

**Plasma Physics Laboratory - ERM/KMS - Brussels - Belgium**

# **Transient versus steady state solutions: a qualitative study**

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## Intro: Philosophy of this talk

2

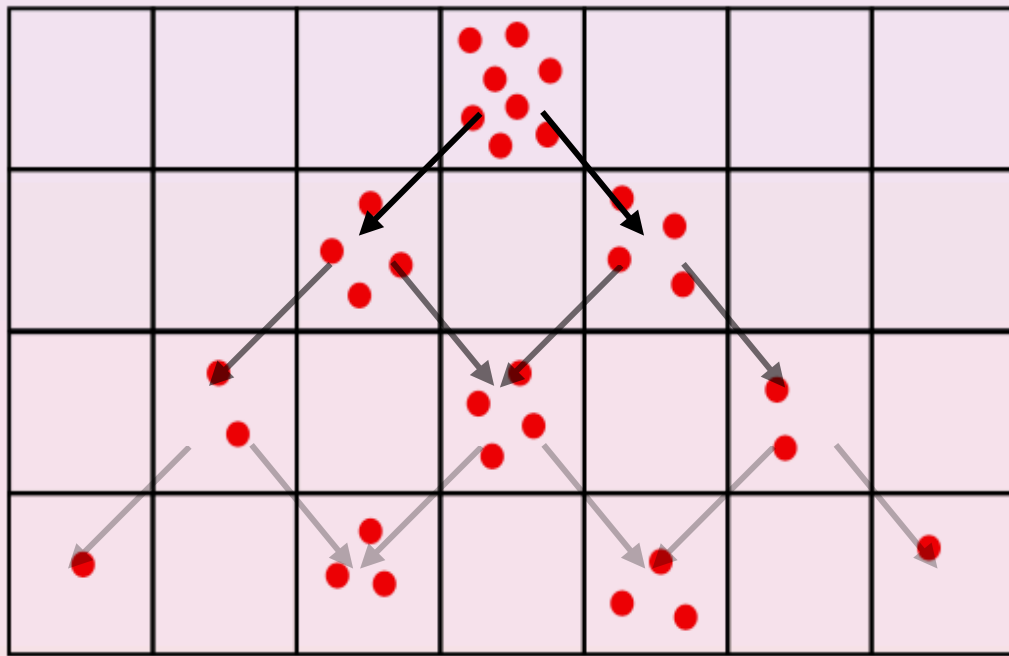
- Steady state often assumed as “given” —> studies often done using **steady state version of equations**.
- “**Snap shot**” approach (using *steady state equations* but plugging in experimentally known or guessed profiles) popular; artificially **hides** effects that may be key ...
- Reaching+holding+leaving* actual steady state **not evident** and deserves attention in its own right.
- Most *present-day* machines do not manage to have very long flat tops. **Transient effects rule rather than exception ...**
- Solving the actual ***time-dependent version of the equations*** increases realism, providing better insight and hence steerability.
- The present talk highlights some transient effects associated with plasma heating using 2 brutally simplified diffusion-convection models.*** More sophisticated models (many of which exist) sidestepped to illustrate bare elementary effects. Hopefully this talk is tickling specialists into looking into shown effects with their better suited models ...

[All examples in this talk are loosely inspired on high-performance JET discharges.]

## Diffusion: the verrrrrry basics

Random walk dynamics for  $D=1\text{m}^2/\text{s}$  &  $V=0\text{m/s}$  (normalised step and time):

$x-x_{\text{ref}}=-3$        $x-x_{\text{ref}}=-2$        $x-x_{\text{ref}}=-1$        $x-x_{\text{ref}}=0$        $x-x_{\text{ref}}=1$        $x-x_{\text{ref}}=2$        $x-x_{\text{ref}}=3$



Diffusion:  $\langle (x-x_{\text{ref}})^2 \rangle = D t$ :  
BI-directional

Convection:  $\langle x-x_{\text{ref}} \rangle = V t$ :  
MONO-directional



Green's function solution:

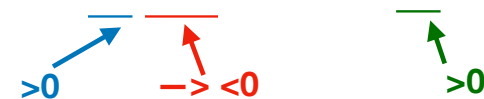
$$\frac{\partial U}{\partial t} = \frac{\partial}{\partial x} \left[ D \frac{\partial U}{\partial x} + V U \right] + S$$

$$S = \delta(t - t_o) \delta(x - x_o)$$

$$U(x) = \frac{1}{[4\pi D(t - t_o)]^{1/2}} \exp - \left[ \frac{[x - x_o + V(t - t_o)]^2}{4D(t - t_o)} \right]$$

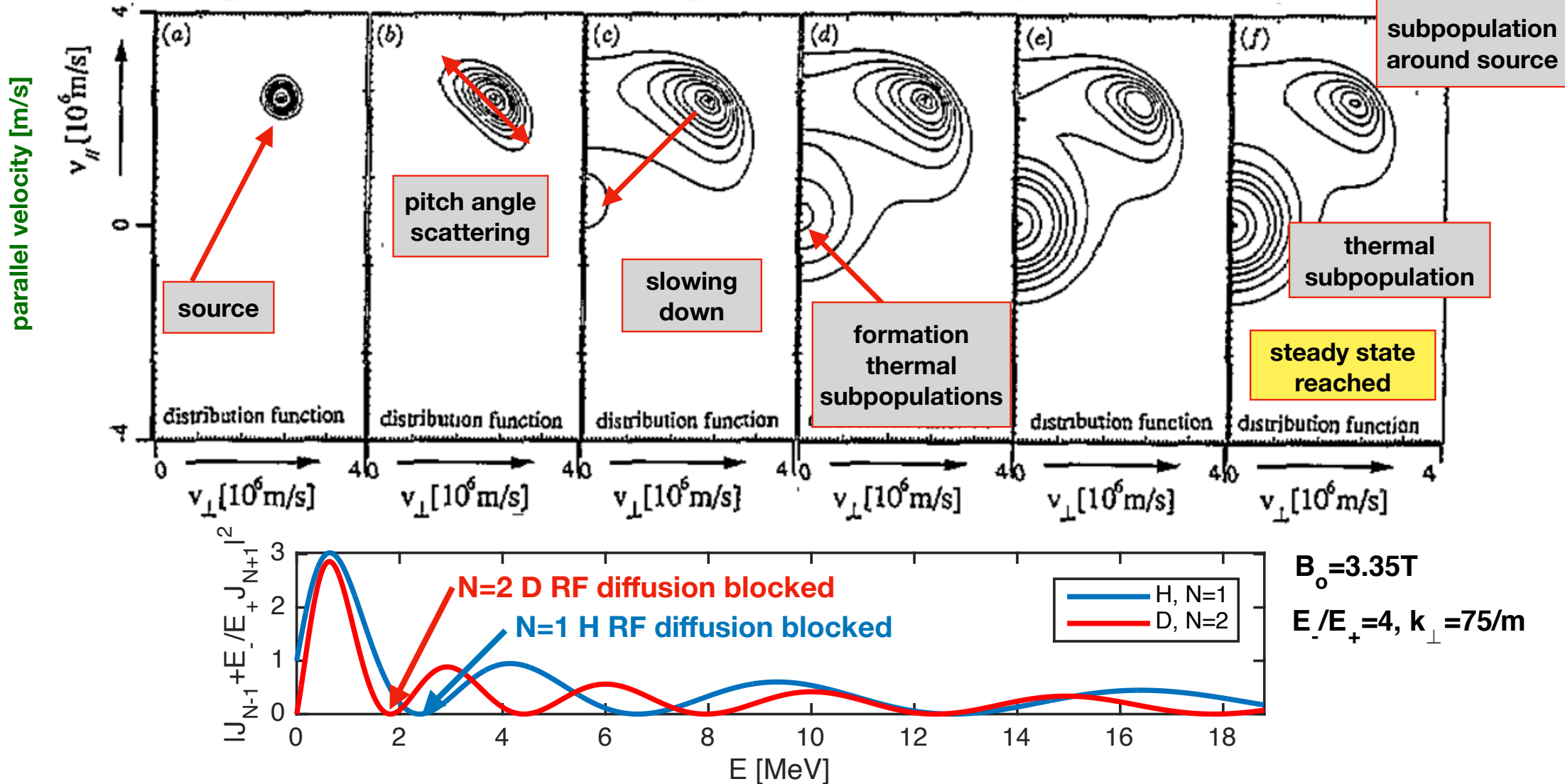
In steady state bi-directionality diffusion *masked*:

$[D dU/dx] + \text{integral } S = 0 \rightarrow \text{net "up-hill" movement}$



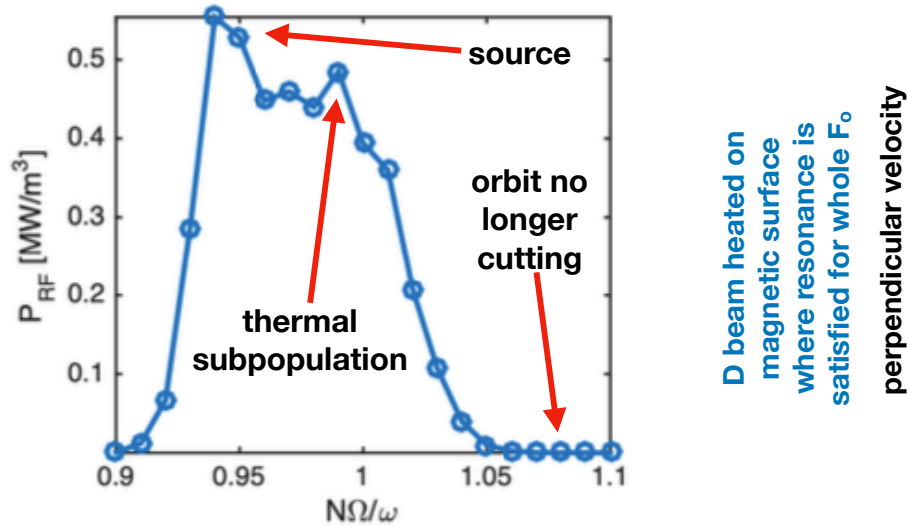
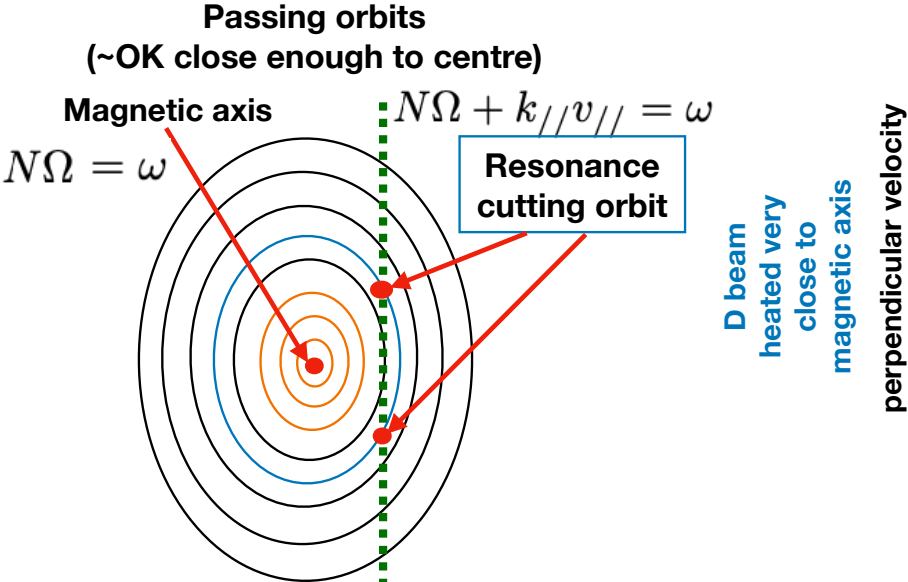
**[Almost simplest-on-earth] Fokker-Planck equation**

# [Almost simplest-on-earth] Fokker-Planck equation 2D Fokker-Planck à la Stix 5

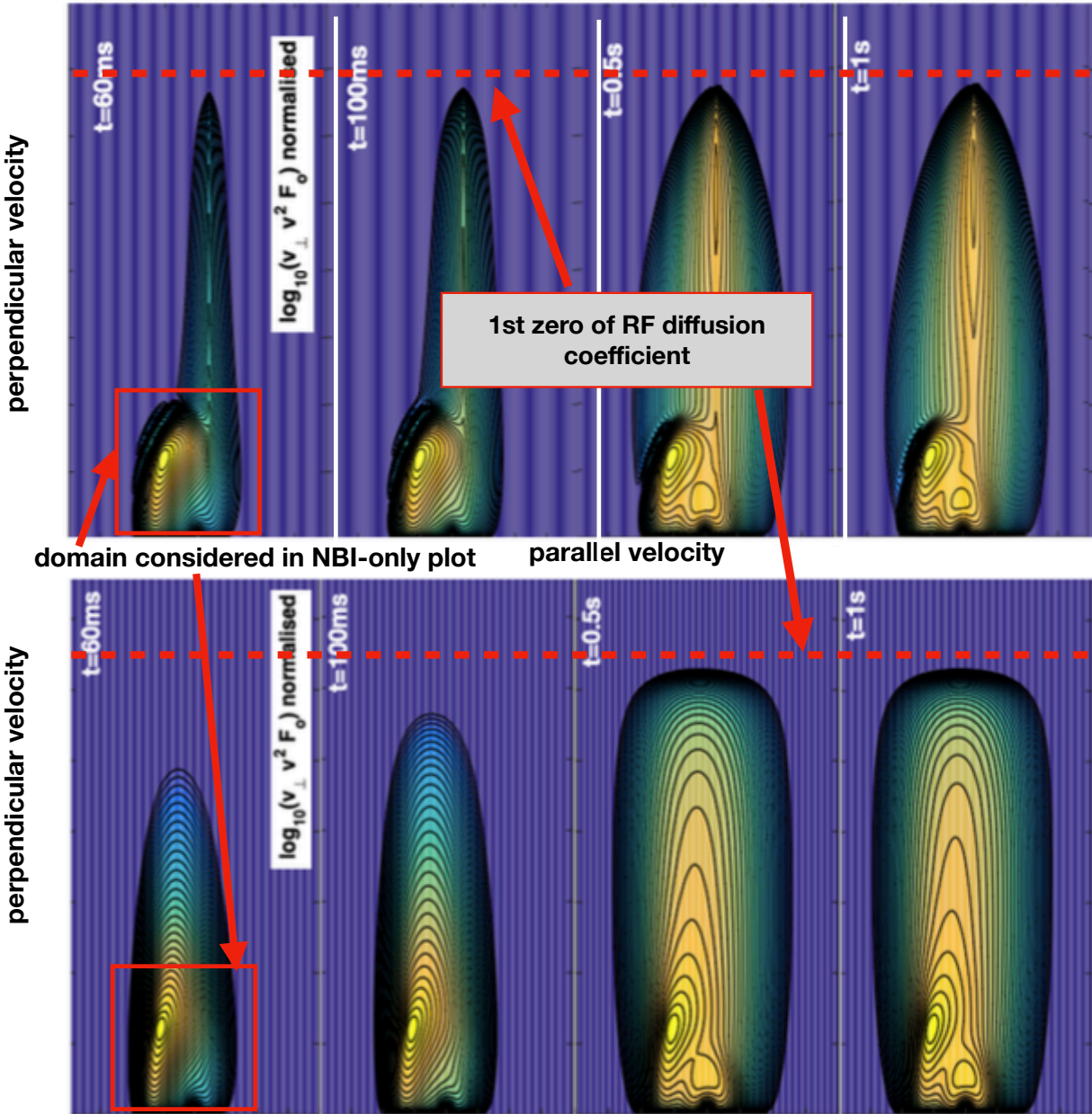


- Pitch angle scattering more efficient than slowing down
- Slowing down brings particles to thermal region; pitch angle scattering makes distribution uniform

# Heating localisation & transient effects



≠ location ≠ tail evolution





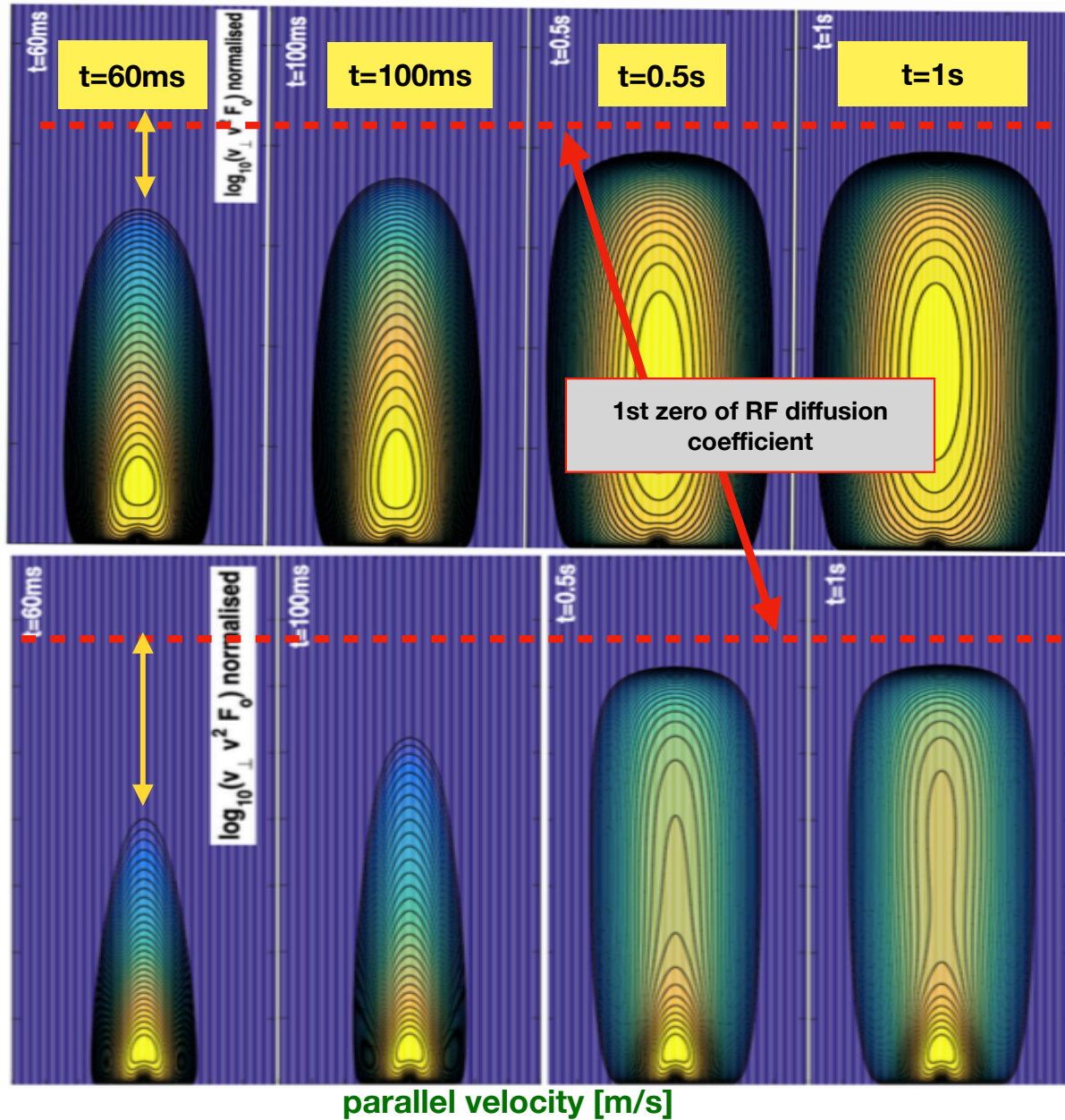
# RF scenario

$\log_{10}$  energy density in terms  
of  $v_{\text{perp}}$  &  $v_{\text{par}}$

D majority N=2 RF heating

H minority N=1 RF heating

perpendicular velocity [m/s]



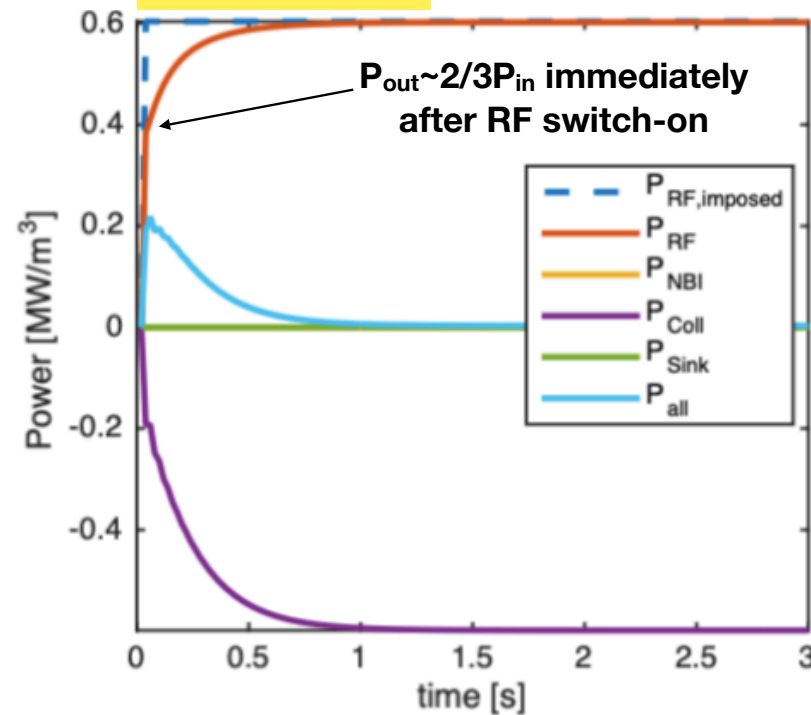
It takes time to build  
a saturated tail; N=1  
more “fat” in //  
direction and more  
populated at low  $v$ .  
Details depend on RF  
scheme as well as on  
type of population

# Convergence towards steady state

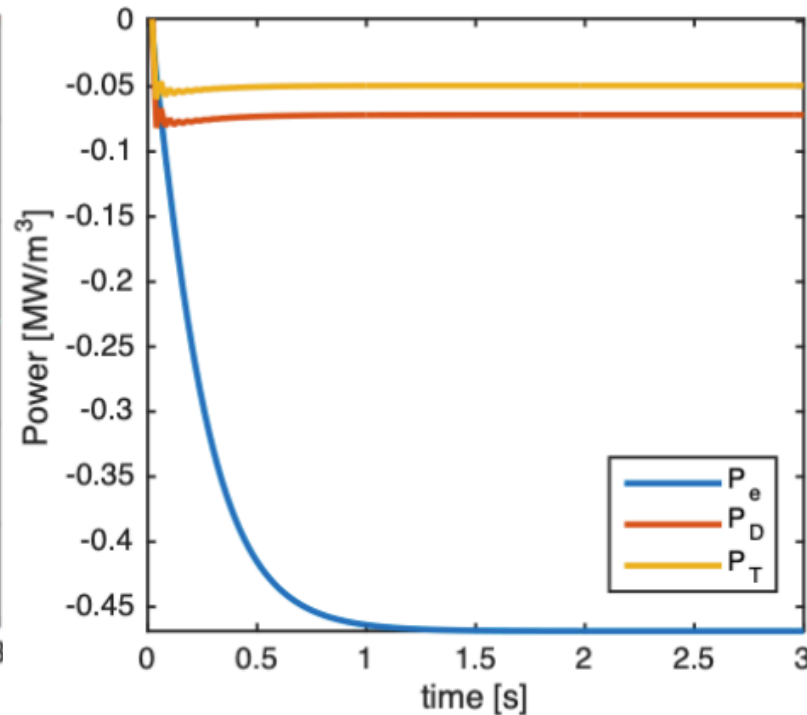
8

Power balance

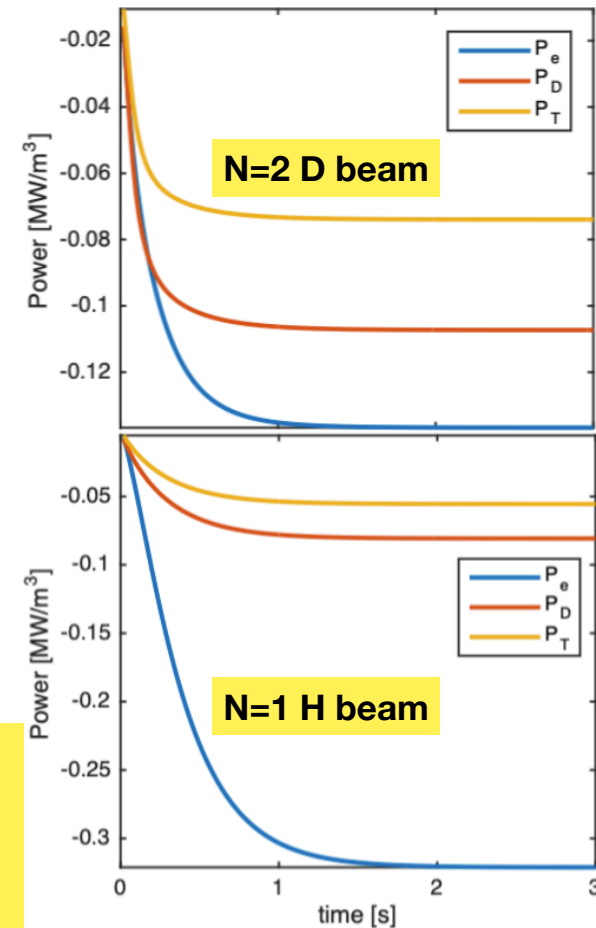
N=1 H minority



Collisional power redistribution



Collisional power redistribution

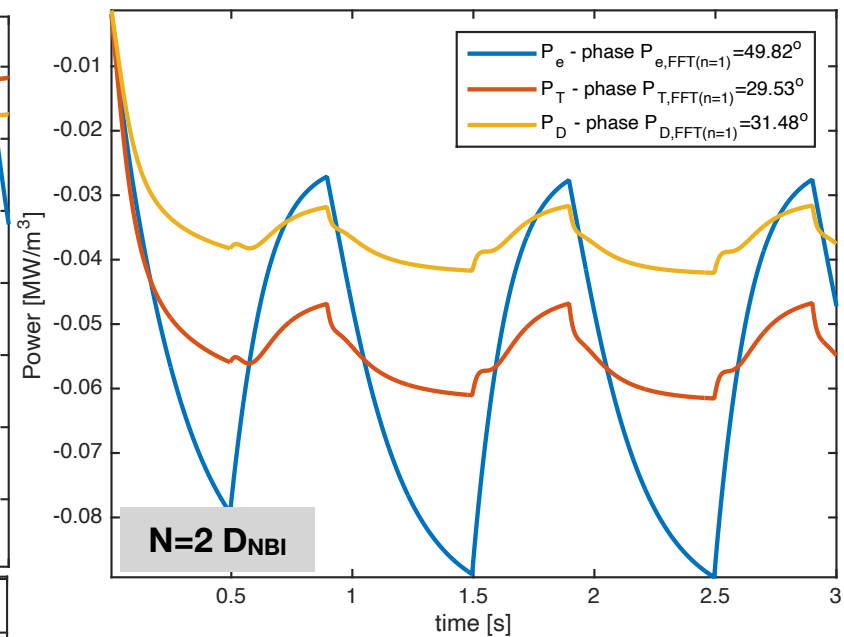
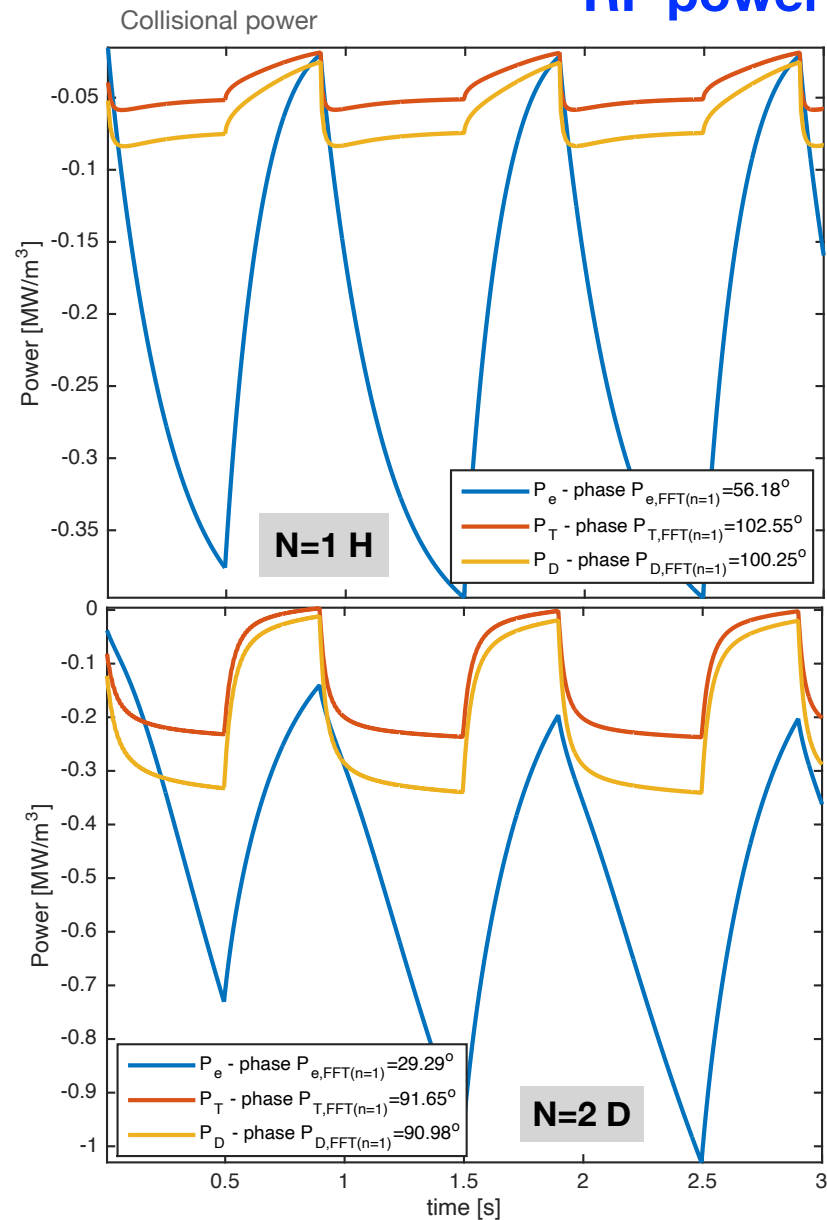


Reaching steady state  $\neq$  for  $\neq$  scenarios and populations:

- balance with ions faster than balance with electrons (if fast ion response desired, heat ions directly!): slower convergence in high energy region where collisions are less efficient
- overall convergence  $\sim 4x$  slowest characteristic time
- N=1 heating quickly reaches  $P_{in} \sim P_{out}$
- N=2 heating steady-state slower although RF faster: fast tail & thermal distribution need to reach equilibrium
- vessel = Faraday cage: prefer N=1 heating even when N=2 is efficient in steady state



## RF power modulation



parameters inspired on JET high density H-mode

full to 10% power -  $f_{mod}=1/s$  - duty cycle 60/40

Phenomena evolving over various modulation periods render interpretation modulation effects more complex ...

**[Definitely simplest-on-earth] Heat and particle transport equations**

## [Definitely the simplest-on-earth] Transport equations

	Diffusion	Convection	Source
<b>Particles</b>	$\frac{\partial N(\rho, t)}{\partial t} = \frac{1}{\rho} \left[ \rho [D_N N' + V_N N] \right]' + S_N$		
<b>Energy</b>	$\frac{\partial E(\rho, t)}{\partial t} = \frac{1}{\rho} \left[ \rho [D_E E' + V_E E] \right]' + S_P$		

Assume a *circular* cross section *without magnetic shift*, omit about toroidal curvature, take known D and V, and simple sources.

1-fluid instead of e, i<sub>bulk</sub>, i<sub>impurities</sub>

BC: flux=0 at axis and (negligible) N & T imposed at edge.

**Steady state conservation equation:**

$$\left[ \rho (D_N N' + V_N N) \right] \Big|_{a_p} = - \int_0^{a_p} d\rho \rho S_N$$

particle flux  $F_N$

$$\left[ \rho (D_E E' + V_E E) \right] \Big|_{a_p} = - \int_0^{a_p} d\rho \rho P$$

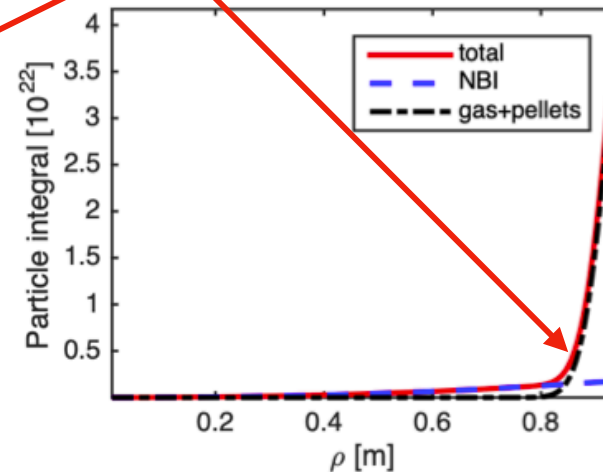
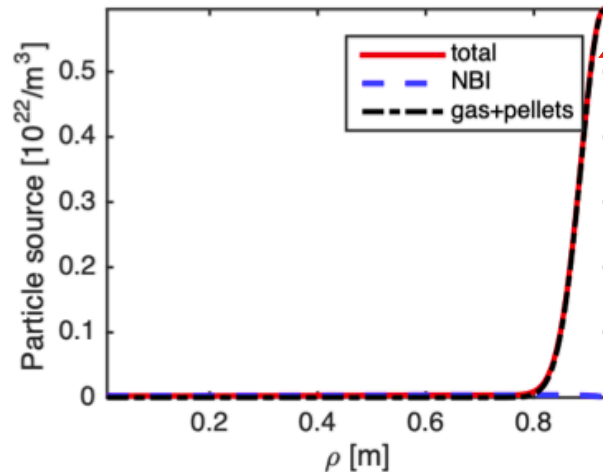
energy flux  $F_E$

Whatever escapes or is brought in has to escape or enter via the edge

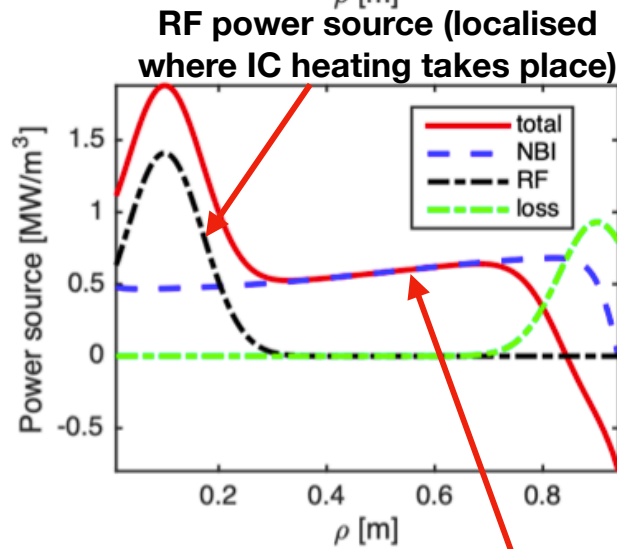
e.g. larger *integrated* source requires larger *edge* D or larger *edge* gradient if V=0

## Source profiles (steady state)

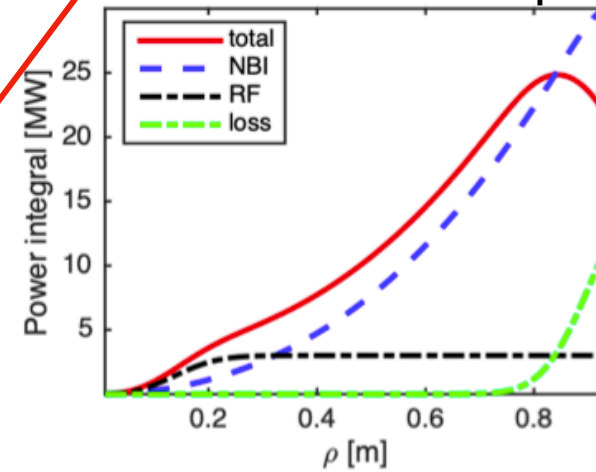
Gas or pellet source (ionisation):  
localised near edge



Beam source small for  
particles w.r.t. gas but  
dominant for energy  
w.r.t. RF



Loss (localised where W-  
dominated radiation takes place)

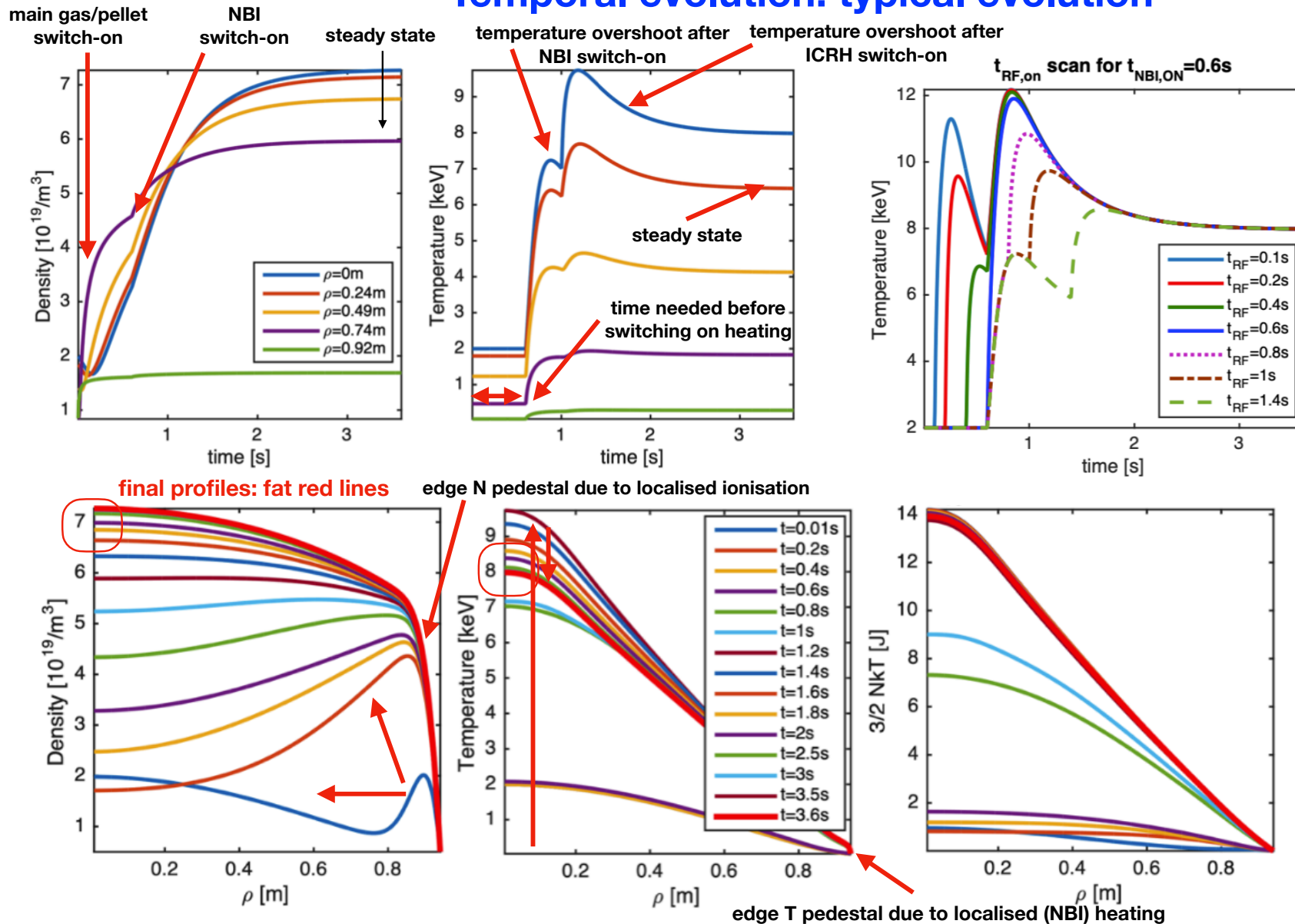


NBI source  
computed from  
ionisation path  
with actual density

Beam power source (broad:  
ionisation & beam intensity)

# Temporal evolution: typical evolution

13



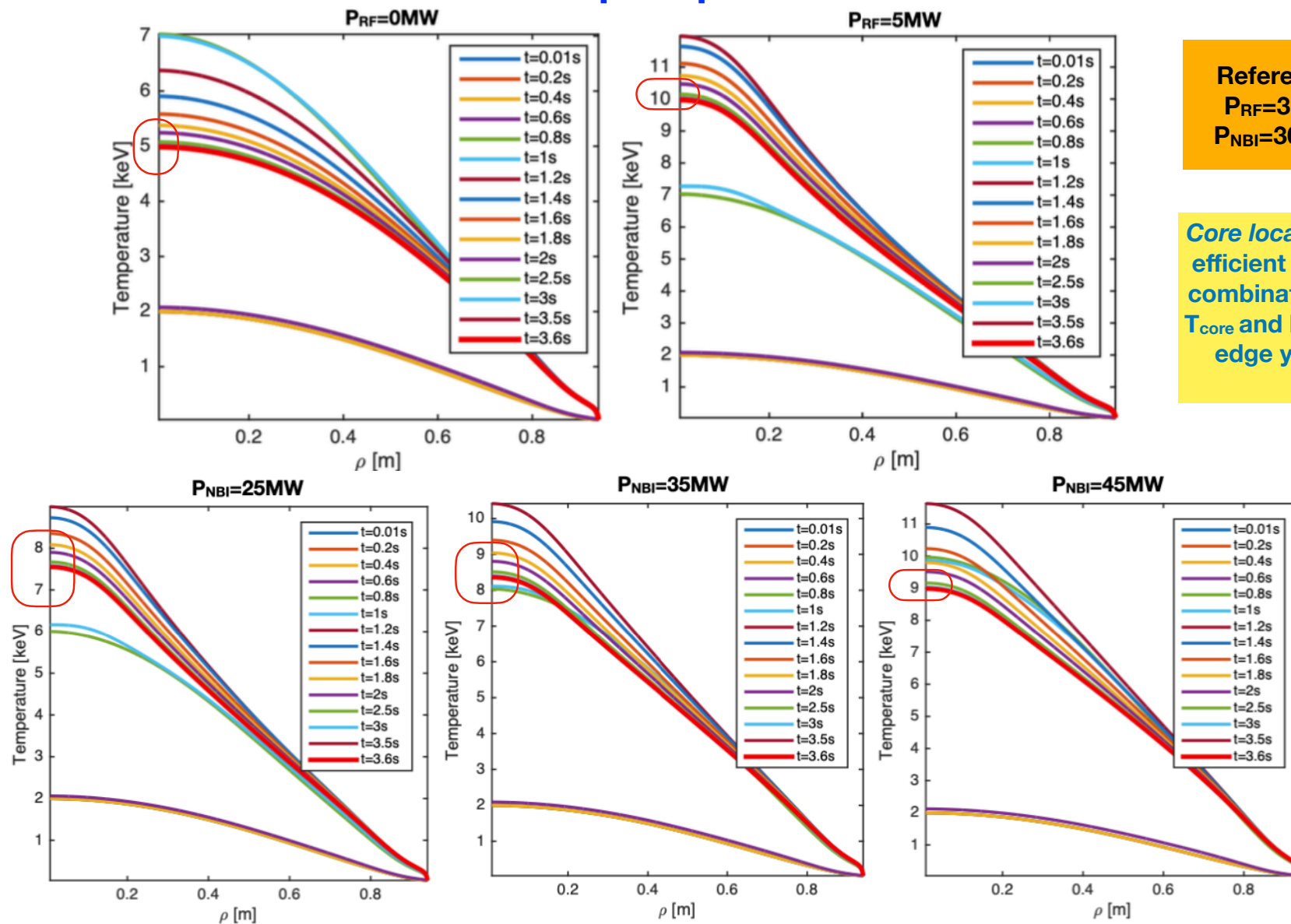
Optimisation timing allows (transiently) boosting performance

Actual start-up phase faked: start when low density profile already set up.  
Strongly inverted density profile at start (with impact on NBI profile ...)

Temperature overshoot, then gradual convergence towards steady state; relative timing switch-on important!

Energy reaches steady state much quicker than N or T

## Impact power sources



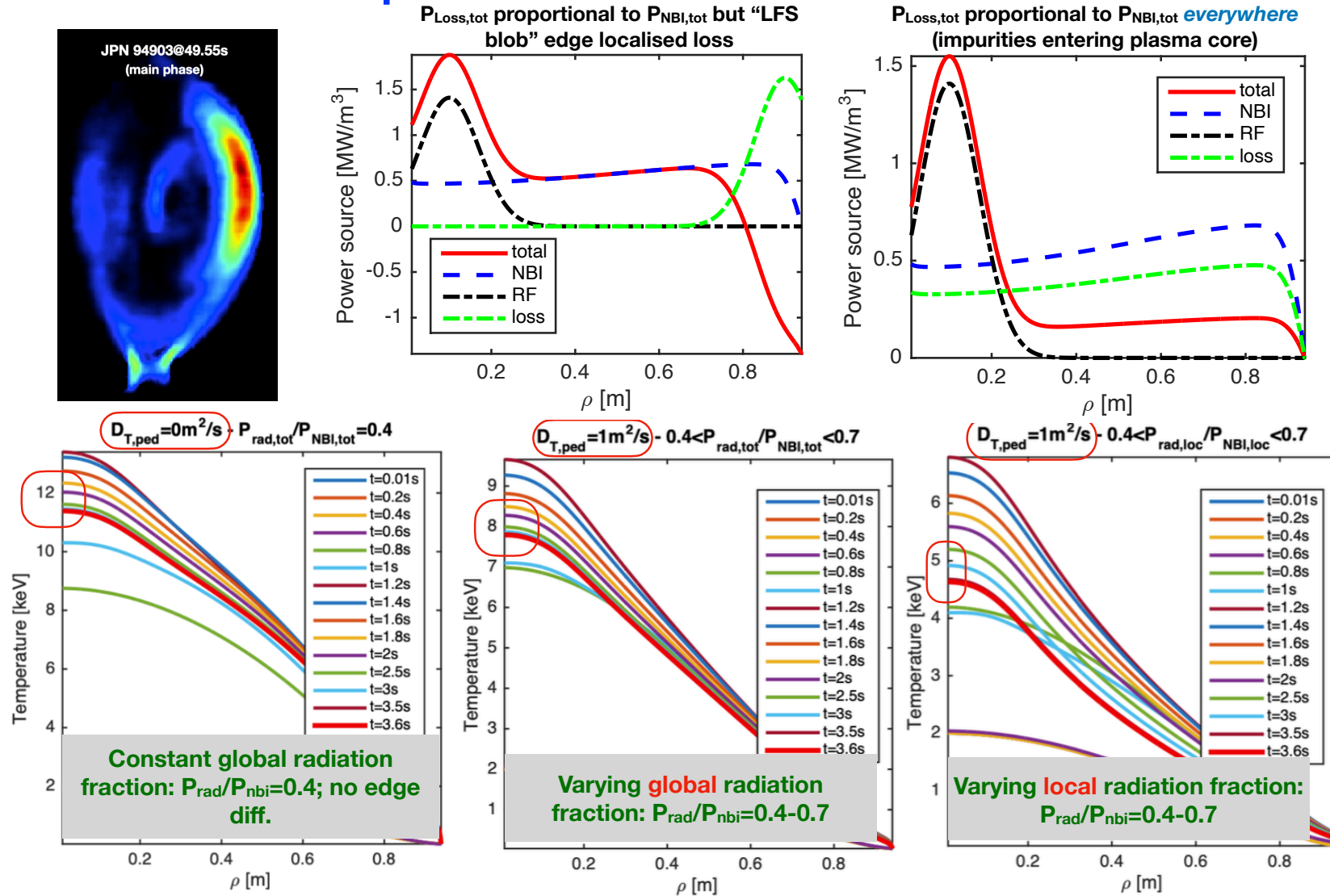
Reference:  
 $P_{RF}=3MW$   
 $P_{NBI}=30MW$

Core localised ICRH is much more efficient than NBI (volume effect); combination RF & NBI allows high  $T_{core}$  and broad T profile but P near edge yields increased loss by diffusion ...



## Temporal evolution: increased losses

15

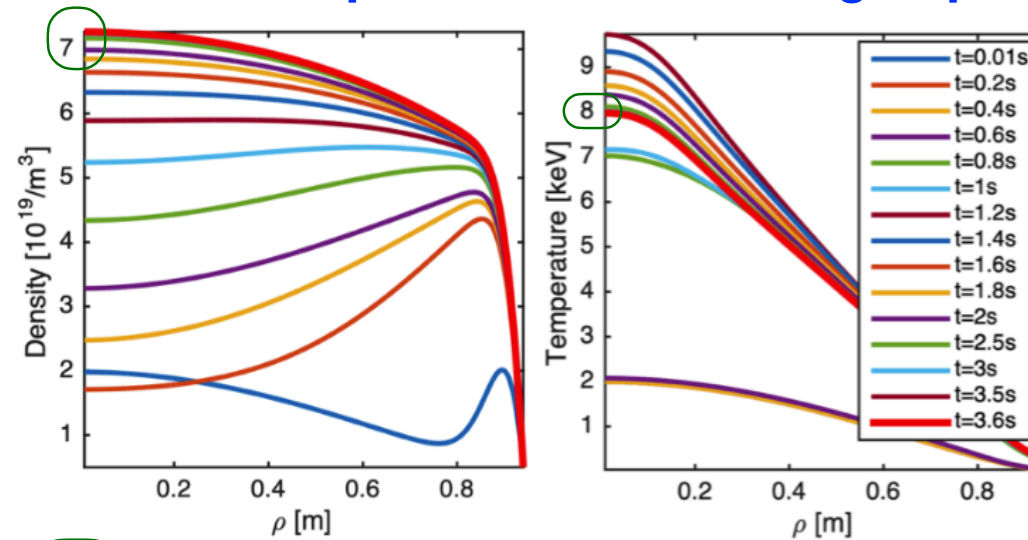


≠ impact of ≠ localisation: ≠ sensitivity relative position sources/losses  
(e.g. core loss much more damaging than edge loss: edge loss first requires transport to make effect sensed ...)

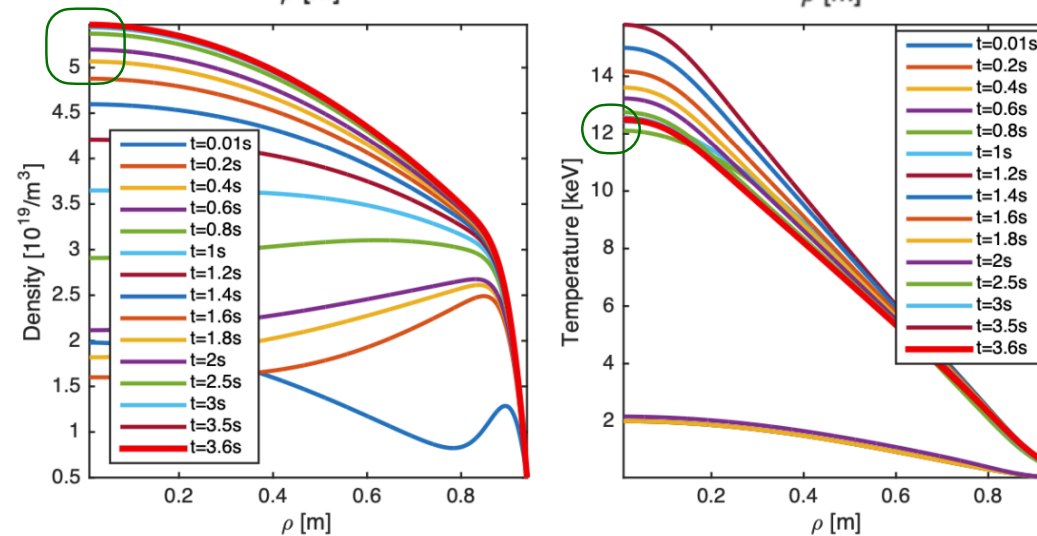
## Forcing increased temperatures: reduced gas/pellets source

16

$S=4 \times 10^{22}/s$



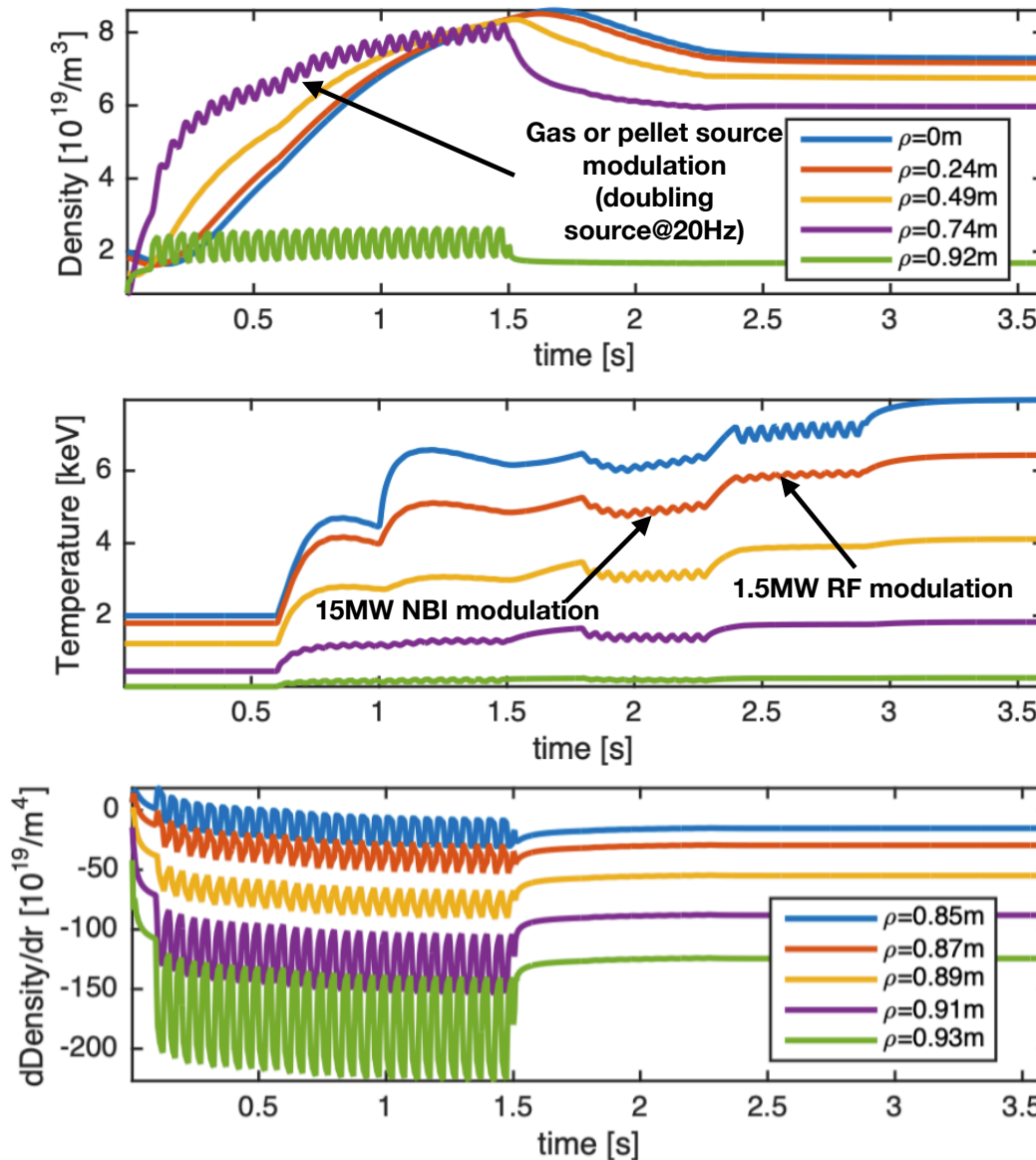
$S=2 \times 10^{22}/s$



Reduced gas or pellets yields lower density and hence allows higher temperatures, transiently as well as in steady state; high T may be profited from in practice to trigger entering  $\neq$  path of the discharge

## Particle or energy source modulation

17



Particle source modulation strongly affects edge (e.g. pedestal gradients cranked up) but weakly affects core (if at all)

NBI source modulation weak w.r.t. RF modulation (due to  $\neq$  localisation)

RF excellent tool for transport analysis, pacing (impacting on ST), ...

Pellets as tool for forcing ELMS: increased spatial gradients

(simple tool describes steepening gradient but cannot provide info on dependence critical gradient to cause ELM crash ...)

# Multiple species interacting: coupled FP & transport equations

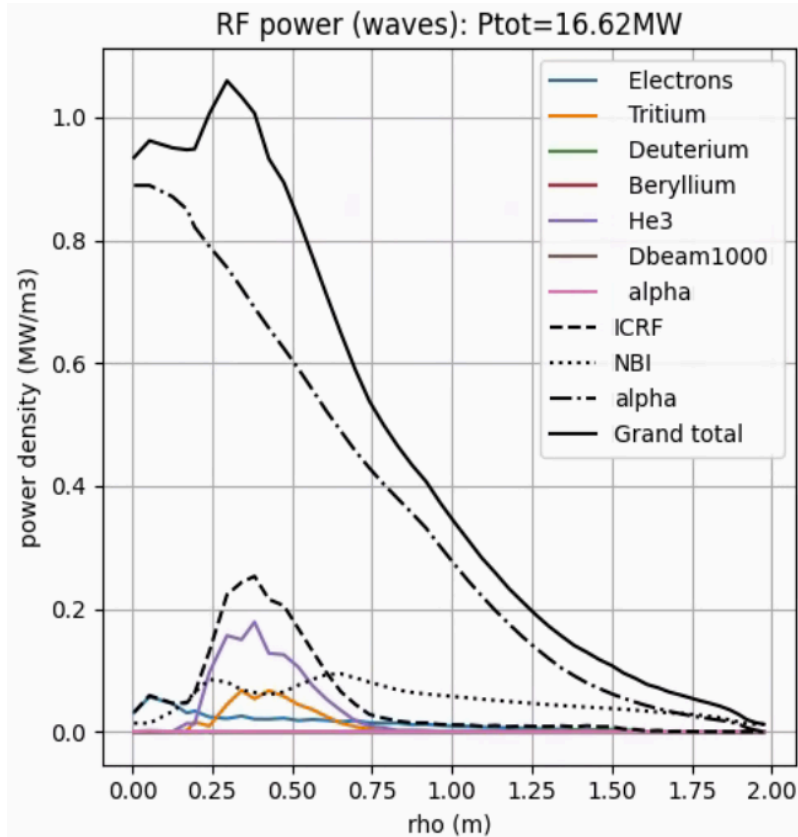
## Proportionality factor cross-talk species Fokker-Planck:

Usual reasoning “heated heavy ions heat (lighter) fuel ions”  
but power partly “stolen” from fuel ions when considering **back-reaction** when accounting from all ion species interacting

$$\Gamma^{a/b} = \frac{n_b q_a^2 q_b^2 \ln \Lambda^{a/b}}{4\pi\epsilon_0 m_a^2}$$

a/alpha: test species  
b/beta: background species

ITER example: (<sup>3</sup>He)-D-T



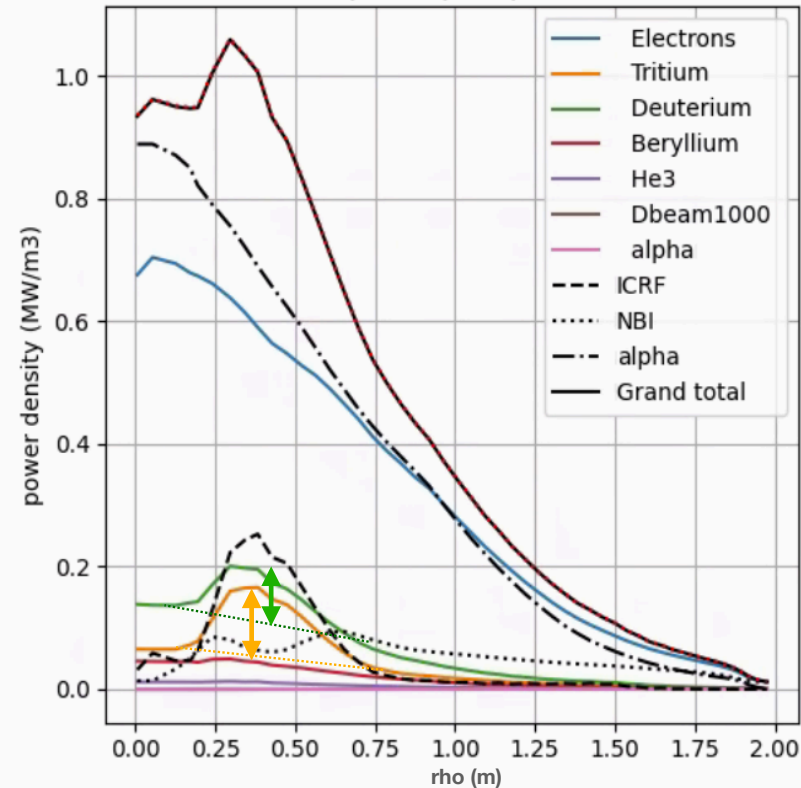
## Equipartition term transport equation

collisional interaction between **e & i** much **slower**  
than between **i & i**; reaching steady state takes time:  
distributions do not at all fill same v-space

$$\bar{\nu}_e^{\alpha/\beta} = 1.8 \times 10^{-19} \frac{(m_\alpha m_\beta)^{1/2} Z_\alpha^2 Z_\beta^2 n_\beta \lambda_{\alpha\beta}}{(m_\alpha T_\beta + m_\beta T_\alpha)^{3/2}} \text{ sec}^{-1}$$

P<sub>RF</sub>=17MW, 2% He<sup>3</sup>, off-axis

Collisional power per species for ETS



Be not directly heated but  
receiving power via collisions

## Conclusions / discussion

- Simple models do not allow to describe physics quantitatively but they allow to qualitatively highlight that transient solutions may differ significantly from steady state solutions, something that can be exploited for shot optimisation.
- Tail formation takes time,  $\neq$  time scales of  $\neq$  mechanisms (N=1 vs N=2; RF vs NBI or vs collisions; pitch angle w.r.t. slowing down; i vs e)
- N=2 in Faraday cage forces  $|E|$  to increase while N=1 yields good absorption; use N=1 minority before switching on N=2 [reason for problematic behaviour initial stage with only N=2?]
- Absorption close to axis yields faster tail formation, higher power density in less particles
- NBI deposition  $\neq$  for  $\neq$  density and  $\neq$  species  $\rightarrow$  has impact on overall performance (e.g. T has more external deposition than D, and D than H)
- Diffusion being 2-directional allows more externally deposited energy “spilled” over the edge sooner
- $d\text{Flux}/d\rho = P_{\text{local}}$  i.e. any outward flux overcompensated by bigger inwards flux if  $P_{\text{local}} > 0$
- Temperature often overshoots and then relaxes to final state [as seen in DTE2]. Heating time matters.
- No steady state can be reached when radiation varies [as seen in DTE2]; exp. effort needed to ensure radiation is kept
- Location sources and losses strongly influences profiles; transient solutions can be very different depending e.g. on relative timings heat sources
- ICRH much more capable than NBI to modify core temperature for given density [as seen in DTE2]; NBI deposition always very broad (broadening T profile); hollow if high density or heavier beam ion
- Particle source modulation (e.g. pellets) mainly visible in edge; allows to force higher gradients (and force ELMs [as used in DTE2])
- Heat source modulation much more efficient for ICRH than for NBI [as used in transport studies]: localised absorption for given integrated power; fraction flowing to e and i is time dependent
- Operating at reduced gas/pellets allows higher temperatures [as seen in many exp. e.g. Hybrid and L-mode] (immediate from energy equation:  $P = dE/dt$ )

# The End

Words & Music by The Doors

Sl...~

(Repeat several times)

with ped.

D 3 C D

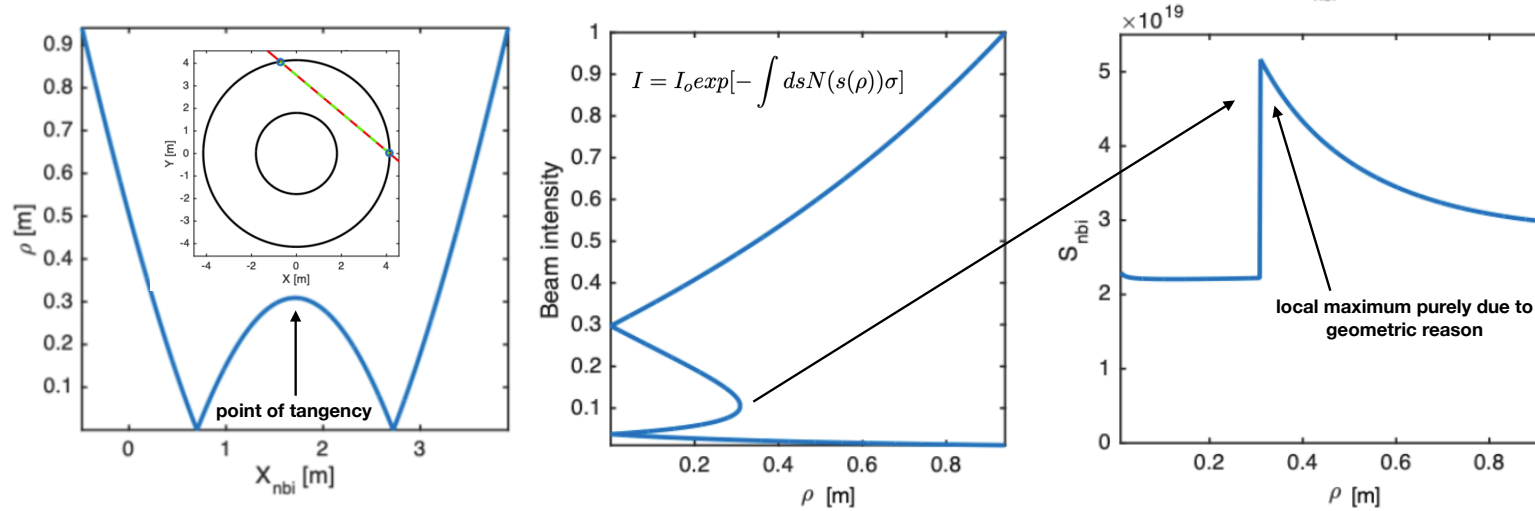
This is the end, beau - ti - ful friend.



**extra slides**

## Introductory notes: beam heating (here for $N_e=ct$ )

22



simplified sigma: [Kazakov, CMSS course on heating schemes]

