



Superconducting thermometers for the direct detection of light dark matter

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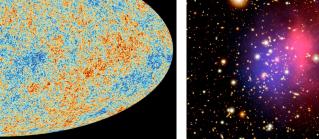
Quantum technologies lecture series, CERN, June 2023

Dark matter

cosmic microwave background (CMB)

bullet cluster (1E 0657-56)

galactic rotation curve (Messier M33) Observation from starlig 100 21 cm hydrogen Velocity (km s-1)



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© X-ray: NASA/CXC/CfA/M.Markevitch, Optical and lensing map: NASA/STScl. Magellan/U.Arizona/D.Clowe, Lensing map: ESO WFI © CC BY-SA 4.0, source: wikimedia

20.000

30.000

Distance (light years)

40.000

10.000

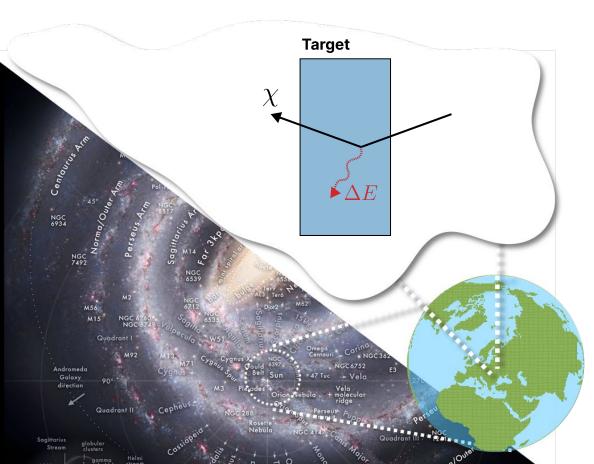
We observe overwhelming gravitational evidence for the existence of non-luminous, stable, cold, non-baryonic "dark" matter (DM) in our universe.

Best estimate from CMB: (83.9 +- 1.5)%* of all matter in our universe is DM.

Many particle candidates: (generalized) WIMPs, Axions, SIMPs, ...

The mass scale of a hypothetical DM particle is not fixed.

Direct detection



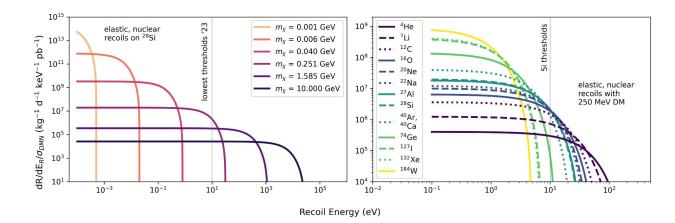
DM velocity distribution in Milky Way can be approximated with Maxwellian distribution.

Density along the solar circle is ~0.4 GeV/c²/cm^{3*}.

Direct detection experiments test DM models via DM-target scattering.

Likelihoods depend on assumed DM mass and interaction cross section σ with target.

Expected scattering rate



Massive DM particles scatter elastically and spin-independently off target nuclei.

Objectives of (ideal) direct DM search: low energy threshold, high target mass.

Light DM and heavy nuclei reduce expected recoil energy \Rightarrow experiments use

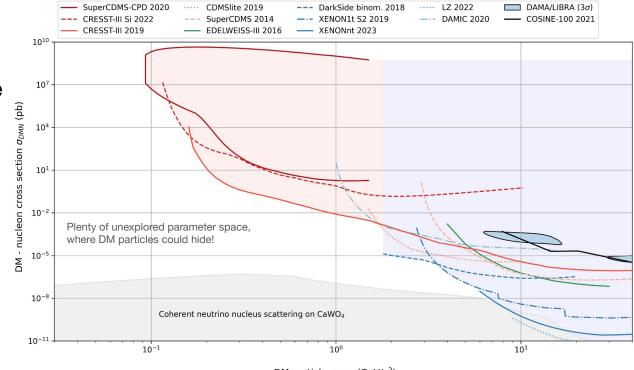
- high target mass and heavy nuclei for heavy DM searches,
- low threshold and light nuclei for light DM searches.

spectra calculated following Phys. Rev. D 31, 3059 (1985) astrophysical parameters from Rev. Mod. Phys. 85, 1561 (2013) Phys. Rev. D 92, 083517 (2015) A&A 523. A83 (2010) Monthly Notices of the Roval Astronomical Society 379, 755 (2007) Monthly Notices of the Roval Astronomical Society 221, 1023 (1986)

Exclusion limits on DM-nucleus elastic scattering

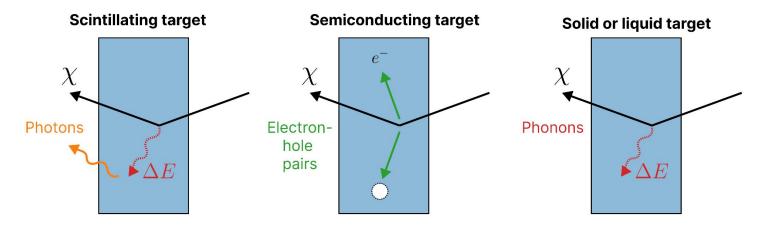
Cryogenic calorimeters with superconducting thermometers reach the best energy thresholds and sensitivity for sub-GeV/c² DM.

For higher DM masses time projection chambers (TPCs) with liquid noble gases are most sensitive.



DM particle mass (GeV/c²)

Detection channels



Signal channels to detect target scattering:

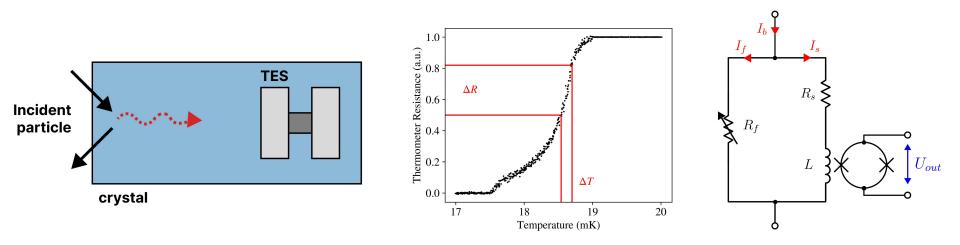
- scintillation (photons) measurable with scintillating crystals, liquids or gases,
- ionisation (electron-hole pairs) measurable with semiconductors, certain liquids, gases,
- lattice vibrations (phonons) measurable (mostly) with single crystals.

Often multiple signal channels used for event discrimination, e.g. scintillating crystals.

Measured recoil energy depends on scattering particle

Particle-independent energy scale!

Superconducting thermometers



Transition edge sensor (TES): superconducting film operated in the transition from its normal to superconducting state. Operation in low-noise SQUID-based readout circuit turns temperature/resistance fluctuations into measurable quantities.

\Rightarrow thermometer sensitive to \approx uK.

Film is thermally coupled to target crystal, collecting phonons from particle recoils in target.

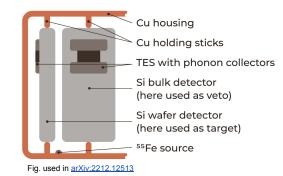
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\Rightarrow recoils down to 10 eV recoil energy detectable**.
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Technology developed in mid 90's* and continuously improved for DM searches by experiments (CRESST, CDMS, et al.).

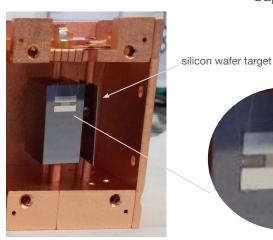
DM detectors with superconducting thermometers



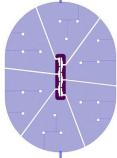
light detector



CRESST-III Si







SuperCDMS-CPD

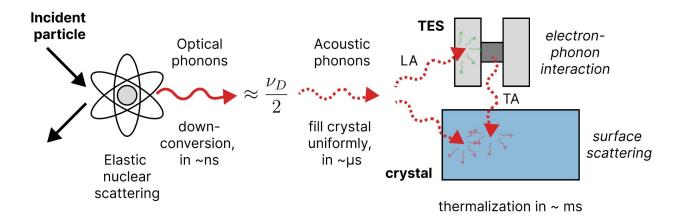
design called "guasiparticle trap assisted electrothermal feedback TES" (QET)

Fig. on the courtesy of Sam Watkins and the SuperCDMS-CPD collaboration, used in Appl. Phys. Lett. 118, 022601 (2021)

TES on main crystal

8

Physics of particle recoils



Particle recoil produces population of high energy phonons that down-convert rapidly.

Crystal is after ~µs uniformly filled, phonons thermalize within ~ms.

Share ε of athermal phonons are collected by the superconducting film and thermalize there, $(1-\varepsilon)$ in the crystal.

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Electrothermal response model

System can be modeled with ordinary differential equations (ODEs), as interaction of thermal and electrical circuits:

thermal

ODE

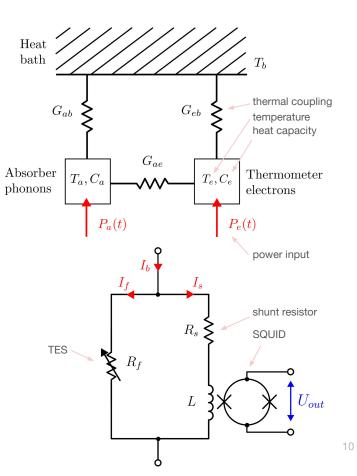
thermal ODEs
$$\begin{array}{l} C_e \frac{dT_e}{dt} + (T_e - T_a) \, G_{ea} + (T_e - T_b) \, G_{eb} = P_e(t), \\ C_a \frac{dT_a}{dt} + (T_a - T_e) \, G_{ea} + (T_a - T_b) \, G_{ab} = P_a(t), \\ \end{array}$$
electrical
$$L \frac{dI_f}{dt} + R_s I_b - (R_f(T_e) + R_s) I_f = 0. \end{array}$$

Model combined from

100

June 7, 2023 Journal of Low Temperature Physics volume 100, pages 69-104 (1995) Topics in Applied Physics, vol 99, Pages 63-150 (2005)

1,



Pulse shape

System can be simplified to describe the temperature rise from a small energy deposition $\Delta E \propto \Delta P$:

$$\frac{d}{dt} \begin{pmatrix} \Delta T_e \\ \Delta T_a \end{pmatrix} = - \begin{pmatrix} \frac{G_{ea} + G_{eff}}{C_e} & -\frac{G_{ea}}{C_e} \\ -\frac{G_{ea}}{C_a} & \frac{G_{ea} + G_{ab}}{C_a} \end{pmatrix} \begin{pmatrix} \Delta T_e \\ \Delta T_a \end{pmatrix} + \begin{pmatrix} \frac{\Delta P_e}{C_e} \\ \frac{\Delta P_a}{C_a} \end{pmatrix}$$

Solutions for temperature rise in thermometer are pulses of \approx 1-100 ms length, with two components and three time constants:



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$$\Delta T_e(t) = \Theta(t) \left[A_n \left(e^{-t/\tau_n} - e^{-t/\tau_{in}} \right) + A_t \left(e^{-t/\tau_t} - e^{-t/\tau_n} \right) \right]$$

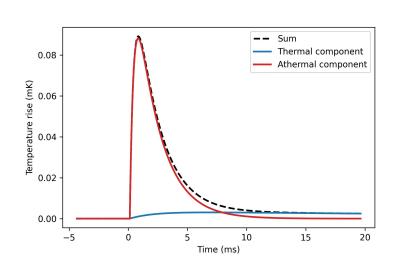
(formulas in appendix)

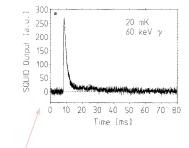




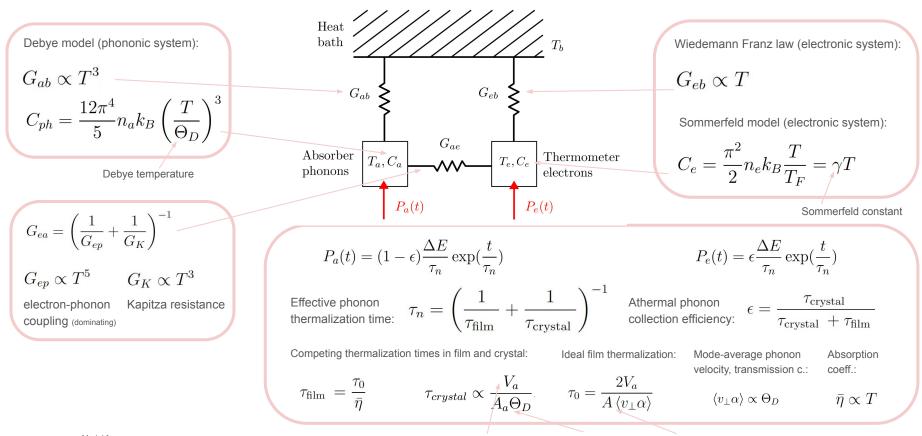


measured pulse for comparison



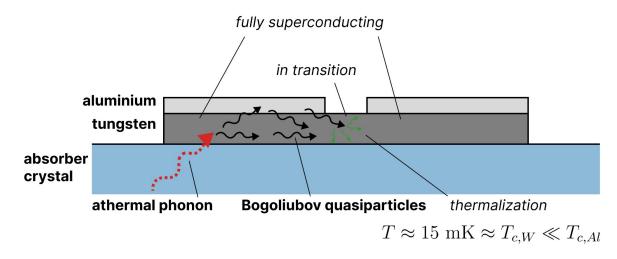


Thermal properties



June 7, 2023 Model from Journal of Low Temperature Physics volume 100, pages 69–104 (1995) Felix Wagner Volume absorber Surface absorber Interface TES

Optimizing athermal phonon collection



For optimized detectors, sensitivity (pulse height and area) is

$$\propto arepsilon \Delta E/C_e$$

Collection efficiency ε can be increased without adding heat capacity, through phonon collectors.

Aluminium/tungsten bilayer is kept through proximity effect in superconducting state (exponentially suppressed heat capacity). Scattering **athermal phonons break cooper pairs** that travel diffusively to the uncovered tungsten film.

Noise model

Similarly to the pulse shape model, a model for noise contributions can be derived, depending on the occurring power and voltage fluctuations in the circuit, and the bias heating P₁:

$$C_e \frac{dT_e}{dt} + (T_e - T_b)G_{eb} = \Delta P_{e,noise} + P_J,$$
$$-L \frac{dI_f}{dt} + R_S I_b - (R_f(T_f) + R_S)I_f = \Delta U_{noise},$$

Solution of the system is a transition matrix S, which translates fundamental fluctuations to observable currents in the readout circuit I_r:

$$\begin{pmatrix} \Delta T_e \\ \Delta I_f \end{pmatrix} \stackrel{\text{def}}{=} \begin{pmatrix} s_{11}(w) & s_{12}(w) \\ s_{21}(w) & s_{22}(w) \end{pmatrix} \begin{pmatrix} \Delta P_e \\ \Delta U \end{pmatrix}$$

(formulas in appendix)

Fundamental sensor noise contributions

1) Electrical, Johnson noise from movement of electrons in shunt with temperature T_s and film in superconducting transition with resistance R_{r_0} :

$$|\Delta U|_{J_s}^2 = 4k_B T_s R_s \qquad |\Delta U|_{J_f}^2 = 4k_B T_e R_{f0}$$

2) Phonon noise from thermal fluctuations between TES and heat bath:

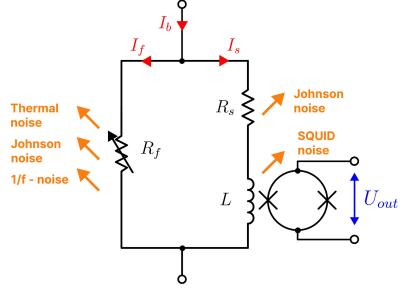
$$|\Delta P_e|_{ph}^2 = 4k_B T_e^2 G_{eb} \frac{2}{5} \frac{1 - \left(\frac{T_b}{T_{e0}}\right)^5}{1 - \left(\frac{T_b}{T_{e0}}\right)^2}$$

3) TES-intrinsic 1/f excess-flicker noise:

$$|\Delta P_e|_{flicker}^2 = \frac{\left(\frac{\Delta R_{f,flicker}}{R_{f0}}\right)^2 R_{f0}^2}{w^{\alpha}} I_{f0}^2$$

4) "White" SQUID noise.





Intrinsic excess noise

Internal Thermal Fluctuation Noise (ITFN): thermometer is spacially extended object, consists therefore of a system of connected heat capacities ⇒ thermal fluctuations exist between them. Higher normal resistance can make it worse.

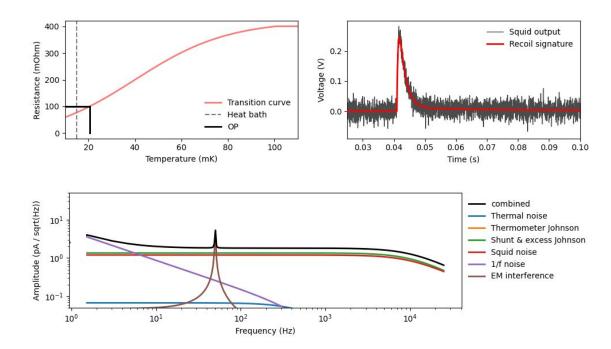
Excess electrical noise*: spectral shape similar to Johnson noise, larger for lower normal resistance and operation low in the transition, affected by homogeneity of film and magnetic field.

1/f-noise*: spectral shape rises steeply towards lower frequencies, cutoff << 1 Hz, often correlated with excess electrical noise.

Telegraph noise*: associated with certain bias values and points in transition curve. Possibly due to structure reordering in metallic film.

Possible sources: non-equilibrium effects, fundamental effects of superconductivity or electron-phonon interaction.

Exemplary detector response



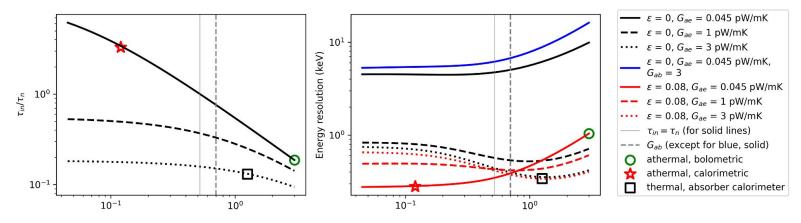
Simulation of detector with silicon target and iridium/gold bilayer TES, based on introduced response and noise models.

Transition curves of modern devices are much steeper, with a width of \approx 1-2 mK.

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Optimizing thermal couplings

variable: G_{eb} , G_{ae} , ε ; fixed: $C_e = 0.001 \text{ pJ/mK}$, $C_a = 0.038 \text{ pJ/mK}$, $G_{ab} = 0.7 \text{ pW/mK}$ (except for blue, solid)

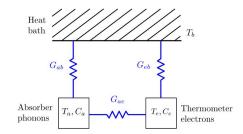


TES-bath coupling G_{eb} (pW/mK)

Ratio of thermometer relaxation time τ_{in} to phonon thermalization time τ_n separates **bolometric** (measuring power) and **calorimetric** (measuring energy) modi.

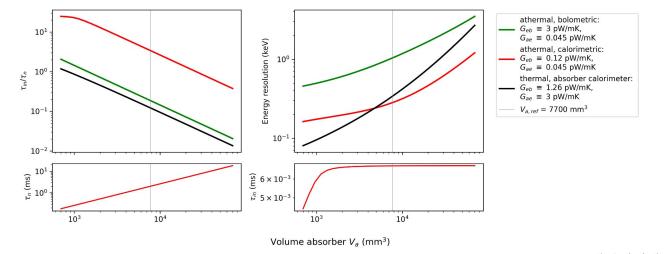
Most modern devices are athermal calorimeters.

Small, thermal detectors can reach comparable thresholds, but scale differently.



Scaling absorber volume

variable: $\tau_n \propto V_a$, $C_a \propto V_a$; fixed: G_{ae} , G_{eb} , $C_e = 0.001$ pJ/mK, $G_{ab} = 0.7$ pW/mK, $C_a(V_{a,ref}) = 0.038$ pJ/mK, $\tau_n(V_{a,ref}) = 2$ ms, $\varepsilon = 0.08$

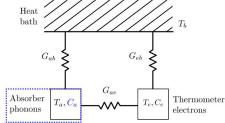


Threshold of **thermal** detector scales $\propto V_a$,

athermal detector scales $\propto V_a^{2/3}$.

 \Rightarrow Operating large target with low threshold challenging,

lower threshold achievable by consequently lowering target mass!



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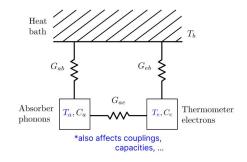
Scaling operation temperature

 $\tau_n(T_{ref}) = 2 \text{ ms}, C_a(T_{ref}) = 0.038 \text{ pJ/mK}, C_e(T_{ref}) = 0.001 \text{ pJ/mK}, G_{ab}(T_{ref}) = 0.7 \text{ pW/mK}, \varepsilon(T_{ref}) = 0.08$ athermal, bolometric: 10¹ $G_{eb}(T_{ref}) = 3 \text{ pW/mK}.$ 10^{0} $G_{ae}(T_{ref}) = 0.045 \text{ pW/mK}$ resolution (keV) athermal. calorimetric: $G_{eb}(T_{ref}) = 0.12 \text{ pW/mK},$ $G_{ae}(T_{ref}) = 0.045 \text{ pW/mK}$ ^u_L/^u₁₀⁻¹ thermal, absorber calorimeter: $G_{eb}(T_{ref}) = 1.26 \text{ pW/mK},$ $G_{ae}(T_{ref}) = 3 \text{ pW/mK}$ Energy 100 $T_{ref} = 20.6 \text{ K}$ 10^{-2} $\begin{array}{c} 2 \times 10^{0} \\ 1.95 \times 10^{0} \\ 1.9 \times 10^{0} \\ 1.85 \times 10^{0} \end{array}$ 2×10^{-1} ω ⊢ 1.8×10° 10^{-1} $1.75 \times 10^{\circ}$ 2×10^{1} 2×10^{1} 3×10^{1} 3×10^{1} 4×10^{1} 6×10^{1} 4×10^{1} 6×10^{1} Operation temperature T (mK)

Thermal couplings scale strongly with temperature: $G_{ae} \propto T^5$.

Calorimetric modus only possible at lowest ≈10-30 mK temperatures!

Heat capacities decrease with temperature \Rightarrow sensitivity improves.



Low energy excess events

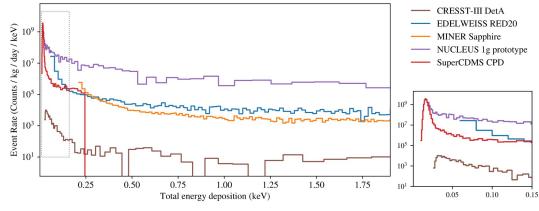


Fig. from EXCESS workshop summary SciPost Phys. Proc. 9, 001 (2022)

Experiments observe steeply rising non-ionising background towards lower energies ≈<100 eV. Event rate ...

- ... apparently uncorrelated with detector mass, surface, surroundings and temperature.
- ... decays after cool down of experiment, on time scale of O(months)*.

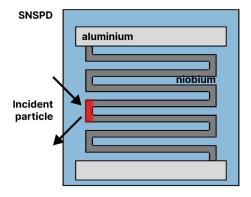
Currently most probable hypothesis: **stress effect**, e.g. on crystal-thermometer interface.

Ongoing community-organized workshop series to discuss the excesses: <u>EXCESS workshop</u>.

Relations to other quantum technologies

Transmon qubit

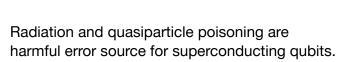
substrate



 $T \approx 1 - 10K$

Superconducting nanowire single-photon detector (SNSPD):

- Wire is operated slightly below T_c.
- Difference to TES: sensor is target.
- Recoil needs to warm wire only locally for measurable effect.
- Can reach lower energy resolution.
- Cannot collect significant exposure.



Phonon population in substrate, induced by radiation or possibly the low energy excess, can break cooper pairs in qubit layer.

Similar effect as in phonon collectors of TES.

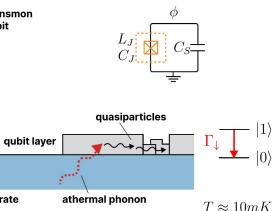
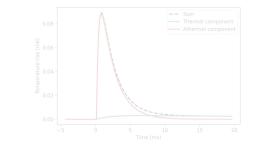


Fig. adapted from Nature Physics volume 18, pages 107-111 (2022)

Qubit errors through radiation discussed in Nature volume 594, pages 369-373 (2021) Nature Communications volume 12. Article number: 2733 (2021) Nature Physics volume 18, pages 107-111 (2022) The European Physical Journal C volume 83. Article number: 94 (2023)

Qubit errors through guasiparticle excess discussed in Nature Physics volume 18, pages 145-148 (2022) arXiv:2208.02790





 Cryogenic detectors with superconducting thermometers are currently the most sensitive devices for sub-GeV/C² DM searches.

Summary

- Lowest achieved nuclear recoil energy threshold is **10 eV**.
- Sensitivity in calorimetric mode is dominated by capability to **collect athermal phonons** from target.
- Similar challenges in fabrication and background mitigation with other superconducting technologies: SNSPDs, transmon qubits, ...
- Devices are continuously improved, new phenomena continue to be discovered and research is done to understand them: **low energy excess** events and excess noise.

Appendices

Formulas for pulse shape model

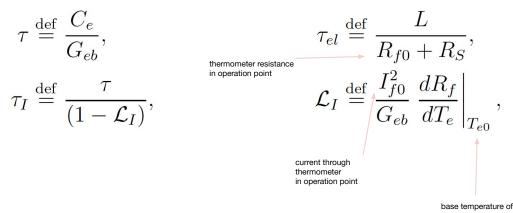
$$\begin{split} &= -I_{f}^{2} \frac{(R_{f0} - R_{s})}{(R_{f0} + R_{s})} \frac{dR_{f}}{dT_{e}} \Big|_{T_{e0}} \qquad G_{eff} = G_{eb} + G_{ETF} \\ &\stackrel{\text{def}}{=} -G_{ETF}. \\ &\lambda_{1,2} = \frac{-a \mp \sqrt{a^{2} - 4b}}{2} \stackrel{\text{def}}{=} -\frac{1}{\tau_{1,2}} \qquad \left(\stackrel{\text{def}}{=} -\frac{1}{\tau_{in,t}} \stackrel{\text{def}}{=} -s_{in,t} \right), \qquad A_{n} \stackrel{\text{def}}{=} \frac{\frac{\Delta E}{\tau_{n}}}{\left(1 - \frac{\alpha_{2}}{\alpha_{1}}\right)} \left(\alpha_{2} \frac{(1 - \epsilon)}{C_{a}} - \frac{\epsilon}{C_{e}} \right) \left(\frac{1}{\tau_{n}} - \frac{1}{\tau_{1}} \right)^{-1} \\ &= -\frac{P_{0}(s_{in} - (G_{ab}/C_{a}))}{\varepsilon(s_{in} - s_{in})(s_{in} - s_{n})} \left(\frac{s_{t} - (G_{ab}/C_{a})}{G_{eff} - (C_{e}/C_{a})G_{ab}} - \frac{\varepsilon}{C_{e}} \right) \\ &\alpha_{1,2} \stackrel{\text{def}}{=} 1 + \frac{G_{ab}}{G_{ea}} + \lambda_{1,2} \frac{C_{a}}{G_{ea}}. \qquad A_{t} \stackrel{\text{def}}{=} -\frac{\frac{\Delta E}{\tau_{n}}}{\left(1 - \frac{\alpha_{1}}{\alpha_{2}}\right)} \left(\alpha_{1} \frac{(1 - \epsilon)}{C_{a}} - \frac{\epsilon}{C_{e}} \right) \left(\frac{1}{\tau_{n}} - \frac{1}{\tau_{2}} \right)^{-1} \\ &= \frac{P_{0}(s_{t} - (G_{ab}/C_{a}))}{\varepsilon(s_{t} - s_{in})(s_{t} - s_{n})} \left(\frac{s_{in} - (G_{ab}/C_{a})}{G_{eff} - (C_{e}/C_{a})G_{ab}} - \frac{\varepsilon}{C_{e}} \right) \\ \end{array}$$

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-1

 $-\frac{\varepsilon}{C_e}$

Definitions for noise matrix elements



thermometer in operation point

Formulas for intrinsic noise matrix elements

$$s_{21}^{int}(w) = \frac{-\frac{G_{eb}\mathcal{L}_{I}}{I_{f0}}}{\left(\frac{1}{(\tau)_{int}} + 2\pi wi\right)C_{e}\left(\frac{1}{\tau_{el}} + 2\pi wi\right)L - \frac{G_{eb}\mathcal{L}_{I}}{I_{f0}}(I_{f0}R_{s} - I_{f0}R_{f0})_{int}}$$

$$= -\frac{1}{I_{f0}}\left[\frac{L\tau}{(\tau)_{int}\tau_{el}\mathcal{L}_{I}} - \frac{(I_{f0}R_{s} - I_{f0}R_{f0})_{int}}{I_{f0}}\right]$$

$$+ 2\pi wi\frac{L\tau}{\mathcal{L}_{I}}\left(\frac{1}{(\tau)_{int}} + \frac{1}{\tau_{el}}\right) - \frac{4\pi^{2}w^{2}\tau L}{\mathcal{L}_{I}}\right]^{-1},$$

$$s_{22}^{int}(w) = \frac{\left(\frac{1}{(\tau)_{int}} + 2\pi wi\right)C_{e}}{\left(\frac{1}{(\tau)_{int}} + 2\pi wi\right)C_{e}\left(\frac{1}{\tau_{el}} + 2\pi wi\right)L - \frac{G_{eb}\mathcal{L}_{I}}{I_{f0}}(I_{f0}R_{s} - I_{f0}R_{f0})_{int}}$$

$$= -s_{21}^{int}(w)I_{f0}\frac{1}{\mathcal{L}_{I}}(1 + 2\pi wi\tau).$$

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Formulas for external noise matrix elements

$$s_{21}^{ext}(w) = \frac{-\frac{G_{eb}\mathcal{L}_{I}}{I_{f0}}}{\left(\frac{1}{(\tau_{I})_{ext}} + 2\pi wi\right)C_{e}\left(\frac{1}{\tau_{el}} + 2\pi wi\right)L - \frac{G_{eb}\mathcal{L}_{I}}{I_{f0}}(-2R_{f0}I_{f0})_{ext}}$$

$$= -\frac{1}{I_{f0}}\left[\frac{L\tau}{(\tau_{I})_{ext}\tau_{el}\mathcal{L}_{I}} - \frac{(-2R_{f0}I_{f0})_{ext}}{I_{f0}}\right]$$

$$+ 2\pi wi\frac{L\tau}{\mathcal{L}_{I}}\left(\frac{1}{(\tau_{I})_{ext}} + \frac{1}{\tau_{el}}\right) - \frac{4\pi^{2}w^{2}\tau L}{\mathcal{L}_{I}}\right]^{-1},$$

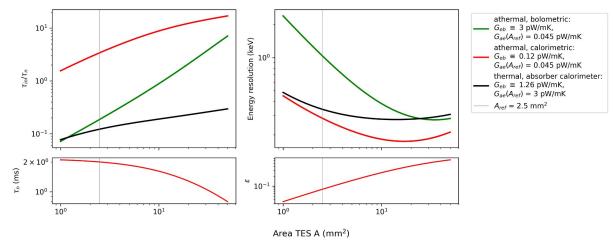
$$s_{22}^{ext}(w) = \frac{\left(\frac{1}{(\tau_{I})_{ext}} + 2\pi wi\right)C_{e}}{\left(\frac{1}{(\tau_{I})_{ext}} + 2\pi wi\right)C_{e}\left(\frac{1}{(\tau_{I})_{ext}} + 2\pi wi\right)L - \frac{G_{eb}\mathcal{L}_{I}}{I_{f0}}(-2R_{f0}I_{f0})_{ext}}$$

$$= s_{21}^{ext}(w)I_{f0}\frac{\mathcal{L}_{I} - 1}{\mathcal{L}_{I}}(1 + 2\pi wi\tau_{I}).$$

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Optimizing thermometer area

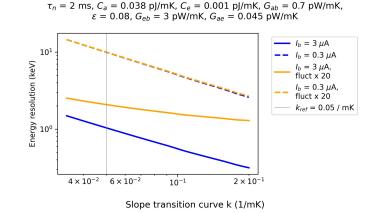
variable: $\varepsilon(A)$, $\tau_n(A)$, $C_e \propto A$, $G_{ae} \propto A$; fixed: G_{eb} , $C_a = 0.038$ pJ/mK, $G_{ab} = 0.7$ pW/mK, $C_e(A_{ref}) = 0.001$ pW/mK, $\tau_n(A_{ref}) = 2$ ms, $\varepsilon(A_{ref}) = 0.08$



Larger thermometer increases TES heat capacity linearly, collection efficiency (slightly) nonlinearly, and TES relaxation time (almost) linearly,

 \Rightarrow for a given detector, **a broad optimum** for it's TES size exists!

Optimizing transition curves



Steeper transition curve improves signal from TES resistance changes against electrical sensor noise.

 \Rightarrow the steeper, the better (with some practical stability constraints ...)

However, if TES-intrinsic noise dominates over electrical noise, a steeper curve cannot improve the sensitivity further.