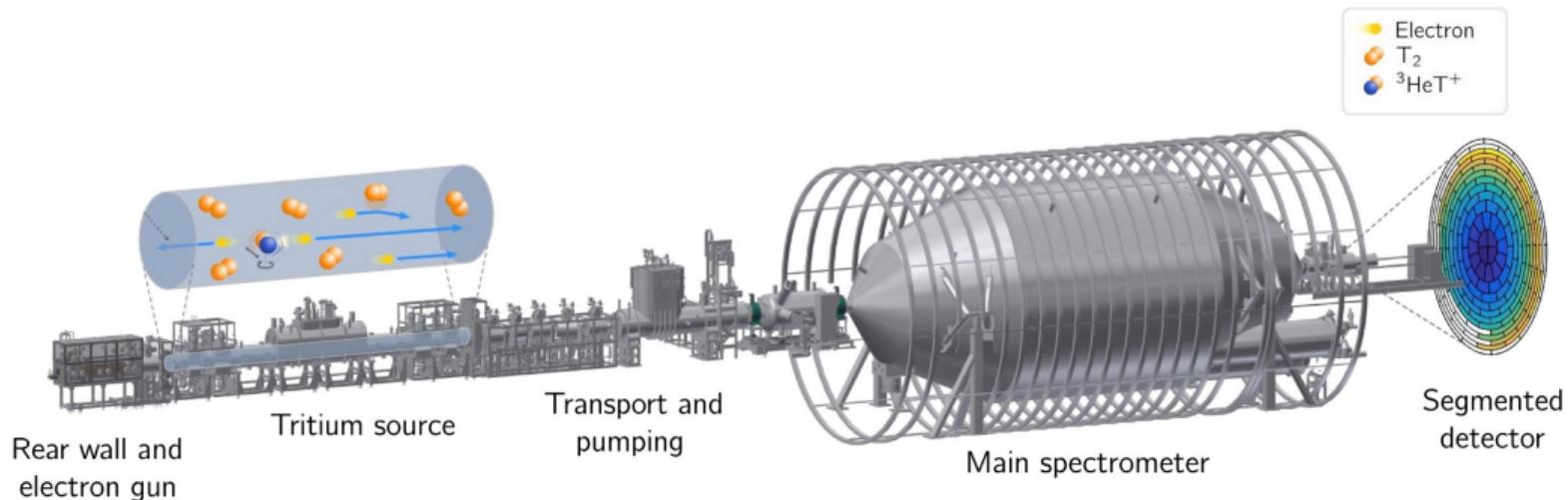


KATRIN: MAGNETIC ADIABATIC COLLIMATION COMBINED WITH AN ELECTROSTATIC FILTER

Adiabatic regime: transverse kinetic energy / $| \vec{B} |$ is constant.

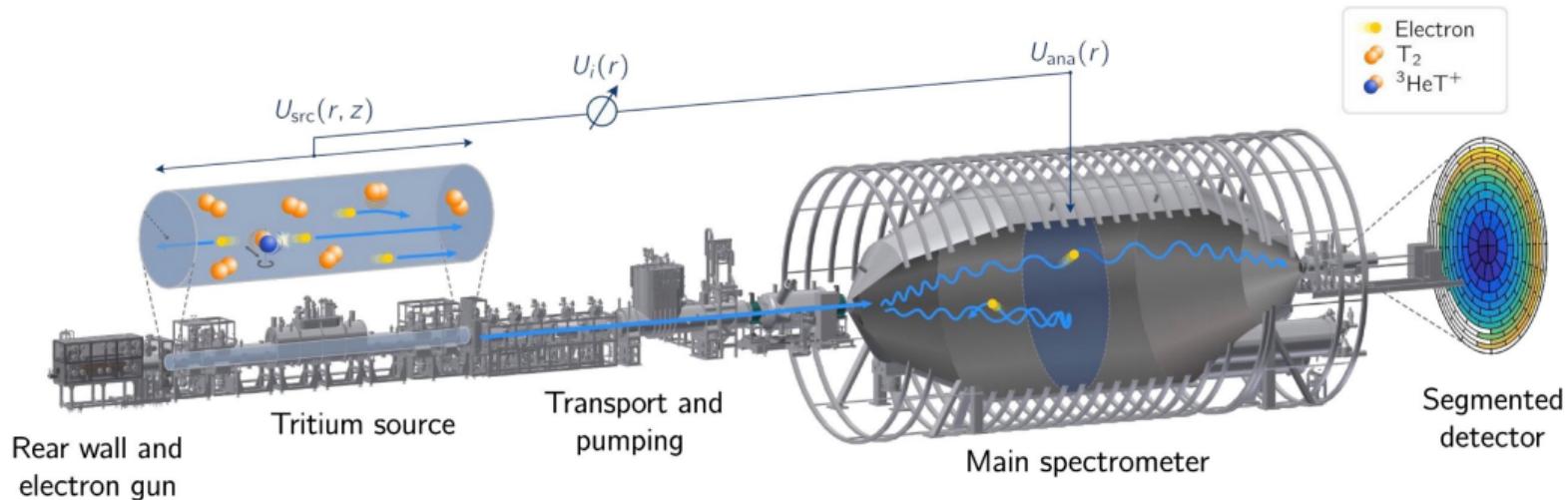


Adiabatic regime: transverse kinetic energy / $| \vec{B} |$ is constant.



KATRIN: MAGNETIC ADIABATIC COLLIMATION COMBINED WITH AN ELECTROSTATIC FILTER

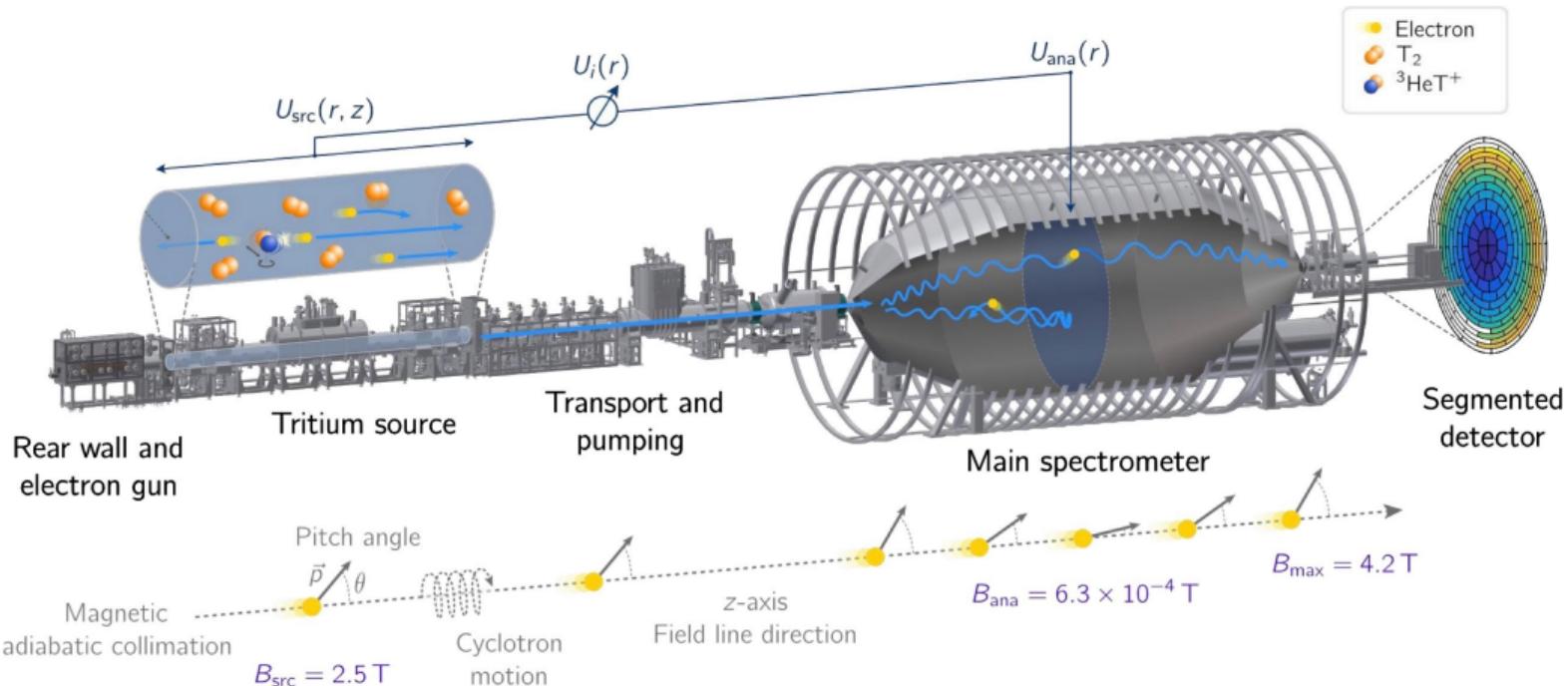
Adiabatic regime: transverse kinetic energy / $| \vec{B} |$ is constant.



Spectrometer acts as an **integrating high-energy pass filter**

KATRIN: MAGNETIC ADIABATIC COLLIMATION COMBINED WITH AN ELECTROSTATIC FILTER

Adiabatic regime: transverse kinetic energy / $| \vec{B} |$ is constant.



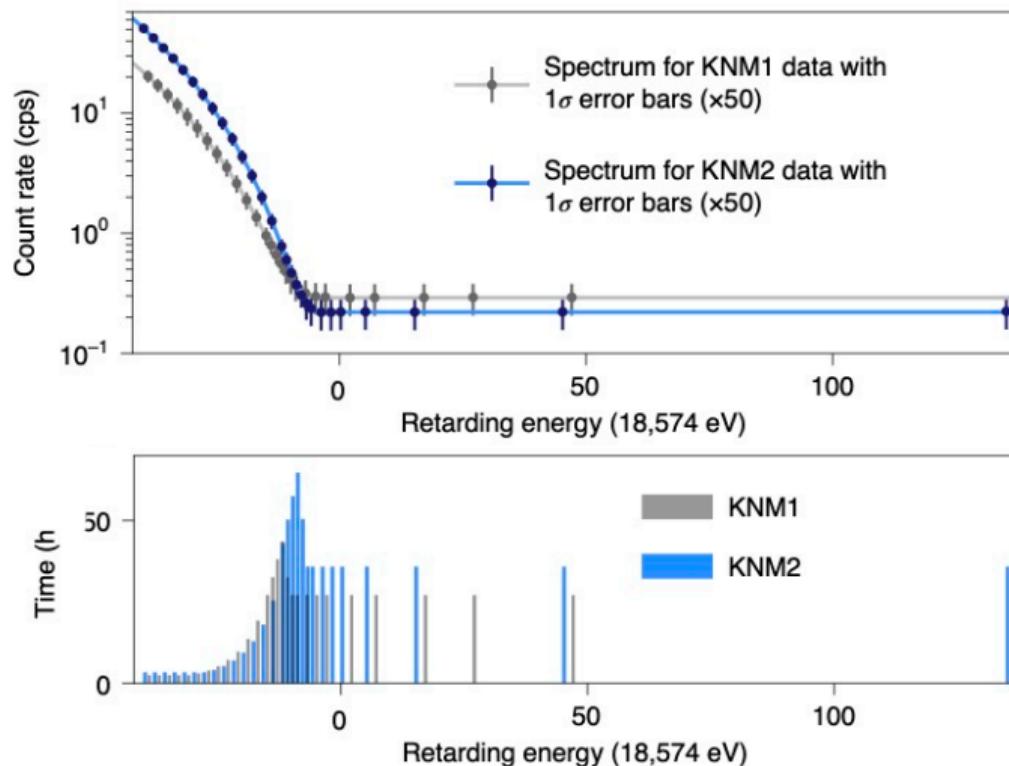
Spectrometer resolution: $\Delta E/E = B_{min}/B_{max}$

KATRIN: THE JOURNEY FROM MUNICH TO KARLSRUHE

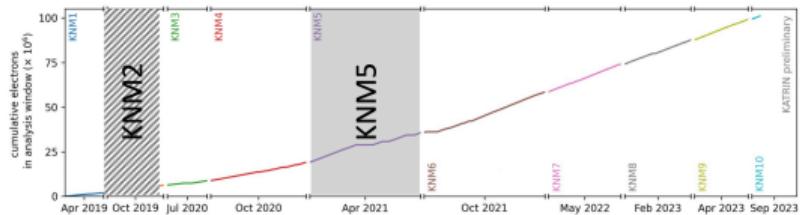
SYMMETRY MAGAZINE



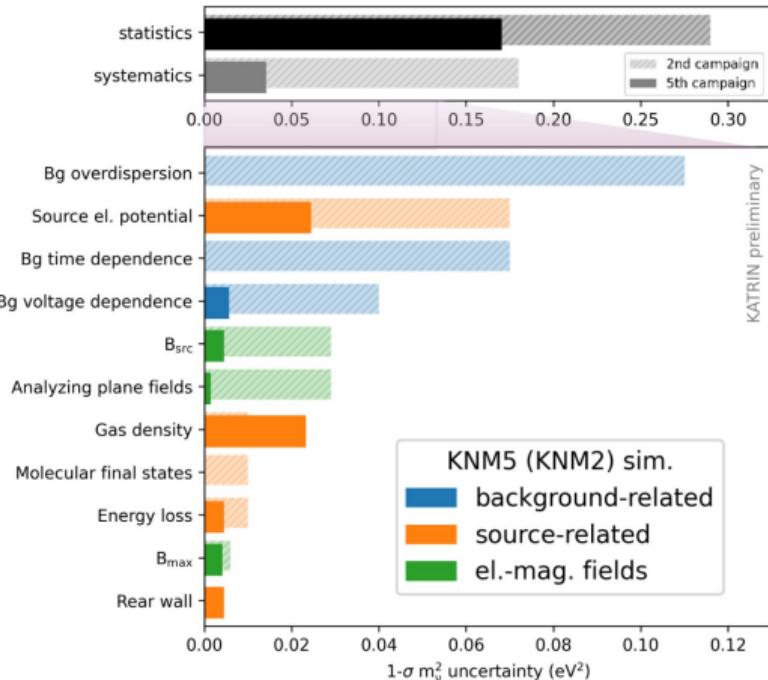
KATRIN: WORLD-LEADING SENSITIVITY TO m_β



- Combination of first two run periods, **less than 100 days of data**.
- Combined result: $m_\beta < 0.8$ eV (90% CL)
- KATRIN Collaboration, Nat. Phys. 18, 160–166 (2022)
- Search for relic- ν overabundance, sterile neutrinos and violation of Lorentz invariance



- A lot more data!
- Reduction in systematic uncertainties.



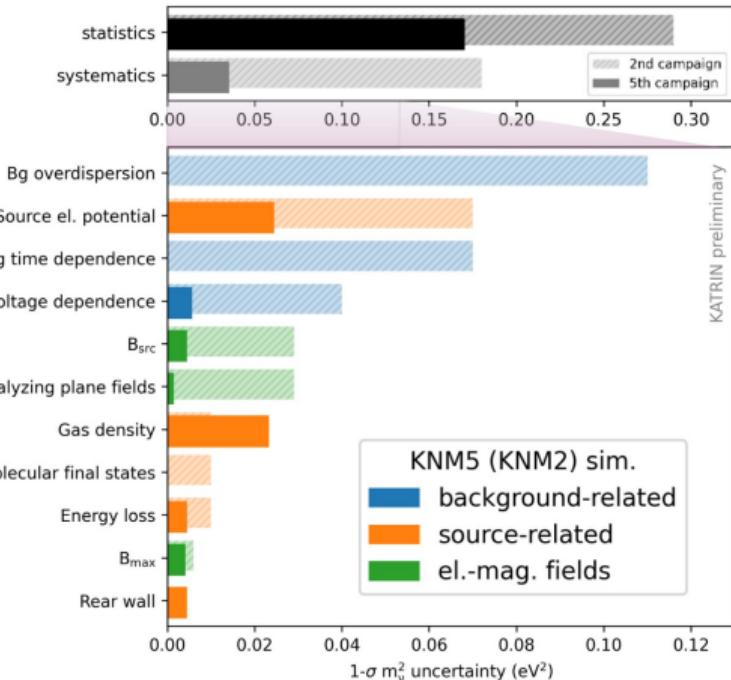
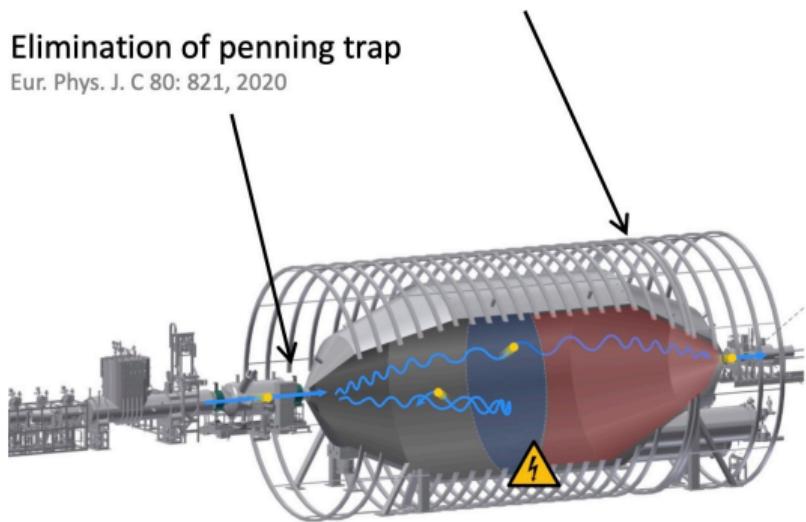
- Reduction in systematic uncertainties.

Shifted analyzing plane

Lokhov et al arXiv:2201.11743 (2022)

Elimination of penning trap

Eur. Phys. J. C 80: 821, 2020



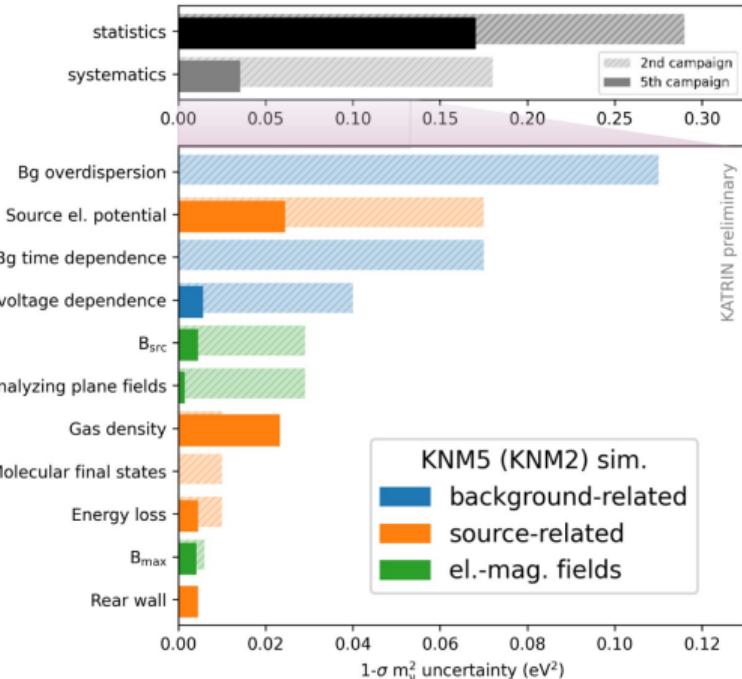
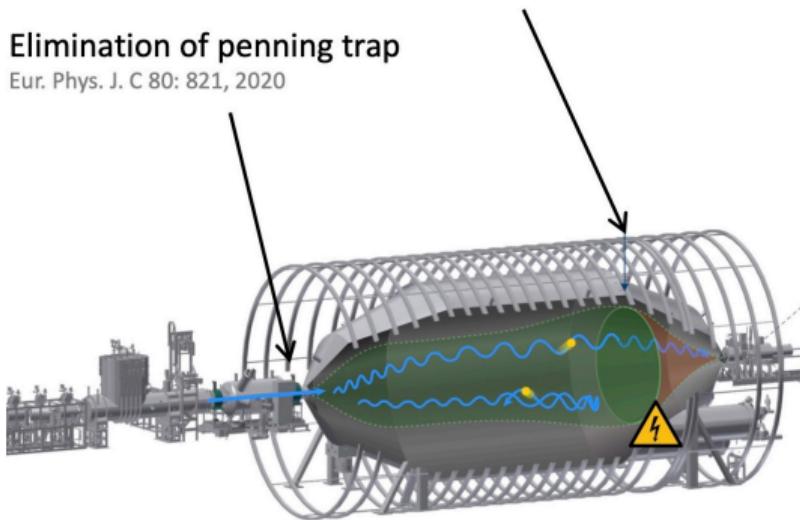
- Reduction in systematic uncertainties.

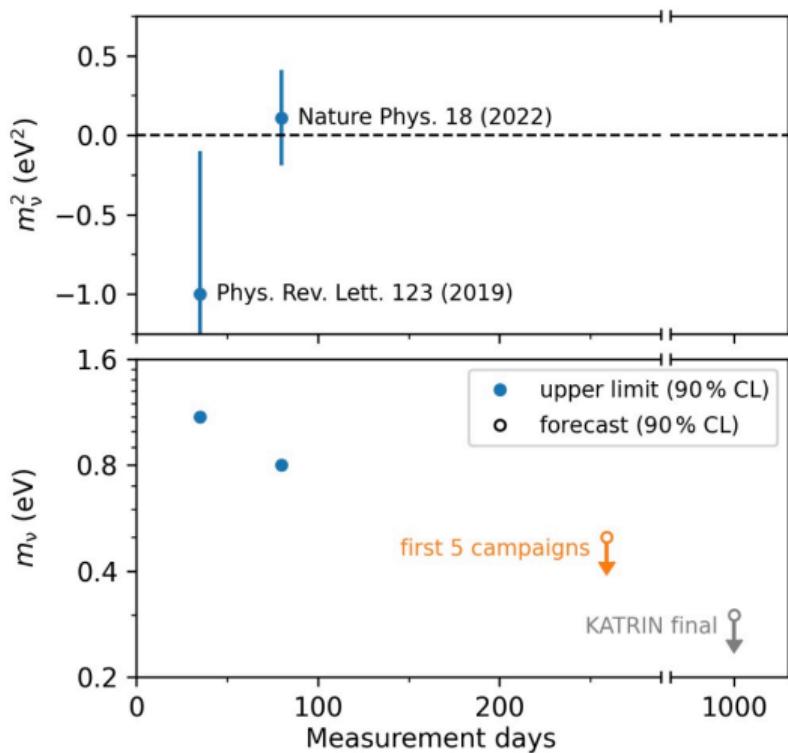
Shifted analyzing plane

Lokhov et al arXiv:2201.11743 (2022)

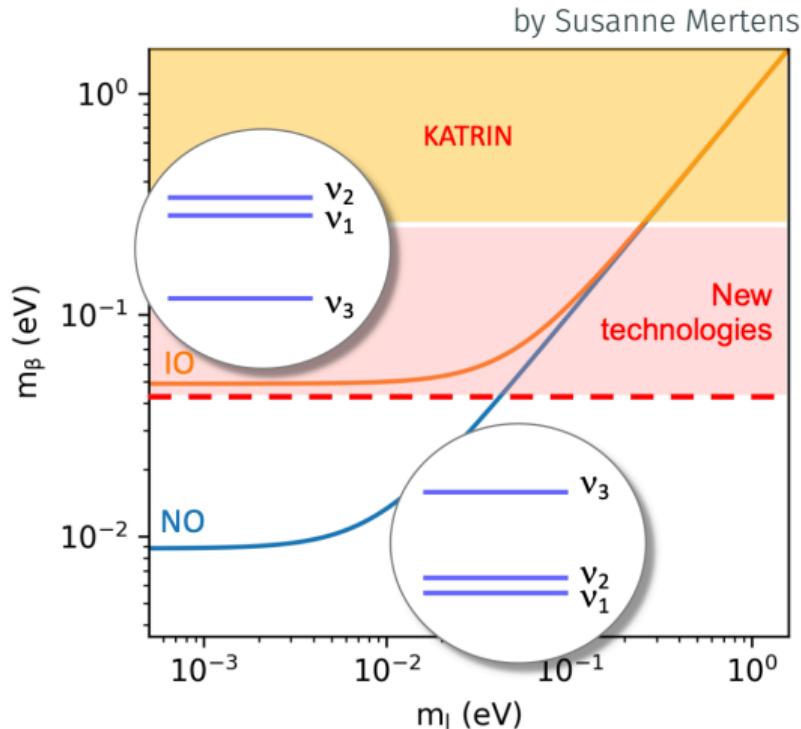
Elimination of penning trap

Eur. Phys. J. C 80: 821, 2020





- Finalising analysis of the first five runs
- Sensitivity: $m_\beta < 0.5 \text{ eV}$ (90% CL)
- Data taking ongoing till 2025 with final sensitivity better than $m_\beta < 0.3 \text{ eV}$.
- Next up: Sterile neutrino search facilitated by silicon-based pixelated e^- detection (TRISTAN).



- Finalising analysis of the first five runs
- Sensitivity: $m_\beta < 0.5$ eV (90% CL)
- Data taking ongoing till 2025 with final sensitivity better than $m_\beta < 0.3$ eV.
- Next up: Sterile neutrino search facilitated by silicon-based pixelated e^- detection (TRISTAN).

New technologies needed to explore the inverted mass ordering!

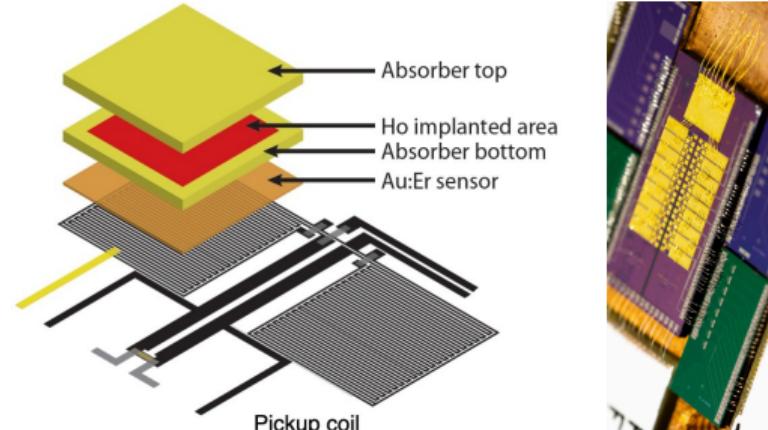
WHAT EXPERIMENTAL TECHNIQUES ARE ON THE TABLE?

| Method | Variable affected by m_ν | Resulting measurements | Isotope |
|----------------------|------------------------------|------------------------|---------|
| MAC-E filter | Electron energy | Counts above threshold | Tritium |
| Calorimetry | Deposited energy | Heat/Phonons | Holmium |
| CRES | Electron energy | Frequency spectrum | Tritium |
| Levitating particles | Recoil spectrum | Scattered light | TBD |



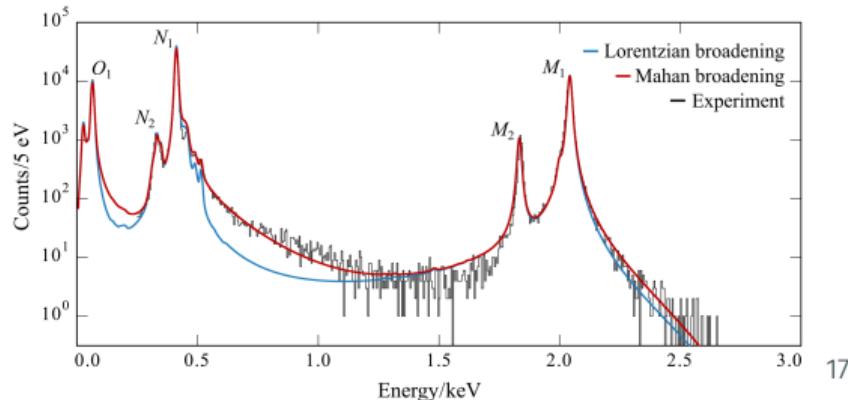
Calorimetric technology

- Magnetic MicroCalorimeter (MMC)
Material magnetisation depends on temperature.



Achievements

- Prototype with 4 pixels: $\nu_\beta < 150 \text{ eV}$.

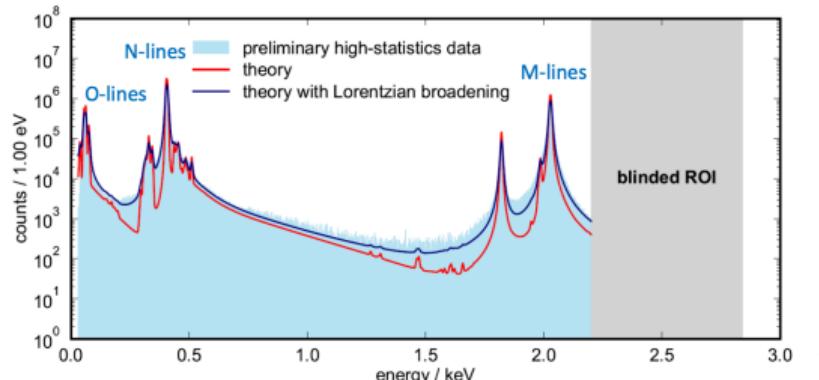
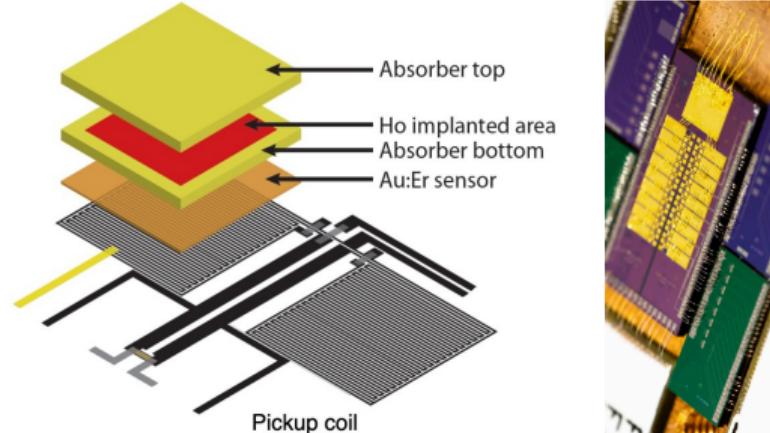


Calorimetric technology

- Magnetic MicroCalorimeter (MMC)
Material magnetisation depends on temperature.

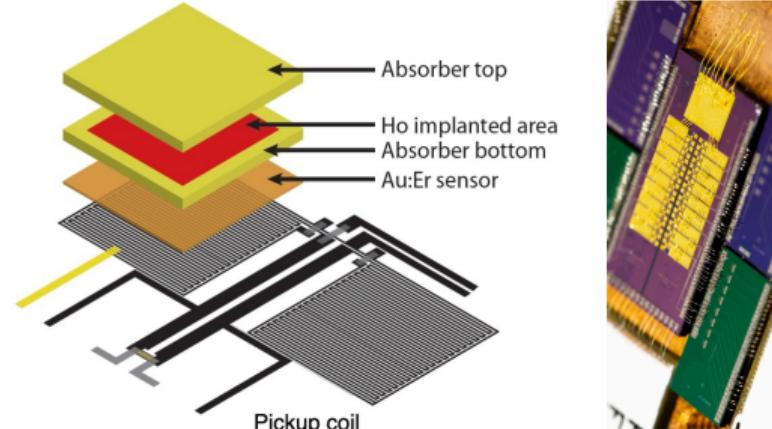
Achievements

- Prototype with 4 pixels: $\nu_\beta < 150 \text{ eV}$.
- ECHo-1k: Systematics in progress
Expected sensitivity: $\nu_\beta < 20 \text{ eV}$.



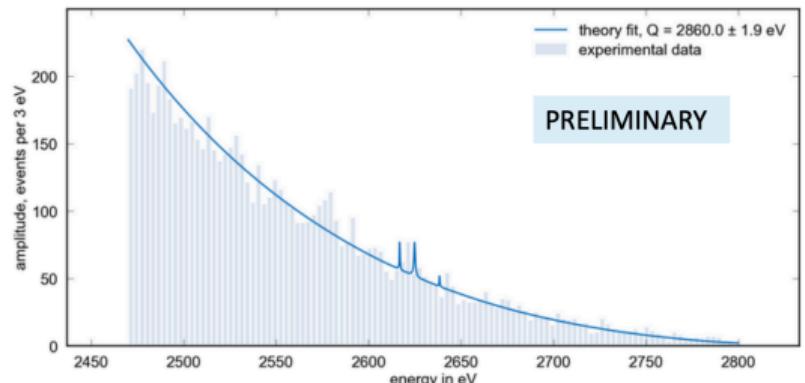
Calorimetric technology

- Magnetic MicroCalorimeter (MMC)
Material magnetisation depends on temperature.



Achievements

- Prototype with 4 pixels: $\nu_\beta < 150 \text{ eV}$.
- ECHo-1k: Systematics in progress
Expected sensitivity: $\nu_\beta < 20 \text{ eV}$.

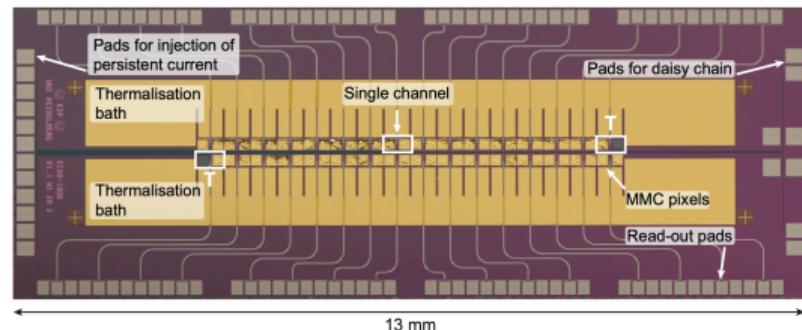
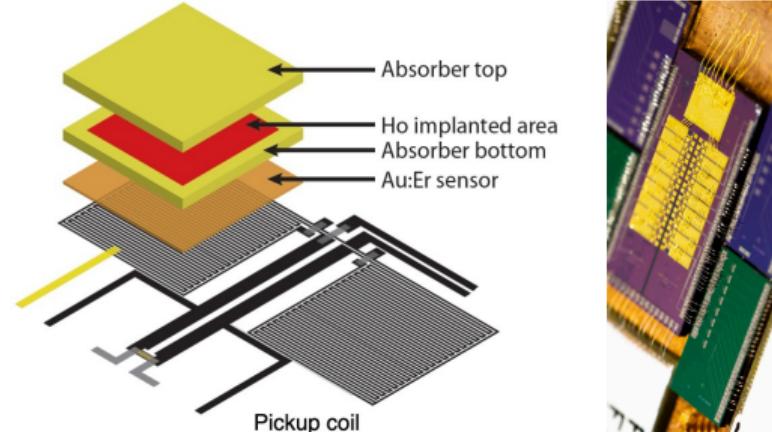


Calorimetric technology

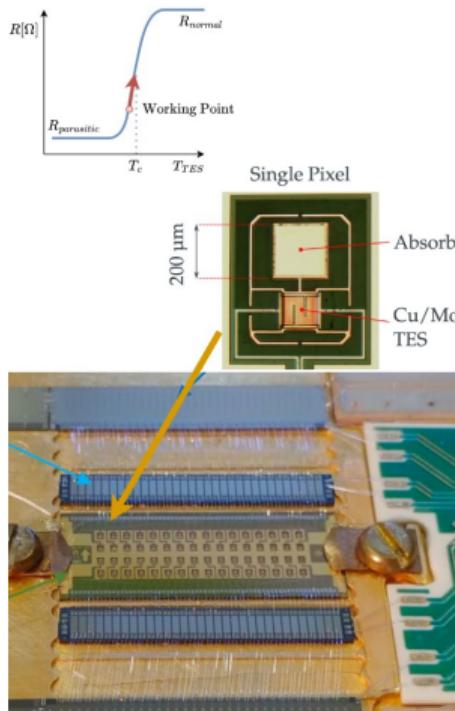
- Magnetic MicroCalorimeter (MMC)
Material magnetisation depends on temperature.

Achievements

- Prototype with 4 pixels: $\nu_\beta < 150 \text{ eV}$.
- ECHO-1k: Systematics in progress
Expected sensitivity: $\nu_\beta < 20 \text{ eV}$.
- ECHO-100k: 1200 multiplexed pixels with $\mathcal{O}(\text{eV})$ sensitivity.



HOLMES NEUTRINO MASS EXPERIMENT

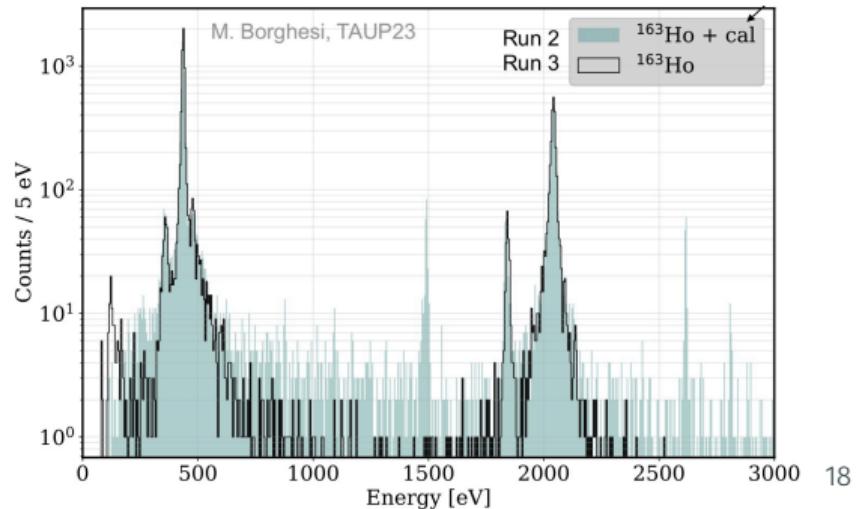


Wouter Van De Pontseele

Calorimetric technology

- Transition Edge Sensors (TES)
Thin film superconductor

First endpoint measurement in progress,
Spectrum from 4 pixels with $\Delta E \approx 6$ eV:



WHAT EXPERIMENTAL TECHNIQUES ARE ON THE TABLE?

| Method | Variable affected by m_ν | Resulting measurements | Isotope |
|----------------------|------------------------------|---------------------------|---------|
| MAC-E filter | Electron energy | Counts above threshold | Tritium |
| Calorimetry | Deposited energy | Heat/Phonons | Holmium |
| CRES | Electron energy | Frequency spectrum | Tritium |
| Levitating particles | Recoil spectrum | Scattered light | TBD |

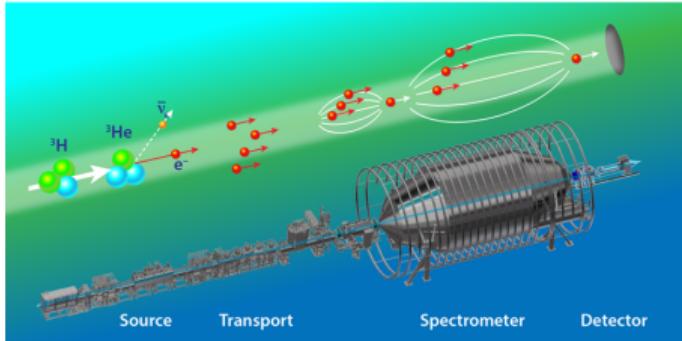
PROJECT 8



Quantum Technologies for
Neutrino Mass

PROJECT 8: A FREQUENCY MEASUREMENT

How to scale down KATRIN while increasing the m_β sensitivity?

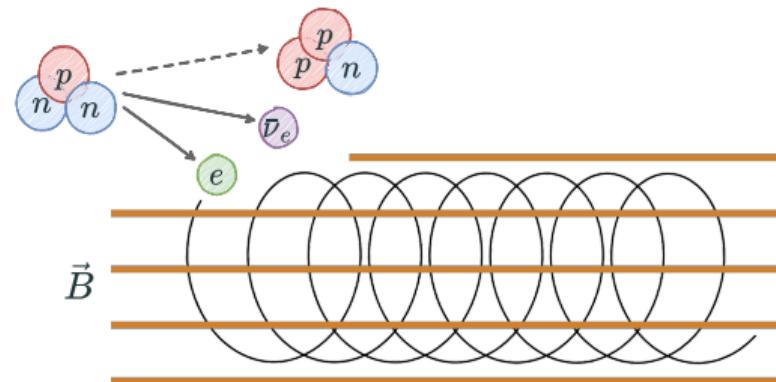


PROJECT 8: A FREQUENCY MEASUREMENT

How to scale down KATRIN while increasing the m_β sensitivity?

1. Spectroscopic instead of integral
2. Use a source inside the detector
Magnetically trapped atomic tritium.
3. Frequency detection of cyclotron radiation emitted by electrons.

Tritium decays inside a uniform \vec{B} -field.



Electron performs cyclotron motion with frequency

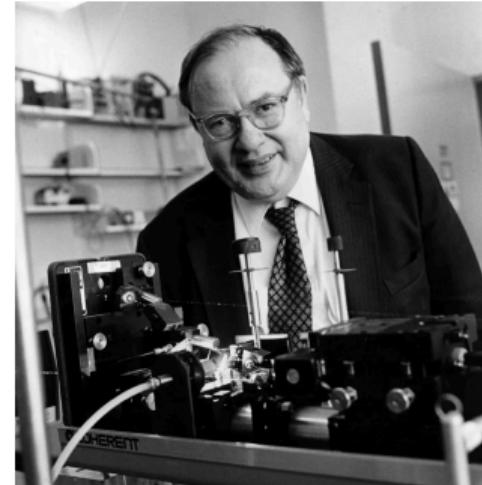
$$f(B, E_{kin}) = \frac{1}{2\pi} \frac{eB}{m_e + E_{kin}/c^2}$$

PROJECT 8: A FREQUENCY MEASUREMENT

How to scale down KATRIN while increasing the m_β sensitivity?

1. Spectroscopic instead of integral
2. Use a source inside the detector
Magnetically trapped atomic tritium.
3. Frequency detection of cyclotron radiation emitted by electrons.

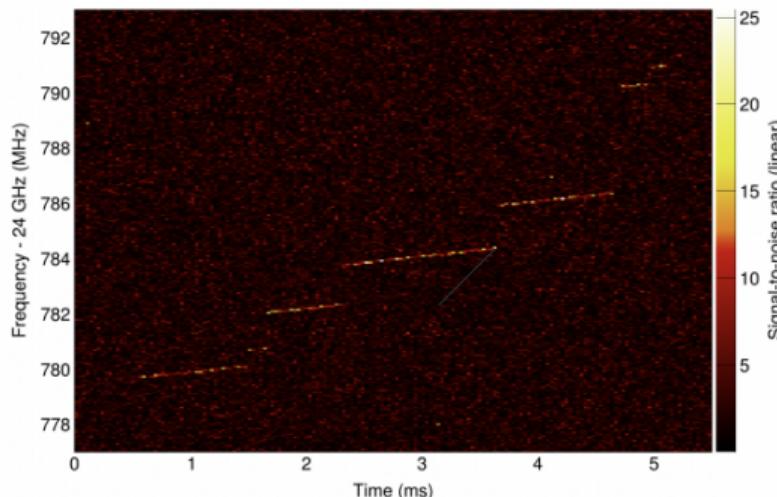
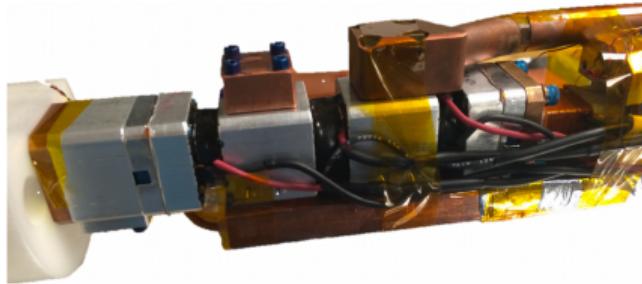
Idea by J. Formaggio and B. Monreal:
Project 8 employs Cyclotron Radiation Emission Spectroscopy (**CRES**)



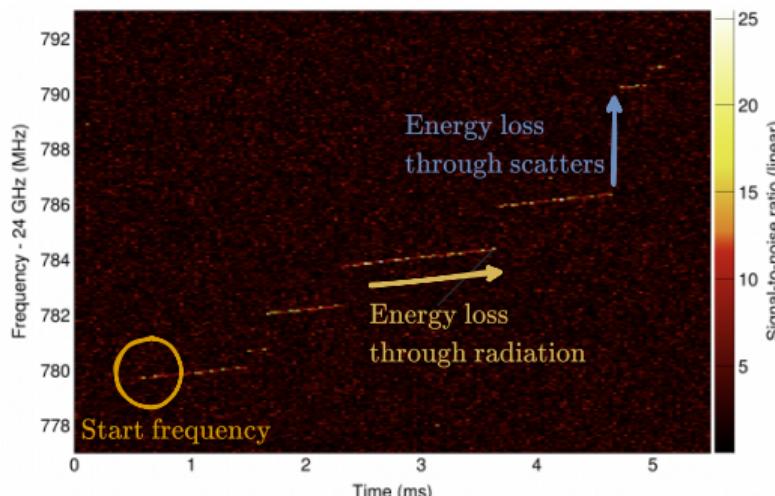
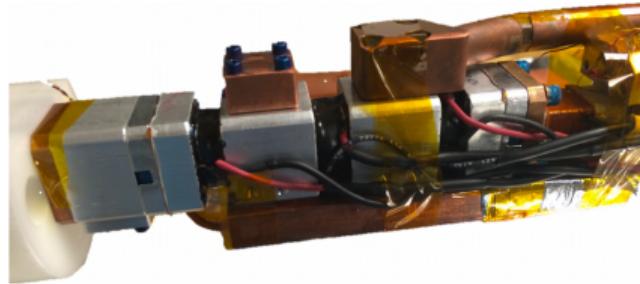
“Never measure anything but frequency”

*– Arthur Schawlow,
co-inventor of the laser
and 1981 Nobel Prize winner*

- Phase I: First detection of cyclotron radiation from a single electron.
Gaseous ^{83m}Kr used as a source.
Phys. Rev. Lett. 114, 1162501 (2015)



- Phase I: First detection of cyclotron radiation from a single electron.
Gaseous ^{83m}Kr used as a source.
Phys. Rev. Lett. 114, 1162501 (2015)



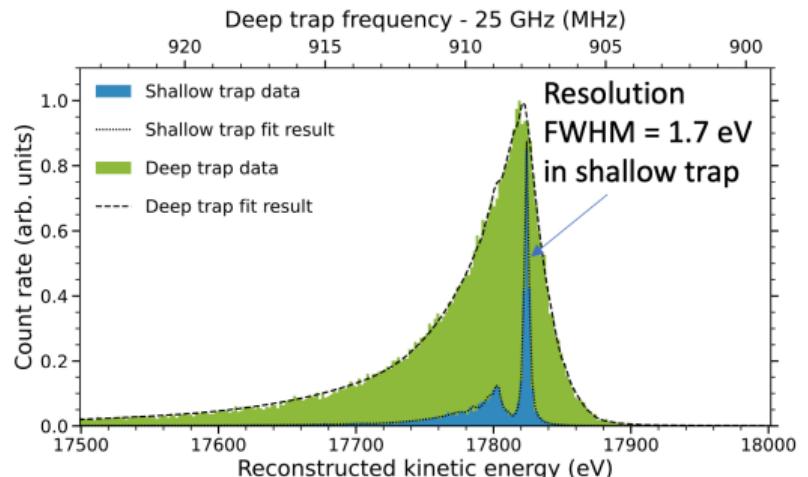
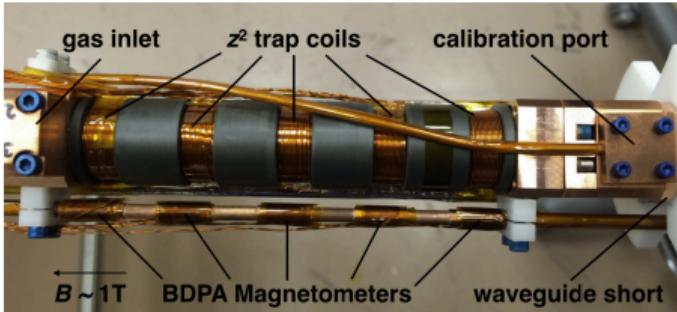
- Phase I: First detection of cyclotron radiation from a single electron.

Gaseous ^{83m}Kr used as a source.

Phys. Rev. Lett. 114, 1162501 (2015)

- Phase II: First limit on the neutrino mass using gaseous tritium and a waveguide antenna.

Phys. Rev. Lett. 131, 102502 (Sep 2023)



- Phase I: First detection of cyclotron radiation from a single electron.

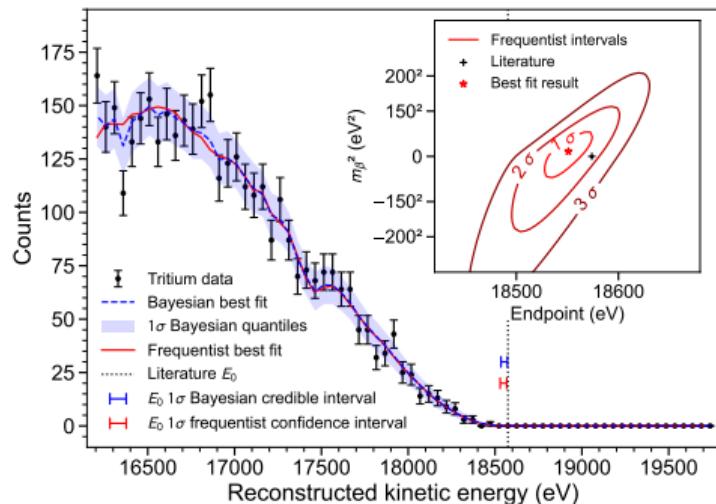
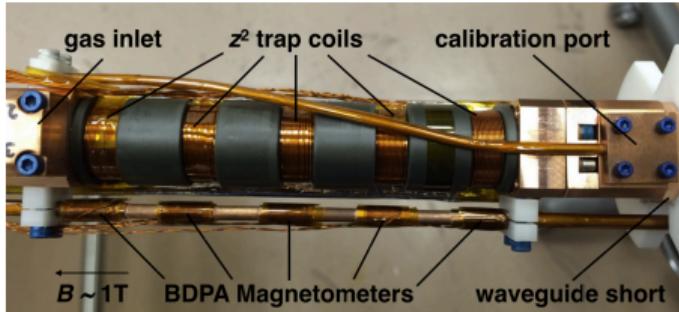
Gaseous ^{83m}Kr used as a source.

Phys. Rev. Lett. 114, 1162501 (2015)

- Phase II: First limit on the neutrino mass using gaseous tritium and a waveguide antenna.

Phys. Rev. Lett. 131, 102502 (Sep 2023)

$$m_\beta < 155 \text{ eV}$$



NOT IMPRESSED?



Project 8 Phase II
Spectrometer to scale

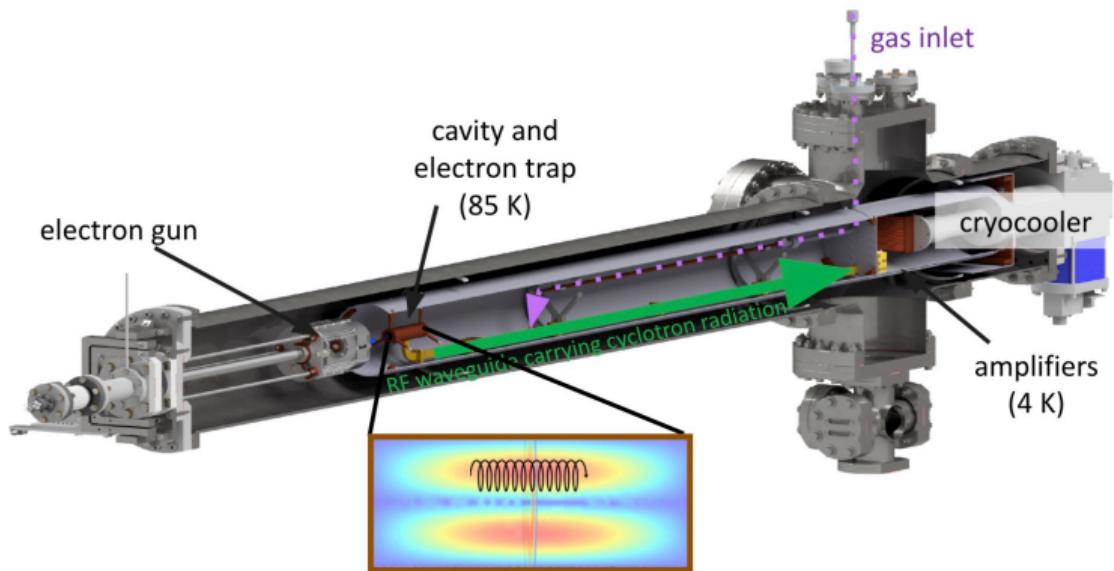
PROJECT 8: WHAT'S NEXT?

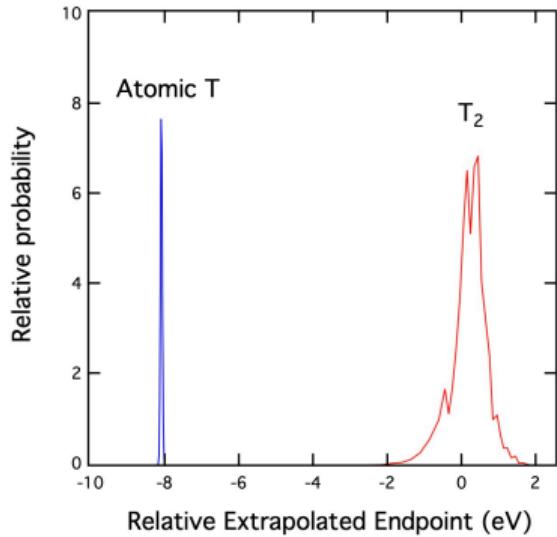
SYMMETRY MAGAZINE



Phase III:

- Resonant cavities

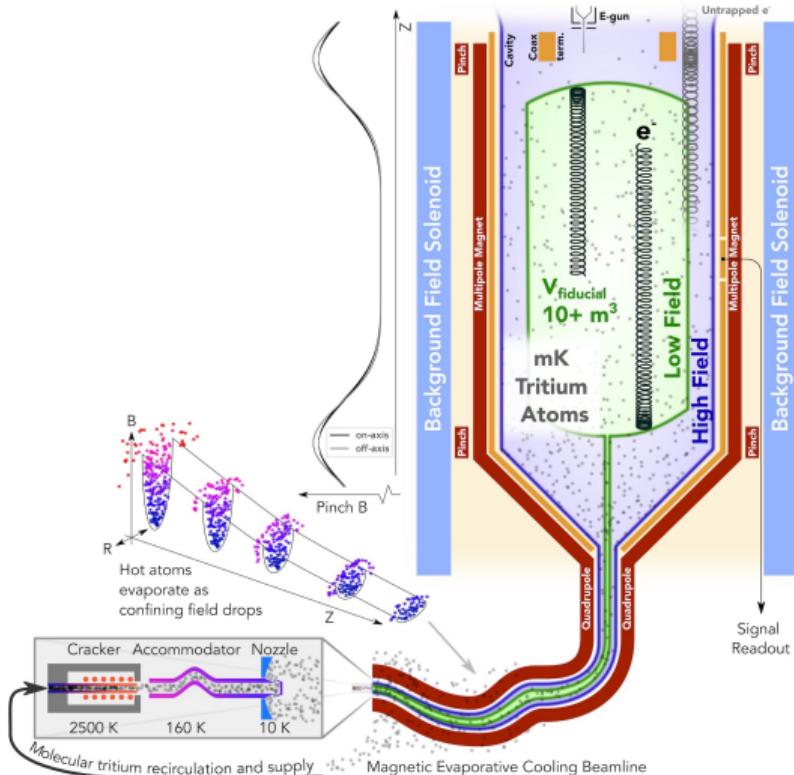




R. G. H. Robertson

Phase III:

- Resonant cavities
- Atomic tritium as a source



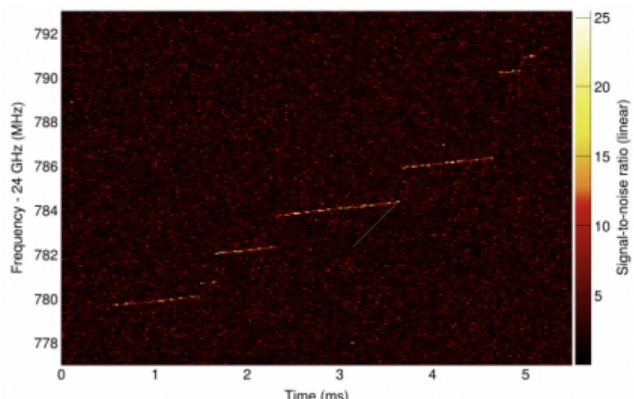
Phase III:

- Resonant cavities
- Atomic tritium as a source
- Scaling to large volumes

Phase IV:

The ultimate neutrino mass experiment probing $m_\beta \approx 40 \text{ meV}$.

- Dropping the magnetic field/frequency to reach $\mathcal{O}(10 \text{ m}^3)$ volume.
- Zeptowatt signals at 325 MHz



Phase III:

- Resonant cavities
- Atomic tritium as a source
- Scaling to large volumes

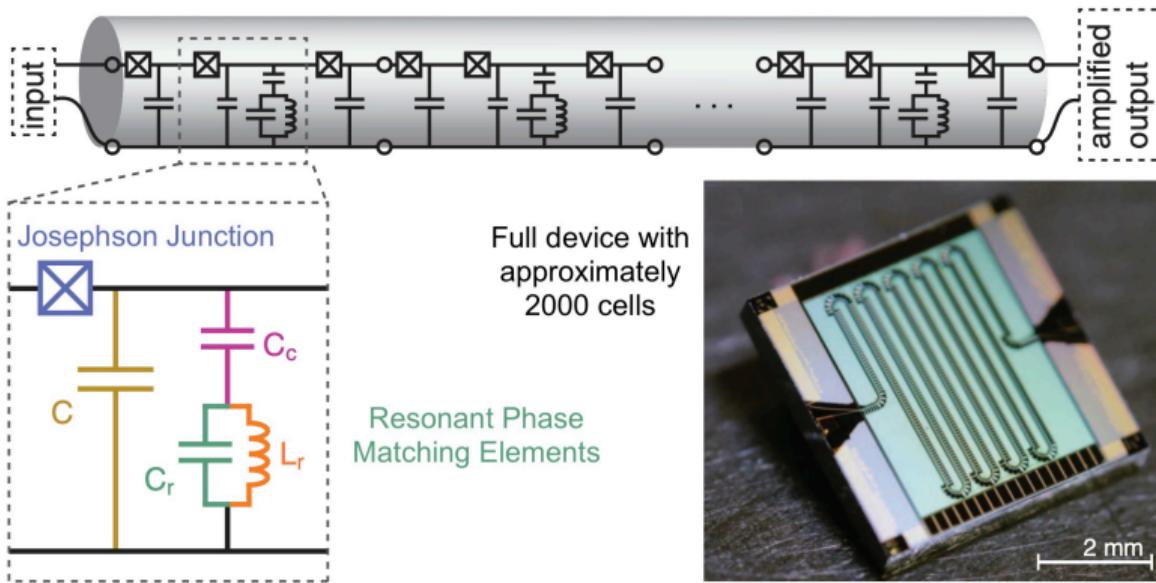
Phase IV:

The ultimate neutrino mass experiment probing $m_\beta \approx 40$ meV.

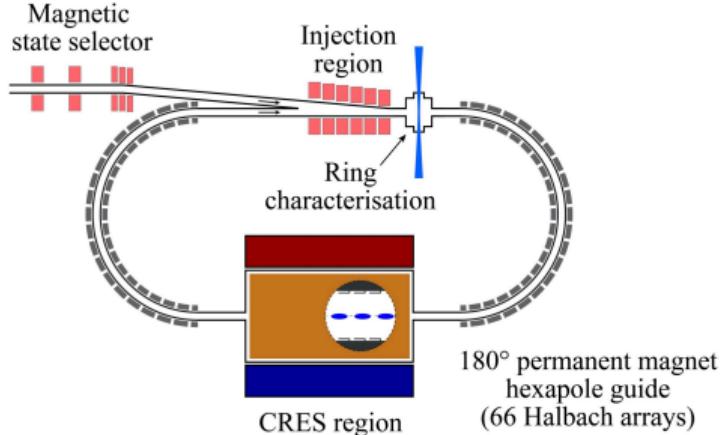
What do we need?

1. A great team!
2. Less noise!

QUANTUM-LIMITED READOUT FOR PROJECT 8: TRAVELLING WAVE PARAMETRIC AMPLIFIERS

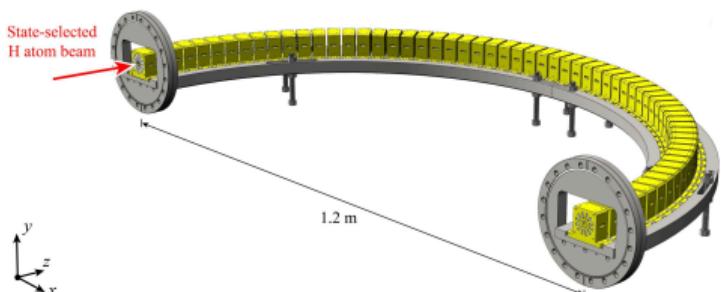


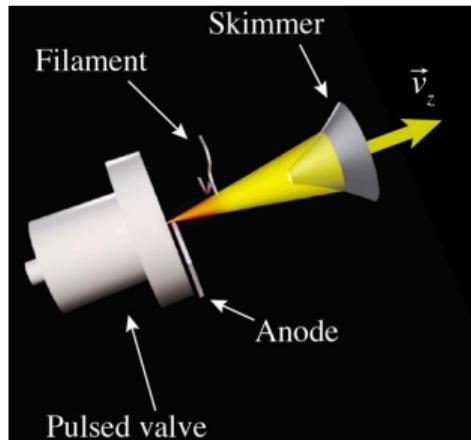
Transmission line-style chain of cells with Josephson Junctions. Bandwidth of $\mathcal{O}(\text{GHz})$.
Designed by Kevin O'Brien's group at MIT, fabricated at Lincoln Laboratories.



Storage ring and free space CRES

- Modular with multiple CRES cells
- Antenna array readout



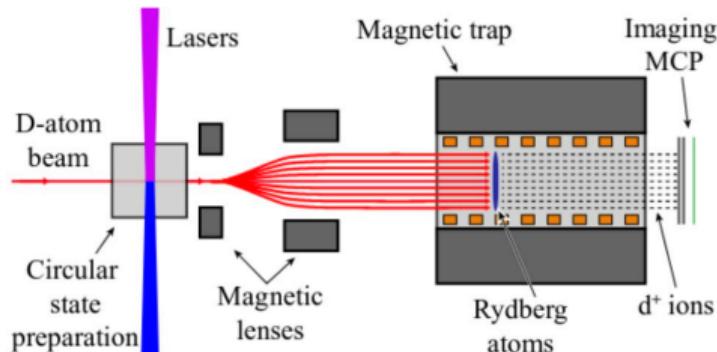


Storage ring and free space CRES

- Modular with multiple CRES cells
- Antenna array readout

Supersonic beam using DC discharge

- Creates beam with temperature of $\mathcal{O}(10 \text{ mK})$ in the moving frame
- Next steps:
higher densities and continuous operation



Storage ring and free space CRES

- Modular with multiple CRES cells
- Antenna array readout

Supersonic beam using DC discharge

- Creates beam with temperature of $\mathcal{O}(10 \text{ mK})$ in the moving frame
- Next steps:
higher densities and continuous operation

Rydberg states magnetometry

- Driven transition inside CRES cell.
- \vec{B} sensitivity $< \mu\text{T}$ with $\approx 1 \text{ mm}$ resolution.

WHAT EXPERIMENTAL TECHNIQUES ARE ON THE TABLE?

| Method | Variable affected by m_ν | Resulting measurements | Isotope |
|----------------------|------------------------------|------------------------|---------|
| MAC-E filter | Electron energy | Counts above threshold | Tritium |
| Calorimetry | Deposited energy | Heat/Phonons | Holmium |
| CRES | Electron energy | Frequency spectrum | Tritium |
| Levitating particles | Recoil spectrum | Scattered light | TBD |

Searches for Massive Neutrinos with Mechanical Quantum Sensors

Daniel Carney, Kyle G. Leach, and David C. Moore
PRX Quantum **4**, 010315 – Published 8 February 2023

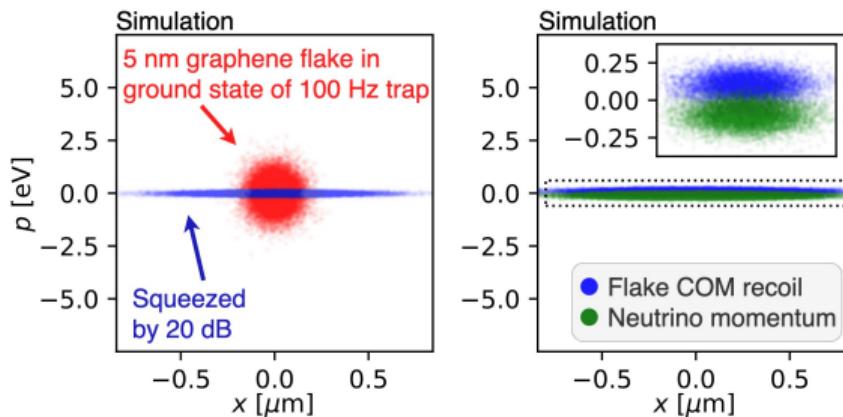


NEUTRINO MASS STUDIES WITH OPTICALLY LEVITATED NANOSPHERES

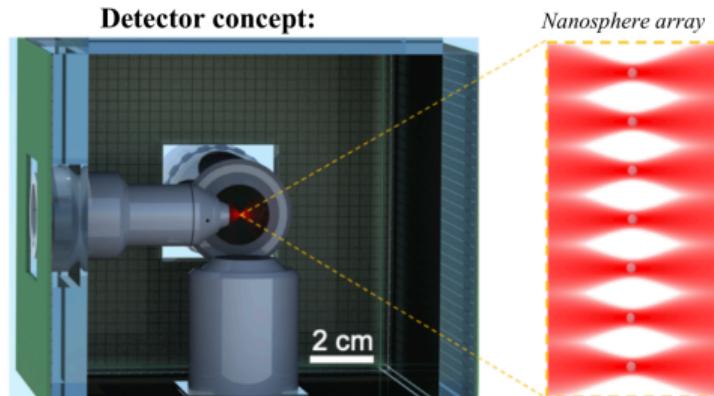
Particles with diameter $\mathcal{O}(100 \text{ nm})$ to $\mathcal{O}(10 \mu\text{m})$ can be optically trapped in a vacuum.

\vec{x}, \vec{p} measured by back-action of readout laser.

Near-future: dark matter limits, search for millicharged particles & sterile neutrinos.



Wouter Van De Pontseele



Squeezing uncertainty into \vec{x} might enable momentum detection $\mathcal{O}(100 \text{ meV})$ in the not-so-near future from graphene spheres doped with ultra-low Q isotopes.

CONCLUSIONS & OUTLOOK

FUTURE OF DIRECT m_β MEASUREMENTS & OUTLOOK

Standing on the shoulders of giants

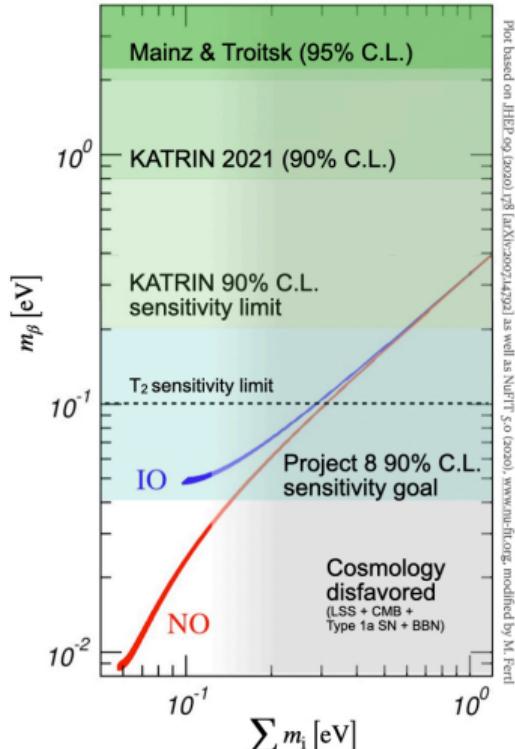
- Neutrino oscillations, theory improvements and collaboration with cosmology & $0\nu\beta\beta$ -decay searches.
- KATRIN's tremendous systematic reduction will lead to $m_\beta < 0.5 \text{ eV}$ limits soon.

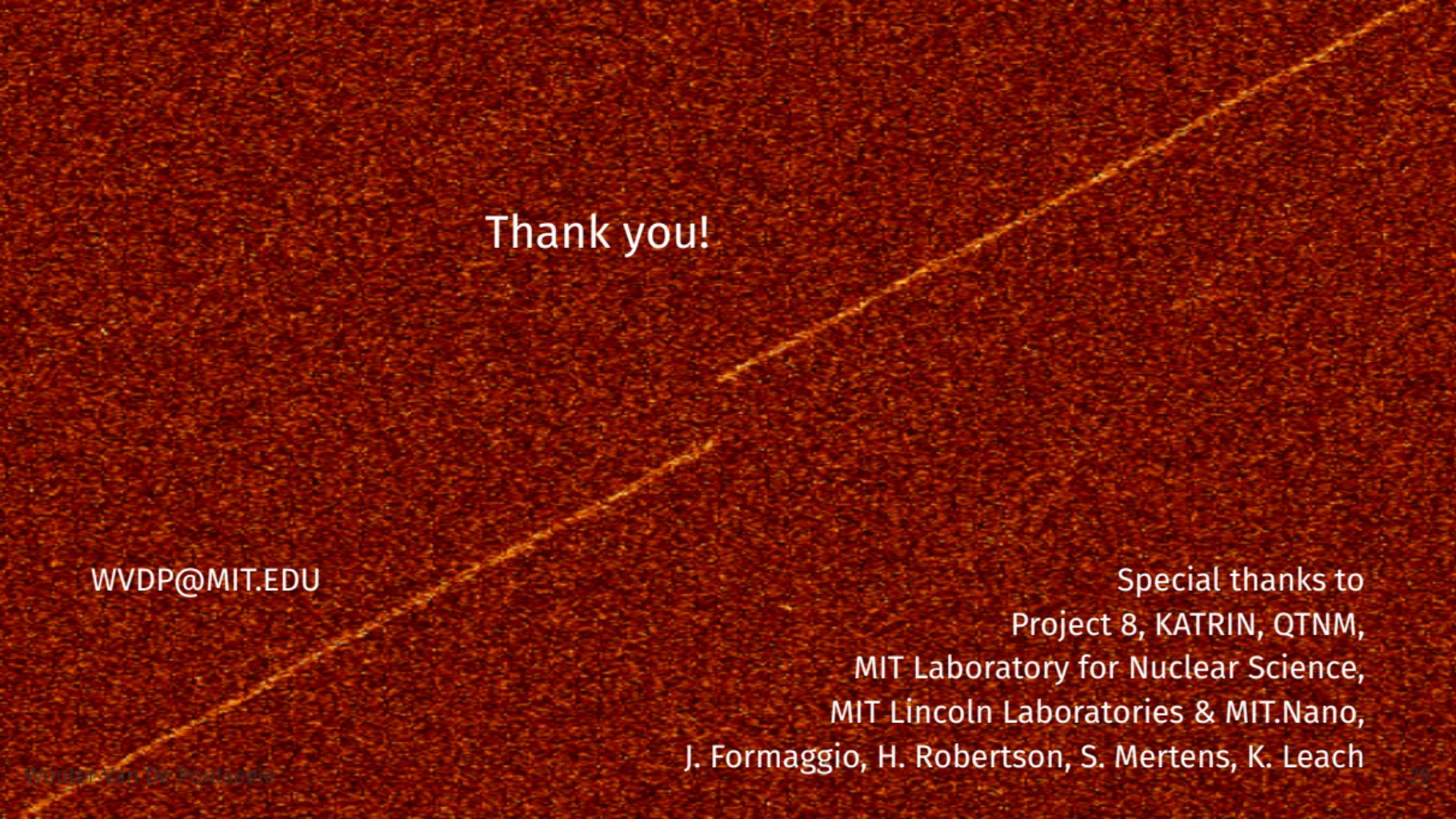
New kids on the block

- Cyclotron radiation, calorimetric spectrum of electron capture, and atomic tritium are progressing quickly.

...And the challenges ahead

- Plenty of demonstrator experiments to look forward to!





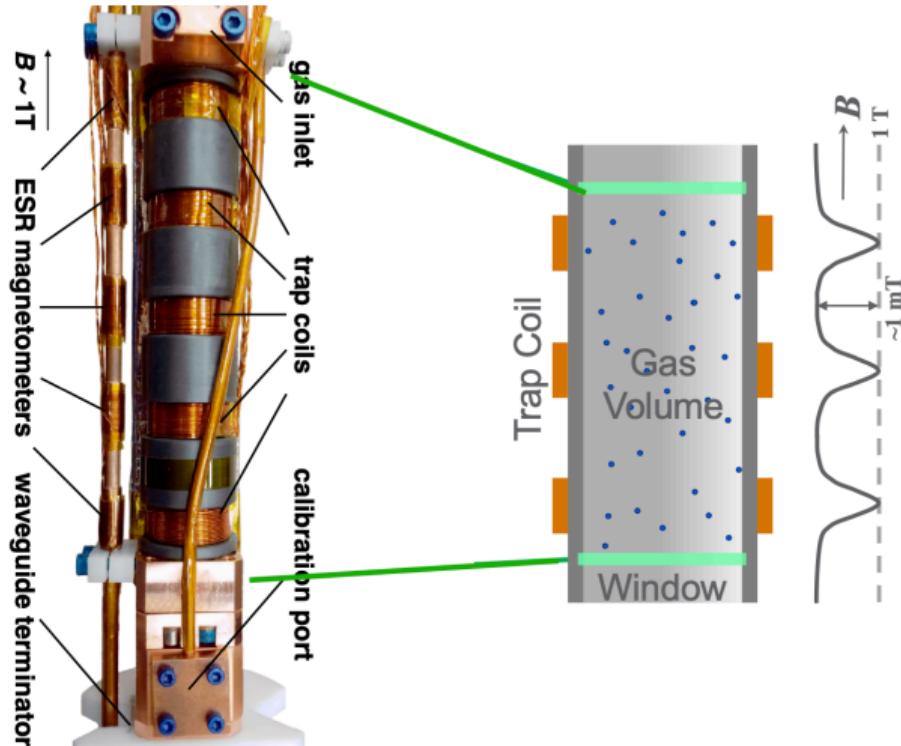
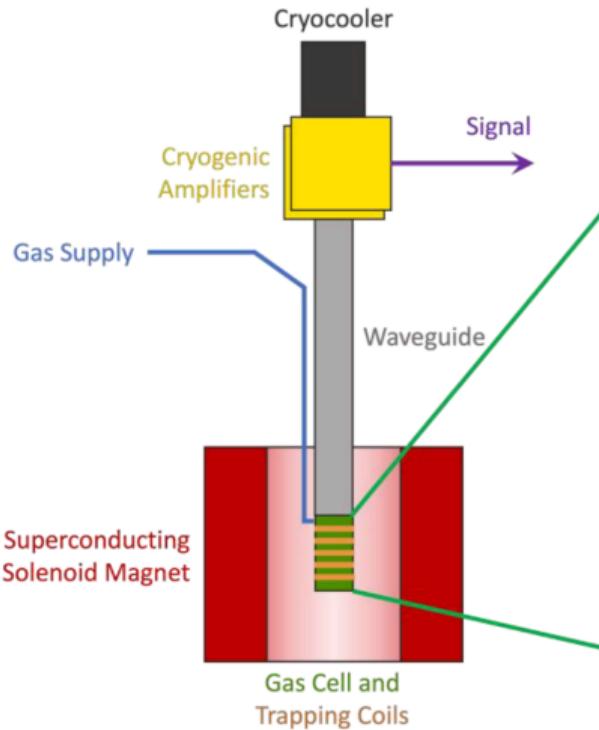
Thank you!

WVDP@MIT.EDU

Special thanks to

Project 8, KATRIN, QTNU,
MIT Laboratory for Nuclear Science,
MIT Lincoln Laboratories & MIT.Nano,
J. Formaggio, H. Robertson, S. Mertens, K. Leach

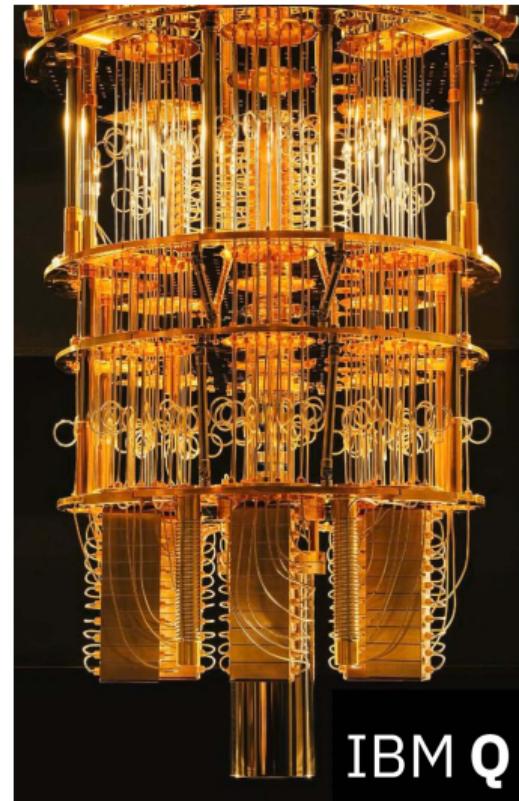
PROJECT 8: PHASE II



QUANTUM AMPLIFIERS FOR PROJECT 8: STATUS AND MOTIVATION

Driven by Superconducting qubit readout

- First stage amplifier limits performance.
- Bandwidth enables multiplexing qubits.



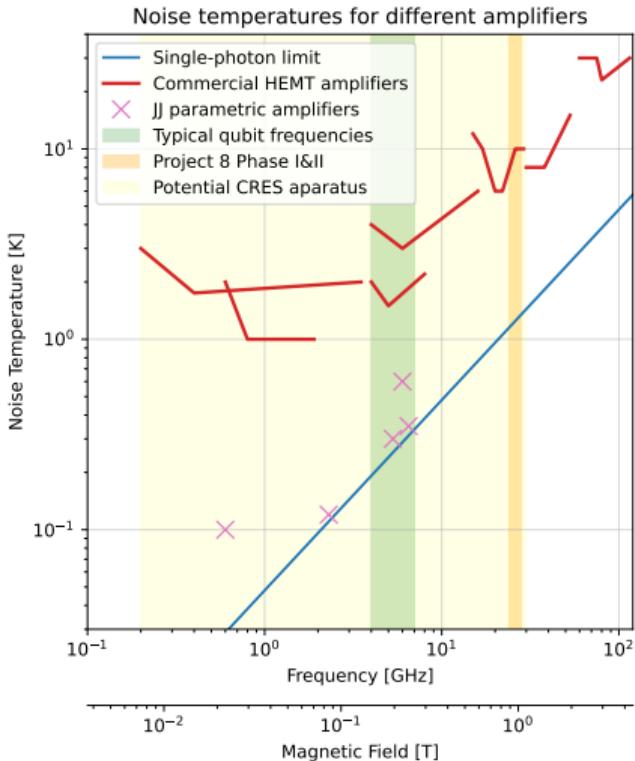
QUANTUM AMPLIFIERS FOR PROJECT 8: STATUS AND MOTIVATION

Driven by Superconducting qubit readout

- First stage amplifier limits performance.
- Bandwidth enables multiplexing qubits.

Essential for experiments detecting microwaves

- Lowest noise to detect CRES signal.
- Wide frequency range for multiplexing arrays.



QUANTUM AMPLIFIERS FOR PROJECT 8: STATUS AND MOTIVATION

Driven by Superconducting qubit readout

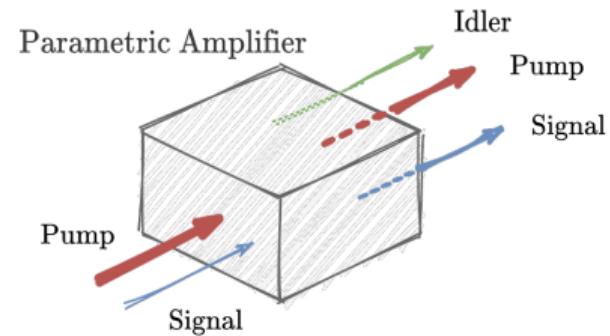
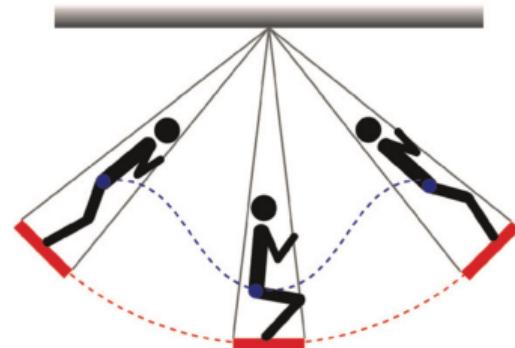
- First stage amplifier limits performance.
- Bandwidth enables multiplexing qubits.

Essential for experiments detecting microwaves

- Lowest noise to detect CRES signal.
- Wide frequency range for multiplexing arrays.

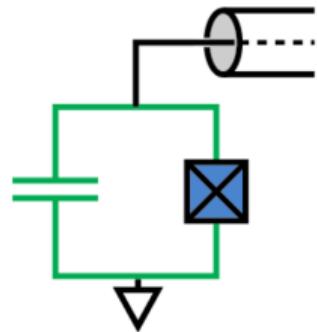
Parametric amplification: $2\omega_p = \omega_s + \omega_i$

- Signal amplification by **exchanging pump power** using the non-linearity of Josephson Junctions.

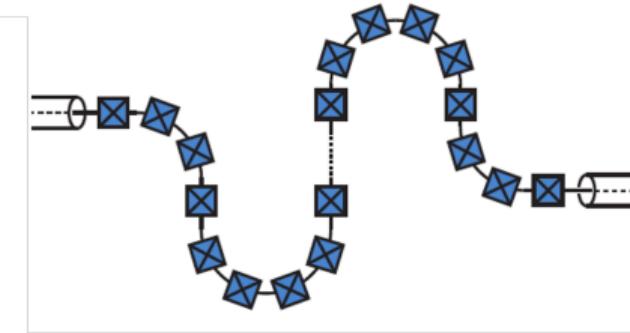


PARAMETRIC AMPLIFICATION: SINGLE CELL VERSUS TRAVELLING WAVES

JPA

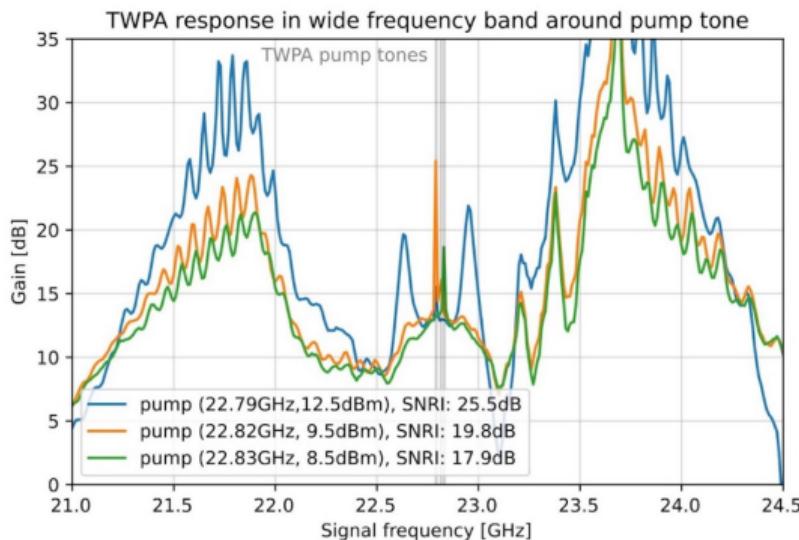


JTWPA



- Few spatial modes – **Cavity** style
- Near **ideal quantum efficiency**
- Small bandwidth of $\mathcal{O}(10 \text{ MHz})$.

- Cell size $< \lambda$
⇒ Create nonlinear meta-material.
- Many spatial modes
⇒ **Transmission line** behaviour.
- Up to $\mathcal{O}(3 \text{ GHz})$ bandwidth.



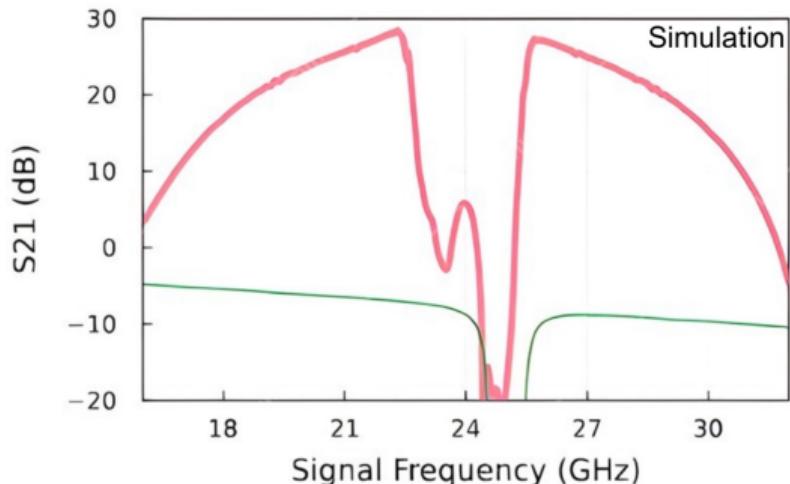
TWPAs fabricated & measured in three frequency bands

| | |
|----------------|--|
| Low Frequency | $\mathcal{O}(0.5 \text{ GHz to } 1 \text{ GHz})$ |
| Mid Frequency | $\mathcal{O}(5 \text{ GHz to } 9 \text{ GHz})$ |
| High Frequency | $\mathcal{O}(17 \text{ GHz to } 26 \text{ GHz})$ |

- At high frequencies, **insertion losses** and package modes matter.
- **Gain ripples** caused by impedance mismatches.
- **Compression power limits** multiplexing applications.

JTWPA improvements on the way

- Packaging improvements at high frequency:
New launch port and connectors.
Redesigned geometry, reducing package modes and reflections.
- Tri-layer aluminium process:
Stable oxides reduce insertion loss.
- Smoother gain profile and better impedance matching conditions.



Next fabrication round being packaged!

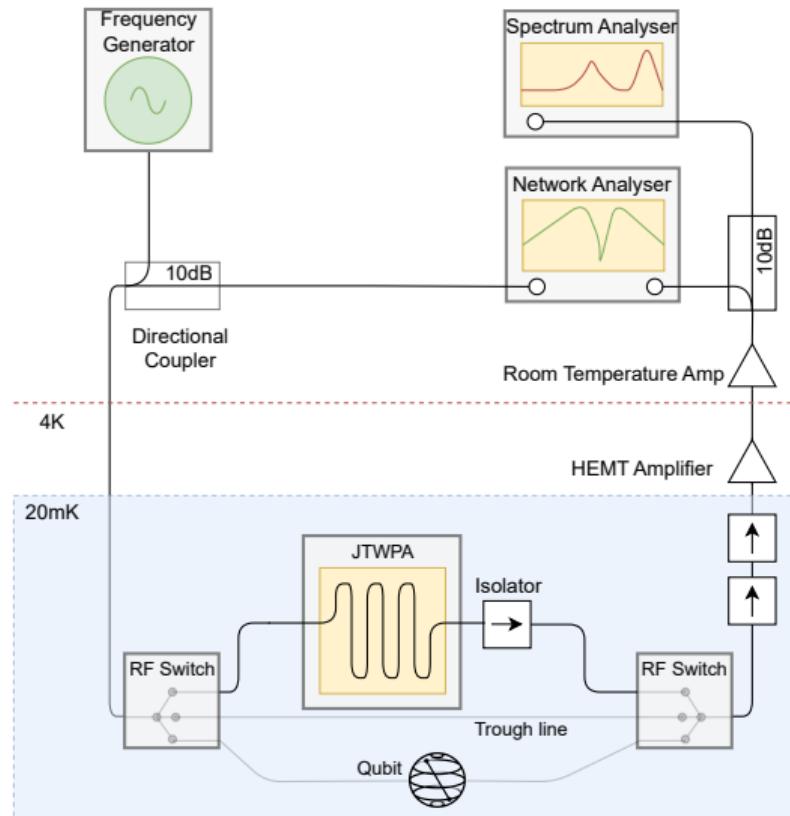
Signal and Noise modelling

- Axion haloscopes and CRES experiment signals propagate through and reflect at microwave components.

Signal and Noise power calibration

Relies on a **cryogenic candle**:

- Hot-Cold thermal noise source measurement
- Waveguide coupled Qubit



Signal and Noise modelling

- Axion haloscopes and CRES experiment signals propagate through and reflect at microwave components.

Signal and Noise power calibration

Relies on a **cryogenic candle**:

- Hot-Cold thermal noise source measurement
- **Waveguide coupled Qubit**

Enables photon flux measurement along the transmission line!

