

BRIEF INTRODUCTION TO MUON COLLIDERS

HPNP 2023 @ Osaka University
June 7, 2023

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Pitt PACC, University of Pittsburgh



A MUON COLLIDER

Why muons?

Although sharing the same EW interactions,
it isn't another electron:

$$m_\mu \approx 207 m_e$$

$$\tau(\mu \rightarrow e\bar{\nu}_e\nu_\mu) \approx 2.2 \mu s$$

$$c\tau \approx 660 m.$$

Once accelerated:

$$E_\mu \sim 1 \text{ TeV} \rightarrow \gamma \sim 10^4 \rightarrow d = c\gamma\tau = 6,600 \text{ km}$$

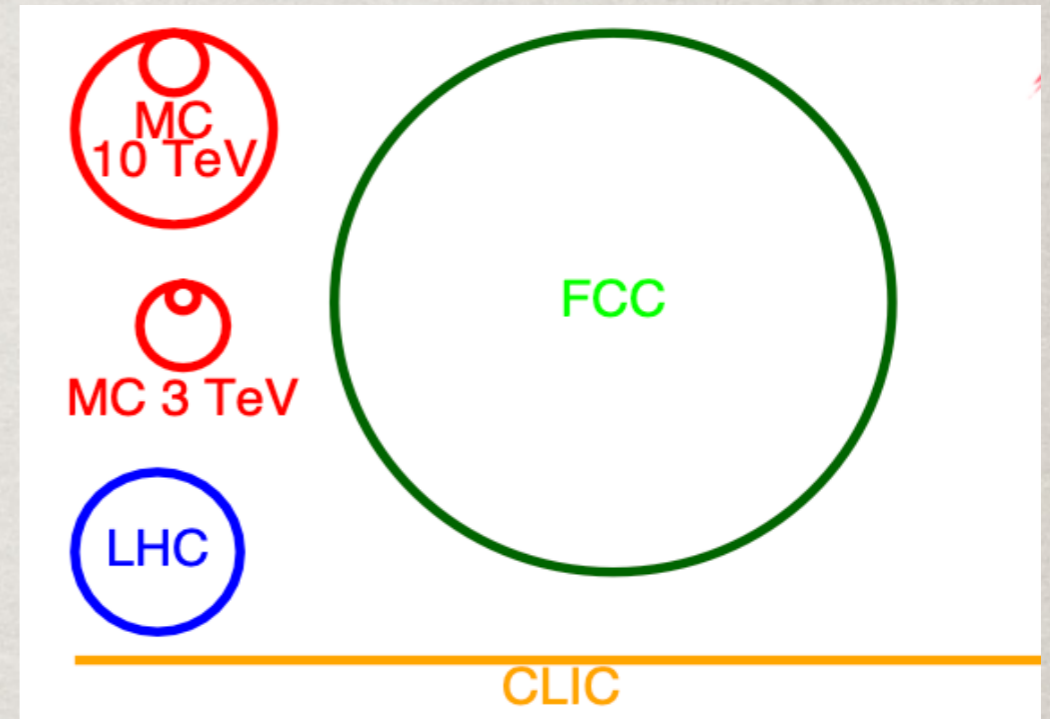
It is these features: heavy mass, short lifetime
that dictate the physics.

- **Advantages of a muon collider**

- Much less synchrotron radiation energy loss than e's:

$$\Delta E \sim \frac{1}{R} \left(\frac{E}{m_\mu} \right)^4$$

which would allow a smaller and a circular machine, thus likely cost-effective:



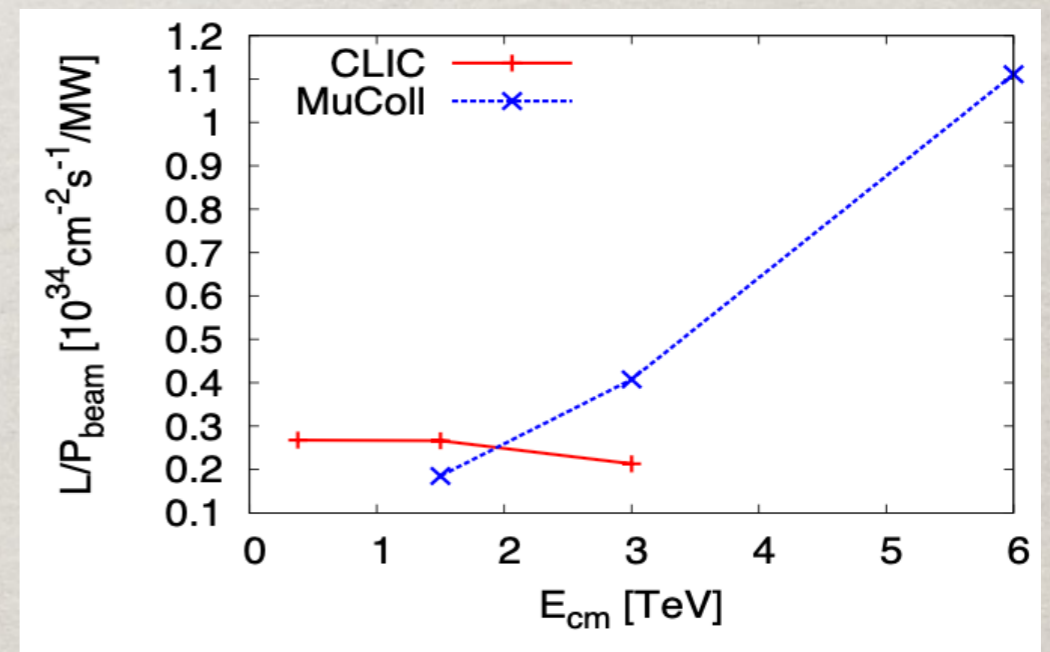
- Luminosity scales with c.m. energy/power, ideally

$$L \sim E^2_{CM}$$

- Smaller beam-energy spread:

$$\Delta E/E \sim 0.1\%$$

potentially $\Delta E/E(m_H) \sim 0.01\% - 0.001\%$



- **Advantages of a muon collider**
- Unlike the proton as a composite particle, E_{CM} efficient in $\mu^+\mu^-$ annihilation, to reach higher new physics threshold $E_{CM} \sim 2 M_{new}$
- Yet, high-energy collisions result in all sort of partons from Initial States Radiation

$$\sigma_{\mu\mu} \sim (1/M_W)^2 \ln^2(E_{CM}/M_W)$$

“Buy one, get one free!”

- Lower (hadronic) background:
 $\sigma_{pp}(\text{total}) \sim 100 \text{ mb}; \quad \sigma_{\mu\mu}(\text{total}) \sim 100 \text{ nb}$

- **Disadvantages of a muon collider**

- Production: Protons on target \rightarrow pions \rightarrow muons:
Require sophisticated scheme for μ capture & transport

- Very short lifetime: in micro-second,

- **Muons cooling in (x,p) 6-dimensions**

- \rightarrow Difficult to make quality beams and a high luminosity

- Beam Induced Backgrounds (BIB)

from the decays in the ring at the interacting point

- Neutrino beam dump (environmental hazard)

$$\sigma_{\nu} \sim G_F^2 E^2 \rightarrow \text{Shielding?}$$

Historically

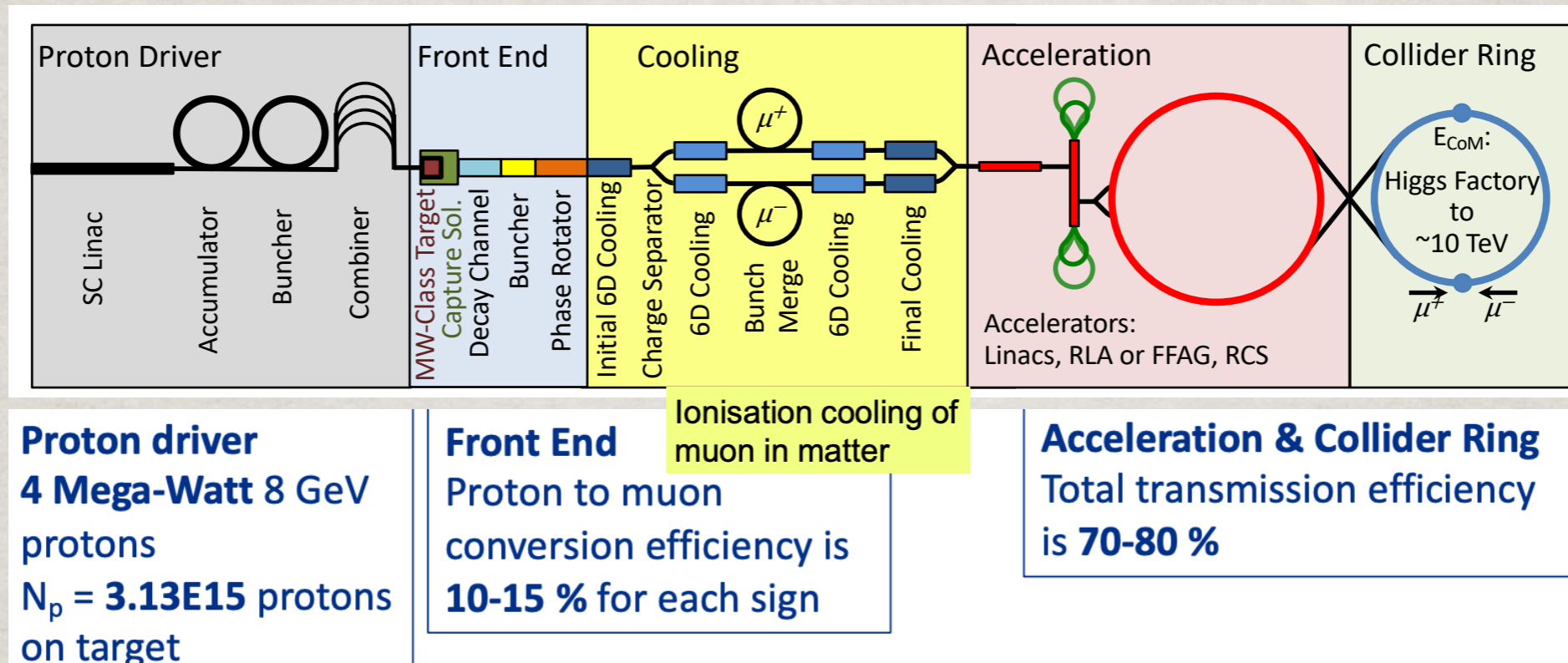
- Concepts mentioned in the 60's
- Early collider design/physics studies in the 90's [*]
- 2011~2016: Muon Accelerator Program formed (MAP): to address key feasibility issues for μC with the proton driver technology
- MAP terminated in 2016, results published in <https://iopscience.iop.org/journal/1748-0221/page/extraproc46>

[*] Some early work:

- *Proceedings of the 1st Workshop on the physics potential and Development of the $\mu\mu$ Colliders*, Napa, California, 1992, Nucl. Inst. Methods. Phys. Res., Sect. A 350, 24 (1994).
- *S-channel Higgs boson production at a muon collider*, Barger et al., PRL75 (1995).
- *$\mu^+ \mu^-$ Collider: Feasibility study*, Muon collider collaboration (July, 1996).
- *Higgs boson physics in the s-channel muon collider*, Barger et al., Phys Rep. 186 (1997).
- *Status of muon collider research*, Muon collider collaboration (Aug., 1999).
- *Recent progress on neutrino factory and muon collider research*, Muon collider collaboration (July, 2003).

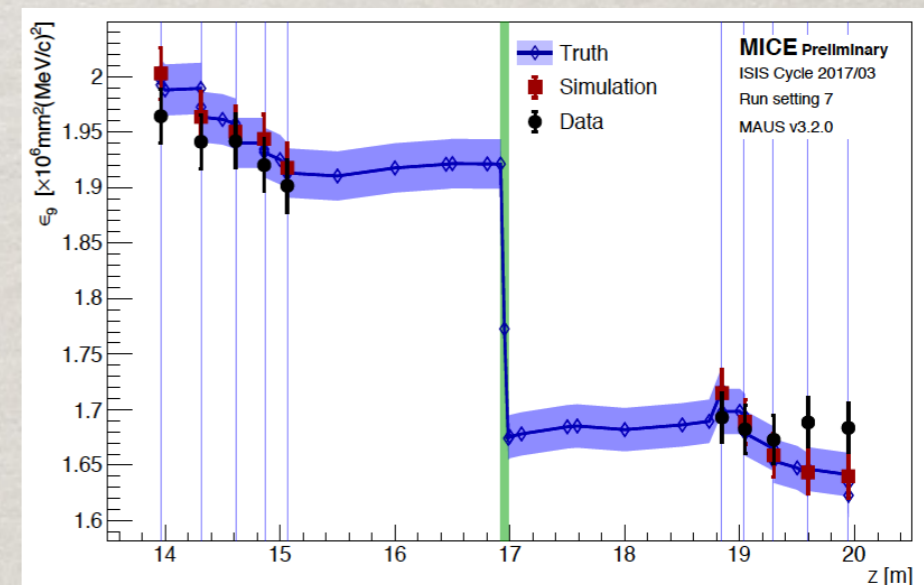
RENEWED INTERESTS

Muon Accelerator Project (MAP)



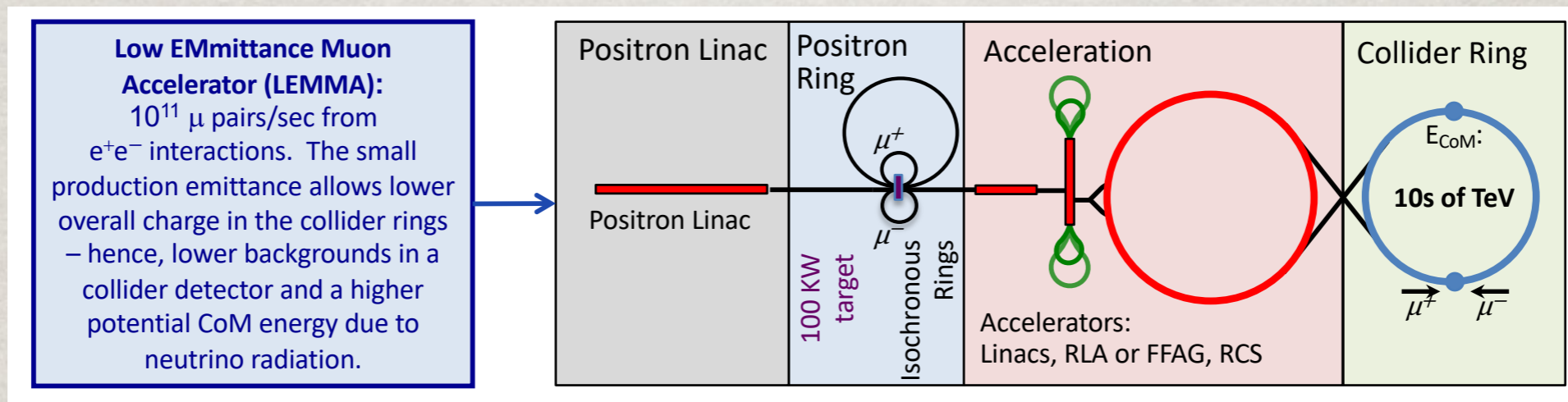
- Protons \rightarrow pions \rightarrow muons
- Transverse ionization cooling achieved by MICE
- Muon emittance exchange demonstrated at FNAL/RAL
- 6D cooling of 5-6 orders needed

Noticeable reduction of 9% emittance



<https://arxiv.org/abs/1907.08562>, J.P. Delahauge et al., arXiv:1901.06150/

LEMMA: e^+e^- (at rest) $\rightarrow \mu^+\mu^-$ (at threshold)



Low EMittance Muon Accelerator
web.infn.it/LEMMA



Cooling is not a problem;
 but high luminosity is challenging:
 large e^+ flux of $O(10^{17}/s)$!

J.P. Delahauge et al., arXiv:1901.06150

New ideas ...

μ TRISTAN

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and Mitsuhiro Yoshida^{2,3}

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²*KEK Accelerator department, Tsukuba 305-0801, Japan*

³*Graduate University for Advanced Studies (Sokendai), Tsukuba 305-0801, Japan*

Abstract

The ultra-cold muon technology developed for the muon $g - 2$ experiment at J-PARC provides a low emittance μ^+ beam which can be accelerated and used for realistic collider experiments. We consider the possibility of new collider experi-

6664v2 [hep-ph] 21 Apr 2022

FRIDAY, 9 JUNE

00:00 → 01:30

Lunch

01:30 → 02:45

Plenary: Plenary 13

01:30

muTRISTAN

Speaker: Ryuichiro Kitano

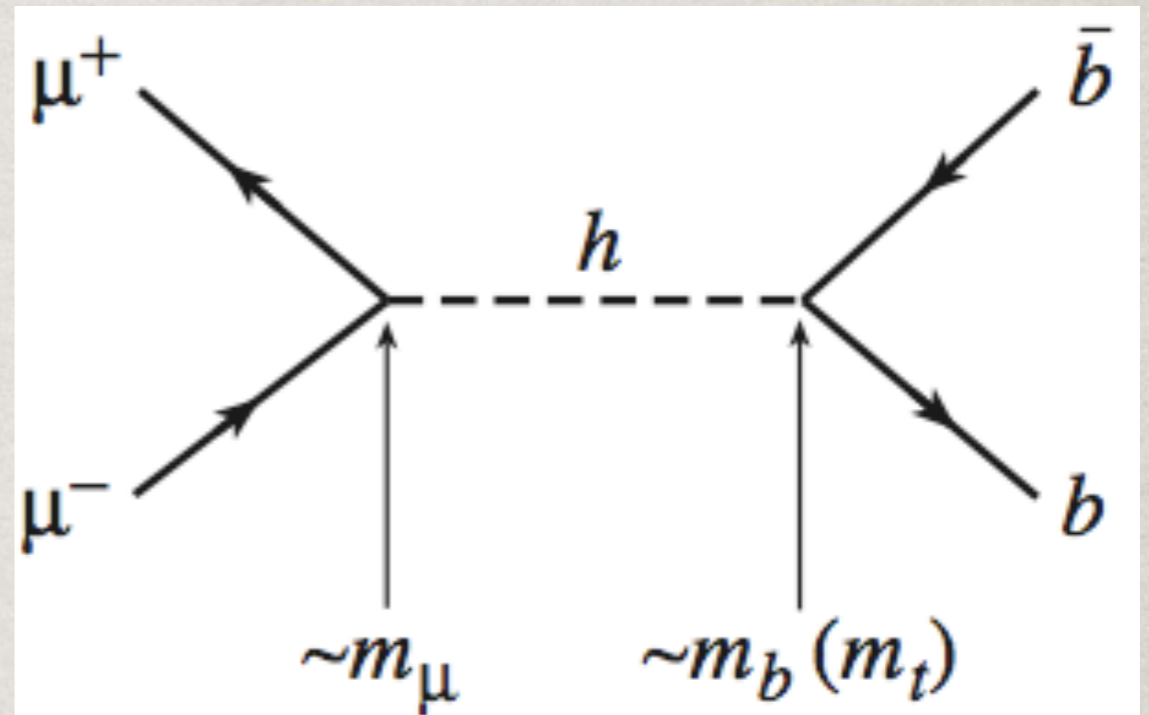
02:00

Probing new physics beyond the Standard Model at multi-TeV muon colliders

Speaker: Dr Adil Jueid (Institute for Basic Science)

PHYSICS POTENTIAL

Higgs factory: Resonant Production:



$$\sigma(\mu^+ \mu^- \rightarrow h \rightarrow X) = \frac{4\pi \Gamma_h^2 \text{Br}(h \rightarrow \mu^+ \mu^-) \text{Br}(h \rightarrow X)}{(\hat{s} - m_h^2)^2 + \Gamma_h^2 m_h^2}.$$

$$\begin{aligned} \sigma_{peak}(\mu^+ \mu^- \rightarrow h) &= \frac{4\pi}{m_h^2} \text{BR}(h \rightarrow \mu^+ \mu^-) \\ &\approx 71 \text{ pb at } m_h = 125 \text{ GeV.} \end{aligned}$$

About **O(70k)** events produced per **fb⁻¹**

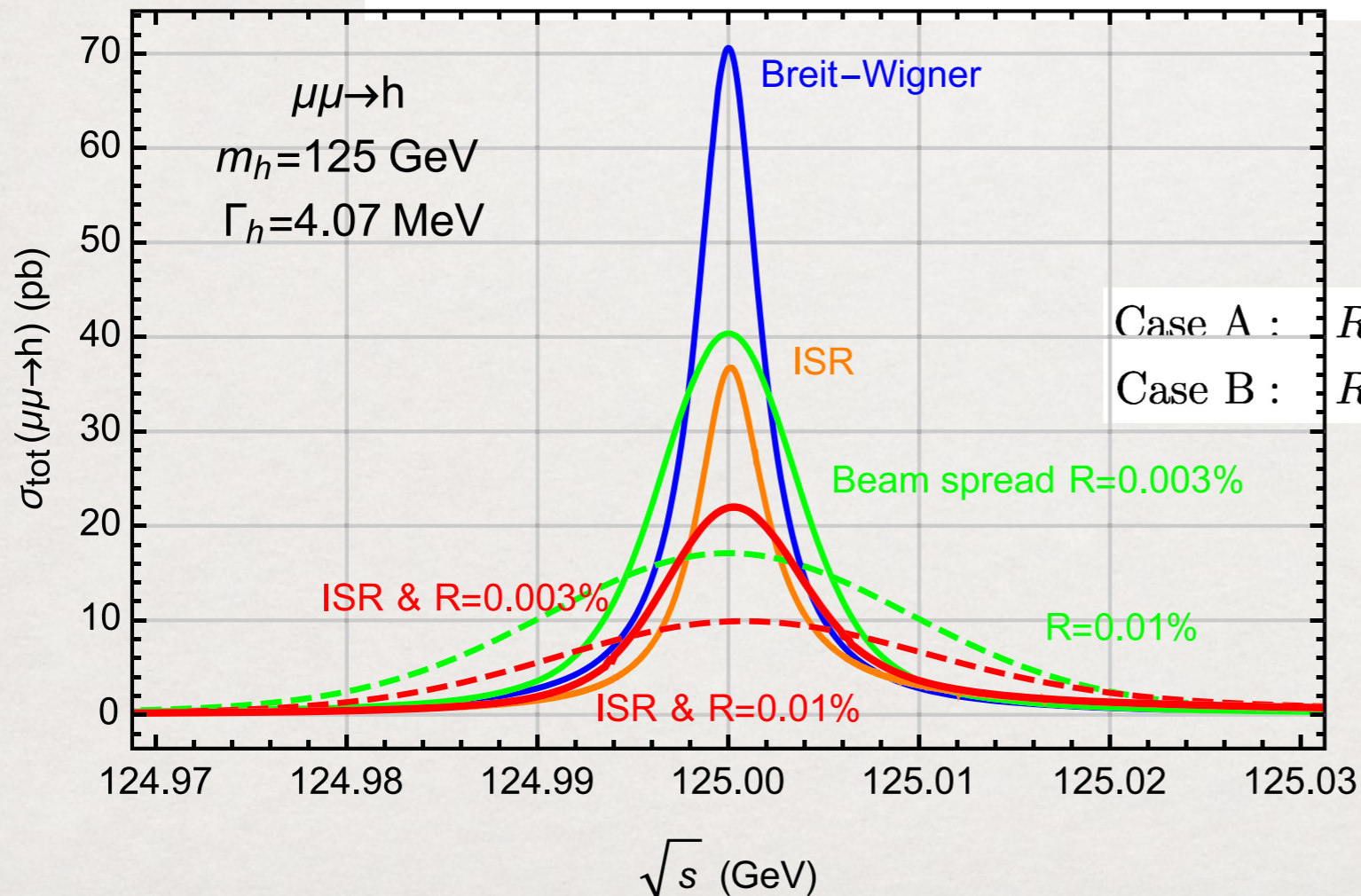
At $m_h=125$ GeV, $\Gamma_h=4.2$ MeV

$$\frac{\exp[-(\sqrt{\hat{s}} - \sqrt{s})^2/(2\sigma_{\sqrt{s}}^2)]}{\sqrt{2\pi}\sigma_{\sqrt{s}}}$$

$$\frac{4\pi\Gamma(h \rightarrow \mu\mu)\Gamma(h \rightarrow X)}{(\hat{s} - m_h^2)^2 + m_h^2[\Gamma_h^{\text{tot}}]^2}$$

$$\sigma_{\text{eff}}(s) = \int d\sqrt{\hat{s}} \frac{dL(\sqrt{s})}{d\sqrt{\hat{s}}} \sigma(\mu^+ \mu^- \rightarrow h \rightarrow X)$$

$$\propto \begin{cases} \Gamma_h^2 B / [(s - m_h^2)^2 + \Gamma_h^2 m_h^2] & (\Delta \ll \Gamma_h), \\ B \exp\left[-\frac{(m_h - \sqrt{s})^2}{2\Delta^2}\right] \left(\frac{\Gamma_h}{\Delta}\right) / m_h^2 & (\Delta \gg \Gamma_h). \end{cases}$$



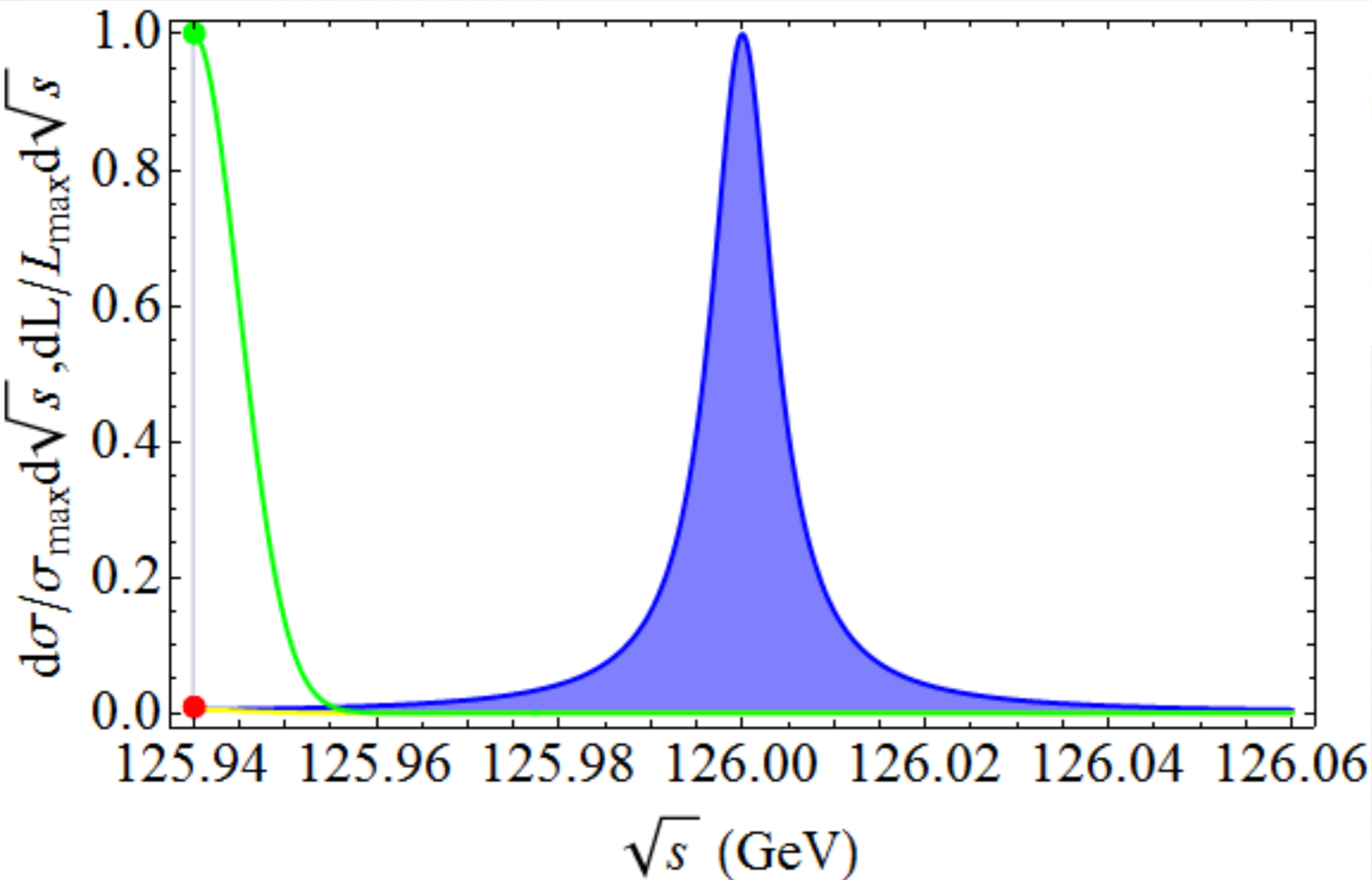
“Muon Collider Quartet”:
 Barger-Berger-Gunion-Han
 PRL & Phys. Report (1995)

Case A : $R = 0.01\%$ ($\Delta = 8.9$ MeV), $L = 0.5$ fb $^{-1}$,
 Case B : $R = 0.003\%$ ($\Delta = 2.7$ MeV), $L = 1$ fb $^{-1}$.

TH, Liu: 1210.7803;
 Greco, TH, Liu: 1607.03210

Ideal, conceivable case:

$$(\Delta = 5 \text{ MeV}, \quad \Gamma_h \approx 4.2 \text{ MeV})$$



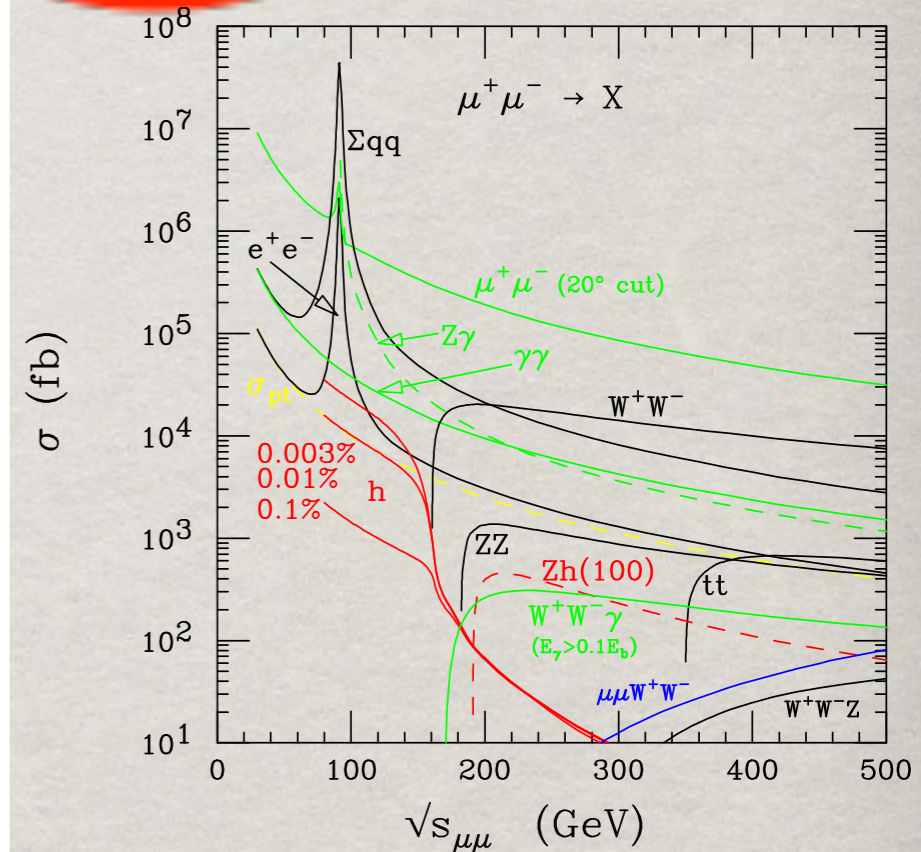
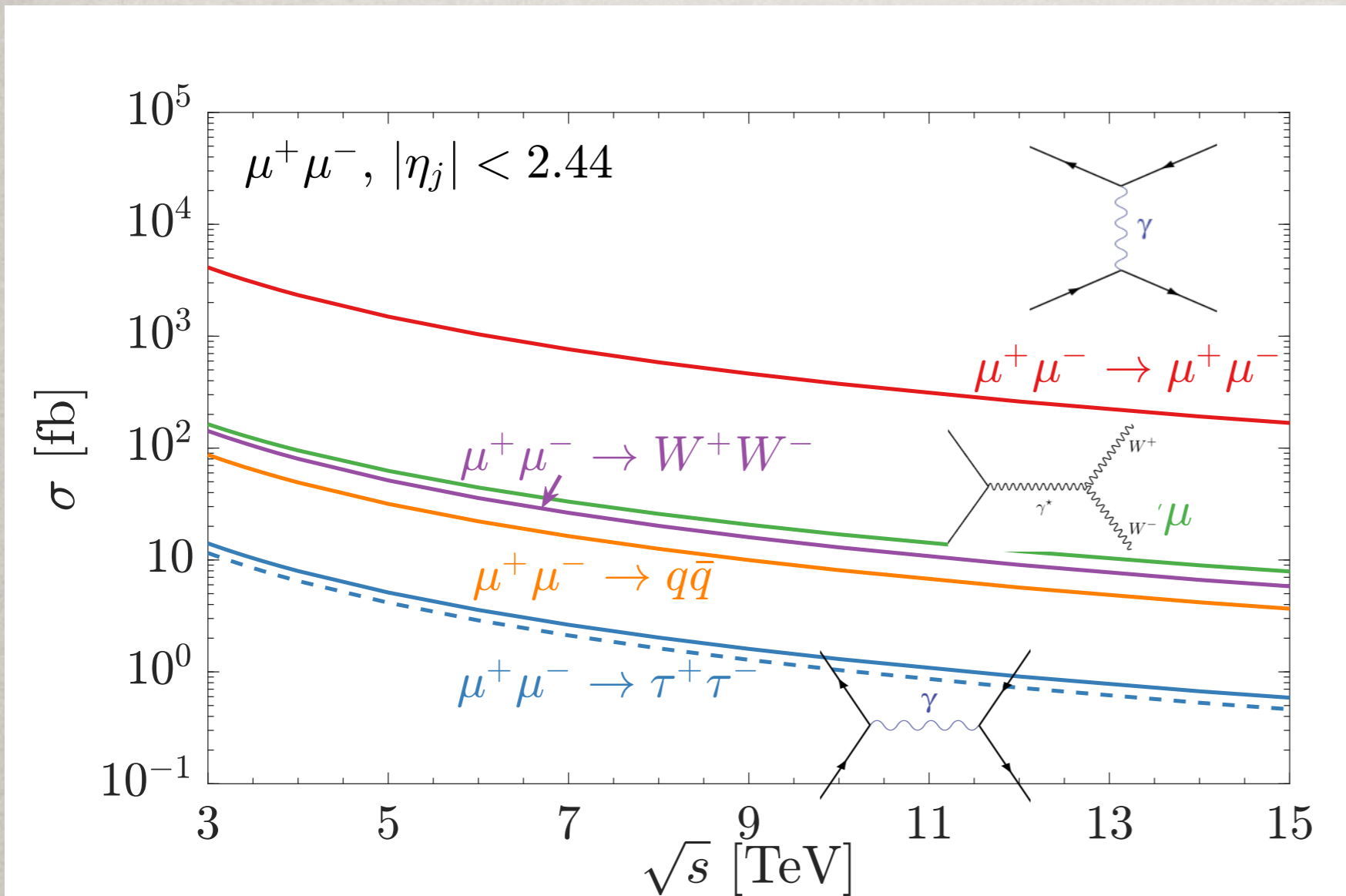
An optimal fitting could reach $\delta\Gamma_h \sim 0.15 \text{ MeV}$, or 3.5%

TH, Liu: 1210.7803; Greco, TH, Liu: 1607.03210

A Multi-TeV Muon Collider

Naïve expectation: leading-order $\mu^+\mu^-$ annihilation:

$$\sigma_{ann} \sim \frac{\alpha^2}{s}$$

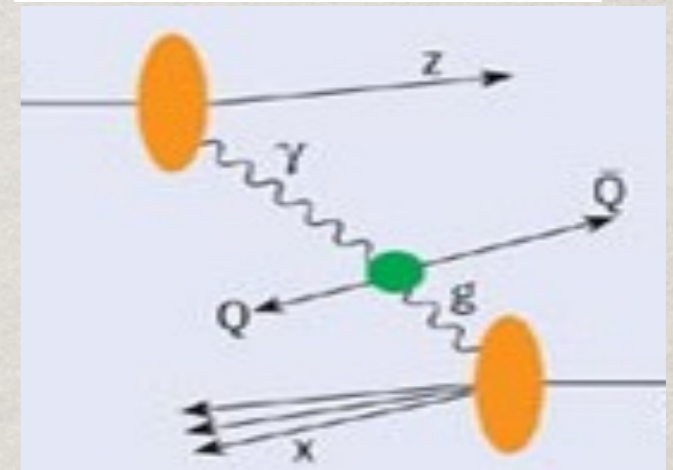
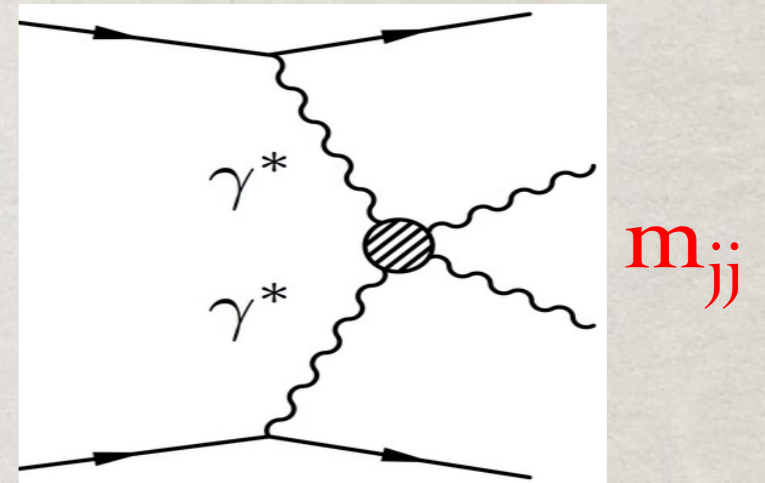
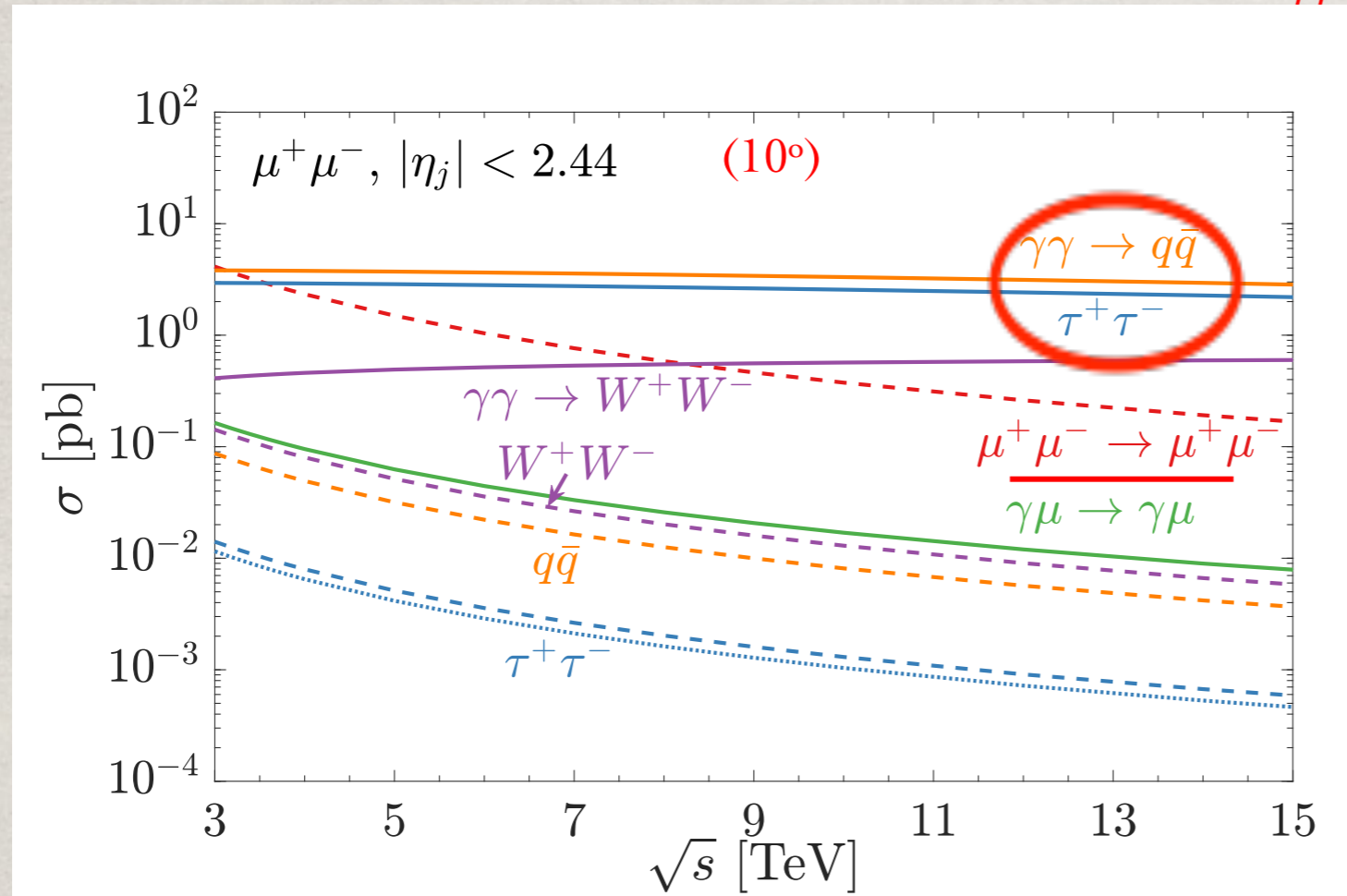


Those aren't what you would first see when you turned on the machine!

• Photon-induced QED cross sections

large rates

$$\sigma_{fusion} \sim \frac{\alpha^2}{m_{ij}^2} \log^2\left(\frac{Q^2}{m^2}\right)$$



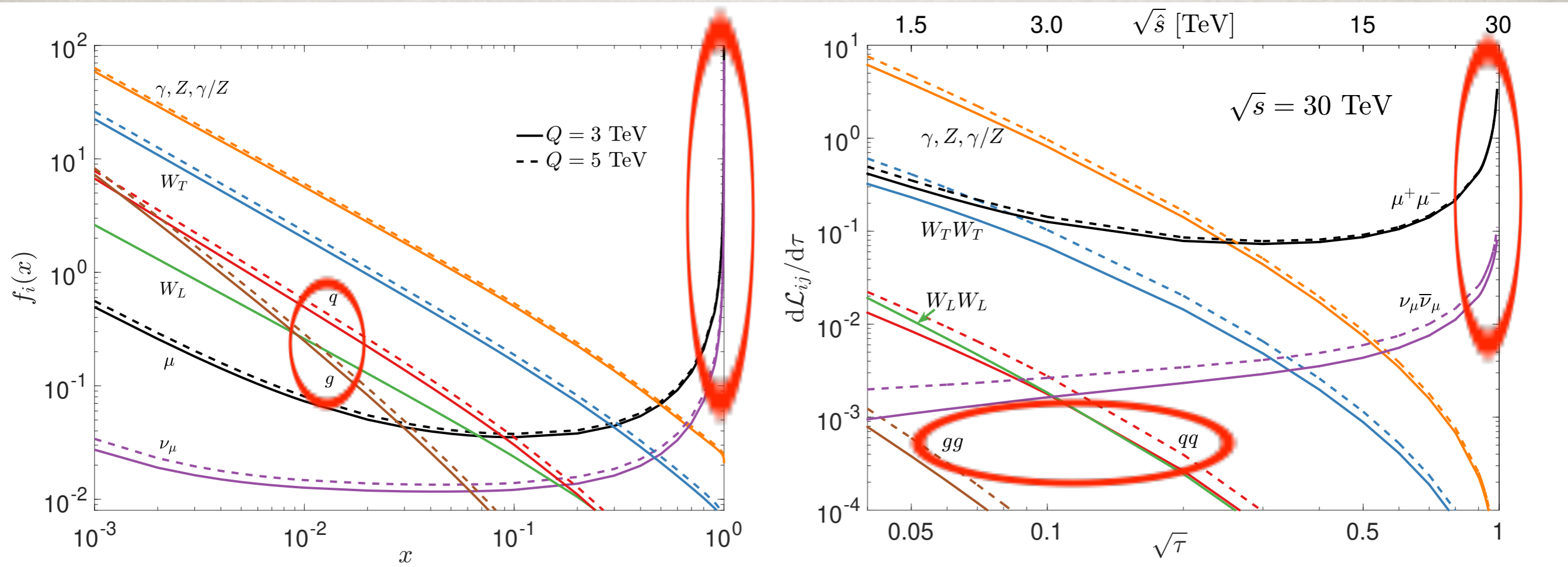
$$p_T^j > \left(4 + \frac{\sqrt{s}}{3 \text{ TeV}}\right) \text{ GeV}, \quad m_{ij} > 20 \text{ GeV}, \quad |\eta_j| < 3.13 \quad (2.44)$$

Quarks/gluons come into the picture via SM DGLAP:

$$\frac{d}{d \log Q^2} \begin{pmatrix} f_L \\ f_U \\ f_D \\ f_\gamma \\ f_g \end{pmatrix} = \begin{pmatrix} P_{ll} & 0 & 0 & 2N_l P_{l\gamma} & 0 \\ 0 & P_{uu} & 0 & 2N_u P_{u\gamma} & 2N_u P_{ug} \\ 0 & 0 & P_{dd} & 2N_d P_{d\gamma} & 2N_d P_{dg} \\ P_{\gamma l} & P_{\gamma u} & P_{\gamma d} & P_{\gamma\gamma} & 0 \\ 0 & P_{gu} & P_{gd} & 0 & P_{gg} \end{pmatrix} \otimes \begin{pmatrix} f_L \\ f_U \\ f_D \\ f_\gamma \\ f_g \end{pmatrix}$$

• **EW PDFs at a muon collider:**

“partons” dynamically generated $\frac{df_i}{d \ln Q^2} = \sum_I \frac{\alpha_I}{2\pi} \sum_j P_{i,j}^I \otimes f_j$



μ^\pm : the valance. ℓ_R, ℓ_L, ν_L and $B, W^{\pm,3}$: LO sea.

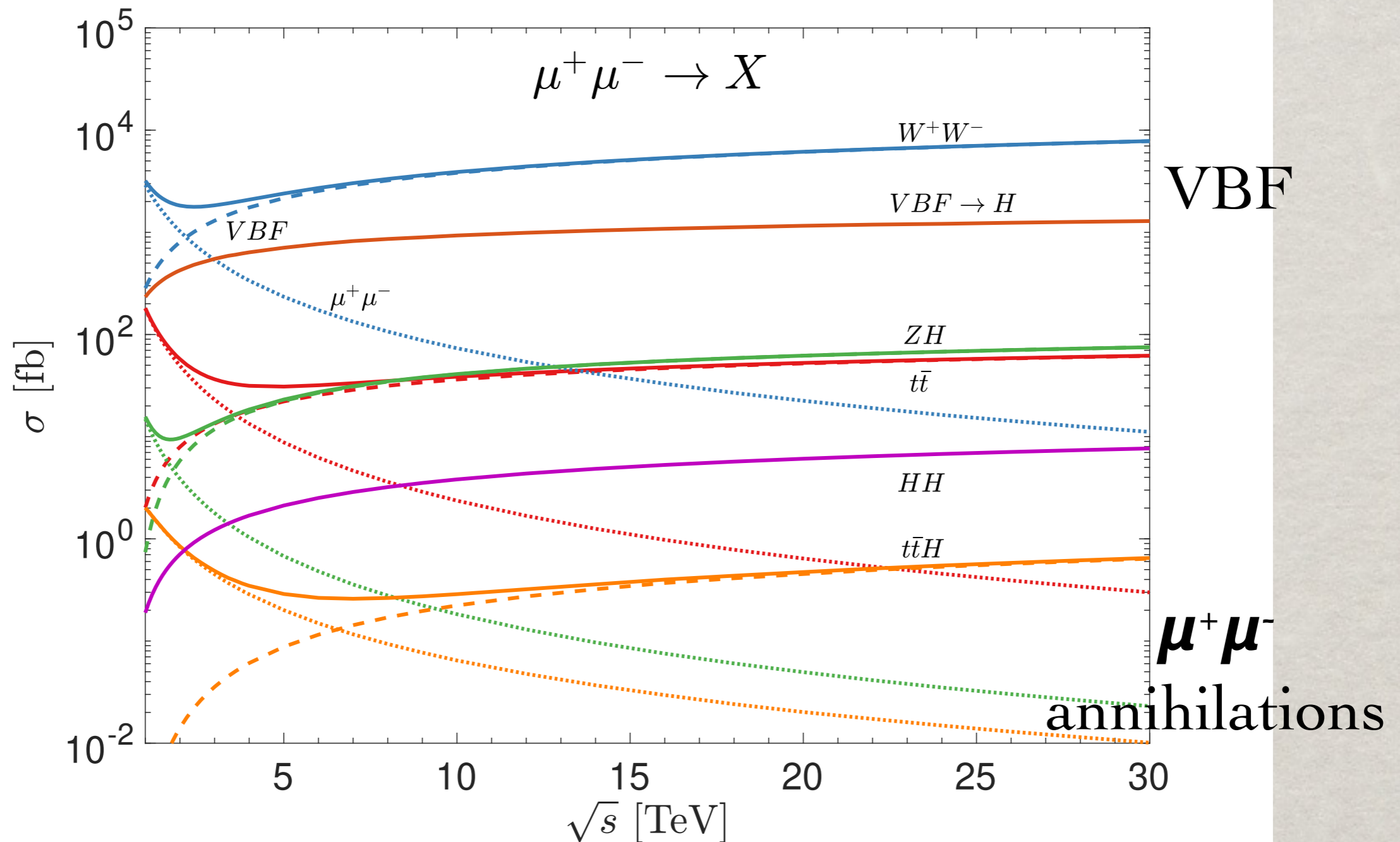
Quarks: NLO; gluons: NNLO.

TH, Yang Ma, Keping Xie, arXiv:2007.14300

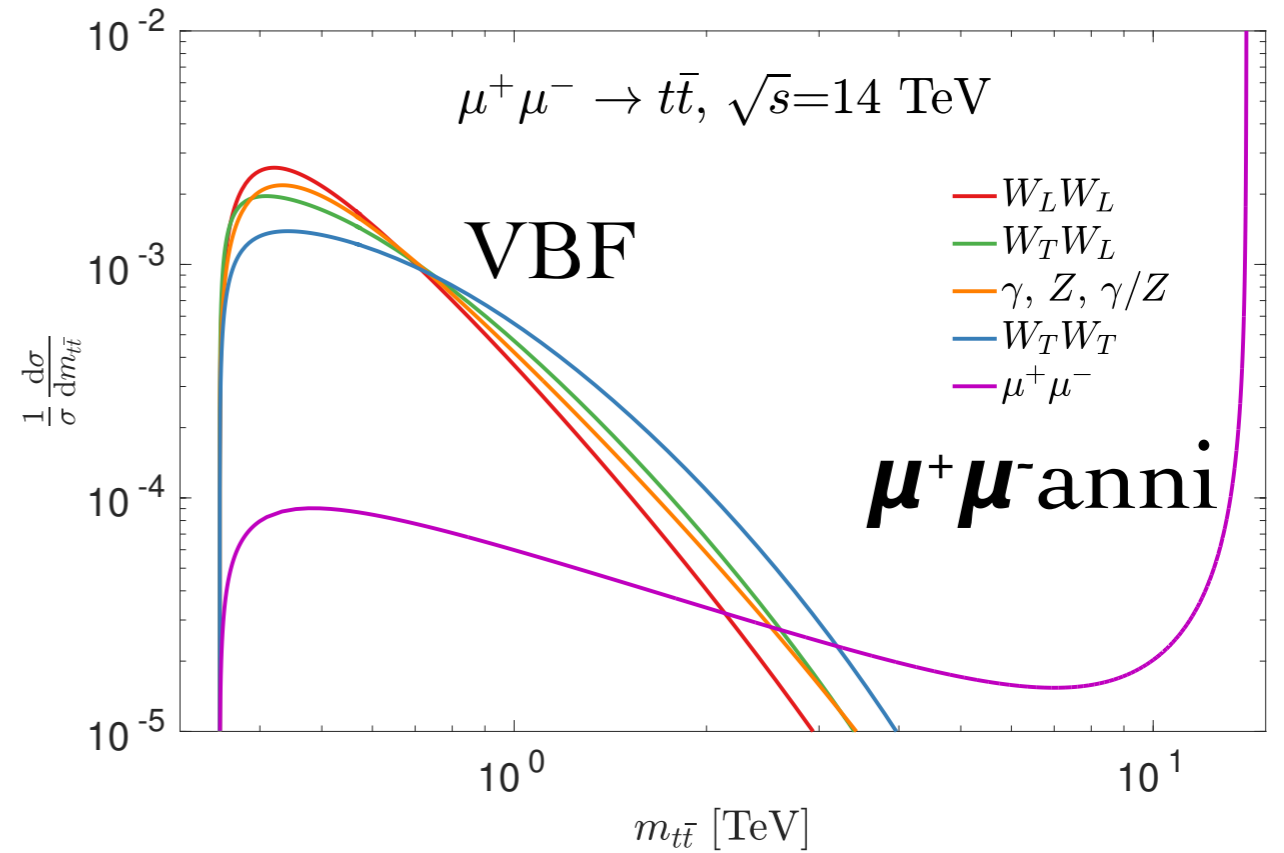
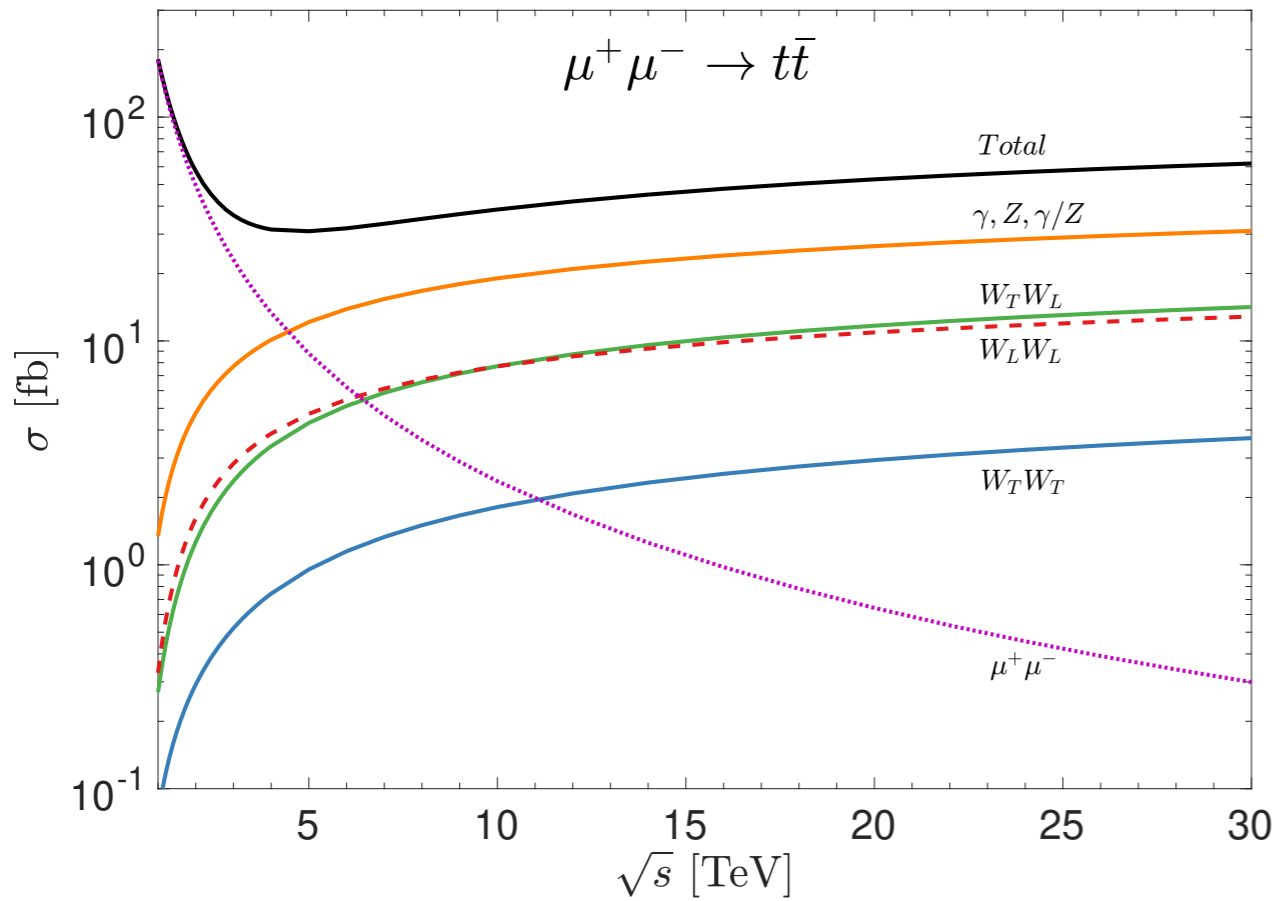
- “Semi-inclusive” processes

Just like in hadronic collisions:

$\mu^+ \mu^- \rightarrow$ exclusive particles + remnants

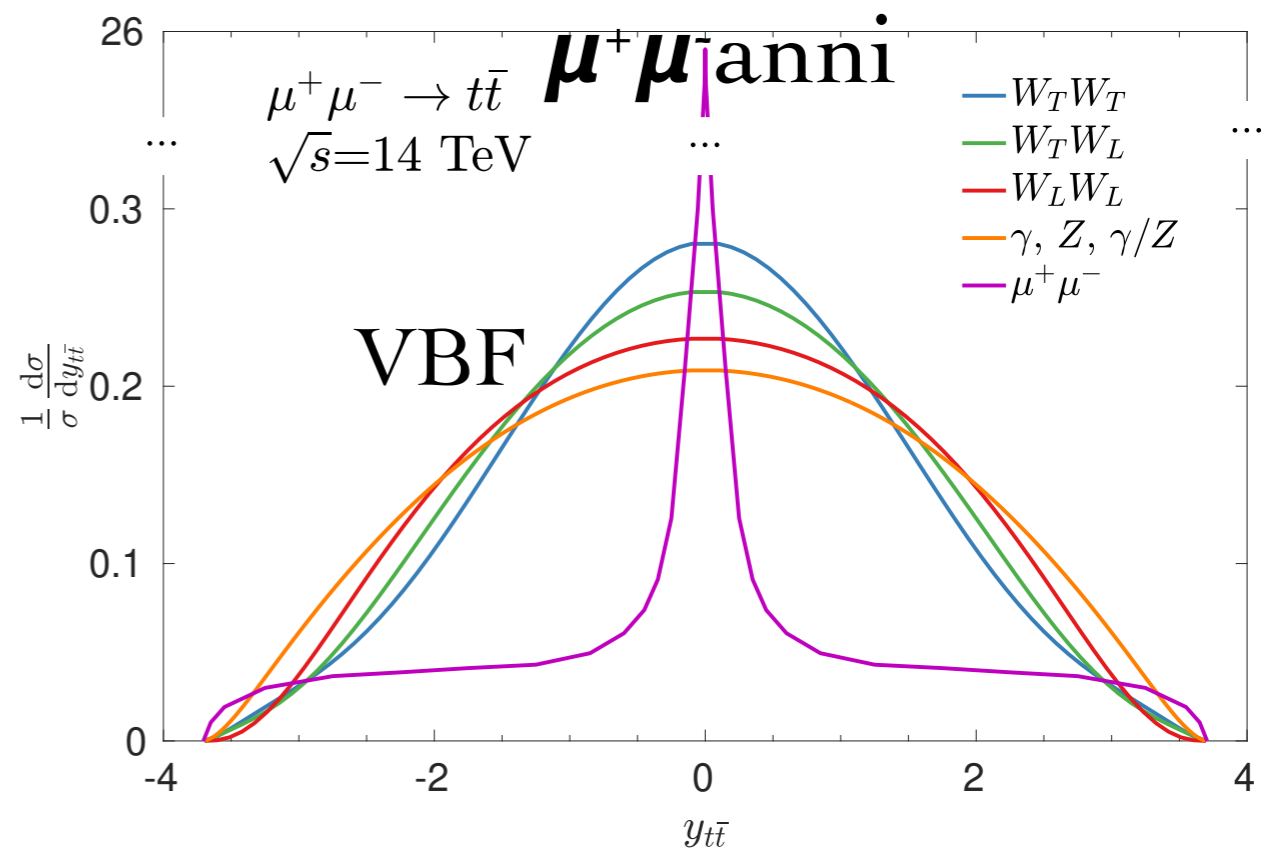


- separable sub-processes:



Partonic contributions

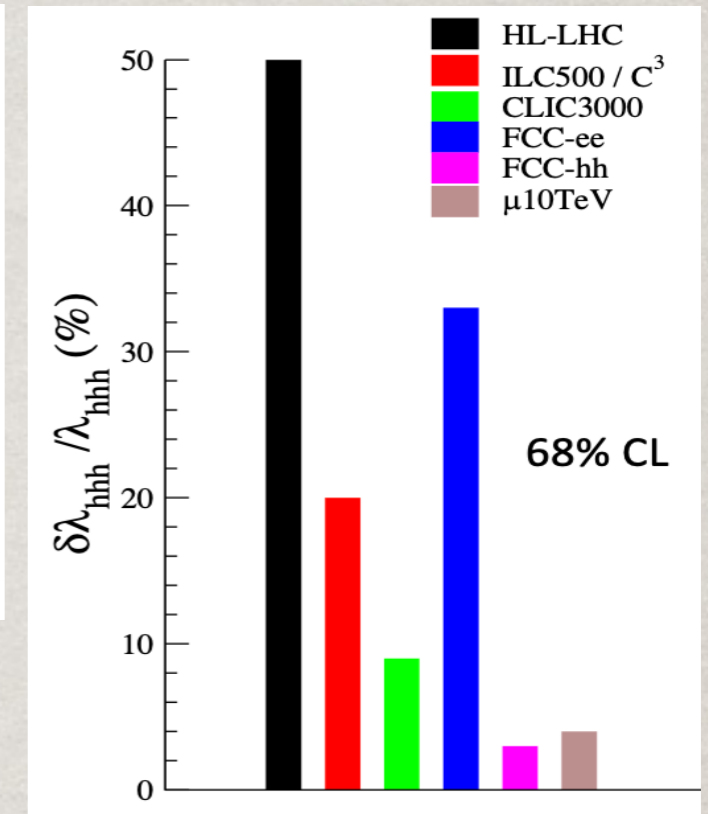
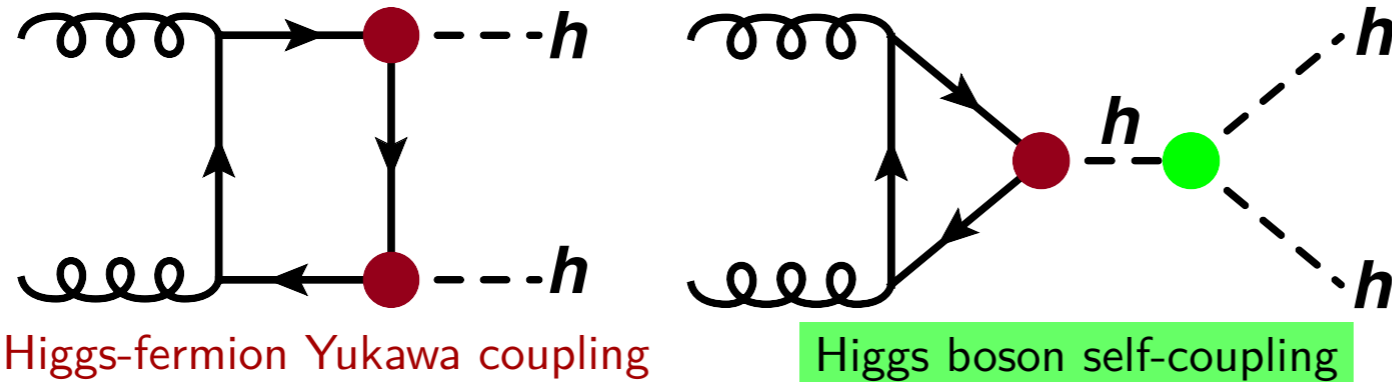
$\mu^+ \mu^-$ Collider:
“Buy one, get one free”
Annihilation + VBF



Higgs pair production & triple coupling:

SM Higgs boson pair production at the LHC

SM Higgs boson pair production (gluon-gluon fusion - ggF):



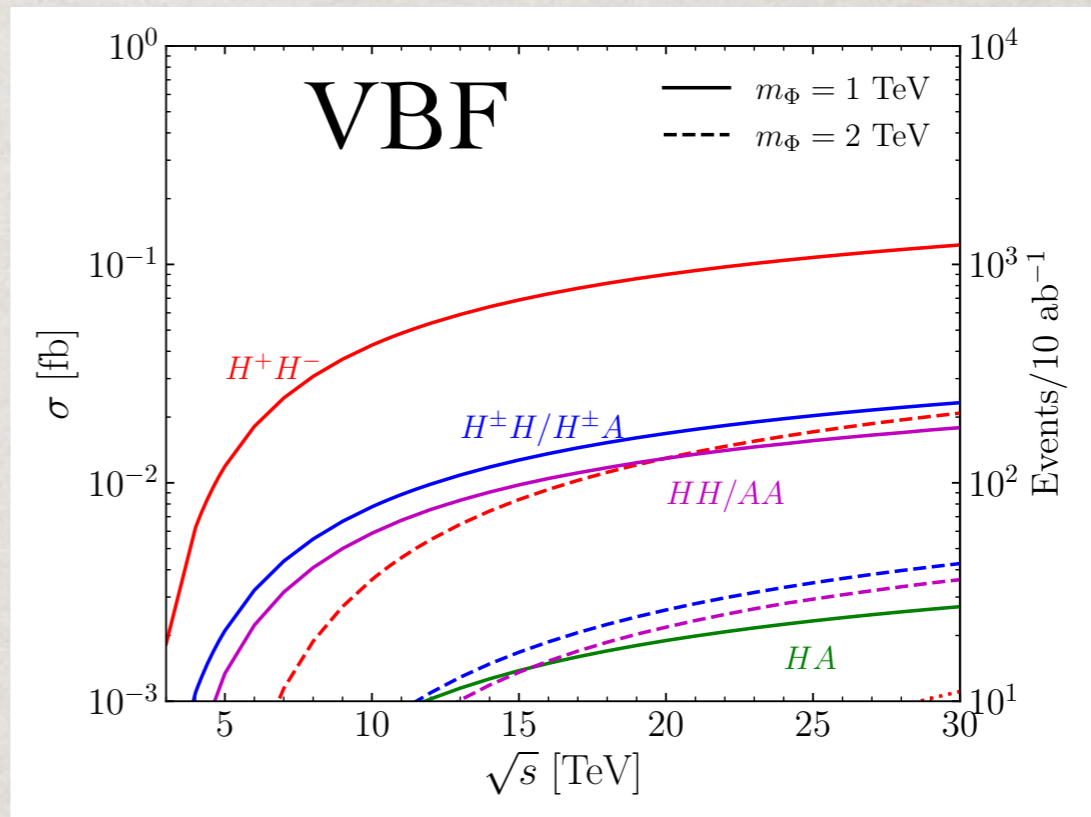
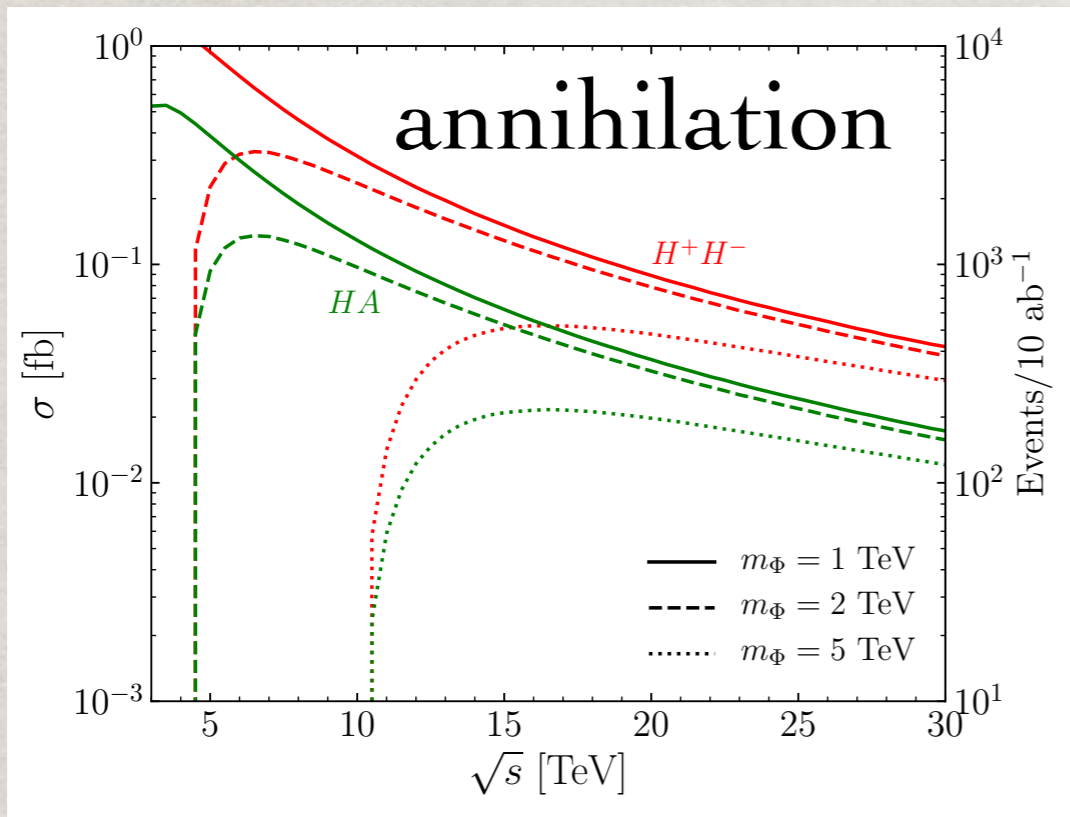
→ dictate EW phase transition & impact on early universe cosmology!

\sqrt{s} (lumi.)	3 TeV (1 ab ⁻¹)	6 (4)	10 (10)	14 (20)	30 (90)	Comparison
WWH ($\Delta\kappa_W$)	0.26%	0.12%	0.073%	0.050%	0.023%	0.1% [41]
$\Lambda/\sqrt{c_i}$ (TeV)	4.7	7.0	9.0	11	16	(68% C.L.)
ZZH ($\Delta\kappa_Z$)	1.4%	0.89%	0.61%	0.46%	0.21%	0.13% [17]
$\Lambda/\sqrt{c_i}$ (TeV)	2.1	2.6	3.2	3.6	5.3	(95% C.L.)
$WWHH$ ($\Delta\kappa_{W_2}$)	5.3%	1.3%	0.62%	0.41%	0.20%	5% [36]
$\Lambda/\sqrt{c_i}$ (TeV)	1.1	2.1	3.1	3.8	5.5	(68% C.L.)
HHH ($\Delta\kappa_3$)	25%	10%	5.6%	3.9%	2.0%	5% [22, 23]
$\Lambda/\sqrt{c_i}$ (TeV)	0.49	0.77	1.0	1.2	1.7	(68% C.L.)

Table 7: Summary table of the expected accuracies at 95% C.L. for the Higgs couplings at a variety of muon collider collider energies and luminosities.

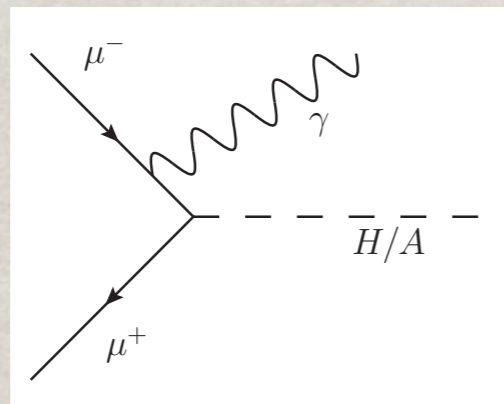
TH, D. Liu, I. Low, X. Wang, arXiv:2008.12204

• Heavy Higgs Bosons Production

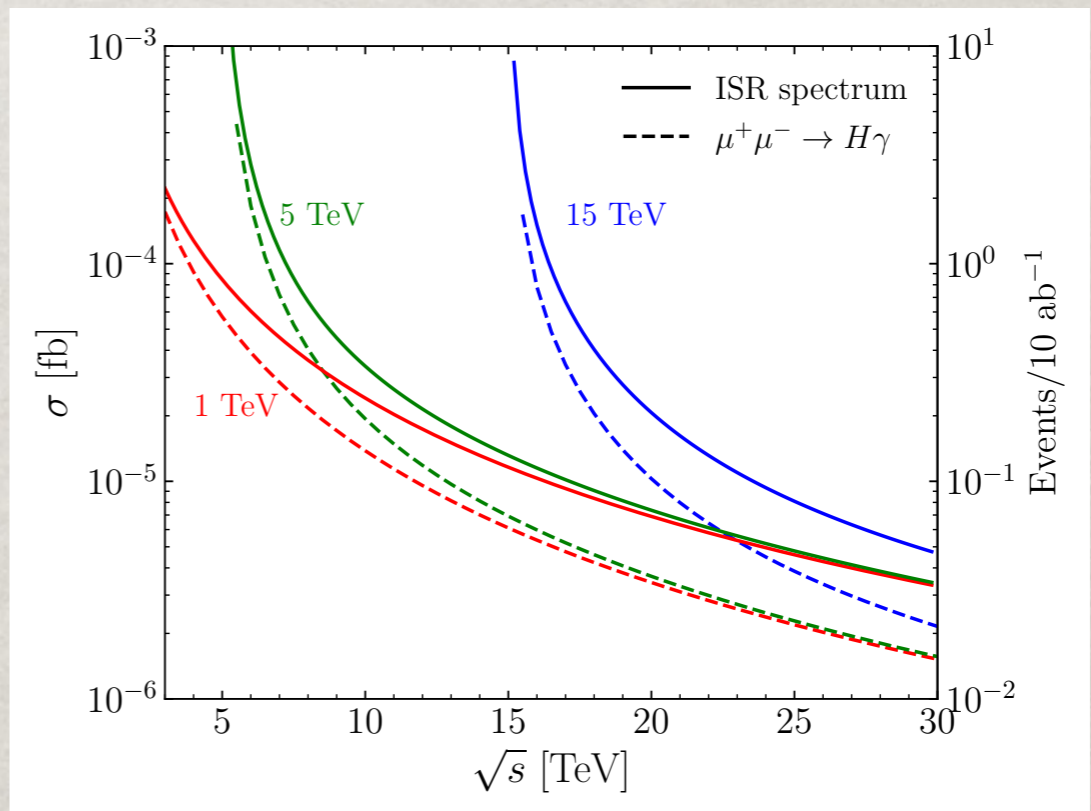


Radiative returns:

$$\hat{\sigma}(\mu^+\mu^- \rightarrow H) = \frac{\pi Y_\mu^2}{4} \delta(\hat{s} - m_H^2) = \frac{\pi Y_\mu^2}{4s} \delta(\tau - \frac{m_H^2}{s})$$



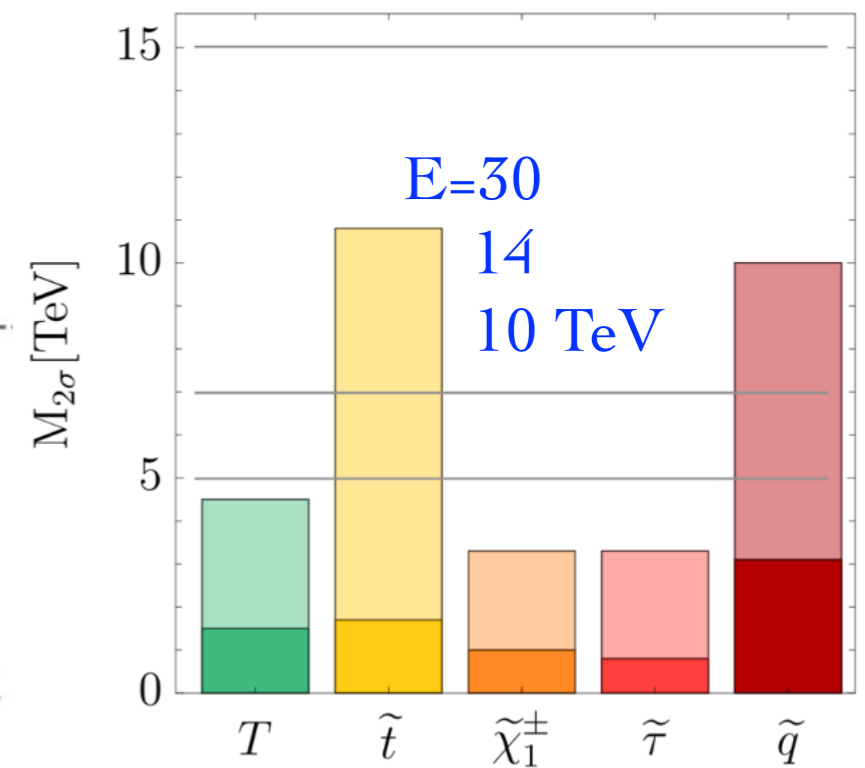
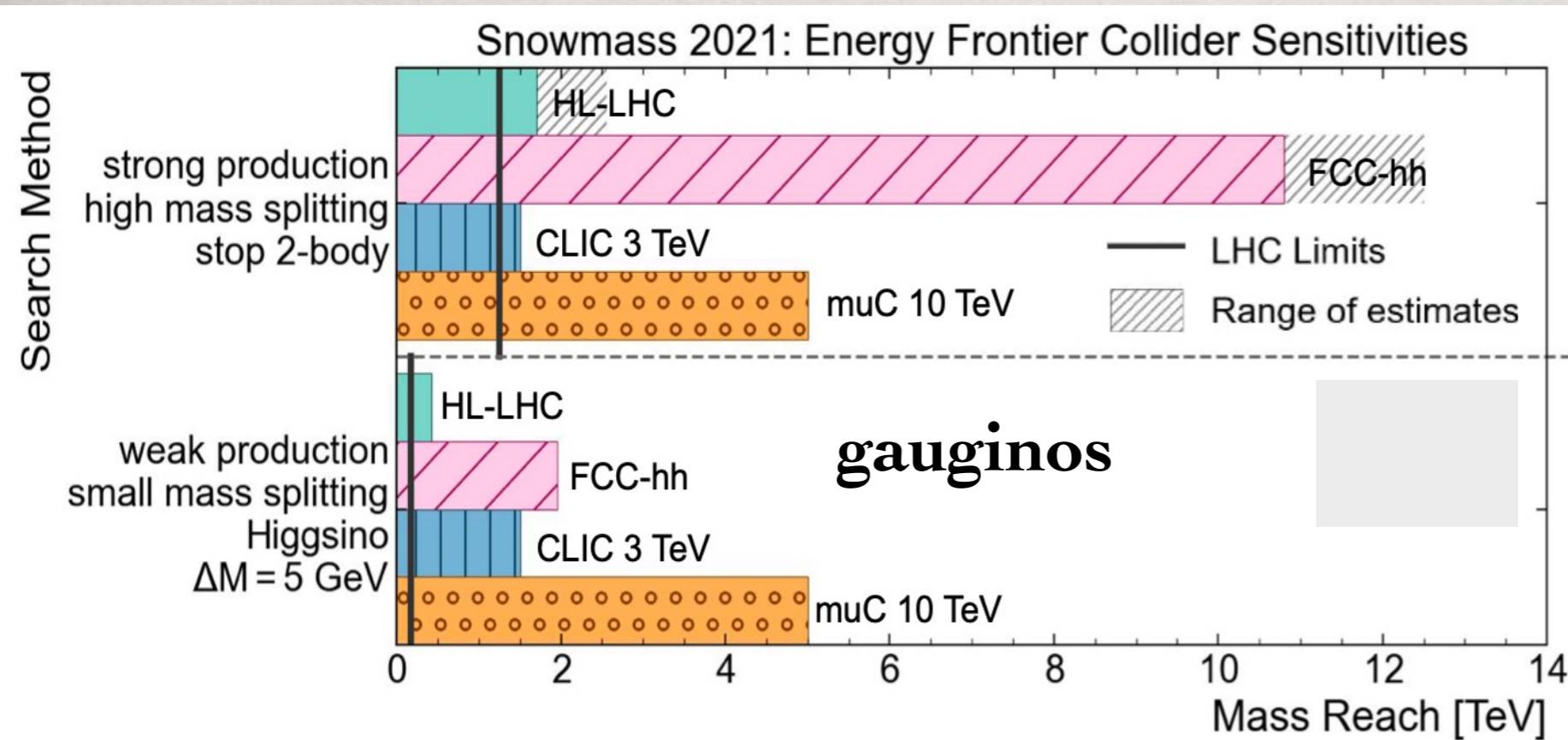
Reach $M \sim E_{cm}$!



TH, S. Li, S. Su, W. Su, Y. Wu, arXiv:2102.08386.

Pushing the “Naturalness” limit

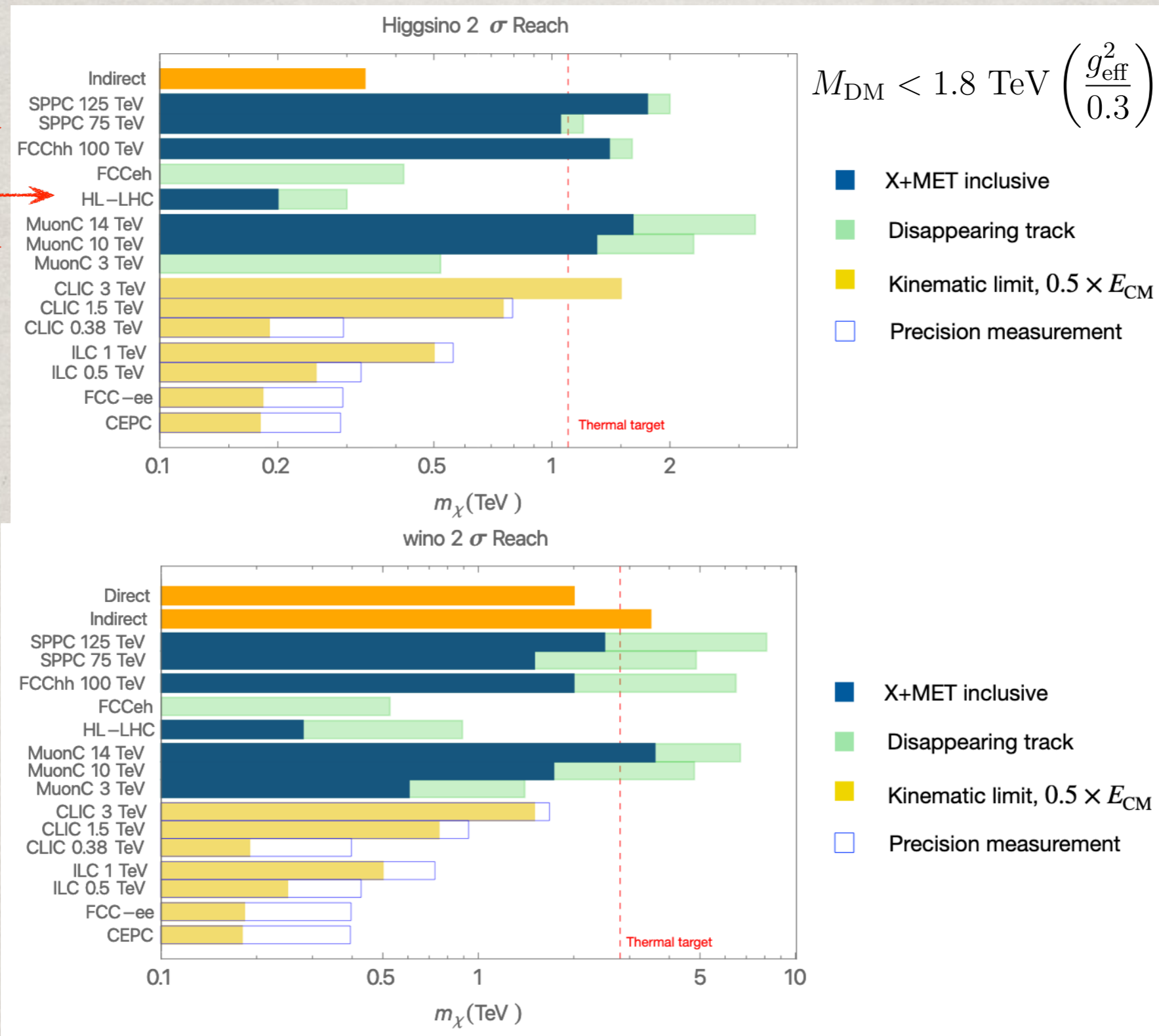
The searches for top quark partners
(most wanted in “naturalness”);
& gluinos, gauginos ...



→ Higgs mass fine-tune: $\delta m_H/m_H \sim 1\% (1 \text{ TeV}/\Lambda)^2$
Thus, $m_{\text{stop}} > 8 \text{ TeV} \rightarrow 10^{-4}$ fine-tune!

WIMP Dark Matter

Covering the thermal target



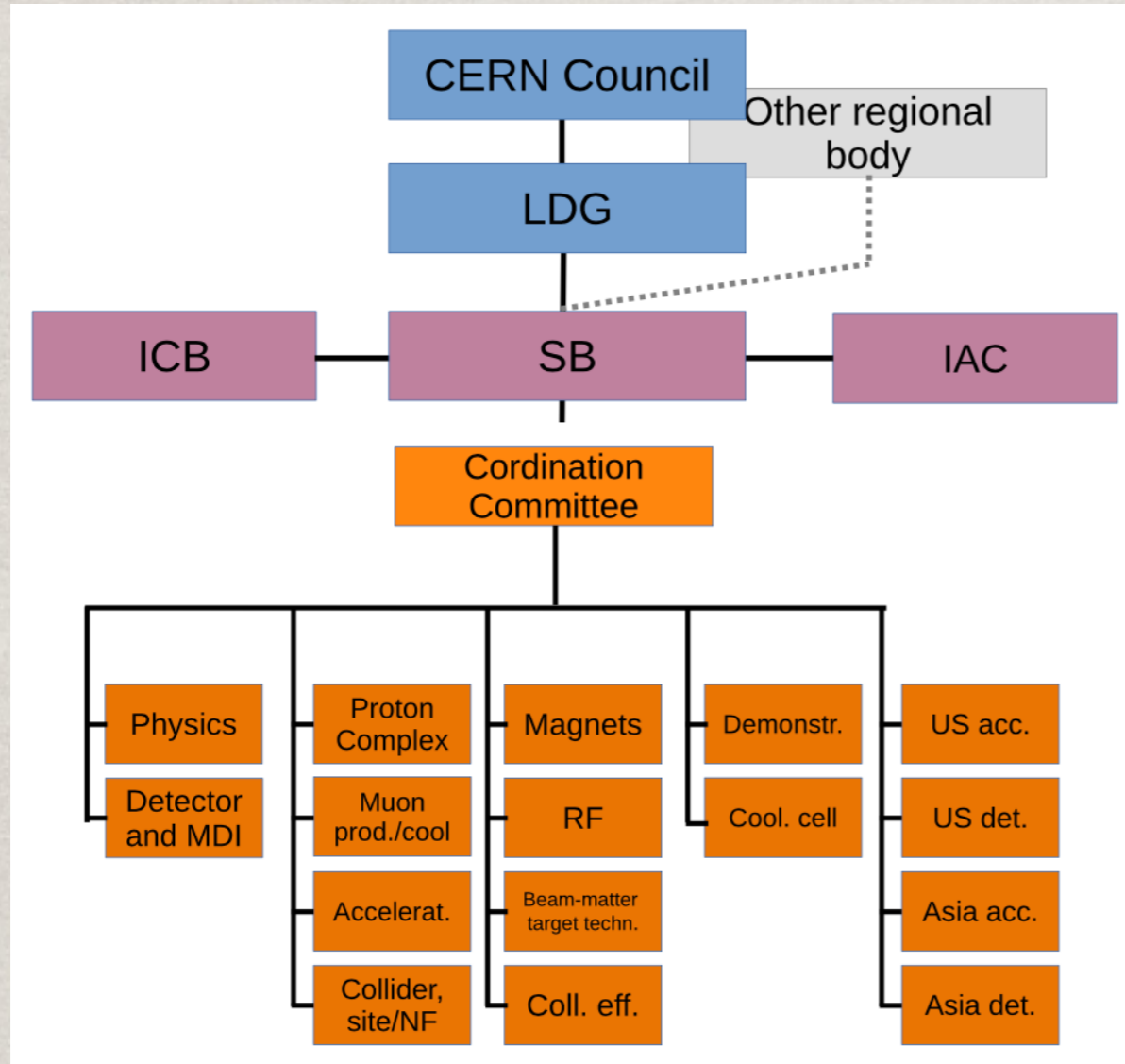
TH, Z. Liu, L.T. Wang, X. Wang: arXiv:2009.11287; arXiv:2203.07351



International
Muon Collider
Collaboration

International Muon Collider Collaboration (IMCC, 2022)

European Commission grant: HORIZON INFRA-DEV



Coordination Committee

- Alexej Grudiev
- Andrea Wulzer
- Antoine Chance
- Anton Lechner
- Chris Rogers
- Christian Carli
- Claude Marchand
- Daniel Schulte (Study Leader)
- Donatella Lucchesi
- Elias Metral
- Jingyu Tang
- Luca Bottura
- Lucio Rossi
- Mark Palmer
- Nadia Pastrone (Collaboration Board Chair)
- Natalia Milas
- Roberto Losito
- Sergo Jindariani
- Steinar Stapnes (Steering Board Chair)

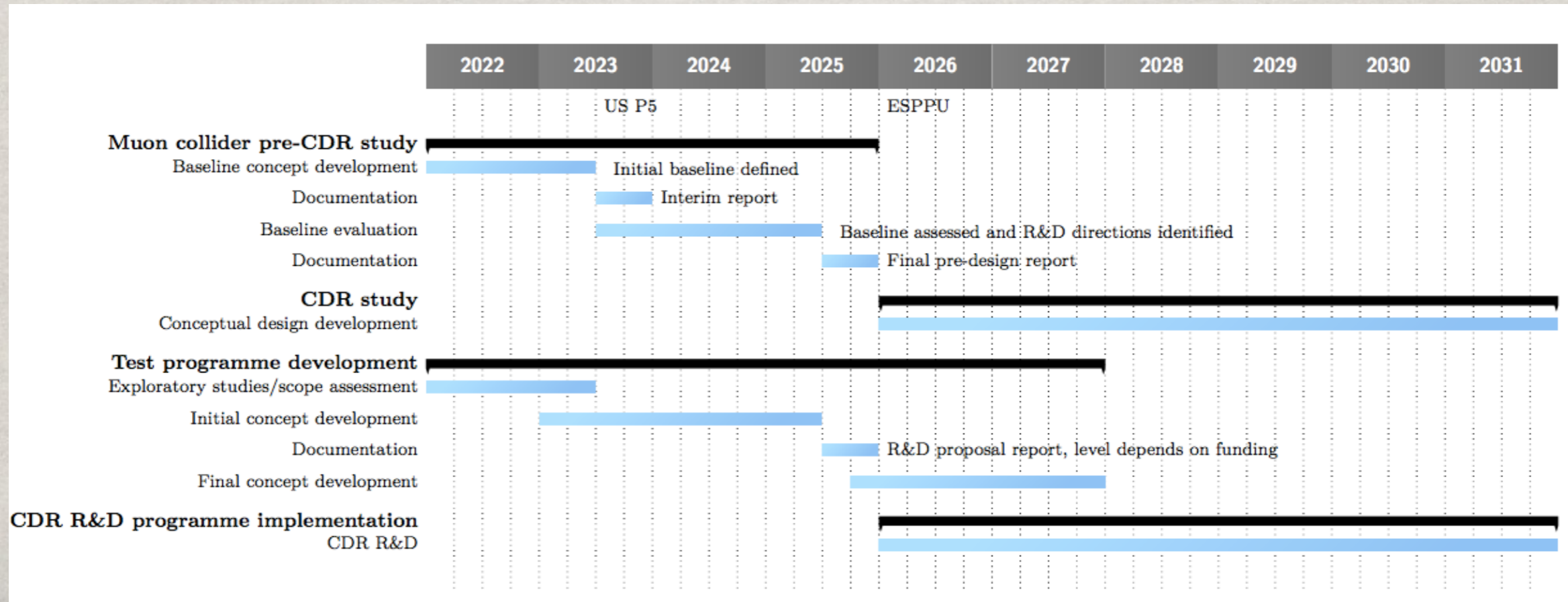
Physics	Andrea Wulzer
Detector and MDI	Donatella Lucchesi

US (detector)	Sergo Jindariani
US (accelerator)	Mark Palmer
Asia (China)	Jingyu Tang

Initial Target Parameters

Parameter	Unit	3 TeV	10 TeV	14 TeV
L	$10^{34} \text{ cm}^{-2}\text{s}^{-1}$	1.8	20	40
N	10^{12}	2.2	1.8	1.8
f_r	Hz	5	5	5
P_{beam}	MW	5.3	14.4	20
C	km	4.5	10	14
$\langle B \rangle$	T	7	10.5	10.5
ϵ_L	MeV m	7.5	7.5	7.5
σ_E / E	%	0.1	0.1	0.1
σ_z	mm	5	1.5	1.07
β	mm	5	1.5	1.07
ϵ	μm	25	25	25
$\sigma_{x,y}$	μm	3.0	0.9	0.63

CLIC at 3 TeV
2 (6)
28



Mark Palmer @ 1st IMCC Conference
CERN, Oct. 11, 2022



Conclusion

The MC option was well-received as a potential Energy Frontier option during Snowmass

- Implementing a US Accelerator R&D Program for Future Colliders would provide a mechanism to:
 - Let US accelerator experts engage with ongoing MC development activities
 - Formally participate in the IMCC
 - Explore potential options for US siting
- Developing critical elements of the design between now and the next Snowmass (and European Strategy Update) is crucial!

Exciting journey ahead!

Lots of recent works!

- D. Buttazzo, D. Redogolo, F. Sala, arXiv:1807.04743 (VBF to Higgs)
A. Costantini, F. Maltoni, et al., arXiv:2005.10289 (VBF to NP)
M. Chiesa, F. Maltoni, L. Mantani, B. Mele, F. Piccinini, and X. Zhao,
arXiv:2005.10289 (SM Higgs)
R. Capdevilla, D. Curtin, Y. Kahn, G. Krnjaic,
arXiv:2006.16277; arXiv:2101.10334 (g-2, flavor)
P. Bandyopadhyay, A. Costantini et al., arXiv:2010.02597 (Higgs)
D. Buttazzo, P. Paradisi, arXiv:2012.02769 (g-2)
W. Yin, M. Yamaguchi, arXiv:2012.03928 (g-2)
R. Capdevilla, F. Meloni, R. Simoniello, and J. Zurita, arXiv:2012.11292 (MD)
D. Buttazzo, F. Franceschini, A. Wulzer, arXiv:2012.11555 (general)
G.-Y. Huang, F. Queiroz, W. Rodejohann,
arXiv:2101.04956; arXiv:2103.01617 (flavor)
W. Liu, K.-P. Xie, arXiv:2101.10469 (EWPT)
Richard Ruiz et al., arXiv:2111.02442 (MadGraph5)

.....
Numerous Snowmass White papers & summary reports

Muon Smasher's Guide: H. Ali, N. Arkani-Hamed, et al, arXiv:2103.14043

Muon Collider Physics Summary: <https://arxiv.org/abs/2203.07256>

Muon Collider Forum Report: <https://arxiv.org/abs/2209.01318>