

# An extended Higgs model as the common origin of $\nu$ mass, dark matter and baryon asymmetry



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(KAIST)

Based on

- Mayumi Aoki<sup>1</sup>, KE, Shinya Kanemura<sup>2</sup>, [arXiv: 2212.14786](https://arxiv.org/abs/2212.14786)  
to be published by PRD

# Introduction

The extended Higgs sector play an important role to explain the unexplained phenomena in the SM

$\nu$  mass ... radiative seesaw models

Dark matter (DM) ... New scalar as WIMP

Baryon asymmetry ... Electroweak baryogenesis (EWBG)

A new physics model with the extended Higgs sector for these 3 problems [Aoki, Kanemura, Seto \(2009\)](#)

- Tiny  $\nu$  masses : Quantum correction via **3-loop diagrams**
- DM :  $Z_2$  symmetry ( New  $Z_2$ -odd neutral particles )
- BAU : **Electroweak baryogenesis** by extended Higgs sector
- Masses of new particles are  $O(100)$  GeV -> **Testable!**

EWBG in this model had not been evaluated

We evaluated it and find **one benchmark scenario** where

**all 3 problems** can be explained

## The model

[Aoki, Kanemura, Seto \(2009\)](#)

[Aoki, KE, Kanemura \(2022\)](#)

## Scalar Bosons

$Z_2$ -even)  $\Phi_1, \Phi_2 : (\mathbf{2}, +1/2)$

$Z_2$ -odd)  $S^+ : (\mathbf{1}, +1), \quad \eta : (\mathbf{1}, 0)$  real scalar

Extension of 2-Higgs doublet model

Type-III 2HDM

$$\mathcal{V} = V_{\Phi}(\Phi_1, \Phi_2) + V_{S\eta}(\Phi_1, \Phi_2, S^+, \eta)$$

## CP-violation

$$\mathcal{V}_{CPV} = \mathbf{Im} \left[ \mu_{12}^2 \Phi_1^\dagger \Phi_2 + (\Phi_1^\dagger \Phi_2) \left\{ \frac{\lambda_5}{2} \Phi_1^\dagger \Phi_2 + \lambda_6 |\Phi_1|^2 + \lambda_7 |\Phi_2|^2 \right\} \right. \\ \left. + \rho_{12} (\Phi_1^\dagger \Phi_2) |S^+|^2 + \frac{\sigma_{12}}{2} (\Phi_1^\dagger \Phi_2) \eta^2 + 2\kappa (\Phi_1^\dagger \Phi_2) S^- \eta \right]$$

$S^\pm$

**6 CP-violating couplings**

## Mass of Neutral Higgs Bosons

Higgs basis

$$\Phi_1 = \begin{pmatrix} G^+ \\ \frac{1}{\sqrt{2}}(v + H_1 + iG^0) \end{pmatrix} \quad \Phi_2 = \begin{pmatrix} H^+ \\ \frac{1}{\sqrt{2}}(H_2 + iH_3) \end{pmatrix}$$

$$M_{\text{neutral}} \propto \begin{pmatrix} H_1 & H_2 & H_3 \\ M_{11} & \text{Re}[\lambda_6] & -\text{Im}[\lambda_6] \\ & M_{22} & -\text{Im}[\lambda_5]/2 \\ \Phi_2 & & M_{33} \end{pmatrix} \begin{matrix} H_1 \\ H_2 \\ H_3 \end{matrix}$$

In the limit

$$\lambda_6 \rightarrow 0 \quad \rightarrow$$

**Mixings vanish [Higgs alignment].**

(Higgs couplings coincide with SM ones)

## The model

[Aoki, Kanemura, Seto \(2009\)](#)

[Aoki, KE, Kanemura \(2022\)](#)

## Higgs alignment scenario

Simple scenario  $\lambda_6 = 0$

[Kanemura, Kubota, Yagyu \(2020\); \(2021\)](#)

[KE, Kanemura, Mura \(2022\); \(2022\)](#)

[Kanemura, Takeuchi, Yagyu \(2022\)](#)

- $H_1, H_2, H_3$  are mass eigenstates w/o mixing

( $H_1$  is 125GeV Higgs boson)

- **3 CPV couplings** in the Higgs potential

$$\mathcal{V}_{CPV} = \mathbf{Im} \left[ \underbrace{\mu_{12}^2 \Phi_1^\dagger \Phi_2}_{\lambda_6 = 0 \text{ (+ Stationary condition)}} + (\Phi_1^\dagger \Phi_2) \left\{ \underbrace{\frac{\lambda_5}{2} \Phi_1^\dagger \Phi_2}_{\Phi_2} + \underbrace{\lambda_6 |\Phi_1|^2}_{\lambda_6 = 0} + \lambda_7 |\Phi_2|^2 \right\} \right. \\ \left. + \rho_{12} (\Phi_1^\dagger \Phi_2) |S^+|^2 + \frac{\sigma_{12}}{2} (\Phi_1^\dagger \Phi_2) \eta^2 + 2\kappa (\Phi_1^\dagger \Phi_2) \underbrace{S^- \eta}_{S^\pm} \right]$$

# Yukawa interaction

Both Higgs doublets couple with the SM fermions.

$$\mathcal{L}_Y = - \frac{m_{fi}}{v} \overline{f_L^i} f_R^i H_1 + \underbrace{(y_2^f)_{ij} \overline{f_L^i} f_R^j (H_2 + iH_3)}_{\text{Non-diagonal } y_2^f} + \text{h.c.}$$

$(i, j = 1, 2, 3)$ 
**SM Yukawa**
**Non-diagonal  $y_2^f$** 
**→ FCNC!**

To avoid FCNC,

(FCNC = Flavor Changing Neutral Current)

■ In [AKS\(2009\)](#): Softly broken  $Z_2$  [Glashow, Weinberg \(1977\)](#)

■ **Current Work: Flavor Alignment**

$$y_2^f = \frac{1}{v} \underbrace{\begin{pmatrix} m_{f1} & 0 & 0 \\ 0 & m_{f2} & 0 \\ 0 & 0 & m_{f3} \end{pmatrix}}_{\text{SM Yukawa}} \begin{pmatrix} \zeta_{f1} & 0 & 0 \\ 0 & \zeta_{f2} & 0 \\ 0 & 0 & \zeta_{f3} \end{pmatrix}$$

$\zeta_f^i \in \mathbb{C}$

For quarks,

$$\zeta_{u^1} = \zeta_{u^2} = \zeta_{u^3} \equiv \zeta_u$$

$$\zeta_{d^1} = \zeta_{d^2} = \zeta_{d^3} \equiv \zeta_d$$

[Pich, Tuzon \(2009\)](#)

## The model

[Aoki, Kanemura, Seto \(2009\)](#)

[Aoki, KE, Kanemura \(2022\)](#)

### Yukawa interaction

$Z_2$ -odd Majorana fermions:  $N_R^a$   
( $a = 1, 2, 3$ )

$$\frac{1}{2} m_{N^\alpha} \overline{(N_R^\alpha)^c} N_R^\alpha$$

Lepton # violating

$$\mathcal{L}_Y = - h_i^\alpha \overline{(N_R^\alpha)^c} \ell_R^i S^+ + \text{h.c.}$$

Lepton flavor violating

### Summary of the model

**New particles:** ( $Z_2$ -even)  $H^\pm, H_2, H_3$  ( $Z_2$ -odd)  $S^\pm, \eta, N_R^a$

**Alignment:**  $\lambda_6 = 0$  &  $(y_2^f)_{ij} \propto m_{fi} \zeta_{fi} \delta_{ij}$   
( $H_1$  is the SM Higgs) (No FCNC)

**CP-violation:**  $\lambda_7, \rho_{12}, \sigma_{12}$  &  $\zeta_u, \zeta_d, \zeta_\tau, \zeta_\mu, \zeta_e, h_i^\alpha$

# Benchmark scenario (BS) [Aoki, KE, Kanemura \(2022\)](#)

## Masses of New particle

$$Z_2 \text{ even: } m_{H^+} = 250 \text{ GeV}, \quad m_{H_2} = 420 \text{ GeV}, \quad m_{H_3} = 250 \text{ GeV}$$

$$Z_2 \text{ odd: } m_S = 400 \text{ GeV}, \quad m_\eta = 63 \text{ GeV}$$

$$(m_{N_1}, m_{N_2}, m_{N_3}) = (3000, 3500, 4000) \text{ GeV}$$

## Scalar couplings

$$\mu_2^2 = (50 \text{ GeV})^2, \quad \mu_s^2 = (320 \text{ GeV})^2, \quad \mu_{12}^2 = 0$$

$$\lambda_2 = 0.1, \quad \lambda_3 \simeq 1.98, \quad \lambda_4 \simeq 1.88, \quad \lambda_5 \simeq 1.88, \quad \lambda_6 = 0, \quad |\lambda_7| = 0.82,$$

$$\rho_1 \simeq 1.90, \quad \sigma_1 = |\sigma_{12}| = 1.1 \times 10^{-3}, \quad \kappa = 2.0, \quad \theta_7 = -0.73, \quad \dots$$

## New Yukawa interactions

$$y_t |\zeta_u| = 0.17, \quad y_b |\zeta_d| = 4.2 \times 10^{-3}, \quad y_e |\zeta_e| = y_\mu |\zeta_\mu| = 2.5 \times 10^{-4},$$

$$y_\tau |\zeta_\tau| = 2.5 \times 10^{-3}, \quad \theta_e = \theta_\mu = \theta_\tau = -2.94, \quad \theta_u = \theta_d = 0.245$$

$$h_i^\alpha \simeq \begin{pmatrix} 1.0 e^{-0.31i} & 0.2 e^{0.30i} & 1.0 e^{-2.4i} \\ 1.1 e^{-1.9i} & 0.21 e^{-1.8i} & 1.1 e^{2.3i} \\ 0.45 e^{2.7i} & 1.3 e^{-0.033i} & 0.10 e^{0.63i} \end{pmatrix}, \quad \dots$$



# Constraints

## Experimental constraints

$H^\pm$  : (Direct)  $H^\pm \rightarrow tb$  [ATLAS \(2021\)](#)

(Flavor)  $B_d \rightarrow \mu^+ \mu^-$  [J. Haller, et al EPJC \(2018\)](#)

$H_{2,3}$  : (Direct)  $H_{2,3} \rightarrow \tau\bar{\tau}$  [ATLAS \(2020\)](#)

$H_{2,3} \rightarrow t\bar{t}$  [ATLAS \(2018\)](#)

$S^\pm$  : (Direct)  $S^\pm \rightarrow H^\pm \eta \rightarrow tb\eta$  (from  $Z^*, \gamma^* \rightarrow S^+ S^-$ ) **Weak constraints**

(Flavor) Lepton flavor violating processes (**Next slides**)

$N_R^\alpha$  : (Direct) too heavy and weak constraints ( $m_{N^\alpha} = 3-4$  TeV)

(Flavor) Lepton flavor violating processes (**Next slides**)

$\eta$  : Dark matter in the model

(DM searches) **3 Pages later**

CP-violating phases : (EDM) **2 Pages later**

**We checked that  
all of these constraints  
can be avoided in the BS**

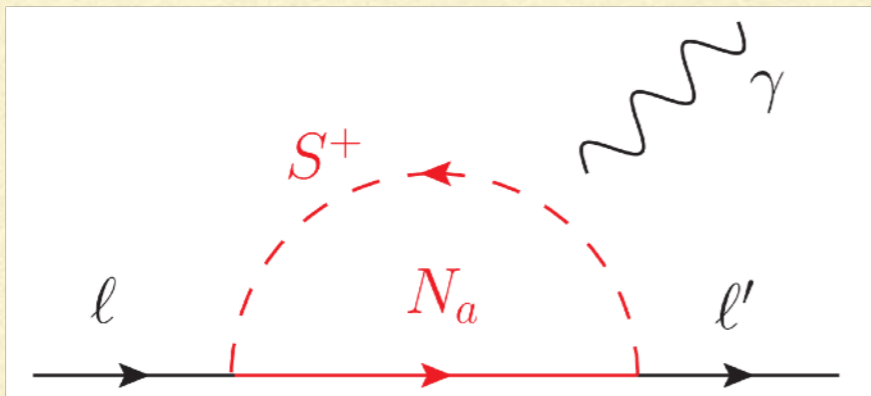
# Lepton flavor violation

$$m_S = 400 \text{ GeV},$$

$$M_N = \{3000, 3500, 4000\} \text{ GeV}$$

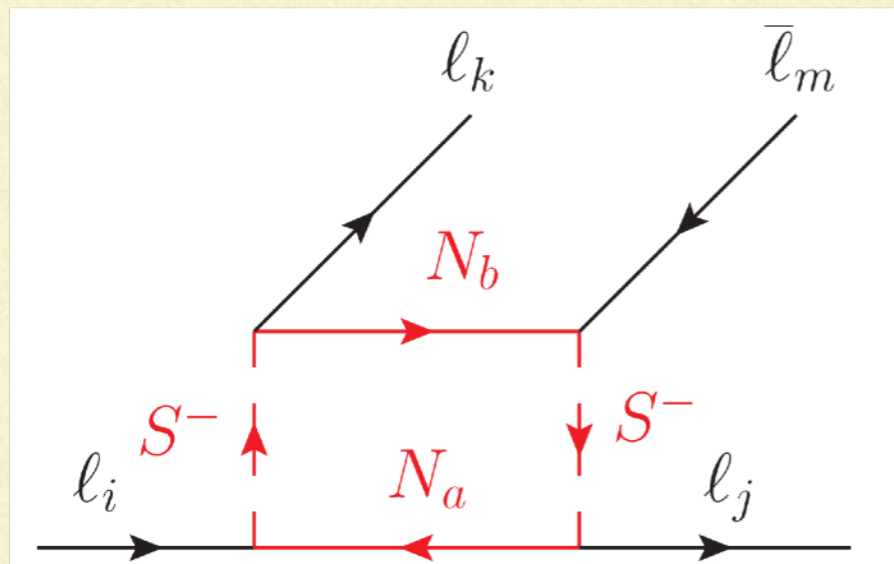
$$h_i^\alpha \simeq \begin{pmatrix} 1.0 e^{-0.31i} & 0.2 e^{0.30i} & 1.0 e^{-2.4i} \\ 1.1 e^{-1.9i} & 0.21 e^{-1.8i} & 1.1 e^{2.3i} \\ 0.45 e^{2.7i} & 1.3 e^{-0.033i} & 0.10 e^{0.63i} \end{pmatrix}$$

■  $\ell \rightarrow \ell' \gamma$



Processes	BR	Upper limits
$\mu \rightarrow e \gamma$	$1.4 \times 10^{-14}$	$4.2 \times 10^{-13}$
$\tau \rightarrow e \gamma$	$5.3 \times 10^{-10}$	$3.3 \times 10^{-8}$
$\tau \rightarrow \mu \gamma$	$1.1 \times 10^{-11}$	$4.4 \times 10^{-8}$

■  $\ell_i \rightarrow \ell_j \ell_k \bar{\ell}_m$

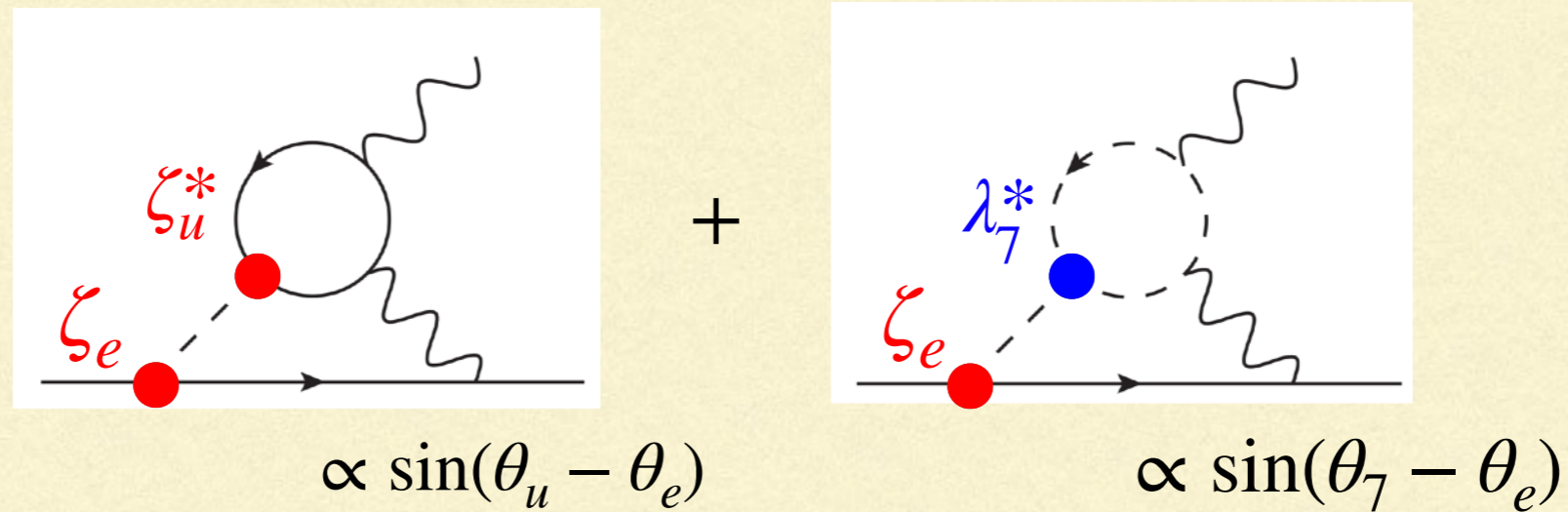


Processes	BR	Upper limits
$\mu \rightarrow 3e$	$1.0 \times 10^{-13}$	$1.0 \times 10^{-12}$
$\tau \rightarrow 3e$	$6.2 \times 10^{-10}$	$2.7 \times 10^{-8}$
$\tau \rightarrow 3\mu$	$2.4 \times 10^{-11}$	$2.1 \times 10^{-8}$
$\tau \rightarrow e \mu \bar{e}$	$5.1 \times 10^{-12}$	$1.8 \times 10^{-8}$
$\tau \rightarrow \mu \mu \bar{e}$	$1.1 \times 10^{-12}$	$1.7 \times 10^{-8}$
$\tau \rightarrow e e \bar{\mu}$	$4.5 \times 10^{-13}$	$1.5 \times 10^{-8}$
$\tau \rightarrow e \mu \bar{\mu}$	$9.6 \times 10^{-11}$	$2.7 \times 10^{-8}$

# Electric dipole moment (EDM)

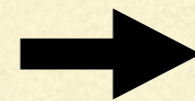
**electron EDM (eEDM)**  $|d_e| < 4.0 \times 10^{-30}$  e cm [Roussy, et al \(2022\)](#)

eEDM can be small by **destructive interference**  
[Kanemura, Kubota, Yagyu \(2020\)](#)



$$m_{H_2} = 420 \text{ GeV}, \quad m_{H_3} = m_{H^\pm} = 250 \text{ GeV}$$

$$\theta_7 = -2.34, \quad \theta_u = 0.245, \quad \theta_e = -2.94$$



$$|d_e| = 0.22 \times 10^{-30} \text{ e cm}$$

**neutron EDM (nEDM)**  $|d_n| < 1.8 \times 10^{-26}$  e cm

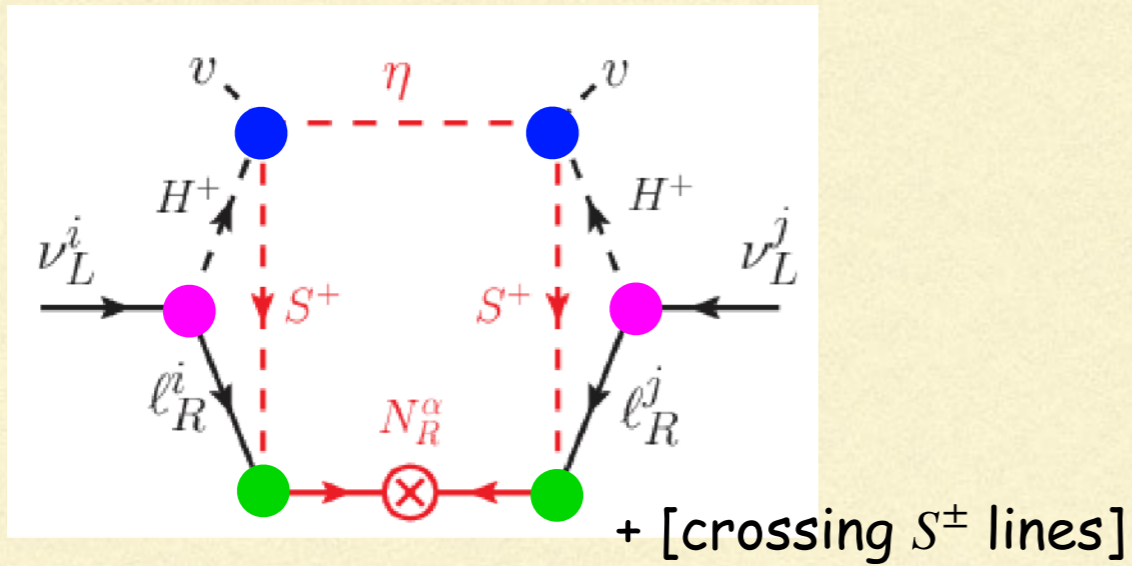
chromo EDM [Barr, Zee \(1990\)](#)

Weinberg ope. [Weinberg \(1989\)](#)

4 fermi interaction [Khatsimovsky, Khriplovich, Yelkhovsky \(1988\)](#)

In the BS,  $|d_n| \sim 10^{-30}$  e cm

## $\nu$ mass



$$\kappa \tilde{\Phi}_1 \Phi_2 S^- \eta \quad h_i^\alpha \overline{(N_R^\alpha)^c} \ell_{iR} S^+ \quad \zeta_{ei} \gamma_{ei} \overline{\ell_R^i} \nu_L^i H^-$$

$$(y_e |\zeta_e|, y_\mu |\zeta_\mu|, y_\tau |\zeta_\tau|) = (0.25, 0.25, 2.5) \times 10^{-3}$$

$$\theta_{\ell i} = -2.94 \quad \kappa = 2.0 \quad h_i^\alpha \simeq \begin{pmatrix} 1.0 e^{-0.31i} & 0.2 e^{0.30i} & 1.0 e^{-2.4i} \\ 1.1 e^{-1.9i} & 0.21 e^{-1.8i} & 1.1 e^{2.3i} \\ 0.45 e^{2.7i} & 1.3 e^{-0.033i} & 0.10 e^{0.63i} \end{pmatrix}$$



Normal ordering  $m_\nu$  w/  $m_{\nu 1} \simeq 0.006$  eV  
 $\delta \simeq 1.36\pi, \quad \alpha_1 \simeq 0, \quad \alpha_2 \simeq -\pi/2$   
 $m_{\beta\beta} \simeq 1$  meV,  $\Sigma m_{\nu i} = 0.067$  eV

$$m_{\beta\beta} < 35 \text{ meV} \quad \Sigma m_{\nu i} < 0.12 \text{ eV}$$

[KamLAND-Zen \(2023\)](#) [Planck \(2018\)](#)

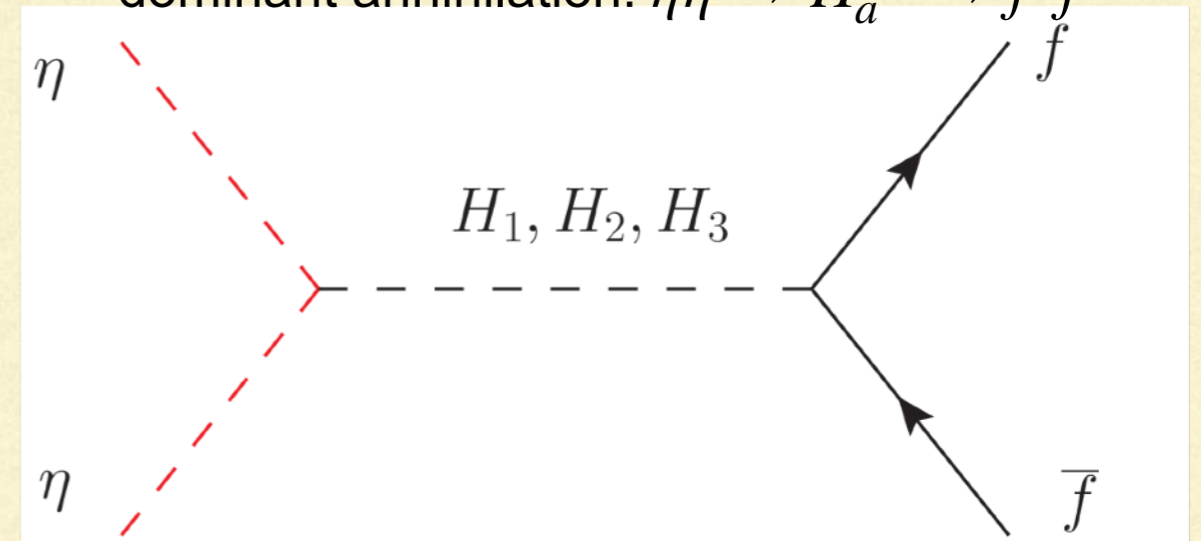
## Dark matter (DM)

$$m_\eta = 63 \text{ GeV}$$

$$(M_{N_1}, M_{N_2}, M_{N_3}) = (3000, 3500, 4000) \text{ GeV}$$

$\eta$ : real scalar DM

dominant annihilation:  $\eta\eta \rightarrow H_a^{(*)} \rightarrow f^i \bar{f}^i$



$$m_\eta = 63 \text{ GeV}, \quad m_{H_2} = 420 \text{ GeV}, \quad m_{H_3} = 250 \text{ GeV}$$

$$\sigma_1 = |\sigma_{12}| = 1.1 \times 10^{-3}, \quad \theta_\rho = -2.94$$



$$\Omega_\eta h^2 \simeq 0.12, \quad \sigma_{\text{SI}} = 2.3 \times 10^{-48} \text{ cm}^2$$

$$\Omega_{\text{DM}} h^2 = 0.120(01)$$

[Planck \(2018\)](#)

$$\sigma_{\text{SI}} \lesssim 10^{-47}$$

[LZ \(2022\)](#)

# Electroweak baryogenesis (EWBG)

## The Sakharov conditions [Sakharov \(1967\)](#)

- |                                       |        |  |
|---------------------------------------|--------|--|
| 1. $B$ -violation                     | ←----- | Sphaleron transition   |
| 2. $C$ and $CP$ violation             | ←----- | CPV phases : $\lambda_7, \rho_{12}, \sigma_{12}, \zeta_u, \zeta_d, \zeta_\ell$ |
| 3. Departure from thermal equilibrium | ←----- | Strongly 1st order electroweak phase transition                                |

## Strongly 1st EWPT (EWPT = ElectroWeak Phase Transition)

**Non-decoupling effect** by  $H_{2,3}, H^\pm, S^\pm$

$$m_{H^+}^2 = \mu_2^2 + \frac{1}{2}\lambda_3 v^2, \quad m_{H_{2,3}}^2 = \mu_2^2 + \frac{1}{2}(\lambda_3 + \lambda_4 \pm \lambda_5)v^2, \quad m_S^2 = \mu_S^2 + \frac{1}{2}\rho_1 v^2$$

$$m_{H^+} = 250 \text{ GeV}, \quad m_{H_2} = 420 \text{ GeV}, \quad m_{H_3} = 250 \text{ GeV}, \quad m_S = 400 \text{ GeV}$$

$$\lambda_3 \simeq 1.98, \quad \lambda_4 \simeq 1.88, \quad \lambda_5 \simeq 1.88, \quad \rho_1 \simeq 1.90$$

We evaluated **one-loop effective potential** in Landau gauge [Coleman, Weinberg \(1973\)](#)  
[Dolan, Jackiw \(1974\)](#)

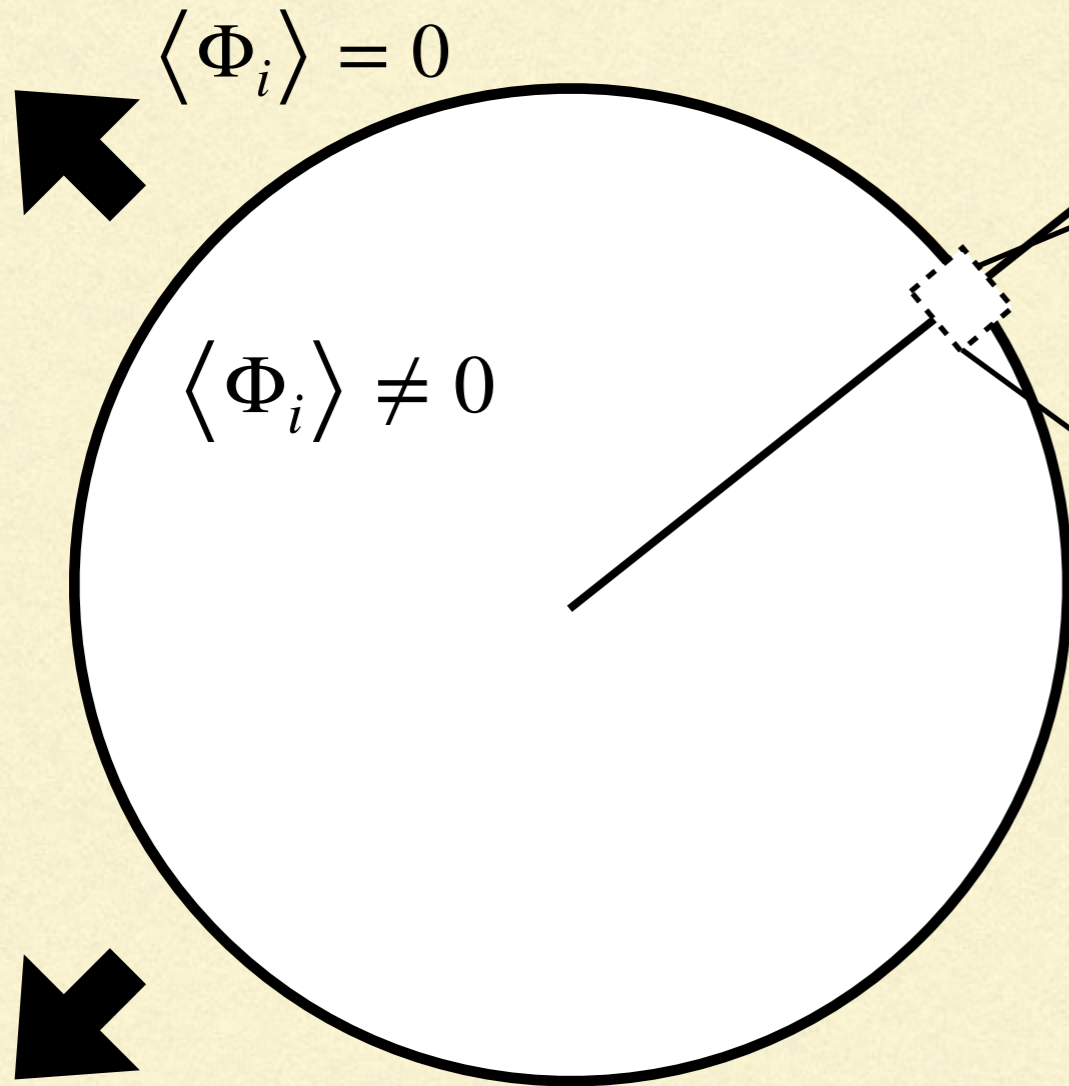
$(T = 0)$  [Kanemura, et al \(2003\)](#) [Kanemura, et al \(2004\)](#)       $(T \neq 0)$  thermal resummation [Parwani \(1992\)](#)

$$\Delta R \equiv \lambda_{hhh} / \lambda_{hhh}^{SM} - 1 = 38 \%$$

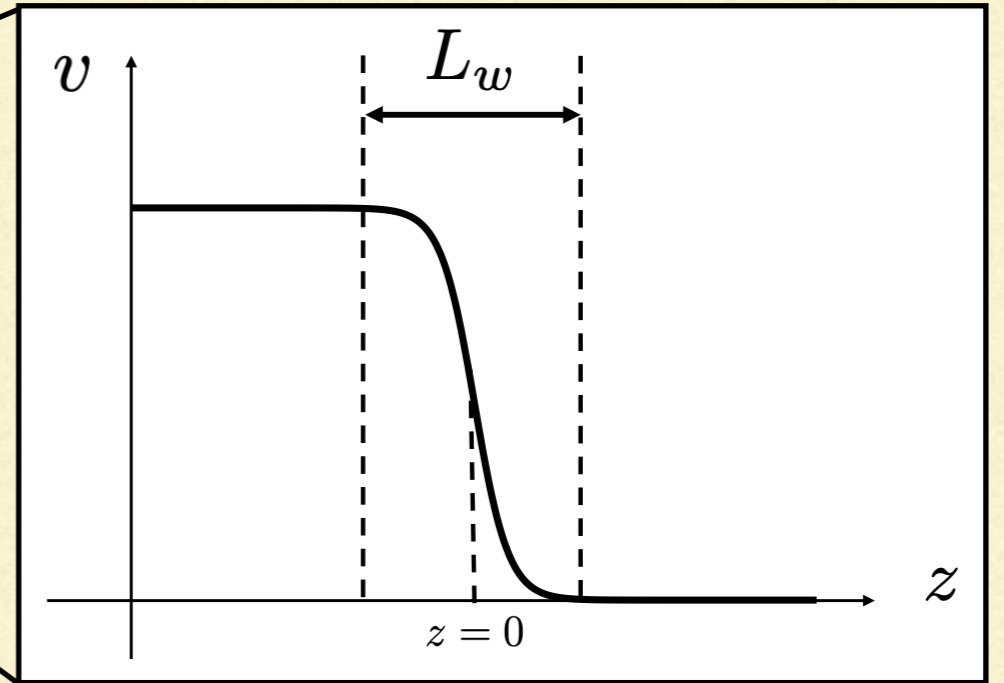
$$v_n / T_n = 1.74 > 1$$

[Kuzmin, Rubakov, Shaposhnikov \(1985\)](#)

# Electroweak baryogenesis



$L_w$ : Wall width

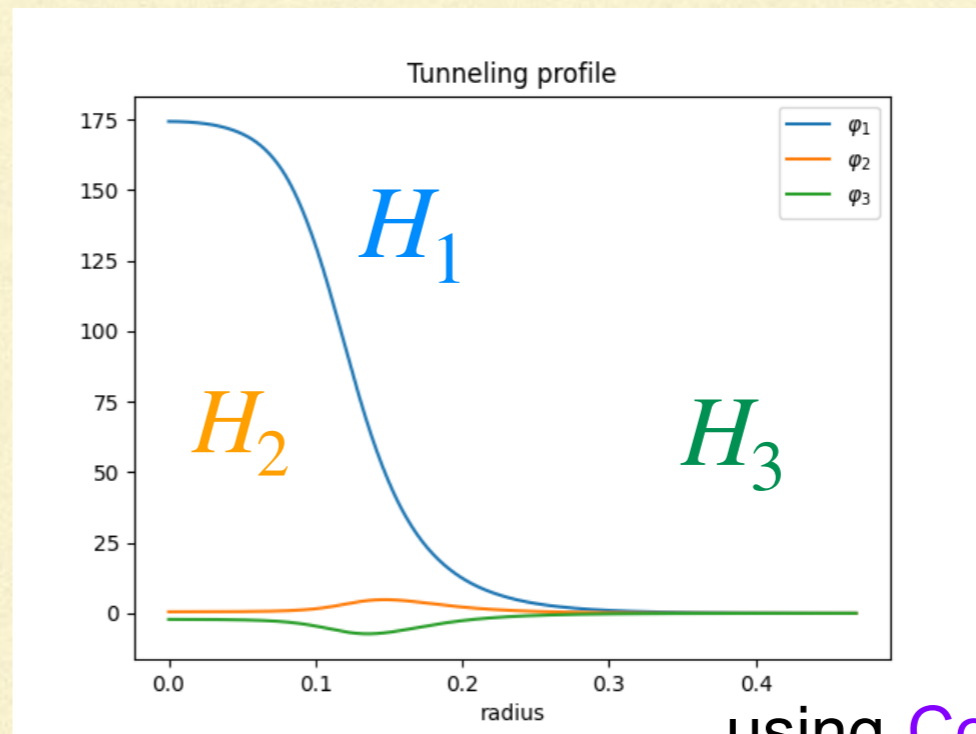


The bubble wall profile is given by **the bounce solution** of the effective potential

$v_w$ : wall velocity

$v_w = 0.1$  is assumed

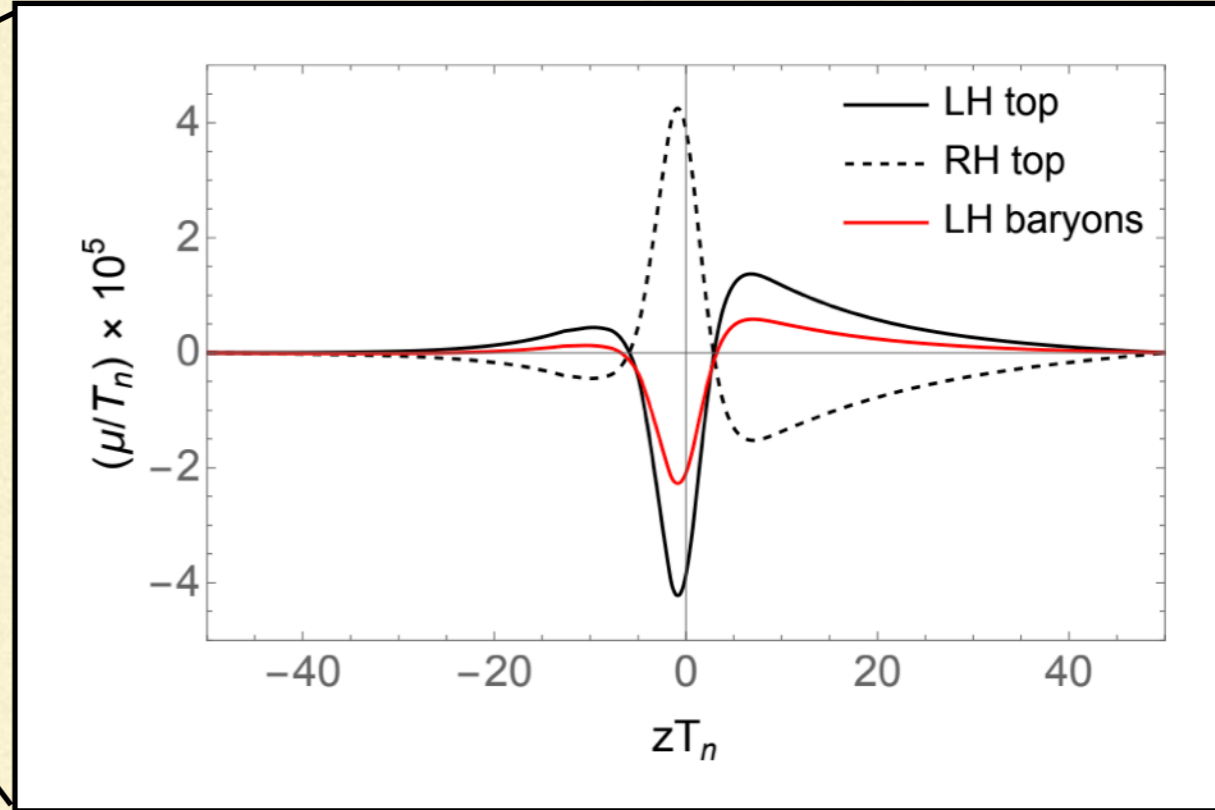
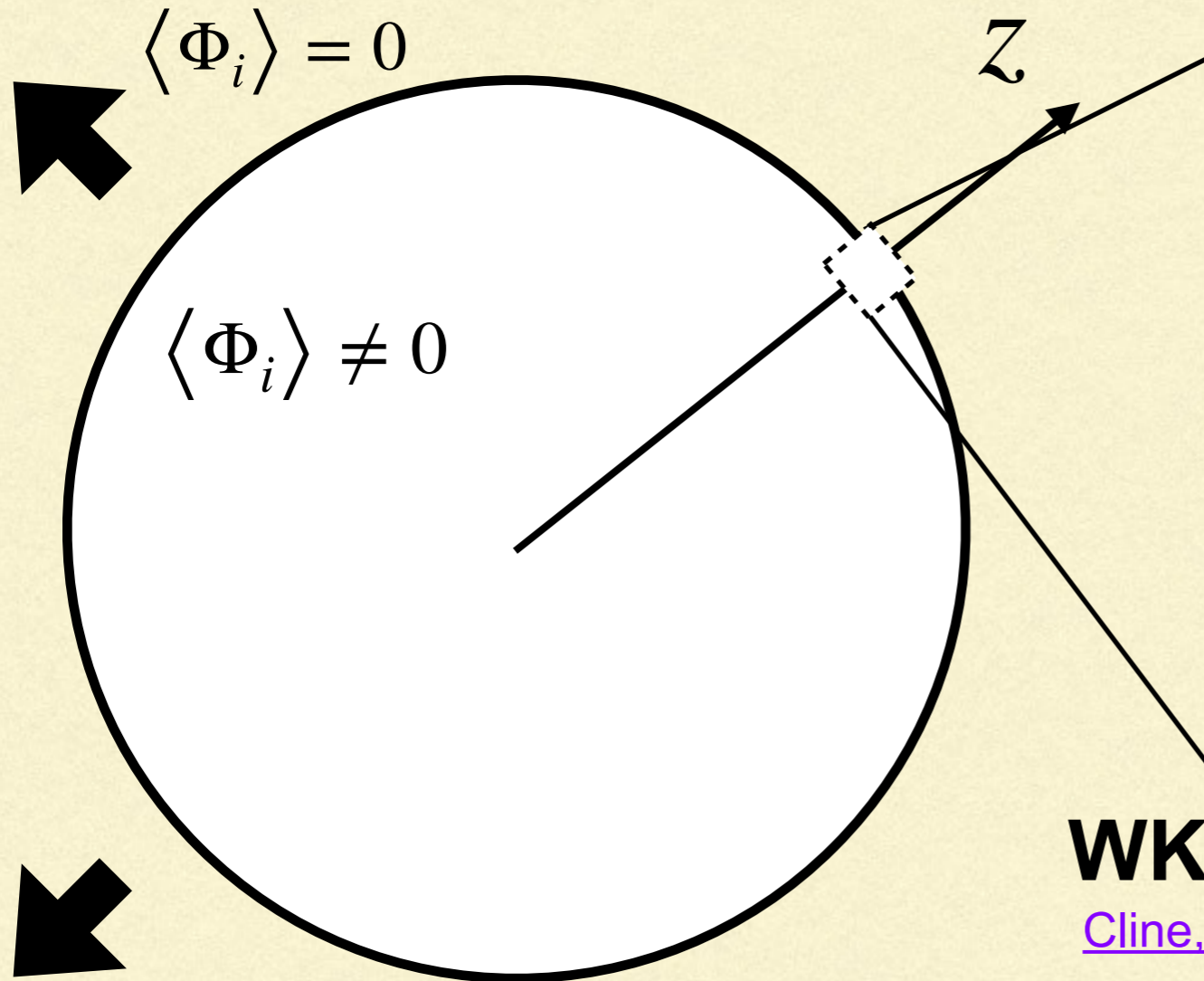
(In the case of the SM plasma,  $v_w = O(0.1)$ )



using [CosmoTransition](#)

# Electroweak baryogenesis

[Aoki, KE, Kanemura \(2022\)](#)



## WKB approximation ( $L_w T_n \gg 1$ )

[Cline, Joyce, Kainulainen \(2000\)](#) [Fromme, Huber \(2007\)](#)

[Cline, Kainulainen \(2020\)](#)

$v_w$ : wall velocity

$v_w = 0.1$  is assumed

**Baryon-to-photon ratio**

$$\eta_B = \frac{n_B}{n_\gamma} \sim \Gamma_{ws} \int_0^\infty dz \mu_{qL} e^{-kz}$$

BBN)  $5.8 \times 10^{-10} \leq \eta_B \leq 6.5 \times 10^{-10}$

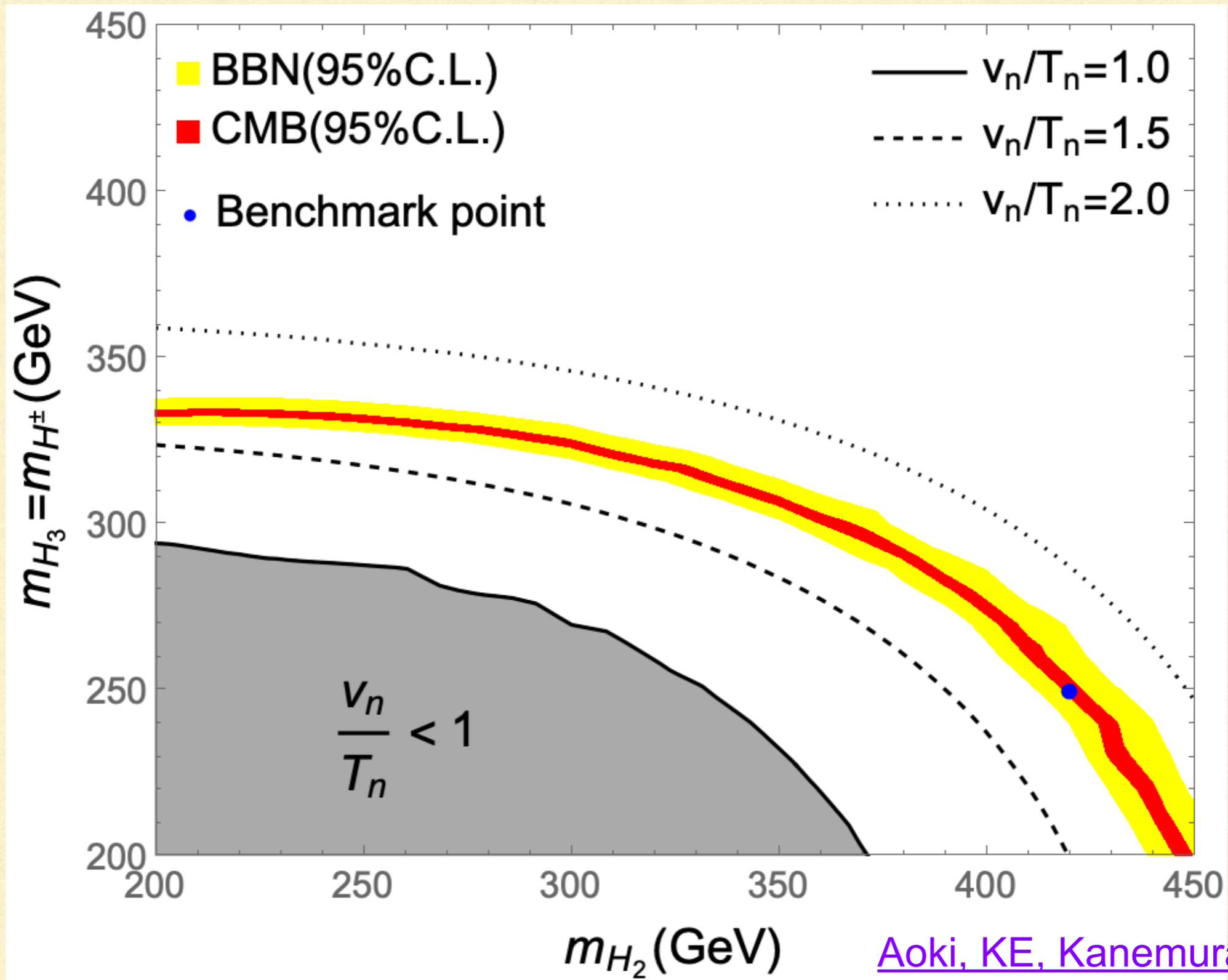
[Fields, et al \(2020\)](#)

CMB)  $6.04 \times 10^{-10} \leq \eta_B \leq 6.20 \times 10^{-10}$

[Planck \(2018\)](#)

$$\eta_B = 6.17 \times 10^{-10}$$

# Electroweak baryogenesis



[Aoki, KE, Kanemura \(2022\)](#)

Other parameters are the same with those in the BS



# How to test the BS

## EDM measurements

- One order improvement is expected in future ACME experiment [ACME\(2018\)](#)

## Flavor experiments

- $B \rightarrow X_s \gamma$  or  $B_d^0 \rightarrow \mu^+ \mu^-$  in Belle-II experiments [E. Kou, et al \[Bell-II\], arXiv:1808.10567 \[hep-ex\]](#)
- CP violation in  $B \rightarrow X_s \gamma$  ( $\Delta A_{CP}$ ) [Benz, Lee, Neubert, Paz \(2011\); Watanuki et al \[Belle\] \(2019\)](#)
- Lepton flavor violating decays  $\mu \rightarrow e \gamma$  [MEG-II](#)  $\mu \rightarrow 3e$ ,  $\tau \rightarrow 3e$  [Belle-II](#)

## Collider experiments

- $gg \rightarrow H_2, H_3$ ;  $gg \rightarrow H^\pm tb$ ;  $q\bar{q} \rightarrow H_{2,3} H^\pm$  [Aiko, Kanemura, Kikuchi, Mawatari, Sakurai, Yagyu \(2021\); S. Kanemura, M. Takeuchi, K. Yagyu \(2021\)](#)
- $q\bar{q} \rightarrow S^+ S^-$ ;  $e^+ e^- \rightarrow S^+ S^-$ ;  $e^+ e^- \rightarrow NN$  [M. Aoki, S. Kanemura, O. Seto \(2009\)](#)
- Higgs triple coupling  $\Delta R = \frac{\Delta \lambda_{hhh}}{\lambda_{hhh}^{SM}} = 38 \%$  **Sensitivity @ ILC** ( $\sqrt{s} = 500$  GeV)  
 $\Delta R = 27 \%$  [K. Fujii, et al, arXiv:1506.05992 \[hep-ph\]](#)
- Azimuthal angle distribution of  $H_{2,3} \rightarrow \tau \bar{\tau}$  at  $e^+ e^-$  collider

[S. Kanemura, M. Kubota, K. Yagyu, JHEP \(2021\)](#)

## Dark matter direct detection

## Observation of gravitational waves

**The detailed study is a work in progress.**

# Summary

- The SM cannot explain some observed phenomena (tiny  $\nu$  masses, DM, BAU), therefore, **we need physics beyond the SM**.
- In the previous work, the authors proposed a model where **tiny  $\nu$  masses**, **DM**, and **BAU** can be explained **simultaneously at TeV-scale**. However, they neglected CPV phases for simplicity.
- We have revisited the model and found a new benchmark scenario **including CPV phases**, where **tiny  $\nu$  masses**, **dark matter**, and **BAU** can be explained under the constraints from the current experiments. (LFV, EDM, ...).
- This benchmark scenario includes **some new particles** at **a few hundred GeV scale**, and they would be testable at various future experiments.

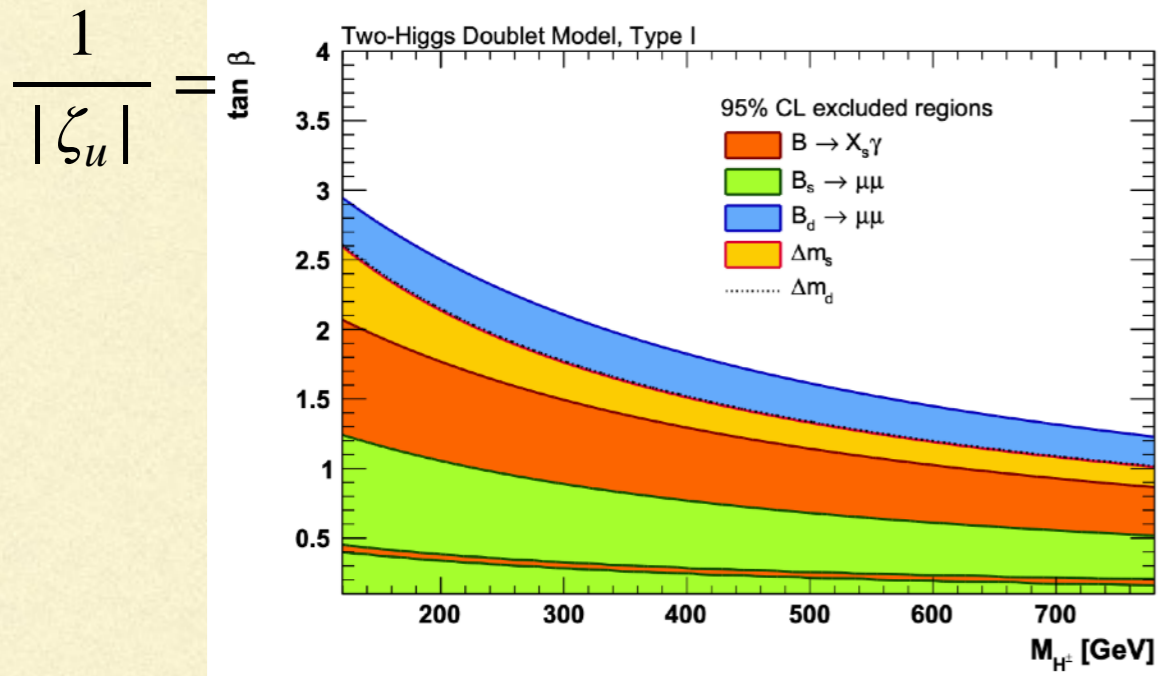
*Thank you for listening!*

# Backup Slides

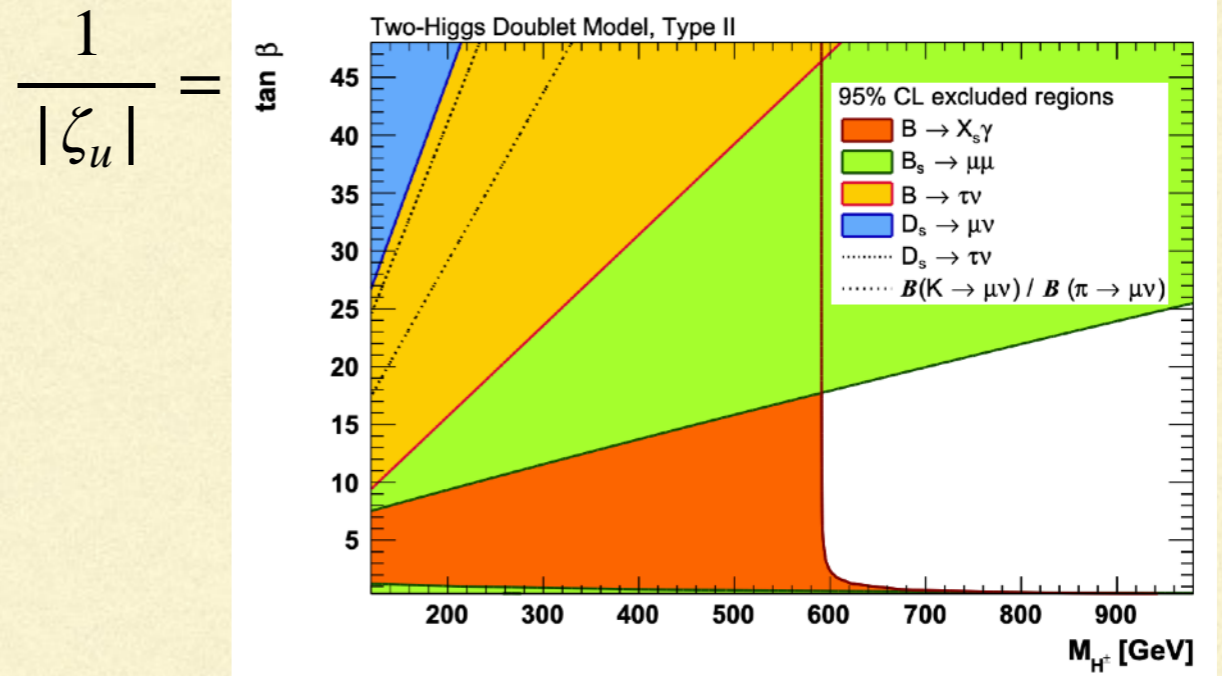
# Constraint from flavor experiments

Figures from [Haller, Hoecker, Kogler, Mooing, Peiffer, Stelzer, EPJC \(2018\)](#)

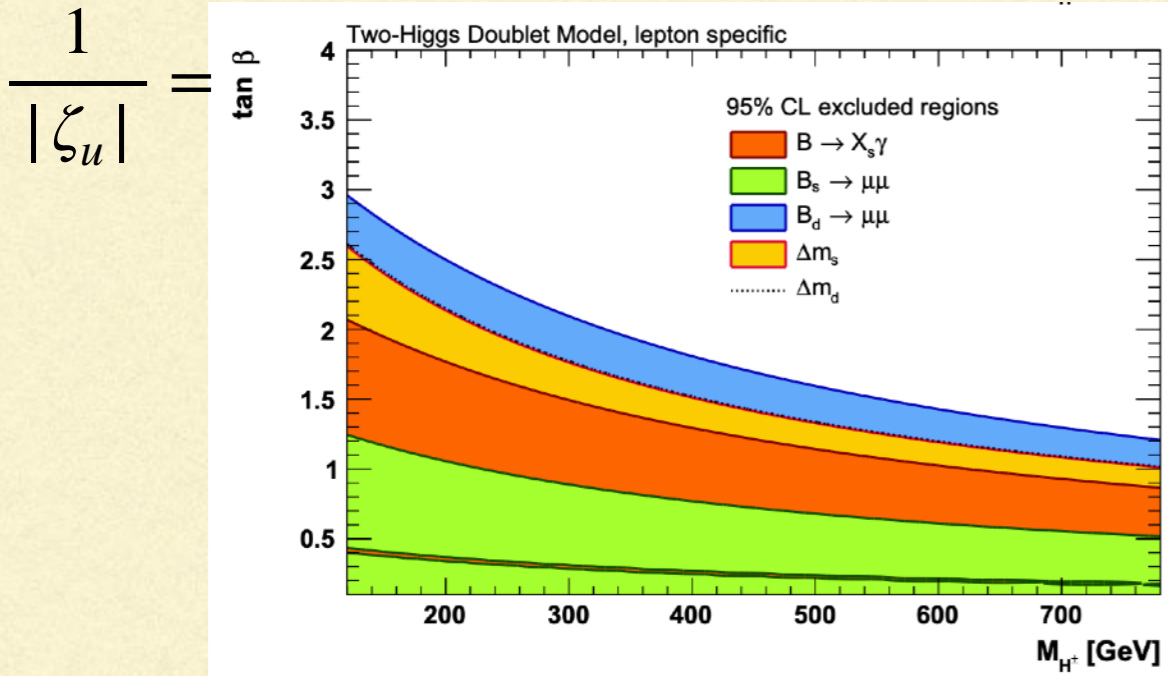
$$|\zeta_u| = |\zeta_d| = |\zeta_{\ell i}|$$



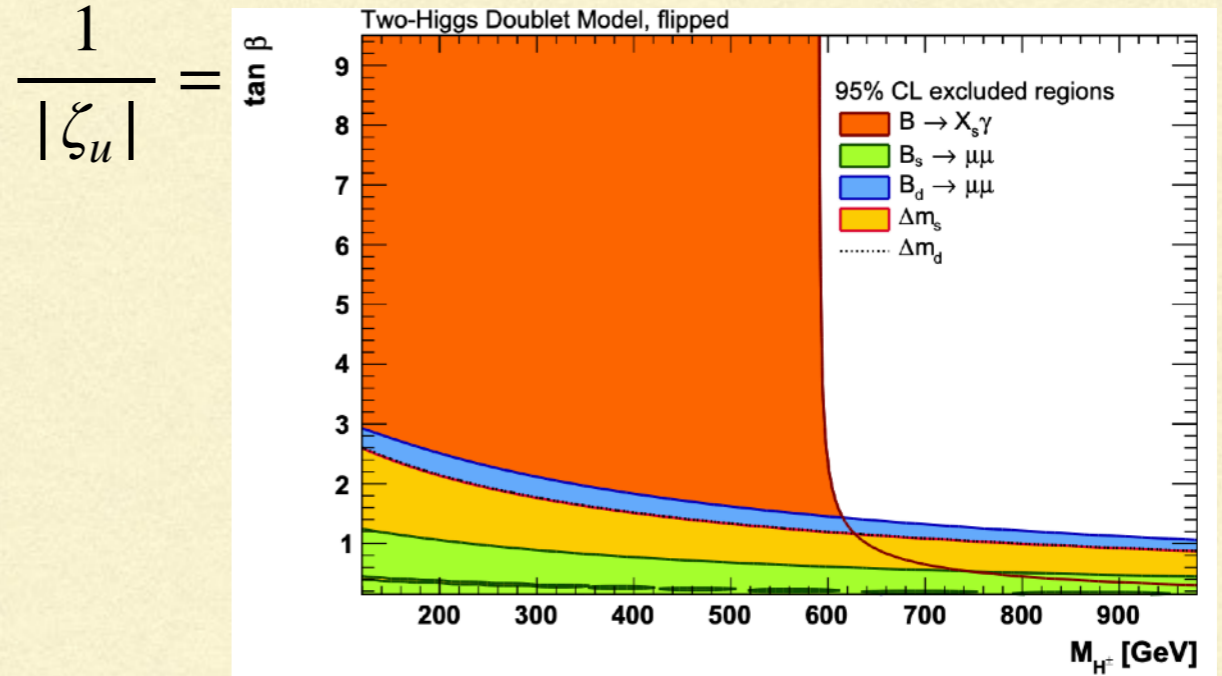
$$|\zeta_u| = 1/|\zeta_d| = 1/|\zeta_{\ell i}|$$



$$|\zeta_u| = |\zeta_d| = 1/|\zeta_{\ell i}|$$



$$|\zeta_u| = 1/|\zeta_d| = |\zeta_{\ell i}|$$

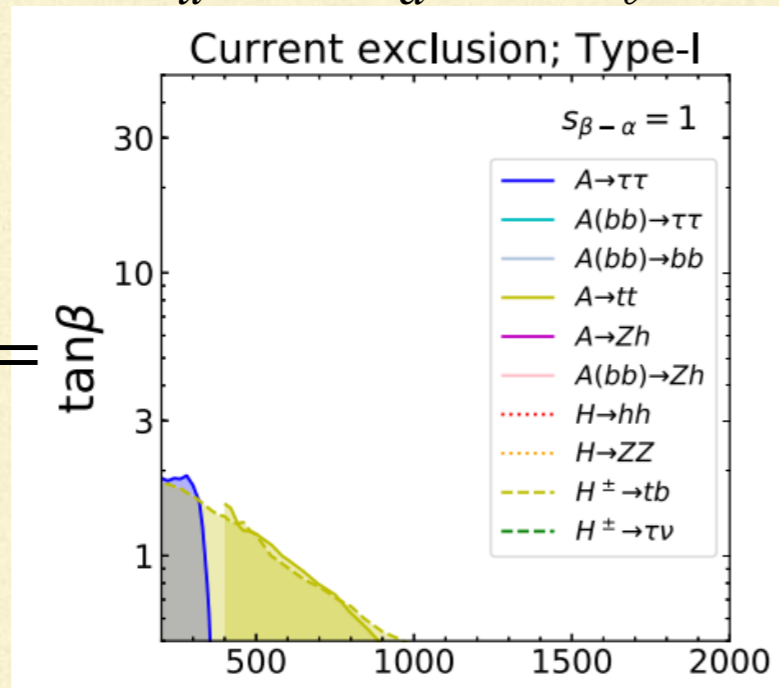


# Constraint from flavor experiments

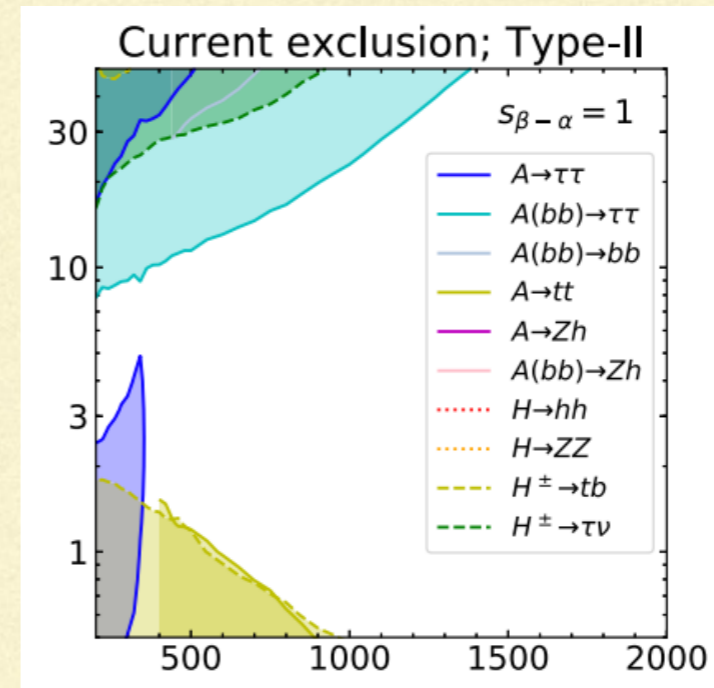
Figures from [Aiko, Kanemura, Kikuchi, Mawatari, Sakurai, Yagyu, NPB \(2021\)](#)

$$\frac{1}{|\zeta_u|} = \tan\beta$$

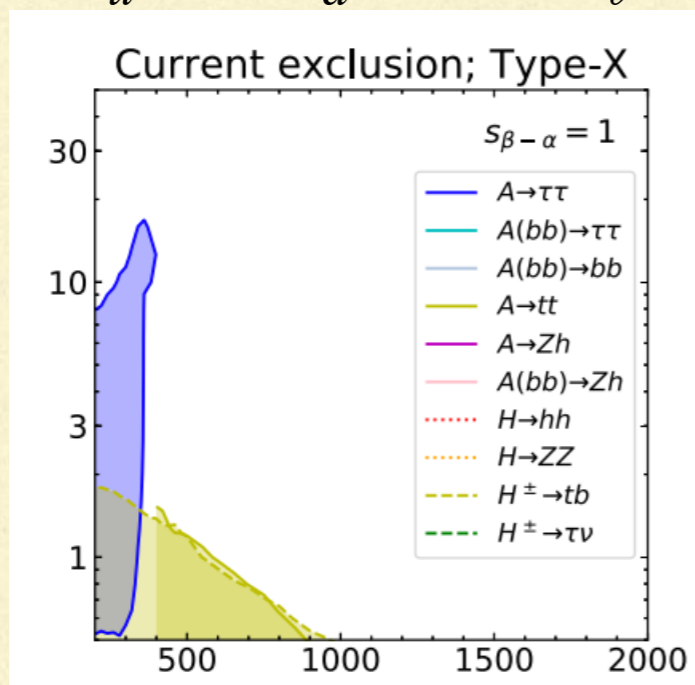
$$|\zeta_u| = |\zeta_d| = |\zeta_{\ell i}|$$



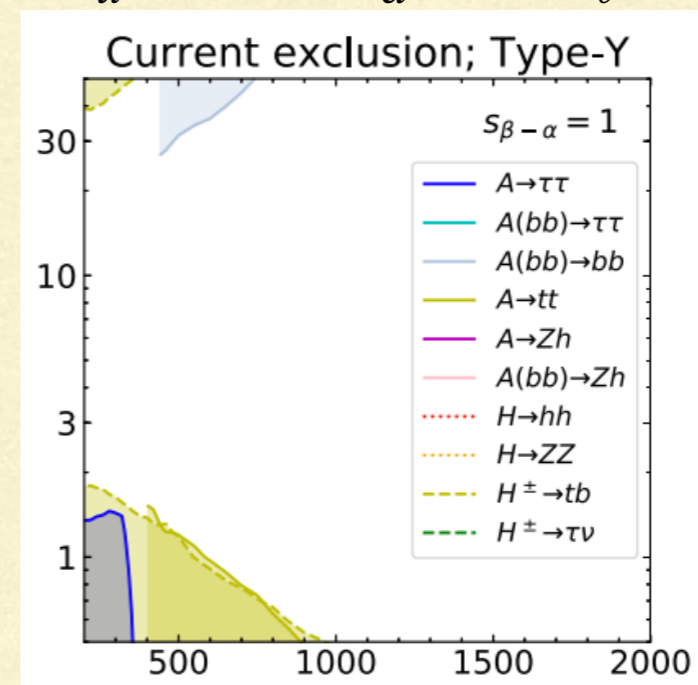
$$|\zeta_u| = 1/|\zeta_d| = 1/|\zeta_{\ell i}|$$



$$|\zeta_u| = |\zeta_d| = 1/|\zeta_{\ell i}|$$



$$|\zeta_u| = 1/|\zeta_d| = |\zeta_{\ell i}|$$

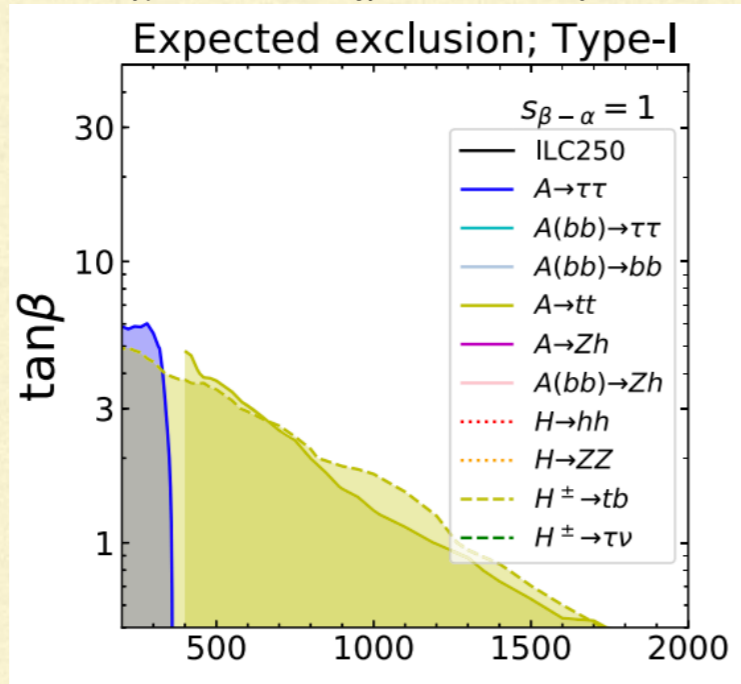


# Future direct search at HL-LHC

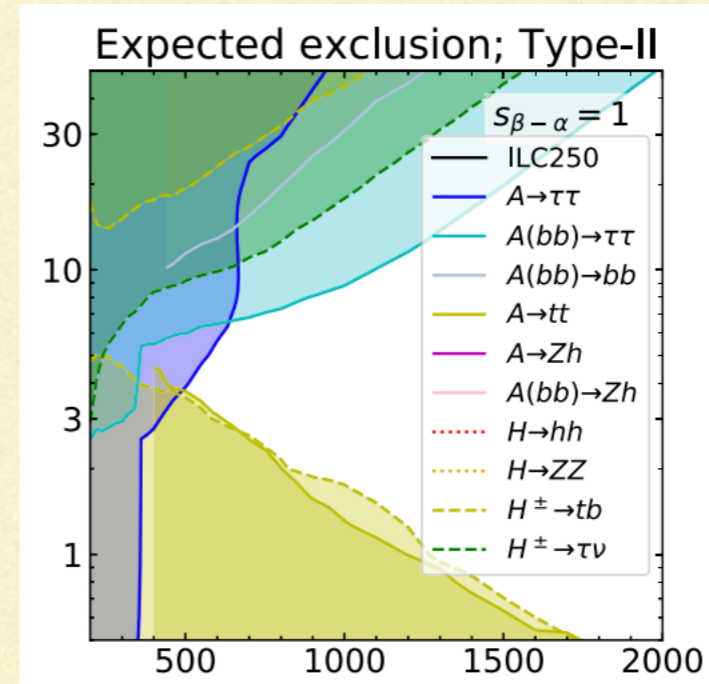
Figures from [Aiko, Kanemura, Kikuchi, Mawatari, Sakurai, Yagyu, NPB \(2021\)](#)

$$\frac{1}{|\zeta_u|} \equiv \tan\beta$$

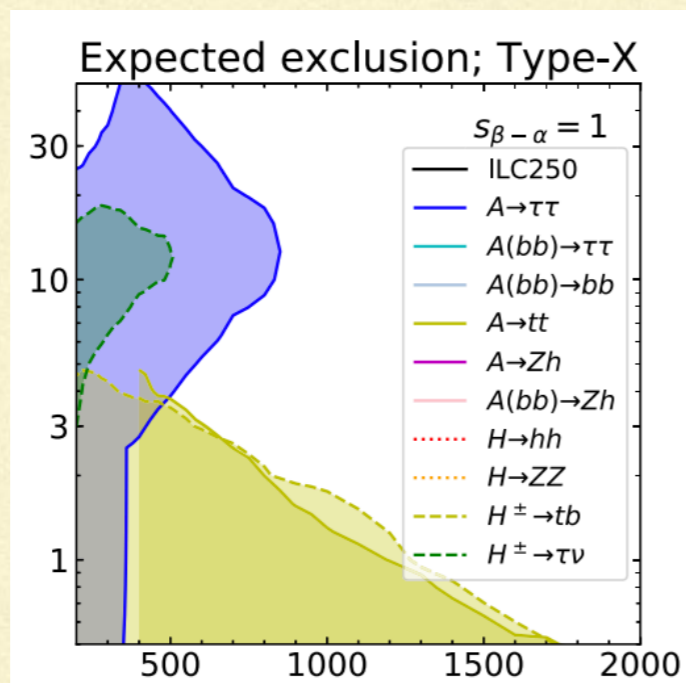
$$|\zeta_u| = |\zeta_d| = |\zeta_{\ell i}|$$



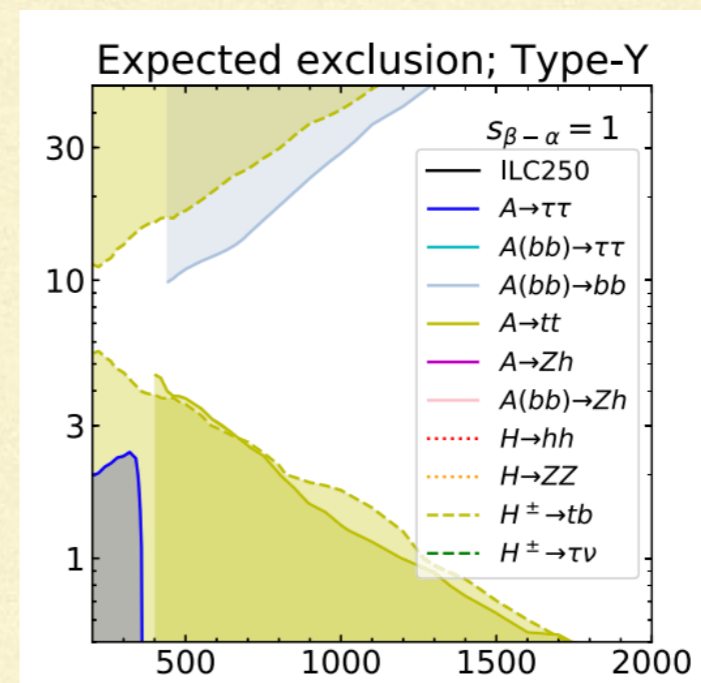
$$|\zeta_u| = 1/|\zeta_d| = 1/|\zeta_{\ell i}|$$



$$|\zeta_u| = |\zeta_d| = 1/|\zeta_{\ell i}|$$



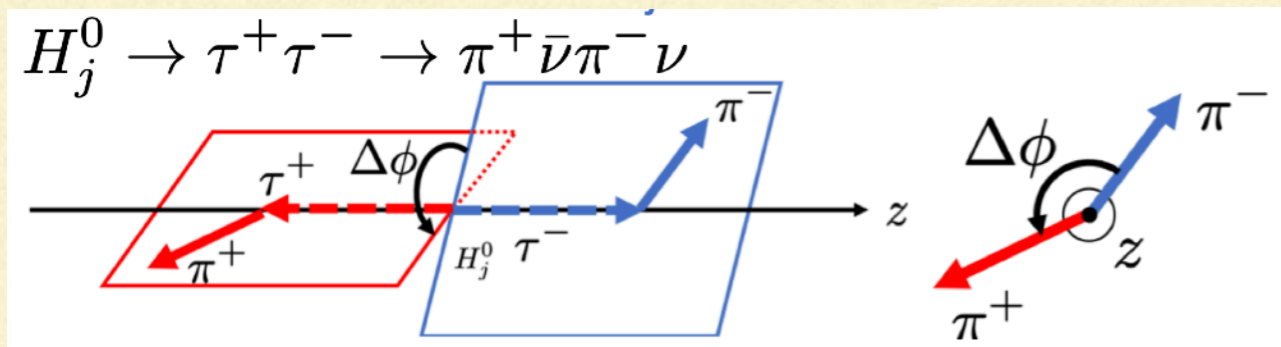
$$|\zeta_u| = 1/|\zeta_d| = |\zeta_{\ell i}|$$



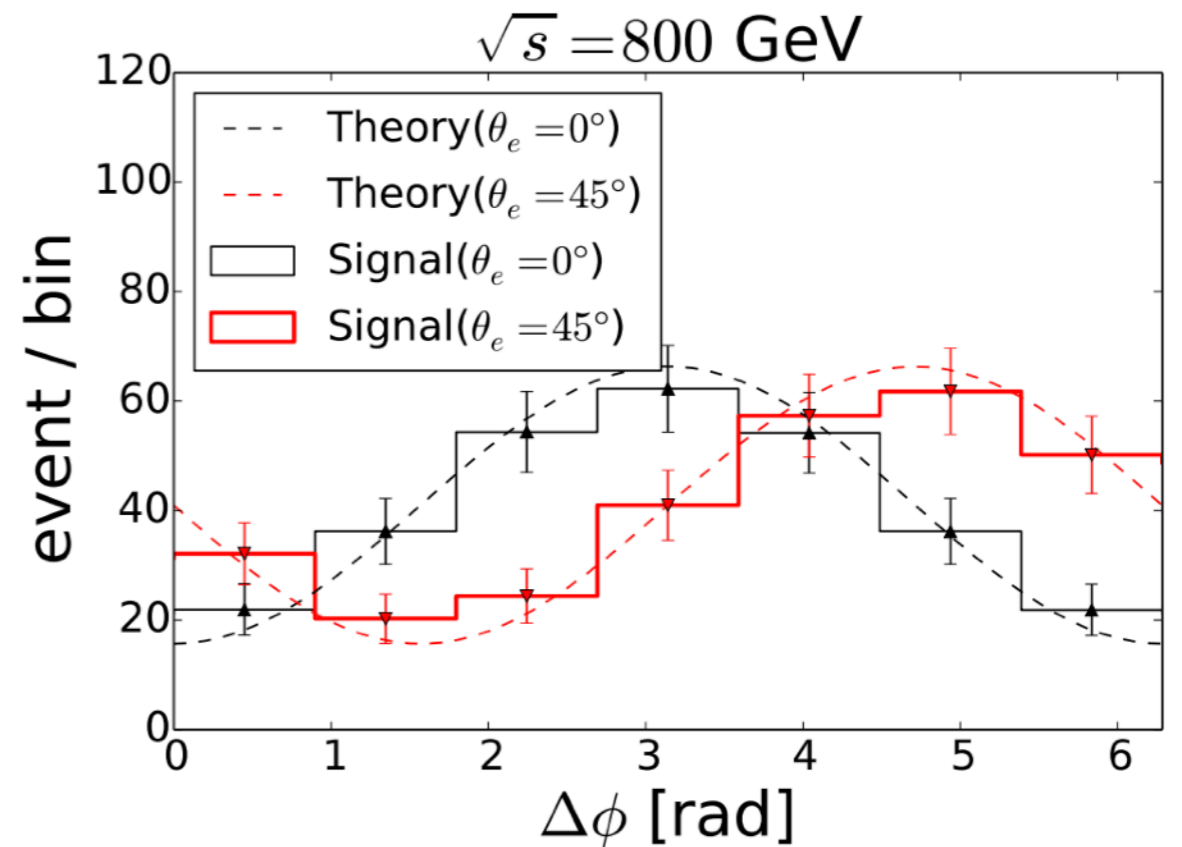
# Future test of CP-violation in $\zeta_\tau$

At  $e^+e^-$  collider

$$e^+e^- \rightarrow H_2H_3 \rightarrow \tau^+\tau^-b\bar{b}$$



[Kanemura, Kubota, Yagyu, JHEP \(2021\)](#)



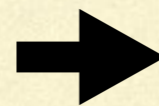
$M = 240,$	$m_{H_2^0} = 280,$	$m_{H_3^0} = 230,$	$m_{H^\pm} = 230$	(in GeV)
$ \zeta_u  = 0.01,$	$ \zeta_d  = 0.1,$	$ \zeta_e  = 0.5,$	$ \lambda_7  = 0.3,$	$\lambda_2 = 0.5$
$\theta_u = 1.2,$	$\theta_d = 0,$	$\theta_e = \pi/4,$	$\theta_7 = -1.8$	(in radian)

# Bubble profiles and nucleation temperature

Euclidean action :  $S_E = \int d^d x \left\{ \frac{1}{2} (\partial_\mu \varphi)^2 + V_{eff}(\varphi) \right\}$  Finite temperature  $d = 3$

Rate of the nucleation per volume :  $\Gamma/V = \omega T^4 e^{-S_E/T}$  ( $\omega = \mathcal{O}(1)$ )

Probability of the bubble nucleation per one Hubble volume is  $\mathcal{O}(1)$



$$\frac{S_E}{T_n} \sim 140$$

$T_n$ : Nucleation temperature

Bubble profile is given by the bounce solution

$$\frac{d^2 \varphi}{d\rho^2} + \frac{2}{\rho} \frac{d\varphi}{d\rho} = \nabla V_{eff}$$

(Boundary)

$$\varphi(\infty) = \varphi_F$$

$$\left. \frac{d\varphi}{d\rho} \right|_{\rho=0} = 0$$

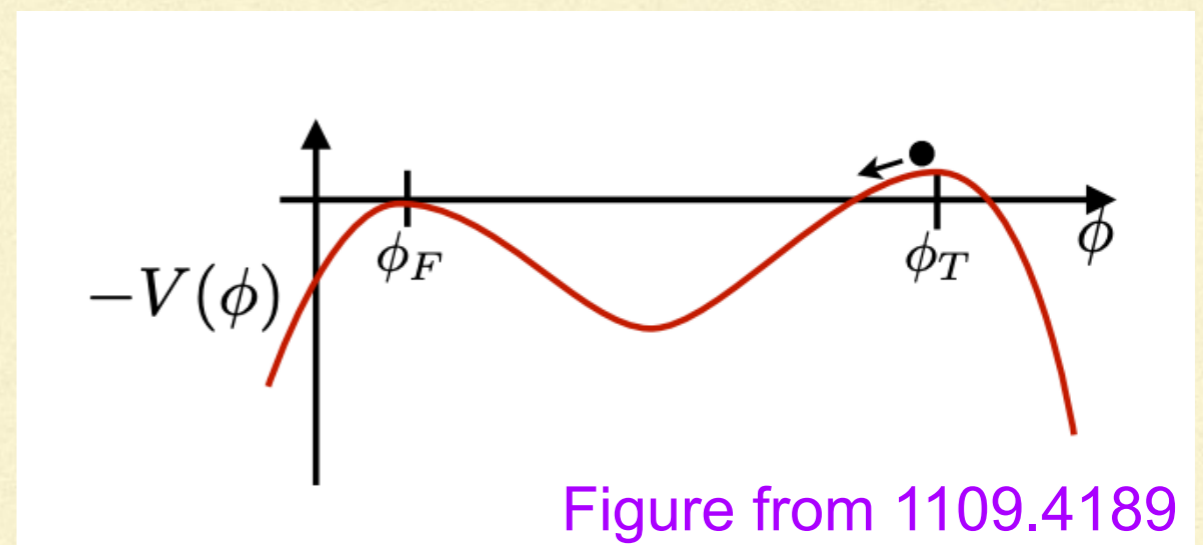


Figure from 1109.4189



# Wall width dependence of BAU

In the WKB method, generated baryon asymmetry is roughly estimated as

$$\eta_B \sim \int_0^\infty dz \frac{S(z)}{T^3} - A \int_{-\infty}^\infty dz \frac{S(z)}{T^3} \quad \text{Cline, Laurent, PRD (2021)}$$

$A$  is a function of  $v_w$  and  $L_w$

$v_w$  : wall velocity

$L_w$  : wall width

When  $A$  has a certain value, the first and second terms are canceled.

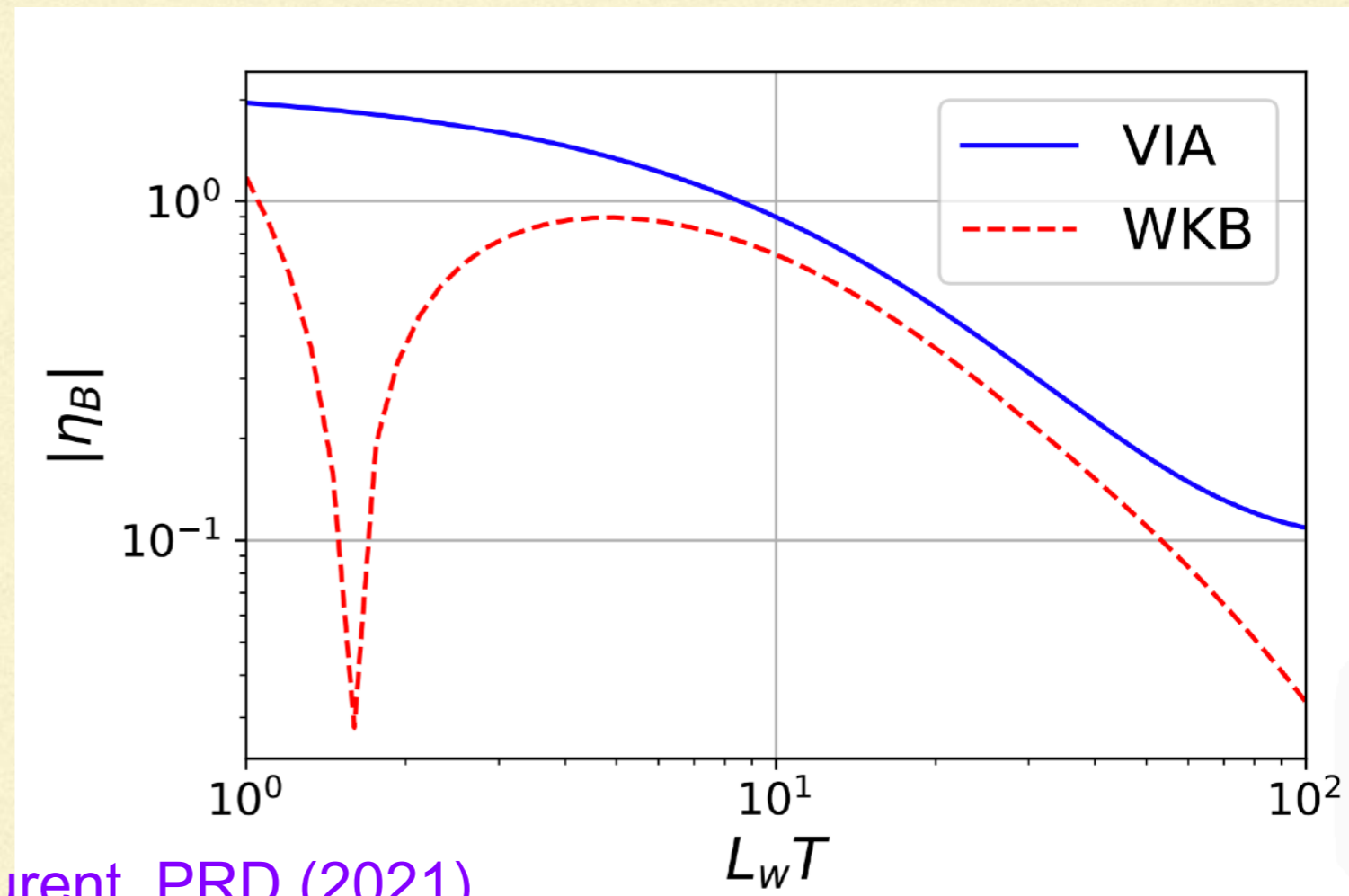


Figure from [Cline, Laurent, PRD \(2021\)](#)

# Relativistic effect in BAU

We used the linear expansion by the wall velocity  $v_w$

Effects of higher order terms : [Cline, Kainulainen, PRD \(2020\)](#)

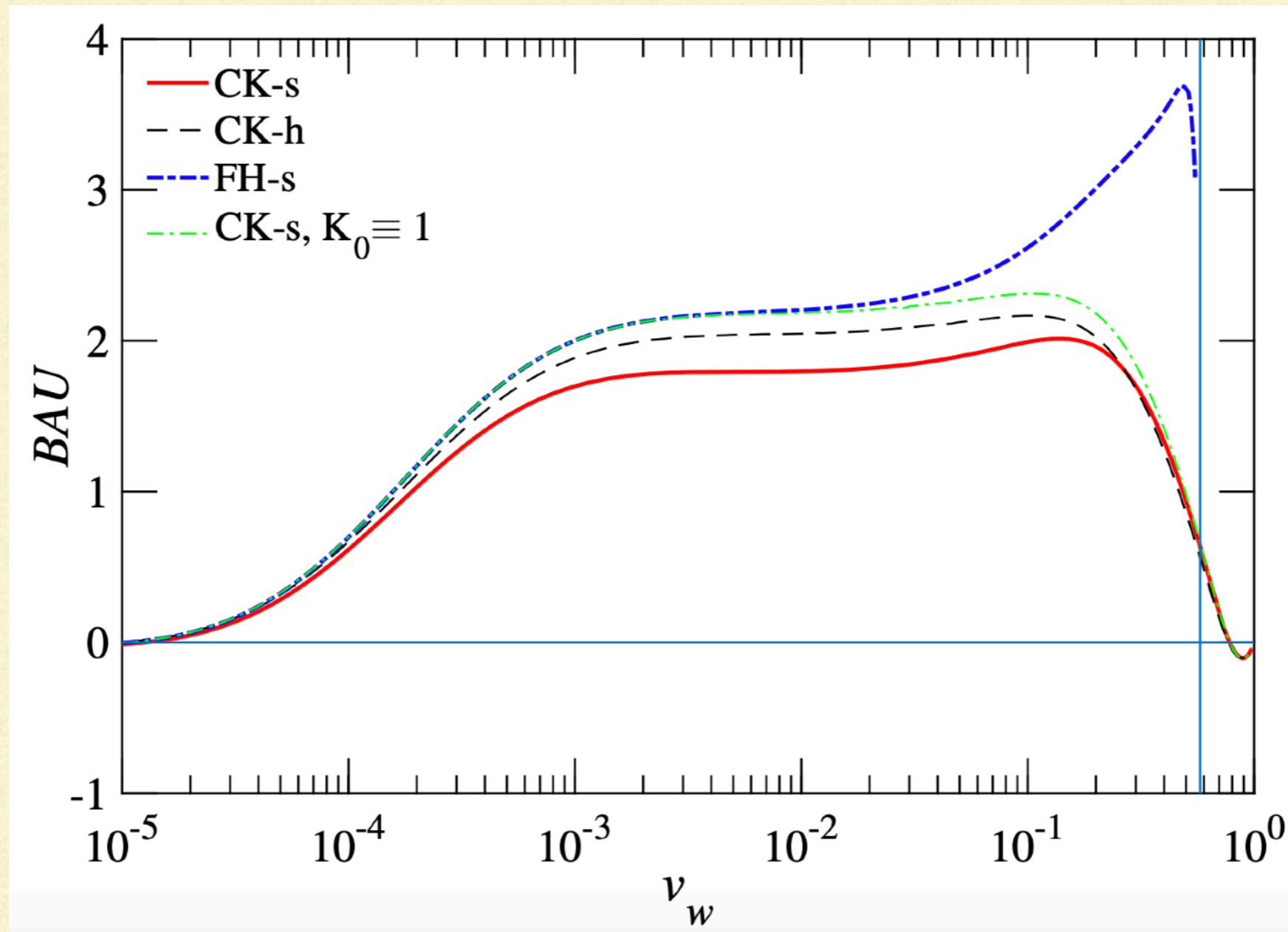


Figure from [Cline, Kainulainen, PRD \(2020\)](#)

# Velocity dependence of $n_B$

$\ell \sim \frac{1}{T}$  : Mean free path

Charge is accumulated within  $\ell$  (Gray region)

Time for accumulation to enter the bubble

$$t = \frac{\ell}{v_w} \sim \frac{1}{v_w T}$$

# of sphaleron tran.

before the charge enters the bubble

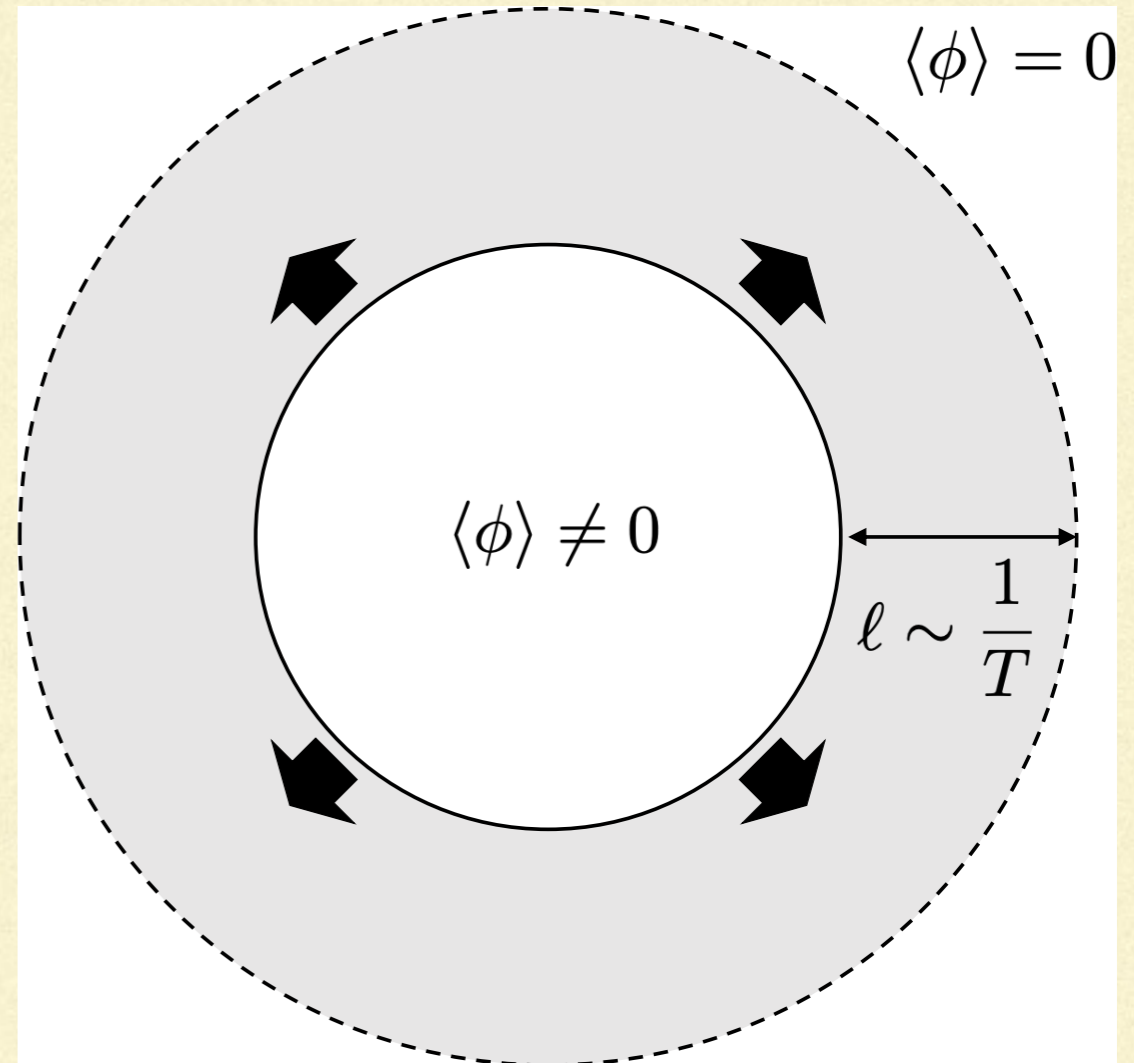
$$N = \Gamma_{sph}^{sym} \times t \sim \frac{\Gamma_{sph}^{sym}}{v_w T}$$

$N$  is too large (small  $v_w$ )

➔ **washed-out**

$N$  is too small (large  $v_w$ )

➔ **too short time**



$$\frac{n_B}{s} \propto \frac{\Gamma_{sph}^{sym}}{v_w T} \int d\hat{z} \frac{\mu_{qL}(\hat{z})}{T} \exp\left(-\frac{\Gamma_{sph}^{sym}}{v_w T} \hat{z}\right)$$

$\hat{z} = zT$

# The benchmark scenario

## Masses of New particle

$$Z_2 \text{ even: } m_{H^+} = 250 \text{ GeV}, \quad m_{H_2} = 420 \text{ GeV}, \quad m_{H_3} = 250 \text{ GeV}$$

$$Z_2 \text{ odd: } m_S = 400 \text{ GeV}, \quad m_\eta = 63 \text{ GeV}$$

$$(M_{N_1}, M_{N_2}, M_{N_3}) = (3000, 3500, 4000) \text{ GeV}$$

## Higgs potential

$$\mu_2^2 = (50 \text{ GeV})^2, \quad \mu_s^2 = (320 \text{ GeV})^2, \quad \mu_\eta^2 \simeq (62.7 \text{ GeV})^2, \quad \mu_{12}^2 = 0$$

$$\lambda_2 = 0.1, \quad \lambda_3 \simeq 1.98, \quad \lambda_4 \simeq 1.88, \quad \lambda_5 \simeq 1.88, \quad \lambda_6 = 0,$$

$$|\lambda_7| = 0.821, \quad \rho_1 \simeq 1.90, \quad |\rho_{12}| = 0.1, \quad \rho_2 = 0.1,$$

$$\sigma_1 = |\sigma_{12}| = 1.1 \times 10^{-3}, \quad \kappa = 2.0, \quad \lambda_S = \lambda_\eta = \xi = 1$$

$$\theta_7 = -2.34, \quad \theta_\rho = -2.94, \quad \theta_\sigma = 0$$

## The benchmark scenario

### Yukawa interactions

$$y_u |\zeta_u| \simeq 2.2 \times 10^{-6}, \quad y_c |\zeta_u| \simeq 1.3 \times 10^{-3}, \quad y_t |\zeta_u| \simeq 0.17,$$

$$y_d |\zeta_d| \simeq 4.7 \times 10^{-6}, \quad y_s |\zeta_d| \simeq 9.3 \times 10^{-5}, \quad y_b |\zeta_d| \simeq 4.2 \times 10^{-3},$$

$$y_e |\zeta_e| \simeq 2.5 \times 10^{-4}, \quad y_\mu |\zeta_\mu| \simeq 2.5 \times 10^{-4}, \quad y_\tau |\zeta_\tau| \simeq 2.5 \times 10^{-3},$$

$$\theta_u = \theta_d = 0.245, \quad \theta_e = \theta_\mu = \theta_\tau = -2.94$$

$$h_i^\alpha \simeq \begin{pmatrix} 1.0 e^{-0.31i} & 0.2 e^{0.30i} & 1.0 e^{-2.4i} \\ 1.1 e^{-1.9i} & 0.21 e^{-1.8i} & 1.1 e^{2.3i} \\ 0.45 e^{2.7i} & 1.3 e^{-0.033i} & 0.10 e^{0.63i} \end{pmatrix}$$

## Messes of the scalar bosons

$$m_{H^+}^2 = \mu_2^2 + \frac{1}{2}\lambda_3 v^2, \quad m_{H_2}^2 = \mu_2^2 + \frac{1}{2}(\lambda_3 + \lambda_4 + \lambda_5)v^2,$$

$$m_{H_3}^2 = \mu_2^2 + \frac{1}{2}(\lambda_3 + \lambda_4 - \lambda_5)v^2,$$

$$m_{S^+}^2 = \mu_s^2 + \frac{1}{2}\rho_1 v^2, \quad m_\eta^2 = \mu_\eta^2 + \frac{1}{2}\sigma_1 v^2$$

$$m_{H^+} = 250 \text{ GeV}, \quad m_{H_2} = 420 \text{ GeV}, \quad m_{H_3} = 250 \text{ GeV}$$

$$m_S = 400 \text{ GeV}, \quad m_\eta = 63 \text{ GeV}$$

$$\mu_2^2 = (50 \text{ GeV})^2, \quad \mu_s^2 = (330 \text{ GeV})^2, \quad \mu_\eta^2 \simeq (62.7 \text{ GeV})^2,$$

$$\lambda_3 \simeq 1.98, \quad \lambda_4 \simeq 1.88, \quad \lambda_5 \simeq 1.88, \quad \rho_1 \simeq 1.90, \quad \sigma_1 = 1.1 \times 10^{-3}$$

# CPV phases in the Yukawa matrix $h$ $h_i^\alpha \overline{(N_R^\alpha)^c} \ell_R^i S^+$

The Yukawa matrix  $h$  includes nine phases.

Three of them can be zero by rephasing lepton fields.

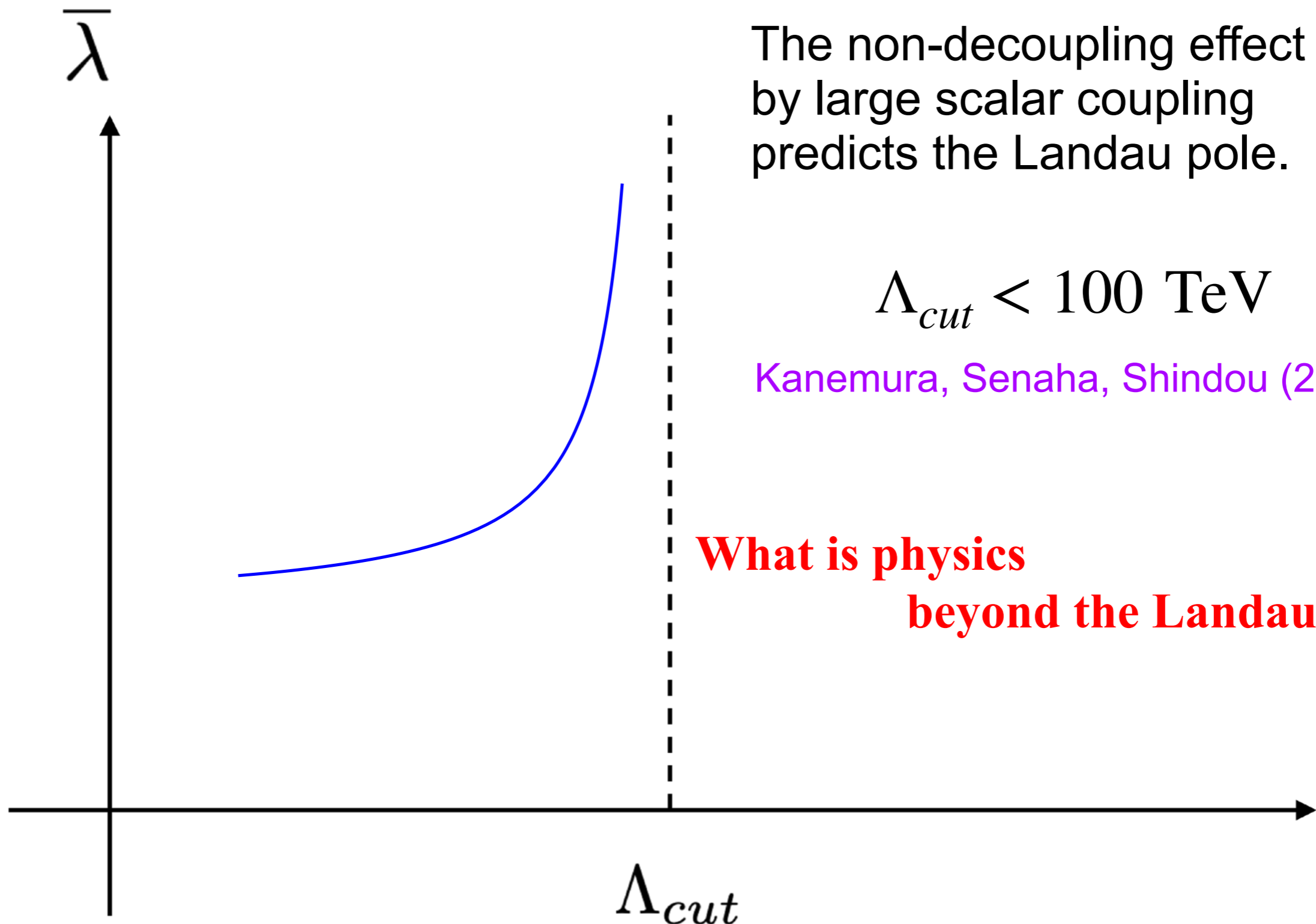
$$\begin{pmatrix} e_{L,R} \\ \mu_{L,R} \\ \tau_{L,R} \end{pmatrix} \rightarrow P_\phi \begin{pmatrix} e_{L,R} \\ \mu_{L,R} \\ \tau_{L,R} \end{pmatrix} \quad \begin{pmatrix} \nu_{eL} \\ \nu_{\mu L} \\ \nu_{\tau L} \end{pmatrix} \rightarrow P_\phi \begin{pmatrix} \nu_{eL} \\ \nu_{\mu L} \\ \nu_{\tau L} \end{pmatrix} \quad P_\phi \equiv \begin{pmatrix} e^{i\phi_e} & 0 & 0 \\ 0 & e^{i\phi_\mu} & 0 \\ 0 & 0 & e^{i\phi_\tau} \end{pmatrix}$$

This rephasing can eliminate **3 phases from the PMNS matrix.**

$$\begin{pmatrix} \nu'_{eL} \\ \nu'_{\mu L} \\ \nu'_{\tau L} \end{pmatrix} = P_\phi \begin{pmatrix} \nu_{eL} \\ \nu_{\mu L} \\ \nu_{\tau L} \end{pmatrix} \quad U_{\text{PMNS}} = P_\phi U'_{\text{PMNS}}$$

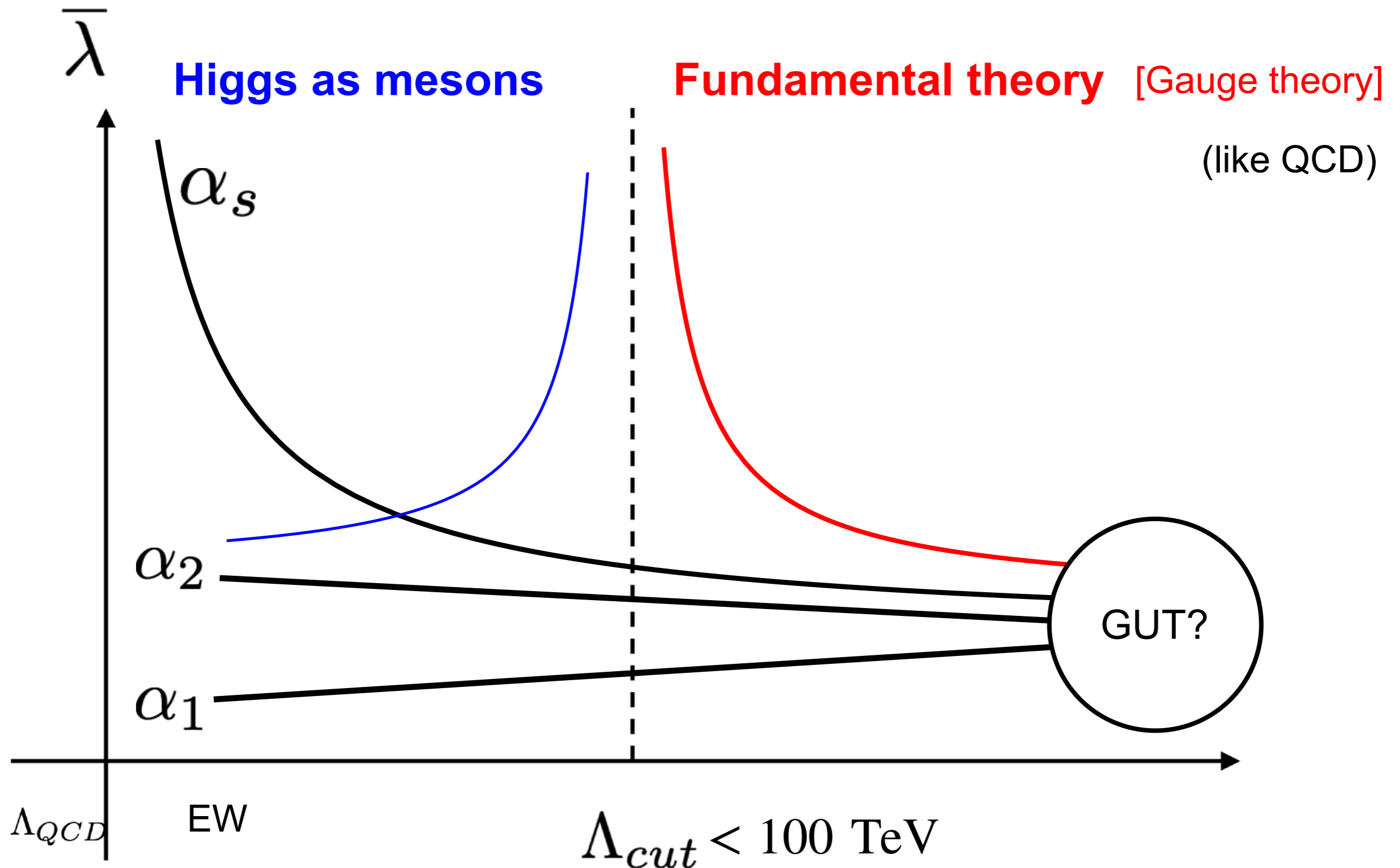
$U_{\text{PMNS}}$  includes only 3 CPV phases:  $\delta_{CP}$ ,  $\alpha_1$ ,  $\alpha_2$

# Landau pole and new physics





# Landau pole and new physics



# Landau pole and new physics

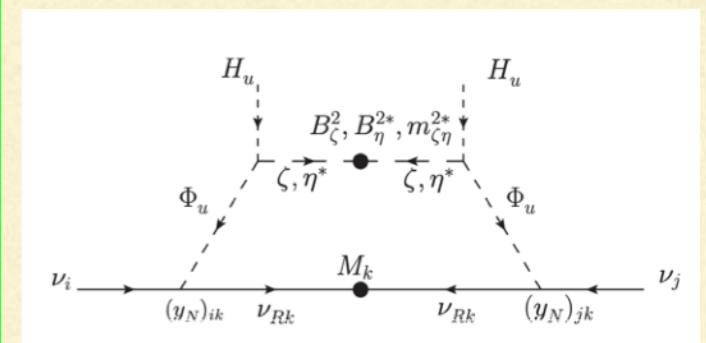
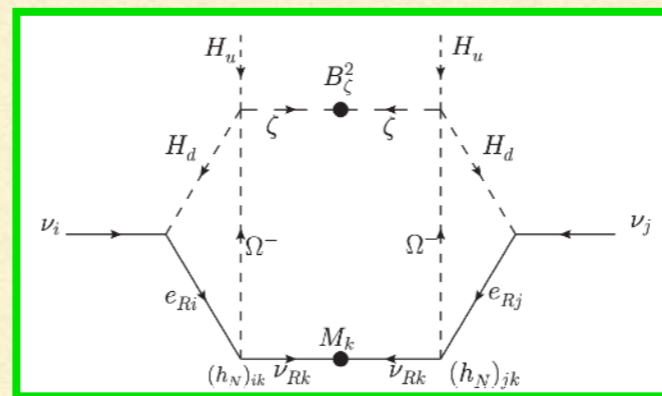
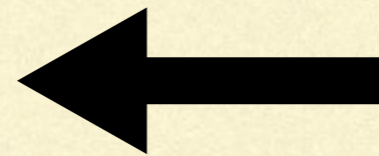
E.g.) SUSY  $SU(2)_H$  gauge theory [Kanemura, Shindou, Yamada, PRD \(2012\)](#)

## Higgs as mesons

Field	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$Z_2$
$H_u$	1	2	+1/2	+1
$H_d$	1	2	-1/2	+1
$\Phi_u$	1	2	+1/2	-1
$\Phi_d$	1	2	-1/2	-1
$\Omega^+$	1	1	+1	-1
$\Omega^-$	1	1	-1	-1
$N, N_\Phi, N_\Omega$	1	1	0	+1
$\zeta, \eta$	1	1	0	-1

## Gauge theory

Superfield	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$Z_2$
$\begin{pmatrix} T_1 \\ T_2 \end{pmatrix}$	1	2	0	+1
$T_3$	1	1	+1/2	+1
$T_4$	1	1	-1/2	+1
$T_5$	1	1	+1/2	-1
$T_6$	1	1	-1/2	-1



**ALL** scalar fields in the model can be included!