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# A quantum generalization of the Cooper-Frye formula

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## Outline

- Notion of freeze-out hypersurface
- The Cauchy problem
- Solution of the Klein-Gordon equation for the initial value problem
- Measurement of the momentum spectrum
- Single-particle spectrum: local thermodynamic equilibrium
- Approximations
- Reduction to the Cooper-Frye formula
- Radiation of particles from a system with a finite lifetime
- Concluding remarks

#### Notion of freeze-out hypersurface

Usually, the sharp freeze-out hypersurface is defined by the parameter P(t, r) taking the critical value  $P_c$  on the hypersurface:

$$P(t, \mathbf{r}) = P_{\rm c}$$

The parameter for defining the hypersurface can be selected as:

• The density of particles:  $n(t, \mathbf{r}) = n_c$ 

D.Adamova (CERES Collaboration), Phys. Rev. Lett. 90, 022301 (2003).

• The energy density:  $\epsilon(t, \mathbf{r}) = \epsilon_{c}$ 

J. Sollfrank, P. Huovinen, and P.V. Ruuskanen, Eur. Phys. J. C 6, 525 (1999).

V.N. Russkikh and Y.B. Ivanov, Phys. Rev. C 76, 054907 (2007).

#### • The temperature: $T(t, \mathbf{r}) = T_c$

H. von Gersdorff, L. McLerran, M. Kataja, and P.V. Ruuskanen, Phys. Rev. D34, 794 (1986).
 P. Huovinen, Eur. Phys. J. A 37, 121 (2008).

#### The freeze-out hypersurfaces in quasi-four-dimensional form: $(x, y) \rightarrow r_{T} = \sqrt{x^2 + y^2}$



[1] D. Anchishkin, V. Vovchenko, and L.P. Csernai, Phys. Rev. C 87, 014906 (2013).

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#### Radiation of quantum fields from space-like hypersurface: the Cauchy problem



Figure: Sketch of a freeze-out hypersurface for a spherically symmetric fireball expansion. Left panel: Constant initial time  $t = t_0$  and constant spatial boundary  $R = R_0$ . Right panel: Dependence of the initial radiation time on the radius and the dependence of the spatial boundary on time (solid curve).

#### Initial value problem

$$(\partial_{\mu}\partial^{\mu} + m^2)\hat{\varphi}(x) = 0,$$

where  $\partial_{\mu}\partial^{\mu} = \partial_t^2 - \vec{\nabla}^2, \hbar = 1, \quad c = 1.$ 

The Cauchy problem or the initial conditions for this equation are specified on a space-like hyper-surface

$$\left. \hat{\varphi}(x^{0}, \mathbf{x}) \right|_{x^{0}=t_{0}} = \left. \hat{\Phi}_{0}(\mathbf{x}) , \quad \left. \frac{\partial \hat{\varphi}(x^{0}, \mathbf{x})}{\partial x^{0}} \right|_{x^{0}=t_{0}} = \left. \hat{\Phi}_{1}(\mathbf{x}) \right.$$

Equation for the evolution of  $\hat{\varphi}(x)$  together with the initial conditions can be written as:

$$(\partial_{\mu}\partial^{\mu}+m^2)\,\hat{\varphi}(\mathbf{x})\,=\,\delta(\mathbf{x}^0-t_0)\hat{\Phi}_1(\mathbf{x})\,+\,\delta'(\mathbf{x}^0-t_0)\hat{\Phi}_0(\mathbf{x})$$

The Green's function:

$$G_{R}(x-y) = -\int \frac{d^{4}k}{(2\pi)^{4}} \frac{e^{-ik \cdot (x-y)}}{(k_{0}+i\delta)^{2} - \mathbf{k}^{2} - m^{2}}$$

#### Solution of the Klein-Gordon equation

Solution:

$$\hat{\varphi}(x) = \int d^4 y \, \delta(y^0 - t_0) \left[ G_R(x - y) \frac{\overleftrightarrow{\partial}}{\partial y^0} \hat{\Phi}(y^0, \mathbf{y}) \right]$$

where we update notations of the initial conditions:

$$\hat{\Phi}_0(\boldsymbol{y}) = \hat{\Phi}(\boldsymbol{y}^0, \boldsymbol{y})\big|_{\boldsymbol{y}^0 = t_0}, \quad \hat{\Phi}_1(\boldsymbol{y}) = \partial \hat{\Phi}(\boldsymbol{y}^0, \boldsymbol{y}) / \partial \boldsymbol{y}^0\big|_{\boldsymbol{y}^0 = t_0}$$

For the arbitrary space-like hyper-surface,  $\sigma(y)$ , solution looks like

$$\hat{\varphi}(x) = \int_{\sigma} d\sigma^{\mu}(y) \left[ G_{R}(x-y) \frac{\overleftrightarrow{\partial}}{\partial y^{\mu}} \hat{\Phi}(y) \right].$$

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#### Measurement of the momentum spectrum

$$\begin{split} \hat{k}\psi_{\mathbf{k}}(\mathbf{r}) \, = \, \mathbf{k}\psi_{\mathbf{k}}(\mathbf{r}) & \rightarrow \qquad \phi(\mathbf{r}) \, = \, \sum_{\mathbf{k}} \langle \psi_{\mathbf{k}} \big| \phi \rangle \, \psi_{\mathbf{k}}(\mathbf{r}) \\ P(\mathbf{k}) \, = \, \left| \langle \psi_{\mathbf{k}} \big| \phi \rangle \right|^2 \end{split}$$

We assume that the detector measures asymptotic momentum eigenstates, i.e. that it acts by projecting the emitted single-particle state onto

$$\phi_{\mathbf{k}}^{\text{out}}(t,\mathbf{r}) = e^{-i\omega_{k}t+i\mathbf{k}\cdot\mathbf{r}}, \quad \text{with} \quad \omega_{k} = \sqrt{m^{2} + \mathbf{k}^{2}}$$
$$\hat{\varphi}(t,\mathbf{r}) = \int \frac{d^{3}k}{(2\pi)^{3} 2\omega_{k}} \left[ b(\mathbf{k}) e^{-i\omega_{k}t+i\mathbf{k}\cdot\mathbf{r}} + b^{+}(\mathbf{k}) e^{i\omega_{k}t-i\mathbf{k}\cdot\mathbf{r}} \right]$$

The single-particle spectrum and two-particle spectrum

 $P_1(\mathbf{p}) = \left\langle \ b^+(\mathbf{p}) \ b(\mathbf{p}) \ 
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angle \ , \quad P_2(\mathbf{p}_1,\mathbf{p}_2) = \left\langle \ b^+(\mathbf{p}_1) \ b^+(\mathbf{p}_2) \ b(\mathbf{p}_2) \ b(\mathbf{p}_1) \ 
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angle$ 

#### Connection of the operators *b* and $b^+$ with the field of the fireball $\hat{\Phi}(x)$

Calculation of operators  $b(\mathbf{k})$  and  $b^+(\mathbf{k})$ :

$$b(\mathbf{k}) = \int d^3 r \, e^{i(\omega_k t - \mathbf{k} \cdot \mathbf{r})} \, i \stackrel{\leftrightarrow}{\partial}_t \hat{\varphi}(t, \mathbf{r}) = -\int d^4 y \, \delta(y^0 - t_0) \left[ e^{-i(\omega_k y^0 - \mathbf{k} \cdot \mathbf{y})} i \frac{\stackrel{\leftrightarrow}{\partial}}{\partial y^0} \hat{\Phi}^+(y) \right]$$

For the arbitrary space-like hyper-surface  $\sigma(x)$ 

$$\begin{split} b(\mathbf{k}) &= i \int_{\sigma} d\sigma^{\mu}(x) \left[ f_{\mathbf{k}}^{*}(x) \frac{\stackrel{\leftrightarrow}{\partial}}{\partial x^{\mu}} \hat{\Phi}(x) \right] \\ b^{+}(\mathbf{k}) &= -i \int_{\sigma} d\sigma^{\mu}(x) \left[ f_{\mathbf{k}}(x) \frac{\stackrel{\leftrightarrow}{\partial}}{\partial x^{\mu}} \hat{\Phi}^{+}(x) \right] \\ , \end{split}$$

where  $f_{\mathbf{k}}(x) = e^{-i(\omega_k x^0 - \mathbf{k} \cdot \mathbf{x})}$ .

#### Single-particle spectrum [2]

$$2E_{p}\frac{dN}{d^{3}p} = \left\langle b^{+}(\mathbf{p}) b(\mathbf{p}) \right\rangle$$
$$2E_{p}\frac{dN}{d^{3}p} = i \int d^{4}x_{1} d^{4}x_{2} \,\delta(x_{1}^{0} - t_{0}) \,\delta(x_{2}^{0} - t_{0}) \left[ f_{\mathbf{p}}(x_{1}) f_{\mathbf{p}}^{*}(x_{2}) \frac{\overleftrightarrow{\partial}}{\partial x_{1}^{0}} \frac{\overleftrightarrow{\partial}}{\partial x_{2}^{0}} G^{<}(x_{2}, x_{1}) \right]$$

Covariant form of the single-particle spectrum

$$2E_{\rho}\frac{dN}{d^{3}\rho}=i\int d\sigma^{\mu}(x_{1})\,d\sigma^{\nu}(x_{2})\,\left[f_{\mathbf{p}}(x_{1})\,f_{\mathbf{p}}^{*}(x_{2})\,\frac{\stackrel{\leftrightarrow}{\partial}}{\partial x_{1}^{\mu}}\frac{\stackrel{\leftrightarrow}{\partial}}{\partial x_{2}^{\nu}}\,G^{<}(x_{2},x_{1})\right]$$

where we have defined the correlation function or the lesser Green's function:

$$i G^{<}(x_2, x_1) = \pm \left\langle \hat{\Phi}^+(x_1) \hat{\Phi}(x_2) \right\rangle$$

Here the plus sign reads for bosons and the minus sign for fermions.

[2] D. Anchishkin, J.Phys.G 49, 055109 (2022)

Concluding remarks

#### Single-particle spectrum: local thermodynamic equilibrium

$$G^{<}(x_2, x_1) = \int \frac{d^4k}{(2\pi)^4} e^{-ik\cdot x} G^{<}(X; k),$$

where  $x = x_1 - x_2$  and  $X = (x_1 + x_2)/2$ .

$$2E_{\rho}\frac{dN}{d^{3}p} = i \int \frac{d^{4}k}{(2\pi)^{4}} d^{4}X d^{4}x \,\delta(X^{0} - t_{0}) \,\delta(x^{0}) \,e^{-i(p-k)\cdot x} \left(p^{0} + k^{0}\right)^{2} G^{<}(X;k) \,,$$
  
where  $p^{0} = E_{\rho} = \sqrt{m^{2} + \mathbf{p}^{2}}$ .

Covariant form of the single-particle spectrum - 1

$$2E_{p}\frac{dN}{d^{3}p} = i\int \frac{d^{4}k}{(2\pi)^{4}} \int_{\sigma} d\sigma^{\mu}(x_{1})d\sigma^{\nu}(x_{2})e^{-i(p-k)\cdot(x_{1}-x_{2})}$$
$$\times (p+k)_{\mu}(p+k)_{\nu}G^{<}(X;k)$$

Here we use: 
$$d^4x_1 \, \delta(x_1^0 \, - \, t_0) p^0 = d\sigma^{\mu}(x_1) p_{\mu}.$$

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#### Approximations

We assumed that the particles are free on the freeze-out hypersurface, this leads to the approximation:

$$G^{<}(X;k) pprox G^{<}_0(X;k)$$

In a system with slowly varying inhomogeneity, the free Green's function can be represented as:

$$i G_0^{<}(X; k^0, \mathbf{k}) = \frac{\pi}{\omega_k} \, \delta(k^0 - \omega_k) \, f_{_{\rm BE}}(X; k^0)$$

We assume that the correlation function G<sup><</sup>(x<sub>1</sub>, x<sub>2</sub>) differs significantly from zero only if |x<sub>1</sub> − x<sub>2</sub>| → 0:

$$\int_{V} d^{4}x \,\delta(x^{0}) \, e^{-i(p-k)\cdot x} = (2\pi)^{3} \delta^{3}(\mathbf{p}-\mathbf{k})$$

$$2E_{\rho}\,\frac{dN}{d^{3}\rho}\,=\,\int d^{4}X\,\delta(X^{0}-t_{0})\,2E_{\rho}\,f_{\rm BE}(X;E_{\rho})$$

In a homogeneous system (no dependence on X) the formula reduces to:

$$rac{dN}{d^3p} = V f_{_{
m BE}}(E_{
ho})$$

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#### $(1) + (2) \rightarrow$ Reduction to the Cooper-Frye formula

$$2E_{\rho}\frac{dN}{d^{3}p} = i\int \frac{d^{4}k}{(2\pi)^{4}}d^{4}Xd^{4}X\,\delta(X^{0}-t_{0})\,\delta(x^{0})\,e^{-i(\rho-k)\cdot x}\left(p^{0}+k^{0}\right)^{2}G^{<}(X;k)$$

With account for the equal-time initial conditions  $dt = dX^0 = 0$  we get:

$$d^4 X \,\delta\left(X^0 - t_0\right) \,p^0 \,=\, d\sigma_\mu(X) \,p^\mu$$



where  $p^0 = E_p = \sqrt{m^2 + p^2}$ , and u(X) is the four-velocity at the point  $X = (X^0, \mathbf{X})$  given on a space-like hypersurface.

#### Taking into account a finite size of the fireball

$$\int_{V} d^{4}x \,\delta(x^{0}) \, e^{-i(p-k)\cdot x} \neq (2\pi)^{3} \delta^{3}(\mathbf{p}-\mathbf{k})$$

A multiparticle system has weak inhomogeneity and is in local thermodynamic equilibrium

$$2E_{p}\frac{dN}{d^{3}p} = \int \frac{d^{3}k}{(2\pi)^{3}2\omega_{k}} \int_{\sigma} d\sigma^{\mu}(x_{1})d\sigma^{\nu}(x_{2})e^{-i(p-k)\cdot(x_{1}-x_{2})}(p+k)_{\mu}(p+k)_{\nu}f(X;k)$$

where 
$$p^0 = E_{\rho} = \sqrt{m^2 + p^2}$$
,  $k^0 = \omega_k = \sqrt{m^2 + k^2}$ .

The case of global thermodynamic equilibrium

$$2E_p \frac{dN}{d^3p} = \int \frac{d^3k}{(2\pi)^3 2\omega_k} f(k) \left| \int_{\sigma} d\sigma^{\mu}(x) (p+k)_{\mu} e^{-i(p-k) \cdot x} \right|^2.$$

where 
$$p^0 = E_p$$
,  $k^0 = \omega_k$ .

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#### Radiation of particles from a system with a finite lifetime $\tau = 1/\gamma$

Dependence of the isotropic pion spectrum on the  $|\mathbf{p}| = p$ ,  $d^3p = \sin\theta d\theta d\varphi p^2 dp$ :

$$\frac{dN}{4\pi p^2 dp} = V \int_0^\infty \frac{dk^0}{\pi} \frac{(E_p + k_0)^2}{2E_p} \frac{m\gamma}{(k_0^2 - E_p^2)^2 + (m\gamma)^2} f_{\rm BE}(k^0)$$

Our goal is to compare this spectrum with the Cooper-Frye formula for the isotropic radiation:

$$\frac{dN}{4\pi p^2 dp} = V f_{\rm BE}(E_{\rho})$$



#### Radiation of particles from a system with a finite lifetime $\tau = 1/\gamma$

A comparison of the pion transverse spectra at midrapidity,

$$(p_x, p_y, p_z) o (\varphi, p_T, y), \ d^3p = \sin \theta d\theta d\varphi p^2 dp, \ m_T = \sqrt{m_\pi^2 + p_T^2} \ \text{and} \ y = 0:$$

$$\frac{dN}{2\pi p_T dp_T dy} = V \int_0^\infty \frac{dk^0}{\pi} (m_T + k_0)^2 \frac{m\gamma}{(k_0^2 - m_T^2)^2 + (m\gamma)^2} f_{\rm BE}(k^0)$$

Our goal is to compare this spectrum with the Cooper-Frye formula for the isotropic radiation:



It is interesting to compare our results with pion spectrum versus transverse momentum at midrapidity in the transverse momentum range 0.6 GeV/c  $< p_T < 12$  GeV/c measured in Pb-Pb collisions at  $\sqrt{s_{NN}} = 2.76$  TeV

B. Abelev, J. Adam, D. Adamová et al. (The ALICE Collaboration), Eur. Phys. J. C 74, 3108 (2014) 🛛 🖌 🖉 🍃 🖌 🚊 🕨 🛓

## Conclusions

We obtained the quantum generalization of the Cooper-Frye for the single-particle spectrum in the framework of the following assumptions:

- The multi-particle system has a weak inhomogeneity;
- The multi-particle system is in local thermodynamic equilibrium;
- Particles are in free states on the freeze-out hypersurface;
- The correlation function G<sup><</sup>(x<sub>1</sub>, x<sub>2</sub>) differs significantly from zero only if the differences of the arguments x<sub>1</sub> - x<sub>2</sub> are close to zero.

The latter assumption (4) cannot be applied if the system is close to a second-order phase transition, when the correlation length becomes sufficiently large. At the same time, this approximation excludes quantum effects.

## Thank you for attention !

#### Acknowledgement

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#### Interference of the radiation from the freeze-out hypersurface

We can argue that formulae derived is a generalization of the Cooper-Frye formula, which takes into account quantum effects. Indeed, one can rewrite

$$2E_{\rho}\frac{dN}{d^{3}\rho} = \int_{\sigma} d\sigma^{\mu}(x_{1}) d\sigma^{\nu}(x_{2}) J_{\mu\nu}(x_{1}-x_{2},\boldsymbol{p}),$$

were we define the tensor  $J_{\mu
u}$ 

$$J_{\mu\nu}(x_1 - x_2, \mathbf{p}) = e^{-ip \cdot (x_1 - x_2)} \int \frac{d^3k}{(2\pi)^3 2\omega_k} e^{ik \cdot (x_1 - x_2)} (\mathbf{p} + k)_{\mu} (\mathbf{p} + k)_{\nu} f(X; k)$$

The last integral includes interference of waves with different momenta k, which are projected onto the out-state with momentum p. There is no such interference in the Cooper-Frye formula, where k = p is taken.