

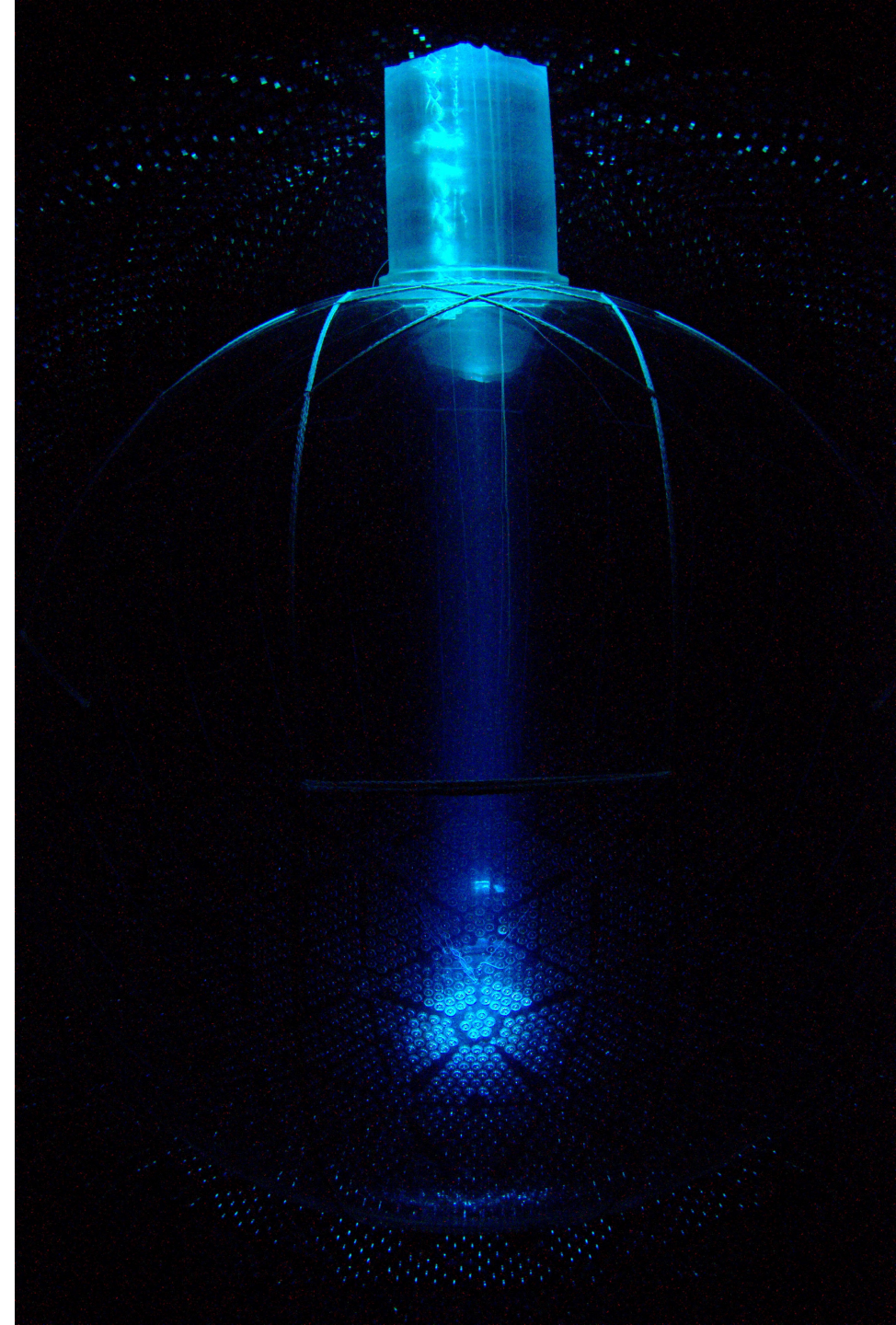


at the SNOLAB – Future of Neutrinoless Double Beta Decay,
2nd International Summit

Mark Chen

Queen's University and CIFAR

April 27, 2023

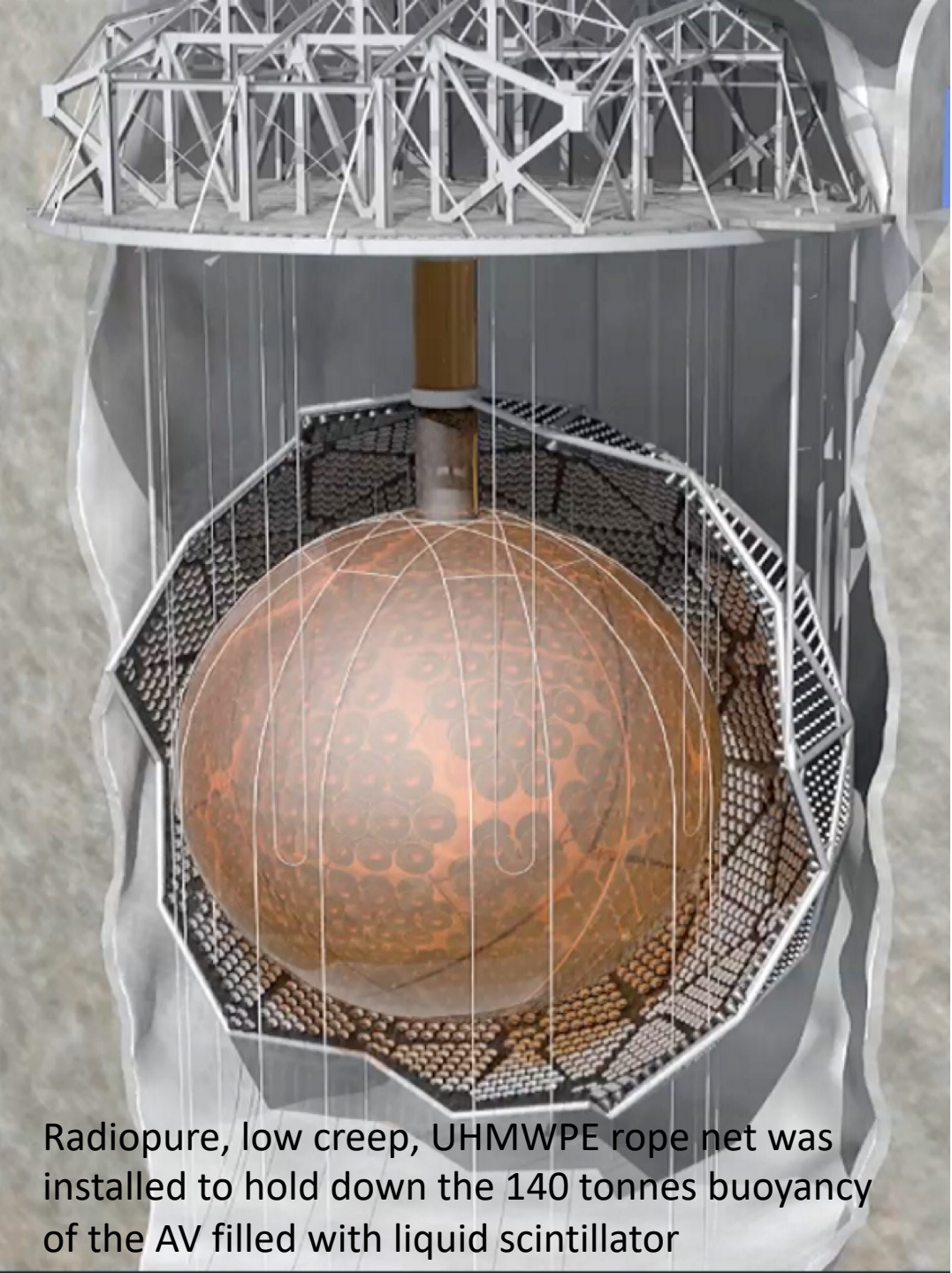
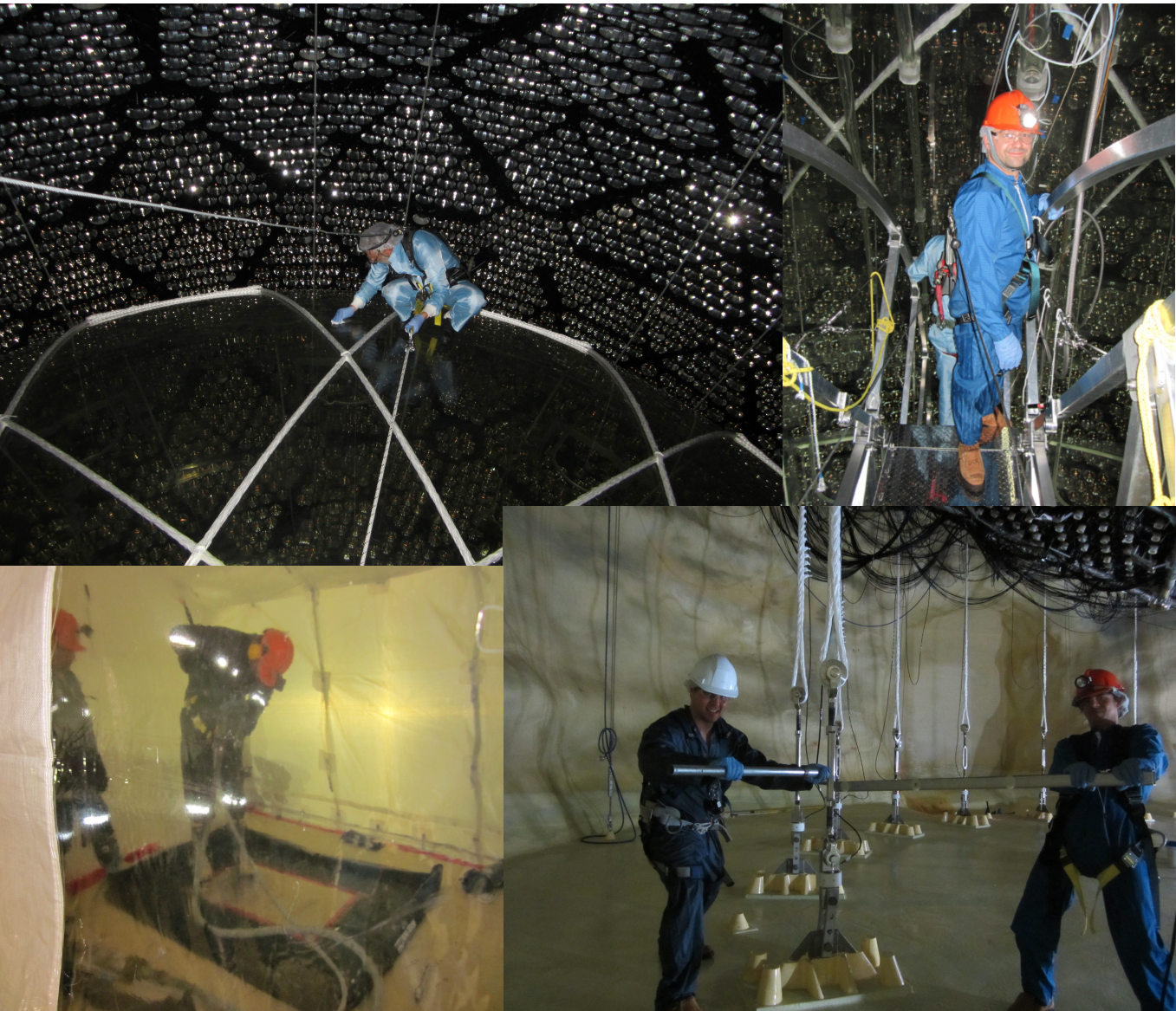


SNO+

August 9, 2022



SNO+ is the Sudbury Neutrino Observatory Filled with Liquid Scintillator



Radiopure, low creep, UHMWPE rope net was installed to hold down the 140 tonnes buoyancy of the AV filled with liquid scintillator

Restorations

SNO was a classic, but even classics need some restorations...

- SNO Cavity floor liner had been badly torn at the end of SNO; had to be remade (during SNO+ hold-down anchor installation)
- Anchor plate installation involved *drilling* into concrete and rock *inside an ultra-low background neutrino detector*
- Submersible pump that drained the SNO AV had self-destructed, covering the inner AV bottom with dirty oil
- SNO Cavity wall liner had many leaks – SNO+ had to find these pinhole leaks paddling around in the Cavity in a raft, in low-light conditions, using multiple leak hunting techniques...many months of effort!
- After all of this, would SNO+ still have low backgrounds?

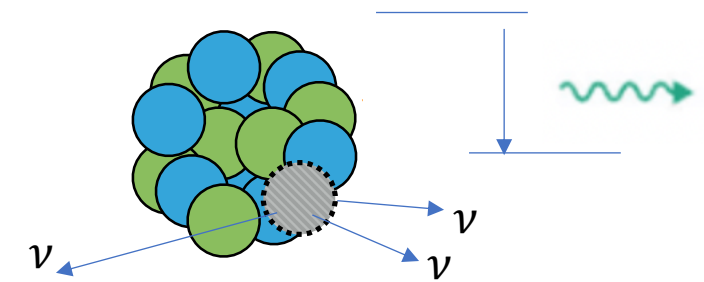


A close-up, high-angle shot of water ripples on a blue surface. The water is a deep, vibrant blue, and the ripples create a complex pattern of light and shadow. The text "SNO+ Water Phase" is overlaid in white, centered horizontally and slightly below the vertical center. The font is a clean, sans-serif typeface. The background shows several concentric ripples, with the largest one in the upper right quadrant and a smaller one in the lower right. The overall effect is a sense of movement and depth.

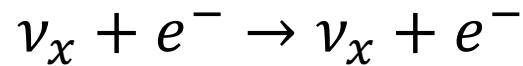
SNO+ Water Phase

SNO+ Water Phase Physics Results

- World's best limits on invisible modes of nucleon decay
 - 2022 update published in *Phys. Rev. D*



- Solar neutrinos
 - detected via neutrino-electron elastic scattering



PHYSICAL REVIEW D **99**, 012012 (2019)

Measurement of the ^8B solar neutrino flux
in SNO+ with very low backgrounds

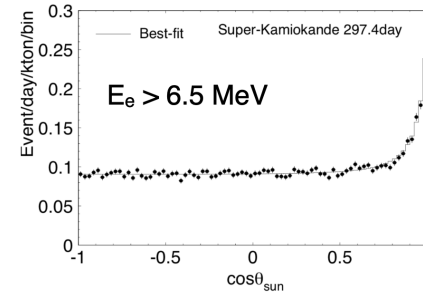
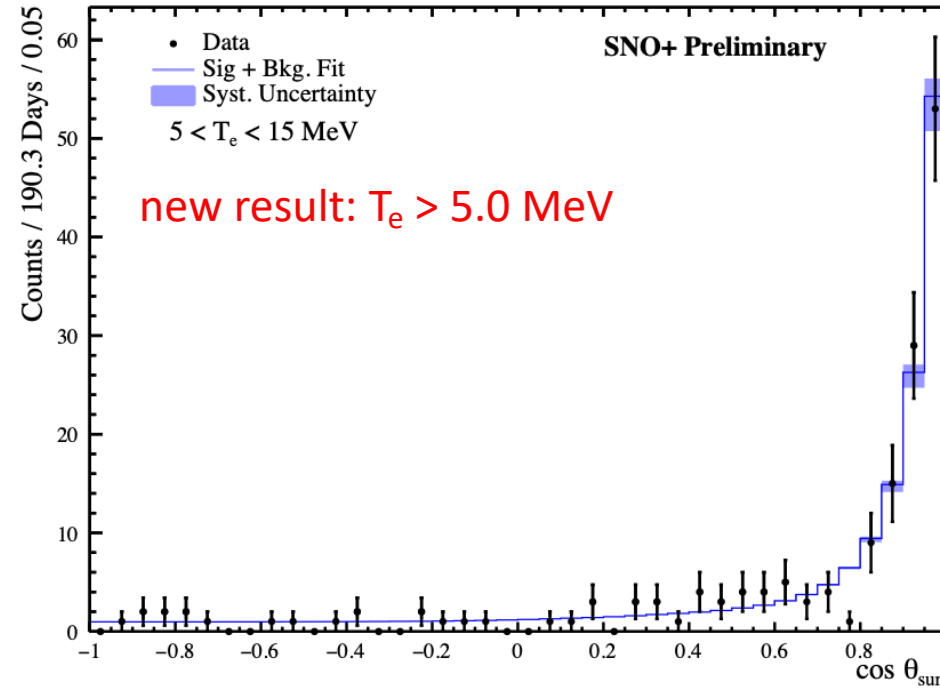
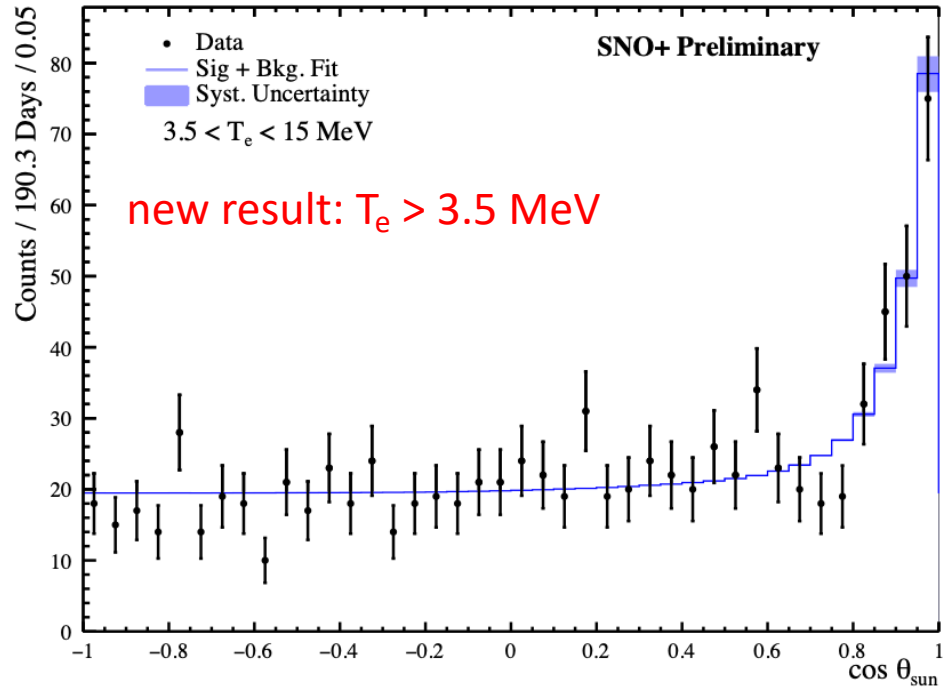
- now with *even lower backgrounds*

- First observation of reactor $\bar{\nu}_e + p \rightarrow e^+ + n$ events using *pure* water (undoped)
 - published in *Phys. Rev. Lett.* on March 1, 2023
 - made possible by $\sim 50\%$ neutron detection efficiency (highest in a water Cherenkov detector)

Decay Mode	Partial Lifetime Limit	Existing Limits
n	9.0×10^{29} y	5.8×10^{29} y [5]
p	9.6×10^{29} y	3.6×10^{29} y [6]
pp	1.1×10^{29} y	4.7×10^{28} y [6]
np	6.0×10^{28} y	2.6×10^{28} y [6]
nn	1.5×10^{28} y	1.4×10^{30} y [5]

world-leading limits set by SNO+

^8B solar neutrinos in SNO+ Water Phase



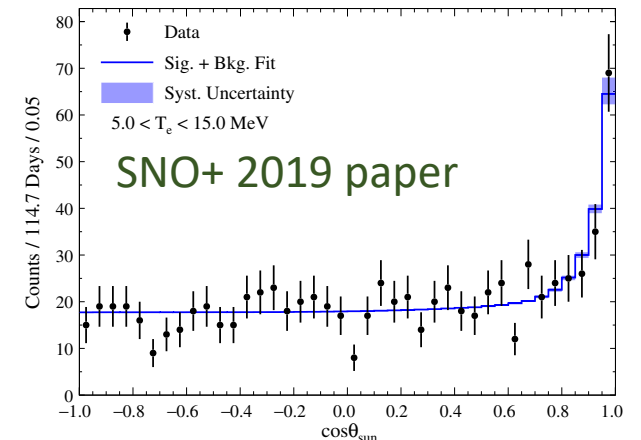
first Super-K
 result, for
 comparison

same 5.0 MeV
 threshold

Our latest results with data from the extended water phase, with $\sim 1/10$ Rn levels (new SNO+ cover gas system)

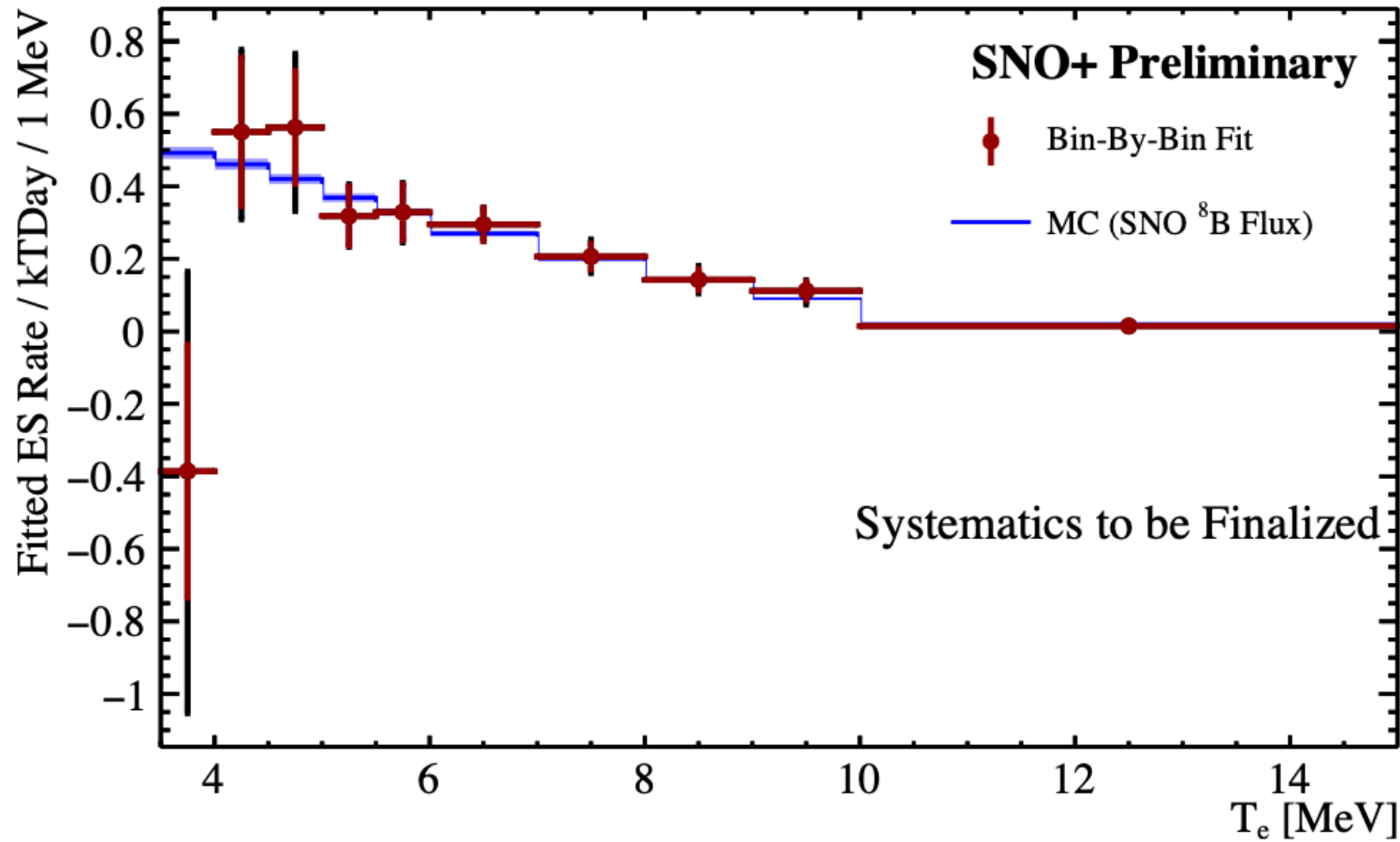
Answering the question: did we succeed to restore SNO+ as a low-background detector? YES!

Deep underground location 6000 m.w.e. greatly suppresses cosmogenic backgrounds.



SNO+ Water Phase ^8B Solar Neutrino Energy Spectrum

result as shown at Neutrino 2022



SNO+ Water Phase list of physics publications

- Set world-leading limits on invisible modes of nucleon decay, [PRD 99, 032008 \(2019\)](#); [PRD 105 112012 \(2022\)](#)
- “Measurement of the ^8B solar neutrino flux in SNO+ with very low backgrounds”, [PRD 99, 012012 \(2019\)](#)
- Highest efficiency (~50%) for neutron detection in a water Cherenkov detector, [PRC 102, 014002 \(2020\)](#)
- Detection of antineutrinos from distant reactors using only pure water, [PRL 130, 091801 \(2023\)](#)

PHYSICAL REVIEW LETTERS **130**, 091801 (2023)

Editors' Suggestion

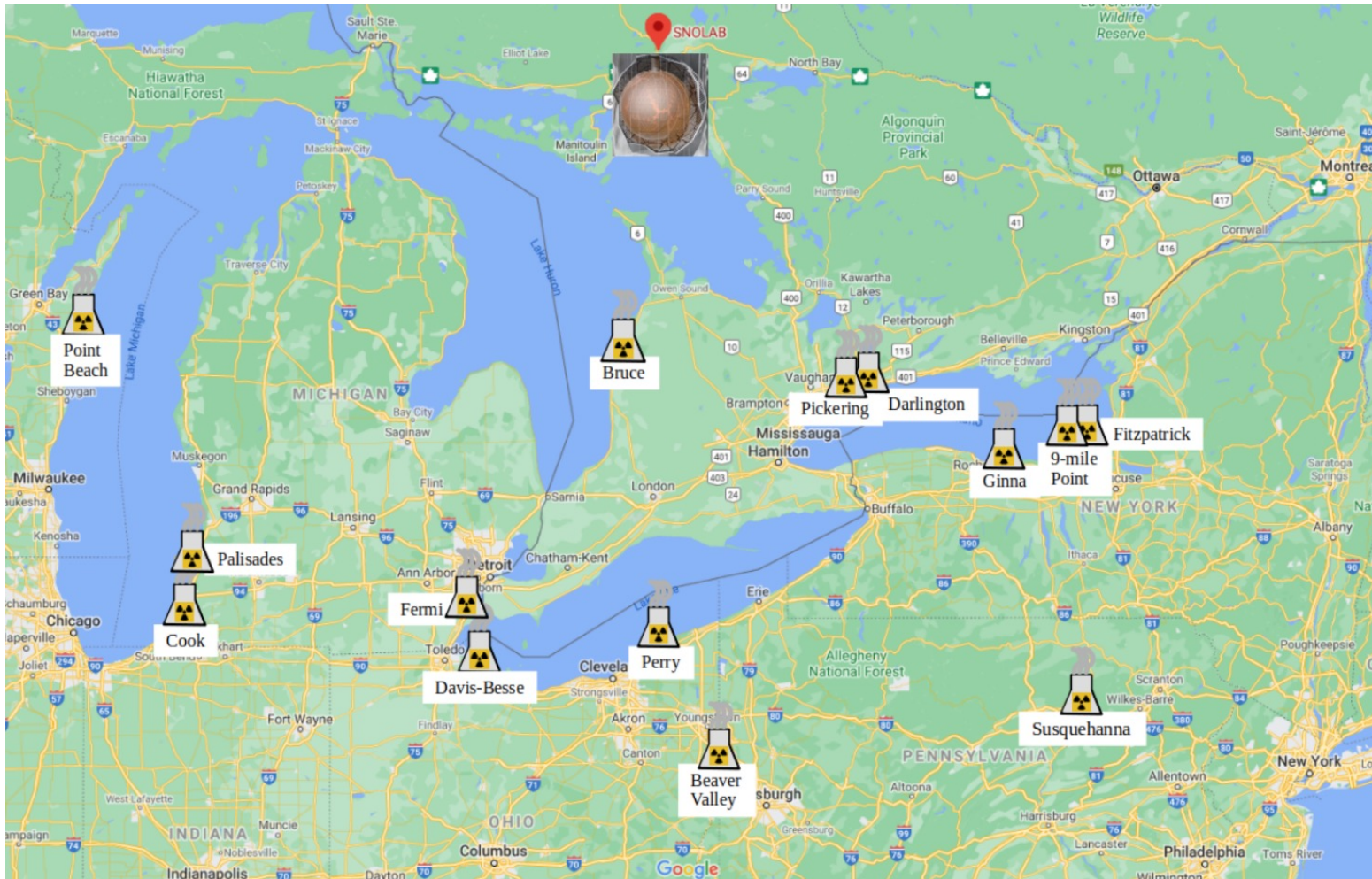
Featured in Physics

Evidence of Antineutrinos from Distant Reactors Using Pure Water at SNO +

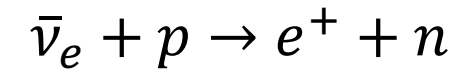
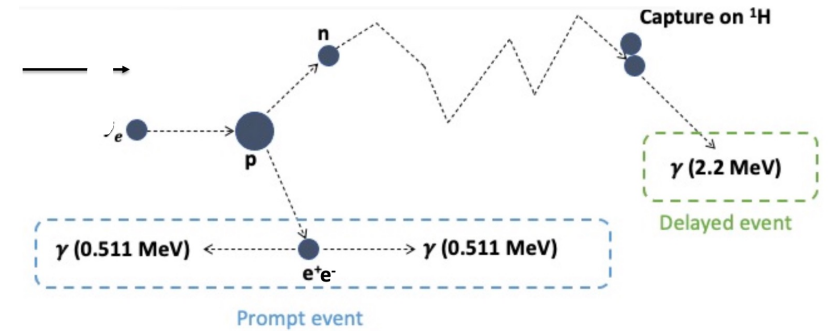
technical papers

- SNO+ “Detector Paper” [JINST 16, P08059 \(2021\)](#)
- SNO+ Scintillator Paper “Development, characterization and deployment of the SNO+ liquid scintillator” [JINST 16, P05009 \(2021\)](#)
- Water Phase optical calibration [JINST 16, P10021 \(2021\)](#)

Reactor Antineutrinos in SNO+



Inverse Beta Decay (IBD)



Coincidence event

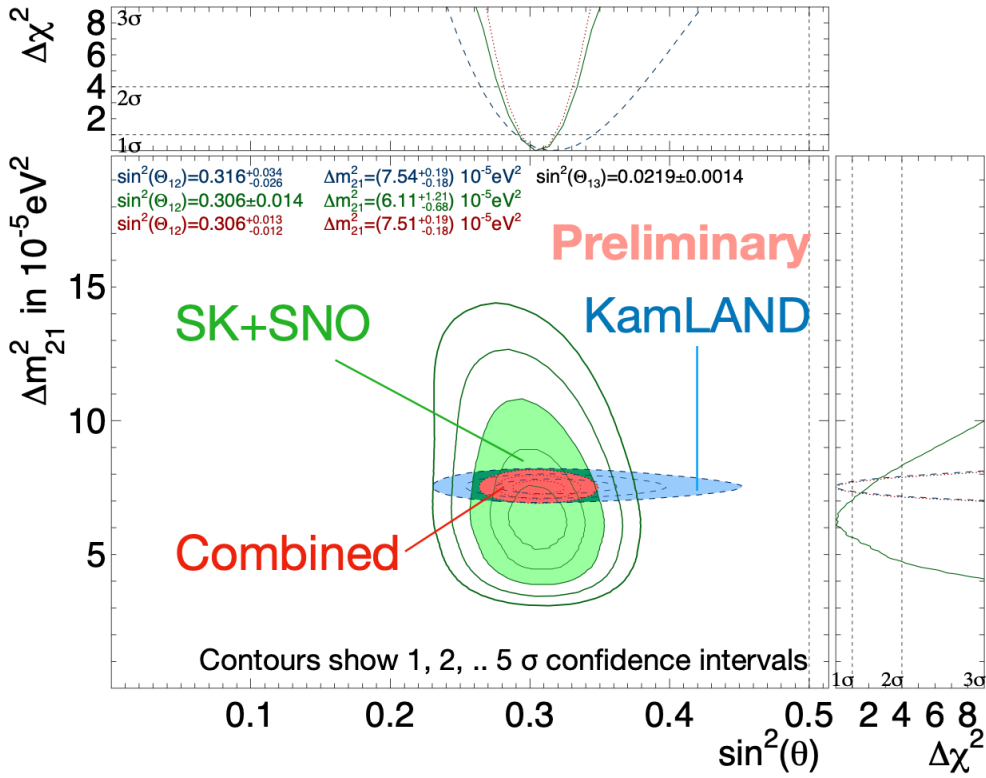
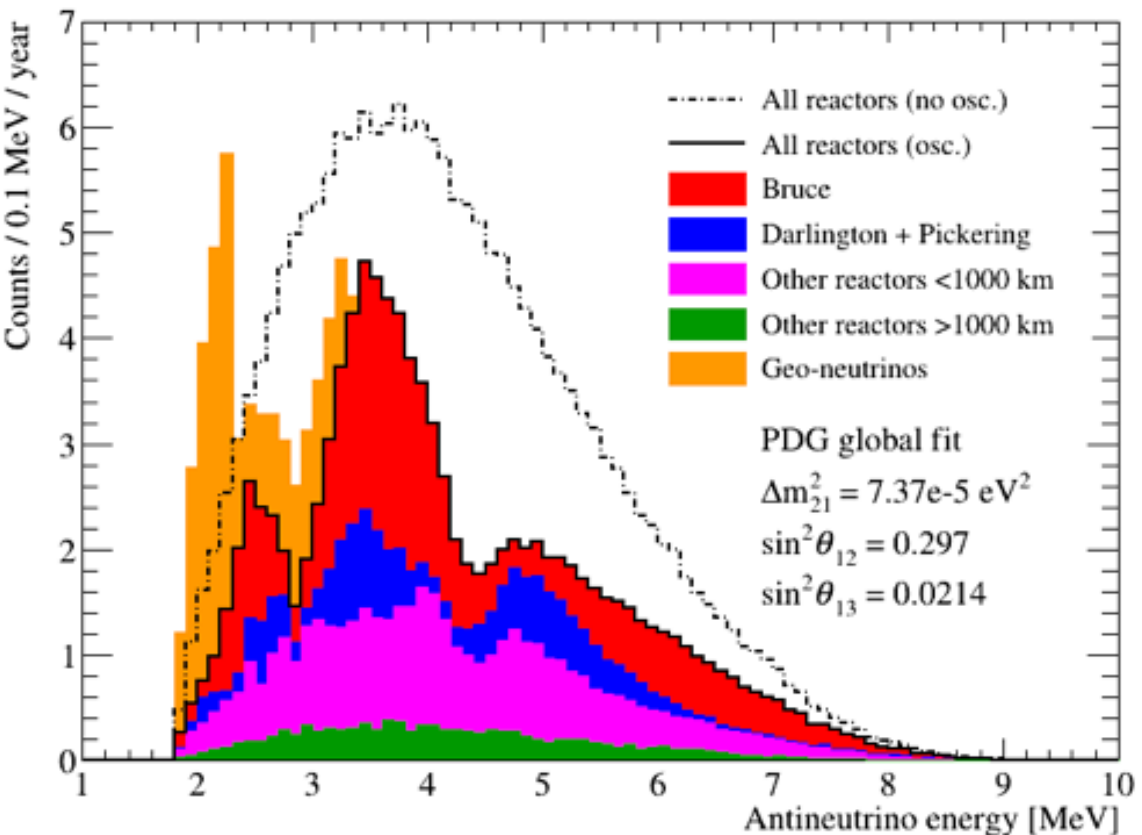
Prompt – positron kinetic energy (several MeV)

plus 1.022 MeV from annihilation γ 's

Delayed – neutron capture 2.2 MeV γ

Antineutrinos in SNO+ Scintillator

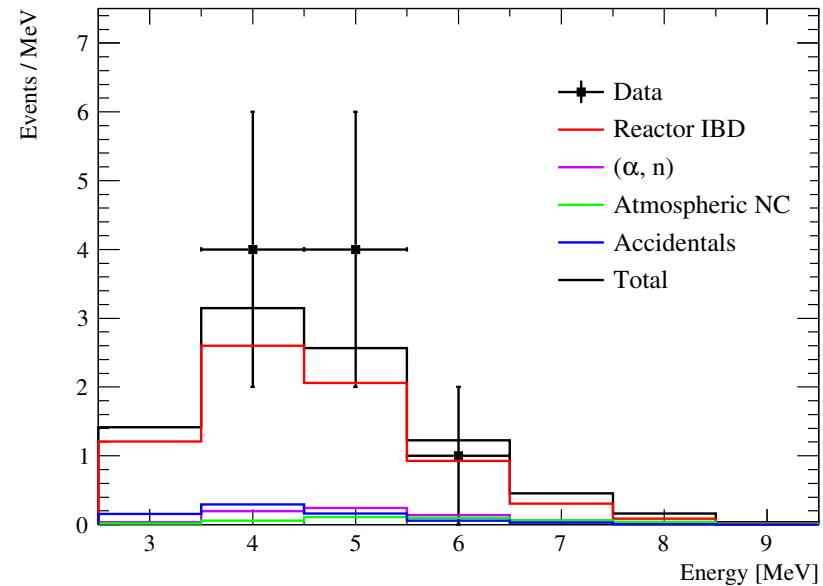
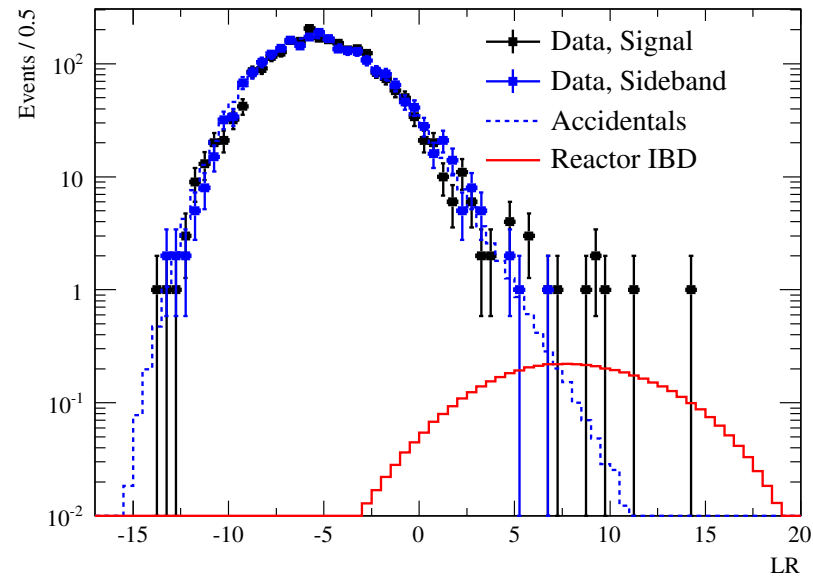
Scintillator Phase – Reactor antineutrino oscillations Δm_{12}^2 (plus geo neutrinos) are one of the main science goals of the Scintillator Phase



Antineutrinos in SNO+ Water? Yes!

Water Phase – Detection of IBD events (reactor antineutrinos) using pure water → **this is a first**

Two independent analyses – likelihood ratio and Boosted Decision Tree – both with 3σ detection significance; using event selection overlap + non-overlap, calculated combined discovery significance of 3.5σ



Science Press

SYNOPSIS

Reactor Neutrinos Detected by Water

March 1, 2023 • Physics 16, s28

Researchers have captured the signal of neutrinos from a nuclear reactor using a water-filled neutrino detector, a first for such a device.



SNO+ Collaboration

NORTHERN ONTARIO
Sudbury SNOLAB makes scientific breakthrough in monitoring nuclear power

Home / Physics / General Physics



APRIL 11, 2023 FEATURE

Editors' notes

The SNO+ collaboration gathers the first evidence of antineutrinos in a water Cherenkov detector

by Ingrid Fadelli, Phys.org

- Facebook 205
- Twitter 37
- LinkedIn Share
- Email



- Facebook
- Twitter
- LinkedIn
- Email
- Print

Reactor antineutrinos detected in pure water in an experimental first

28 Mar 2023



Reactor reactions: the SNO+ detector has seen antineutrinos from distant reactors when it was filled with pure water. (Courtesy: SNO+)

SNO+ captures first reactor neutrinos detected by water

SNO+ Scintillator Purification Plant

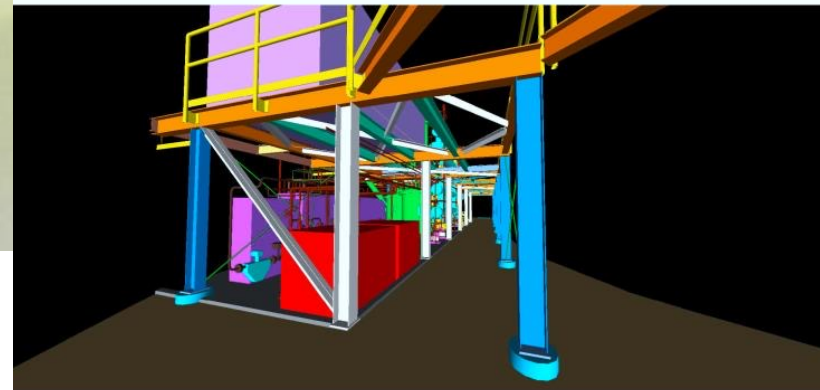
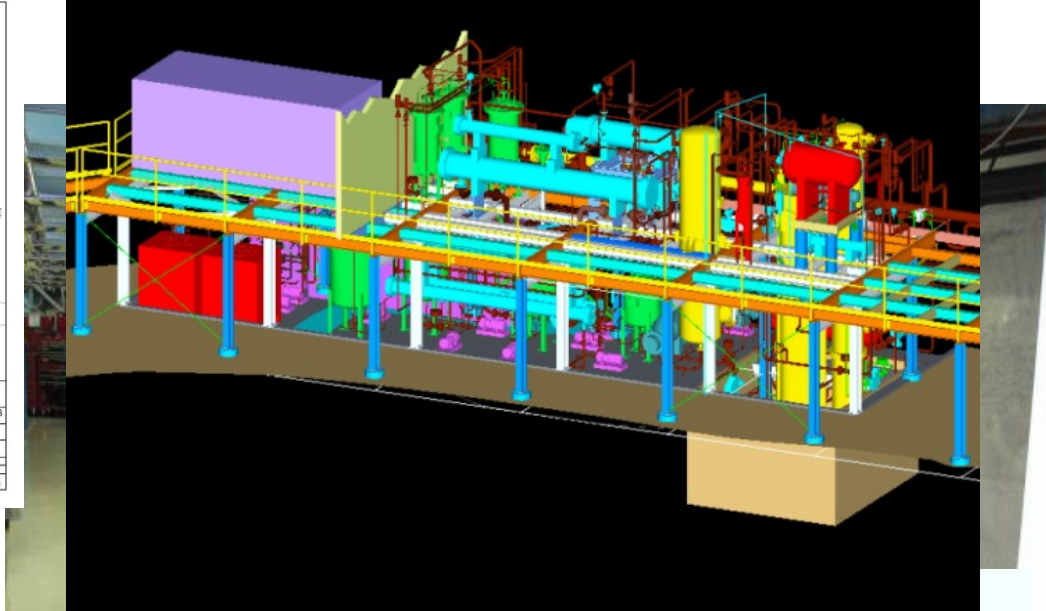
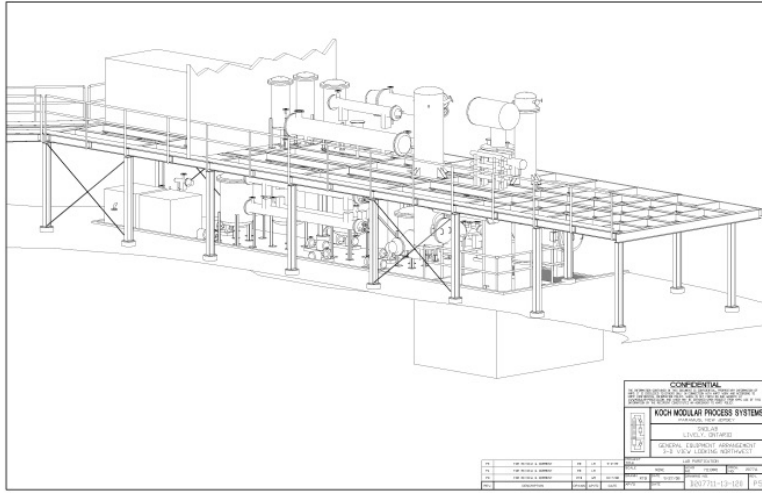
- reinforced mezzanine steel
- made D2O pit deeper “mining in a clean room”
- installed columns, vessels, heat exchangers, tank, pumps, valves, high-grade sanitary piping (orbital-welded, electropolished stainless steel tubing)
- utility plumbing (cooling water, compressed air, vent, boil-off nitrogen)
- process control, wiring, instrumentation, electrical
- firewalls, fire detection and suppression



the SNO heavy water purification system was here

SNO+ Scintillator Purification Plant

- rein
- ma
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- stainless steel tubing)
- utility plumbing (cooling water, compressed air, vent, boil-off nitrogen)
- process control, wiring, instrumentation, electrical
- firewalls, fire detection and suppression





SNO+ upgrades also included

- Refurbishing the electronics
- Repair of many “dead” PMT bases
- All-new DAQ
- New cover gas system
- New calibration systems capable of deploying in LAB scintillator
- New *in-situ* injected LED/laser light calibration system
- Calibration system cameras (for photogrammetry)

...in addition to the hold-down ropes and the scintillator plant

SNO+ Scintillator Fill



Started in mid-late 2019 and was proceeding smoothly (post-commissioning) when the pandemic struck, halting all activities for >6 months. At 365 tonnes filled (~45%), SNO+ **partial-fill** benefited from a quiet period with no operations, allowing radon backgrounds to decay and background levels in the LS to be measured.

SNO+ Partial Fill

- LS backgrounds measured at

$^{214}\text{BiPo}$ delayed coincidences for U chain

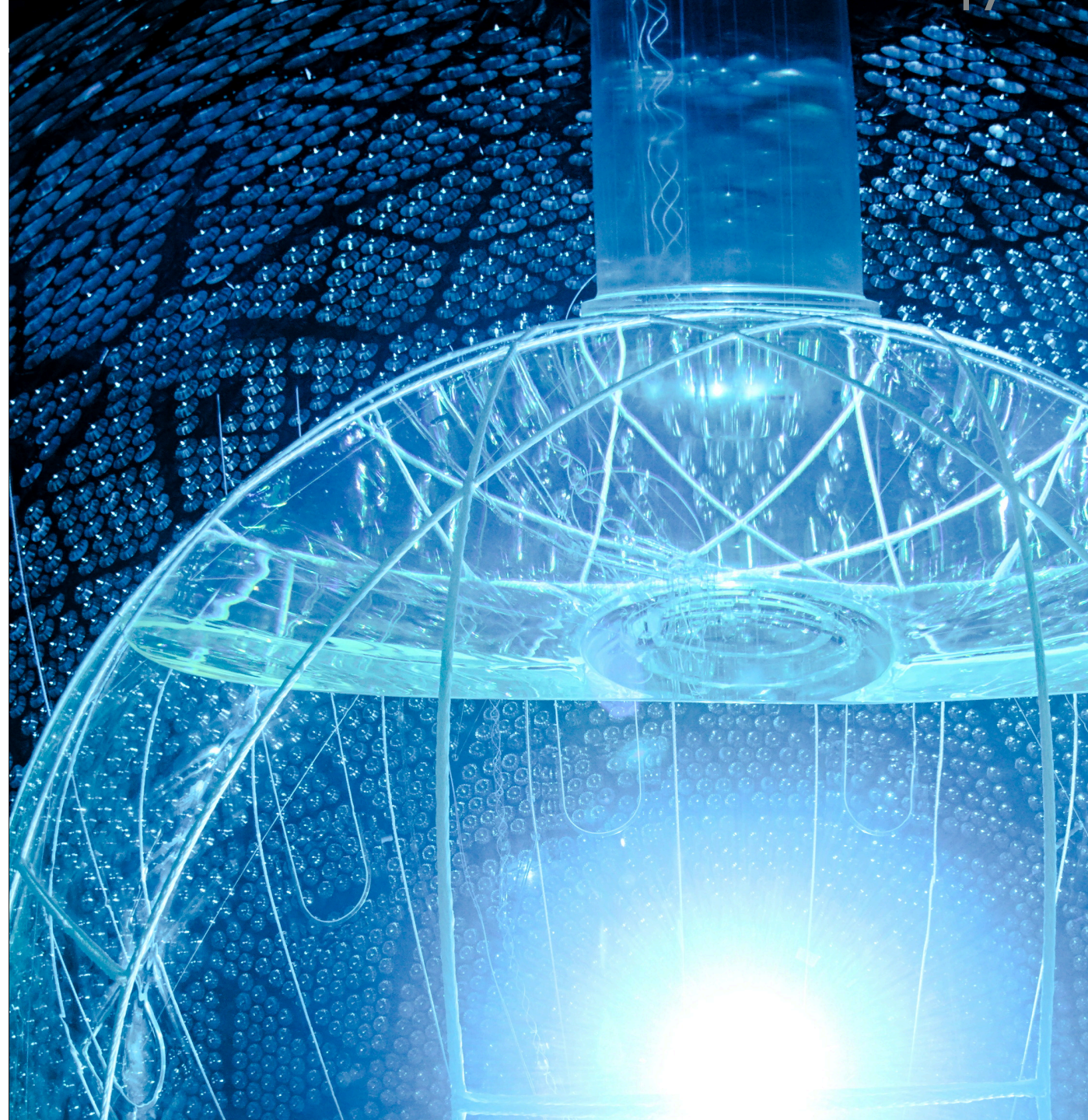
$$(4.7 \pm 1.2) \times 10^{-17} \text{ g}_U / \text{g}_{\text{LAB}}$$

$^{212}\text{BiPo}$ delayed coincidences for Th chain

$$(5.3 \pm 1.5) \times 10^{-17} \text{ g}_{\text{Th}} / \text{g}_{\text{LAB}}$$

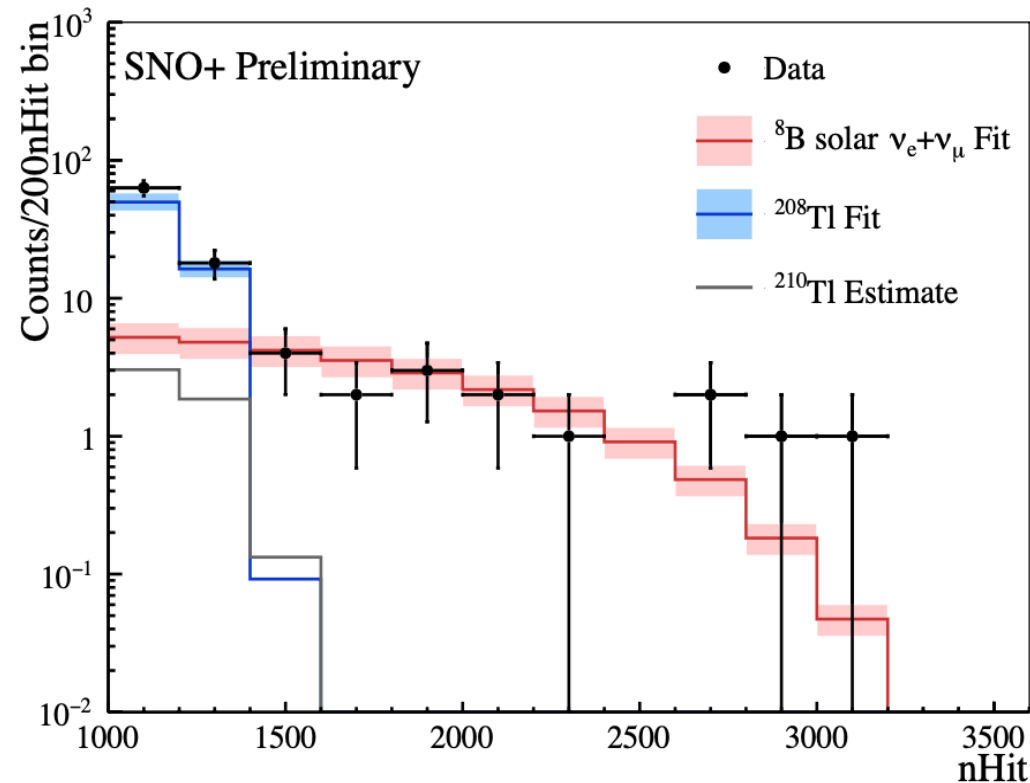
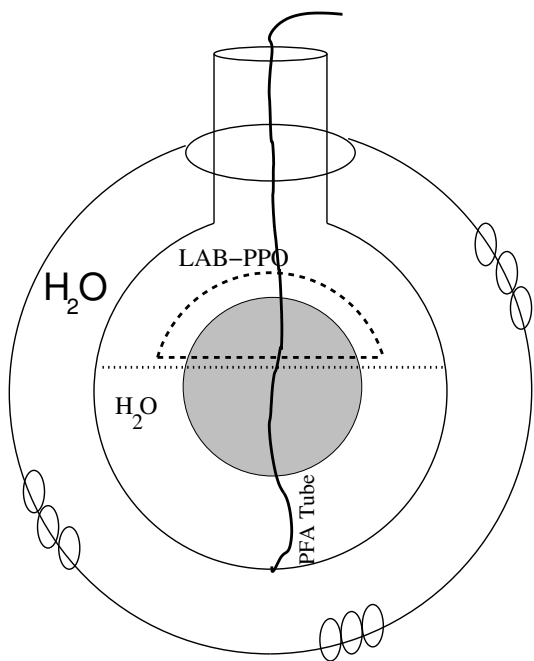
meeting SNO+ background targets (for double beta decay)

- Optical properties of LS 👍
- Also physics from SNO+ partial fill...



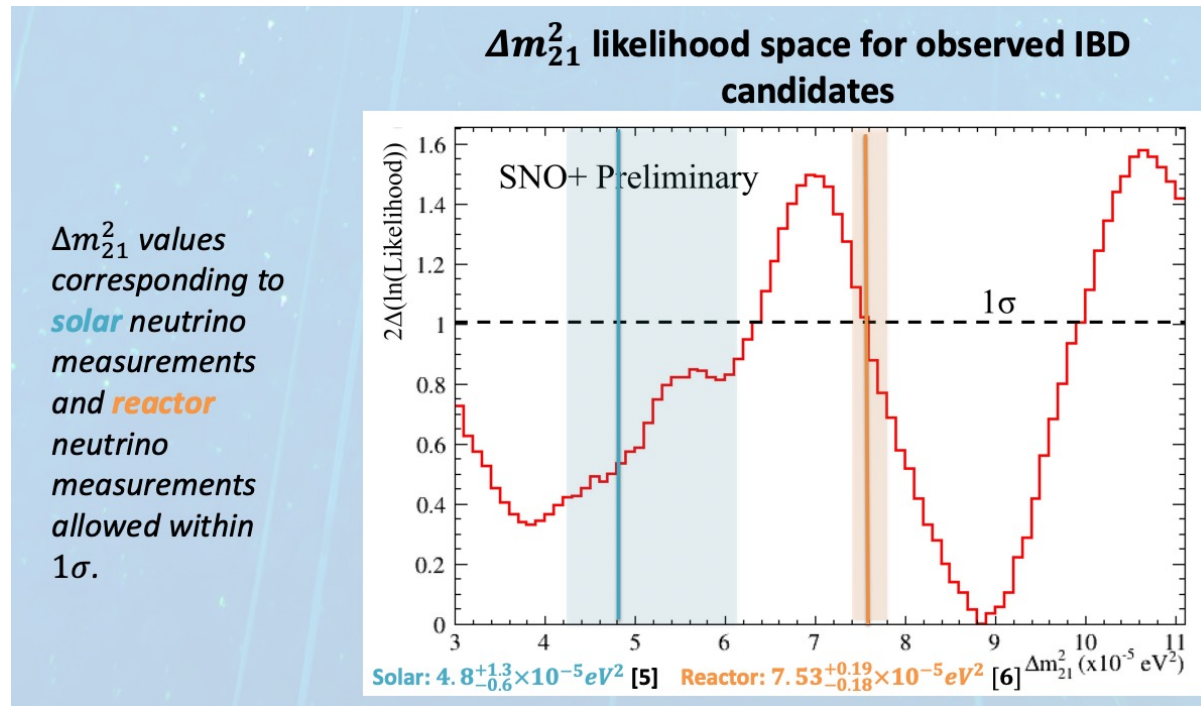
Physics with Partial-Fill Scintillator

- ^8B solar neutrinos in partial-fill scintillator
 - demonstrates SNO+ LS solar neutrino detection, even in a sub-optimal detector configuration



SNO+ reactor antineutrinos in partial-fill

- Publication in preparation; the result won't challenge our understanding of Δm_{12}^2 ; but draws attention to upcoming SNO+ measurements with full LS that will

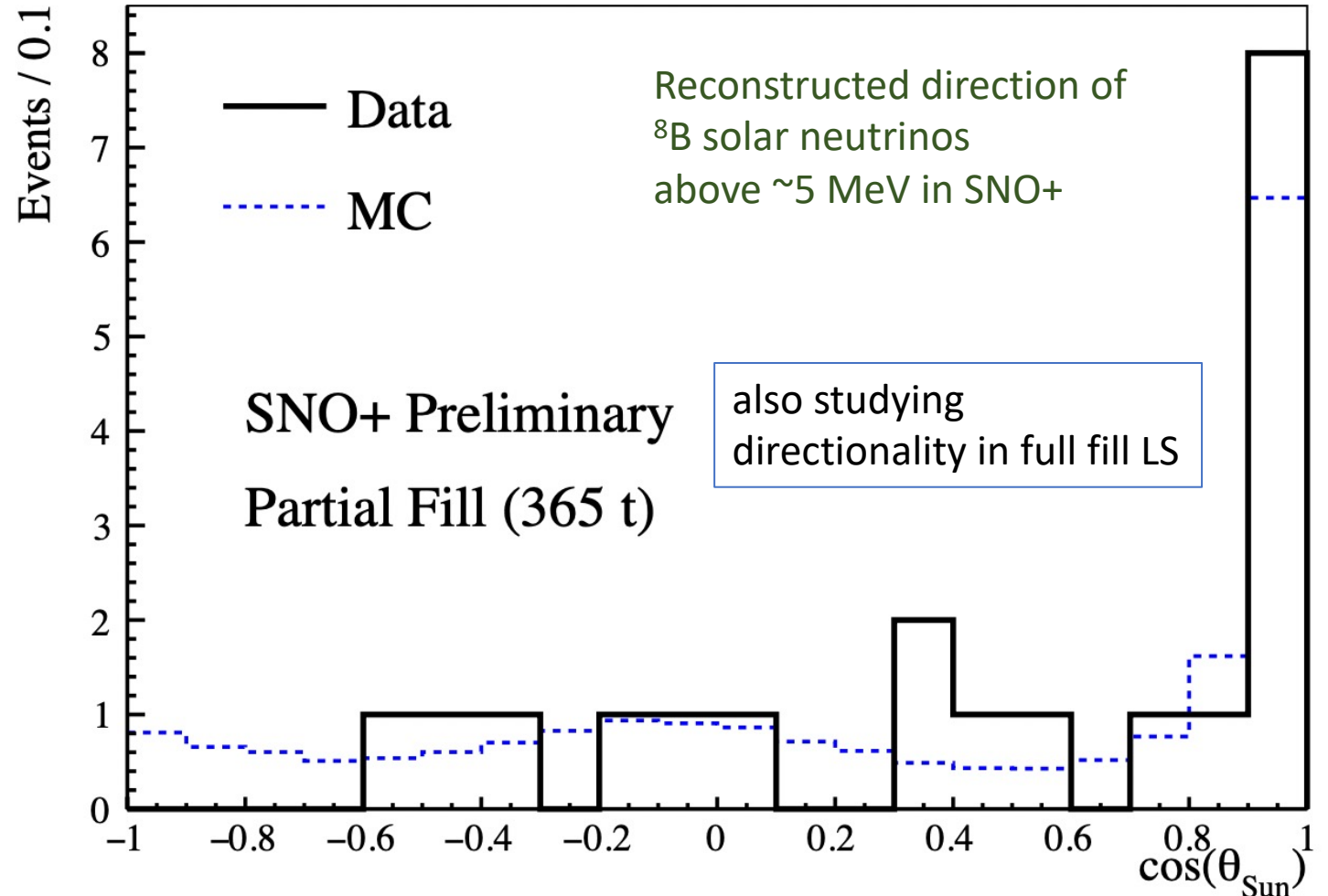


Event-by-Event Direction Reconstruction of Solar Neutrinos in SNO+ Liquid Scintillator

- Borexino has published the observation of a correlation between early PMT hits in the forward direction caused by the Cherenkov light produced by ^7Be solar neutrinos in liquid scintillator
- new SNO+ result: each recoil electron event's direction can be reconstructed by fitting with the combined Cherenkov+scintillation pdf

This is a first – **event-by-event direction reconstruction** of MeV events **in liquid scintillator**!

Publication being prepared

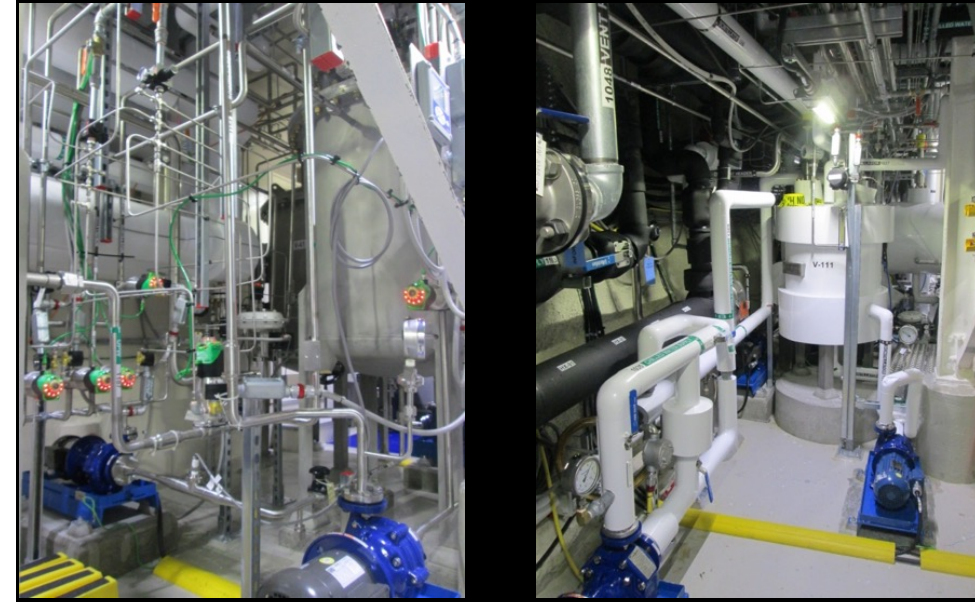


SNO+ Scintillator Fill Completed (during the pandemic)

- Deliveries of LAB from CEPISA (Bécancour, QC) to SNOLAB
- Transportation of LAB from surface to underground, coordinating with Vale, shipping railcars underground
- Distillation of LAB
- Water extraction and secondary distillation of PPO
- Nitrogen stripping
- Simultaneous filling of AV with purified LS and draining of water
- Nearly 5,000 QA samples analyzed (with lots of assistance from the SNOLAB Scientific Support Group) to verify the quality of the process to approve it before sending purified LS to the AV
- After completion of “bulk” fill, topped up the PPO concentration in the detector LS to 2.2 g/L

**Truly a monumental effort by SNO+ and SNOLAB during the pandemic!
Many thanks!**

Almost exactly 1 year ago, scintillator operations concluded and we started the...



The background of the slide is a dark blue, textured surface with a grid of faint, glowing lines, resembling a scintillator or a detector array. The text "SNO+ Scintillator Phase" is centered in white. There are some bright, glowing spots on the background, possibly representing light or data points.

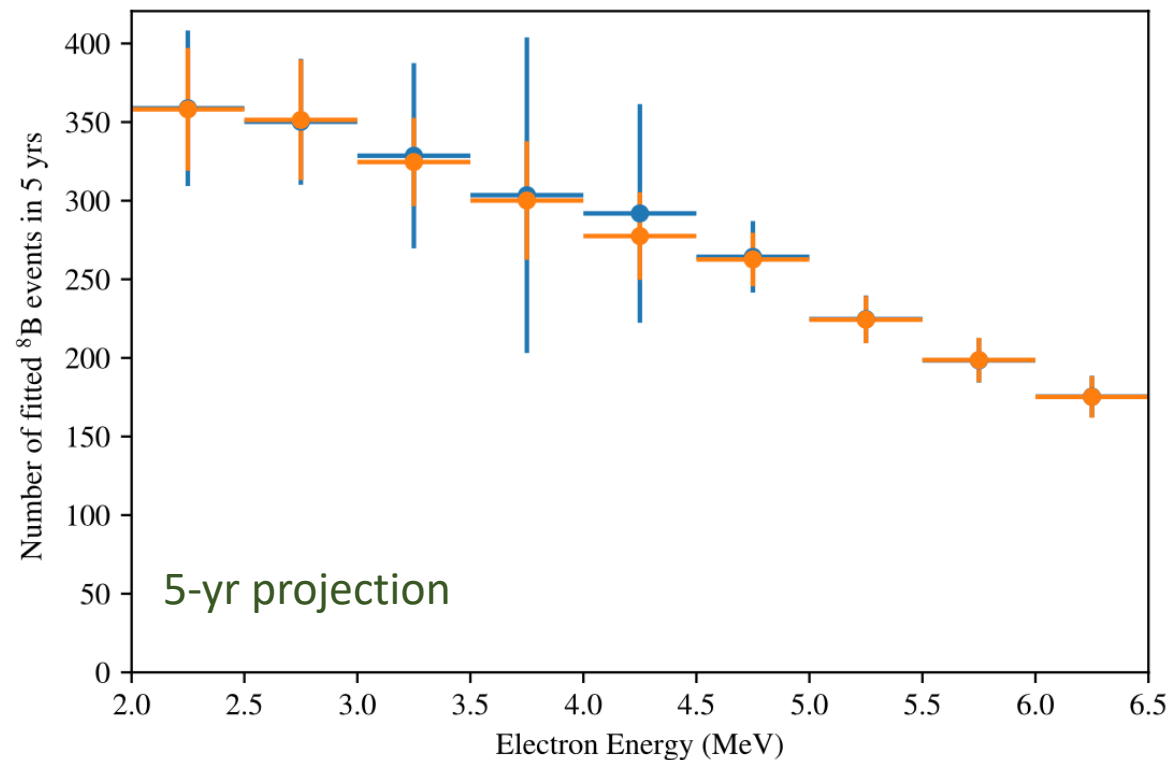
SNO+ Scintillator Phase

Physics Goals in the Scintillator Phase

- Solar neutrinos at lower energies
- Reactor antineutrinos flux, spectrum, oscillations (Δm_{12}^2 , in particular)
- Geo neutrinos
- We are supernova neutrino live
- and other physics (e.g. MIMP dark matter searches, DSNB – diffuse supernova neutrino background, nucleon decay)

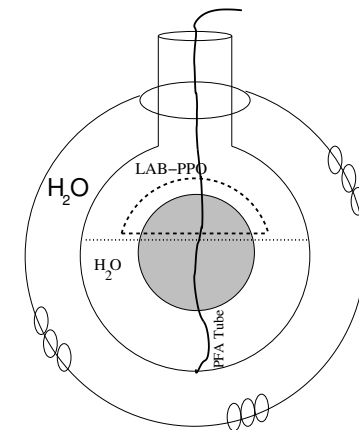
Objectives for SNO+ Scintillator Phase (Full): ^8B Solar Neutrinos at Low Energies

- See if we can measure below 3 MeV (hasn't been done before)
 - larger fiducial volume than Borexino
 - cosmogenic backgrounds much lower than KamLAND (e.g., no ^{10}C , ^{11}C)



Blue 5×10^{-17} g ^{238}U /g $_{\text{LAB}}$, 5×10^{-17} g ^{232}Th /g $_{\text{LAB}}$
Orange 5×10^{-18} g ^{238}U /g $_{\text{LAB}}$, 5×10^{-18} g ^{232}Th /g $_{\text{LAB}}$

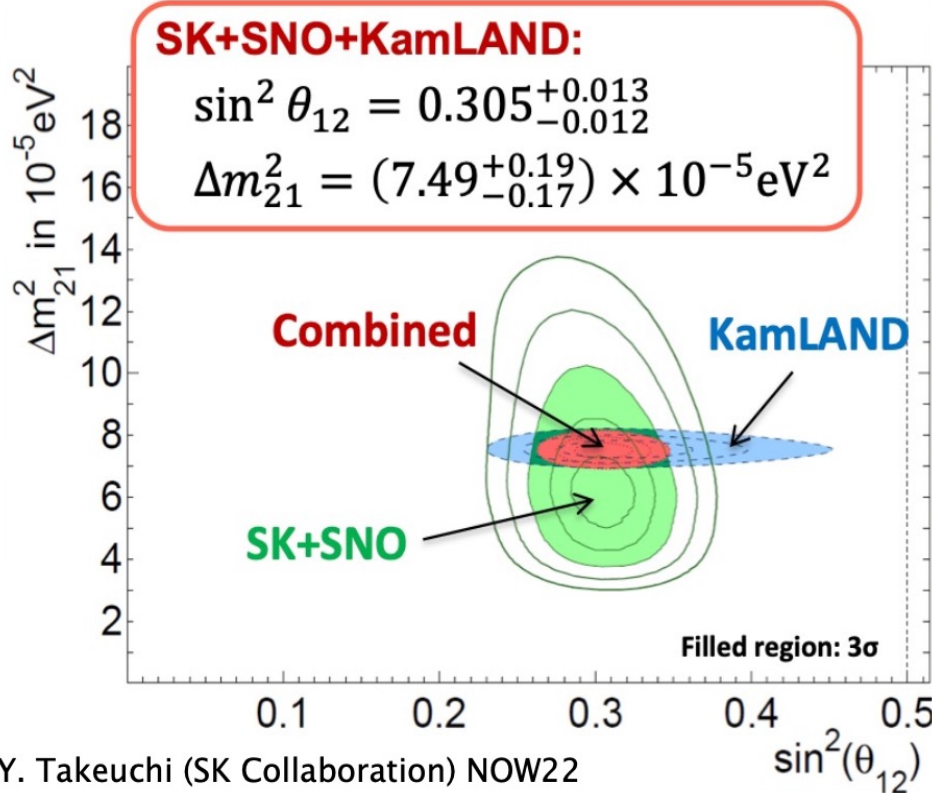
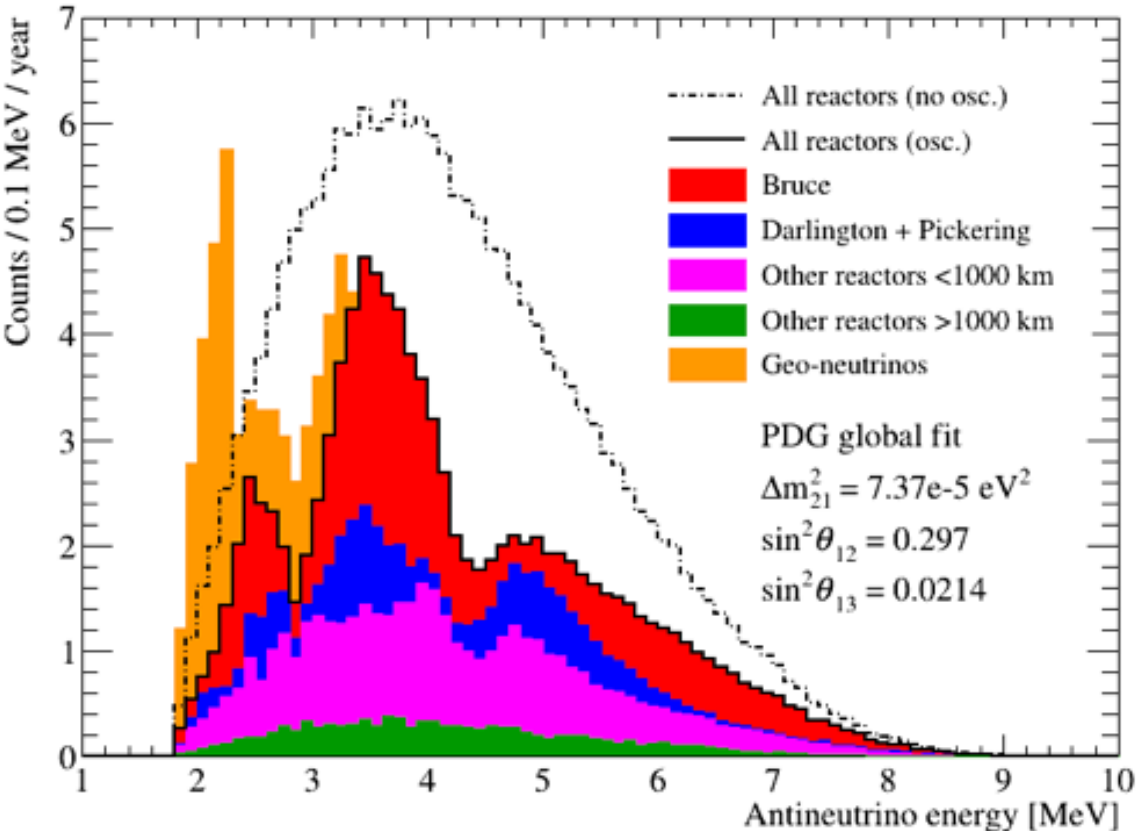
Blue U and Th at partial-fill levels
Orange U and Th below 10^{-17} g/g



Reminder:
partial-fill was
sub-optimal
configuration

Antineutrinos in SNO+ Scintillator

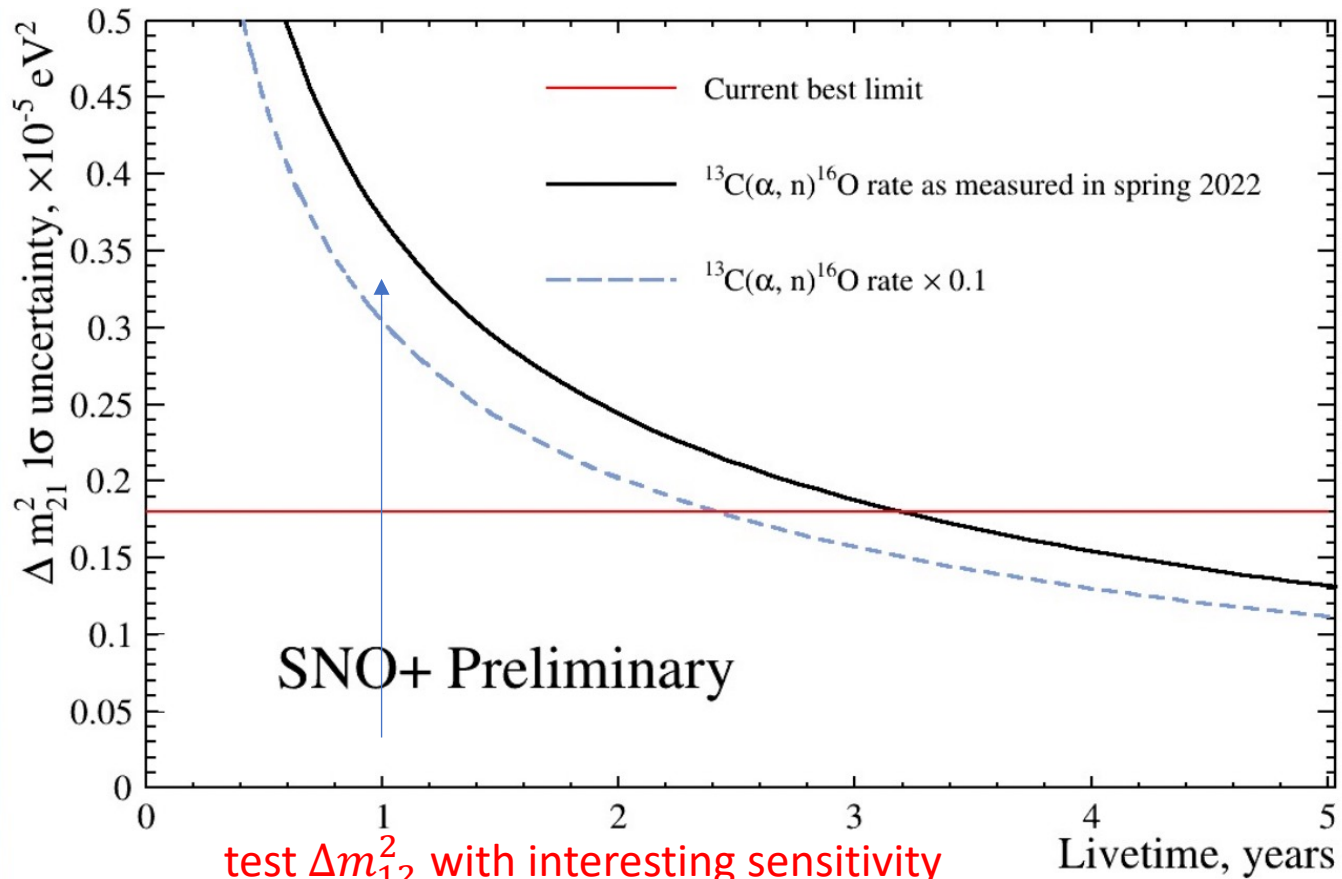
Scintillator Phase – Reactor antineutrino oscillations Δm_{12}^2 (plus geo neutrinos) are one of the main science goals



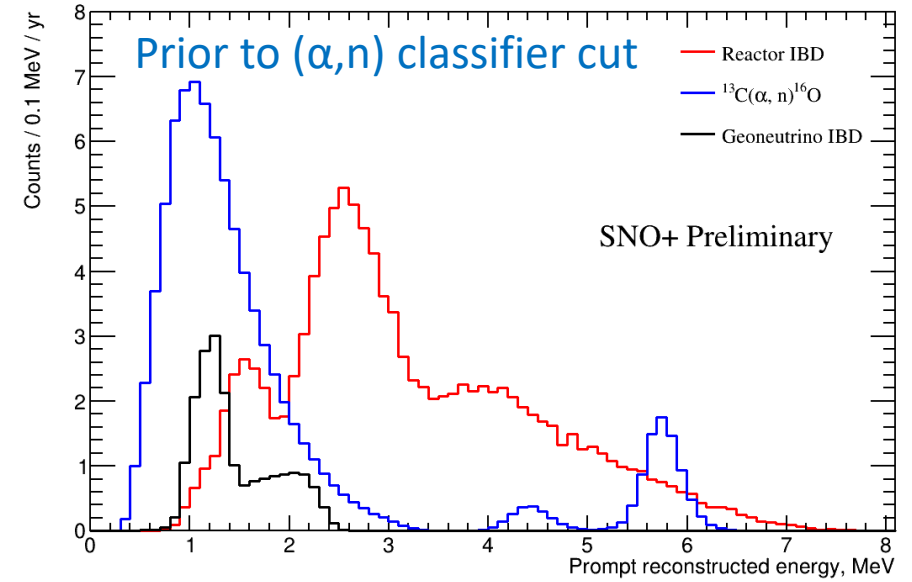
Y. Takeuchi (SK Collaboration) NOW22

physics motivation for reactor neutrino oscillation studies

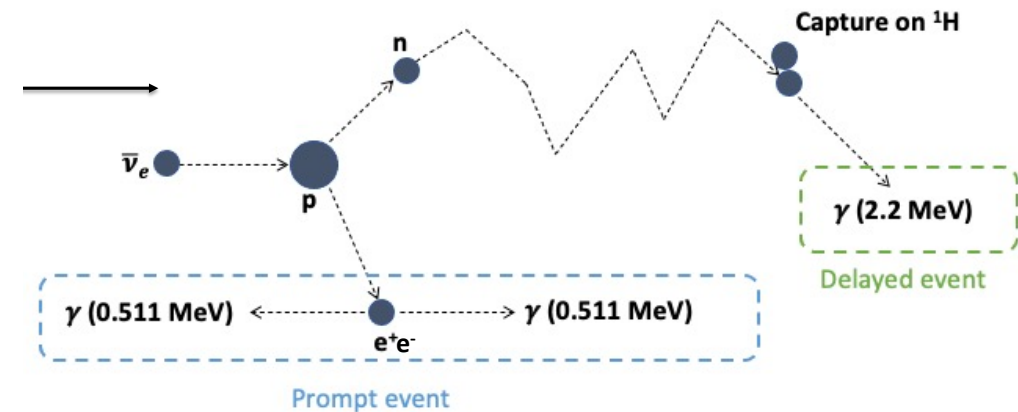
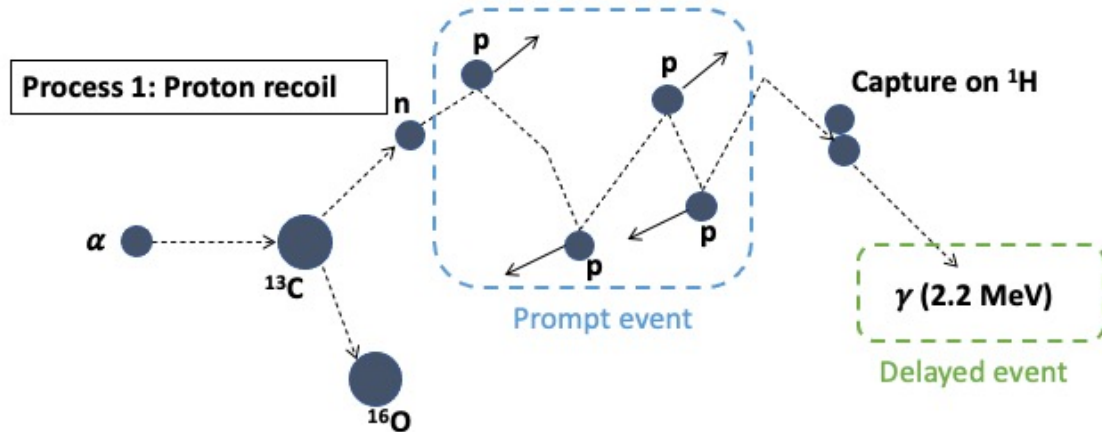
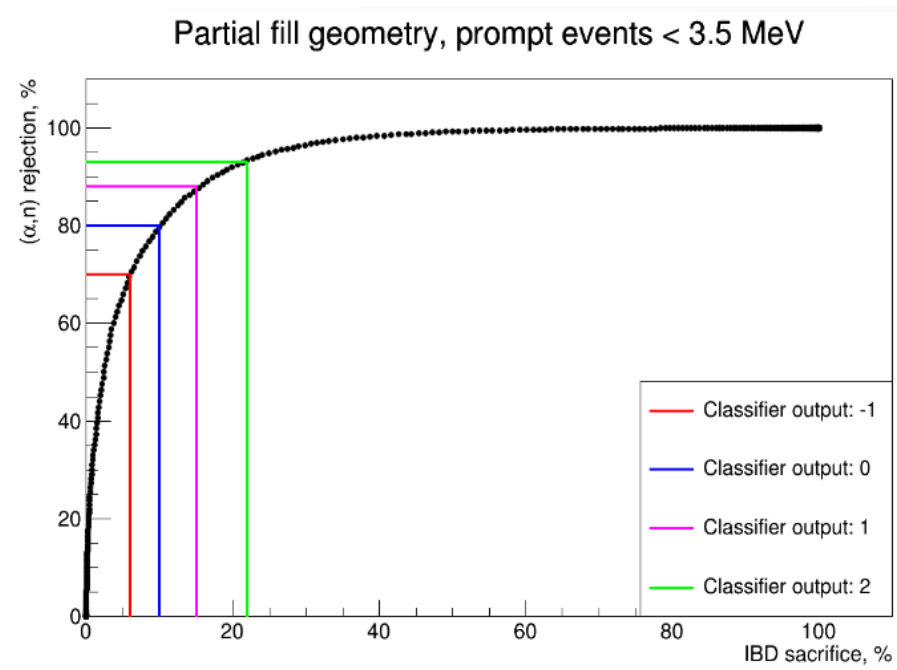
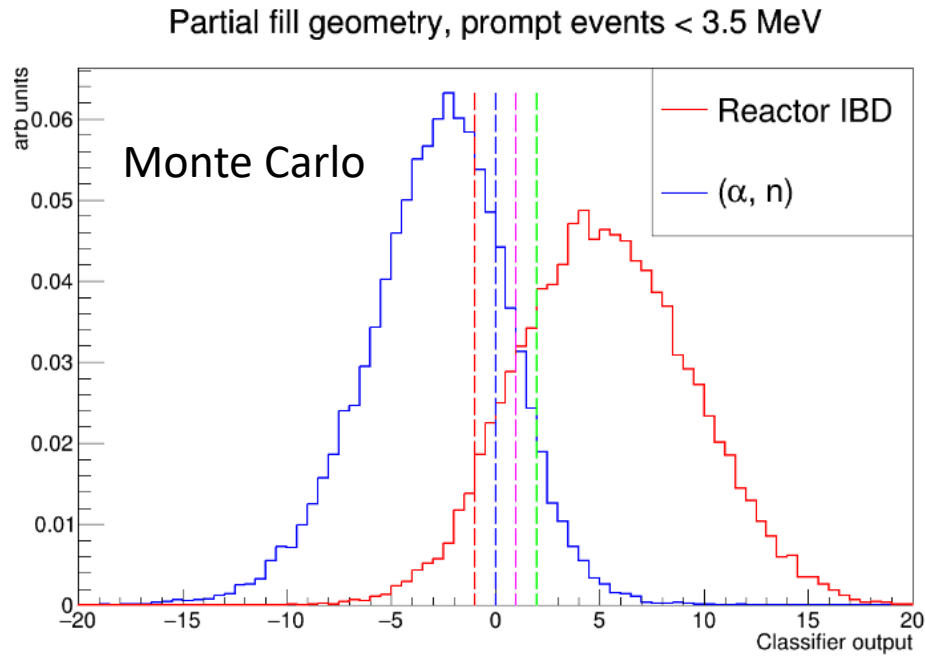
Objectives for SNO+ Scintillator Phase: Reactor Antineutrinos Δm_{12}^2



test Δm_{12}^2 with interesting sensitivity
already after 1 year (shape is more
important than rate) to distinguish
5.1 from 7.5 ($\times 10^{-5} \text{ eV}^2$)



(α, n) Classifier (modelled in partial-fill) – new SNO+ development



For some exotic dark matter models, SNO+ has some capability to probe further than others have or can...

MIMPs in SNO+

Multiply-Interacting Massive Particles (dark matter) at the Planck scale

SNO+ mass reach

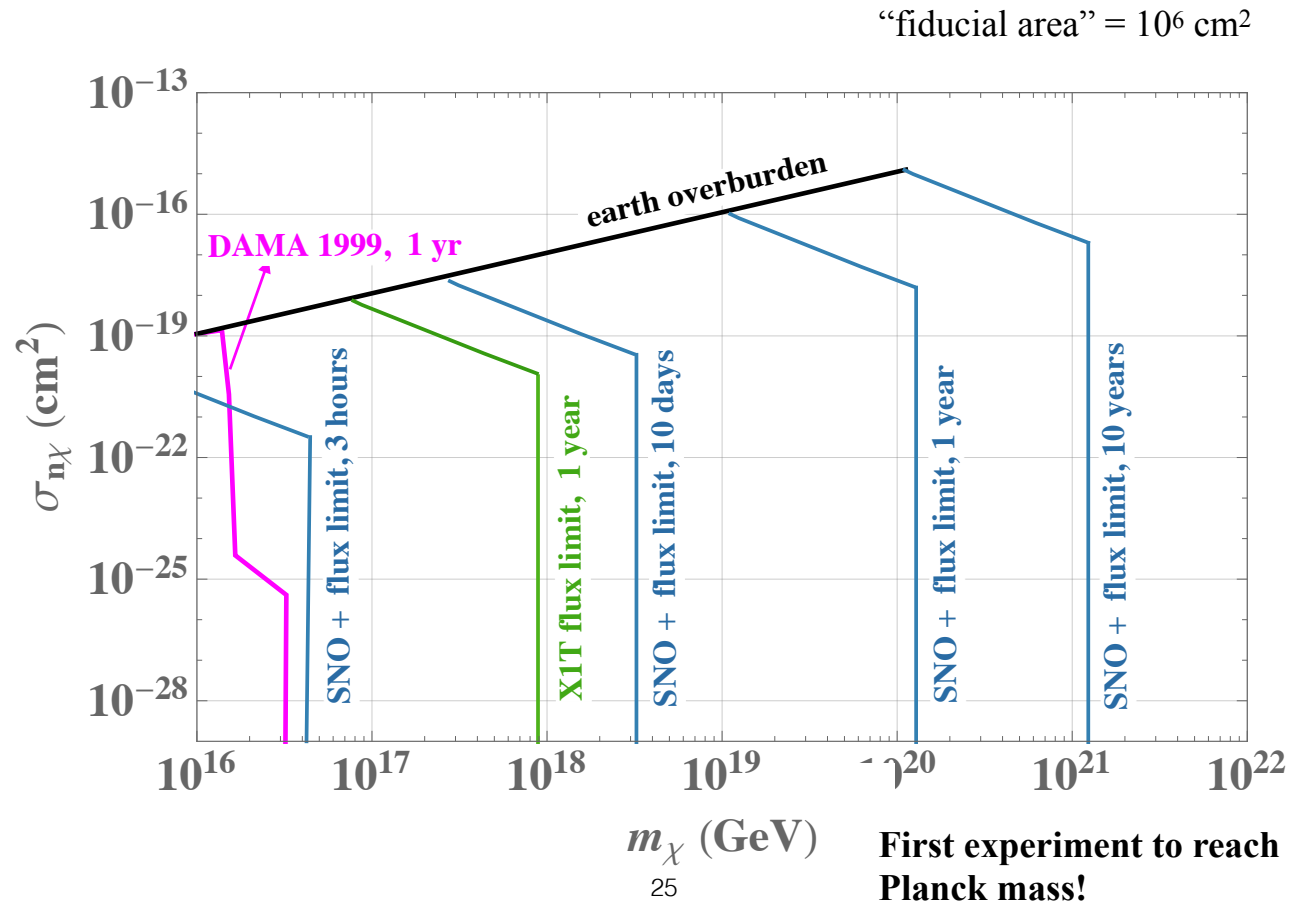
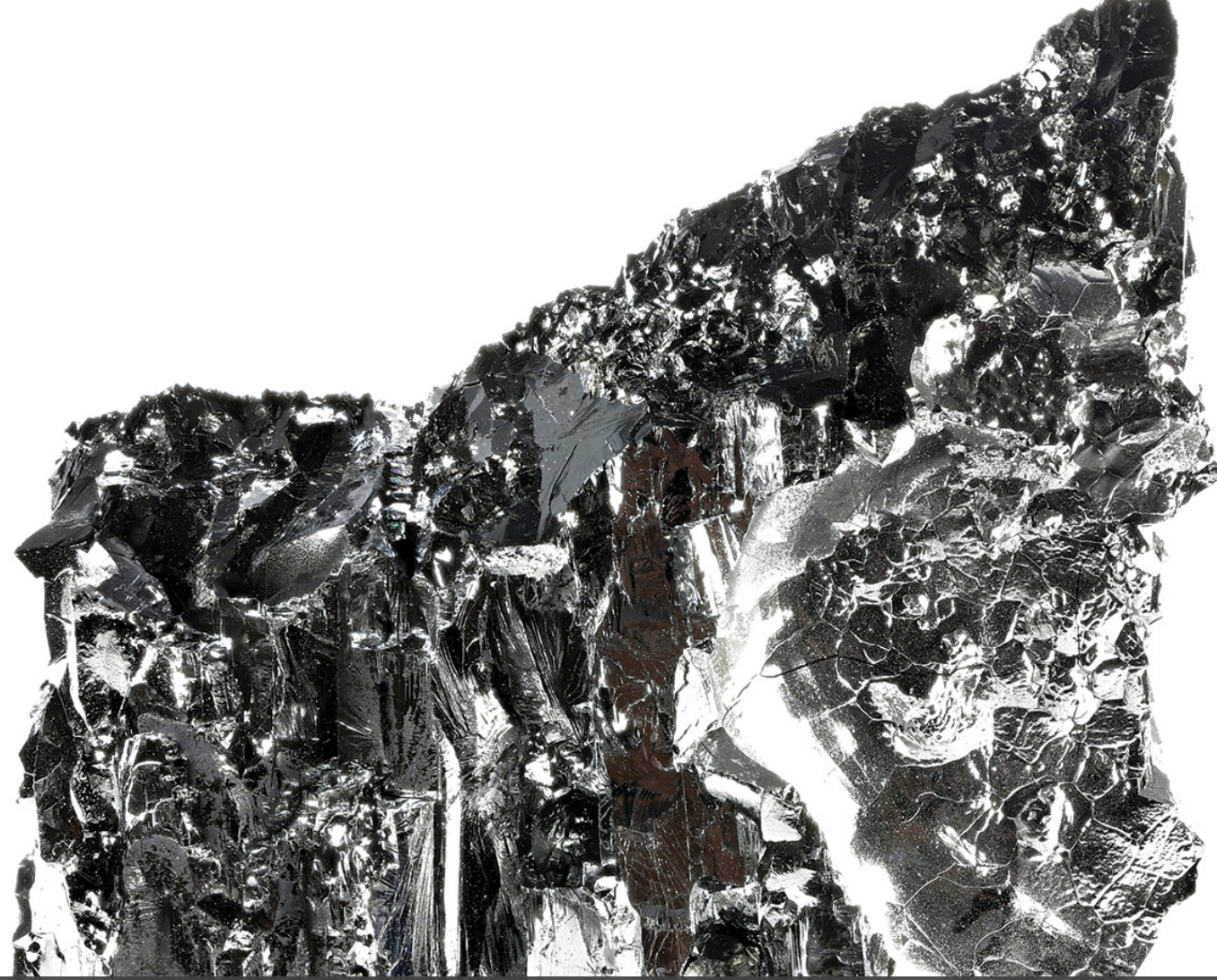


figure from N. Raj talk

The advantages of a well-understood detector with very low backgrounds

- are being demonstrated!
- SNO+ has a diverse physics program that is being pursued.
- With the detector performing well; with all background components being measured and constrained (most coming in at or below target levels), it looks promising for the final phase of SNO+...

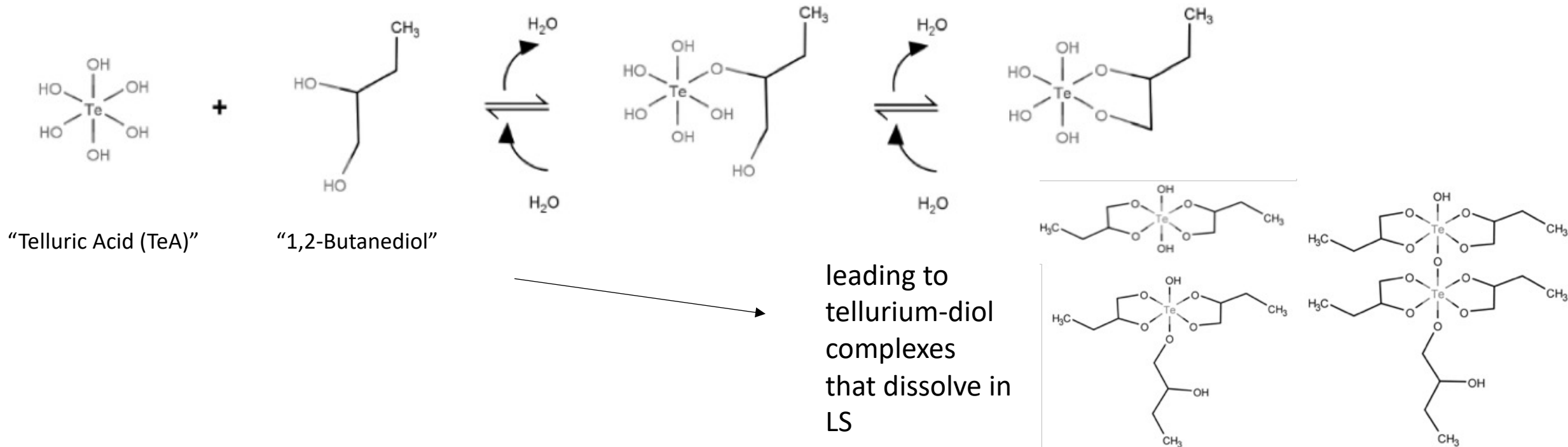


SNO+ Tellurium Double Beta Decay Phase

Neutrinoless Double Beta Decay in SNO+ with Tellurium-Loaded Liquid Scintillator

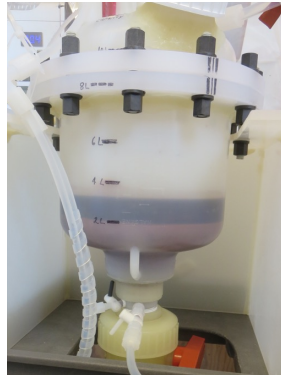
Principal goal: economical, scalable approach to $0\nu\beta\beta$; achieving sensitivity to $m_{\beta\beta}$ in the parameter space corresponding to the Inverted Neutrino Mass Ordering...and beyond

^{130}Te has 34% natural abundance = no costly or logistically difficult isotopic enrichment required

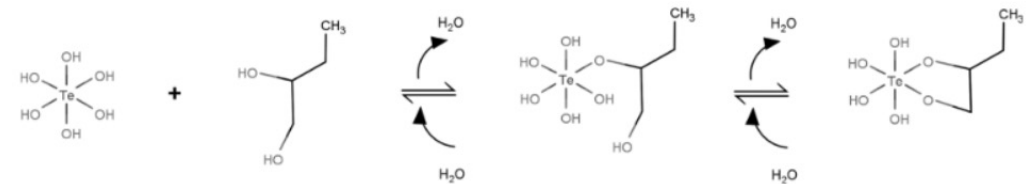


Novel Tellurium Purification and Tellurium Loading Techniques Pioneered by SNO+

Te purification technique established





Practical, stable Te loading method established



Nuclear Instruments and Methods in Physics
Research Section A: Accelerators, Spectrometers,
Detectors and Associated Equipment
Volume 795, 21 September 2015, Pages 132-139

Purification of telluric acid for SNO+ neutrinoless double-beta decay search

S. Hans^{a,1}, R. Rosero^{a,1}, L. Hu^{a,1}, O. Chkvorets^b, W.T. Chan^{a,1}, S. Guan^{a,1},
W. Beriguete^{a,1}, A. Wright^d, R. Ford^c, M.C. Chen^d, S. Biller^e, M. Yeh^{a,1}  



Contents lists available at [ScienceDirect](https://www.sciencedirect.com)
Nuclear Inst. and Methods in Physics Research, A

journal homepage: www.elsevier.com/locate/nima

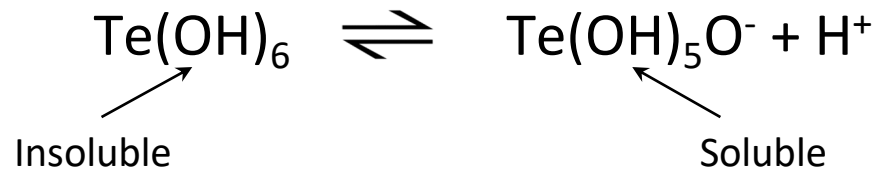
Full Length Article

A method to load tellurium in liquid scintillator for the study of neutrinoless double beta decay

D.J. Auty^a, D. Bartlett^b, S.D. Biller^{c,*}, D. Chauhan^{f,b}, M. Chen^b, O. Chkvorets^d, S. Connolly^b, X. Dai^b, E. Fletcher^b, K. Frankiewicz^e, D. Gooding^e, C. Grant^e, S. Hall^f, D. Horne^b, S. Hansⁱ, B. Hreljac^b, T. Kaptanoglu^{g,h}, B. Krar^b, C. Kraus^{d,f}, T. Kroupová^{c,g}, I. Lam^b, Y. Liu^b, S. Maguireⁱ, C. Miller^b, S. Manecki^{f,b}, R. Roseroⁱ, L. Segui^c, M.K. Sharma^a, S. Tacchino^f, B. Tam^b, L. Tian^b, J.G.C. Veinot^a, S.C. Walton^d, J.J. Weigand^j, A. Wright^{b,k}, M. Yehⁱ, T. Zhao^b

Tellurium Purification Process via pH Selective Telluric Acid Recrystallization

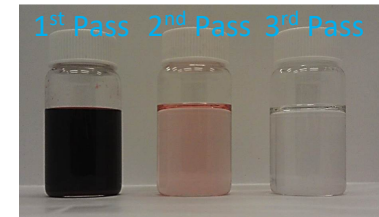
Telluric acid obeys the following equilibrium:



pH determines the equilibrium state

Purification basics:

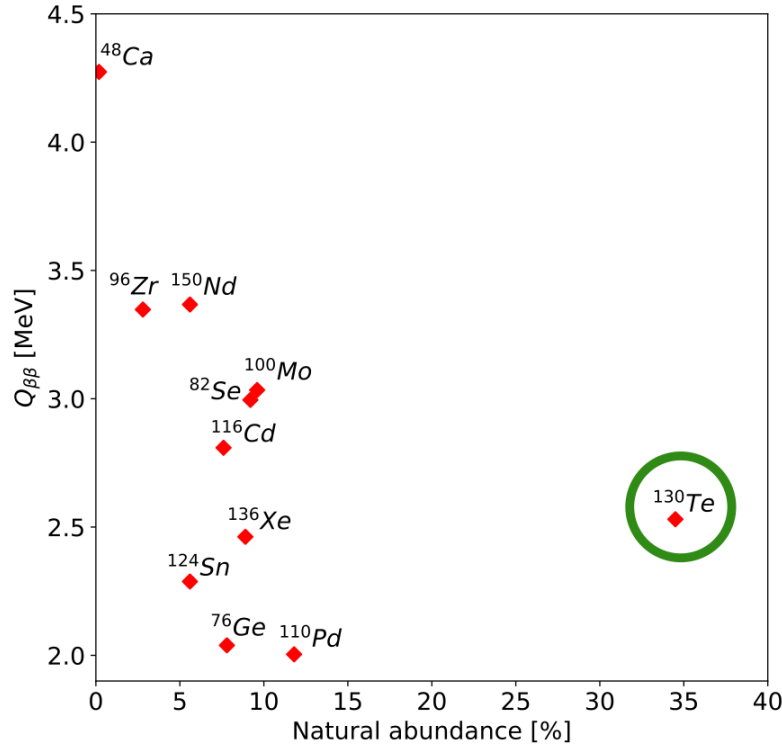
1. Dissolve telluric acid in water and filter it
 - Removes water insoluble impurities
2. Add nitric acid to force the telluric acid to recrystallize/precipitate, pump away the liquid and rinse
 - Removes acid soluble impurities



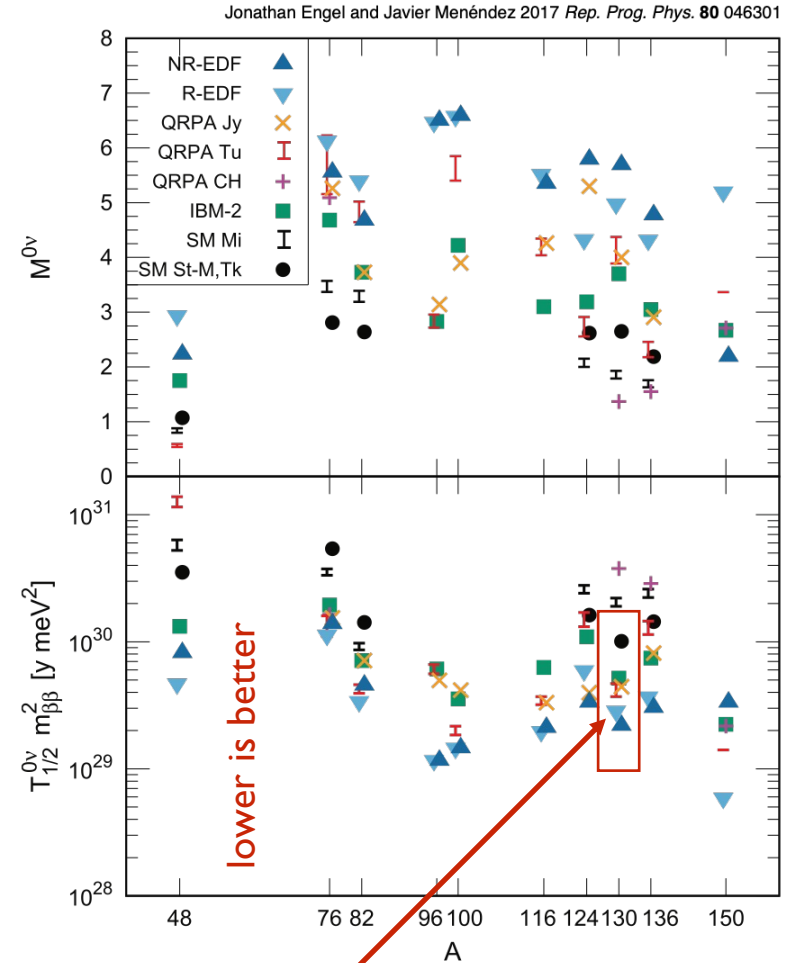
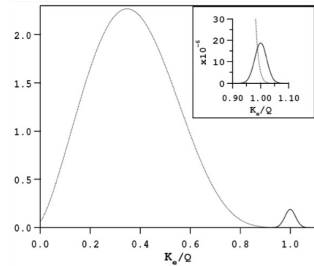
50kg pilot-scale

By “tuning” the process pH’s, this can be quite specific to telluric acid – most other chemicals are removed with high efficiency.

Tellurium for Double Beta Decay



^{130}Te & ^{136}Xe have the smallest $2\nu\beta\beta/0\nu\beta\beta$ ratio



Favourable $G_{0\nu}(Q, Z) |M_{0\nu}|^2$

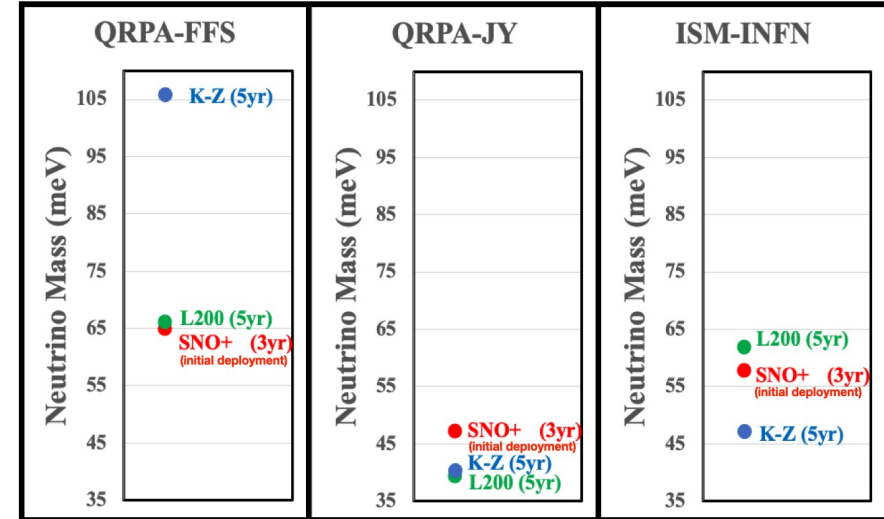
Loaded Scintillator Approach to DBD

- Previous slide: “why tellurium?”
 - This slide: “why in liquid scintillator?”
1. very low backgrounds: 5×10^{-7} counts/keV/kg_{fiducial detector}/yr
 2. homogeneous detector volume – reliable background model
 3. “target out” – ability to measure/constrain backgrounds *before* isotope added
 4. “sideband analysis” – not just counts in a bin but distributions in position and energy verify detector response and background model
 5. liquid detector permits: assays, chemistry; liquid medium can be modified *in situ* (e.g., adding more Te, more fluor)

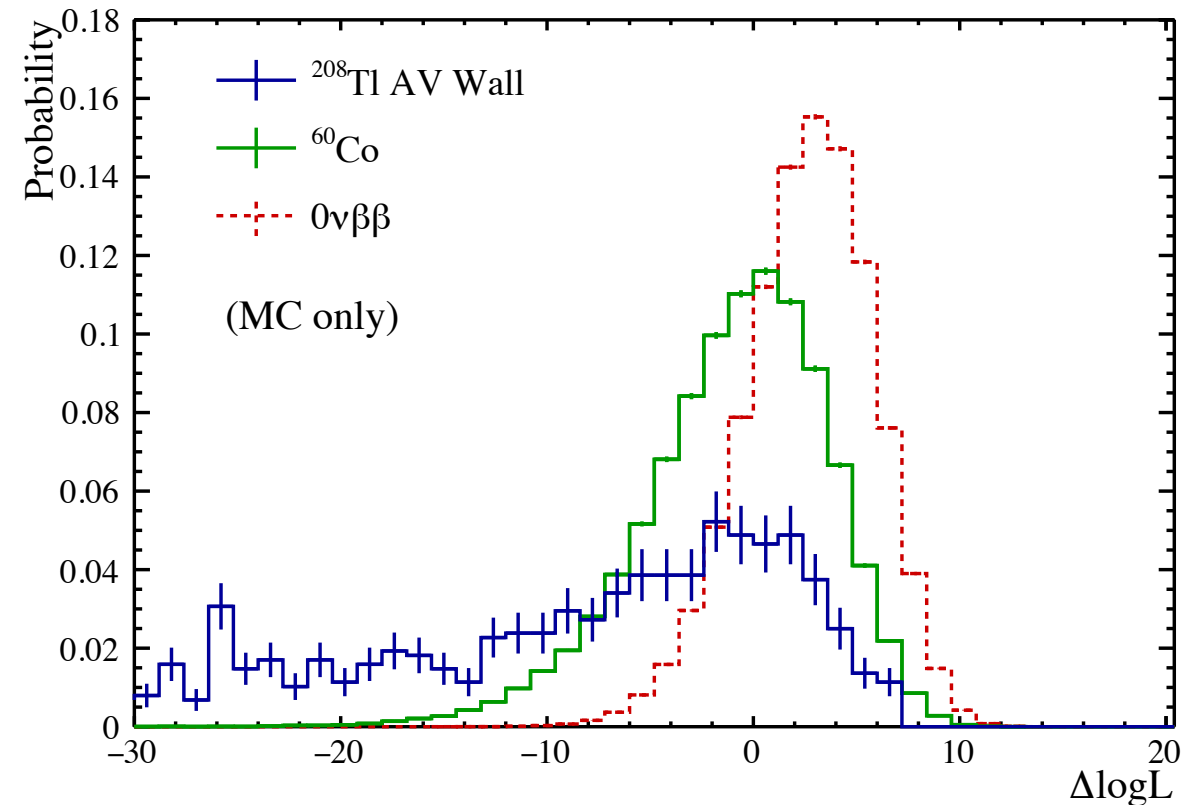
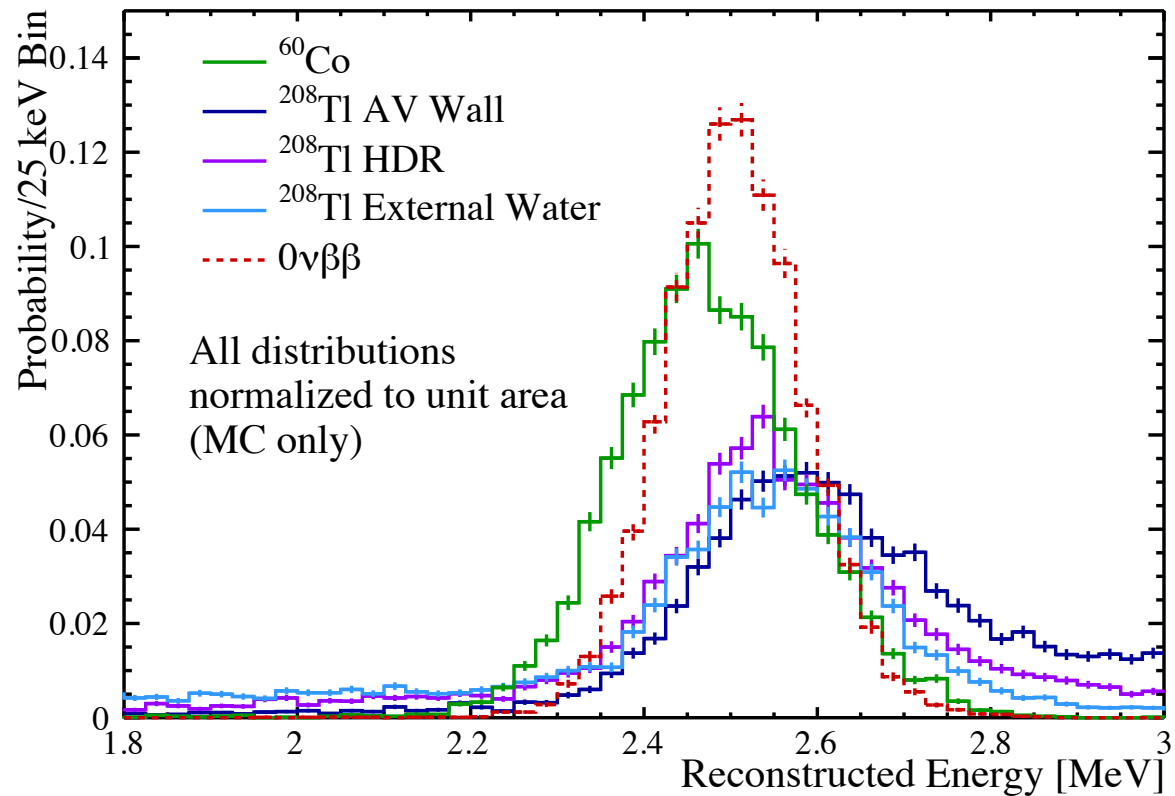
The dependence of a putative signal with amount of isotope would be a strong confirmation!

SNO+ Te DBD Additional Considerations

- ^{130}Te DBD is **scalable**, cost effective, **unimpacted by logistics difficulties of isotope enrichment**
- KL-Z 800 has world-leading sensitivity (upper limit 36-156 meV) and highlights the **strength of the loaded LS DBD approach**
- Complementarity of isotope
 - NME model dependencies
 - SNO+ sensitivity at %Te loading “fills the gap”, before larger experiments come online
 - early Te deployment would already be a competitive measurement, ready to test any hints of a positive signal
 - purification of Te underground is **novel technology**
 - **“target out” analysis** is a strong and unique feature; all non-Te backgrounds constrained prior to adding any Te
- SNO+ also has single-site/multi-site background constraining power



SNO+ Multi-site Background Likelihood Constraint



Thoughts on Future NLDBD Sensitivity

(general, for all experiments)

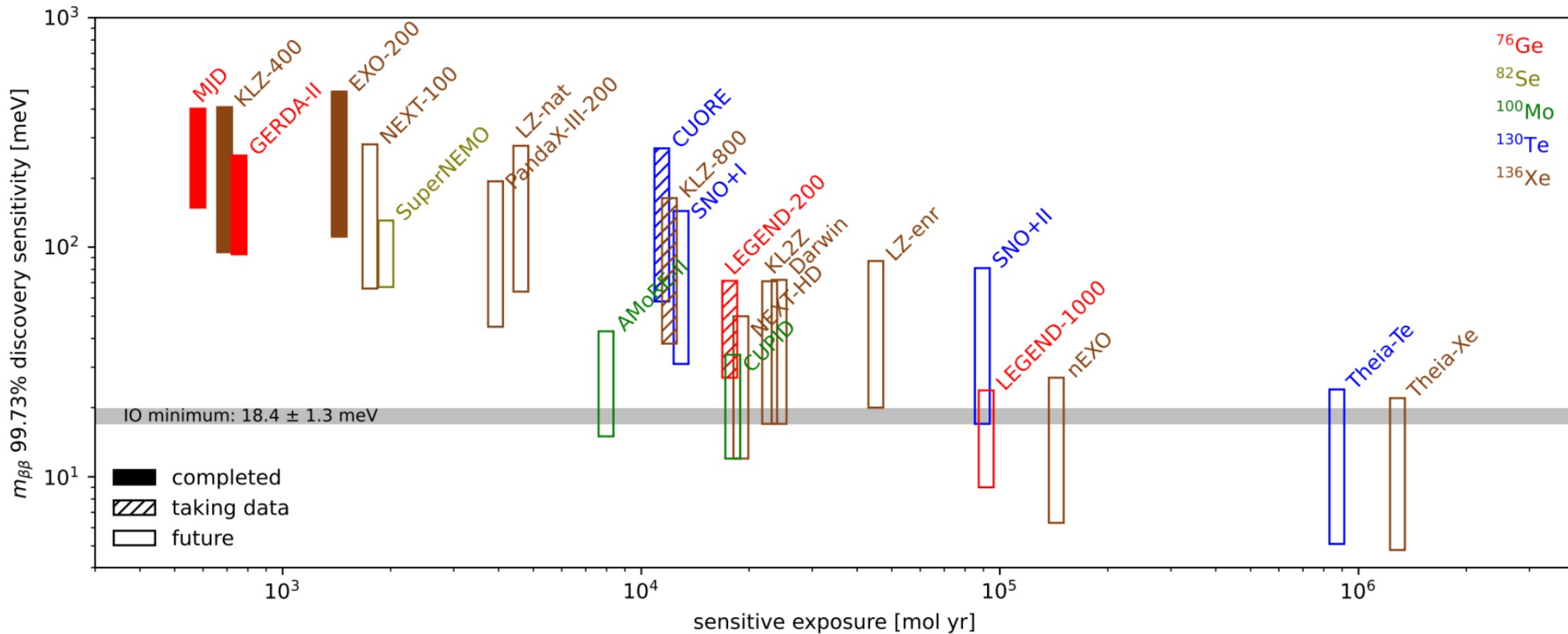
S/B determines sensitivity

- Signal / Background
 - an affordable way to deploy a *large* quantity of isotope is required to reach non-degenerate “Normal Mass Ordering” sensitivity
- “Background Index” = counts/keV/kg/yr
 - to improve backgrounds, one can improve
 - “keV” energy resolution – SNO+ has **increased the light yield of the TeLS** in recent R&D; photon detection can be improved in future experiment
 - “counts/kg-yr” – low radioactivity techniques have been developed by SNO+ with **underground purification of tellurium** just getting started → potential for future improvement
- Final State “Identification”
 - two-electron (single-site) signal topology suppresses backgrounds – many experiments have event classifiers like **single-site/multi-site discrimination, including SNO+** (also Cherenkov/scintillation separation in SNO+ R&D)
 - tagging the DBD daughter nucleus – an interesting capability being developed by nEXO and NEXT

Including background rejection techniques;
Including cosmogenic backgrounds

Sensitivity

from NSAC NLDBD White Paper



Status of SNO+ Te DBD

Tellurium systems are built and ready for operation!

TeA purification test batch (at full-scale) being prepared for beginning in the next few months



Telluric acid purification



Te-diol synthesis

Summary

- SNO+ is an **operating liquid scintillator neutrino detector** filled with LAB + 2.2 g/L PPO and taking data
- **Diverse** program of neutrino (and other) physics is underway
- **Already-built underground tellurium plants represent novel technology** in the field of low-radioactivity techniques and are beginning full-scale, test batch operations in the next few months
- **Operating the plants and demonstrating their capabilities** is the next step towards preparing to load SNO+ with Te for the $0\nu\beta\beta$ phase