# CONUS experiment: recent results and prospects for reactor CEvNS research

Kaixiang Ni – On behalf of the CONUS collaboration



NuPhys2023: Prospects in Neutrino Physics King's College London, Dec. 2023

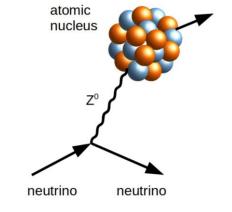
### Coherent elastic neutrino nucleus scattering (CEvNS)

"Coherent": neutrino interacts with the nucleus as a whole

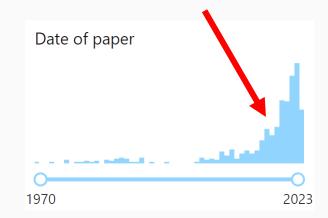
- Standard model predicted: *D. Freedman, PRD 9 1389 (1974)*
- Relatively large cross section:  $\sim 10^{-16}$  barn
- Low energy scale: several keV of recoil energy

$$\frac{d\sigma}{dT} = \frac{G_f^2}{4\pi} \left( N - \left( 1 - 4\sin^2\theta_w \right) Z \right)^2 F^2(q^2) M(1 - \frac{MT}{2E_v^2})$$

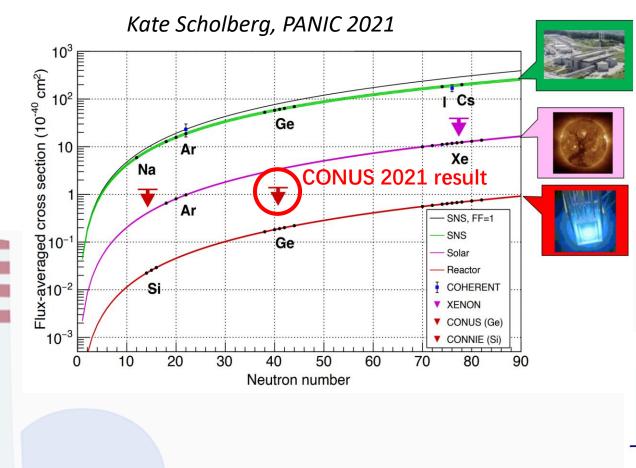
- ✓ Verification of Standard model (e.g. Weinberg angle) at low energy
- ✓ BSM searches for new neutrino interactions: magnet moments, millicharged, etc.
- ✓ Insight of nuclear structure (nuclear matrix, etc.)
- $\checkmark\,$  Supernova neutrinos, "neutrino floor" in dark matter search



#### Detection of CEvNS by COHERENT! Science 357 (2017) 6356, 1123-1126



### Neutrino sources



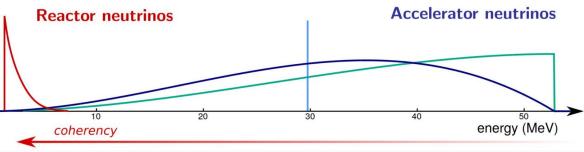
- Different neutrino flavors tected by 000 2000  $E_{v} \sim 20-50 \text{ MeV} (F < 1)$ , less coherency Pulsed flux, higher S/N ratio r v (<sup>8</sup>B v):

#### Solar v (<sup>8</sup>B v):

- By-product of dark matter experiments
- No on-off comparison

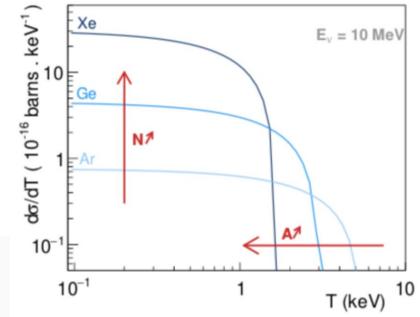
#### Reactor v

- Inside fully coherent regime
- **Complementary** to the accelerator experiments ٠



### Reactor experiments worldwide

reactor	CONUS	Brokdorf: 3.9 GW	17 m	4 MeV	Ge	4 kg
	CONUS+	Leibstadt: 3.6 GW	21 m	4 MeV	Ge	4 kg (upgraded)
	VGEN <sup>O</sup>	Kalinin: 3.1 GW	11-12 m	4 MeV	Ge	1.4 kg→6 kg
	RED-100 <sup>0</sup>	Kalinin: 3.1 GW	19 m	4 MeV	Xe	160 kg
	NCC-1701 T	Dresden-II: 3 GW	10.4 m	4 MeV	Ge	3 kg
	NCC-1701 <sup>P</sup>	Ringhals: 3.6 GW	23 m	4 MeV	Ge	3 kg (upgraded)
	RECODE	Sanmen: 3.4 GW	25 m	4 MeV	Ge	2x5 kg
	RELICS	Sanmen: 3.4 GW	25 m	4 MeV	Xe	30 kg
	N.N. <sup>P</sup>	Taishan: 4.6 GW	35 m	4 MeV	Ar	100 kg
	CONNIE	Angra: 3.9 GW	30 m	4 MeV	Si	50 g
	NEON	Hanbit: 2.8 GW	24 m	4 MeV	Nal	15kg
	TEXONO	Kuosh.: 2.9 GW	25 m	4 MeV	Ge	1-2 kg
	MINER	TAMU: 1 MW	2-3 m	4 MeV	Ge/Si	1 kg
	Richochet AP	MIT R: 5.5 MW	4 m	4 MeV	Ge, Zn	5 kg, 5 kg
	Richochet BP	ILL: 58 MW	8.8 m	4 MeV	Ge	1 kg
	Basket	Chooz: 8.6 GW	70-400 m	4 MeV	Li2WO4	1000
	ν-CLEUS <sup>P</sup>	Chooz: 8.6 GW	70-100 m	4 MeV	CaWO <sub>4</sub> . Al <sub>2</sub> O <sub>3</sub>	$1 g \rightarrow 10 kg$ $1 g \rightarrow 10 kg$



O/T: operational/terminated, P: in preparation, R= R&D

Choose of target materials:

- Larger atomic number: higher cross section, but larger quenching and lower threshold
- Technologies applied to detect energy at sub-keV level
- Target mass as kg scale is enough to see CEvNS.

### The CONUS/CONUS+ collaboration







Max-Planck-Institut für Kernphysik (MPIK), Heidelberg: N. Ackermann, S. Armbruster, A. Bonhomme, H. Bonet, C. Buck, J. Hakenmüller, J. Hempfling, G. Heusser, M. Lindner, W. Maneschg, K. Ni, T. Rink, E. Sanchez-Garcia, H. Strecker
Former collaborators: T. Schierhuber, E. Van der Meeren, J. Henrichs, T. Hugle, J. Stauber
Preussen Elektra GmbH, Kernkraftwerk Brokdorf (KBR), Brokdorf: K. Fülber, R. Wink
Kernkraftwerk Leibstadt AG (KKL), Leibstadt: J. Wönckhaus, M. Rank

#### Scientific cooperations with:

Physikalisch-Technische Bundesanstalt (PTB), Braunschweig: R. Nolte, E. Pirovano, M. Reginatto, M. Zboril, A. Zimbal Paul-Scherrer-Institut (PSI), Villigen: E. Hohmann, R. Gaalev

### CONUS @ KBR

#### **Detector:**

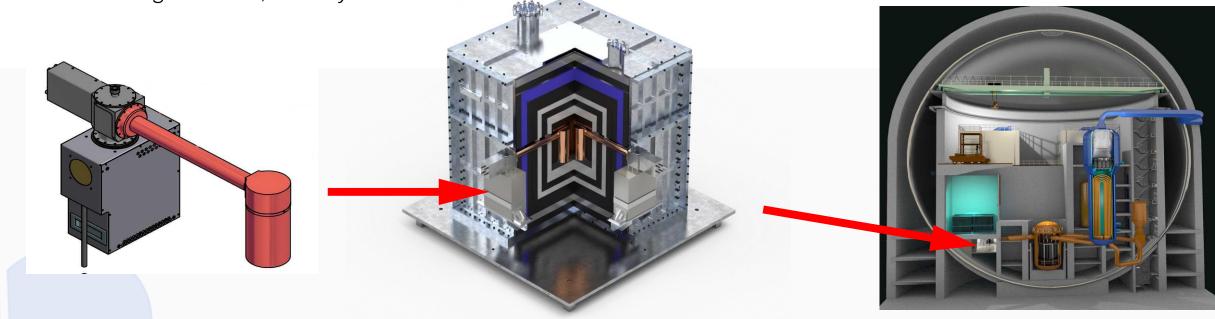
- Point-contact high purity Germanium crystal
- Electrical cryo-cooled
- Energy threshold: 200-300eV
- Four 1-kg modules, in array

#### Shield:

- Lead + Polyethylene + active muon veto scintillators
- Volume: 1.65m<sup>3</sup>, mass: 11 tons
- Total bkg suppression:  $>10^4 \text{ x}$

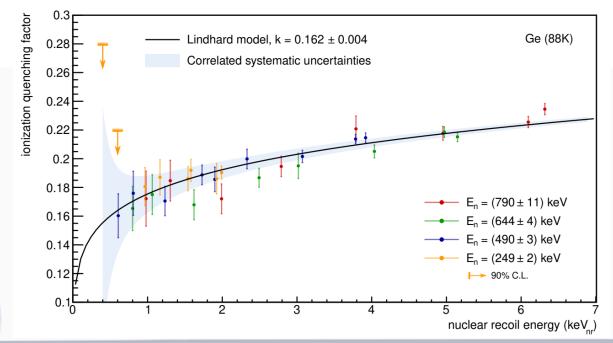
#### **Experiment site:**

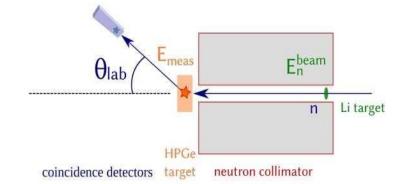
- 3.9GW thermal power
- 17m distance to the reactor core, 2.3 x  $10^{13} v/s/cm^2$
- 24m w.e. overburden

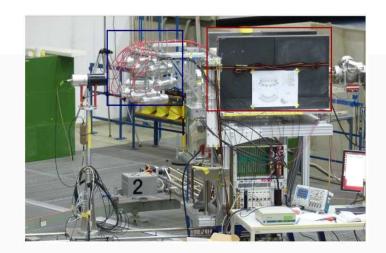


## Quenching measurement

- Only part of the nuclear recoil energy could turn into detectable signal
   → quenching
- Most commonly used model: Lindhard model, with unknown parameter k
- Auxiliary measurement done with neutron beams at PTB, Germany
- k=0.162+-0.004 (stat.+syst.)







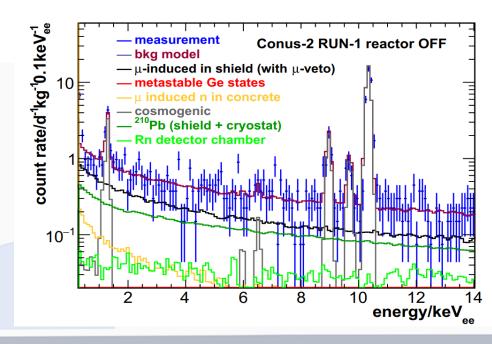
Eur. Phys. J. C 82, 815 (2022)

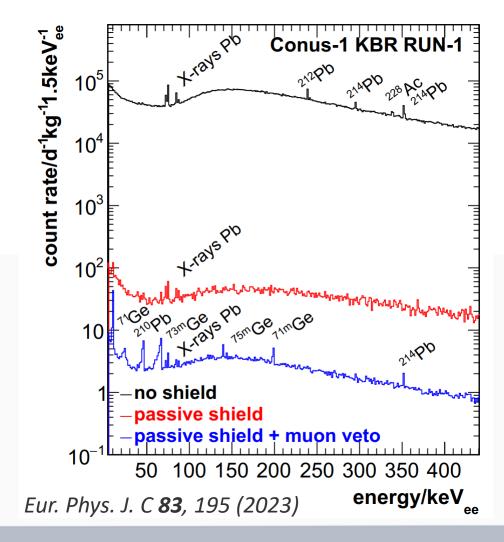
2023/12/20

#### NuPhys2023: Prospects in Neutrino Physics

### **Background estimation**

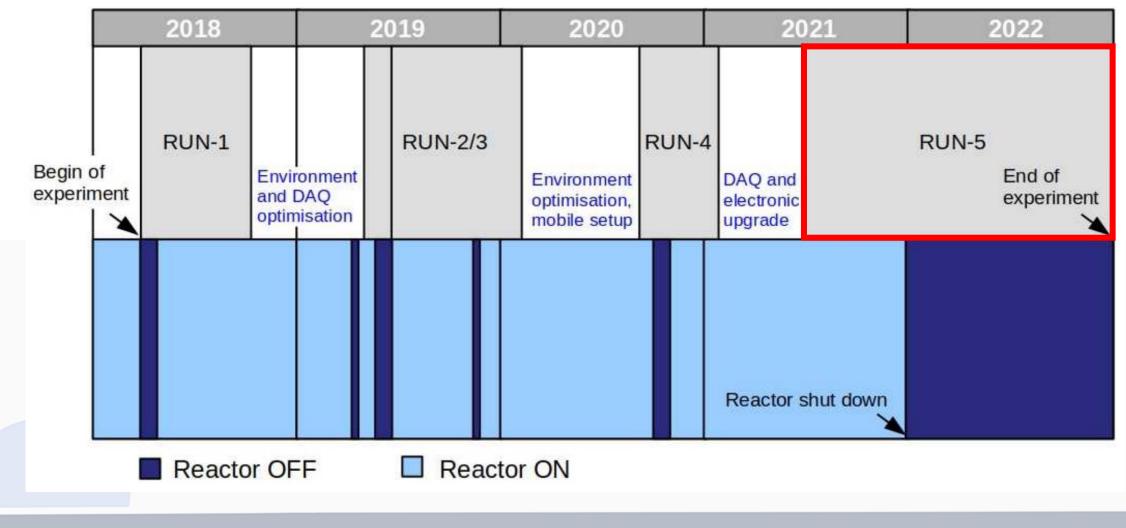
- Suppression factor by shield: >10<sup>4</sup>
- Remaining bkg rate in ROI: O(10) cts/d/kg
- Bkg is dominated by muon-induced events and <sup>210</sup>Pb events
  - Reactor neutron/activation negligible



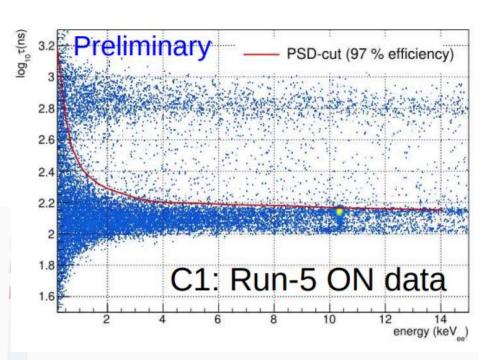


#### NuPhys2023: Prospects in Neutrino Physics

### Runs

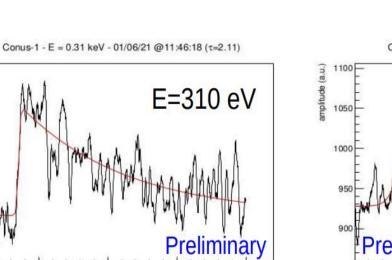


### Run 5 with Pulse Shape Discrimination (PSD)



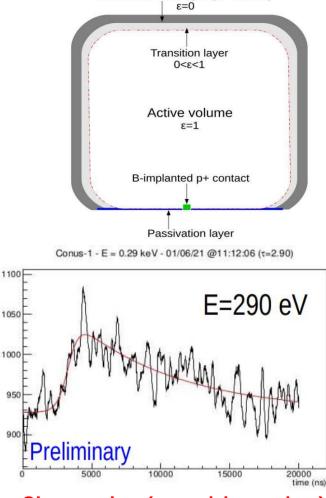
**Efficiency:** remove ~50% of the surface events at ~300eV with >90% bulk event acceptance

- Energy deposition near the transition layer contributes to a slow signal
- Removing slow pulses could reduce surface background, while losing a little effective mass



#### Normal (fast) pulse

10000



Li-diffused dead layer (n+ contact)

Slow pulse (transition edge)

#### 2023/12/20

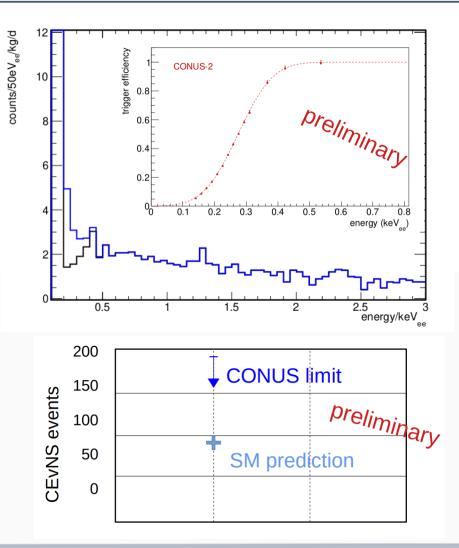
5000

1050

### New limits from Run5

- Upper limit: factor ~2 above SM prediction
- Strongest limit on reactor CEvNS count rate! (Assuming Lindhard quenching)
- Publication in preparation

Detector	Exposure (ON/OFF, kg-d)	Threshold (eV)	Anticipated Signals (k=0.16)	Likelihood fit
C1	142/40	210	42	<59
C2	146/130	210	(26	<75
C4	139/102	210 210 mina relimina	24	<90
Total	426/272	oreli	92	<163
		K		



### Comparison with other experiments

#### Current results from reactor CEvNS experiments:

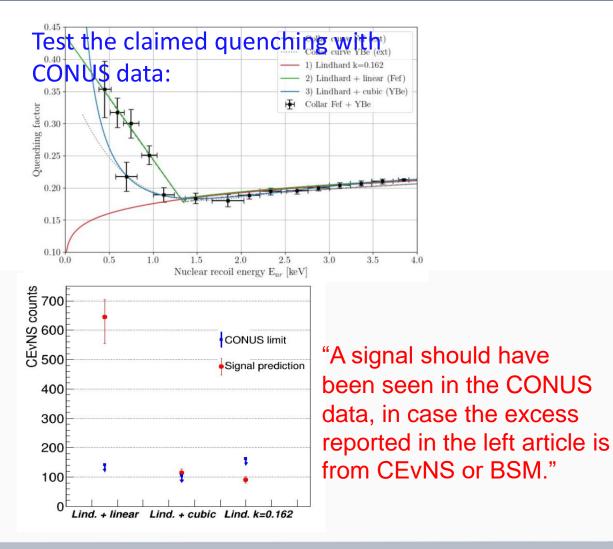
- constraints from vGen, CONNIE,...
- strong signal preference with NCC-1701 at Dresden-II reactor US:

#### Abstract of Phys. Rev. Lett. 129, 211802 (2022)

The 96.4 day exposure of a 3 kg ultralow noise germanium detector to the high flux of antineutrinos from a power nuclear reactor is described. A very strong preference  $(p < 1.2 \times 10^{-3})$  for the presence of a coherent elastic neutrino-nucleus scattering (CE $\nu$ NS) component in the data is found, when compared to a background-only model. No such effect is visible in 25 days of operation during reactor outages. The best-fit CE $\nu$ NS signal is in good agreement with expectations based on a recent characterization of germanium response to sub-keV nuclear recoils. Deviations of order 60% from the standard model CE $\nu$ NS prediction can be excluded using present data. Standing uncertainties in models of germanium quenching factor, neutrino energy spectrum, and background are examined.

#### Abstract of Phys. Rev. D 103, 122003 (2021)

Germanium is the detector material of choice in many rare-event searches looking for low-energy nuclear recoils induced by dark matter particles or neutrinos. We perform a systematic exploration of its quenching factor for sub-keV nuclear recoils, using multiple techniques: photoneutron sources, recoils from gamma-emission following thermal neutron capture, and a monochromatic filtered neutron beam. Our results point to a marked deviation from the predictions of the Lindhard model in this mostly unexplored energy range. We comment on the compatibility of our data with low-energy processes such as the Migdal effect, and on the impact of our measurements on upcoming searches.



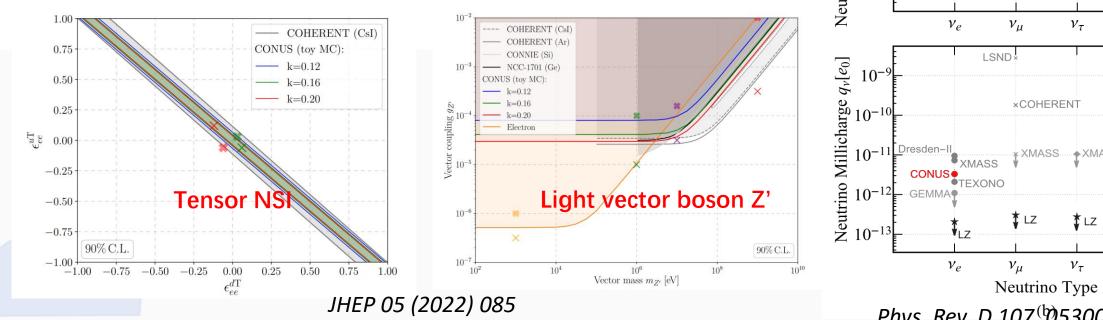
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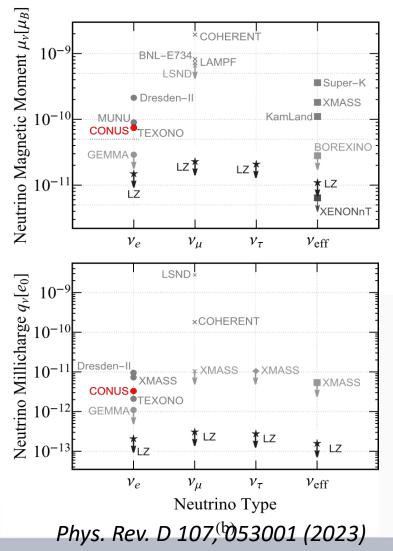
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### **BSM** results

**Tensor/Vector NSI** (non-standard interactions): limits the coupling parameter space

**Light vector boson**: limits the mass-coupling parameter space **Neutrino millicharged**:  $|q_v| < 3.3 \times 10^{-12} e_0$ **Neutrino magnetic moment**:  $\mu_v < 7.5 \times 10^{-11} \mu_B$ 





NuPhys2023: Prospects in Neutrino Physics

## From CONUS to CONUS+

- Nuclear power plants in Germany are shut-down…
- Our new home: Kernkraftwerk Leibstadt (KKL), Switzerland
  - Experiment hall: ~21m from 3.6 GW reactor core, 1.45 x  $10^{13} v/s/cm^2$
  - New environmental background characterized, large differences observed



Parameter	Method	CONUS+ vs. CONUS	
Gamma-radiation (>3 MeV)	Low bg. Ge spectr. CONRAD	25x smaller	
Cosmic muon flux	Liquid scintillator cells	2.2x larger (critical)	
Neutron spectrometry	Bonner Spheres from MPIK + PSI	30x larger (still subdominant)	
Temperture, Radon conc	diff. sensors	similar	

COvUS+

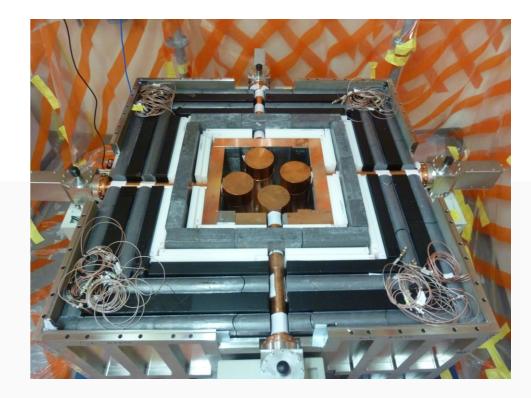
### Upgraded experiment

#### • Ge detector refurbishment:

- Reduced point-contact size
- Higher trigger efficiency ASIC
- Water-cooled system to reduce vibration and microphonic noise

#### • 2<sup>nd</sup> muon veto

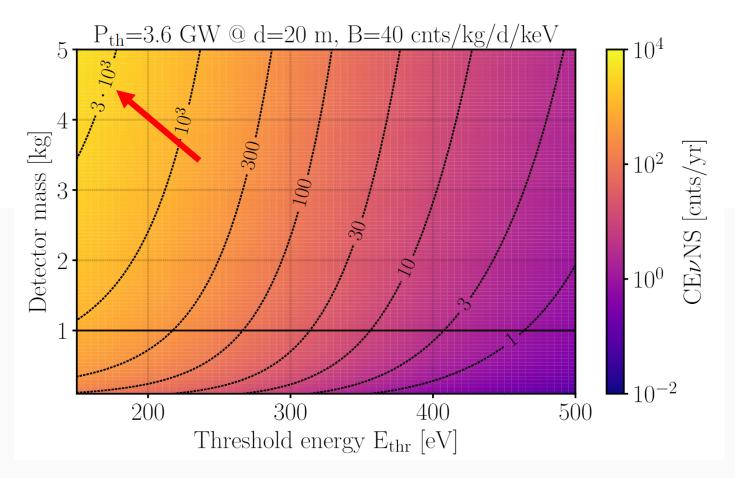
- Higher coverage, less gamma interference
- Independent trigger on each scintillator, coincidence available
- Real-time remote control
- NOW ongoing physical data taking!



### Future prospects

For CONUS+, a threshold down to <200eV will hopefully provide >10<sup>3</sup> CEvNS events per year!

Global pursuit of reactor CEvNS detection has just begun, aiming for higher detector mass and lower threshold.

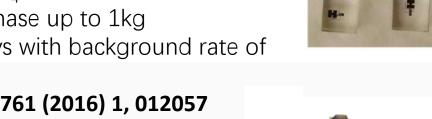


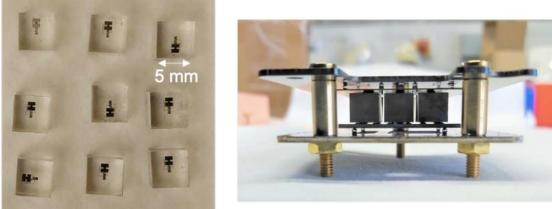
### Towards lower threshold

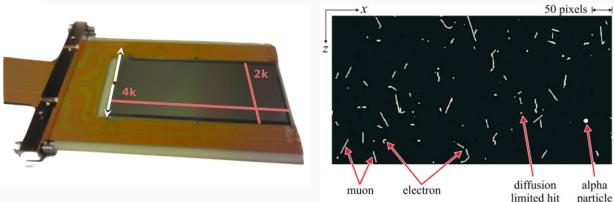
#### With new detection technology!

#### NUCLEUS: transition edge sensors(TES) J.Low Temp.Phys. 199 (2019) 1-2, 433-440

- Energy threshold: O(10 eV)
- Target material: Al<sub>2</sub>O<sub>3</sub>/CaWO<sub>4</sub>
- Target mass: 10g, second phase up to 1kg
- ightarrow Reach 5 $\sigma$  in less than 40 days with background rate of
- 100 counts/(keV·kg·day) CONNIE: CCD J.Phys.Conf.Ser. 761 (2016) 1, 012057
- Energy threshold: 50eV
- Target material: Si
- Target mass: 50g
- ightarrow Additional tracking ability
- → Reached limit: <551 counts/keV/kg/day, 39 times larger than the standard model expectation







### Towards larger target mass

At O(100)kg level, ordinary crystals may not be available. The alternative is to use noble liquid detectors.

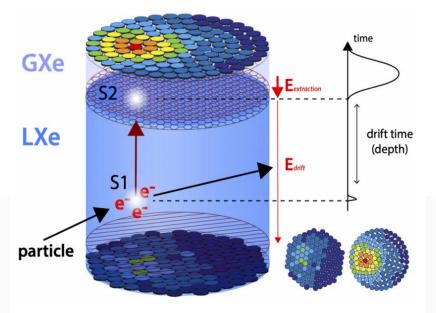
- Well developed technique thanks to the dark matter community
- Happens both scintillation/ionization, helps to discriminate electron/nuclear recoils
- Delayed single electron signal might be the dominant background

Experiments:

- Currently running: RED-100.
- More proposals/R&Ds came out recently: RELICS, NUXE, etc.

Signal prediction of a xenon TPC with 30kg-yr exposure @  $10^{13}$   $v/s/cm^2$  flux: >5000

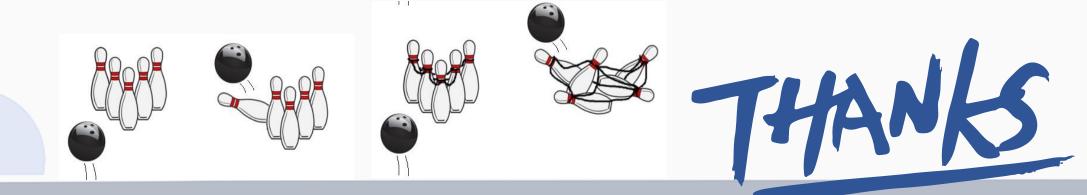
Also stay tuned to solar neutrino CEvNS detection by the dark matter experiments (PandaX-4T, LZ, XENONnT)!



the two-phase (LXe/GXe) time projection chamber (TPC), XENON Collaboration

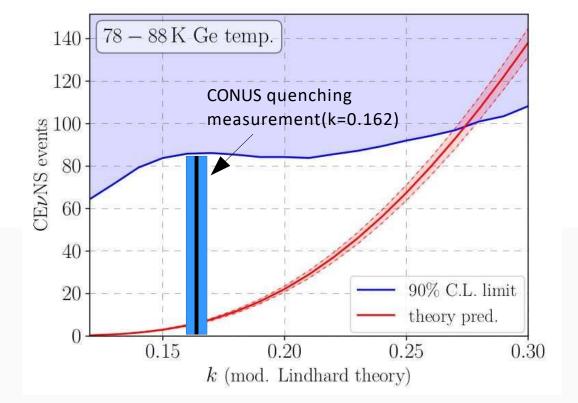
### Summary

- Reactor CEvNS is a validation for SM neutrino interaction at low energy limit, but is still not observed.
- The CONUS experiment was running in KBR to detect CEvNS until the end of 2022. Analysis of the full data set is completed and the preliminary result shows that *we are at the edge of making discovery.*
- The next phase of the experiment, CONUS+, is moved to KKL with improved Ge detector performance and adaptions to the new background composition.
- There is a global effort to search for CEvNS where various detection technologies beside HPGe are being developed. We are glad to see a more and more active CEvNS community!

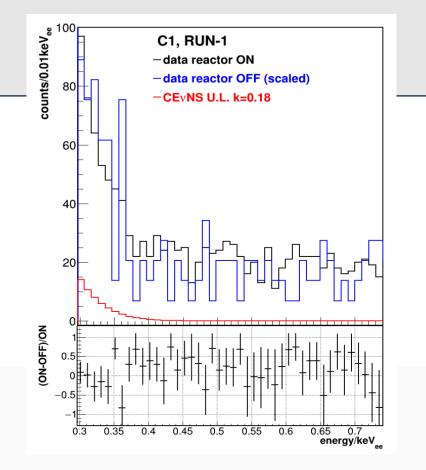








<85 events are detected Upper limit is still factor 17 higher than SM!

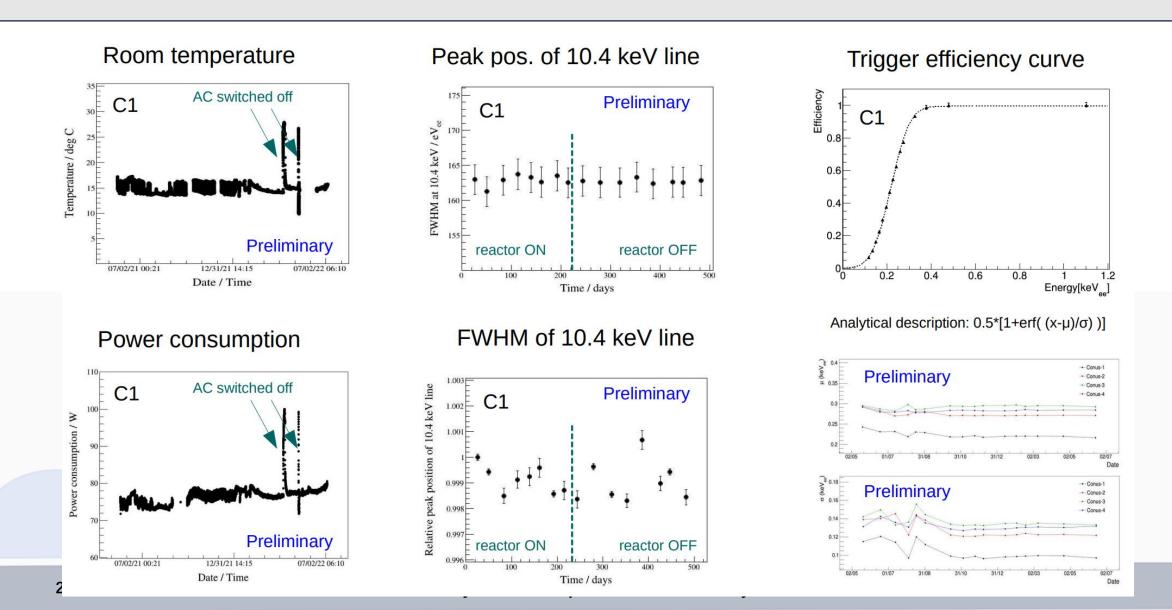


- Data: 248.7kg-d ON, 58.8kg-d OFF
- Threshold: ~300eV
- Binned Likelihood:
  - Simultaneously fit ON/OFF data
  - Poisson distribution in each bin

#### 2023/12/20

#### Phys.Rev.Lett. 126 (2021) 4, 041804 NuPhys2023: Prospects in Neutrino Physics

### Run stability (Run5)



### Reactor correlated neutron

- Neutron spectrometry with NEMUS detectors by PTB
- → Highly thermalized (>80%) and correlated with reactor thermal power
- Muon induced neutron takes the major role in CONUS background, instead of reactor neutron



