

**MoEDAL**

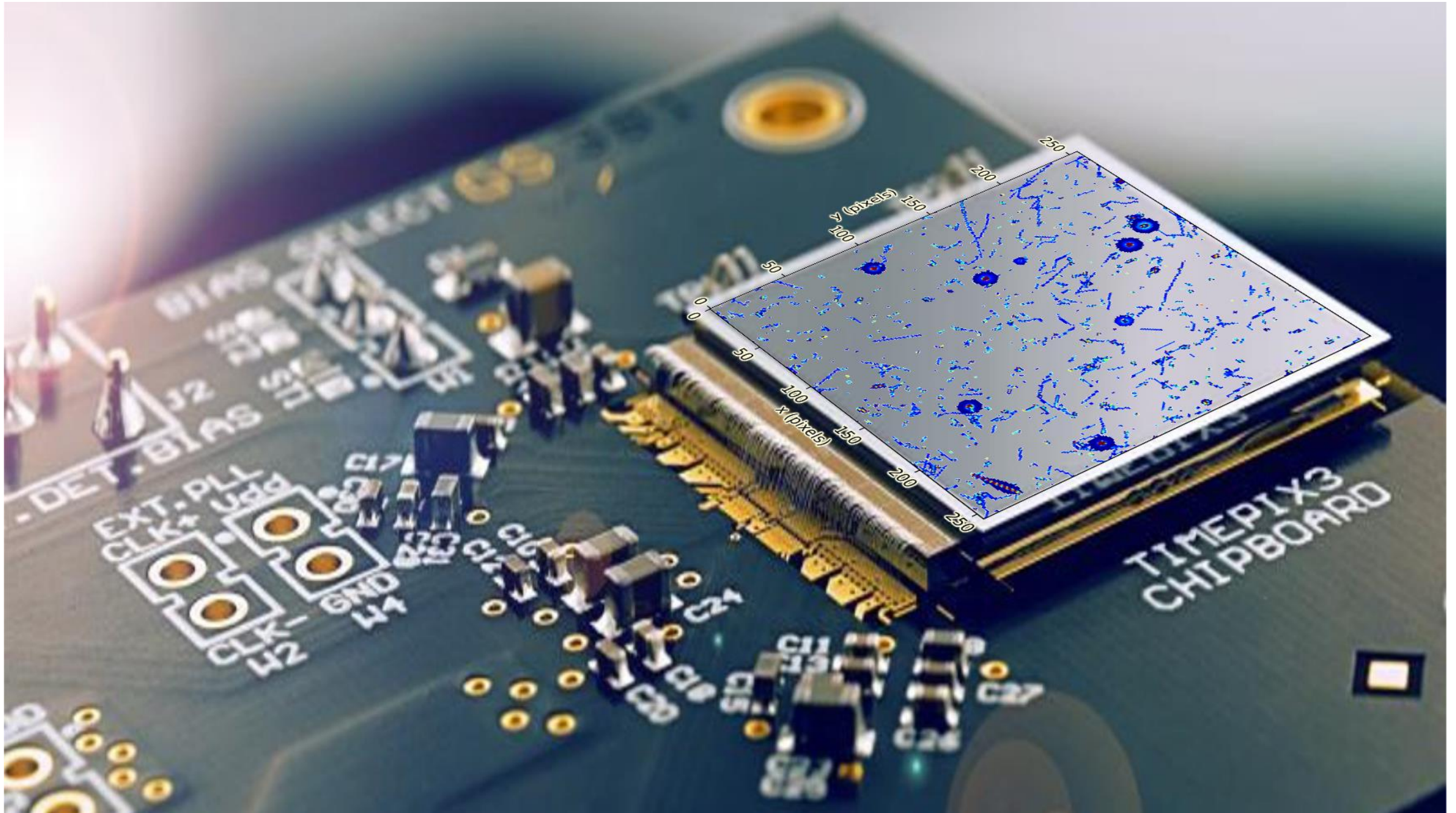
# **Application of single-layer particle tracking for radiation field decomposition and interaction point reconstruction at MoEDAL**

**Declan Garvey**, Benedikt Bergmann, Petr Mánek, Stanislav Pospíšil,  
Petr Smolyanskiy,

**On behalf of MoEDAL**

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# Timepix3: Radiation Imaging Detector



# Timepix and Timepix3

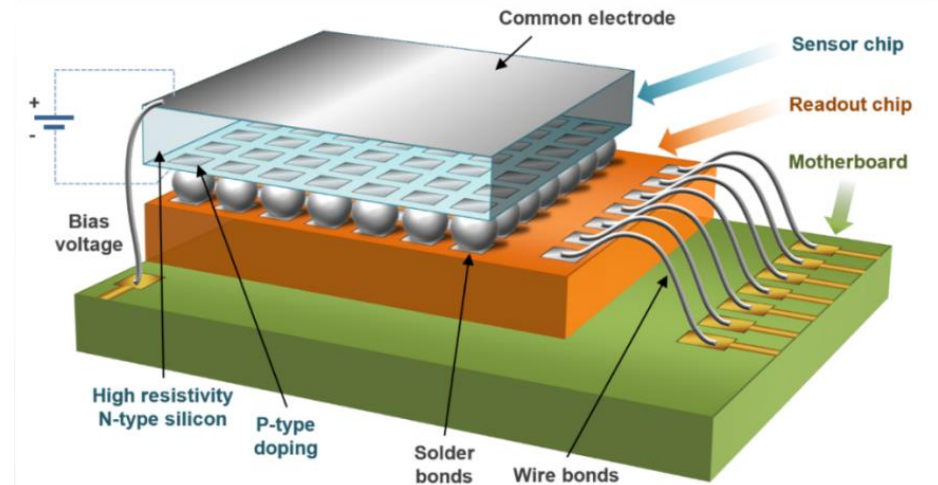
- 256 x 256 pixels with 55  $\mu\text{m}$  pitch (1.98  $\text{cm}^2$ )
- Sensor layer (silicon, GaAs, CdTe, ...) flip-chip bump bonded to the ASIC

## Timepix

- Frame-based readout (92 fps) - dead time  $> 11$  ms
  - ➔ Information is read out on a frame-by-frame basis
- Measurement of energy **or** time ( $\Delta t = 20.8$  ns)
- Minimal detectable energy:  $\sim 3.5$  keV

## Timepix3

- Data-driven readout (max. count rate 40 Mpix/s)
  - ➔ Pixels are continuously read out throughout the entire measurement
- Per pixel dead time: 475 ns
- Measurement of energy **and** time ( $\Delta t = 1.56$  ns)
- Minimal detectable energy per pixel: 3 keV



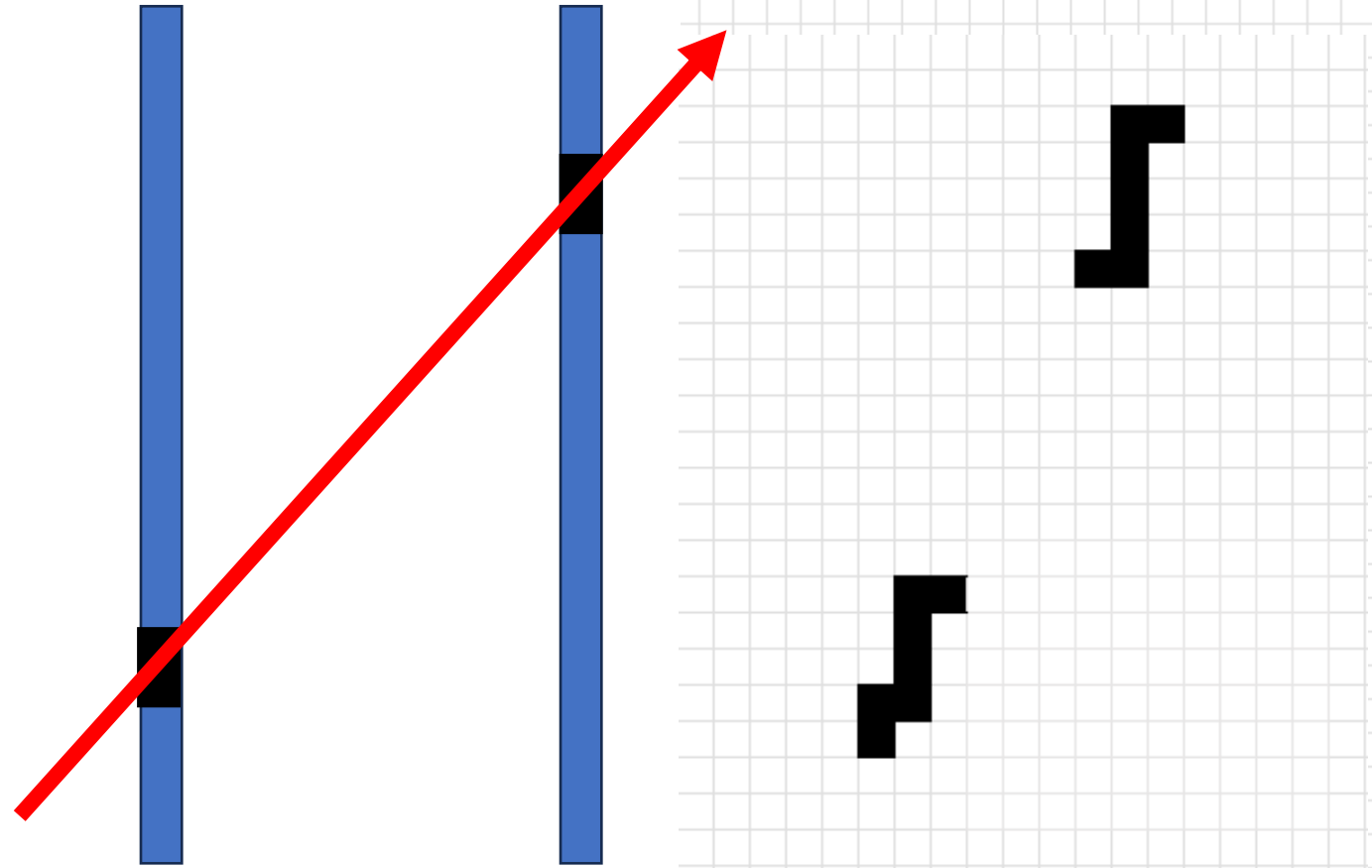
Timepix3 with chipboard

# Particle Tracking by Connecting the Dots

- Traditionally, particle tracking is performed using multiple layers of thin detectors
- Tracking information is then calculated by recording coincidence measurements in each detector and "connecting the dots"

So, the question is:

***Can we achieve similar results using only a single-layer detector?***

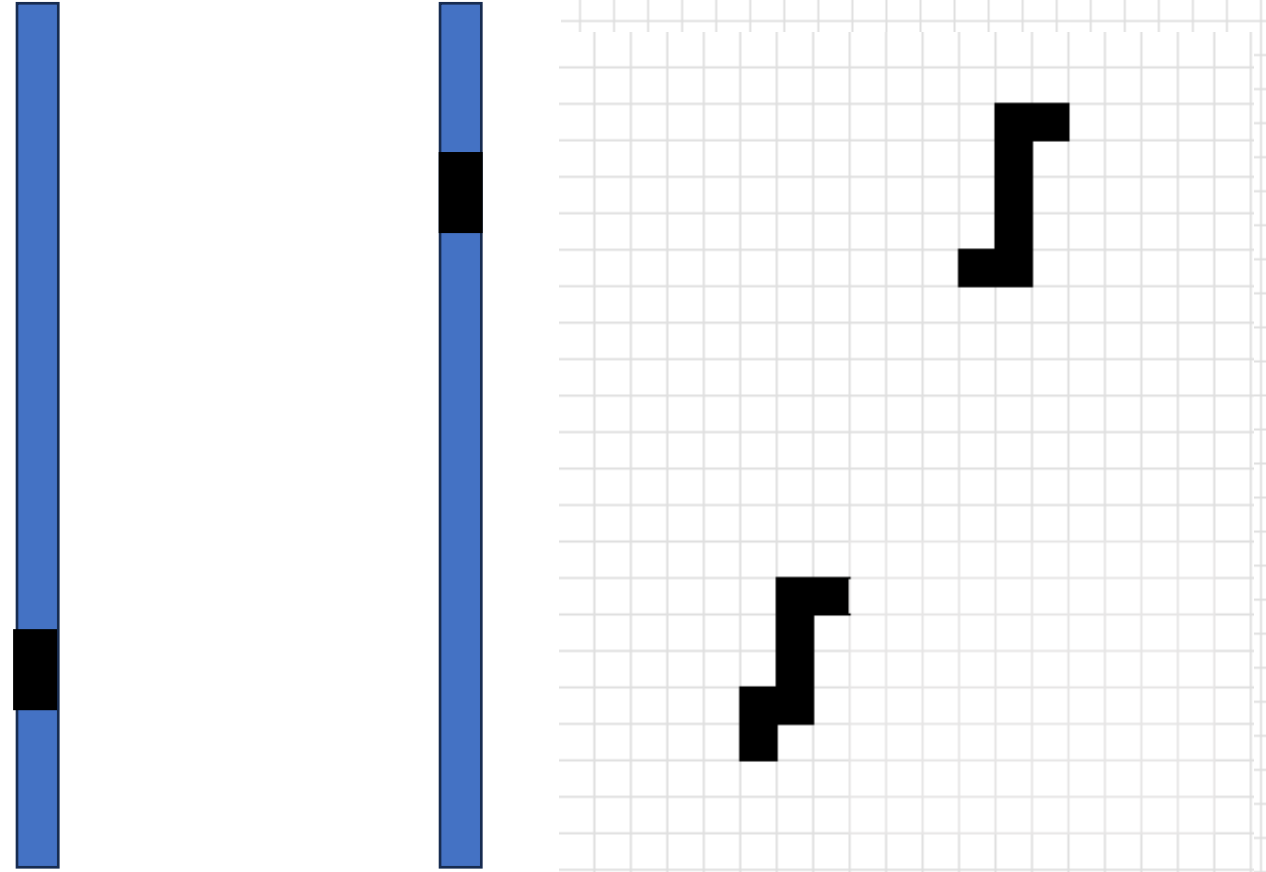


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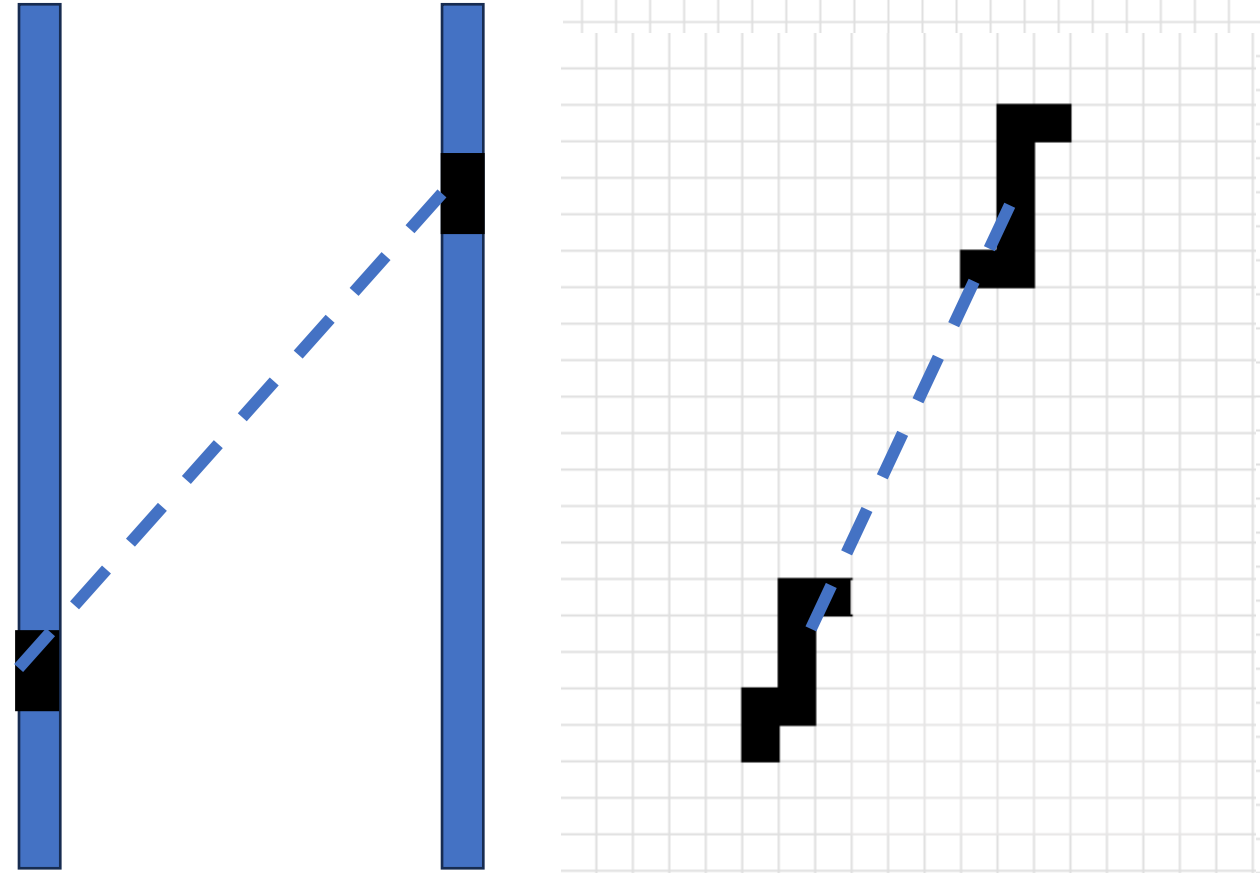


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# 3D Reconstruction of Particle Tracks



Charge carrier drift motion:

$e^-$  and  $h^+$  drift described by

$$v_e = -\mu_e \times E(z)$$

$$v_h = \mu_h \times E(z)$$

$\mu_{e/h}$ : Mobility of  $e^-/h^+$

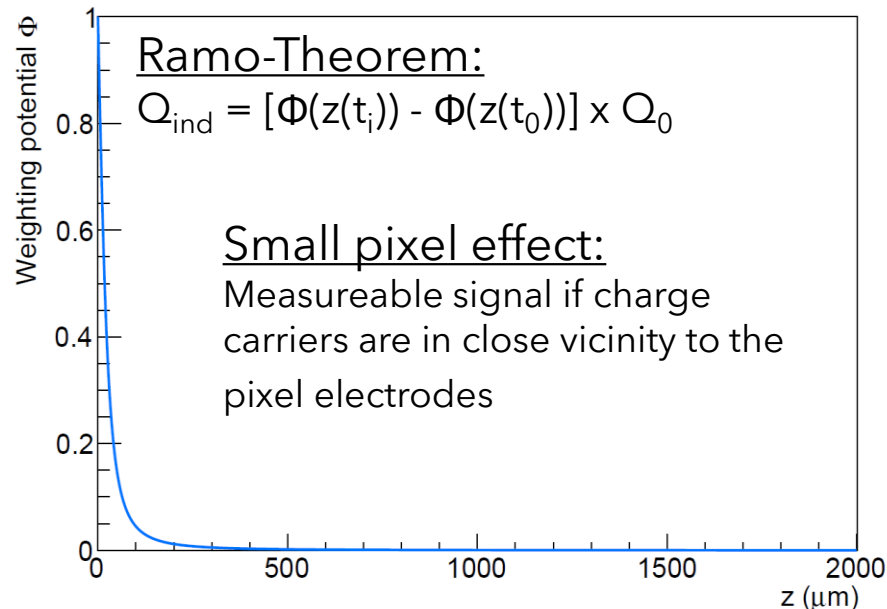
Electric field parametrisation:

Si:  $\vec{E}(z) = \frac{U_B}{d} \vec{e}_z + \frac{2U_{dep}}{d^2} \left(\frac{d}{2} - z\right) \vec{e}_z ;$

CdTe:  $\vec{E}(z) = \frac{U_B}{d} \vec{e}_z$

$U_B$ : Bias voltage;  $U_{dep}$ : Depletion voltage;  $d$ : Sensor thickness

2000  $\mu\text{m}$ , pixel pitch: 55  $\mu\text{m}$

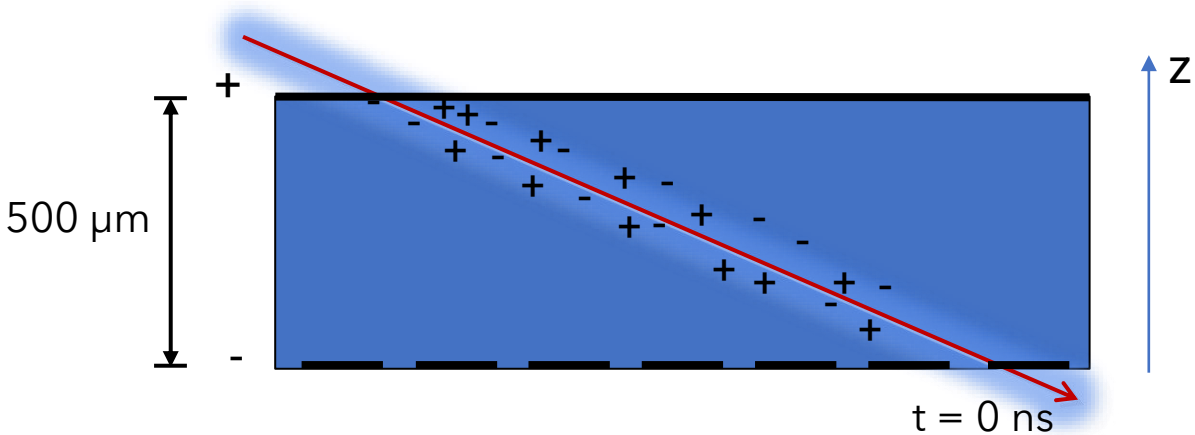


→ Look up table:  $z(t_{meas.}, E_{meas.})$

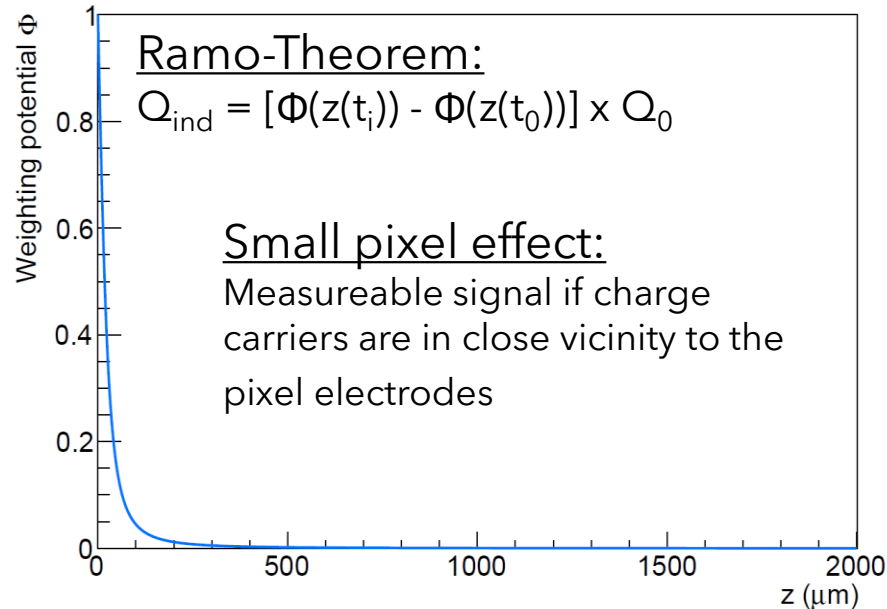
Bergmann et al. Eur. Phys. J. C (2017) 77: 421. <https://doi.org/10.1140/epjc/s10052-017-4993-4>

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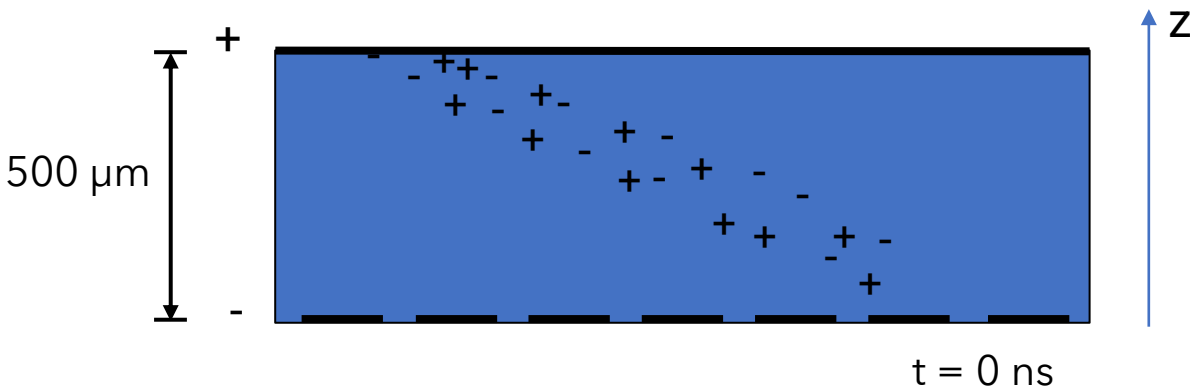
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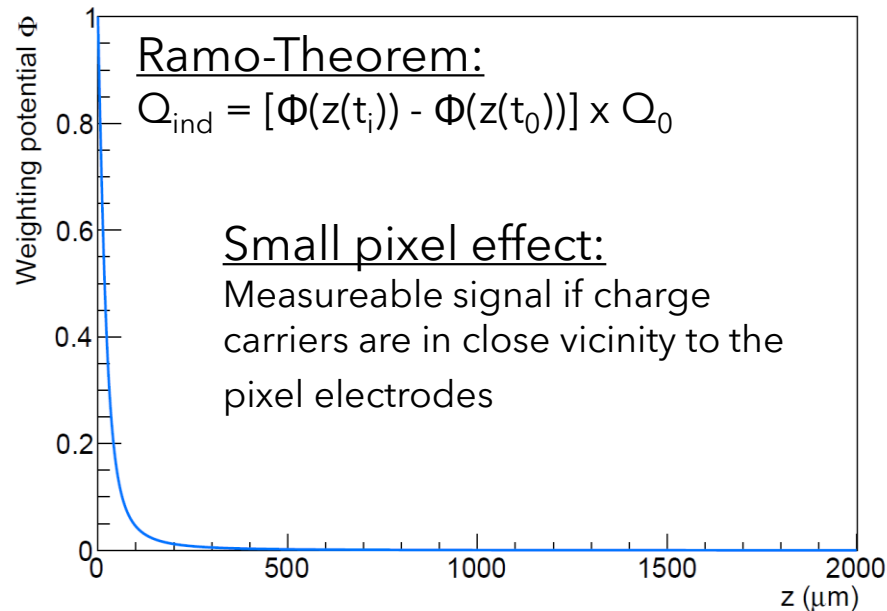
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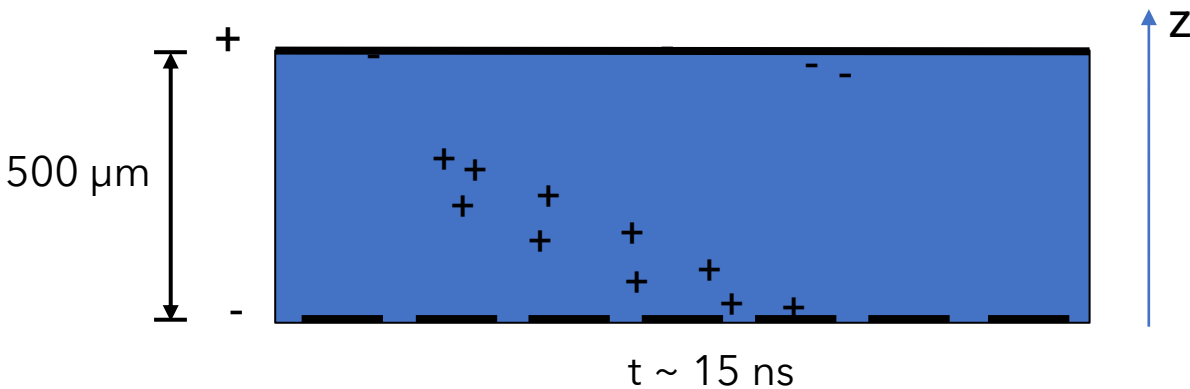


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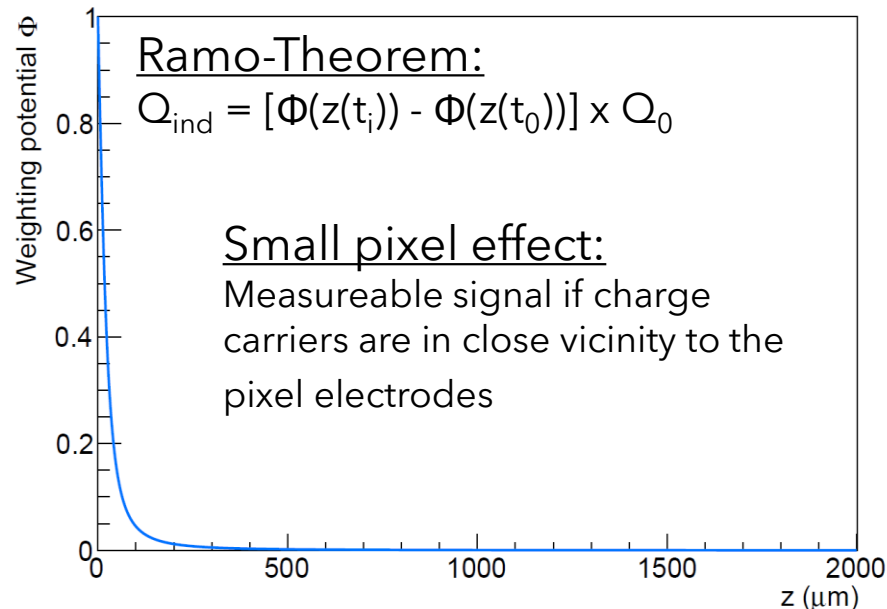
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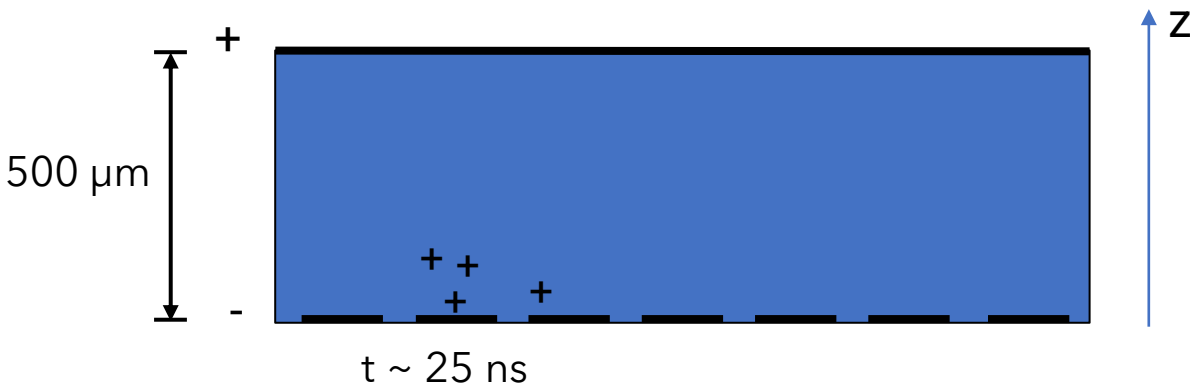
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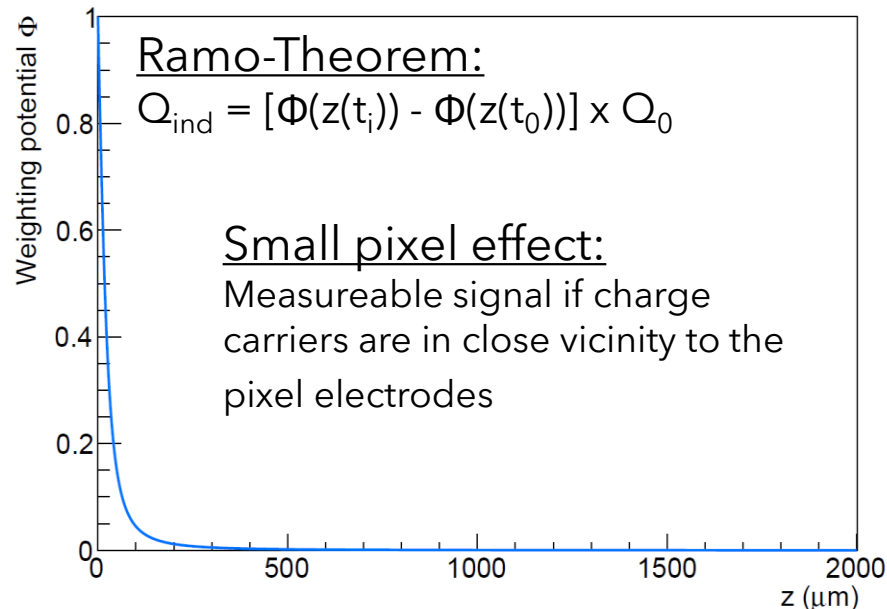
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2000 μm, pixel pitch: 55 μm



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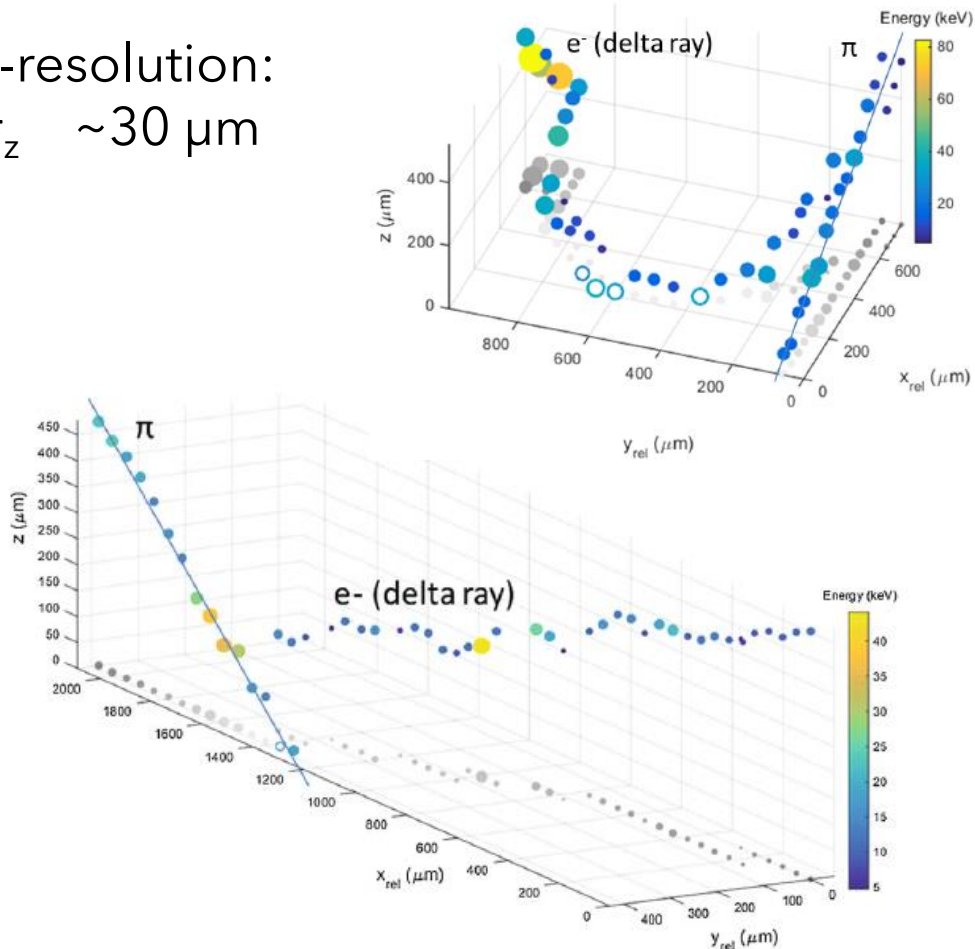
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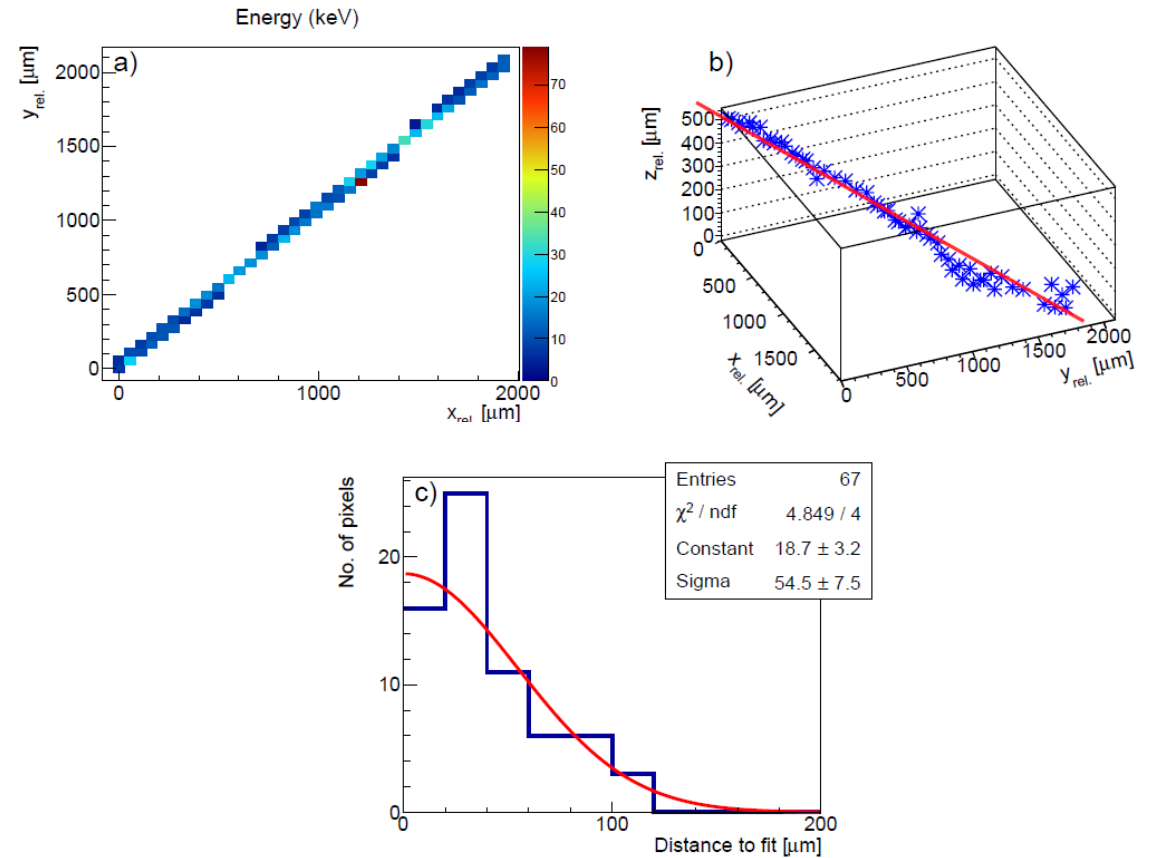
# 3D Tracking in 500 $\mu\text{m}$ Si Timepix3

120 GeV/c pion tracks accompanied by  $\delta$ -rays measured at SPS:

z-resolution:  
 $\sigma_z \sim 30 \mu\text{m}$

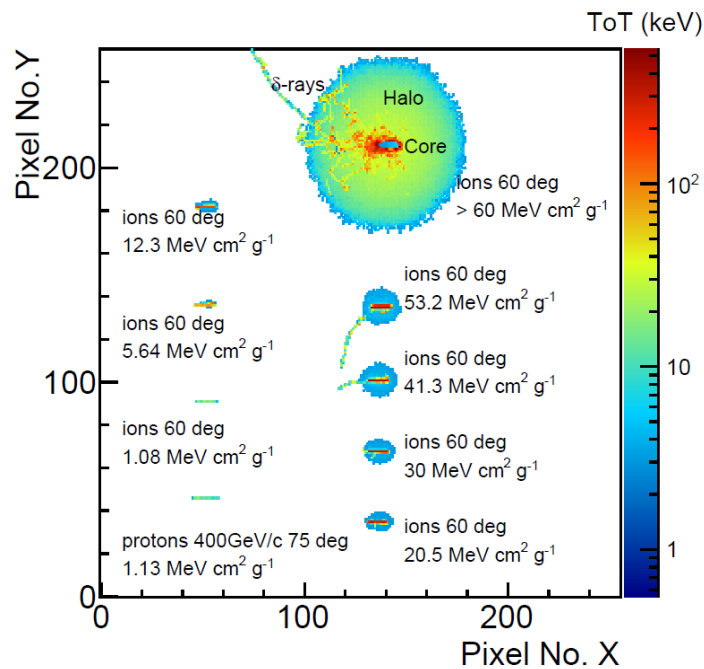


A cosmic  $\mu$  measured in the Prague laboratory:



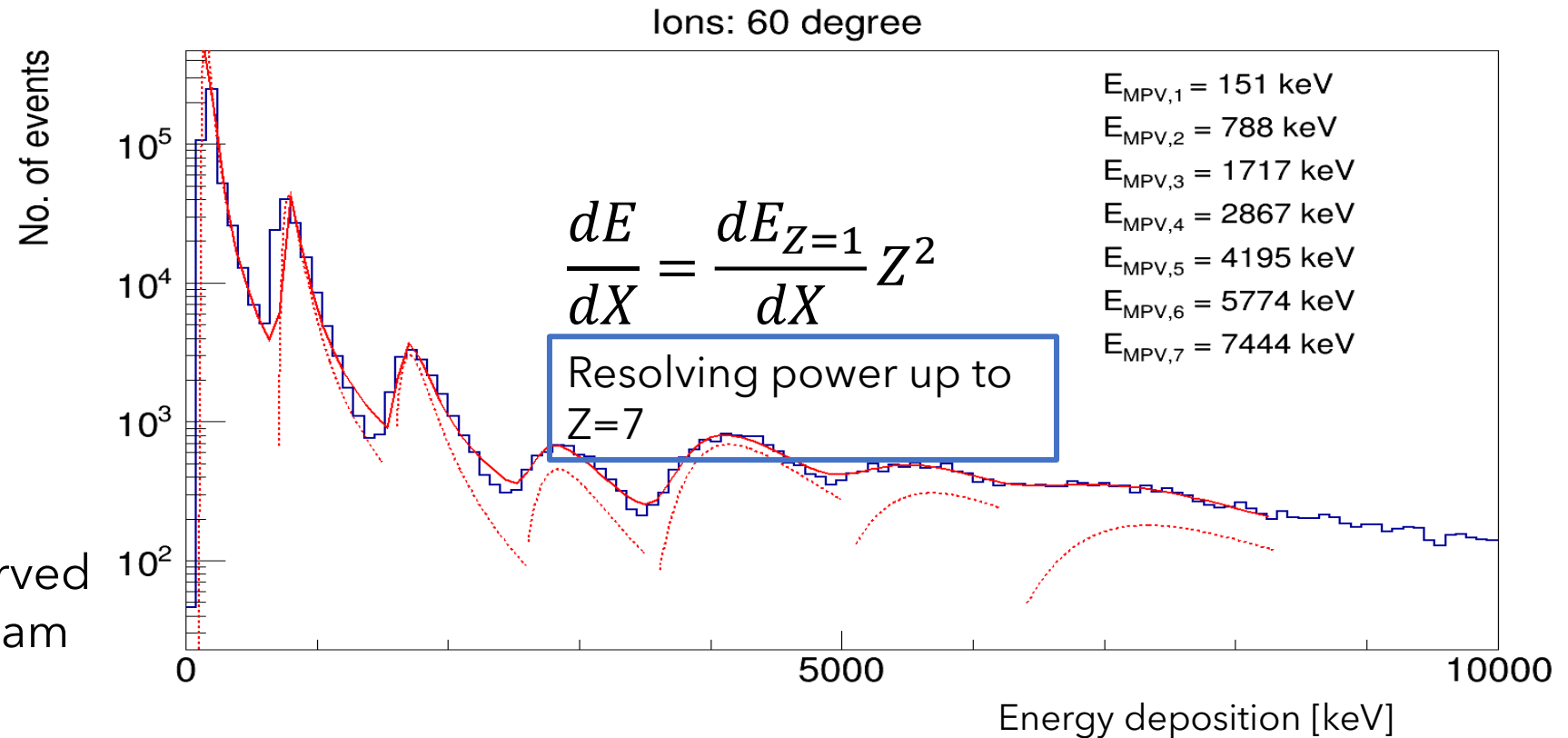
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# Energy Deposition Spectra of Charged Energetic Particles



Different track shapes observed in a mixed relativistic ion beam 330 GeV/c Pb on target

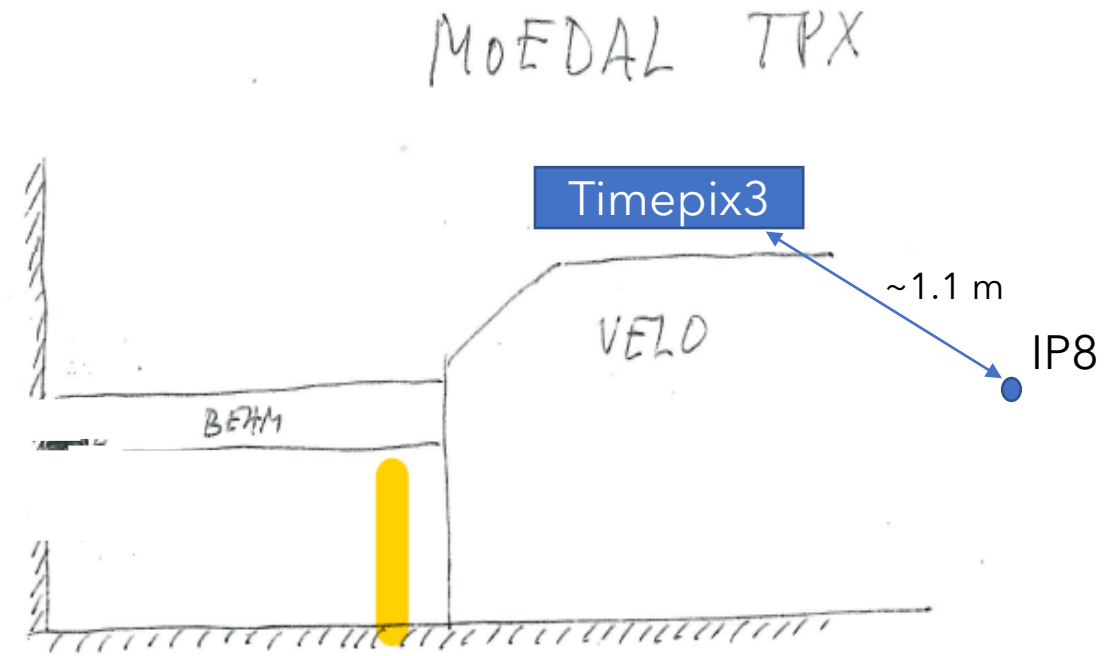
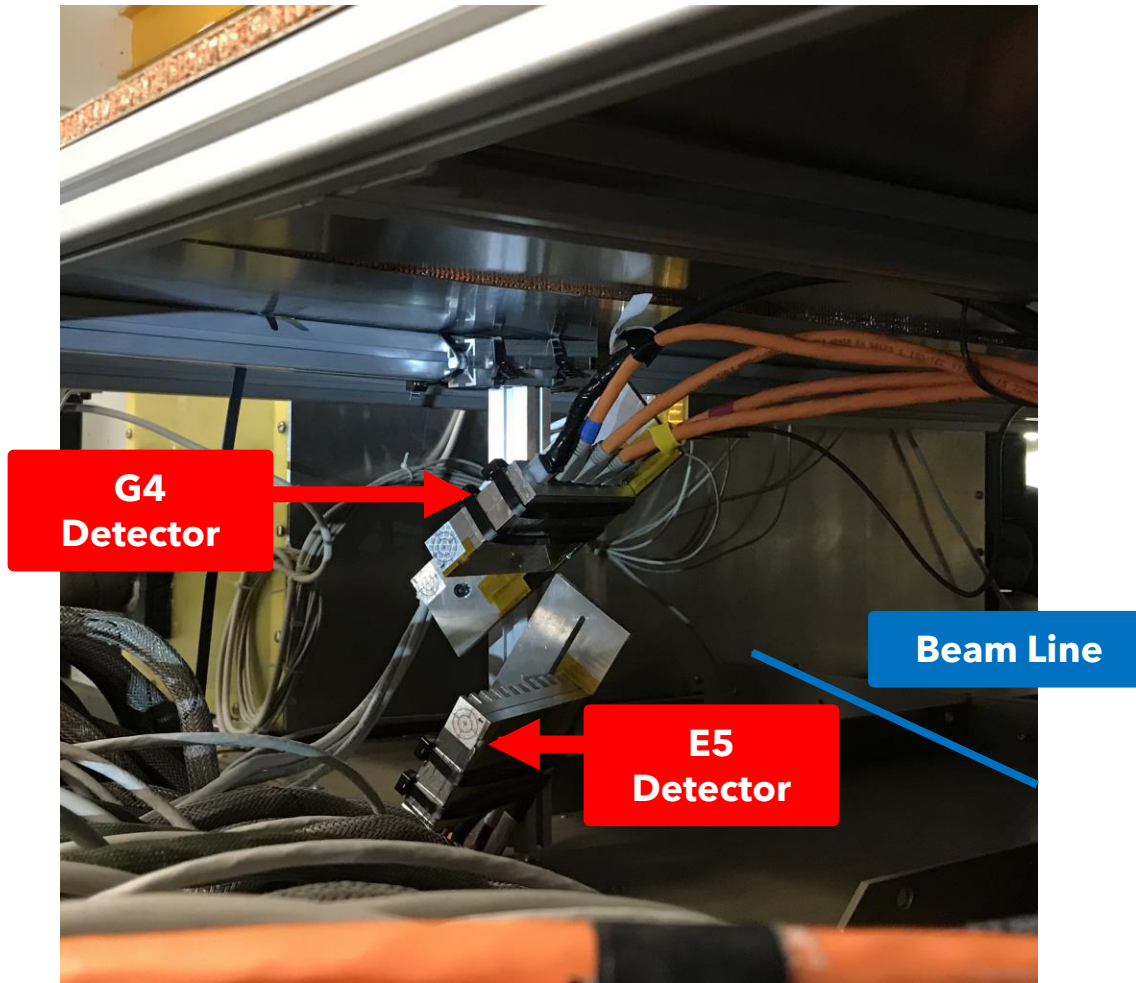
Energy loss of energetic charged particles described by Landau-Vavilov distribution with Gaussian smearing



# **Single-Layer Particle Tracking in MoEDAL**

# Timepix3 in MoEDAL at Run 2 - Feasibility Study

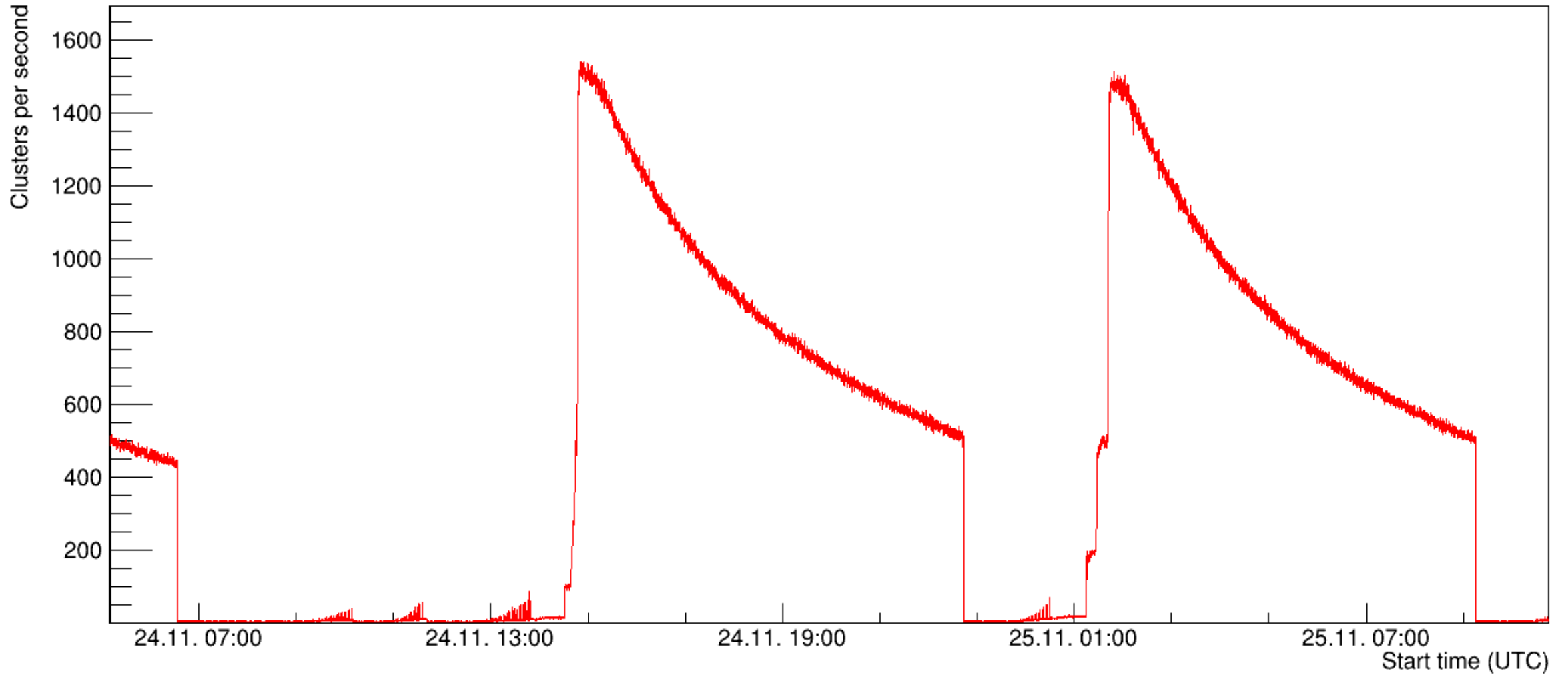
Installation of two Timepix3 detectors in MoEDAL in **September 2018**. Timepix3 are placed at 1.1 m distance to IP8 with a relatively unobstructed view



Continuous quasi dead-time free measurement (in real time) keeping a permanent record of **all particle traces**

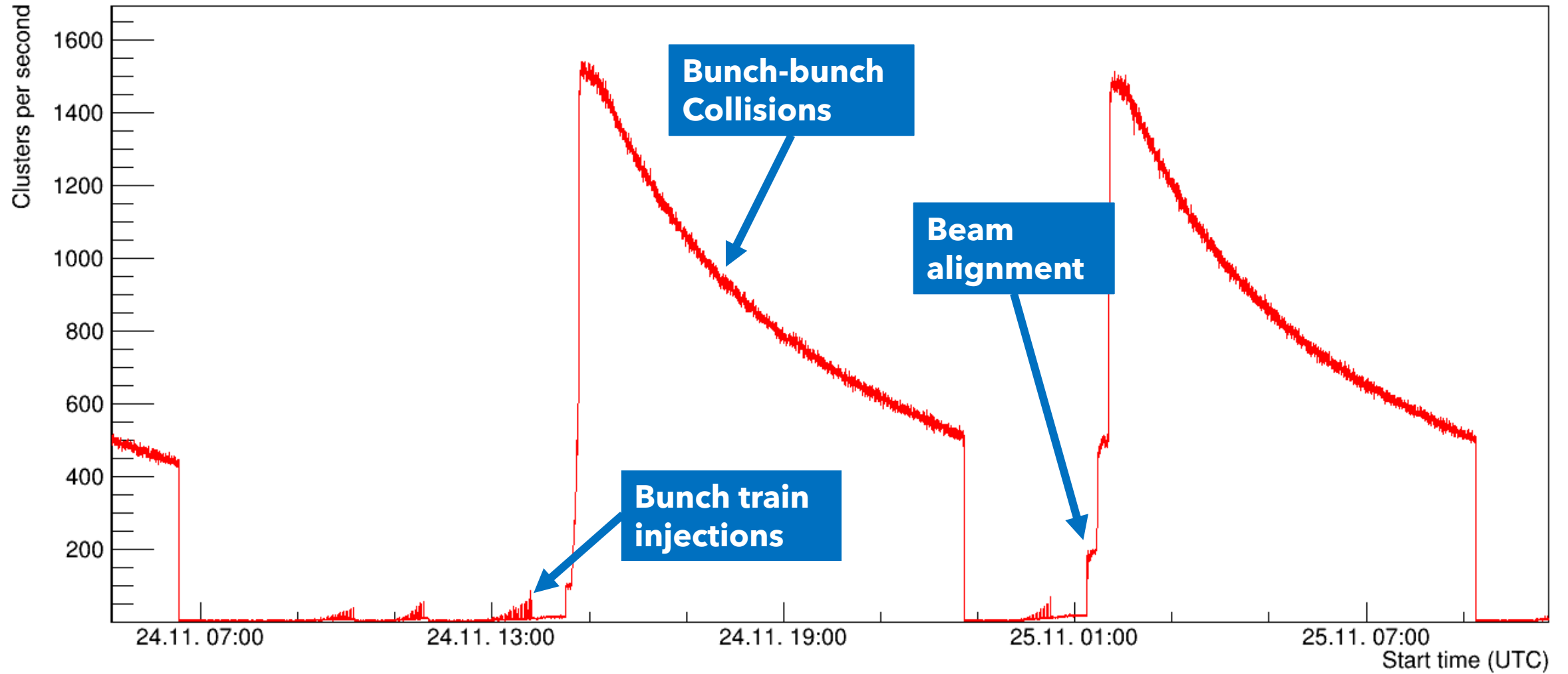
- Tracking and identification of **all** particles
- Online outlier detection to search for exotics (highly ionising events)
- Bunch-by-bunch luminosity measurement
- Search for exotic signatures, e.g., "soup" events requiring timing information

# Radiation Levels and Radiation Field

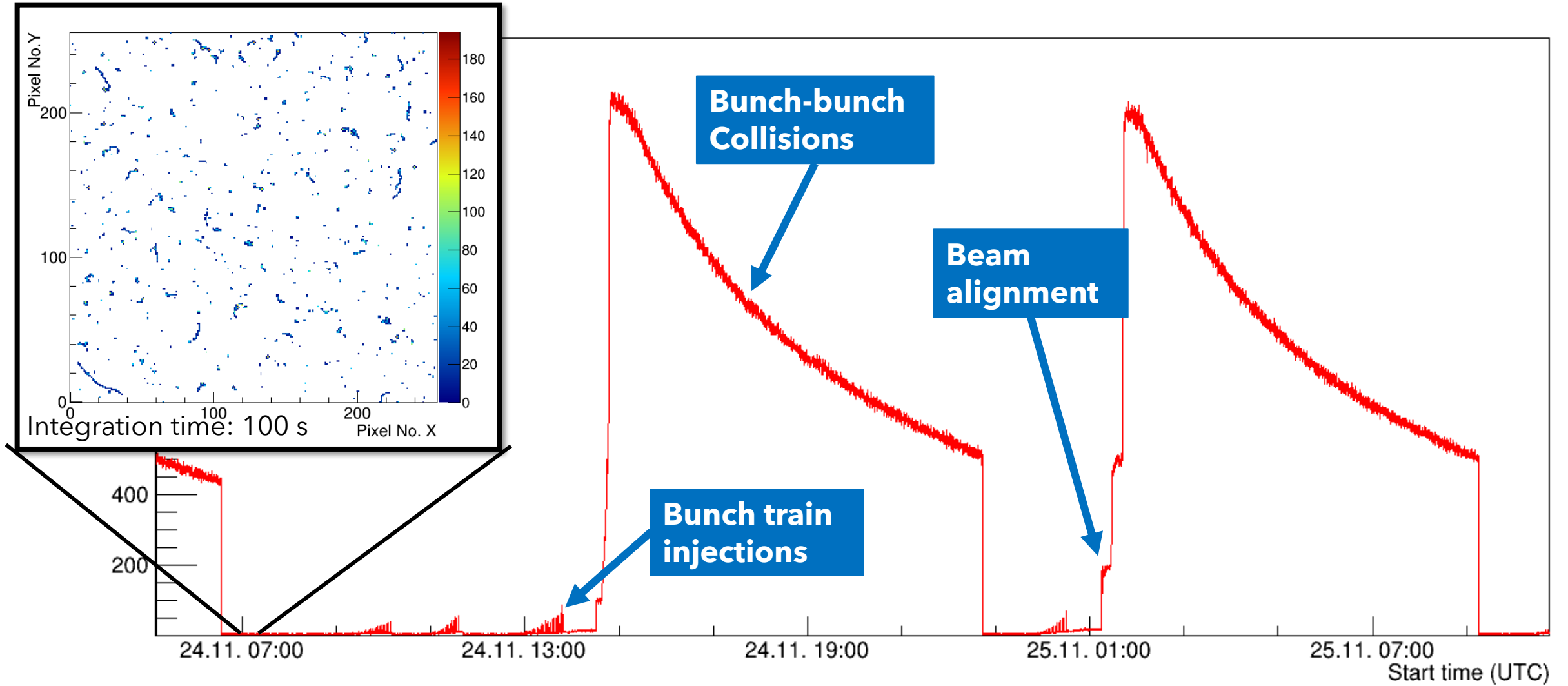




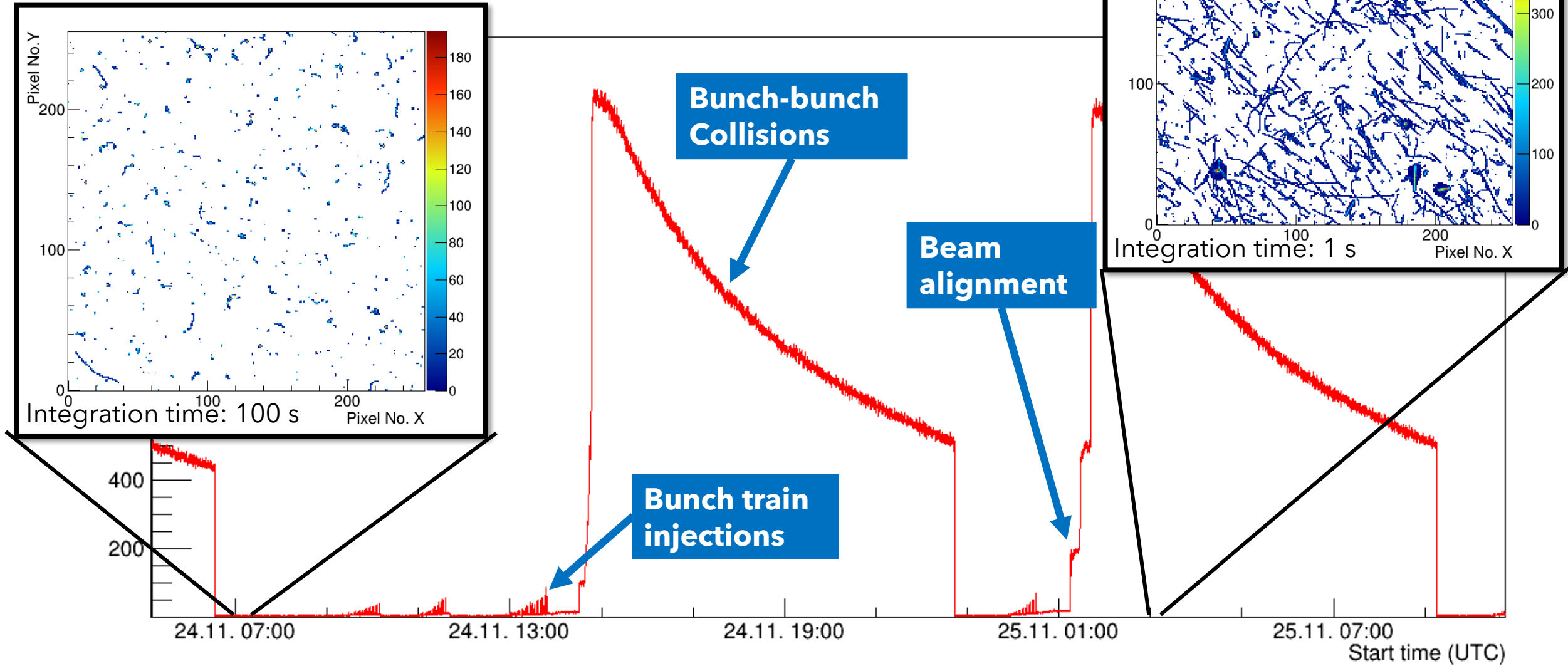
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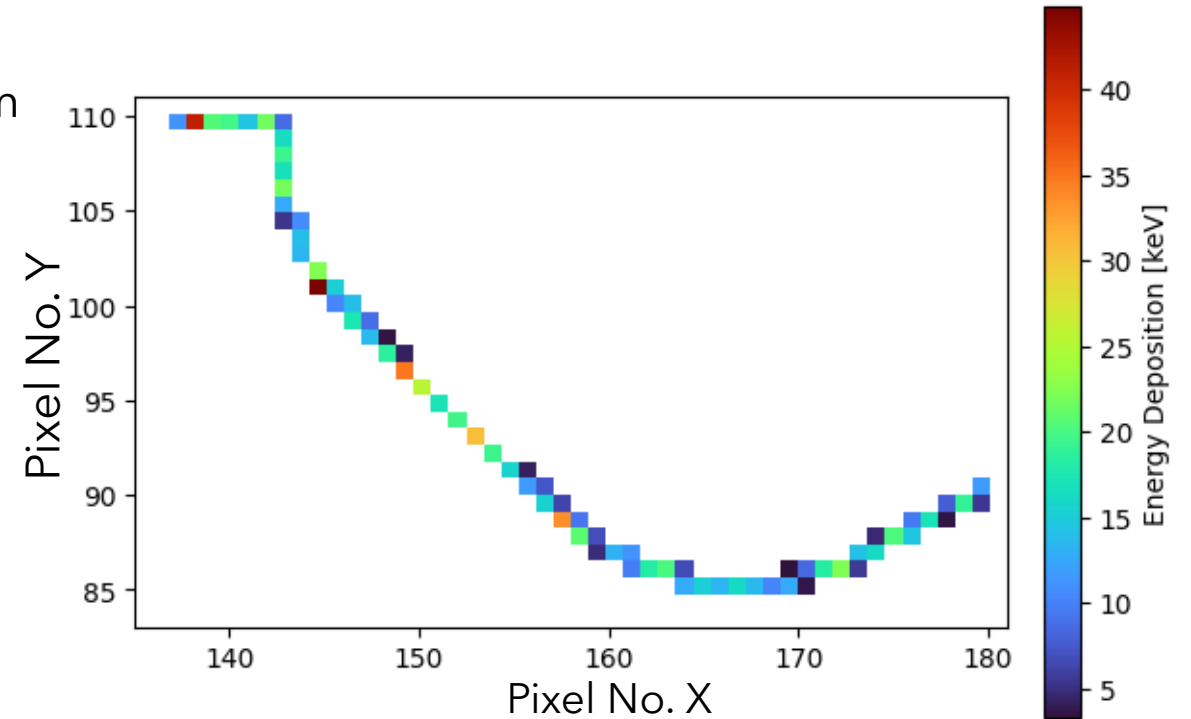
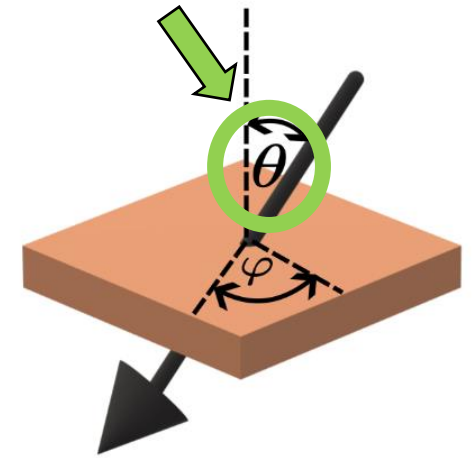
# Radiation Levels and Radiation Field



# Theta Predictor Model

- To reduce the complexity of the problem, only the following two incidence angles are predicted: Azimuthal and perpendicular. Reducing it to regression problem
- **Datasets used for model development:**
  - ⇒ 0.01-10 MeV electrons (curly tracks)
  - ⇒ 10-500 MeV protons (thick, straight tracks)
  - ⇒ 40 GeV pions (thin, straight tracks)
- It was found through an extensive study that a Random Forrest Regressor with a selection of input features produced optimal results

## Input Features:

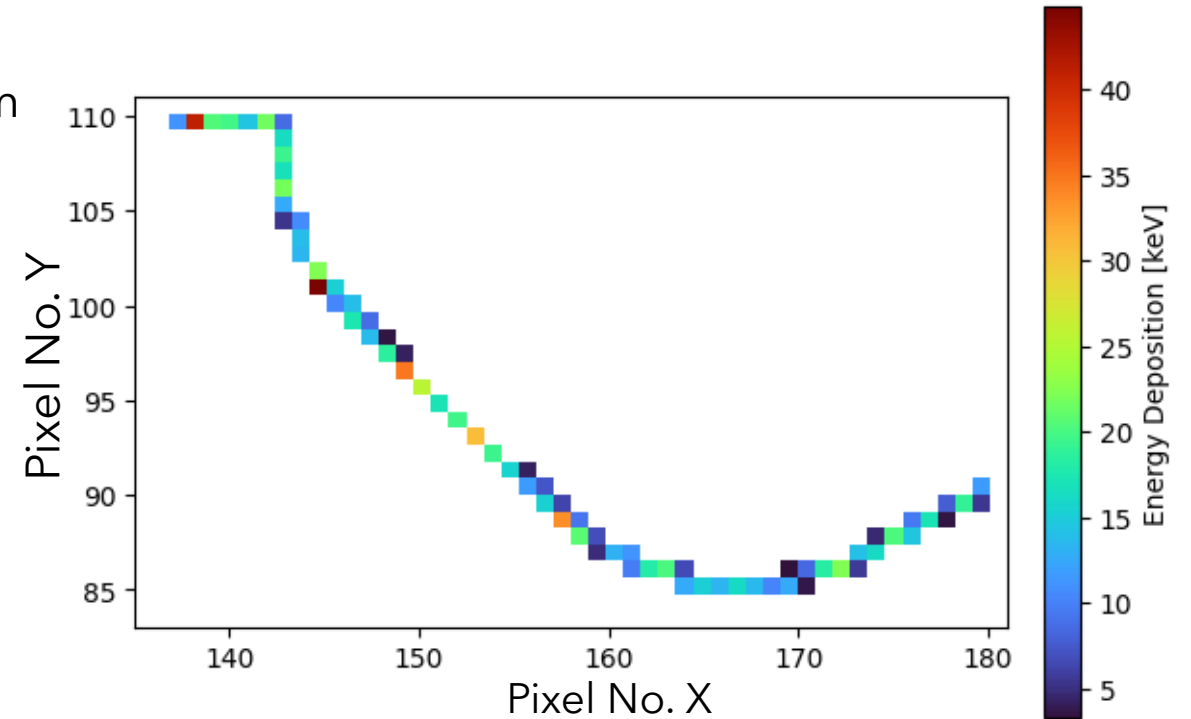
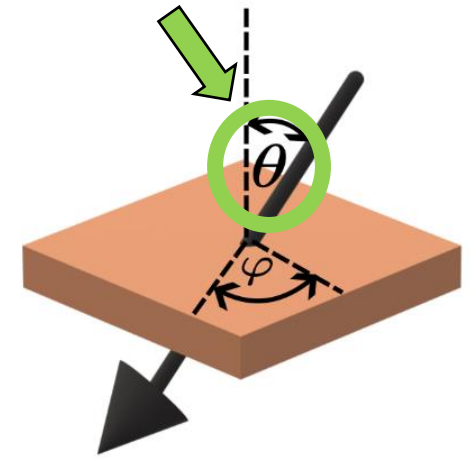


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## Input Features:

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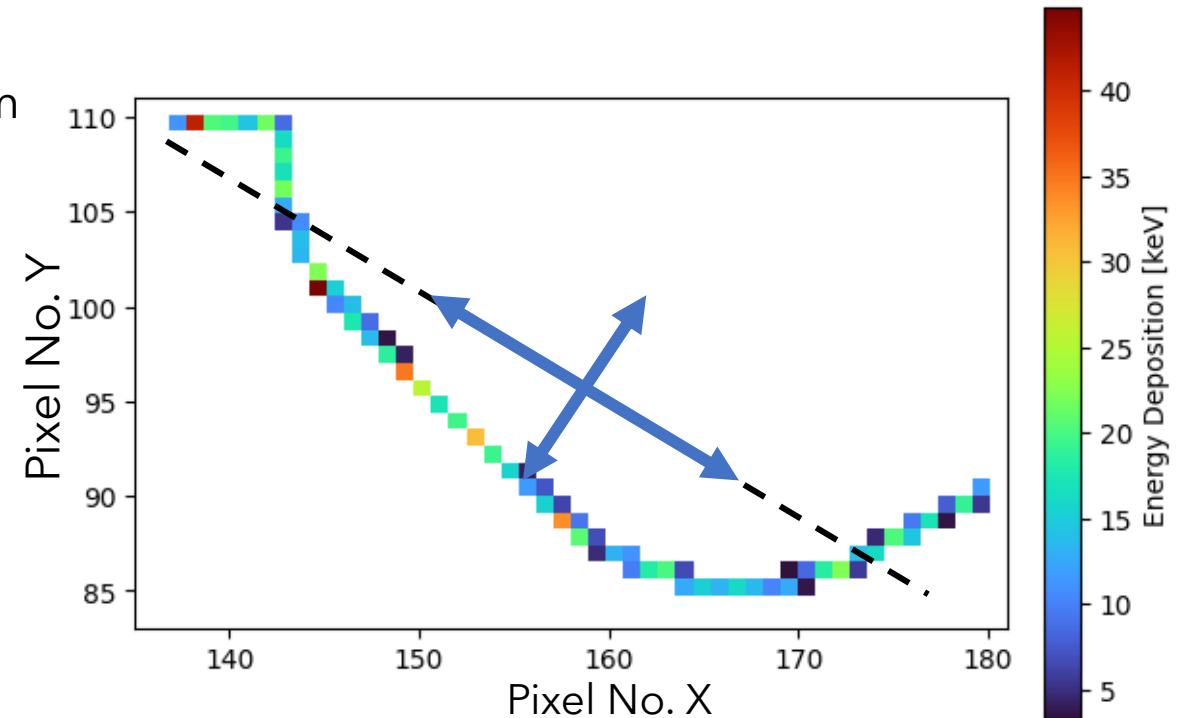
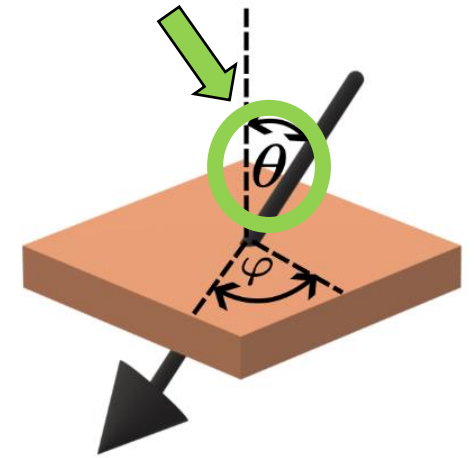


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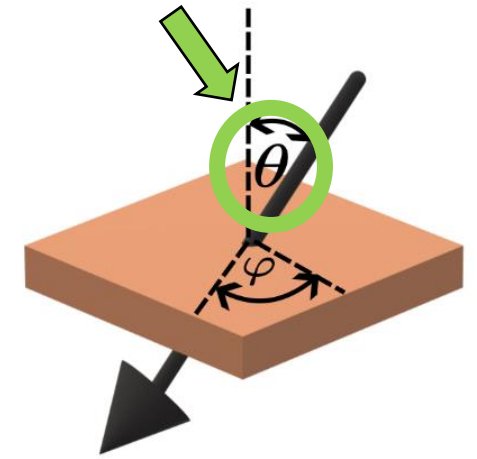
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2. Line standard deviation



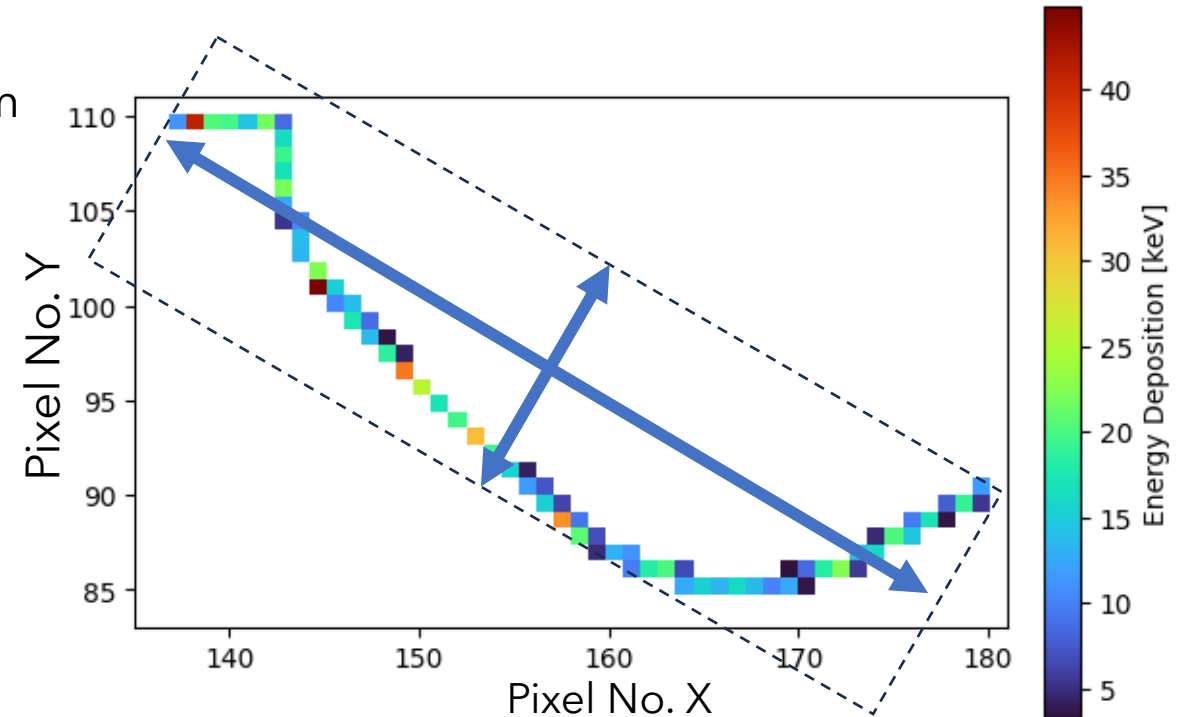
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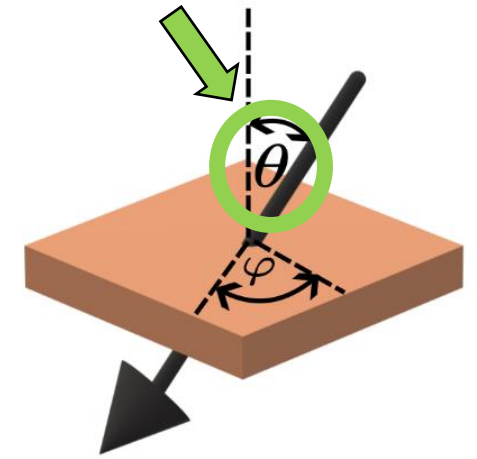
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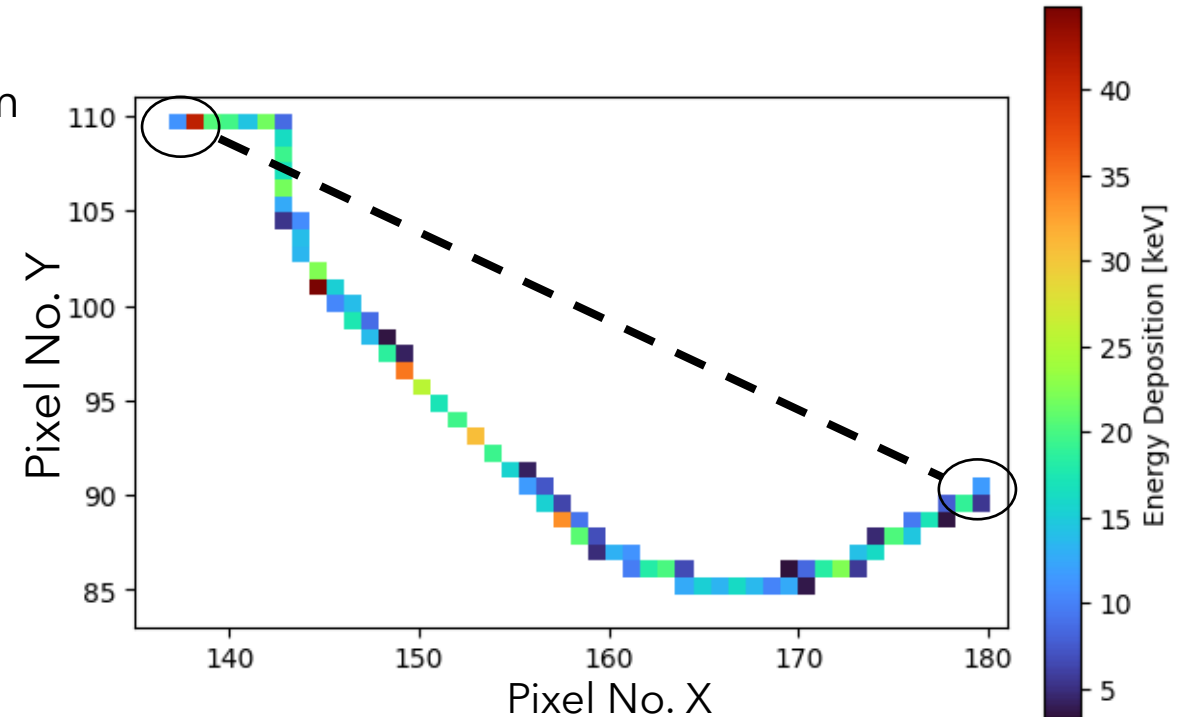
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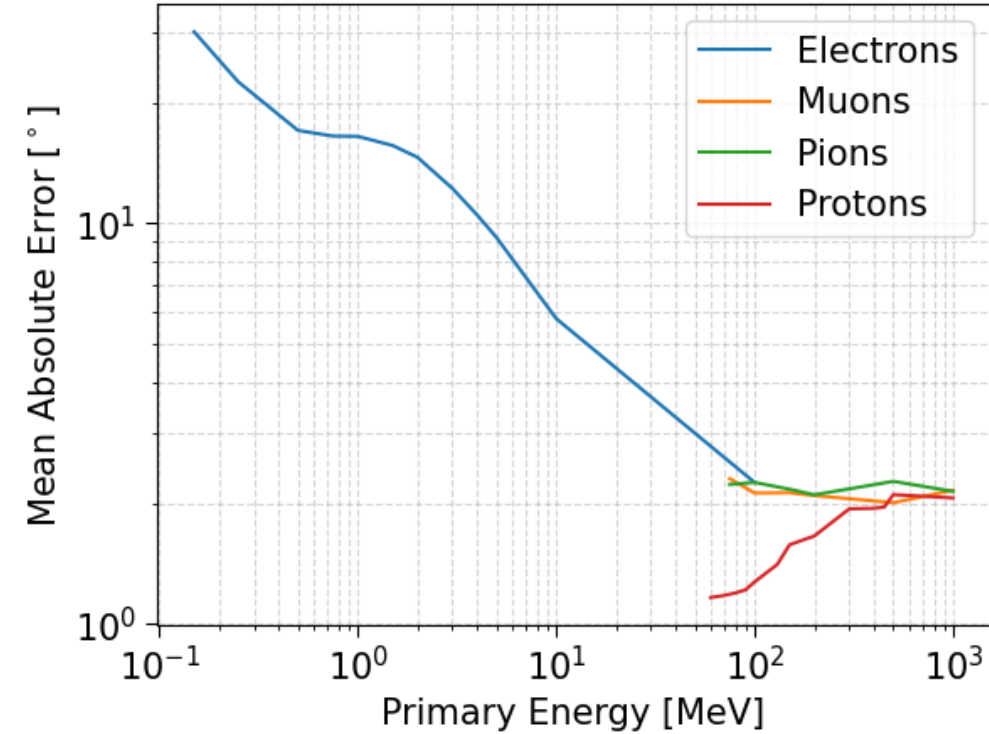
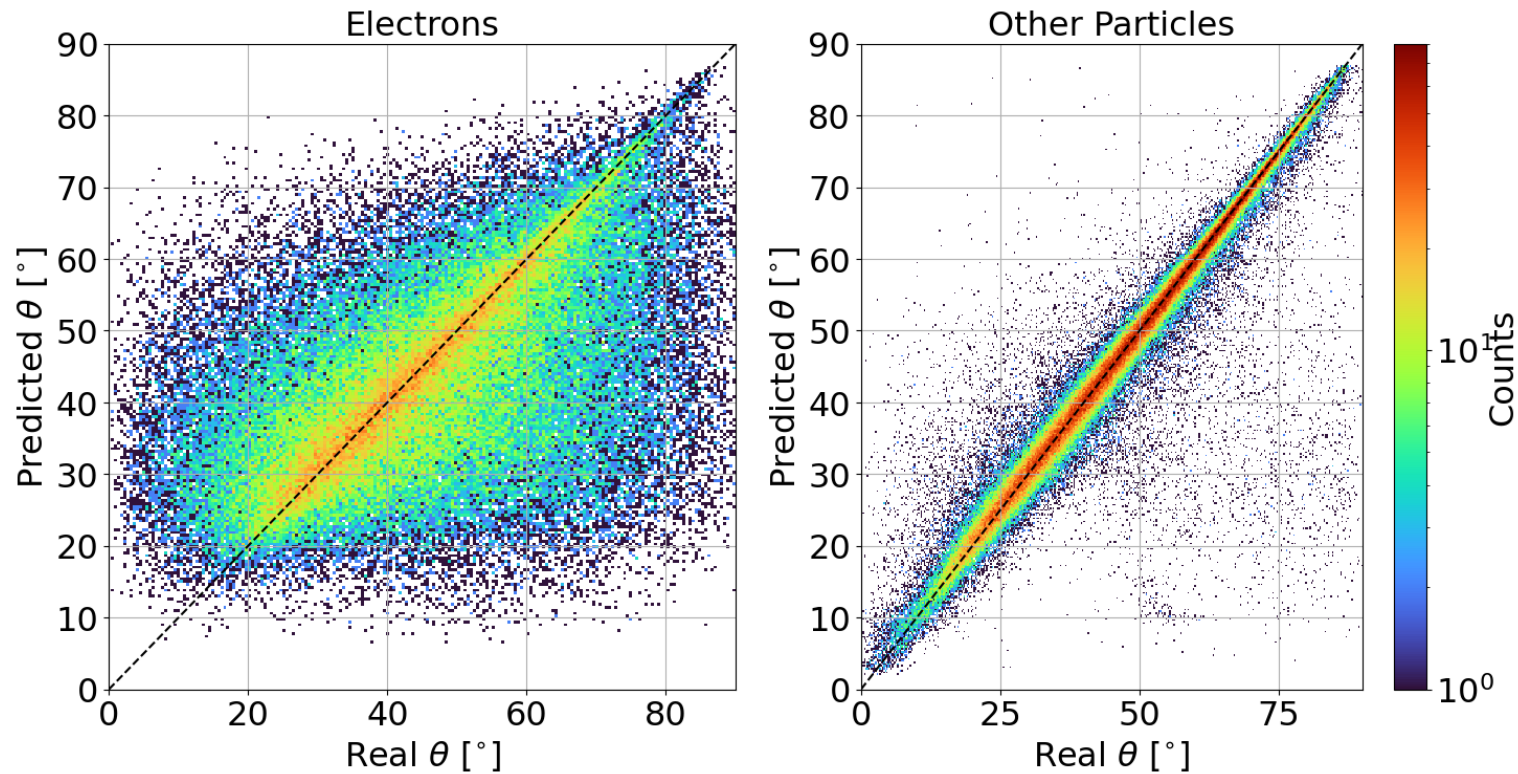
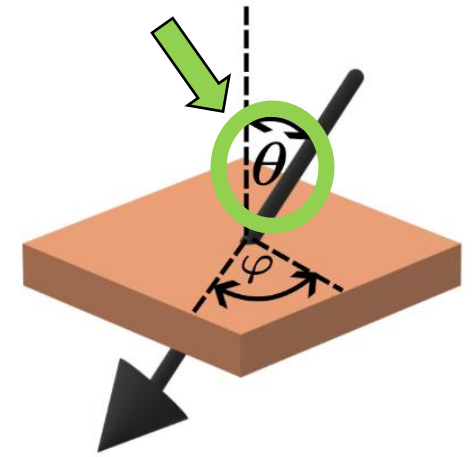
1. Size [no. of pixels]
2. Line standard deviation
3. Box dimensions
4. End point distance





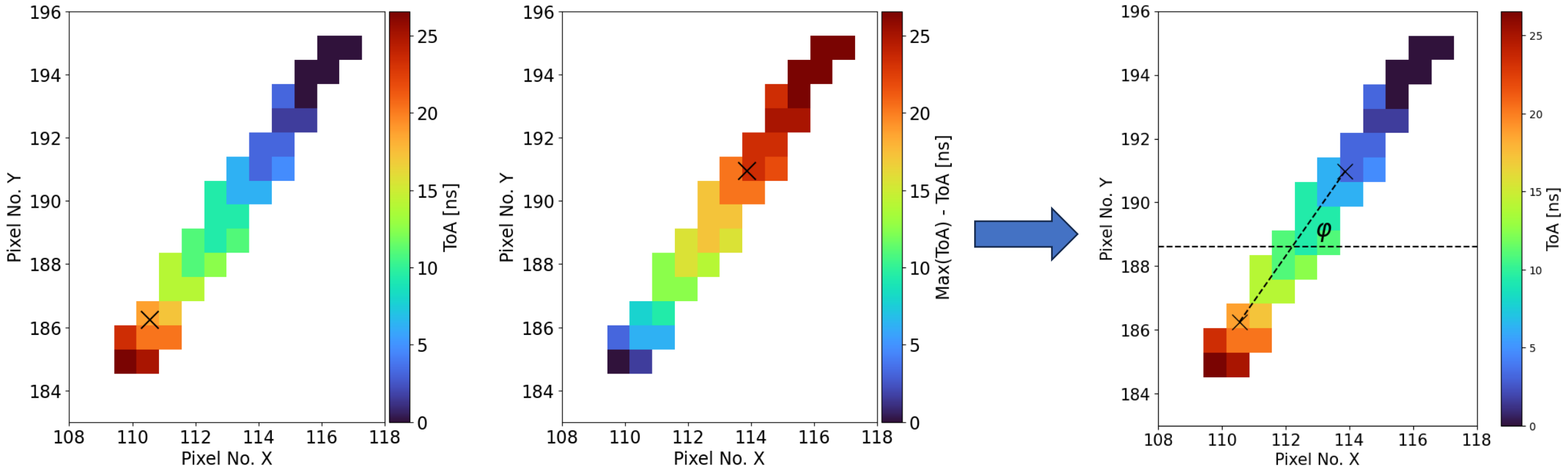
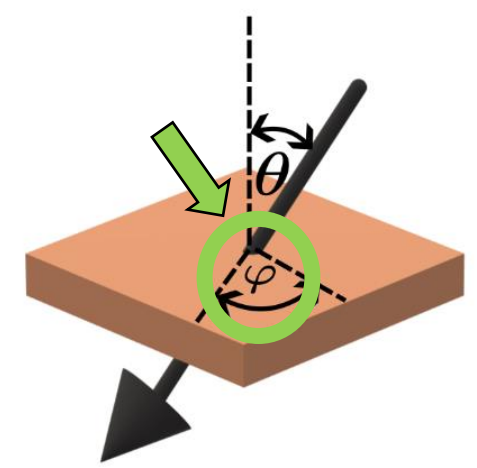
# Theta Prediction Results

For testing separate simulation datasets were created



# Phi Determination

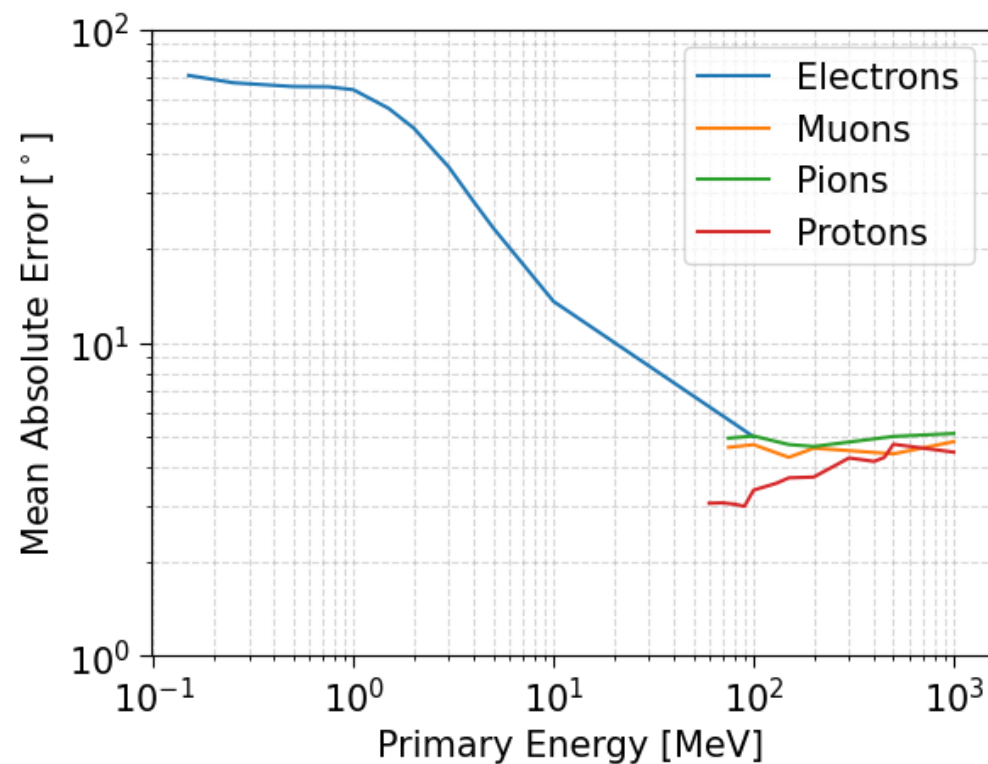
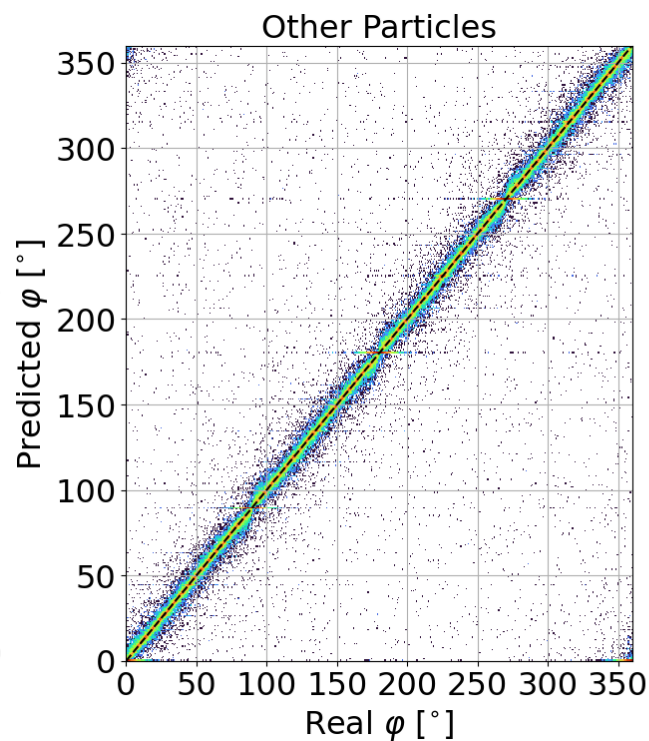
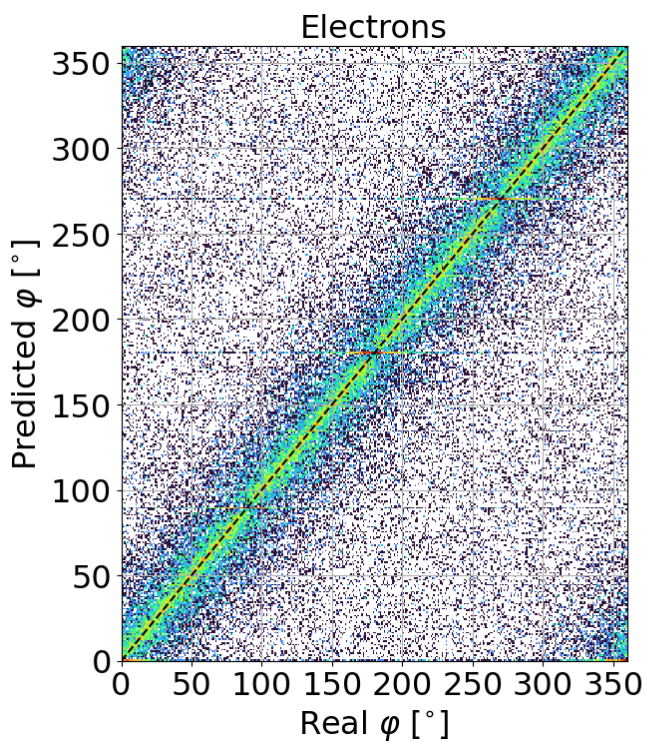
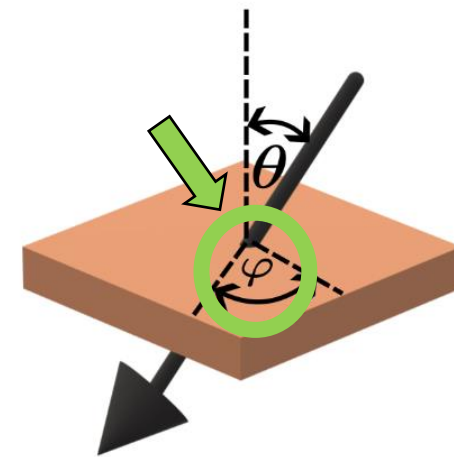
- An analytic approach was used as no improvement using machine learning algorithms was achieved
- Following an extensive study, Time of Arrival [ToA] weighted averaging was found to produce the most accurate results\*
- In this method, we utilise the effect that drift time within the sensor has on ToA to approximate the "mean" entry and exit points



\*P. Mánek et al. In: Journal of Instrumentation (2022), p. C01062. <https://dx.doi.org/10.1088/1748-0221/17/01/C01062> <sup>LO</sup>

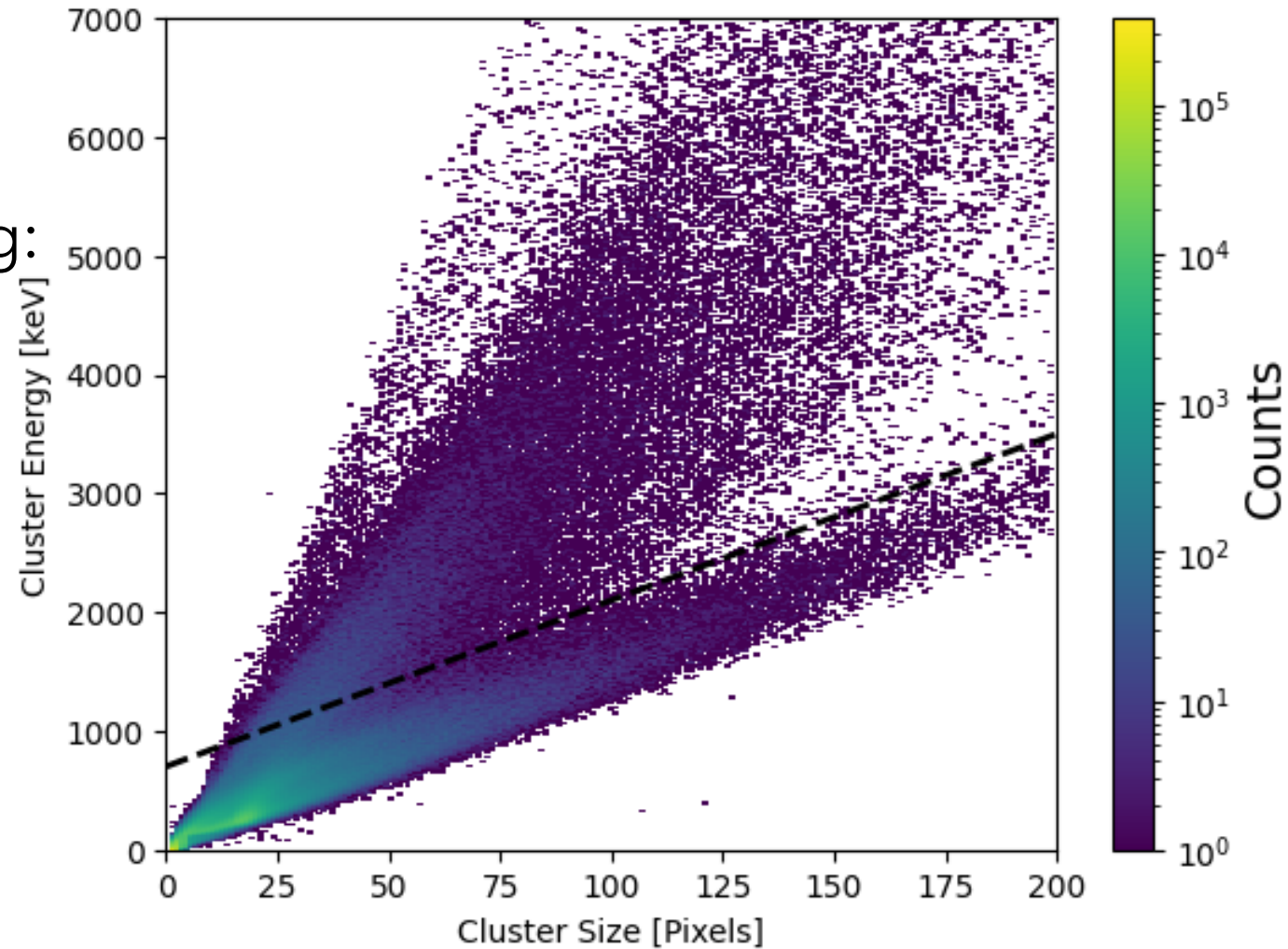
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# Rudimentary Classification

Particle tracks can be divided into two groups, with variable linearity and scattering:

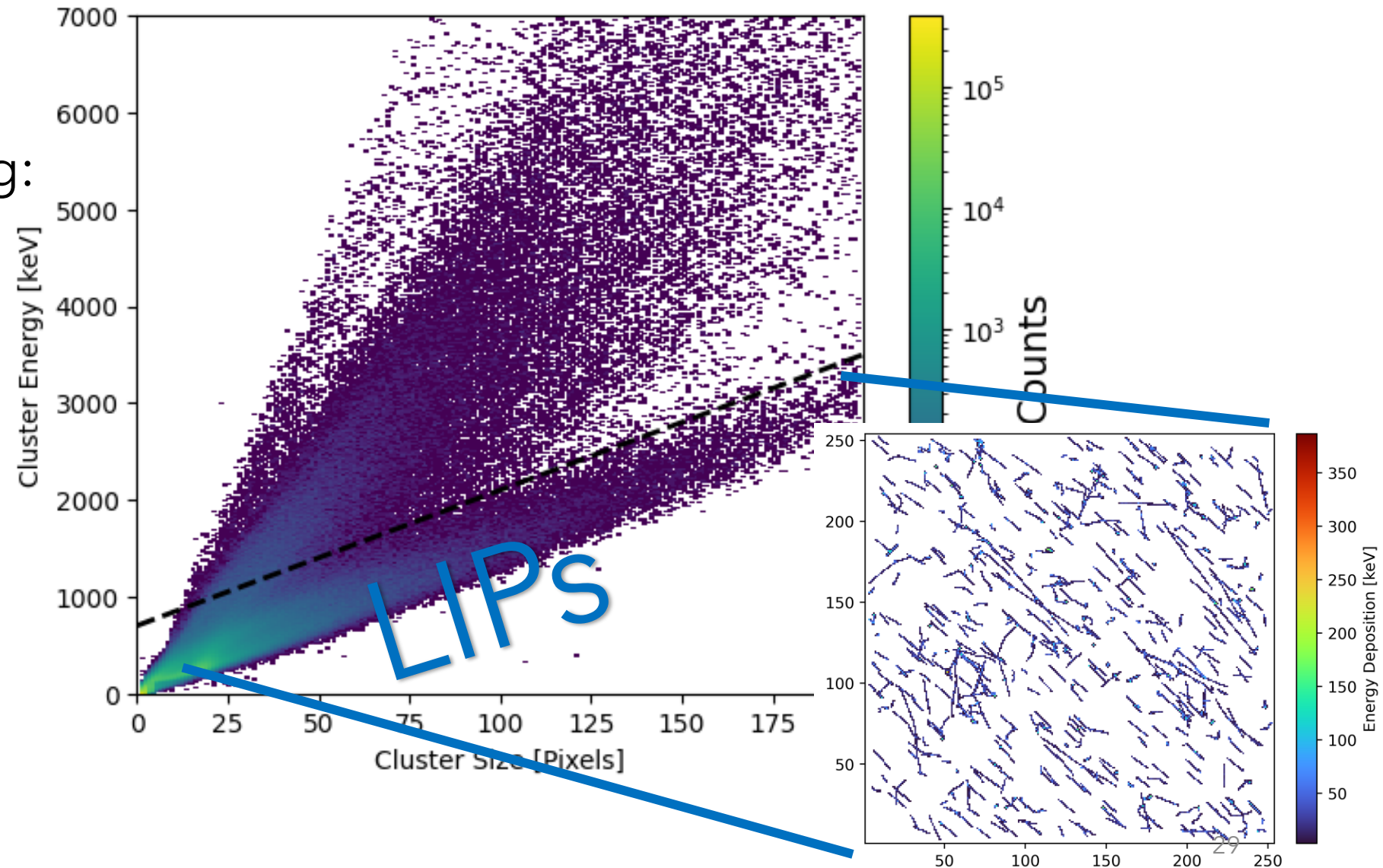


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## 1. Lowly Ionising Particles (LIPs):

- ➔ Electrons
- ➔ Photons
- ➔ Relativistic singly charged particles



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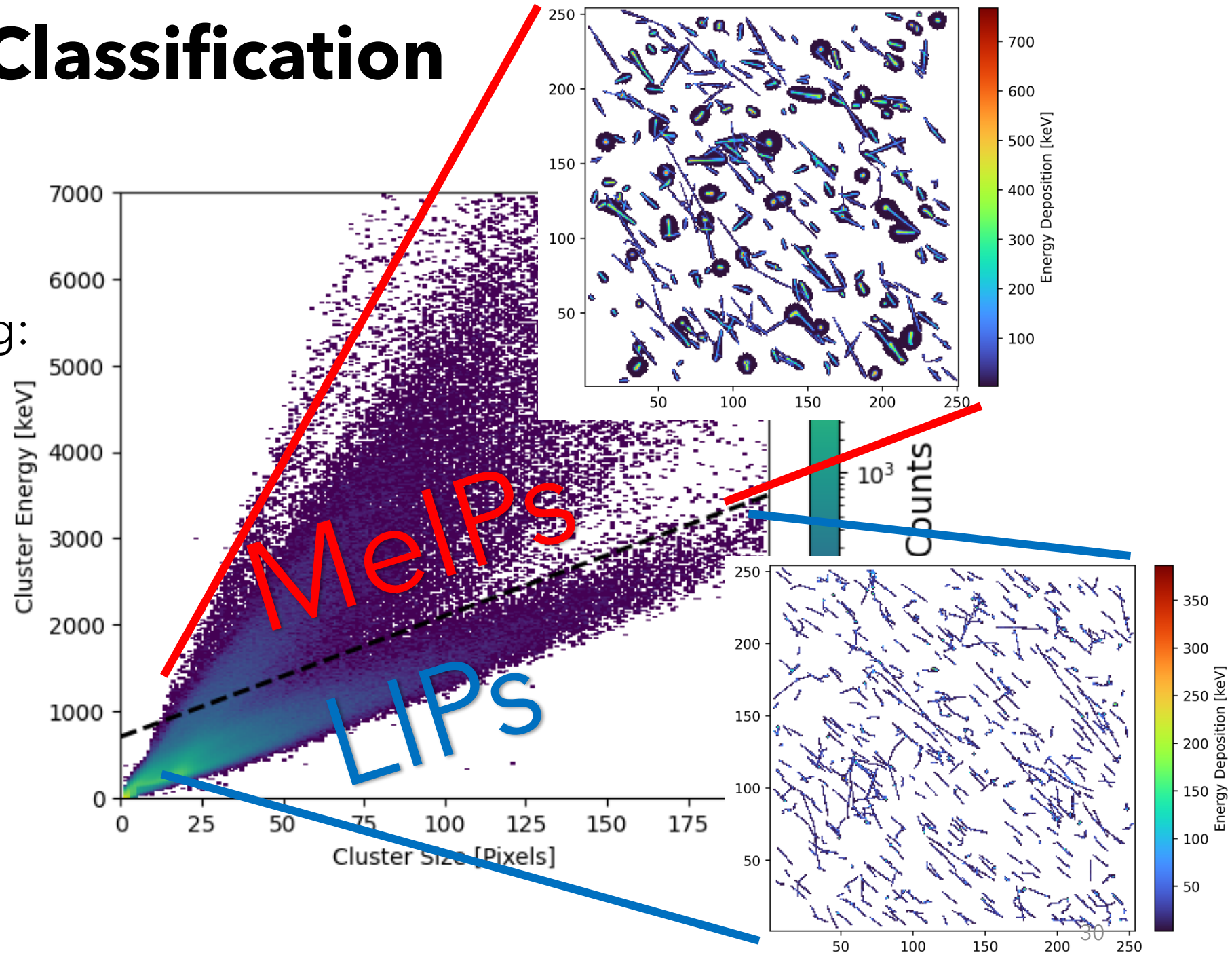
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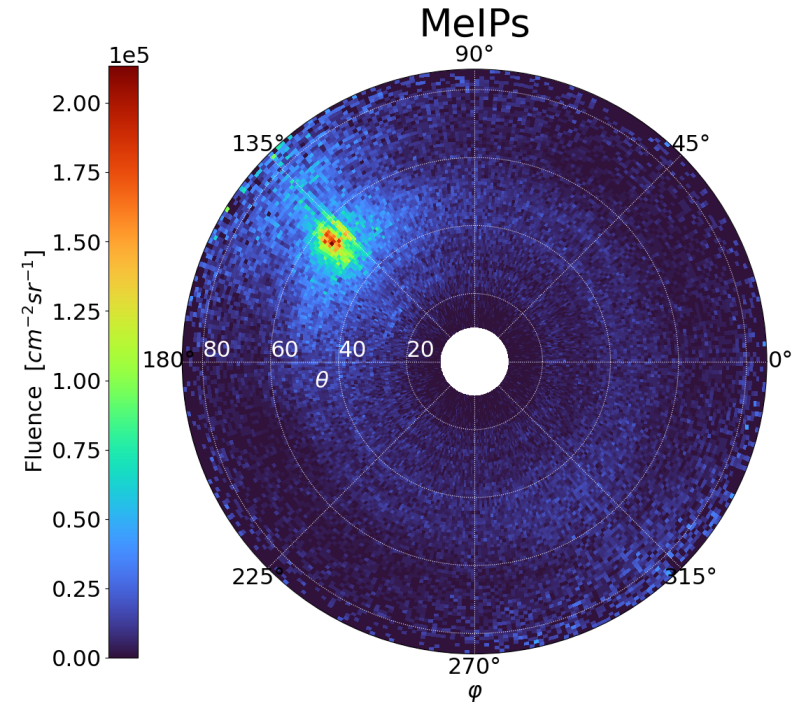
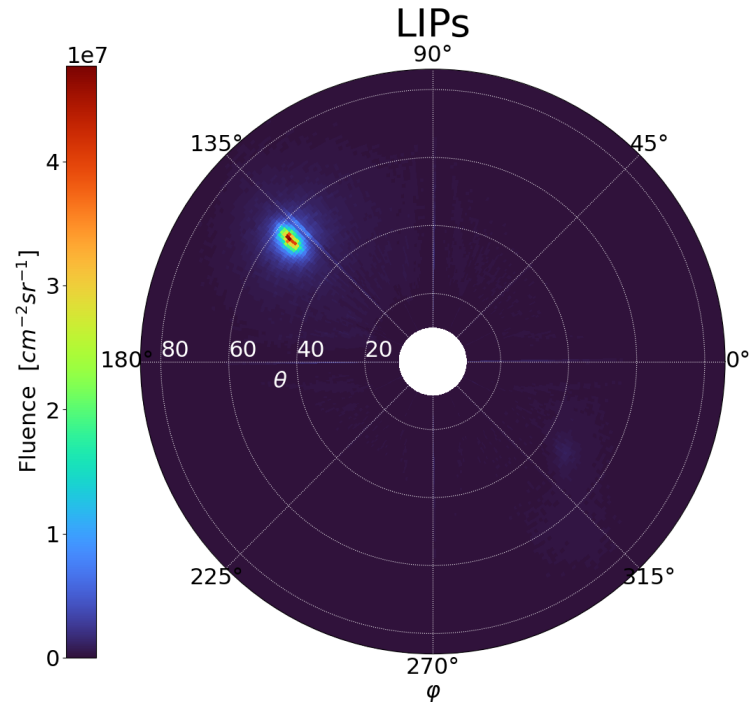
## 2. Medium Ionising Particles (MeIPs):

- Protons < 500 MeV
- Pions < 200 MeV
- Muons < 300 MeV
- $Z > 1$  particles

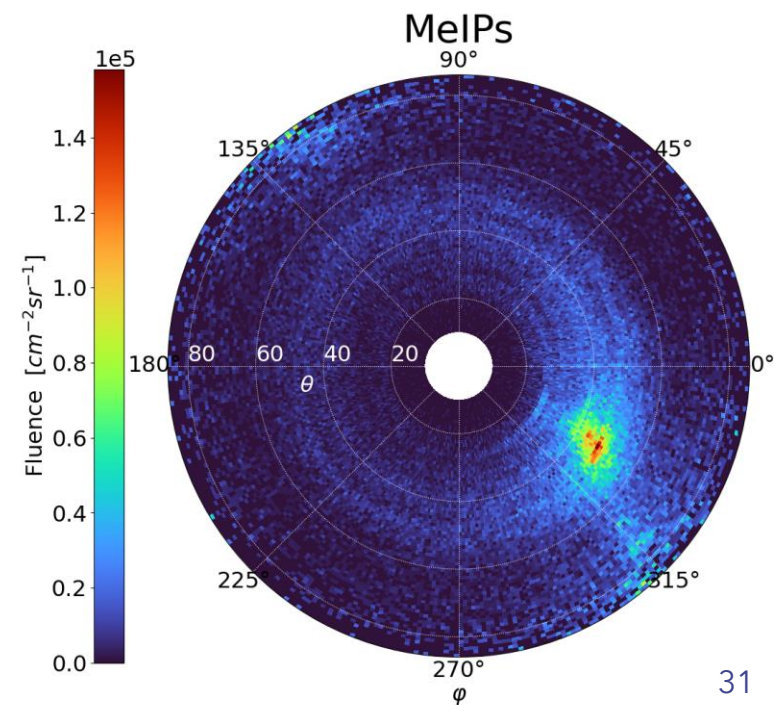
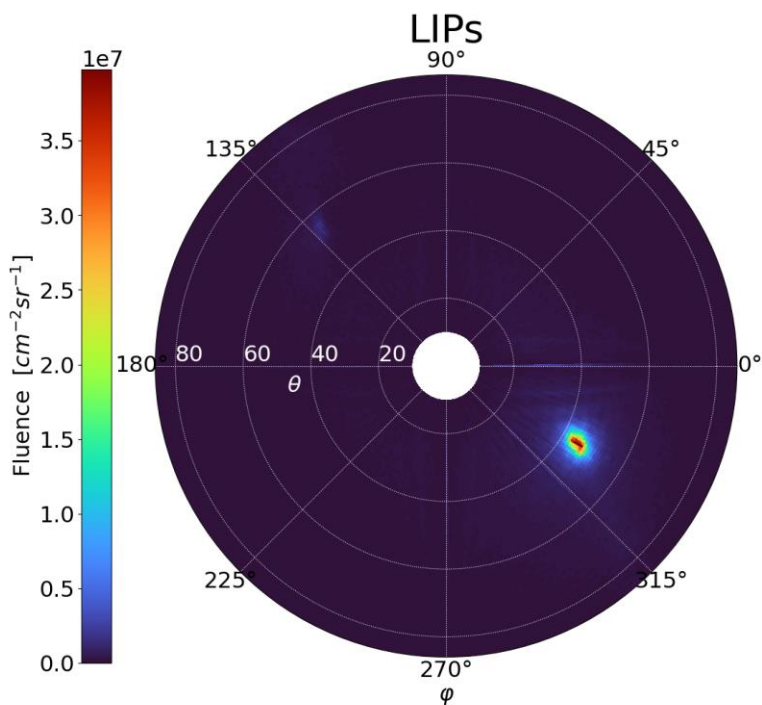


# Field Directionality During Proton- Proton Collisions

**E5 Detector**

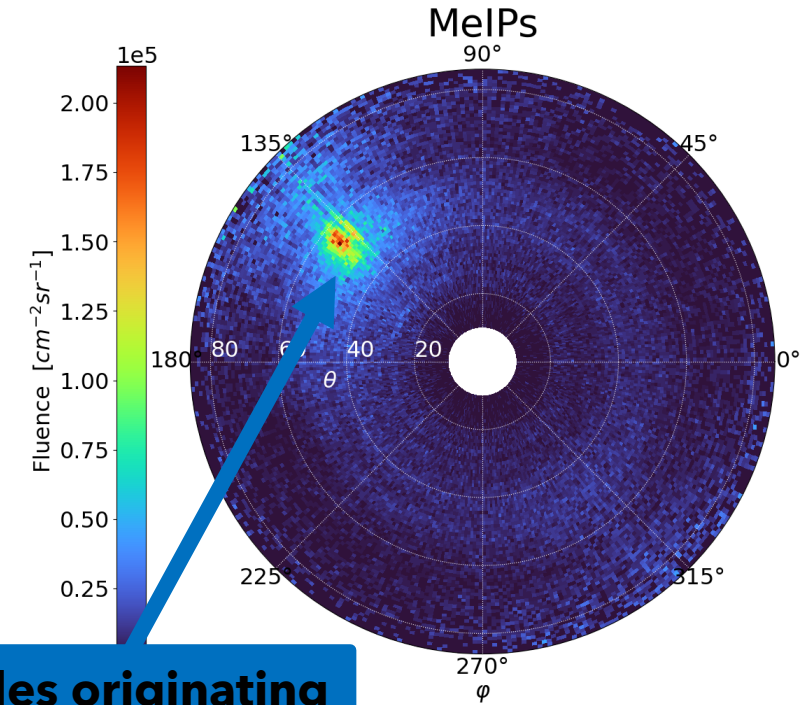
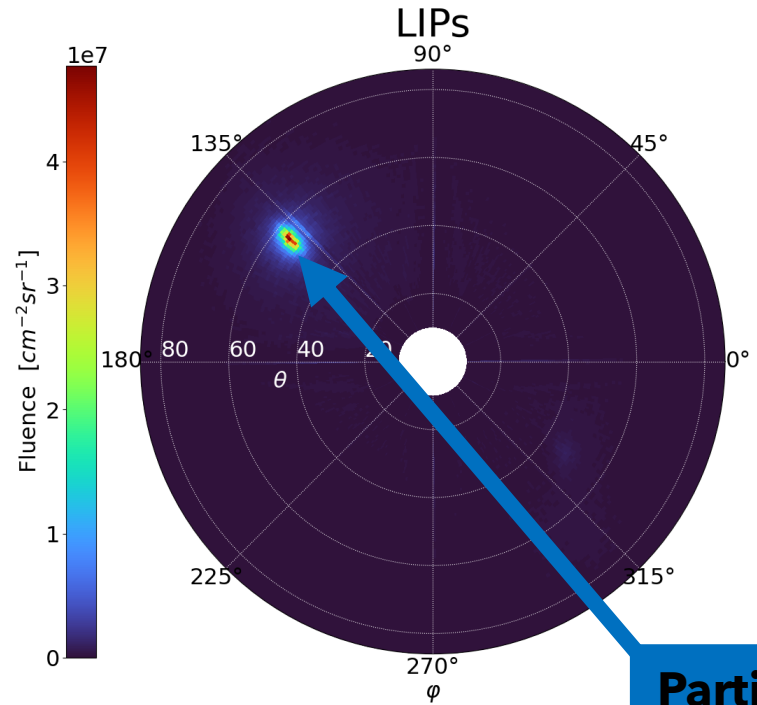


**G4 Detector**



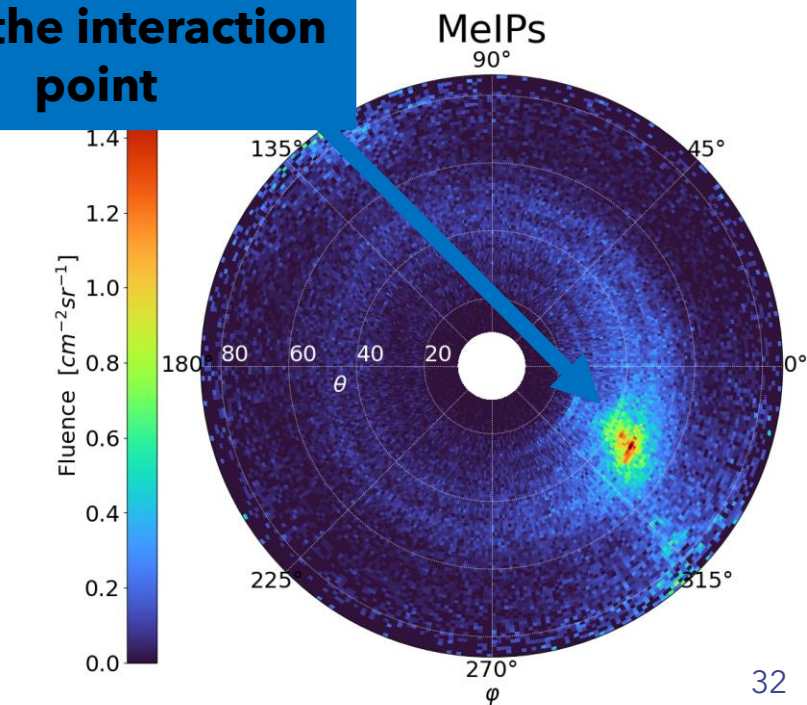
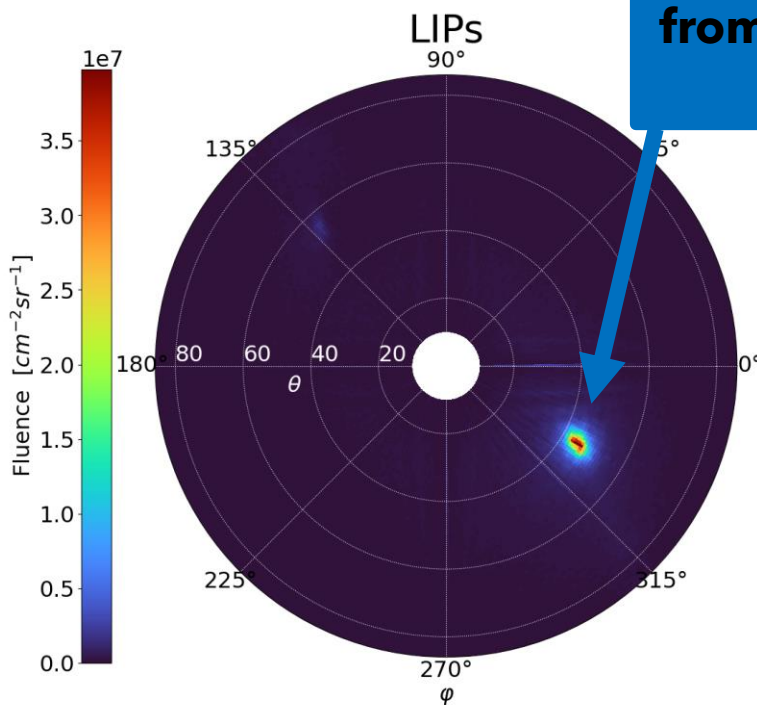
# Field Directionality During Proton- Proton Collisions

**E5 Detector**



**Particles originating  
from the interaction  
point**

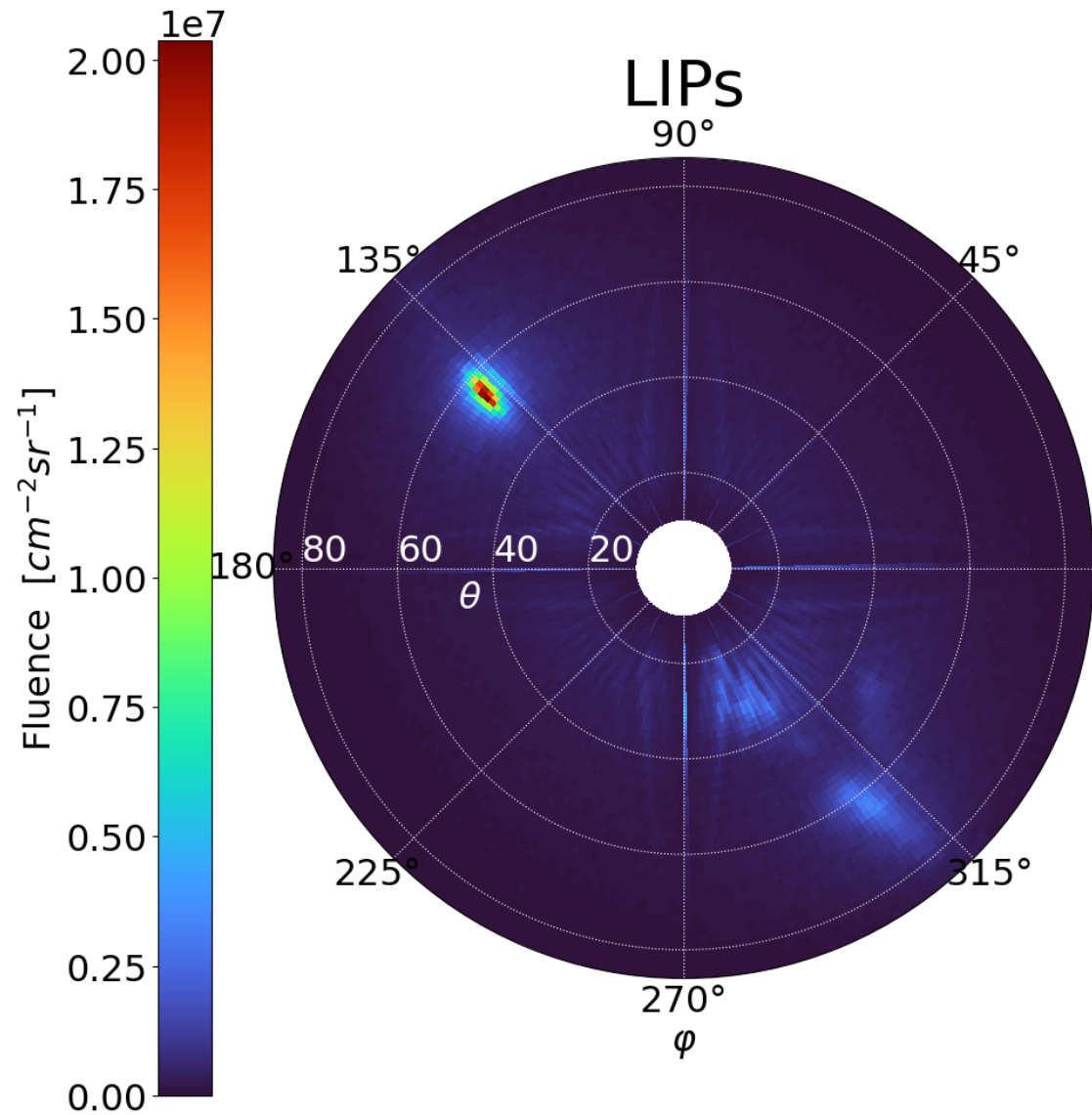
**G4 Detector**



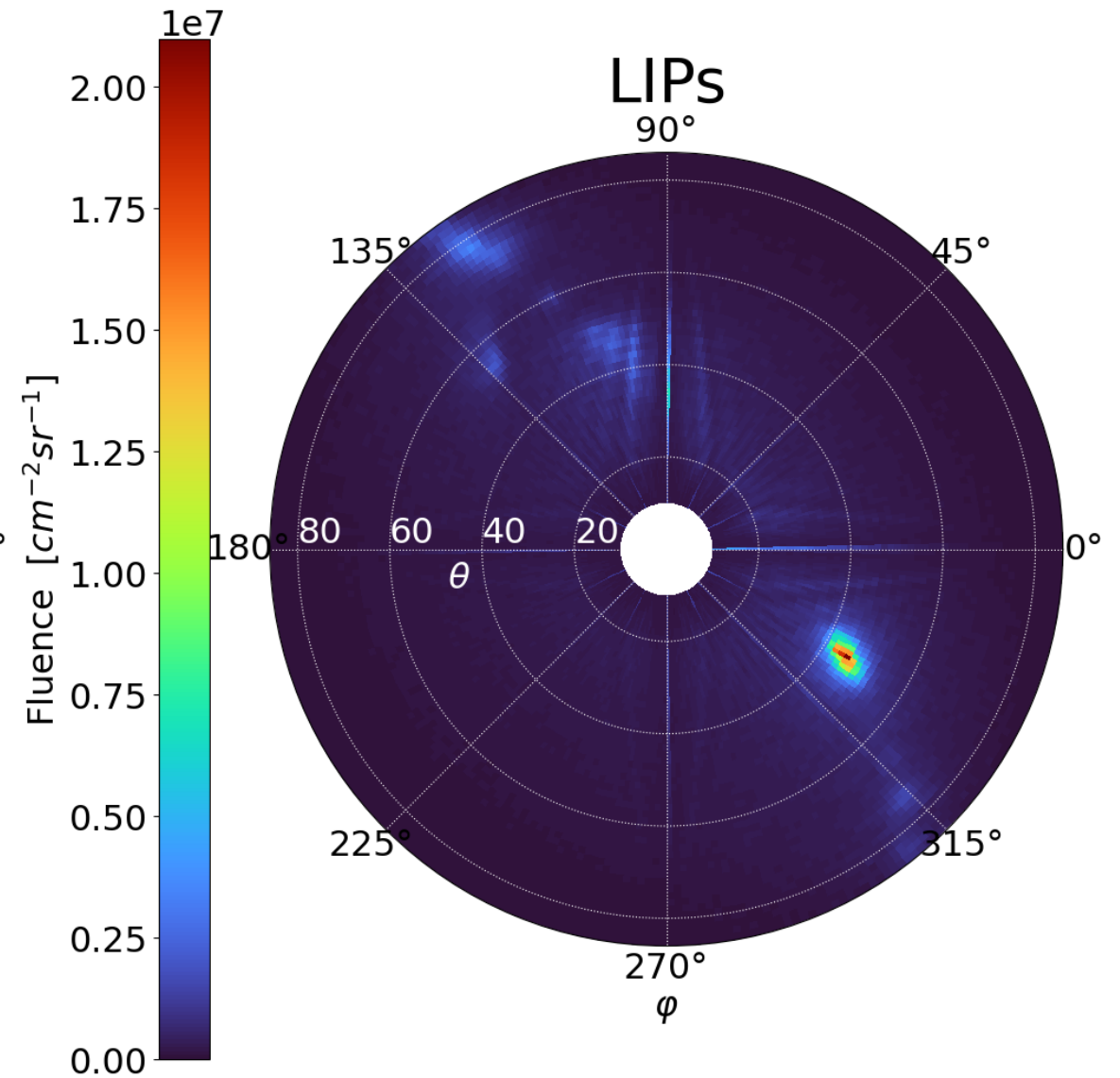


# Field Directionality During Pb-Pb Collisions

## E5 Detector

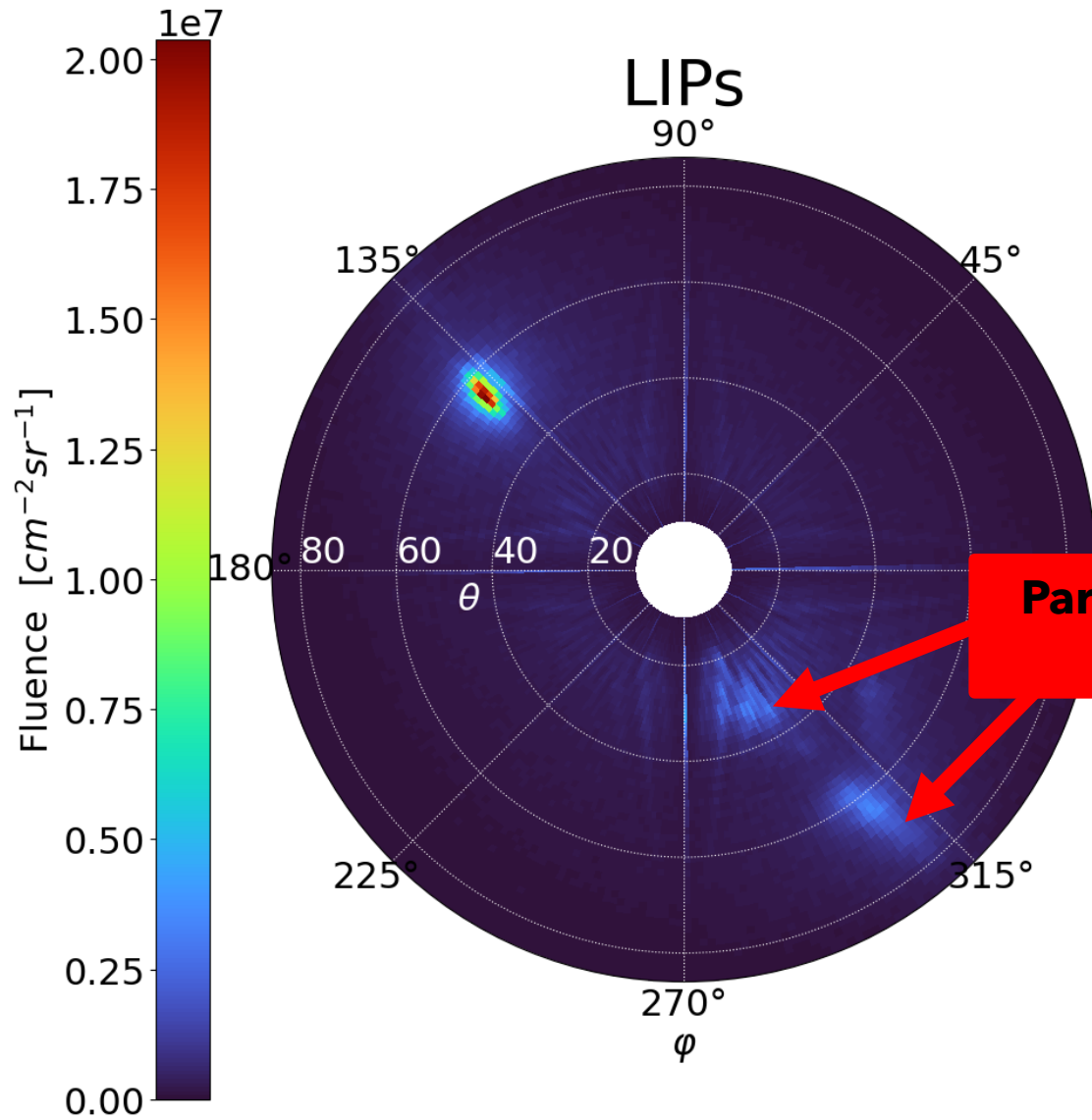


## G4 Detector

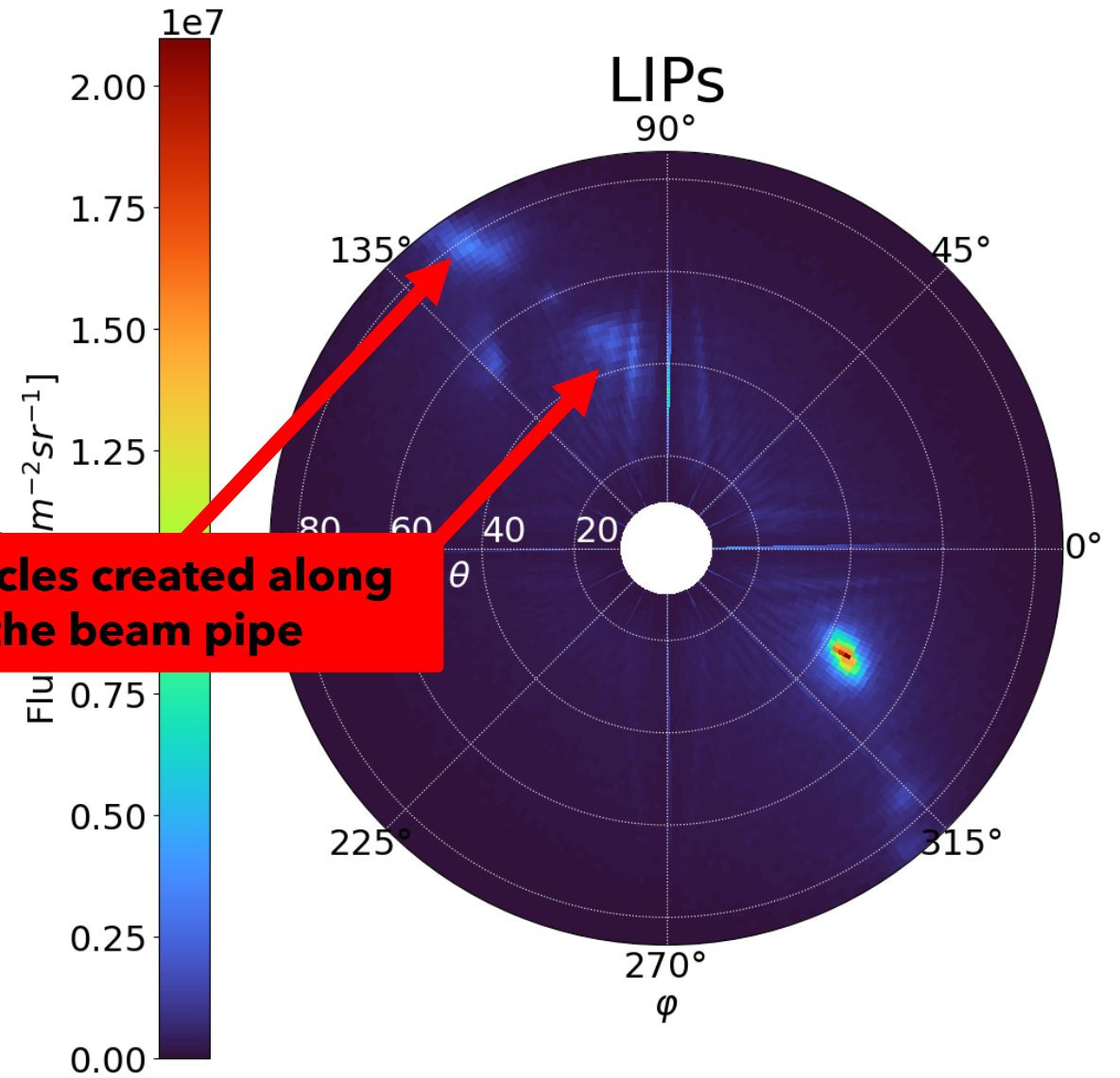


# Field Directionality During Pb-Pb Collisions

## E5 Detector



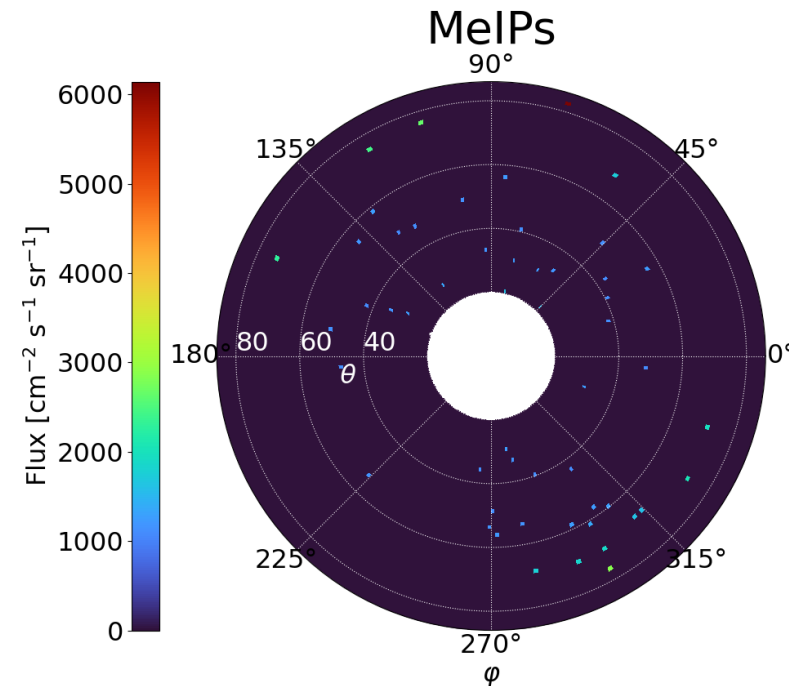
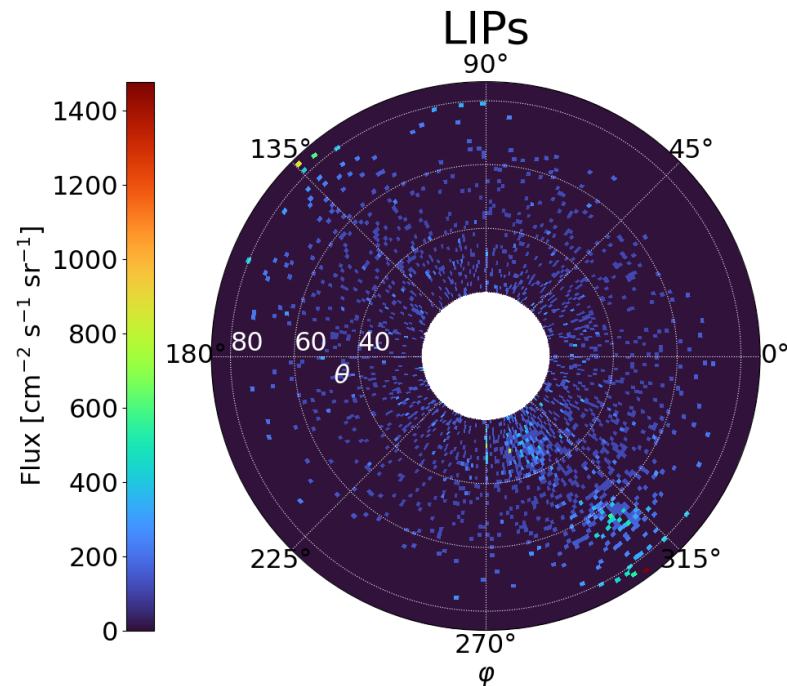
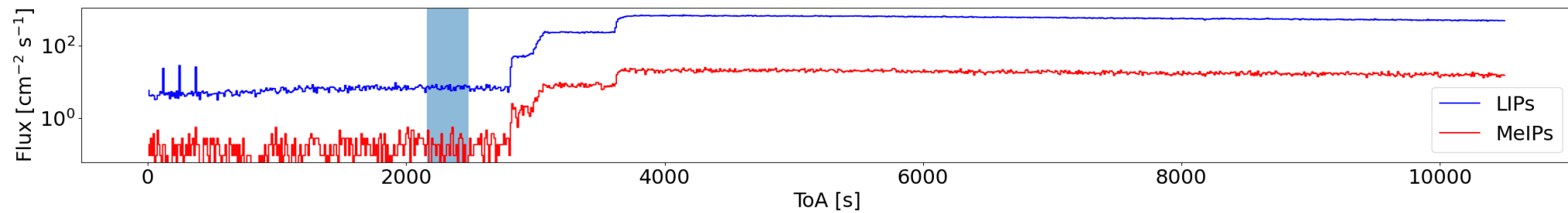
## G4 Detector



**Particles created along  
the beam pipe**

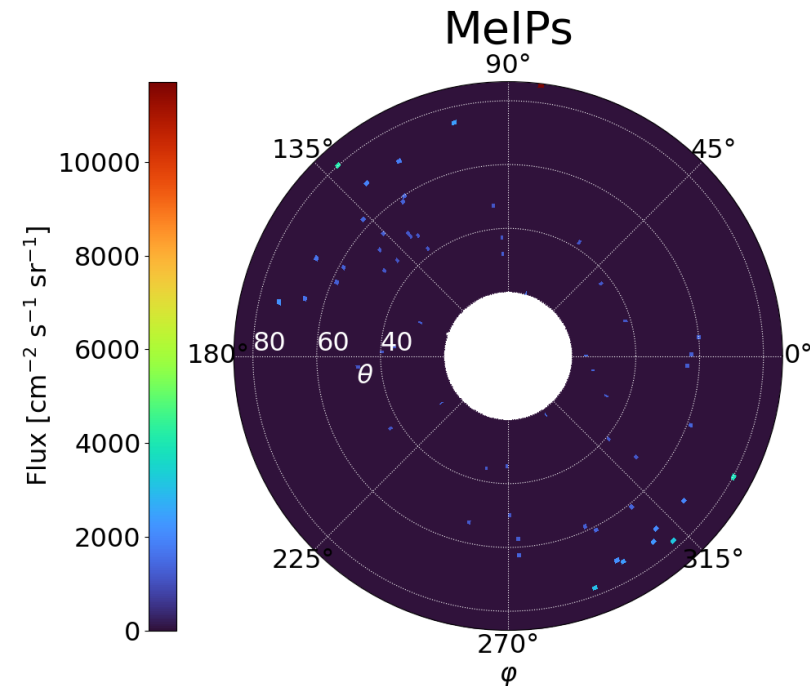
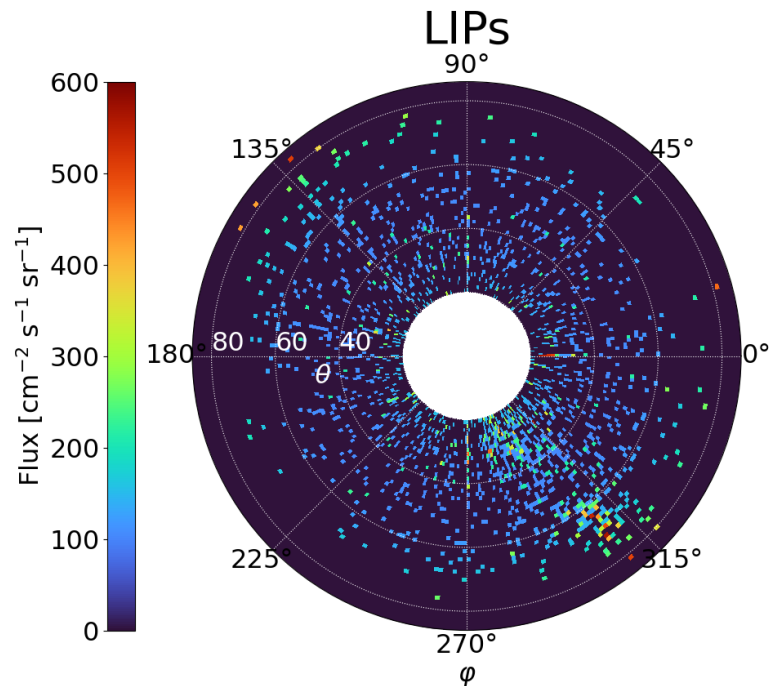
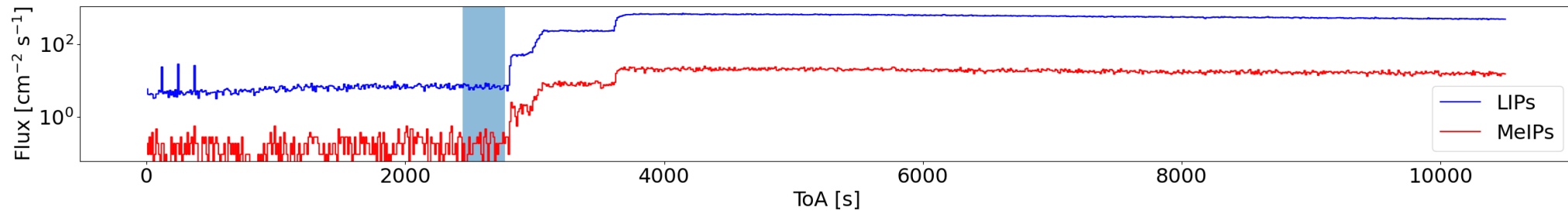
# Time-Resolved Measurement of the Directionality Map Pb-Pb Collision Period

For increased accuracy only 5 minutes of data acquisition is needed to produce reasonable statistics



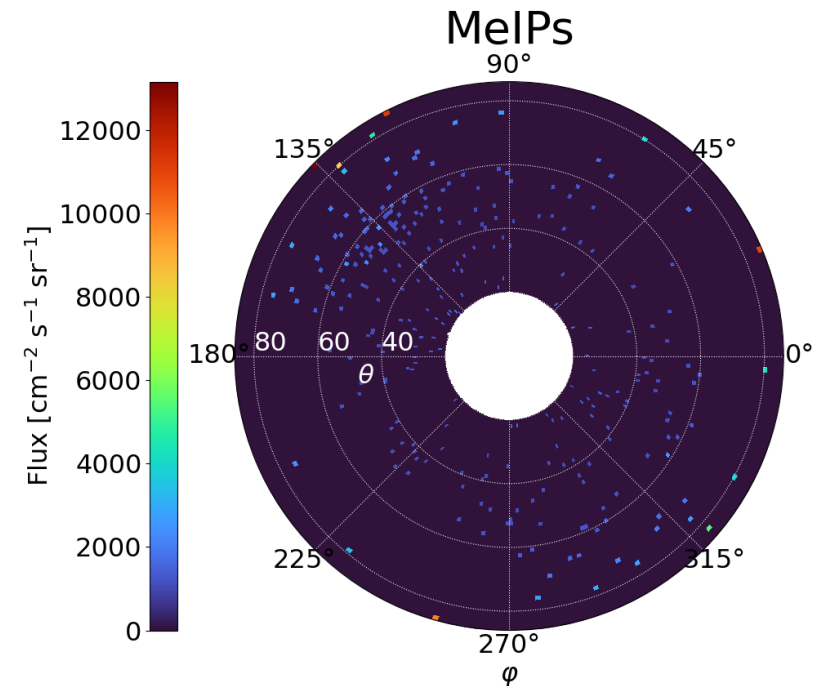
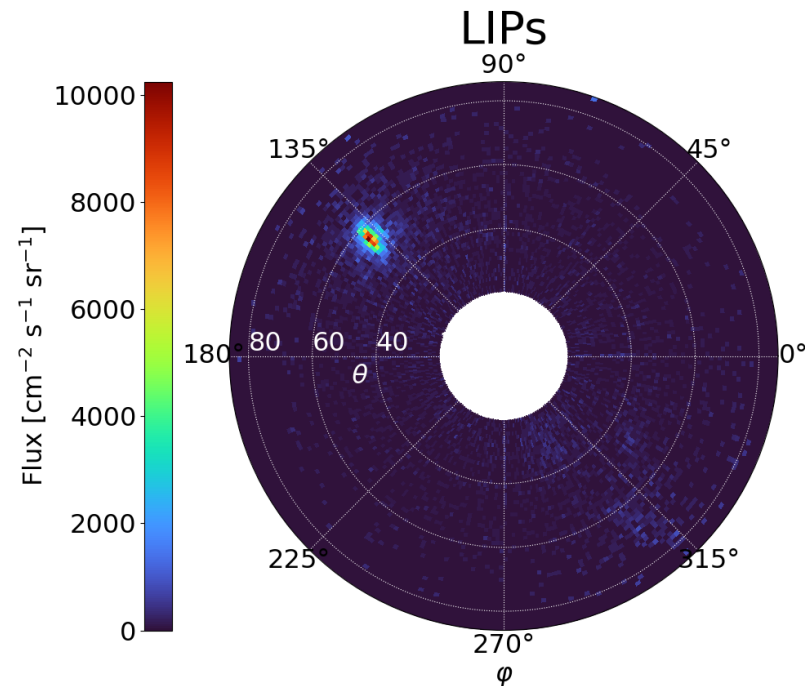
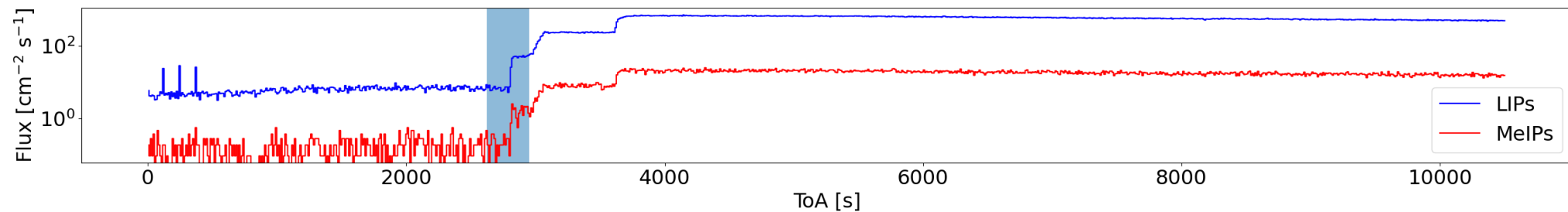
# Time-Resolved Measurement of the Directionality Map Pb-Pb Collision Period

For increased accuracy only 5 minutes of data acquisition is needed to produce reasonable statistics



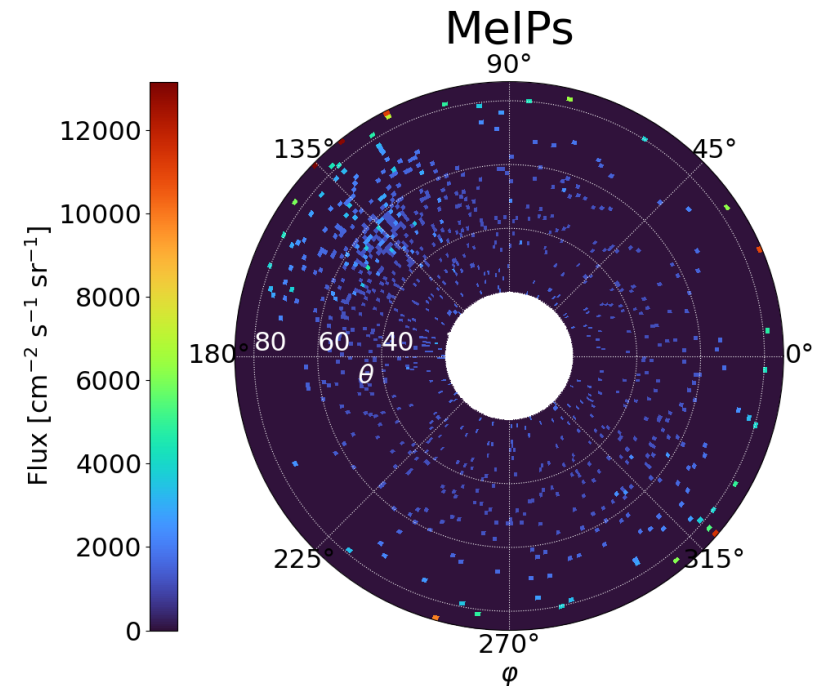
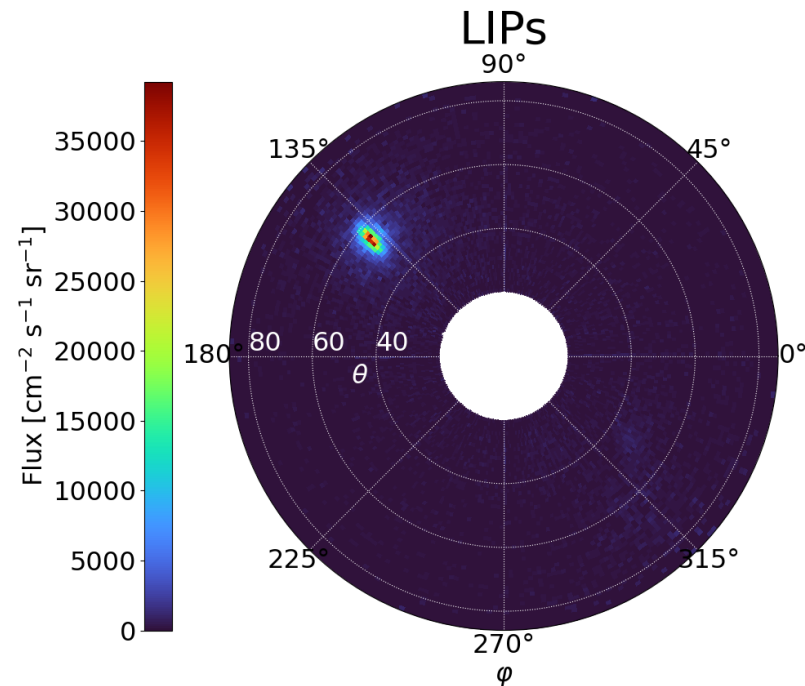
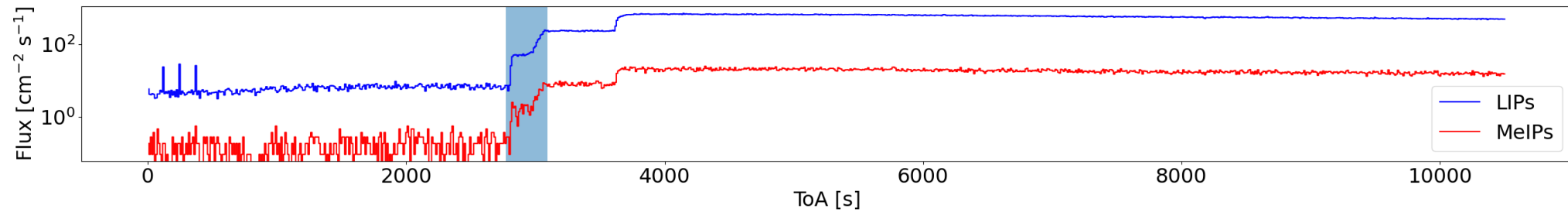
# Time-Resolved Measurement of the Directionality Map Pb-Pb Collision Period

For increased accuracy only 5 minutes of data acquisition is needed to produce reasonable statistics



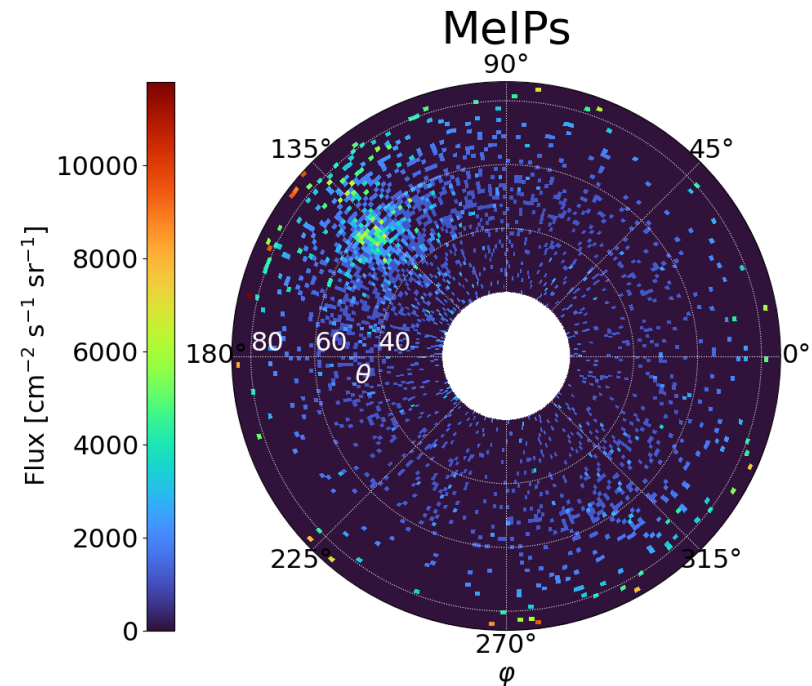
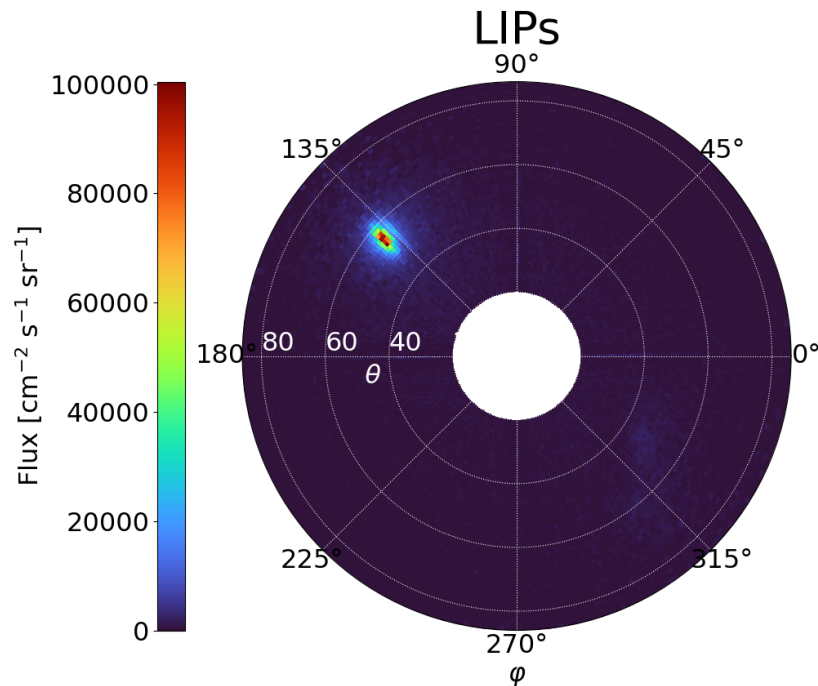
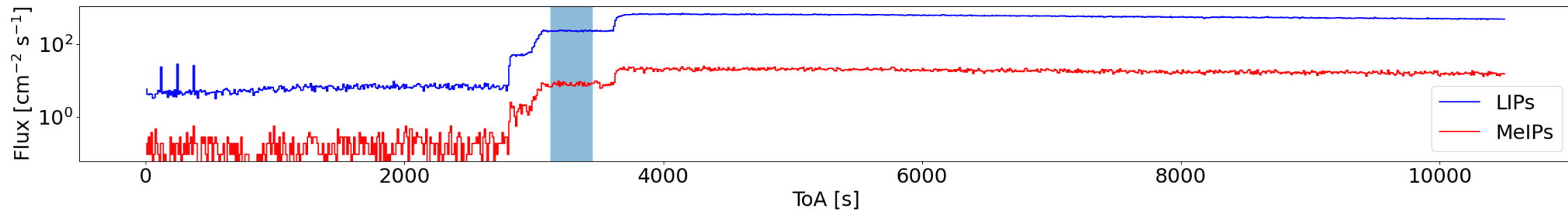
# Time-Resolved Measurement of the Directionality Map Pb-Pb Collision Period

For increased accuracy only 5 minutes of data acquisition is needed to produce reasonable statistics



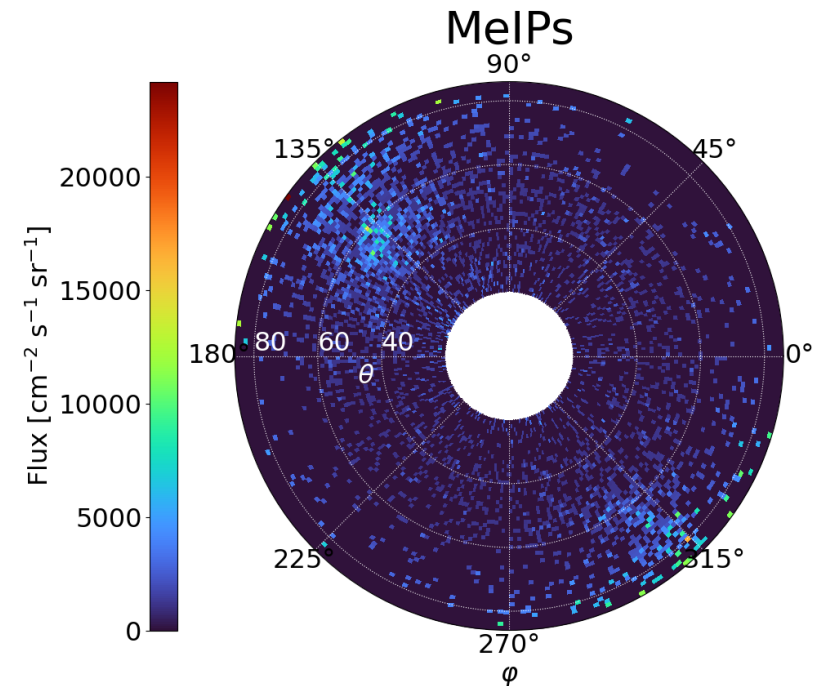
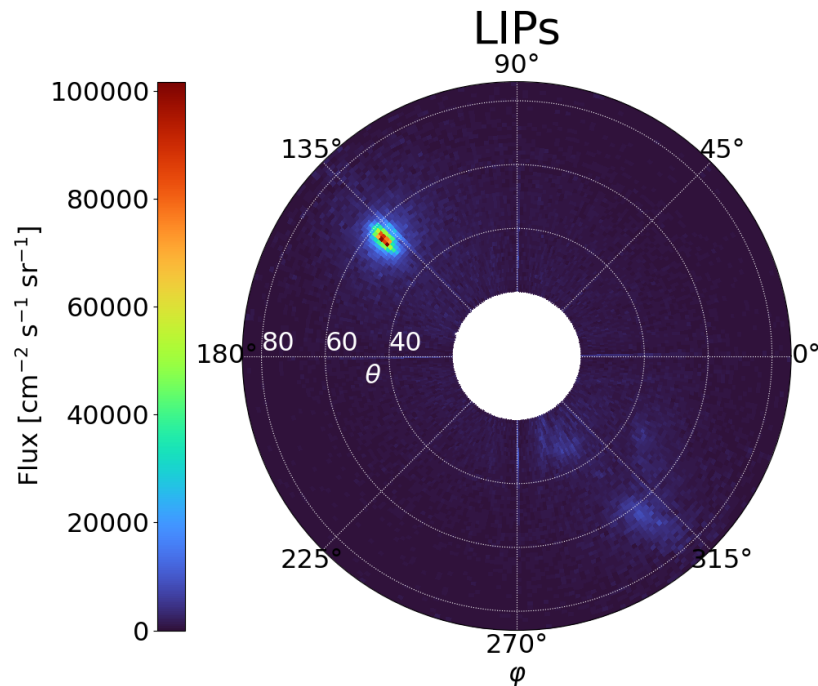
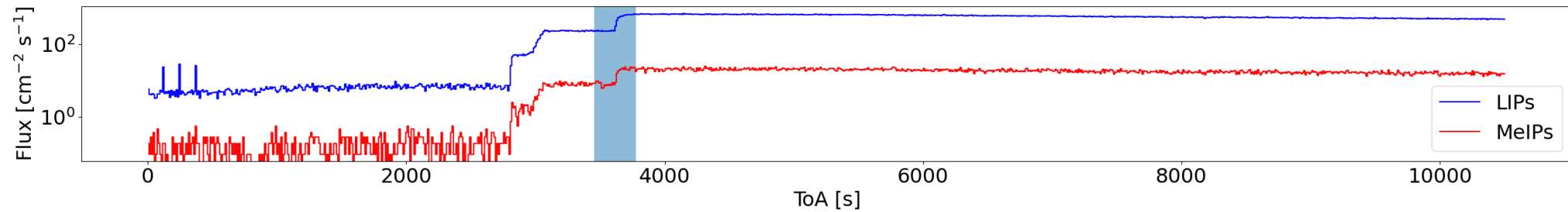
# Time-Resolved Measurement of the Directionality Map Pb-Pb Collision Period

For increased accuracy only 5 minutes of data acquisition is needed to produce reasonable statistics



# Time-Resolved Measurement of the Directionality Map Pb-Pb Collision Period

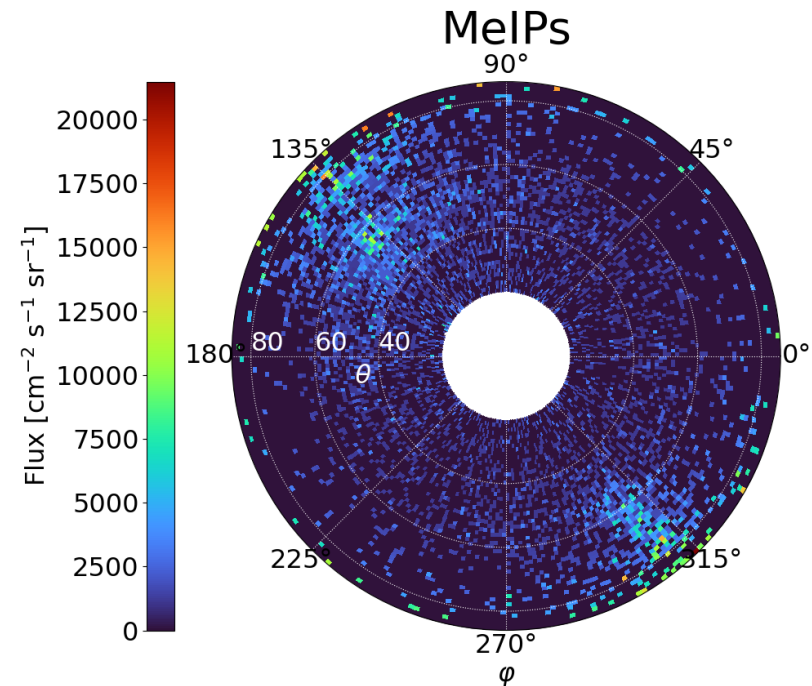
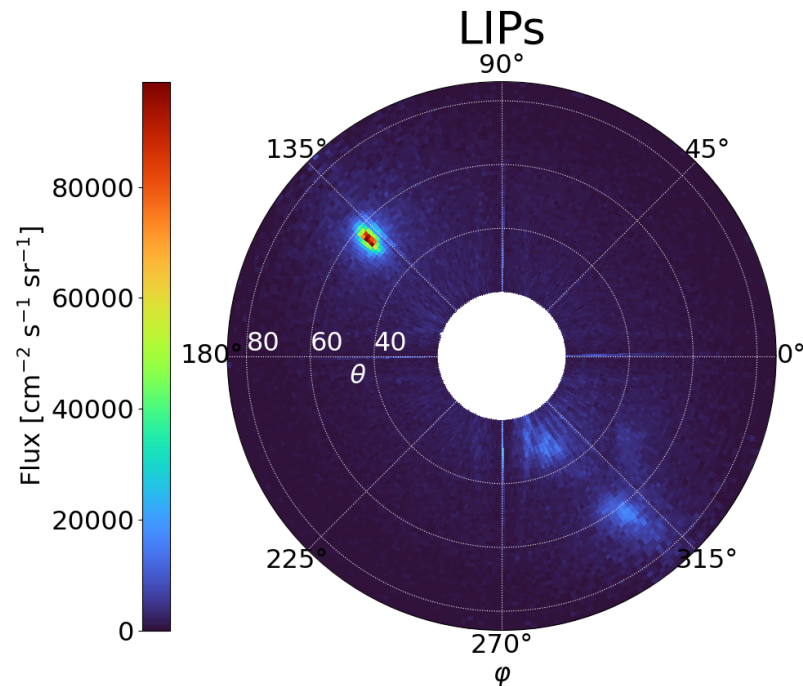
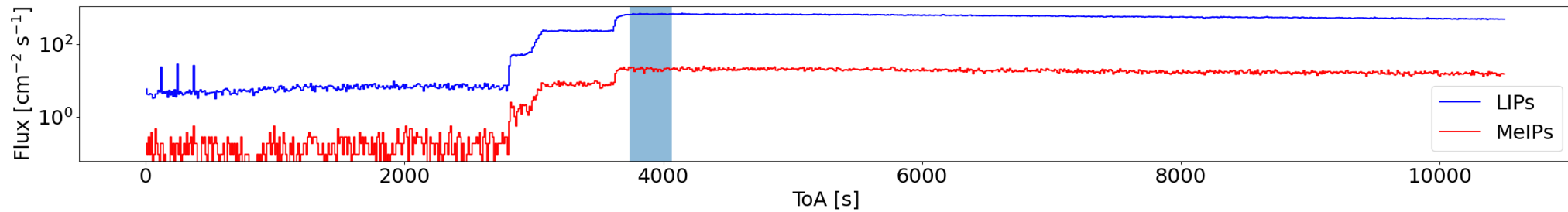
For increased accuracy only 5 minutes of data acquisition is needed to produce reasonable statistics





# Time-Resolved Measurement of the Directionality Map Pb-Pb Collision Period

For increased accuracy only 5 minutes of data acquisition is needed to produce reasonable statistics



# **Radiation Field Decomposition According to Stopping Power**

SATRAM And MOEDAL

# A Brief Excursion:

## Radiation Field Decomposition in Space

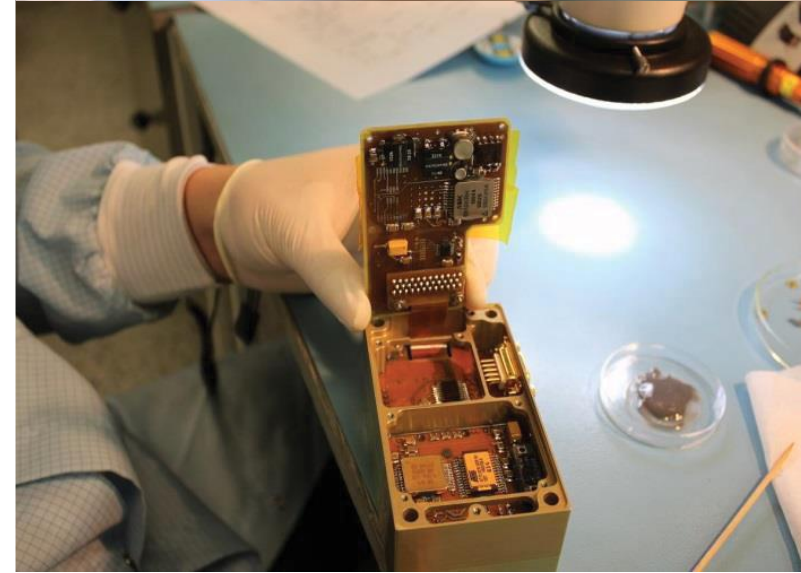
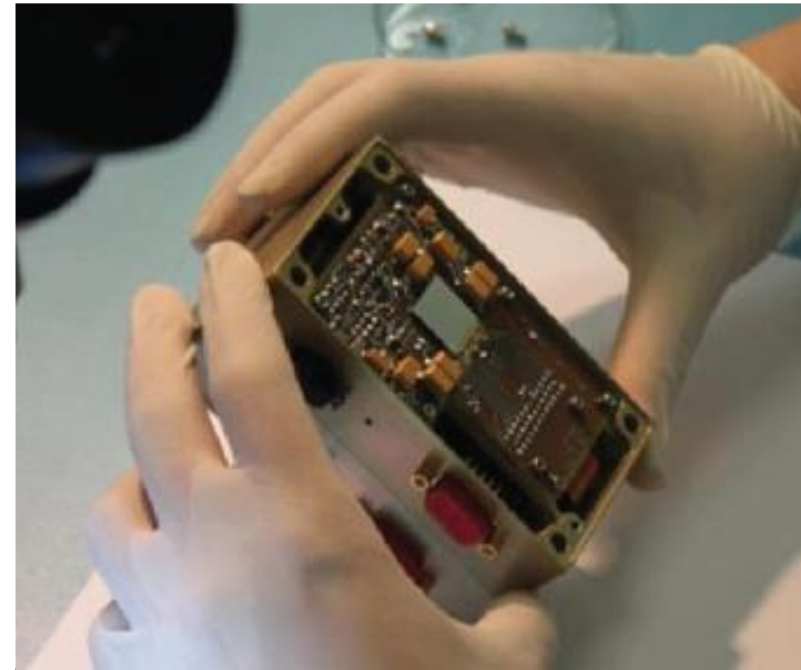
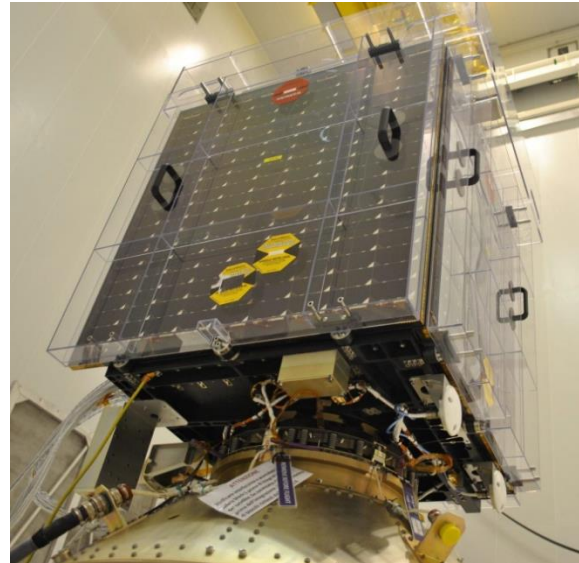
- Why the detour?
  - The mixed radiation field present in space has much lower complexity allowing us to more effectively develop and test algorithms for classification
- The data that will be used will come from the **S**pace **A**pplication **T**imepix **R**adiation **M**onitor (**SATRAM**)
- The acquired knowledge can then be applied to the MoEDAL experiment

# Space Application of Timepix Radiation Monitor (SATRAM)

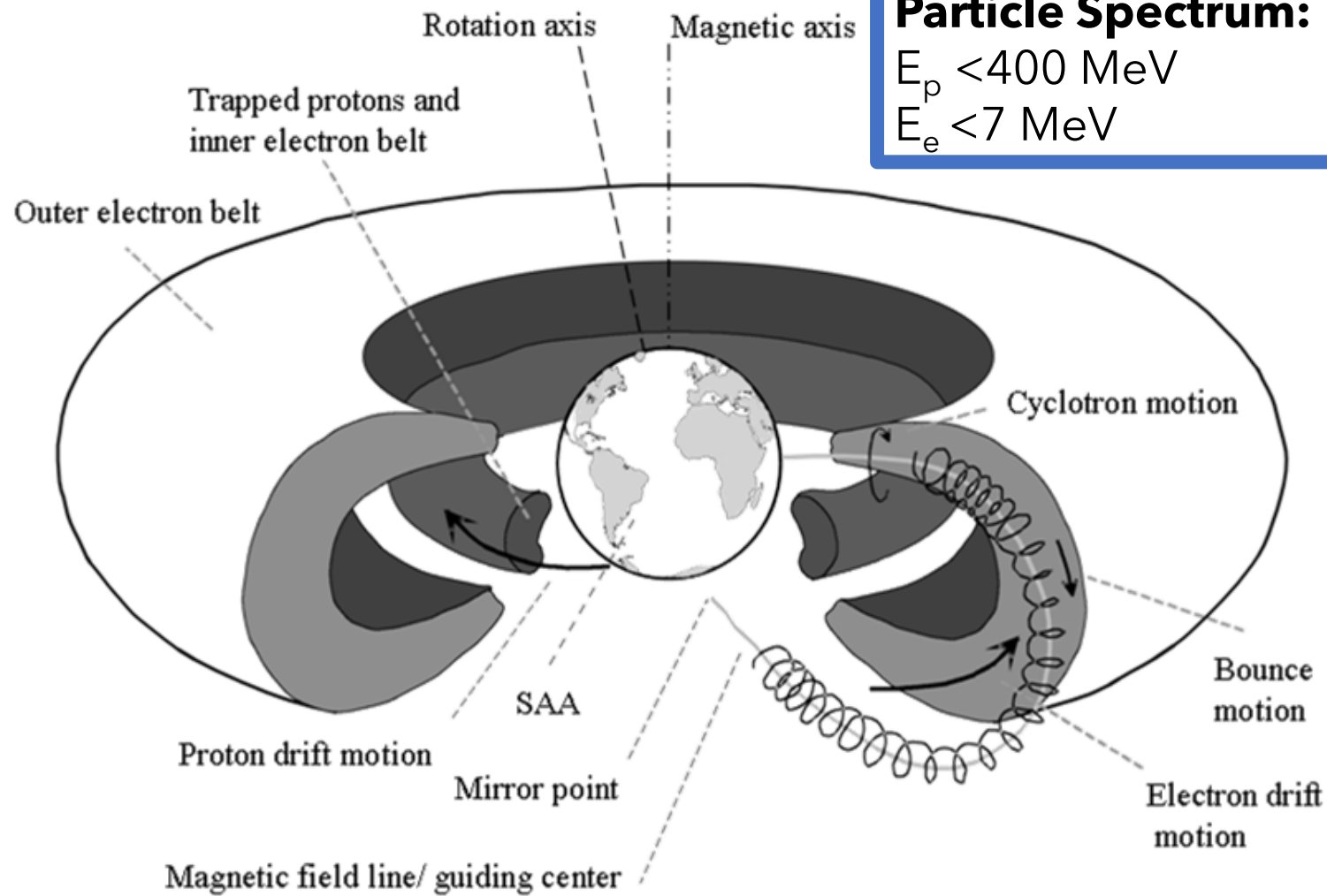
- First Timepix in open space
- Power consumption of **2.5 W**
- Total mass **380 g** (107 x 70 x 55 mm)

## Proba-V

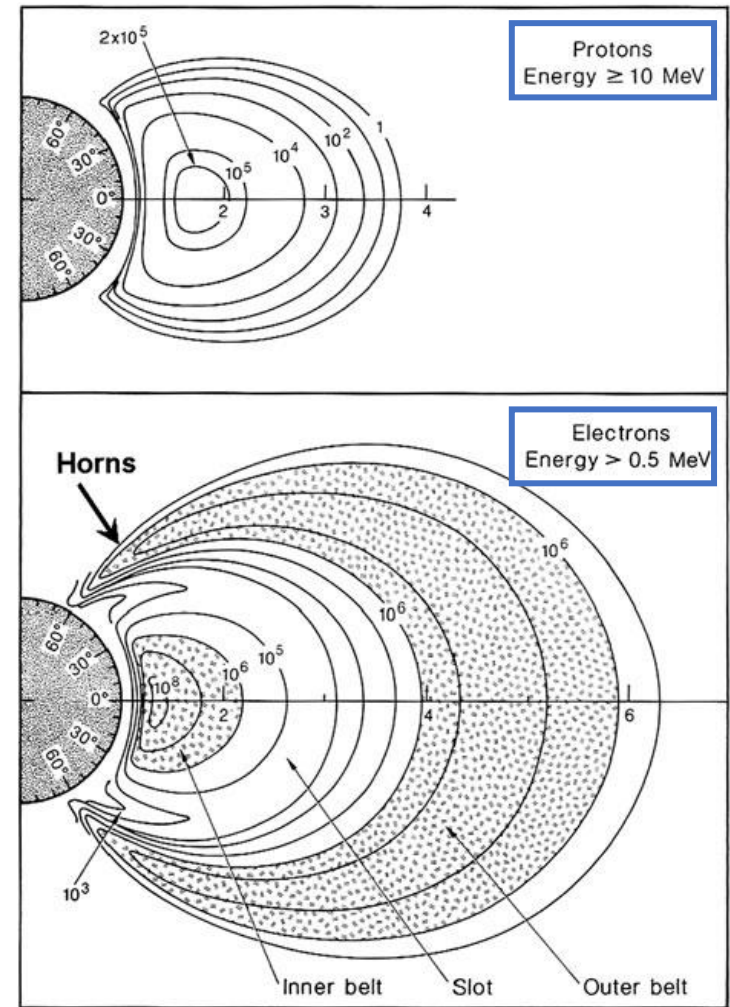
- Minisatellite (158 kg)
- Altitude ~820 km (LEO)
- 101.21 minutes orbit duration
- Inclination 98.6°
- Sun-synchronous
- Launched 7th March 2013



# The Radiation Environment

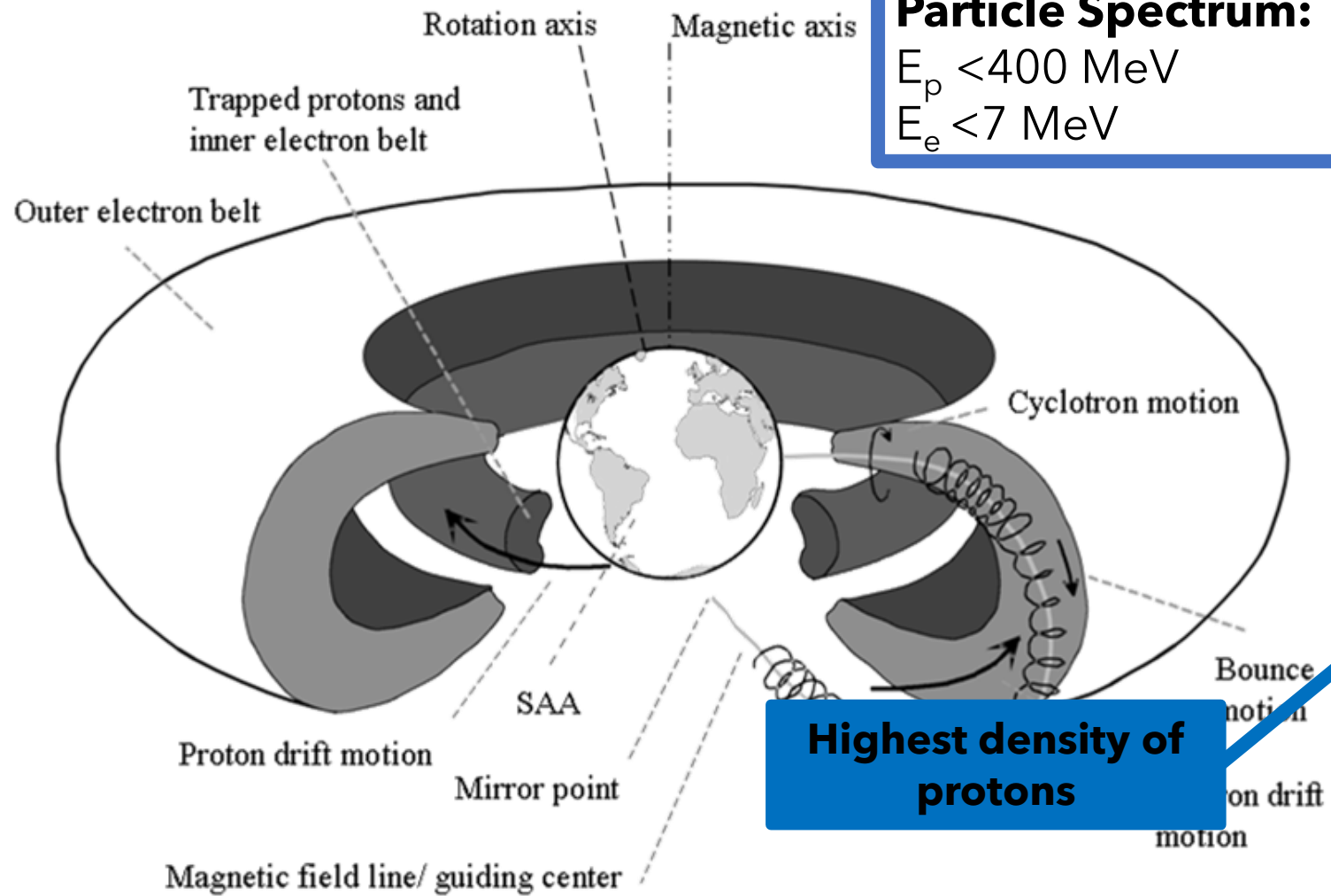


**Particle Spectrum:**  
 $E_p < 400 \text{ MeV}$   
 $E_e < 7 \text{ MeV}$



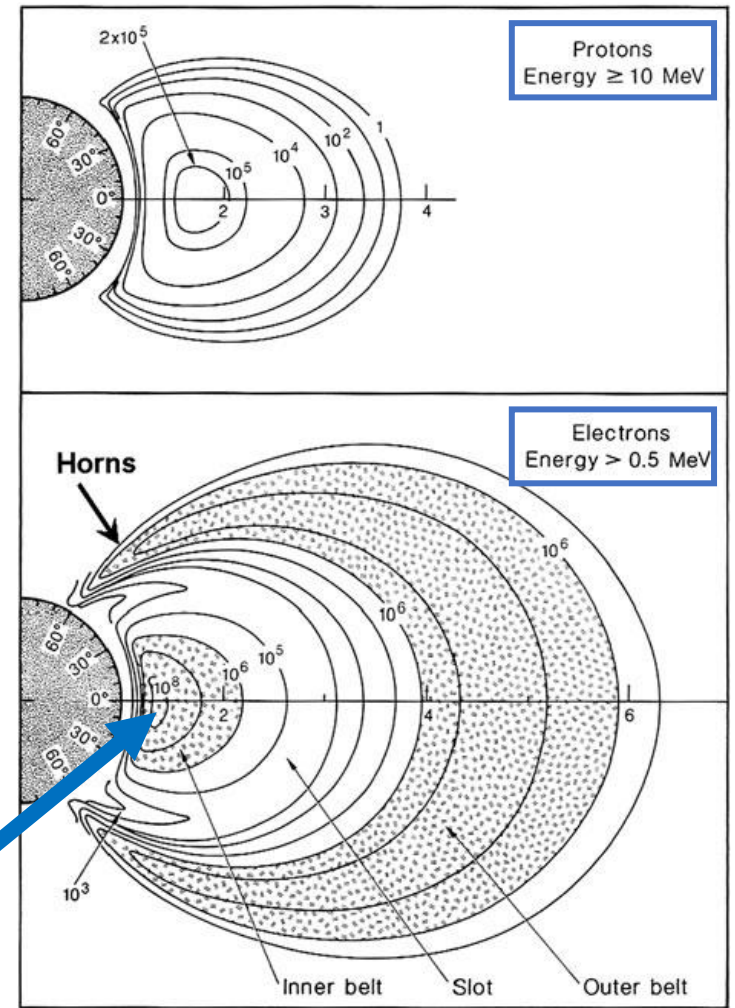
Mauk, B.H., Fox, N.J., Kanekal, S.G. et al. Space Sci Rev (2013) 179: 3.

# The Radiation Environment

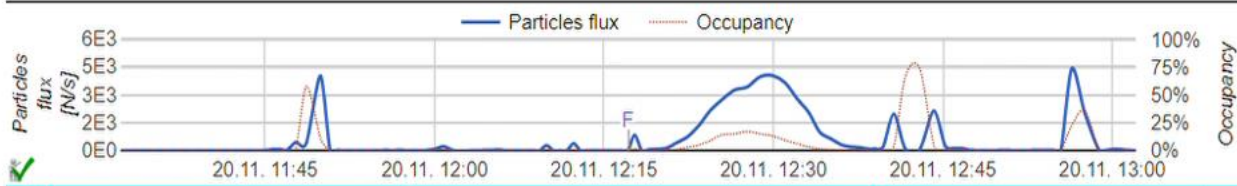


**Particle Spectrum:**  
 $E_p < 400 \text{ MeV}$   
 $E_e < 7 \text{ MeV}$

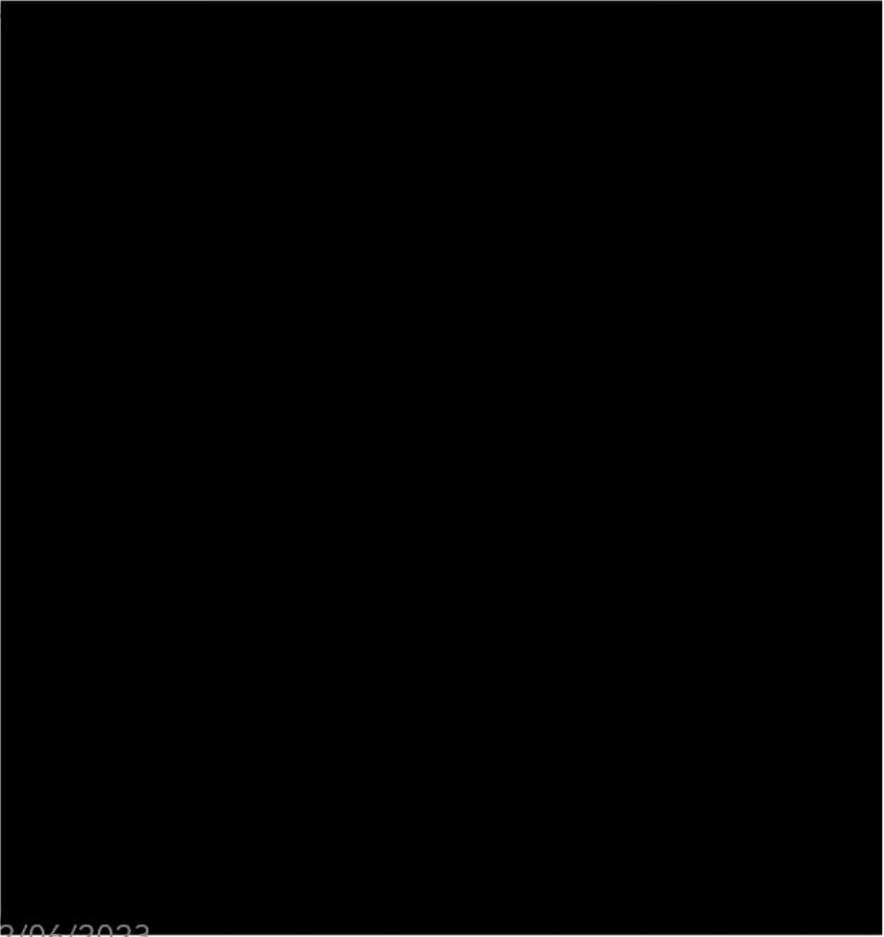
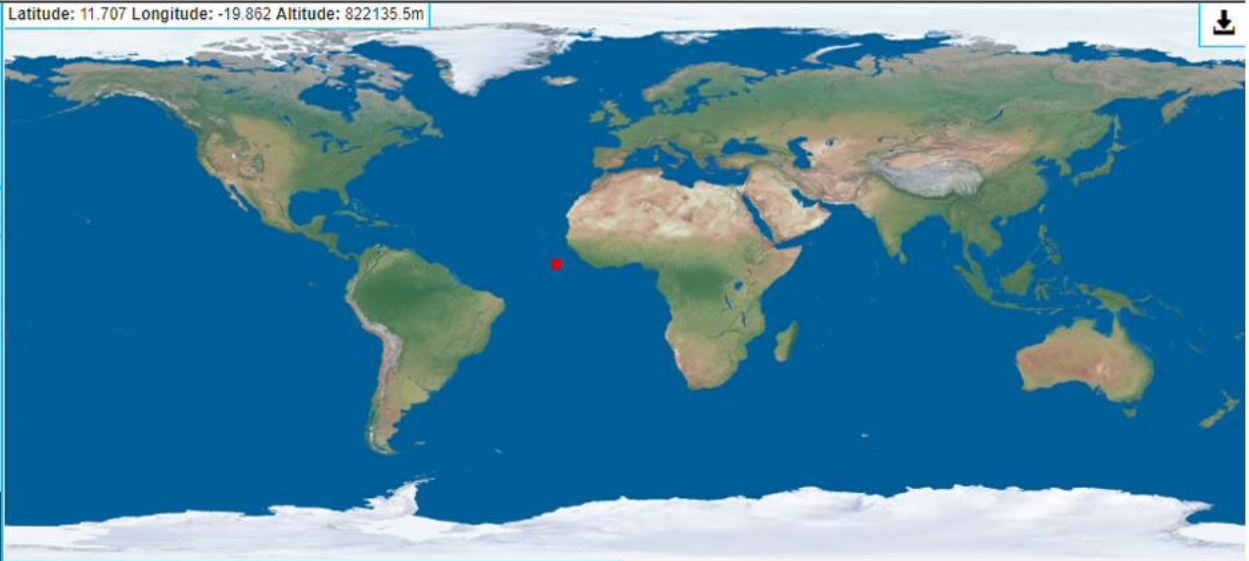
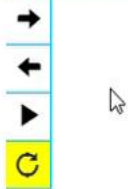
**Highest density of protons**



Mauk, B.H., Fox, N.J., Kanekal, S.G. et al. Space Sci Rev (2013) 179: 3.



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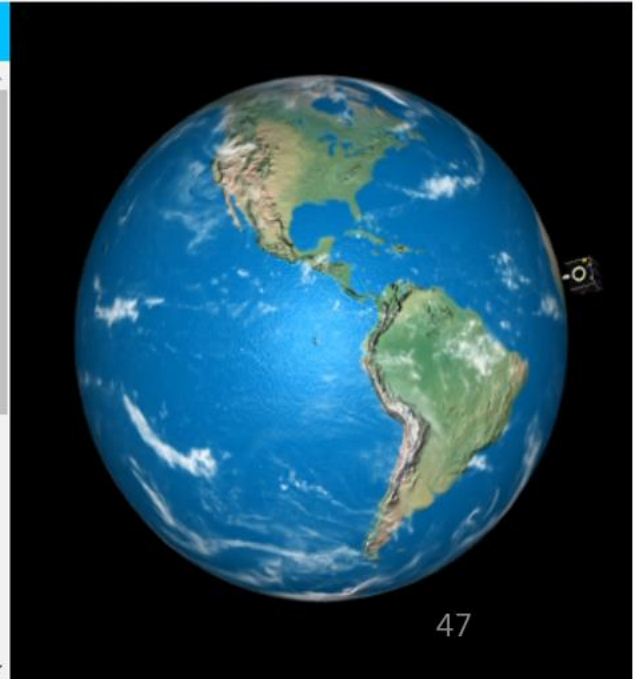
Energy [keV]

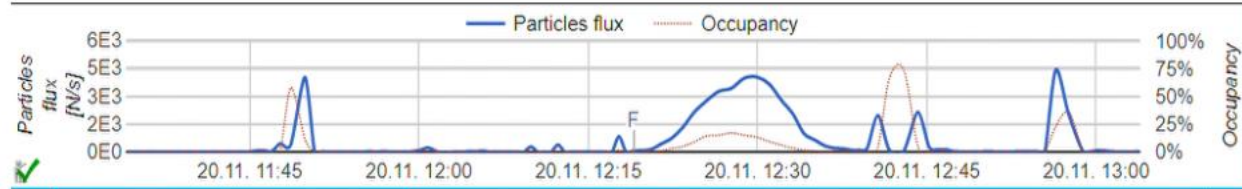
Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	0	0	0	
Small blob	0	0	0	
Heavy blob	0	0	0	
Heavy track	0	0	0	
Straight track	0	0	0	
Curly track	0	0	0	
<b>Sum:</b>	<b>0</b>	<b>0</b>	<b>0</b>	

Histograms

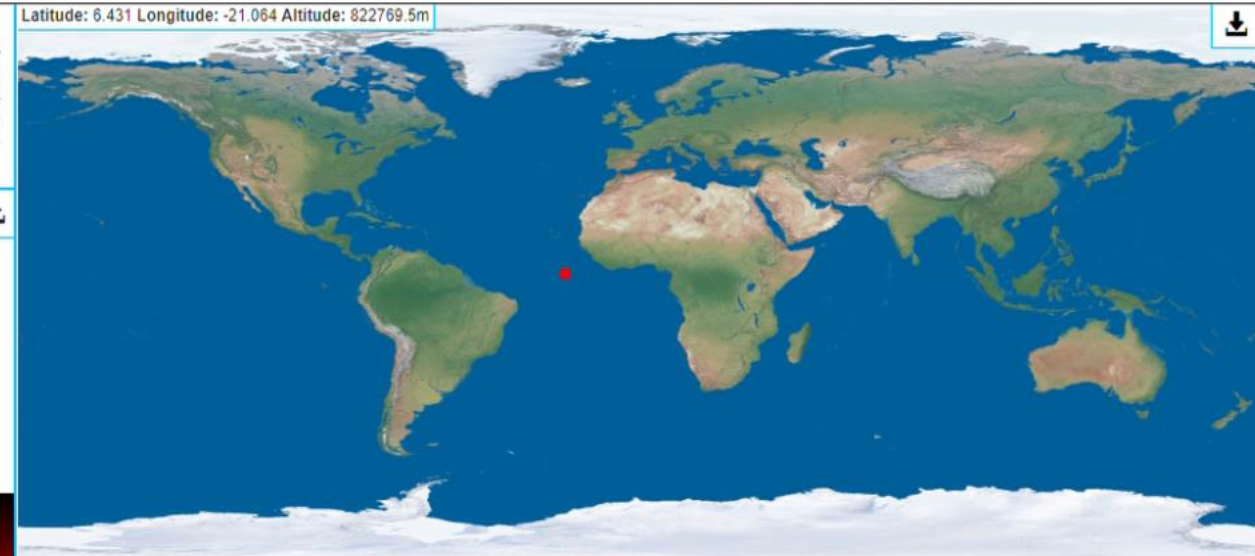
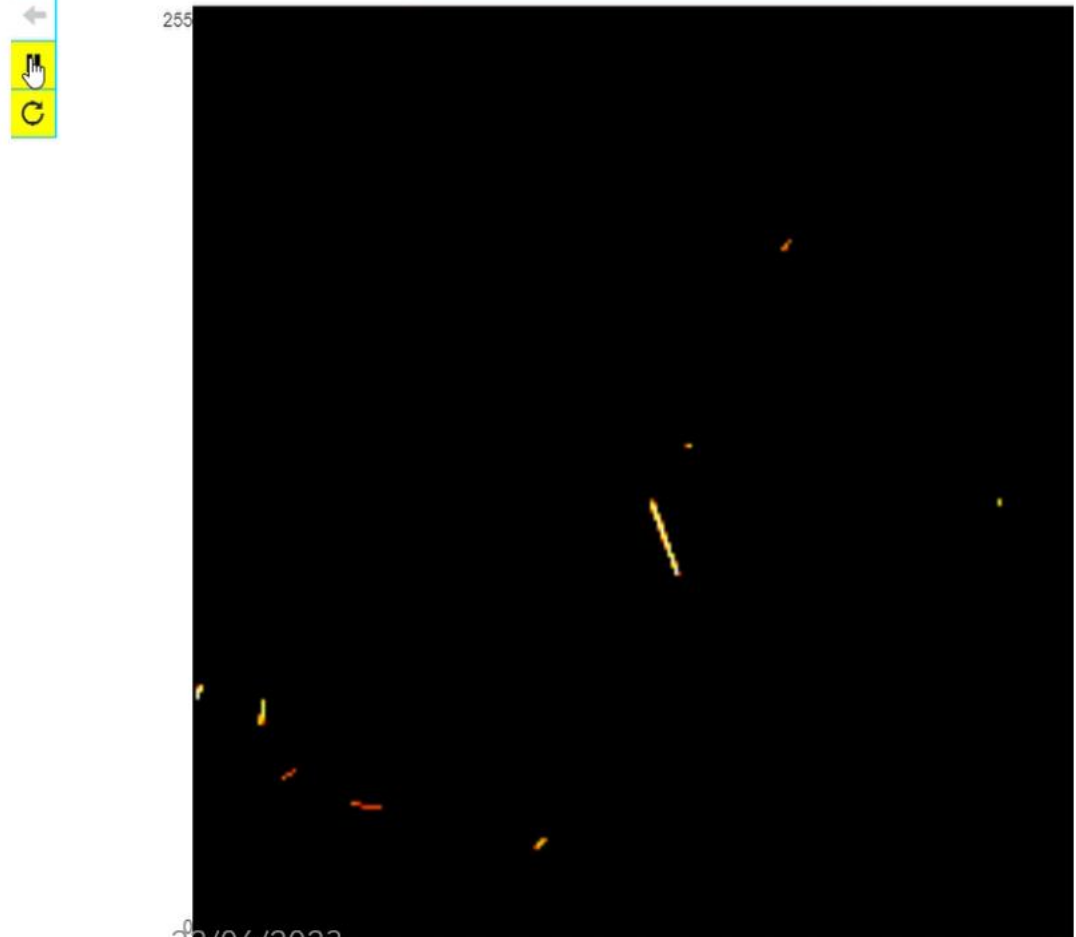
Type of histogram:  Max number of bins:

Min. value:  Max. value:





© 2014-11-20 12:18:59 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	0	0	0	
Small blob	2	10	0.706	
Heavy blob	0	0	0	
Heavy track	0	0	0	
Straight track	0	0	0	
Curly track	7	35	21.772	
<b>Sum:</b>	<b>9</b>	<b>45</b>	<b>22.479</b>	

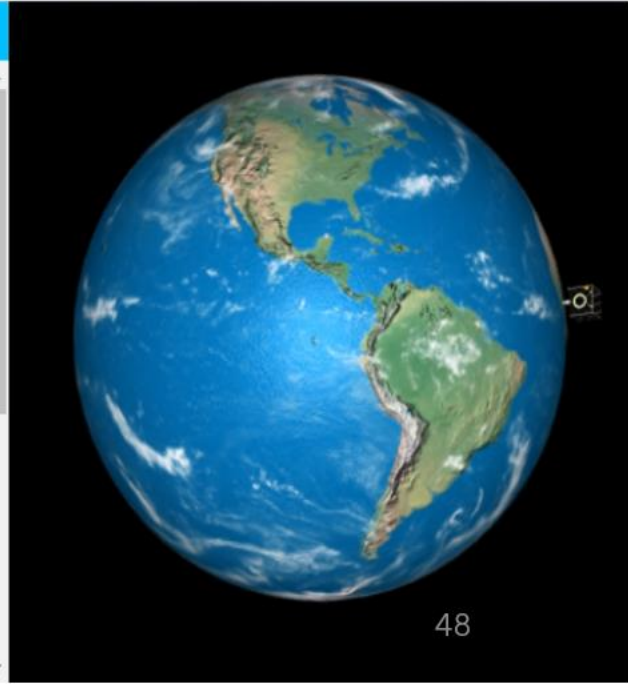
Histograms

Type of histogram: Volume [keV]

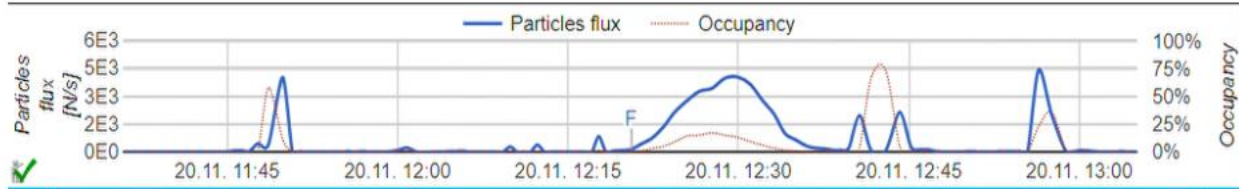
Max number of bins:

Min. value:

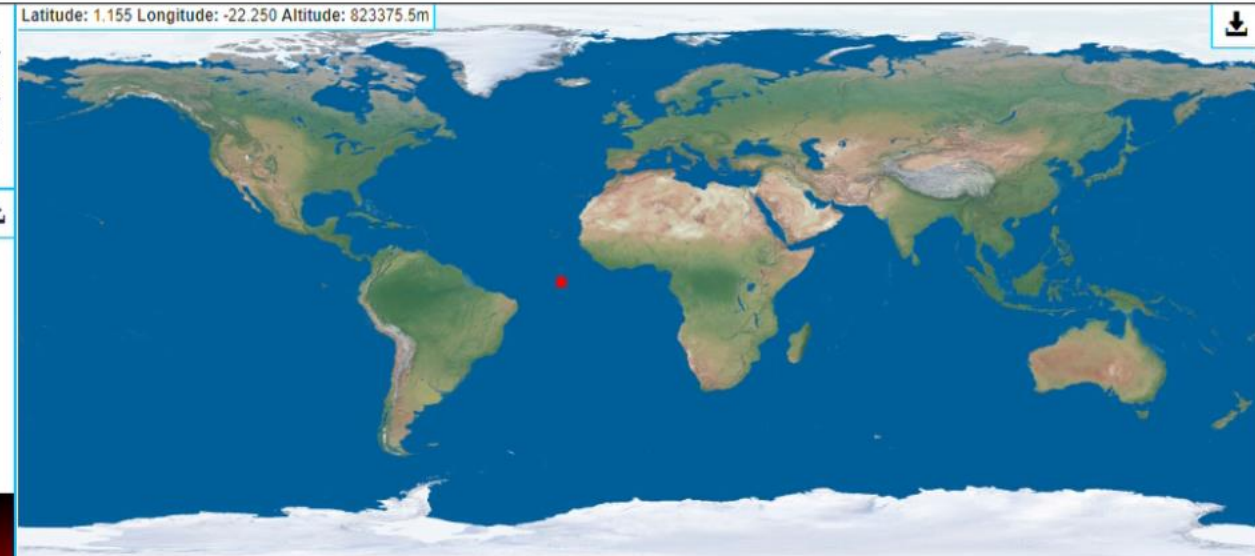
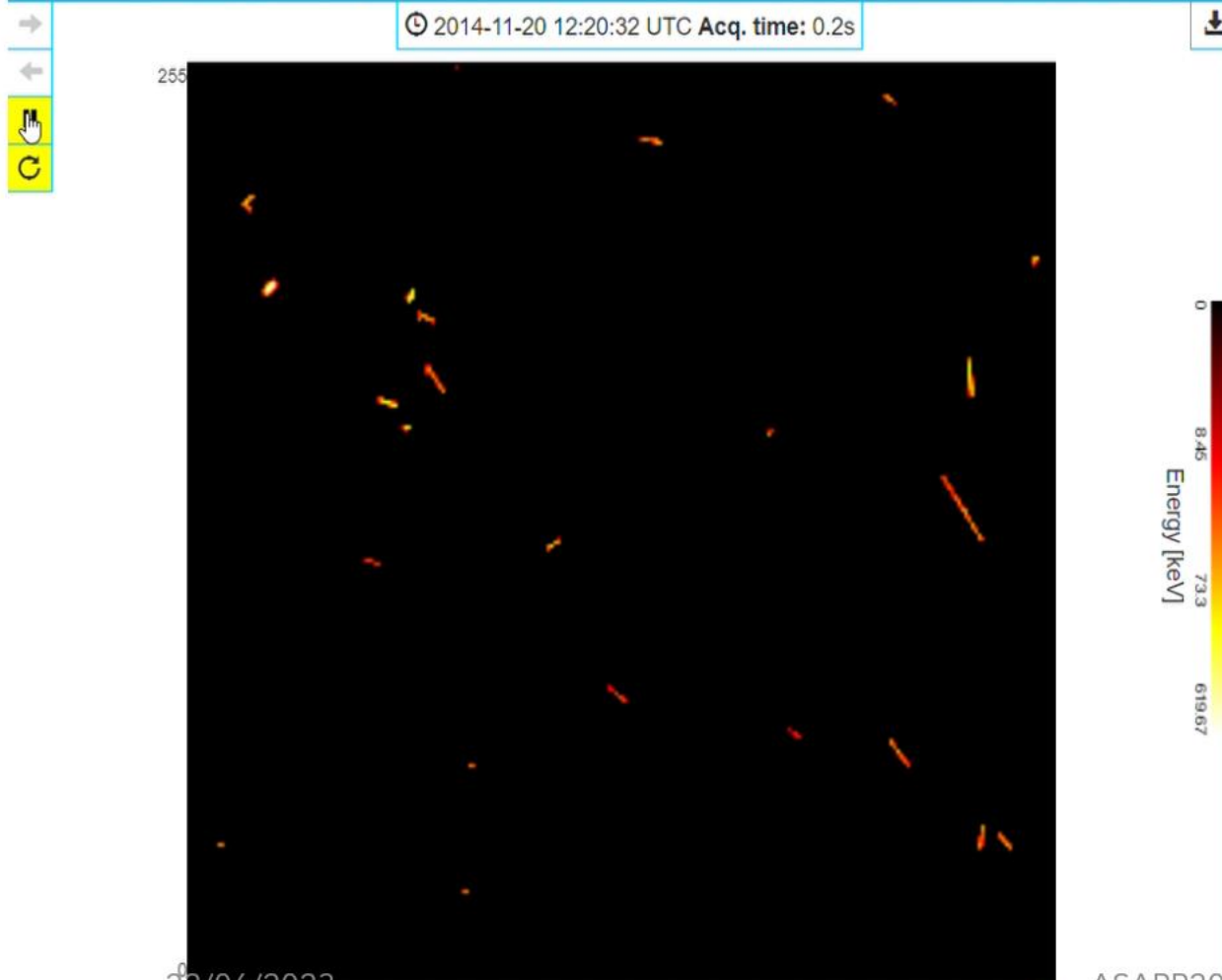
Max. value:







© 2014-11-20 12:20:32 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	0	0	0	
Small blob	4	20	1.08	
Heavy blob	1	5	4.243	
Heavy track	1	5	12.367	
Straight track	2	10	3.933	
Curly track	16	80	25.729	
<b>Sum:</b>	<b>24</b>	<b>120</b>	<b>47.353</b>	

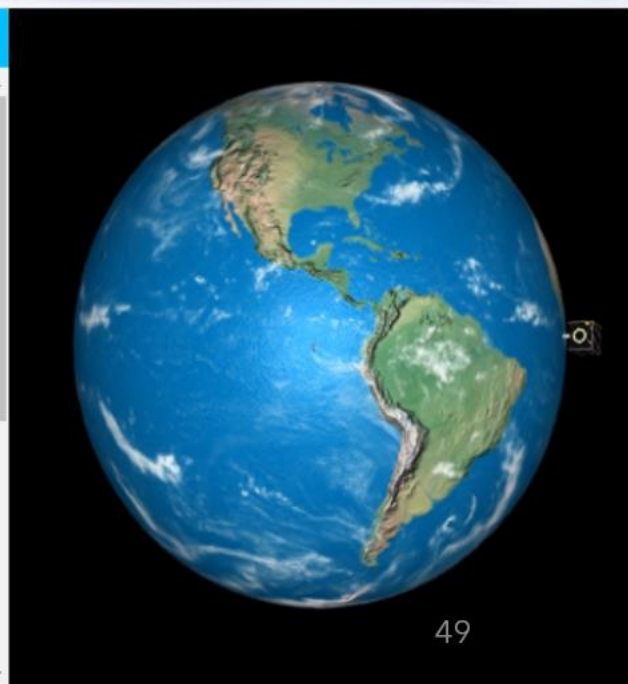
Histograms

Type of histogram: Volume [keV]

Max number of bins:

Min. value:

Max. value:

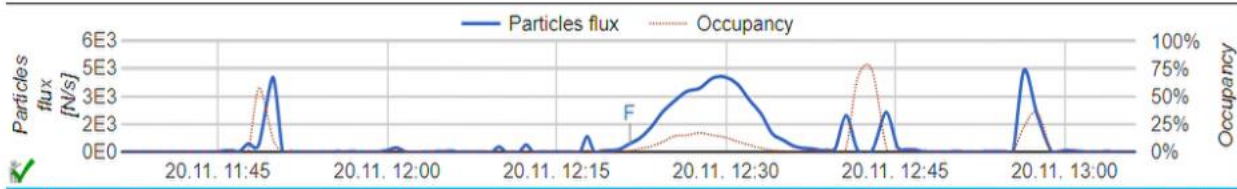


23/06/2023

255

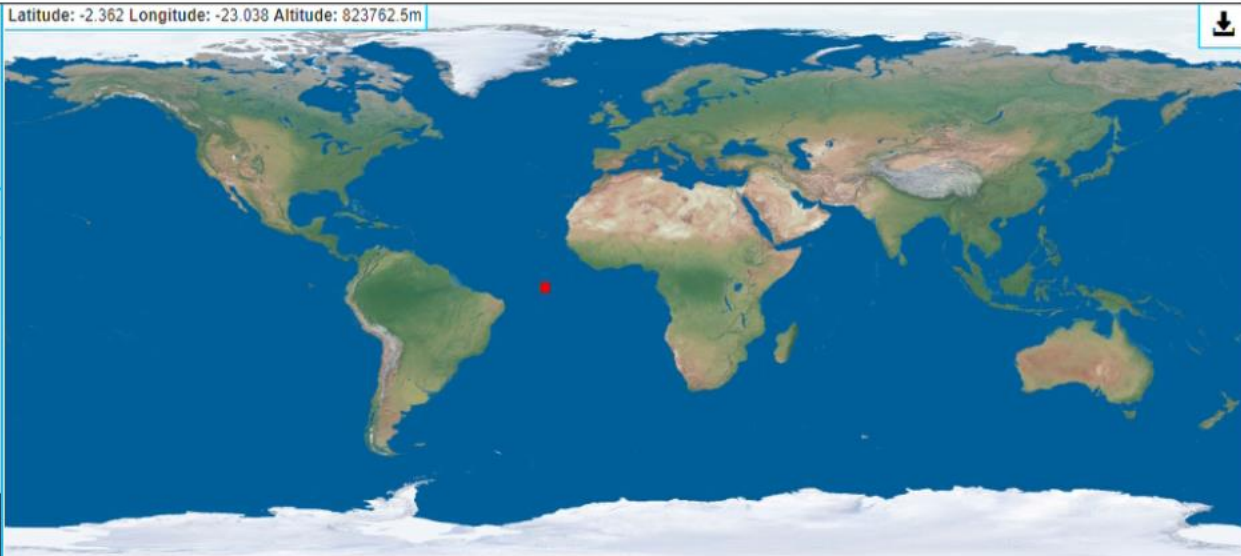
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© 2014-11-20 12:21:27 UTC Acq. time: 0.2s

Navigation icons: checkmark, right arrow, left arrow, hand cursor, refresh.



Analysis tools: histogram, filter, zoom, pan, info.

Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	2	10	0.111	
Small blob	6	30	3.703	
Heavy blob	6	30	51.891	
Heavy track	1	5	6.305	
Straight track	1	5	3.802	
Curly track	67	335	90.875	
<b>Sum:</b>	<b>83</b>	<b>415</b>	<b>156.688</b>	

Histograms

Type of histogram: Volume [keV] (dropdown)

Max number of bins:

Min. value:

Max. value:

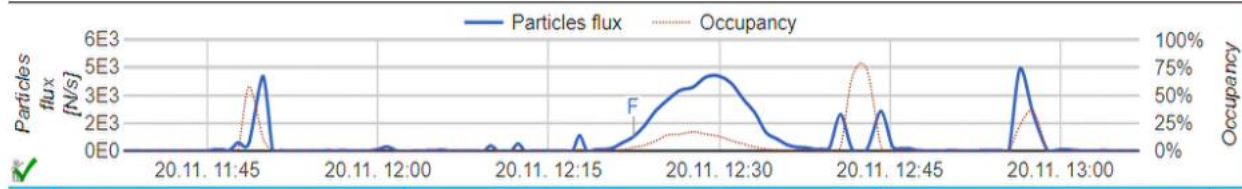


22/06/2023

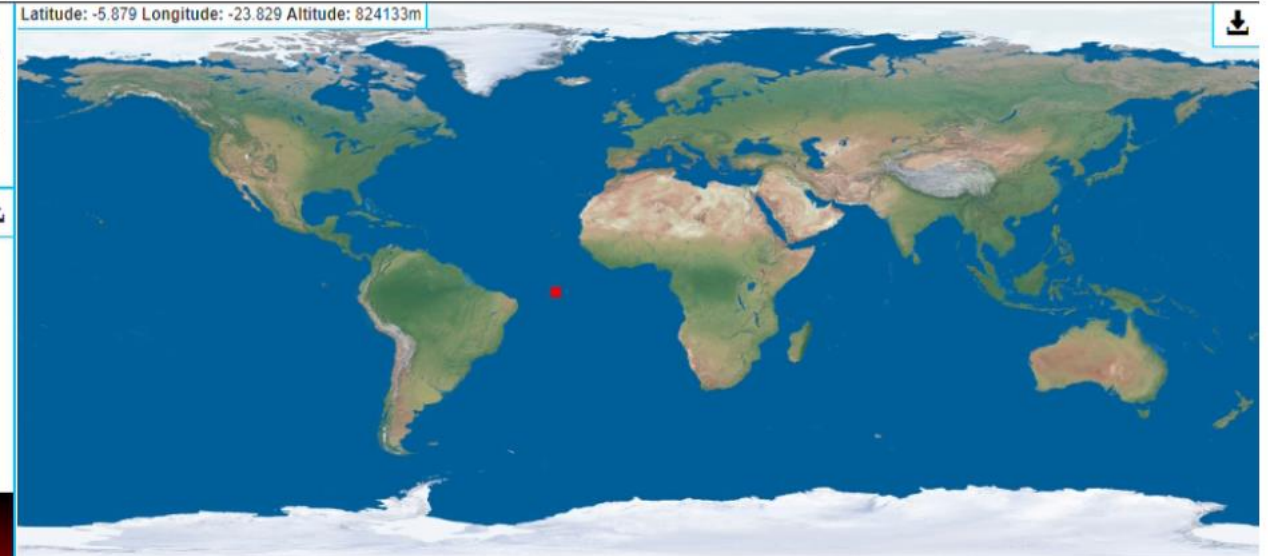
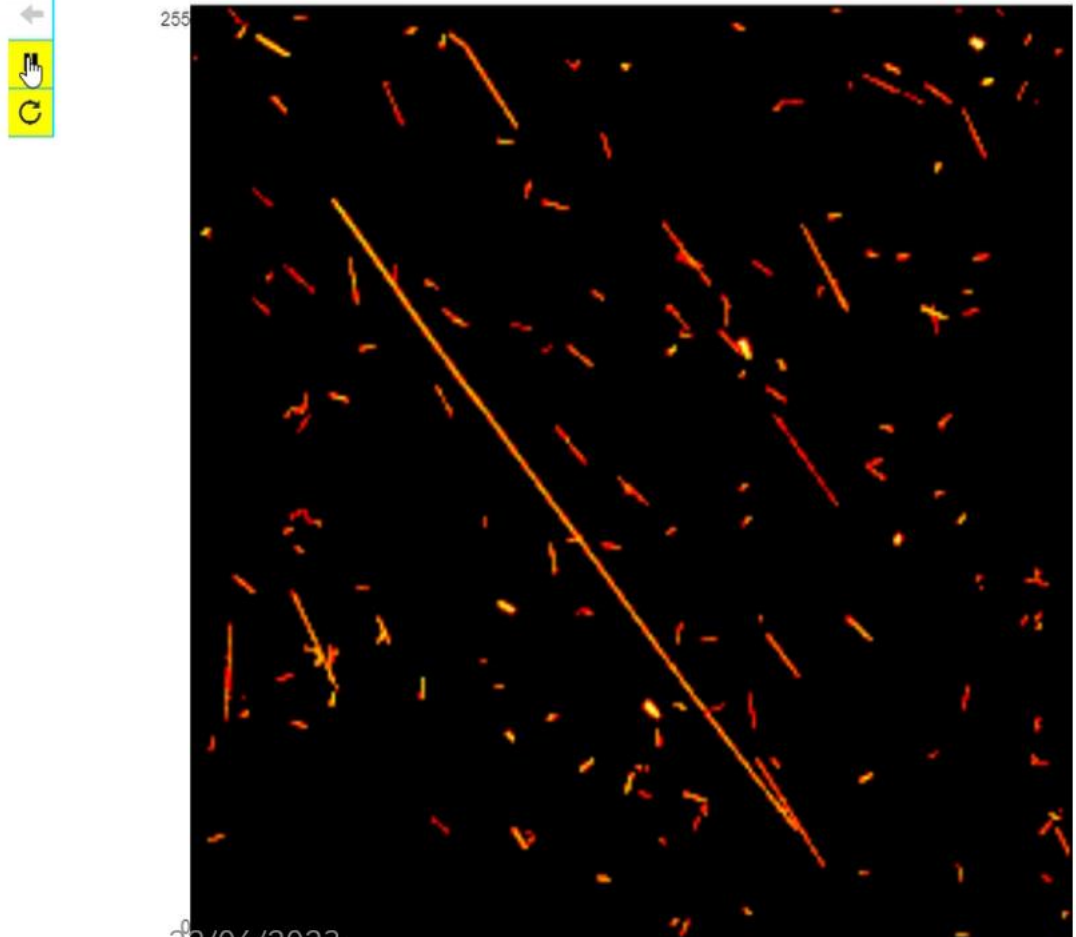
255

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© 2014-11-20 12:22:24 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	4	20	0.311	
Small blob	9	45	3.489	
Heavy blob	4	20	38.797	
Heavy track	2	10	26.582	
Straight track	4	20	20.761	
Curly track	123	615	339.531	
<b>Sum:</b>	<b>146</b>	<b>730</b>	<b>429.472</b>	

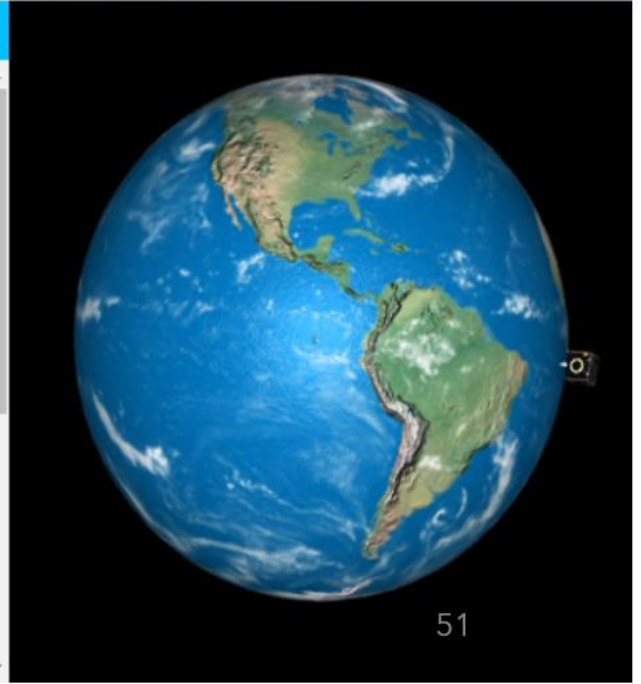
Histograms

Type of histogram: Volume [keV]

Max number of bins:

Min. value:

Max. value:

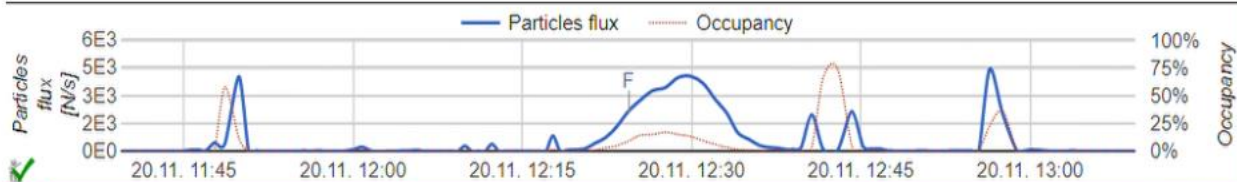


20/06/2023

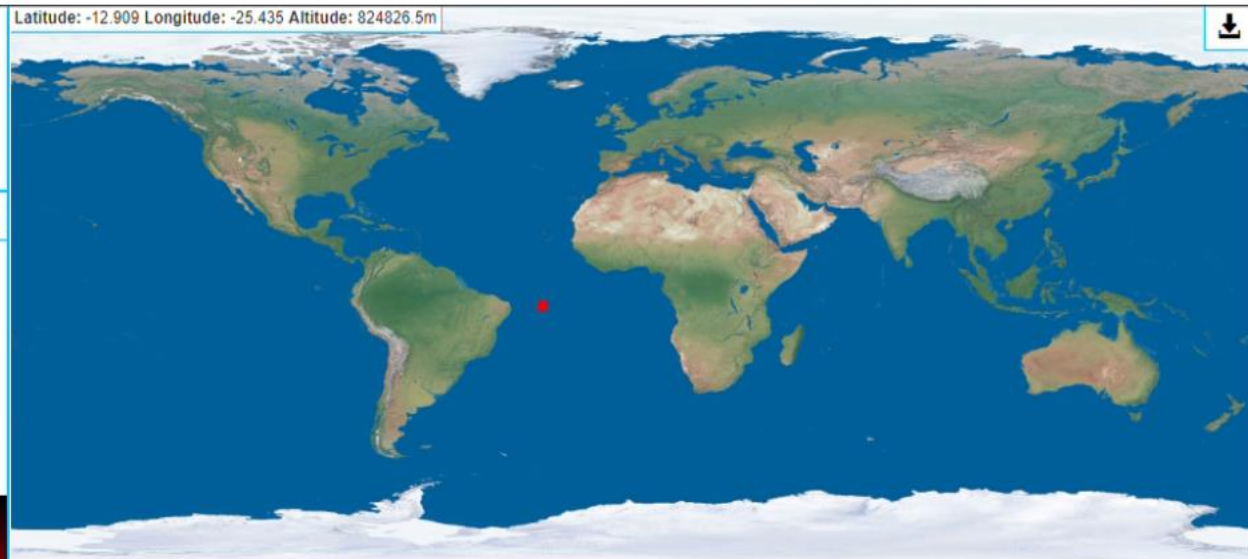
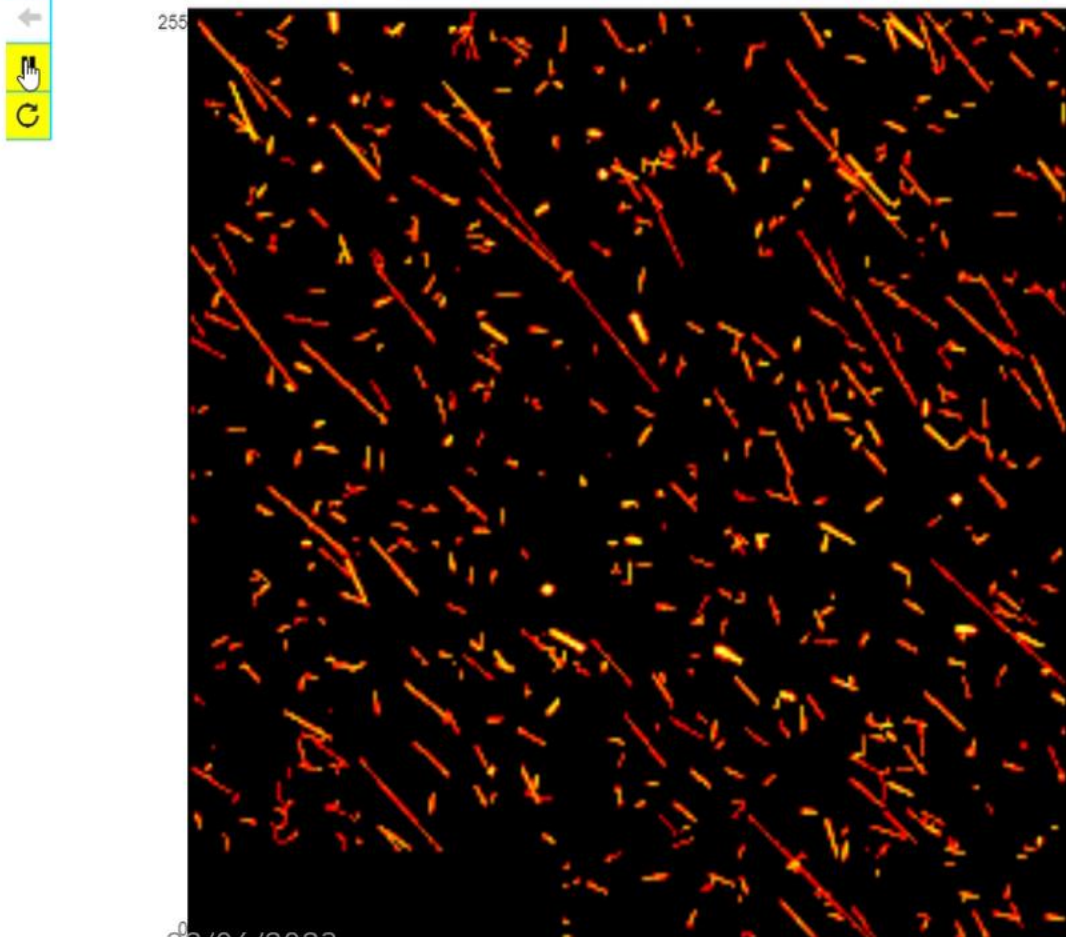
255

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© 2014-11-20 12:24:23 UTC Acq. time: 0.2s



Navigation icons: histogram, filter, zoom, pan, info.

Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	16	80	1.286	
Small blob	33	165	21.799	
Heavy blob	17	85	92.73	
Heavy track	9	45	155.187	
Straight track	11	55	54.947	
Curly track	338	1690	942.583	
<b>Sum:</b>	<b>424</b>	<b>2120</b>	<b>1.269e+3</b>	

Histograms

Type of histogram: Volume [keV]

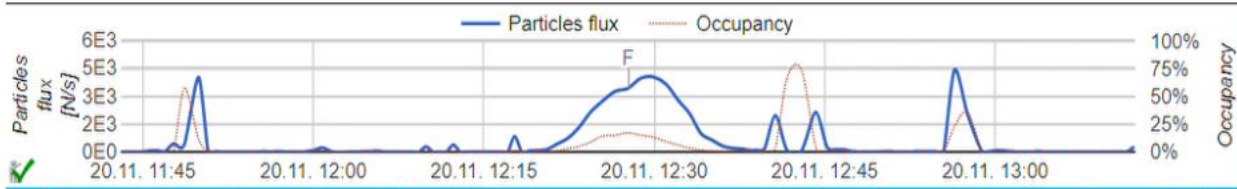
Max number of bins:

Min. value:

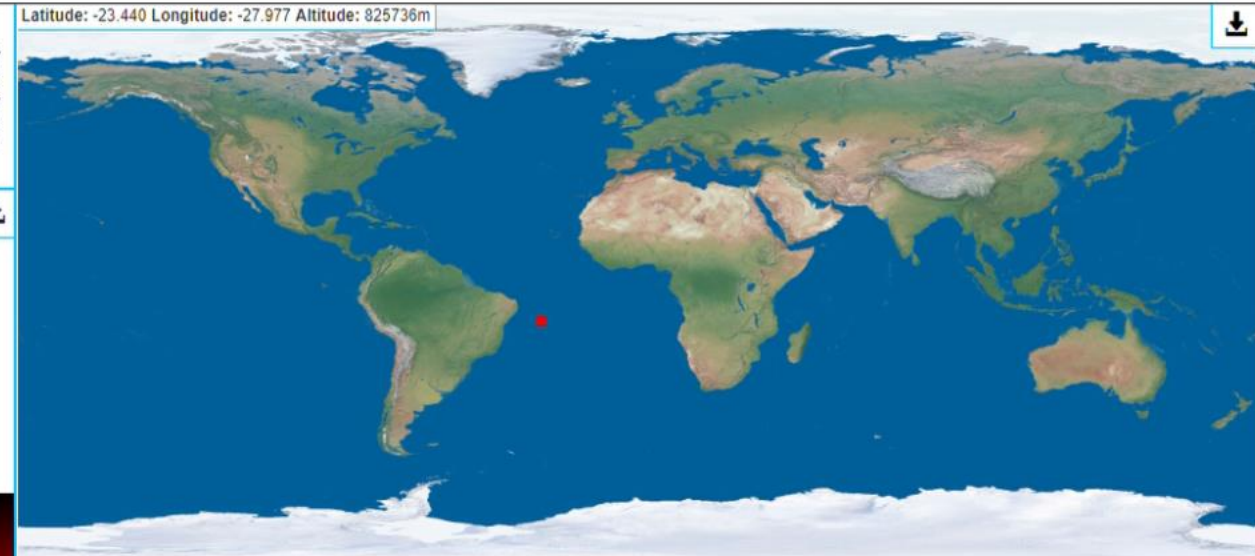
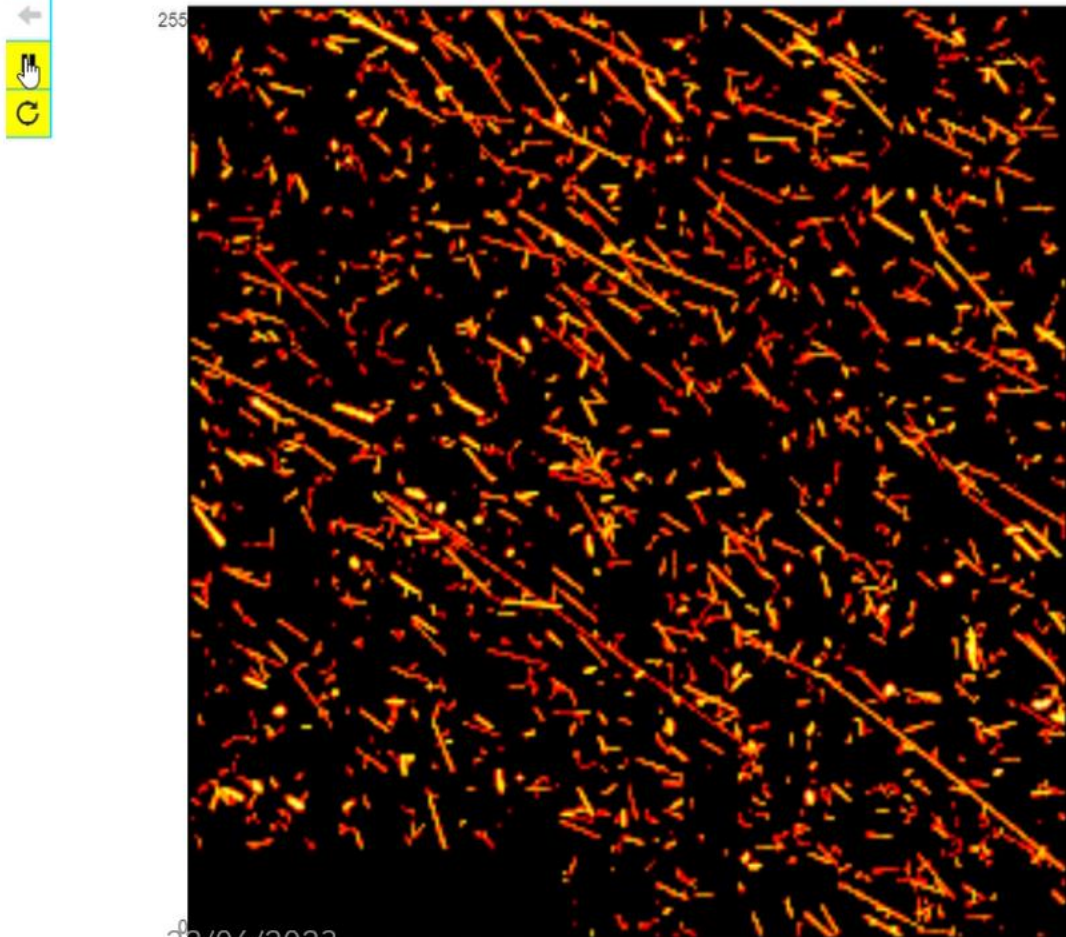
Max. value:



22/06/2023



© 2014-11-20 12:27:41 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	101	505	7.997	
Small blob	86	430	34.32	
Heavy blob	31	155	175.737	
Heavy track	25	125	534.847	
Straight track	4	20	25.526	
Curly track	433	2165	2.071e+3	
<b>Sum:</b>	<b>680</b>	<b>3400</b>	<b>2.850e+3</b>	

**Histograms**

Type of histogram: Volume [keV] (dropdown menu)

Max number of bins:

Min. value:

Max. value:

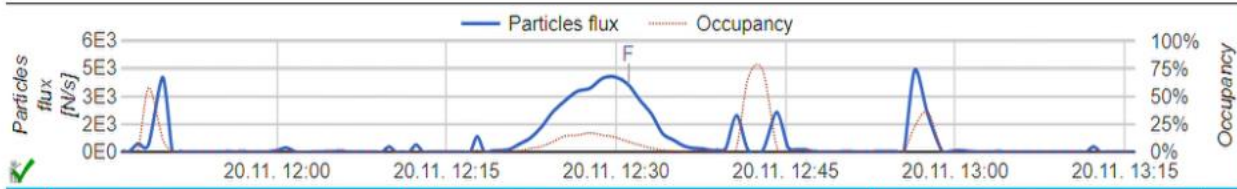


23/06/2023

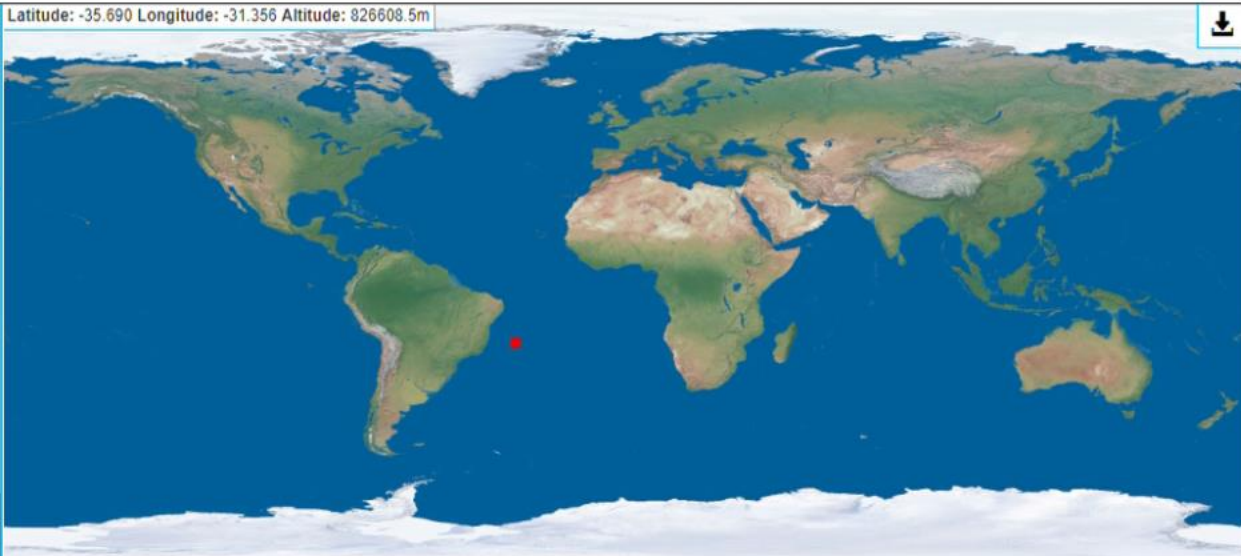
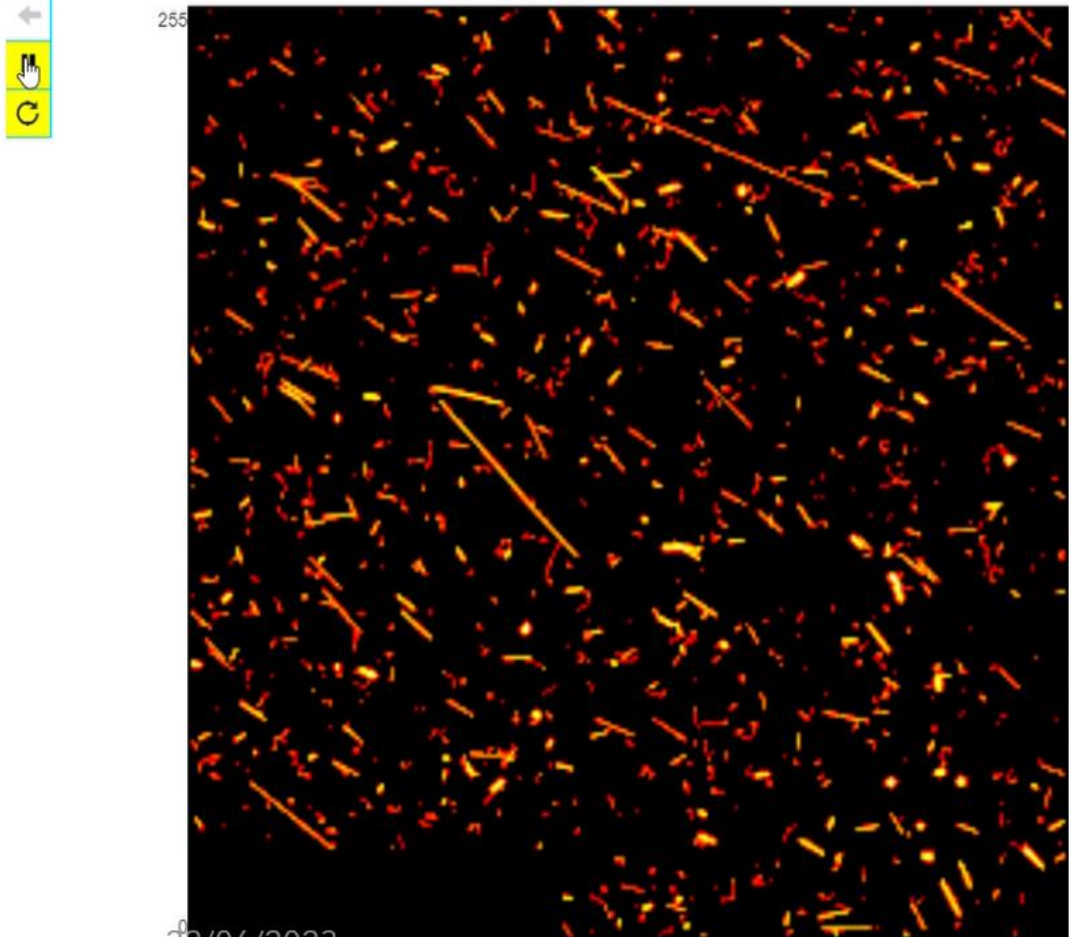
255

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© 2014-11-20 12:31:03 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	165	825	13.769	
Small blob	132	660	51.519	
Heavy blob	42	210	204.791	
Heavy track	25	125	358.957	
Straight track	2	10	7.83	
Curly track	358	1790	922.126	
<b>Sum:</b>	<b>724</b>	<b>3620</b>	<b>1.559e+3</b>	

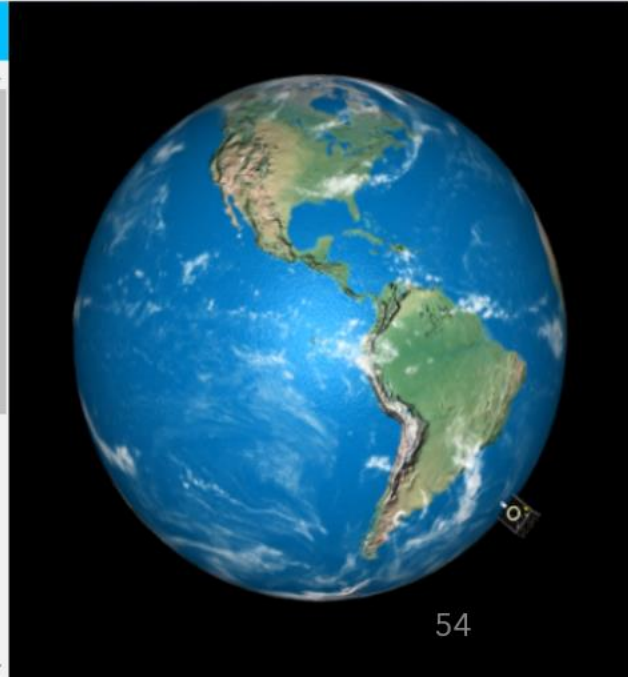
Histograms

Type of histogram: Volume [keV]

Max number of bins:

Min. value:

Max. value:

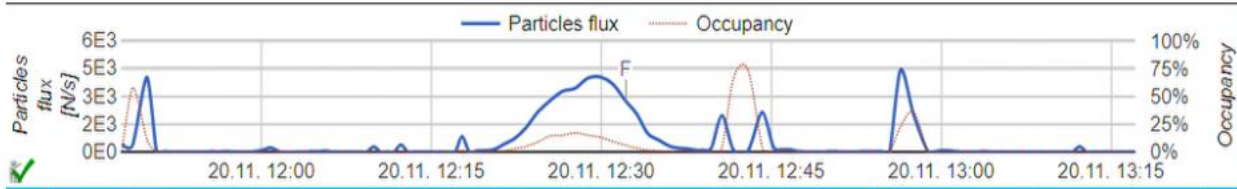


23/06/2023

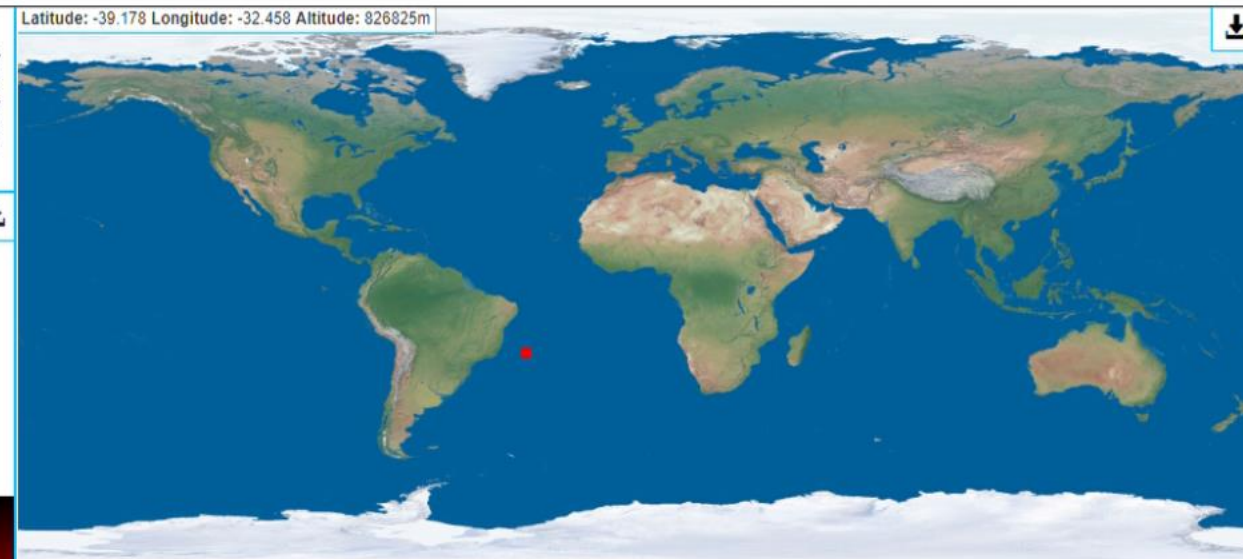
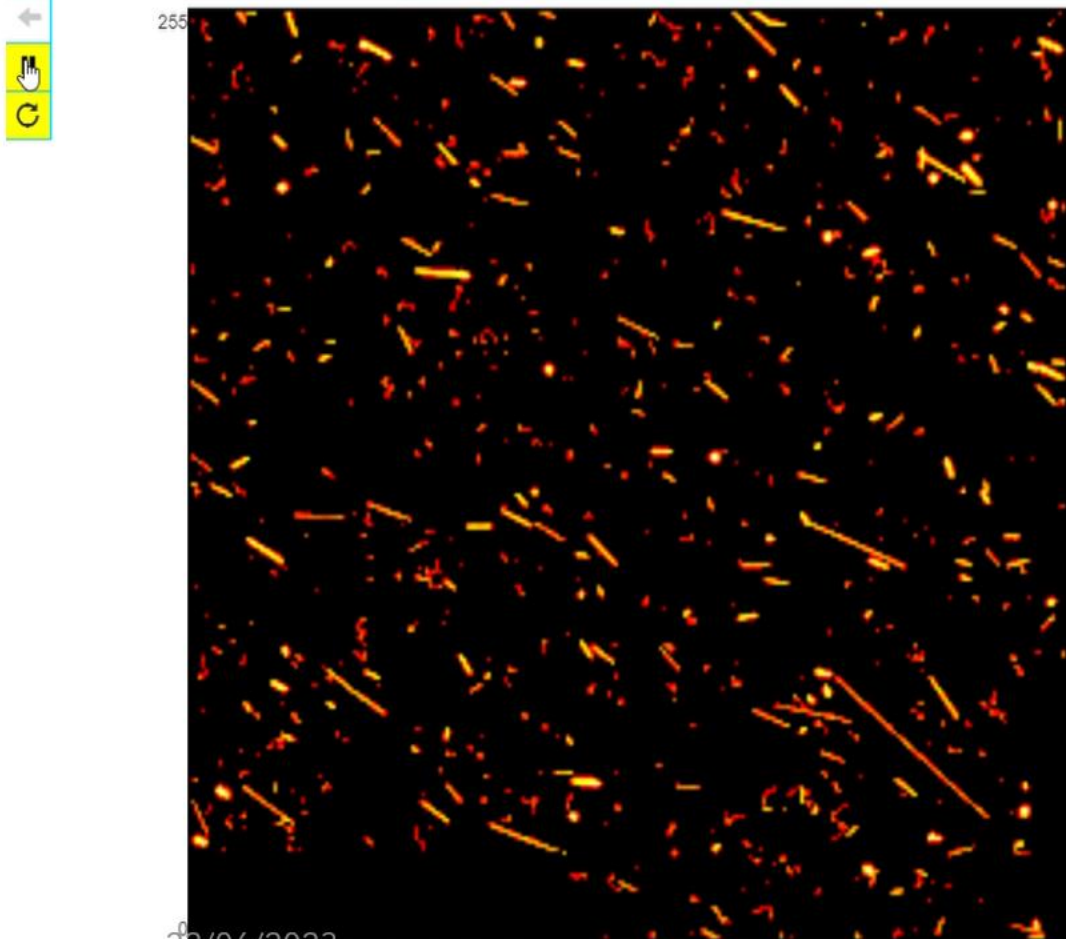
255

ASAPP2023 Perugia

54



© 2014-11-20 12:32:07 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	153	765	12.06	
Small blob	113	565	40.135	
Heavy blob	31	155	175.231	
Heavy track	22	110	335.718	
Straight track	7	35	55.771	
Curly track	226	1130	527.607	
<b>Sum:</b>	<b>552</b>	<b>2760</b>	<b>1.147e+3</b>	

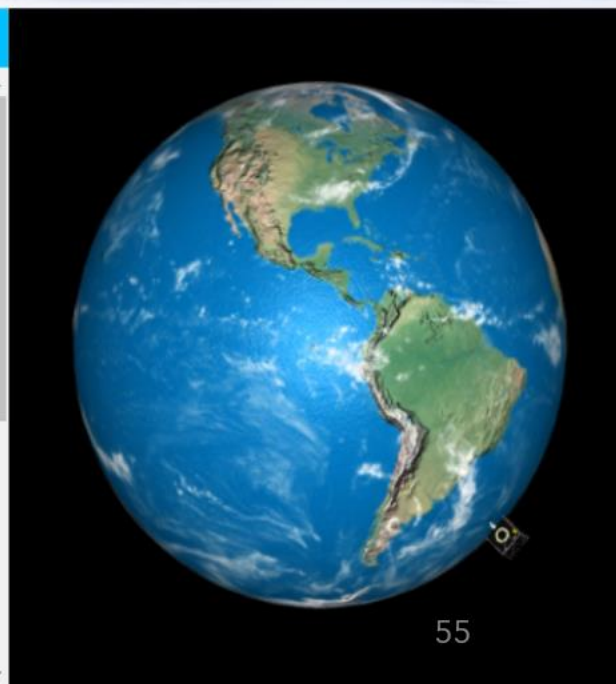
Histograms

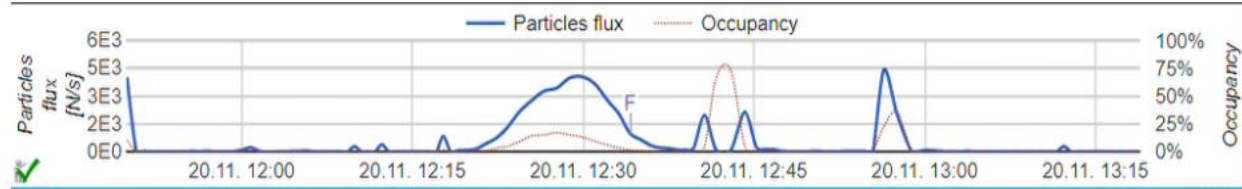
Type of histogram: Volume [keV]

Max number of bins:

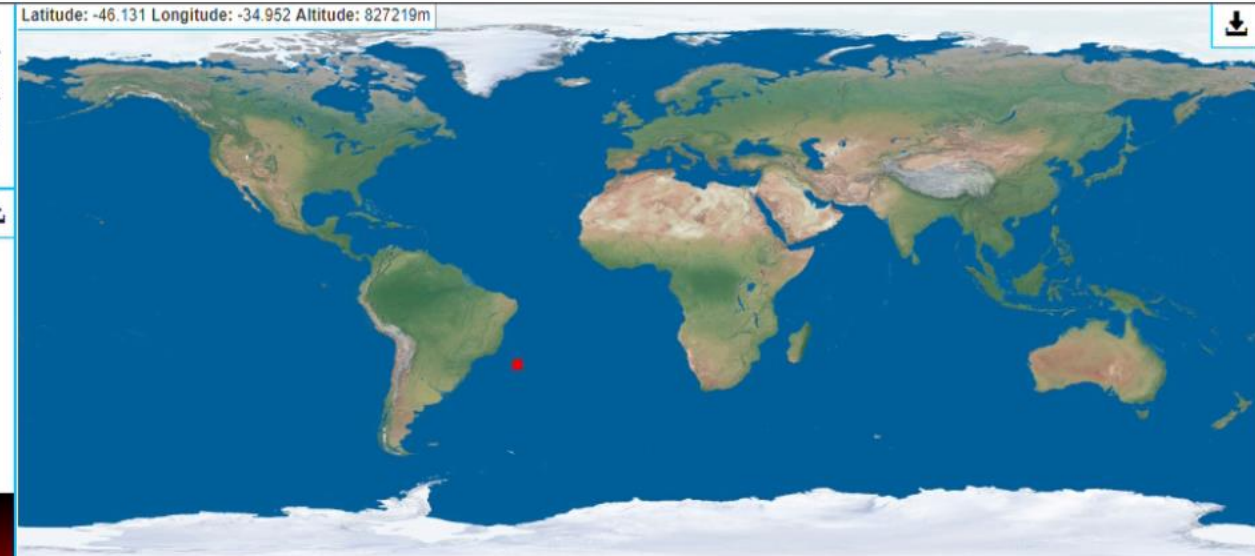
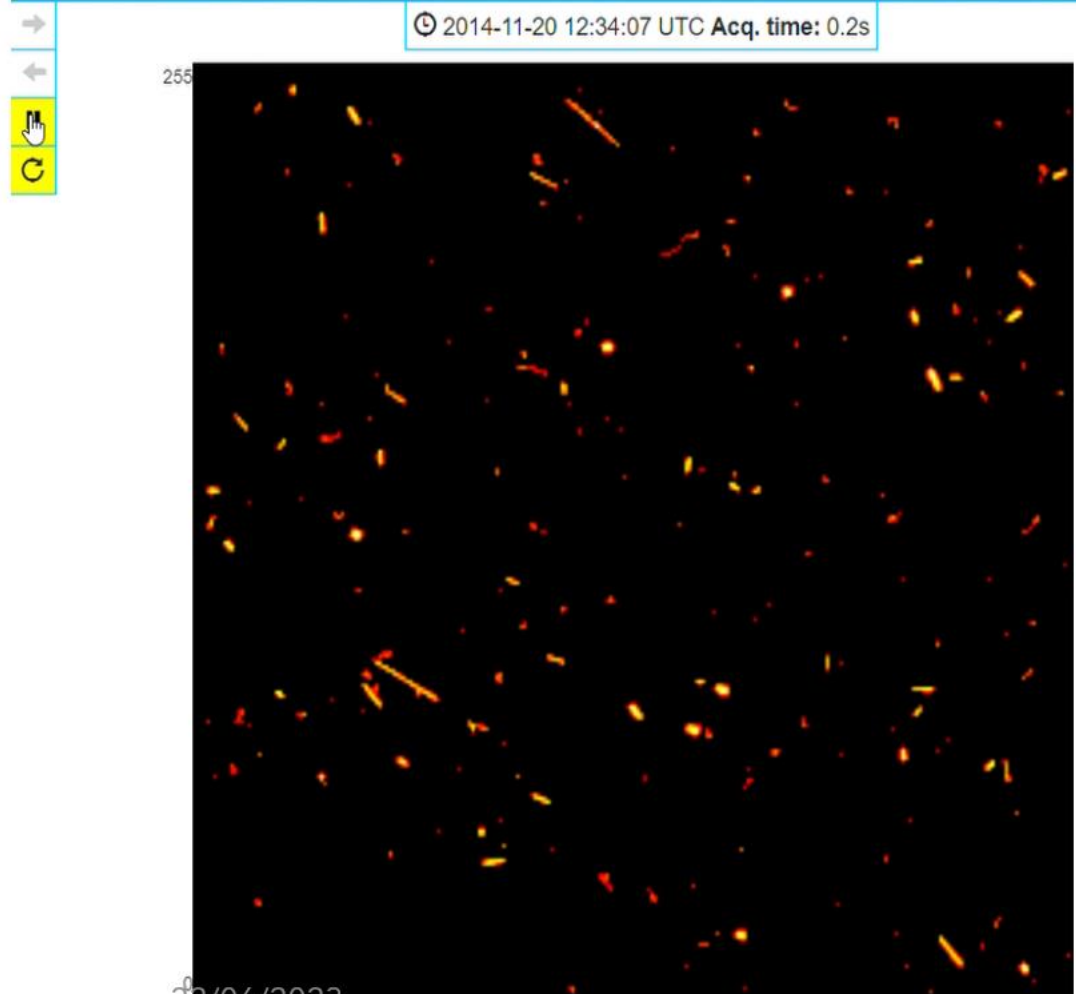
Min. value:

Max. value:





© 2014-11-20 12:34:07 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	74	370	5.94	
Small blob	34	170	9.809	
Heavy blob	17	85	110.327	
Heavy track	6	30	57.049	
Straight track	1	5	11.34	
Curly track	63	315	121.343	
<b>Sum:</b>	<b>195</b>	<b>975</b>	<b>315.807</b>	

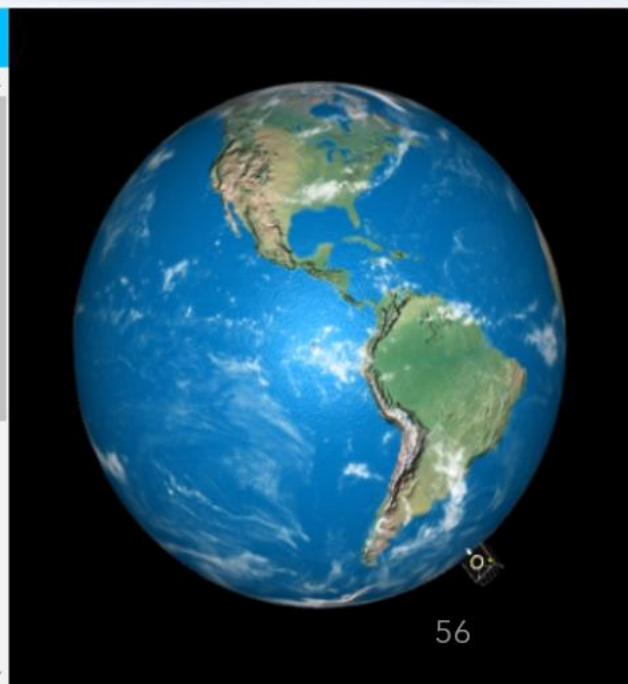
**Histograms**

Type of histogram: Volume [keV] (dropdown menu)

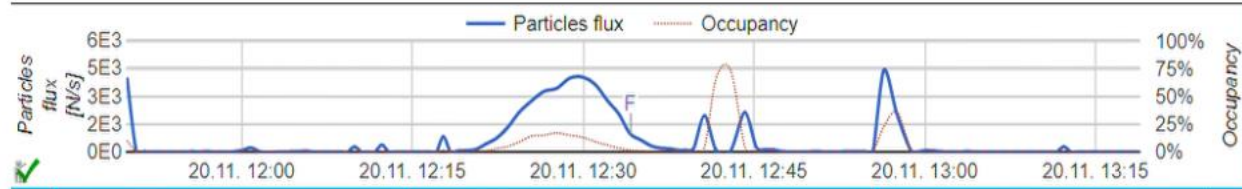
Max number of bins:

Min. value:

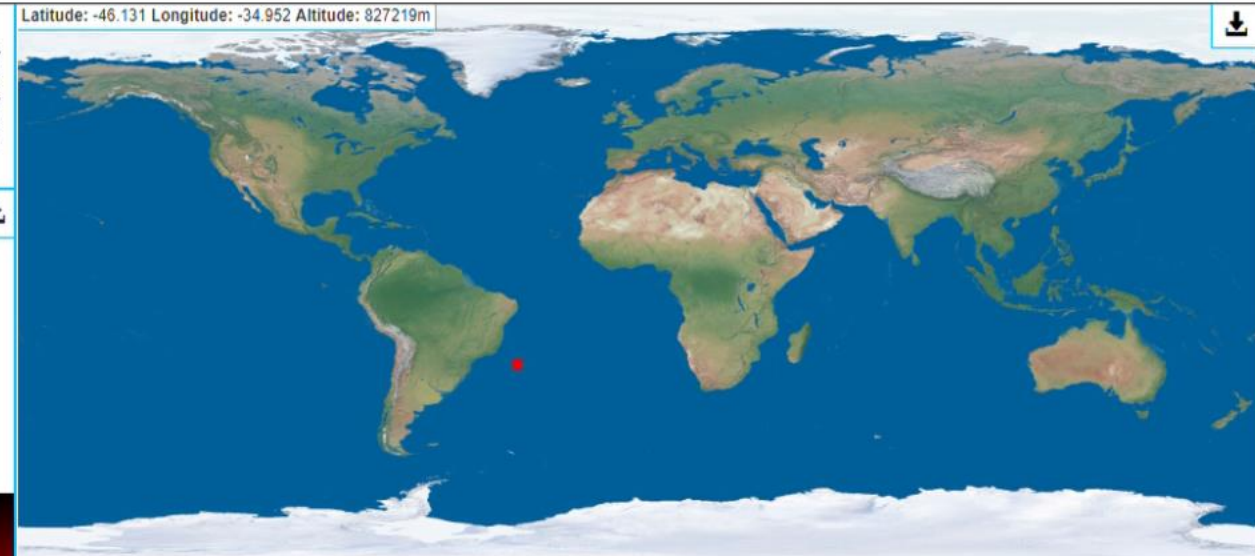
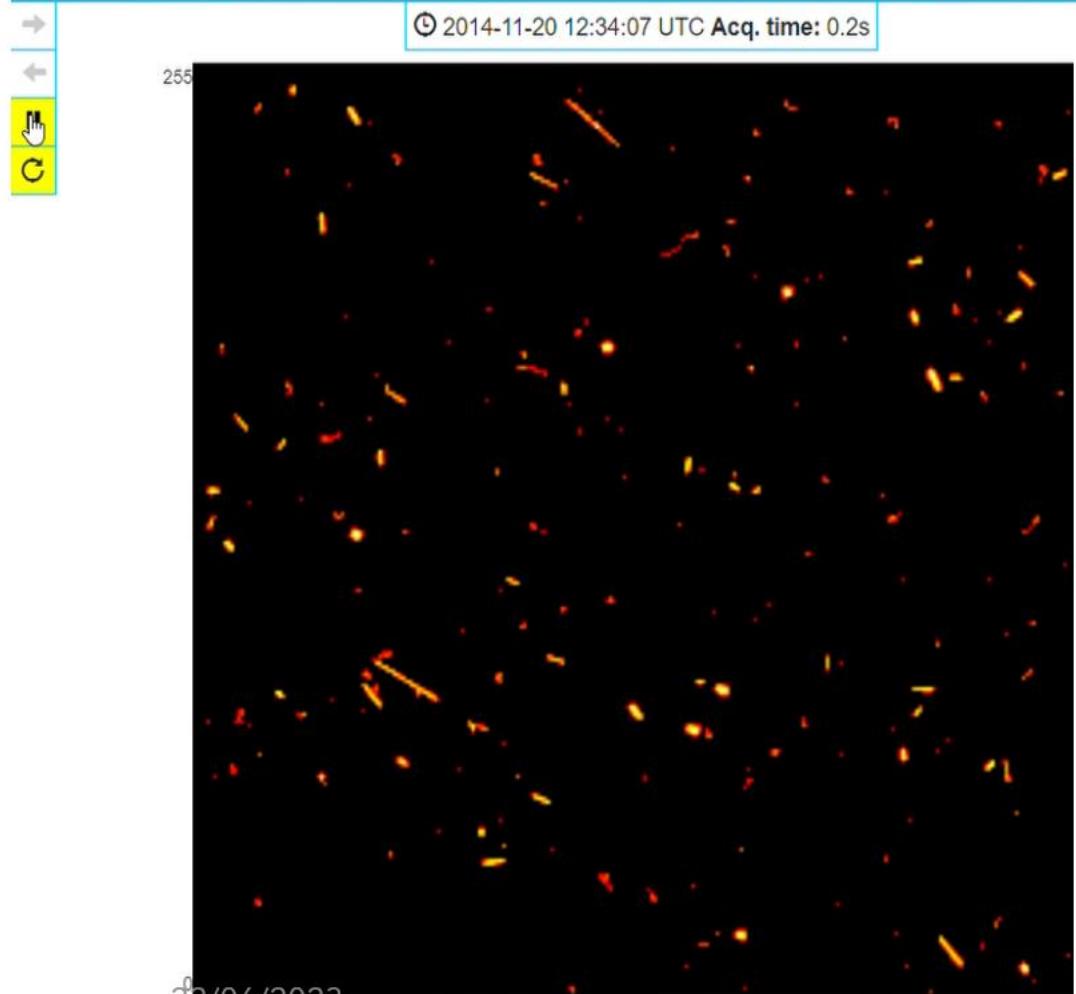
Max. value:







© 2014-11-20 12:34:07 UTC Acq. time: 0.2s



Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	74	370	5.94	
Small blob	34	170	9.809	
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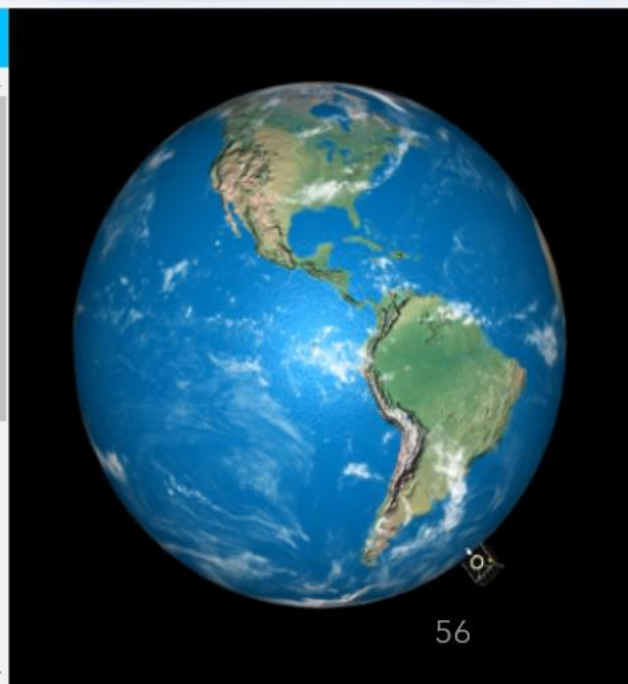
**Histograms**

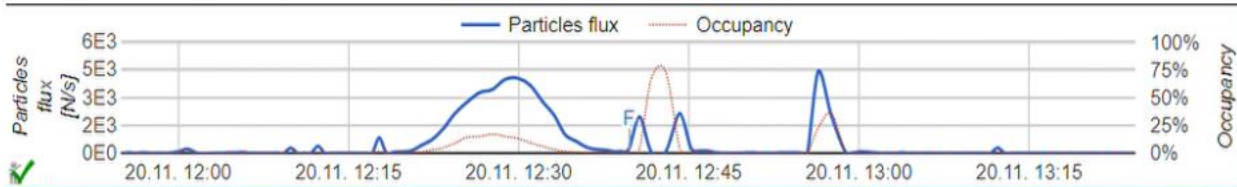
Type of histogram: Volume [keV]

Max number of bins:

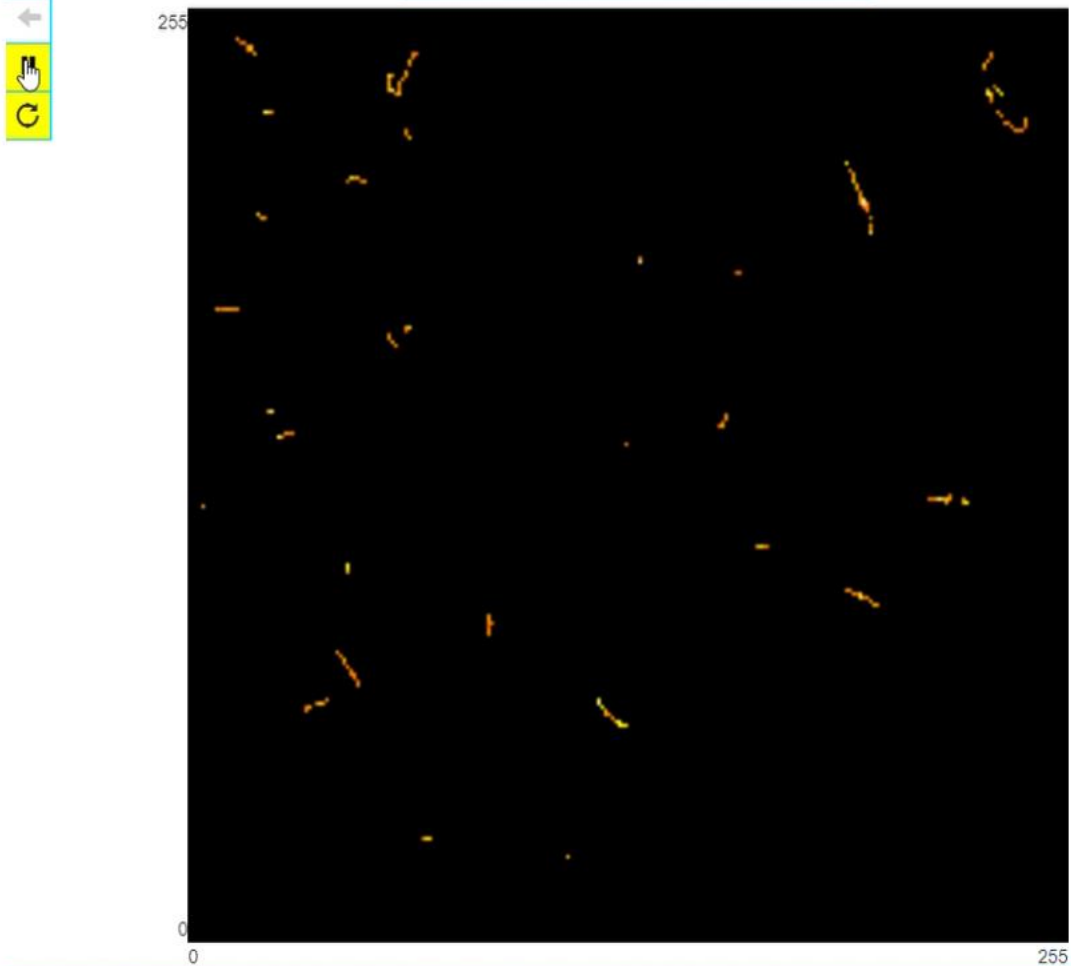
Min. value: Perugia

Max. value:

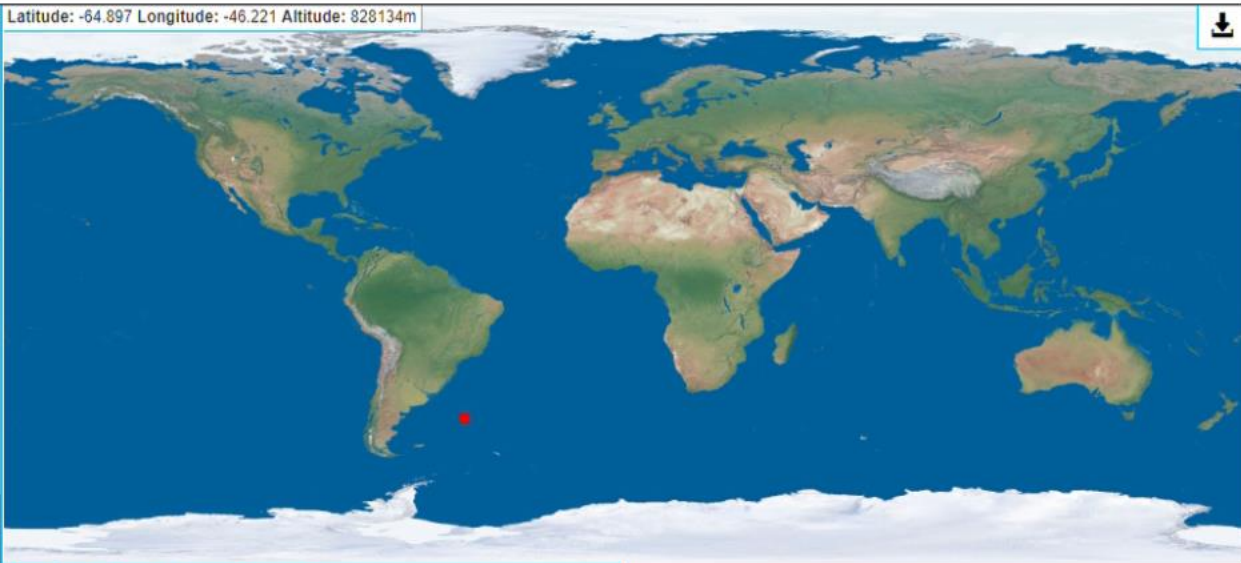




© 2014-11-20 12:39:43 UTC Acq. time: 0.2s



Status: ✓ All server queries done. (29ms)



Energy [keV] color scale: 0 (red) to 63.8 (yellow)

Cluster type	Sum	Particles flux	Energy flux [MeV]	H
Dot	5	25	0.347	
Small blob	7	35	1.44	
Heavy blob	0	0	0	
Heavy track	0	0	0	
Straight track	0	0	0	
Curly track	26	130	11.952	
<b>Sum:</b>	<b>38</b>	<b>190</b>	<b>13.74</b>	

Histograms

Type of histogram: Volume [keV] (dropdown menu)

Max number of bins:

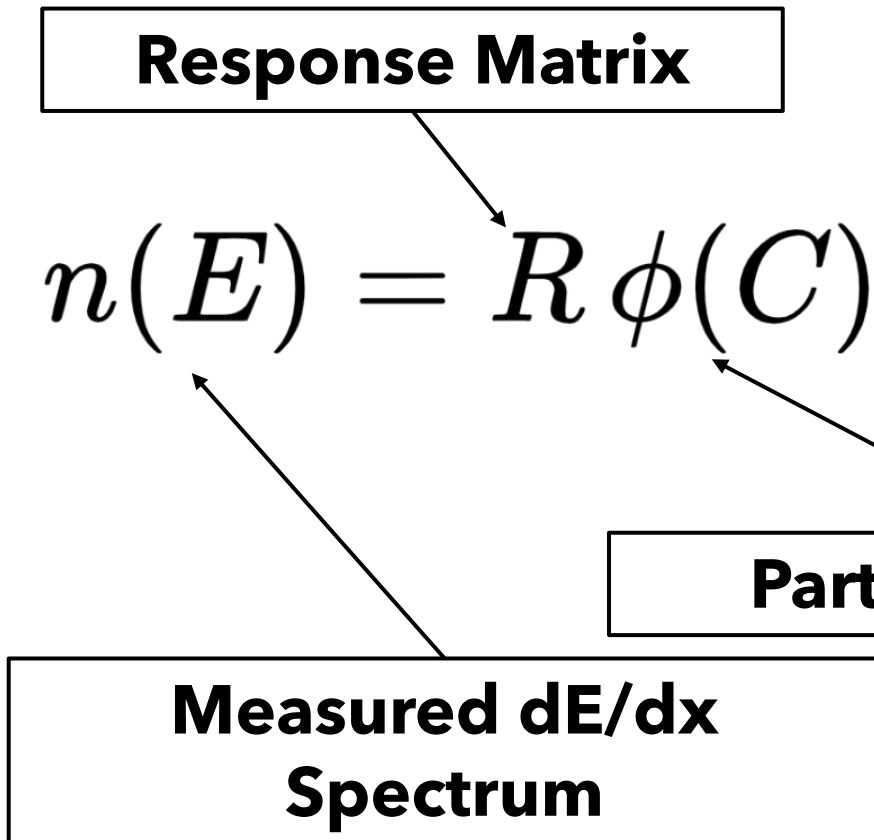
Min. value:

Max. value:

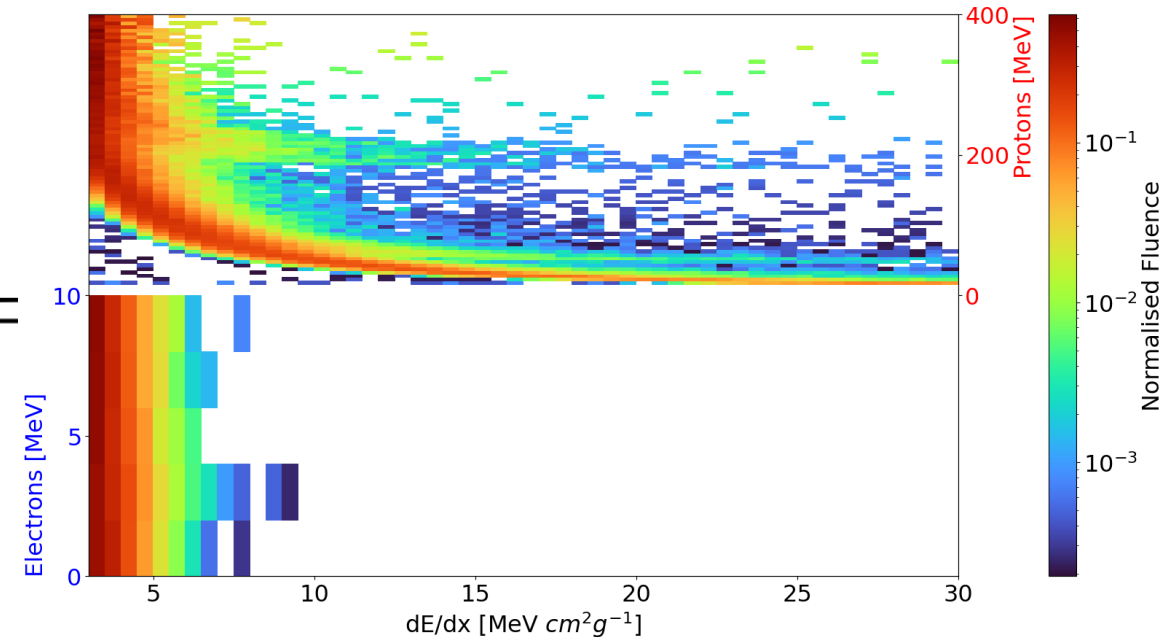


# Particle Classification - Bayesian Deconvolution

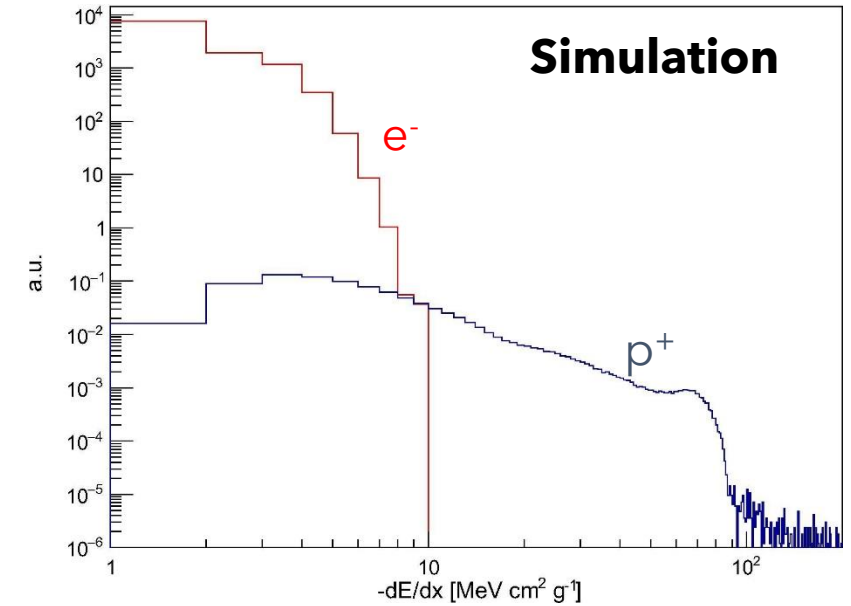
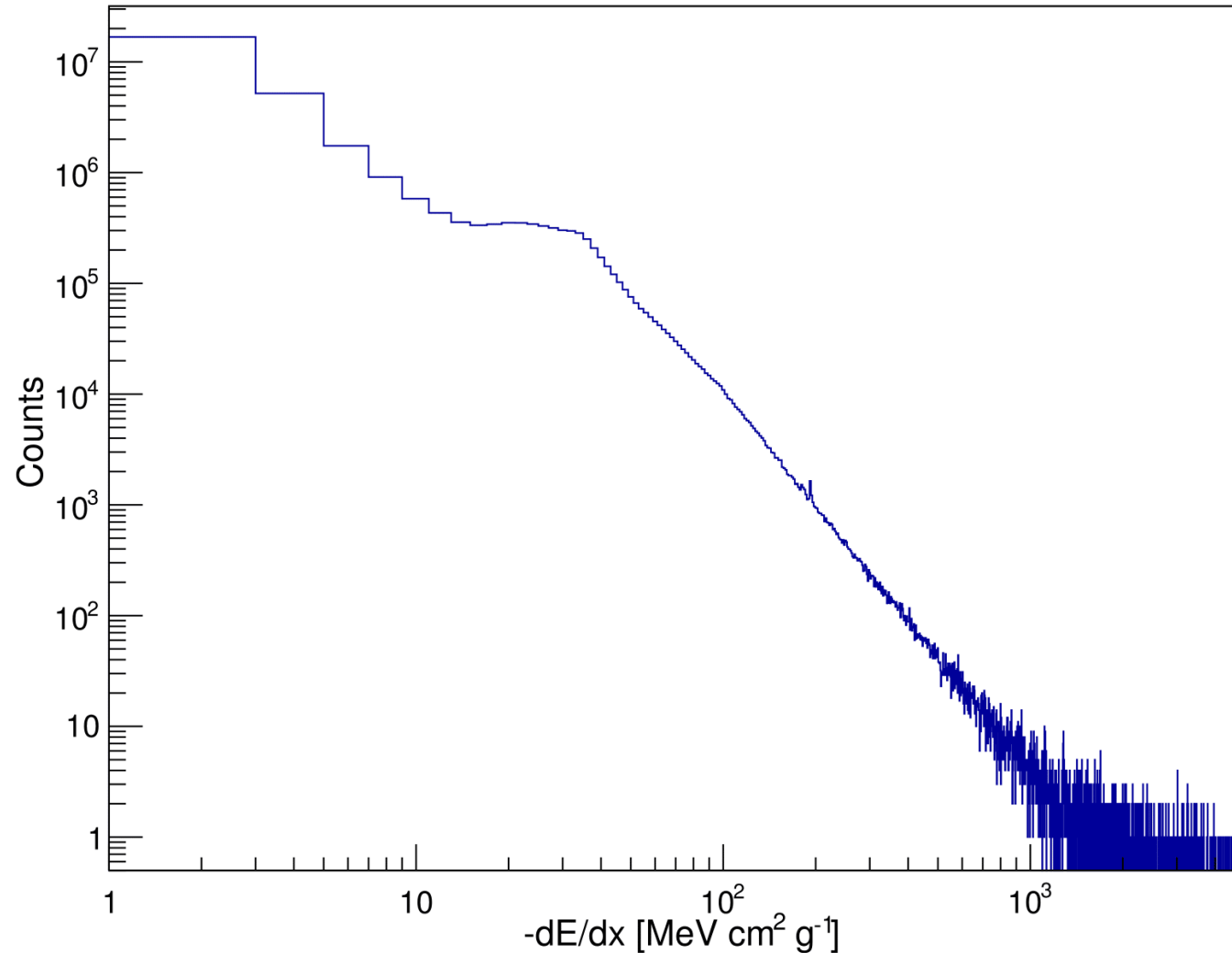
This method works by decomposing the stopping power "signal" of the field into its contributing particle signals, from which the particle's distributions can be inferred



$$R^T =$$

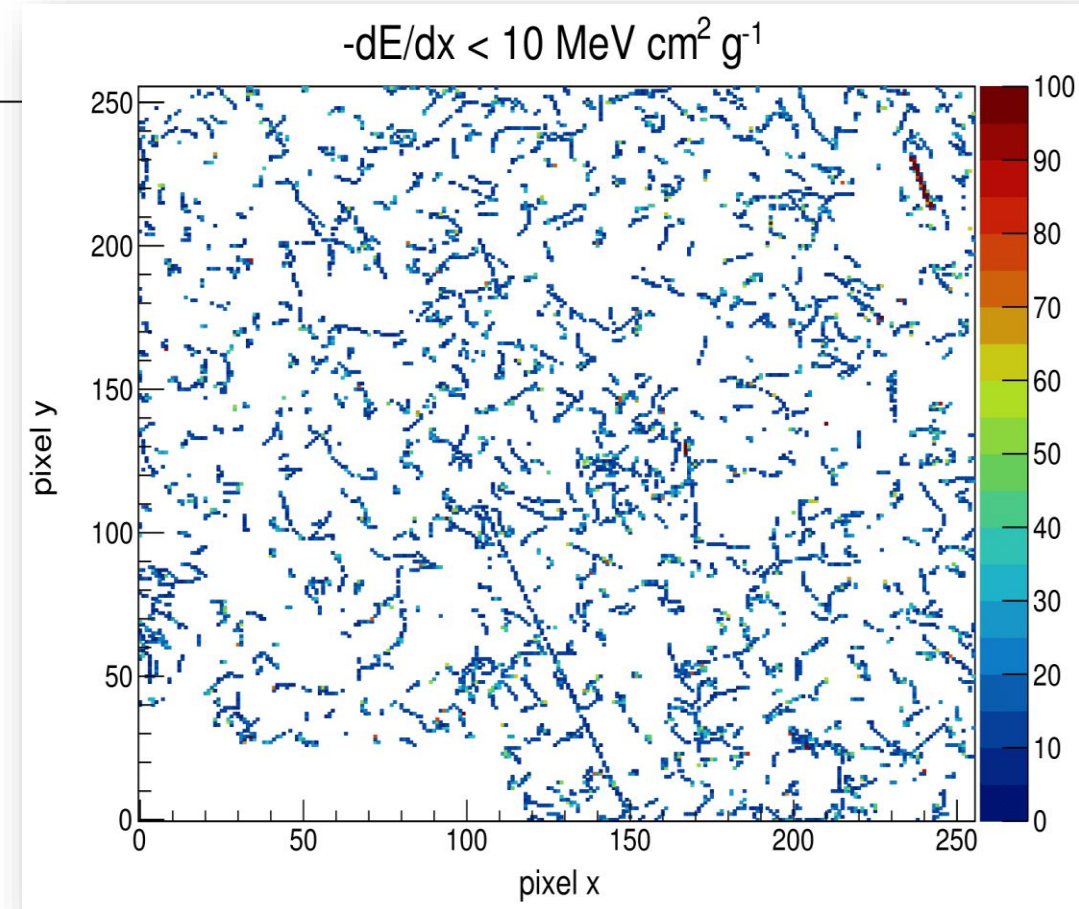
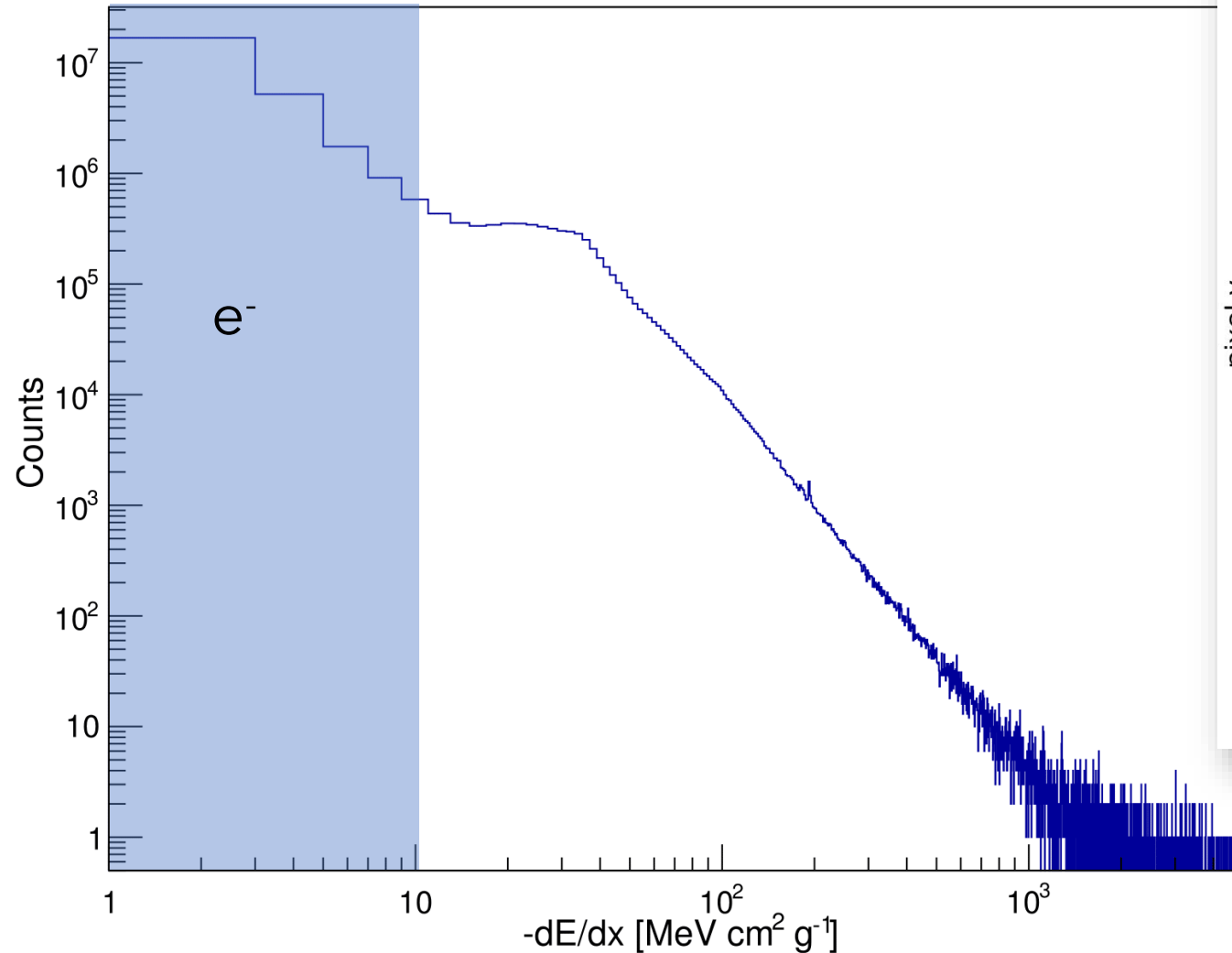


# dE/dX and Particle Classification



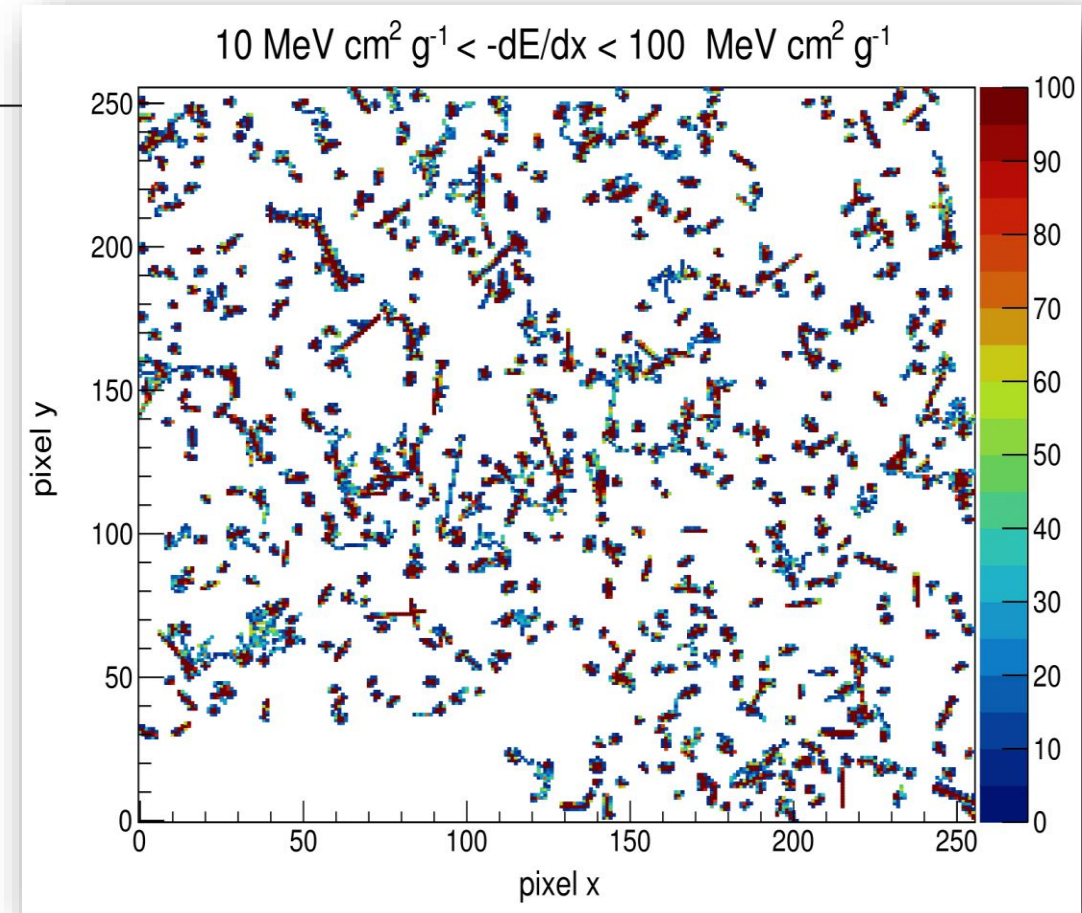
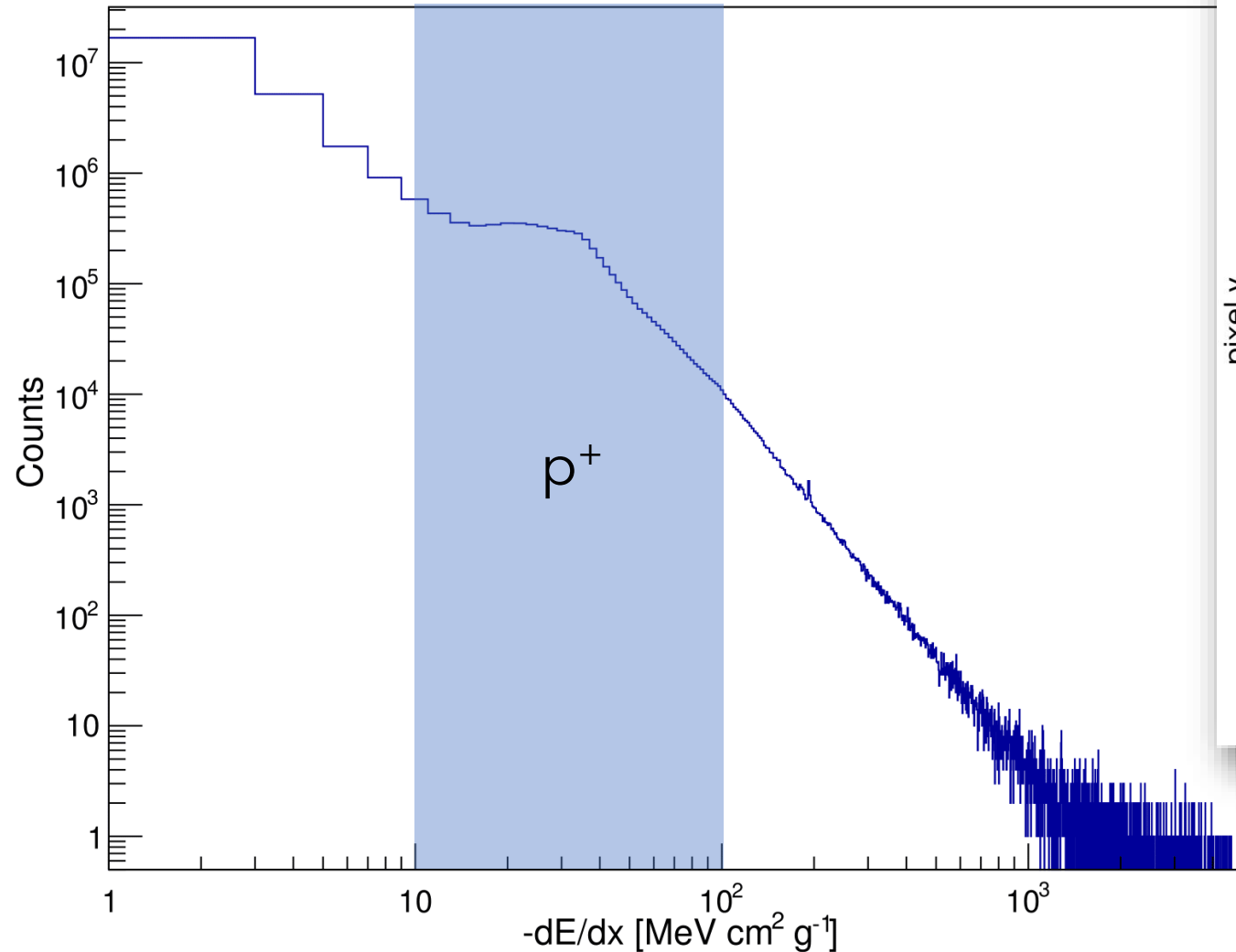
St. Gohl et al., "Study of the radiation fields in LEO with the Space Application of Timepix Radiation Monitor (SATRAM)", *Advances in Space Research* **63**, Issue 5, pp. 1646-1660, (2019).

# dE/dX and Particle Classification



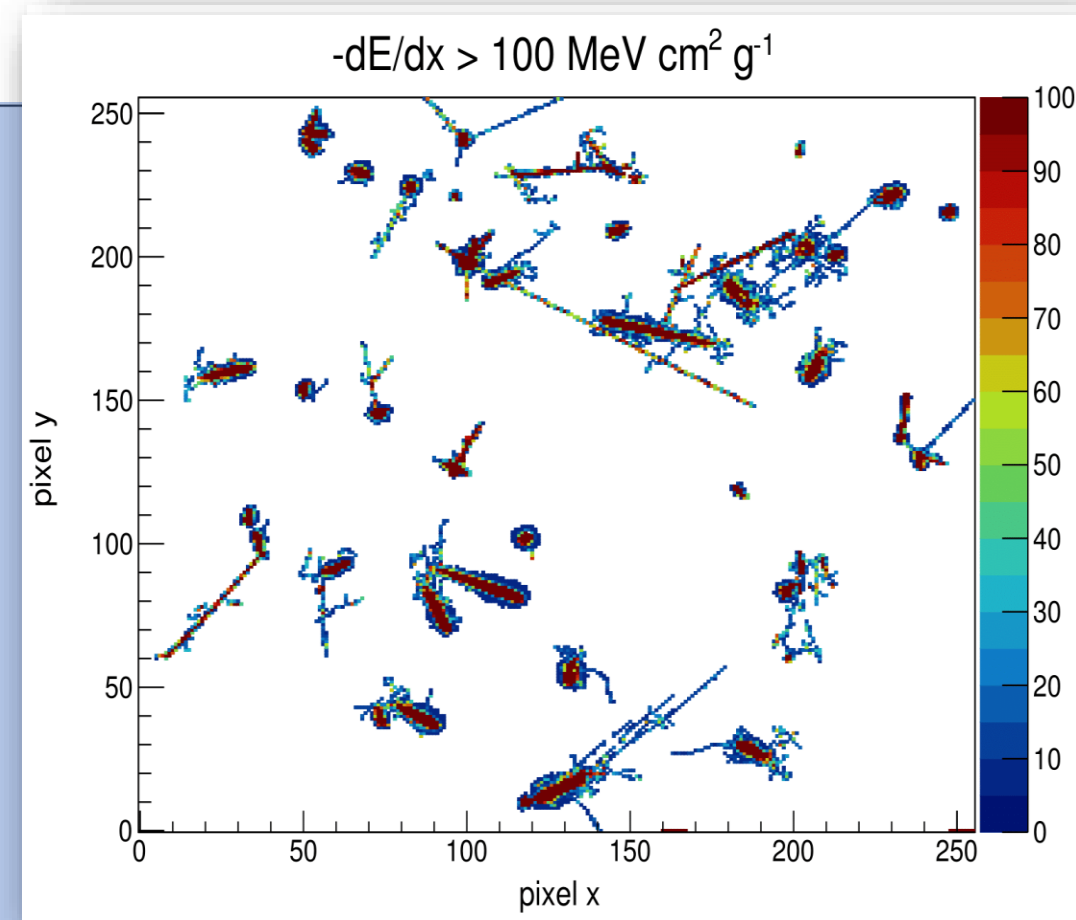
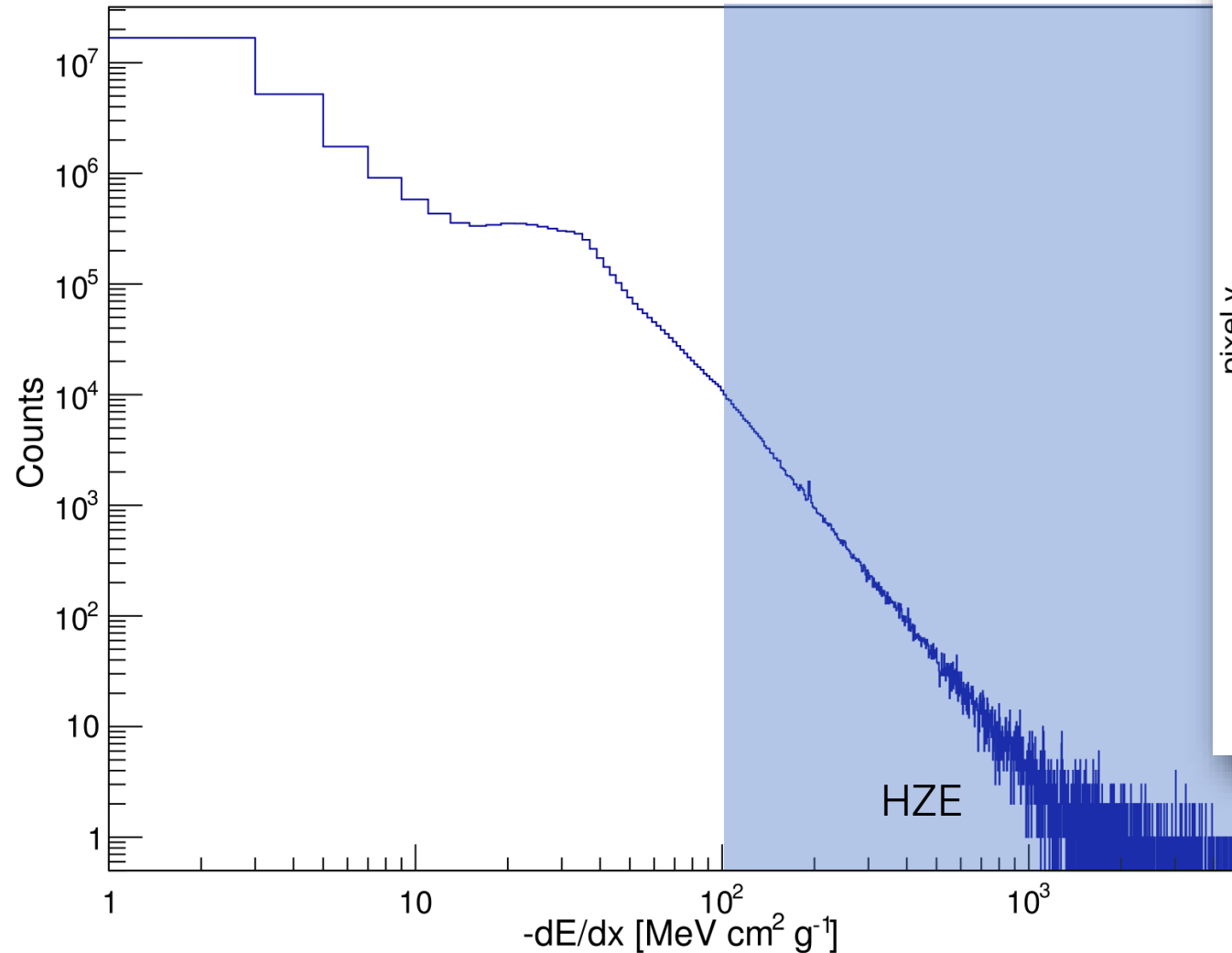
St. Gohl et al., "Study of the radiation fields in LEO with the Space Application of Timepix Radiation Monitor (SATRAM)", *Advances in Space Research* **63**, Issue 5, pp. 1646-1660, (2019).

# dE/dX and Particle Classification



St. Gohl et al., "Study of the radiation fields in LEO with the Space Application of Timepix Radiation Monitor (SATRAM)", *Advances in Space Research* **63**, Issue 5, pp. 1646-1660, (2019). 62

# dE/dX and Particle Classification



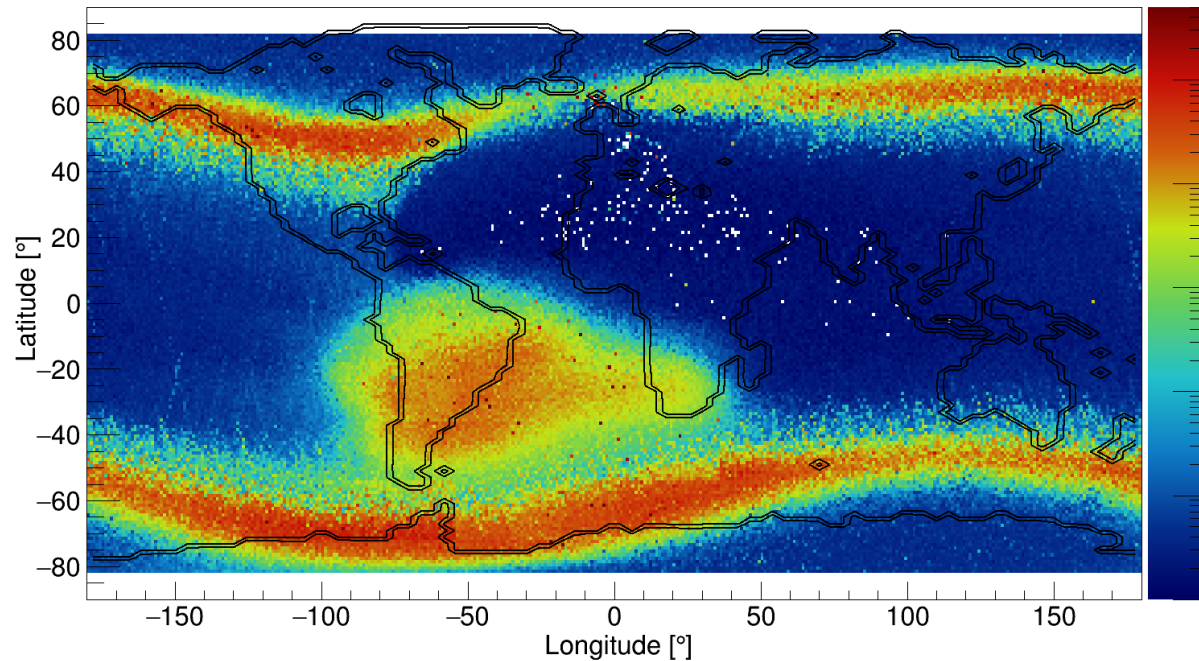
St. Gohl et al., "Study of the radiation fields in LEO with the Space Application of Timepix Radiation Monitor (SATRAM)", *Advances in Space Research* **63**, Issue 5, pp. 1646-1660, (2019).

# Electron and Proton Flux Maps

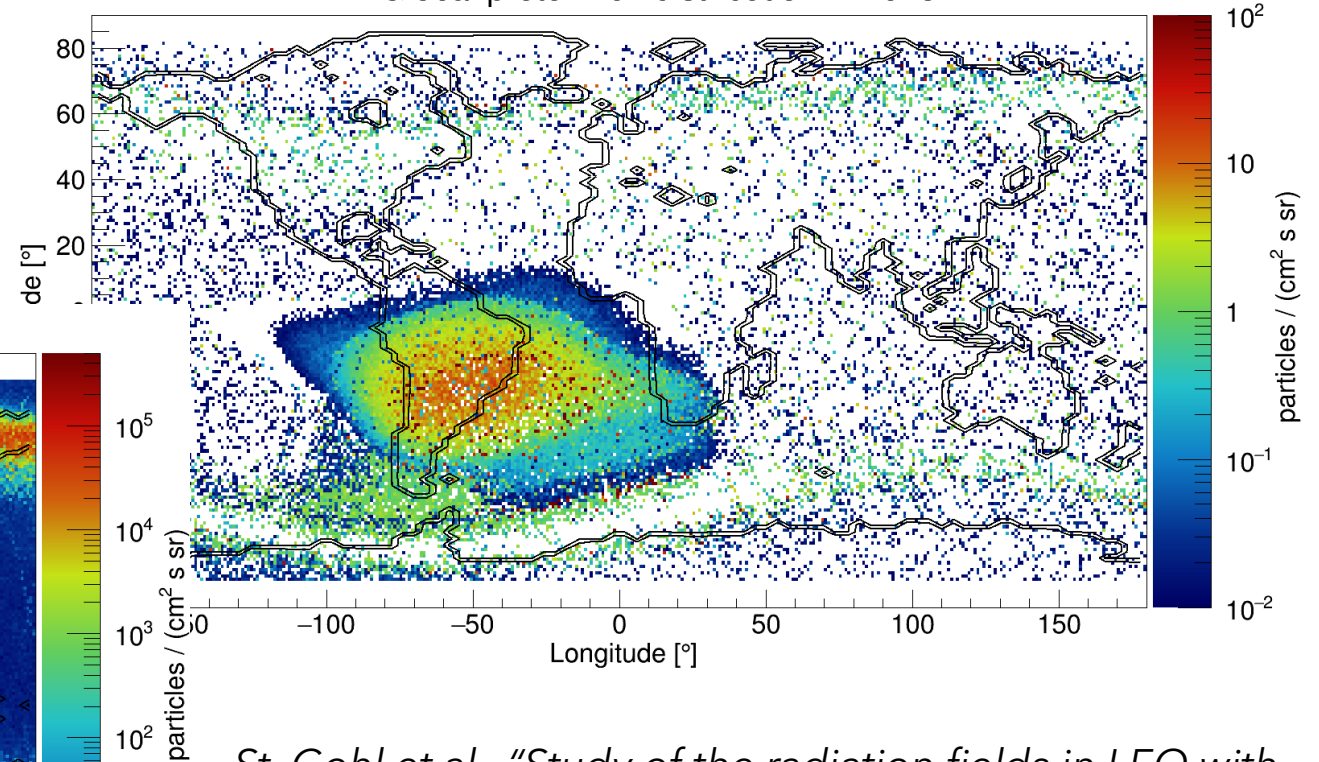
$e^-$  fluxes 3 orders of magnitude larger than  $p^+$  fluxes

→ Even small  $e^-$  misclassification distorts  $p^+$  flux measurement

Global electron flux distribution in 2015



Global proton flux distribution in 2015

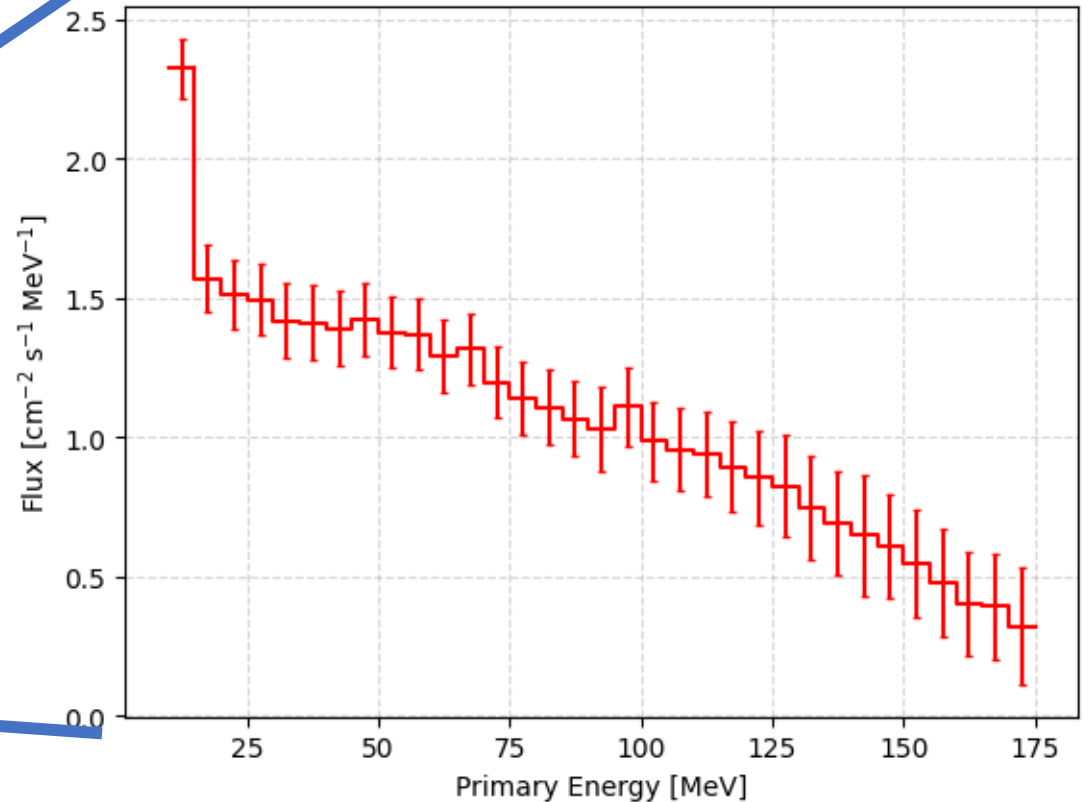
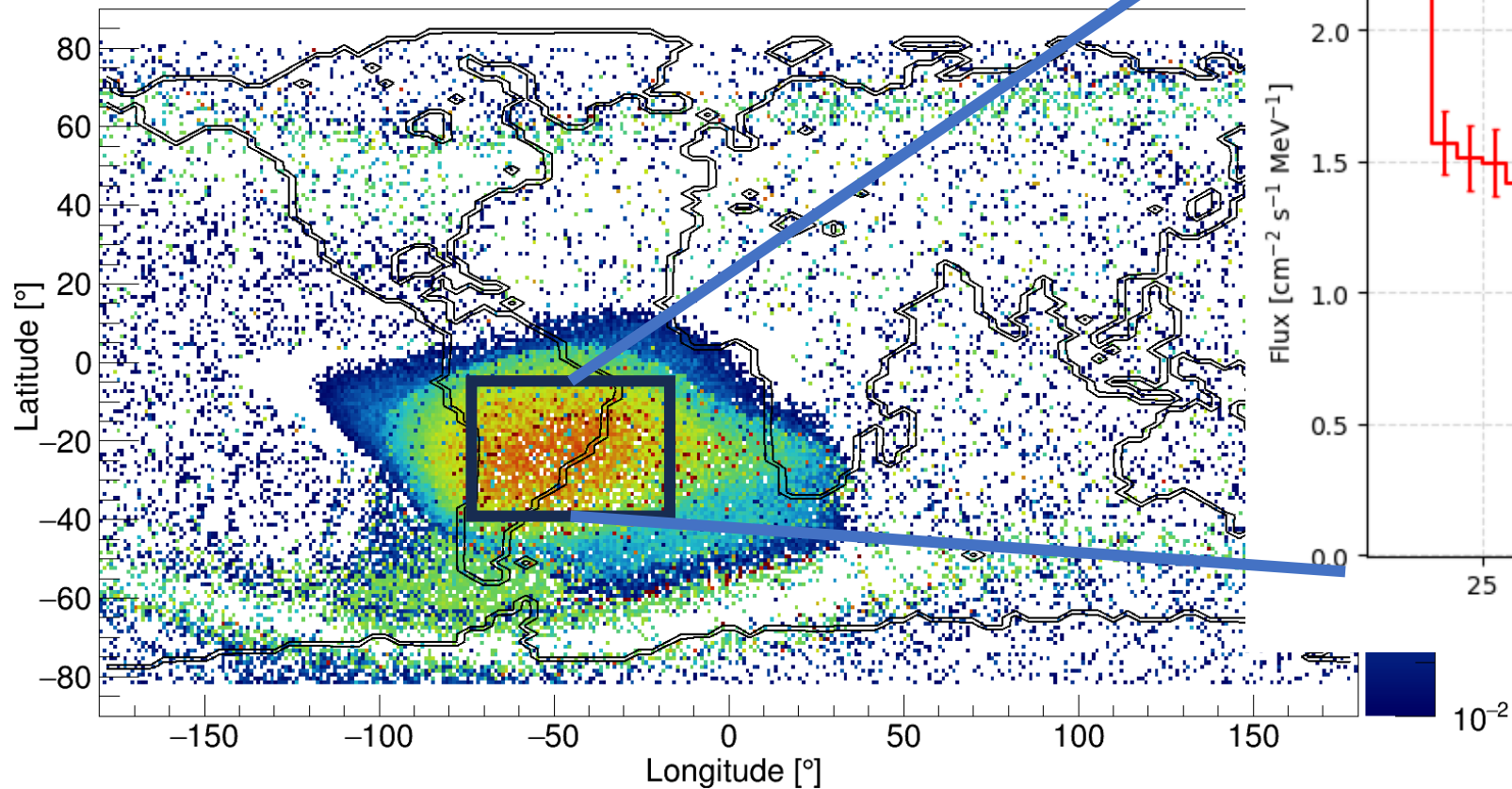


St. Gohl et al., "Study of the radiation fields in LEO with the Space Application of Timepix Radiation Monitor (SATRAM)", *Advances in Space Research* **63**, Issue 5, pp. 1646-1660 (2019).



# First Measurement of the Trapped Proton Energy Spectrum with a Single-Layer Detector

Global proton flux distribution in 2015



# **Stopping Power Classification at MoEDAL**

# A Closer look at Stopping Power

## Region 1:

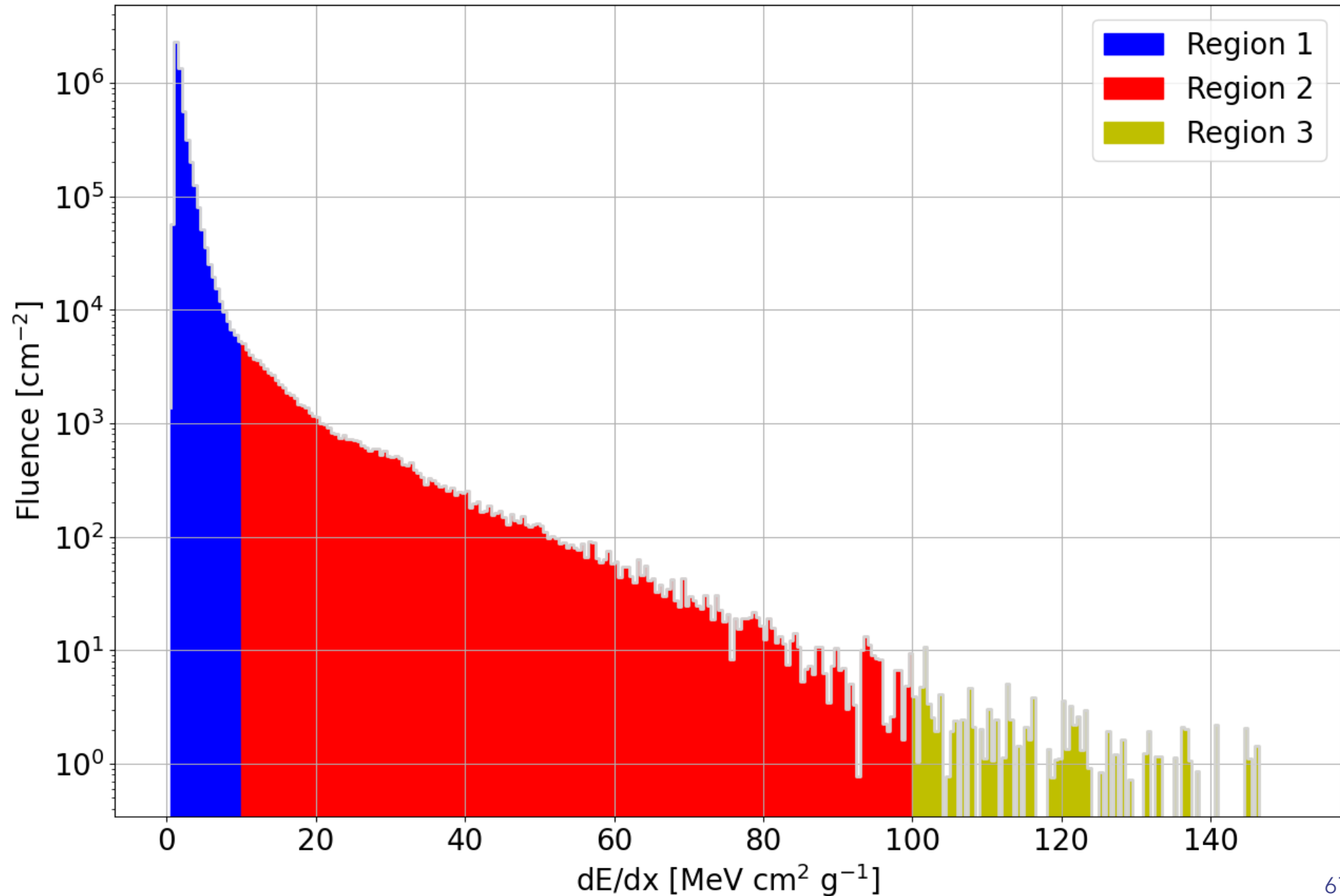
- Electrons
- Gammas
- Singly charged relativistic particles

## Region 2:

- Protons <250 MeV
- Pions <15 MeV
- Muons <10 MeV

## Regions 3:

- Fragmentations
- $Z > 1$  particles



# A Closer look at Stopping Power

## Region 1:

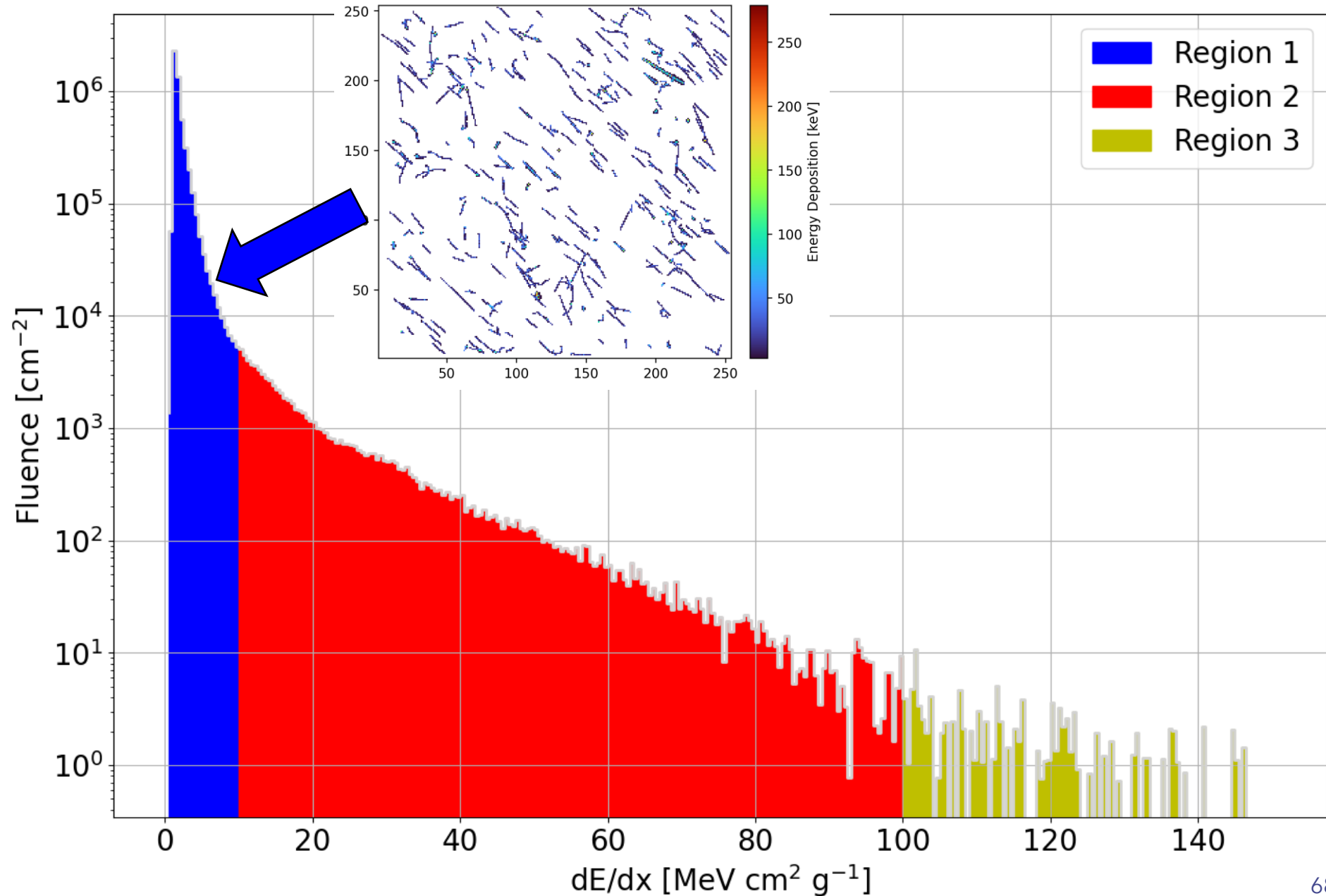
- Electrons
- Gammas
- Singly charged relativistic particles

## Region 2:

- Protons <250 MeV
- Pions <15 MeV
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# A Closer look at Stopping Power

## Region 1:

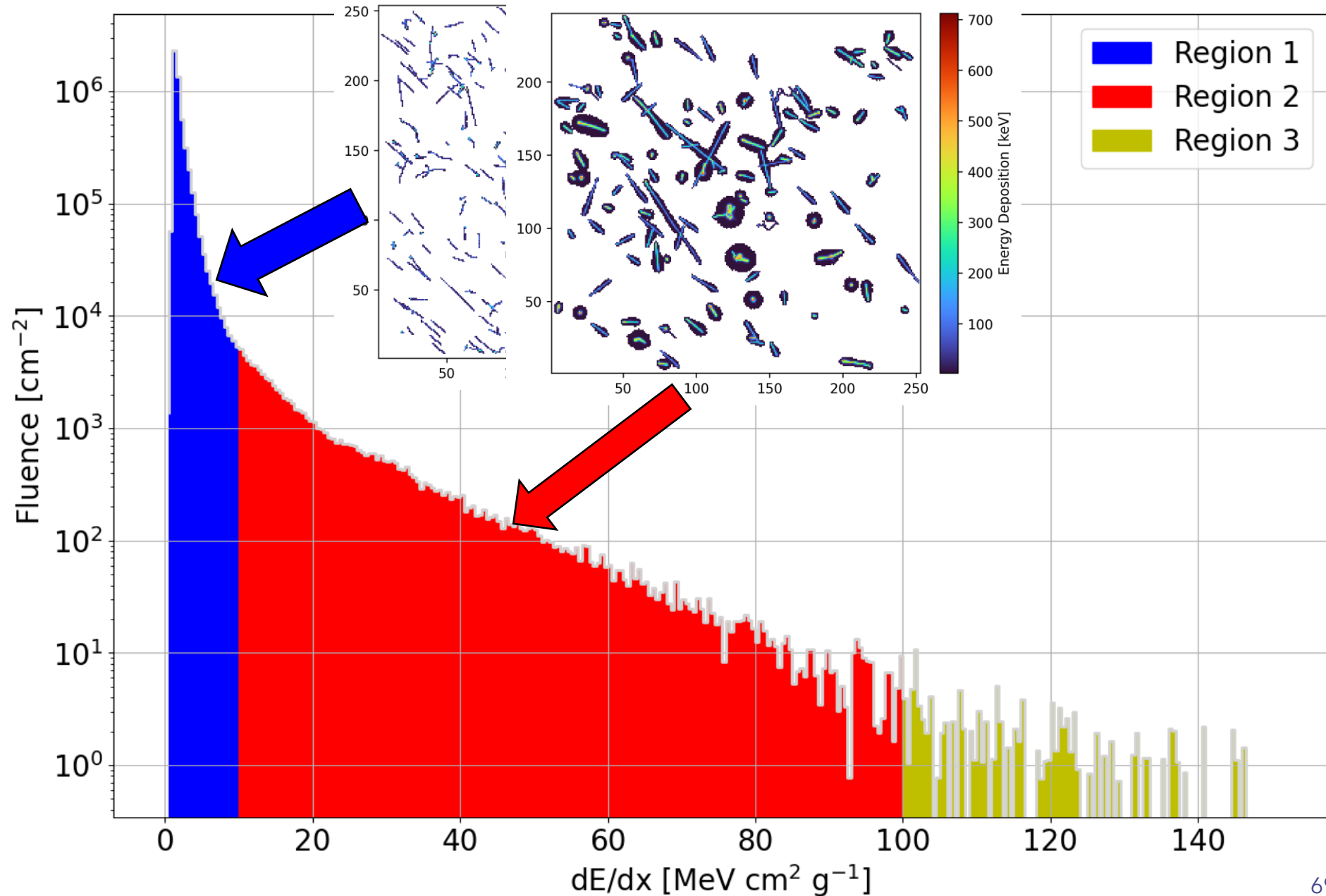
- Electrons
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# A Closer look at Stopping Power

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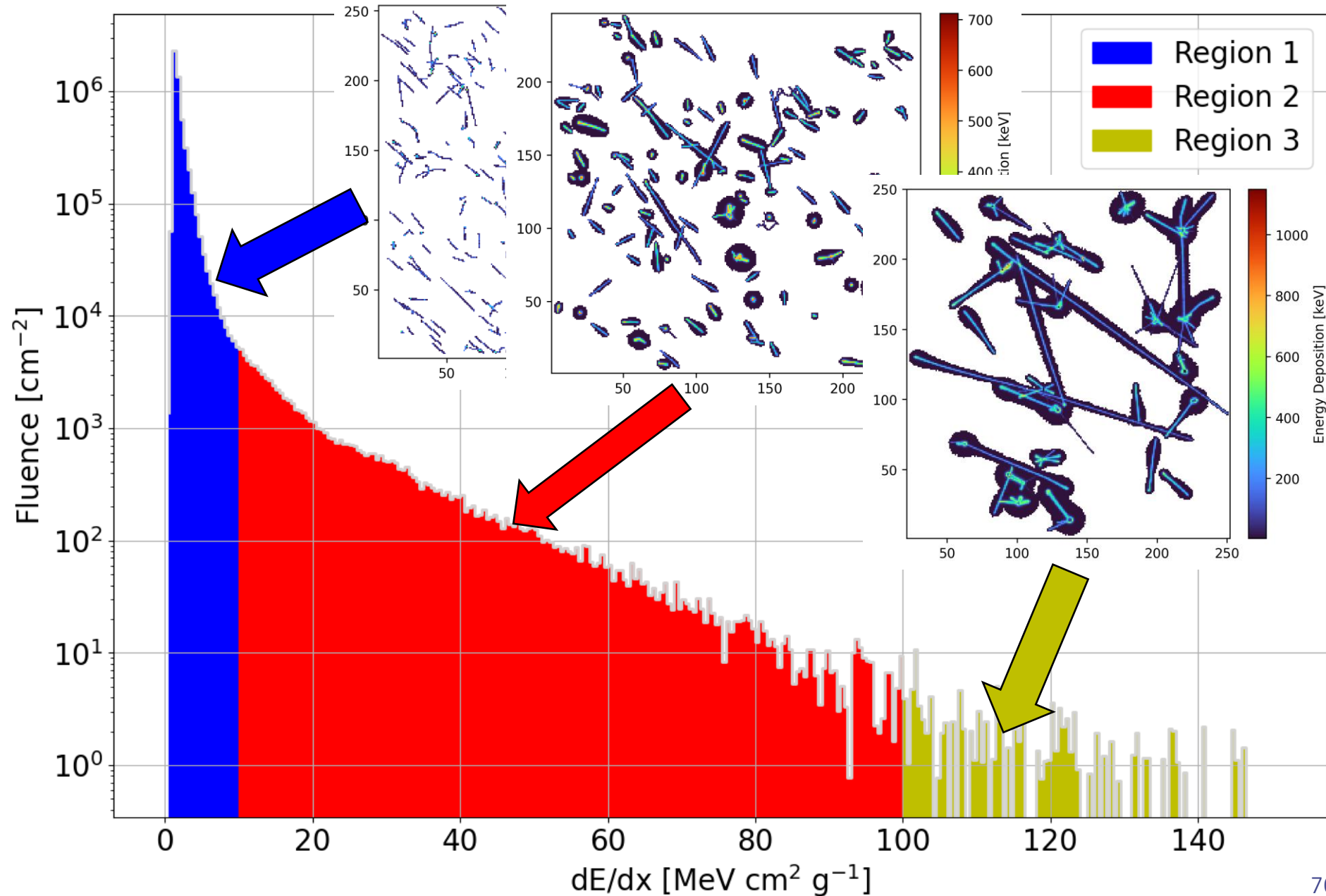
- Electrons
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- Singly charged relativistic particles

## Region 2:

- Protons <250 MeV
- Pions <15 MeV
- Muons <10 MeV

## Regions 3:

- Fragmentations
- $Z > 1$  particles



# Conclusions

- 3D tracking
  - ➔ The capabilities of reconstructing particle trajectories from IP8 in real time and with single layer Timepix3 detectors have been demonstrated
- Stopping power and radiation field decomposition
  - ➔ The current are state-of-the-art algorithms in particle classification were shown and validated in the environment found in low earth orbit
  - ➔ Future work will be to improve particle classification by using advanced machine learning algorithms

# Can we use Timepix3 in Search for the Avatars of New Physics?

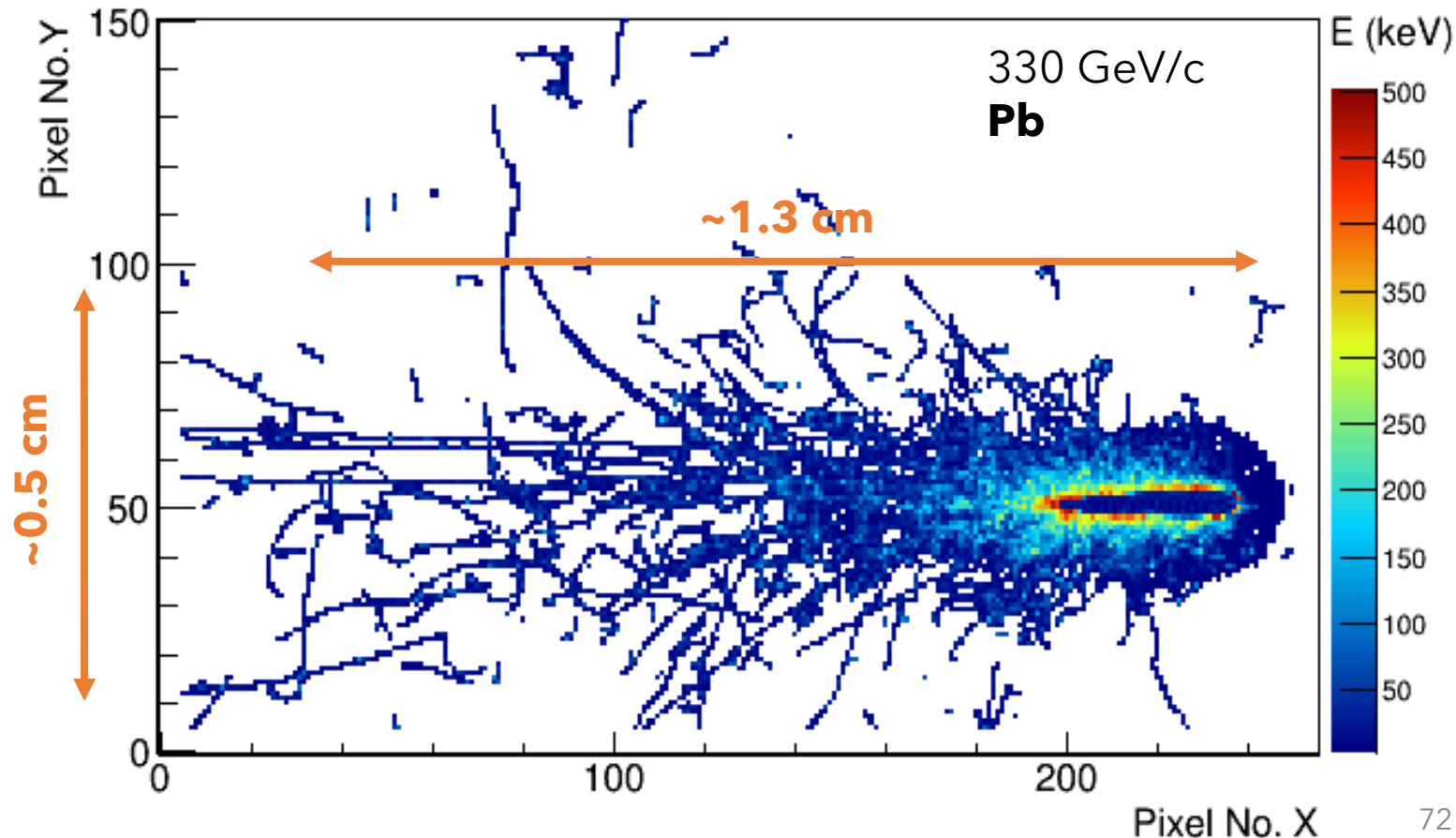


$$\left\langle -\frac{dE}{dx} \right\rangle = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[ \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 W_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$$

Highly charged energetic particles are known to produce distinct cluster pattern within Timepix3

Associated with:

- ➔ High Stopping Power
- ➔ Large number of delta rays





**Thank you for your attention**

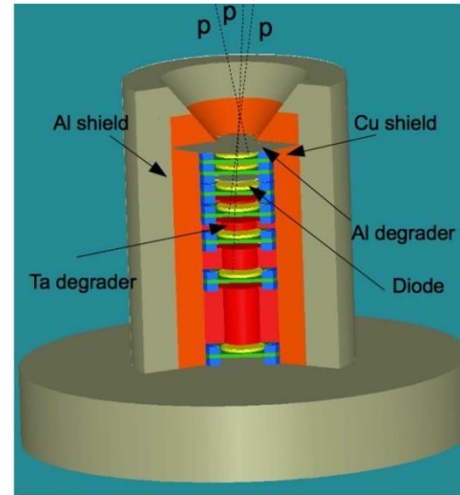
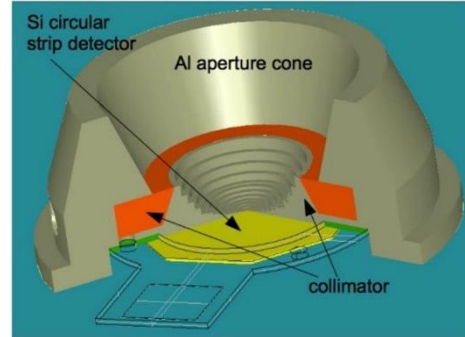
**This research was funded by the Czech Science Foundation grant number  
GM23-04869M.**

# Instruments for measurements in LEO



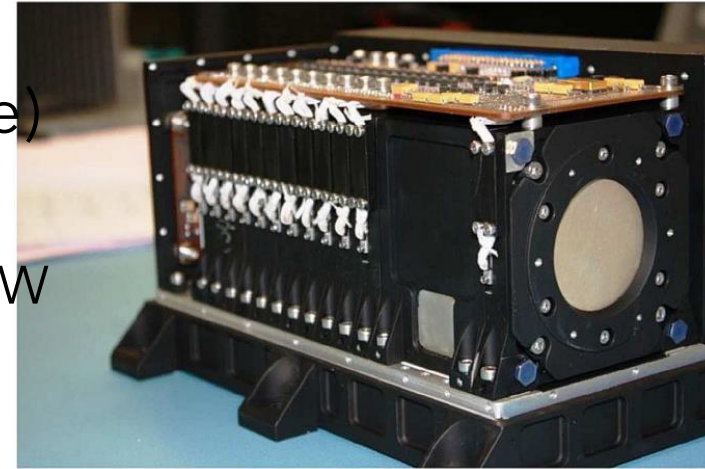
Next Generation Radiation Monitor (NGRM)

- Mass ~ 1 kg
- Consumption ~ 1-2 W



EPT (Energetic Particle Telescope)

- Mass: **4.6 kg**
- Consumption: 5.6 W



ICARE-NG:

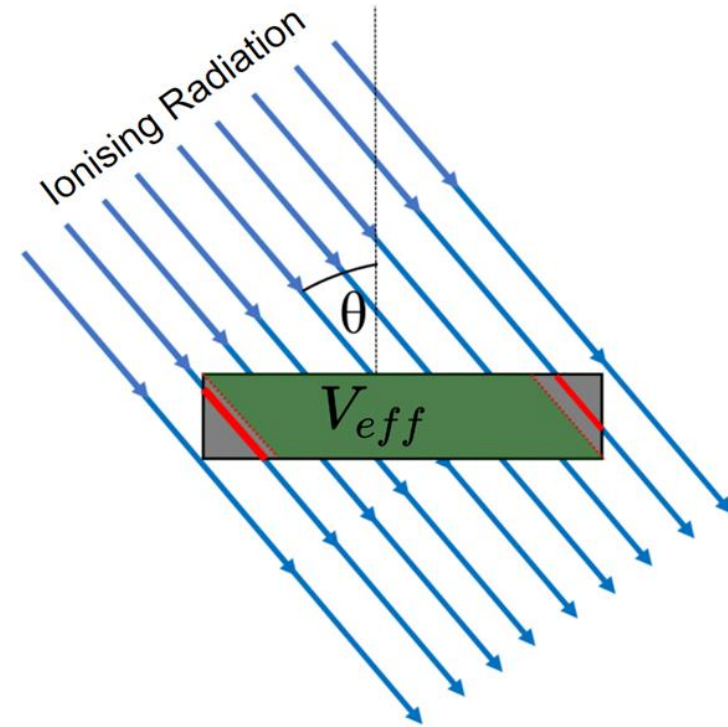
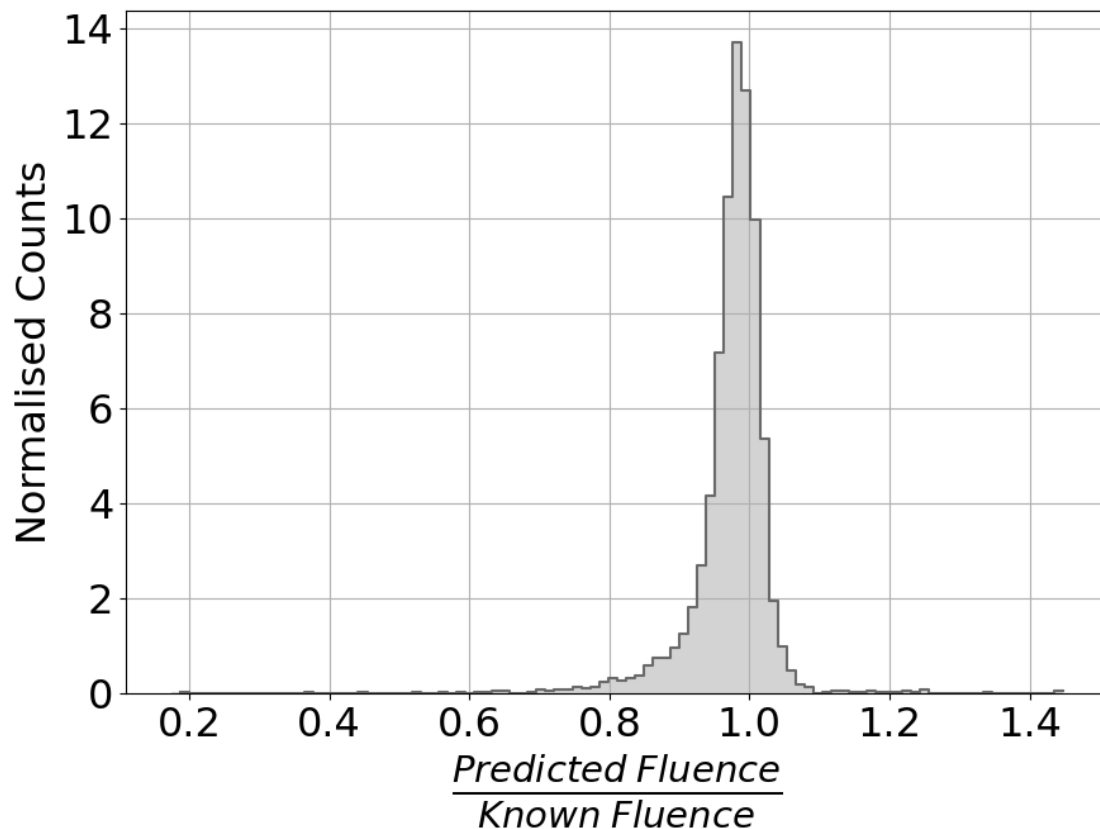
- Mass: **2.4 kg**
- Consumption: 3 W



# Particle Fluence

The formula was derived analytically to account for the removal of edge clusters

$$F = \sum_{i \in \{\text{particles}\}} \frac{d / \cos(\theta_i)}{l \cdot d \cdot (l - d \cdot \tan(\theta_i))}$$

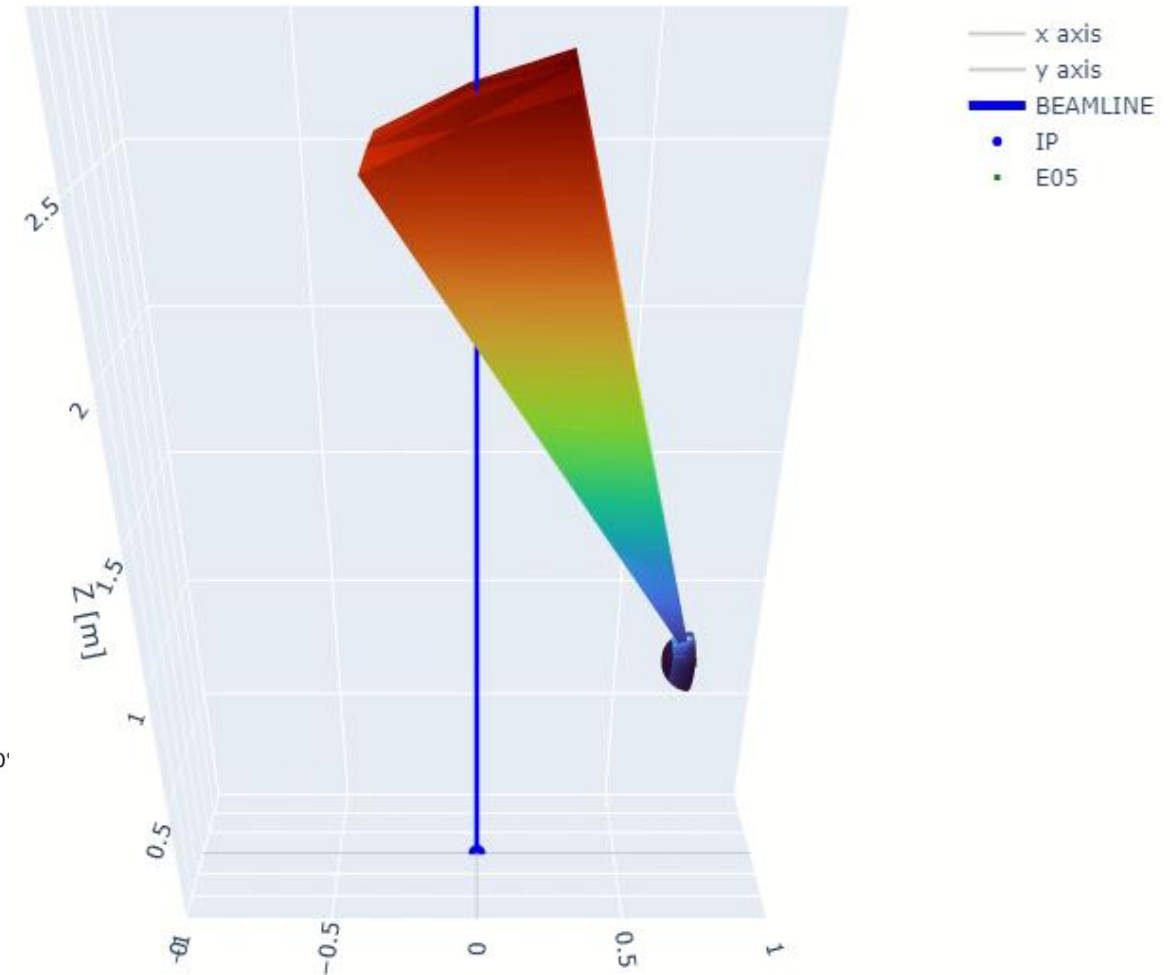
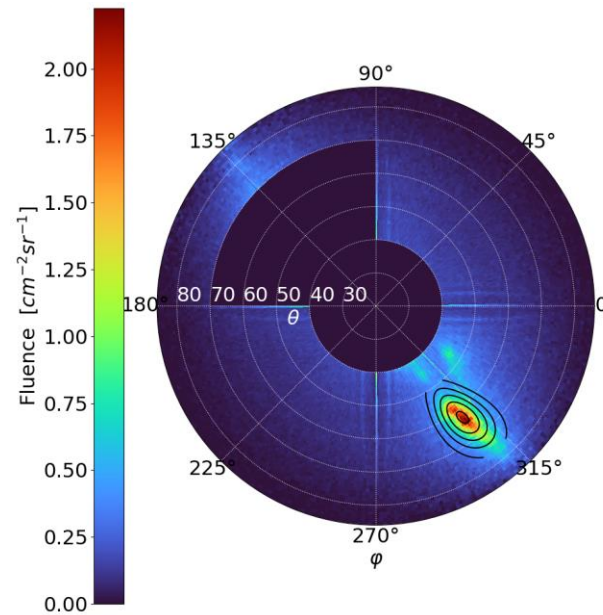


The new algorithm was tested in a Geant4 omnidirectional field simulation.

Testing Accuracy: 97%

# Second Peak Localisation

- Follow correct alignment of the interaction of point
- It becomes possible to project the second peak in 3D space to determine its origin
- Projection shows the origin along the beam line further the theory of scattering beam due to wider beam



# Bayesian Deconvolution Mathematical Background

Response Matrix

$$n(E) = R \phi(C)$$

Particle Spectrum

Measured dE/dx Spectrum

$$M \approx R^{-1} \Rightarrow \phi(C) = M n(E)$$

Bayesian Formula  $\Rightarrow P(C_\mu | E_j) = \frac{P(E_j | C_\mu) P(C_\mu)}{\sum_\nu^{n_C} P(E_j | C_\nu) P(C_\nu)}$

Probability of a cause (particle class present)  
given an effect (measurement)

$$\Rightarrow \phi(C_\mu) = \sum_{i=1}^{n_E} P(C_\mu | E_i) n(E_i)$$