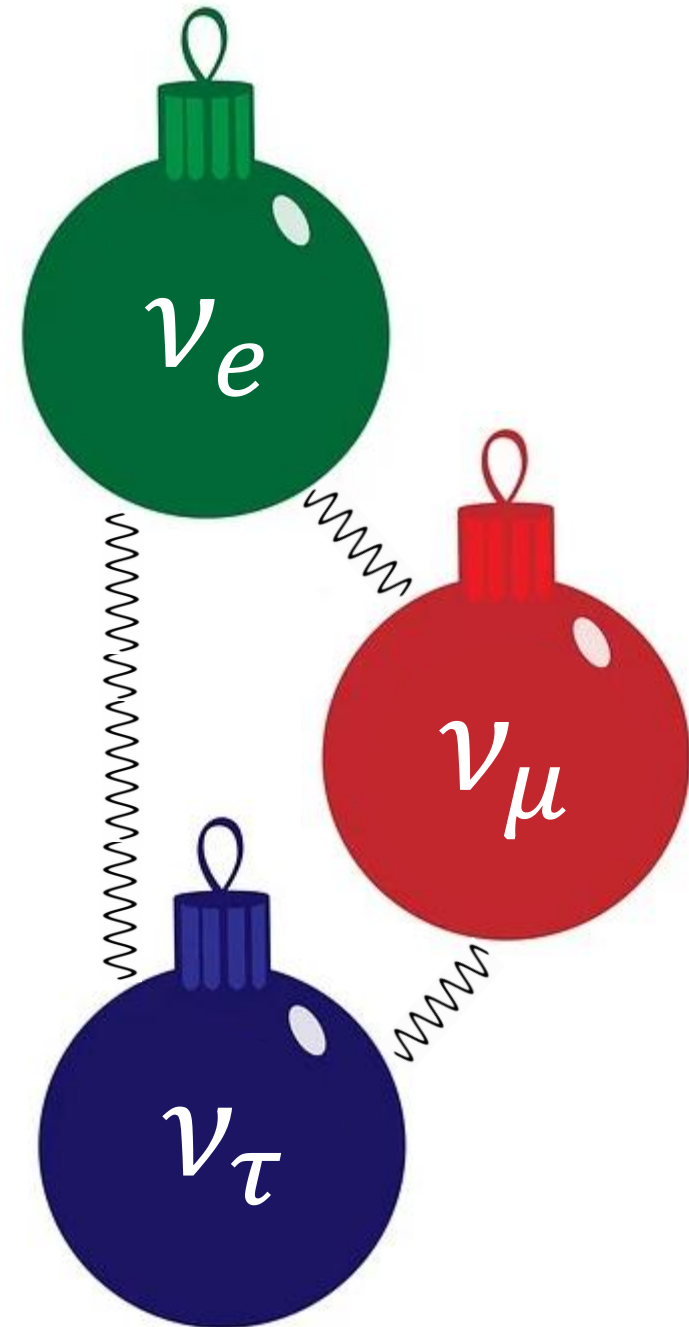


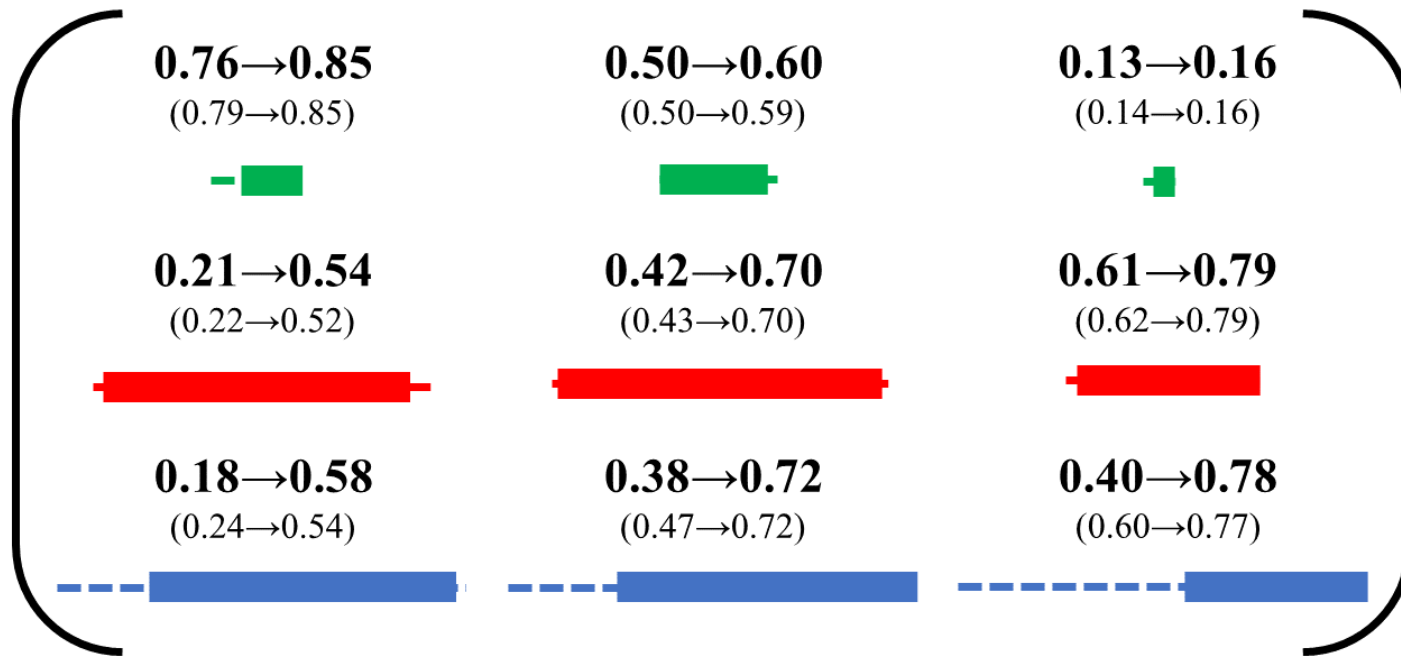
Improving Non-Unitary Limits on the Tau Row Matrix Elements Using Tau Neutrino Appearance

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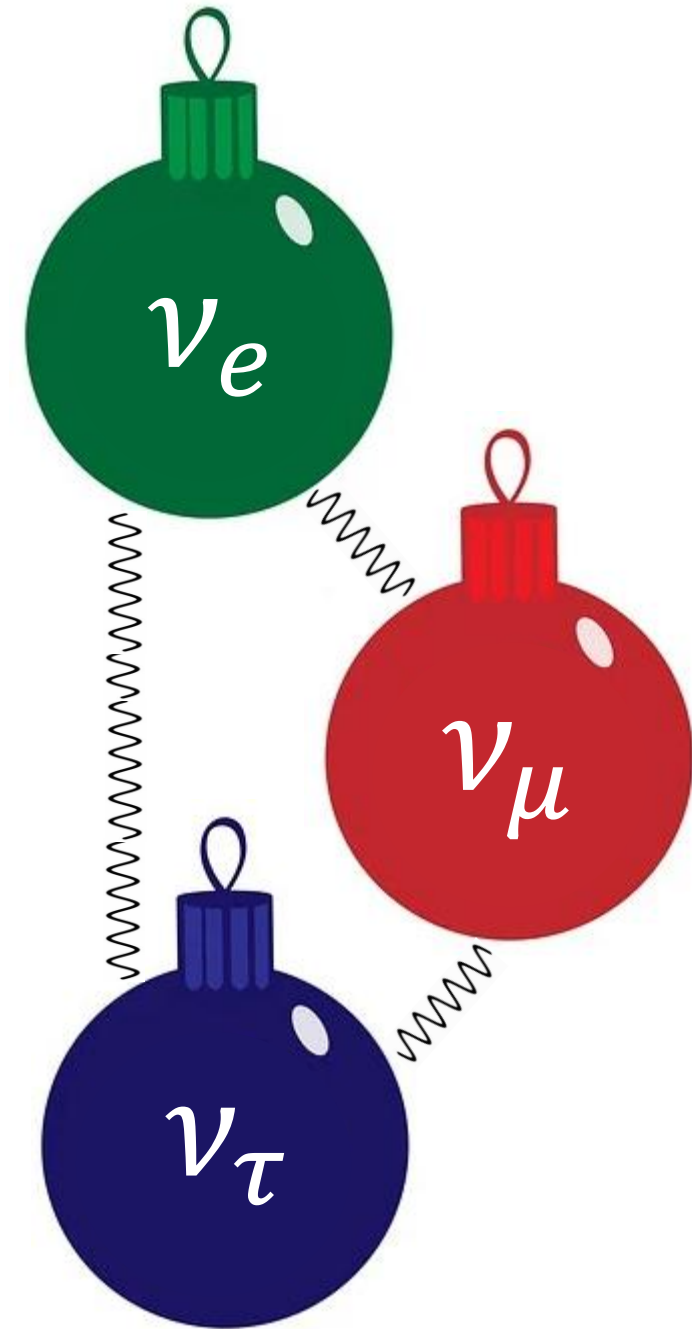


PMNS Matrix Limits

$$|U_{PMNS}|_{3\sigma} \begin{matrix} \text{w/o unitarity} \\ \text{(with unitarity)} \end{matrix} =$$

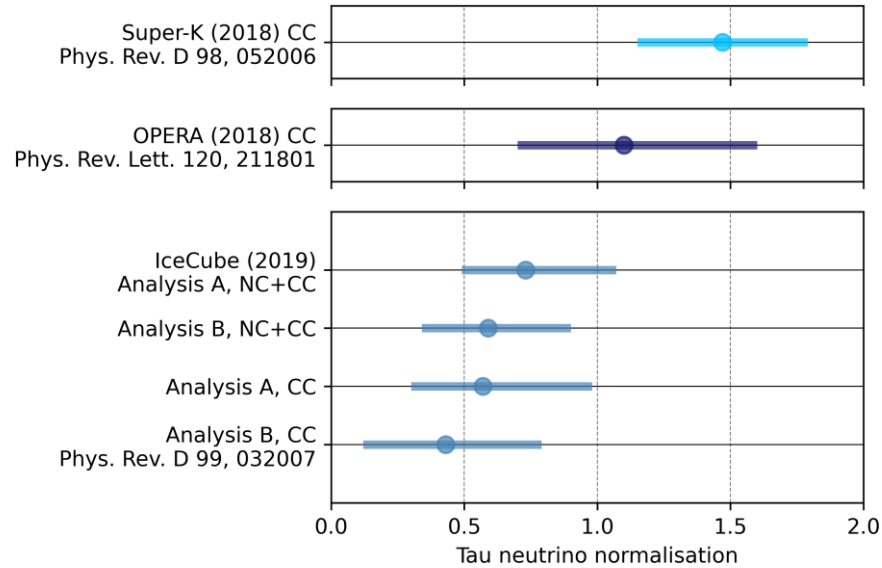
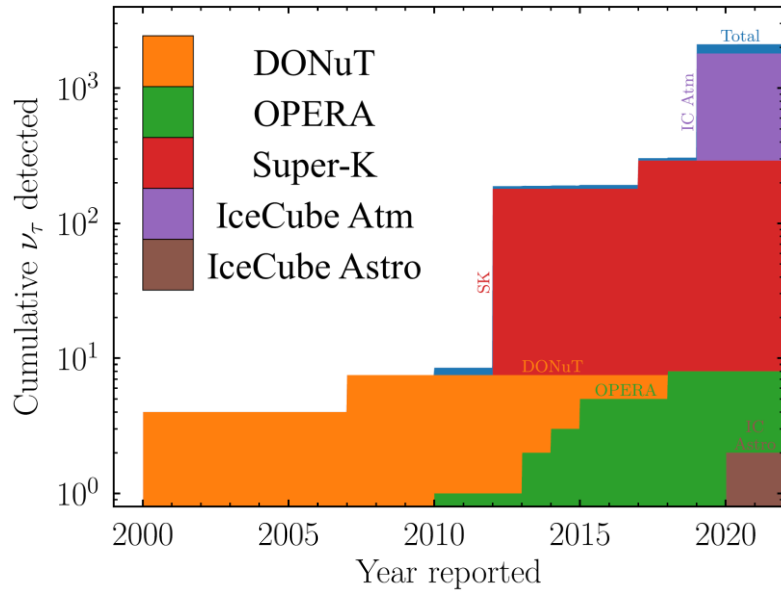


Best fit [1] limits for tau without unitarity constraint are poor



[1] Parke, S. and Ross-Lonergan, M. *Phys. Rev. D* 93, 113009 (2016) arXiv:1508.05095 [hep-ph]

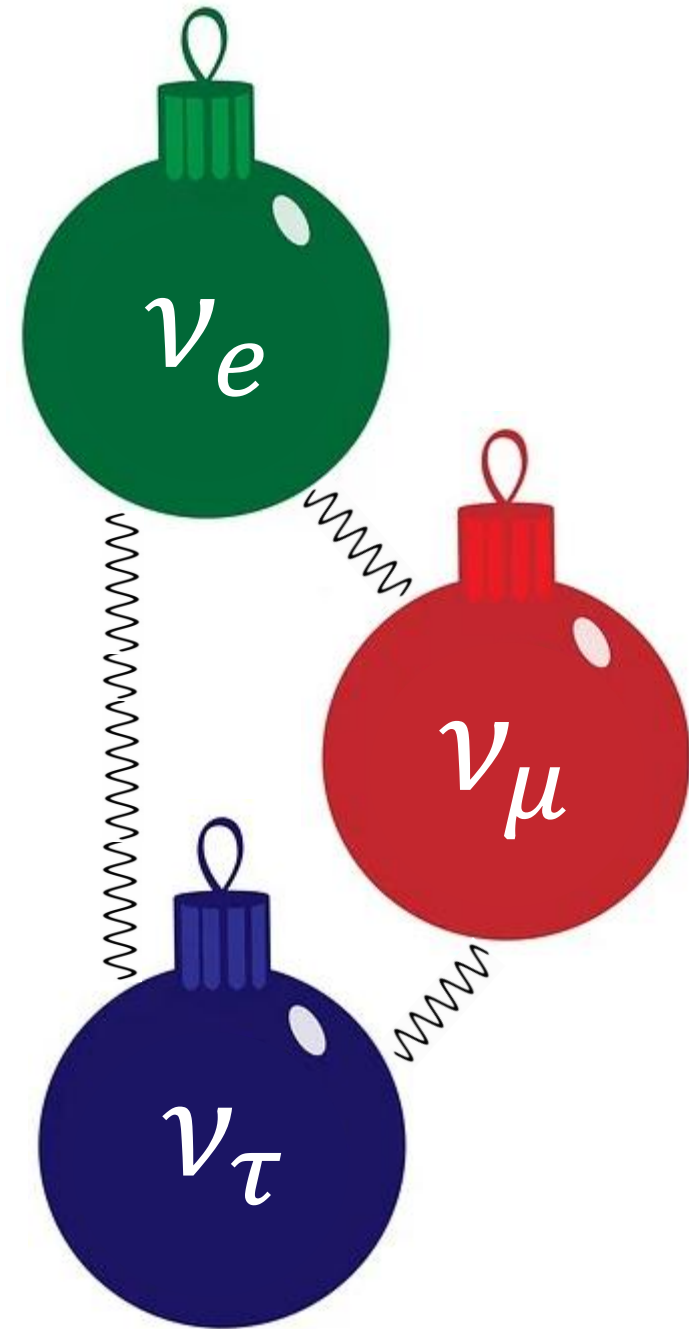
Tau Neutrinos Measured in Experiments



Measuring more and more tau neutrinos over the years [2]

[2] R. Mammen Abraham *et al.* *J. Phys. G* **49** (2022) 110501 [2203.05591].

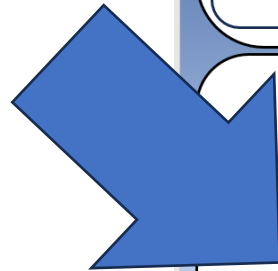
But they do not always agree with predictions...



What can we do?

Come find out at
EX-25!

Hint: it does not
interact with
Standard Model
forces



Improving Non-Unitary Limits on the Tau Row Matrix Elements Using Tau Neutrino Appearance

Rory Ramsden, King's College London (rory.ramsden@kcl.ac.uk)

Supervision by Francesca Di Lodovico



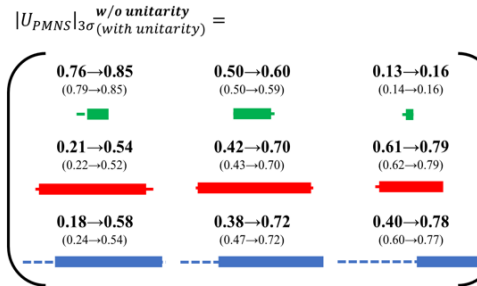
MATRIX ELEMENTS WITHOUT UNITARITY

The PMNS matrix, U_{PMNS} , governs neutrino mixing,

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}.$$

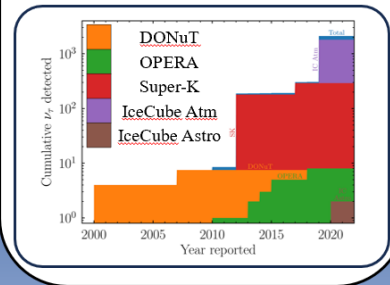
We can measure the elements from oscillation experiments and normally constrain them with unitarity; without unitarity, the constraints are poor.

Limits on the PMNS matrix elements at 3σ from [1].



TAU NEUTRINO SOURCES

Tau neutrinos are one of the least studied particles in the Standard Model due to difficult detection signatures. Only ~2000 ν_τ total have been observed at a few experiments [2].

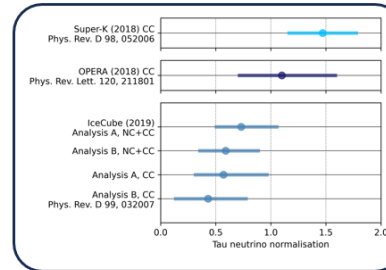


TAU NEUTRINO NORMALISATION

There are tensions in the number of tau neutrinos measured versus the number expected between different experiments,

$$N_\tau = \frac{\phi_{\nu_\tau}^{meas}}{\phi_{\nu_\tau}^{theo}}.$$

Sterile ν could possibly resolve high and low energy regimes.



REFERENCES

- [1] S. Parke, M. Ross-Lonergan. *Phys. Rev. D* **93** (2016) 113009 [1508.05095].
- [2] R. Mammen Abraham *et al.* *J. Phys. G* **49** (2022) 110501 [2203.05591].

