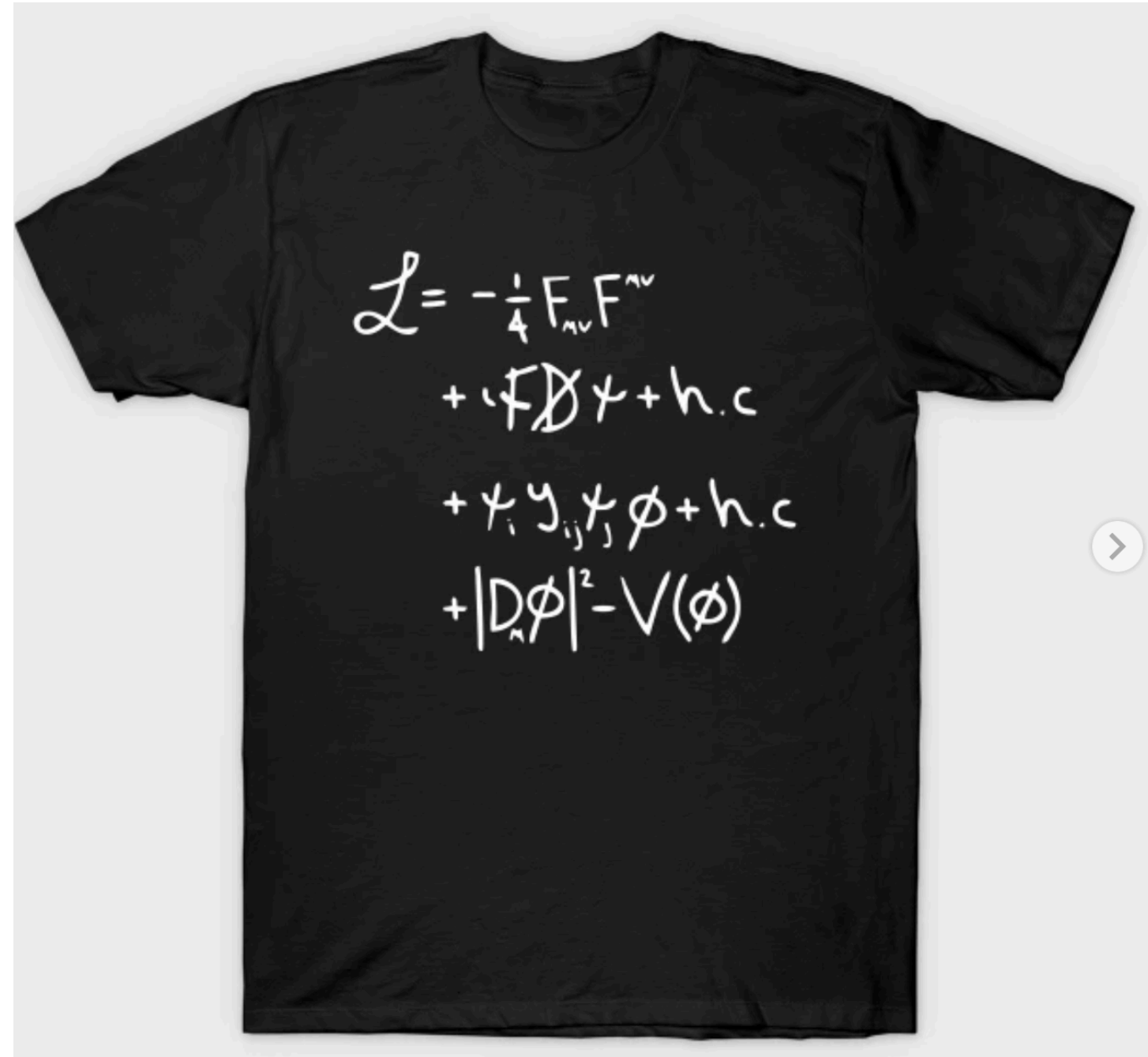


# Neutrinoless double beta decay and new physics

David McKeen





Works great



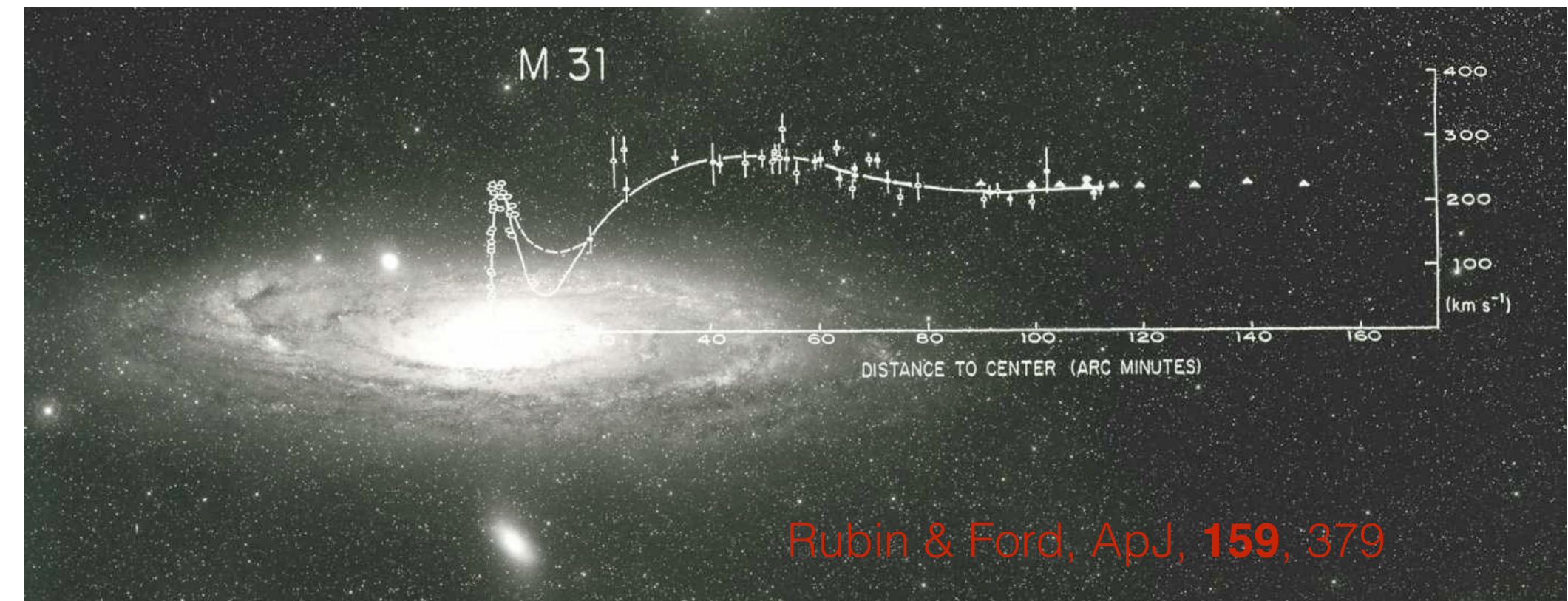
Works great

...but can't explain...



Works great

...but can't explain...

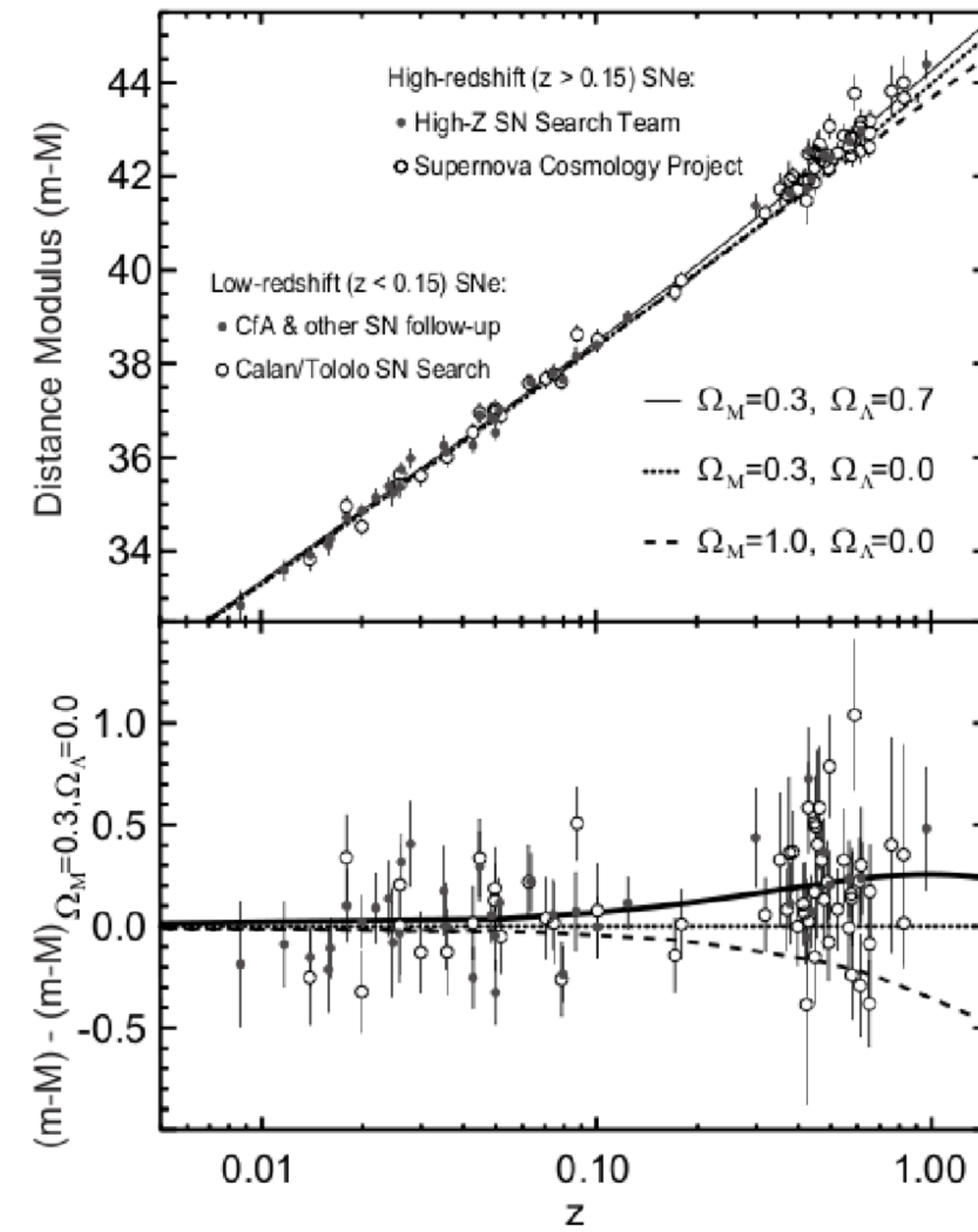


dark matter



Works great

...but can't explain...

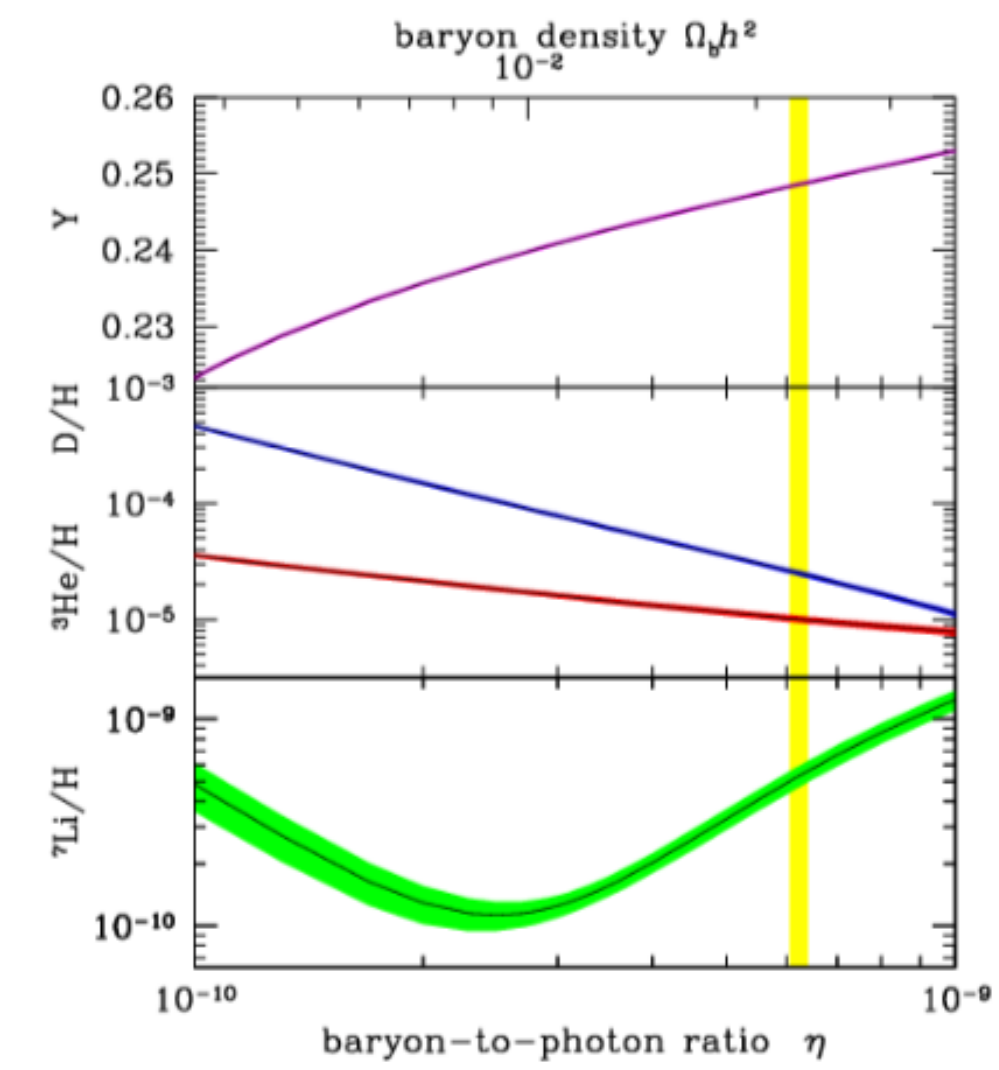


dark energy



Works great

...but can't explain...

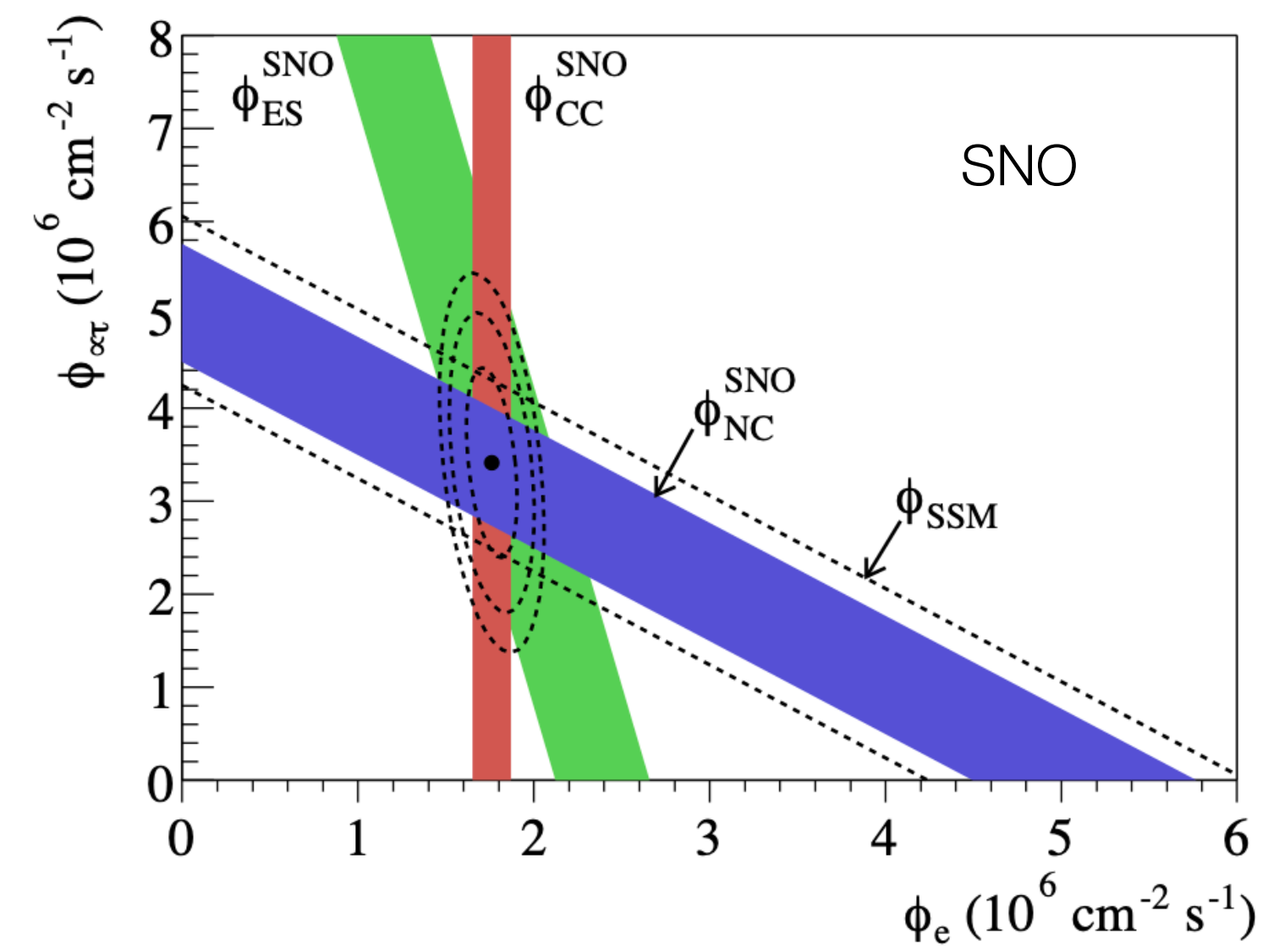


matter asymmetry



Works great

...but can't explain...

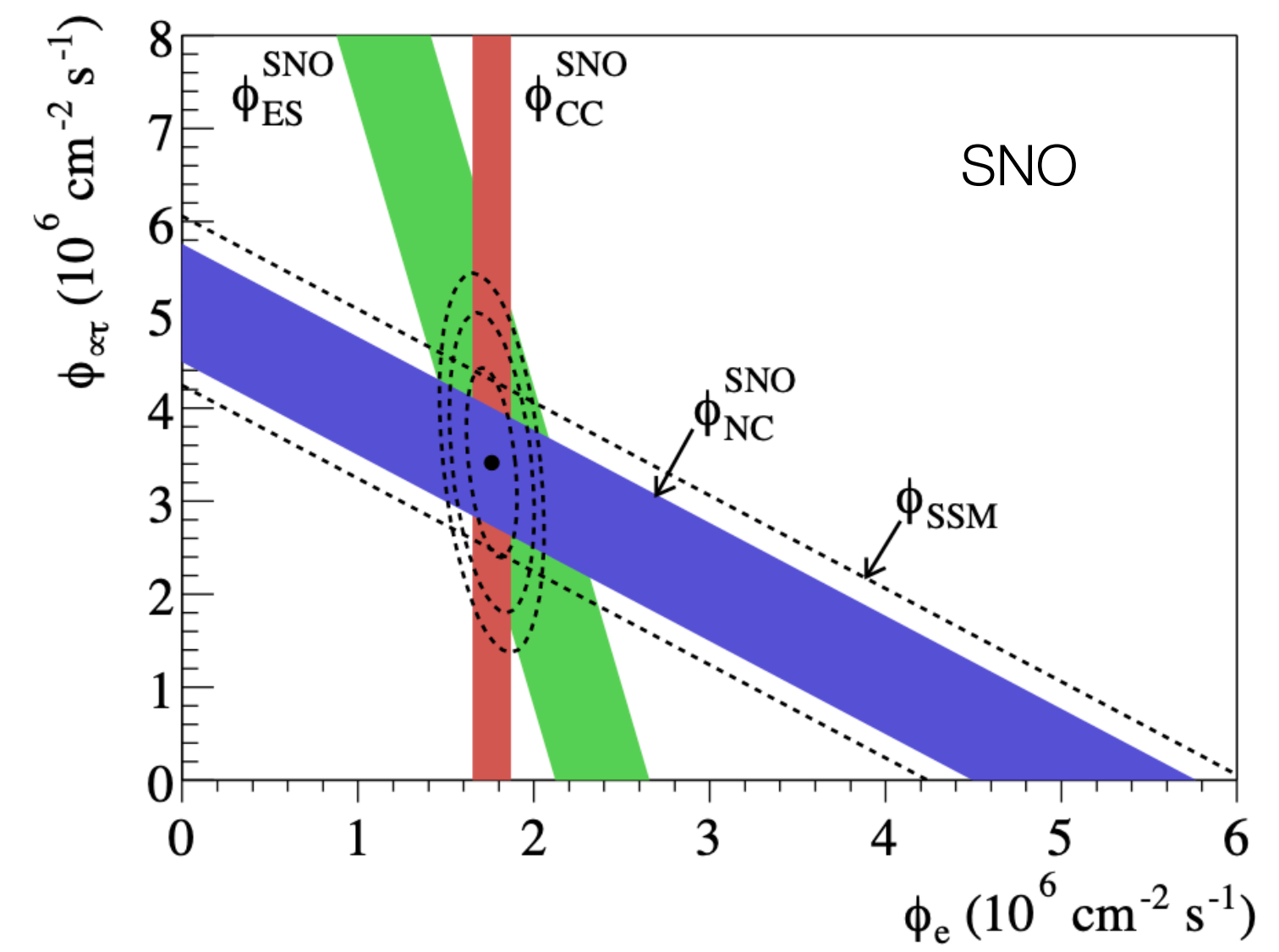


neutrino masses!



Works great

...but can't explain...



neutrino masses!

How can we tackle (some of) these questions?



# Plan

Neutrinos are massive—their masses are **different**

Probing the nature of neutrino masses with  $0\nu\beta\beta$

Current status, outlook, and other opportunities

# Neutrinos change flavour

Produce them in association with one type of charged lepton, see them in association with another

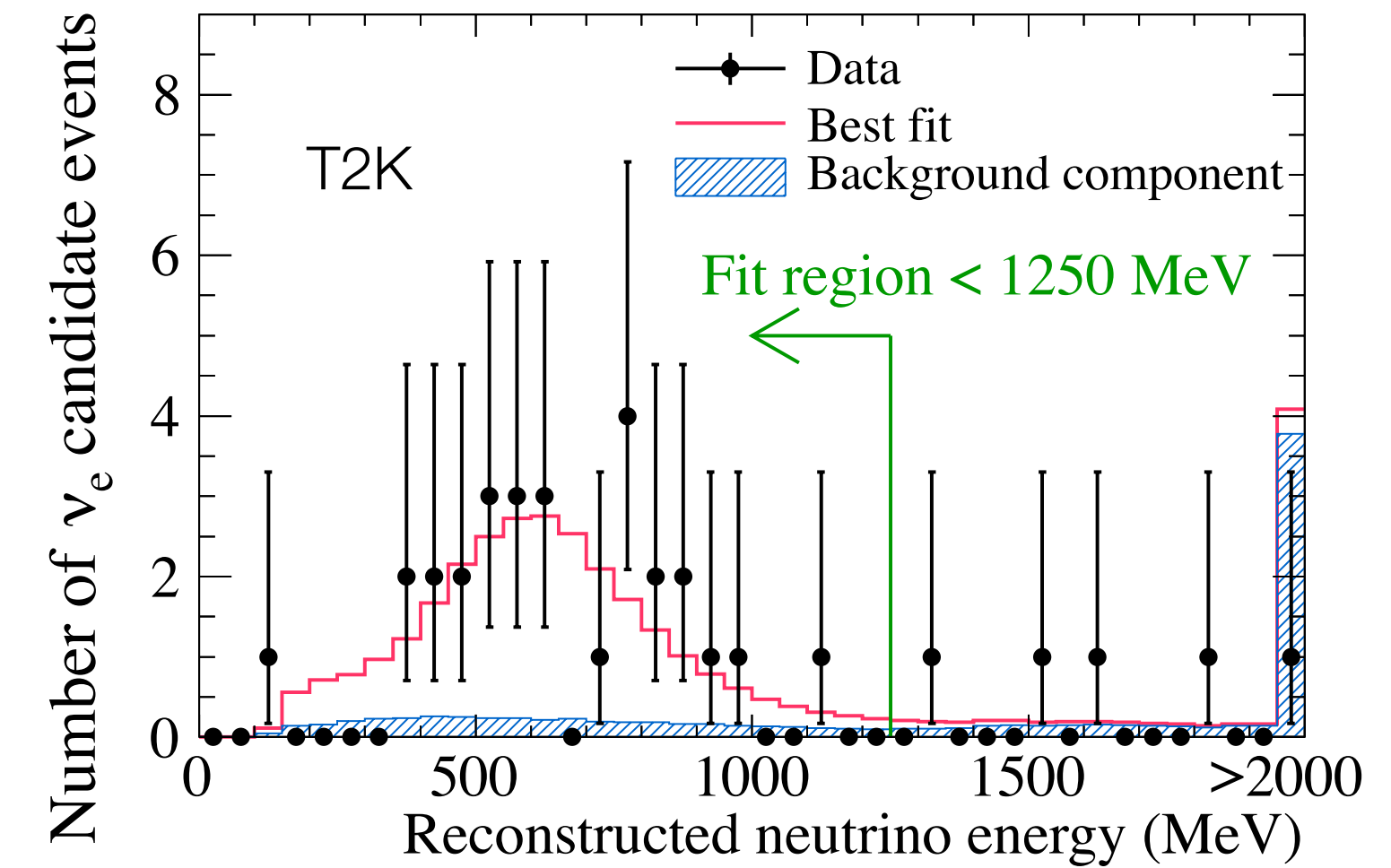
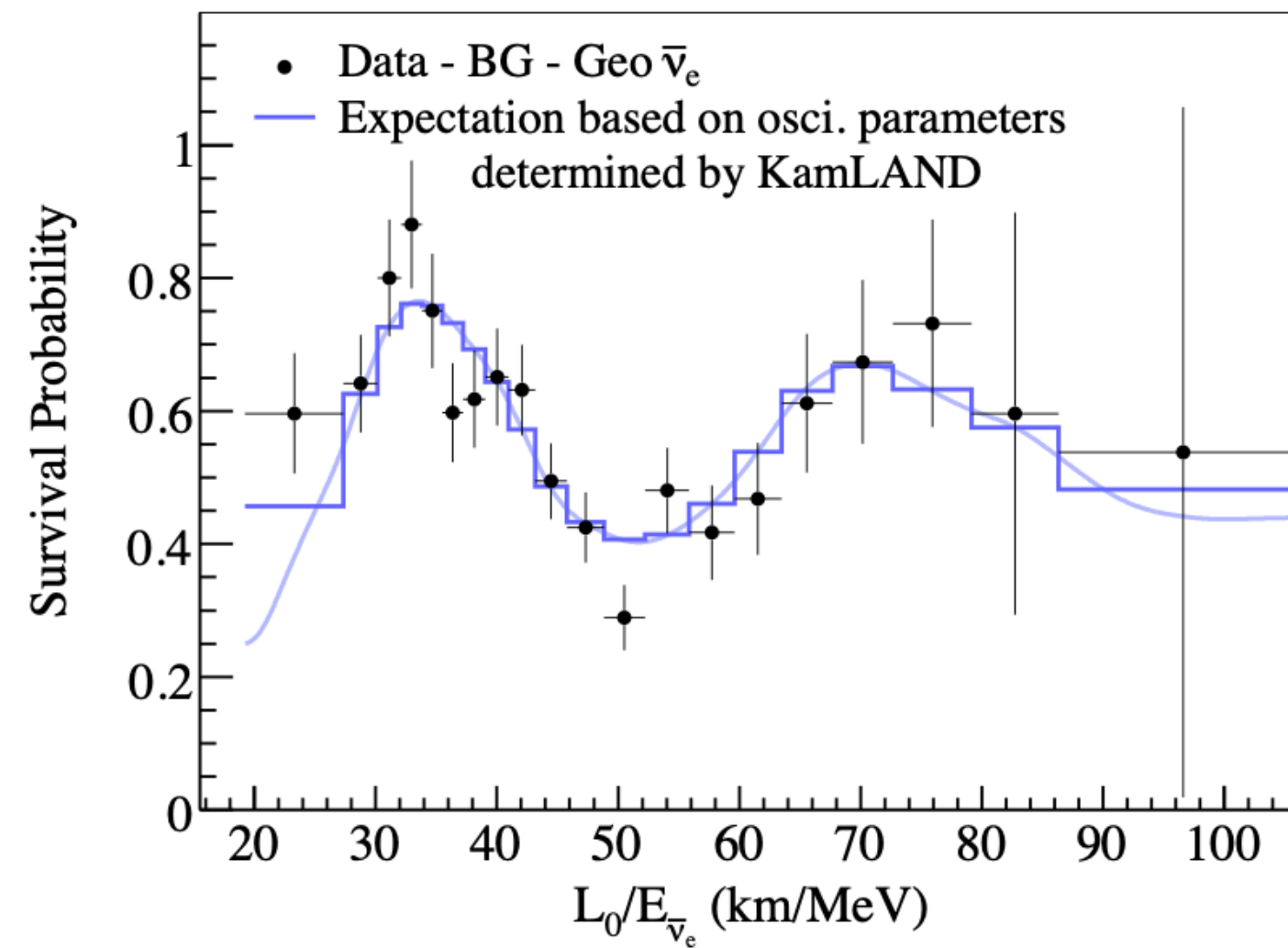
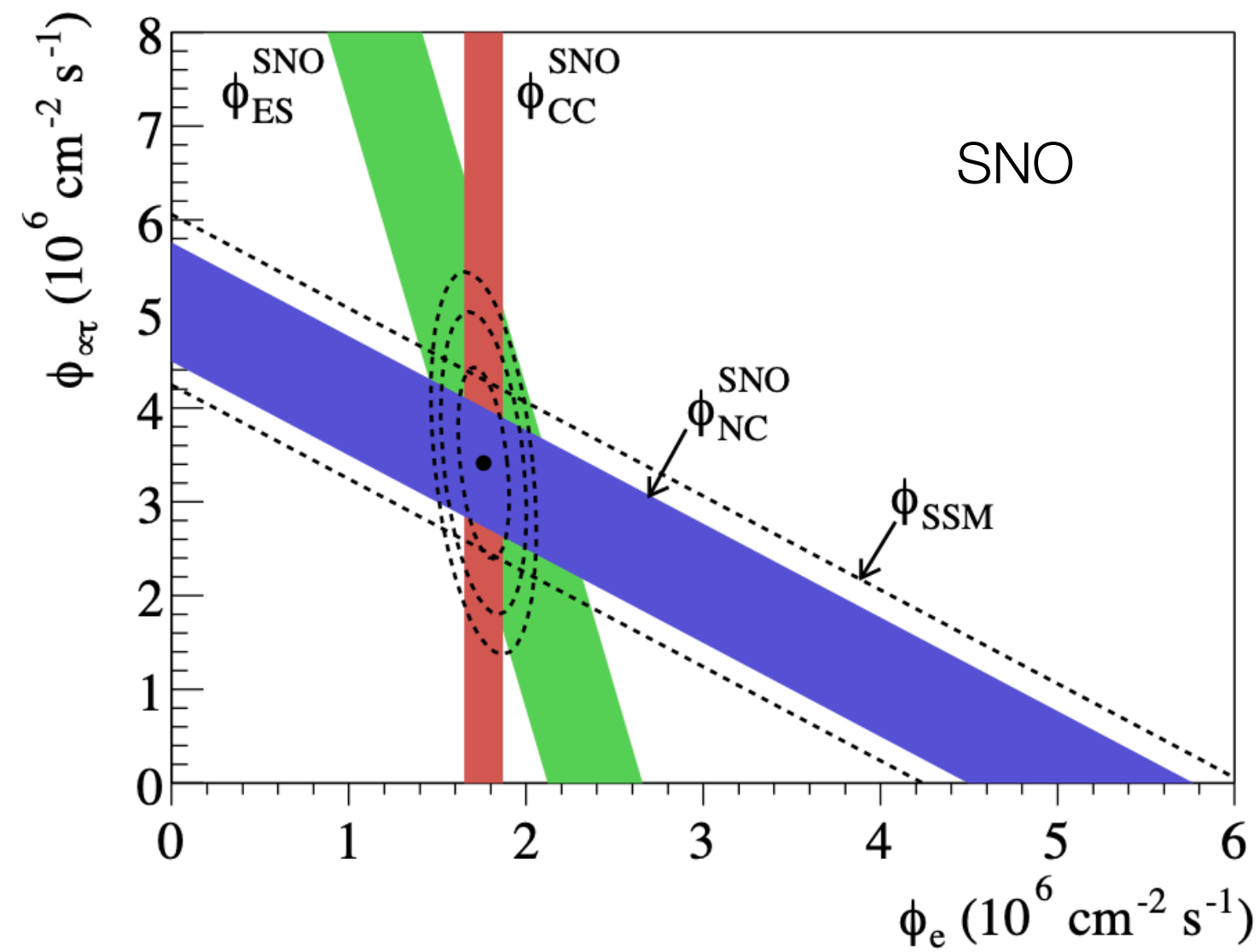
$$P_{\alpha \rightarrow \beta} \sim \left| \sum_i U_{\alpha i}^* U_{\beta i} e^{i \frac{m_i^2 L}{2E}} \right|^2$$

In the simple 2 neutrino case, relation between flavor and mass eigenstates:  $U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$

$$\Rightarrow P_{\alpha \rightarrow \beta} = 1 - P_{\alpha \rightarrow \alpha} = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

Seeing neutrino oscillations  $\Rightarrow$  they have mass

# Neutrinos have mass

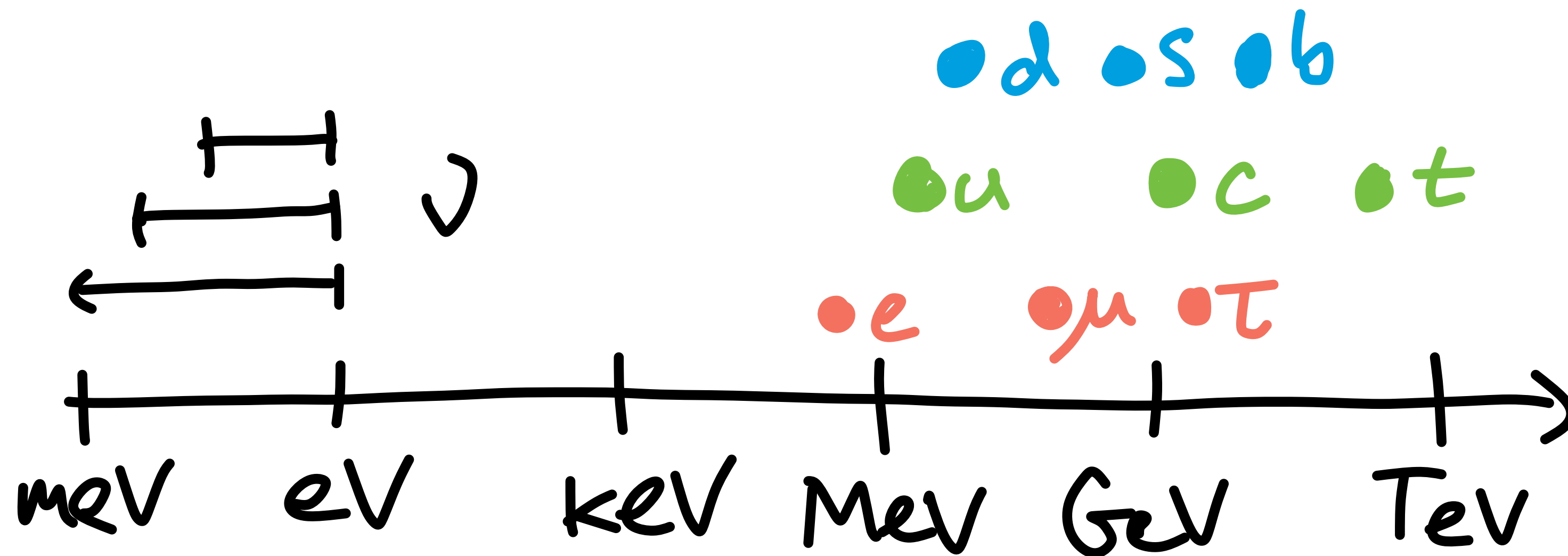


Have seen neutrino oscillations in a number of experimental setups

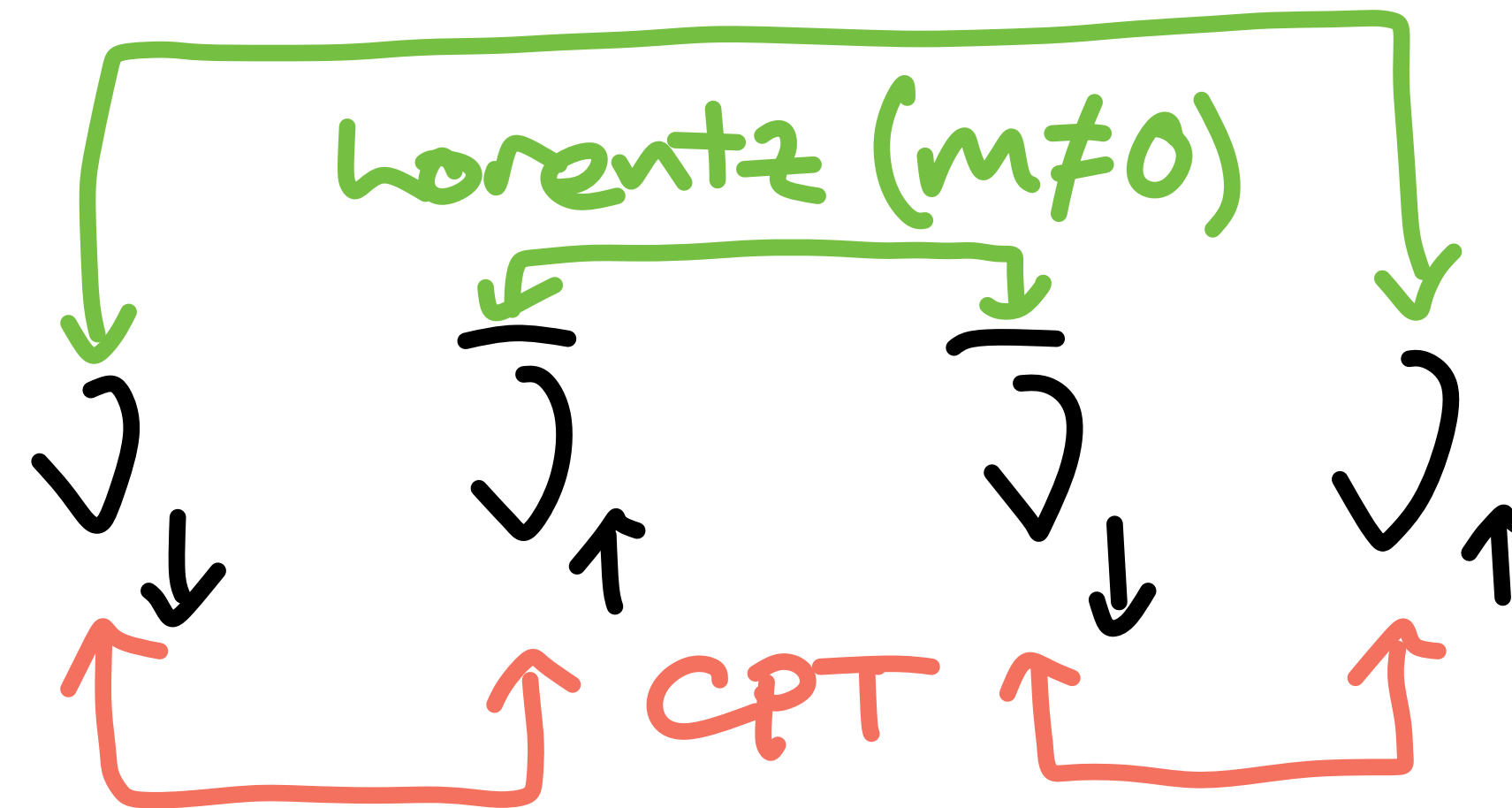
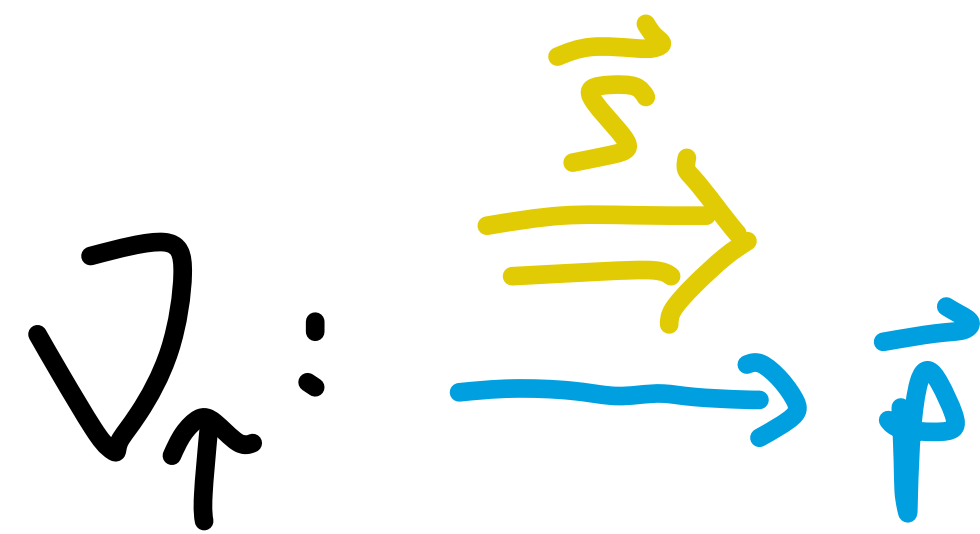
$$\text{Observed } \Delta m_{\text{atm}}^2 = \left| \Delta m_{13}^2 \right| \sim 10^{-3} \text{ eV}^2 \text{ and } \Delta m_{\odot}^2 = \Delta m_{12}^2 \sim 10^{-5} \text{ eV}^2$$

# Neutrinos have mass

But they are **very** light

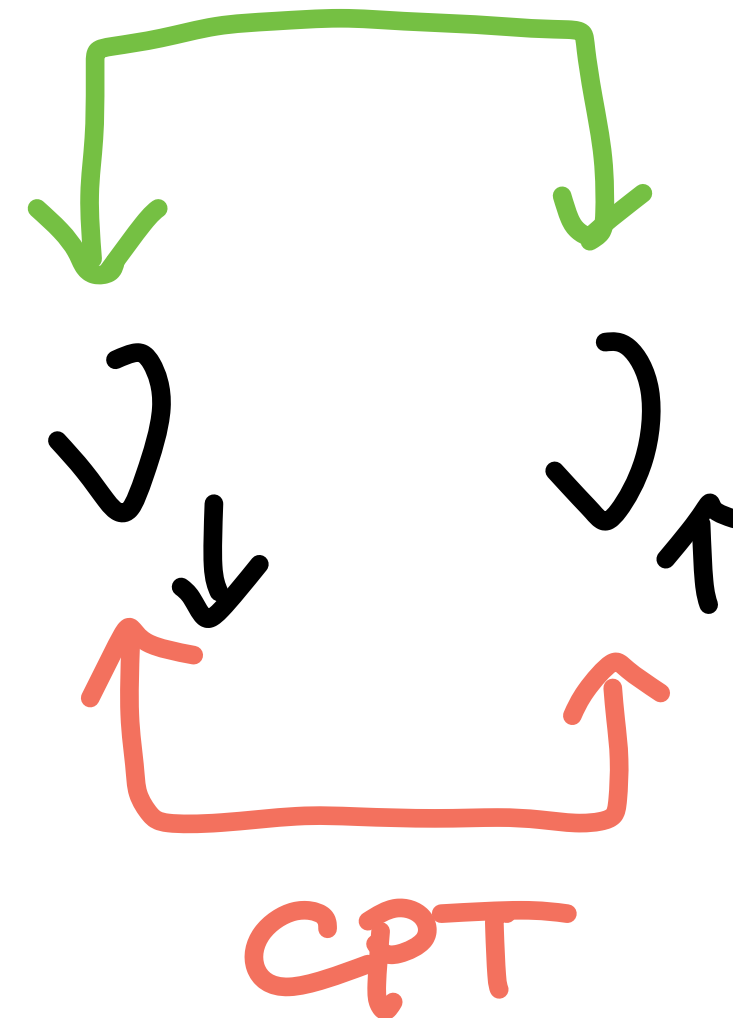


# Neutrino masses: Dirac or Majorana?



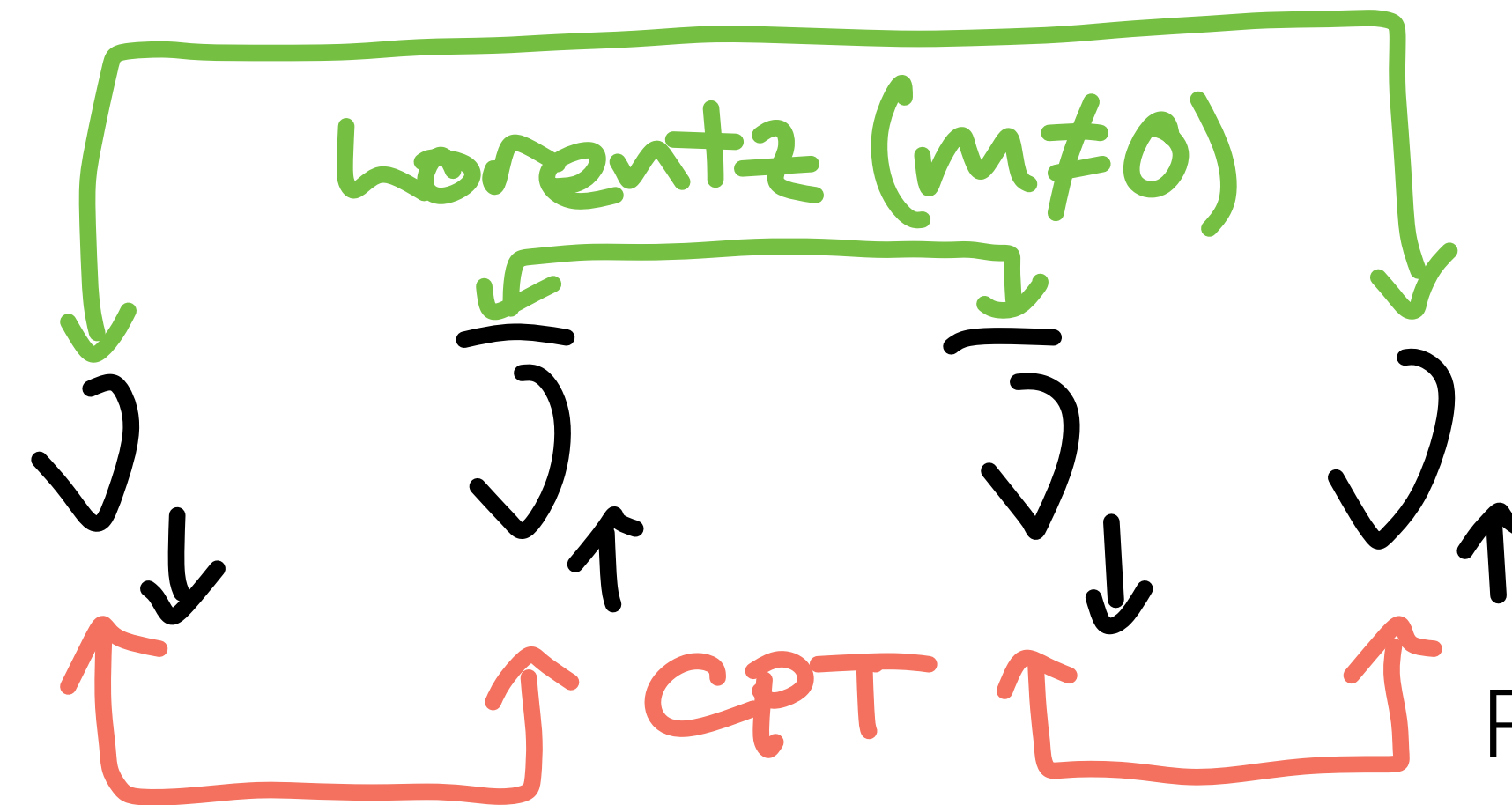
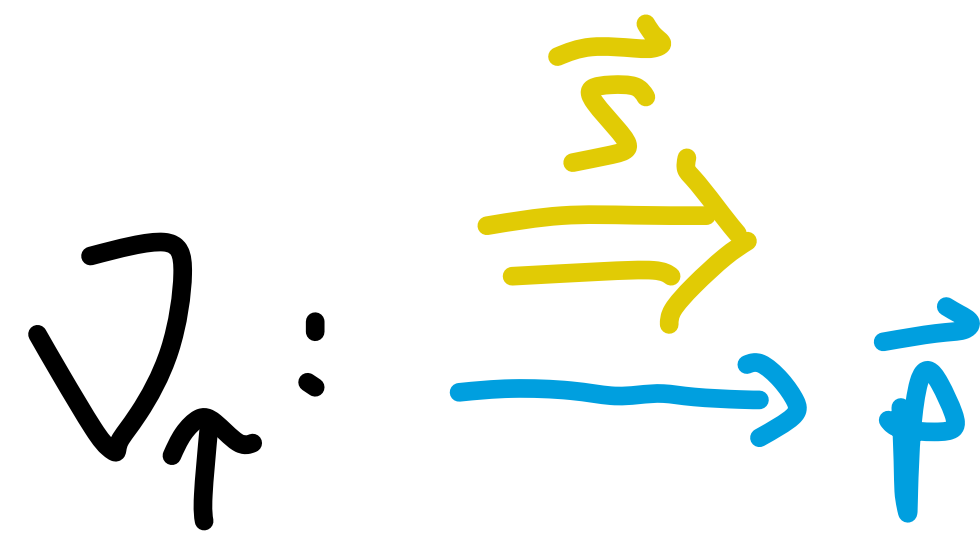
Dirac  
4 states

Lorentz ( $m \neq 0$ )

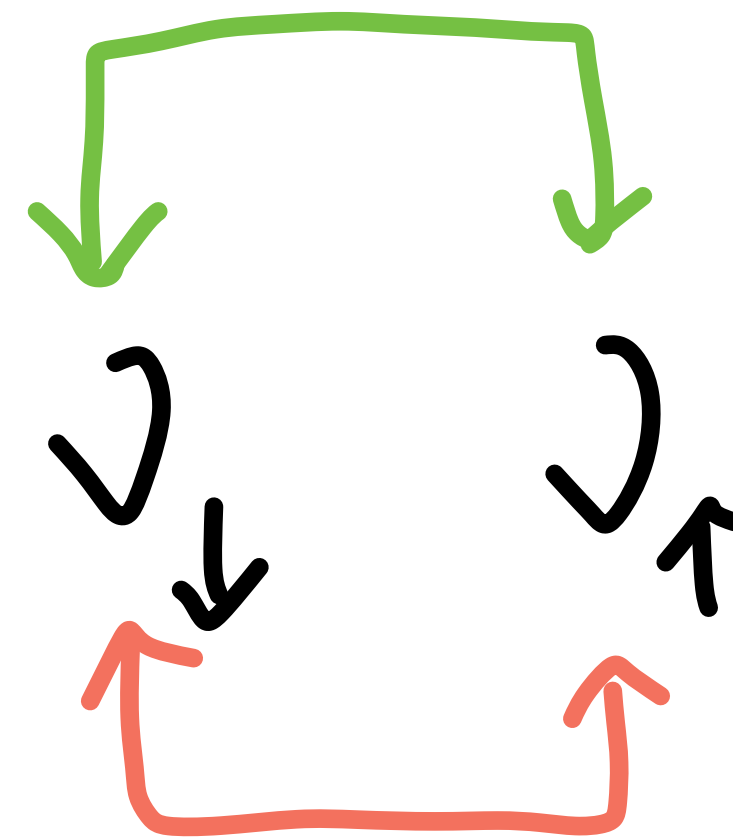


Majorana  
2 states

# Neutrino masses: Dirac or Majorana?



Lorentz ( $m \neq 0$ )



CPT

Dirac  
4 states

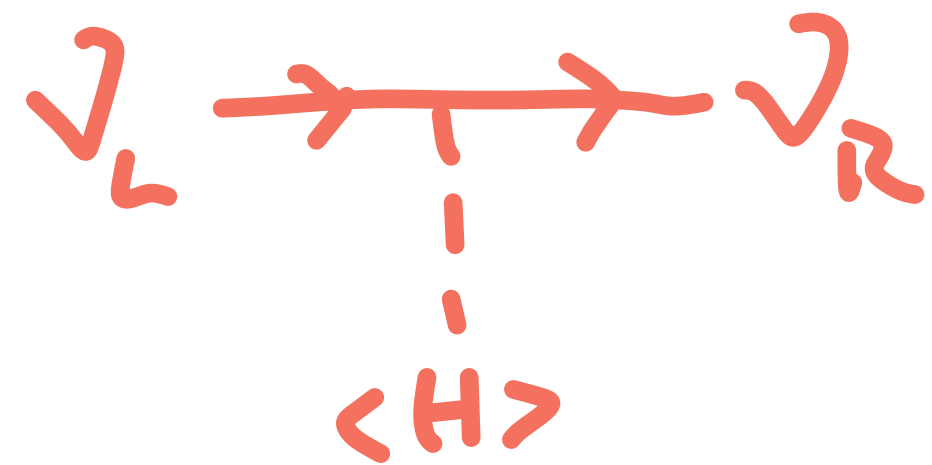
Pair of degenerate Weyl spinors  
Degeneracy enforced by  $U(1)$

$$\nu \rightarrow e^{i\theta}\nu, \bar{\nu} \rightarrow e^{-i\theta}\bar{\nu}$$

Majorana  
2 states

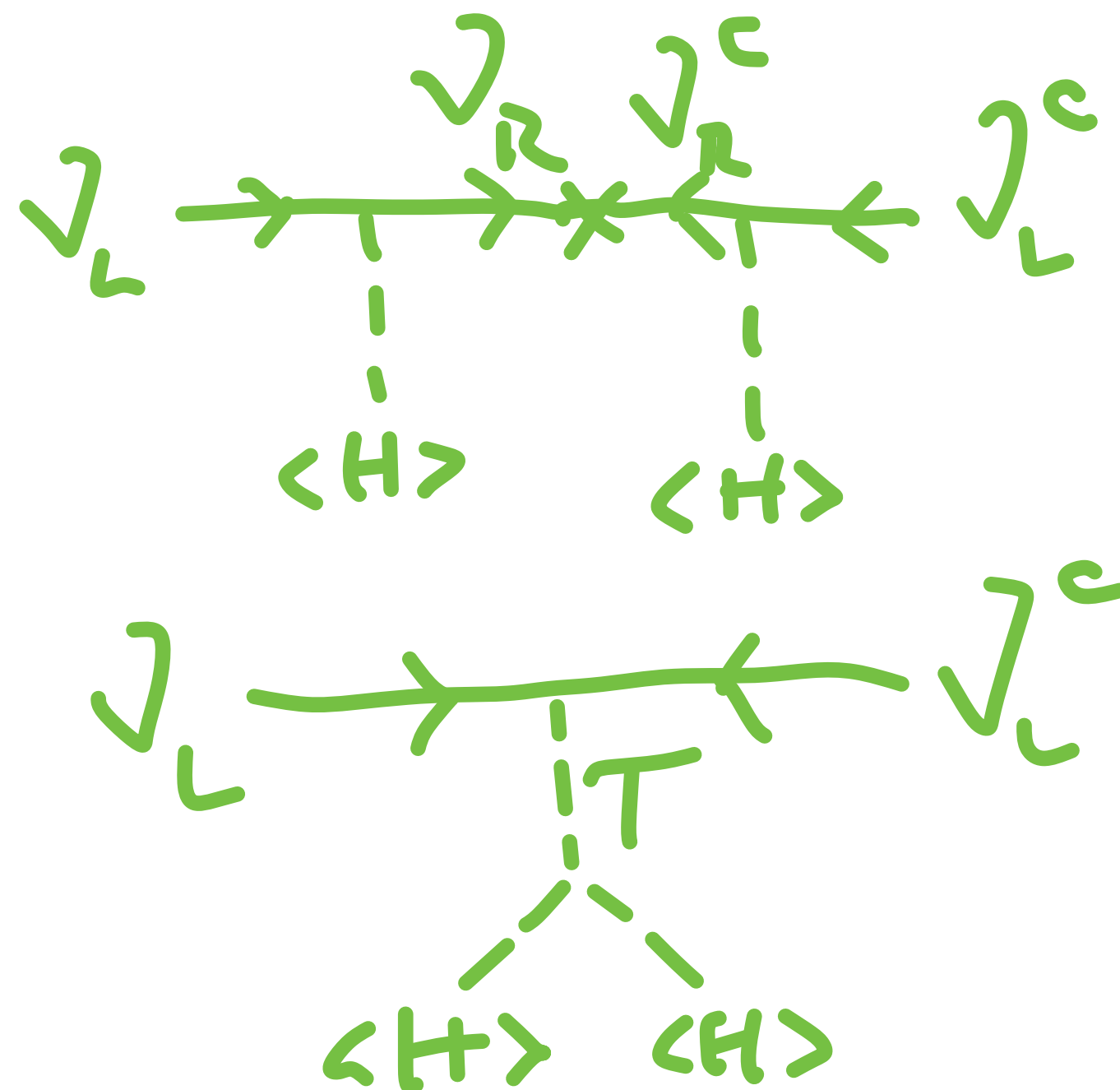
Single Weyl spinor

# Neutrino masses: Dirac or Majorana?



$$-\mathcal{L} \supset \underbrace{m \bar{\nu}_R \nu_L}_{\Delta L = 0} + \text{h.c.}$$

Dirac  
4 states



$$-\mathcal{L} \supset \underbrace{m \bar{\nu}_L^c \nu_L}_{\Delta L = 2} + \text{h.c.}$$

Majorana  
2 states

Each option requires degrees of freedom not seen in SM  
Neutrino masses are **qualitatively different**

# Lepton number

Lepton number is an accidental symmetry of the renormalizable standard model  
(it is a consequence of choice of gauge group and charges)

Individual lepton numbers  $L_e, L_\mu, L_\tau$  violated in neutrino oscillation experiments  
[similar story for quarks — only total baryon number (accidentally) conserved]

Lepton number is a global symmetry  $\Rightarrow$  no gauge field associated with it

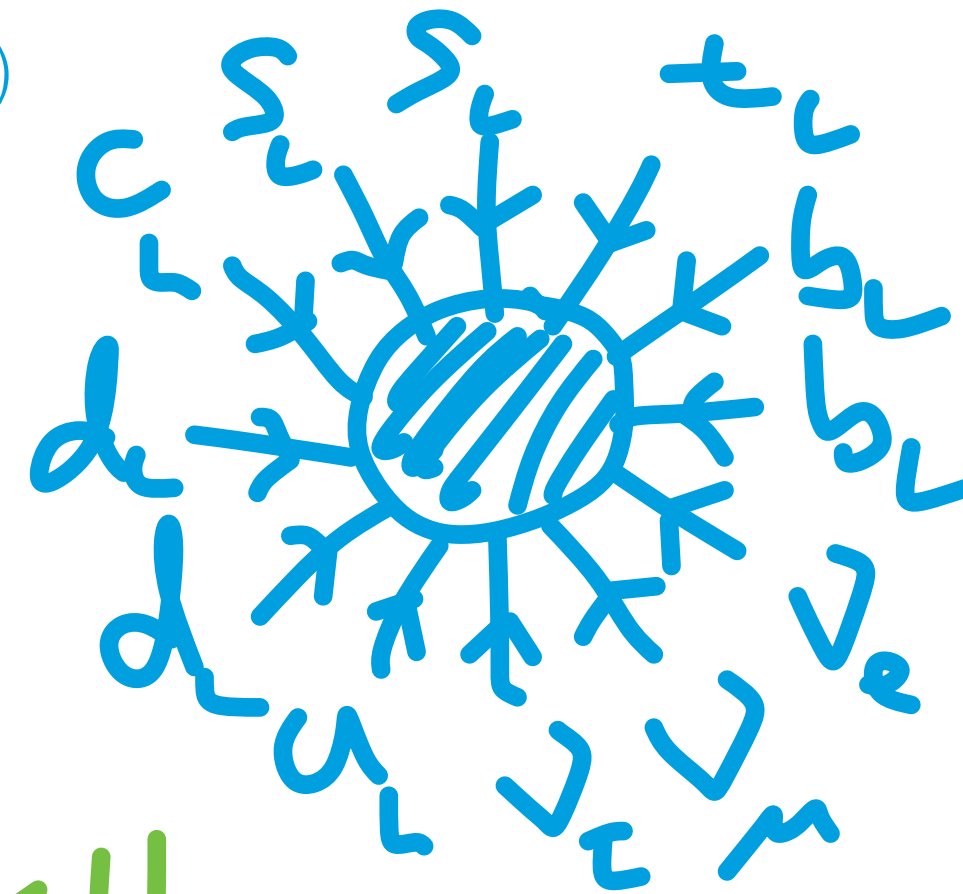
Should we expect lepton number to be conserved?



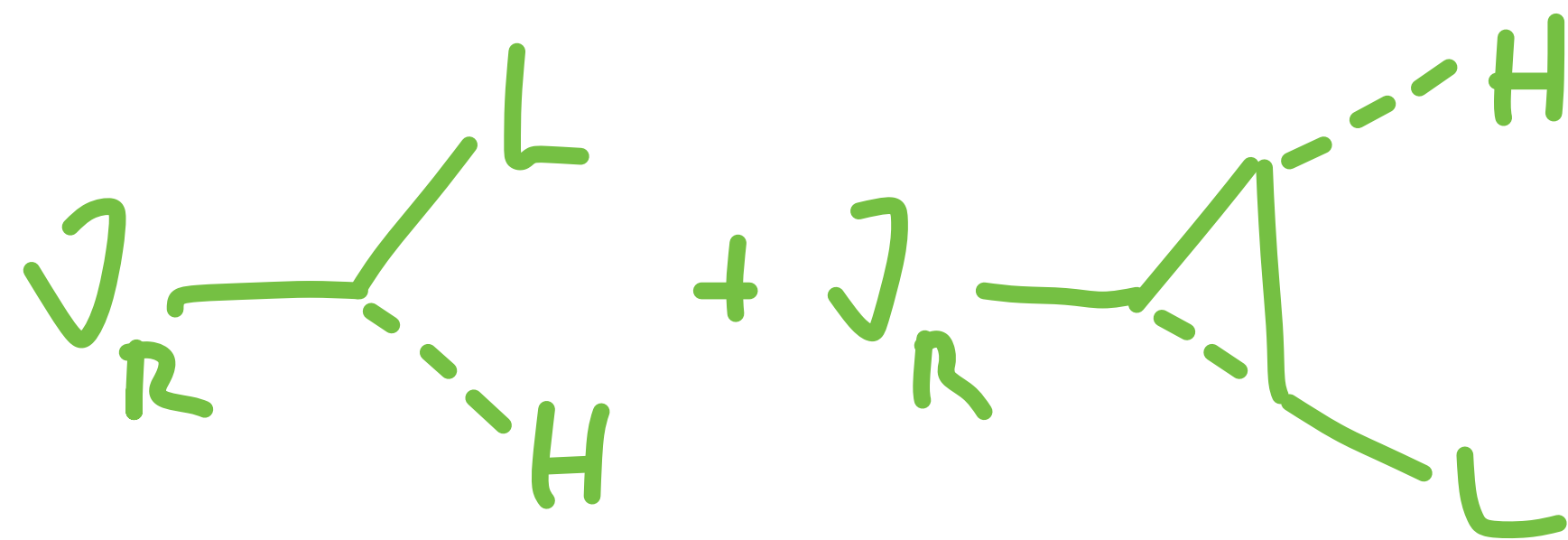
# Lepton number violation

(Kuzmin, Rubakov, Shaposhnikov)

Lepton number is violated by the standard model nonperturbatively  
( $B - L$  conserved)



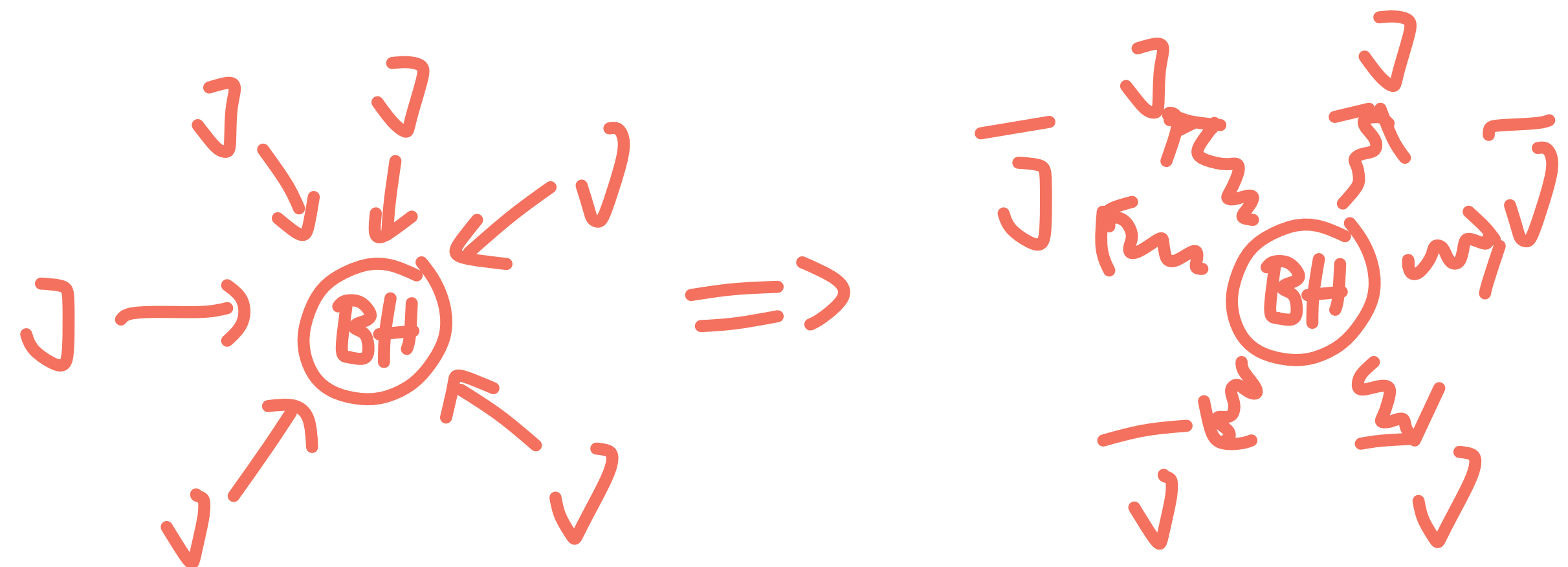
and at dim-5  $(HL)^2$  [Weinberg]



(Fukugita, Yanagida)

Lepton number violation could be intimately related to the baryon asymmetry of the universe

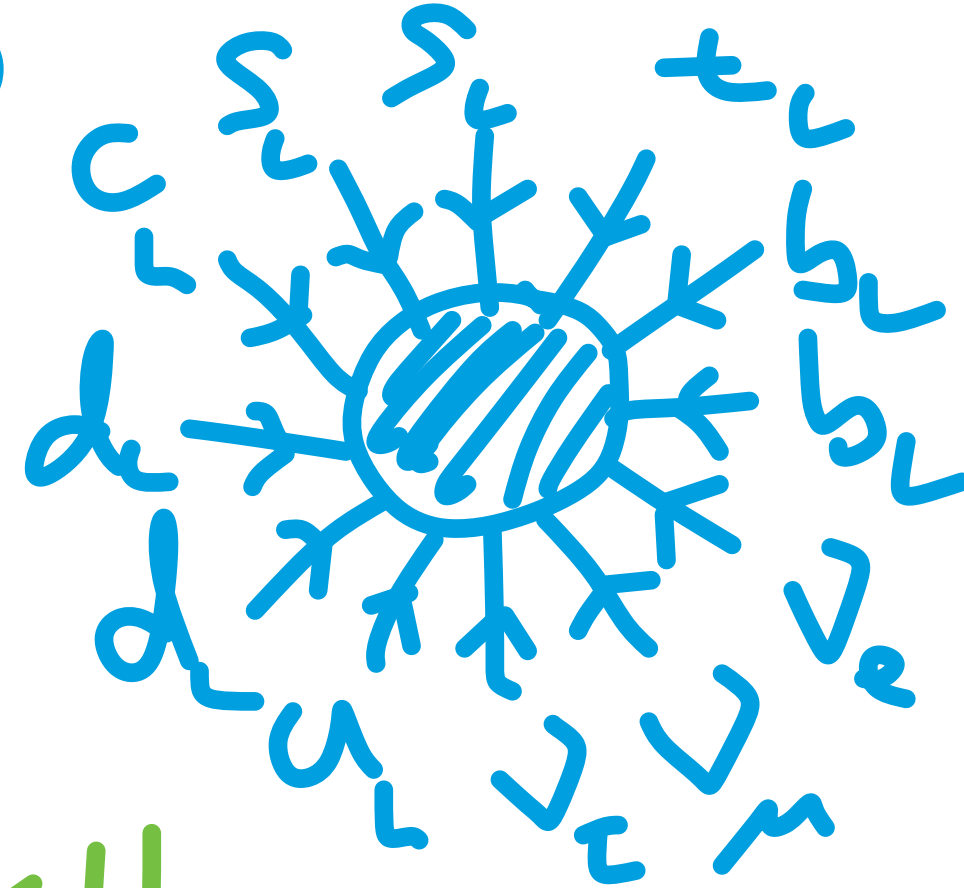
Gravity does not respect global symmetries



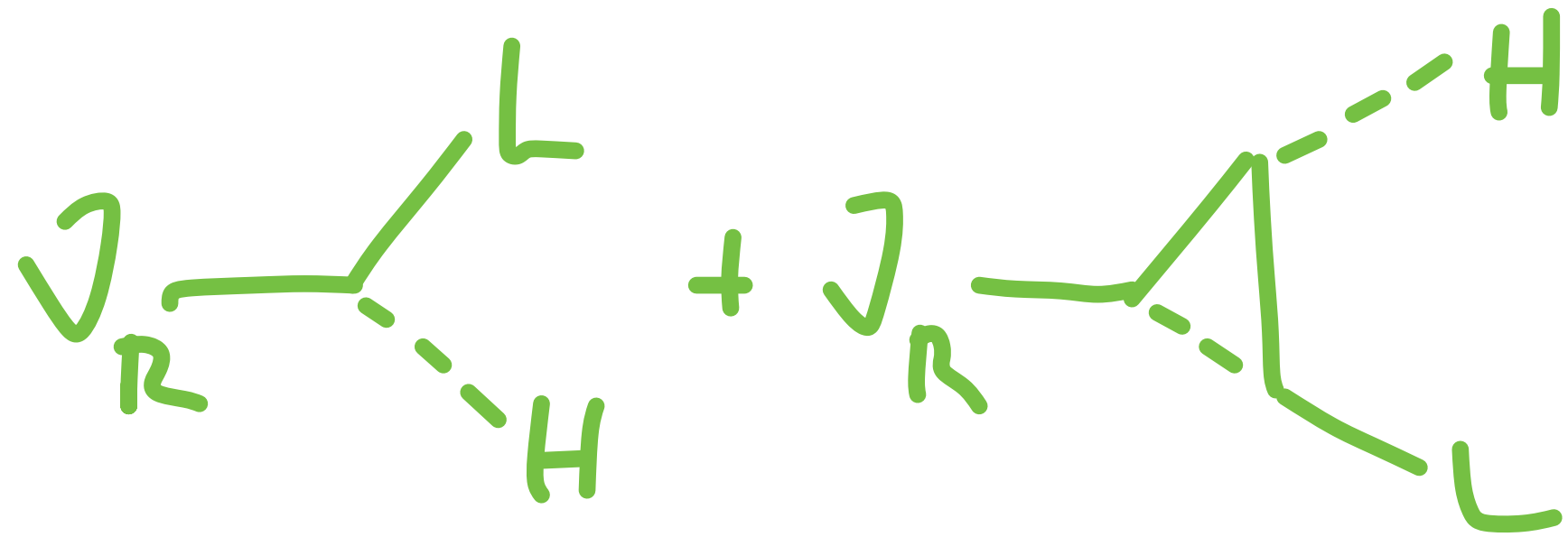
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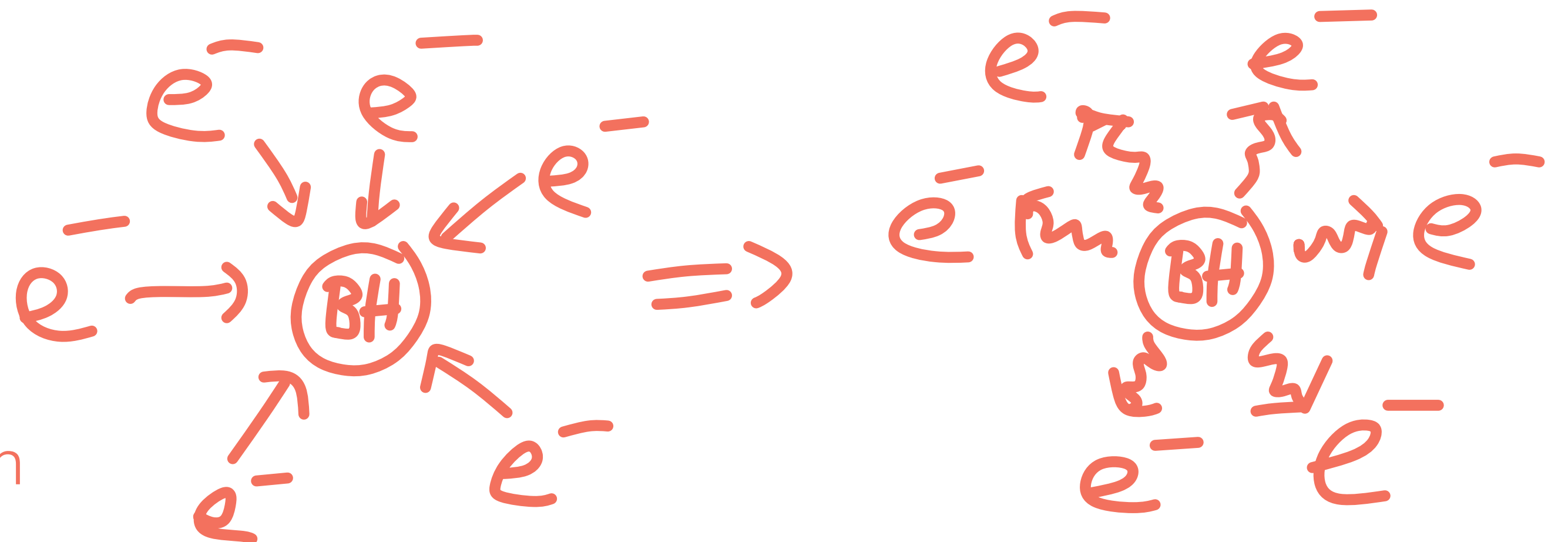


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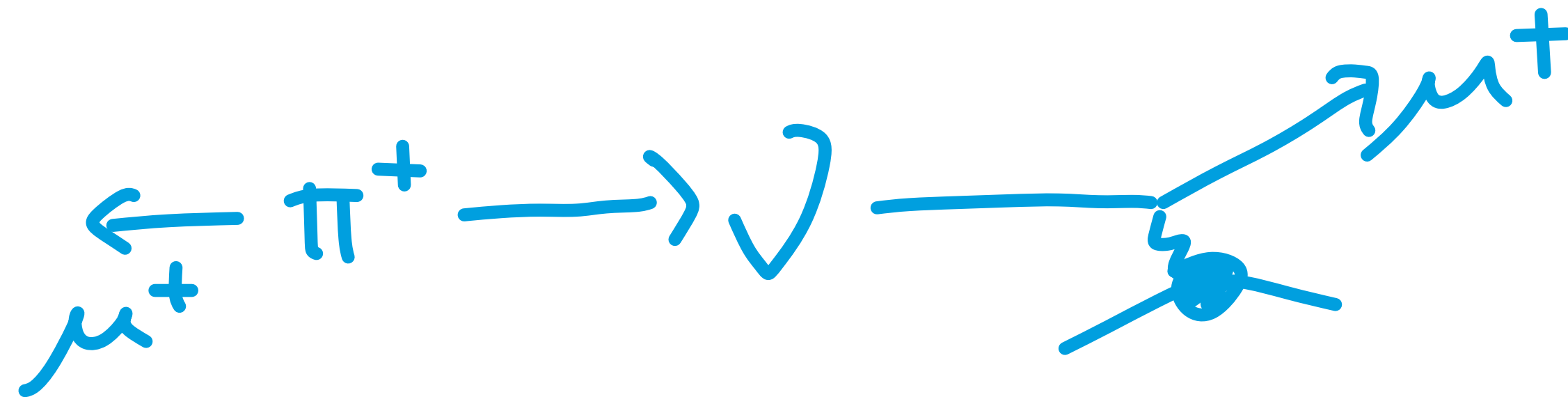
Contrast with



# How to find lepton number violation

How would a particle theorist do it? Kayser: use  $\pi^+ \rightarrow \mu^+ \nu$  to create  $\nu$  beam.

Check if this beam creates  $\mu^+$  in scattering on target

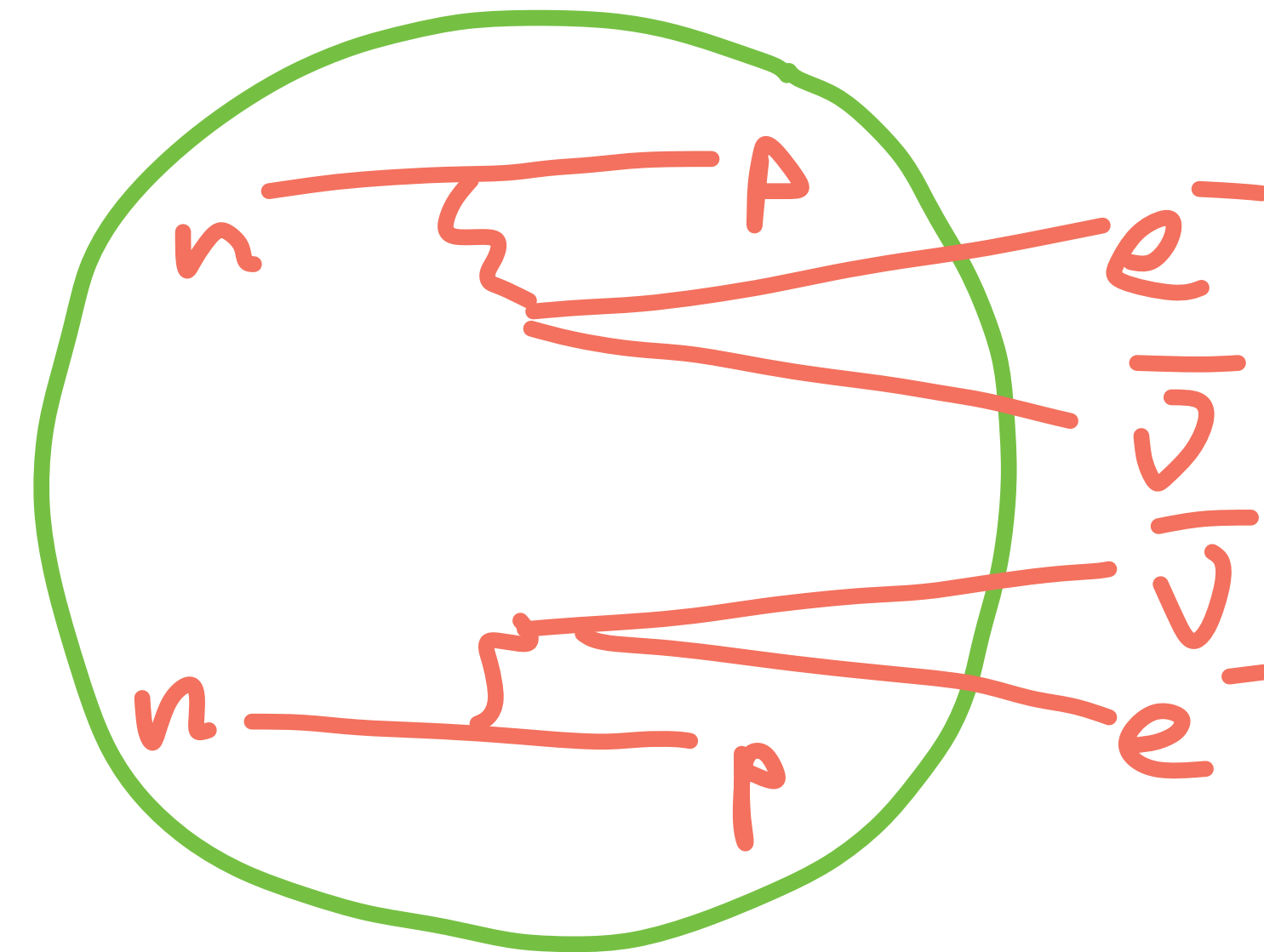
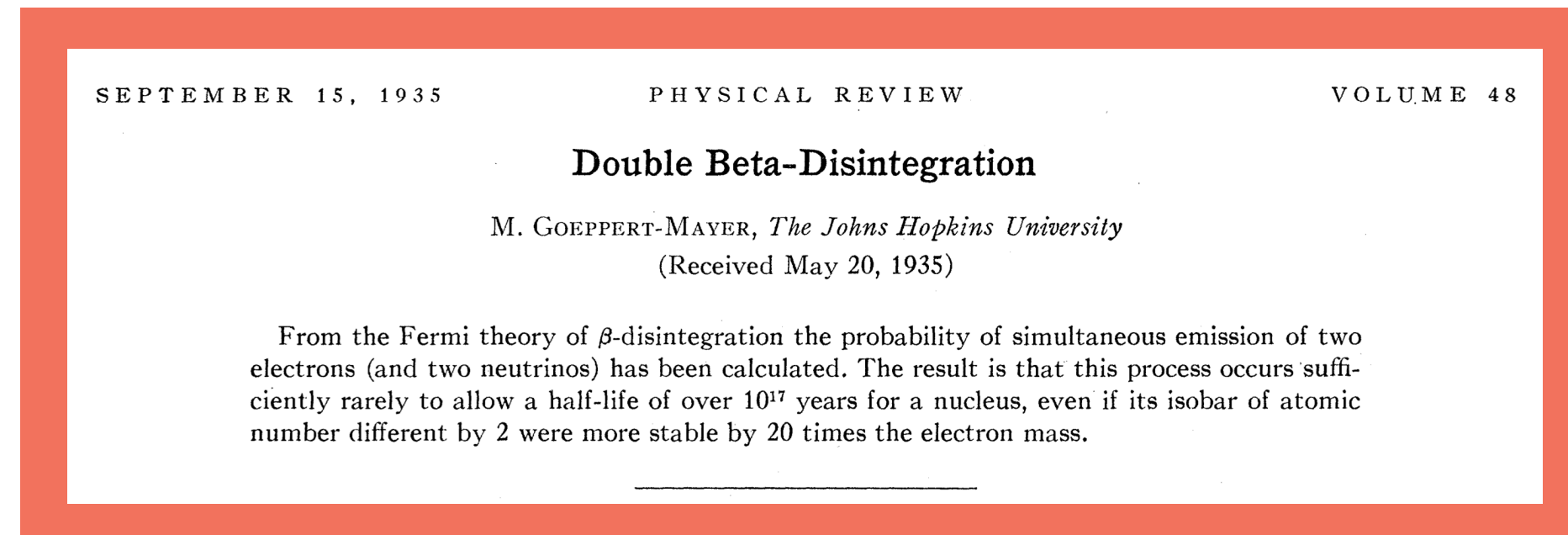
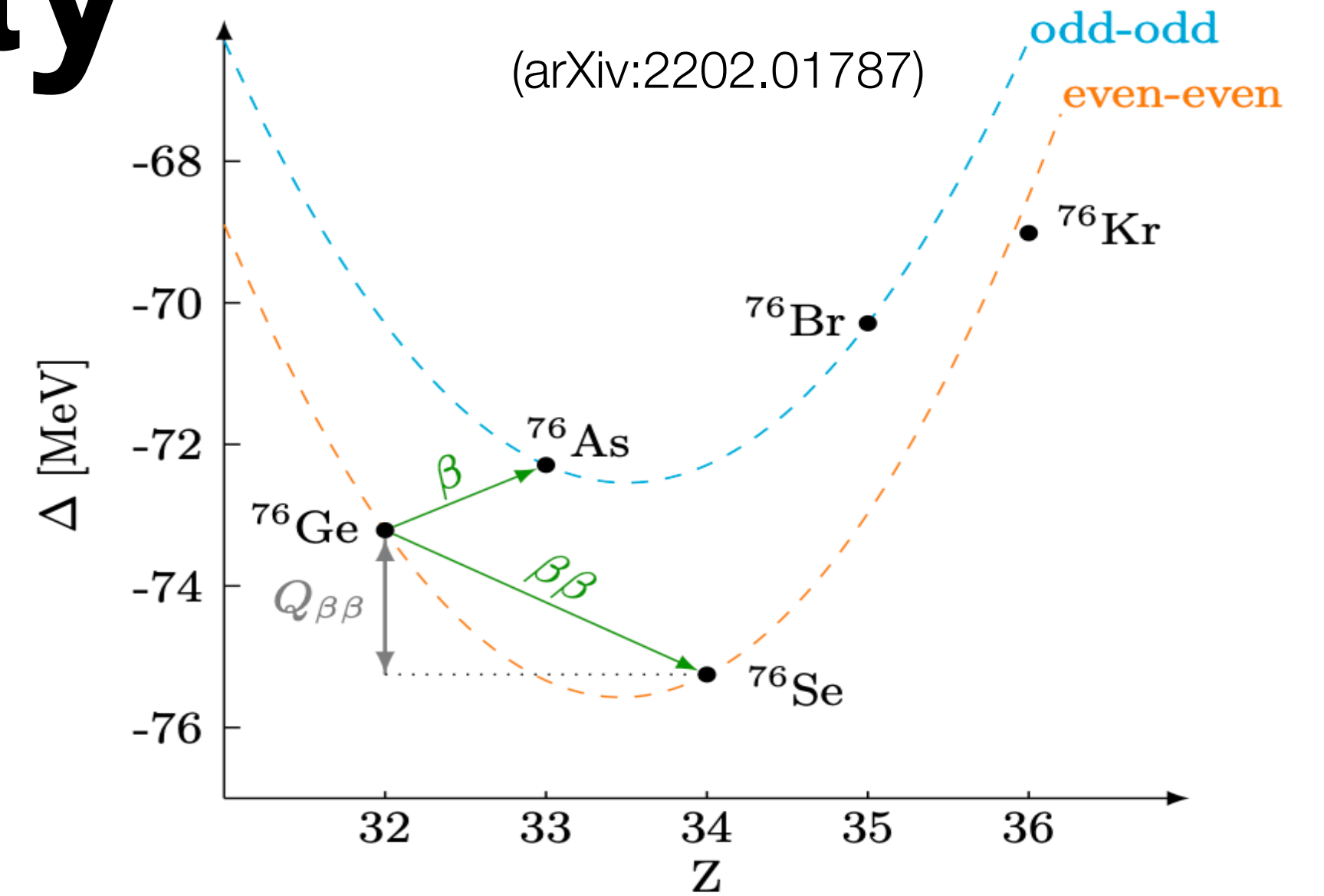


Problem: need helicity flip, rate  $\propto \left( \frac{m_\nu}{100 \text{ MeV}} \right)^2 \sim 10^{-16}$

Solution: use nuclei and let Avogadro help!

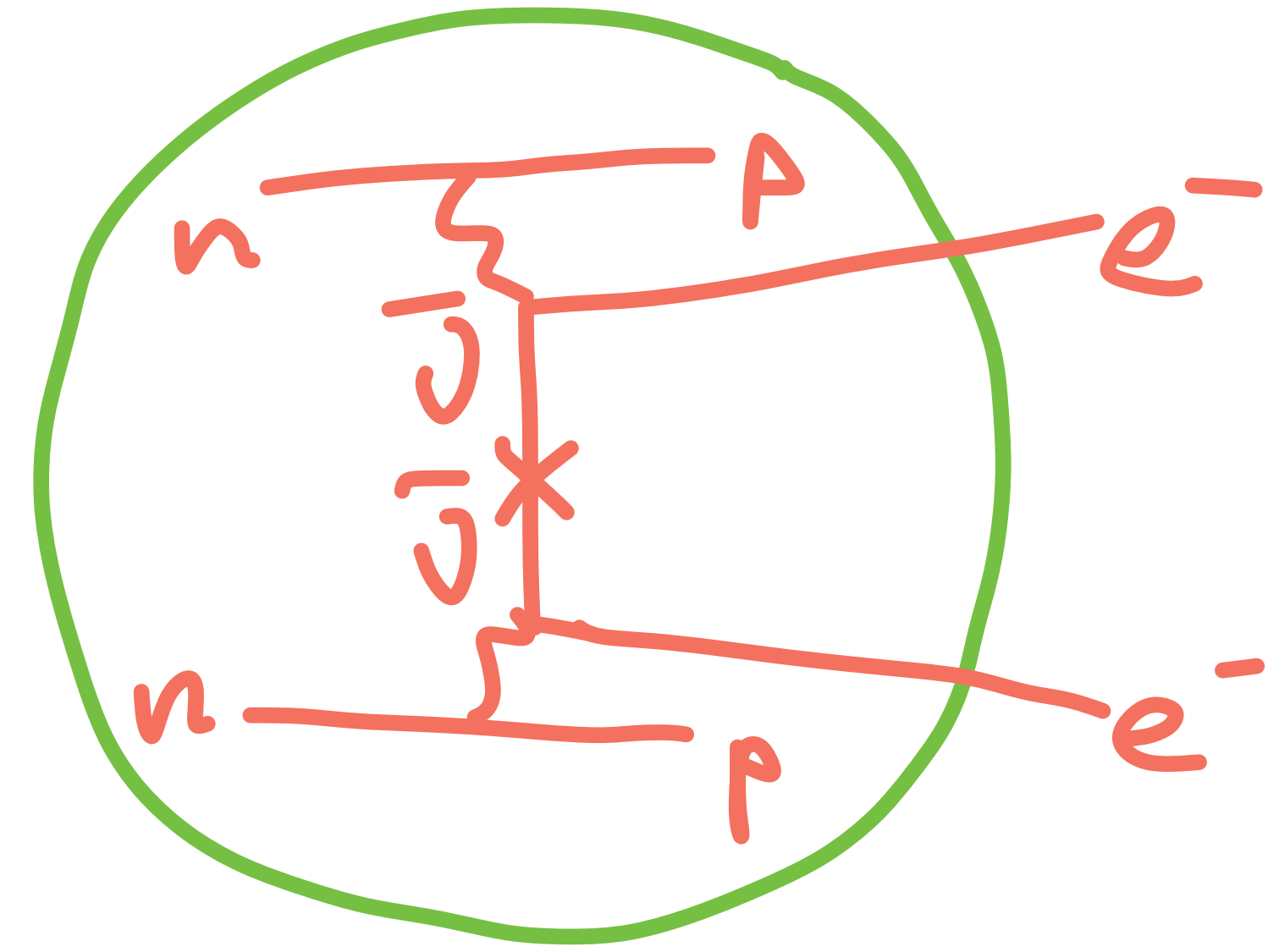
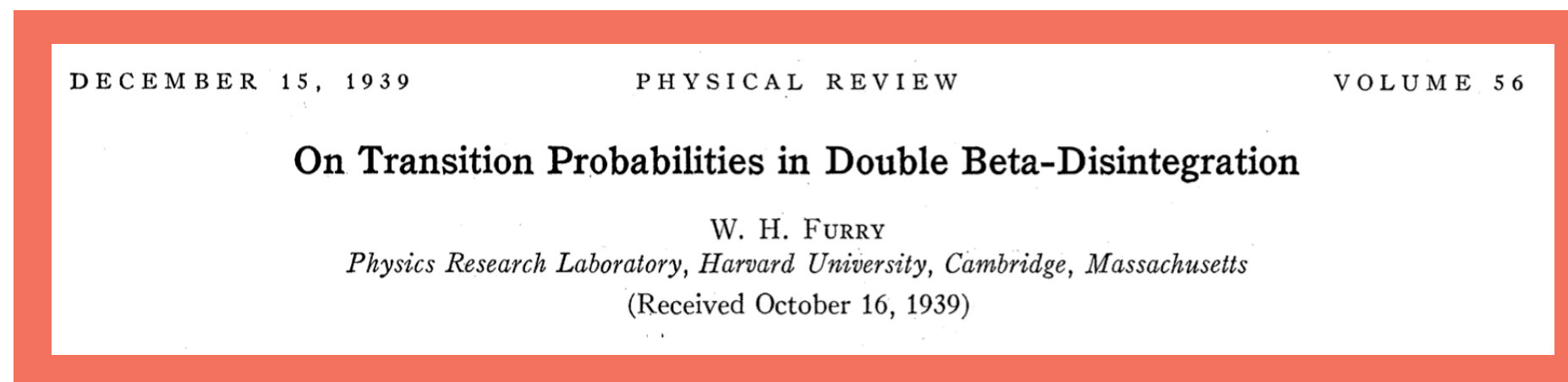
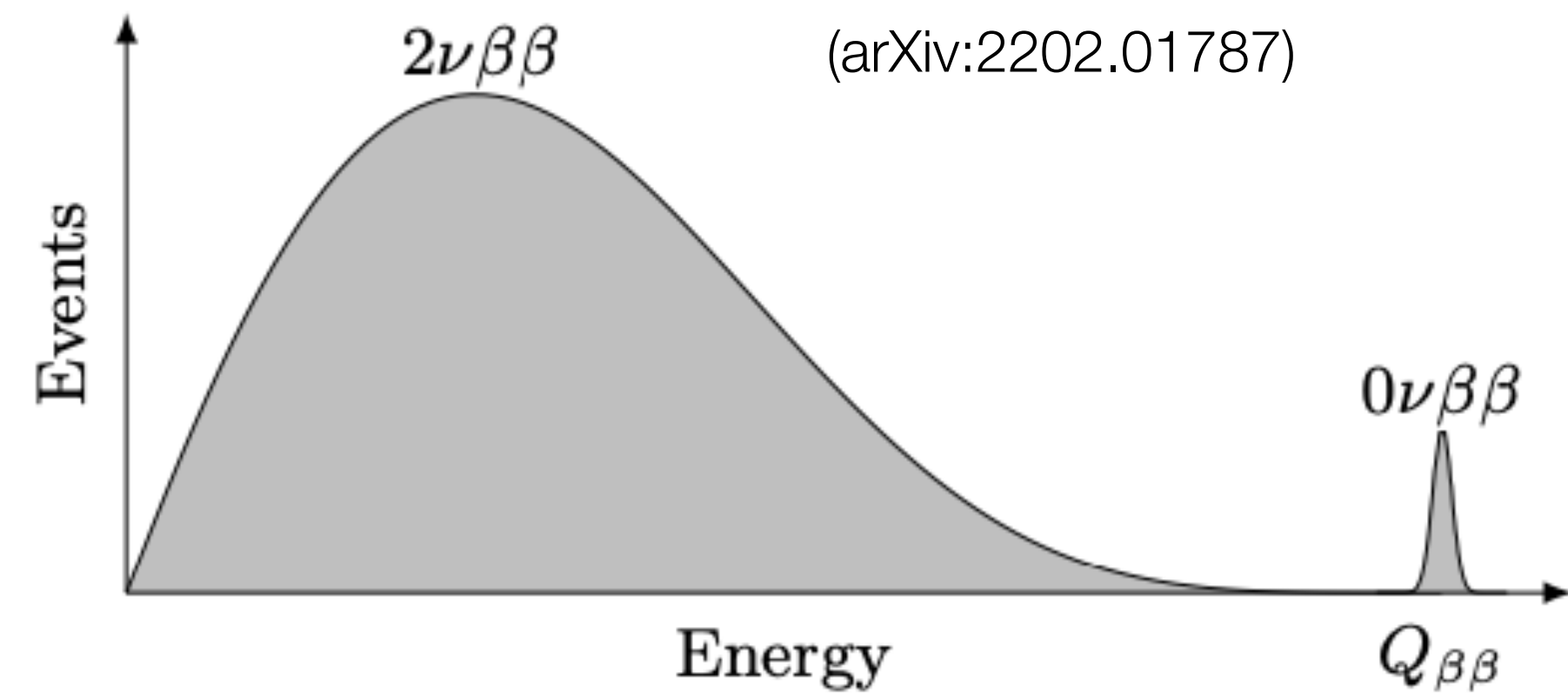
# $\beta\beta$ decay

Some even-even nuclei are energetically forbidden from  $\beta$  decay (or it is highly suppressed), have to undergo  $\beta\beta$  decay

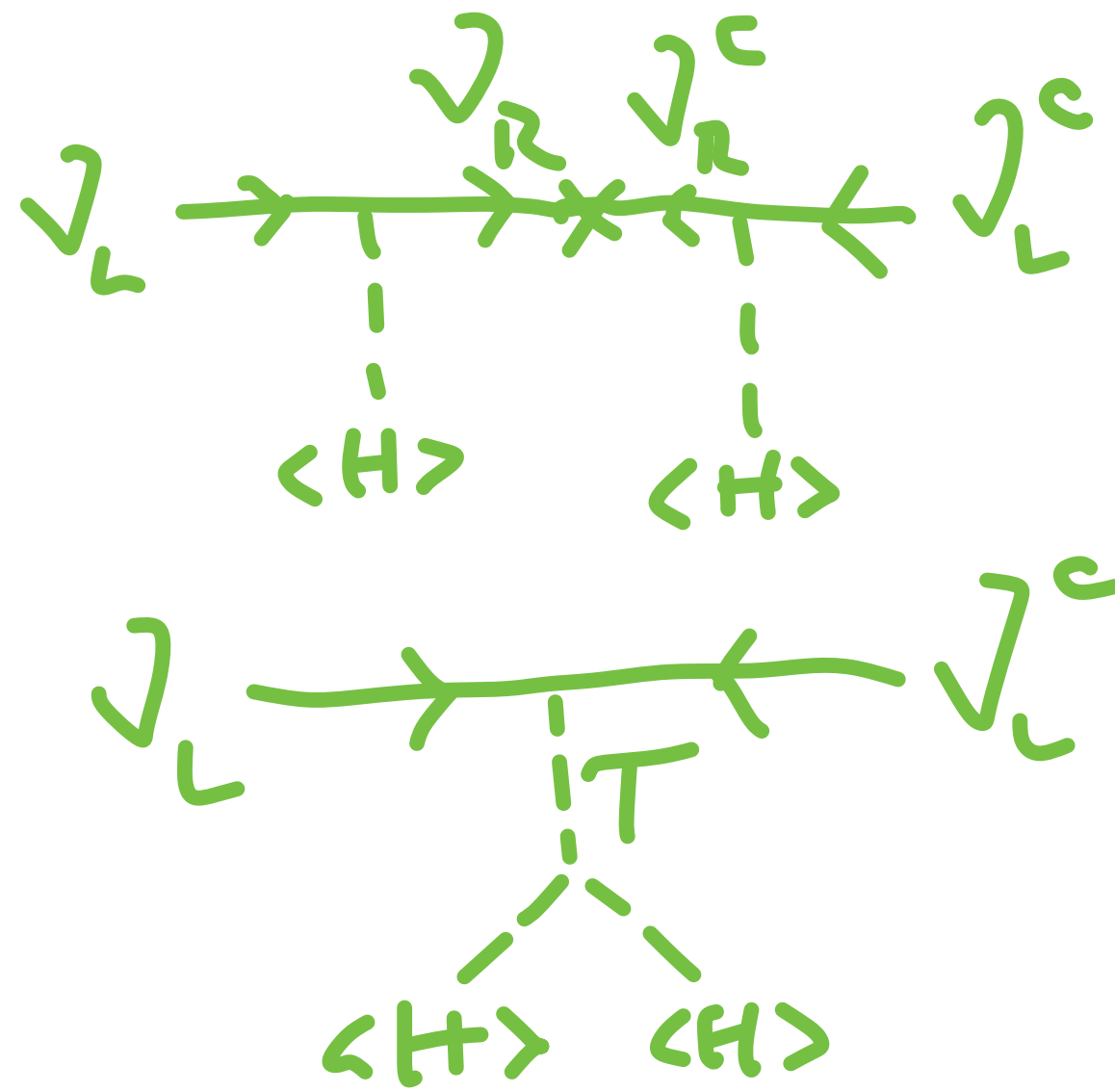


# $0\nu\beta\beta$ decay

If lepton number is violated can also have process with no neutrinos emitted



# High-scale seesaw



$$-\mathcal{L} \supset \underbrace{m \bar{\nu}_L^c \nu_L}_{\Delta L = 2} + \text{h.c.}$$

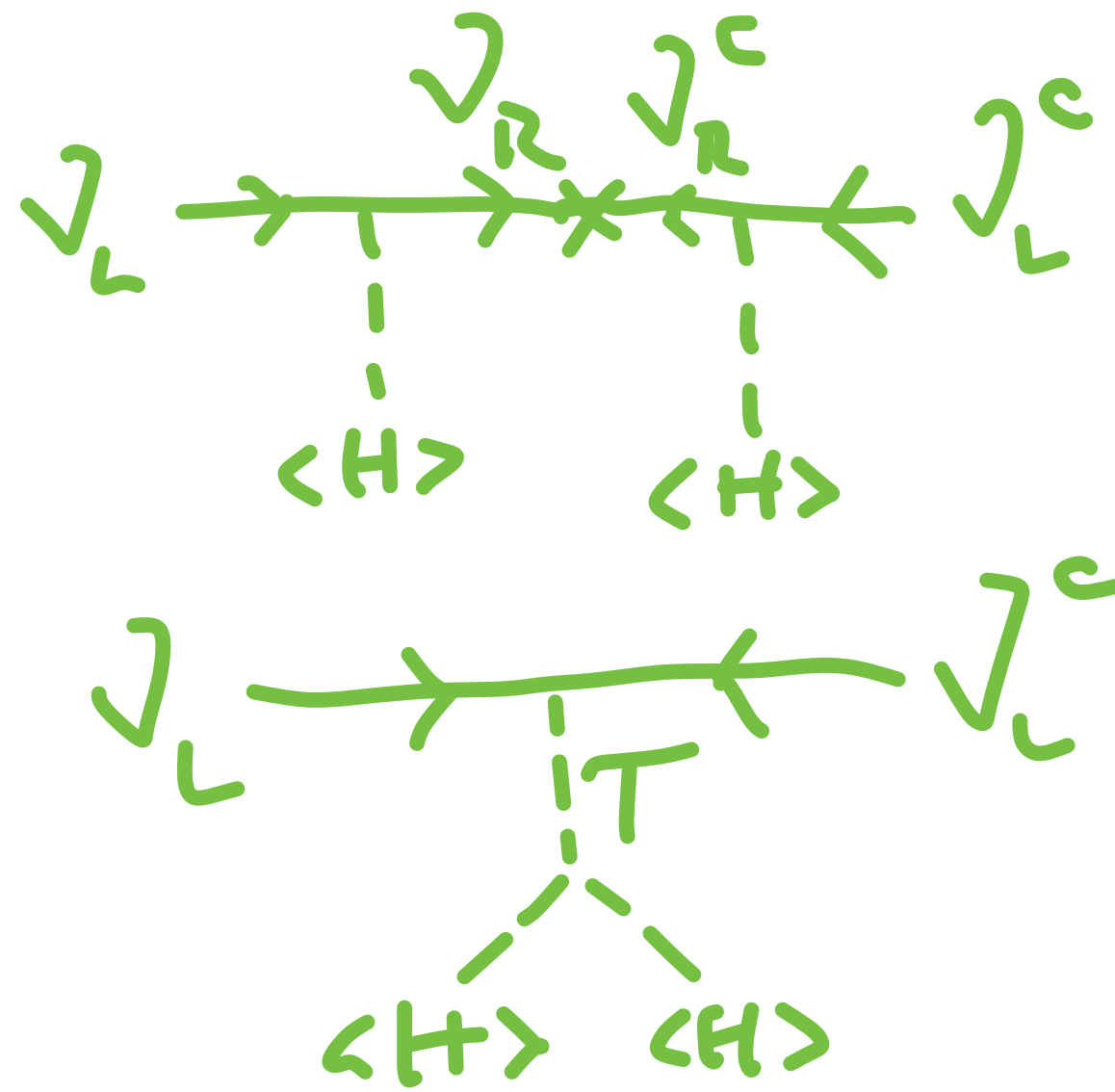
Neutrino mass comes from coupling to heavy SM singlets

(can explain smallness of  $\nu$  masses,

$$m \sim \frac{y^2 \langle H \rangle^2}{M}, \text{ another reason to expect LNV})$$

Key prediction of leptogenesis as an explanation of baryon asymmetry of the universe

# High-scale seesaw



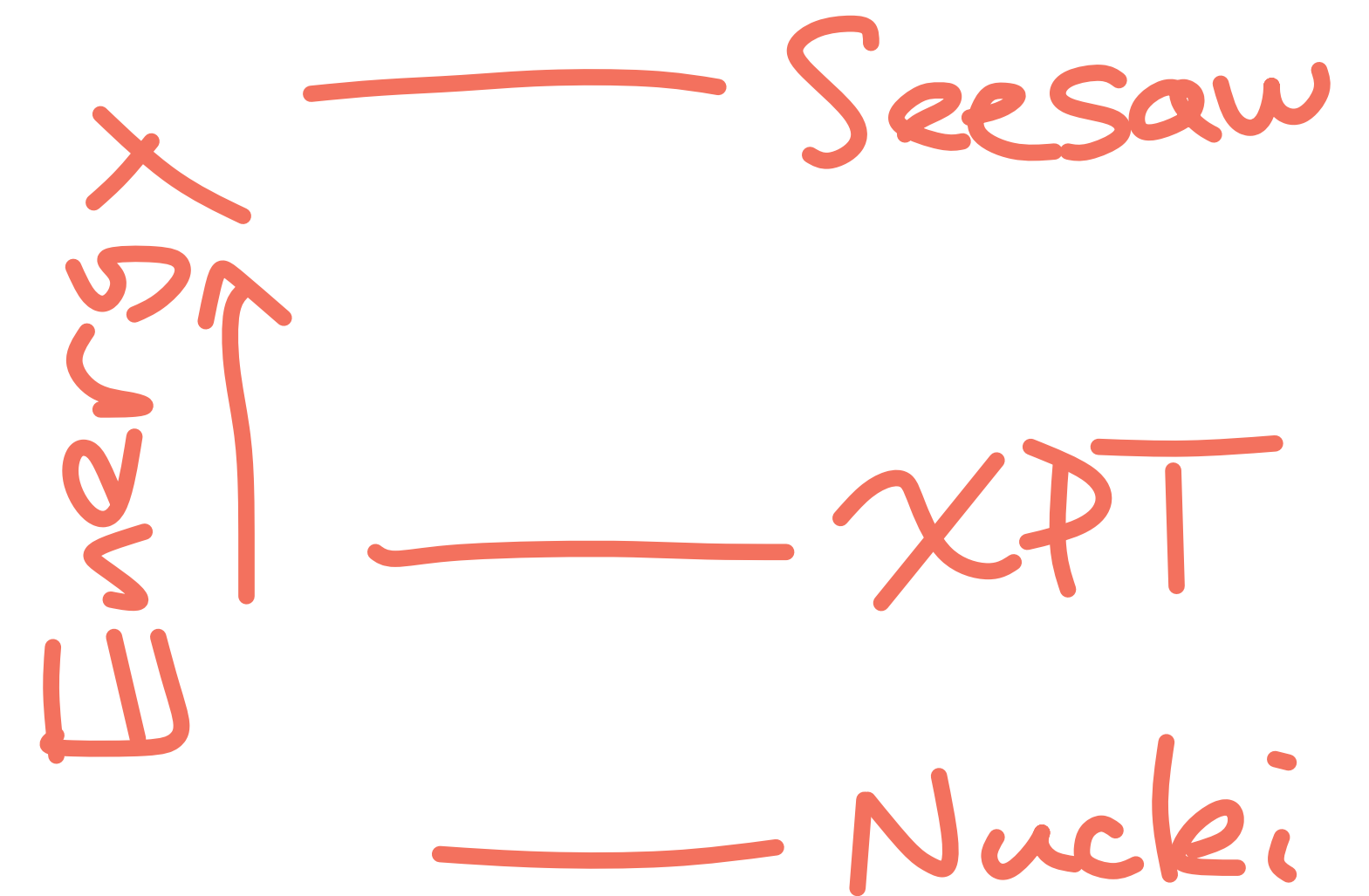
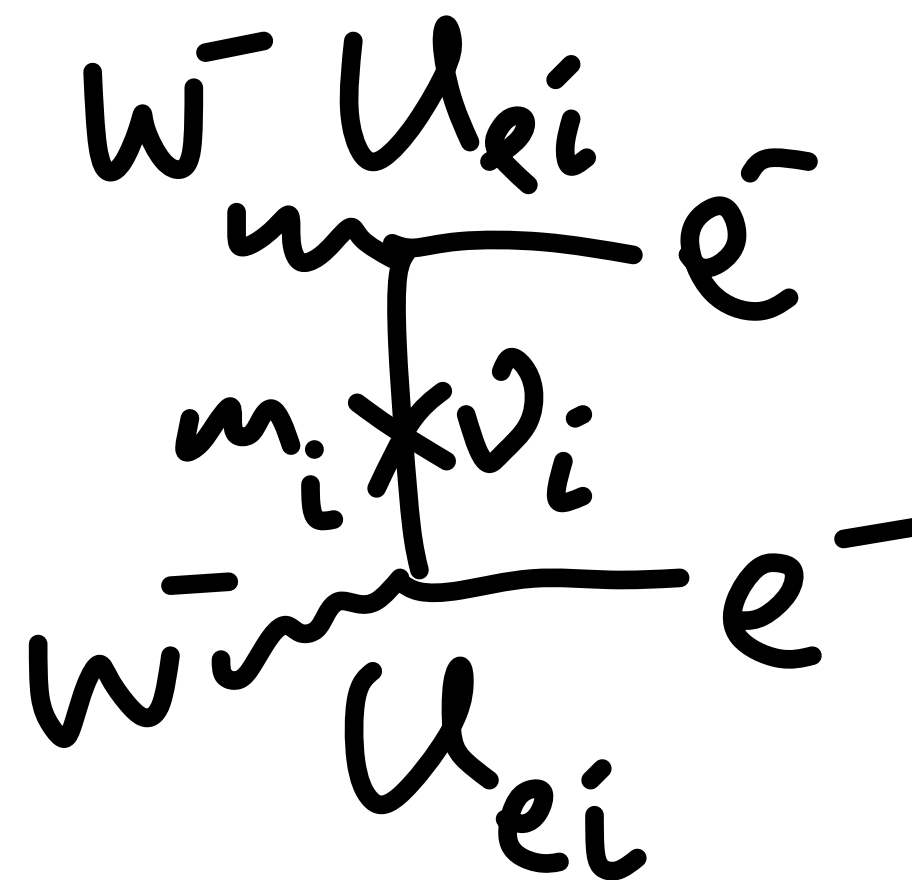
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Neutrino mass comes from coupling to heavy SM singlets (can explain smallness of  $\nu$  masses, another reason to expect LNV)

Short-distance LNV physics

captured by  $m_{\beta\beta} = \left| \sum_i U_{ei}^2 m_i \right|$

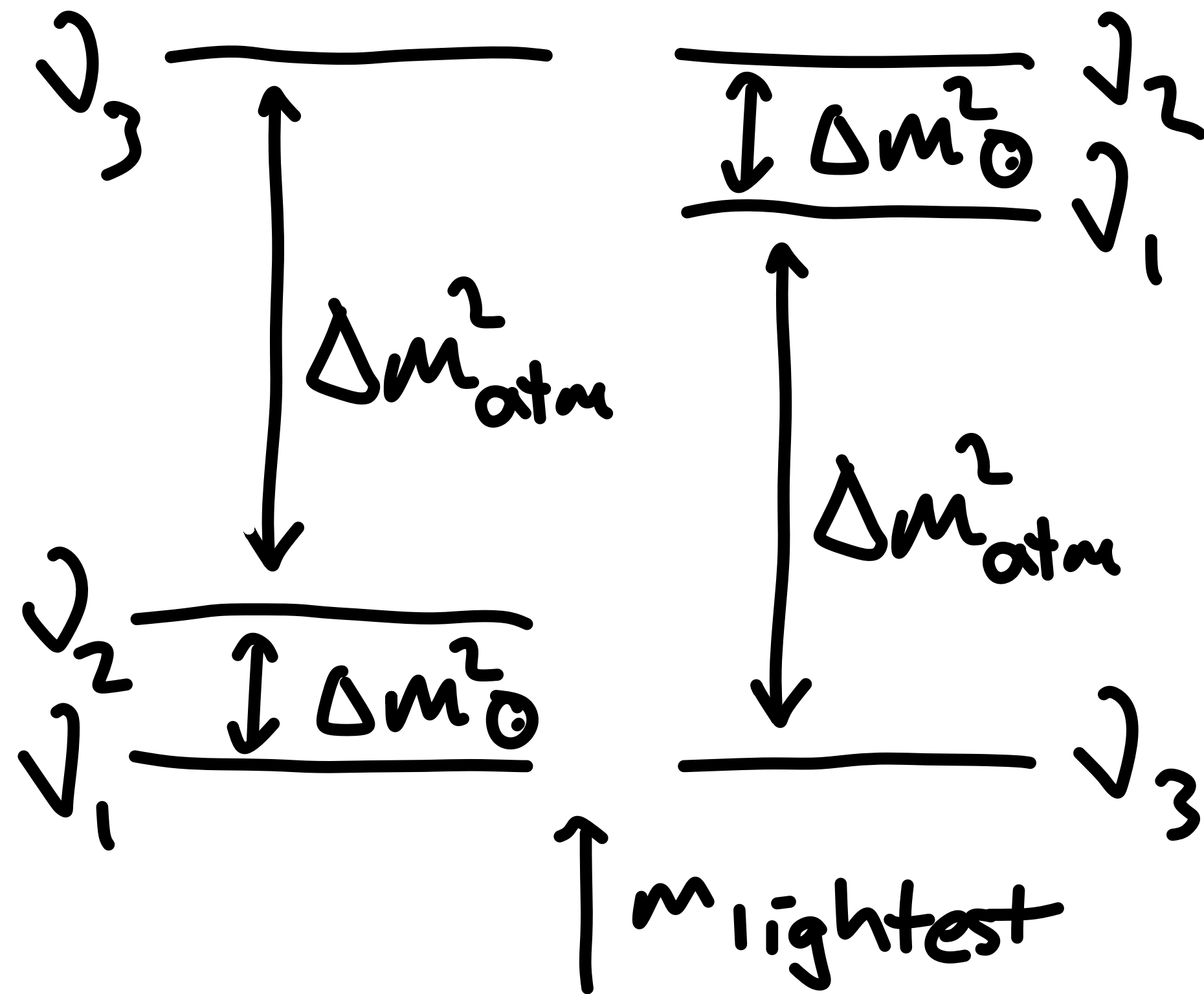
and matched onto successively longer scales



# High-scale seesaw

At low energies only remnants are Majorana  $\nu$  masses

$m_{\beta\beta}$  can be related to lightest  $\nu$  mass, depends on mass hierarchy,  $U_{\text{PMNS}}$



[NuFit 5.2]

$$|U|_{3\sigma}^{\text{with SK-atm}} = \begin{pmatrix} 0.803 \rightarrow 0.845 & 0.514 \rightarrow 0.578 & 0.143 \rightarrow 0.155 \\ 0.244 \rightarrow 0.498 & 0.502 \rightarrow 0.693 & 0.632 \rightarrow 0.768 \\ 0.272 \rightarrow 0.517 & 0.473 \rightarrow 0.672 & 0.623 \rightarrow 0.761 \end{pmatrix}$$

3 angles (known)  
1+2 phases (unknown)



# High-scale seesaw

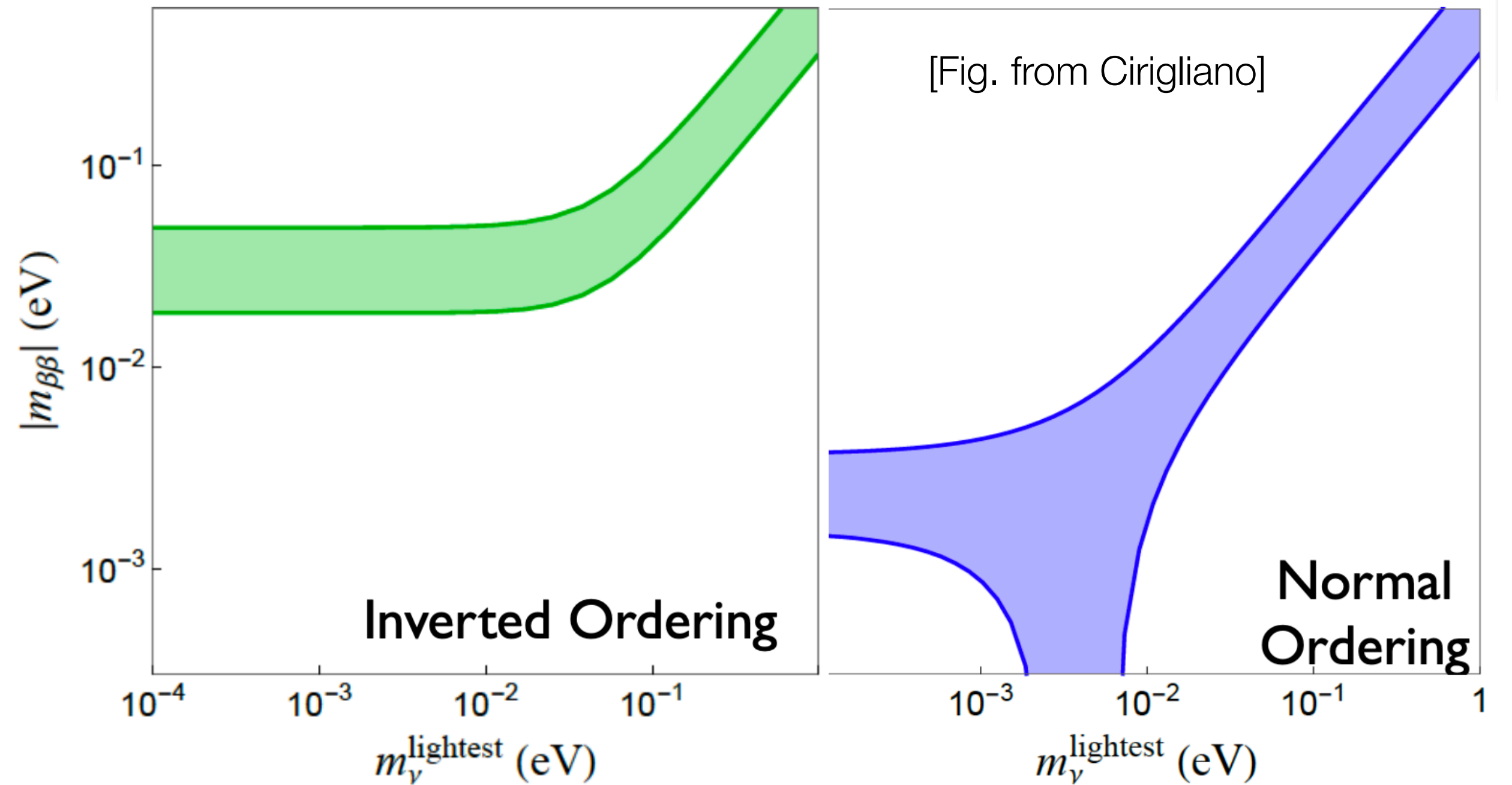
At low energies only remnants are Majorana  $\nu$  masses

$m_{\beta\beta}$  can be related to lightest  $\nu$  mass, depends on mass hierarchy,  $U_{\text{PMNS}}$

Half-life is related to  $m_{\beta\beta}$ :

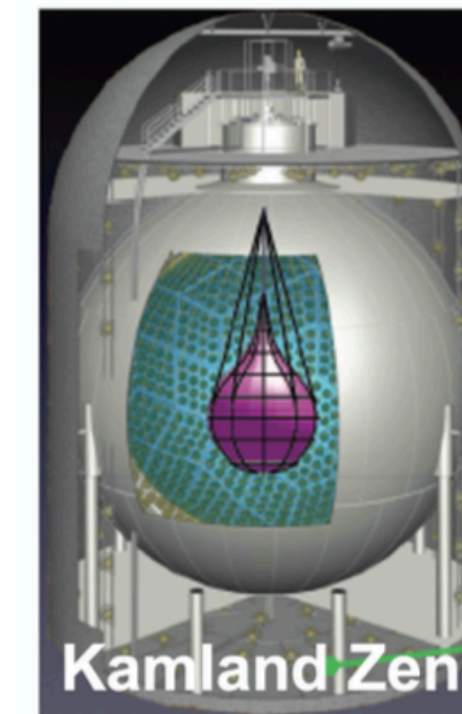
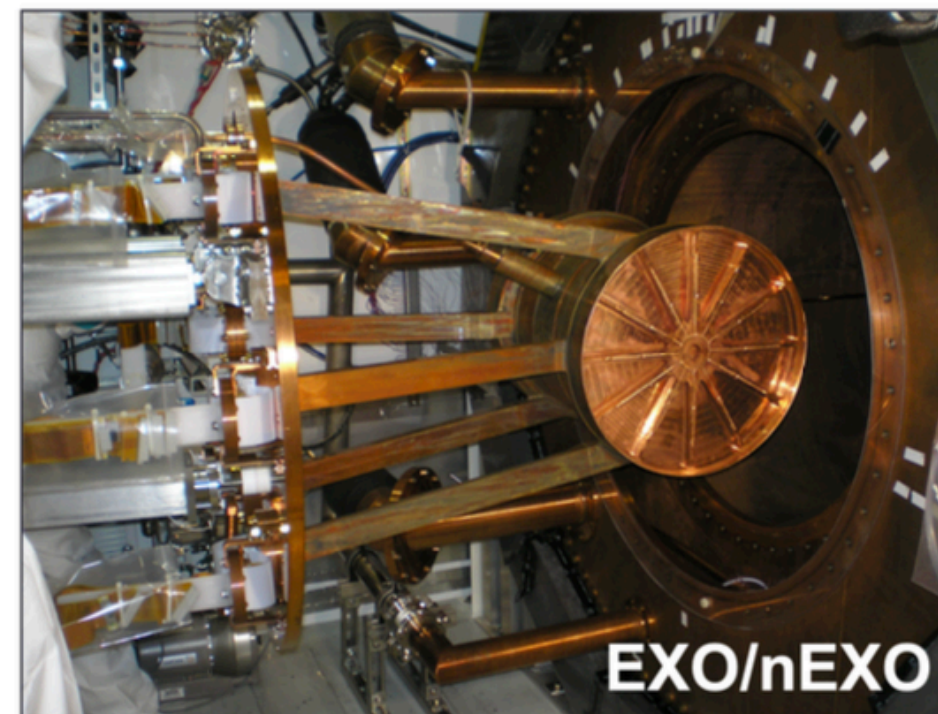
$$\frac{1}{T_{1/2}^{0\nu\beta\beta}} = m_{\beta\beta}^2 \times (\text{nuc. matrix element}) \times (\text{phase space})$$

increasing  $T_{1/2}^{0\nu\beta\beta}$

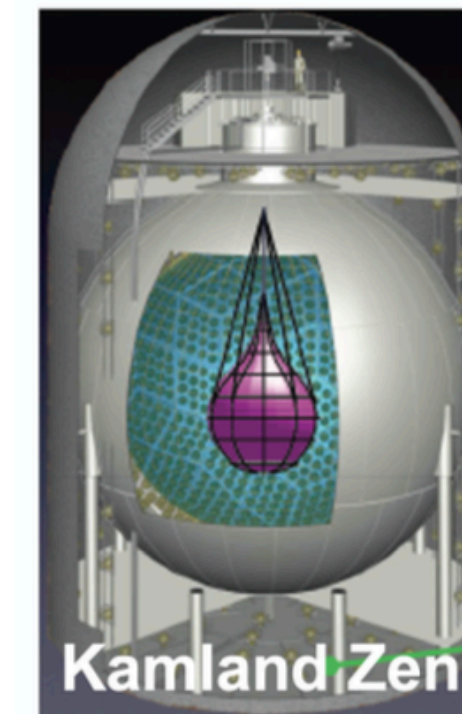
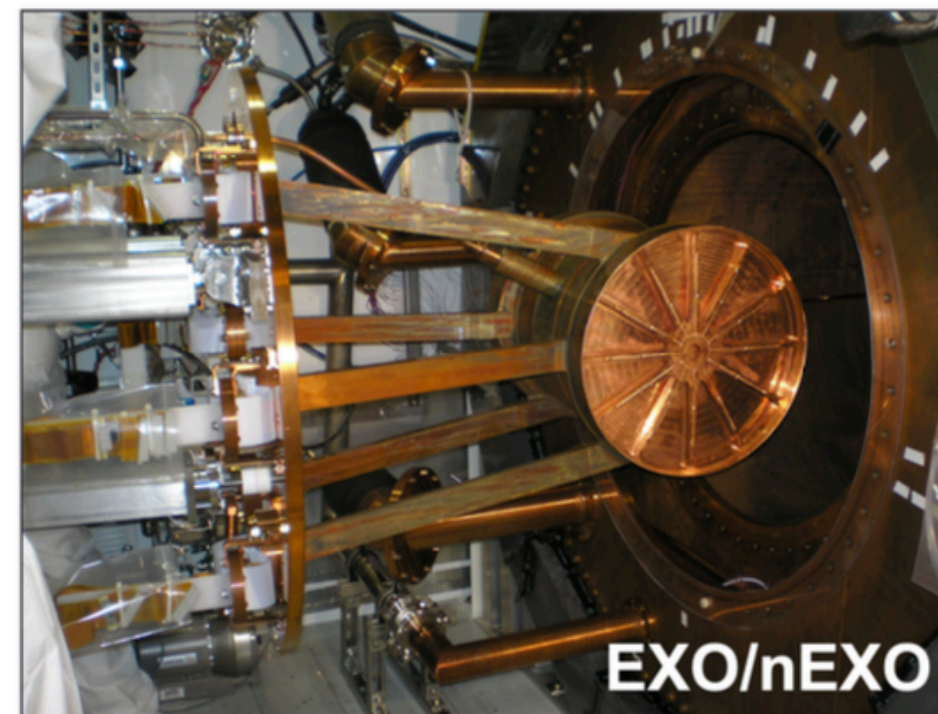
# $0\nu\beta\beta$ decay searches are on!

Worldwide effort to search for  $0\nu\beta\beta$   
with a number of isotopes is underway



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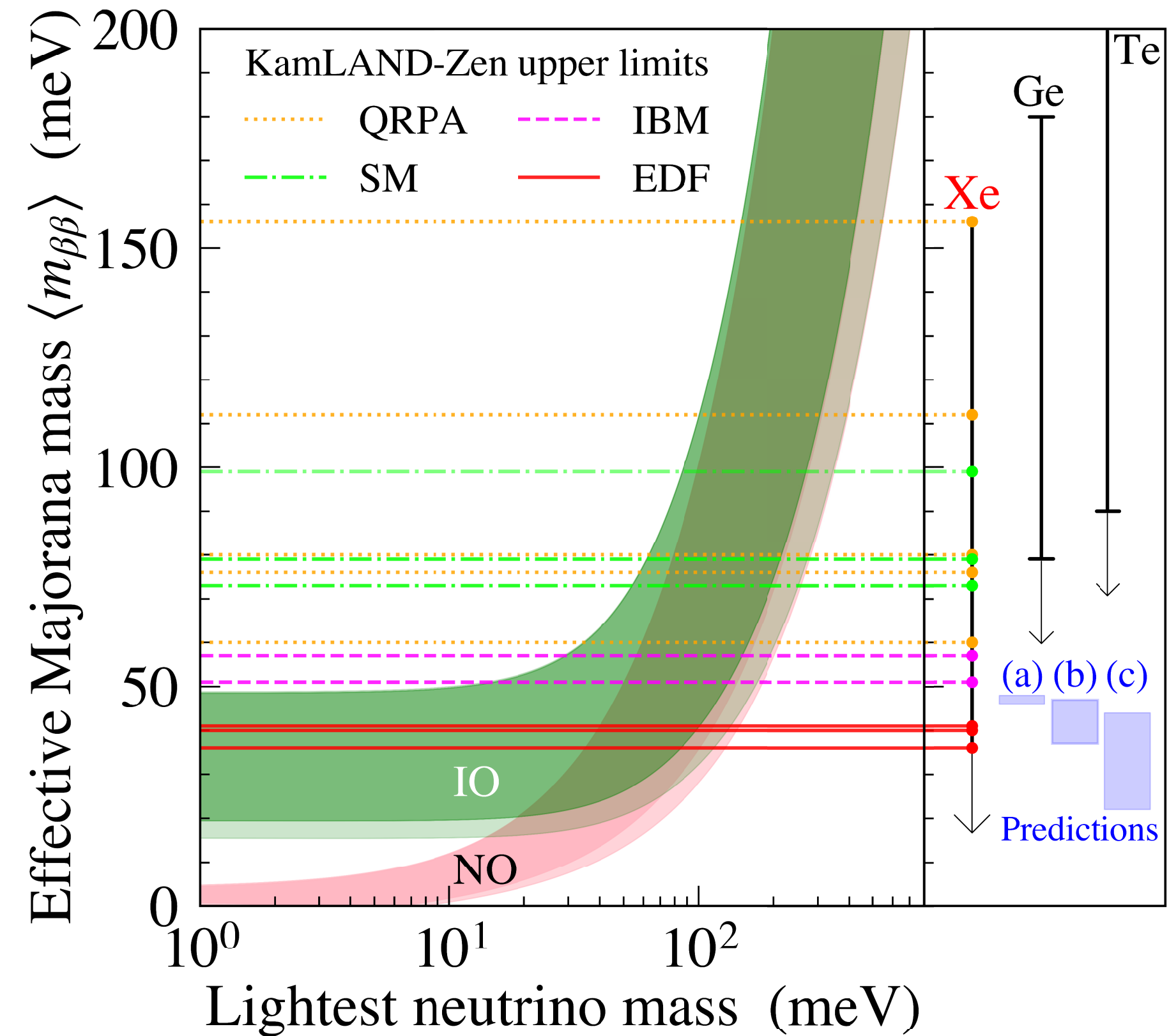


See excellent talks yesterday

# $0\nu\beta\beta$ decay searches are on!

arXiv: 2203.02139

KamLAND-Zen  
 $T_{1/2}^{0\nu\beta\beta}(^{136}\text{Xe}) > 2.6 \times 10^{26} \text{ yr}$

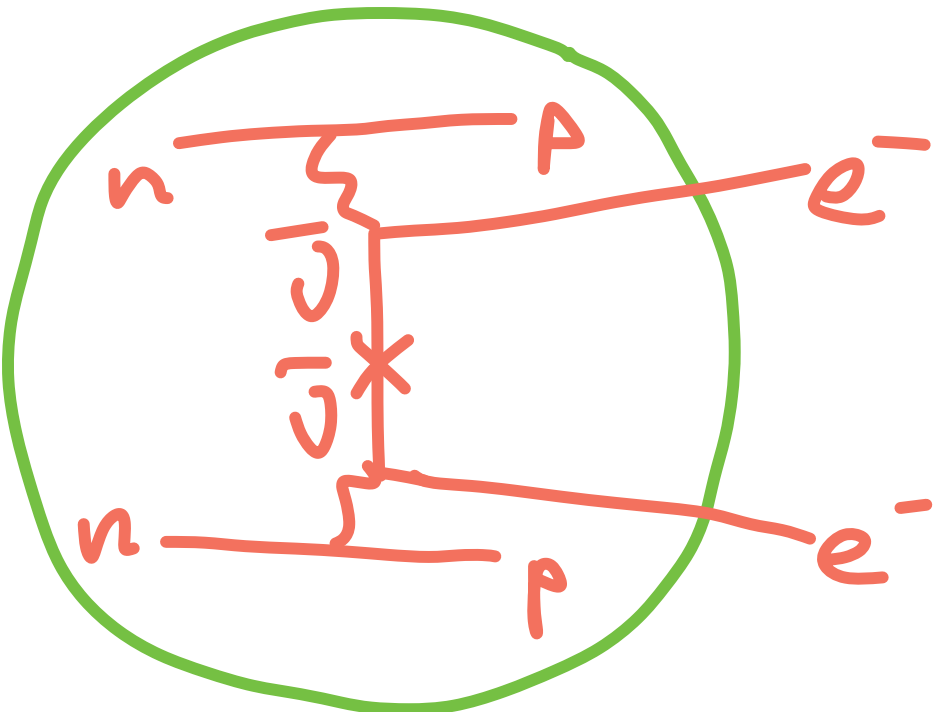


# Nuclear matrix elements

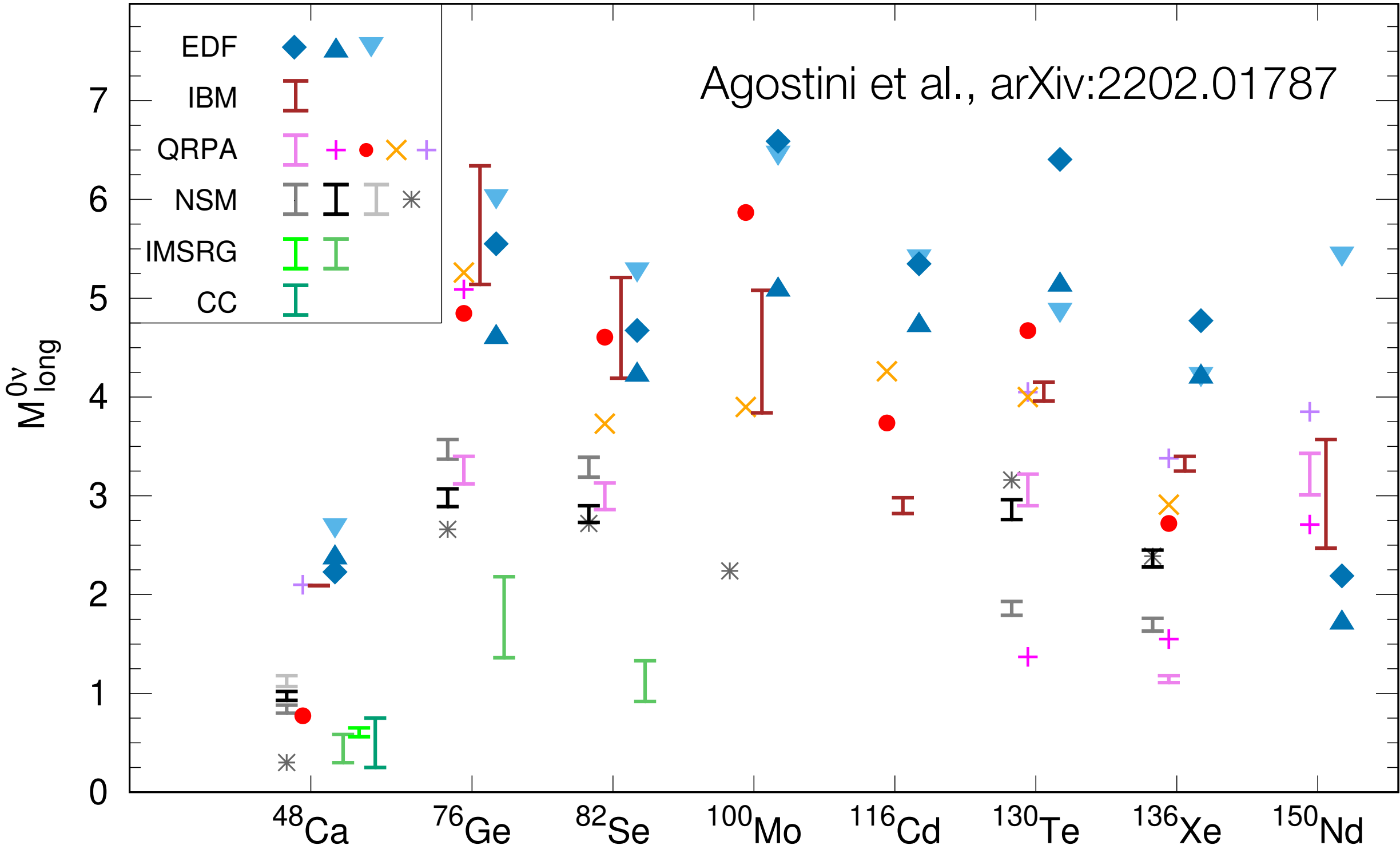
Half-life requires nuclear physics:

$$\frac{1}{T_{1/2}^{0\nu\beta\beta}} = m_{\beta\beta}^2 \times (\text{nuc. matrix element}) \times (\text{phase space})$$

Crucial for prediction



nuclear matrix elements  
can vary between  
different approaches

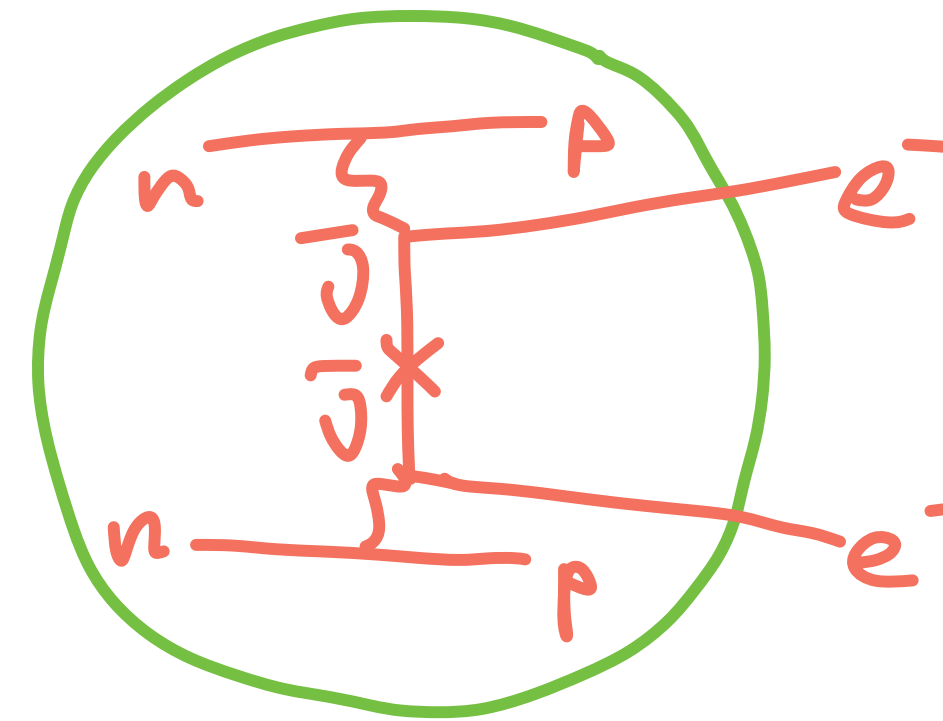


# Nuclear matrix elements

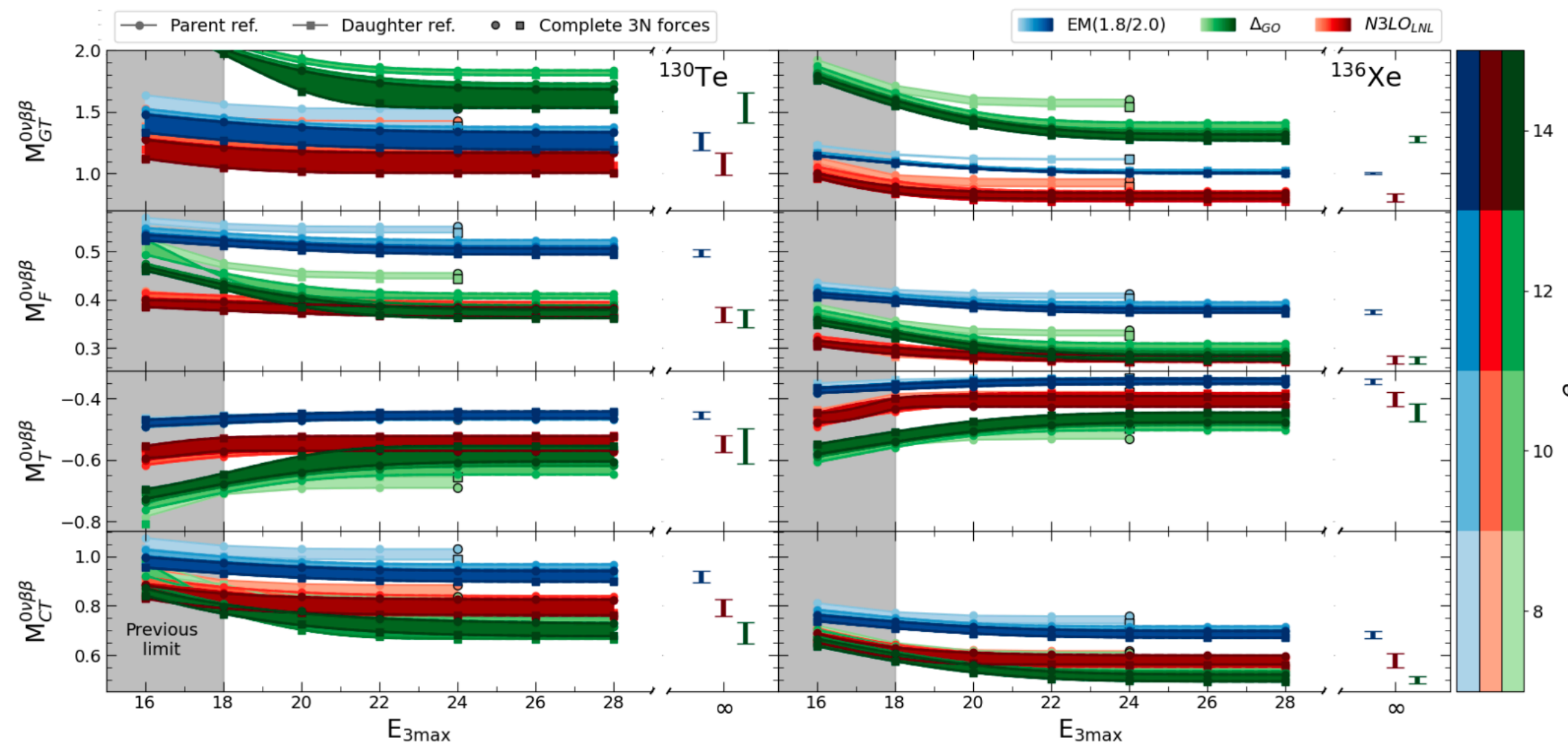
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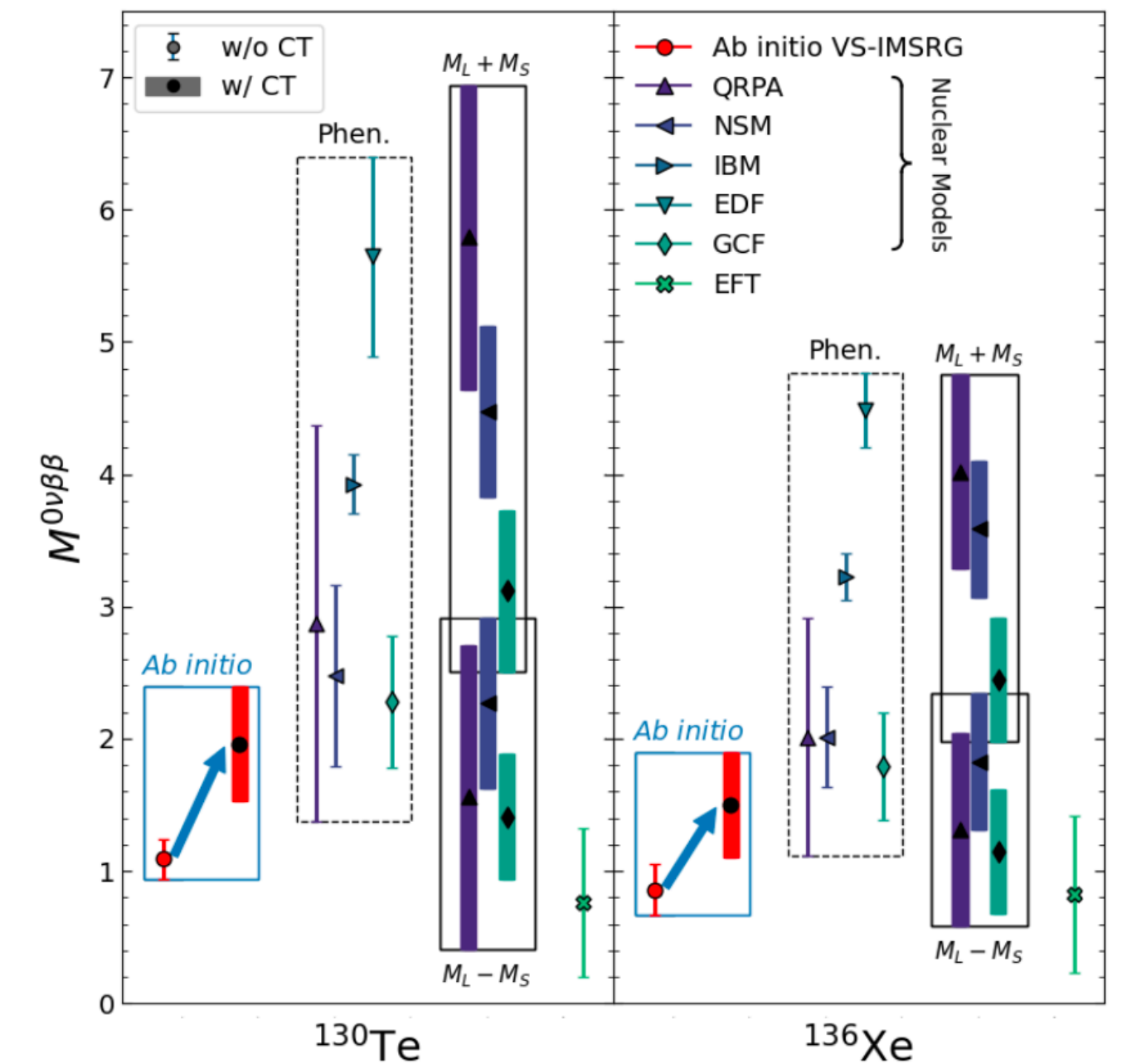
Crucial for prediction



Lots of recent work on *ab initio* predictions of experimentally relevant matrix elements



Belley, Miyagi, Stroberg, Holt, in prep.

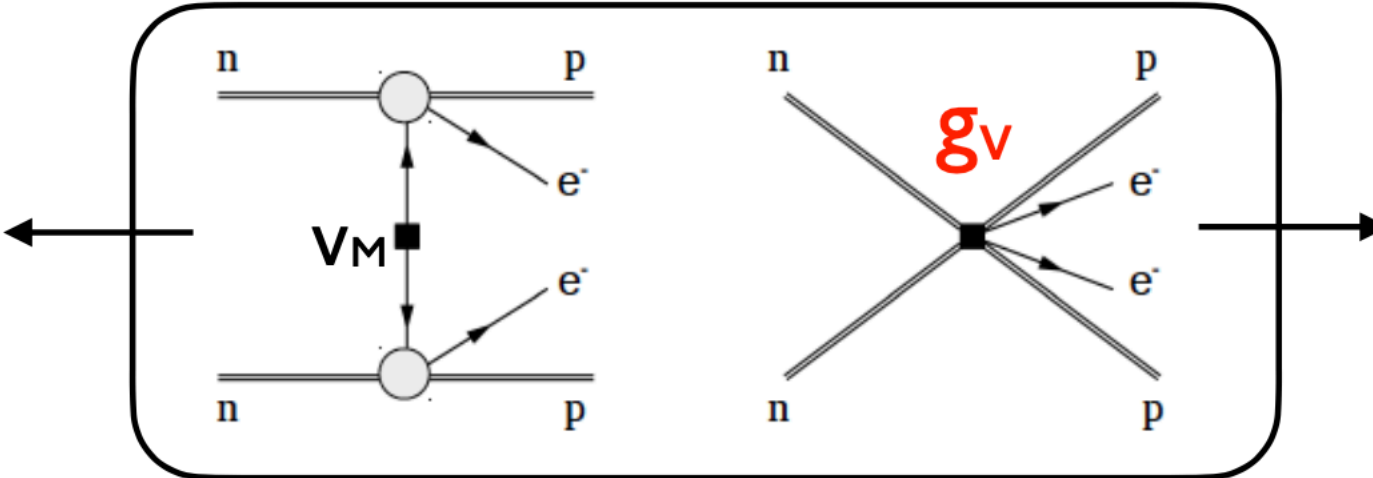


# Nuclear matrix elements

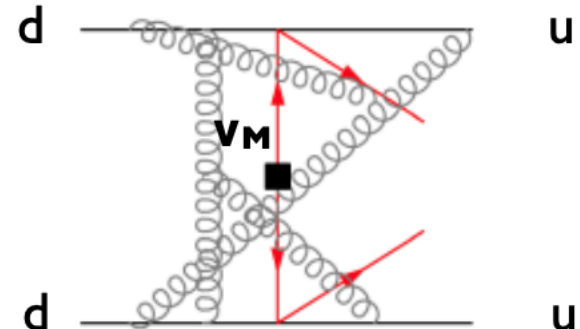
Cirigliano et al., 1802.10097

Additional complication in  $\chi$ PT – short distance contribution that is difficult to relate to data

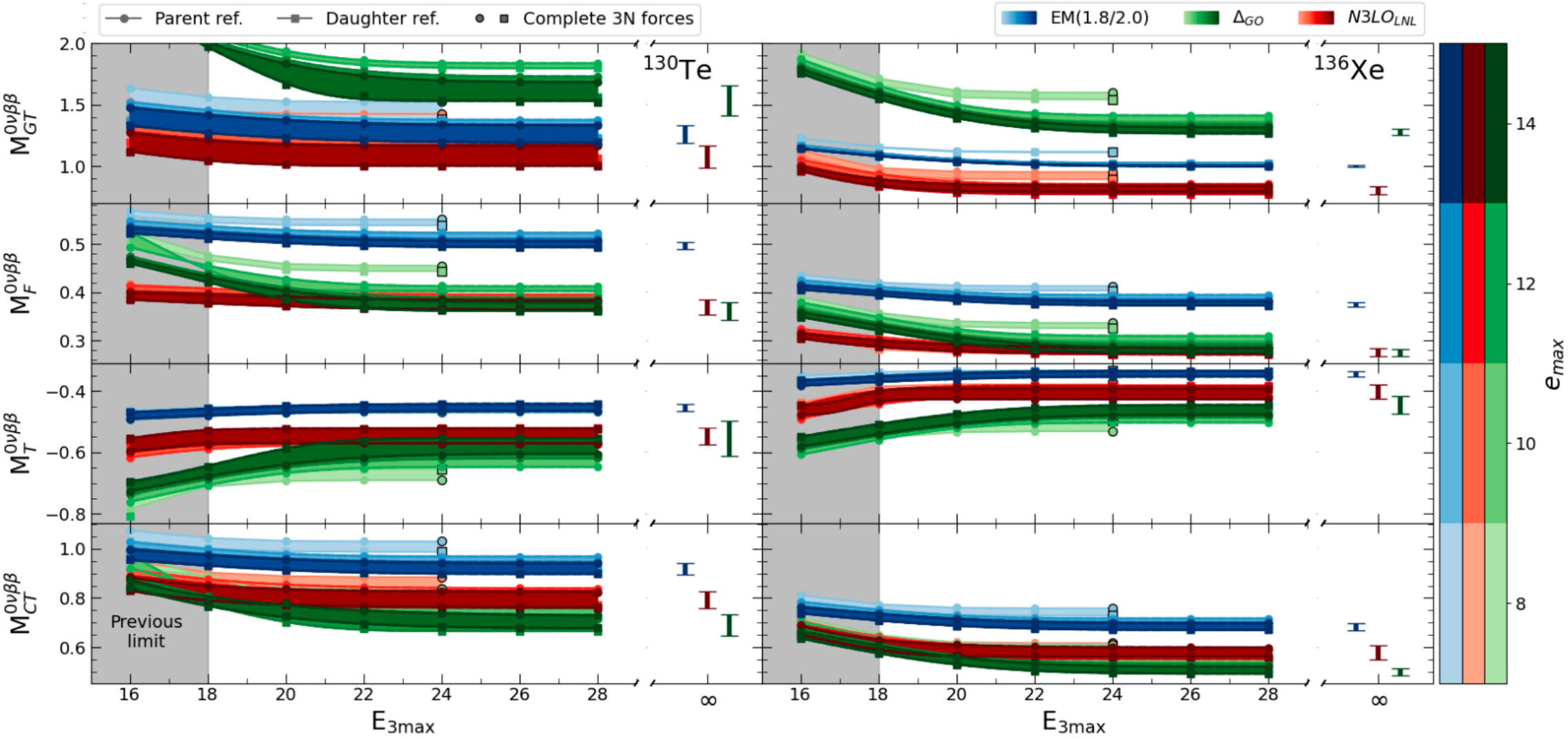
'Usual'  $V_M$  exchange  
 $\sim 1/k_F^2 \sim 1/Q^2$   
 Coulomb-like potential



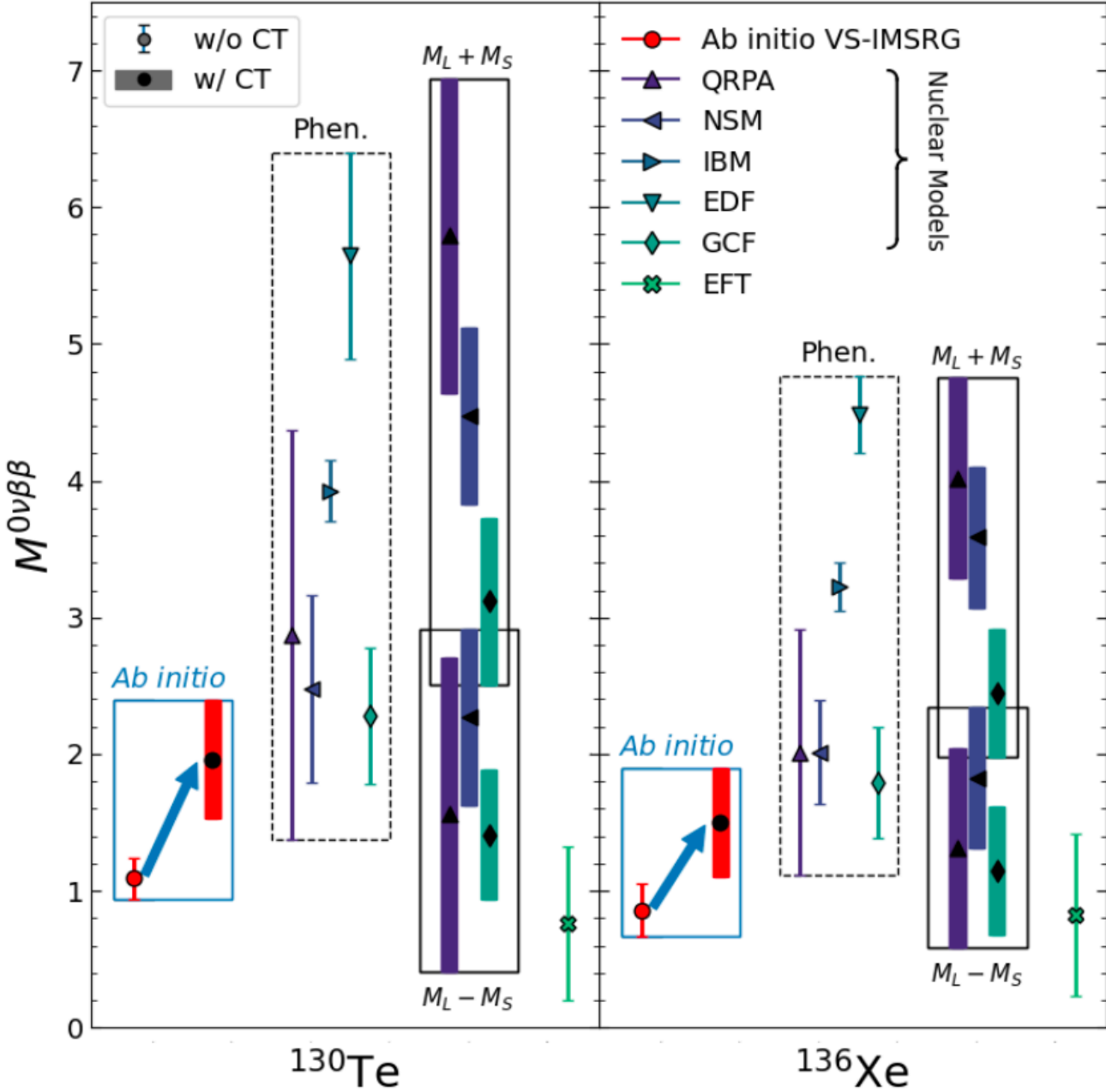
'New': short-range coupling  $g_v \sim 1/Q^2$



Contact term is an important ingredient in predictions



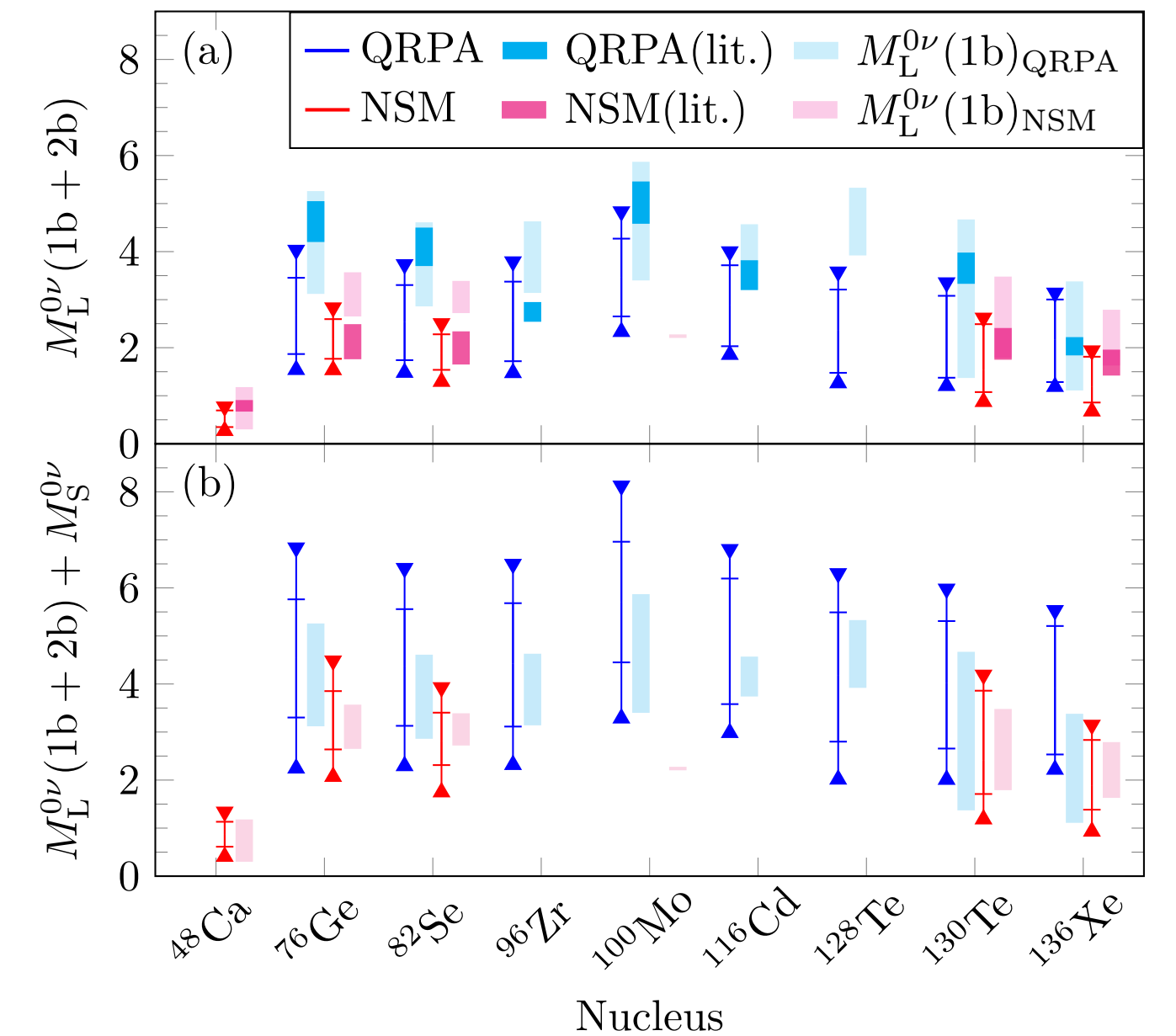
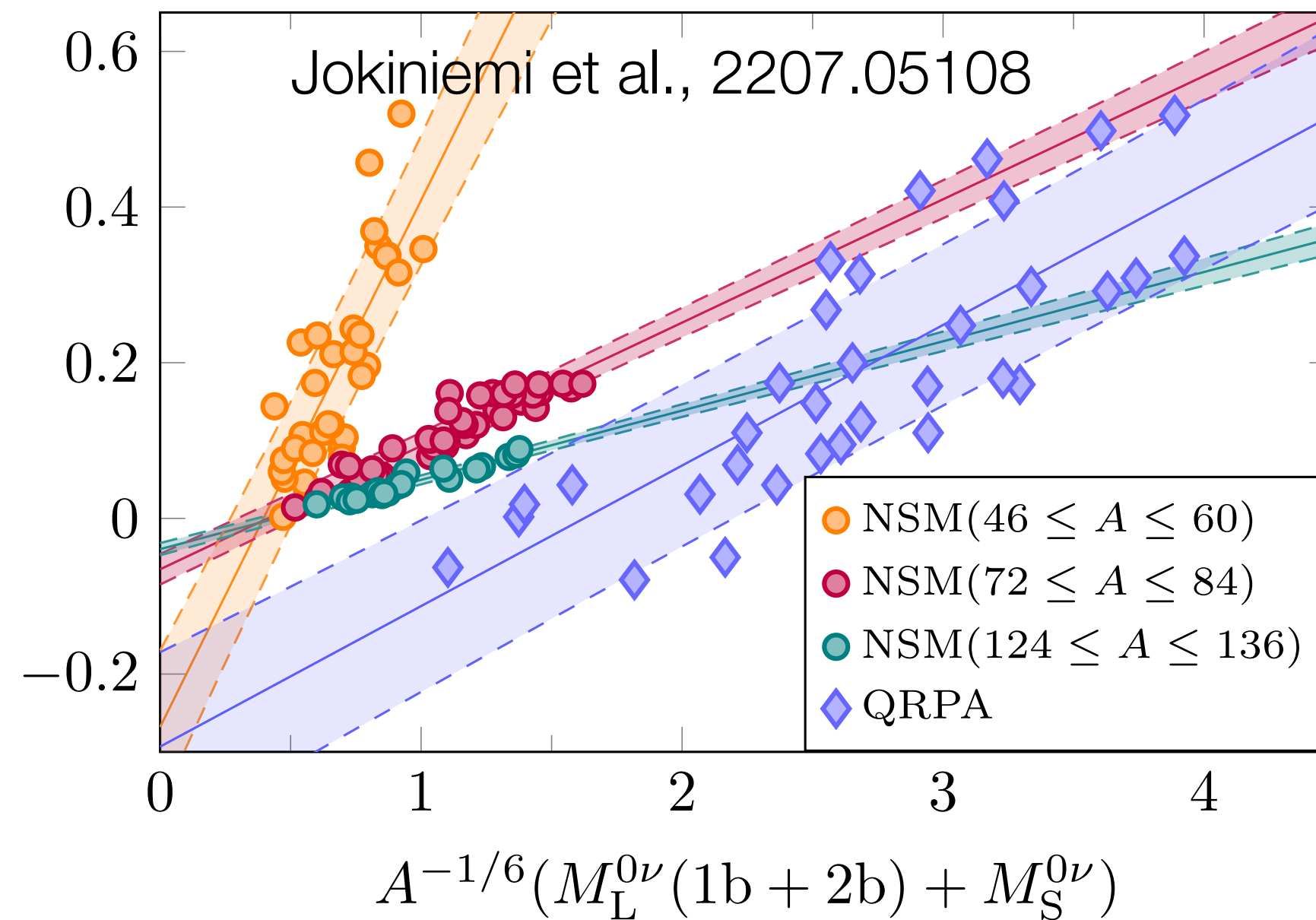
Belley, Miyagi, Stroberg, Holt, in prep.



# Nuclear matrix elements

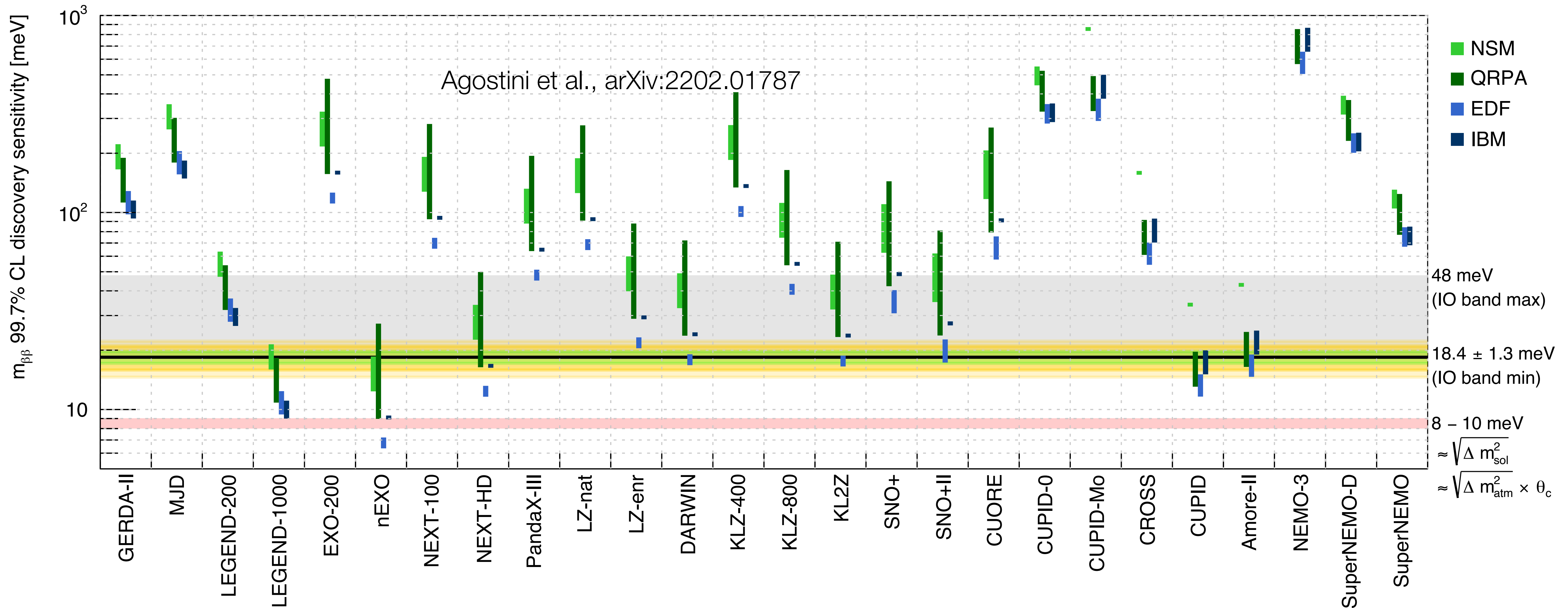
Other approaches to characterizing and understanding nuclear matrix elements

correlation between  $0\nu\beta\beta$  &  $2\nu\beta\beta$



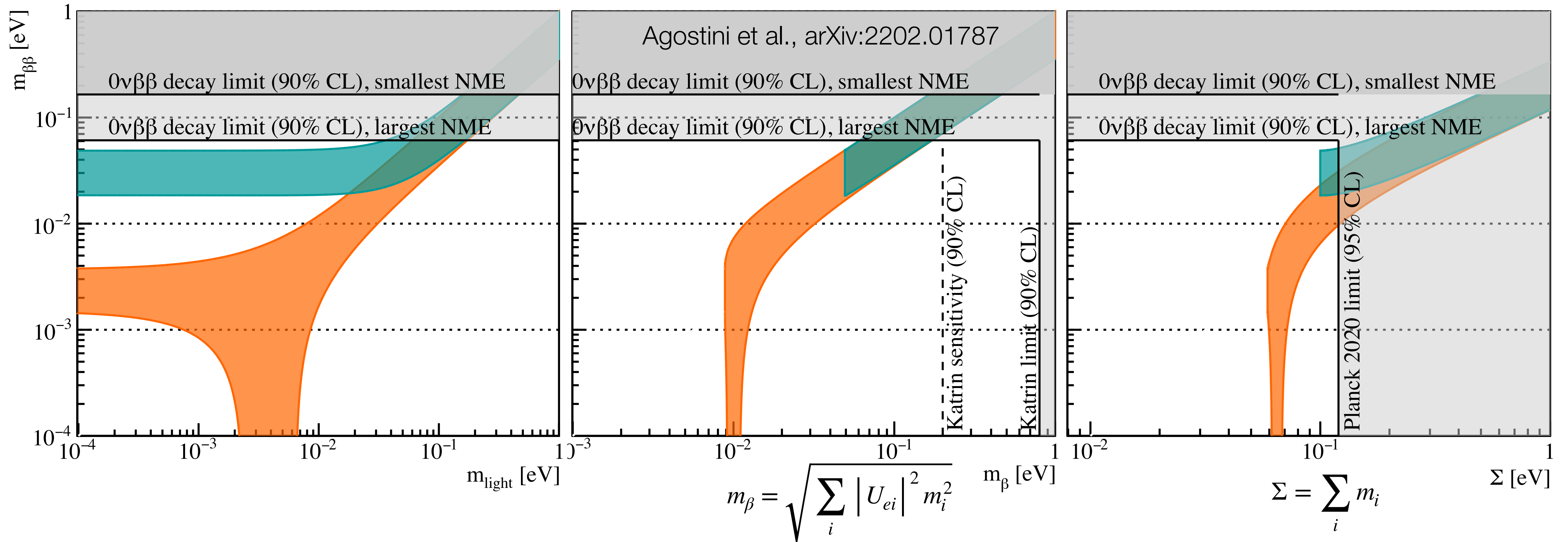


# Experimental outlook



Tonne-scale experiments will probe much of the inverted hierarchy region of  $m_{\beta\beta}$

# Relation to other probes of neutrino mass



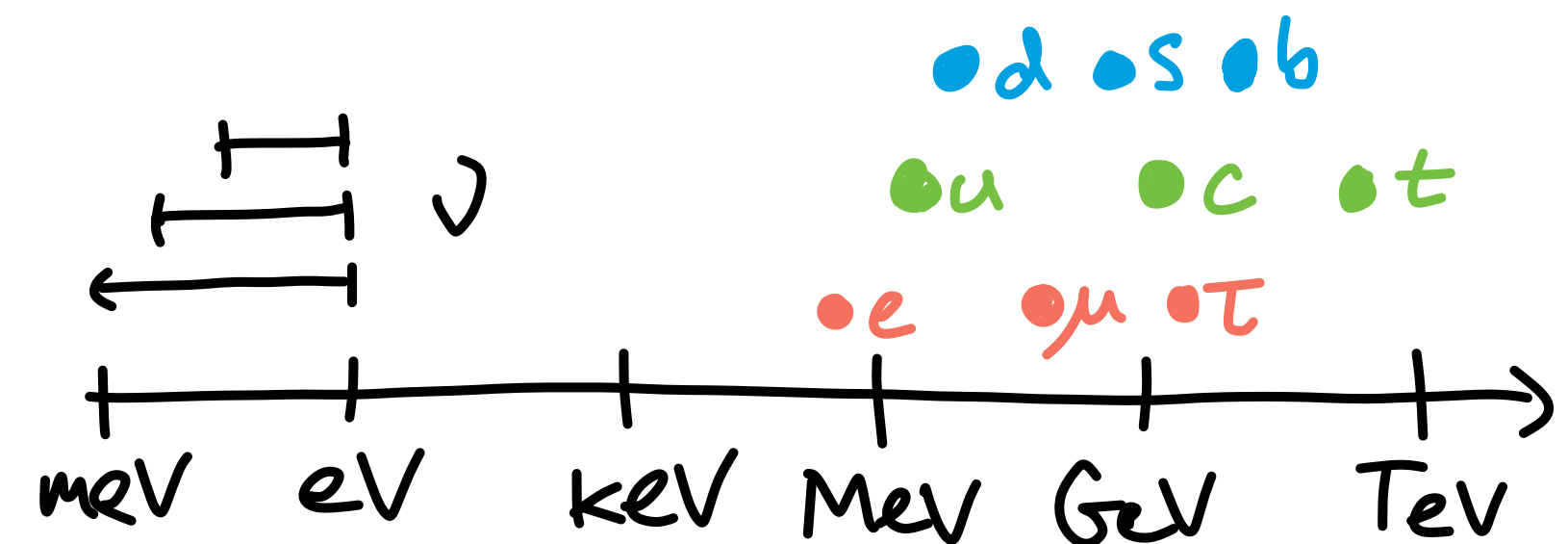
$0\nu\beta\beta$  is complementary to  $\beta$ -decay endpoint and cosmology but is **uniquely** sensitive to  $\Delta L \neq 0$ !

# Other new physics: light $\nu_R$

Large  $\nu_R$  Majorana mass could explain light neutrino masses

$$-\mathcal{L} \supset M\nu_R\nu_R + \text{h.c.} \Rightarrow m \sim \frac{y^2 \langle H \rangle^2}{M}$$

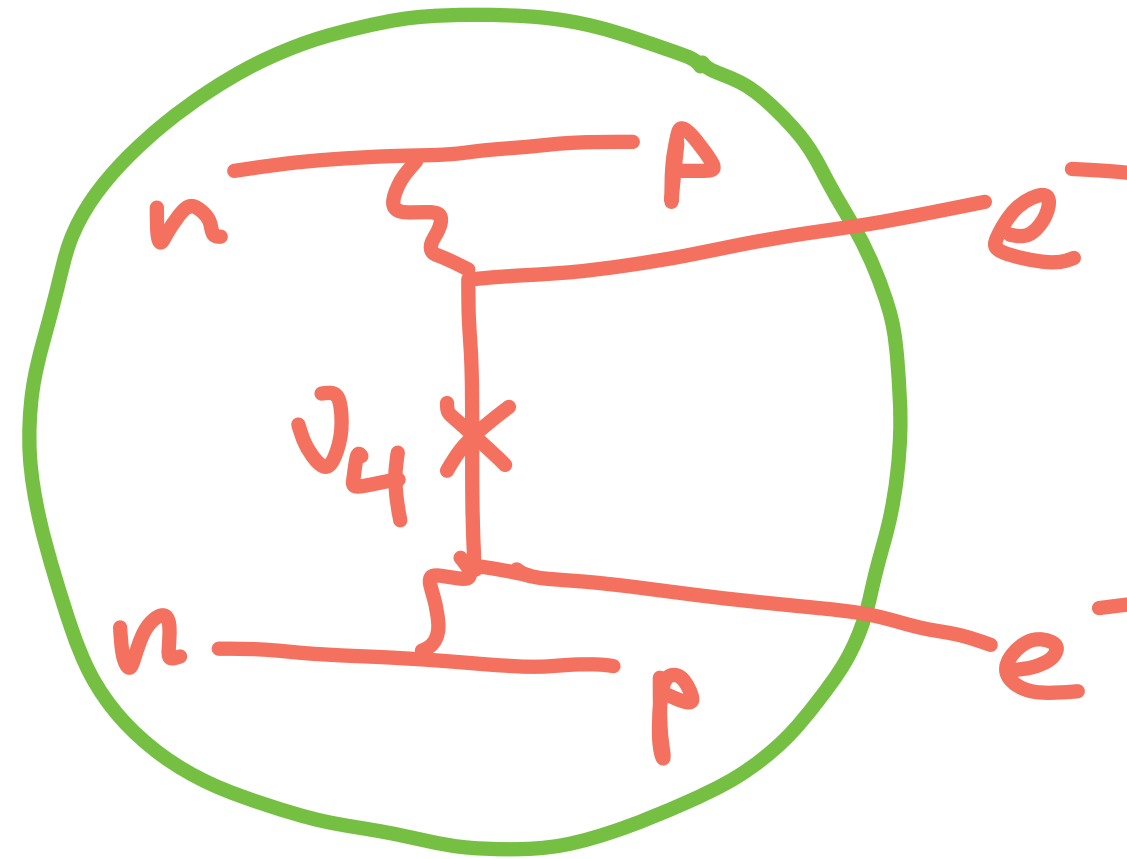
But Yukawa couplings could be small (we see  $\sim 10^5$  variation in the case of charged fermions)



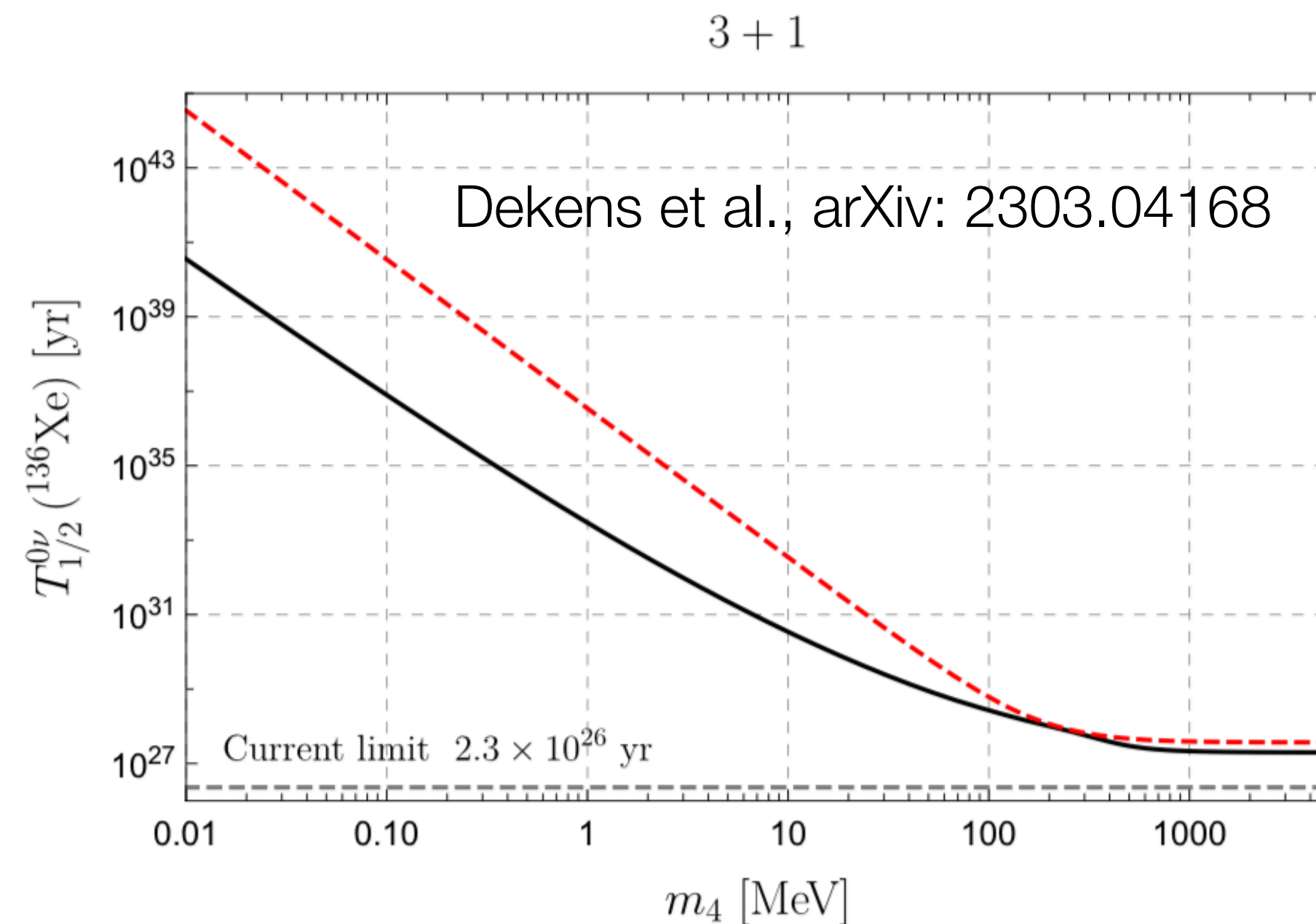
$\Rightarrow$  (mostly) sterile neutrinos may not be so heavy, with  $M \sim k_F$   
Can affect nuclear matrix elements

# Other new physics: light $\nu_R$

Heuristically  $\mathcal{M} \sim \frac{1}{p^2 - m_4^2}$



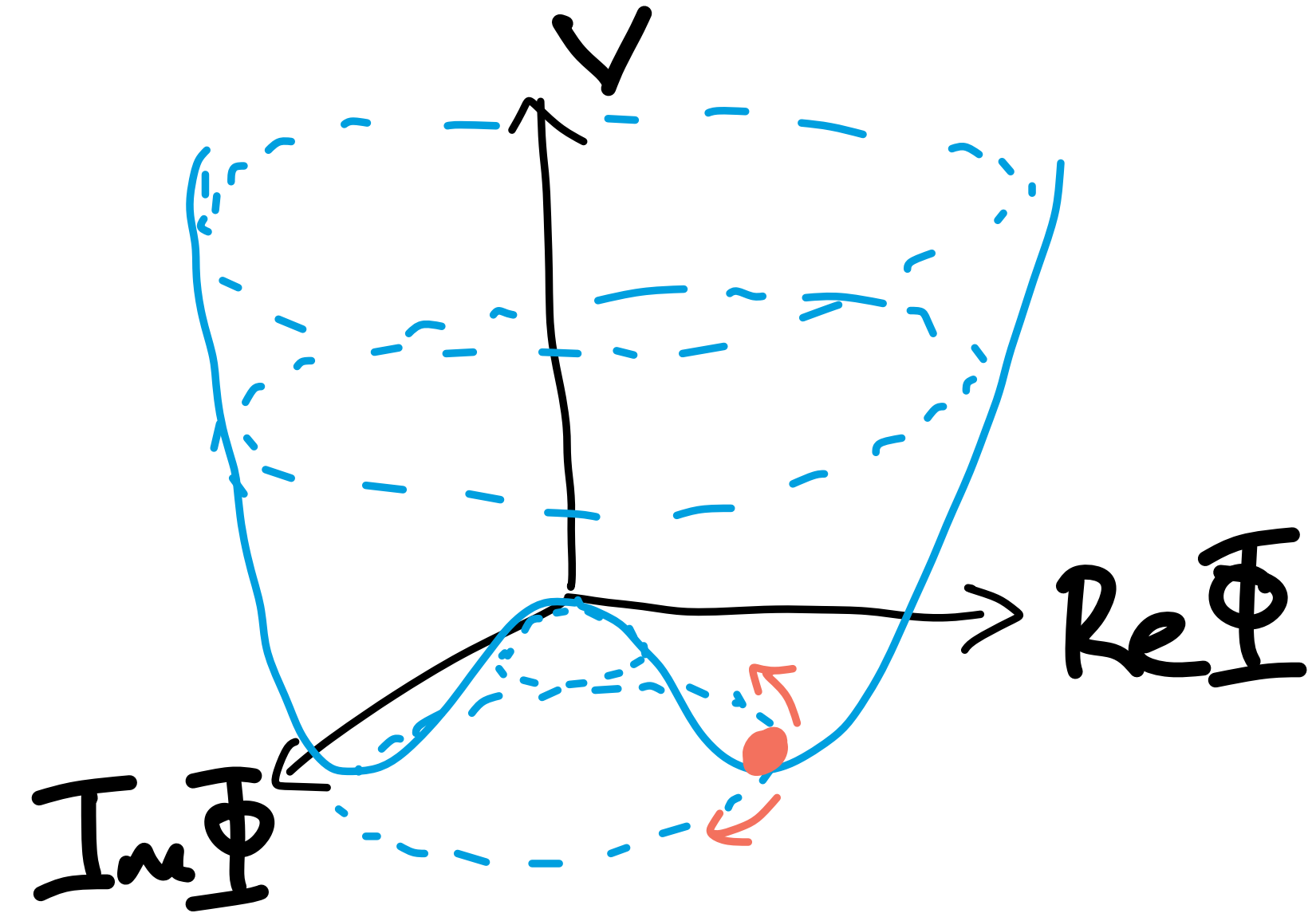
Detailed study shows rates can be impacted



# Other new physics: Majorons

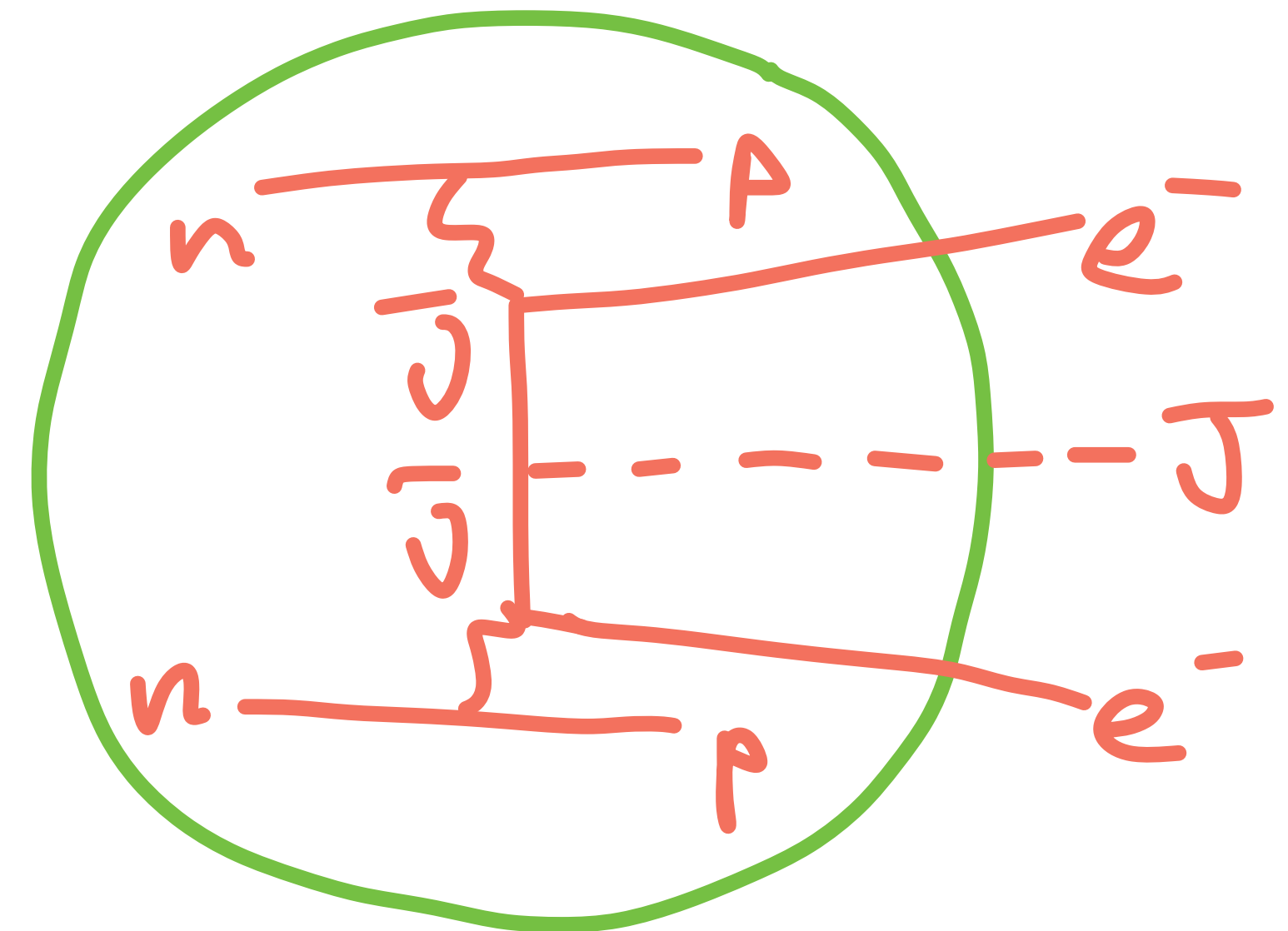
$\nu_R$  Majorana mass could be associated with vev of a scalar that spontaneously breaks  $U(1)_L$

$$-\mathcal{L} \supset g\Phi\nu_R\nu_R + \text{h.c.} = Me^{iJ/f}\nu_R\nu_R + \text{h.c.}$$



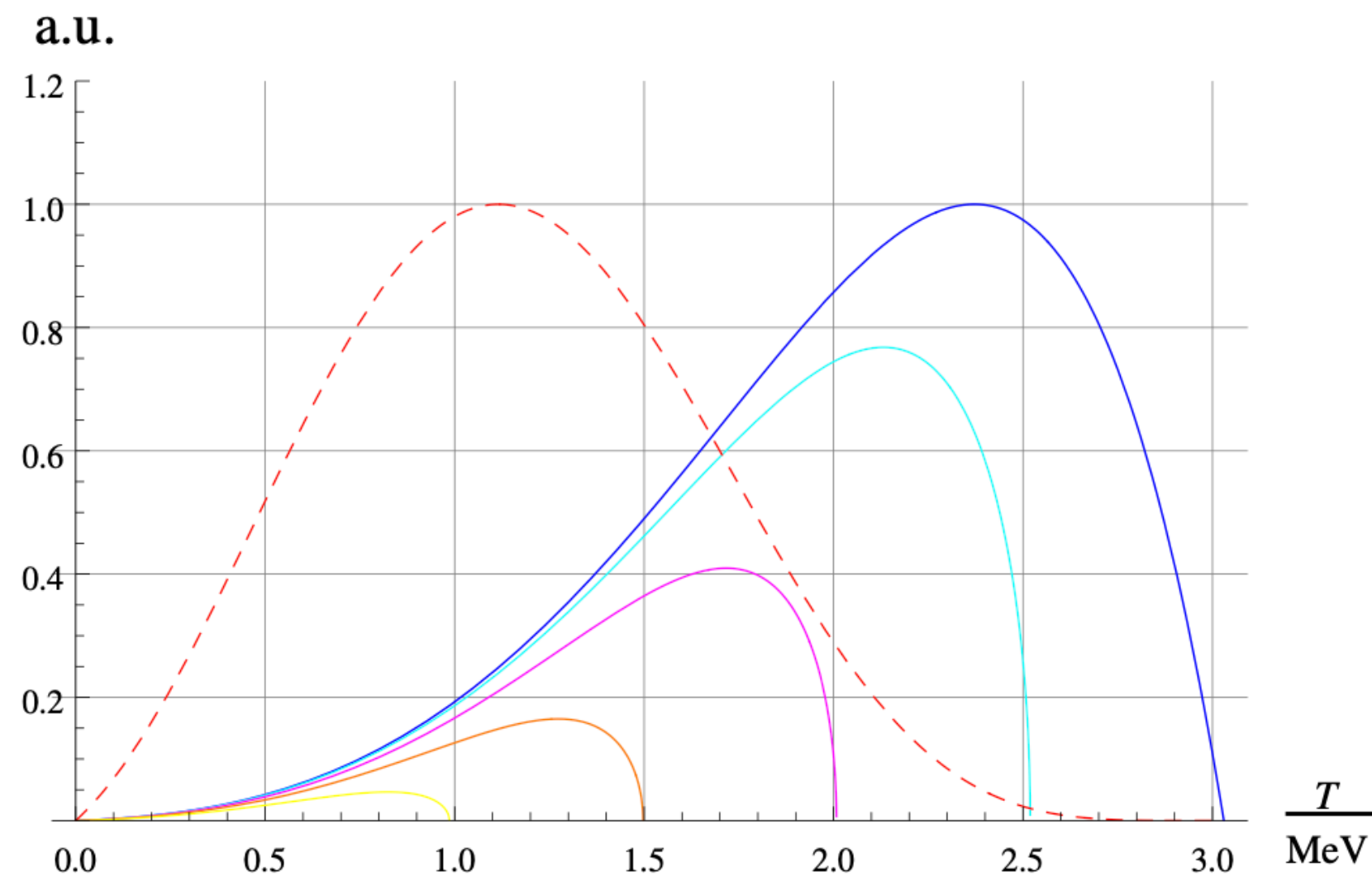
Light pseudo-Nambu-Goldstone boson could then be emitted in  $\beta\beta$  decay

(Georgi, Glashow, Nussinov)

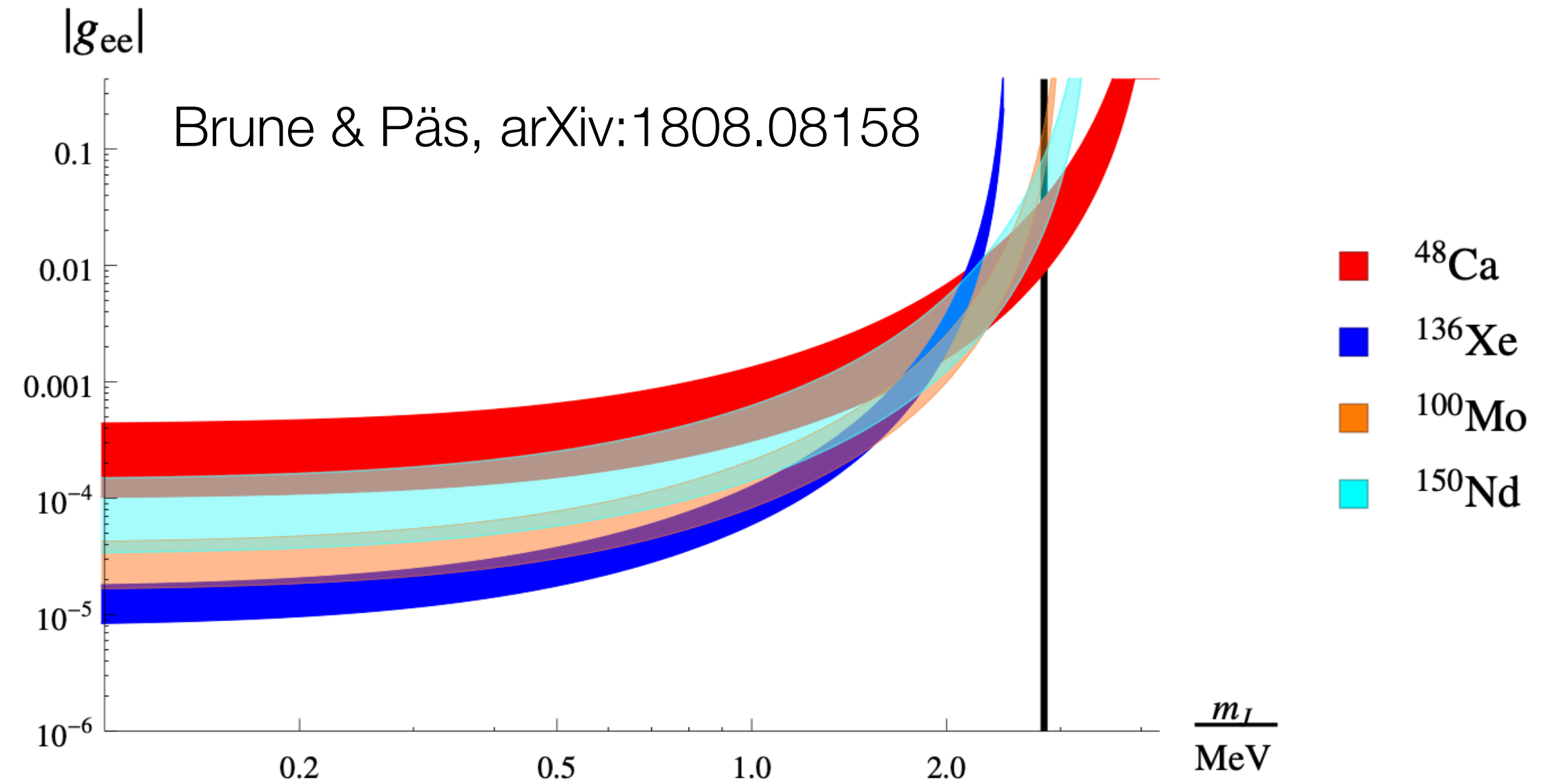


# Other new physics: Majorons

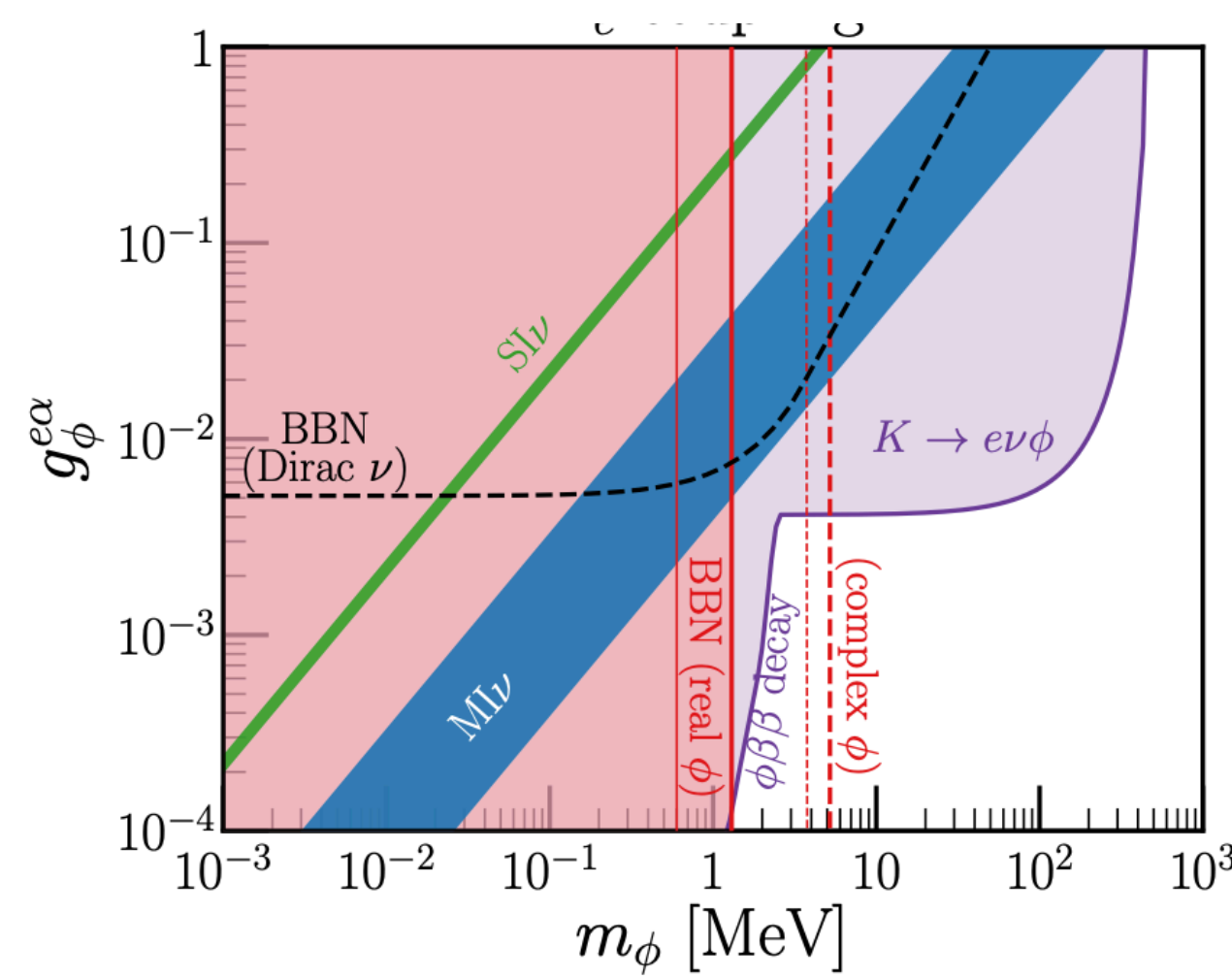
Can be constrained with existing data/searches



- $0\nu\beta\beta J, m_J = 0$
- $0\nu\beta\beta J, m_J = m_e$
- $0\nu\beta\beta J, m_J = 2m_e$
- $0\nu\beta\beta J, m_J = 3m_e$
- $0\nu\beta\beta J, m_J = 4m_e$
- - -  $2\nu\beta\beta$



Interesting because of impact on cosmology tensions!



Blinov et al. arXiv:1905.02727

# Wrap up

Neutrino masses are only terrestrial evidence of physics beyond SM

Adding neutrino masses to SM is qualitatively different—connected to existence of a global symmetry

Ongoing effort to study the nature of neutrino masses with  $0\nu\beta\beta$  decay searches

This effort is profoundly important—impacts our understanding of  $\nu$  masses themselves, the matter asymmetry of the universe, cosmology, quantum gravity, ...

It's vitally important that we push these searches forward

# Backup: back of the t-shirt

