

# SEARCHES FOR THEORETICAL DM SIGNATURES WITH LATTICE FIELD THEORY

YANNICK DENGLER, FABIAN ZIERLER, AXEL MAAS

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ACHT

SEPTEMBER 27



# CONTENT

- Dark Matter
- Lattice Field Theory
- Dark Matter in Neutron Stars
- Scattering in  $Sp(4)$

Bertone et al.: Phys.Rept.405 (2005)

Rubin et al.: Ap.J.L. 225 (1978)

Chandra X-ray Observatory

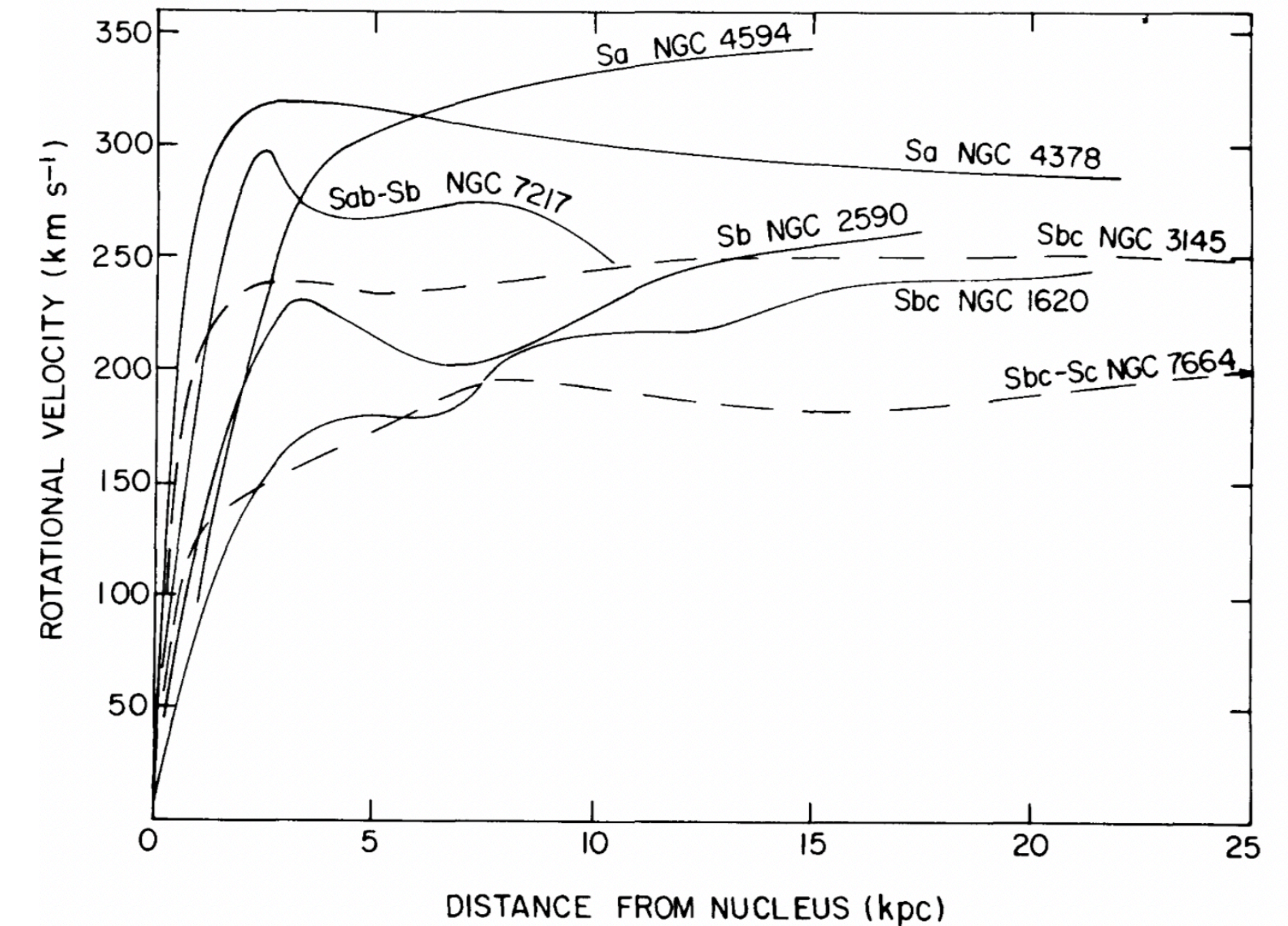
Tulin, Yu: arXiv:1705.02358 (2017)

Kaplinghat et al: Phys. Rev. Lett. 116 (2016)

# INTRODUCTION

# DARK MATTER

- Collection of astrophysical phenomena
- Rotational Curves
- Modified Gravity?

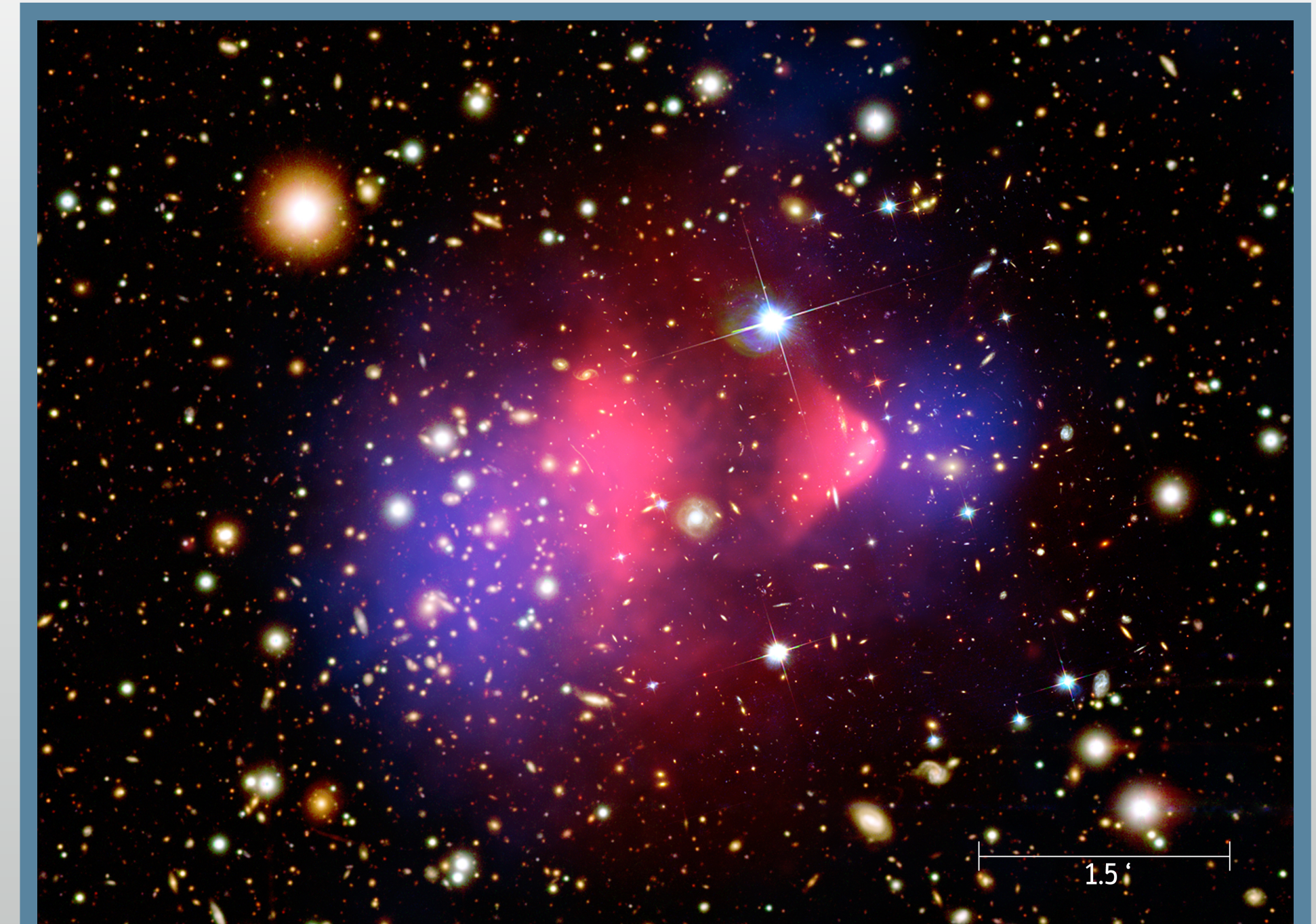


Rotational curves

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# DARK MATTER

- Collection of astrophysical phenomena
- Rotational Curves
- Particle picture preferred
  - No candidate from standard model

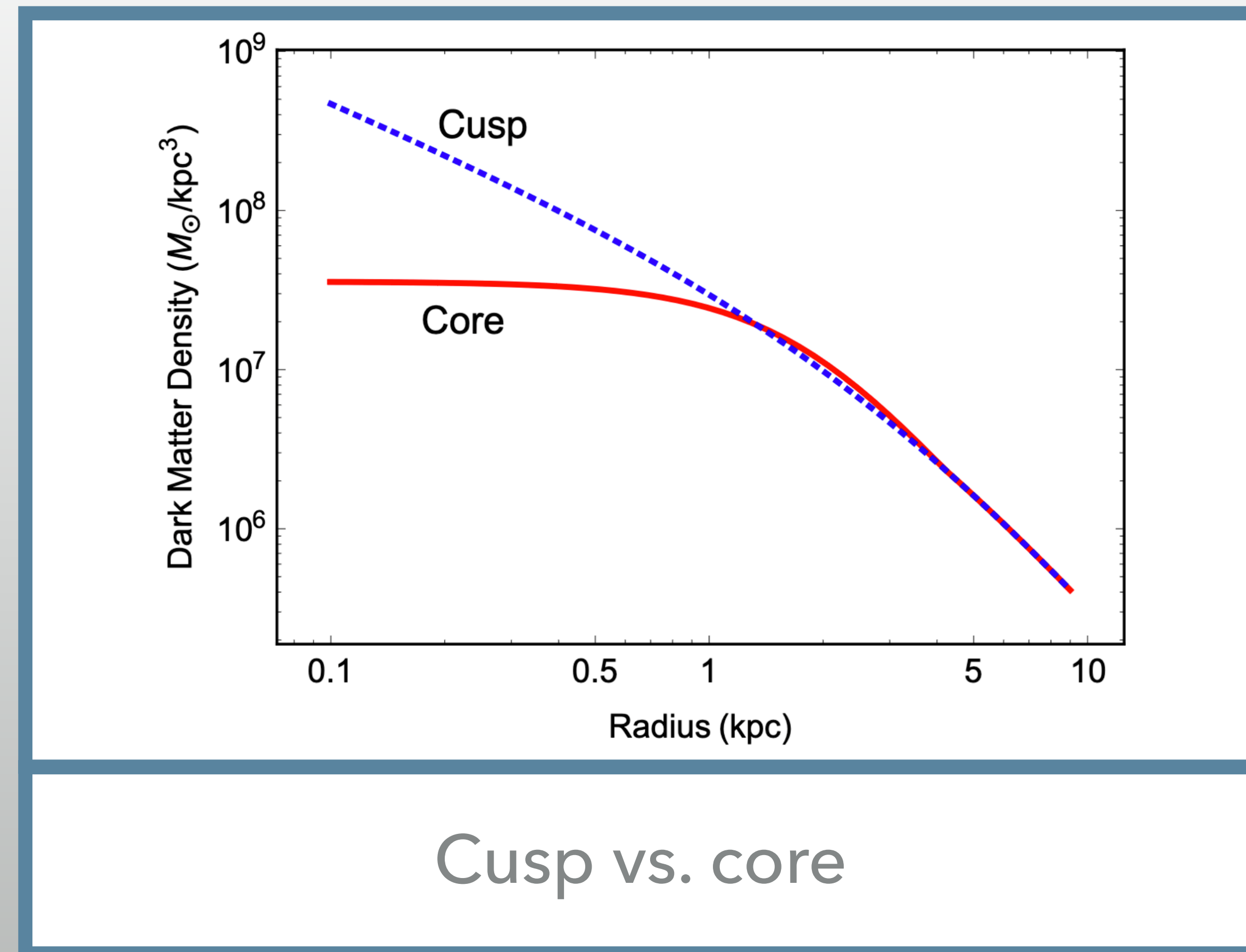


„Bullet“ cluster

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# DARK MATTER

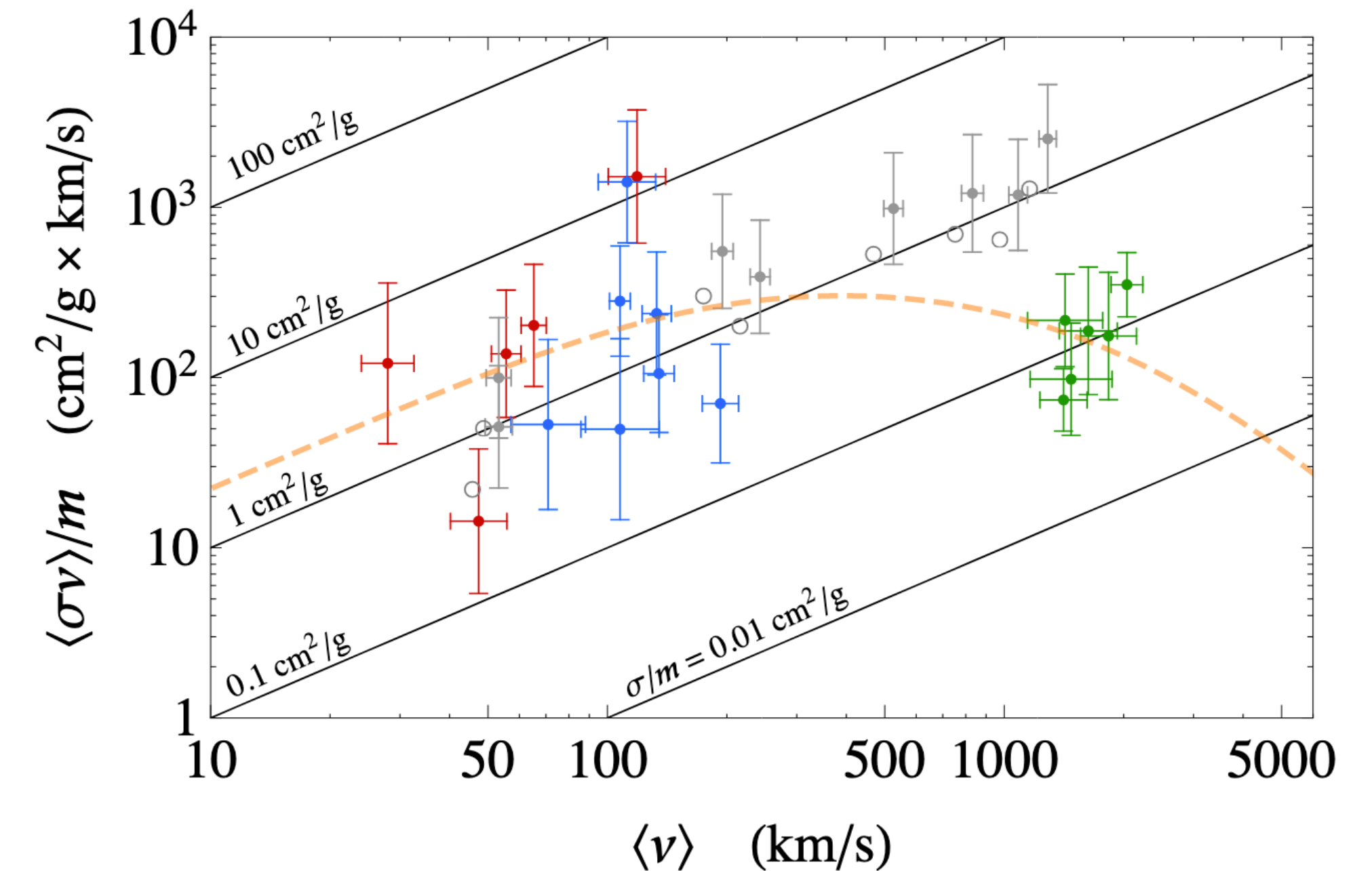
- Collection of astrophysical phenomena
- Rotational Curves
- Particle picture preferred
  - No candidate from standard model
- Large scale simulations hint towards dark matter self-interaction



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# DARK MATTER

- Collection of astrophysical phenomena
- Rotational Curves
- Particle picture preferred
  - No candidate from standard model
- Large scale simulations hint towards dark matter self-interaction
- Velocity dependence?

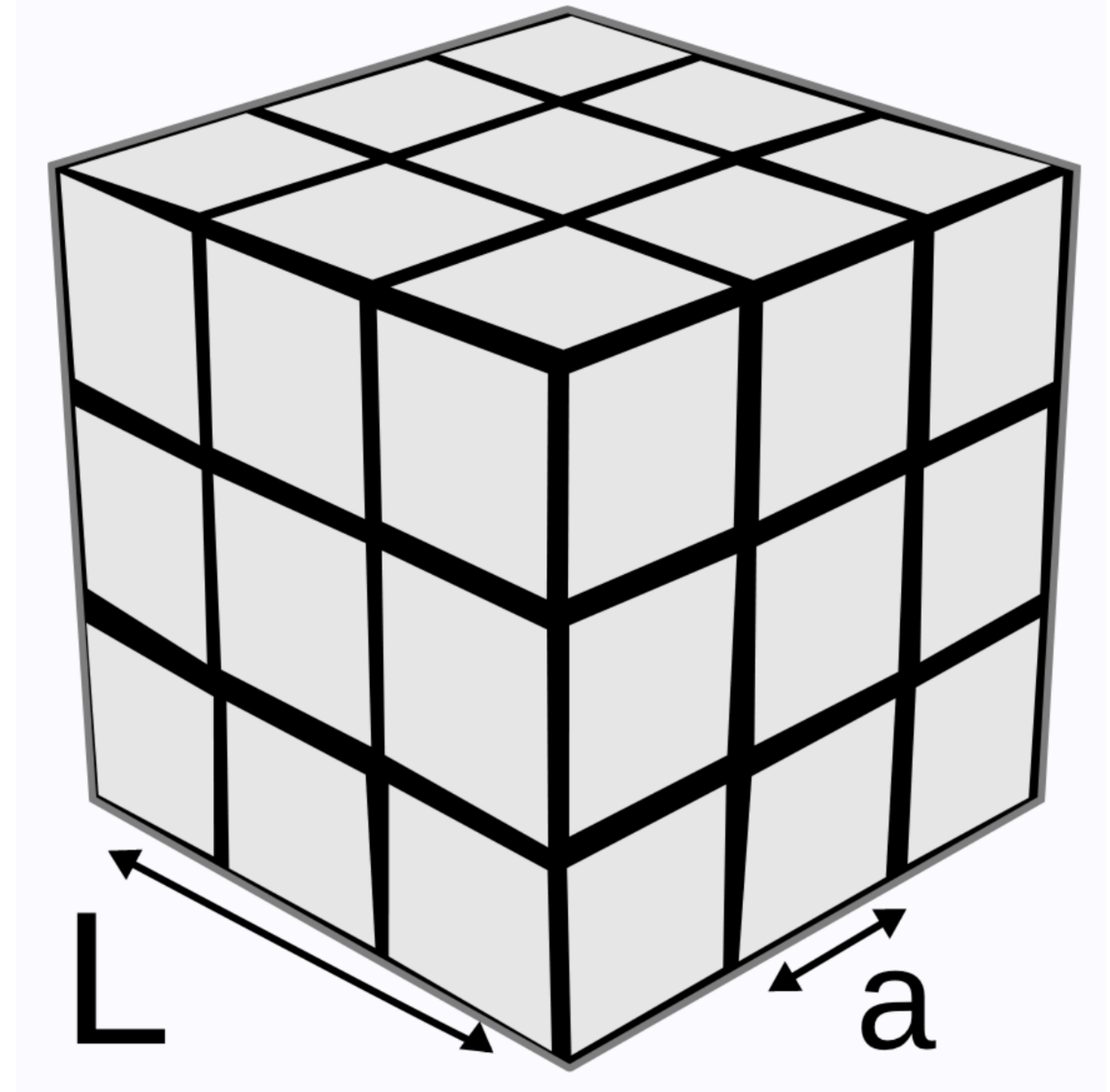


„DM Halos as particle colliders“

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# LATTICE FIELD THEORY

- Discretize action on a 4D euclidean lattice
- Montecarlo sampling
- UV- ( $a$ ) and IR-cutoff ( $L$ )



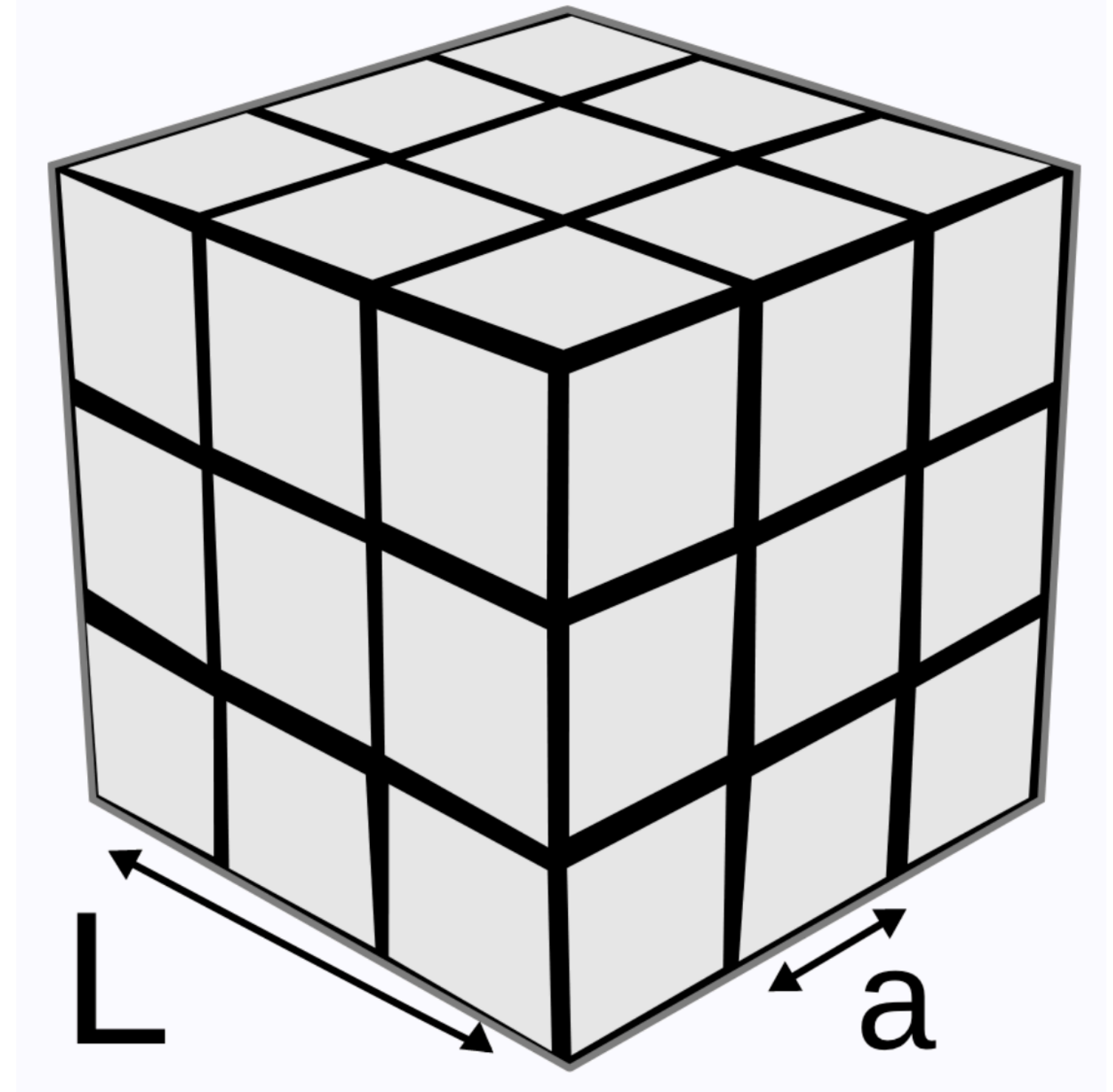
Lattice Field Theory

Leinweber: arXiv:0004025 (2000)  
Gattringer, Lang (2010)



# LATTICE FIELD THEORY

- Discretize action on a 4D euclidean lattice
- Montecarlo sampling
- UV- ( $a$ ) and IR-cutoff ( $L$ )
- Calculate correlation functions
  - Energy levels depend on lattice size
- Extrapolate to infinite volume



Lattice Field Theory

Leinweber: arXiv:0004025 (2000)  
Gattringer, Lang (2010)

# MOTIVATION

- Nature of dark matter is still elusive
- What can theory/lattice do?
- 2 approaches:
  1. Indirect detection via astronomical objects
    - DM in compact objects
  2. Constraints from cosmological objects
    - Self-scattering in galaxy clusters etc.

# DARK MATTER IN NEUTRON STARS

# NEUTRON STARS – TOV EQUATION

- Pressure gradient in Schwarzschild metric
- Mass and radius of a star

$$\frac{dP}{dr} = - \frac{(\epsilon + P)4\pi r^3 + m}{r(r - 2m)}$$

$$\frac{dm}{dr} = 4\pi\epsilon r^2 dr$$

TOV-equations

Bertone et al.: Phys.Rept.405 (2005)

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Chandra X-ray Observatory

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# NEUTRON STARS – TOV EQUATION

- Pressure gradient in Schwarzschild metric
- Mass and radius of a star
- Equation of state (EoS) as input
  - Contains the information about the matter

$$\frac{dP}{dr} = - \frac{(\epsilon + P)4\pi r^3 + m}{r(r - 2m)}$$

$$\frac{dm}{dr} = 4\pi\epsilon r^2 dr$$

TOV-equations

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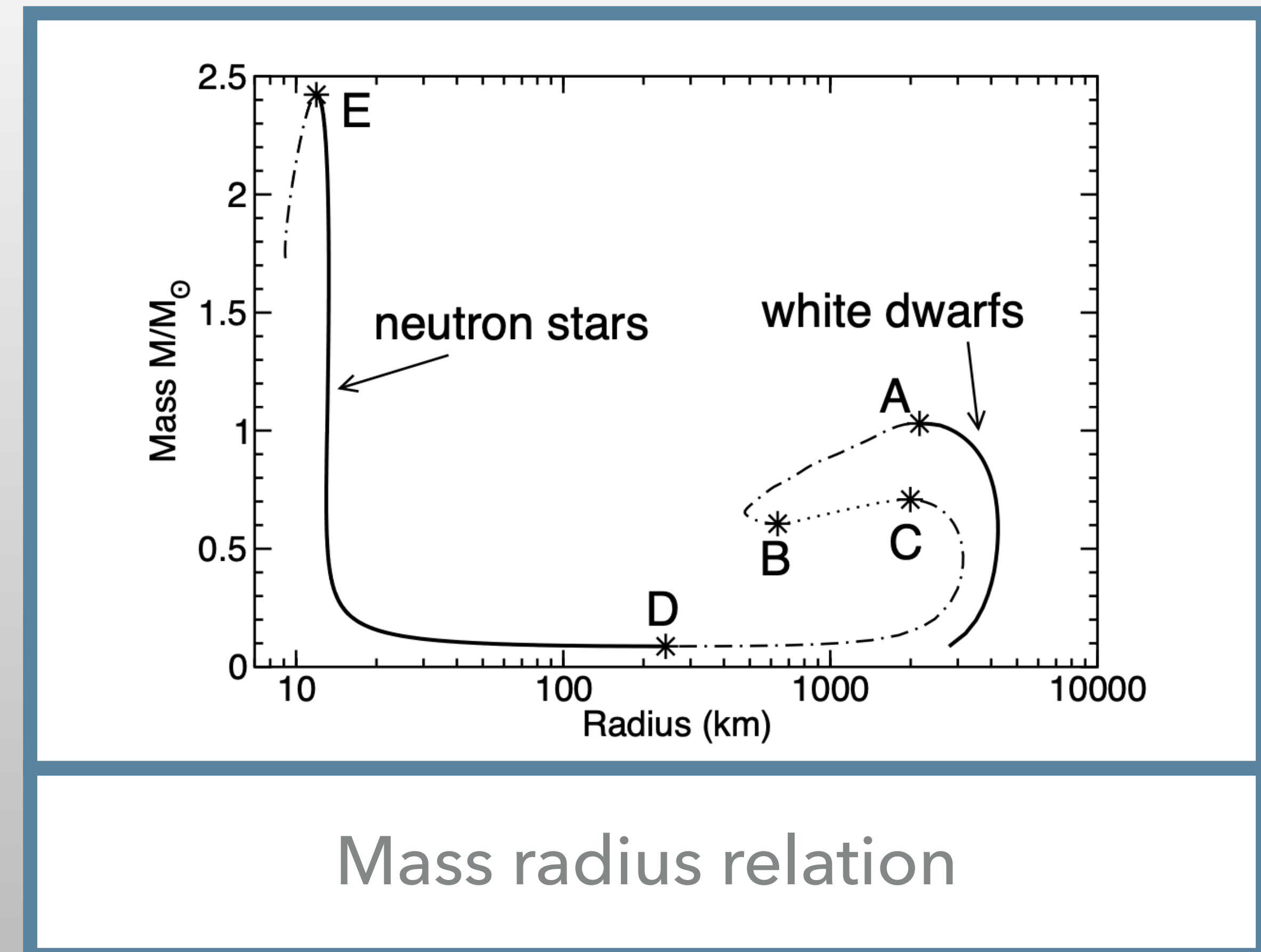
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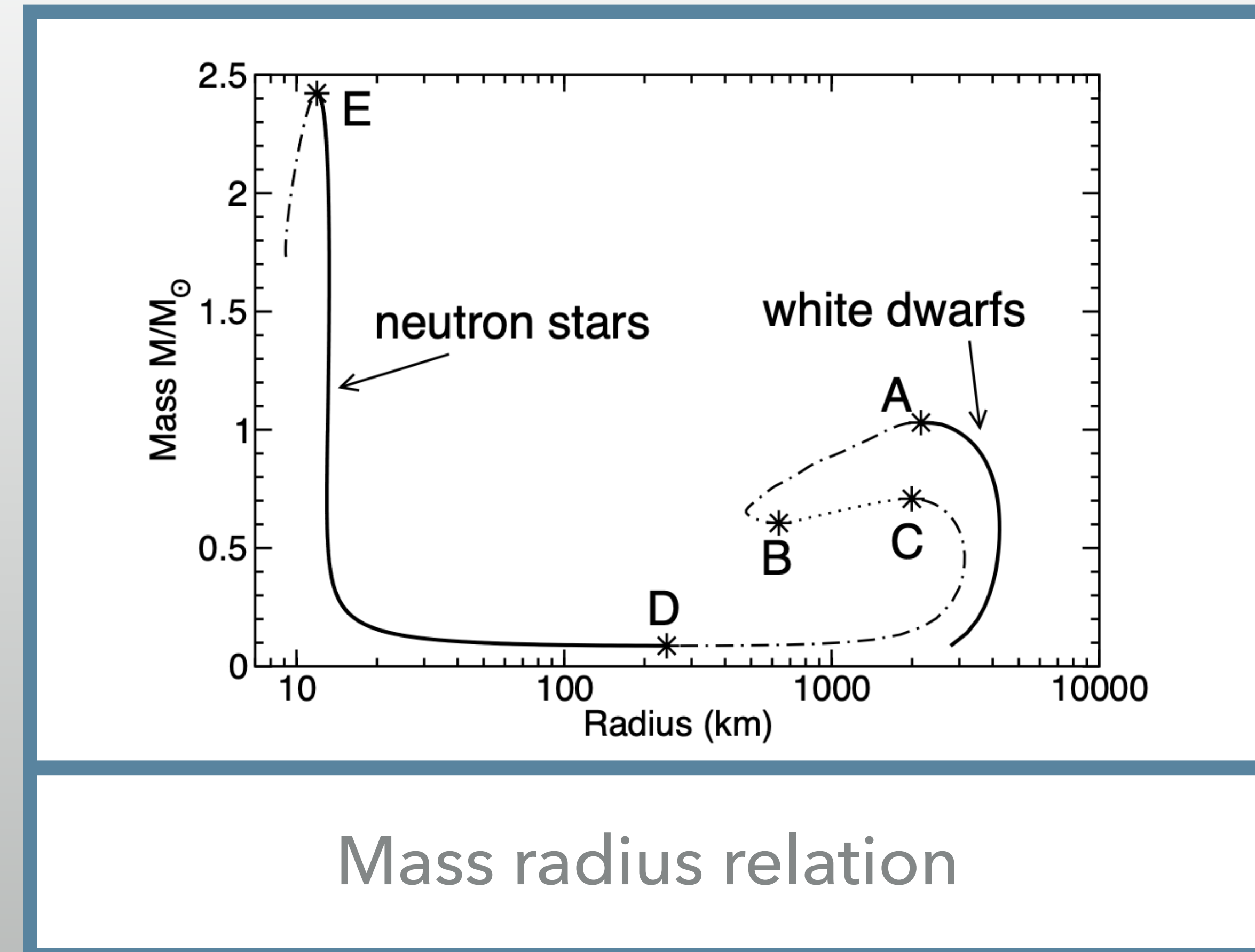
# NEUTRON STARS – TOV EQUATION

- Iteration over central pressures
- Result: Mass radius relation
  - Neutron stars and white dwarfs



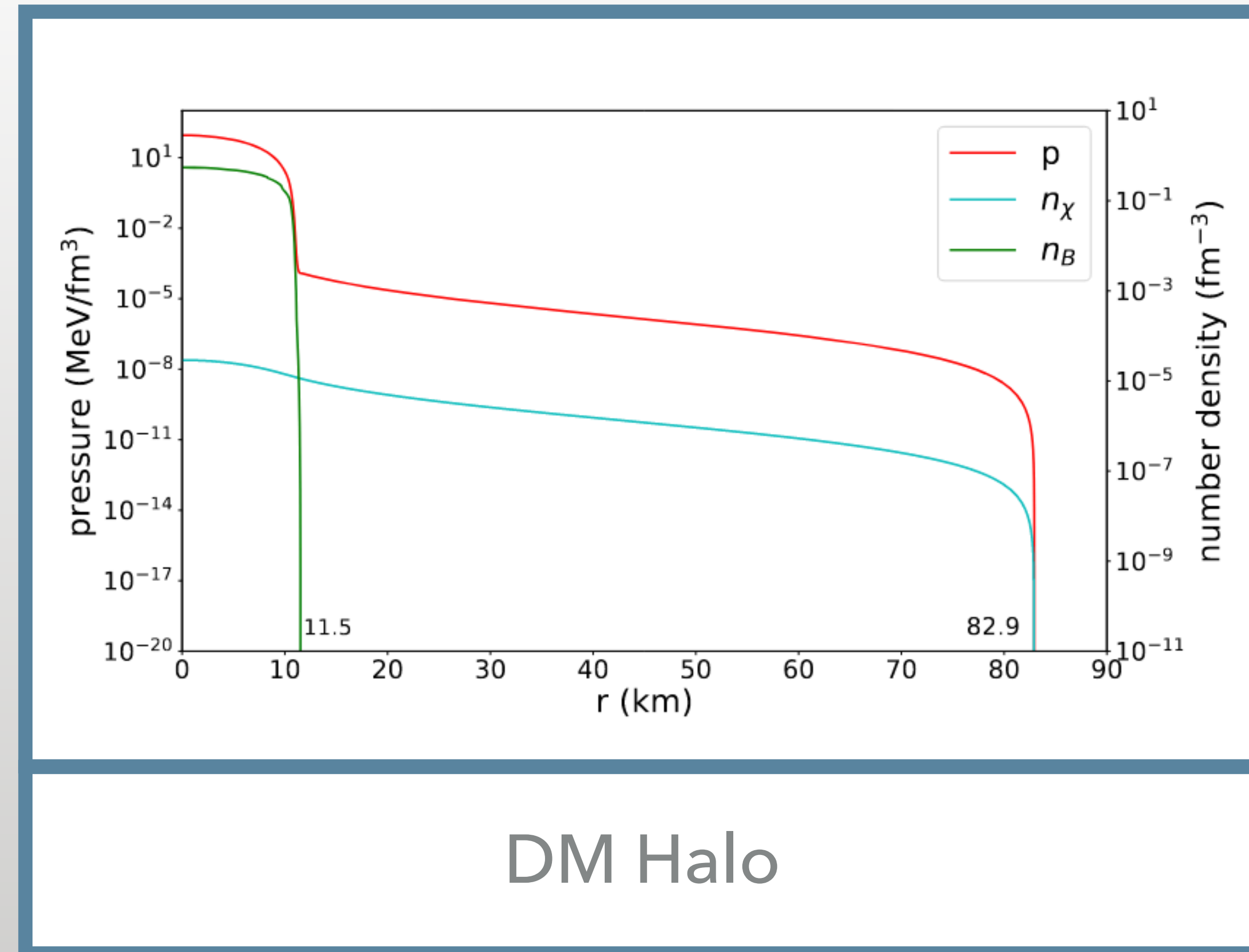
# NEUTRON STARS – TOV EQUATION

- Iteration over central pressures
- Result: Mass radius relation
  - Neutron stars and white dwarfs
- Also: Information about gravitational wave properties
  - Tidal deformability



## DM IN NEUTRON STARS

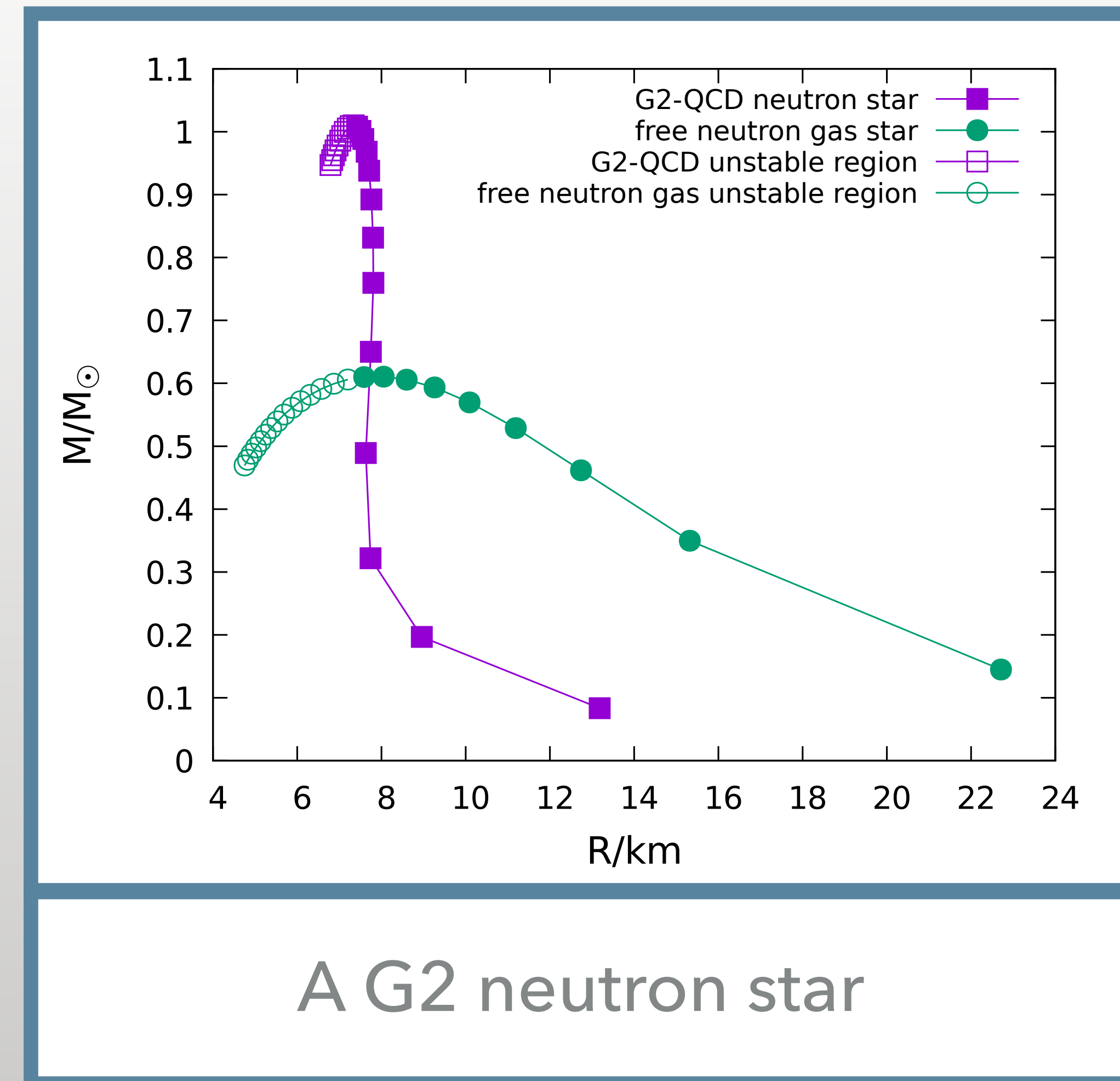
- TOV equations can be solved for two „kinds“ of fluid (SM, DM)
- New parameters:
  - DM EoS
  - $r = \frac{P_{0,DM}}{P_{0,OM}}$





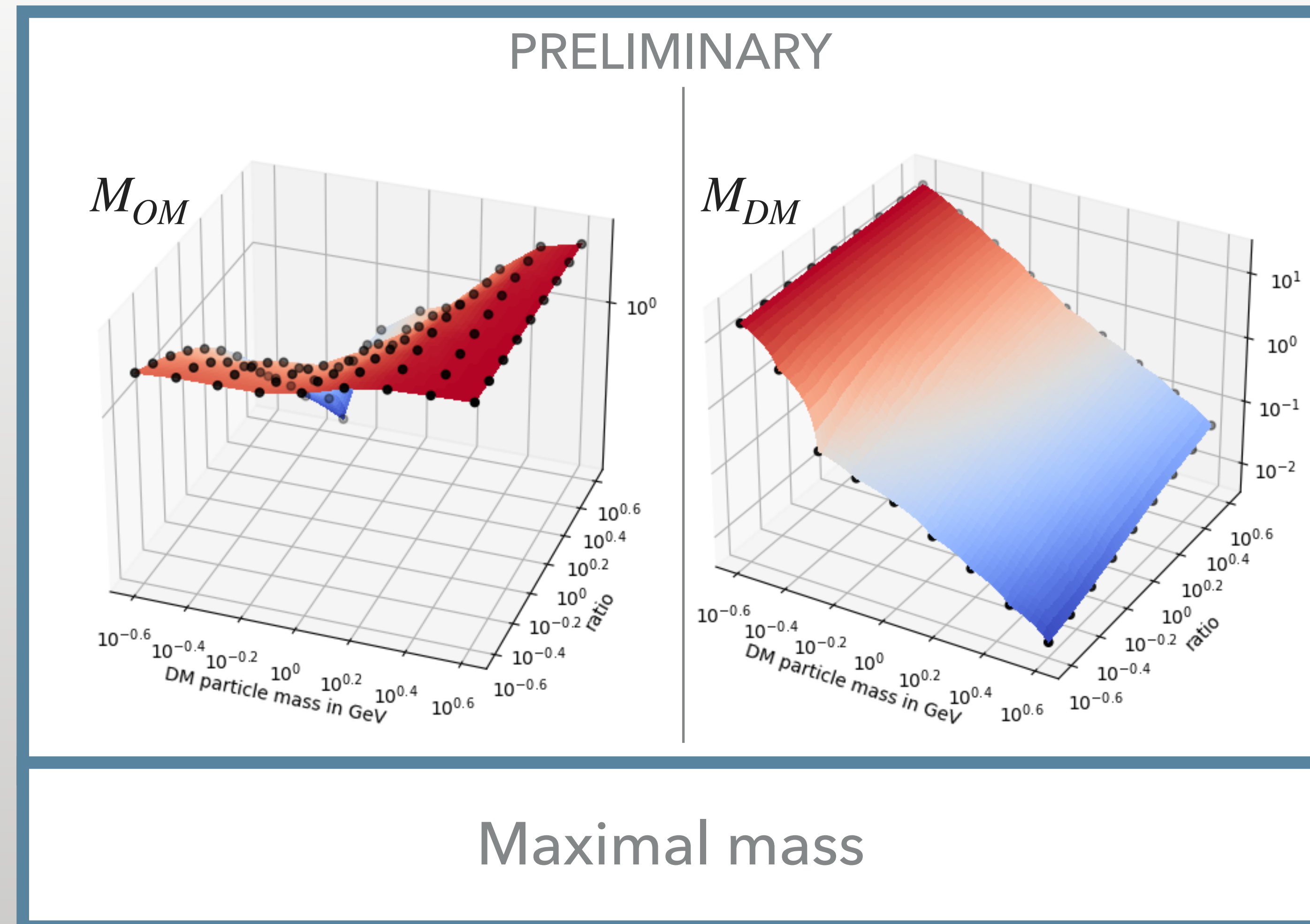
# LATTICE FOR NEUTRON STAR

- Sign problem?
- There are gauge groups without a sign problem  $SU(2N)$ ,  $Sp(2N)$ ,  $G2$  ...
- Lattice results for the exceptional group  $G2$  to mimic a neutron star
  - Too low masses



# RESULTS

- OM: „Standard“
- DM: Lattice results for G2
- Implications for Mass Radius Relation and Tidal Deformability
- Work in progress

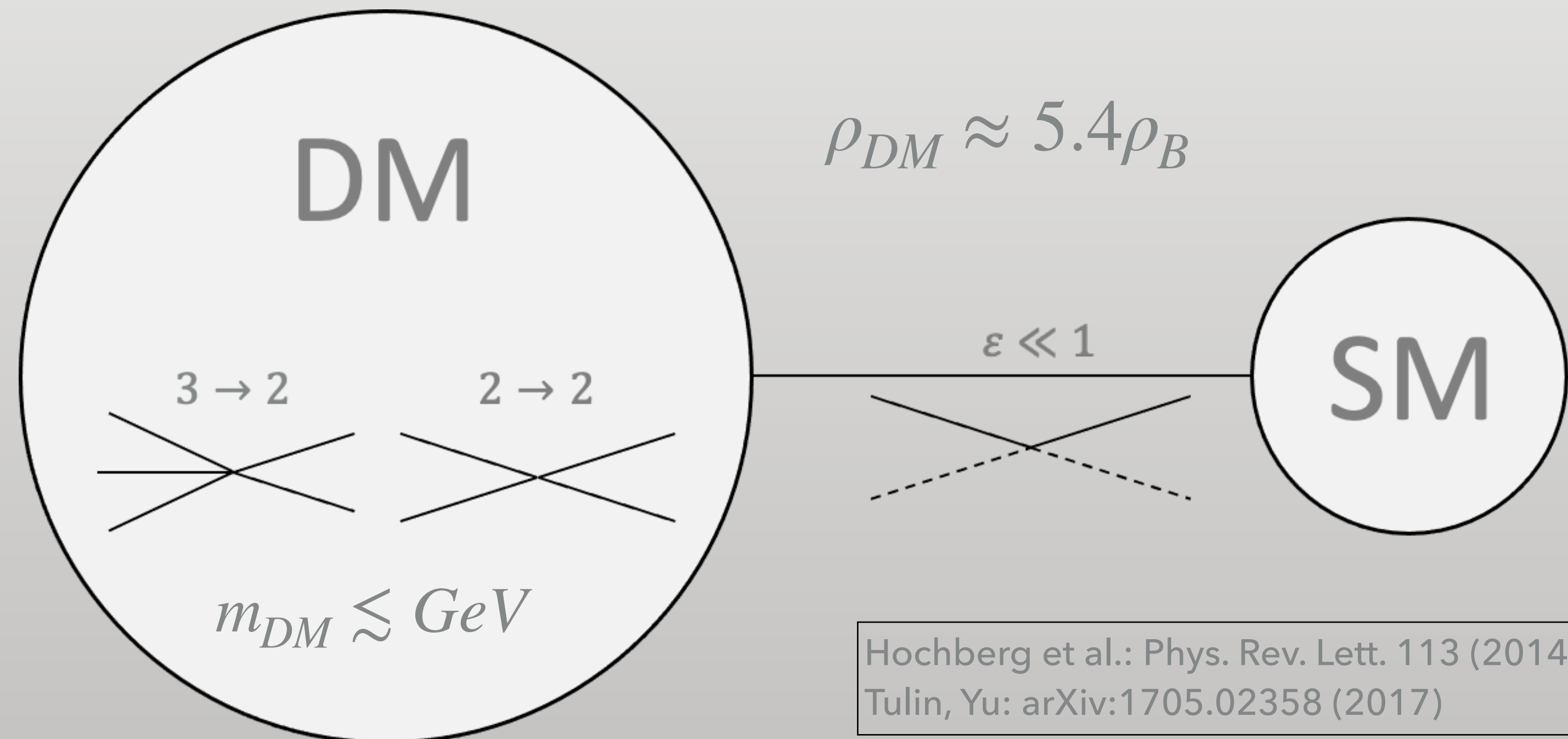


Kevin Radl (Bachelor student)

# SCATTERING IN $SP(4)$ DARK MATTER

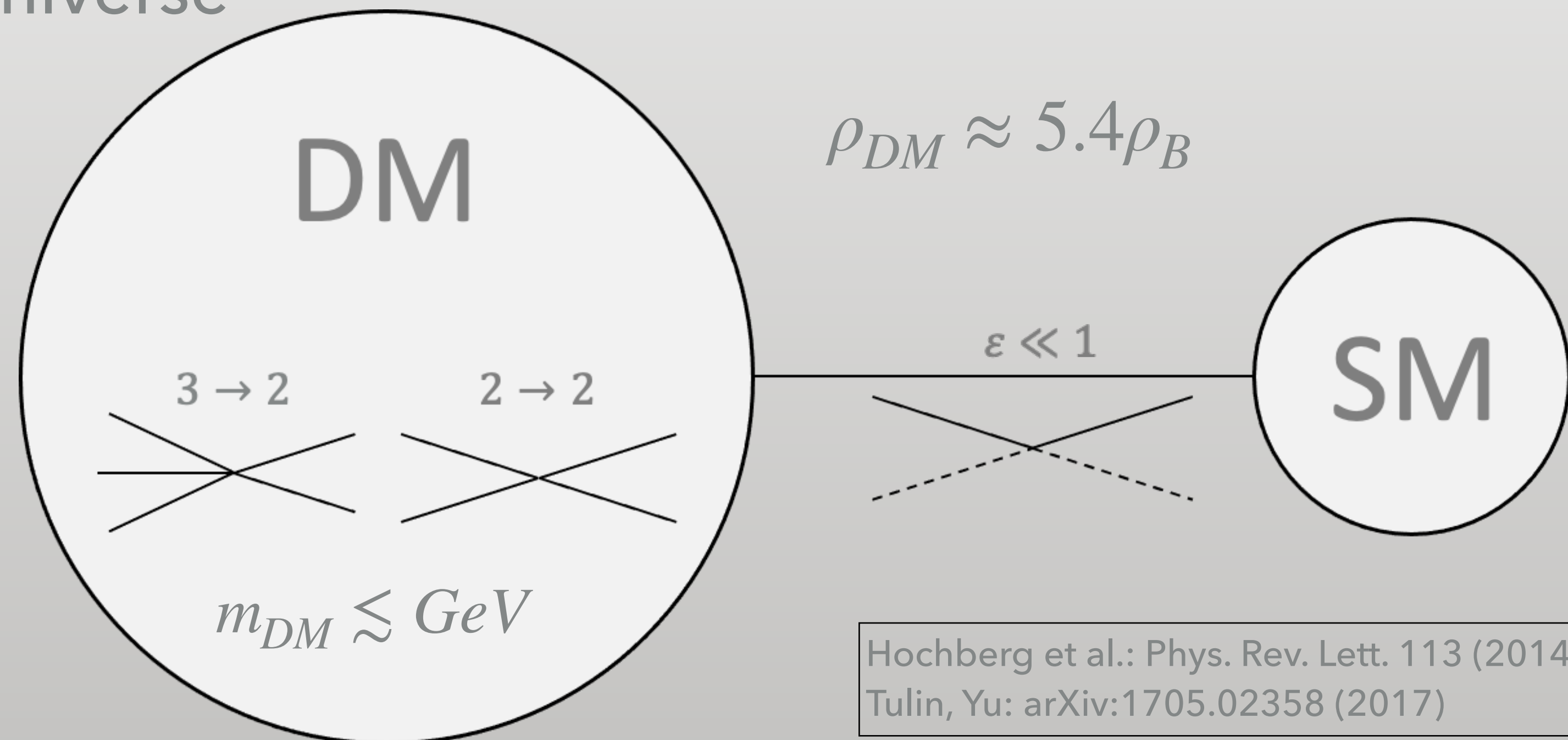
# OUR MODEL - SIMP (STRONGLY INTERACTING MASSIVE PARTICLES)

- Paradigm for DM as a thermal relic from early universe via freeze-out



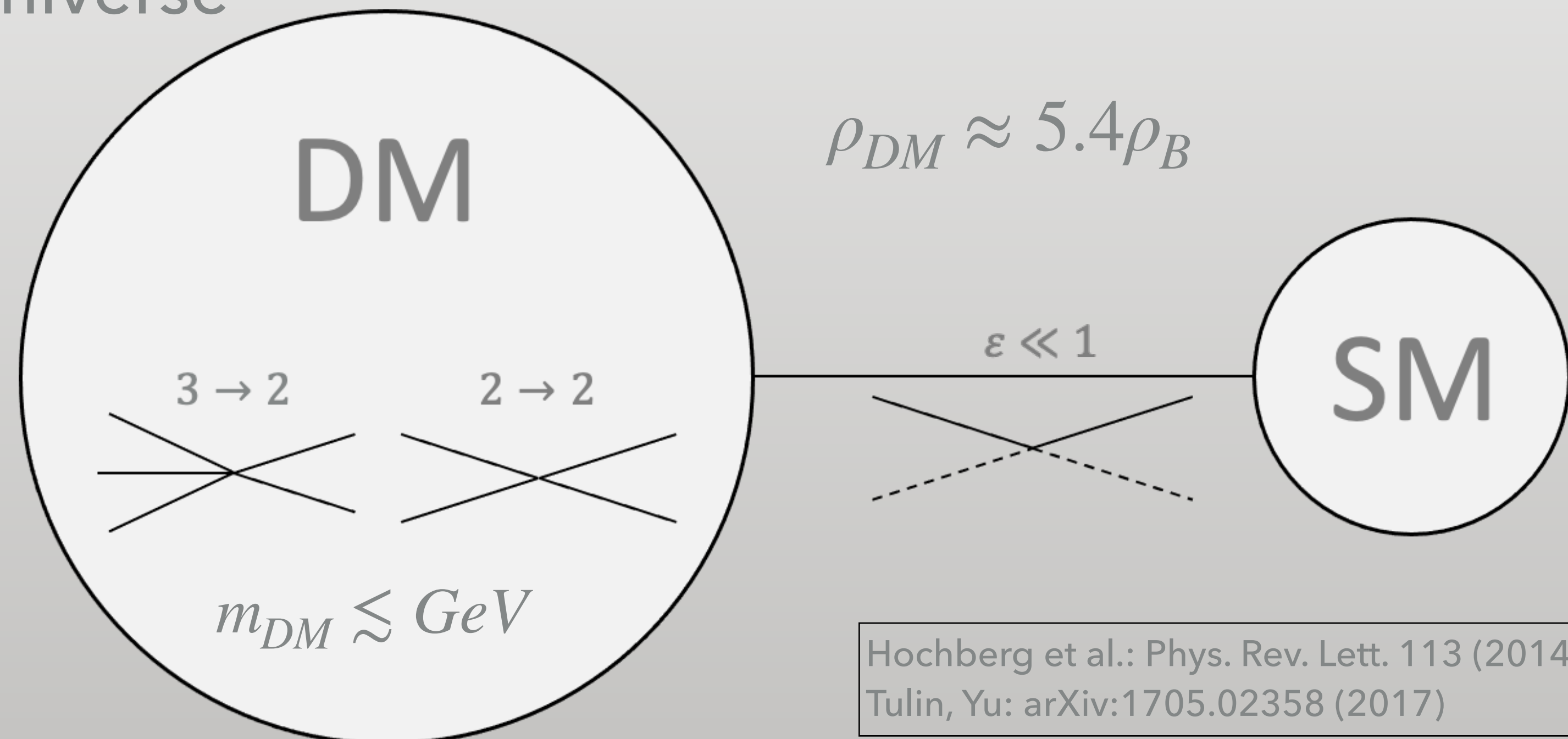
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- Mediator with the standard model
  - Thermal equilibrium in early universe
- Number changing process in dark sector ( $3 \rightarrow 2$ )



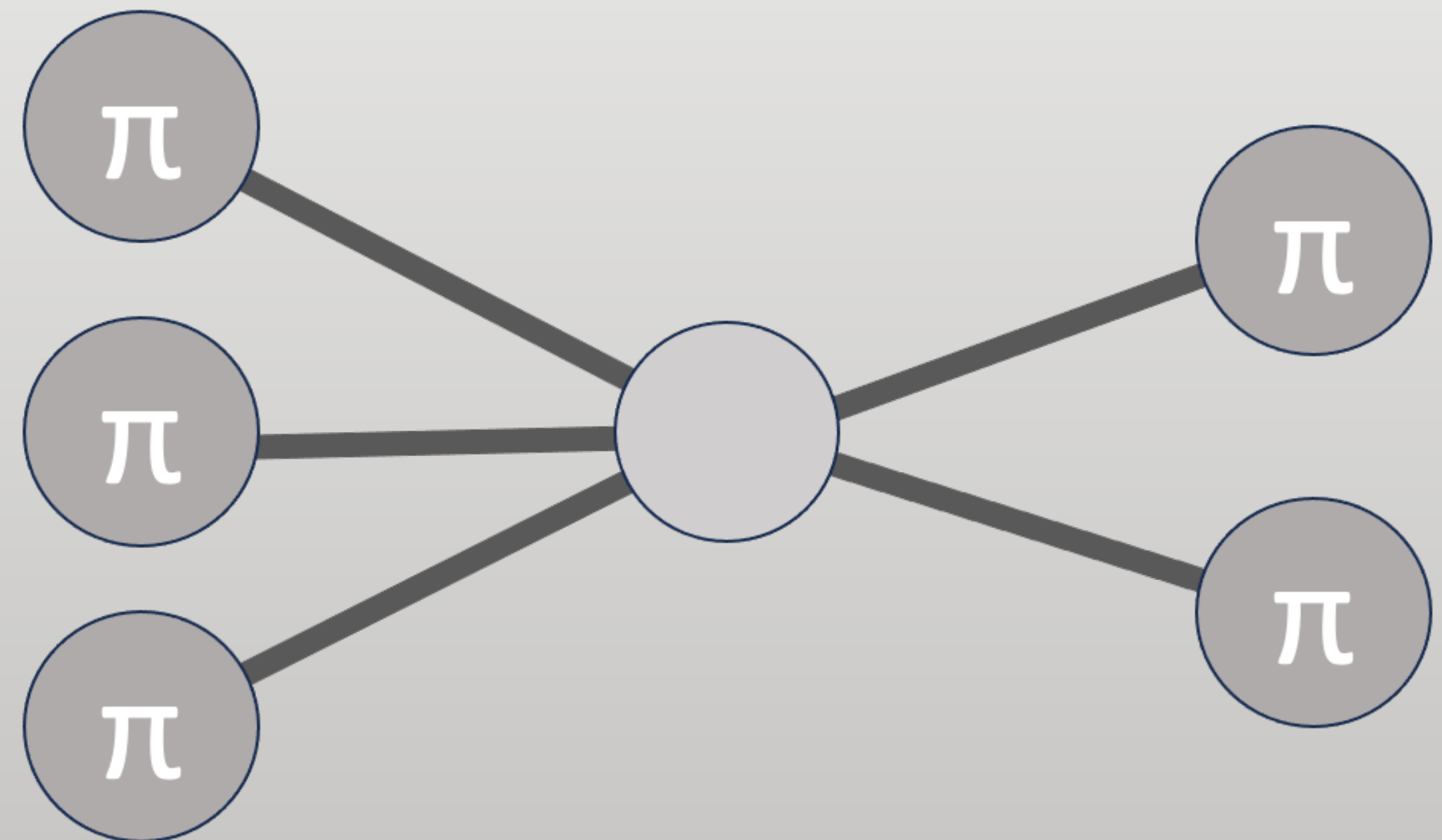
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- Paradigm for DM as a thermal relic from early universe via freeze-out
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- Number changing process in dark sector ( $3 \rightarrow 2$ )
- Self interaction addresses structure formation
- Mediator enables direct detection



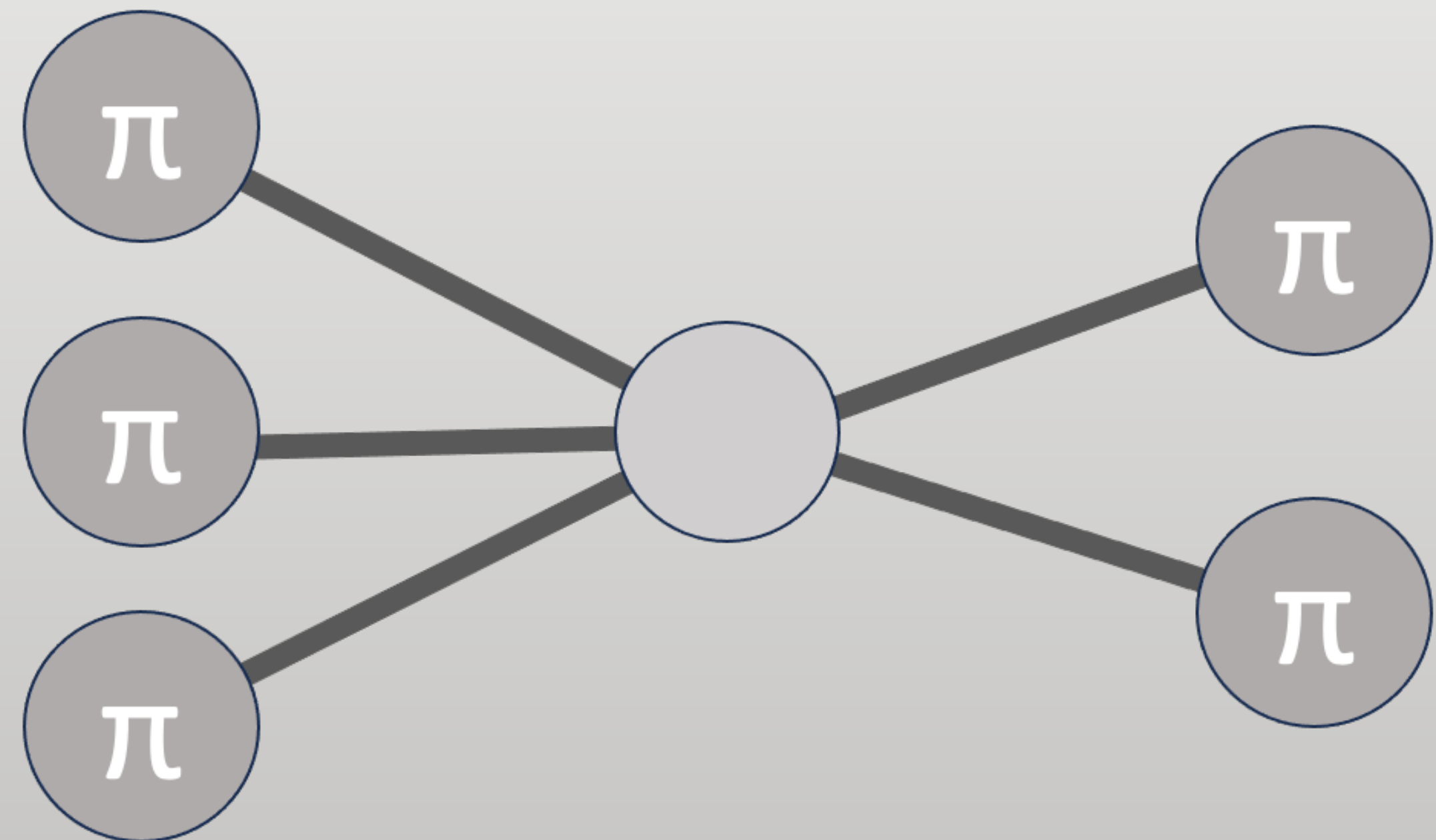
## WHY SP(4)

- A minimal realisation of SIMP:
  - Sp(4) gauge with  $N_f=2$  flavours („dark quarks“)



## WHY SP(4)

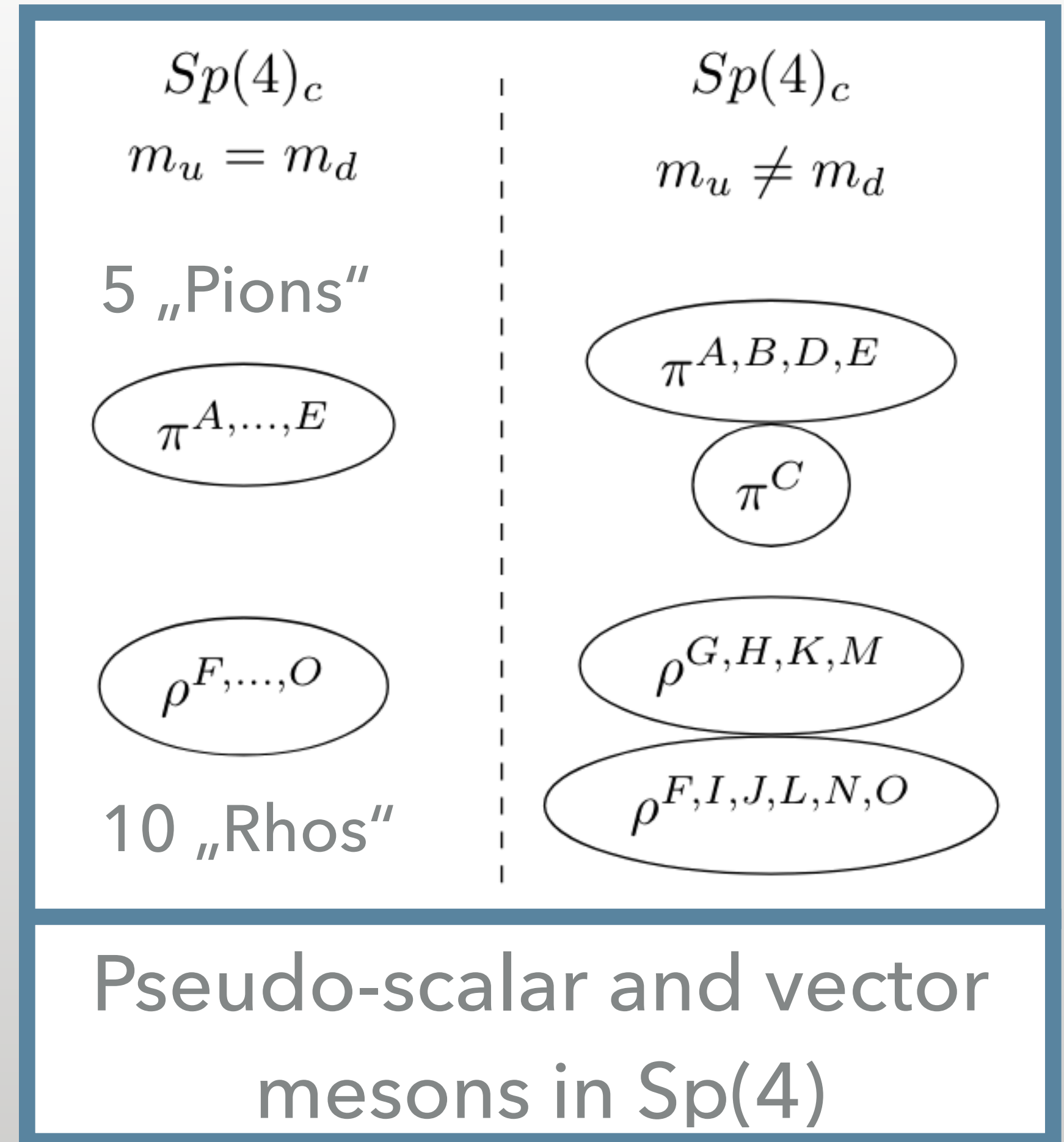
- A minimal realisation of SIMP:
  - Sp(4) gauge with  $N_f=2$  flavours („dark quarks“)
- DM candidate: pNGB from chiral symmetry breaking
  - 5 „dark Pions“
  - $3 \rightarrow 2$  process possible





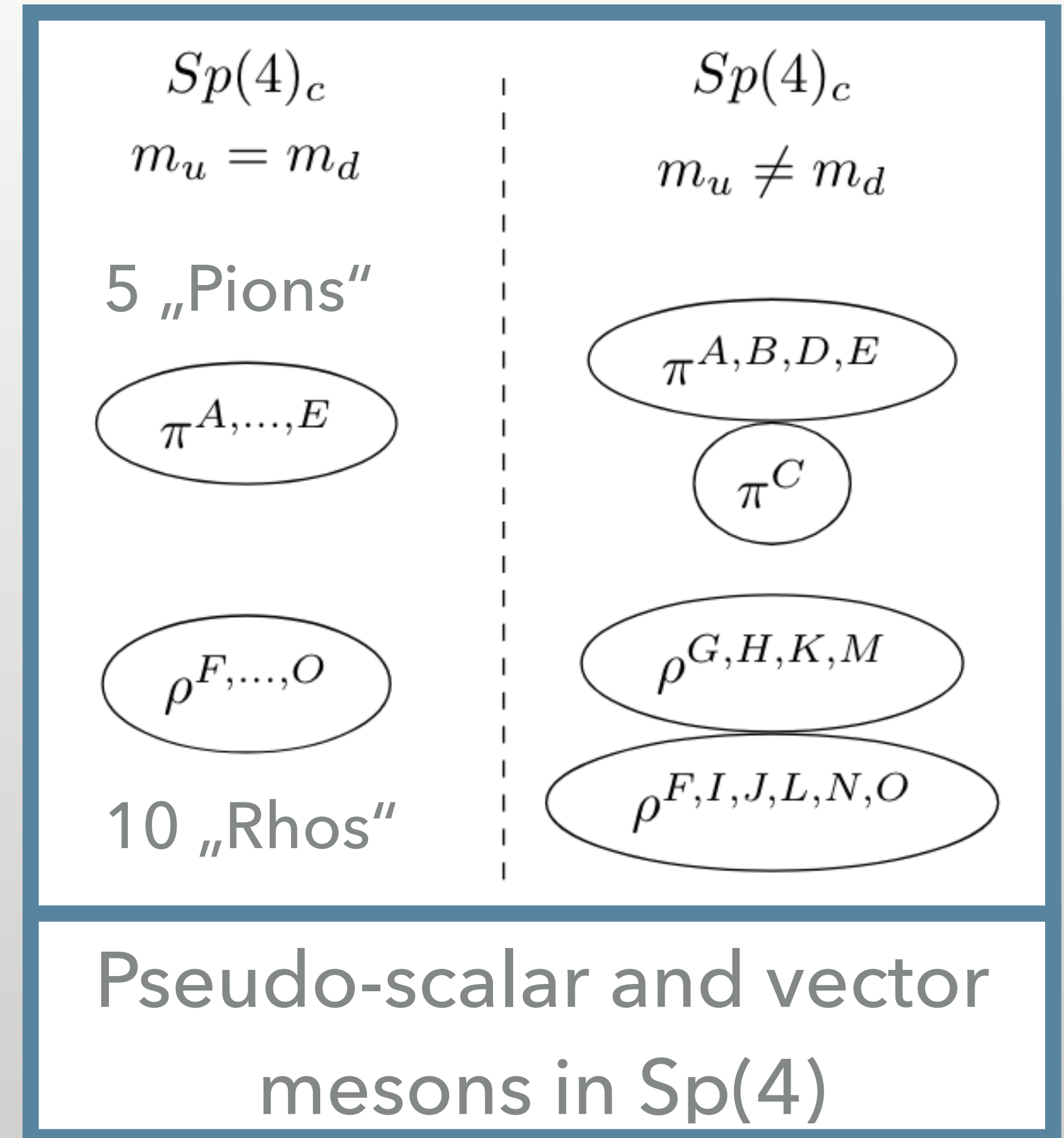
# SP(4)

- Fermions in fundamental representation (pseudo-real)
- 5 „dark Pions“
  - 3 Pions like in QCD
- 2 „diquarks“



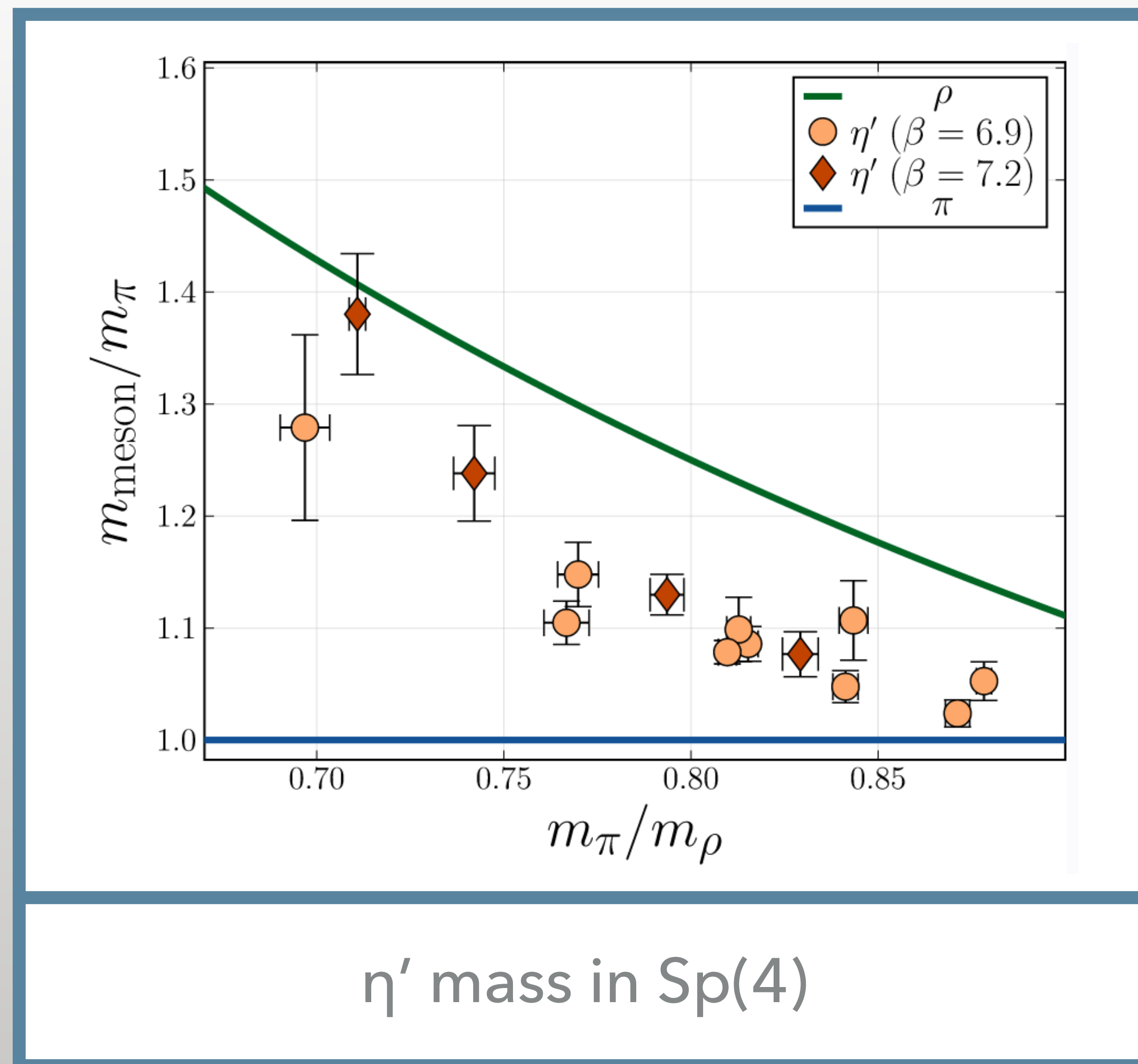
# SP(4)

- Fermions in fundamental representation (pseudo-real)
- 5 „dark Pions“
  - 3 Pions like in QCD
  - 2 „diquarks“
- No fermionic bound states (even number of colours)
- Rich hadron sector like QCD



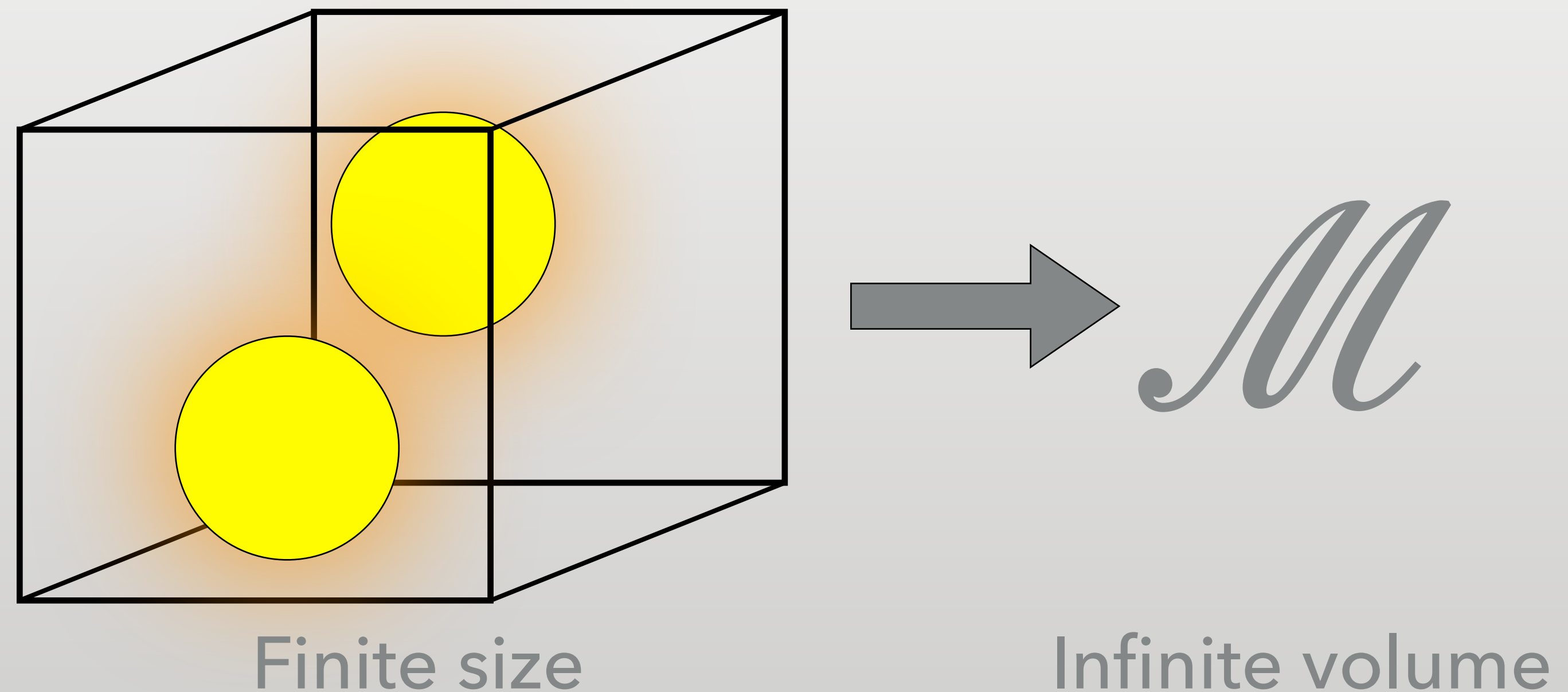
# MASSES

- Mass spectrum of mesons
  - For details: [2304.07191]
- Spectrum differs from QCD
- Example:
  - Light  $\eta'$
  - Effects scattering:  $\pi\pi \leftrightarrow \eta'$



# SCATTERING ON THE LATTICE

- Relate finite size effects to infinite volume scattering properties

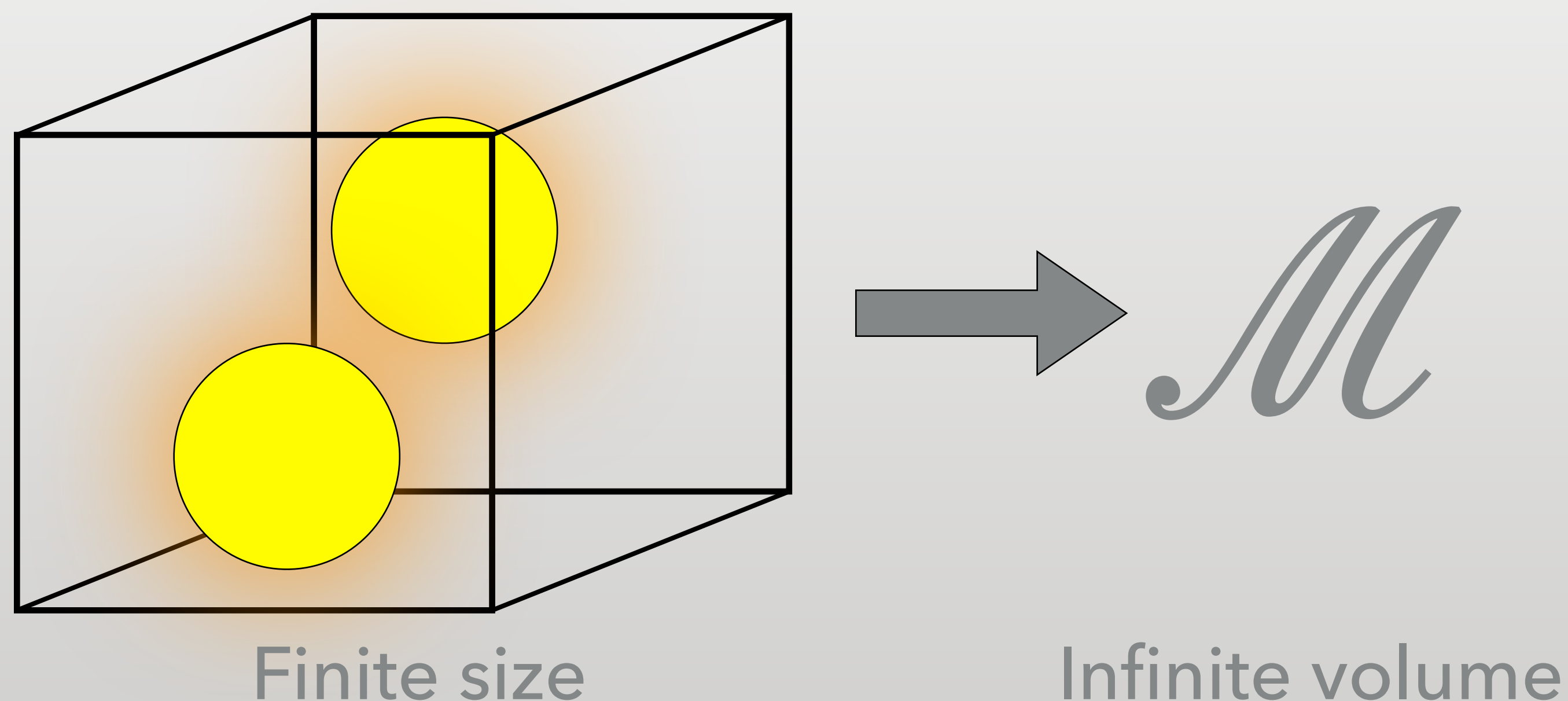


Lüscher et al.: Commun. Math. Phys. 104/105 (1986)  
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Hansen et al.: arXiv:2101.10246 (2021)

# SCATTERING ON THE LATTICE

- Relate finite size effects to infinite volume scattering properties
- Phase shift from  $\pi\pi$ -energy levels

$$\mathcal{O}_{\pi\pi} = \mathcal{O}_{\pi^+} \mathcal{O}_{\pi^+} = (\bar{d}\gamma_5 u) (\bar{d}\gamma_5 u)$$



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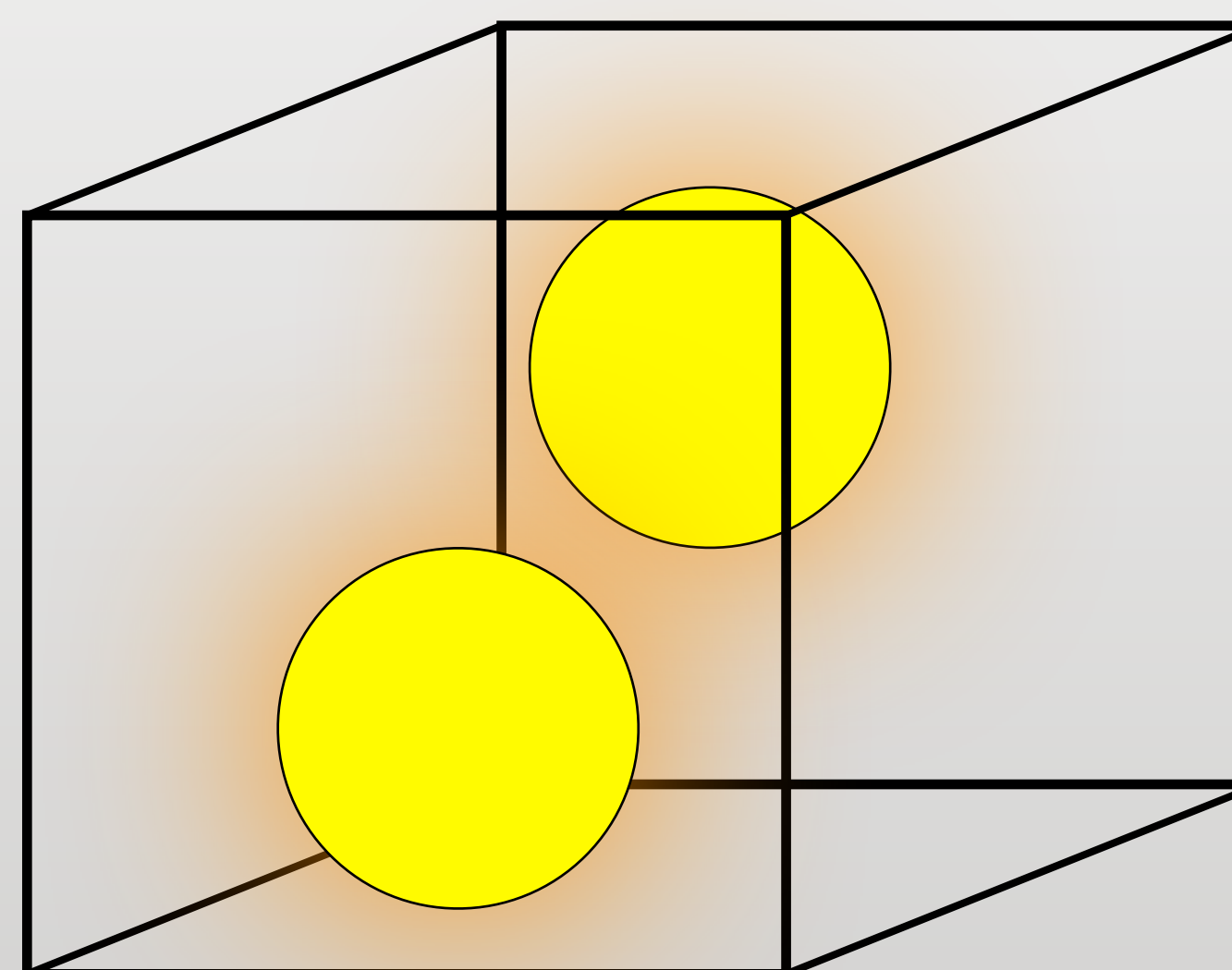
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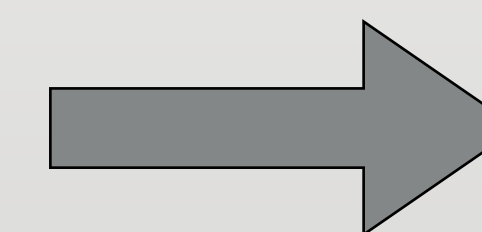
- $m_u = m_d$  & maximal isospin ( $I=2$ )

- Lüscher's finite size method:

- $\tan(\delta_0(q)) = \frac{\pi^{\frac{3}{2}} q}{\mathcal{L}_{00}^0(1, q^2)}$



Finite size

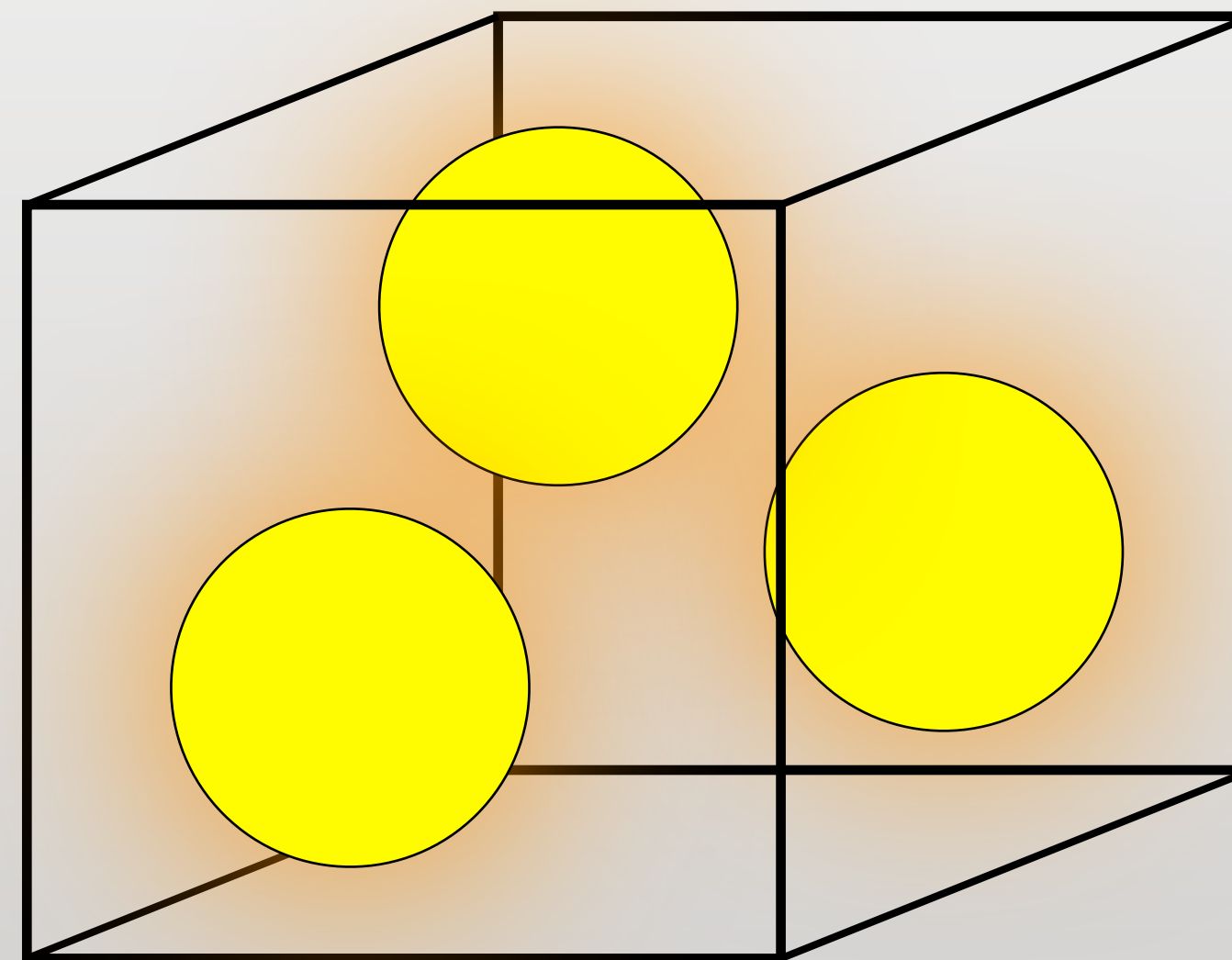

 $\mathcal{M}$ 

Infinite volume

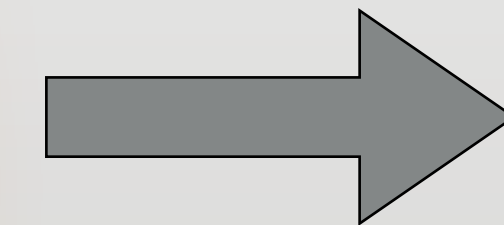
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# SCATTERING ON THE LATTICE

- For  $3 \rightarrow 2$  process
  - Full  $2 \rightarrow 2$  information needed
- 3 particle quantization condition
  - $\det[F_3^{-1} + \mathcal{K}_3] = 0$
  - Hansen, Romero-López, Sharpe  
arXiv:2101.10246 [hep-lat]



Finite size

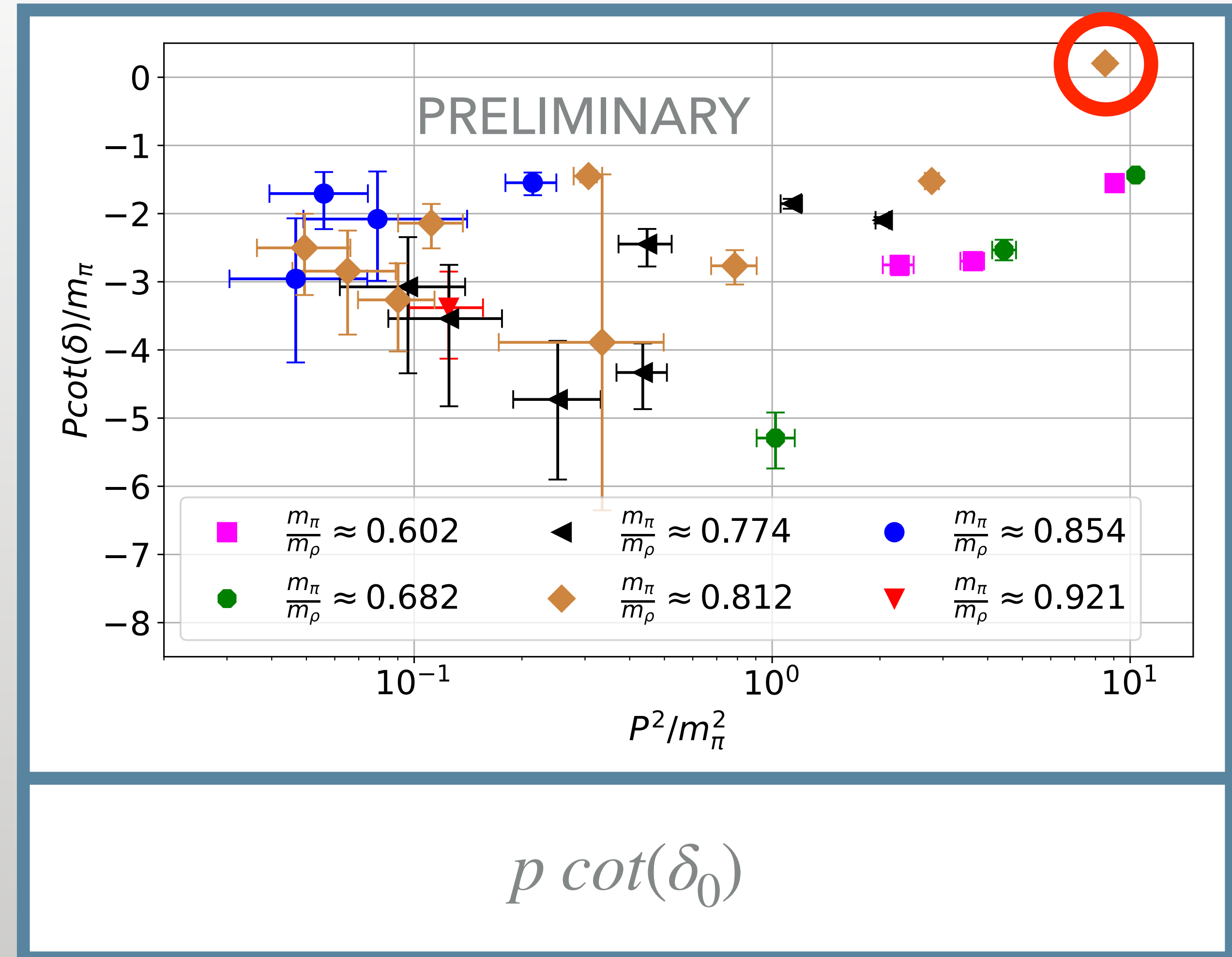

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# RESULTS - PHASE SHIFT $\delta_0$

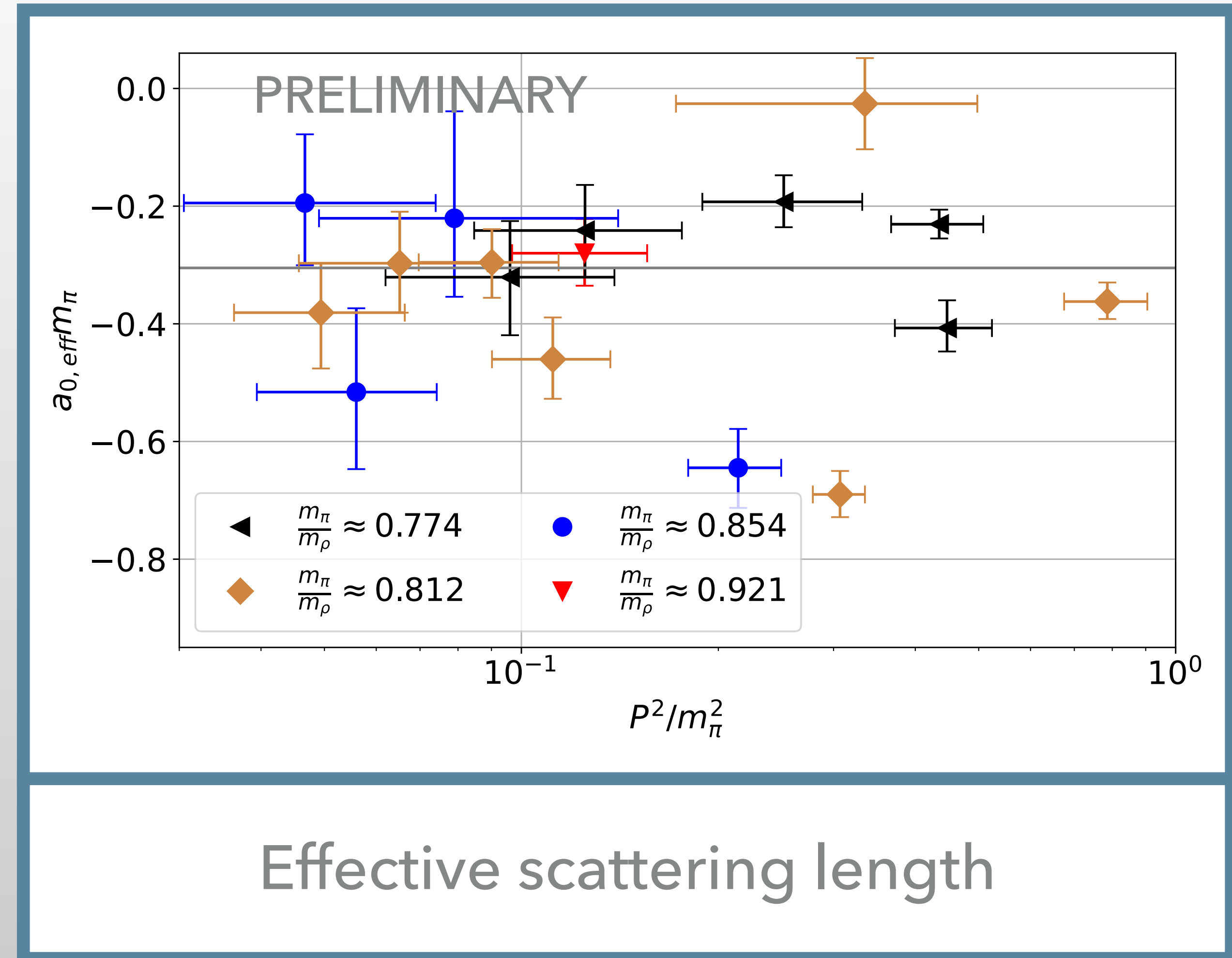
- $|\vec{p}| \cot(\delta_0(q)) = \frac{1}{a_0} + \mathcal{O}(|\vec{p}|^2)$
- Threshold ( $|\vec{p}| \rightarrow 0$ ):
  - Information about a potential bound state
- Zero crossing:
  - Information about a potential resonance





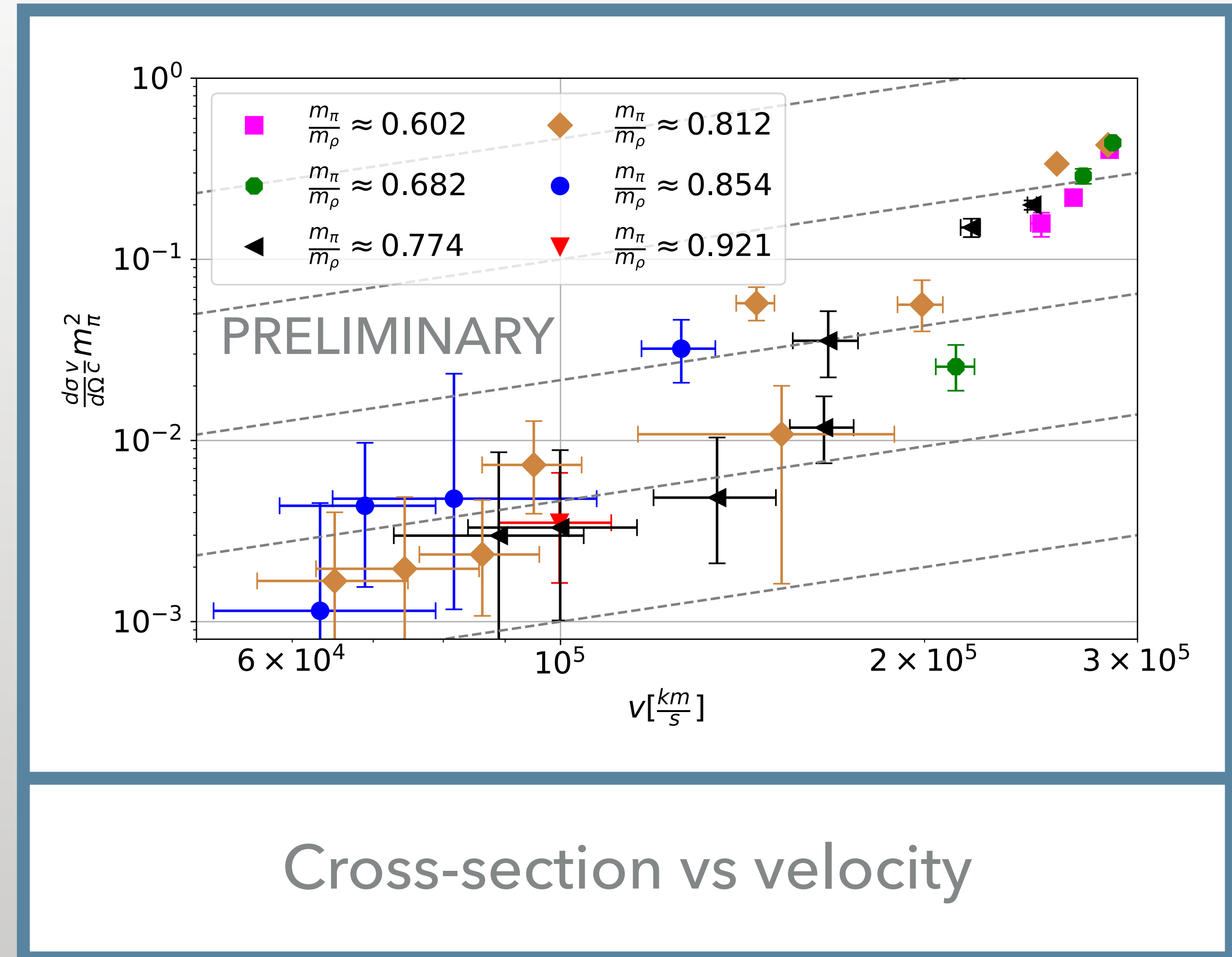
## RESULTS - EFFECTIVE $a_0$

- $|\vec{p}| \cot(\delta_0(q)) = \frac{1}{a_0} + \mathcal{O}(|\vec{p}|^2)$
- $a_{0,eff} = \lim_{|\vec{p}| \rightarrow 0} \frac{\tan(\delta_0)}{|\vec{p}|}$
- Constraint from density profiles of galaxy clusters:
  - $m_{DM} > 75 MeV$



# RESULTS - CROSS SECTION

- $\frac{d\sigma}{d\Omega} \propto \sigma$  (s-wave)
- $\frac{d\sigma}{d\Omega} \frac{v}{c} m_\pi^2$  vs velocity
- Do we see a hint for velocity dependence?
- Yes, but not the one from the motivation



# SUMMARY & OUTLOOK

- Framework for DM in neutron stars
- Mass spectrum of  $Sp(4)$
- First results for the  $2 \rightarrow 2$  scattering phase shift
- Constraint on dark matter particle mass
- $3 \rightarrow 2$  scattering in  $Sp(4)$

## SUMMARY & OUTLOOK

- Mass spectrum of  $Sp(4)$
- Full Lüscher Analysis of „dark Pion“  $2 \rightarrow 2$  scattering
- Constraint on dark matter particle mass
  
- Solve GEVP with relative momenta
- All isospin channels
- $3 \rightarrow 2$  scattering

**THANK YOU!**

## BACKUP - SCATTERING (LATTICE MOMENTA)

● Scattering phase-shift from „lattice momenta“

● Finite volume effects correspond to scattering properties

●  $q = |\vec{p}| \frac{L}{2\pi}, \quad \cosh\left(\frac{E_{\pi\pi}}{2}\right) = \cosh(m_\pi) + 2\sin^2\left(\frac{|\vec{p}|}{2}\right)$

● Valid for  $2m_\pi < E_{\pi\pi} < 4m_\pi$

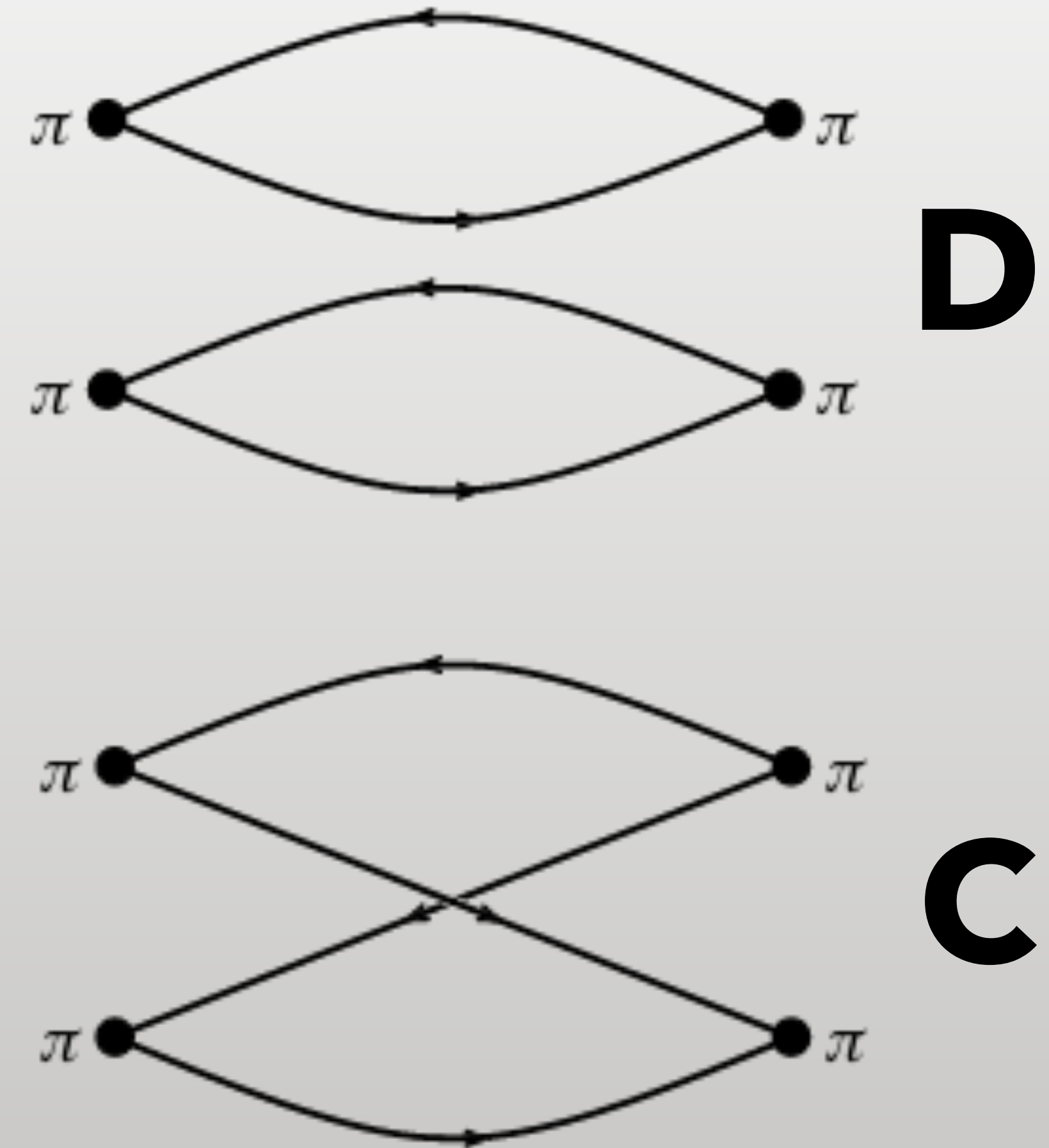
●  $\tan(\delta_0(q)) = \frac{\pi^{\frac{3}{2}} q}{\mathcal{L}_{00}^0(1, q^2)}$

## BACKUP - WICK-CONTRACTIONS - $2-\pi$ -SCATTERING

- ▶ In  $Sp(4)_F$ 
  - ▶ 5-plet of pseudo scalar (5 dark „Pions“)
  - ▶  $\pi^+, \pi^-, \pi^0, \Pi_{ud}, \Pi_{\bar{u}\bar{d}}$
- ▶ 3 Isospin channels:
  - ▶  $5 \otimes 5 = 1 \oplus 10 \oplus 14$  (Isospin  $I=0,1,2$ )

## BACKUP - WICK-CONTRACTIONS - I=2

- ▶ For I=2: One Operator sufficient
  - ▶  $\pi^+$  for example:
- ▶  $\mathcal{O}_{\pi\pi} = \mathcal{O}_{\pi^+}\mathcal{O}_{\pi^+} = (\bar{d}\gamma_5 u) (\bar{d}\gamma_5 u)$
- ▶  $C_{\pi\pi} = \langle \mathcal{O}_{\pi\pi} \mathcal{O}_{\pi\pi}^\dagger \rangle = 2D - 2C$
- ▶ Same as in QCD



## BACKUP - WICK-CONTRACTIONS - I=0

- ▶ For I=0: Contraction of all 5 „Pions

- ▶ 
$$\mathcal{O}(\pi\pi, I = 0) = \frac{1}{\sqrt{5}}(\pi^+\pi^- + \pi^-\pi^+ - \pi^0\pi^0 + \Pi_{ud}\Pi_{\bar{u}\bar{d}} + \Pi_{\bar{u}\bar{d}}\Pi_{ud})$$

- ▶ 25 Terms with 4  $\psi$ 's and 4  $\bar{\psi}$ 's

- ▶ 6 similar groups

- ▶ 344 terms

- ▶ 
$$C_{\pi\pi}^{I=0} = 2D + 3C - 10R + 5V$$



## BACKUP - WICK-CONTRACTIONS - I=0

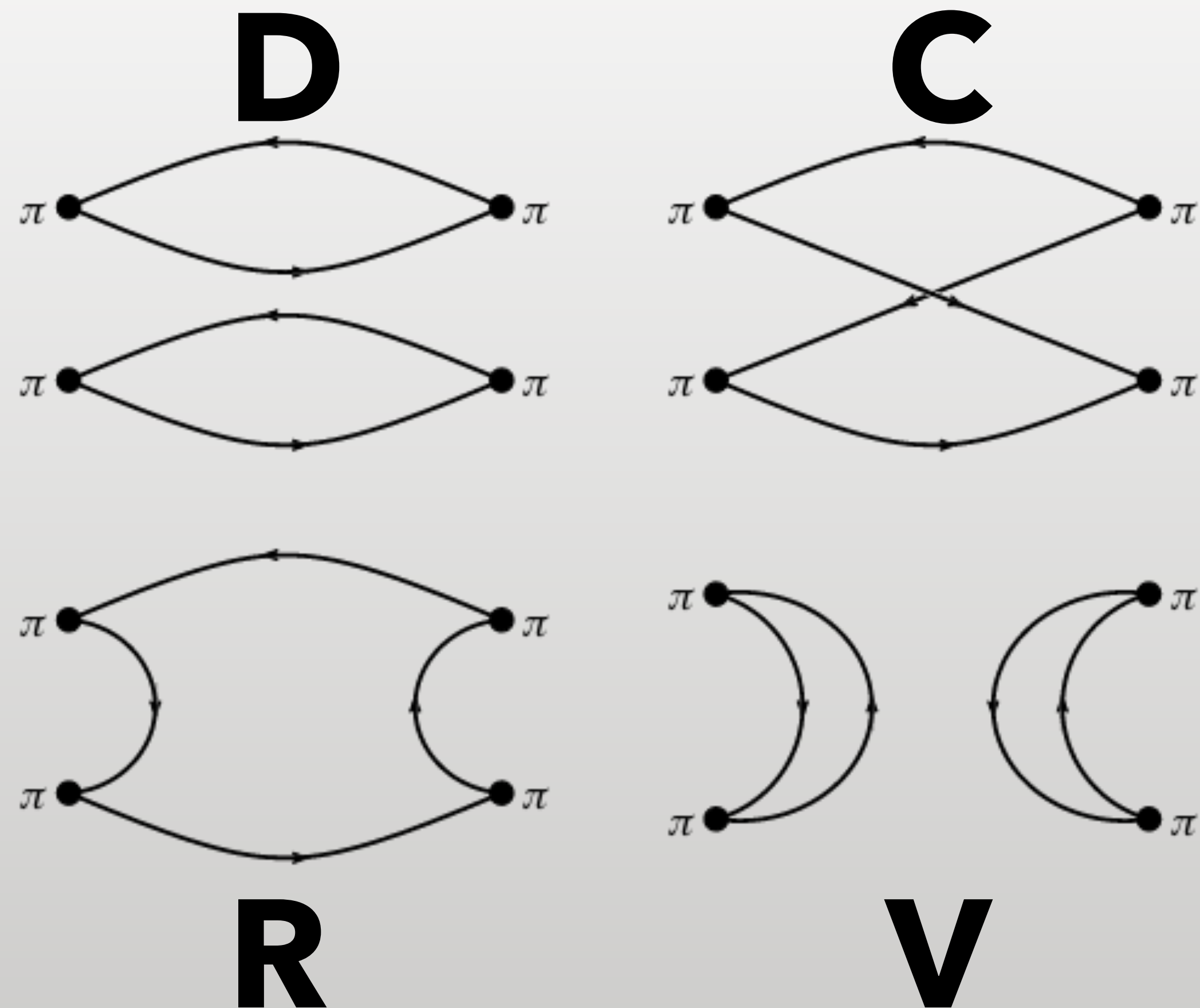
▶ We confirm the result of [2107.09974]

▶  $C_{\pi\pi}^{I=0} = 2D + 3C - 10R + 5V$

▶ QCD case:

▶  $C_{\pi\pi, QCD}^{I=0} = 2D + C - 6R + 3V$

▶ „Noisy“ vacuum term V



## BACKUP - MOMENTUM ON THE LATTICE

- Momentum via Fourier transform:

$$\cdot C(t) = \sum_{\vec{x}\vec{y}} e^{i(\vec{x}\vec{p}_x - \vec{y}\vec{p}_y)} \langle \mathcal{O}(\vec{x}) \mathcal{O}^\dagger(\vec{y}) \rangle$$

- In 2->2 Scattering 4  $\pi$  involved
  - -> 4 momenta ( $P_{in} = P_{out}$ )
  - 3 independent momenta

## BACKUP - MOMENTUM ON THE LATTICE

- Total momentum  $\sum \vec{p} \neq 0$ 
  - Probes different energy levels
  - Introduces noise
- Relative momenta between  $\pi \sum \vec{p} = 0$ 
  - Probes the same energy
  - Might yield higher overlap with states

## BACKUP - VARIATIONAL ANALYSIS

- Extraction of energy states via Generalized Eigenvalue Problem (GEVP)
- Calculate cross-correlator matrix from an operator basis:

- $C_{ij}(t) = \left\langle \mathcal{O}_i(t) \mathcal{O}_j^\dagger(0) \right\rangle$

- $\mathcal{O}_i = \mathcal{O}_{\pi\pi}(t, \vec{p} = \vec{p}_i)$

- $\vec{p}_0 = (0,0,0), \vec{p}_1 = (1,0,0), \vec{p}_2 = (1,1,0) \dots$

- $\lambda_k(t) \propto e^{-E_k t} (1 + \mathcal{O}(e^{-\Delta E t}))$ , for Eigenvalues of  $C_{ij}$

## BACKUP - DARK EFT (CHI PT)

- Low energy description of dark sector
  - „Dark Pions“ are the fundamental degrees of freedom ( $\chi$ PT)
- Low energy constants from lattice
  - Masses, decay constants, etc.
- 3- $\rightarrow$ 2 interaction via Wess-Zumino-Witten Term

## BACKUP - LATTICE

▶ Two sets of parameters:

▶  $\beta = 6.9, m_u = m_d = -0.9$

▶  $\frac{m_\pi}{m_\rho} \approx 0.81$

▶  $\beta = 7.2, m_u = m_d = -0.78$

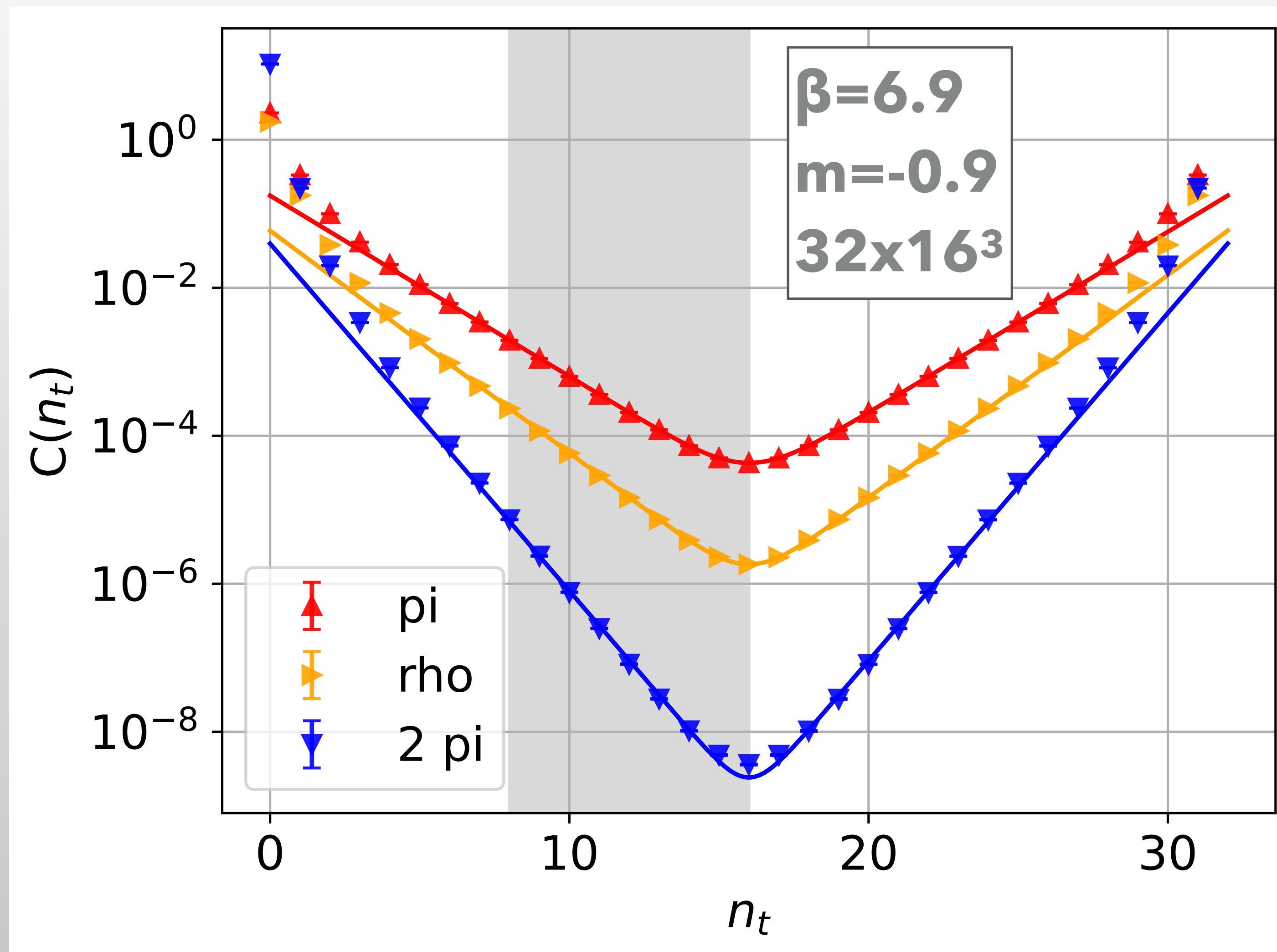
▶  $\frac{m_\pi}{m_\rho} \approx 0.78$

$L \times T \setminus \beta$	6.9	7.2
$20 \times 10^3$	1273	195
$24 \times 12^3$	2904	150
$24 \times 14^3$	942	425
$32 \times 16^3$	546	265
$48 \times 12^3$	251	X
$64 \times 12^3$	94	X

▶ Number of uncorrelated configurations per choice of parameters

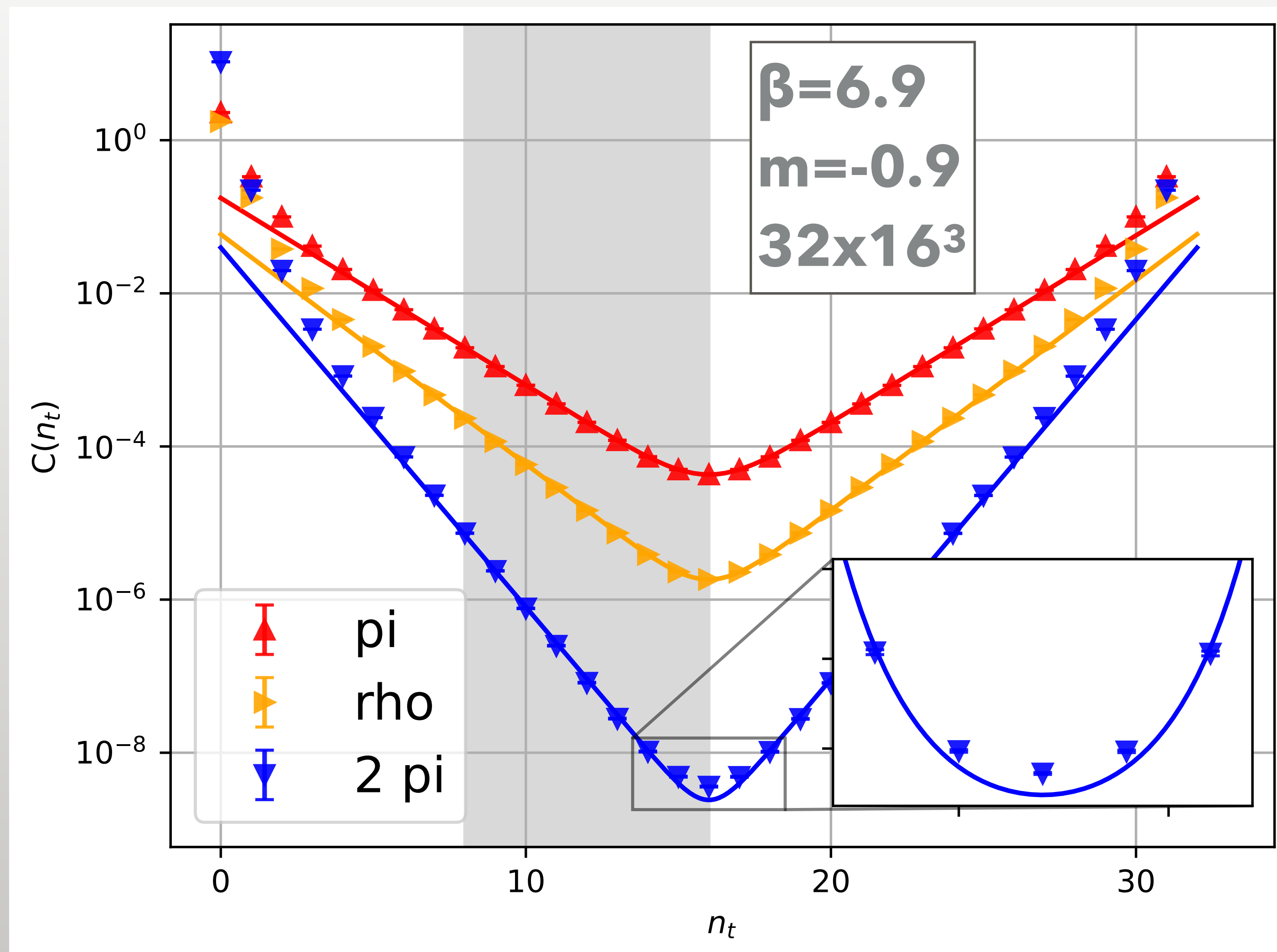
## BACKUP - CORRELATOR

- ▶ Correlation function
  - ▶ Only for the largest lattice ( $32 \times 16^3$ )
- ▶ Fit works fine for  $\pi$  and  $\rho$  but not for  $\pi\pi$
- ▶ Correlator of  $\pi\pi$  has lowest values



## BACKUP - CORRELATOR

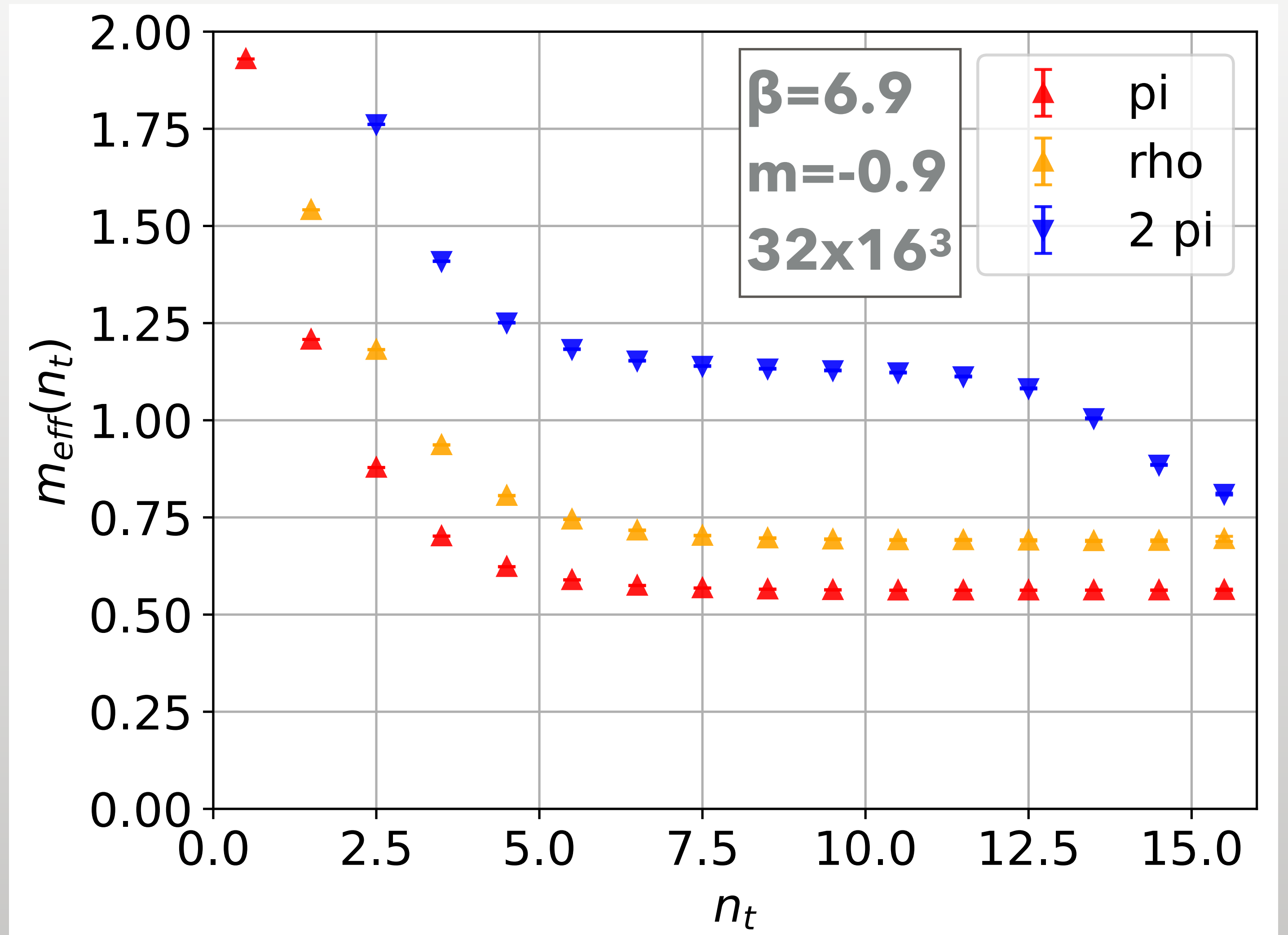
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## BACKUP - EFFECTIVE MASS

- ▶ Effective mass should show plateau when ground state dominates
- ▶ Second drop off closer to the center of the lattice
- ▶ Low values of the correlator
- ▶ Needs further investigation



## BACKUP - DERIVATIVE METHOD

- ▶ Redefinition of  $C$  with its derivative
- ▶ Constant cancels

$$\begin{aligned}\tilde{C}(n_t + 1) &= C(n_t) - C(n_t + 2) \\ \tilde{C}(n_t + 1) &\propto \sinh\left(\left(\frac{N_T}{2} - n_t\right)E_0\right)\end{aligned}$$

- ▶ Downside: Loss of two time-steps

## BACKUP - EFFECTIVE MASS

- ▶ Now: Nice plateaus for every correlator

