# A model with two scalars in the view of effective theory Based on: arXiv:2108.03639 [hep-ph]; PTEP 2022 1, 013B01 (2022)

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## Outline

- Introduction
- Model
- Effective Potential
- RG Improved Effective Potential
- Summary and Outlook

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# Research Background

## (1) Higgs Boson

- It has been discovered in 2012 which confirms its mass of around 125 GeV
- A scalar boson with spin 0 and charge 0
- Responsible for fermions (quark and lepton) and gauge boson mass generations.



#### **Problems**

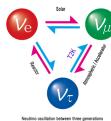
- Experiment results of Higgs do not tell us clearly how Spontaneous symmetry breaking occurs and how many Higgs particles are included in the process.
- Is there only one Higgs Boson?
  - → BSM with multiple Higgs is possible.
- Is there a heavier mass than the energy range of the current experiment?
  - → The other (second) Higgs boson has not been observed yet.



# Research Background...

#### (2) Neutrino

- Three flavors exist and oscillate with each other. (electron neutrino  $\nu_e$ , muon neutrino  $\nu_\mu$ , tau neutrino  $\nu_\tau$ )
- Charge is zero with a very tiny mass.
- Only interacting weakly with matter.



https://www.secretsofuniverse.in/8-

interesting-facts-about-neutrinos/

#### **Problems**

- Mass Hierarchy problem (Normal or Inverted hierarchies),
   Either is Majorana or Dirac particles, etc.
- In order to explain the small mass of the neutrinos with the Yukawa couplings of neutrinos, we must assume a quite small of Yukawa couplings.

#### Motivation

#### Motivation

- If a second Higgs boson (assuming with heavy mass) exists, we would like to verify its existence → direct detection or indirect detection.
- a Two Higgs doublet model (THDM) [1] can explain the small mass of (Dirac) neutrinos without assuming a small Yukawa coupling.
  - [1] S.M. Davidson and H.E. Logan, Phys. Rev. D80 (2009).

#### Purpose

 To study the THDM with features like [1] with the viewpoint of the effective theory.

#### Method

 We construct an effective field theory of THDM in the low-energy region by integrating the heavy particles which are not observed on the low energy.

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# Effective Field Theory

- Effective field theory (EFT) is one of the useful tools in the research of physical systems.
- In EFT, the theory on the energy scale  $\mu$  is derived by integrating out particles with a mass higher than the energy scale  $\mu$ .
- Example of EFT:  $\beta$  decay:  $d \rightarrow u + e^- + \bar{\nu}_e$

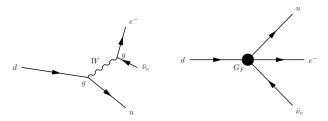


Figure: Feynman diagram for beta decay at quark level is shown (left) and the Feynman diagram in the weak EFT in which the W boson is integrated (right).

#### The Model

- We consider a simple toy model that can explain the small mass of Dirac neutrinos by the small vacuum expectation value of the second scalar.
- Action (in terms of bare (neutral) fields, bare mass, bare coupling and neutrino )

$$S = \int d^d x \left( -\frac{1}{2} \sum_{i=1}^2 \rho_{0i} \left( \Box + \frac{m_{0i}^2}{2} + \frac{\lambda_{0i}}{2} \rho_{0i}^2 \right) \rho_{0i} - \frac{\lambda_{03}}{4} \rho_{01}^2 \rho_{02}^2 - (y_0 \bar{n}_0 n_0 + m_{012}^2 \rho_{01}) \rho_{02} + \sum_{i=1}^2 h_{0i} m_{0i}^4 + h_{012} m_{012}^4 + 2h_{03} m_{01}^2 m_{02}^2 \right)$$

with  $m_{012}^2$  is the bare mixing mass and  $y_0$  is the Yukawa coupling of the neutrino and the second scalar.

• We assume the following hierarchy for the mass parameters:

$$m_2^2 \gg -m_1^2 \simeq \epsilon \ m_{12}^2 > 0, \quad \epsilon := \frac{m_{12}^2}{m_2^2} \ll 1$$



# Symmetry of Model

- We impose the two symmetries ( $Z_2$  and  $Z_2'$  symmetries) on the model.
- In order to forbid Majorana mass terms such as  $\overline{(n_R)^c}n_R$  and  $\overline{(n_L)^c}n_L$ , we impose the symmetry under the transformation,  $(n_L,n_R) \to e^{i\frac{\pi}{2}}(n_L,n_R)$ .

Symmetry	$\rho_1$	$\rho_2$	$n_L$	$n_R$
$Z_2$	_	_	+	_
$Z_2'$	_	+	+	+

Table: The charge assignment under  $Z_2$  and  $Z_2'$  symmetries.

## Action in terms of Renormalized Quantities

• The relations between bare quantities and renormalized ones are:

$$\rho_{0i} = \sqrt{Z_i}\rho_i,$$

$$n_0 = \sqrt{Z_n}n,$$

$$m_{0i}^2 Z_i = \sum_{j=1}^2 Z_{mij}m_j^2,$$

$$m_{012}^2 \sqrt{Z_1 Z_2} = m_{12}^2 Z_{12},$$

$$\lambda_{0i} Z_i^2 = \sum_{I=1}^3 Z_{\lambda_{iI}} \lambda_I \mu^{2\eta},$$

$$\lambda_{03} Z_1 Z_2 = \sum_{I=1}^3 Z_{\lambda_{3I}} \lambda_I \mu^{2\eta},$$

$$y_0 Z_n \sqrt{Z_2} = Z_y y \mu^{\eta},$$

 $\mu$ : renormalization scale and  $\eta=2-\frac{d}{2}.$ 

## Action in terms of Renormalized Quantities

$$S[\rho_{1}, \rho_{2}, n]$$

$$= -\frac{1}{2} \int d^{d}x \sum_{i=1}^{2} \left( Z_{i} \rho_{i} \Box \rho_{i} + \rho_{i}^{2} Z_{mij} m_{j}^{2} + \frac{\mu^{2\eta}}{2} \sum_{I=1}^{3} \left( \rho_{i}^{4} Z_{\lambda_{iI}} \lambda_{I} + \rho_{1}^{2} \rho_{2}^{2} Z_{\lambda_{3}I} \lambda_{I} \right) \right)$$

$$- \int d^{d}x \left( Z_{y} y \mu^{\eta} \bar{n} n + Z_{12} m_{12}^{2} \rho_{1} \right) \rho_{2}$$

$$+ \int d^{d}x \mu^{-2\eta} \left( \sum_{i=1}^{2} Z_{hi} h_{i} m_{i}^{4} + 2 Z_{h3} h_{3} m_{1}^{2} m_{2}^{2} + Z_{h12} h_{12} m_{12}^{4} \right).$$

## Action in terms of Renormalized Quantities

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We do not include the kinetic term for neutrino and its quantum correction.

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# Low-Energy Effective Action

ullet Introducing the following generating functional  $W[J_1,n]$ ,

$$e^{iW[J_1,n]} = \int d\rho_1 \int d\Delta_2 e^{iS[\rho_1,\Delta_2,n] + i \int \rho_1 J_1 d^4 x}$$

• We can define the effective action  $\Gamma_{\rm eff}$ , a functional of  $\bar{\rho}_1$  by Legendre transformation of  $W[J_1,n]$  as,

$$\begin{split} \Gamma_{\text{eff}}[\bar{\rho}_{1},n] &= W[J_{1},n] - \int J_{1}\bar{\rho}_{1}d^{4}x \\ &= -i\log\int d\Delta_{1}\int d\Delta_{2}e^{iS[\bar{\rho}_{1}+\Delta_{1},\Delta_{2},n]-i\int\Delta_{1}\frac{\delta\Gamma_{\text{eff}}[\bar{\rho}_{1},n]}{\delta\bar{\rho}_{1}(x)}d^{4}x} \\ \Delta_{1} &\equiv \rho_{1} - \bar{\rho}_{1}, \quad \Delta_{2} = \rho_{2}, \quad J_{1}(x) = -\frac{\delta\Gamma_{\text{eff}}[\bar{\rho}_{1},n]}{\delta\bar{\rho}_{1}(x)} \end{split}$$

•  $\Delta_1$  is the quantum fluctuation from the expectation value  $\bar{\rho}_1$ .



## Low-Energy Effective Action...

ullet  $ar{
ho}_1$  is the expectation value of  $ho_1$  defined by,

$$\bar{\rho}_1|_{J_1} = \frac{\delta W[J_1,n]}{\delta J_1} = \frac{\int d\rho_1 \int d\Delta_2 \rho_1 e^{iS[\rho_1,\Delta_2,n] + i \int \rho_1 J_1 d^4 x}}{\int d\rho_1 \int d\Delta_2 e^{iS[\rho_1,\Delta_2] + i \int \rho_1 J_1 d^4 x}}$$

 Next, we define the following quantity by subtracting the classical action from the effective action,

$$\tilde{\Gamma}_{\mathrm{eff}}[\bar{\rho}_{1},n] = \Gamma_{\mathrm{eff}}[\bar{\rho}_{1},n] - S[\bar{\rho}_{1},0,0],$$

$$\begin{split} &e^{i\tilde{\Gamma}_{\text{eff}}[\bar{\rho}_{1},n]} \\ &= \int d\Delta_{1} e^{i\left\{\frac{1}{2}\int d^{d}x \int d^{d}y \Delta_{1}(x) \frac{\delta^{2}S[\bar{\rho}_{1},0,n]}{\delta\bar{\rho}_{1}(x)\delta\bar{\rho}_{1}(y)} \Delta_{1}(y) + S_{\text{lint}}(\Delta_{1},\bar{\rho}_{1}) - \int d^{d}x \Delta_{1}(x) \left(\frac{\delta\tilde{\Gamma}_{\text{eff}}[\bar{\rho}_{1},n]}{\delta\bar{\rho}_{1}(x)}\right)\right\}} \\ &\times e^{iW_{2}[J_{2},\Delta_{1},n]}. \end{split}$$

• The last factor summarizes the contribution from the heavy field to the effective action.

 $\bullet$  Next, the contribution from the heavy field to EA,  $e^{iW_2[J_2,\bar{\rho}_1,\Delta_1]}$  , is

$$e^{iW_{2}[J_{2},\bar{\rho}_{1},\Delta_{1}]} = \left(\det \frac{\delta^{2}S[\rho_{1},\rho_{2},n]}{\delta\rho_{2}(x)\delta\rho_{2}(y)}\Big|_{\rho_{1}=\bar{\rho}_{1},\rho_{2}=0}\right)^{-\frac{1}{2}} \times \langle e^{iS_{2}\inf(\Delta_{2})+iS_{12}\inf(\Delta_{1},\Delta_{2},\bar{\rho}_{1})}\rangle_{0} e^{iW_{2}^{c}[J_{2},\bar{\rho}_{1},\Delta_{1}]},$$

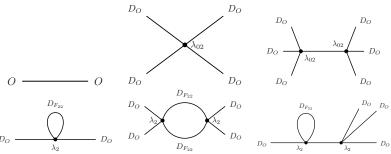
$$e^{iW_2^c[J_2,\bar{\rho}_1,\Delta_1]} = \frac{\left\langle e^{iS_{2\,\mathrm{int}}(\Delta_2) + iS_{12\,\mathrm{int}}(\Delta_1,\Delta_2,\bar{\rho}_1) + i\int d^dx \Delta_2(x)J_2(x)}\right\rangle_0}{\left\langle e^{iS_{2\,\mathrm{int}}(\Delta_2) + iS_{12\,\mathrm{int}}(\Delta_1,\Delta_2,\bar{\rho}_1)}\right\rangle_0}$$

- ullet The first factor is the vacuum graph contribution from the quadratic part of  $\Delta_2$  and the second part corresponds to that from the interaction.
- The third factor is the connected Green function contribution of  $\Delta_2$  with the source term  $J_2$ :

$$J_2(x) = -Z_{12}m_{12}^2(\Delta_1 + \bar{\rho}_1) - \mu^{\eta} y Z_y \bar{n} n.$$

that consists of the fields linearly coupled to the heavy scalar  $\Delta_2$ .



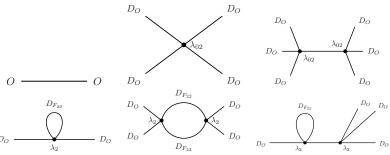


ightarrow All the propagators are heavy scalars and the external source terms denoted by O correspond to the background field  $\bar{\rho}_1$  and the bilinear neutrino  $\bar{n}n$ .

$$iW_2^c[J_2, \bar{\rho}_1, \Delta_1] = iW_2^{c(\text{tree})}[J_2, \bar{\rho}_1, \Delta_1] + i\bar{W}_2^{c(1\text{loop})}[\bar{\rho}_1, n].$$

$$D_O(x) = \int d^d x_i D_{F22}(x, x_i) iO(x_i),$$

$$-J_2(\Delta_1 = 0) = O(x) = Z_{12} m_{12}^2 \bar{\rho}_1 + \mu^n y Z_y \bar{n} n$$



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$$D_O(x) = \int d^d x_i D_{F22}(x, x_i) iO(x_i),$$

$$-J_2(\Delta_1 = 0) = O(x) = Z_{12}m_{12}^2 \bar{\rho}_1 + \mu^{\eta} y Z_u \bar{n} n$$

I do not show how we integrate the lighter scalar field in this presentation.

#### • Substituting the vacuum expectation value $v_1$ for $\bar{\rho}_1$ , we obtain

$$\begin{split} V_{\rm eff}(v_1) = & V_{\rm cosmo} + \frac{m_{1\rm eff}^2}{2} v_1^2 - \frac{m_{1\rm 2\rm eff}^2}{2} \epsilon v_1^2 + \frac{\lambda_{1\rm eff}}{4} v_1^4 + \frac{\lambda_{3\rm eff}}{4} \epsilon^2 v_1^4 - y_{\rm eff} \epsilon \bar{n} n v_1 \\ & - \frac{\lambda_3^2}{8m_2^2} \epsilon^2 v_1^6 + \frac{\lambda_3}{2m_2^2} \epsilon (\bar{n}n) v_1^3 - \frac{y^2}{2m_2^2} (\bar{n}n)^2. \\ V_{\rm cosmo} = & -h_{1\rm eff} m_1^4 - h_{2\rm eff} m_2^4 - h_{1\rm 2\rm eff} m_{12}^4 - 2h_{3\rm eff} m_1^2 m_2^2, \\ m_{1\rm eff}^2 = & m_1^2 \left( 1 - \frac{3\lambda_1}{16\pi^2} \left( \frac{3}{2} - \log \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right) - \frac{\lambda_3 m_2^2}{32\pi^2} \left( 1 - \log \frac{m_2^2}{\mu^2} \right), \\ m_{1\rm 2\rm eff}^2 = & m_{12}^2 \left\{ 1 + \frac{3\lambda_1}{16\pi^2} \log \frac{m_1^2 + 3\lambda_1 v_1^2}{m_2^2} + \frac{3\lambda_2}{16\pi^2} \left( 1 - \log \frac{m_2^2}{\mu^2} \right) - \frac{\lambda_3}{8\pi^2} \left( \frac{5}{4} - \log \frac{m_2^2}{\mu^2} \right) \right\}, \\ \frac{\lambda_1 {\rm eff}}{4} = & \frac{\lambda_1}{4} \left\{ 1 - \frac{9\lambda_1}{16\pi^2} \left( \frac{3}{2} - \log \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) + \frac{1}{64\pi^2} \frac{\lambda_3^2}{\lambda_1} \log \frac{m_2^2}{\mu^2} \right\} \\ \frac{\lambda_3 {\rm eff}}{4} = & \frac{\lambda_3}{4} \left\{ 1 + \frac{3\lambda_2}{16\pi^2} \left( 2 - \log \frac{m_2^2}{\mu^2} \right) - \frac{\lambda_3}{4\pi^2} \left( \frac{25}{16} - \log \frac{m_2^2}{\mu^2} \right) + \frac{3\lambda_1}{16\pi^2} \left( \log \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} + 5 \log \frac{m_1^2 + 3\lambda_1 v_1^2}{m_2^2} - 6 \frac{\lambda_1}{\lambda_3} \log \frac{m_1^2 + 3\lambda_1 v_1^2}{m_2^2} \right) \right\}, \\ h_{1\rm eff} = & h_1 + \frac{1}{64\pi^2} \left( \frac{3}{2} - \log \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right), \quad h_{2\rm eff} = h_2 + \frac{1}{64\pi^2} \left( \frac{3}{2} - \log \frac{m_2^2}{\mu^2} \right), \\ h_{3\rm eff} = h_3 + \frac{\epsilon^2}{32\pi^2} \log \frac{m_1^2 + 3\lambda_1 \bar{p}_1^2}{\mu^2}, \quad h_{12\rm eff} = h_{12} + \frac{1}{32\pi^2} \left( 1 - \log \frac{m_2^2}{\mu^2} \right), \\ y_{\rm eff} = & y \left( 1 + \frac{3\lambda_2 - \lambda_3}{16\pi^2} \left( 1 - \log \frac{m_2^2}{\mu^2} \right) \right). \end{aligned}$$

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# Low energy theory without heavy scalar

 Integrate heavy scalars in the original theory and build a low-energy theory without heavy scalars (including its counter terms)

$$\begin{split} V_{\text{eff}}^{\text{Low}} &= \frac{m_1^2}{2} \left\{ 1 - \frac{3\lambda_1}{16\pi^2} \left( \frac{3}{2} - \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right) \right\} v_1^2 \\ &- \frac{m_{12}^2}{2} \left\{ 1 - \frac{3\lambda_1}{16\pi^2} \left( 1 - \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right) \right\} \epsilon v_1^2 - \epsilon y \bar{n} n v_1 - \frac{1}{2m_2^2} \left( \frac{\lambda_3 \epsilon}{2} v_1^3 - y \bar{n} n \right)^2 \\ &+ \frac{\lambda_1}{4} \left\{ 1 - \frac{9\lambda_1}{16\pi^2} \left( \frac{3}{2} - \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right) \right\} v_1^4 \\ &+ \frac{\lambda_3}{4} \left\{ 1 + \frac{9}{8\pi^2} \frac{\lambda_1 (\lambda_1 - \lambda_3)}{\lambda_3} + \frac{9}{8\pi^2} \frac{\lambda_1 (\lambda_3 - \lambda_1)}{\lambda_3} \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right\} \epsilon^2 v_1^4 \\ &- m_1^4 \left\{ h_1 + \frac{1}{64\pi^2} \left( \frac{3}{2} - \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right) \right\} \\ &+ 2\epsilon^2 m_1^2 m_2^2 \left\{ -\frac{h_3}{\epsilon^2} + \frac{1}{64\pi^2} \left( 1 - \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right) \right\} \\ &- \epsilon^2 m_{12}^4 \left\{ \frac{h_{12}}{\epsilon^2} - \frac{1}{64\pi^2} \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{\mu^2} \right) \right\} - h_2 m_2^4. \end{split}$$

# RG Improved Effective Potential

 Since we can improve the low energy effective potential by RG, we are able to obtain the RG improved effective potential defined by,

$$V_{\mathrm{eff}}^{\mathrm{Improved}} = V_{\mathrm{eff}}(\mu = m_2) - V_{\mathrm{eff}}^{\mathrm{Low}}(\mu = m_2) + V_{\mathrm{eff}}^{\mathrm{Low \ RGimproved}},$$

 $\bullet~V_{\rm eff}^{\rm Low~RGimproved}$  is obtained by using the RG improved coupling and masses (after solving the RG equations)

$$\lambda'_{1}(\mu_{0}) = \frac{\lambda'_{1}(\mu)}{1 + \frac{9\lambda'_{1}(\mu)}{16\pi^{2}} \log \frac{\mu^{2}}{\mu_{0}^{2}}},$$

$$m'_{1}^{2}(\mu_{0}) = \frac{m'_{1}^{2}(\mu)}{\left(1 + \frac{9\lambda'_{1}(\mu)}{16\pi^{2}} \log \frac{\mu^{2}}{\mu_{0}^{2}}\right)^{\frac{1}{3}}},$$

$$\bar{h}(\mu_{0}) = \bar{h}(\mu) - \frac{1}{64\pi^{2}} \log \frac{\mu_{0}^{2}}{\mu^{2}}.$$

$$\begin{split} V_{\text{eff}}^{\text{Improved}} = & V_{\text{cosmo}}^{\text{(Imp.)}} + \frac{m_{\text{1eff}}^{2,\text{(Imp.)}}}{2} v_{1}^{2} - \frac{m_{12\text{eff}}^{2,\text{(Imp.)}}}{2} \epsilon v_{1}^{2} + \frac{\lambda_{\text{1eff}}^{\text{(Imp.)}}}{4} v_{1}^{4} + \frac{\lambda_{3\text{eff}}^{\text{(Imp.)}}}{4} \epsilon^{2} v_{1}^{4} \\ & - \left(1 + \frac{3\lambda_{2} - \lambda_{3}}{16\pi^{2}}\right) y \epsilon \bar{n} n v_{1} - \frac{1}{2m_{2}^{2}} \left(\frac{\lambda_{3}}{2} \epsilon v_{1}^{3} - y \bar{n} n\right)^{2} \end{split}$$

Cosmological term of RG improved effective potential

$$\begin{split} V_{\text{cosmo}}^{\text{(Imp.)}} &= -\,m_1^4 \left\{ h_1(m_2) + \frac{1}{64\pi^2} \left( \frac{3}{2} - \log \frac{m_1^2 + 3\lambda_1 v_1^2}{m_2^2} \right) \right\} \\ &- m_2^4 \left( h_2(m_2) + \frac{3}{128\pi^2} \right) - m_{12}^4 \left( h_{12}(m_2) + \frac{1}{32\pi^2} \right) \\ &- 2m_1^2 m_2^2 \left\{ h_3(m_2) + \frac{\epsilon^2}{64\pi^2} \log \left( \frac{m_1^2 + 3\lambda_1 v_1^2}{m_2^2} \right) \right\} \end{split}$$

$$\begin{split} V_{\text{eff}}^{\text{Improved}} = & V_{\text{cosmo}}^{\text{(Imp.)}} + \frac{m_{\text{1eff}}^{2,\text{(Imp.)}}}{2} v_{1}^{2} - \frac{m_{12\text{eff}}^{2,\text{(Imp.)}}}{2} \epsilon v_{1}^{2} + \frac{\lambda_{\text{1eff}}^{\text{(Imp.)}}}{4} v_{1}^{4} + \frac{\lambda_{3\text{eff}}^{\text{(Imp.)}}}{4} \epsilon^{2} v_{1}^{4} \\ & - \left(1 + \frac{3\lambda_{2} - \lambda_{3}}{16\pi^{2}}\right) y \epsilon \bar{n} n v_{1} - \frac{1}{2m_{2}^{2}} \left(\frac{\lambda_{3}}{2} \epsilon v_{1}^{3} - y \bar{n} n\right)^{2} \end{split}$$

Mass squared term of RG improved effective potential

$$\begin{split} m_{1\text{eff}}^{2,(\text{Imp.})} = & m_1^2(m_2) \left[ -\frac{\lambda_3}{32\pi^2} \frac{m_2^2}{m_1^2} + \frac{1 - \frac{9\lambda_1(m_2)}{32\pi^2}}{\left(1 + \frac{9\lambda_1(m_2)}{16\pi^2} \log\left(\frac{m_2^2}{m_1^2 + 3\lambda_1 v_1^2}\right)\right)^{\frac{1}{3}}} \right] \\ m_{12\text{eff}}^{2,(\text{Imp.})} = & m_{12}^2(m_2) \left[ -\frac{5\lambda_3}{32\pi^2} + \frac{3\lambda_2}{16\pi^2} + \frac{3\lambda_1}{16\pi^2} + \frac{1 - \frac{3\lambda_1(m_2)}{16\pi^2}}{\left(1 + \frac{9\lambda_1(m_2)}{16\pi^2} \log\left(\frac{m_2^2}{m_1^2 + 3\lambda_1 v_1^2}\right)\right)^{\frac{1}{3}}} \right] \end{split}$$

$$\begin{split} V_{\text{eff}}^{\text{Improved}} = & V_{\text{cosmo}}^{\text{(Imp.)}} + \frac{m_{\text{1eff}}^{2,\text{(Imp.)}}}{2} v_{1}^{2} - \frac{m_{12\text{eff}}^{2,\text{(Imp.)}}}{2} \epsilon v_{1}^{2} + \frac{\lambda_{\text{1eff}}^{\text{(Imp.)}}}{4} v_{1}^{4} + \frac{\lambda_{3\text{eff}}^{\text{(Imp.)}}}{4} \epsilon^{2} v_{1}^{4} \\ & - \left(1 + \frac{3\lambda_{2} - \lambda_{3}}{16\pi^{2}}\right) y \epsilon \bar{n} n v_{1} - \frac{1}{2m_{2}^{2}} \left(\frac{\lambda_{3}}{2} \epsilon v_{1}^{3} - y \bar{n} n\right)^{2} \end{split}$$

Quartic term of RG improved effective potential

$$\begin{split} \lambda_{1\text{eff}}^{\text{(Imp.)}} = & \lambda_{1}(m_{2}) \left[ \frac{1 - \frac{27\lambda_{1}(m_{2})}{32\pi^{2}}}{1 + \frac{9\lambda_{1}(m_{2})}{16\pi^{2}} \log\left(\frac{m_{2}^{2}}{m_{1}^{2} + 3\lambda_{1}v_{1}^{2}}\right)} \right] \\ \lambda_{3\text{eff}}^{\text{(Imp.)}} = & \lambda_{3}(m_{2}) \left[ \frac{6\lambda_{2}}{16\pi^{2}} - \frac{25\lambda_{3}}{64\pi^{2}} + \frac{9\lambda_{1}}{8\pi^{2}} \left(1 - \frac{\lambda_{1}}{\lambda_{3}}\right) \right. \\ \left. + \frac{1 - \frac{9\lambda_{1}(m_{2})}{8\pi^{2}} \left(1 - \frac{\lambda_{1}(m_{2})}{\lambda_{3}(m_{2})}\right)}{1 + \frac{9\lambda_{1}(m_{2})}{8\pi^{2}} \left(1 - \frac{\lambda_{1}(m_{2})}{\lambda_{3}(m_{2})}\right) \log\left(\frac{m_{2}^{2}}{m_{1}^{2} + 3\lambda_{1}v_{1}^{2}}\right)} \right] \end{split}$$

# Heavy mass dependence on VEV

We study the stationary condition of the improved effective potential,

$$\frac{\partial V_{\text{eff}}^{\text{Improved}}}{\partial v_1^2} = 0.$$

The solution satisfies,

$$\begin{split} \frac{v_1}{v_{10}} &= \sqrt{\left[-\frac{\lambda_3 m_2^2}{32\pi^2 m_1^2} + \frac{1}{\left\{1 - \frac{9\lambda_1}{16\pi^2}\log\left(\frac{m_1^2 + 3\lambda_1 v_1^2}{m_2^2}\right)\right\}^{\frac{1}{3}}}\right] \left[1 - \frac{9\lambda_1}{16\pi^2}\log\left(\frac{m_1^2 + 3\lambda_1 v_1^2}{m_2^2}\right)\right]}, \\ v_{10} &= \sqrt{-\frac{m_1^2}{\lambda_1}}. \end{split}$$

• We keep the leading logarithmic correction and the correction proportional to the heavy scalar mass squared.

# Illustration for heavy mass dependence on VEV

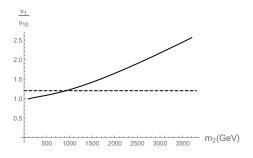


Figure: The vev's ratio as a function of heavy scalar mass  $m_2$ . The dashed line corresponds to  $\frac{v_1}{v_{10}}=1.2$ . We fix the parameters as  $m_1^2=-(100)^2(\text{GeV})^2$  and  $\lambda_1=\lambda_3=1$  which corresponds to  $v_{10}=100$  (GeV).

• The correction to the vev increases as the heavy scalar mass increases.

# Illustration for heavy mass dependence on VEV

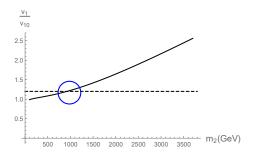


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- The correction to the vev increases as the heavy scalar mass increases.
- If we require the correction to be within 20% compared to the vev without the radiative correction, the upper bound on the heavy scalar mass is about 1000 (GeV).

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# Summary and Outlook

- We have constructed a low-energy effective potential by integrating heavy and light scalar particles.
- We found that the effective Yukawa coupling constant for neutrino mass is inversely proportional to the heavy scalar mass squared as  $y_{\rm eff} \simeq y \frac{m_{12}^2}{m_2^2}$  and is naturally suppressed.
- The effective coupling constant of the quartic interaction of Higgs boson and the quartic interaction term of neutrino is given up to the dimension six operators.
- As an outlook, we will apply the present analysis in the low-energy region for Davidson and Logan model (Ongoing work). In addition, we will discuss the effects that experiments can verify.

# Thank you!

