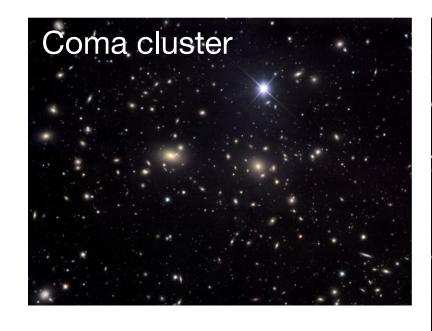
# Dark Matter Direct Detection Experiments & Machine Learning

**Aaron Higuera** 

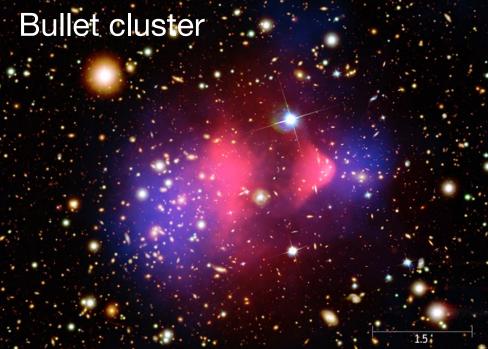


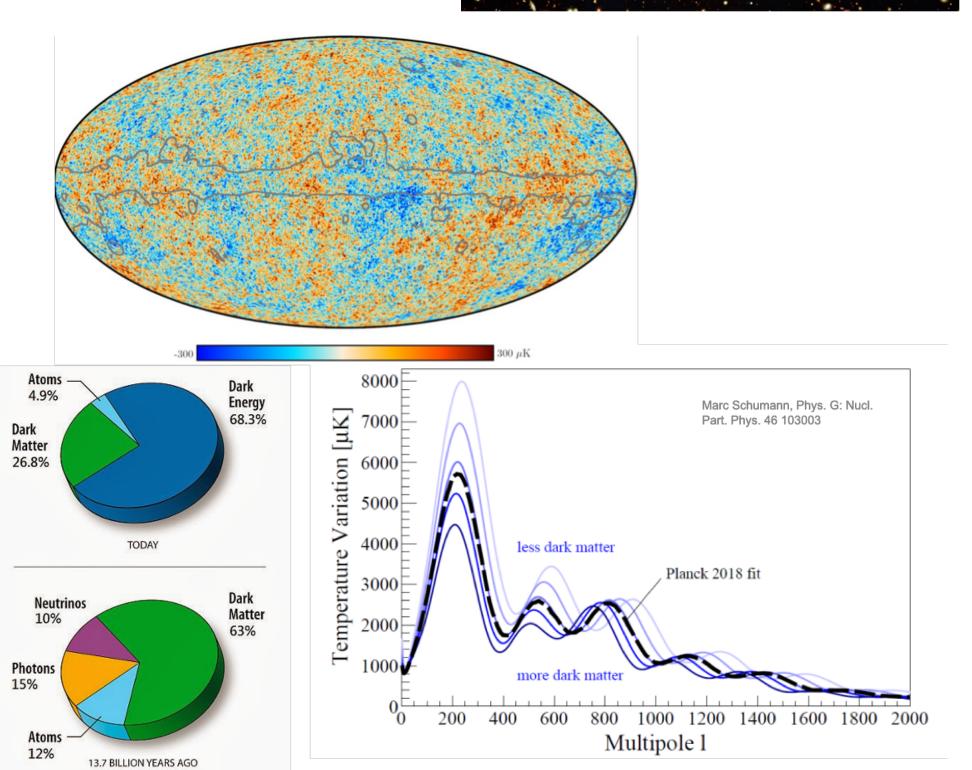
### Dark Matter

- Decades of astronomical evidence
  - From galaxies and clusters
    - Rotation curves
    - Cluster movement, collisions and gravitational lensing
  - Cosmic scale
    - Structure of the universe
    - Cosmic microwave background
- What do we know about dark matter?
  - Electrically neutral
  - Weakly interacting (weaker than the SM weak interaction)
  - Either long-lived/stabled or produced
  - Should be among us



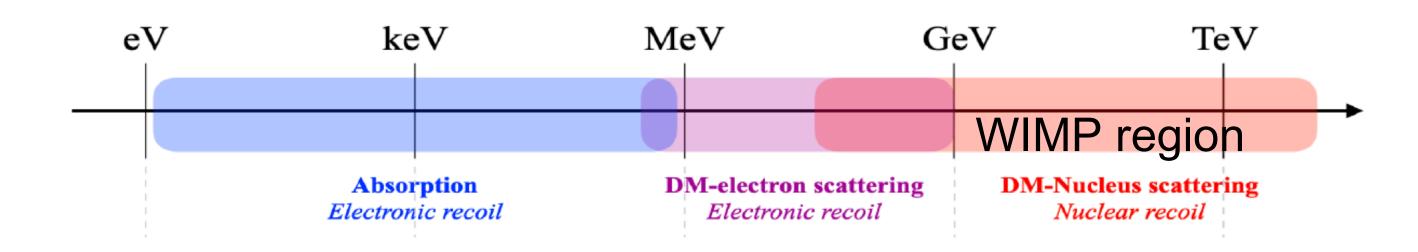
(Universe 380,000 years old)

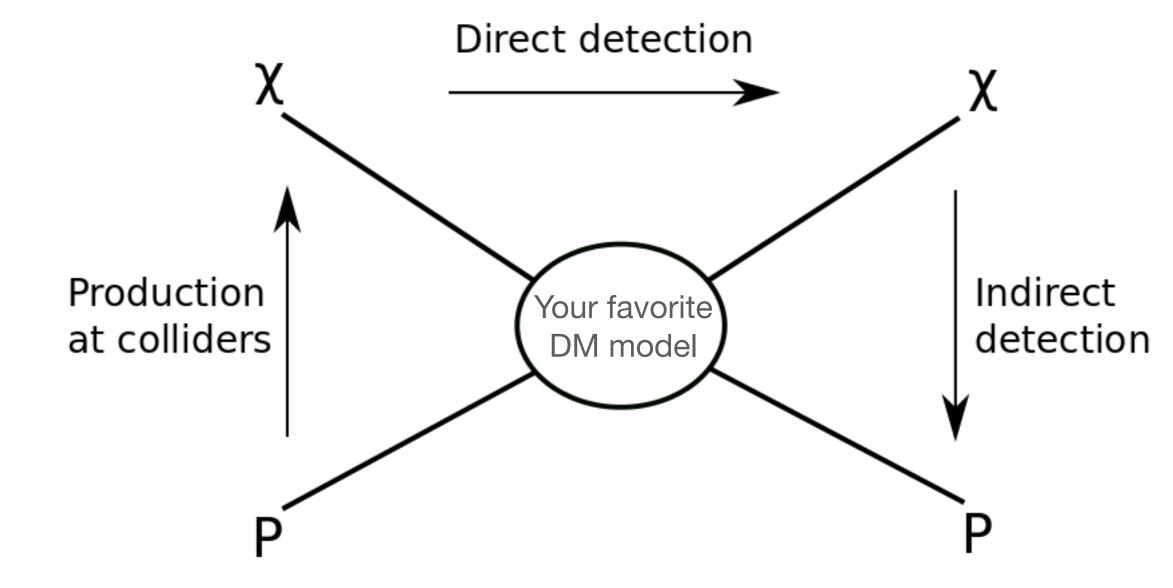


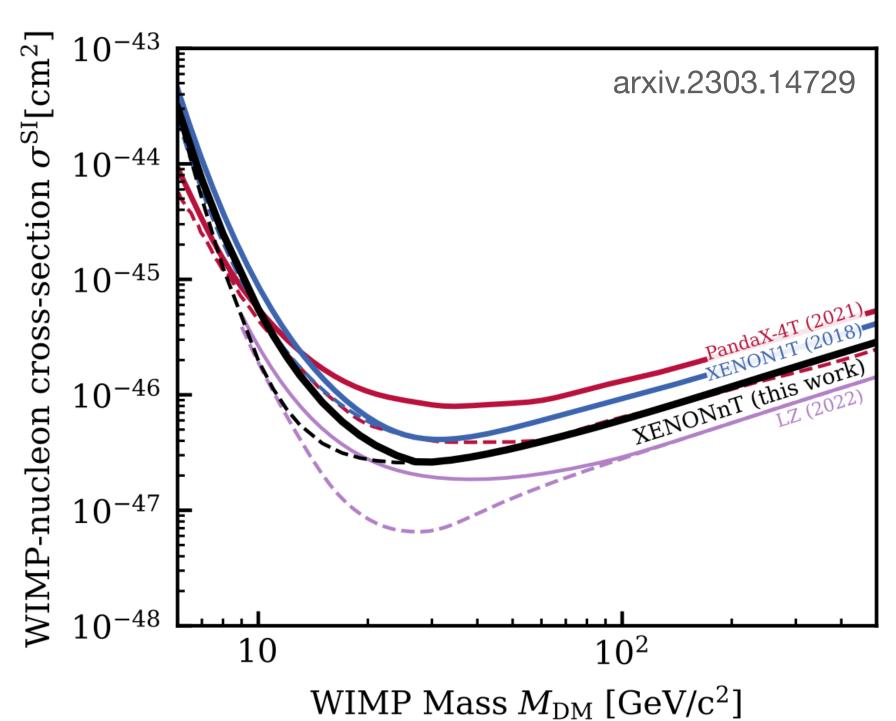


### Dark Matter

- How do we search for dark matter?
  - In direct detection
    - Dark matter particle-nucleus scattering
    - Ultra-low background experiments
    - Low thresholds

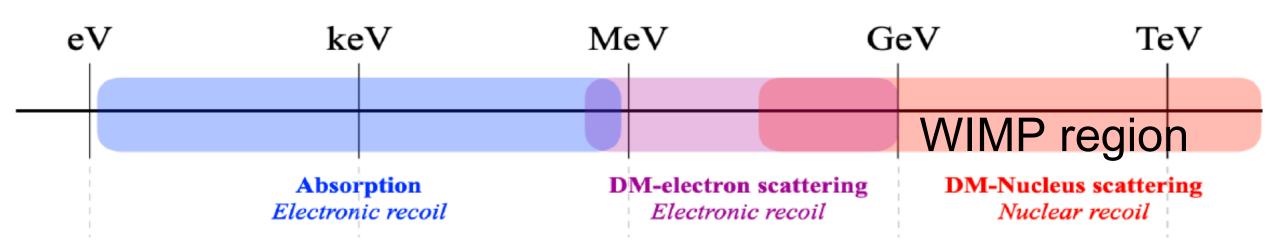


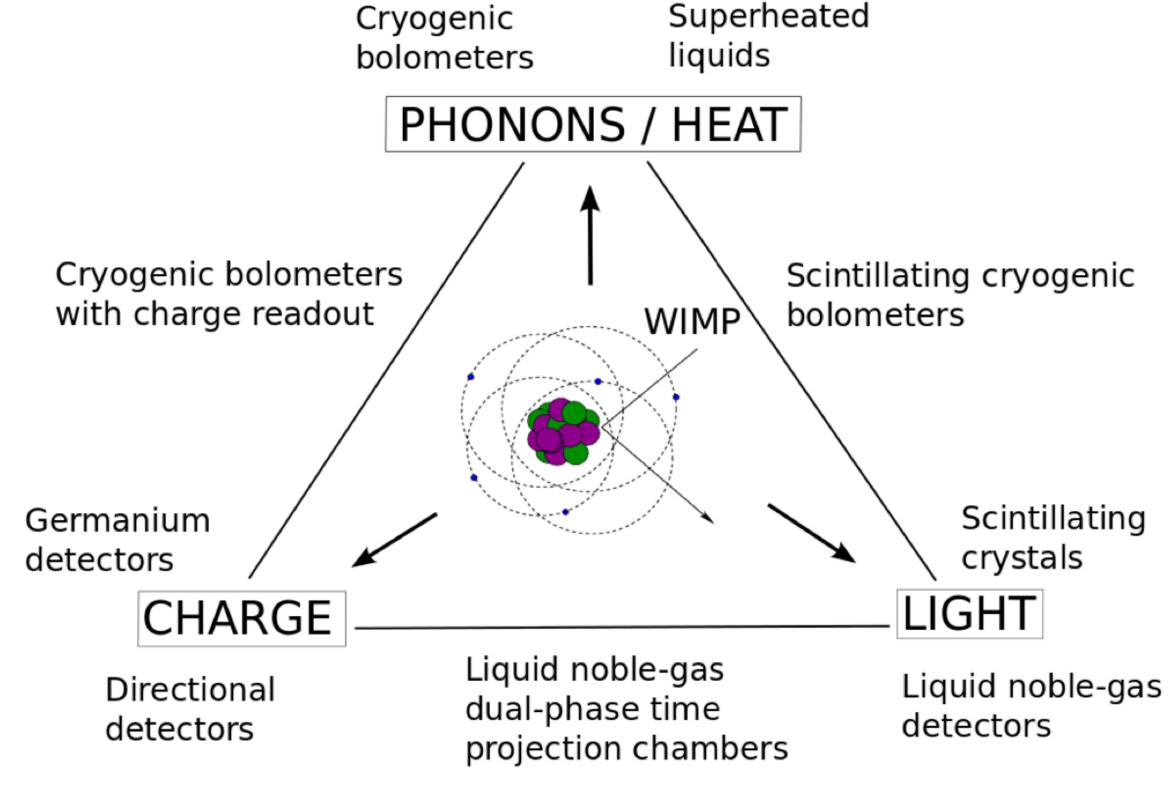




### Dark Matter

- How do we search for dark matter?
  - In direct detection
    - Dark matter particle-nucleus scattering
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    - Low thresholds





Scintillating cryogenic

Scintillating

bolometers

Superheated

liquids

WIMP

PHONONS / HEAT

Cryogenic

Cryogenic bolometers

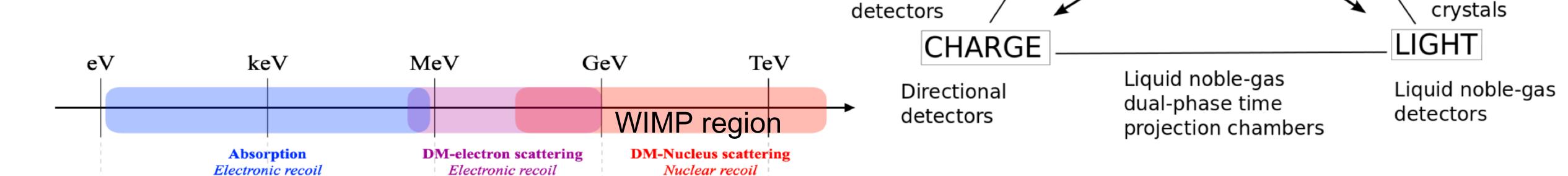
with charge readout

Germanium

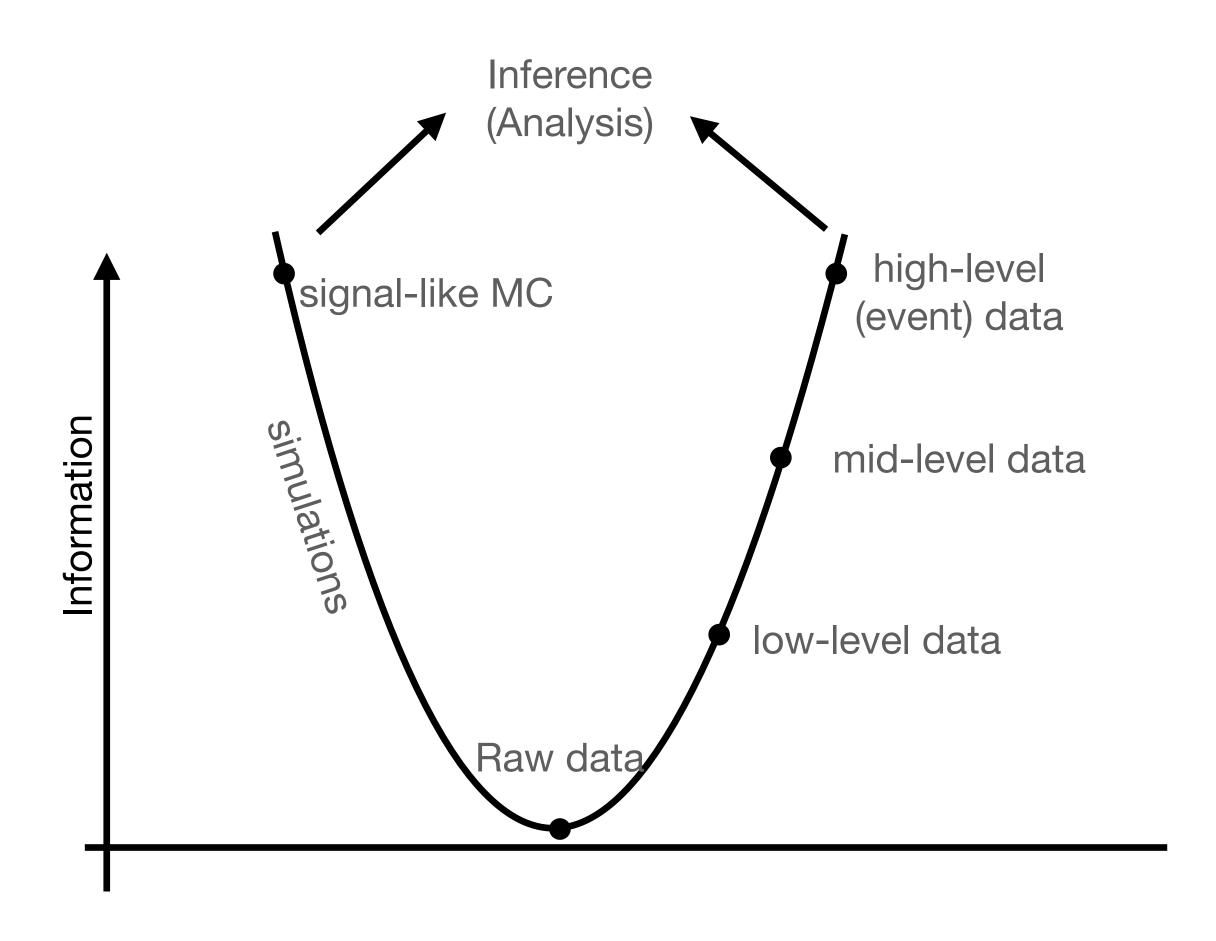
bolometers

### Dark Matter

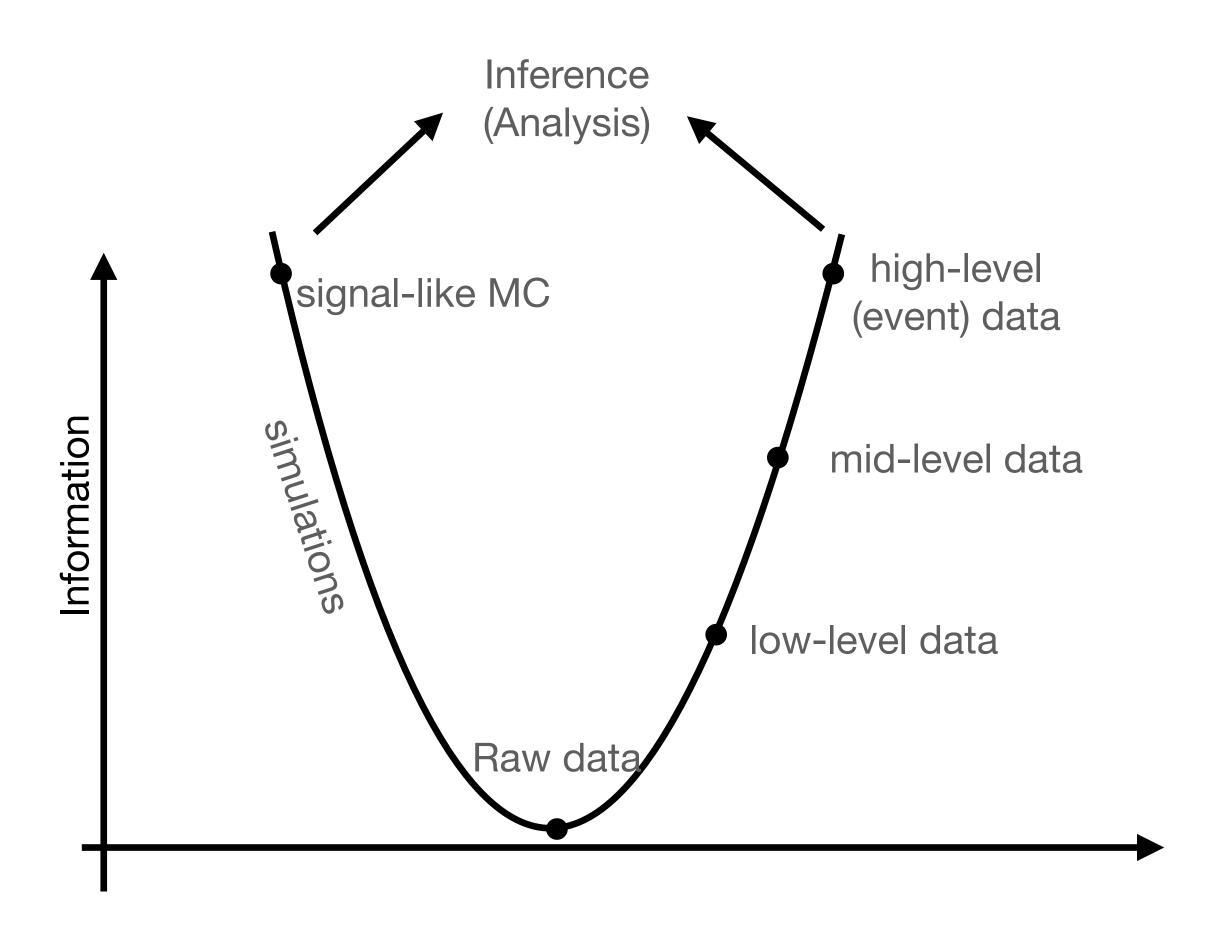
- How do we search for dark matter?
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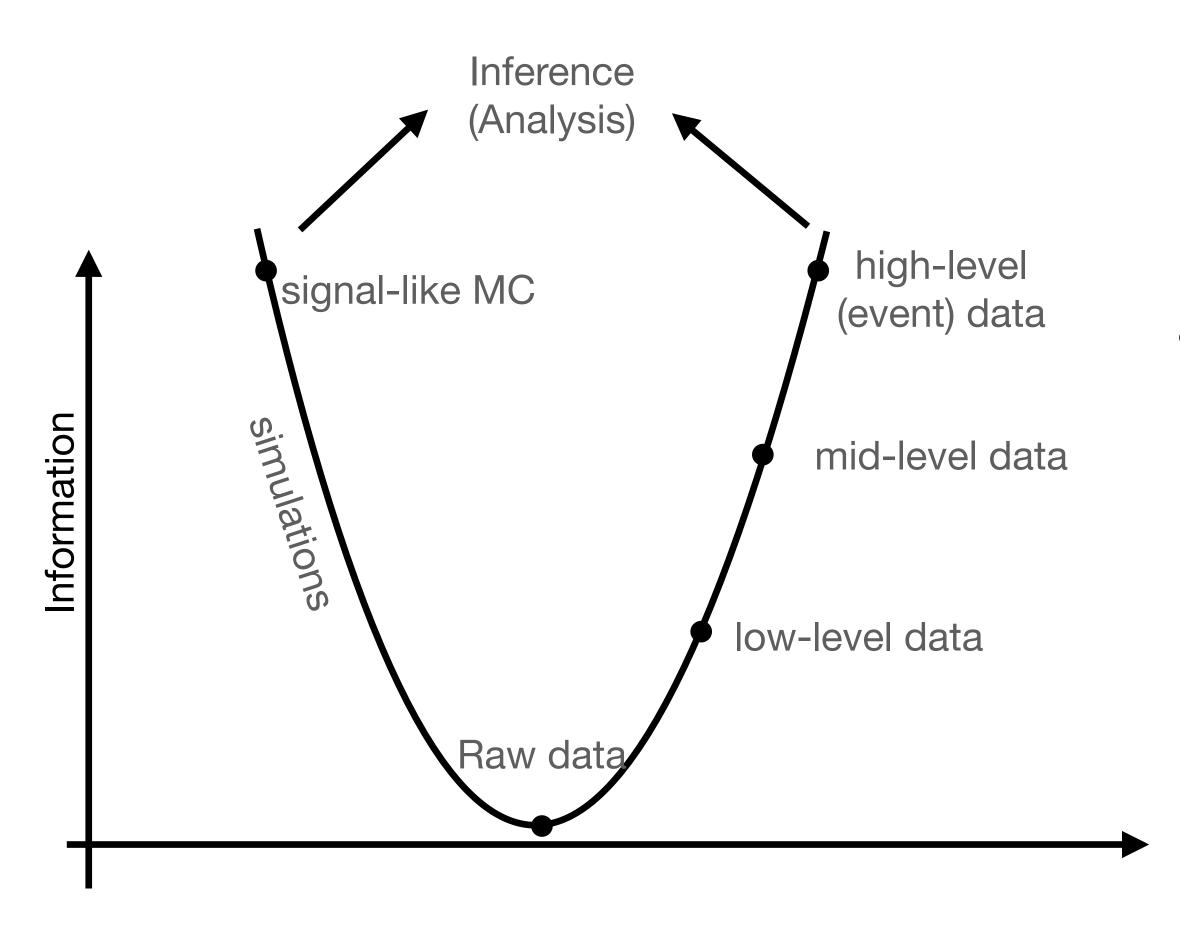
#### This talk will discuss liquid-noble TPC detectors



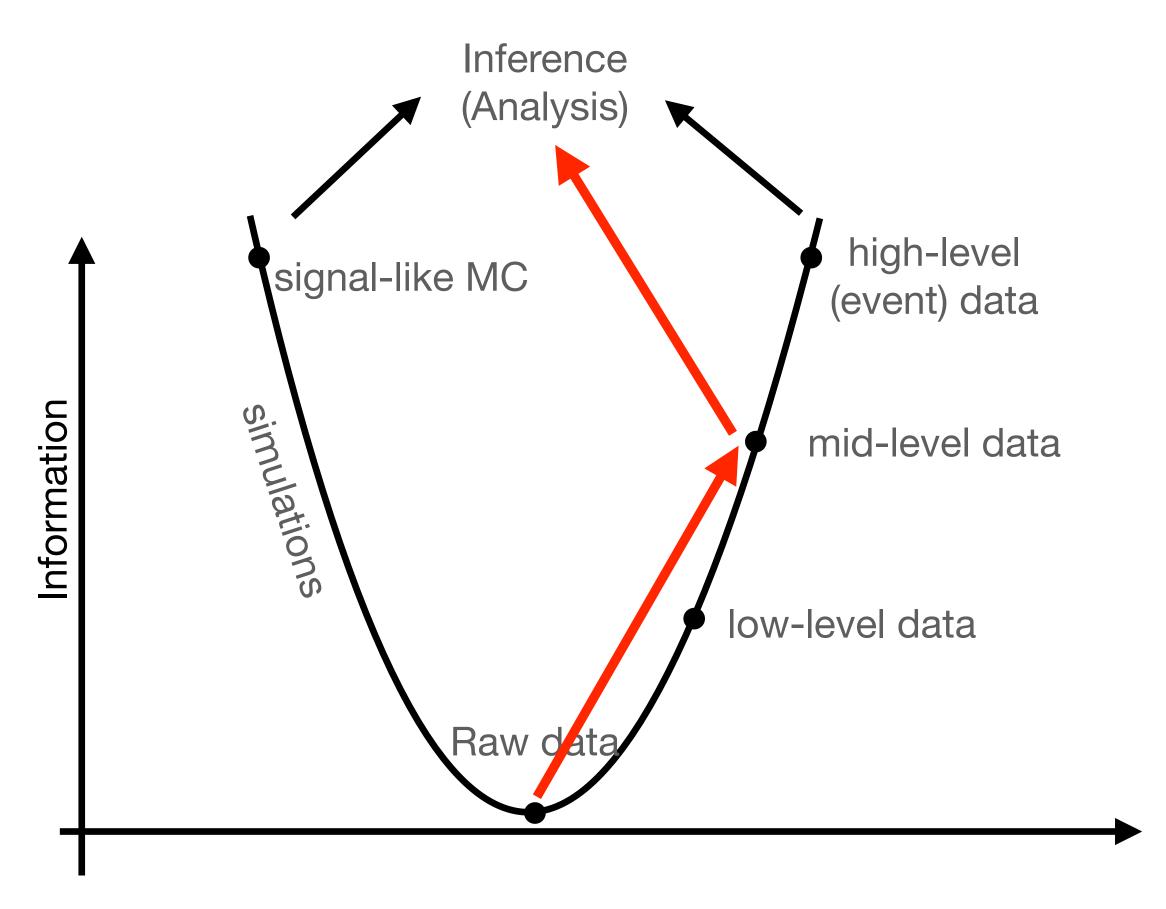
- Forward model and solving the inverse problem
- Detailed knowledge of the physical processes are encoded in the forward model



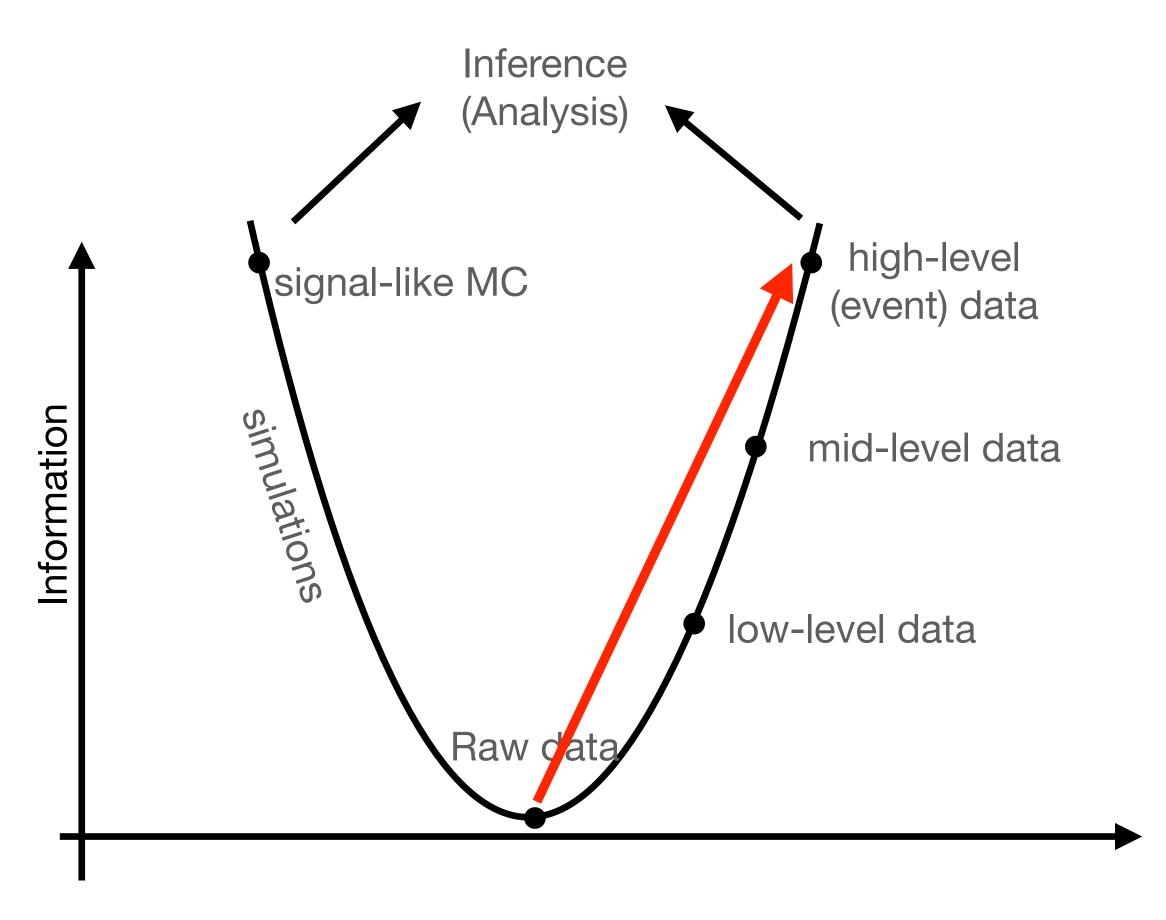
- Forward model and solving the inverse problem
- Detailed knowledge of the physical processes are encoded in the forward model
  - Model dependency
  - Pushing the limits, low-thresholds
    - Hardware
    - Software (high-rates ~50 MB/s in search mode and up to 500 MB/s in calibration)



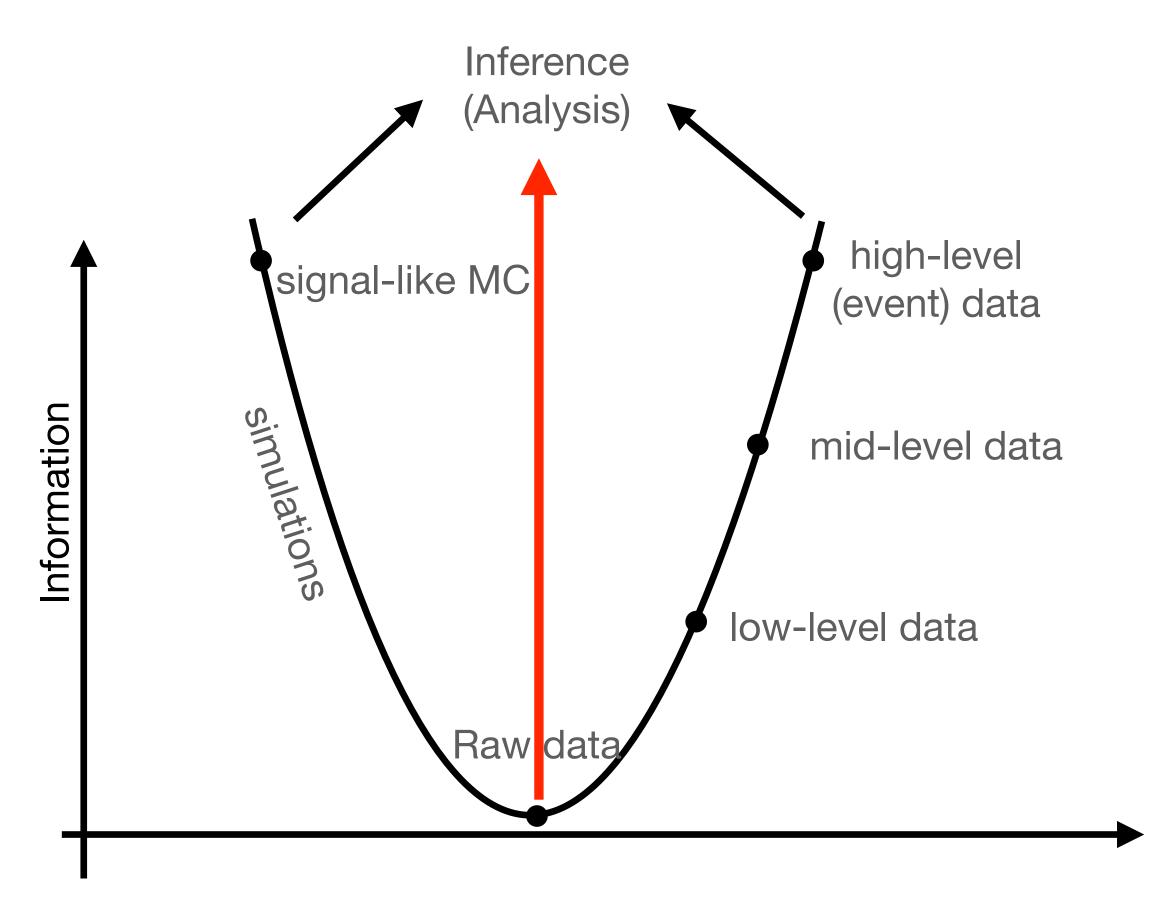
- Forward modeling and solving the inverse problem
- Detailed knowledge of the physical processes are encoded in the forward model
- How can we optimize analysis such that we can improve the science?



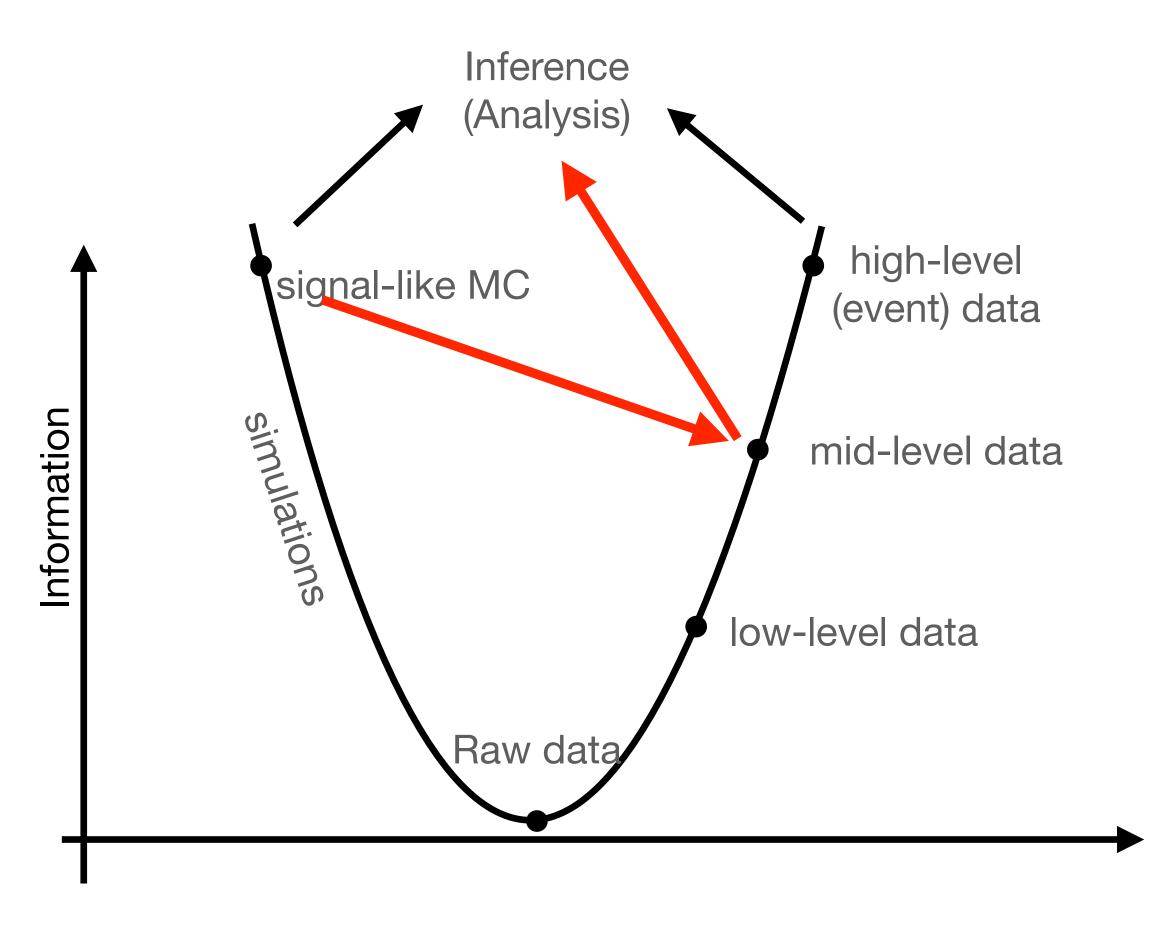
- Forward modeling and solving the inverse problem
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- How can we improve "reconstruction" to advance the science?



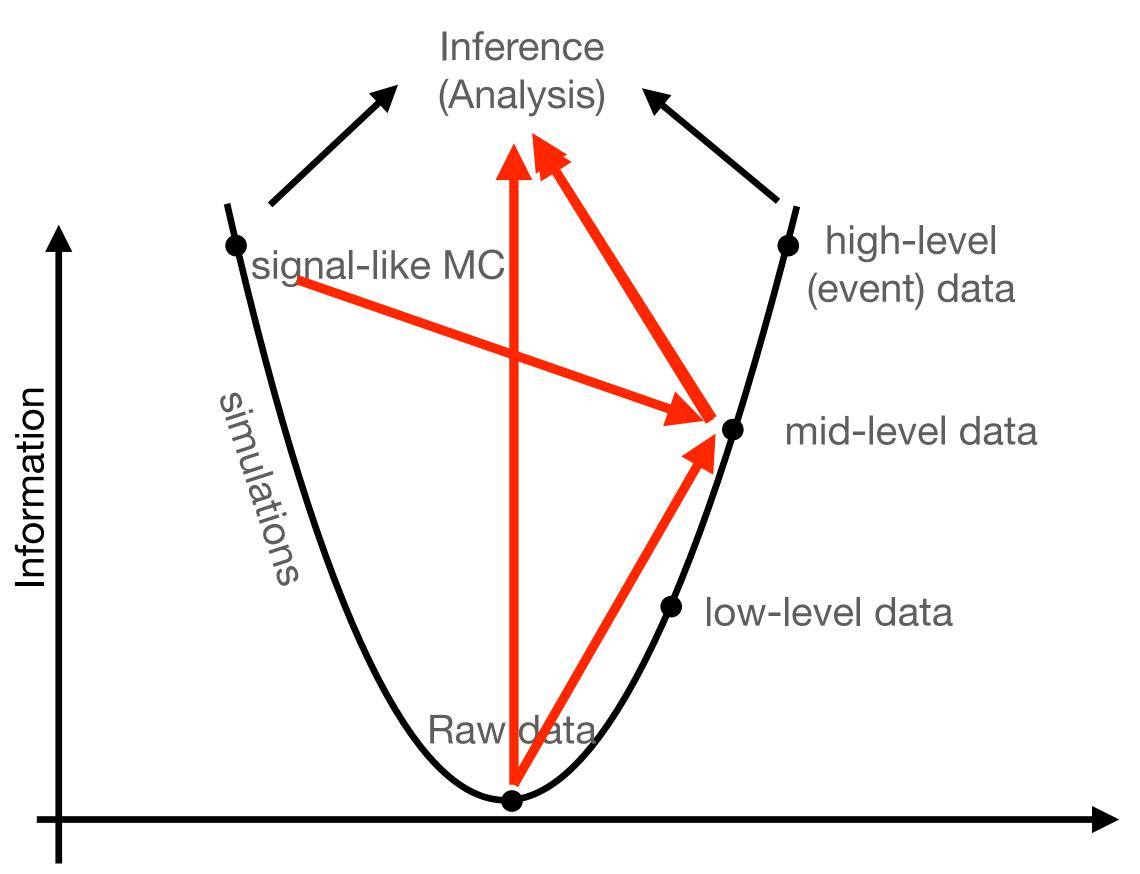
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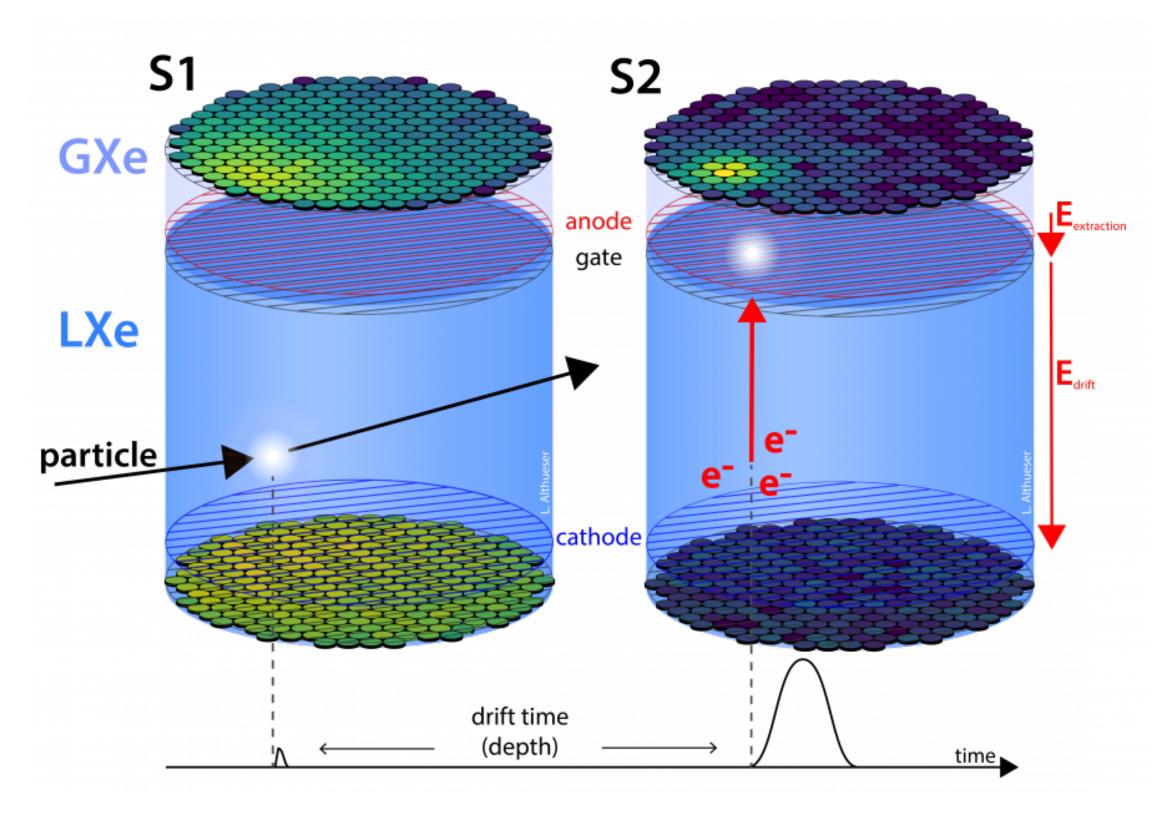
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- Forward modeling and solving the inverse problem
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- Forward modeling and solving the inverse problem
- Detailed knowledge of the physical processes are encoded in the forward model
- How can we improve "reconstruction" to advance the science?
  - Problems
    - Regression (energy, position)
    - Classification (PID, anomaly detection)
    - Inference

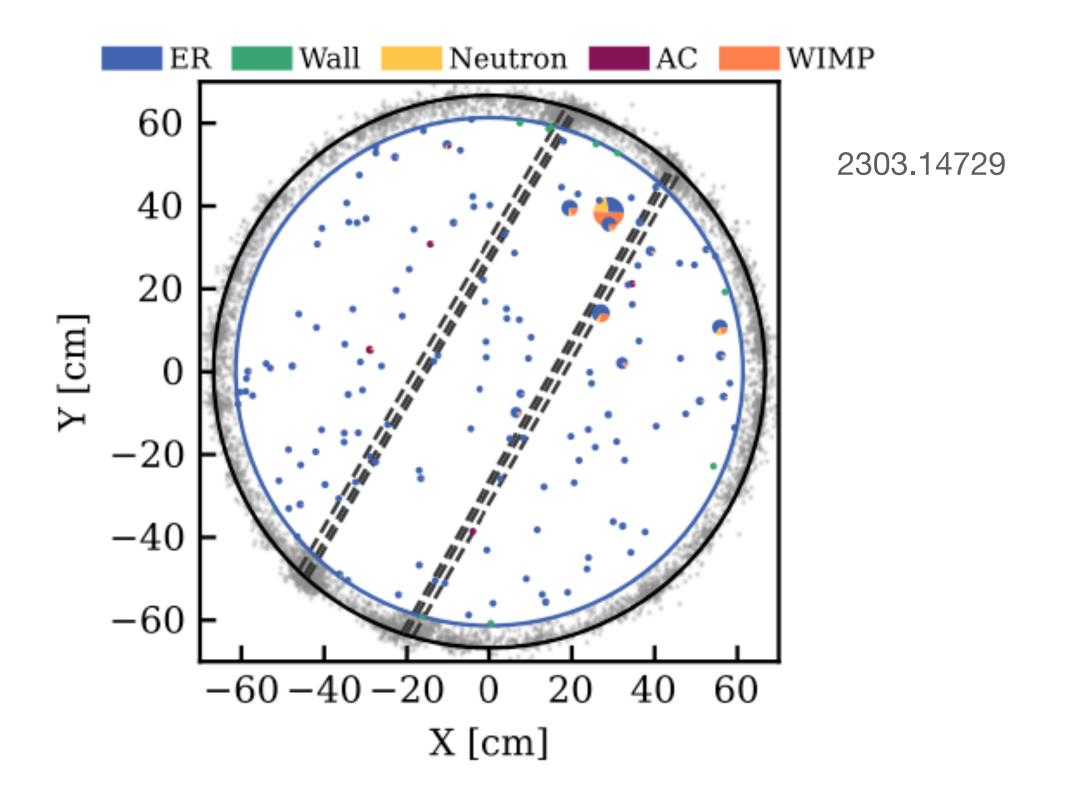


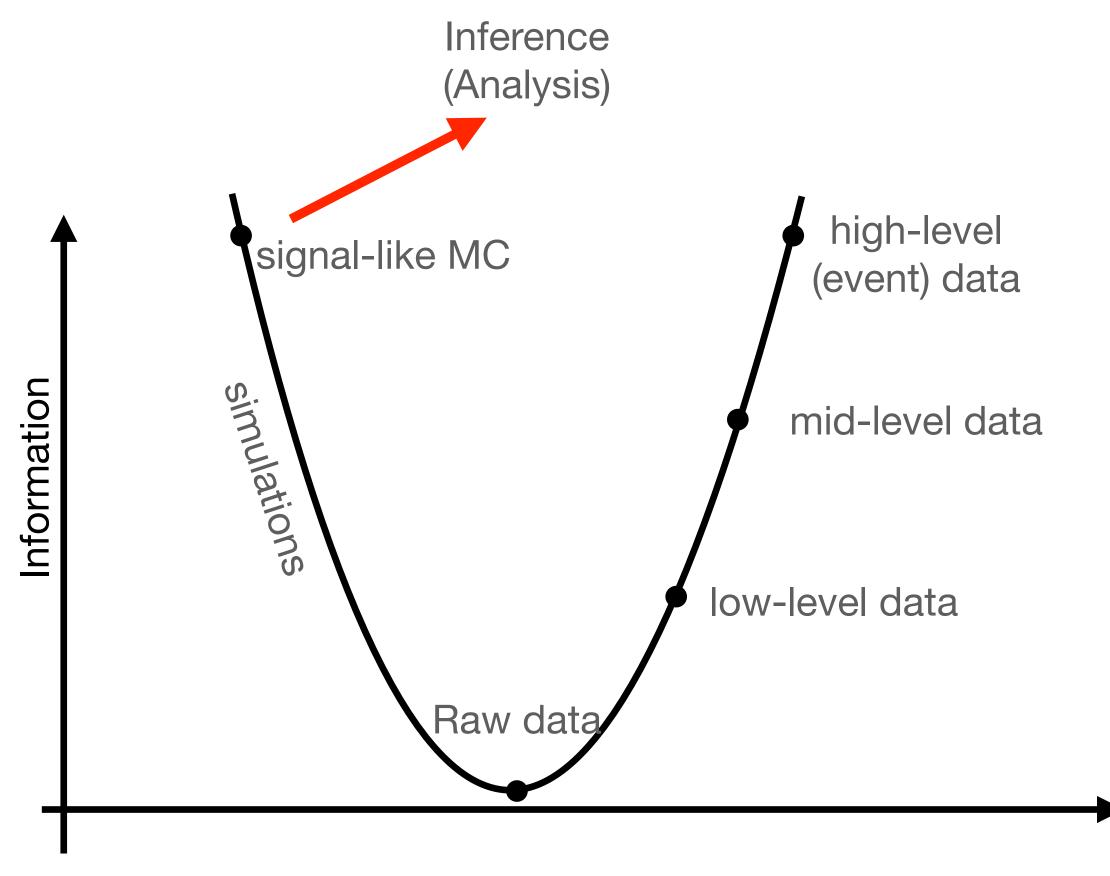
$$E = W \left( \frac{cS_1}{g_1} + \frac{cS_2}{g_2} \right)$$

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#### **Interaction vertex**

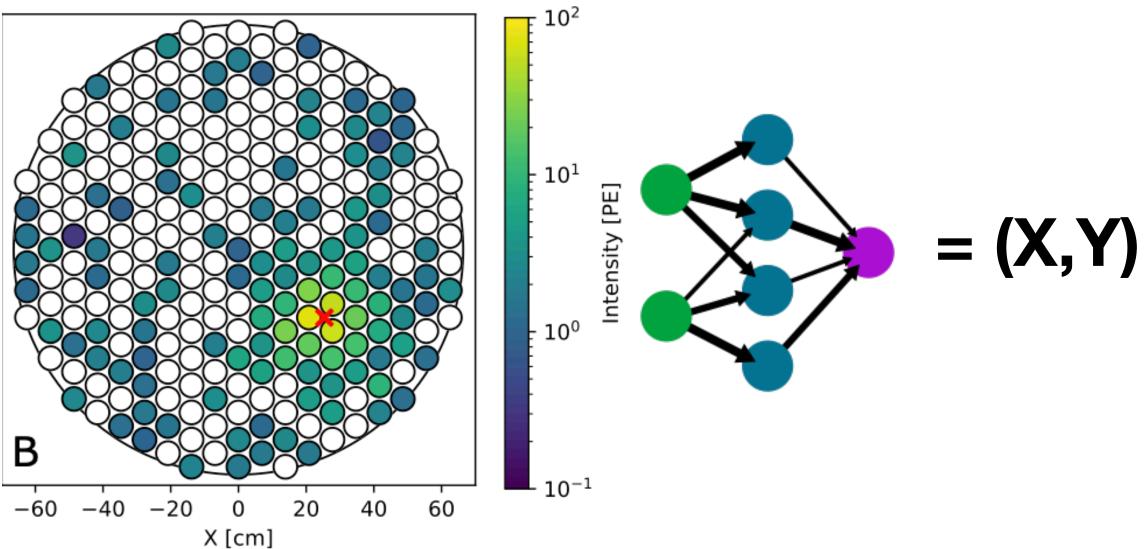
- 2D position given by hit pattern (bottom array)
- XENONnT, LZ, Panda-X (dual-phase LXe-TPC)
- Darkside -20k (dual-phase LAr-TPC)





#### Interaction vertex

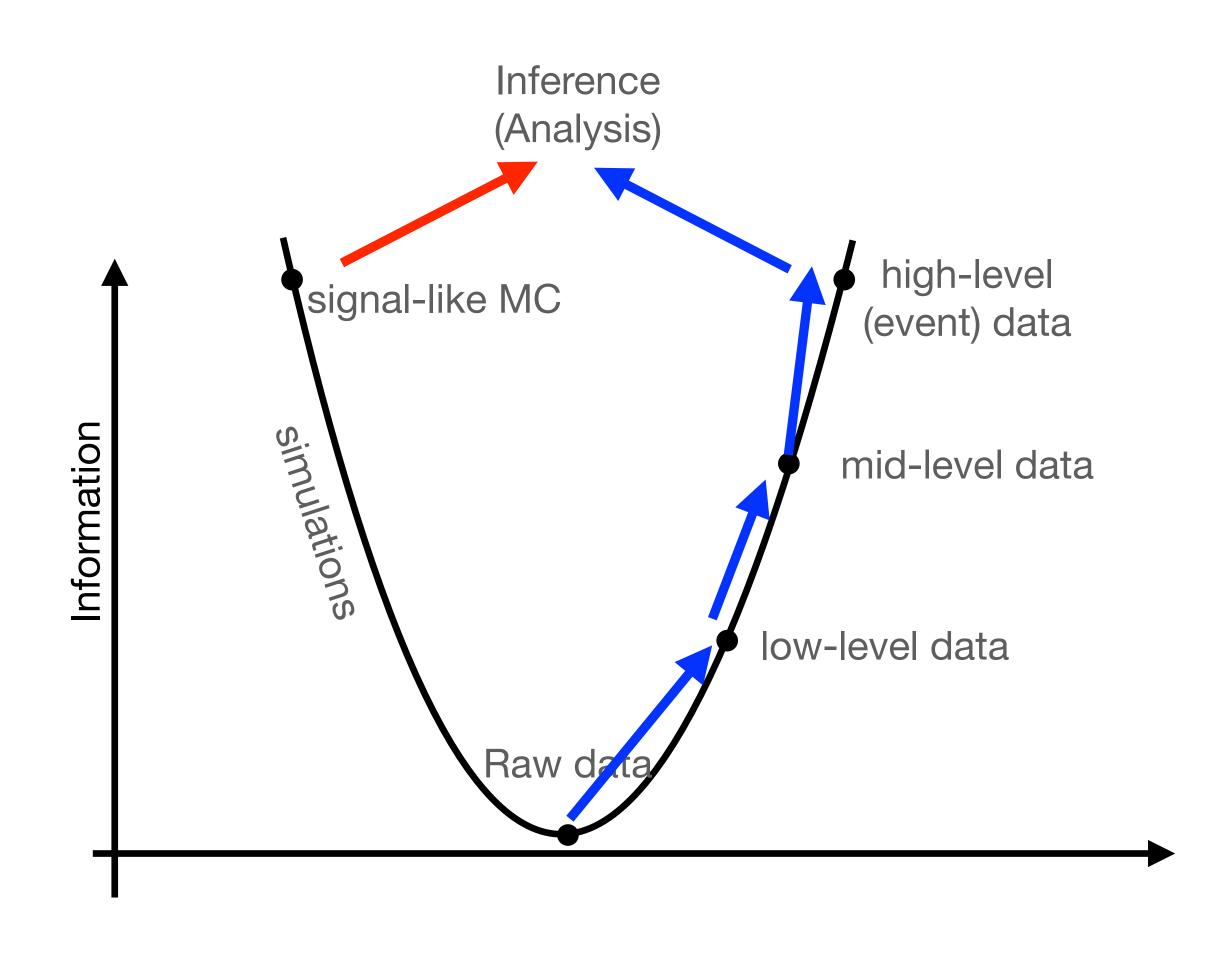
- 2D position given by hit pattern (bottom array)
- Train on simulation (MC)



Classic approach: MLP

GCN Front. Artif. Intell. 5 (2022) 832909

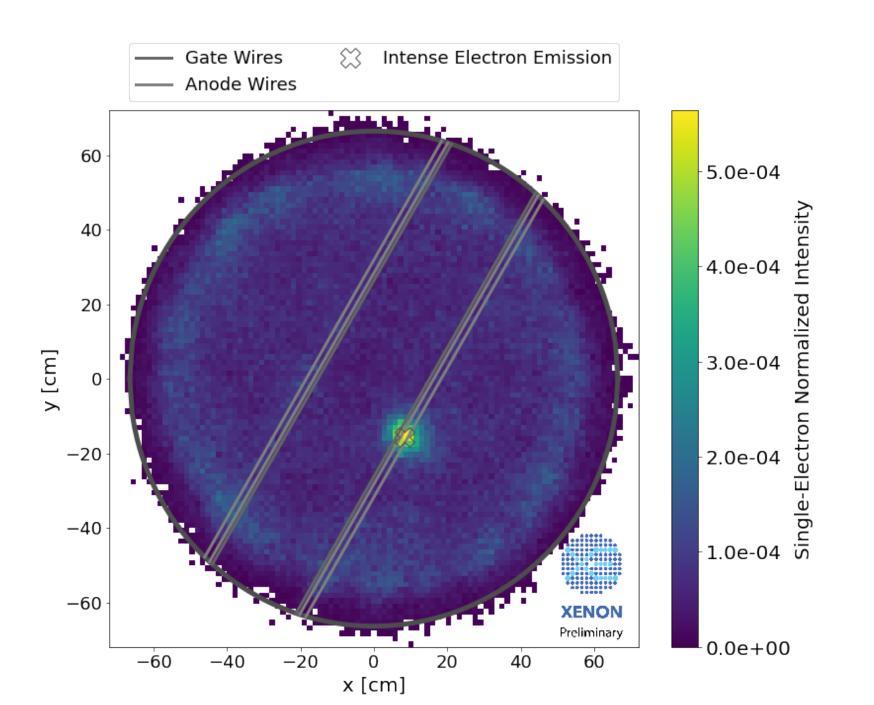
arXiv:2112.07995

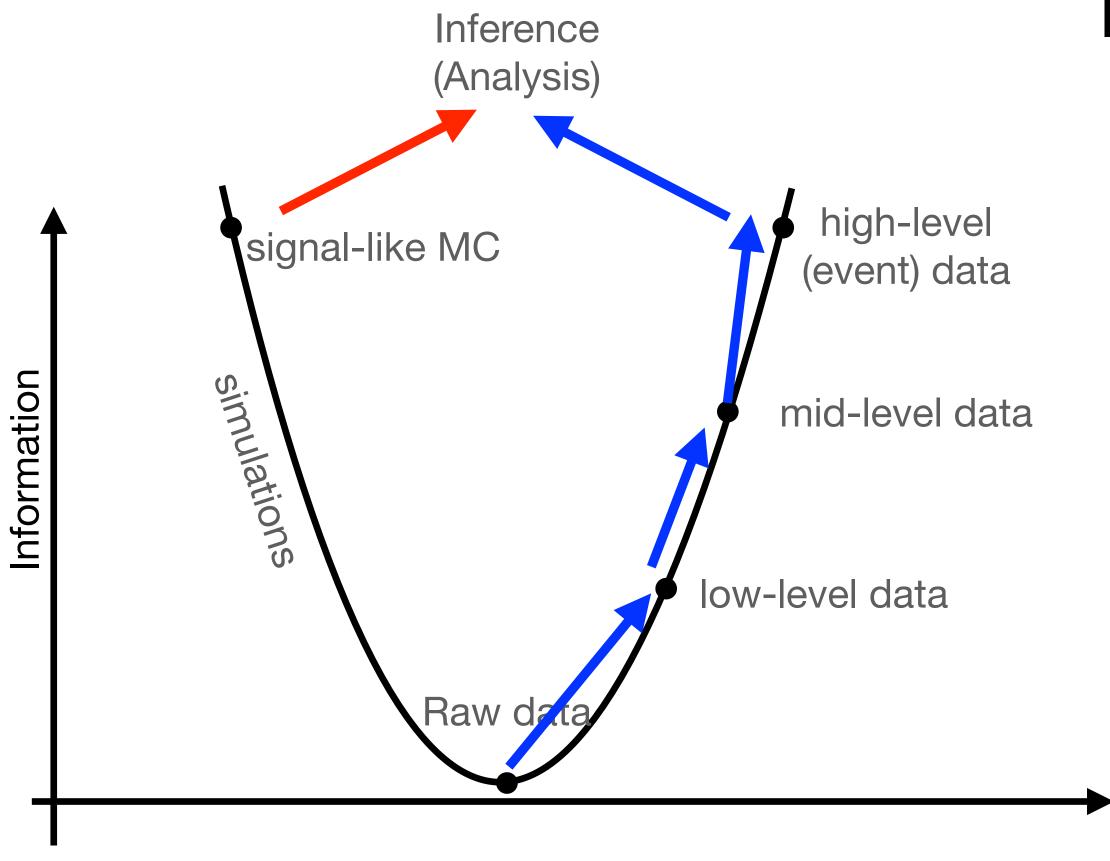


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#### Interaction vertex

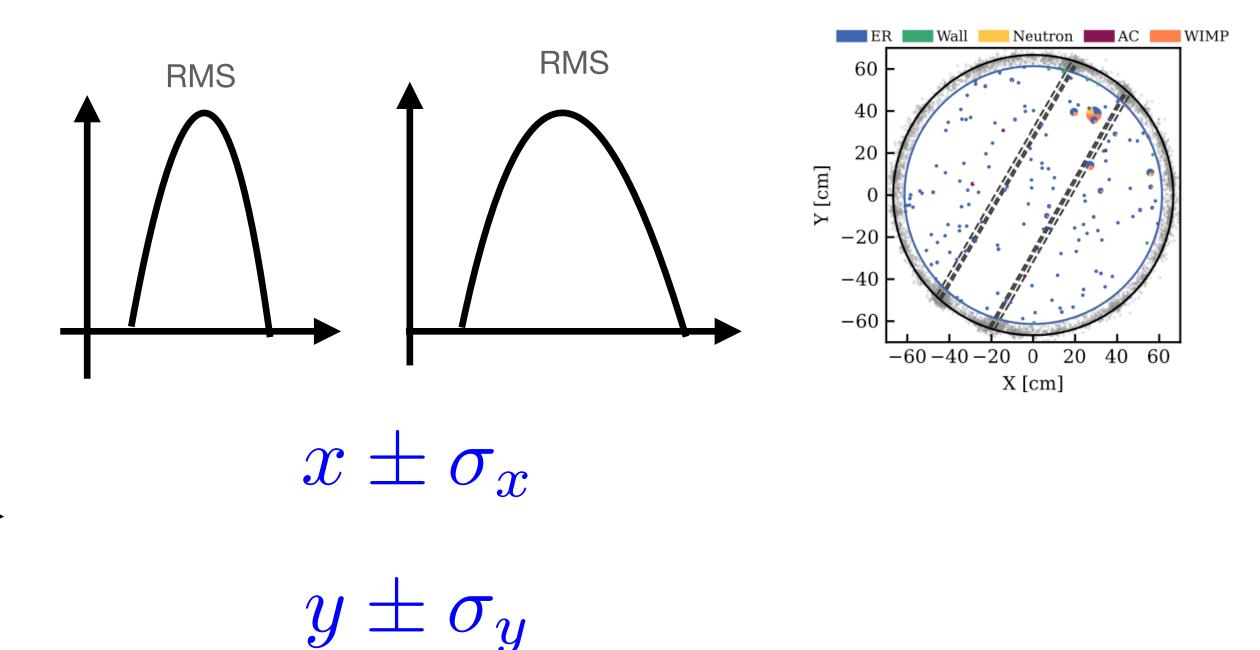
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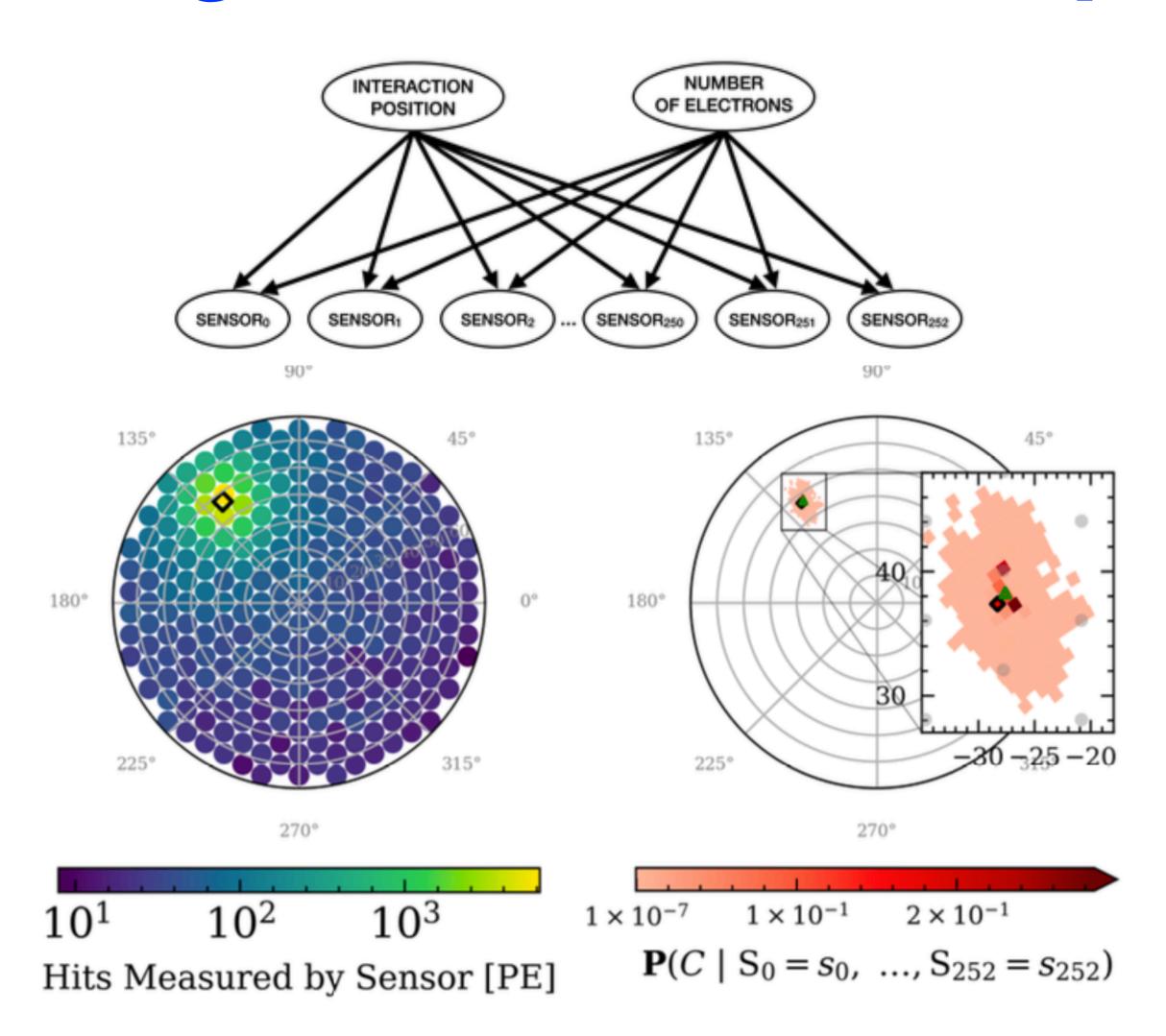


#### **Interaction vertex**

- 2D position given by hit pattern
- What is the uncertainty on the inference?



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#### **Interaction vertex**

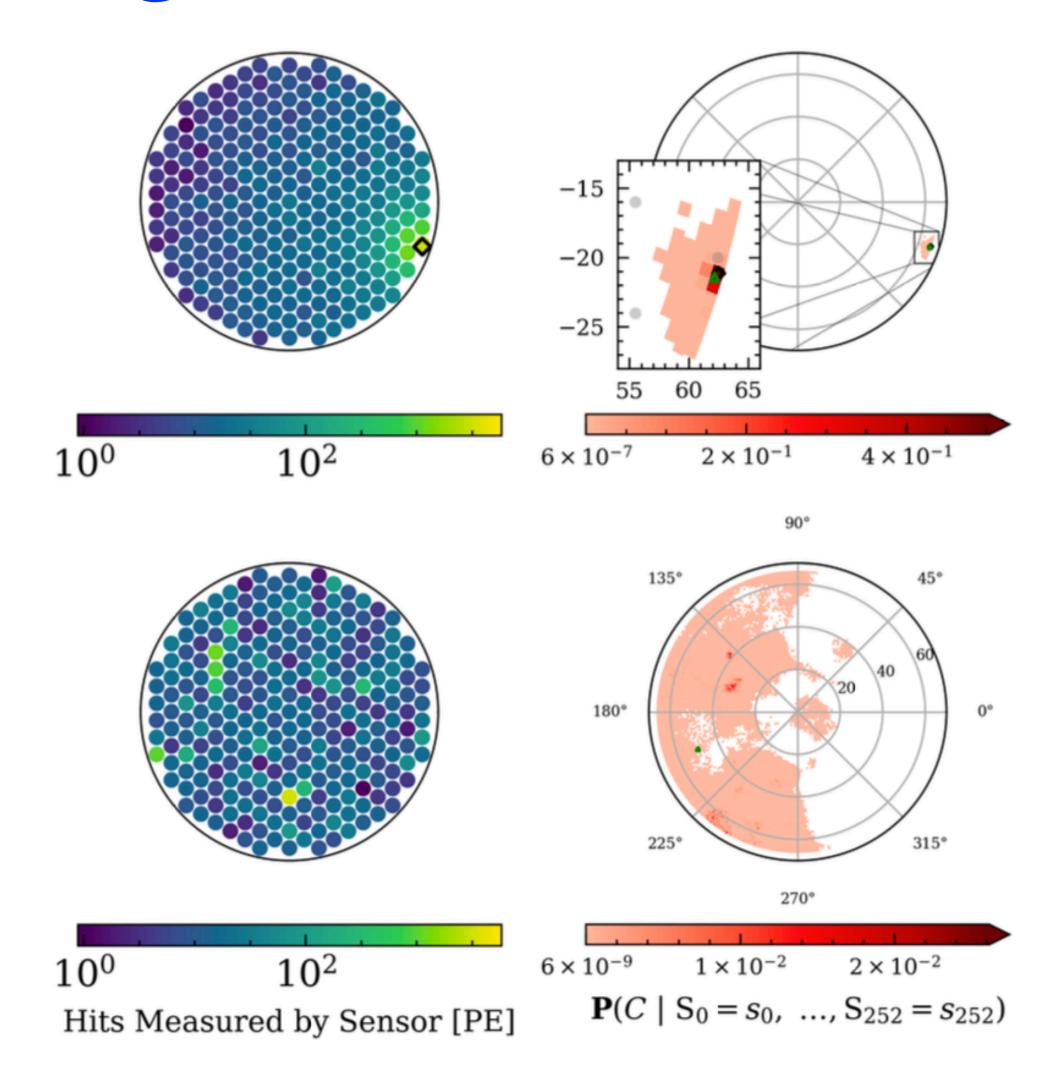
- 2D position given by hit pattern
- What is the uncertainty on the inference?

Graphical Models are All You Need: Per-interaction reconstruction uncertainties in a dark matter detection experiment

NeurIPS 2022 by C. Peters arxiv:2205.10305

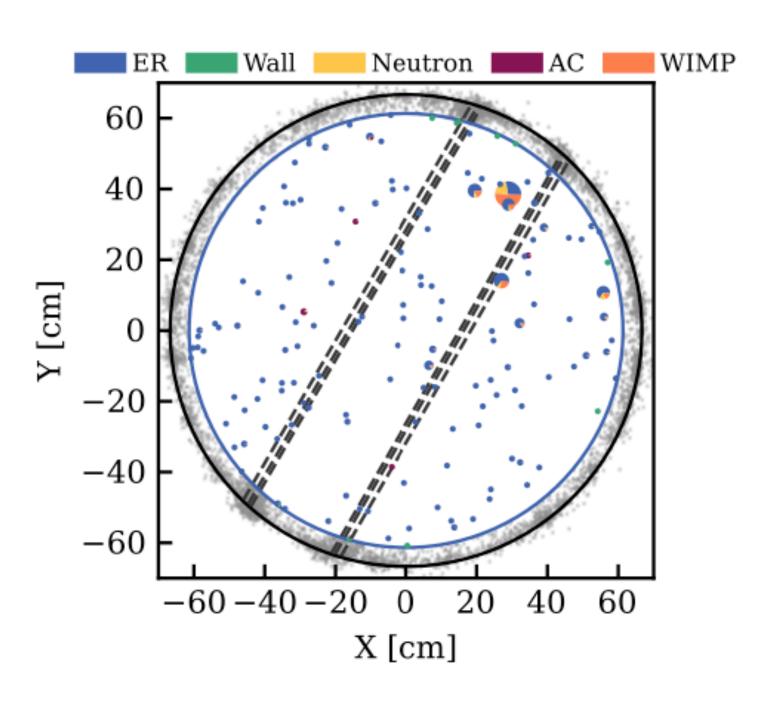
Provides per-event uncertainty on the position

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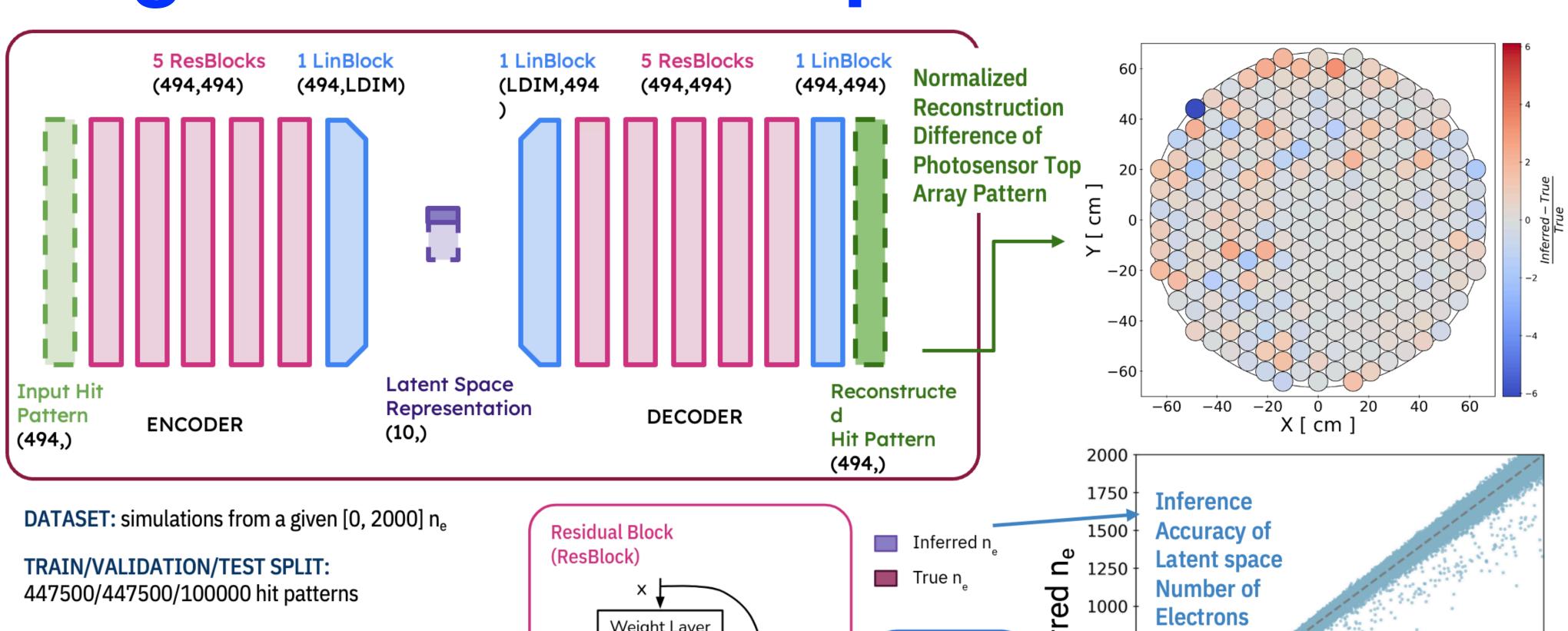
Graphical Models are All You Need: Per-interaction reconstruction uncertainties in a dark matter detection experiment

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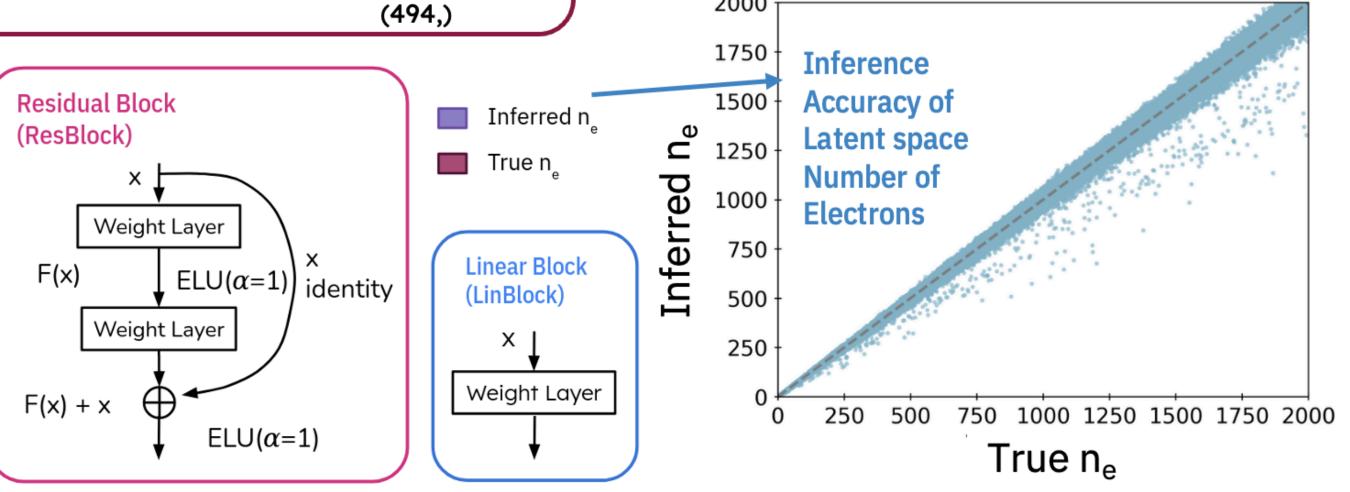
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Ivy Li, Rice



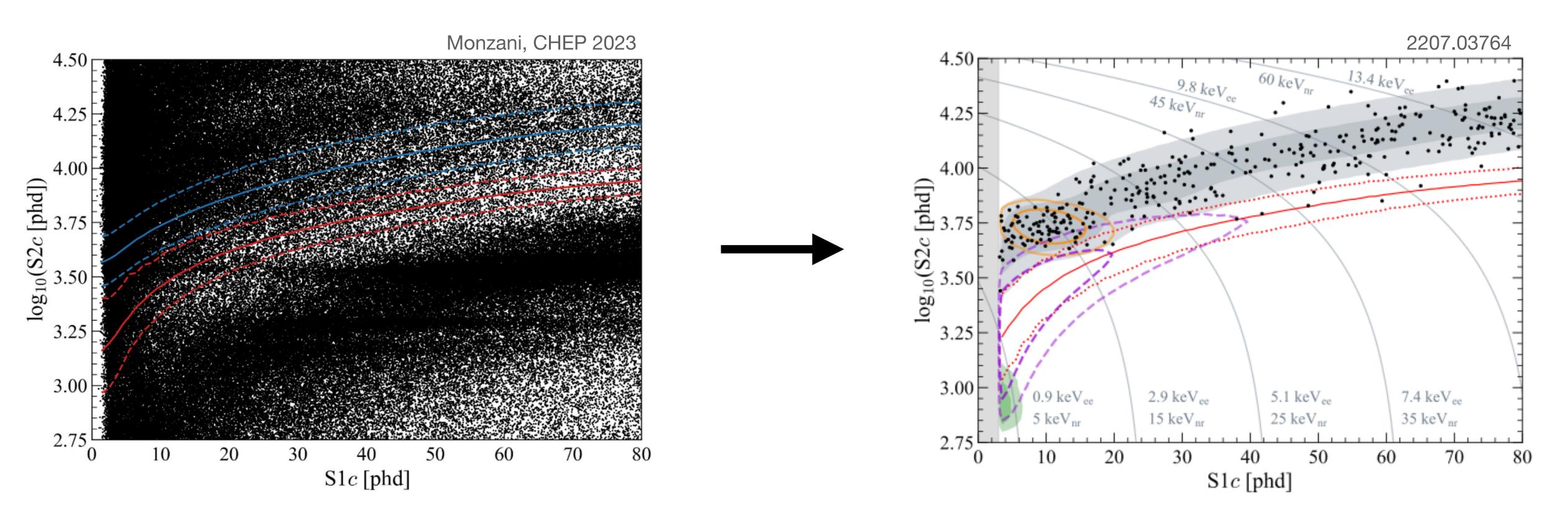
**LATENT SPACE REPRESENTATION:** constrain one value **as number of electrons** but allow the others to evolve freely

**LOSS FUNCTION:** weighted sum of the MSE loss of reconstruction of hit pattern and MSE loss of inferred number of electrons, trained with cyclic annealing

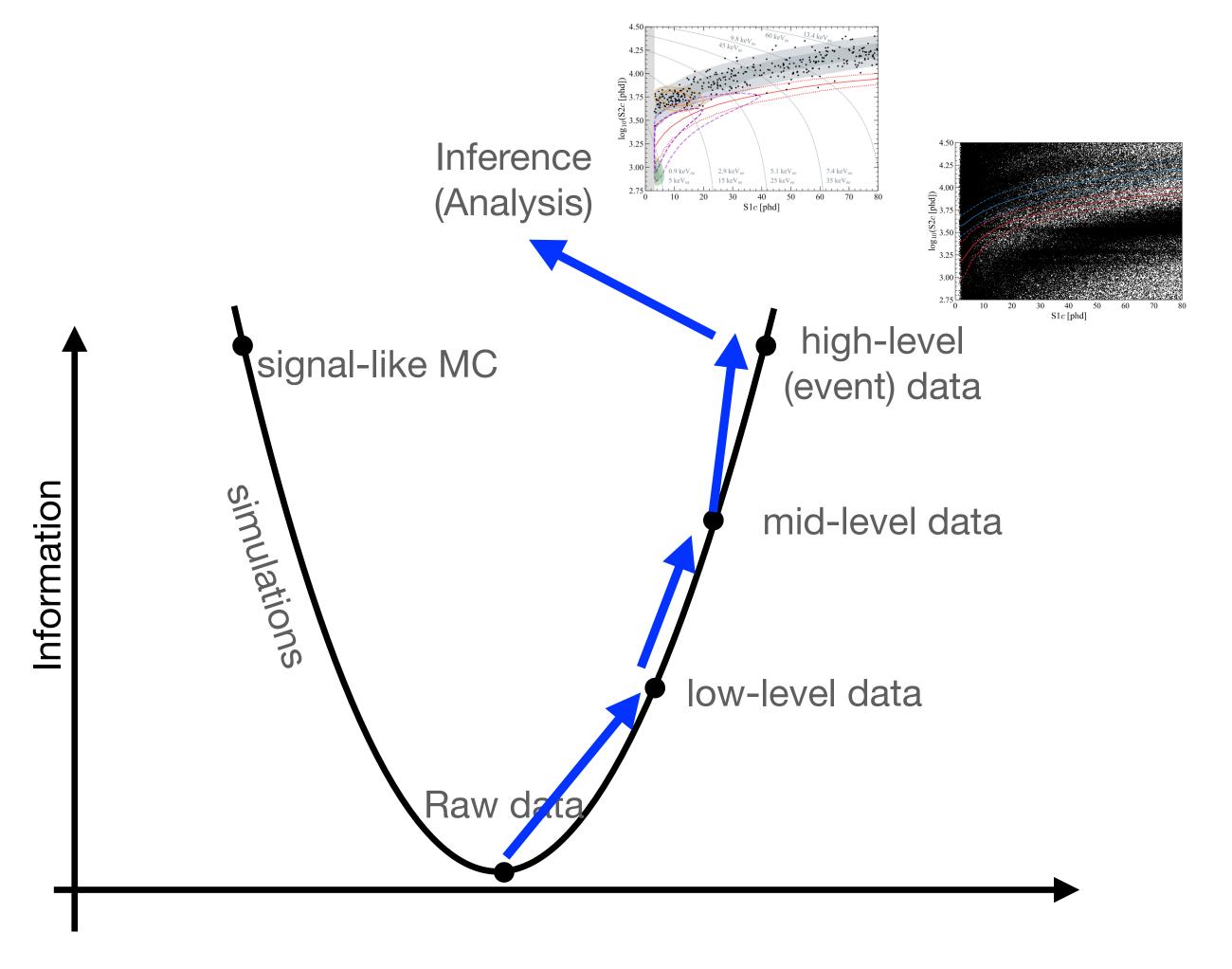


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#### **Everything is background!**



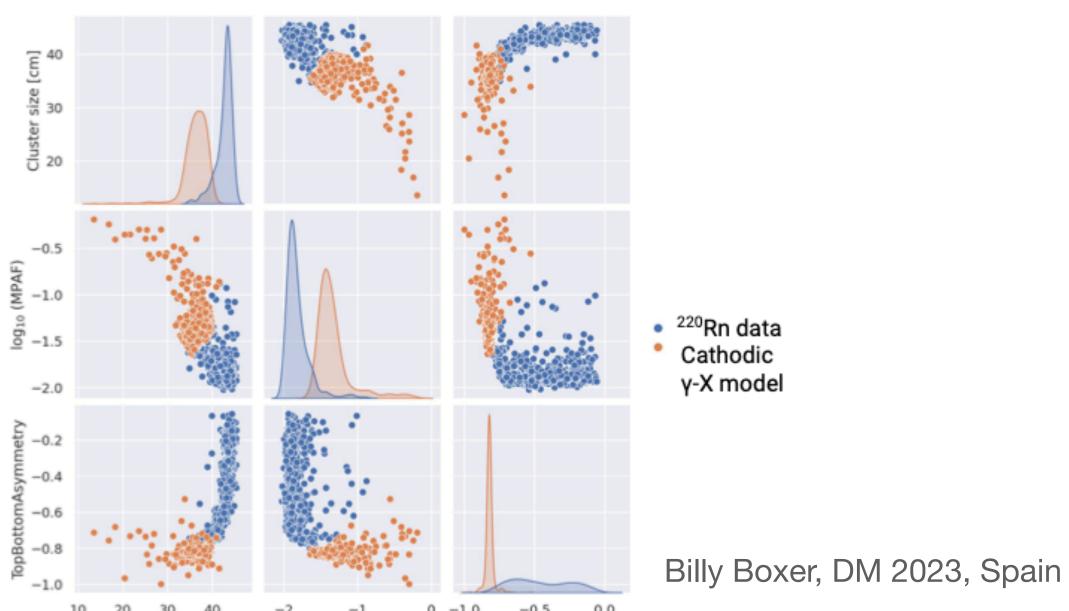
A. Higuera, MODE workshop, Princeton University, July 2023



Need to classify something? You can never go wrong with

#### **Boosted Decision Trees!!!**

- Easy to understand how the algorithm uses input data
- Select high-level features

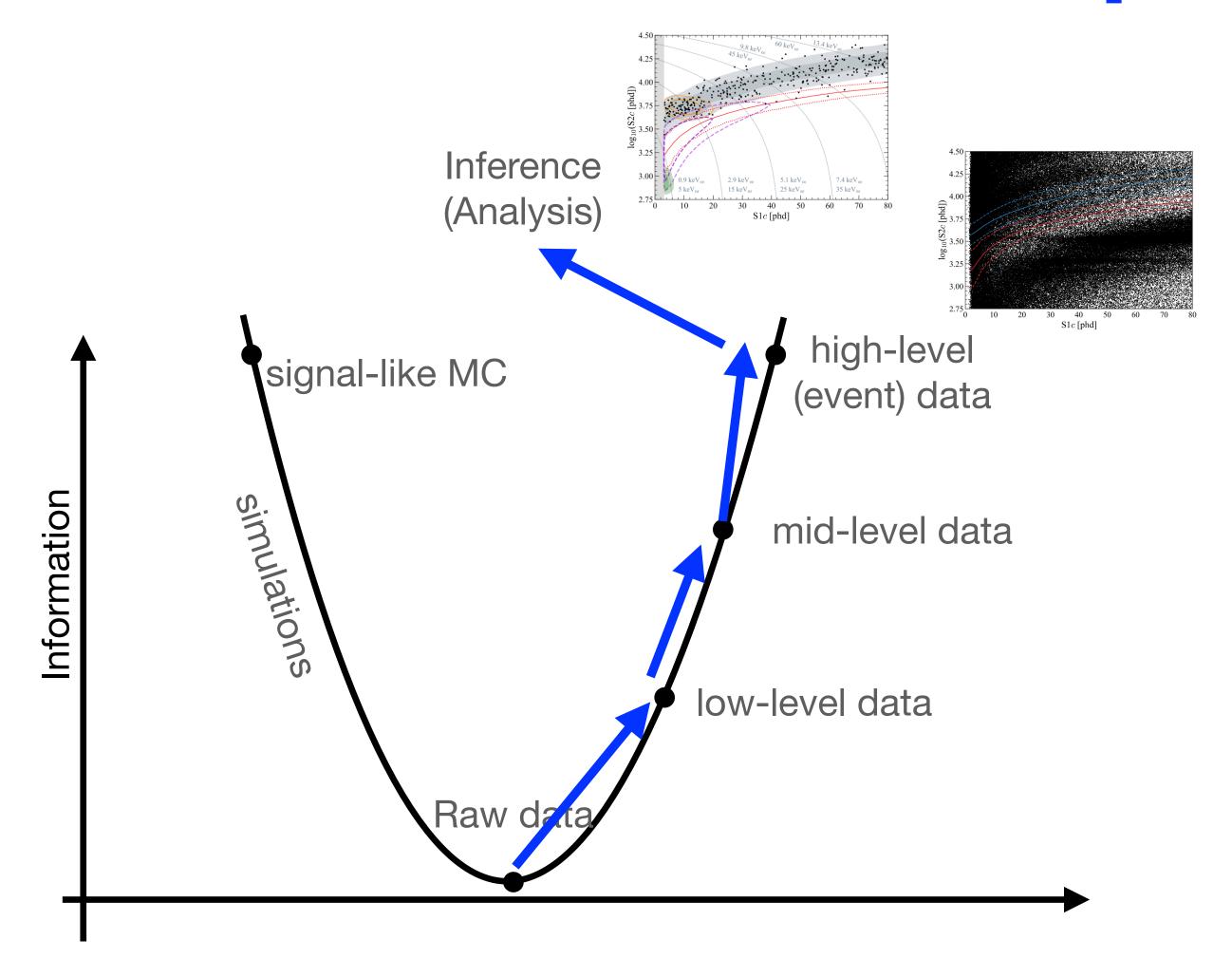


log<sub>10</sub> (MPAF)

TopBottomAsymmetry

Cluster size [cm]

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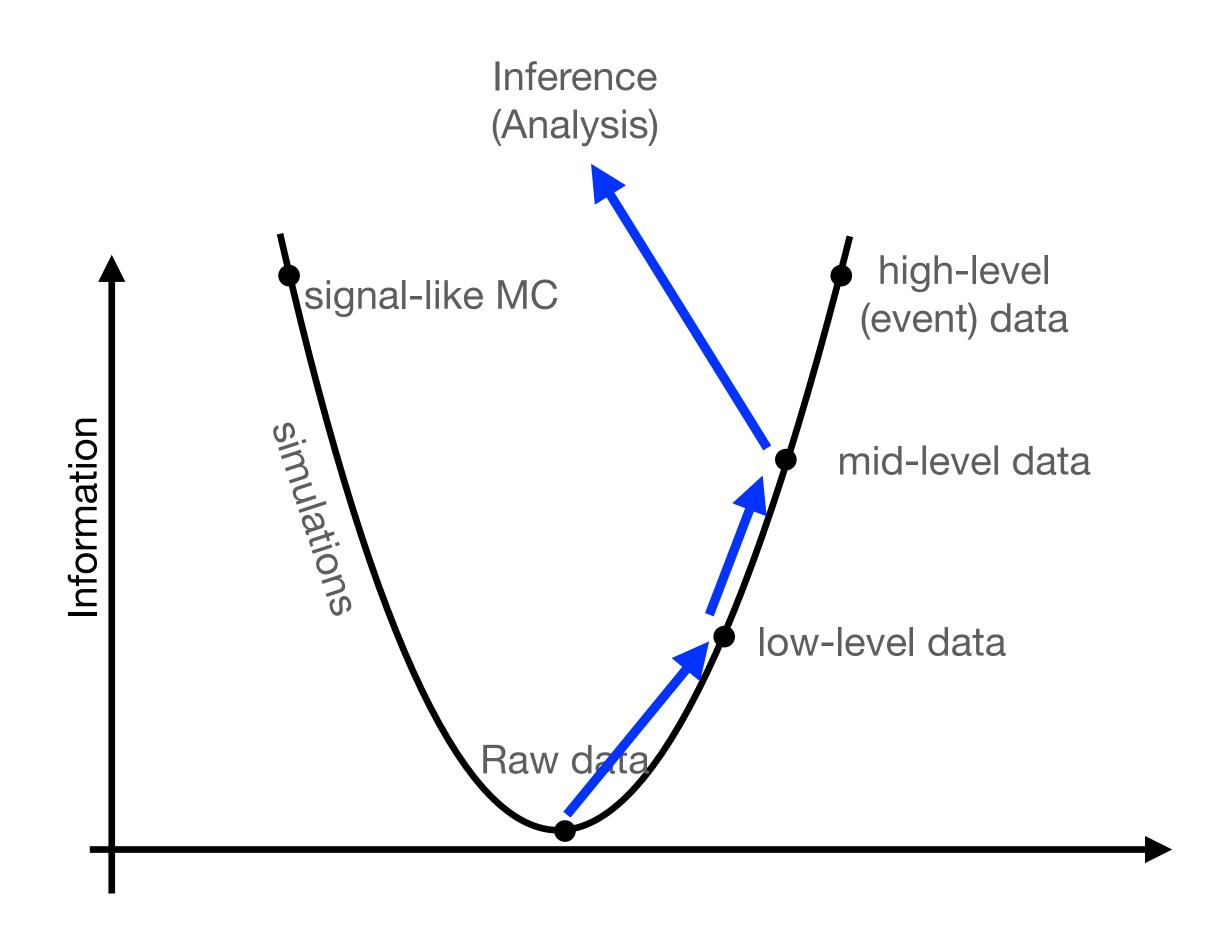


Need to classify something? You can never go wrong with

#### **Boosted Decision Trees!!!**

Why do tree-based models still outperform deep learning on tabular data? (arxiv.2207.08815)

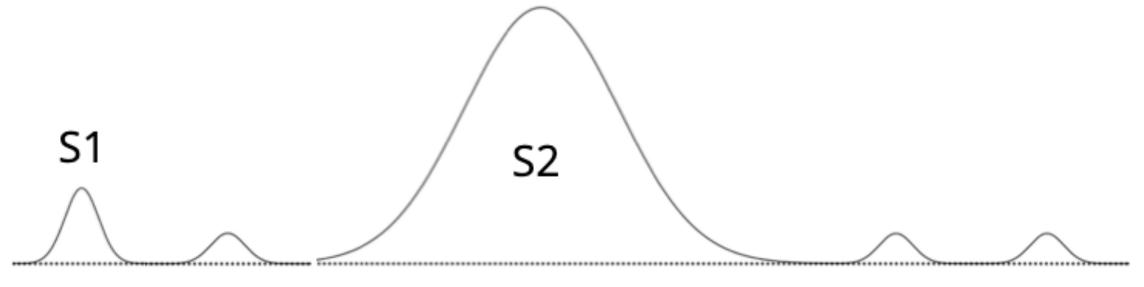
- Improving sensitivity to low-mass dark matter in LUX using an novel electrode background mitigation technique (arxiv:2011.0960)
- Study of background from accidental coincidence signals in the PandaX-II experiment (arxiv:2204.11175)
- Boosted decision trees approach to neck alpha events discrimination in DEAP-3600 experiment (arxiv.2009.00895)

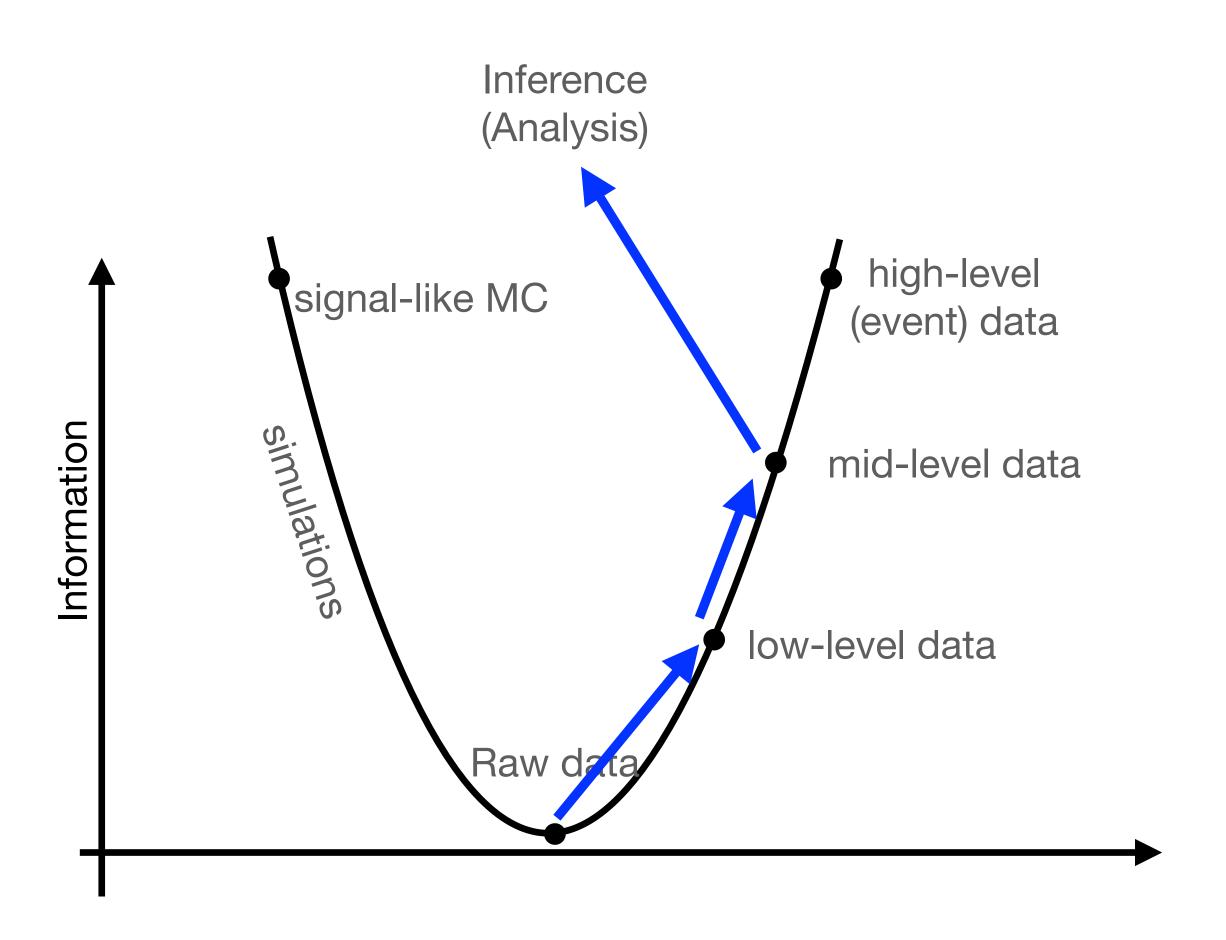


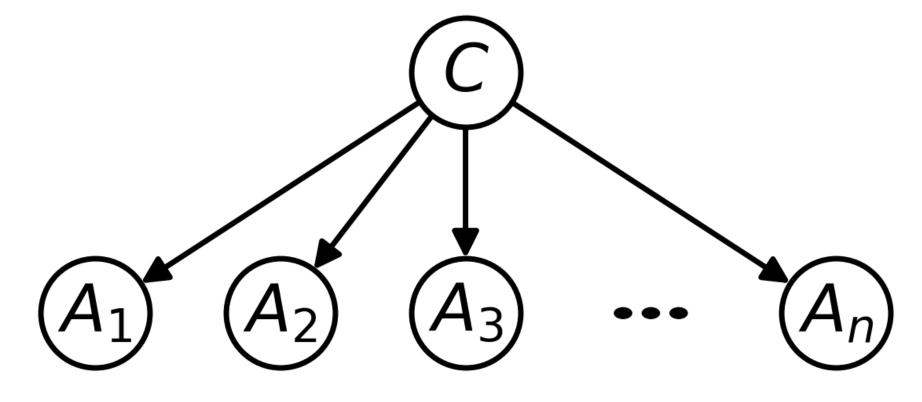
Why do tree-based models still outperform deep learning on tabular data? (arxiv.2207.08815)

What if you are not dealing with tabular data (event-level data)?

Pulse data (waveform)



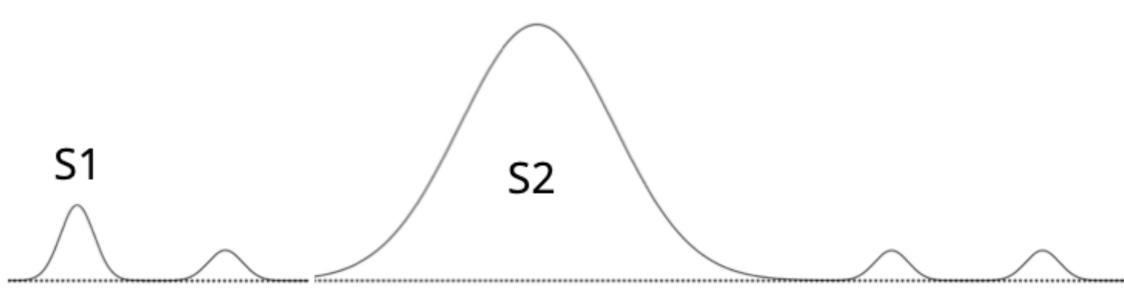


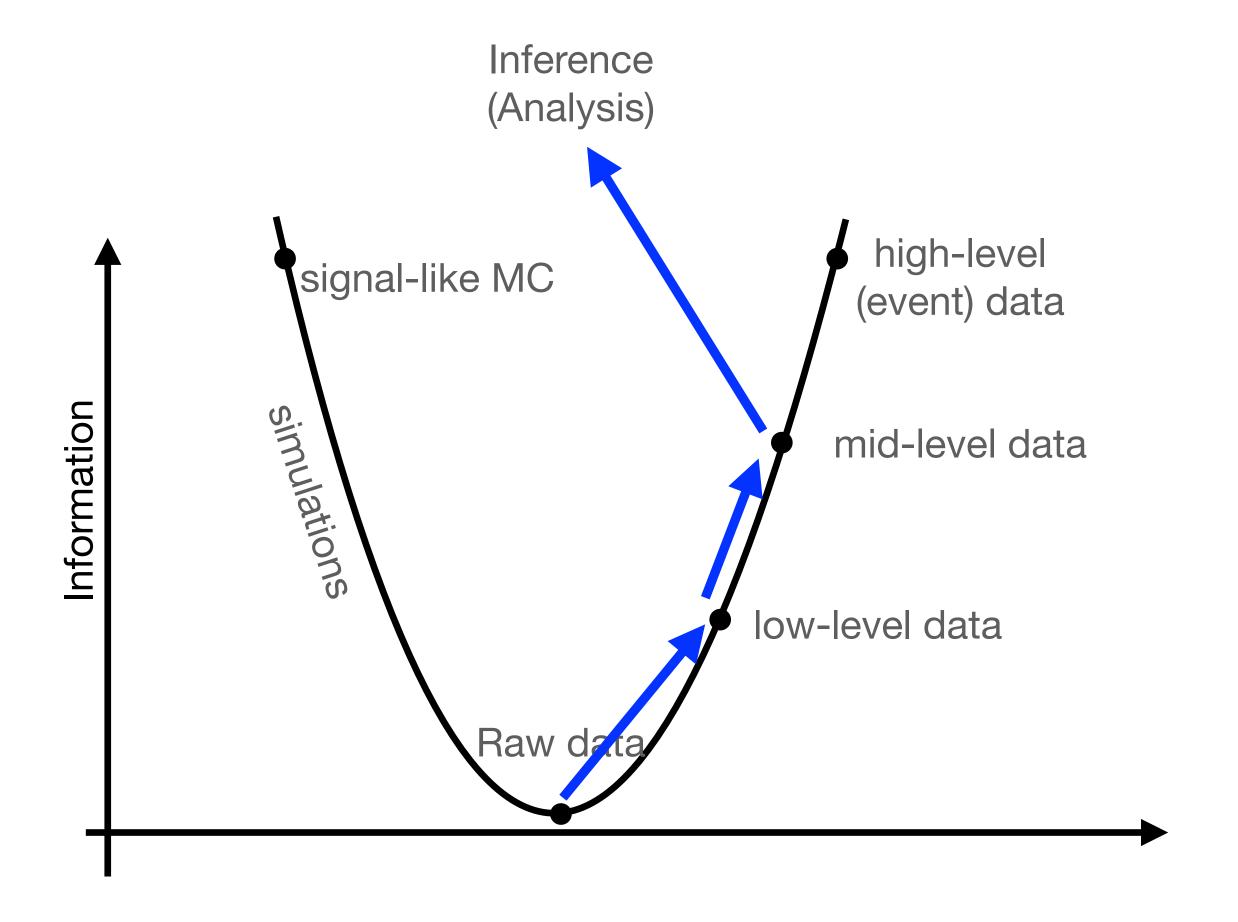


**Bayesian Network** 

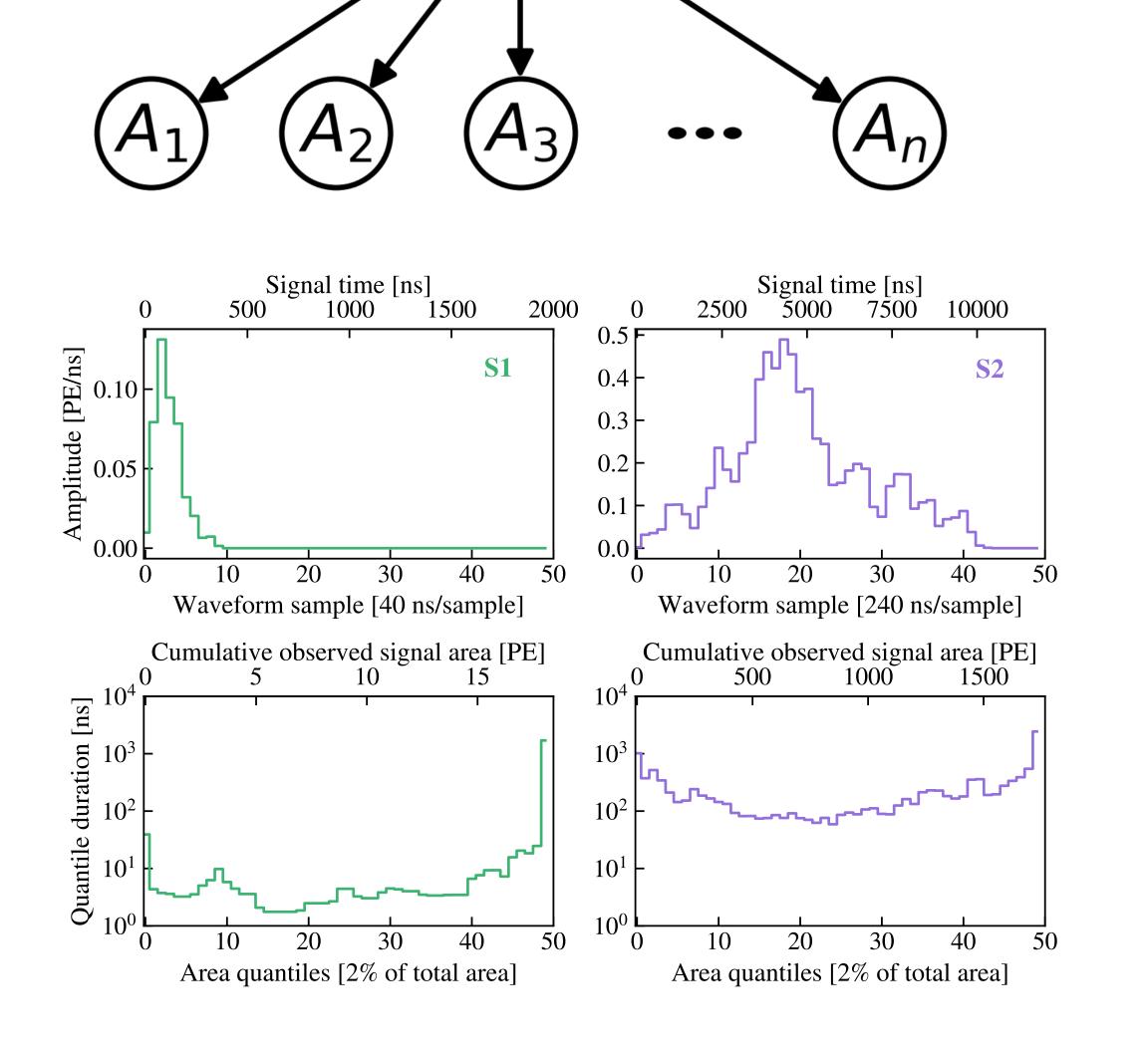
**Input: waveform attributes** 

Output: Pseudo probability of being signal-like





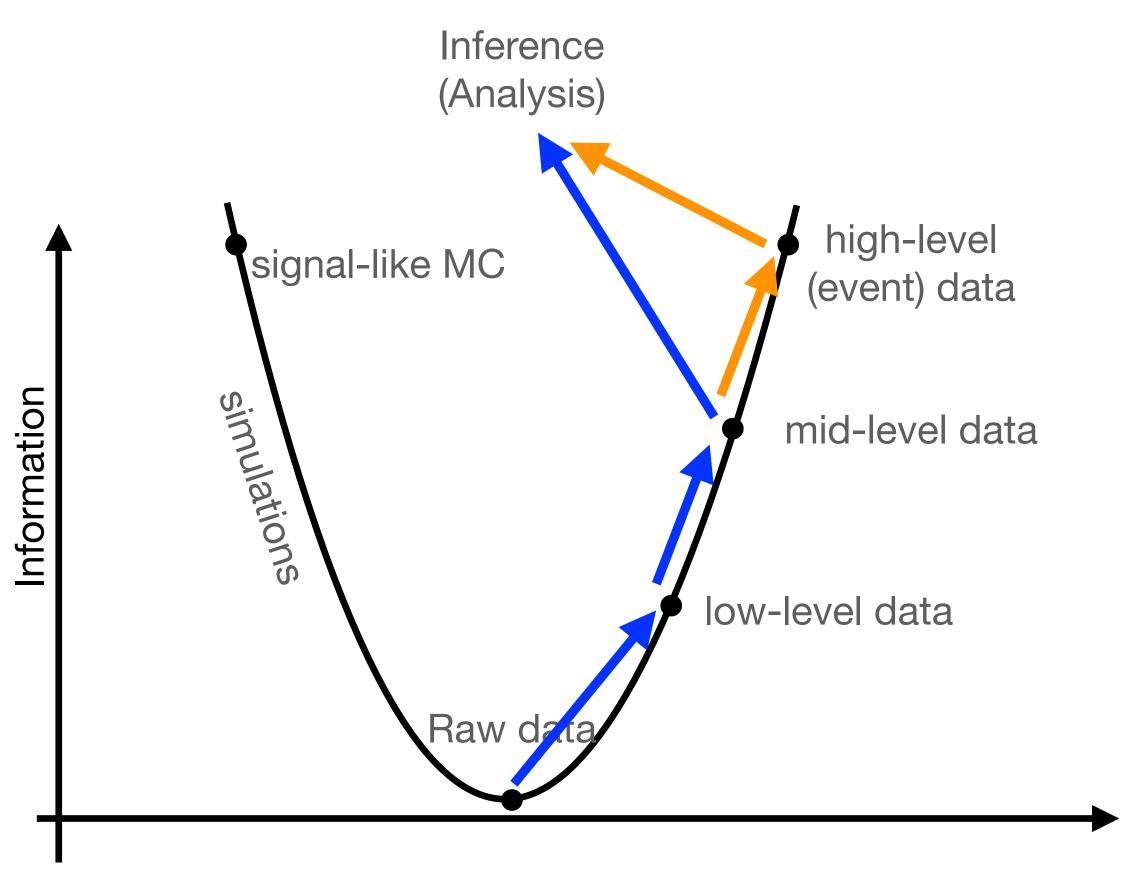


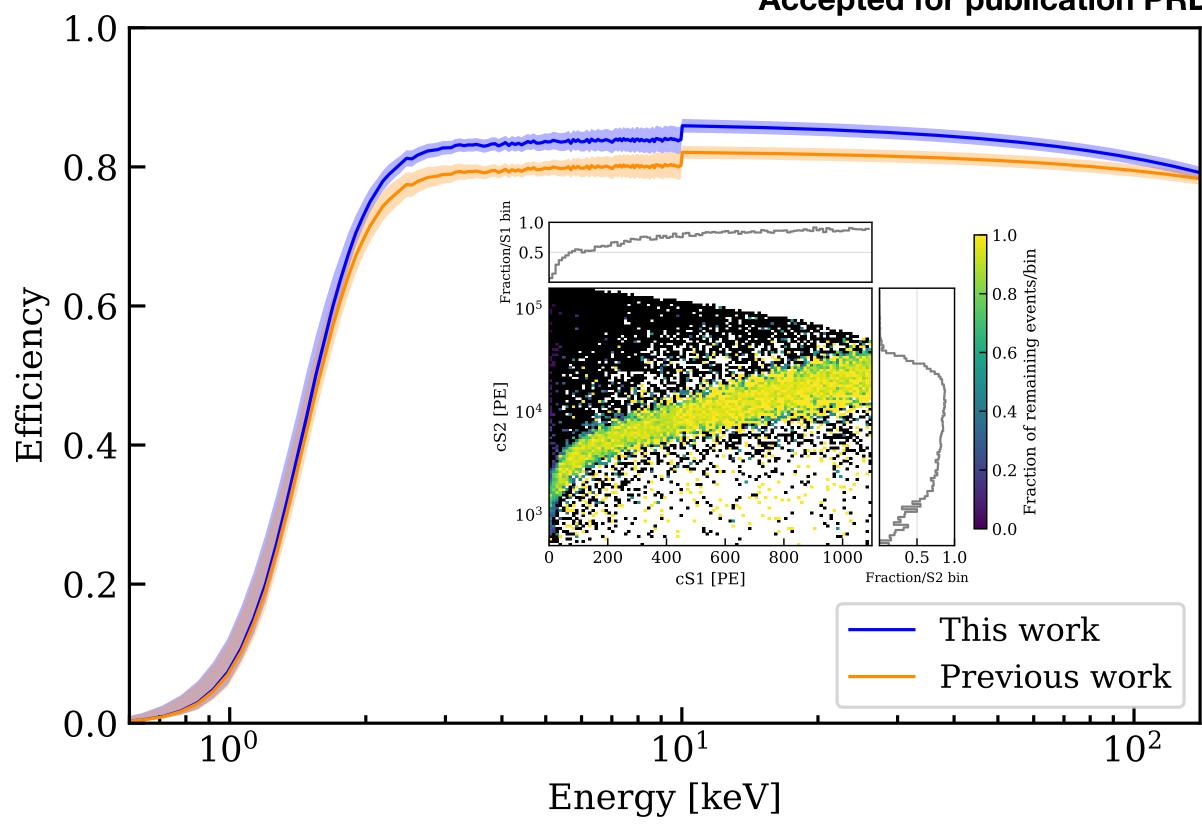


arxvi:2304.05428

**Accepted for publication PRD** 

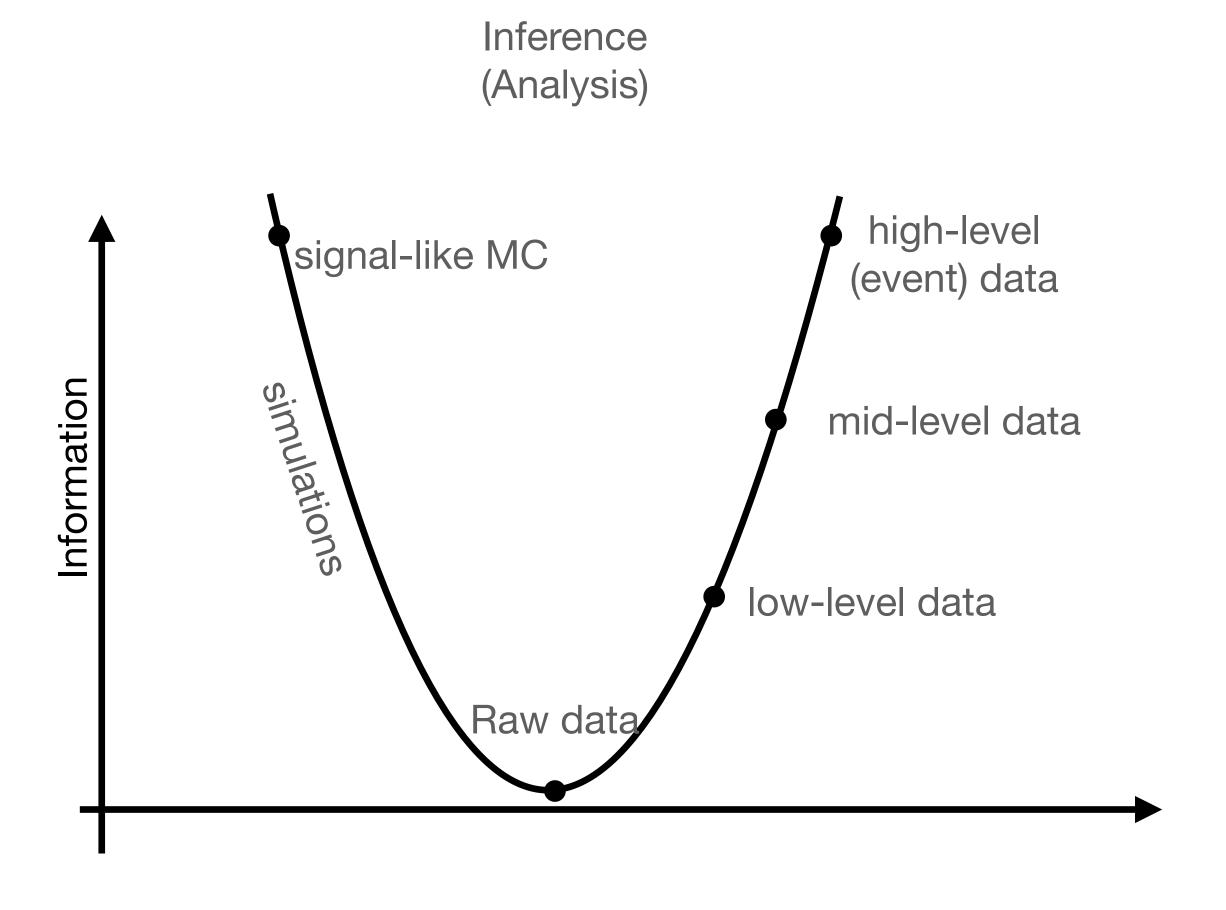
arxvi:2304.05428
Accepted for publication PRD





Using NBC based metric in event selection reduces significantly detector signals outside of the ROI and solves the need for additional selection criteria based on detector signal quality

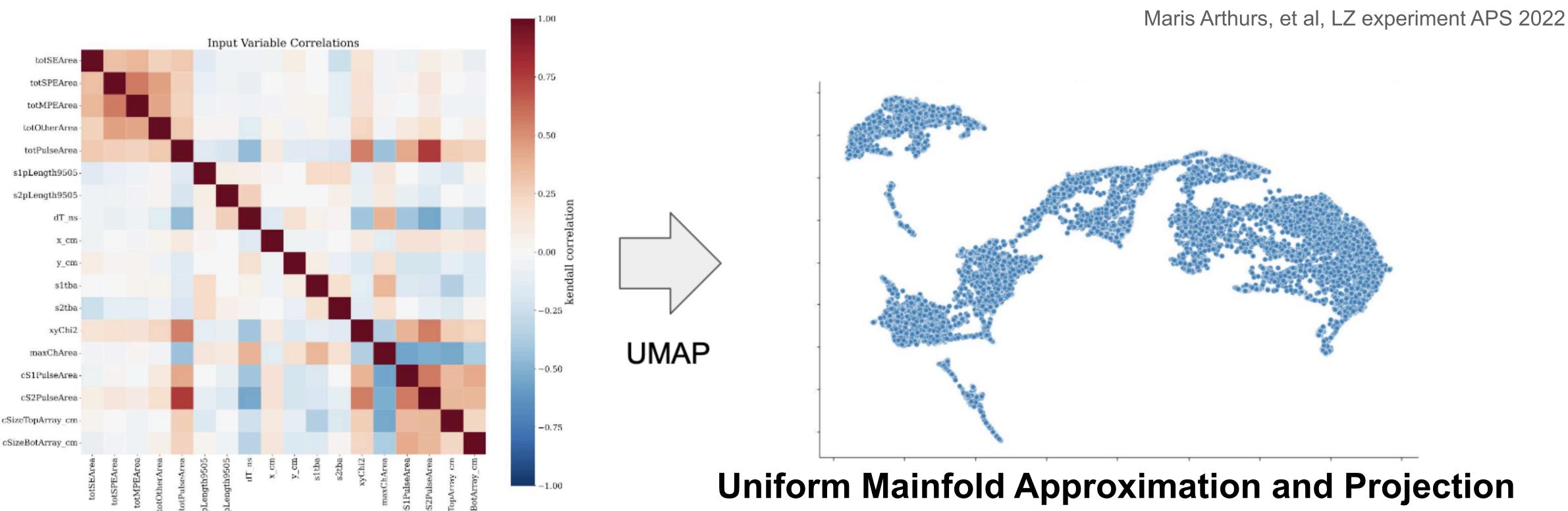
Increase the selection efficiency from previous method



#### Identify and interpret anomalous data

Using high-dimension spaces with a combination of high-level data and mid-level data

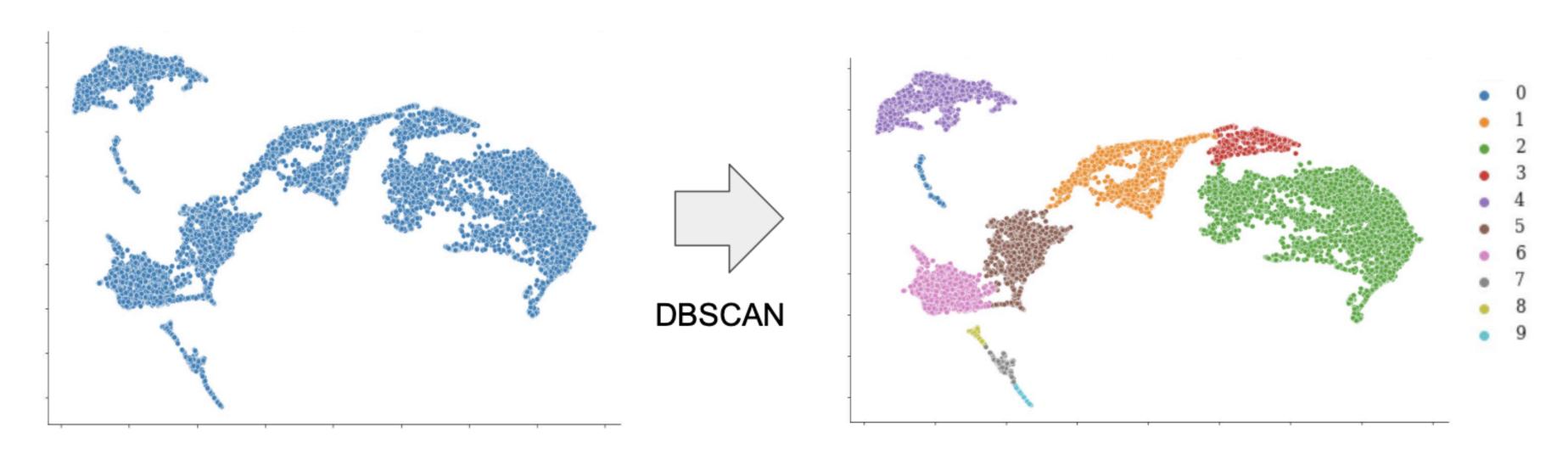
With dimensional reduction and/or clustering



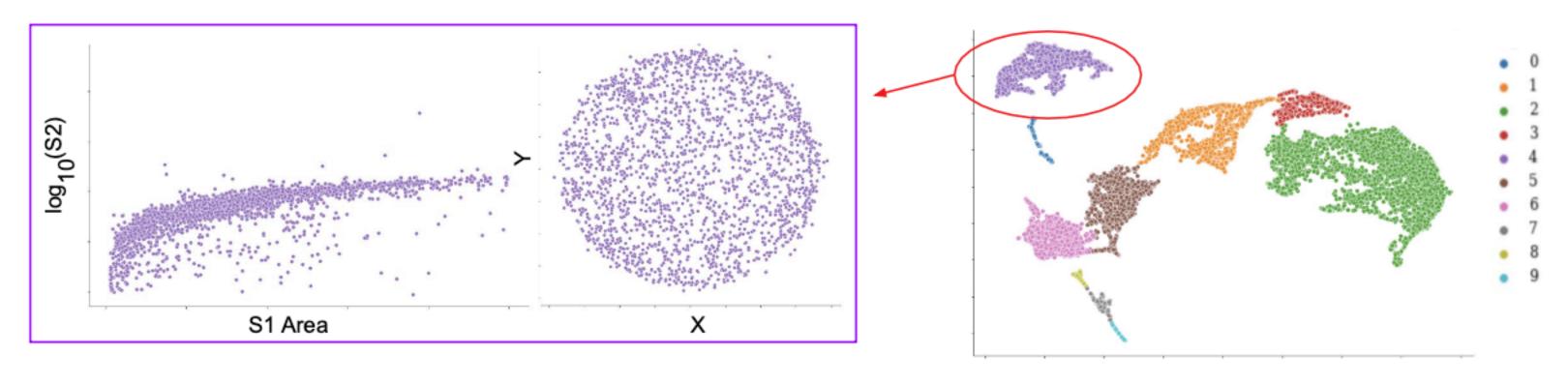
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A novel manifold learning technique for dimension reduction.

UMAP is an algorithm for dimension reduction based on manifold learning techniques and ideas from topological data analysis, it preserves more of the global structure with superior run time performance (arxiv:1802.03426)



Identify different topologies, e.g. gas events (detector effects)



Maris Arthurs, et al, LZ experiment APS 2022

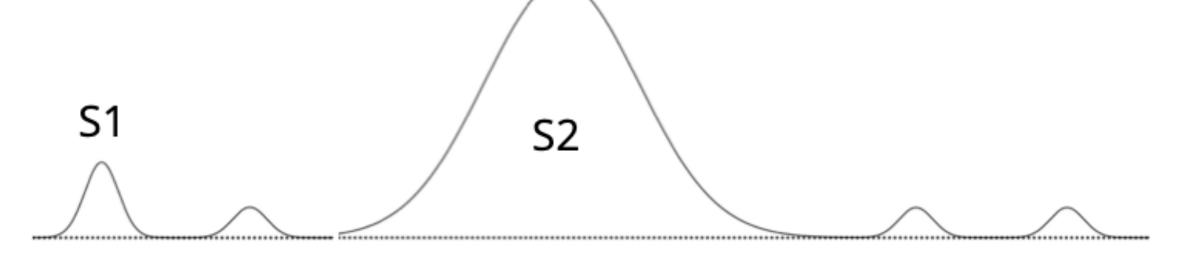
Luis Sanchez & Sanya Arora, Rice

#### Identify and interpret anomalous data

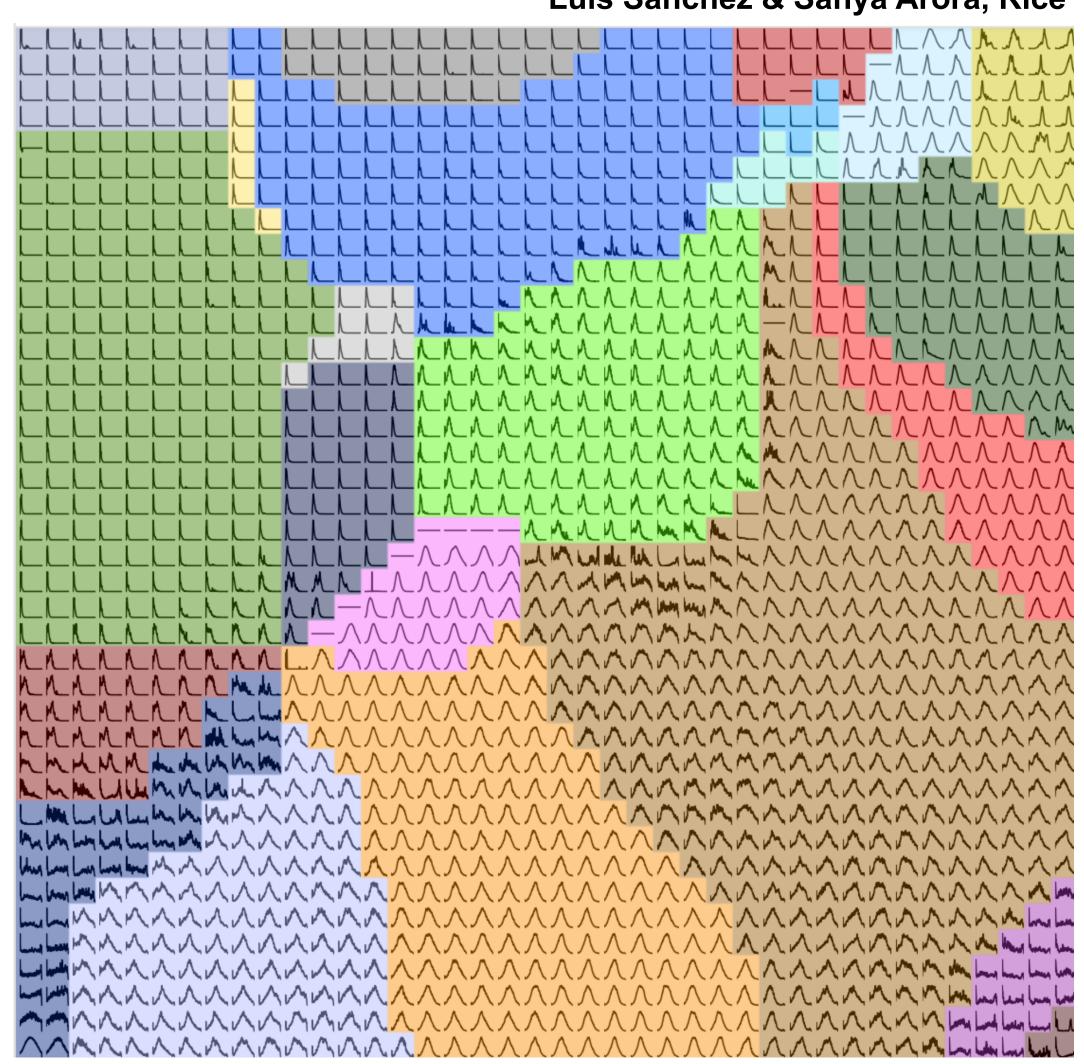
#### **Self Organizing Maps**

An unsupervised technique using competitive training to produce a low-dimensional representation of a higher dimensional data set while preserving the topological structure of the data

Input data: deciles + fraction area per sensor at the top array and total areas

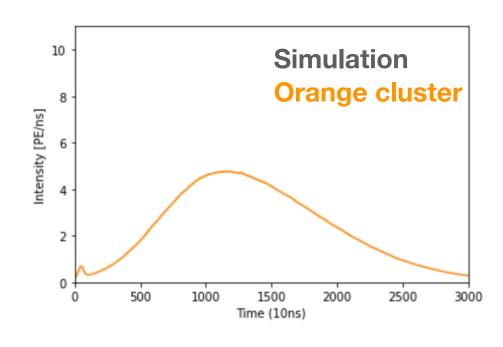


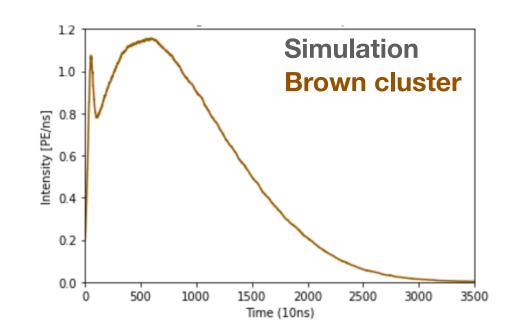
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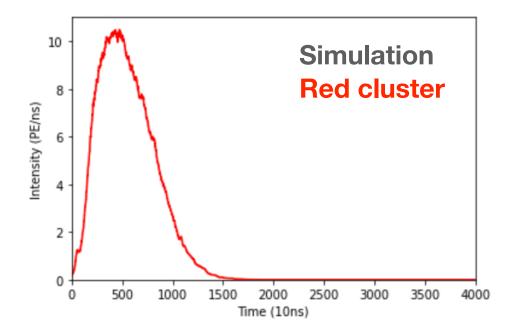


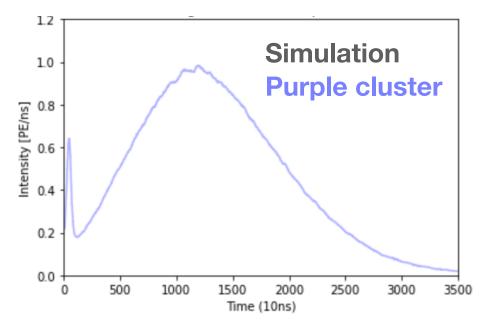
Luis Sanchez & Sanya Arora, Rice

### Identify and interpret anomalous data Self Organizing Maps

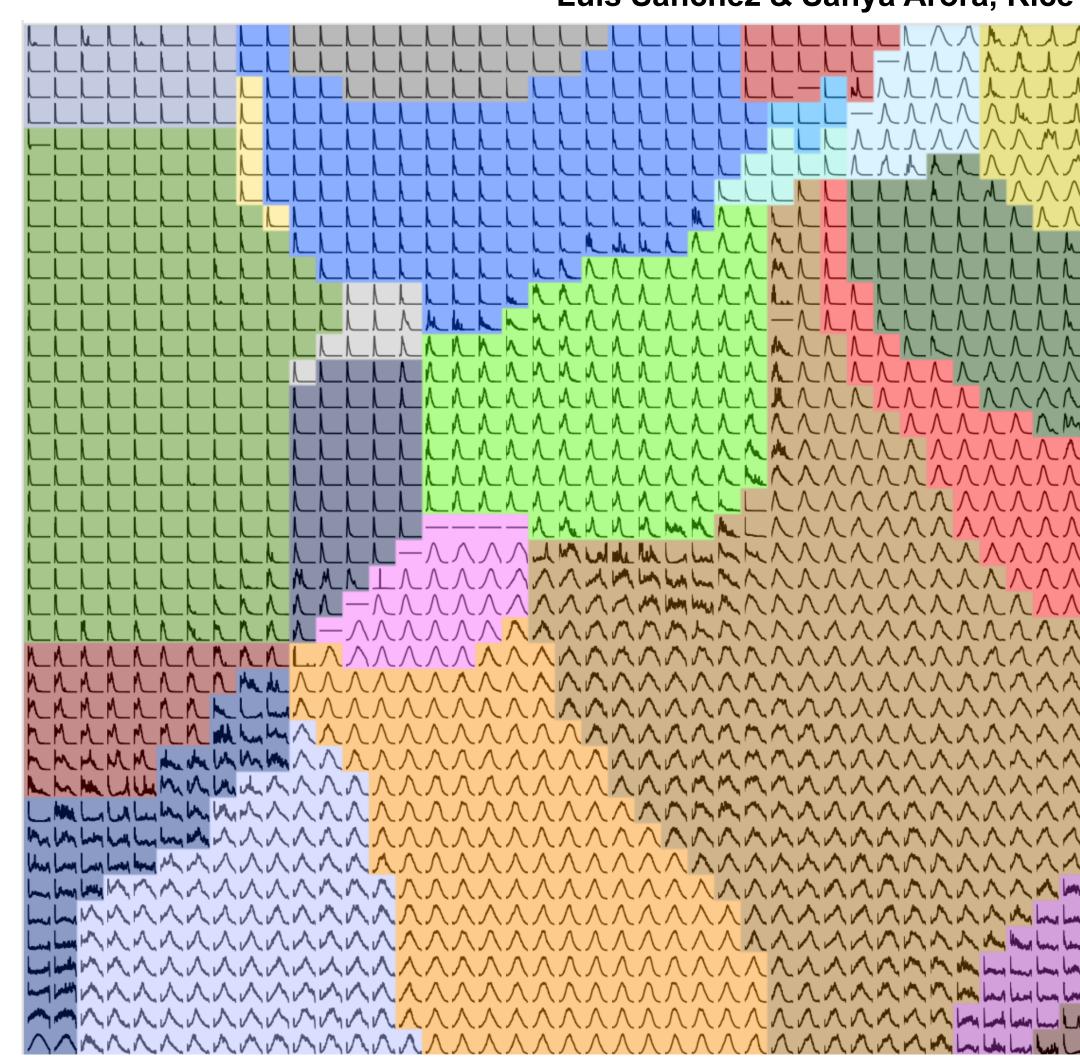






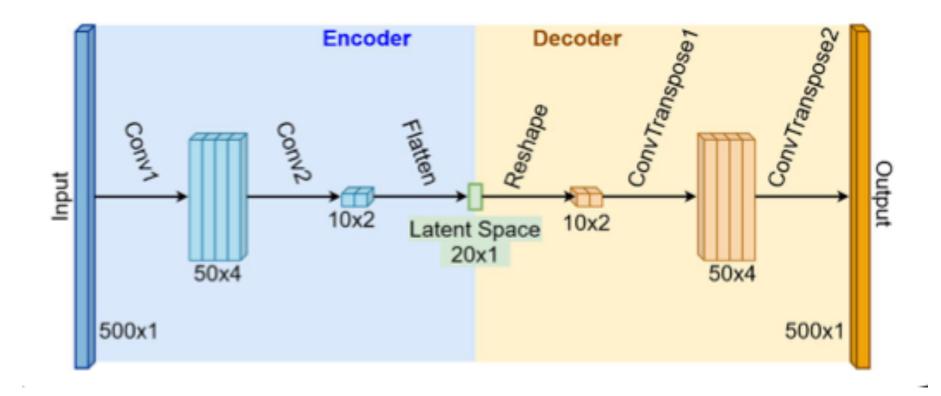


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#### Identify and interpret anomalous data

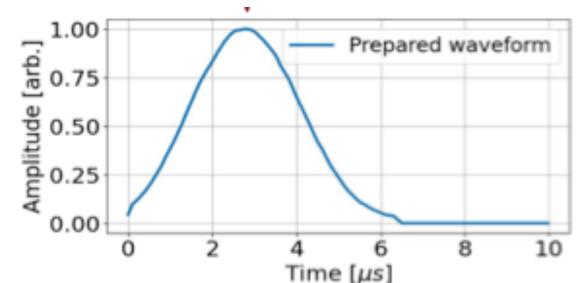
#### Autoencoders

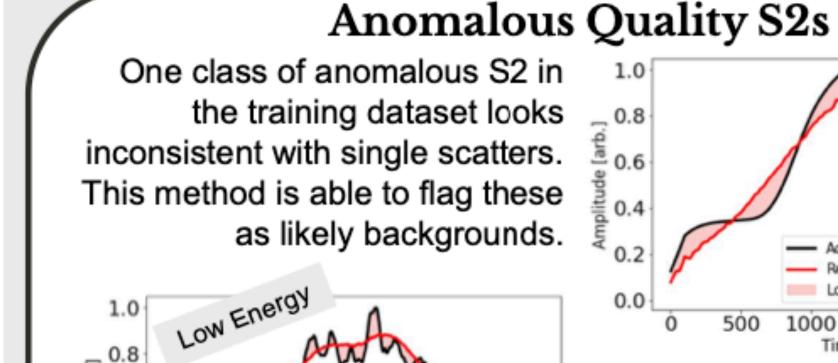


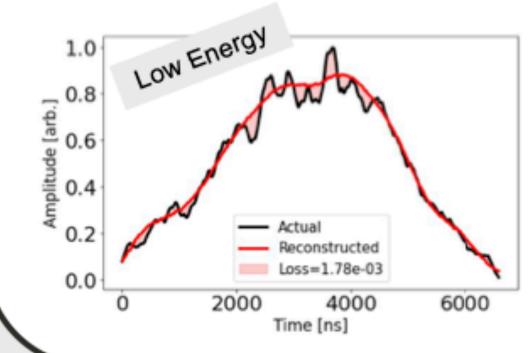
Waveform preparation:

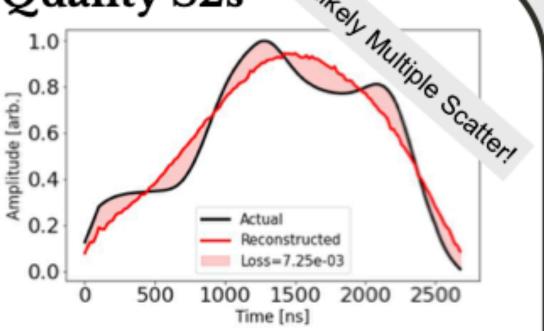
- Downsampled to 200 MHz
- Normalized by max amplitude
- Smoothed
- Zero-padded to 500 samples





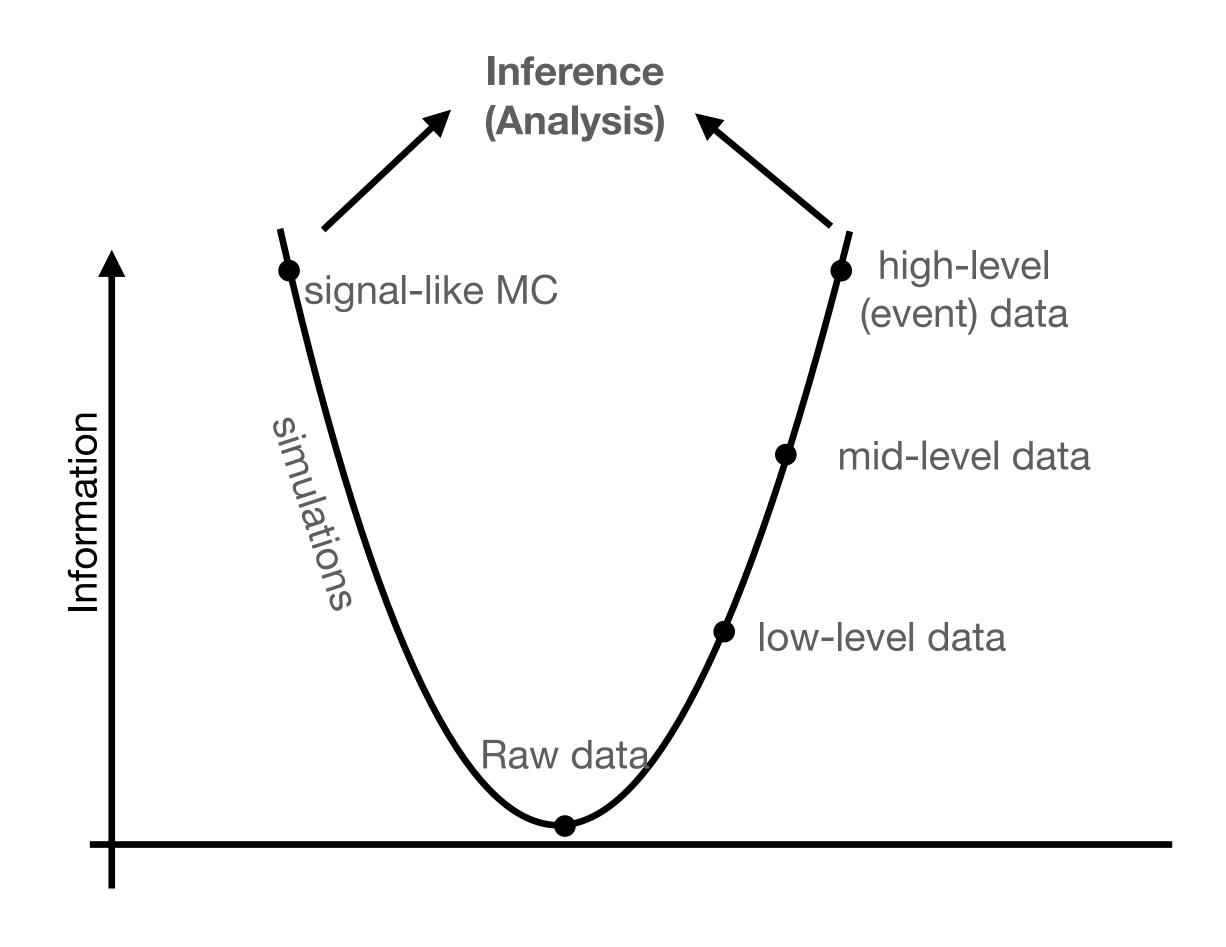






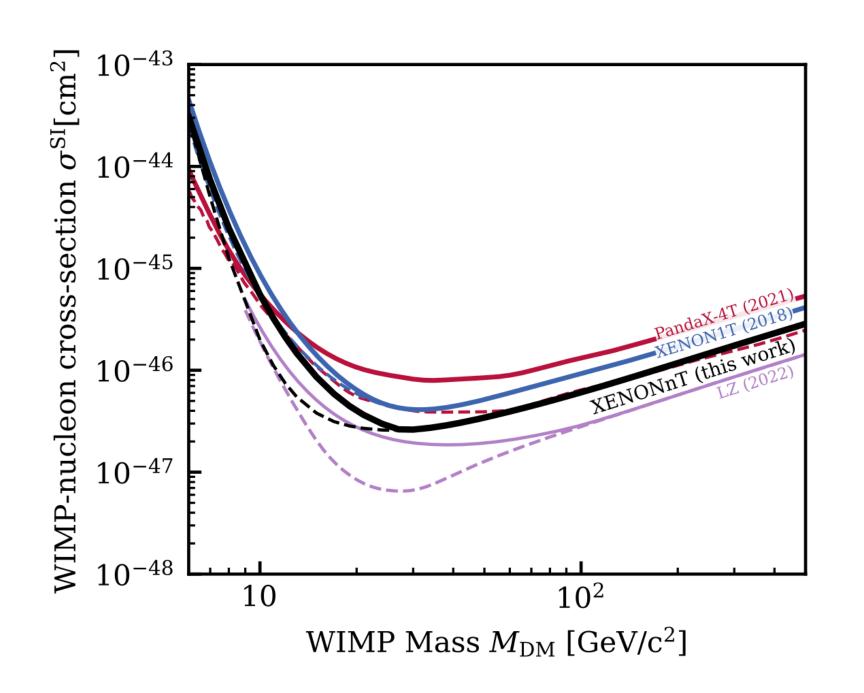
The second class comprises jagged low energy S2s. These are not necessarily backgrounds and show the limitations of this method with low energy S2s.

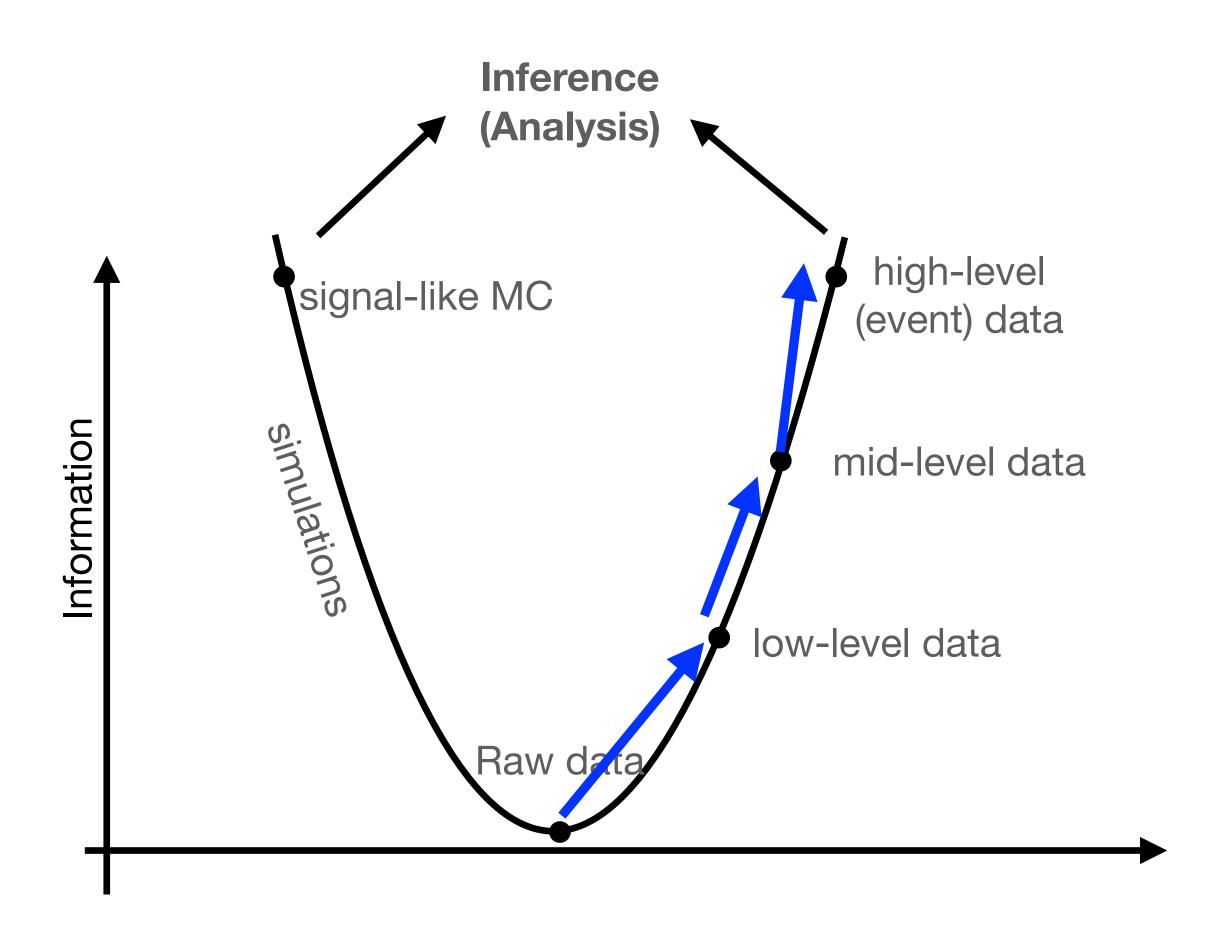
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#### A. Higuera, MODE workshop, Princeton University, July 2023

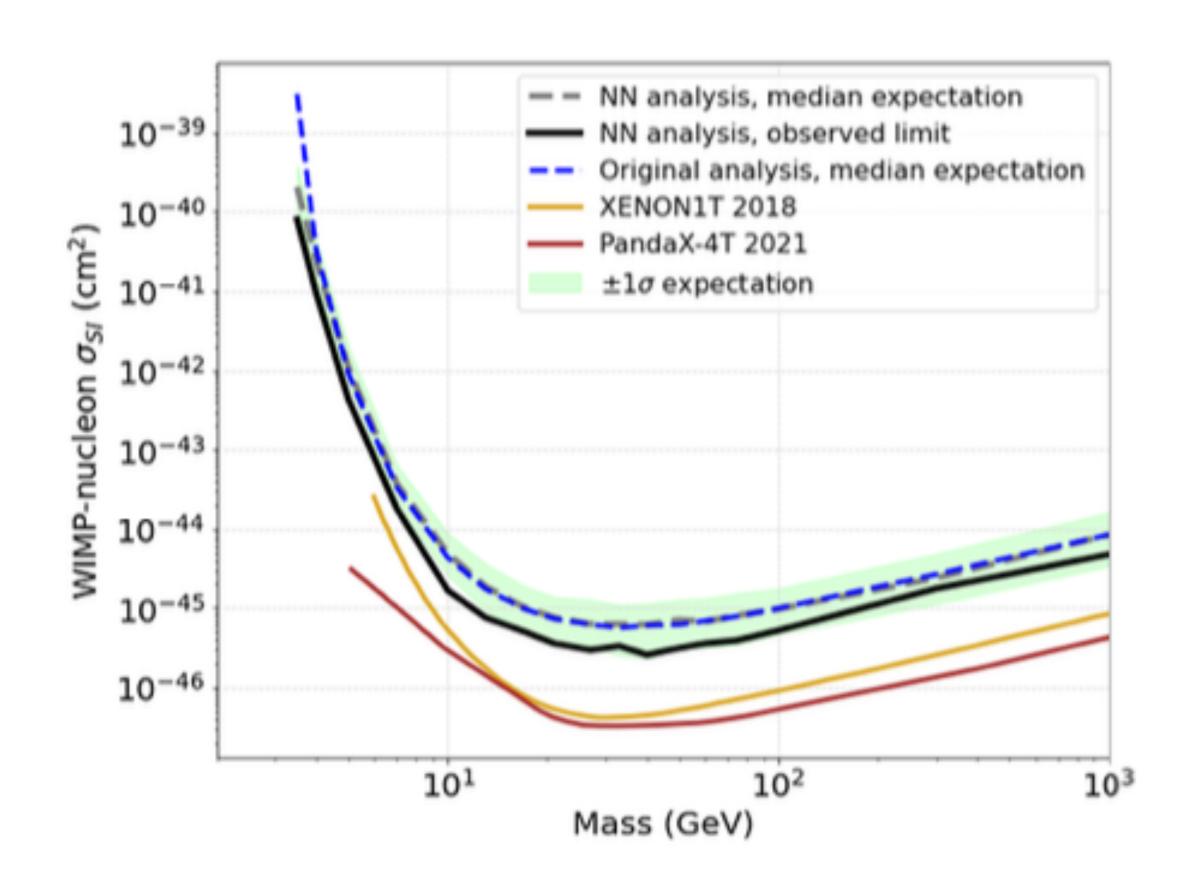
- We are searching for processes that have been predicted by a model but not yet seen
- If no observation we report the expected significance (sensitivity)





- We are searching for processes that have been predicted by a model but not yet seen
- If no observation we report the expected significance (sensitivity)
- The most common approach is calculating a profile likelihood ratio ... significant computational burden

arxiv:2201.05734

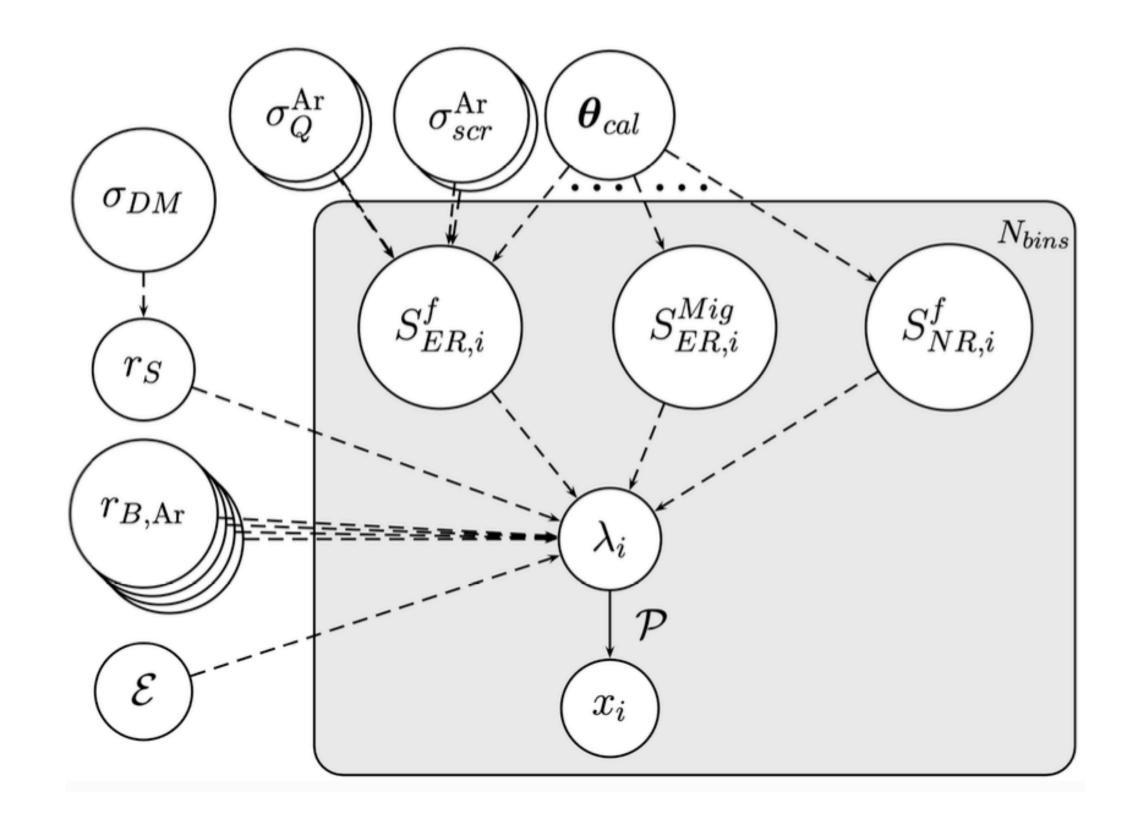


# Fast and flexible analysis of direct dark matter search data with machine learning

- Combining ML with profile likelihood fit procedure
- ML is more efficient at capturing information to create models in the form of PDFs
- Faster computation time: ML collapses multiple variables for signal/background discrimination into only one 1D discriminant, taking much of the work out of generating PDFs
- Evaluation of systematic uncertainties through the variation of nuisance parameters in the fit is made more feasible due to the faster calculation and lower statistics requirements

A. Higuera, MODE workshop, Princeton University, July 2023

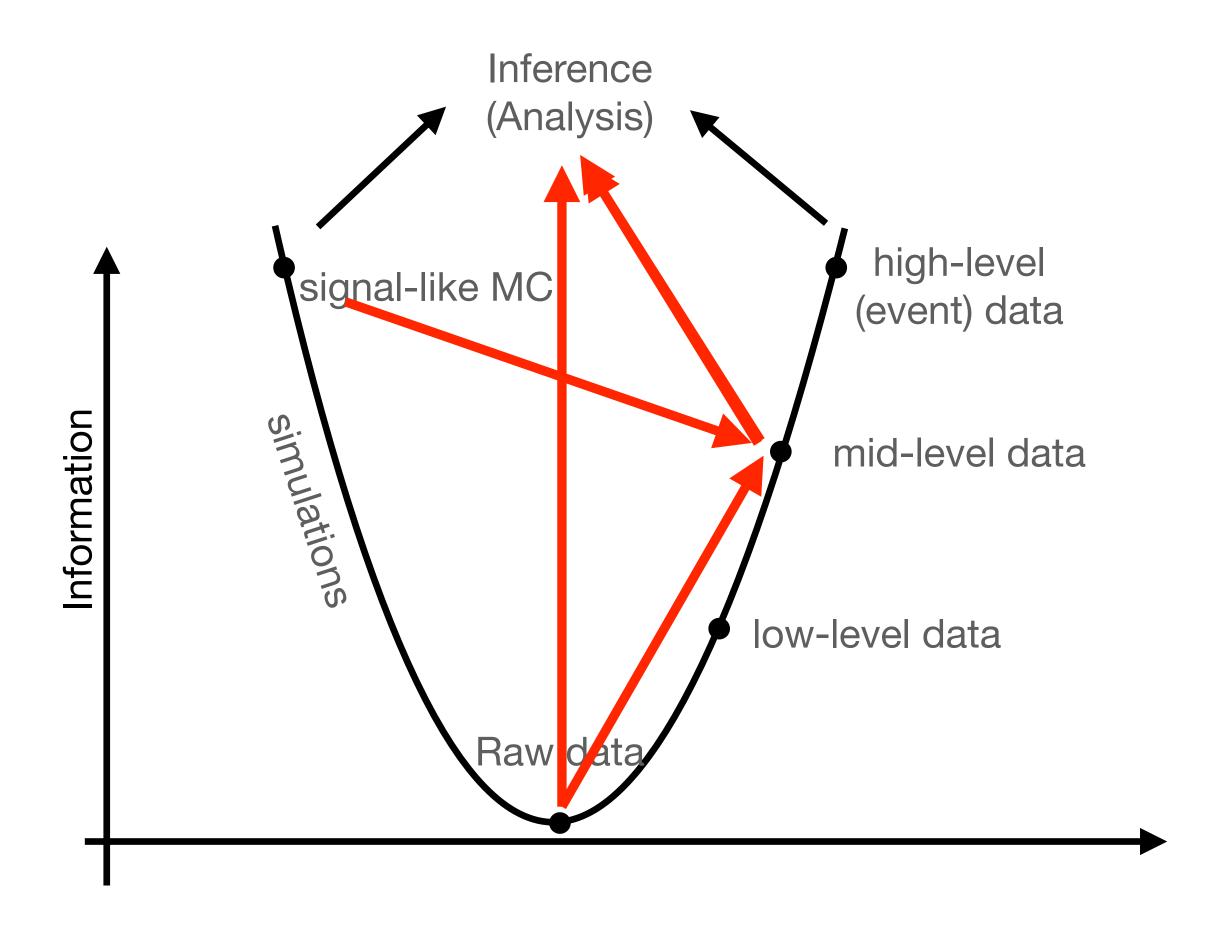
arxiv:2302.018302



# Search for low mass dark matter in DarkSide-50: the bayesian network approach

- The method incorporates the detector response model into the likelihood function
- No assumptions about the linearity of the problem or the shape of the probability distribution functions are required
- This method keeps the dependence of the parameter of interest on the response model parameters, allowing also a possible constraint on the response model

### Discussion



- There is a strong science case to search for dark matter
- The DM community is pushing the boundaries(low-background & lowthreshold experiments)
- Data analysis are become more complex and ML is the tool to tackle complex problems
- We physicists have domain knowledge and should leverage that when using ML

### Extas

2203.08084

### Neutrinos

 Neutrinos will become a dominant "background"

