# The MAGO cavity and prospects for HFGW searches

Krisztian Peters CERN, 4. Dec. 2023

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CLUSTER OF EXCELLENCE QUANTUM UNIVERSE





## **GW interaction with SRF cavities**

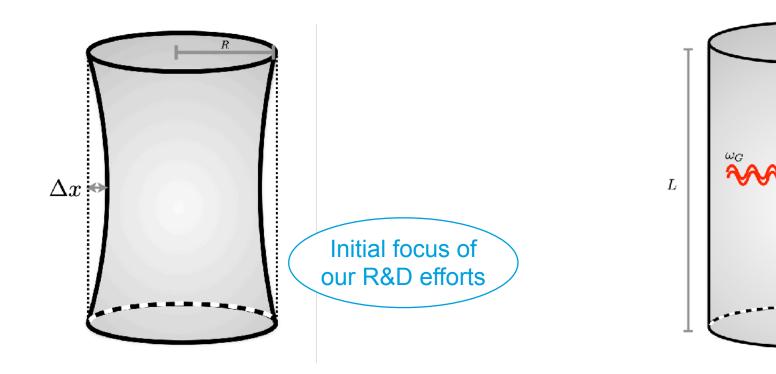
**Two detection principles** 

#### **Mechanical coupling**

• GWs perturb detector, spreading power in frequency space

#### **EM-coupling**

Graviton - photon mixing (Gertsenshtein effect)
 GWs induce in a magnetic field an effective current

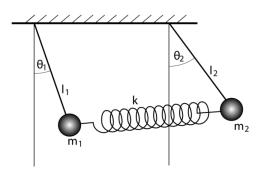


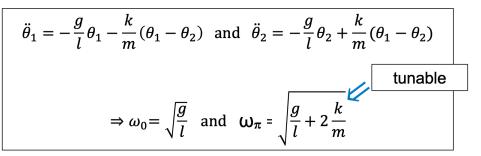
## **Heterodyne detection**

#### GWs induce energy transfer between two levels of an EM resonator

632 mm







#### Two EM levels achieved by coupling identical cavities

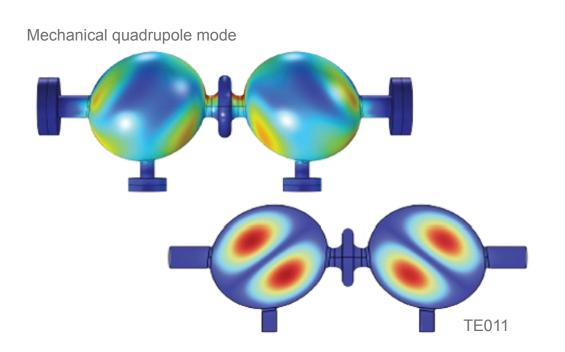
• Each resonant mode of the individual cavities is split in two modes of the coupled resonator with different spacial field distribution ( $\omega_0$  and  $\omega_{\pi}$ , symmetric and anti-symmetric modes)

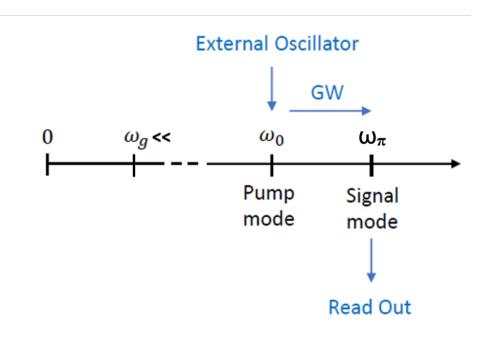
#### $\Delta \omega = \omega_{\pi} - \omega_0$ is tunable by "spring constant" k, which is given by the geometry of the tuning cell

• Original plans for this cavity were to tune the distance between the two cells with piezo elements

## **Heterodyne detection**

#### **Further requirements**





**GW couples to modes with quadruple symmetry:** mechanical I=2 mode and slightly deformed spheres to induce orthogonal field polarisations such that two EM levels in TE011 mode have I=2 difference GW with  $\omega_g$  transfers energy from  $\omega_0$  to  $\omega_0+\omega_g$ . Resonantly enhanced if  $\omega_g = \omega_{\pi} - \omega_0$ 

## The MAGO proposal

... and the PACO cavities

#### Initial idea from the 70s, which led to the MAGO proposal

JETP LETTERS	VOLUME 13, NUMBER 11	5 JUNE 197
HIGH-FREQUENCY DETECTION OF GRAVITATIONAL WAVES		
V. B. Braginskii a Physics Department Submitted 18 March	t, Moscow State University	
	3, No. 11, 585 - 587 (5 June 1971)	

J. Phys. A: Math. Gen., Vol. 11, No. 10, 1978. Printed in Great Britain

On the operation of a tunable electromagnetic detector for gravitational waves

F Pegoraro<sup>†</sup>, E Picasso<sup>‡</sup> and L A Radicati<sup>‡</sup>§ <sup>†</sup>Scuola Normale Superiore, Pisa, Italy <sup>‡</sup>CERN, Geneva, Switzerland

Received 6 December 1977, in final form 20 April 1978

#### ELECTROMAGNETIC DETECTOR FOR GRAVITATIONAL WAVES

F. PEGORARO, L.A. RADICATI Scuola Normale Superiore, Pisa, Italy

and

Ph. BERNARD and E. PICASSO CERN, Geneva, Switzerland

Received 29 June 1978

MAGO was a proposal for a scaled-up experiment with 500 MHz cavities (not funded)

During the R&D activities 3 superconducting cavities were built

#### 1. PACO-3GHz-pillbox

- 2-cell cylindrical pillbox-cavity @ 3GHz as proof-of-principle experiment
- Low Q, test of RF system, excitation of signal mode



## The MAGO proposal

... and the PACO cavities

#### PACO-2GHz-fixed

- 2-cell cavity with optimised geometry
- Underwent chemistry and cold test to obtain Q<sub>0</sub>(U) for TE011

#### PACO-2GHz-variable

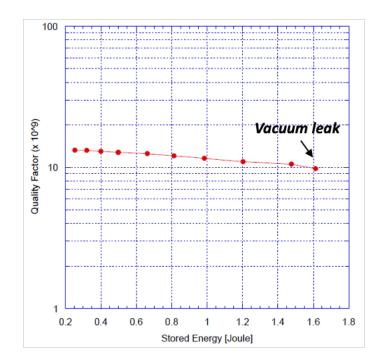
- 2-cell cavity with same geometry but variable coupling
- Never treated nor tested on shelf for >15y @ INFN Genova



In the following, denoted as "MAGO cavity"



Figure 5. Niobium spherical cavities (fixed coupling).



## **Continuation of R&D efforts**

**DESY/UHH - FNAL - INFN collaboration** 

**Cavity at DESY** 

June

2023

Mid.

2024

End

2024

Today –

 Mechanical characterisation and RF measurements at room temperature (done)

#### **Cavity at FNAL**

 Treatment of cavity, construction of a support structure and RF antennas, first cryogenic characterisation

#### Cavity back at DESY

Cryogenic test with (initial) LLRF system

#### Cavity back at FNAL

• First GW search in existing cryostats at Fermilab

In parallel, work started on an LLRF system to drive and read-out the cavity



## **Towards GW experiments with SRF cavities**

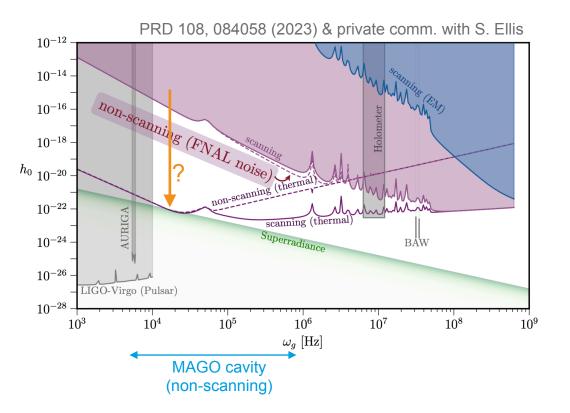
**DESY/UHH - FNAL - INFN collaboration** 

#### Main pillars (and plans) for a future experiment

- Based on the experience with MAGO, fabricate new cavities with optimised geometry
- Design **LLRF system** to drive and read-out the cavity with highest possible sensitivity
- Design a **suspension system** to eliminate mechanical noise (from environment and ground motion)
- Design a cryostat with required thermal properties and which minimises acoustic noise in the He bath

Coordinated and **synchronised observatories** at DESY, FNAL and possible other locations for different frequencies

This talk with focus on DESY activities and plans, Bianca's talk with focus on FNAL

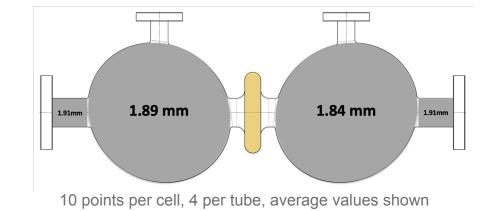


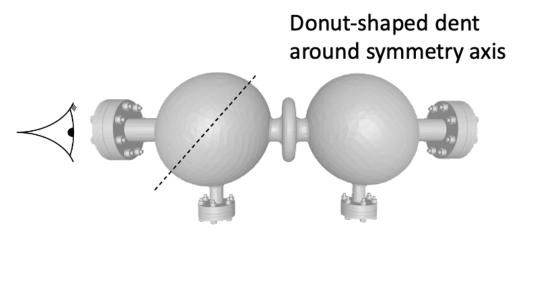
## The MAGO cavity

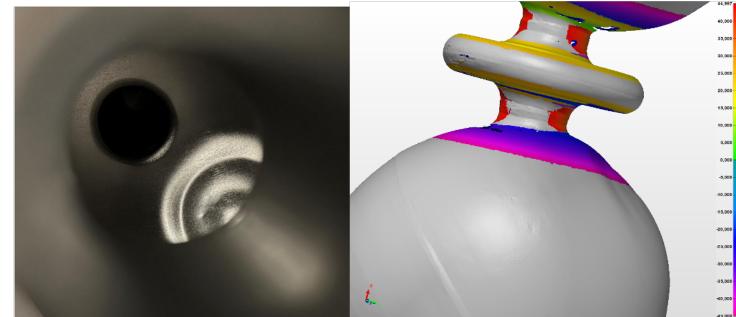
## **Cavity is out of shape**

**Mechanical survey** 

Measurement of the cavity geometry at arrival at DESY with a laser line scan and an ultrasonic wall thickness measurement (expected wall thickness 2mm)

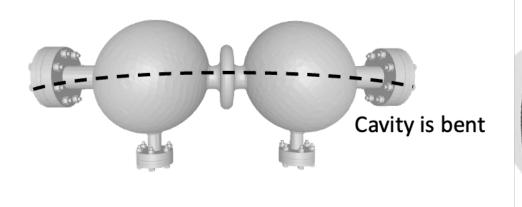






## **Cavity is out of shape**

**Mechanical survey** 

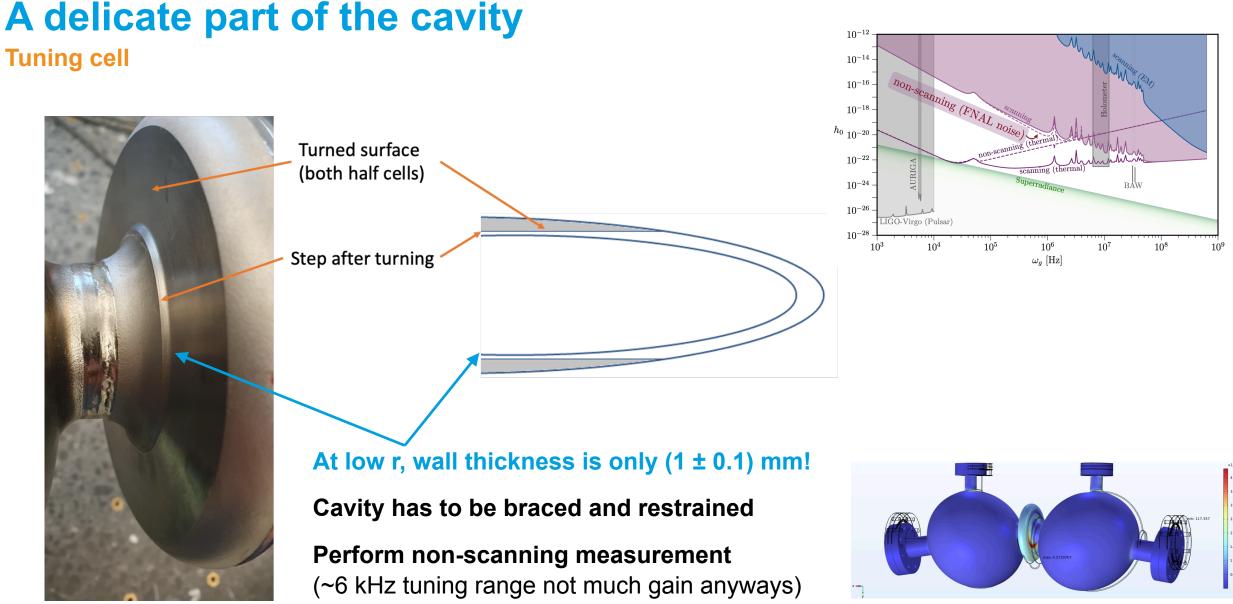


Colour coding shows deviations from technical drawings





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#### **Tuning cell**

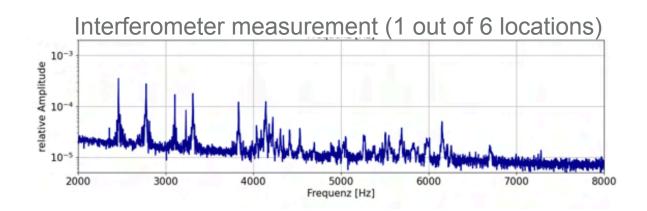
## No damage to flanges or sealing surface

Leaktight at room temperature

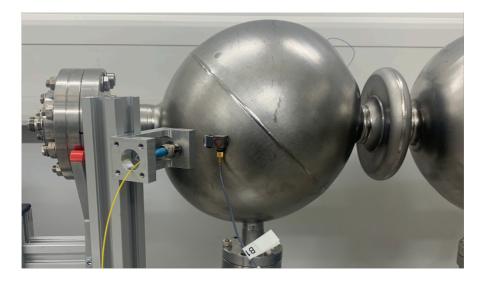


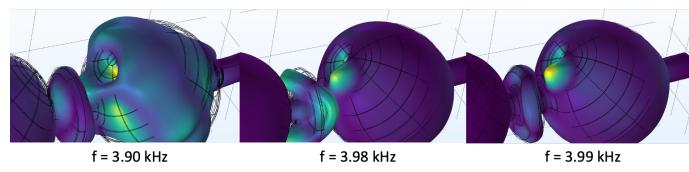
## **Mechanical resonances**

Measurement with accelerometers and interferometric sensors



## Results difficult to interpret, better understanding with simulations

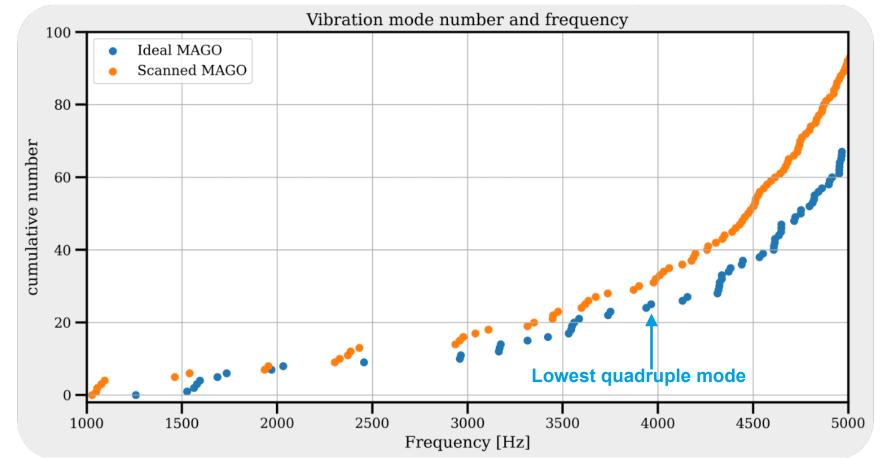




Dent in cell causes oscillation maximum in vibration Eigenmodes

## **Mechanical resonances**

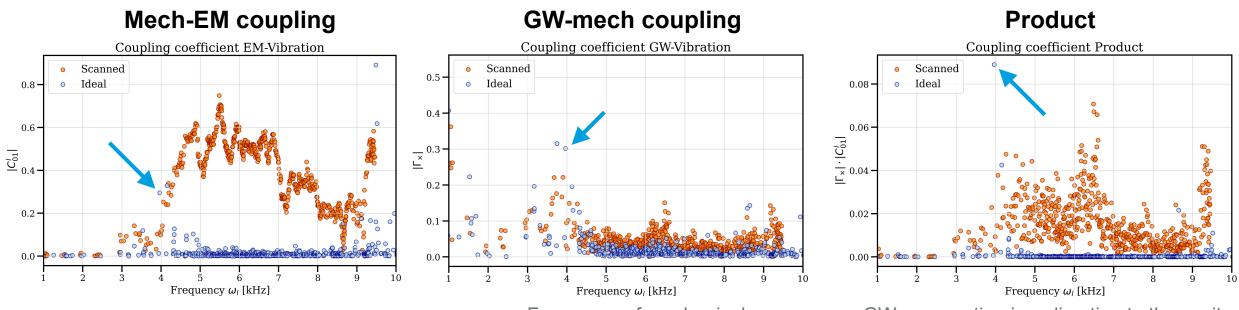
Many more and densely populated mechanical normal modes with distorted detector



Ideal MAGO = cavity from technical drawings

## **Couplings to mechanical resonances**

Simulation with ideal and measured cavity



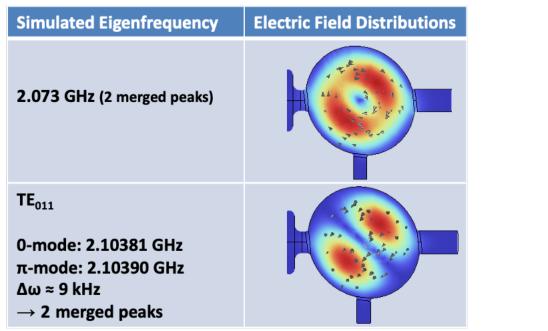
Frequency of mechanical resonances, GW propagating in z-direction to the cavity

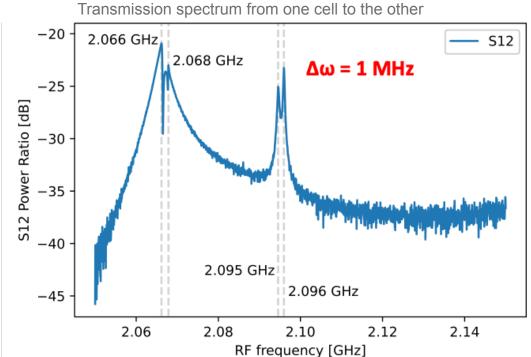
#### With distorted geometry

- Several multipoles with quadrupole fractions (otherwise mech-EM coupling zero)
- Not a clear lowest lying dominating quadruple resonance like with a clean (ideal) geometry

## **RF** measurements worrisome

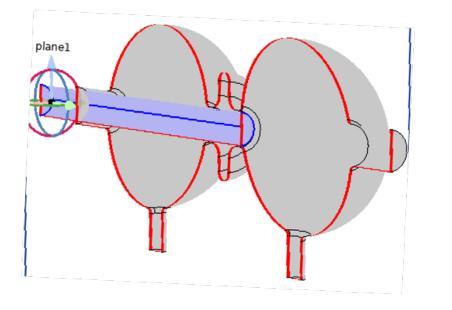
Not what expected from ideal geometry at room temperature

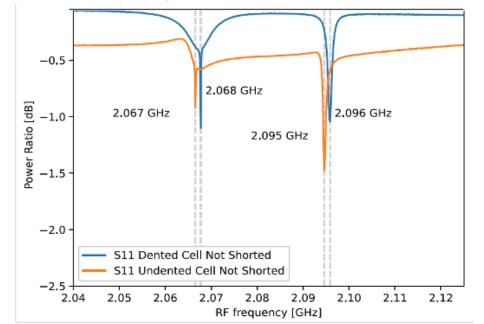




## **Eigenfrequencies of single-cells do not match**

Idea: Short one cell and measure the other





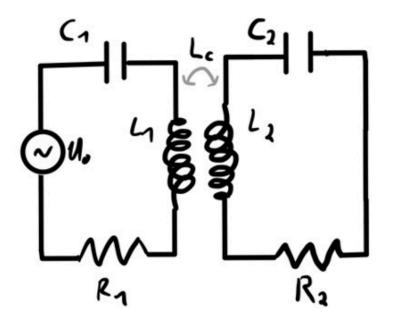
Absorption spectra for both cells

Not wanted nor expected! Remember the coupled pendulum?

## **Shed light on RF measurements**

Model two MAGO cells as inductively coupled RLC circuits

Driven by an external oscillator U(t) on one side



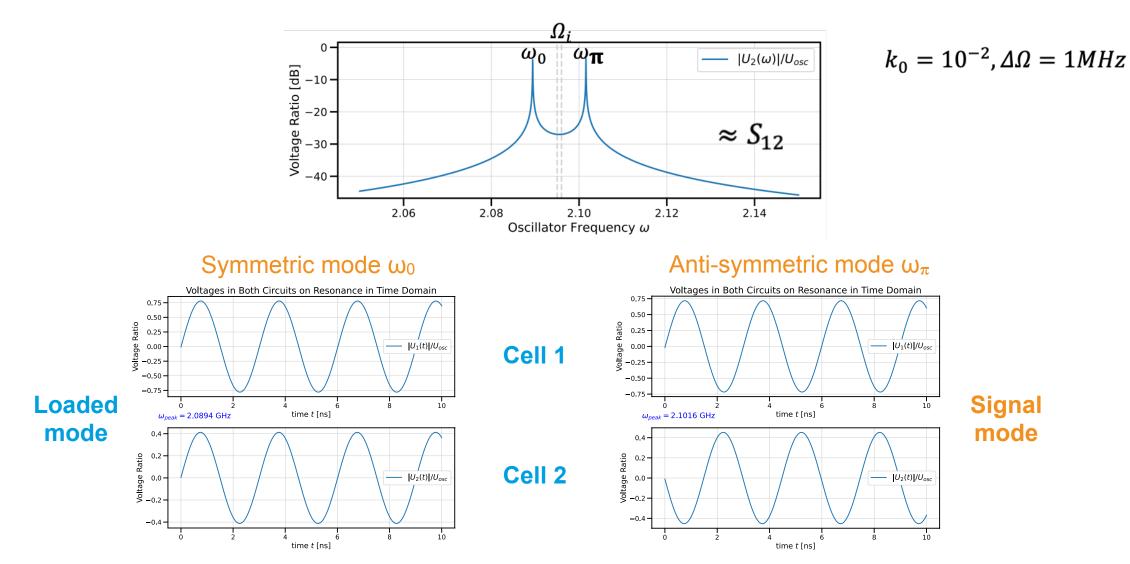
$$L_1 \ddot{I}_1 + R_1 \dot{I}_1 + \frac{1}{C_1} I_1 = -L_c \ddot{I}_2 + \dot{U}$$
$$L_2 \ddot{I}_2 + R_2 \dot{I}_2 + \frac{1}{C_2} I_2 = -L_c \ddot{I}_1$$

Define parameters:

$$k = L_c / \sqrt{L_1 L_2}$$
 Coupling strength  
 $\Delta \Omega = \Omega_2 - \Omega_1, \ \Omega_i = \frac{1}{\sqrt{L_i C_i}}$  Single cell  
eigenfrequencies

## Solution of the coupled e.o.Ms

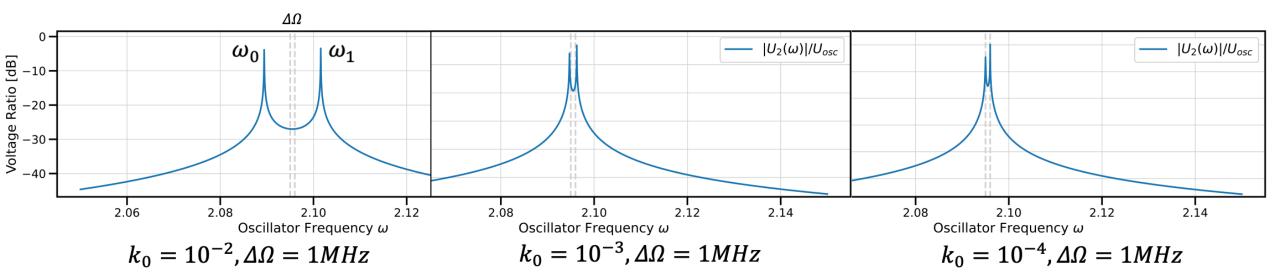
#### In frequency and time domain



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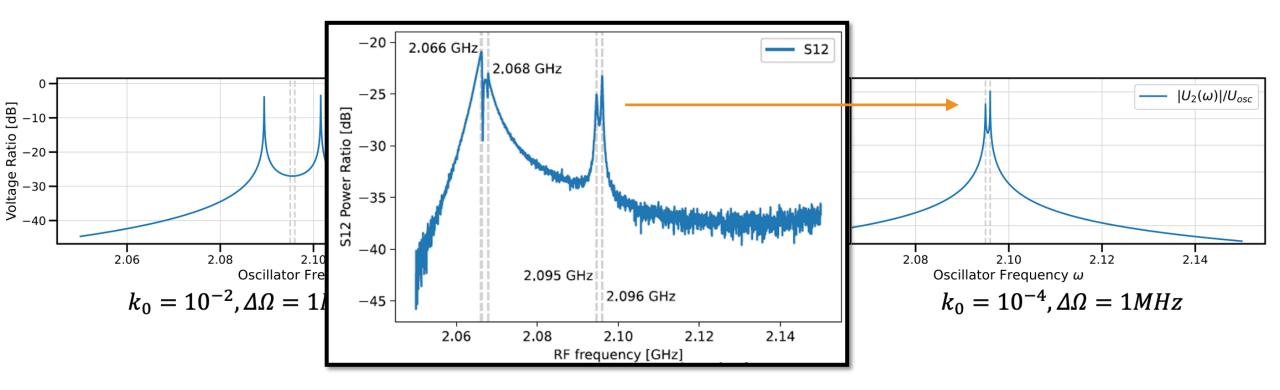
## Varying the coupling parameter

Leads to different peak sizes



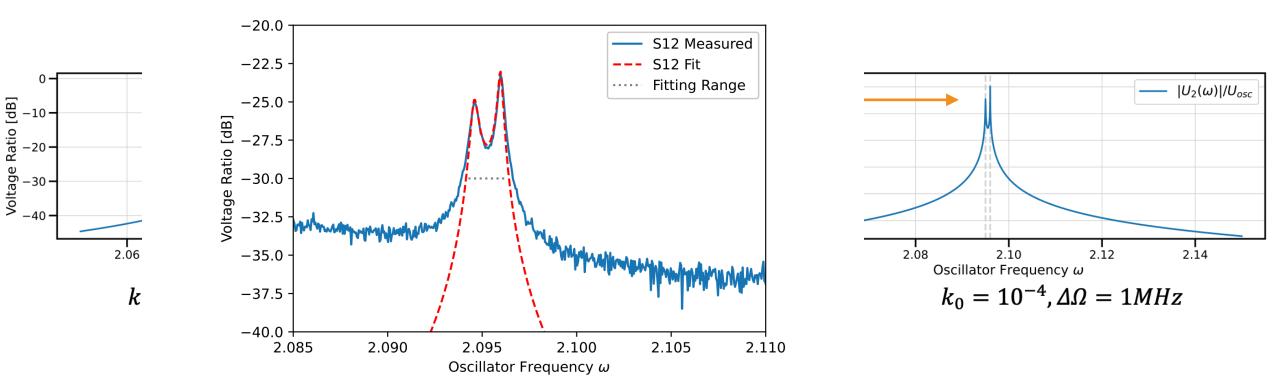
## Varying the coupling parameter

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## Varying the coupling parameter

#### Leads to different peak sizes



tentative results: (S12 and  $|U_2/U_0|$  are not exactly the same)

 $k = 0.5 \cdot 10^{-5}$ ,  $\Delta \Omega = 1.4$  MHz,  $Q_1 = 0.5 \cdot 10^4$ ,  $Q_2 = 0.8 \cdot 10^4$ 

## Weak coupling is inherent

Recover by bringing  $\Delta\Omega$  closer together

∆**Ω=1MHz; k₀=10**-4

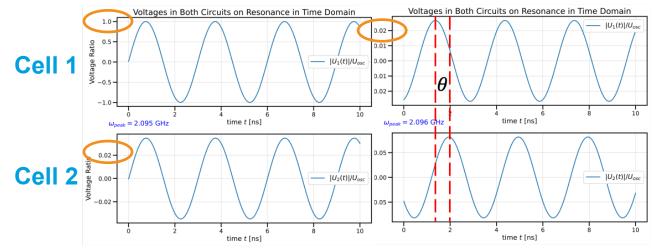
$$k = 2\frac{\omega_{\pi} - \omega_0}{\omega_{\pi} + \omega_0} \approx 2\frac{1MHz}{4GHz} = 5 \times 10^{-4}$$

#### Phase shift which vanishes for high Q!

•  $Q_{1,2} \gg 10^9$  to avoid arbitrary phase shifts

#### Low amplitude ratio limits sensitivity

- Mech-EM coupling  $\sim E_0 * E_{\pi}$
- Largest amplitude limits operation power



#### Symmetric mode $\omega_0$ Anti-symmetric mode $\omega_{\pi}$

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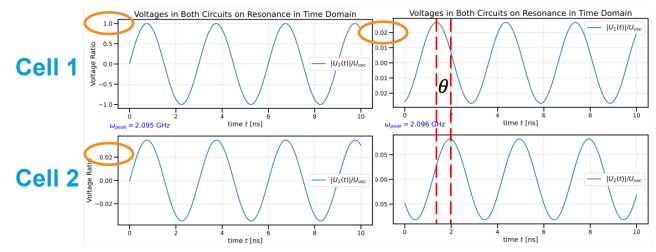
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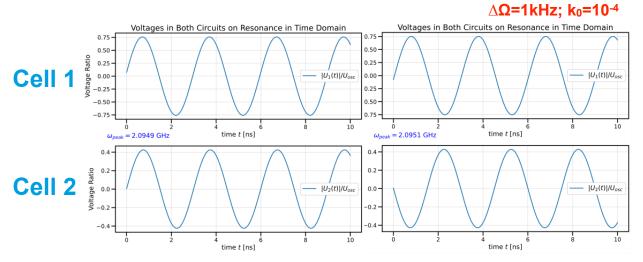
- Mech-EM coupling  $\sim E_0 * E_{\pi}$
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#### Need to tune cavity/cells to achieve wanted $\Delta \Omega$

- Working on simulations to identify origin and possible "turning knobs"
- Mechanical or dielectric tuning?



#### Symmetric mode $\omega_0$ Anti-symmetric mode $\omega_{\pi}$

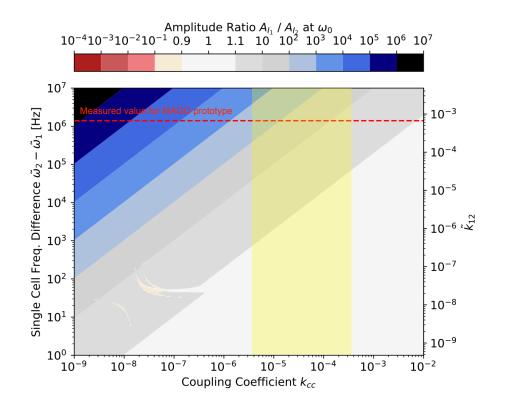


## Some tuning practically unavoidable

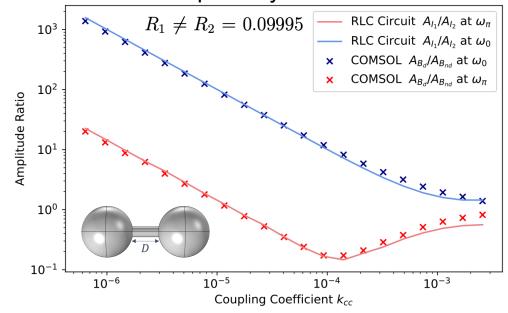
Single cell eigenfrequency differences also expected for close to ideal geometries

For GHz cells, a relative difference in the radii 0.1 % will already lead to  $\Delta \omega = 1$  MHz (This is also what we roughly get from current wall thickness measurement)

$$R_1, R_2 = R_1 + \Delta R_1$$
  
 $rac{\Delta \tilde{\omega}}{\tilde{\omega}_1} = rac{\Delta R}{R_1}$ 



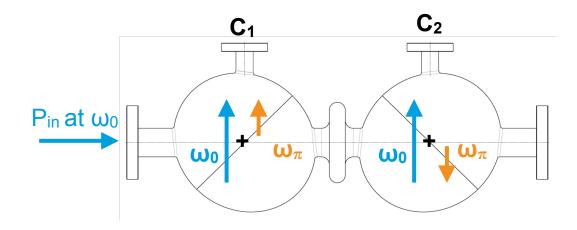
Compare to a simplified MAGO geometry Good description by RLC circuit



# How can we control the cavity and detect the signal?

## **MAGO's readout**

#### Two similar concepts



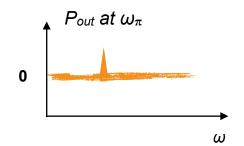
#### 

#### Readout with magic-tee

- Shift signal phase of one cell (with magic-tee) by  $\pi$
- Loaded mode cancels, signal mode amplified, works well with similar amplitudes

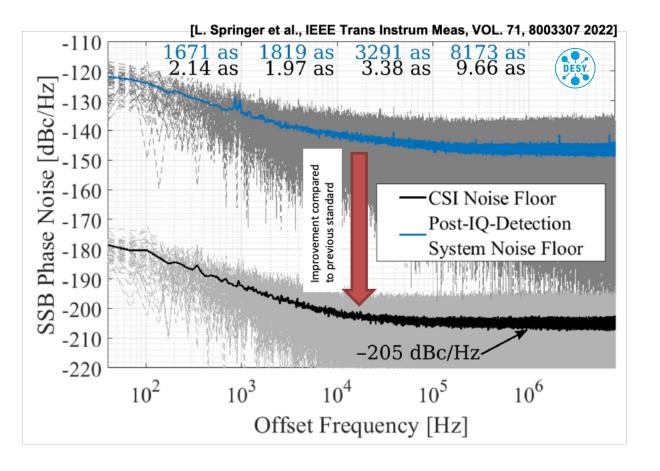
#### **Carrier suppression interferometer (CSI)**

- Shift driving signal phase by  $\pi$
- Loaded mode cancels with driving signal, less sensitive to amplitude differences



## **Detecting signals with unprecedented sensitivity**

**Carrier suppression interferometer** 



Established a 60 dB improvement of the detection noise floor at 1.3 GHz in a laboratory-controlled environment

-205 dBc/Hz (Δω=1 MHz), -180 dBc/Hz (Δω=100 Hz)

#### Matches MAGO requirements and conditions

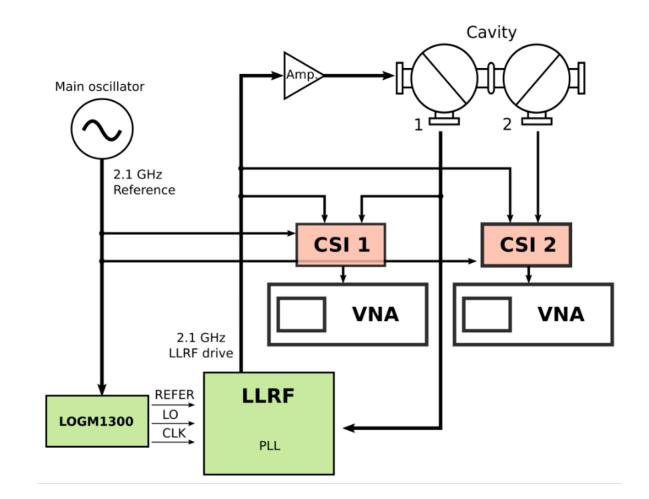
- High RF field amplitude for better sensitivity to GW
- Excited signal by GW is mixed to driving signal, with  $\Delta \omega$  in the 10-100 kHz range

## **Preliminary design of our LLRF architecture**

#### **Current R&D**

- Use FPGA based cavity simulator to develop
   LLRF/CSI system
- Set up CSI at 2.1 GHz (operational frequency of MAGO)
- Integrate CSI with µTCA LLRF system

Test design with 1 CSI at cold RF test in 2024



# How could a future experiment look like?

## **Cryoplatform at HERA North hall**

Possible realisation of dedicated HFGW experiments

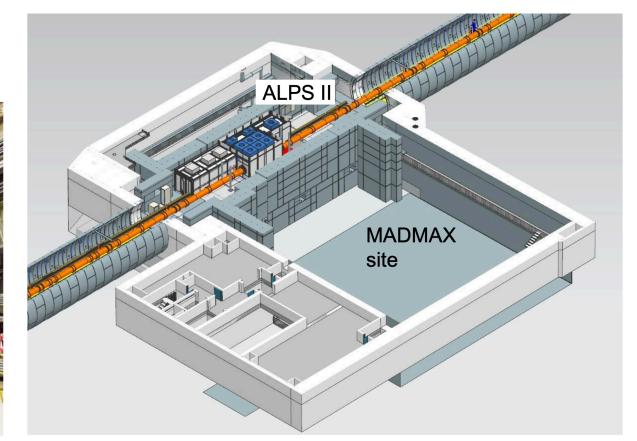
Distribution system of liquid helium (at ~ 4.5 K)

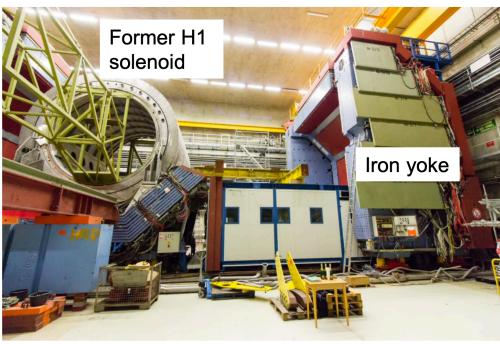
Supply for up to three experiment

In construction, operation from 2026

Located 30 m underground

Dimensions: height = 16 m, area =  $1222 \text{ m}^2$ 





## A dedicated cryostat

**Main requirements** 

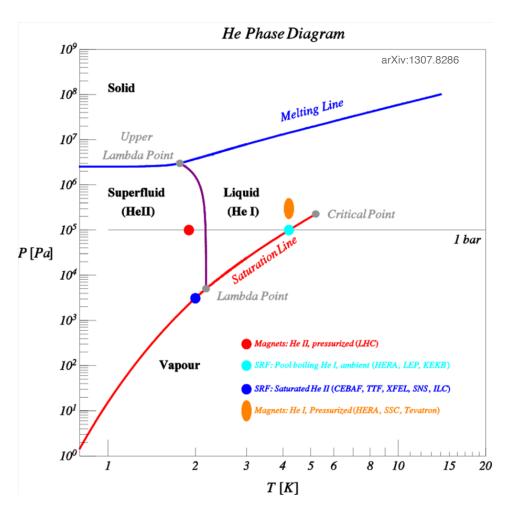
#### Exploit advantages of RF superconductivity

#### Avoid acoustic noise in the helium bath

- From thermal dissipation
- Pressure fluctuations
- Seismic and environmental noise coupled to the liquid

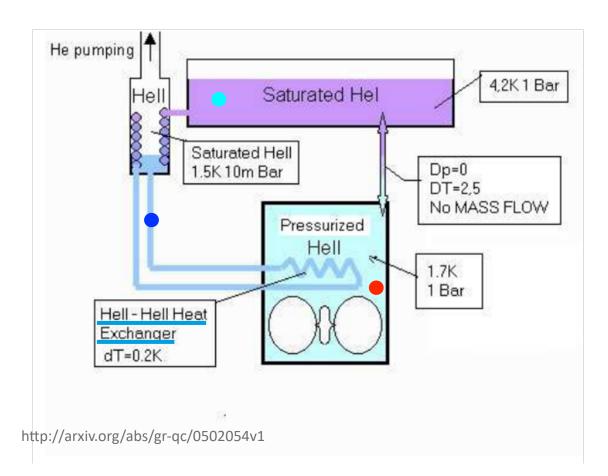
## Pressurised superfluid Helium (1 bar, 1.8K) gives best thermal properties and minimised noise from He bath

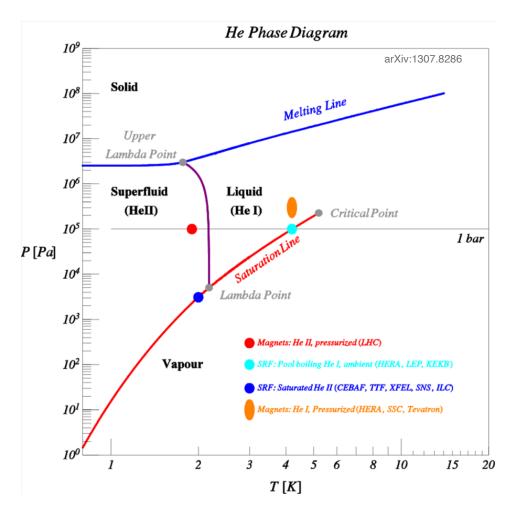
- High specific heat and thermal conductivity
- Low sound wave speed
- High mechanical quality factor
- At 1 bar avoids bubble creation



## **Concept from original MAGO proposal**

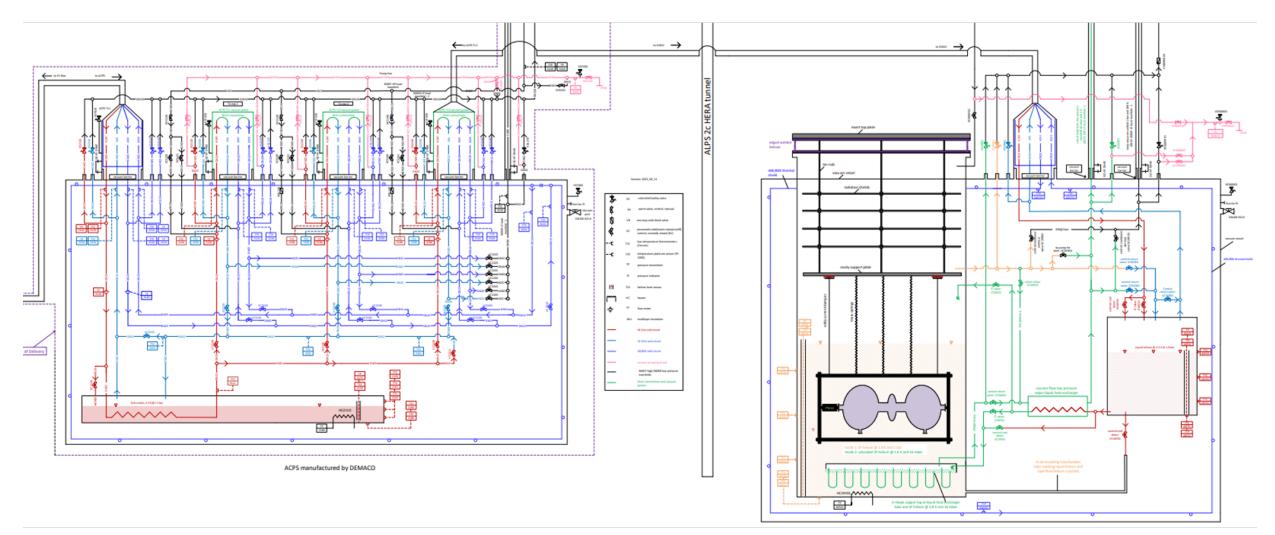
Use superfluid helium with a heat exchanger





## **Design of cryostat started**

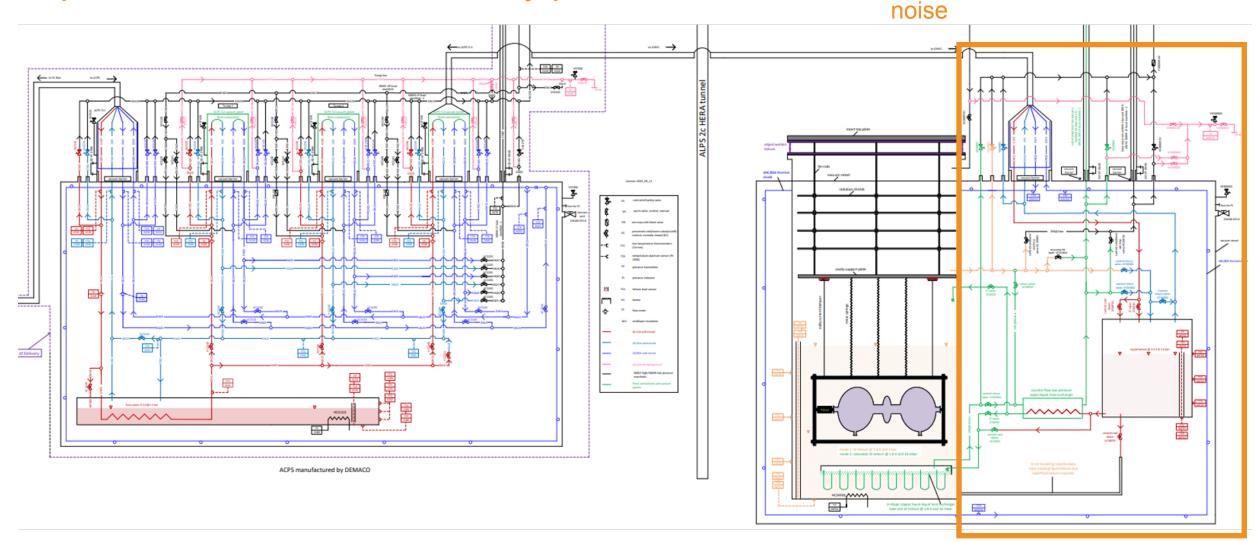
A possible flow scheme to host in DESY cryoplatform



## **Design of cryostat started**

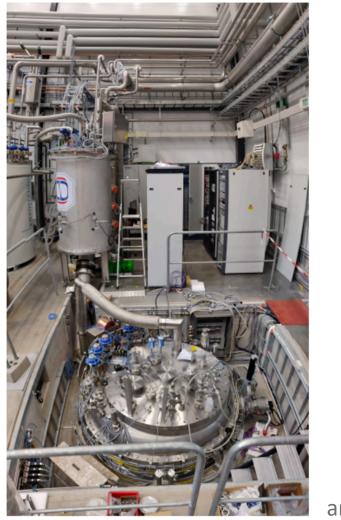
#### A possible flow scheme to host in DESY cryoplatform

#### Elevate and outsource this part to maintain pressure and minimise

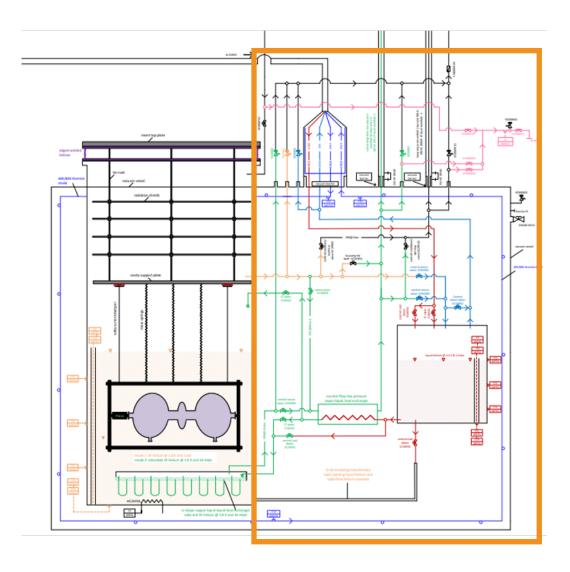


## **Gersemi vertical cryostat at FREIA Laboratory**

#### Similar design concepts



#### arXiv:2103.05265v1



## **Suspension system**

Inspired from Ligo/Virgo concepts

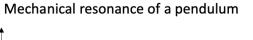
**Requirements not as strict as for interferometers** Pendulum resonance ~Hz, measurement in kHz - MHz range (strong natural damping and high Q factor of superfluid helium)

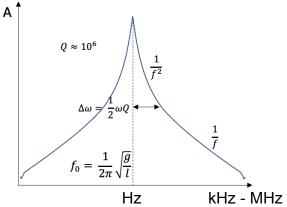
1 or 2 levels of **pendulum** with **leaf springs** for vertical damping

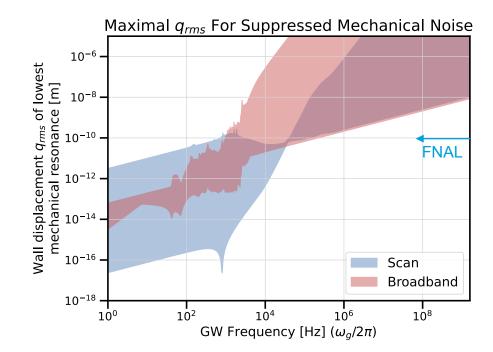
Minimise mechanical shortcut of connections (Vacuum, RF)

**Measure ground motion** in cryoplatform and **vibration** at cavity insert (accelerometer op. 77 K) to estimate required damping

**In general:** working on a more realistic estimate of all the different noise sources to gauge noise suppression requirements







Upper limit: wall displacement rms such that mech. noise < thermal noise or amp noise

### **Conclusions**

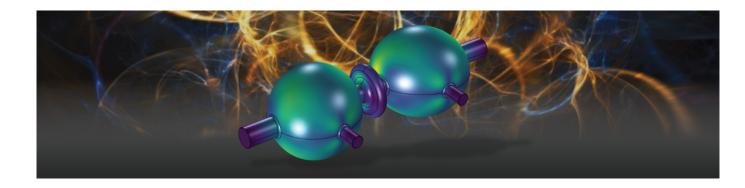
With the discovery of GWs, renewed interest in the heterodyne detection method with SRF cavities Especially the broad band sensitivity is tantalising

Use the existing MAGO prototype cavity to gain experience and design further optimised cavities

LLRF system development on the way, cryostat and suspension design addressed next

Still a long way ahead to reach projected sensitivities, but could aim for a CDR in 2-3 years for a future experiment

Moderate size of the experiment very attractive, allows also for several synchronised observatories



## Thank you

#### Thanks to

Asher Berlin (FNAL), Thorsten Büttner (DESY), Sergio Callatroni (CERN), Ralph Doermann (DESY), Sebastian Ellis (U Geneva), Erika Garutti (UHH), Gianluca Gemme (INFN), Bianca Giaccone (FNAL), Michael Grefe (UHH), Roni Harnik (FNAL), Beate Heinemann (DESY), Axel Lindner (DESY), Frank Ludwig (DESY), Norbert Meyners (DESY), Cornelius Martens (UHH), Andrea Muhs (DESY), Sam Posen (FNAL), Yuri Pischalnikov (FNAL), Jörn Schaffran (DESY), Tobias Schnautz (DESY), Patrik Wiljes (DESY) and many more to come

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