Long-Lived Particles to Probe the EWPT

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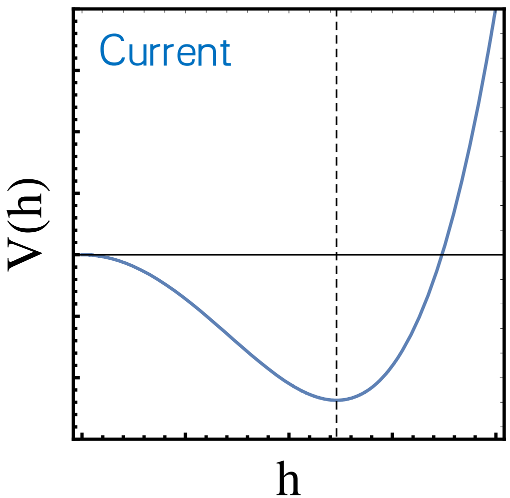
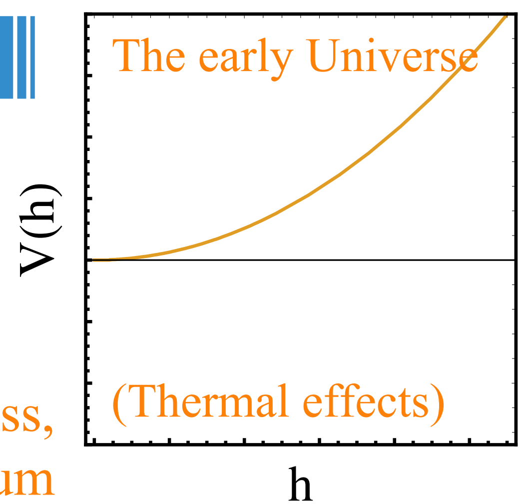
Phase transition in electroweak theory

EW symmetry restoration in the early Universe



W & *Z* bosons are massive;
Photon is massless,
Mexican-hat like

Phase transition

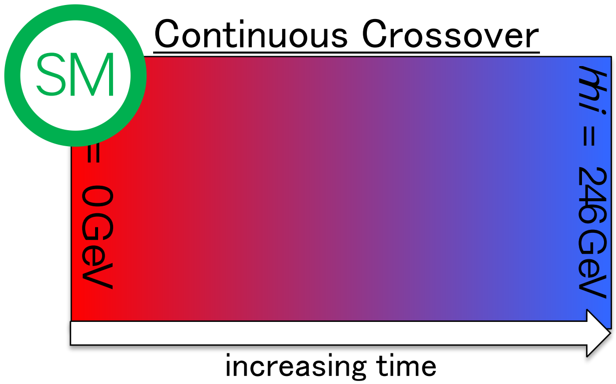


$SU(2)_L$ & $U(1)_Y$ bosons are massless,
True vacuum

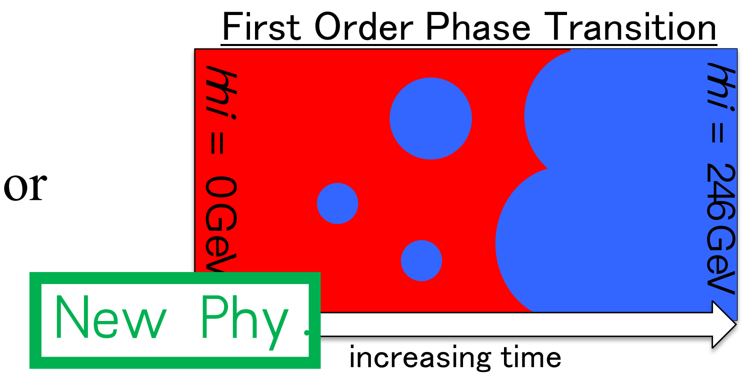
What is the pattern of EW phase transition (PT)?

It could be –

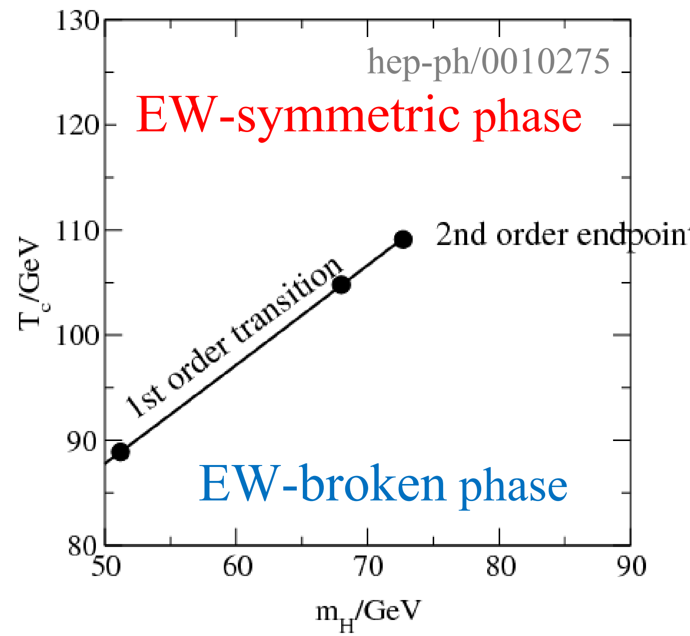
Figure from L.-T. Wang's talk in IHEP workshop



or



Lattice calculation shows the phase diagram ==>



Thus in the SM it is a crossover, since $M_h = 125$ GeV > 75 GeV;

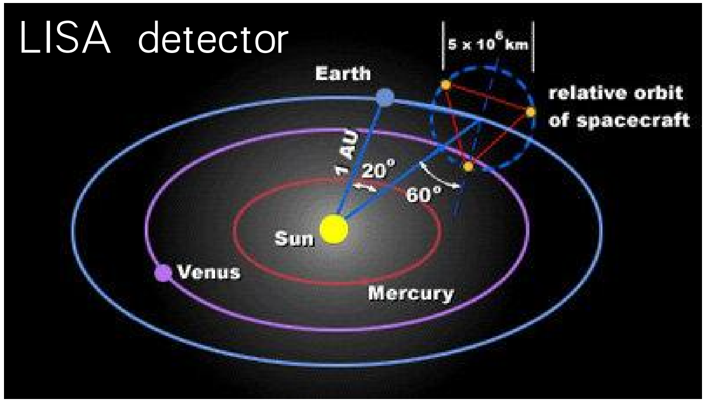
However, a 1st-order EWPT is more interesting.
(Needs new physics)

Why is a 1^{st}-order EWPT interesting?

- It's the essential ingredient of the EW baryogenesis.
- Acting as the background of very rich **dark matter** mechanisms
- Sources of the stochastic GWs：

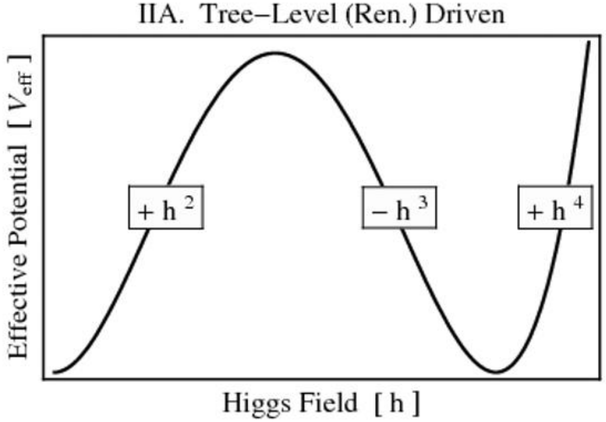
- Collision of the bubbles
- Sound waves in plasma
- Turbulance in plasma

EWPT GWs typically peak in mHz.



How to achieve a 1st-order EWPT?

Adding a barrier for the Higgs potential via new physics!
The decay between two vacua separated by a barrier.
The VEV of the Higgs field *jumps*.



Getting a barrier via the help of additional scalar field(s):

- SM + real singlet (xSM);
- 2HDM;
- Georgi-Machacek model;
- ……

We choose the **xSM** as the benchmark model.

- It's simple, but has captured the most important feature of EWPT;
- It can be treated as the prototype of many new physics EWPT models.

EWPT in the xSM (SM + real singlet)

We choose the **xSM** as the benchmark model.
It's simple, but has captured the most important feature of EWPT.

The scalar potential of the xSM

$$V = -\mu^2|H|^2 + \lambda|H|^4 + \frac{a_1}{2}|H|^2 S + \frac{a_2}{2}|H|^2 S^2 + b_1 S + \frac{b_2}{2}S^2 + \frac{b_3}{3}S^3 + \frac{b_4}{4}S^4$$

8 input parameters:
1 unphysical, 2 fixed by Higgs mass & VEV; 5 *free* parameters.

Expansion around the VEV

$$H = \frac{1}{\sqrt{2}}\begin{pmatrix} 0 \\ v+h \end{pmatrix}, \quad S = v_s + s, \quad \begin{pmatrix} h \\ s \end{pmatrix} = \begin{pmatrix} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix}$$

Mass eigenstates & the mixing angle.

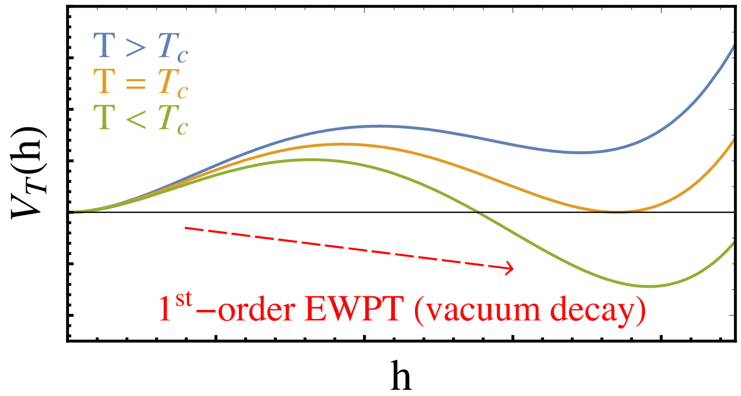
1^{st}-order EWPT in the xSM

At finite temperature:

$$V = -(\mu^2 - c_H T^2)|H|^2 + \lambda |H|^4 + \frac{a_1}{2}|H|^2 S + \frac{a_2}{2}|H|^2 S^2$$
$$+ (b_1 + m_1 T^2) S + \frac{b_2 + c_S T^2}{2} S^2 + \frac{b_3}{3} S^3 + \frac{b_4}{4} S^4$$

$$c_H = \frac{3g^2 + g'^2}{16} + \frac{y_t^2}{4} + \frac{\lambda}{2} + \frac{a_2}{24}, \quad c_S = \frac{a_2}{6} + \frac{b_4}{4}, \quad m_1 = \frac{a_1 + b_3}{12}$$

An Illustration --

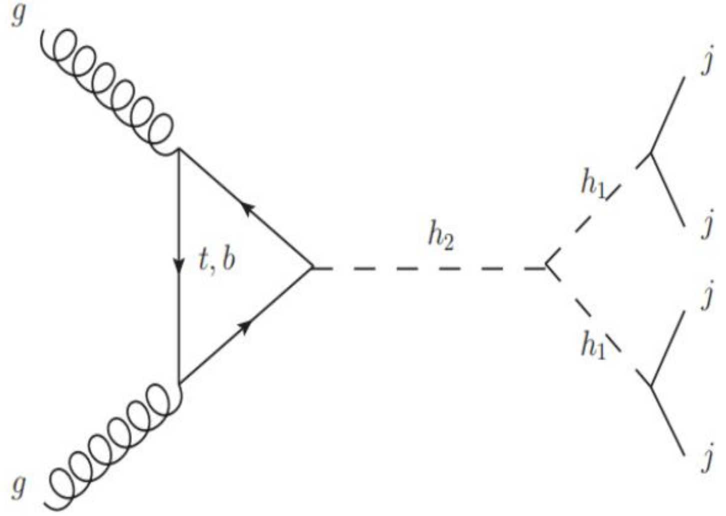


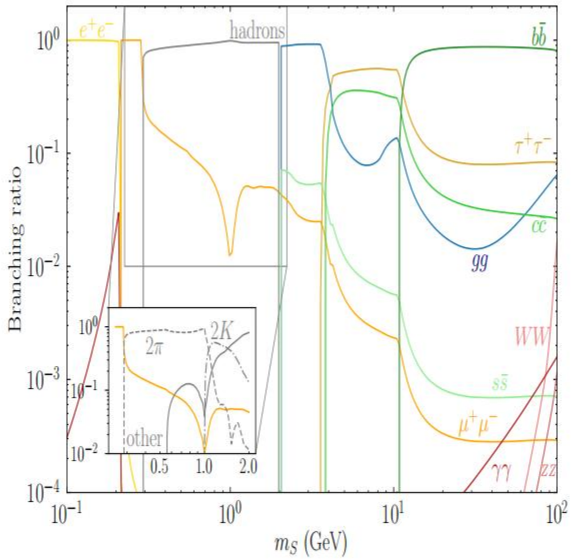
Question:

Can collider experiments probe the 1^{st}-order EWPT parameter space?

Production processes

$$gg \to h_{2(SM)} \to h_1 h_1 \to 4j$$



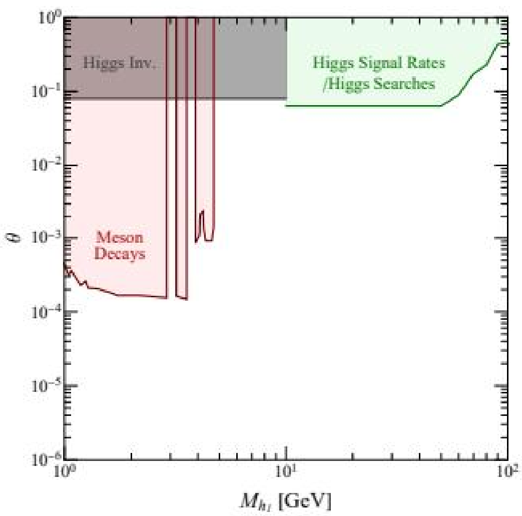


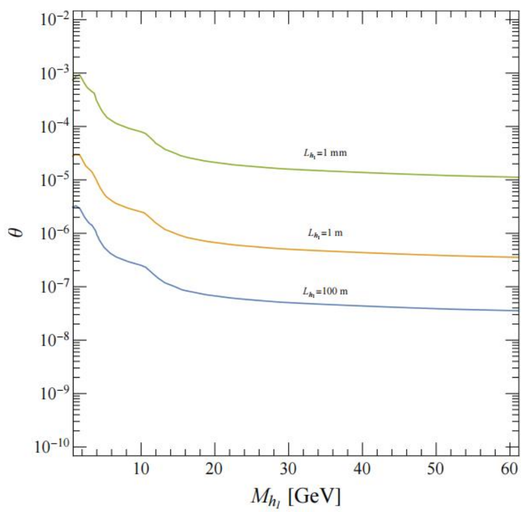
1st-order EWPT leads to large $BR(h_2 \to h_1 h_1)$.

h_1 decays into jets dominantly.

Long-lived Particles

The current limits on (M_{h_1},θ)





For M_{h_1}< 10 GeV, the current limits from rare meson decays at the LHCb, leads to h_1 as a long-lived particle (LLP).

LLP EWPT

From J. Kozaczuk, M. Ramsey-Musolf, J. Shelton, Phys.Rev.D 101 (2020) 11, 115035

The Higgs mixing is small

Approximate **Z_2 symmetry**

EWPT can happen as **two-step transition**

$$(h = 0, s \simeq 0) \to (h = 0, s \neq 0) \to (h \neq 0, s \simeq 0),$$

There are analytical bounds from two-step transition

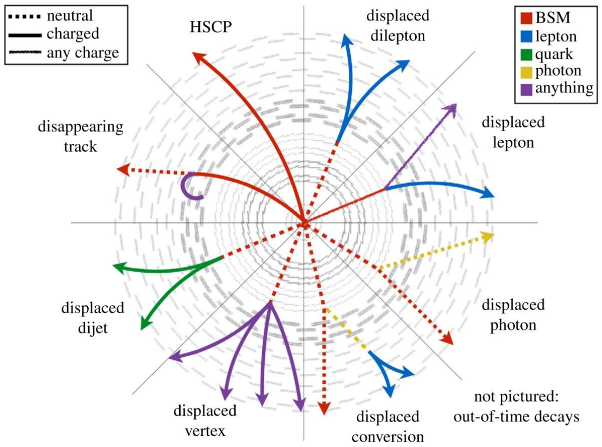
$$a_2 \gtrsim \frac{m_{h_1}^2}{4v^2}\frac{\Delta}{1-\Delta},$$

$$|b_3| > \sqrt{\frac{9}{4}b_4(2m_{h_1}^2 - a_2 v^2 + 2T_{\rm EW}^2\beta)},$$

$$b_4 \gtrsim \frac{m_{h_1}^4 \Delta}{4\lambda v^4(1-\Delta)},$$

Long-lived Particles

LLP is widely searched, of great interests experimentally and theoretically.



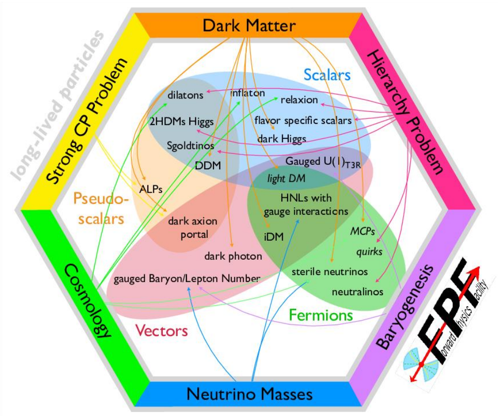
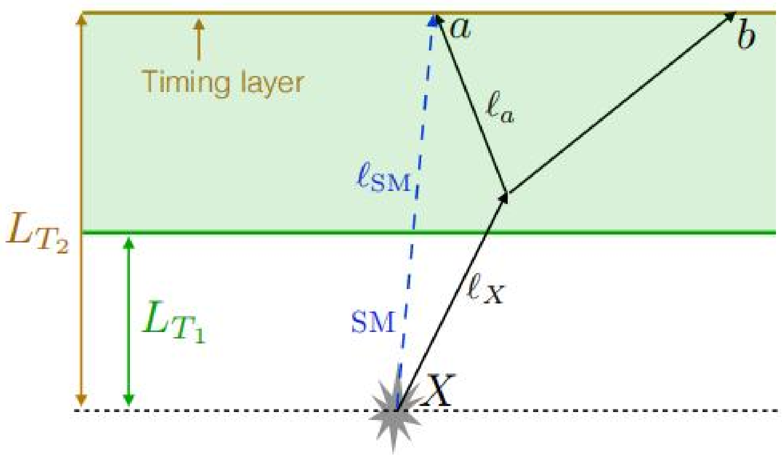


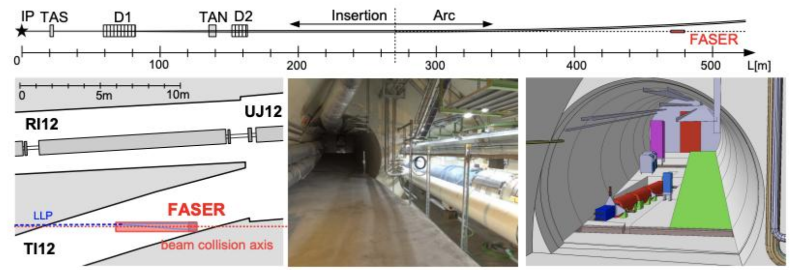
Figure from Albert De Roeck. Figure from 2203.05090.

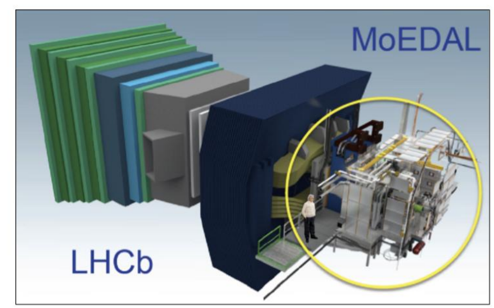
Light weakly-coupled particles as LLPs, are strongly motivated, including the light scalar responsible for 1st-order EWPT in the xSM.

Detectors for LLPs

CMS-Timing, FASER, MoEDAL-MAPP are to be operated at Run 3.







Many others, e.g. MATHUSLA and CODEX-b are in discussions.

CMS-Timing detector using the **time-delayed leptons/jets** as signals, while the other detectors using displaced vertex.

Connects to 1st-order EWPT

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Link between the number of events and 1st-order EWPT

$$N_{signal} = \sigma_{pp\to h_2} \times L \times BR_{h_2\to h_1 h_1}(a_2, M_{h_1}) \times BR_{h_1\to jj}{}^2(M_{h_1}) \times \epsilon_{kin}(M_{h_1}) \times \epsilon_{geo}(M_{h_1}, \theta)$$

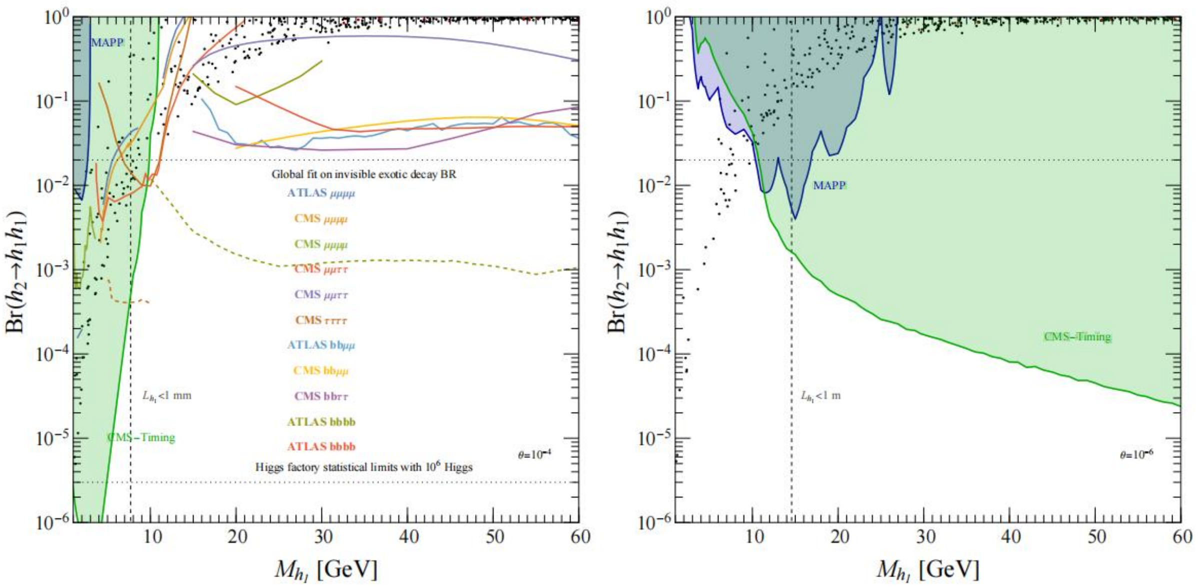
$$BR_{h_2\to h_1 h_1}(a_2, M_{h_1}) = \frac{\Gamma_{h_2\to h_1 h_1}}{\Gamma_{h_2}{}^{SM} + \Gamma_{h_2\to h_1 h_1}}$$

$$\Gamma_{h_2\to h_1 h_1} \propto (a_2 v)^2$$

LLP events are sensitive to $|H|^2 S^2$ couplings.

Sensitivity

Fixed $\theta = 10^{-4}(left),\ 10^{-6}(right)$.

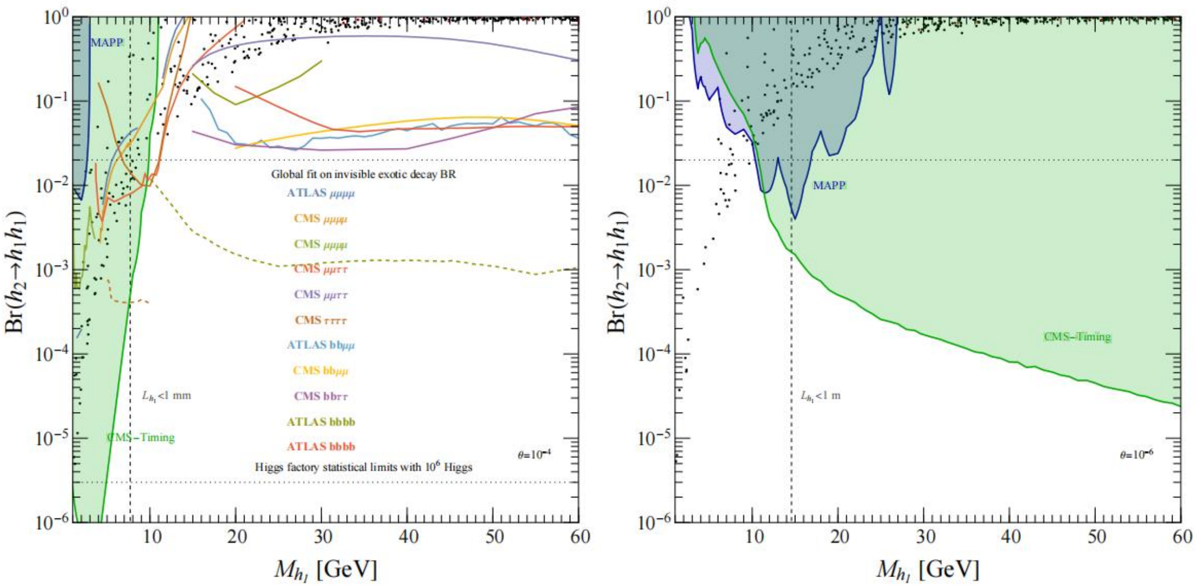


CMS-Timing can probe large parameter space where the **searches for promptly exotic Higgs decays can not reach**.

MAPP can only probe small parameter space, while none for FASER.

Sensitivity

Fixed $\theta = 10^{-4}(left),\ 10^{-6}(right)$.



H→hh **distribute at the transverse direction.**

And Meson decay cannot produce H

So no FASER sensitivity

Signatures at Colliders

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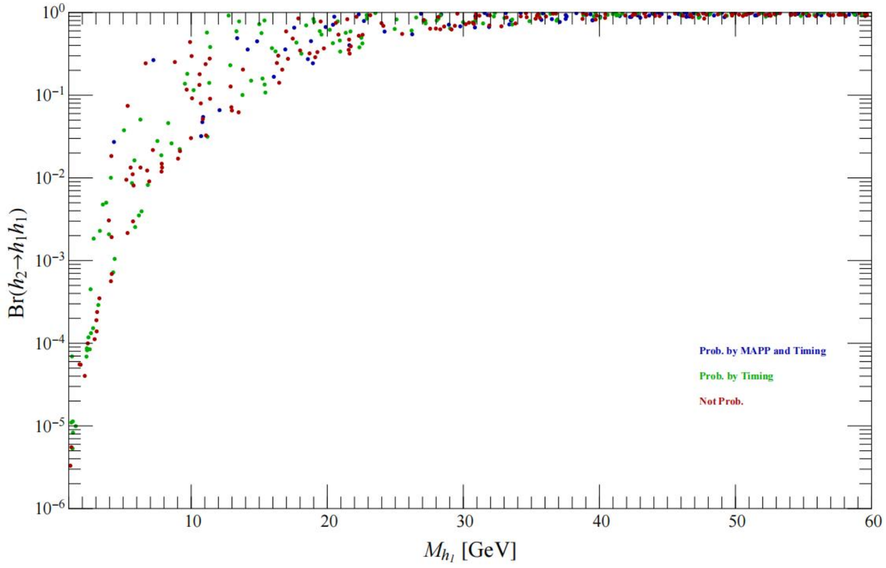
Running θ



Green points are probed by CMS-Timing, but not by MAPP.

CMS-Timing can probe **a lot more** 1st-order EWPT points. There are still appreciable points not probed by any of LLP detectors.

Conclusion

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1st-order EW phase transition is interesting:

- Theoretically, it is the essential ingredient of EW baryogenesis, and can trigger very rich dark matter mechanisms;
- Experimentally, it yields detectable gravitational waves.

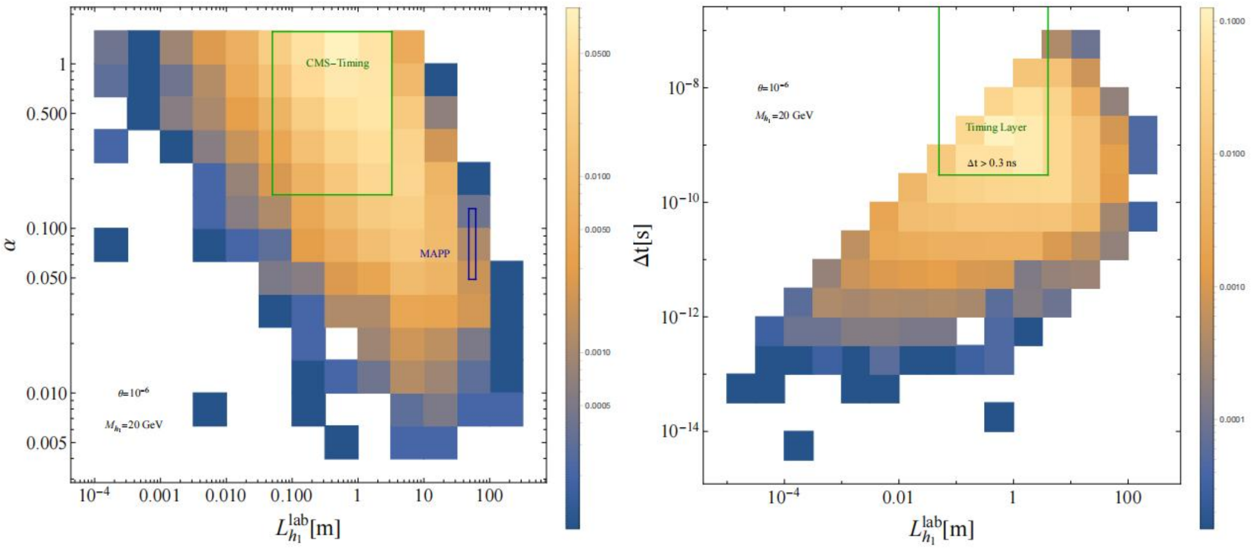
We propose strategies to probe 1st-order EWPT from LLP signatures at the HL-LHC.

For light, weakly coupled scalar corresponding to 1st-order EWPT, it can lead to LLP signatures, but no detectable gravitational waves.

LLP search is complementary to the searches for promptly exotic Higgs decays!

Detector efficiency

Detector efficiency is a function of geometrical coverage,



and resolution in time for timing detector.
CMS-Timing has large coverage, and good resolution.
MAPP has small coverage,
while negligible for FASER.

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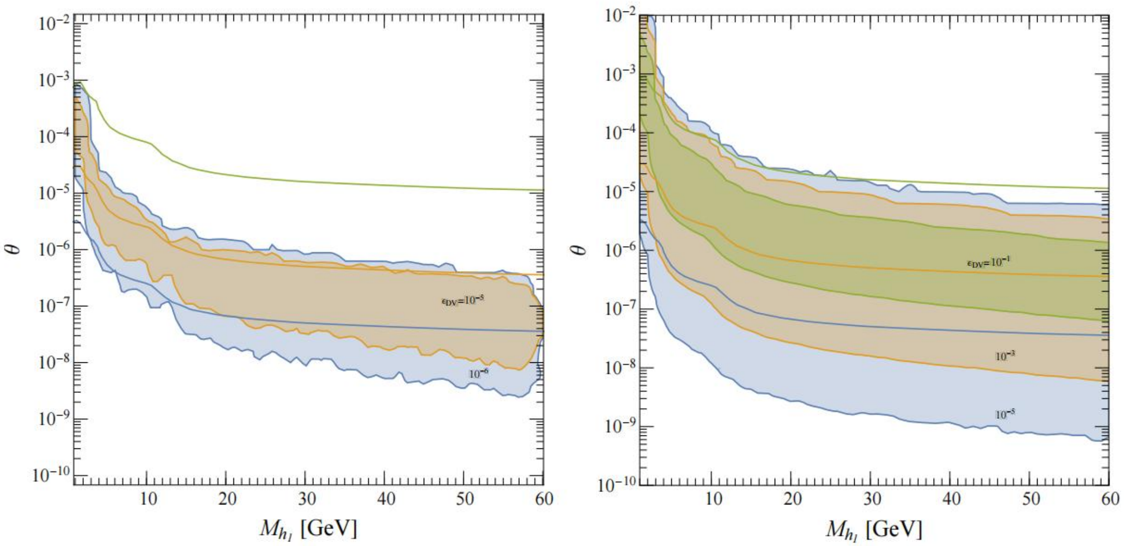
Detector efficiency

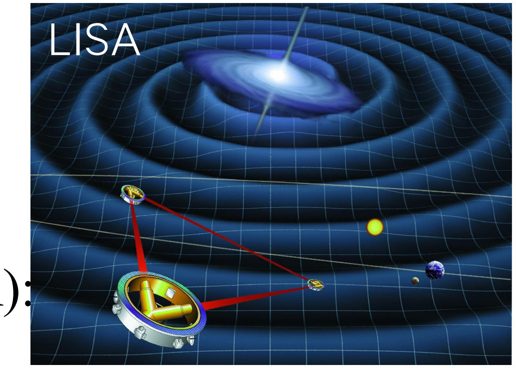
Detector efficiency is a function of geometrical coverage,



CMS-Timing has efficiency up to 10^{-1}.
MAPP has 10^{-4},
while negligible for FASER.

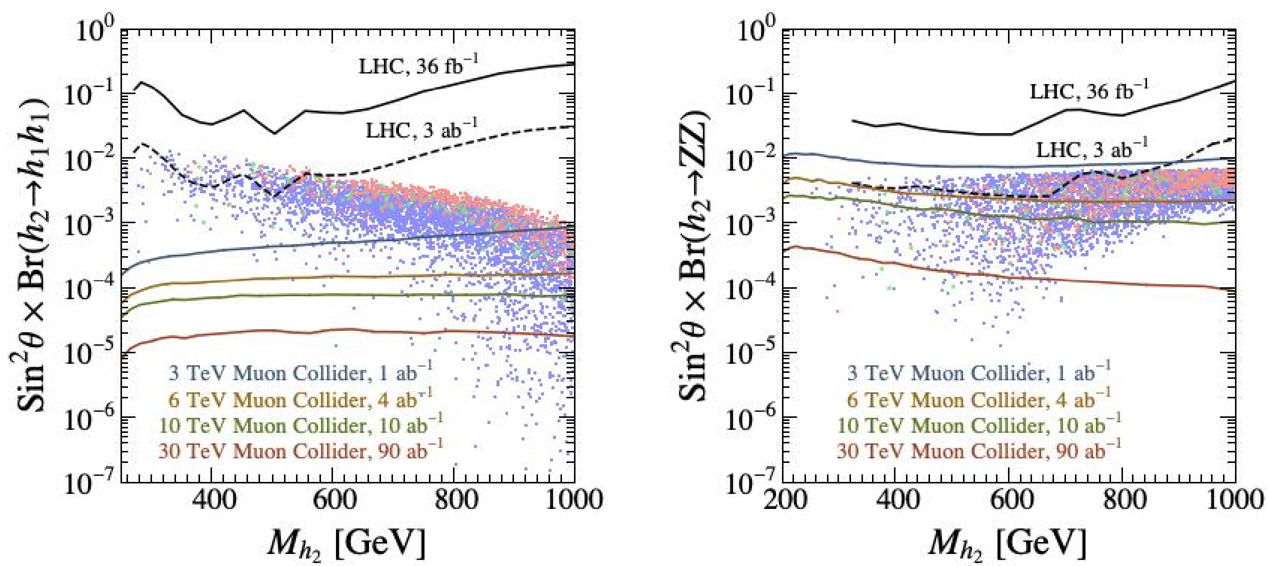
What if there is a TeV h_2?

1st-order EWPT can lead to **gravitational wave**, and diHiggs & diboson final states especially at **muon colliders**!



For the LISA detector, signal-to-noise ratio (SNR):

$$\mathrm{SNR} = \sqrt{\mathcal{T} \int_{f_{\min}}^{f_{\max}} df \left(\frac{\Omega_{\mathrm{GW}}(f)}{\Omega_{\mathrm{LISA}}(f)} \right)^2}$$



Wei Liu, Ke-pan Xie, *JHEP* 04 (2021) 015.