

NON-PERTURBATIVE EFFECTS ACROSS DARK MATTER MODELS:
FROM THE RELIC DENSITY TO INDIRECT DETECTION

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in collaboration with S. Vogl and J. Bollig (2308.14594)

PARTICLE INTERPRETATION OF DM AND FREEZE-OUT

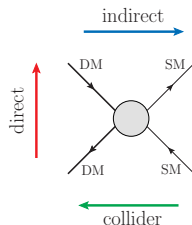
- DM from many compelling (gravitational) observations

- DM as a particle: many candidates (Bertone and Hooper [1605.04909])

- Any model has to comply with

$$\Omega_{\text{DM}} h^2(M_{\text{DM}}, M_{\text{DM}'}, \alpha_{\text{DM}}, \alpha_{\text{SM}}) = 0.1200 \pm 0.0012$$

- ◇ from CMB anisotropies with Λ CDM *Planck Collab. Results 2018*



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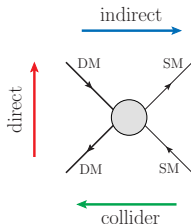
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THERMAL FREEZE-OUT GONDOLO AND GELMINI (1991)

- Boltzmann equation for DM (χ)

$$\frac{dn_\chi}{dt} + 3Hn_\chi = -\langle\sigma v\rangle(n_\chi^2 - n_{\chi,\text{eq}}^2)$$

- relevant processes $\chi\chi \leftrightarrow \text{SM SM}$, $\chi\chi \leftrightarrow \chi'\chi'$
- decoupling from $H \sim n_{\text{eq}} \langle\sigma_{\text{ann}} v_{\text{rel}}\rangle$

$$H \simeq \frac{T^2}{M_{\text{Pl}}}, \quad \langle\sigma_{\text{ann}} v_{\text{rel}}\rangle \simeq \frac{\alpha^2}{M^2}, \quad \frac{T}{M} \approx \frac{1}{25}$$

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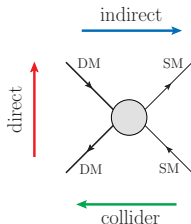
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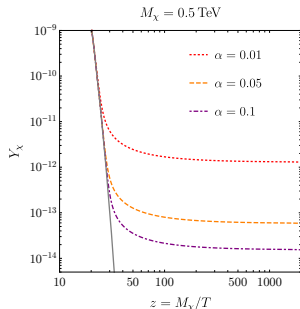
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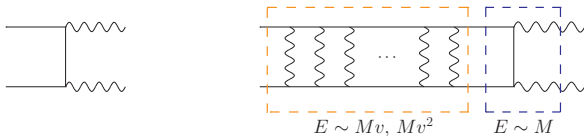
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DM MODELS WITH MEDIATORS

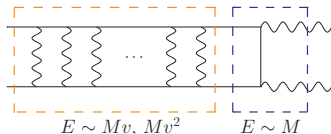
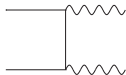
- DM and/or coannihilating partners interact with gauge bosons and scalars



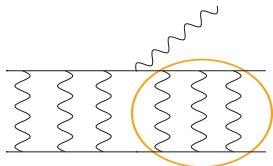
- repeated soft interactions: Sommerfeld enhancement and bound states Hisano, Matsumoto, Nojiri [hep-ph/0212022], [hep-ph/0307216]; B. von Harling and K. Petraki [1407.7874]; Beneke, Hellmann, Ruiz-Femenia [1411.6924] ...

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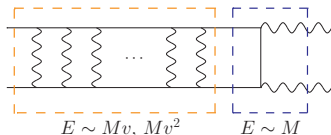
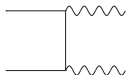
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$$\langle\sigma_{\text{eff}} v_{\text{rel}}\rangle = \langle\sigma_{\text{ann}} v_{\text{rel}}\rangle + \sum_n \langle\sigma_{\text{bsf}}^n v_{\text{rel}}\rangle \frac{\Gamma_{\text{ann}}^n}{\Gamma_{\text{ann}}^n + \Gamma_{\text{bsd}}^n}$$

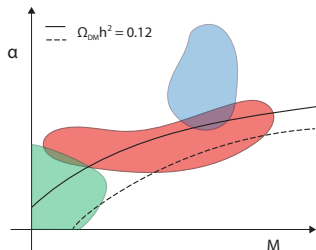
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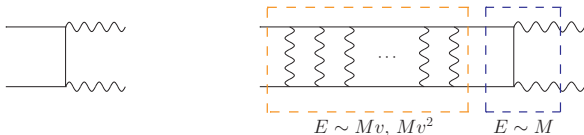
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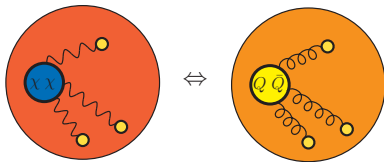
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HEAVY QUARKONIUM IN QGP

- $(Q\bar{Q})$ dynamics in a thermal environment
- NRQCD and pNRQCD

Matsui and Satz (1986); Laine, Philipsen, Romatschke and Tassler

[hep-ph/0611300]; Brambilla, Ghiglieri, Petreczky, Vairo [0804.0993]

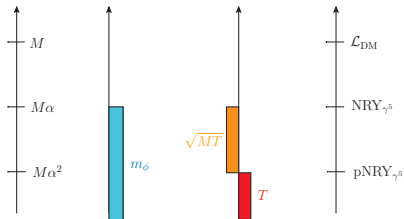
PNREFT FOR (PSEUDO)SCALAR MEDIATORS

SIMPLIFIED MODEL

$$\mathcal{L} = \bar{X}(i\cancel{\partial} - M)X + \frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{1}{2}m^2\phi^2 - \frac{\lambda}{4!}\phi^4 - \bar{X}(g + ig_5\gamma_5)X\phi + \mathcal{L}_{\text{portal}},$$

M. B. Wise and Y. Zhang [1407.4121]; K. Kainulainen, K. Tuominen and V. Vaskonen [1507.04931]

- g_5 allows for pair annihilations $X\bar{X} \rightarrow \phi\phi$ at $\mathcal{O}(v^0)$
- pair annihilations $\rightarrow \Omega_{\text{DM}}h^2$ and signals for indirect detection



- Derive $\sigma_{v_{\text{rel}}}$ and Γ in the unified framework of NREFTs

- Include near-threshold effects for $\Omega_{\text{DM}}h^2$

$$M \gg M_\alpha \gg \pi T \gtrsim M_\alpha^2, \quad M_\alpha \gtrsim m_\phi$$

- Include near-threshold effects for **indirect detection** [prospects for CTA]

NREFTs FOR PSEUDO-SCALAR MEDIATORS

- integrate out the hard scale $E \sim M$ [$T = 0$ matching] M. E. Luke and A. V. Manohar (hep-ph/9610534)

$$\mathcal{L}_{\text{NRY}\gamma_5} = \mathcal{L}_\psi^{\text{bilinear}} + \mathcal{L}_\chi^{\text{bilinear}} + \mathcal{L}_{4\text{-fermions}} + \mathcal{L}_{\text{scalar}}$$

- interactions between NR fermions/antifermions and scalar mediators

$$\mathcal{L}_\psi^{\text{bilinear}} = \psi^\dagger \left(i\partial_0 - g\phi + g_5 \frac{\sigma \cdot [\nabla\phi]}{2M} - g_5^2 \frac{\phi^2}{2M} + \frac{\nabla^2}{2M} \right) \psi$$

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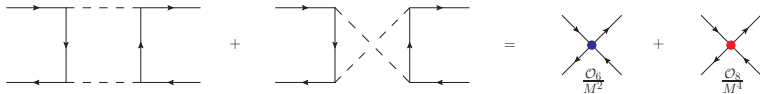
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- four-fermion interactions encode the annihilations into light degrees of freedom [$2^{S+1}L_J$]

G. T. Bodwin, E. Braaten and G. P. Lepage (hep-ph/9407339)



$$(\mathcal{L}_{4\text{-fermions}})_{d=6} = \frac{f(^1S_0)}{M^2} \mathcal{O}(^1S_0) + \frac{f(^3S_1)}{M^2} \mathcal{O}(^3S_1), \quad \mathcal{O}(^1S_0) = \psi^\dagger \chi \chi^\dagger \psi, \quad \mathcal{O}(^3S_1) = \psi^\dagger \sigma \chi \cdot \chi^\dagger \sigma \psi$$

$$\text{Im}[f(^1S_0)] = 2\pi\alpha\alpha_5 \mathcal{F}(m_\phi/M), \quad \mathcal{F}(m_\phi/M) = \sqrt{1 - \frac{m_\phi^2}{M^2}} \left(1 - \frac{m_\phi^2}{2M^2}\right)^{-2}, \quad \text{Im}[f(^3S_1)] = 0$$

PNREFT FOR (PSEUDO)SCALAR MEDIATORS

- integrating out the scale $M\alpha$ and m_ϕ : EFT for DM pairs and **ultrasoft/thermal scalar mediators**

SB and V. Shtabovenko (2106.06472 and 2112.10145)

$$\mathcal{L}_{\text{pNRY}_{\gamma 5}} = \int d^3\mathbf{r} \varphi^\dagger(\mathbf{r}, \mathbf{R}, t) \left\{ i\partial_0 - H(\mathbf{r}, \mathbf{p}, \mathbf{P}, \mathbf{S}_1, \mathbf{S}_2) - 2g\phi(\mathbf{R}, t) - g\frac{r^j}{4} \left[\nabla_R^i \nabla_R^j \phi(\mathbf{R}, t) \right] - g\phi(\mathbf{R}, t) \frac{\nabla_R^2}{M^2} \right\} \varphi(\mathbf{r}, \mathbf{R}, t) + \frac{1}{2}(\partial^\mu \phi(\mathbf{R}, t))^2 - \frac{m^2}{2}\phi(\mathbf{R}, t)^2 - \frac{\lambda}{4!}\phi(\mathbf{R}, t)^4$$

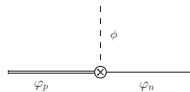
- soft and ultrasoft pseudoscalar interactions are suppressed; Hamiltonian and potential

$$H(\mathbf{r}, \mathbf{p}, \mathbf{P}, \mathbf{S}_1, \mathbf{S}_2) = 2M + \frac{\mathbf{p}^2}{M} + \frac{\mathbf{P}^2}{4M} - \frac{\mathbf{P}^4}{4M^3} + V(\mathbf{r}, \mathbf{p}, \mathbf{P}, \mathbf{S}_1, \mathbf{S}_2) + \dots$$

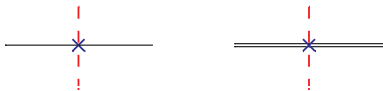
$$V(\mathbf{r}, \mathbf{p}, \mathbf{P}, \mathbf{S}_1, \mathbf{S}_2) = V^{(0)} + \frac{V^{(1)}}{M} + \frac{V^{(2)}}{M^2} + \dots, \quad \boxed{V^{(0)} = -\frac{\alpha}{r} e^{-m_\phi r}}$$

- account for **above-threshold** and **below-threshold states** (for heavy quarkonium X. Yao and T. Mehen [1811.07027])

$$\varphi_{ij}(t, \mathbf{r}, \mathbf{R}) = \int \frac{d^3\mathbf{P}}{(2\pi)^3} \left[\sum_n e^{-iE_n t + i\mathbf{P}\cdot\mathbf{R}} \Psi_n(\mathbf{r}) S_{ij} \varphi_n(\mathbf{P}) + \int \frac{d^3\mathbf{p}}{(2\pi)^3} e^{-iE_p t + i\mathbf{P}\cdot\mathbf{R}} \Psi_p(\mathbf{r}) S_{ij} \varphi_p(\mathbf{P}) \right]$$



ANNIHILATIONS AND DECAYS



$$\text{Im}[f(^1S_0)] = 2\pi\alpha\alpha_5\mathcal{F}(m_\phi/M)$$

ANNIHILATION SCATTERING STATES

$$(\sigma_{\text{ann}} v_{\text{rel}})(\mathbf{p}) = \frac{1}{M^2} \left\{ \left(\text{Im}[f(^1S_0)] + \text{Im}[g(^1S_0)] \frac{v_{\text{rel}}^2}{4} \right) S_{\text{ann}}^0(\zeta, \xi) + \frac{\text{Im}[f(^3P_0)] + 5\text{Im}[f(^3P_2)]}{12} v_{\text{rel}}^2 S_{\text{ann}}^1(\zeta, \xi) \right\}$$

BOUND-STATE DECAY

$$\Gamma_{\text{ann}}^{nS} = \frac{|R_{nS}(0)|^2}{\pi M^2} \text{Im}[f(^1S_0)], \quad \Gamma_{\text{ann}}^{nP_J} = \frac{|R'_{nP}(0)|^2}{\pi M^4} \text{Im}[f(^3P_J)], \quad E_{nl} = -\gamma_{nl}^2(\xi) \frac{M\alpha^2}{4n^2}$$

- $\zeta = \frac{\alpha}{v_{\text{rel}}}$, $\xi = \frac{M\alpha}{2m_\phi}$; let us keep in mind the dependence of $\Gamma_{\text{ann}}^{1S} \propto \alpha_5\alpha^4$

BOUND-STATE FORMATION AND DISSOCIATION

BOUND-STATE FORMATION

$$\sigma_{\text{BSF}}^n v_{\text{rel}} = \frac{\alpha}{120} [(\Delta E_n^p)^2 - m_\phi^2]^{\frac{5}{2}} [|\langle \mathbf{p} | \mathbf{r}^2 | n \rangle|^2 + 2|\langle \mathbf{p} | r^i r^j | n \rangle|^2] + 2\alpha [(\Delta E_n^p)^2 - m_\phi^2]^{\frac{1}{2}} \left| \langle \mathbf{p} | \frac{\nabla_{\mathbf{r}}^2}{M^2} | n \rangle \right|^2 - \frac{\alpha}{3} [(\Delta E_n^p)^2 - m_\phi^2]^{\frac{3}{2}} \text{Re} \left[\langle \mathbf{p} | \frac{\nabla_{\mathbf{r}}^2}{M^2} | n \rangle \langle n | \mathbf{r}^2 | \mathbf{p} \rangle \right] [1 + n_B(\Delta E_n^p)]$$



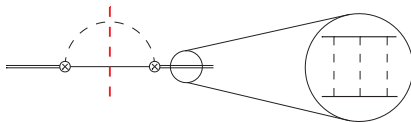
BOUND-STATE DISSATIATION

$$\Gamma_{\text{BSD}}^n = \int_{|\mathbf{k}| > k_{\min}} \frac{d^3 k}{(2\pi)^3} n_B(|\mathbf{k}|) \sigma_{\text{ion}}^n(|\mathbf{k}|), \quad \sigma_{\text{ion}}^n(|\mathbf{k}|) = \frac{g_X^2}{g_\phi g_n} \frac{M^2 v_{\text{rel}}}{4|\mathbf{k}|^2} \sigma_{\text{BSF}}^n$$

BOUND-STATE FORMATION AND DISSOCIATION

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$$\begin{aligned} \sigma_{\text{BSF}}^n v_{\text{rel}} &= \frac{\alpha}{120} \left[(\Delta E_n^p)^2 - m_\phi^2 \right]^{\frac{5}{2}} \left[|\langle \mathbf{p} | r^2 | n \rangle|^2 + 2 |\langle \mathbf{p} | r^i r^j | n \rangle|^2 \right] + 2\alpha \left[(\Delta E_n^p)^2 - m_\phi^2 \right]^{\frac{1}{2}} \left| \langle \mathbf{p} | \frac{\nabla_{\mathbf{r}}^2}{M^2} | n \rangle \right|^2 \\ &\quad - \frac{\alpha}{3} \left[(\Delta E_n^p)^2 - m_\phi^2 \right]^{\frac{3}{2}} \text{Re} \left[\langle \mathbf{p} | \frac{\nabla_{\mathbf{r}}^2}{M^2} | n \rangle \langle n | r^2 | \mathbf{p} \rangle \right] [1 + n_B(\Delta E_n^p)] \end{aligned}$$

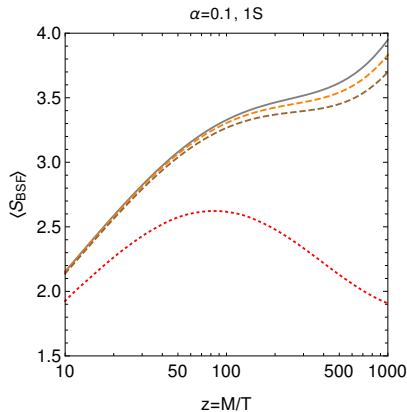
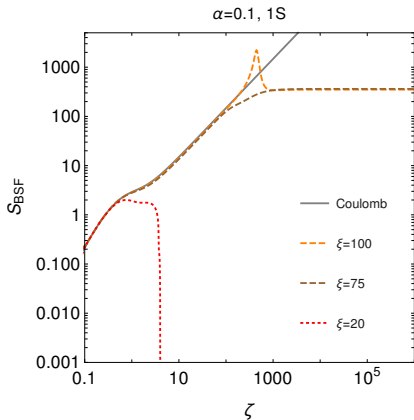


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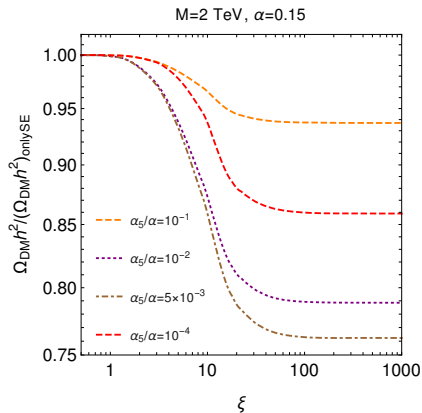
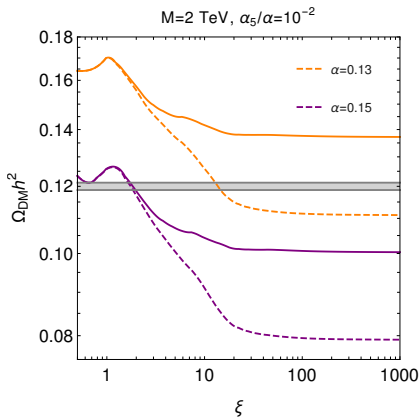
THERMAL BOUND-STATE FORMATION

$$\zeta = \frac{\alpha}{v_{\text{rel}}}, \quad \xi = \frac{M\alpha}{2m_\phi}, \quad \sigma_{\text{BSF } v_{\text{rel}}}^{1\text{S}} \equiv \underbrace{\frac{\pi\alpha^4}{M^2}}_{\sigma_0} S_{\text{BSF}}^{1\text{S}}(\zeta, \xi), \quad \langle \sigma_{\text{BSF } v_{\text{rel}}}^{1\text{S}} \rangle = \frac{2\sigma_0}{\sqrt{\pi}} \int_0^\infty du S_{\text{BSF}}^{1\text{S}} \left(\alpha \sqrt{\frac{z}{4u}} \right) \frac{\sqrt{ue} \frac{|E_b|}{T}}{e^{u+\frac{|E_b|}{T}} - 1}$$



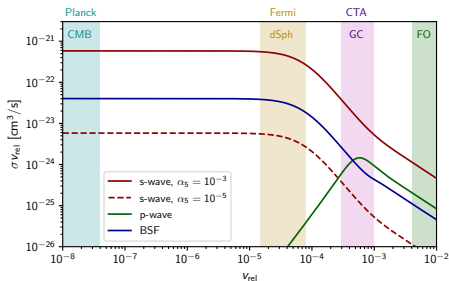
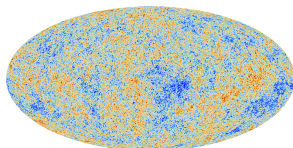
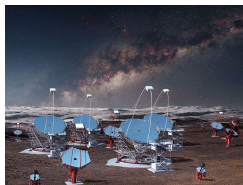
RELIC DENSITY

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INDIRECT DETECTION

- look for **DM residual annihilations** on the flux of cosmic γ -rays



- photons from cascade decays

$$\mathcal{L}_{\text{portal}} = -\mu_{\phi h} \phi \left(H^\dagger H - \frac{v^2}{2} \right) - \frac{1}{2} \lambda_{\phi h} \phi^2 \left(H^\dagger H - \frac{v^2}{2} \right),$$

- $(X\bar{X} \rightarrow \phi\phi) \times (\phi \rightarrow \text{SMSM}')$
- [left plot] $M = 1 \text{ TeV}$, $m_\phi = 1 \text{ GeV}$, $\alpha = 0.1$

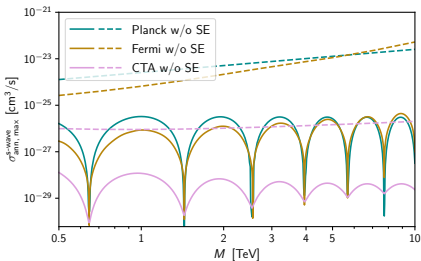
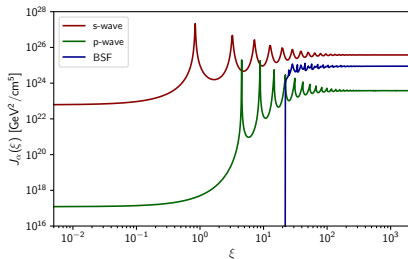
GENERALIZED J FACTORS

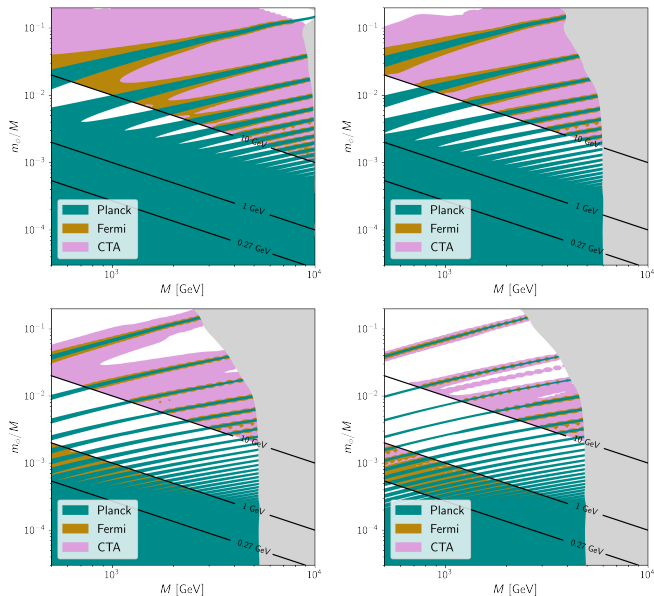
- The contribution of DM annihilation to the differential photon flux is F. Ferrer and D. R. Hunter (1306.6586)

$$\frac{d\Phi_\gamma}{dE_\gamma} = \frac{1}{16\pi M^2} \frac{dN_\gamma}{dE_\gamma} \int_{\Delta\Omega} d\Omega \int_0^\infty d\psi \int d^3 v_1 \int d^3 v_2 f_X(r(\psi, \Omega), \vec{v}_1) f_X(r(\psi, \Omega), \vec{v}_2) \sigma_{\text{rel}},$$

- $\sigma_{\text{ann}}^{\text{s-wave}} \equiv \sigma_{\text{ann}}^{(0,0)} = \text{Im}[f(^1S_0)]/M^2$ and $\sigma_{\text{ann}}^{\text{p-wave}} \equiv \sigma_{\text{ann}}^{(1,0)} = (\text{Im}[f(^3P_0)] + 5\text{Im}[f(^3P_2)])/12M^2$

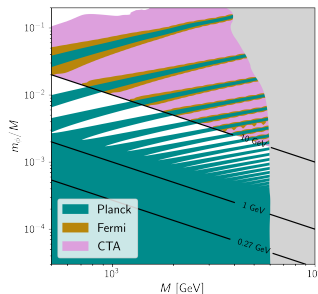
$$\frac{d\Phi_\gamma}{dE_\gamma} = \frac{1}{16\pi M^2} \left(\sum_f B_f \frac{dN_\gamma^{(f)}}{dE_\gamma} \right) \left(\sigma_{\text{ann}}^{\text{s-wave}} J_{\text{ann},\alpha}^{(0,0)}(\xi) + \sigma_{\text{ann}}^{\text{p-wave}} J_{\text{ann},\alpha}^{(1,0)}(\xi) + \sigma_{\text{BSF}} J_{\text{BSF},\alpha}(\xi) \right).$$



EXCLUSION LIMITS FOR $\alpha_5/\alpha = 10^{-1}, 10^{-2}, 10^{-3}, 10^{-4}$ 

CONCLUSIONS

- DM particles and thermal freeze-out in the early universe:
non-relativistic particles in a thermal environment
- Focus on models with scalar force carriers between DM particles
⇒ develop NREFTs similar to heavy quarkonium at $T = 0$ and finite temperature



- Developed NREFTs and pNREFT techniques for determining $\sigma_i v_{\text{rel}}$ and Γ_i
- Include the leading effects of bound-state formation for both $\Omega_{\text{DM}} h^2$ and indirect detection
- J -factors for Sommerfeld effects and BSF
- Assessed CTA prospects for the model

- Outlook: include excited states and inspect their effects on $\Omega_{\text{DM}} h^2$ and indirect detection

[see T.Binder et al (2308.01336) for vector-mediator models and excited bound states on the thermal freeze-out]

$$\operatorname{Im}[f(^1S_0)] = 2\pi\alpha\alpha_5\mathcal{F}(\tilde{r}),$$

$$\operatorname{Im}[g(^1S_0)] = -\frac{8\pi\alpha\alpha_5}{3} \frac{\left(1 - \frac{13}{8}\tilde{r}^2 + \frac{5}{8}\tilde{r}^4 - \frac{3}{32}\tilde{r}^6\right)}{(1 - \tilde{r}^2/2)^2(1 - \tilde{r}^2)} \mathcal{F}(\tilde{r})$$

$$\operatorname{Im}[f(^3P_0)] = \frac{\pi}{6} \left[3\alpha \left(2 - \frac{\mathcal{G}(\tilde{r})}{3} \right) - \alpha_5\mathcal{G}(\tilde{r}) \right]^2 \mathcal{F}(\tilde{r}),$$

$$\operatorname{Im}[f(^3P_2)] = \frac{\pi}{15} (\alpha + \alpha_5)^2 \mathcal{G}(\tilde{r})^2 \mathcal{F}(\tilde{r})$$

- auxiliary functions with $\tilde{r} \equiv m_\phi/M$

$$\mathcal{F}(\tilde{r}) = \frac{\sqrt{1 - \tilde{r}^2}}{\left(1 - \frac{\tilde{r}^2}{2}\right)^2}, \quad \mathcal{G}(\tilde{r}) = \frac{1 - \tilde{r}^2}{1 - \frac{\tilde{r}^2}{2}}.$$

$$\int d^3 v f_X(\vec{r}, \vec{v}) = \rho_X(\vec{r}) \quad J_0 = \int d\Omega \int d\psi \rho_X(\vec{r}(\psi, \Omega))^2$$

- In the following, we assume that we can split up the cross section as

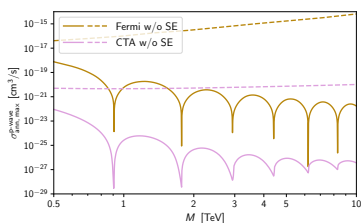
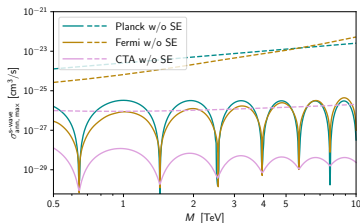
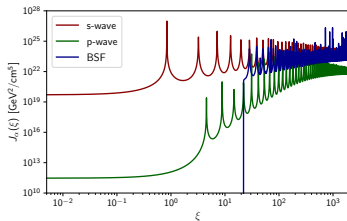
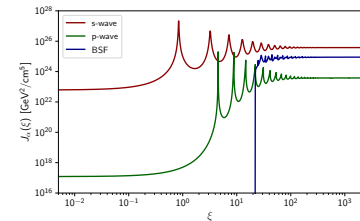
$$\sigma_{\text{rel}} = \sum_{l,j=0}^{\infty} \sigma_{\text{ann}}^{(l,j)} v_{\text{rel}}^{2(l+j)} S_{\text{ann}}^{(l)}(\zeta, \xi) + \sum_{n>l} \sigma_{\text{BSF}}^{(n,l)} S_{\text{BSF}}^{(n,l)}(\zeta, \xi).$$

- This motivates the definition of a generalized J-factor

$$J_{\text{ann},\alpha}^{(l,j)}(\xi) \equiv \int_{\Delta\Omega} d\Omega \int_0^\infty d\psi \int d^3 v_1 f_X(r(\psi, \Omega), \vec{v}_1) \int d^3 v_2 f_X(r(\psi, \Omega), \vec{v}_2) v_{\text{rel}}^{2(l+j)} S_{\text{ann}}^{(l)}(\alpha/v_{\text{rel}}, \xi)$$

- for f_X we use methods for the analysis of galactic dynamics that hold in collisionless systems in a quasi-equilibrium state J. Binney and S. Tremaine (Galactic Dynamics 2008)

OTHER LIMITS (dSPHS)

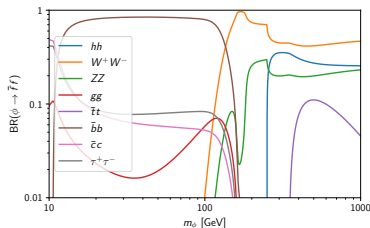


- [UP] Galactic center and Dwarf Spheroidal Galaxies
- [DOWN] Limits on the s-wave and p-wave dominated DM annihilation scenarios, $m_{\phi} = 10$ GeV, $\alpha = 0.1$

PHOTON SPECTRUM

- $\frac{dN_\gamma}{dE_\gamma}$ number of photons produced by annihilating dark matter per photon energy
- 1) Br_ϕ into various SM final states
- 2) γ produces by a given SM final state (decay or hadronization)
- 3) prescription for the boost of the decaying ϕ in the galactic rest frame
- high-mediator mass range: $m_\phi \geq 10$ GeV; $\phi \rightarrow \{hh, W^+W^-, ZZ, gg, \bar{t}t, \bar{b}b, \bar{c}c, \tau^+\tau^-\}$

A. Djouadi (hep-ph/0503172)



- low-mediator mass range $2m_{\pi^0} \leq m_\phi \leq 1$ GeV: decays into pion pairs play a large role;
 $\bar{X} + X \rightarrow 2\phi \rightarrow 4\pi_0 \rightarrow 8\gamma$ [we use here M.W. Winkler (1809.01876)]