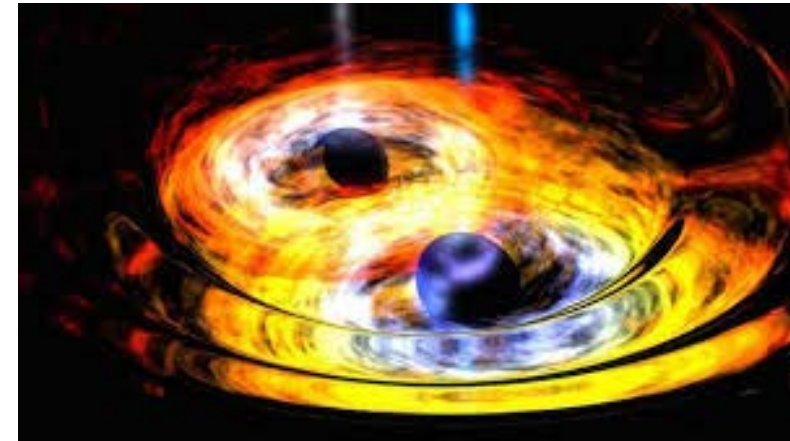
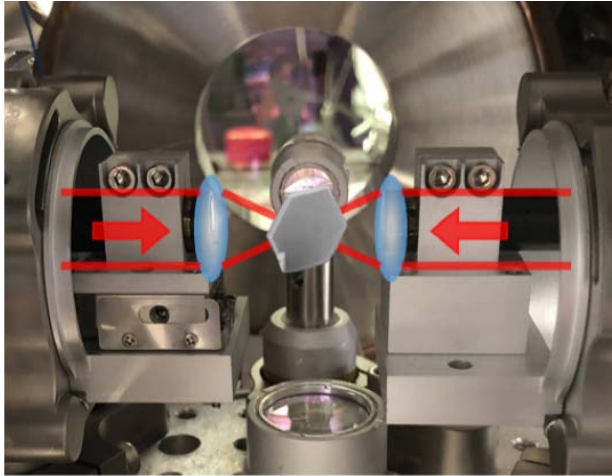


The Levitated Sensor Detector for high-frequency gravitational wave detection



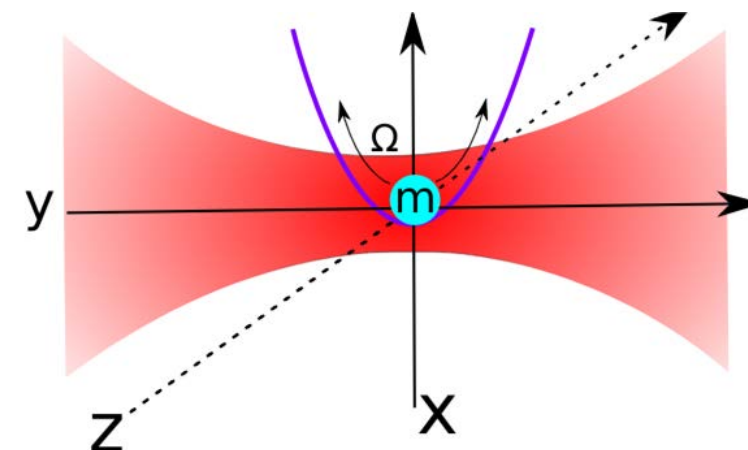
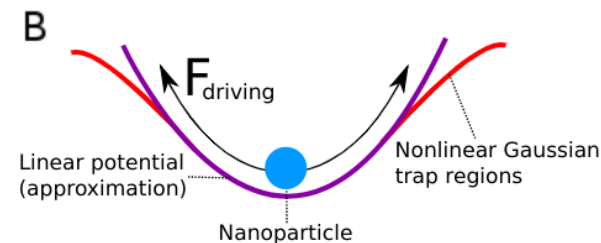
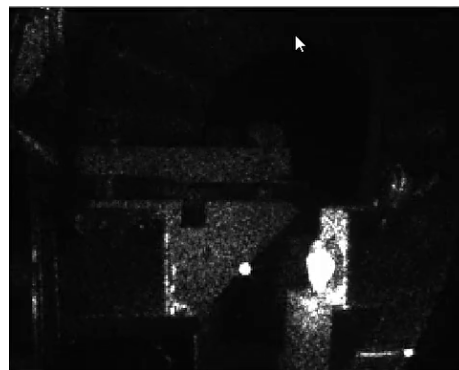
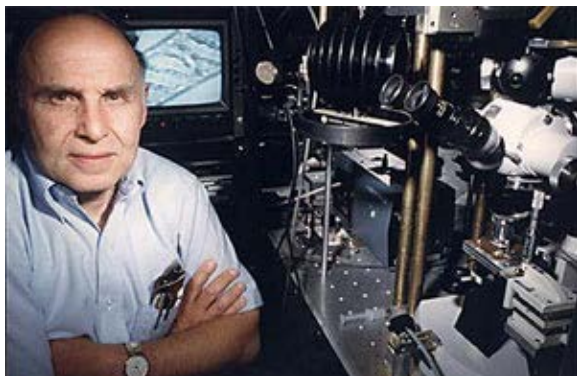
A. Geraci, Northwestern University
Center for Fundamental Physics (CFP)



Ultra high-frequency gravitational waves:
where to next? CERN Dec. 4-8, 2023

Levitated optomechanics - brief introduction

- Ashkin, Bell Labs, 1970s Optical tweezers → biology, biophysics
- Ashkin (76) Levitation in high vacuum



1. Polarisable neutral glass particles act as high field seekers - trapped in laser focus
2. When particle is approximately spherical, acts as 3 decoupled harmonic oscillators.
3. Characteristic length scale 10's nanometers to 10's microns.
4. High Q factors $\sim 10^{12}$ → excellent force sensors
5. Can reduce the effective motional 'temperature' of such objects through feedback schemes

(Ground) state of the art:

Dynamical backaction:

Measurement based cooling:

Science

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REPORT



Cooling of a levitated nanoparticle to the motional quantum ground state

Uroš Delić^{1,2,*}, Manuel Reisenbauer¹, Kahan Dare^{1,2}, David Grass^{1,†}, Vladan Vuletić³, Nikolai Kiesel¹, ...

+ See all authors and affiliations

Science 21 Feb 2020:
Vol. 367, Issue 6480, pp. 892-895
DOI: 10.1126/science.aba3993

Article

Figures & Data

Info & Metrics

eLetters

PDF

A nanoparticle trapped and cooled

Cooling massive particles to the quantum ground state allows fundamental tests of quantum mechanics to be made; it would provide an experimental probe of the boundary between the classical and quantum worlds. Delić *et al.* laser-cooled an optically trapped solid-state object (a ~150-nanometer-diameter silic a nanoparticle) into its quantum ground state of motion starting from room temperature. Because the object is levitated using optical forces, the experimental configuration can be switched to free fall, thereby providing a test bed for several macroscopic quantum experiments.

Science, this issue p. 892

nature

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Article | [Published: 14 July 2021](#)

Quantum control of a nanoparticle optically levitated in cryogenic free space

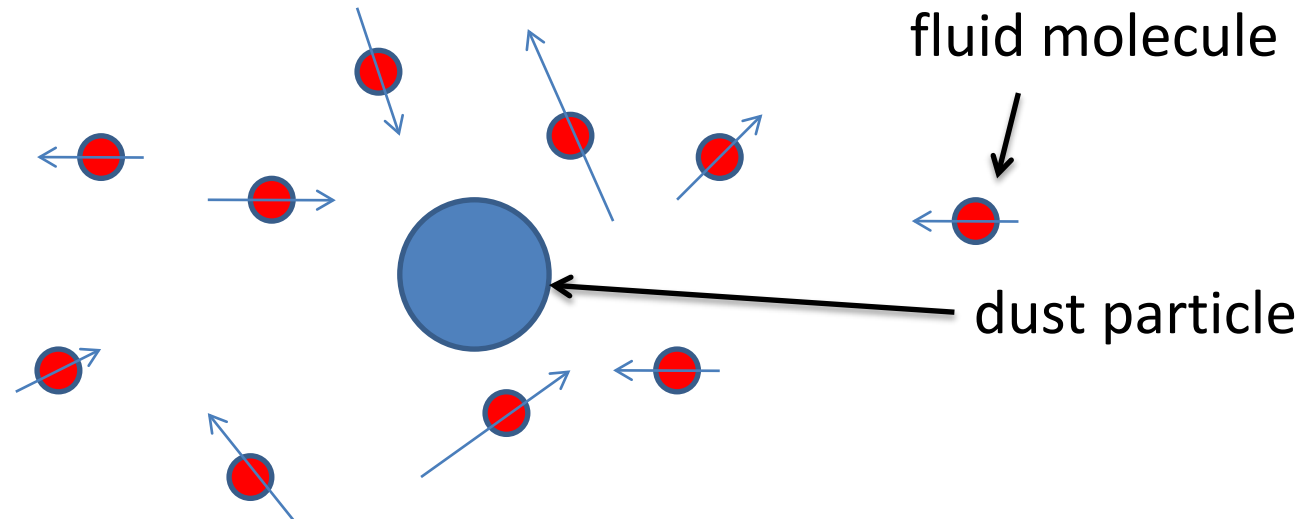
[Felix Tebbenjohanns](#), [M. Luisa Mattana](#), [Massimiliano Rossi](#), [Martin Frimmer](#) & [Lukas Novotny](#)

[Nature](#) 595, 378–382 (2021) | [Cite this article](#)

5736 Accesses | 2 Citations | 262 Altmetric | [Metrics](#)

Fundamental limitation: thermal noise

Brownian motion – random “kicks” given to particle due to thermal bath



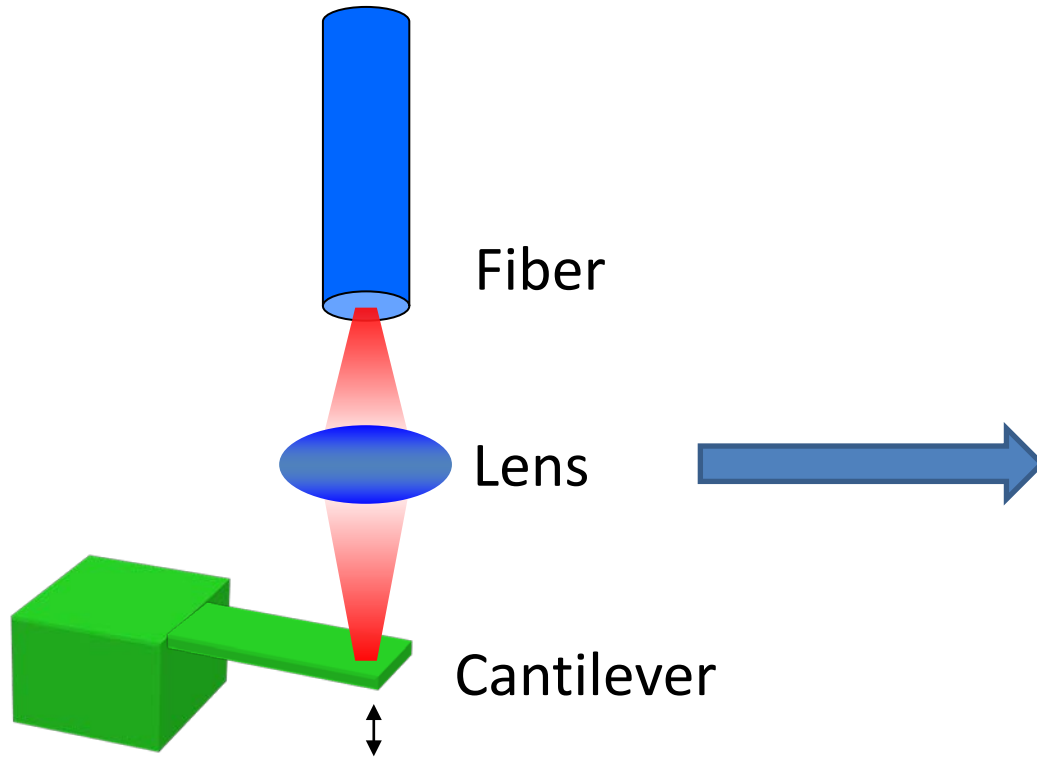
- Random “kicks” are given to sensor due to finite T of oscillator

$$\frac{1}{2}k\langle x^2 \rangle = \frac{1}{2}k_B T$$



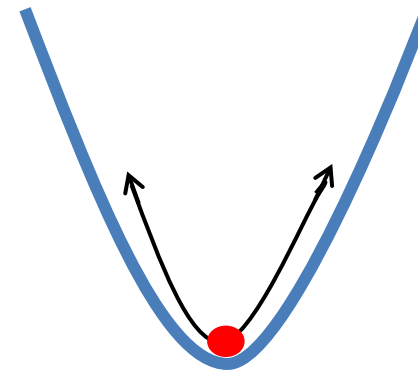
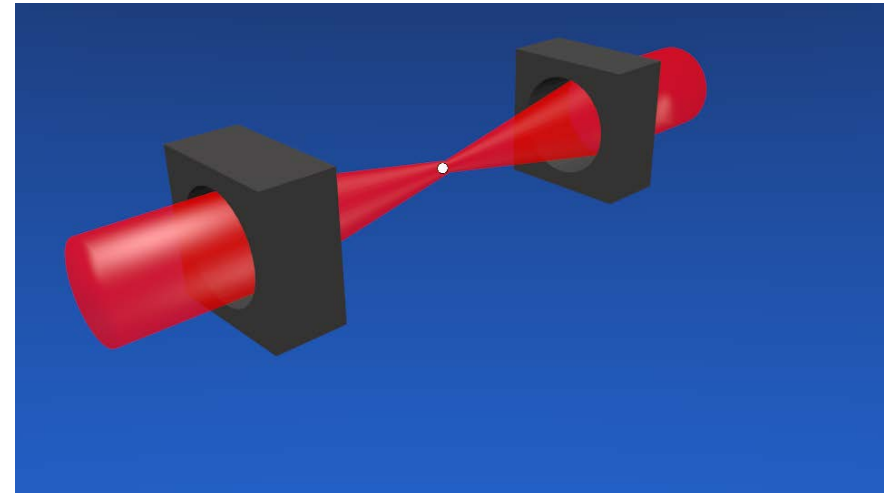
$$F_{\min} = \left(\frac{4kk_B T b}{Q\omega_0} \right)^{1/2}$$

Improving sensitivity



Limitations on Q: Clamping, surface imperfections, internal materials losses

Levitate the force sensor!



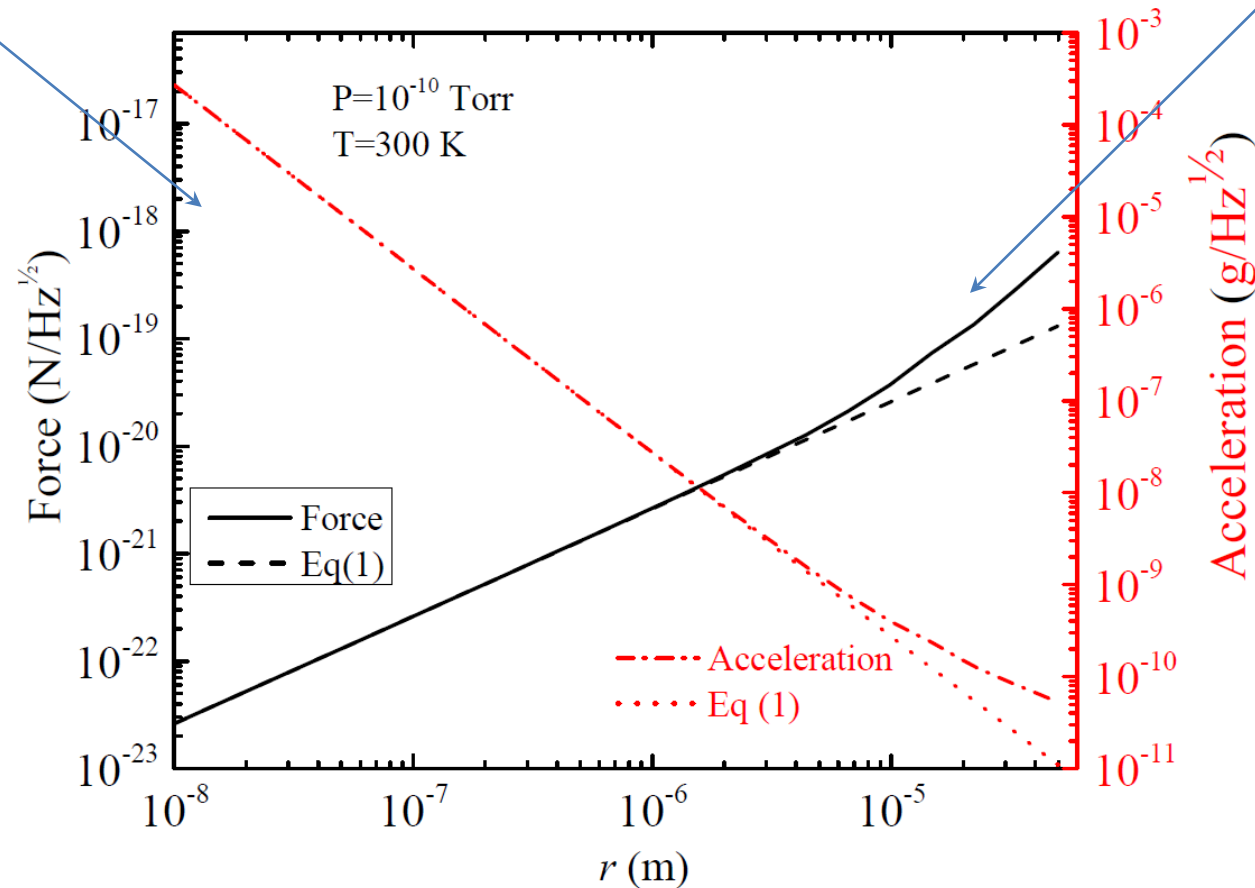
CM motion decoupled from environment – no clamping, materials losses

Projected Force sensitivity

Cantilever sensors

$$F_{\min} = (4k_B T \gamma m)^{1/2} \quad (1)$$

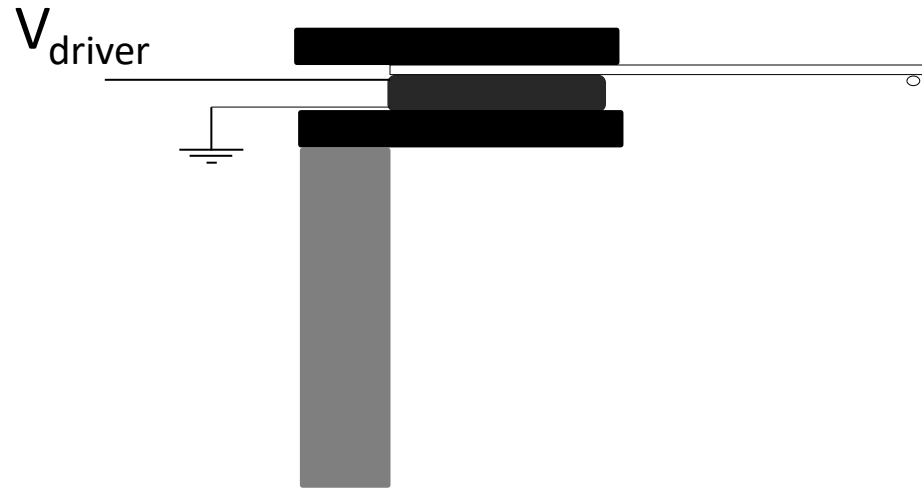
Photon recoil heating
Seen recently by
Novotny group
V. Jain et. al.,
PRL 116, 243601
(2016)



20 zN/Hz^{1/2} Gieseler, Novotny, Quidant (Nature Phys. 2013)

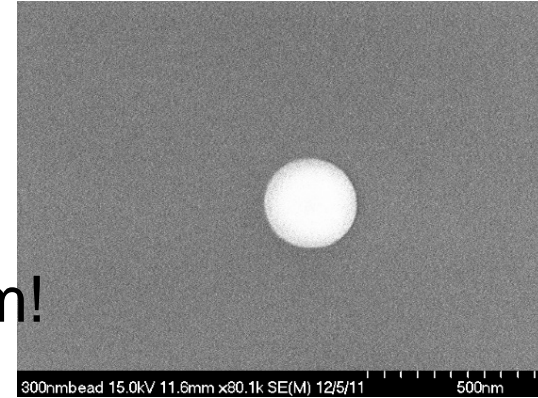
Z. Yin, A. Geraci, T. Li, Int. J. Mod. Phys. B 27,1330018 (2013).

Trap loading

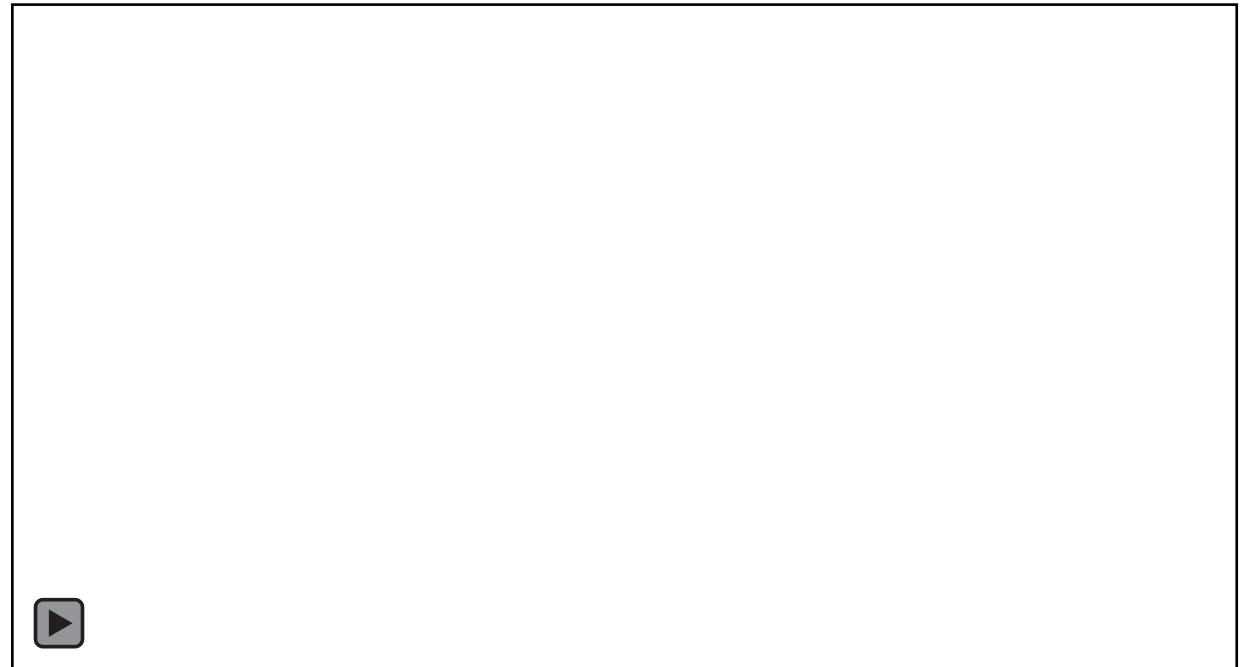
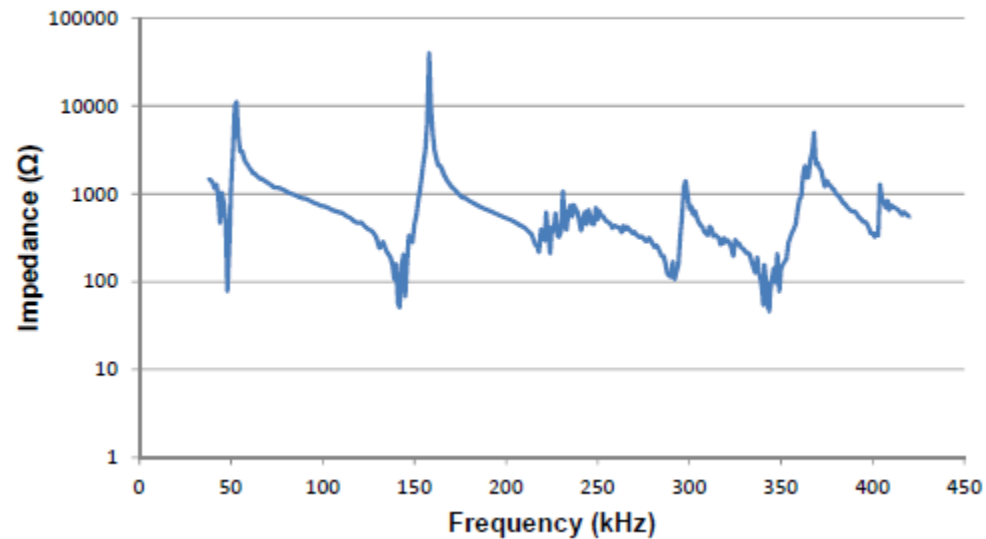


$$a \propto \frac{1}{R^2}$$

$\sim 10^7$ g for $R=150$ nm!



Impedance of piezoelectric ceramic ring



Trapping instabilities

- **Radiometric forces**

Trap instabilities arise from uneven heating of the sphere surface

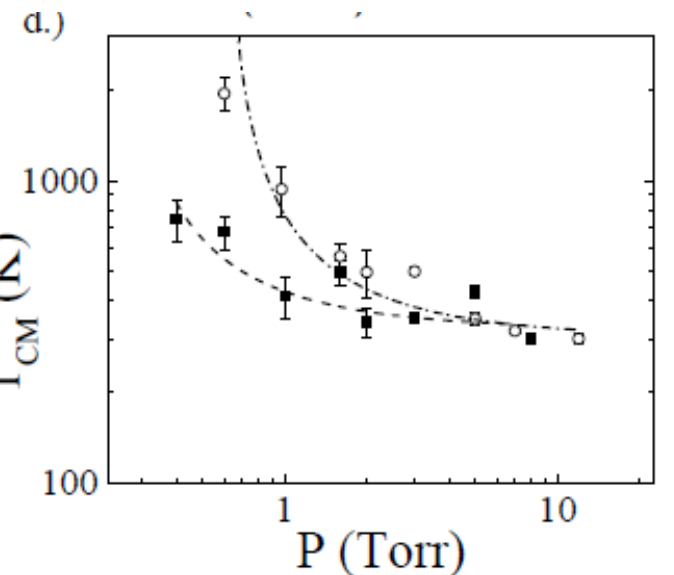
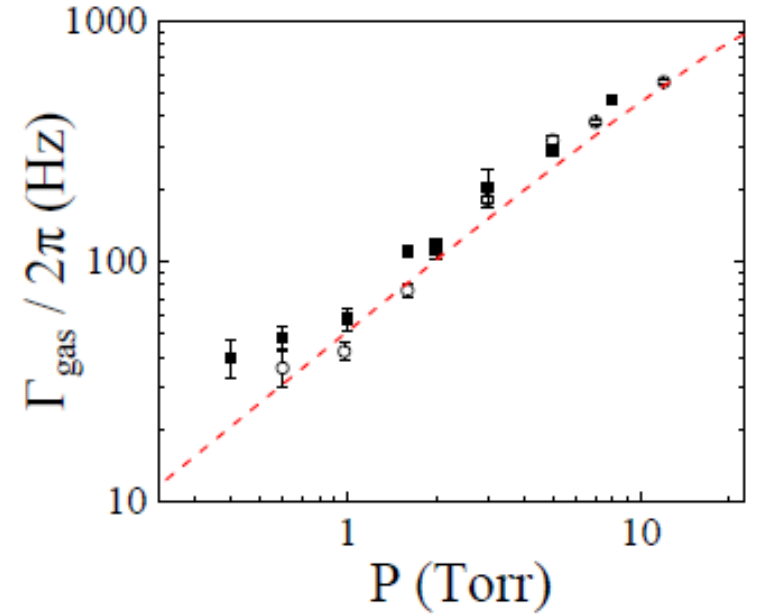
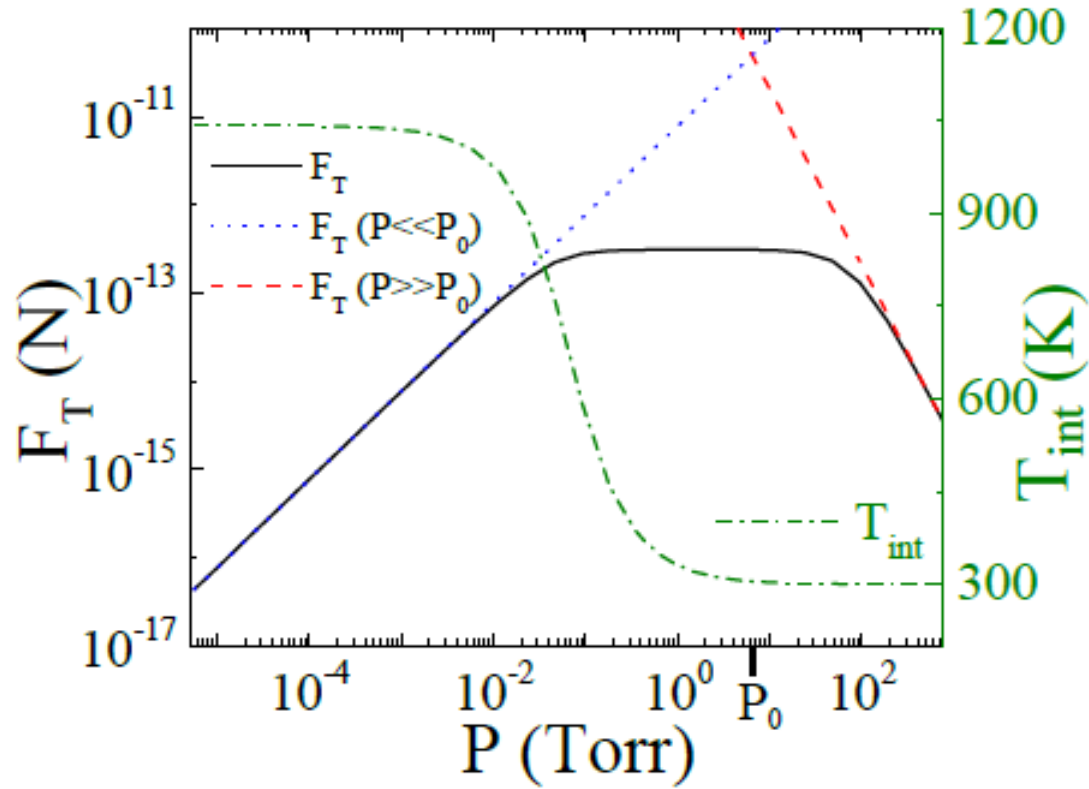
Important when mean free path \sim object size



Crooke's Radiometer

Radiometric forces

$$F_T = - \frac{\pi r^2 \eta \sqrt{\frac{\alpha R_g}{MT}} \Gamma_i}{\frac{P}{P_0} + \frac{P_0}{P}}$$



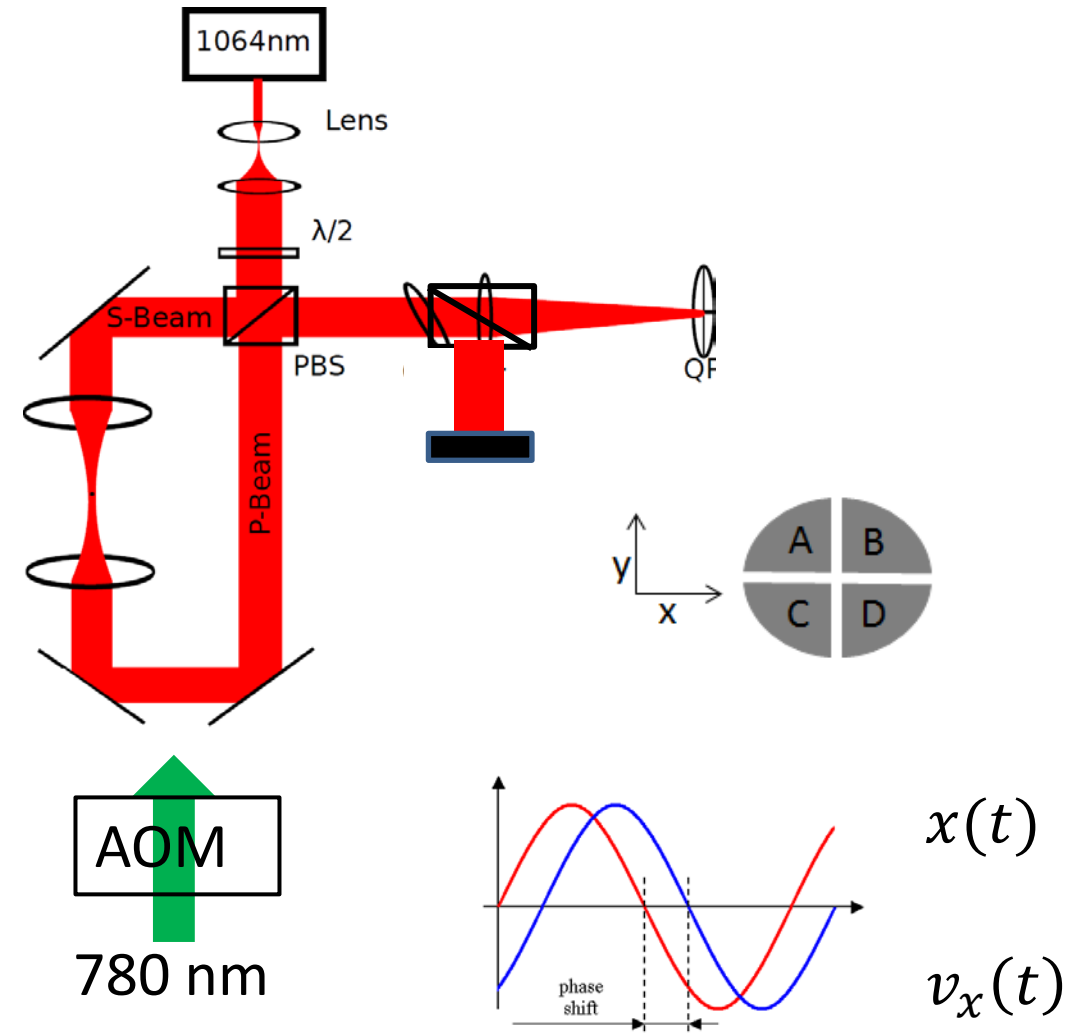
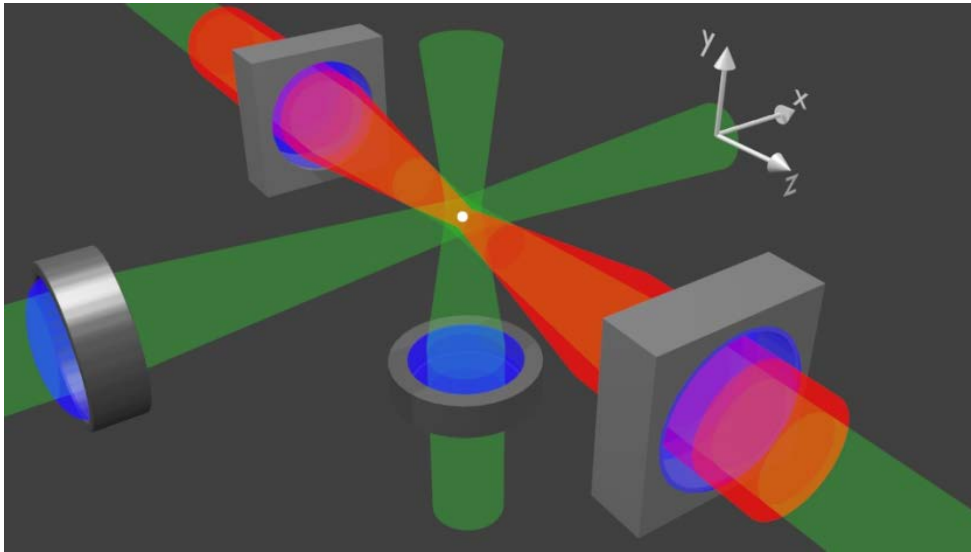
1% temp gradient across surface
 $R=1.5 \mu\text{m}$, $l=2 \times 10^9 \text{ W/m}^2$

**Ranjit et.al., PRA 91, 051805(R)
 (2015).**

Heating rate > gas damping rate
 \rightarrow Particle loss \rightarrow Need feedback!

3D feedback cooling of a nanosphere

Needed to stabilize the particle, damp and cool it
Mitigate photon recoil heating

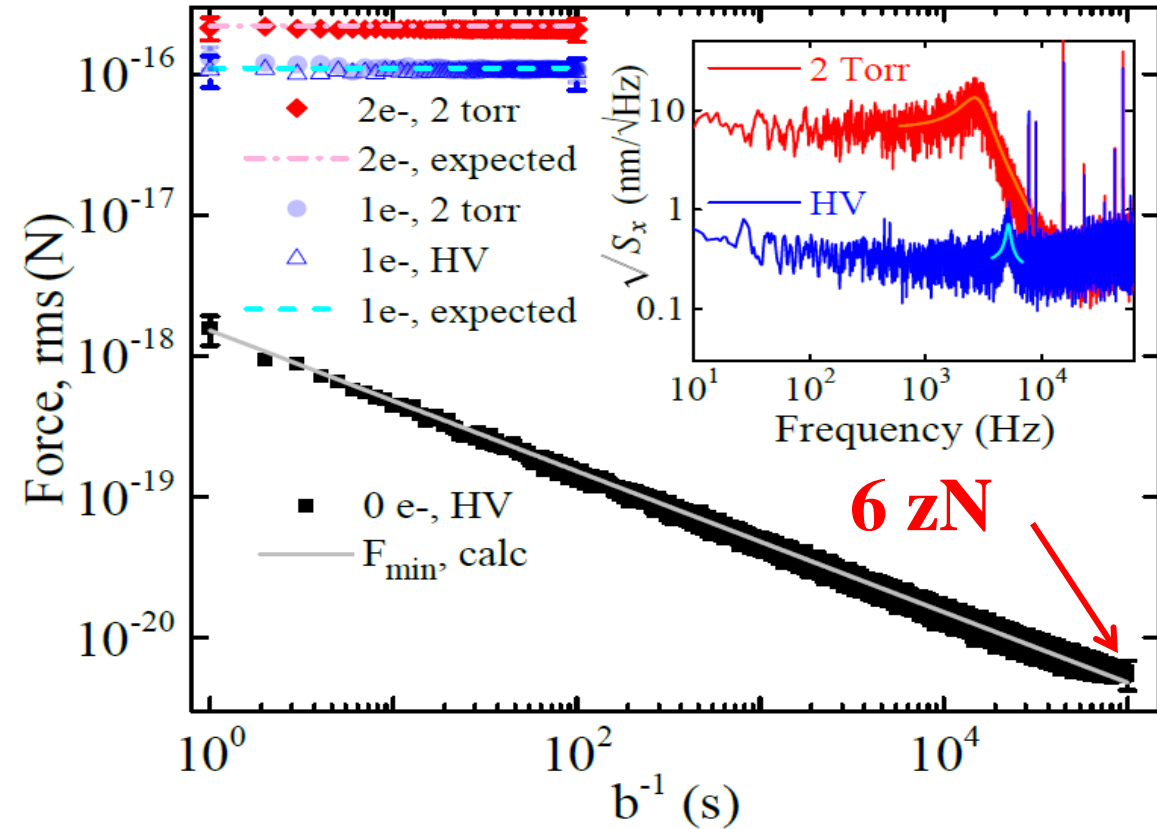
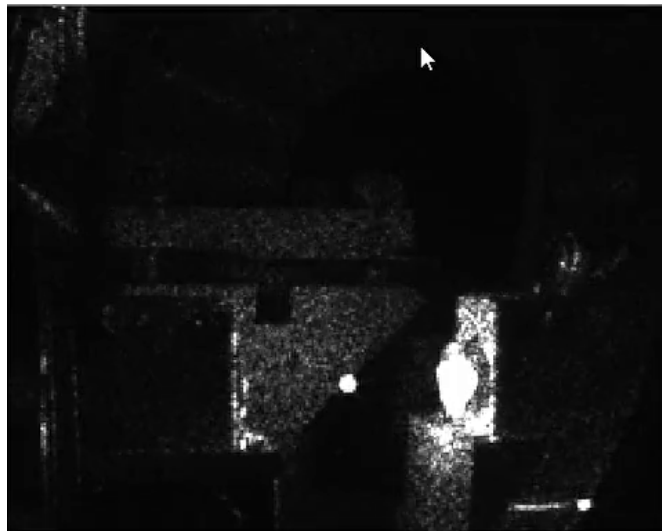
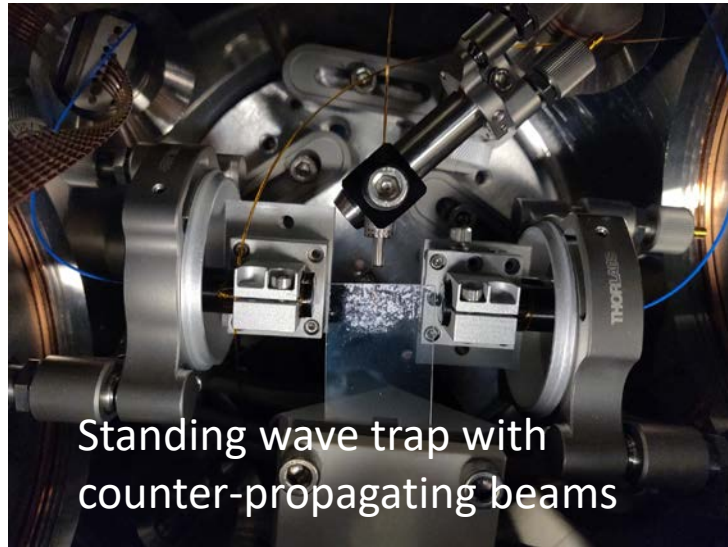


$$F_{\min} = \sqrt{\frac{4kK_B T B}{\omega_0 Q}}$$

$$Q_{\text{eff}} = \frac{Q_0 \Gamma_0}{\Gamma_0 + \Gamma_{\text{cool}}}$$

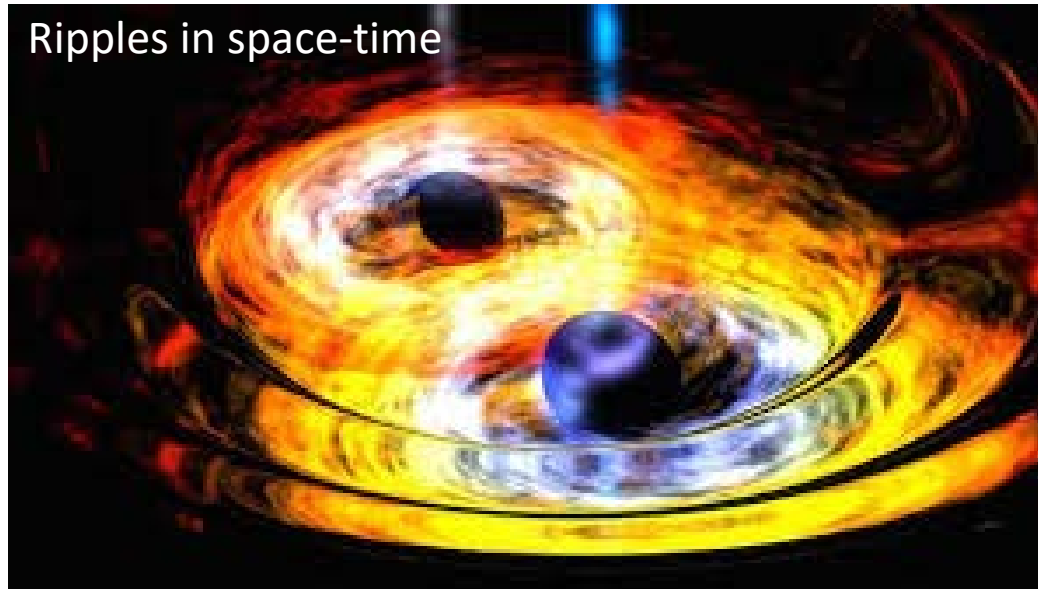
$$T_{\text{eff}} = \frac{T_0 \Gamma_0}{\Gamma_0 + \Gamma_{\text{cool}}}$$

Calibrated zN force sensitivity



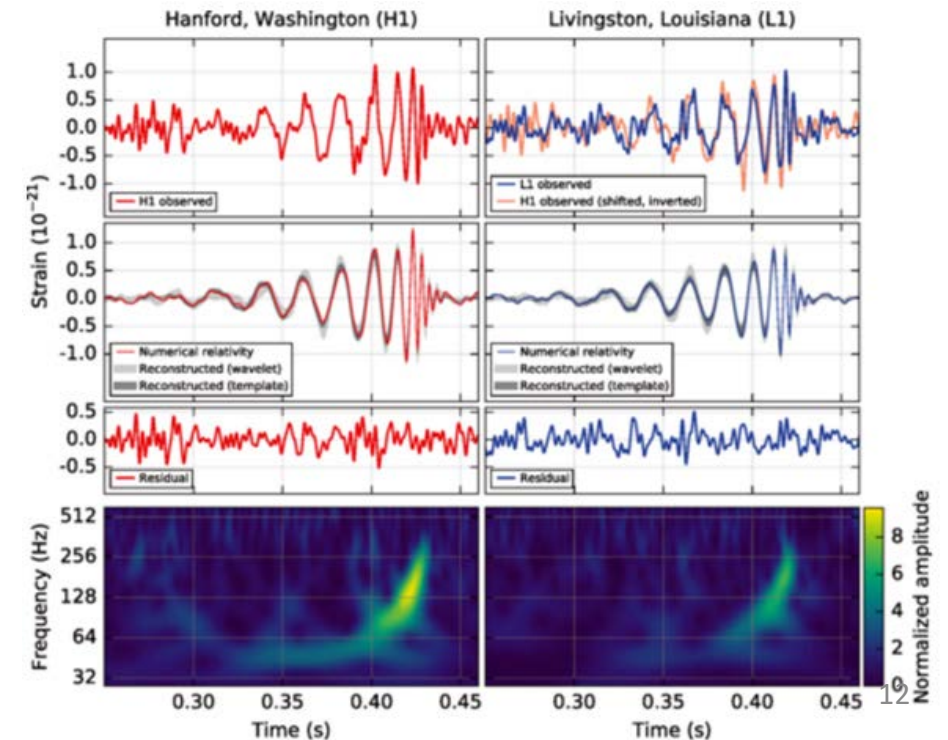
zeptonewton sensitivity in a standing wave trap
w/ 3-D laser feedback cooling –

Gravitational waves



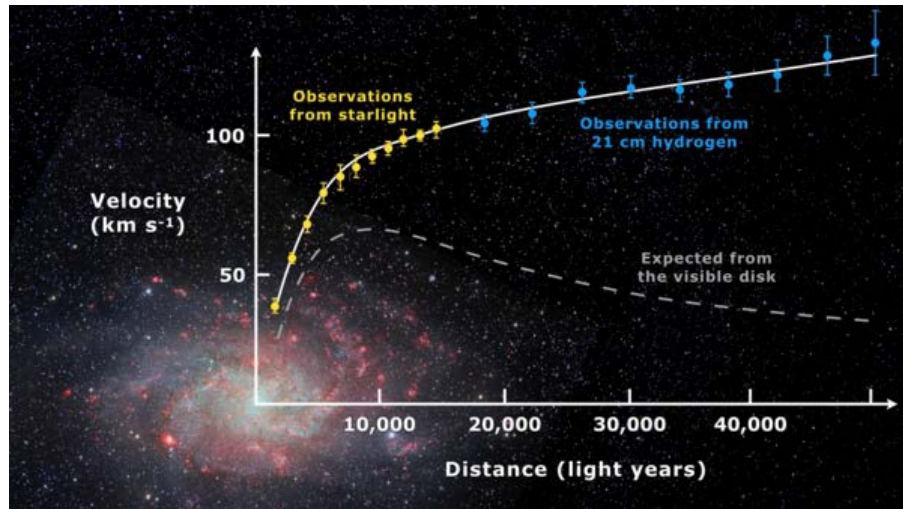
- One of last predictions of GR to be tested
- Discovered by LIGO Sep 2015 !!
- Sources:
 - Inspirals of astrophysical objects
 - Inflation, Phase transitions, etc.

B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration)
Phys. Rev. Lett. **116**, 061102 (2016).



Dark Matter

- Zwicky detected its influence on cluster motions in the mid-1930s, Rubin detected it in galaxy rotation curves in the 1960s
- Its nature and origin is one of the most vexing problems in physics and astronomy
- One of strongest hints of physics beyond the Standard Model



Galactic Rotation Curves: much more gravity than from ordinary matter

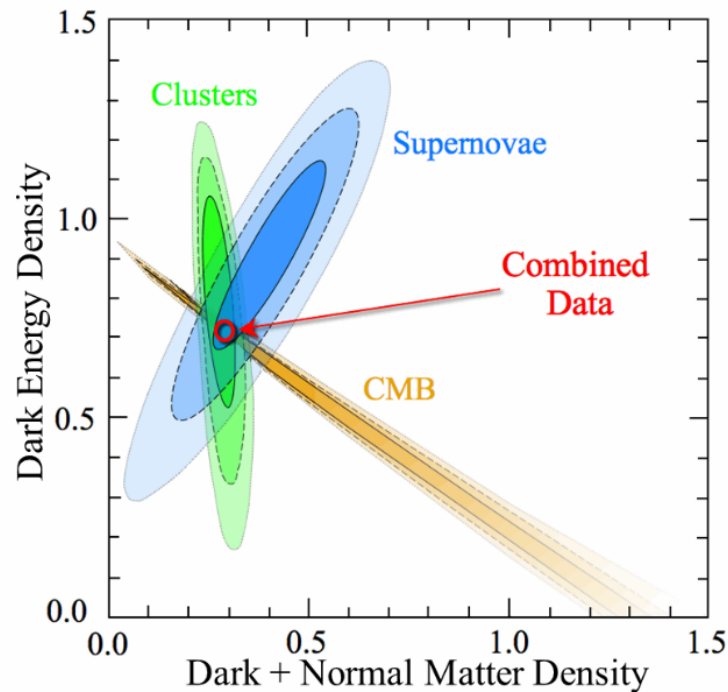


Gravitational Lensing: “Bullet” cluster – separation of lensing from visible matter

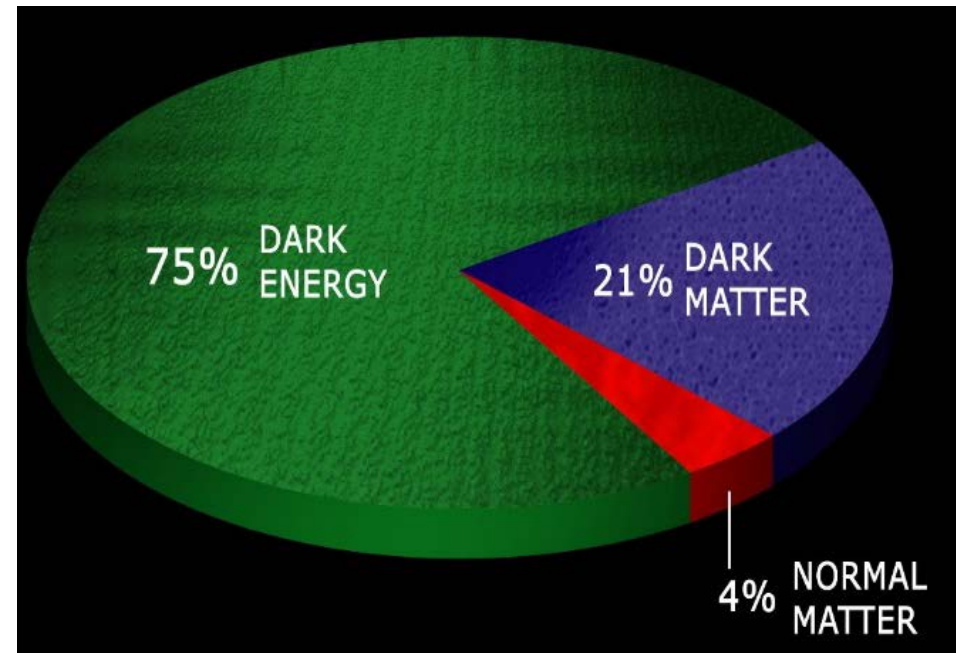
Also: Structure formation, Cosmic Microwave background, ...

The Dark Sector

- New particles, fields, and forces may be lurking undetected because of their extremely weak coupling to ordinary matter
- ~95% of the energy content of our universe is unknown to us!

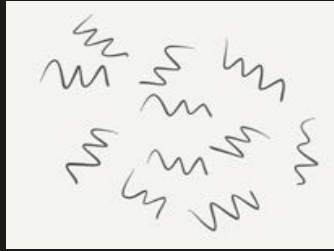


Evidence supporting a universe of dark matter and energy. Image: Kowalski *et al*, *Astrophys. J.* **686**, 794 (2008).

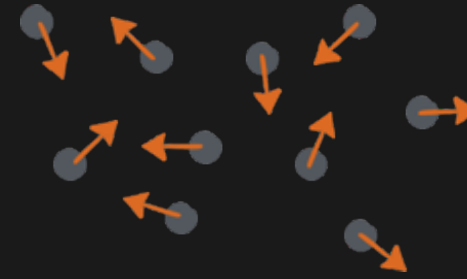


Our best estimate for composition of universe. Image credit: ADMX

Dark Matter

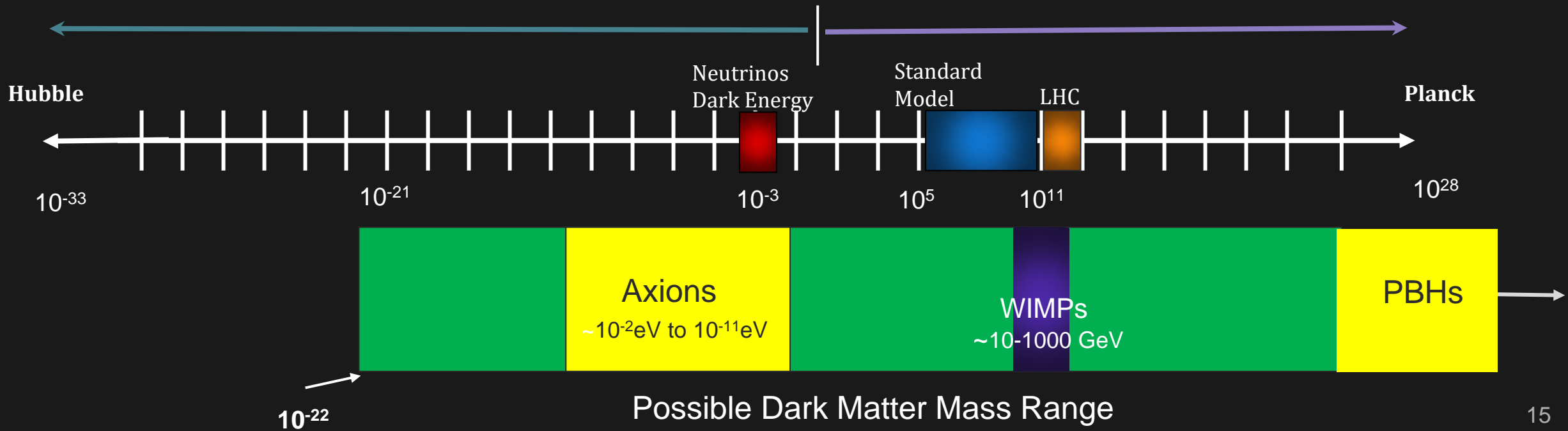


Wave-like

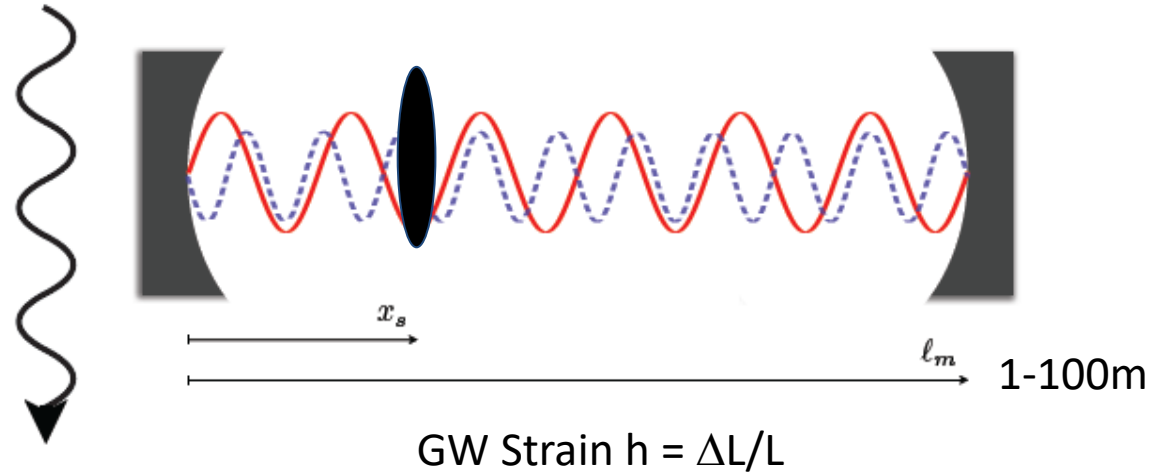
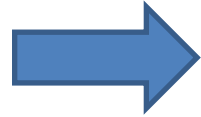


Particle-like

$1\text{eV}/c^2$



Novel gravitational wave detector for > 10 kHz



- Laser intensity changed to match trap frequency to GW frequency
- For a 10m cavity, $h \sim 10^{-22} \text{ Hz}^{-1/2}$ at high frequency (100kHz)
- Limited by thermal noise in sensor (not laser shot noise) \rightarrow much better at high frequency!

Frequency landscape for gravitational waves

COSMIC CREATORS OF GRAVITATIONAL WAVES

SUPERMASSIVE BINARY BLACK HOLES

EXPLOSIONS

BINARY MERGERS

AXIONS

STARS FALLING INTO SUPERMASSIVE BLACK HOLES

PRIMORDIAL BLACK HOLES

COMPACT BINARIES

WAVE FREQUENCY (HZ)

Extreme Low Frequency

High Frequency

10^{-16} 10^{-14} 10^{-12} 10^{-10} 10^{-8} 10^{-6} 10^{-4} 10^{-2} 1 10^2 10^4 10^6

DETECTION METHODS

CMB Observations

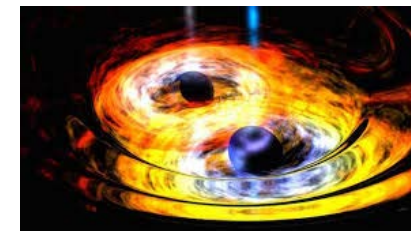
Pulsar Timing Arrays

LISA

LIGO Virgo

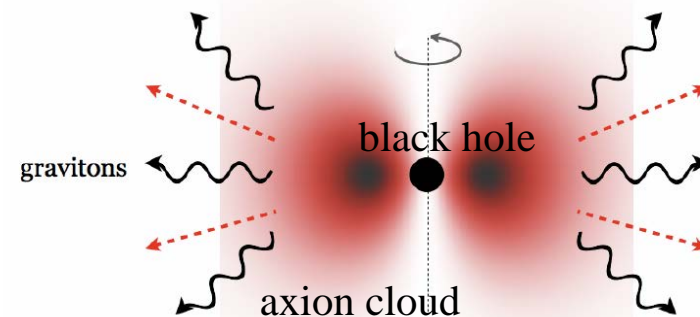
Levitated Sensors

PBHs:

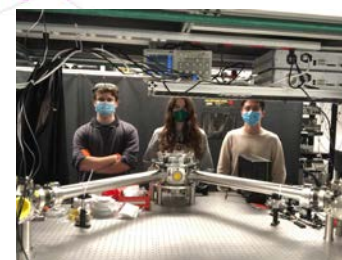
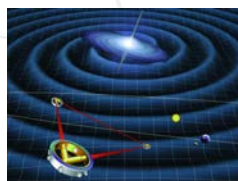


Distance to source: 1 kpc (within our galaxy)

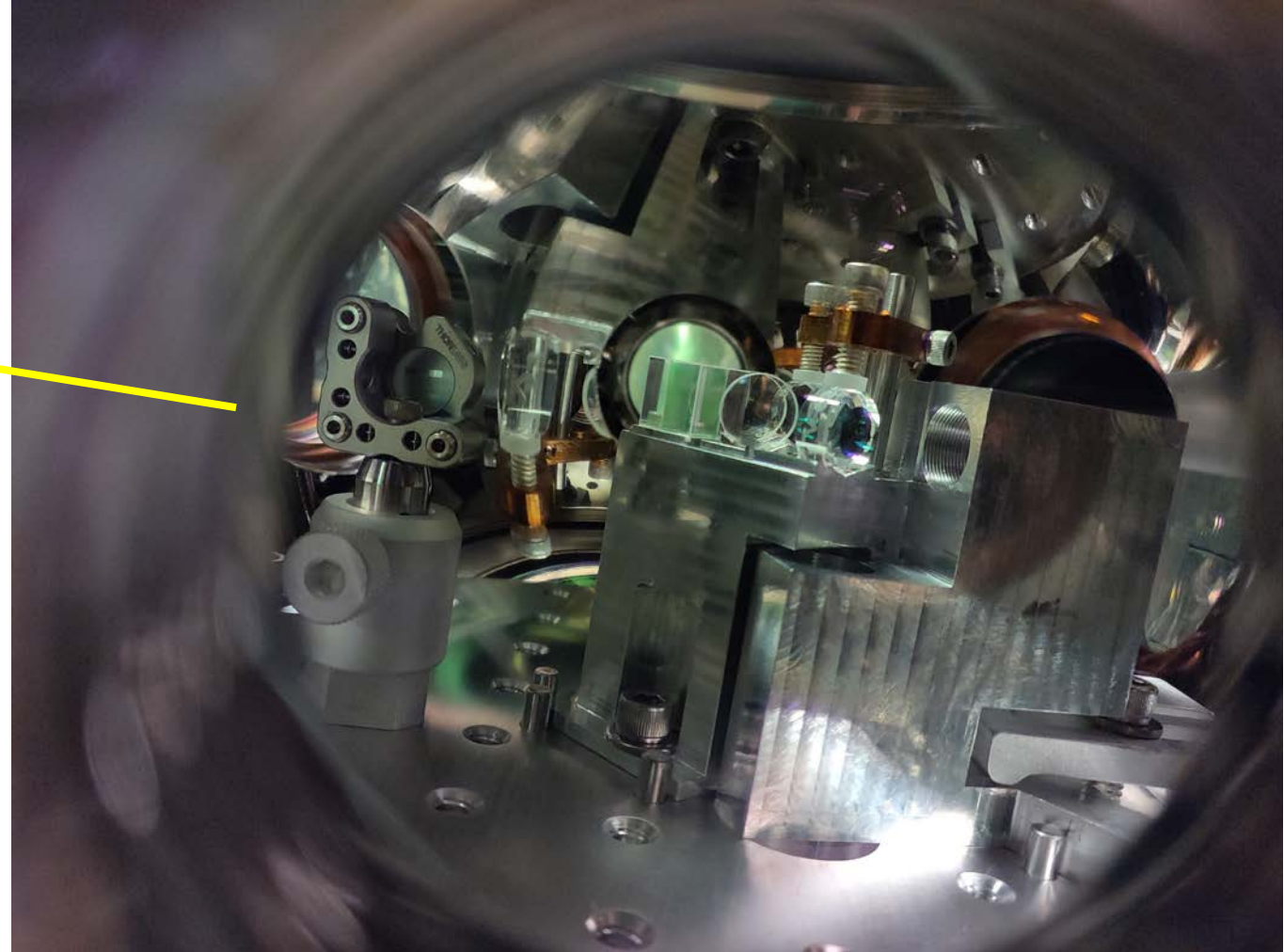
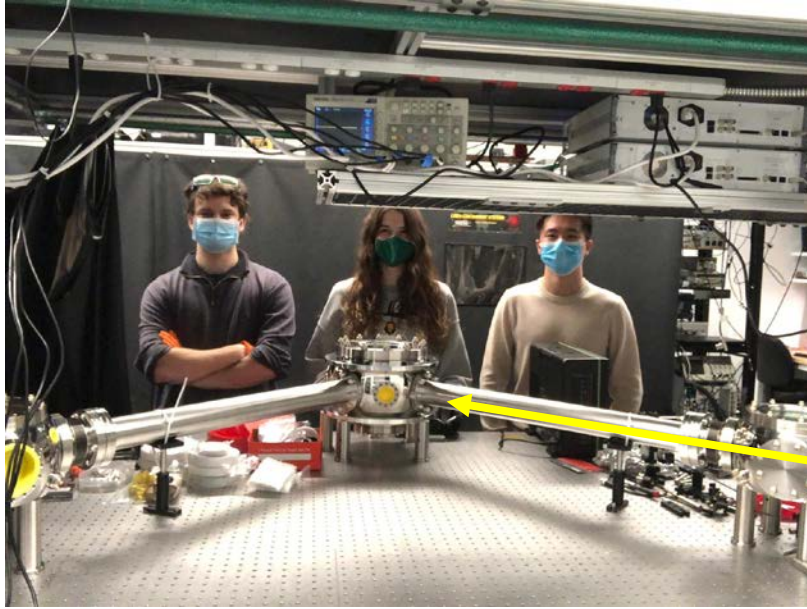
Axions:



Distance to source: 10 kpc (within our galaxy)



1-meter Levitated Sensor Detector (LSD) prototype

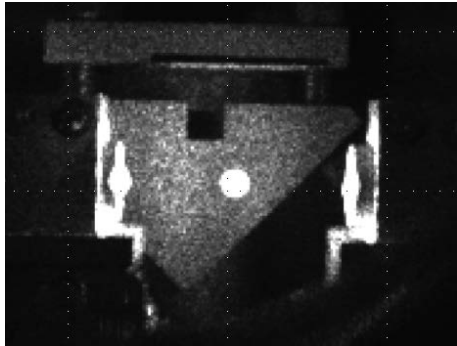


The LSD Collaboration -- Experiment: AG(PI), Nancy Aggarwal (co-I), Geroge Winstone (PD), Shelby Klomp (G), Aaron Wang (G), Peter J. Pauzauskie (UW), Greg Felsted (UW, G) Theory: Jacob Sprague (G), Shane Larson (co-I), Vicky Kalogera (co-I)

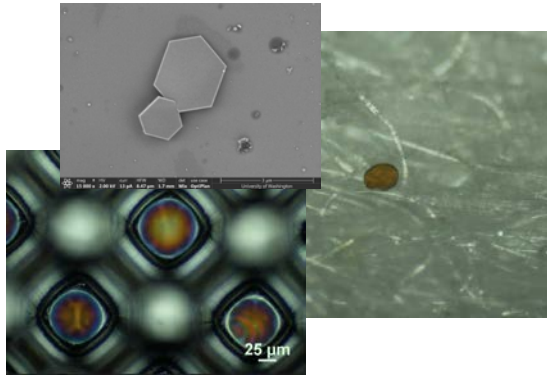
Levitated object optimization: disc/plate vs sphere

Ideal sensor: high mass, high frequency, low photon recoil

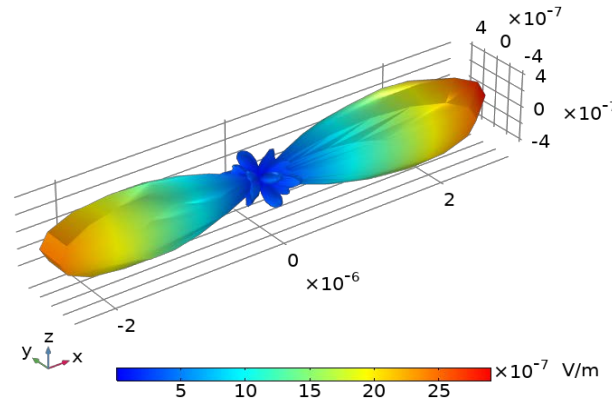
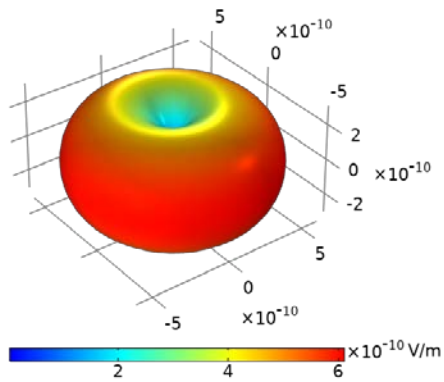
SiO₂ spheres



SiO₂ disks & NaYF hexagons



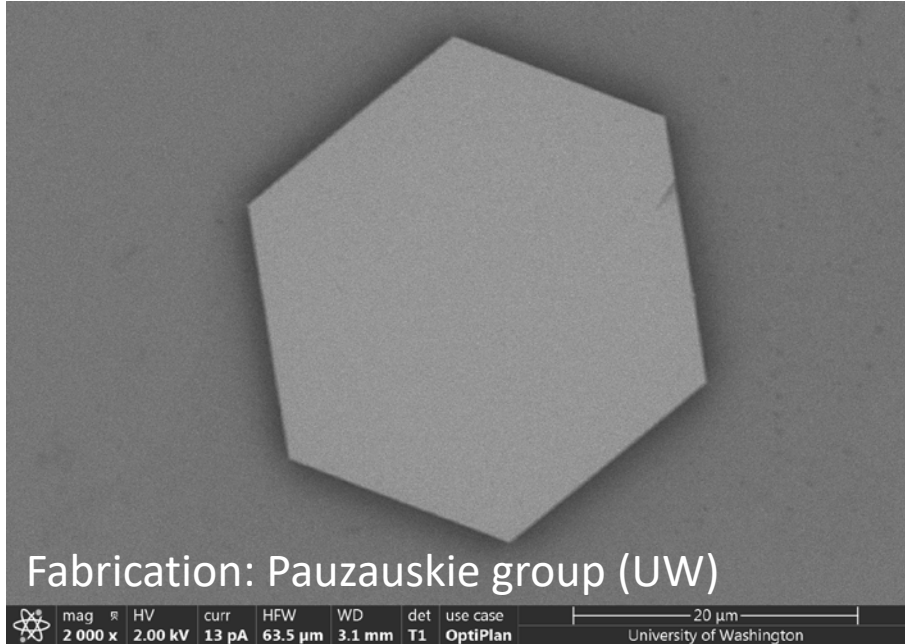
SiN/SiO₂/SiN stacks



- High-index end caps with low-index spacer
- Increased mass and internal reflections

- Less isotropic scattering reduces photon recoil noise

Optical trapping of NaYF hexagonal prisms



Hexagonal plate “stands up” with normal vector to face along the axis of the standing wave trap

Axial frequency (z-) highest due to intensity gradient along optical lattice

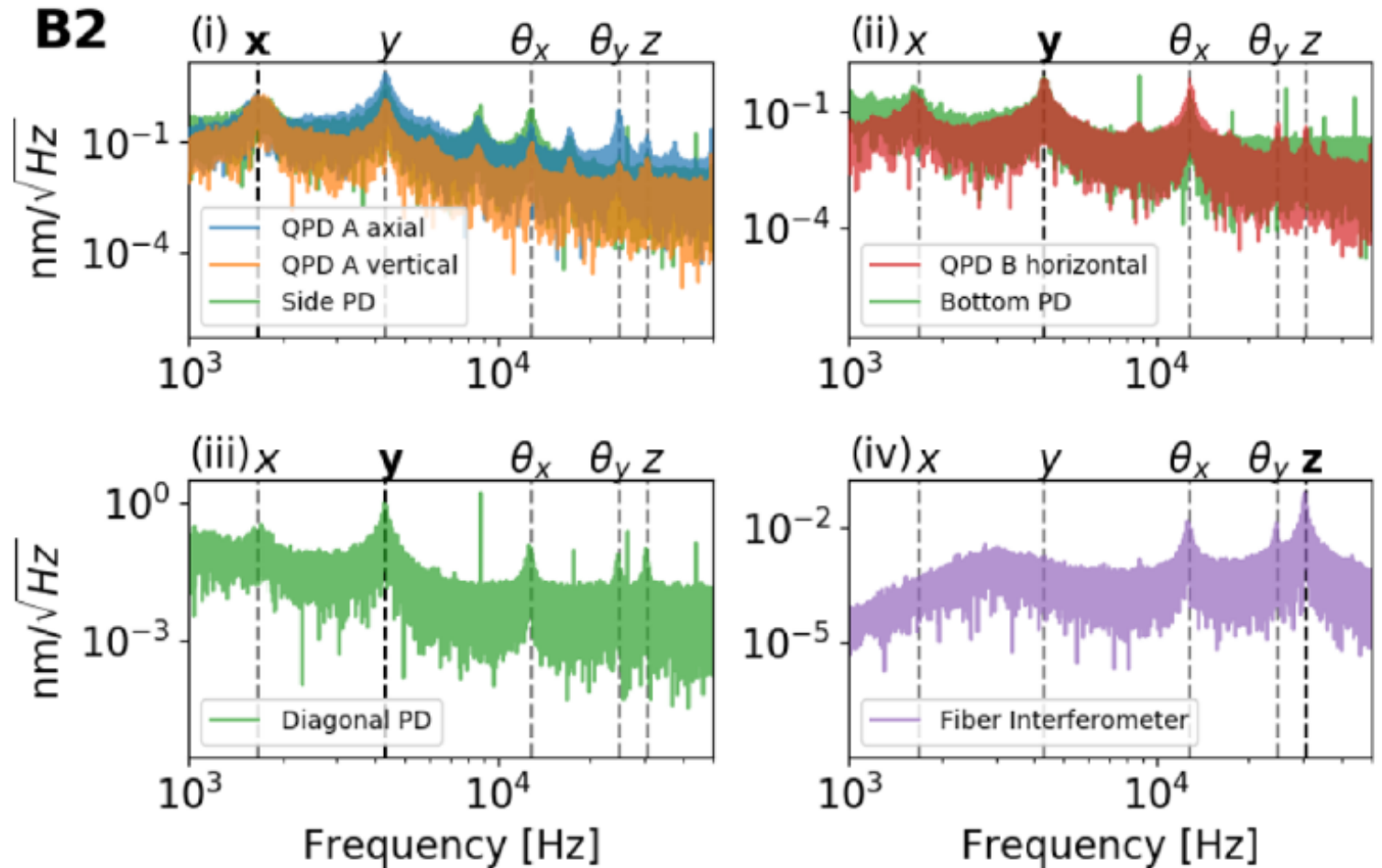


Figure of merit η for GW detection

$$h_{\text{limit}} = \frac{4}{\omega_0^2 L} \sqrt{\frac{k_B T_{\text{CM}} \gamma_g b}{M} \left[1 + \frac{\gamma_{\text{sc}}}{N_i \gamma_g} \right]} H(\omega_0)$$

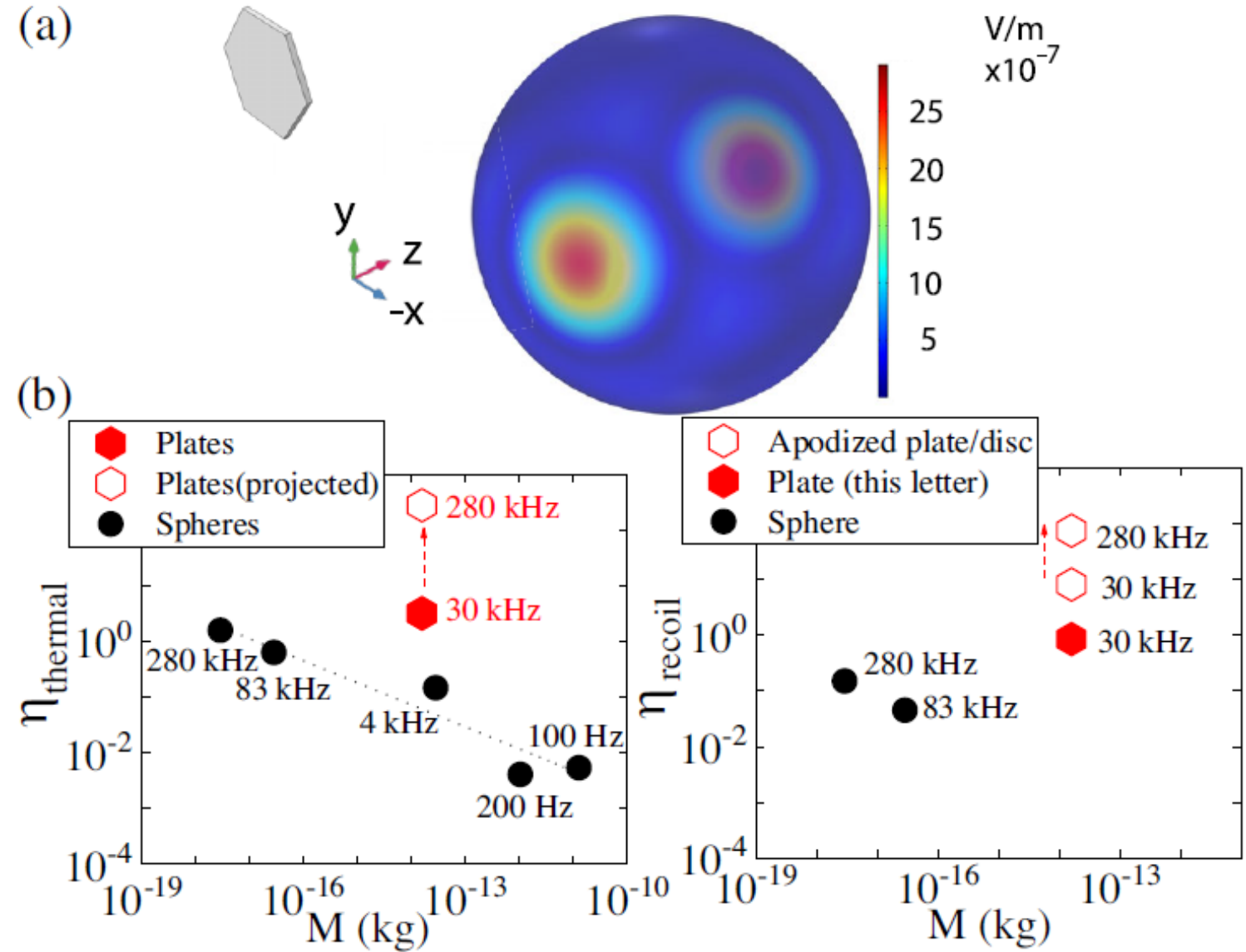
Ideal sensor: high mass, high frequency, low photon recoil

Thermal noise dominated

$$\eta_{\text{thermal}} \equiv \omega_0^2 \sqrt{M \rho l}$$

Photon recoil noise dominated

$$\eta_{\text{recoil}} \equiv \omega_0^{3/2} \sqrt{M / \gamma_{\text{sc}}}$$



Future goal: Laser refrigeration of NaYF material

- Particles could withstand higher intensities
→ Pathway for MHz trapping frequencies → high bandwidth accelerometers and gravitational wave detectors

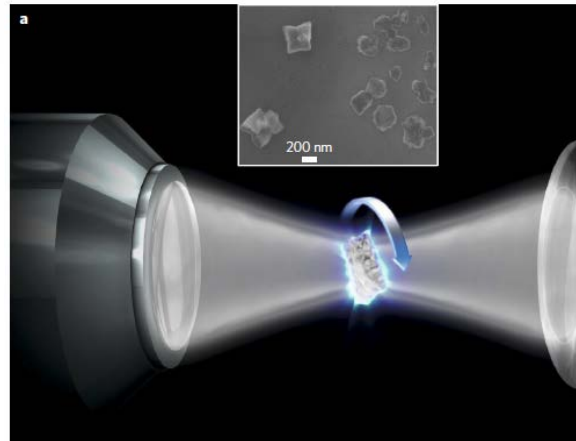
LETTERS

DOI: 10.1038/s41566-017-0005-3

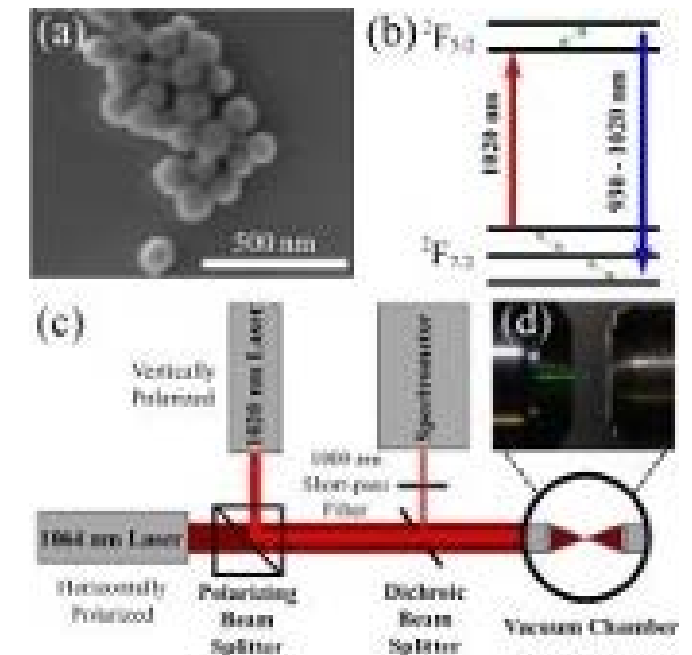
nature
photonics

Laser refrigeration, alignment and rotation of levitated $\text{Yb}^{3+}:\text{YLF}$ nanocrystals

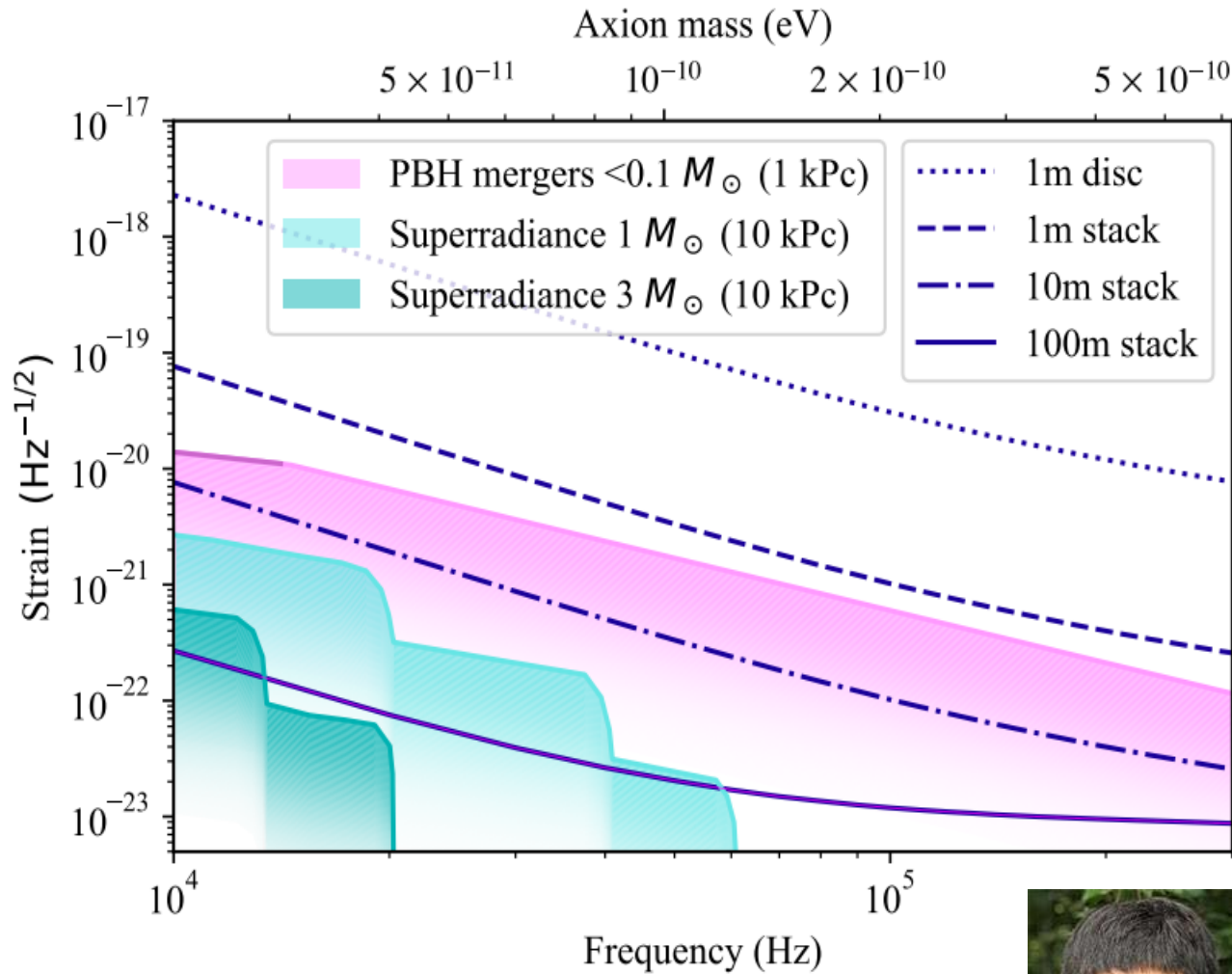
A. T. M. Anishur Rahman and P. F. Barker*



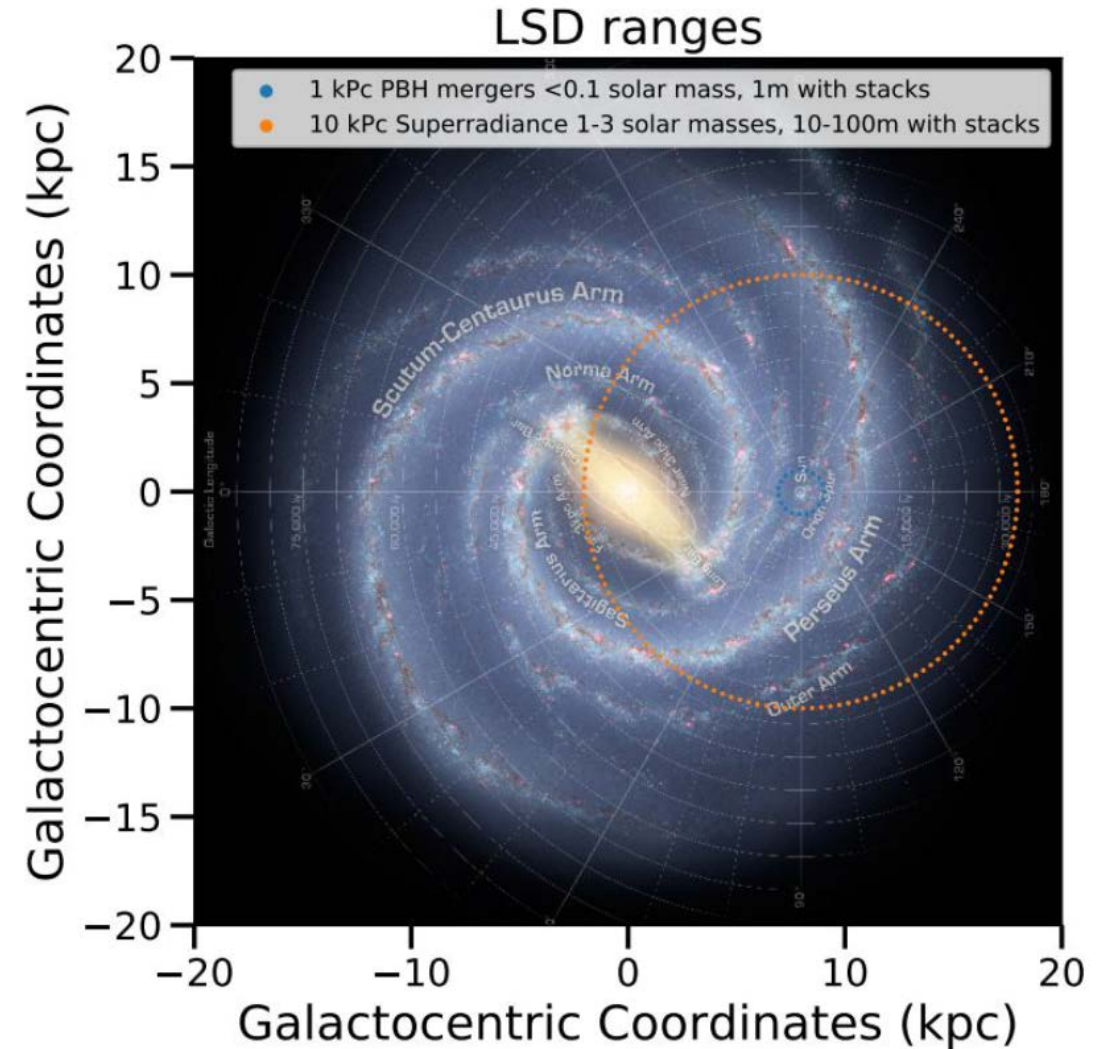
Danika R. Luntz-Martin, R. Greg Felsted, Siamak Dadras, Peter J. Pauzauskie, and A. Nick Vamivakas, "Laser refrigeration of optically levitated sodium yttrium fluoride nanocrystals," *Opt. Lett.* **46**, 3797-3800 (2021)



Search Range for the Levitated Sensor Detector (LSD)



Source modeling: Next talk by Jacob Sprague:



Summary

- Calibrated zeptonewton force sensing with optically levitated nanospheres
→ new approach for high frequency gravitational waves
Geometric methods to mitigate recoil heating, improve sensitivity
Materials allow solid-state cooling → higher frequency detectors
- Source modeling in process - PBHs, BH superradiance (axions, vector bosons)
- Future outlook: network of detectors planned: GOLDEN (Gravitational wave Observatory Levitated DETector Network (UC Davis – N. Aggarwal, University College London – P. Barker, others?))

Acknowledgements



From left to right: Alexey Grinin (PD), Nancy Aggarwal (UC Davis), William Eom (G), Nia Burrell (G), Aaron Zhiyuan Wang (G), Chethn Galla (G), Andrew Geraci (PI), Cris Montoya (PD), George Winstone (PD), Chloe Lohmeyer (G), Shelby Klomp (G), Evan Weisman (PD), Andrew Poverman (G), Eduardo Alejandro (G), Andrew Laeuger (UG)

Collaborators (GW):
Nancy Aggarwal (Davis)
Peter Barker (UCL)
Peter Pauzauskie (UW)

Astro Theory:
Asimina Arvanitaki (Perimeter)
Masha Baryakhtar (Perimeter)
Mae Hwee Teo (Stanford)
Shane Larson (NU)
Vicky Kalogera (NU)