

Dinesh Loomba on behalf of the MIGDAL collaboration University of New Mexico 8th CYGNUS Workshop on Directional Recoil Detection 12th December 2023, Sydney

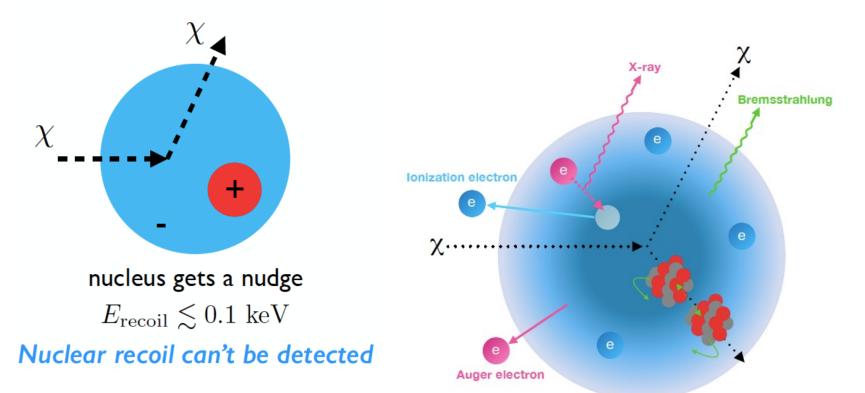




Arkady Migdal

The Migdal Effect

A.B. Migdal, Ionization of atoms accompanying α - and β -decay , J. Phys. USSR 4 (1941) 449



Christopher McCabe Dark Matter at the dawn of discovery? -Heidelberg, 11th April 2018

XENON1T collab arXiv:1907.12771

Recent application to low mass DM has led to huge interest in Migdal

Migdal effect calculations reformulated by **M. Ibe et al.** with ionisation probabilities for atoms and recoil energies relevant to Dark Matter searches.



Published for SISSA by D Springer

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Migdal effect in dark matter direct detection experiments

Masahiro Ibe,^{a,b} Wakutaka Nakano,^a Yutaro Shoji^a and Kazumine Suzuki^a

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 ^b Kavli IPMU (WPI), UTIAS, The University of Tokyo, Kashiwa, Chiba 277-8583, Japan

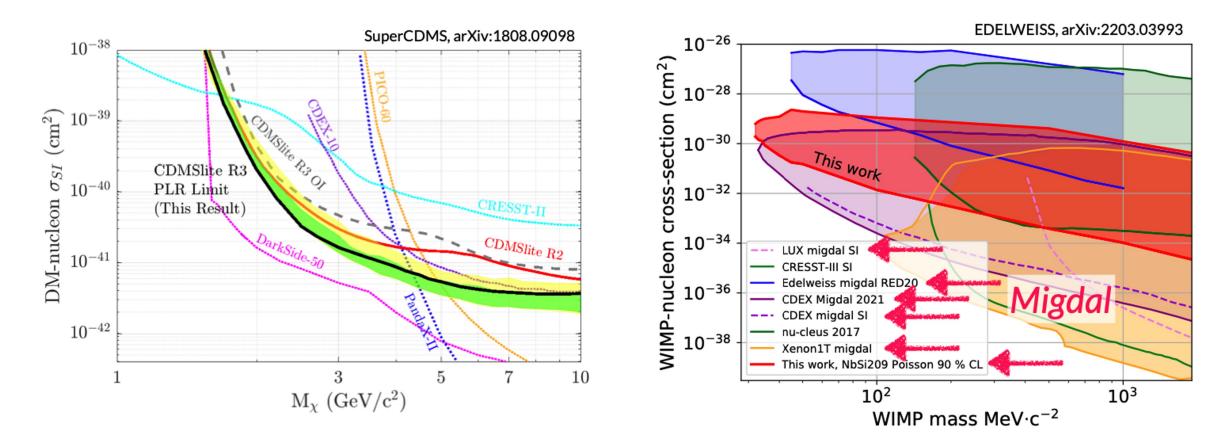
E-mail: ibe@icrr.u-tokyo.ac.jp, m156077@icrr.u-tokyo.ac.jp, yshoji@icrr.u-tokyo.ac.jp, ksuzuki@icrr.u-tokyo.ac.jp

So far ~ 200 citations of Ibe et al. Experiments using Migdal to enhance sensitivity to LDM:

LUX, XENON1T, EDELWEISS,CDEX-1B, SENSEI

Including targets: Ge, Si, Xe and Ar

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Pre-2018 No Migdal limits

Migdal effect in dark matter direct detection experiments, Ibe et al arXiv:1707.07258

Today Dominated by Migdal

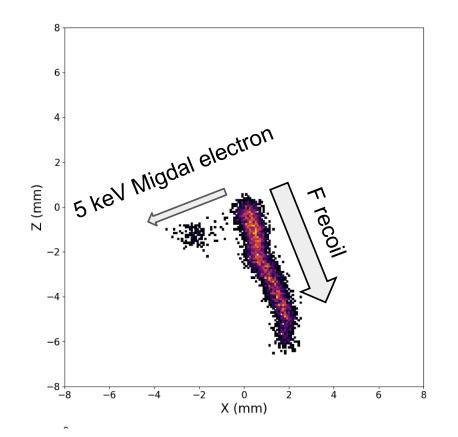
Detecting the Migdal Effect

• The Migdal effect has not been measured in this regime

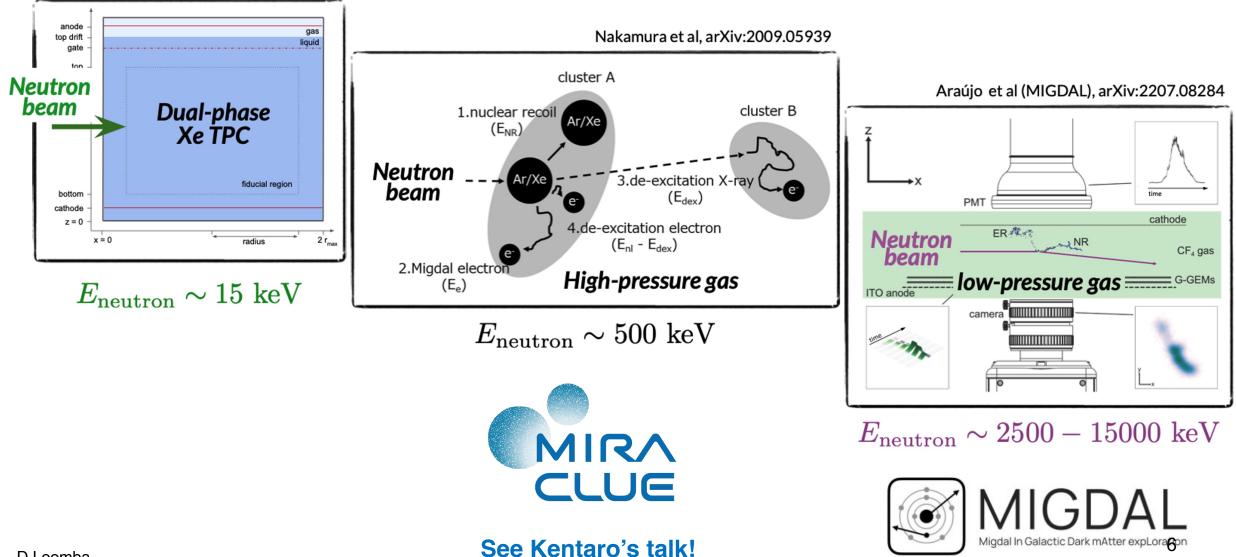
→ needs validation!

- Many efforts/techniques underway (see Kentaro's talk!)
- Here we focus on detecting the electron and nuclear recoils – the Migdal topology

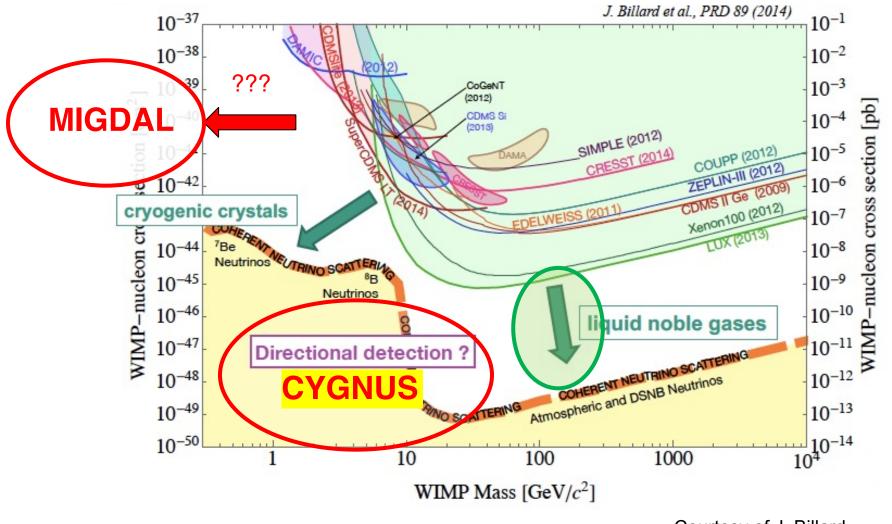
 Challenging!



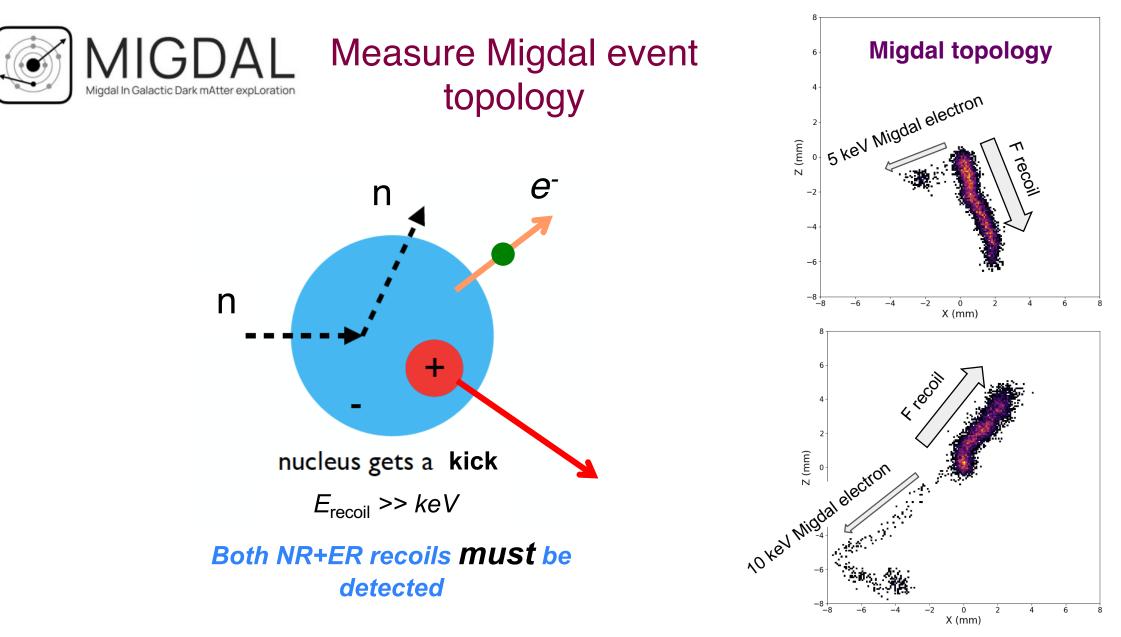
Bell et al, arXiv:2112.08514



Directional Detectors



Courtesy of J. Billard

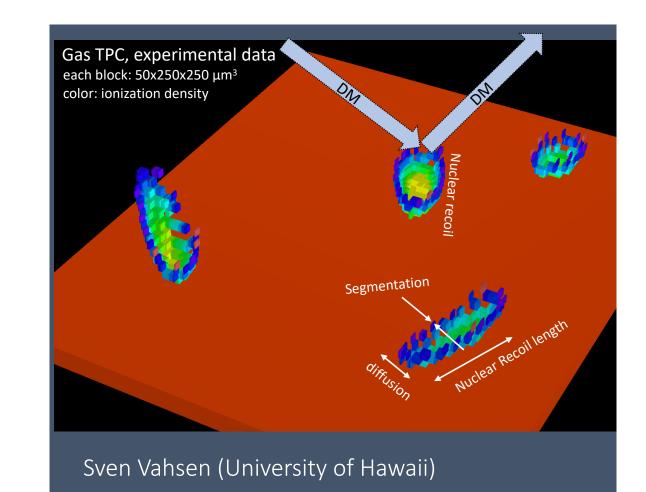


P. Majewski, "OTPC for the observation of the Migdal effect in nuclear scattering", New Horizons in Time Projection Chambers, 2020

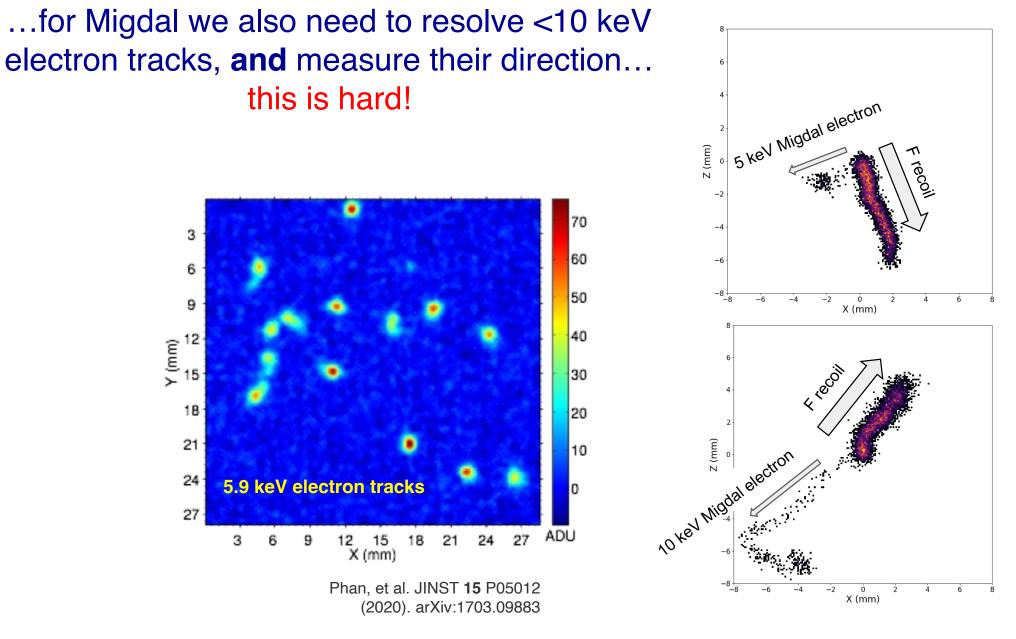
8

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The Directional DM Community has pointed the way **for NRs**



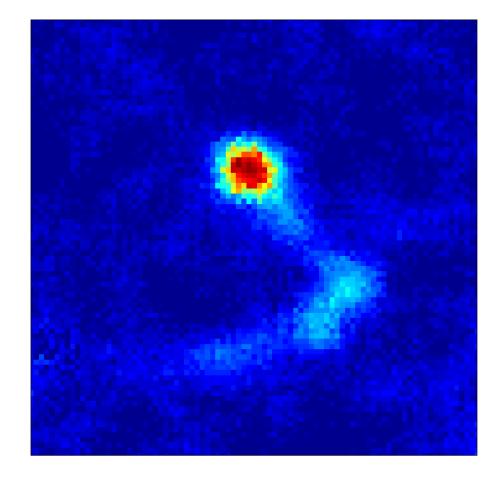
Well resolved NRs:

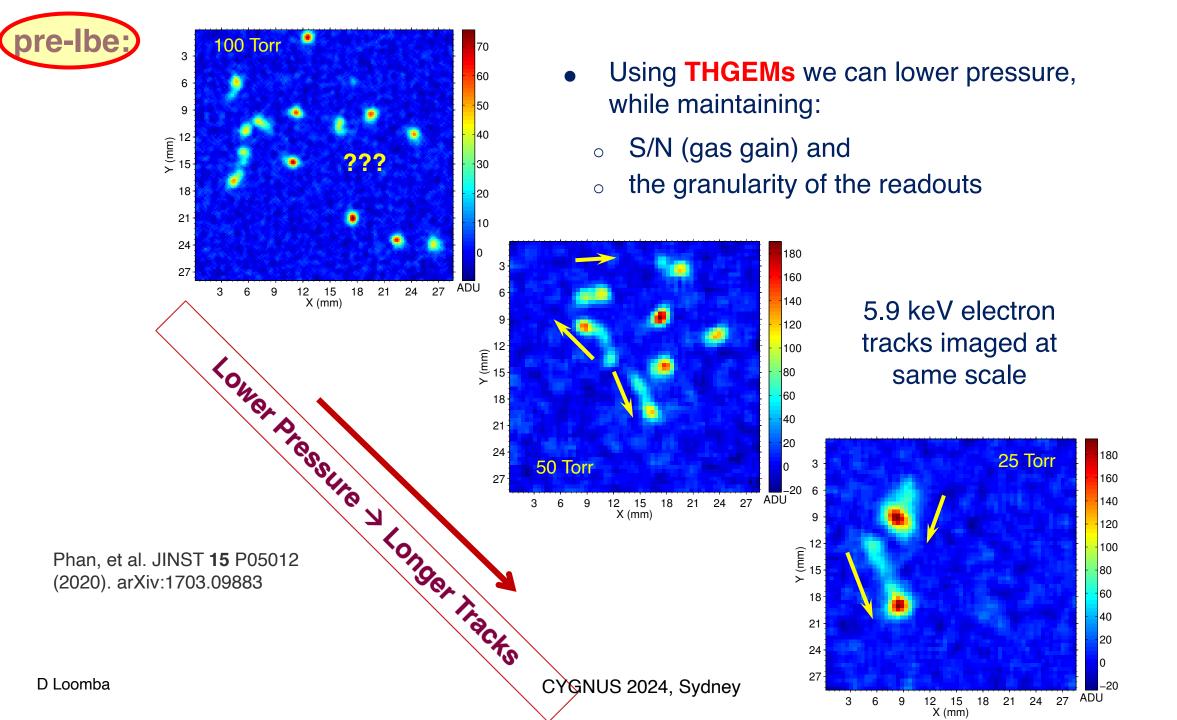


The challenge: electron tracks have **low dE/dx**, large fluctuations:

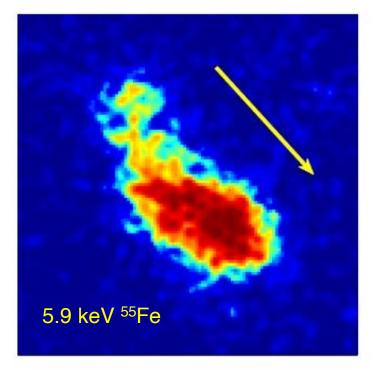
We need:

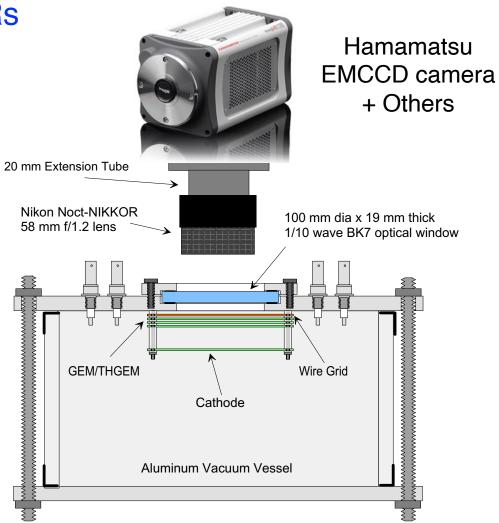
- low pressures
- fine granularity readouts
- high S/N, and
- Large dynamic range!



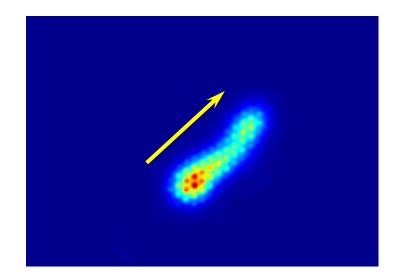








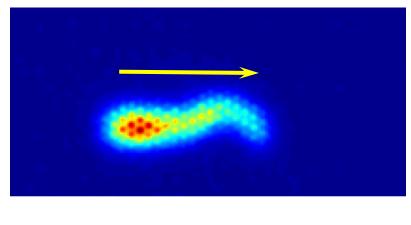
Results:

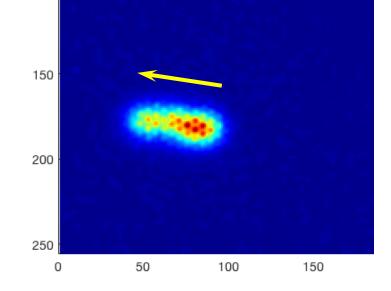


5.9 keV 55Fe

5.9 keV ER tracks

~100-300 keV NR tracks





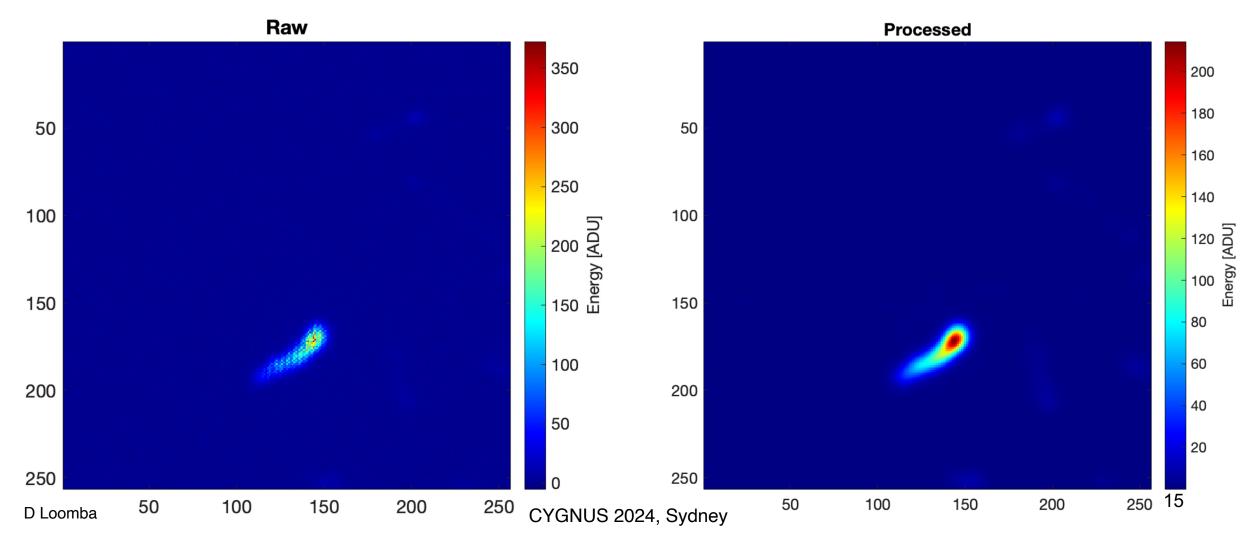
At 30 Torr, these tracks are well resolved with clear directionality (arrows)

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Plus we demonstrated we could cover the huge dynamic range:

• With GEM gain set such that DD NR's are stable:

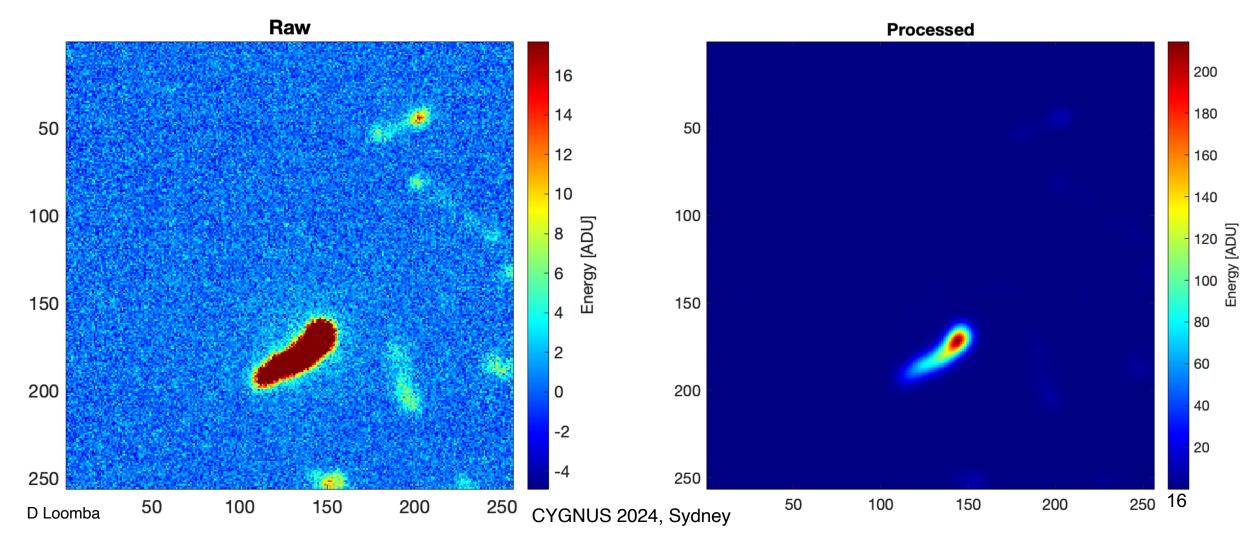
(A. Mills, RD51, Nov 2021)



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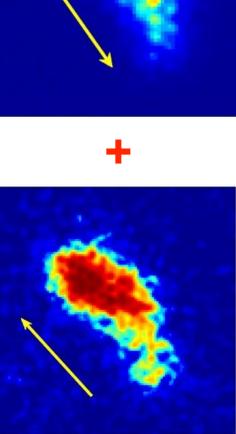
(A. Mills, RD51, Nov 2021)



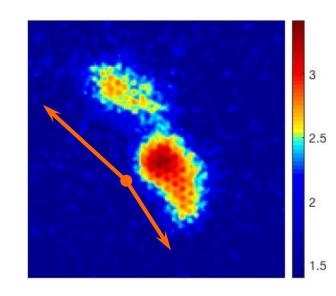
"Migdal Events"

Nuclear Recoil

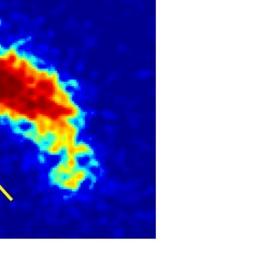
Electron

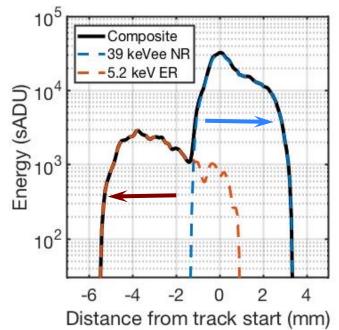


E = 39 keVee



Migdal





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The MIGDAL Collaboration

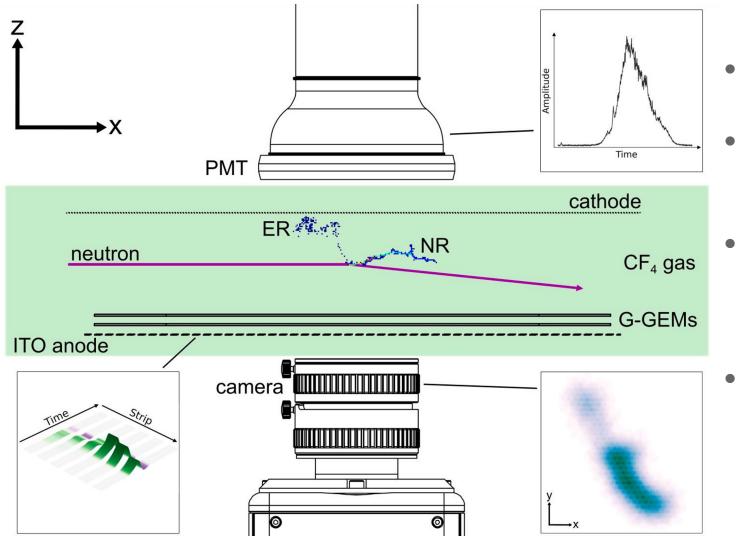


- Over 35 physicists and engineers from 14 institutions across 6 countries
- At Rutherford Appleton Lab at the NILE facility





The MIGDAL experiment



Low-pressure gas: 50 Torr of CF₄

- To extend particle tracks
- Minimize gamma interactions

TPC signal amplification

• 2 x glass-GEMs (Cu or Ni cladded)

Readout

- Optical : Camera + photomultiplier tube
- Charge: GEMs + 120 ITO anode strips

• High-yield neutron generator

- D-D: 2.47 MeV (10⁹n/s)
- D-T: 14.7 MeV (10¹⁰n/s)
- Defined beam, "clear" through TPC

Electron and nuclear recoil tracks

- Migdal: NR+ER tracks, common vertex
- Cover dynamic range of NRs and ERs
- 5 keV electron threshold

5.9 keV X-rays from Fe-55 for calibration at threshold (5.2 keV photoelectron)

H.M. Araújo *et al* 2023 *Astropart. Phys.* **151** 102853 https://doi.org/10.1016/j.astropartphys.2023.102853

Expected Migdal signal/backgrounds per 1 million DD-induced nuclear recoils with E > 100 keV

Component	Topology	D–D neutrons	
		>0.5	5–15 keV
Recoil-induced δ -rays	Delta electron from NR track origin	≈0	0
Particle-Induced X-ray Emission (PIXE)			
X-ray emission	Photoelectron near NR track origin	1.8	0
Auger electrons	Auger electron from NR track origin	19.6	0
Bremsstrahlung processes ^a			
Quasi-Free Electron Br. (QFEB)	Photoelectron near NR track origin	112	≈ 0
Secondary Electron Br. (SEB)	Photoelectron near NR track origin	115	≈0
Atomic Br. (AB)	Photoelectron near NR track origin	70	≈0
Nuclear Br. (NB)	Photoelectron near NR track origin	≈ 0	≈0
Neutron inelastic γ -rays	Compton electron near NR track origin	1.6	0.47
Random track coincidences			
External γ - and X-rays	Photo-/Compton electron near NR track	≈ 0	≈0
Trace radioisotopes (gas)	Electron from decay near NR track origin	0.2	0.01
Neutron activation (gas)	Electron from decay near NR track origin	0	0
Muon-induced δ -rays	Delta electron near NR track origin	≈ 0	≈0
Secondary nuclear recoil fork	NR track fork near track origin	-	≈ 1
Total background	Sum of the above components		1.5
Migdal signal	Migdal electron from NR track origin		32.6

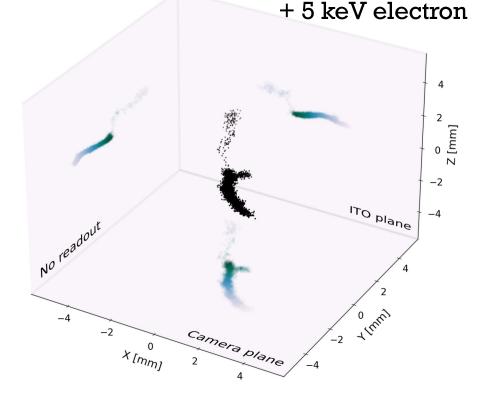
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End-to-end Simulations:

- DEGRAD (electron track)
- TRIM (NR cascade and electronic dE/dx)
- Magboltz (drift properties)
- Garfield++ (GEMs)
- Gmsh/Elmer & ANSYS (ITO and E-field)

Migdal event

150 keV F recoil



Anode strip readout

Camera readout

25 75 150 300

2 mm

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Diffusion + GEMs + noise

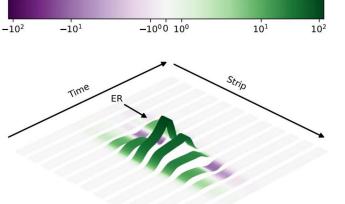
Intensity [ADU]

600

1200

2400

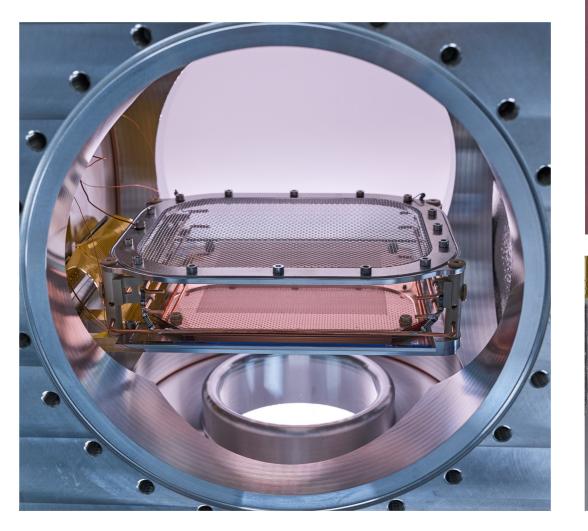
Induction/collection (electronics deconvolved)

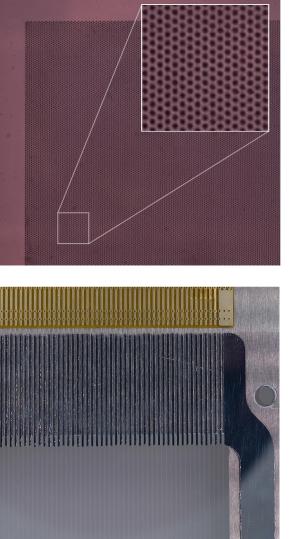


Current [fC/ns]

Image analysis Deconvolution + RidgeFinder 1.2 Intensity [ADU \times 10⁴] 1.0 0.8 0.5 0.2 0.0 1.0 2.0 3.0 4.0 5.0 6.0 Distance along track [mm] 21

The MIGDAL optical-TPC





Two glass GEMs one Cuand one Ni-cladded :

- thickness: 550 µm
- OD /pitch: 170/280 μm
- active area: 10x10 cm²
- total gain ~105

ITO strips wire bonded to readout:

- 120 strips
- width/pitch: 0.65/0.83 mm

Two field shaping copper wires

- TPC inside aluminium cube vacuum vessel
- Drift gap: 3 cm between woven mesh cathode and cascade of two glass-GEMs (E_{DRIFT}=200 V/mm for min. electron diffusion)
- Transfer and signal induction gaps : 2 mm
- Low outgassing materials; vacuum before fill 2x10⁻⁶ mbar; vacuum unchanged several days after fill

Charge and Light readout



ITO anode strips

Post-GEM ionisation

Readout of (x,z) plane

Pitch: 833 µm Digitised at 2 ns/sample (Drift velocity: 130 µm/ns)

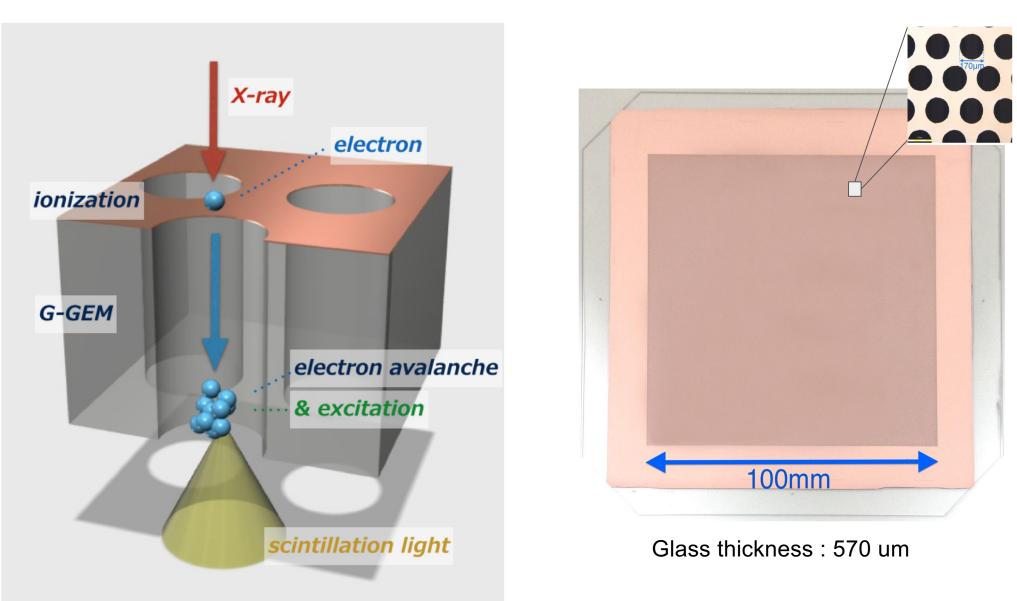
qCMOS camera (Hamamatsu ORCA - QUEST)

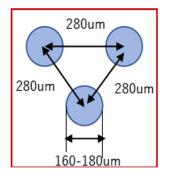
GEM scintillation through glass viewport behind ITO anode **Readout of (x,y) plane**

Exposure: 8.33 ms/frame (~120 Hz) (continuous) Pixel scale: 39 µm (2 × 2 binning) Lens: EHD-25085-C; 25mm f/0.85

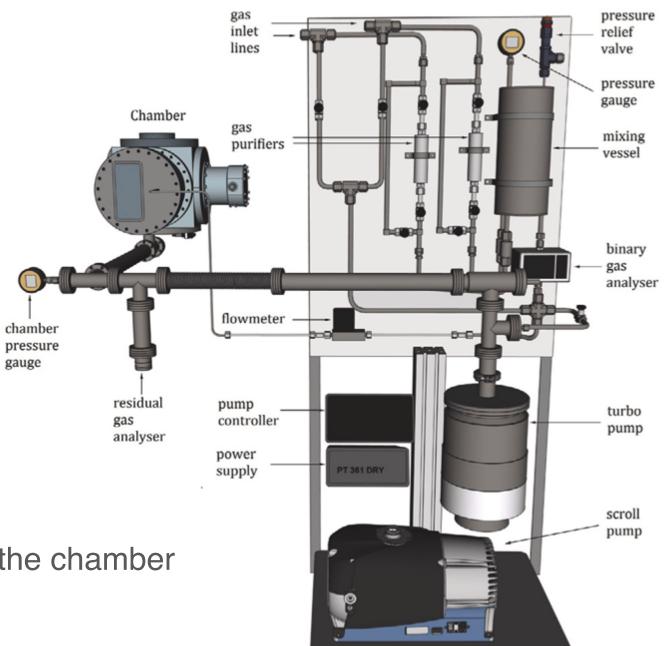
VUV PMT (Hamamatsu R11410)

Detects primary (S1) and secondary (S2 from GEM) scintillation Absolute depth (z) coordinate Digitised at 2 ns/sample → Event Trigger **Glass-GEMs**



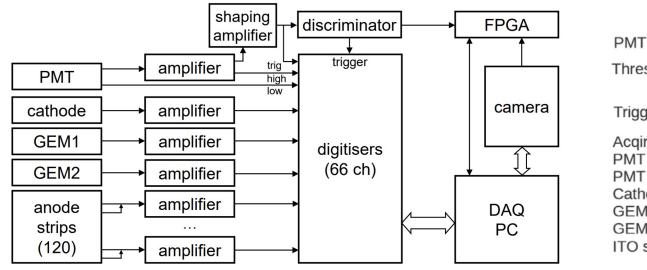


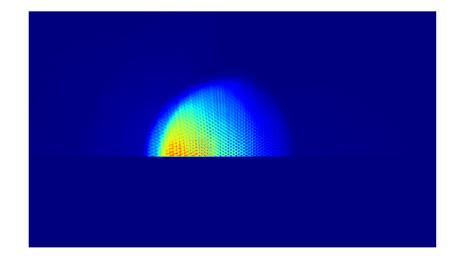
Gas System

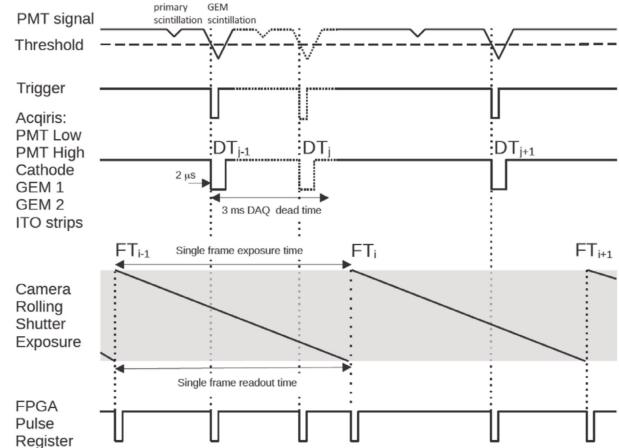


Initial vacuum in the chamber ~1E-06 mbar

DAQ







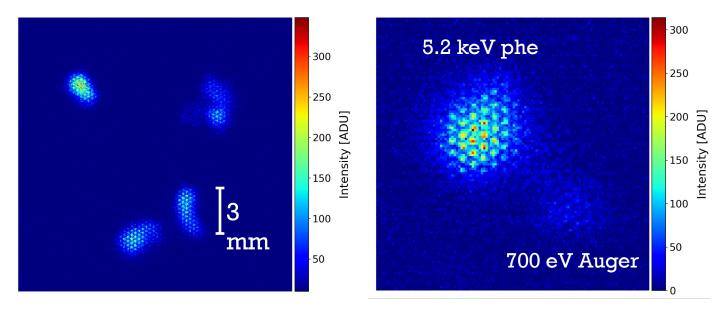
Synchronisation with LED pulse Image cut due to a rolling shutter

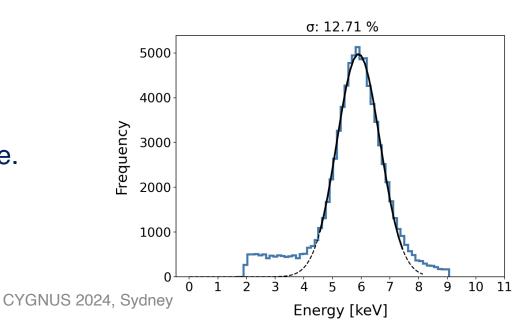
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Pre-run Calibrations

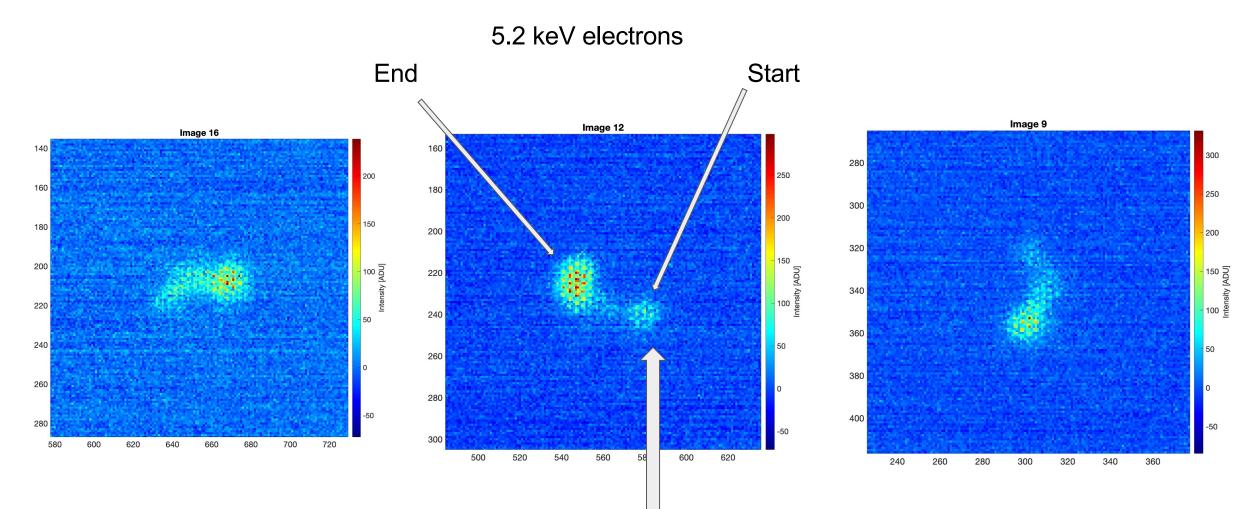
Calibration with ⁵⁵Fe – Pure CF₄

- Tests were performed with ⁵⁵Fe (5.9 keV x-ray).
- The gain was maximum achievable with ⁵⁵Fe.
- Achievable energy resolution is excellent $(\sigma/\mu \sim 12.7\%)$.
- Head & tail is clearly resolved.
- 700 eV Auger electron from fluorine is visible.





Example images from 5.9 keV X-ray conversions using Fe-55 source

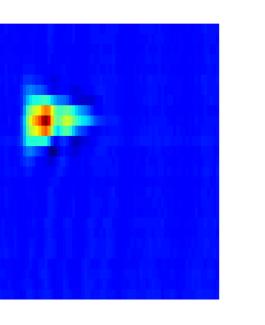


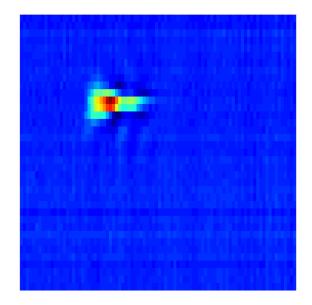
Extra charge at the start from Fluorine 650 eV Auger Electron

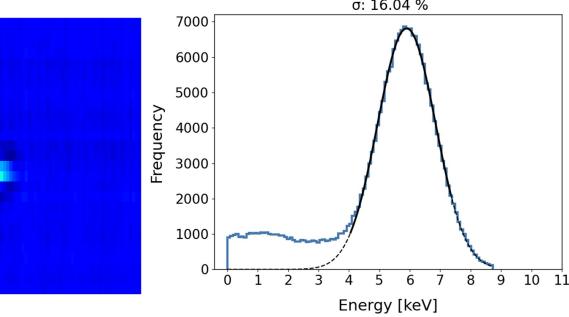
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ITO (Pure CF_4)

- Very good signal to noise.
- Spatial resolution is not as good as camera (~0.83 mm pitch).
- Good energy resolution even with no flat fielding correction.
- Analysis of ITO images is ongoing, methods are still being refined.



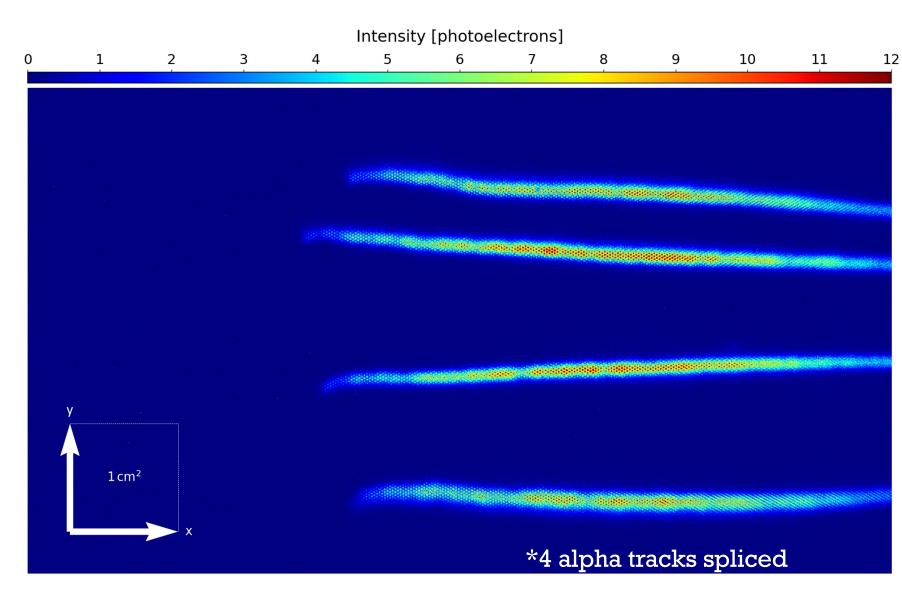




σ: 16.04 %

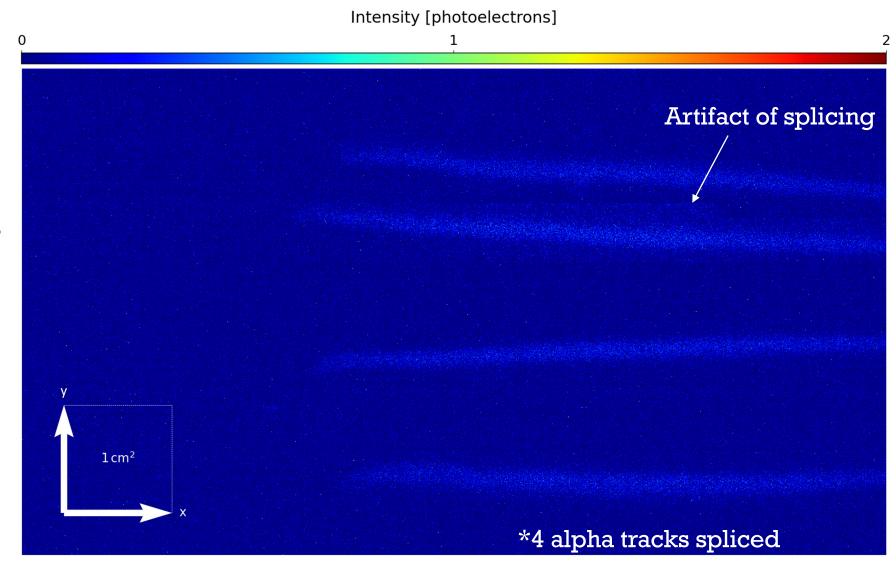
Alphas:

- Testing operational stability with 37 Bq ²⁵²Cf source in 50 Torr CF₄.
- The new camera produces very good-looking images!
- The optical distortion and lens field curvature are visible towards the edges of the image.



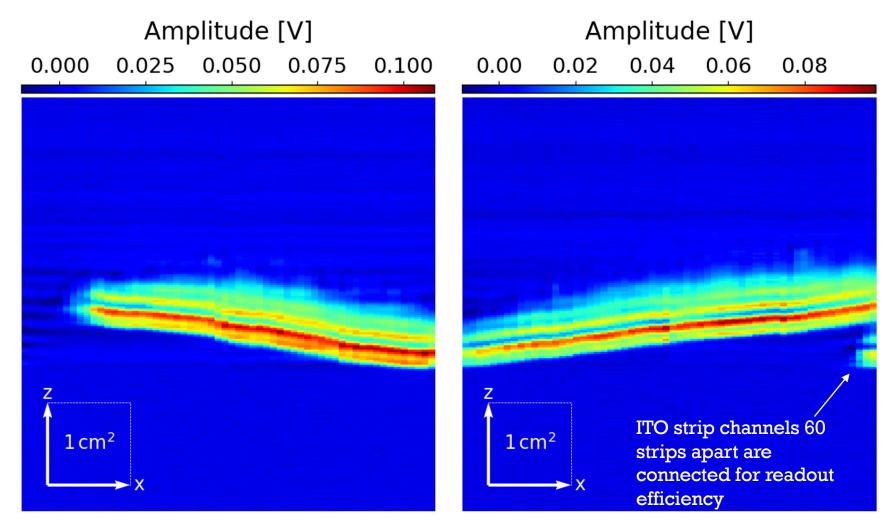
Afterglow (Ghosts!) with ORCA Quest

- In frames following alpha tracks (any bright track), we see an afterglow of ~1 photoelectron in many pixels.
- This does not seem to vary with exposure time.
- We are in contact with Hamamatsu.



Alphas in the ITO strips

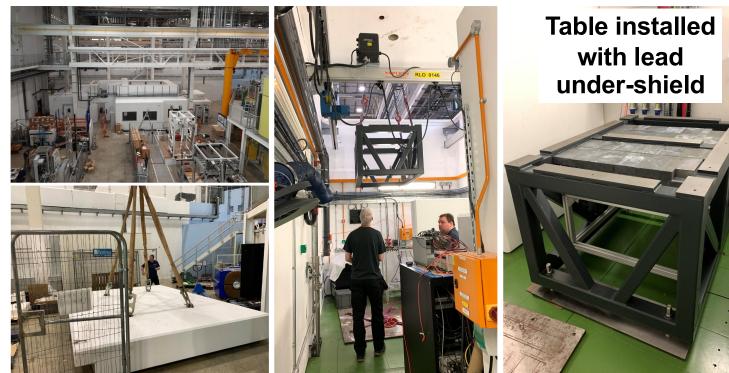
- The signals from alpha tracks create a 'ripple' in the ITO strips. (under investigation)
- ITO strips 1 & 61, 2 & 62 etc. are connected. This is ok for nuclear recoils as no tracks will be longer than 5 cm.



First Science/Commissioning Run at NILE

NILE facility at RAL, UK

- NILE facility is at TS2, ISIS
- Chamber packed and moved from lab to NILE mid-May.





Chamber driven over to NILE





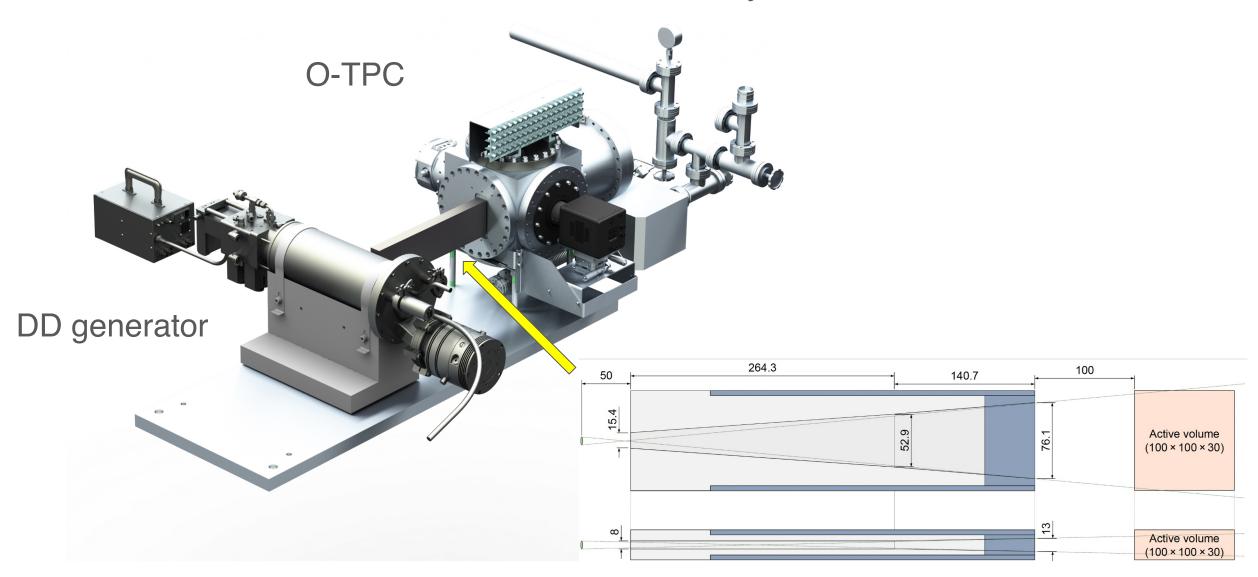


Assembling at NILE



Assembling at NILE

To Gas System



30 cm long collimator

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Experiment installation in the NILE bunker

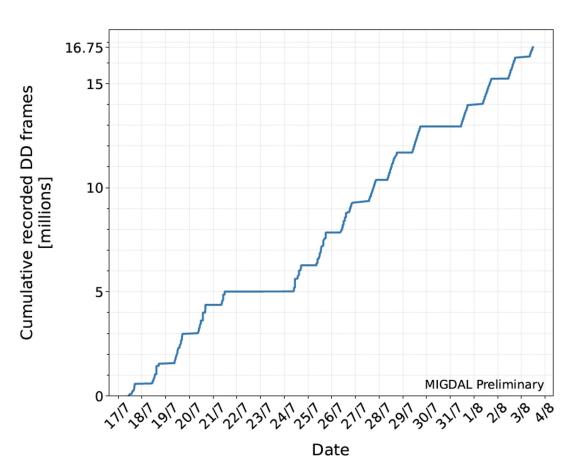


MIGDAL experiment fully assembled at NILE

- Lead shield : 10 cm
- Borated HDPE shield : 20 cm
- Collimator HDPE+Clead^S: 230⁻ chinitiong

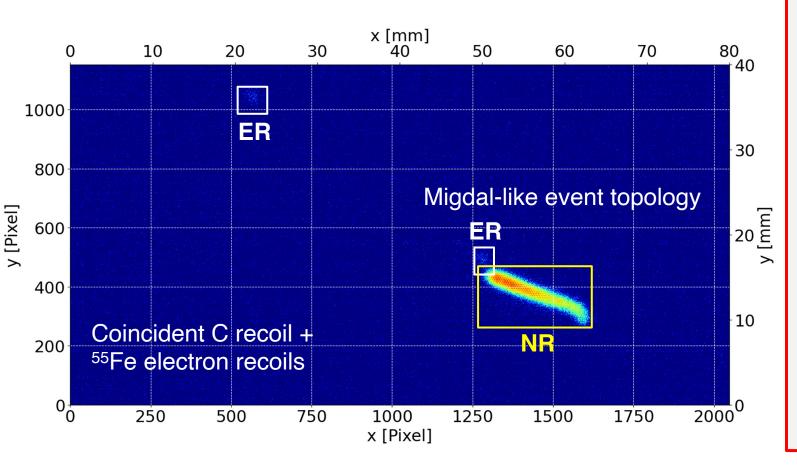
First Science Run – Summary

- The First Science run took place from the 17th of July to the 3rd of August.
- Data taken using D-D neutron generator recorded continuously during 10 hour long shifts. Significant fraction of empty frames due to low DD rate.
- Frames taken with 20 ms exposure time. Longer than planned (8 ms) due to problems with camera's Linux firmware.
- Data taking interspersed with regular calibration runs (⁵⁵Fe) to monitor the gain of the detector.
- GEM dV increased by a small amount each day to keep constant gain.
- Total gain in GEMs tuned to a threshold required to fully resolved ⁵⁵Fe peak.
- Average spark rate ~ 7/min due to need to cover high dynamic range.
- Half the data is blinded.



Real-time analysis with YOLO (You only look once)

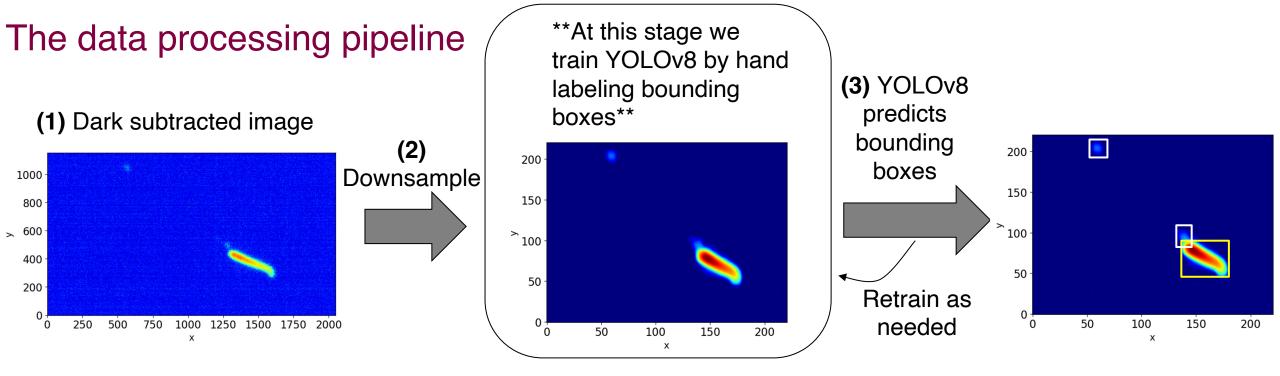
YOLOv8 is a state of the art *object detection* algorithm that simultaneously locates (draws a bounding box) and identifies objects of interest in an image

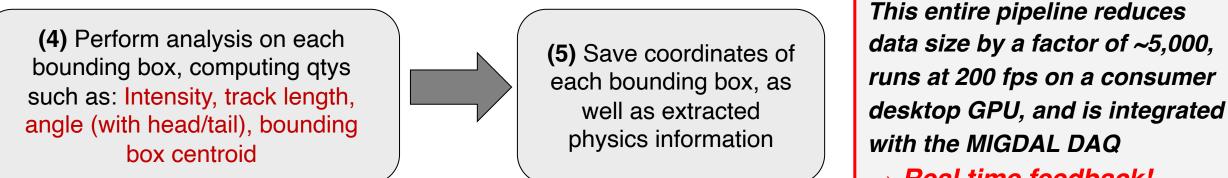


We train YOLOv8 on measured data to identify ERs, NRs, protons, alphas, sparks, camera afterglow, rolling shutter, etc.

Benefits:

- 1. Enables real time ⁵⁵Fe calibrations and ER/NR event rate counting
- 2. Can identify multiple particle species within a continuous cluster
- 3. Not trained specifically to find Migdal candidates →robust and doesn't need to be trained on simulation!
- 4. Single-shot identification and analysis of tracks

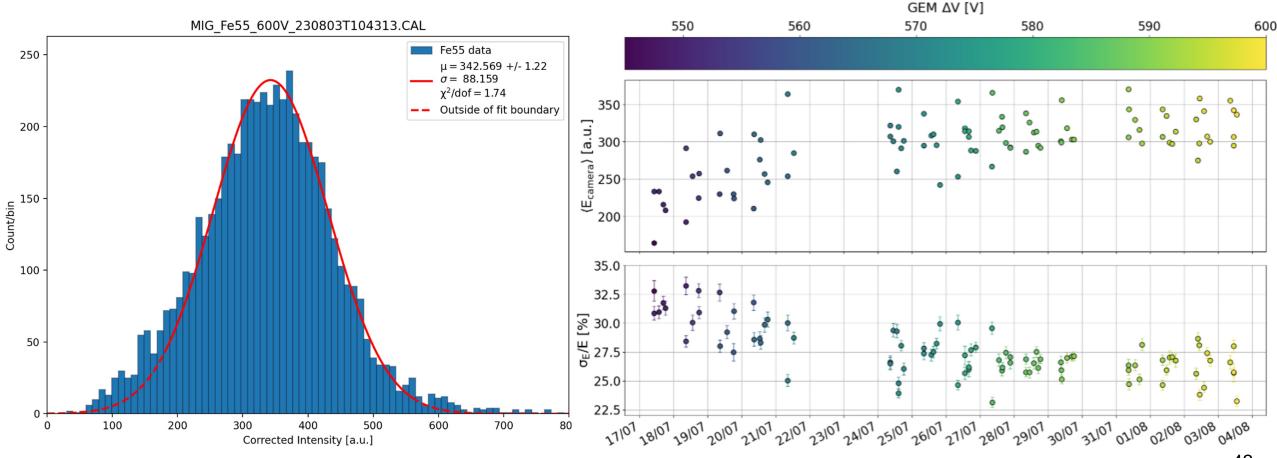




→ Real time feedback!

Real-time gain calibration with YOLO

⁵⁵Fe spectra automatically processed at the end of calibration runs Summary of gain and gain resolution over the course of our science run from July 17th - Aug. 4th, 2023

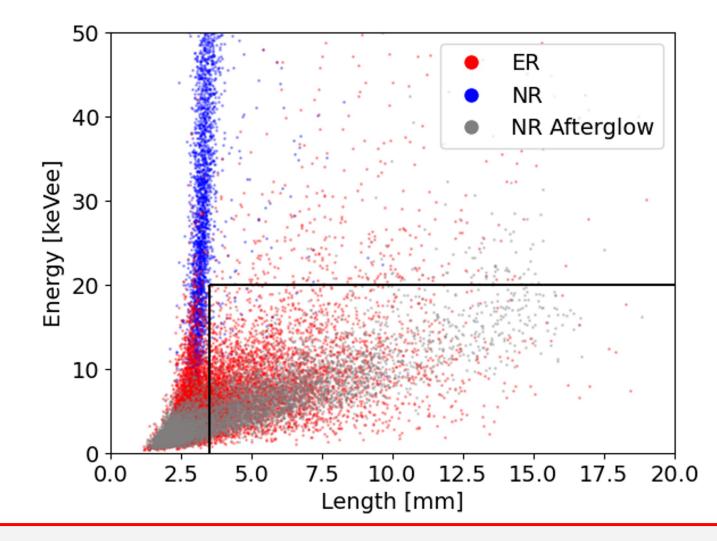


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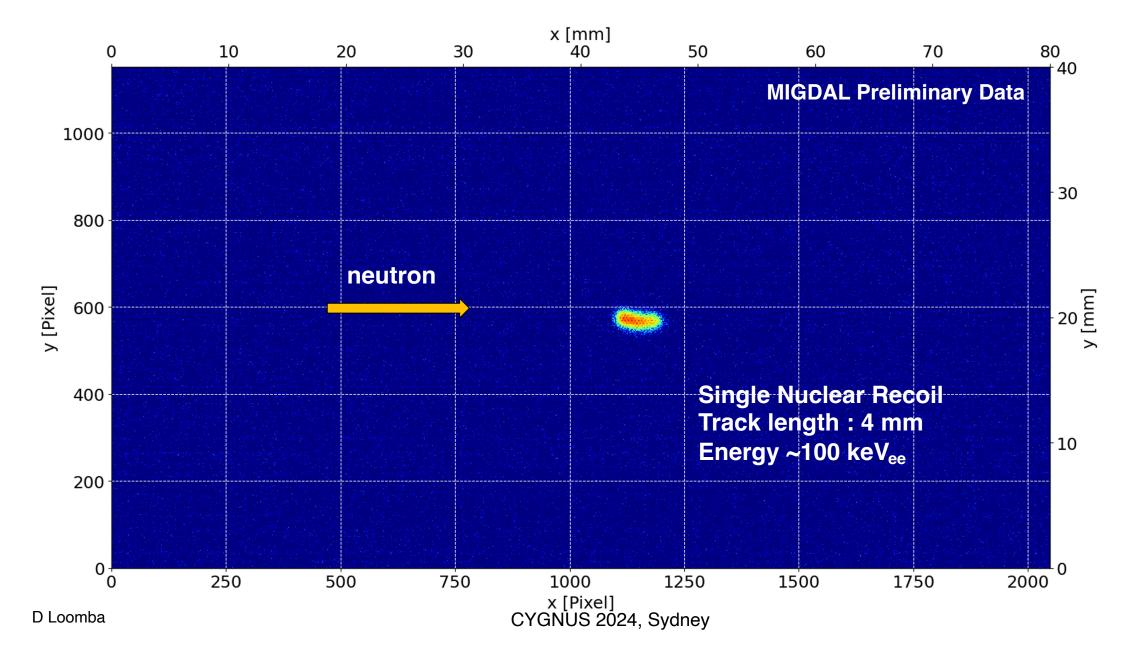
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Real-time E vs R2 from YOLO



Particle species rates and dE/dx distributions automatically generated from YOLO pipeline → Detector performance monitoring

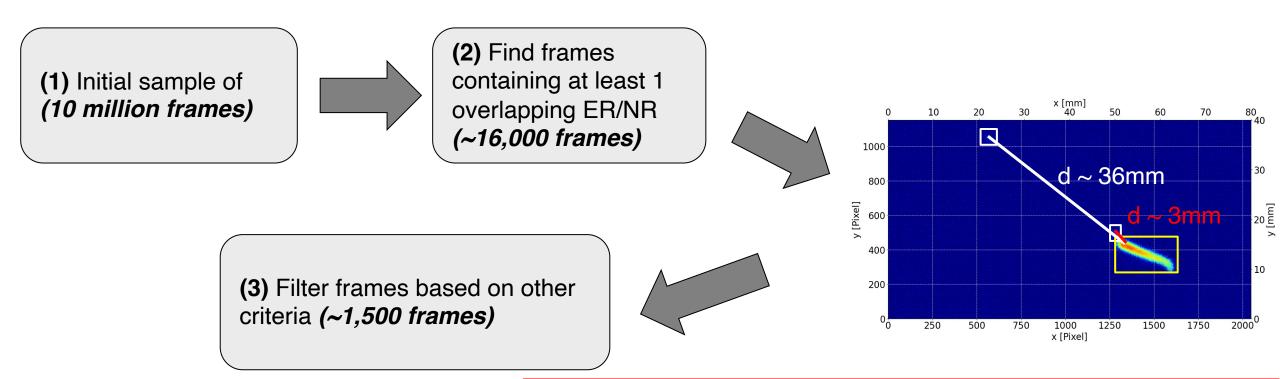
Example events - single Nuclear Recoil



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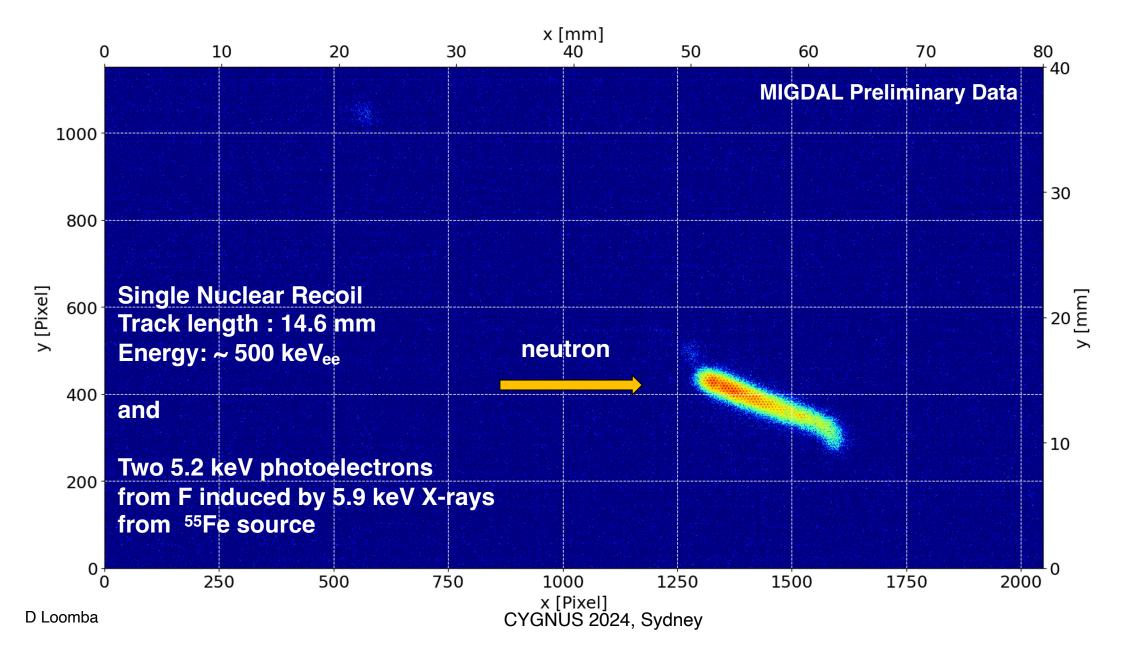
YOLO applied to Migdal search

- Initial science run recorded from July 17th 2023 August 4th 2023
- Collected an unblinded dataset consisting of 10 million 2,048 x 1,152 images

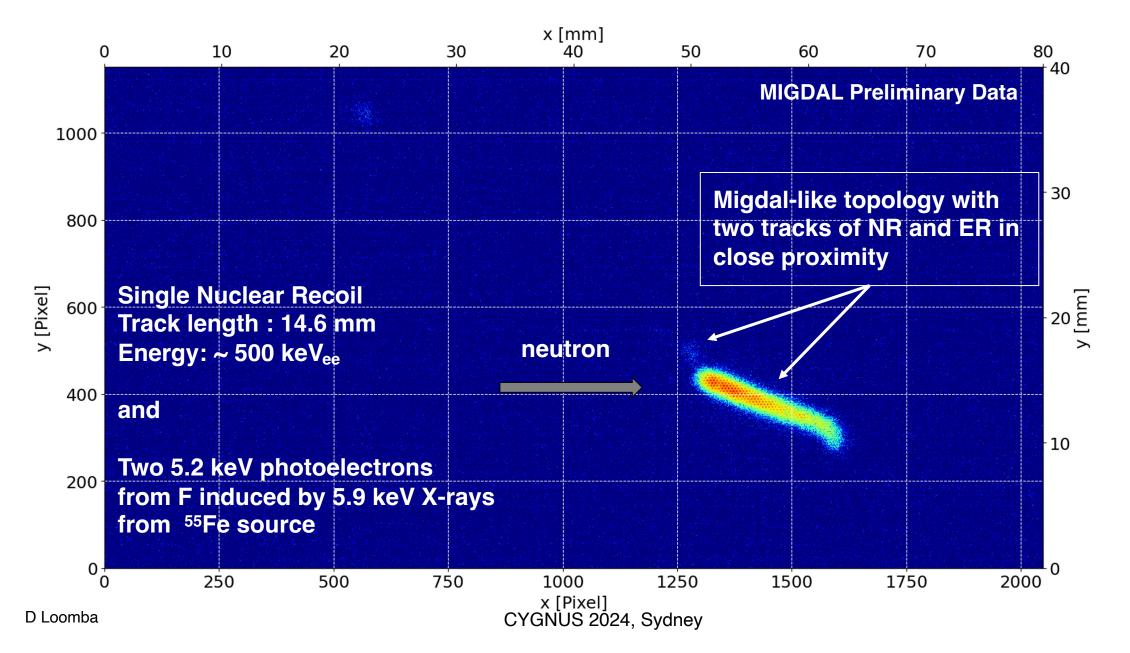


YOLOv8 bounding box analysis takes us from 10's of millions of frames to a few 1000 frames \rightarrow No longer a rare event search! We can use other techniques to optimize signal purity

Example event - Migdal-like

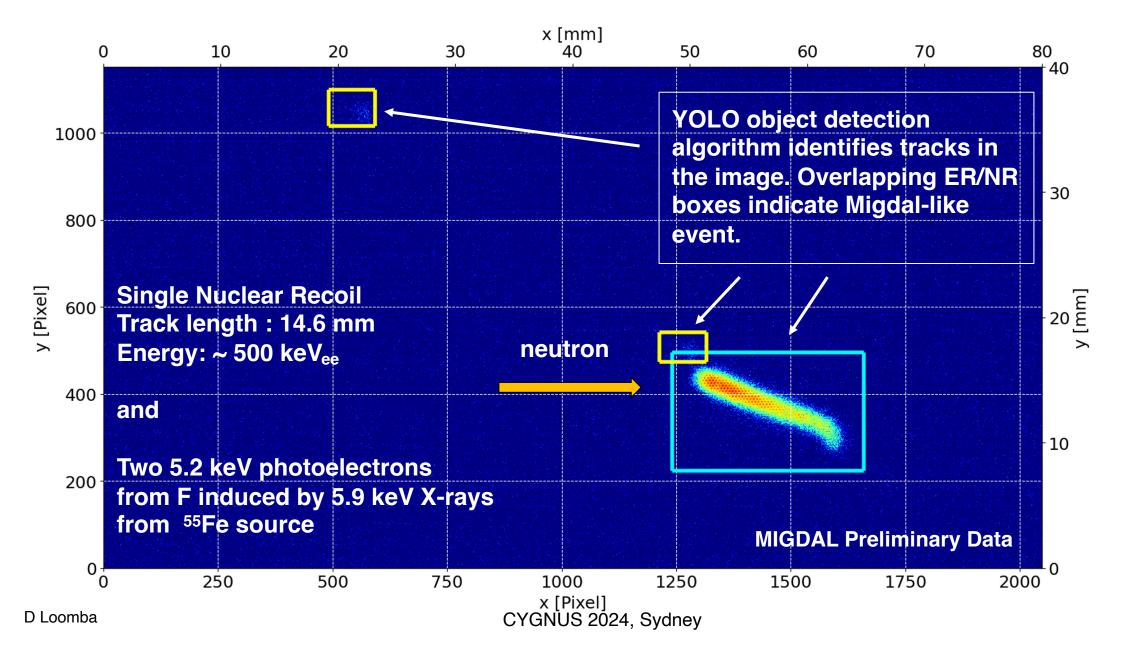


Example event - Migdal-like



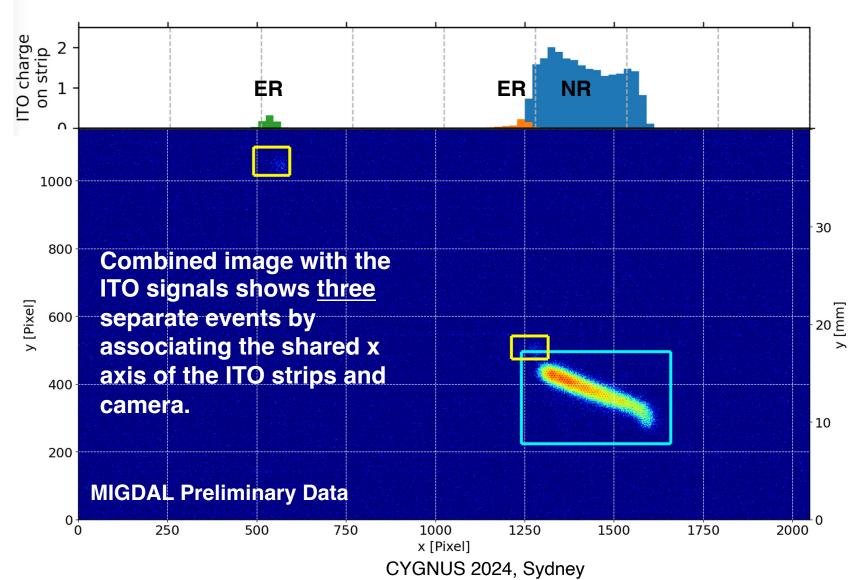
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Example event - Migdal-like



Migdal-like event

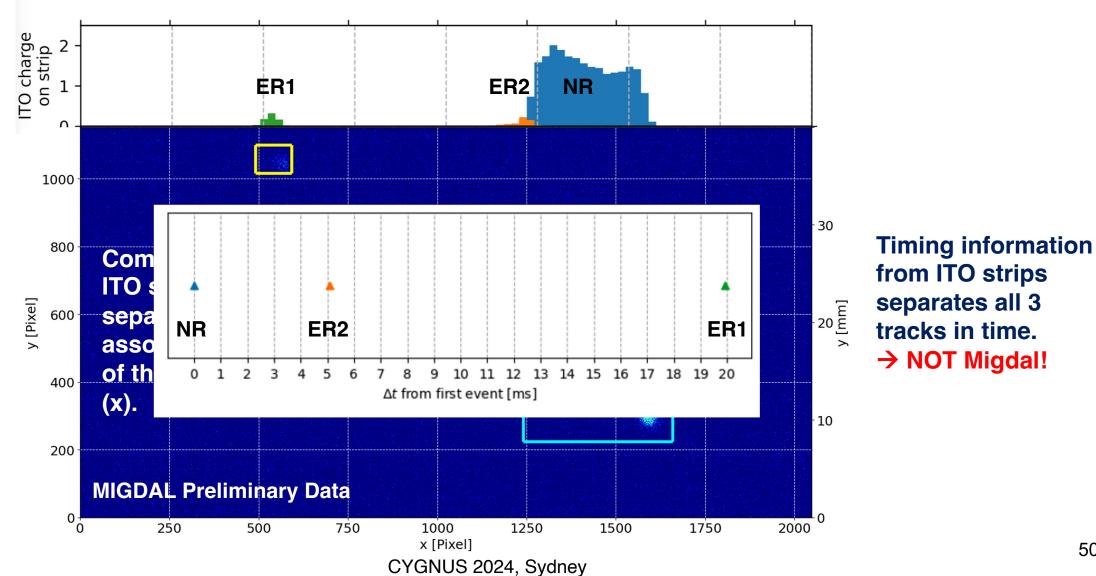
Synchronizing with ITO signals:



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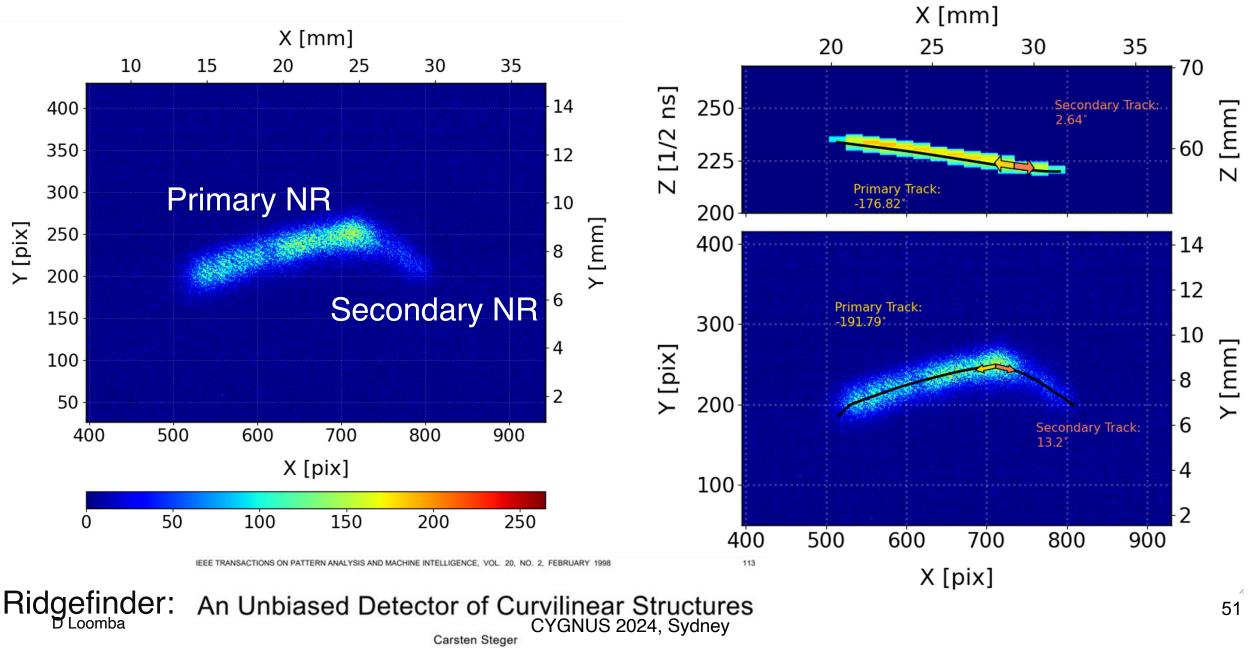
Migdal-like event



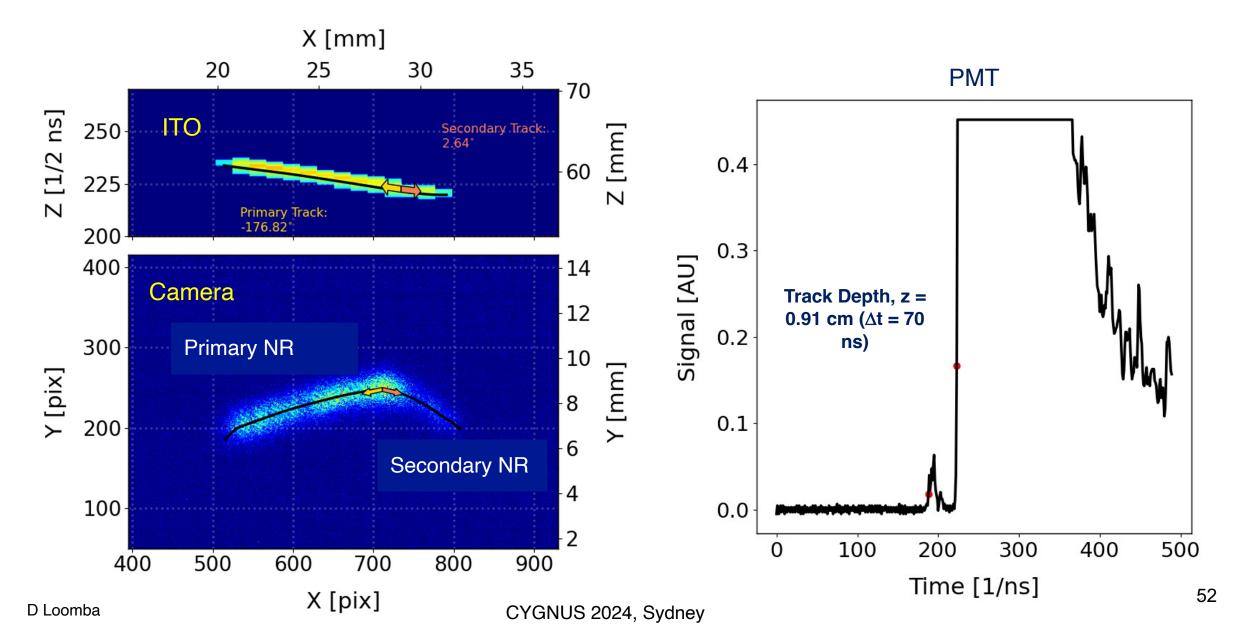


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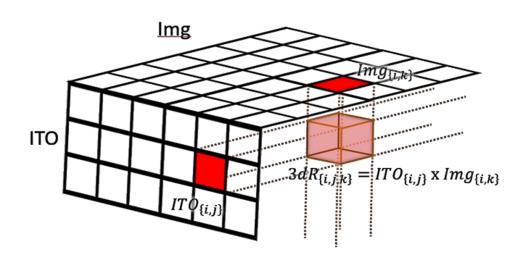
3D track reconstruction: Camera + ITO + PMT



3D track reconstruction: Camera + ITO + PMT



3D track reconstruction: 3D voxels



3D track reconstruction of low-energy electrons in the MIGDAL low pressure optical time projection chamber

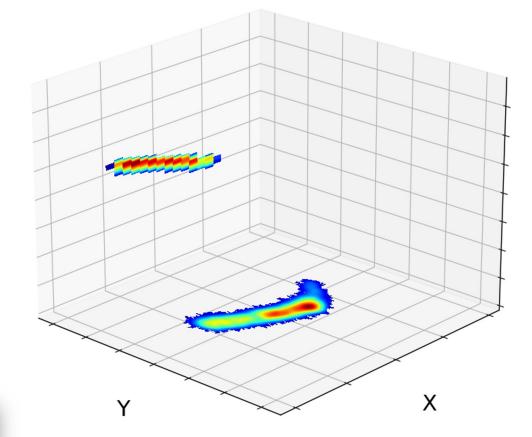
E. Tilly¹ and M. Handley^{2,3} on behalf of the MIGDAL collaboration Published 17 July 2023 • © 2023 IOP Publishing Ltd and Sissa Medialab

Journal of Instrumentation, Volume 18, July 2023

Citation E. Tilly et al 2023 JINST 18 C07013

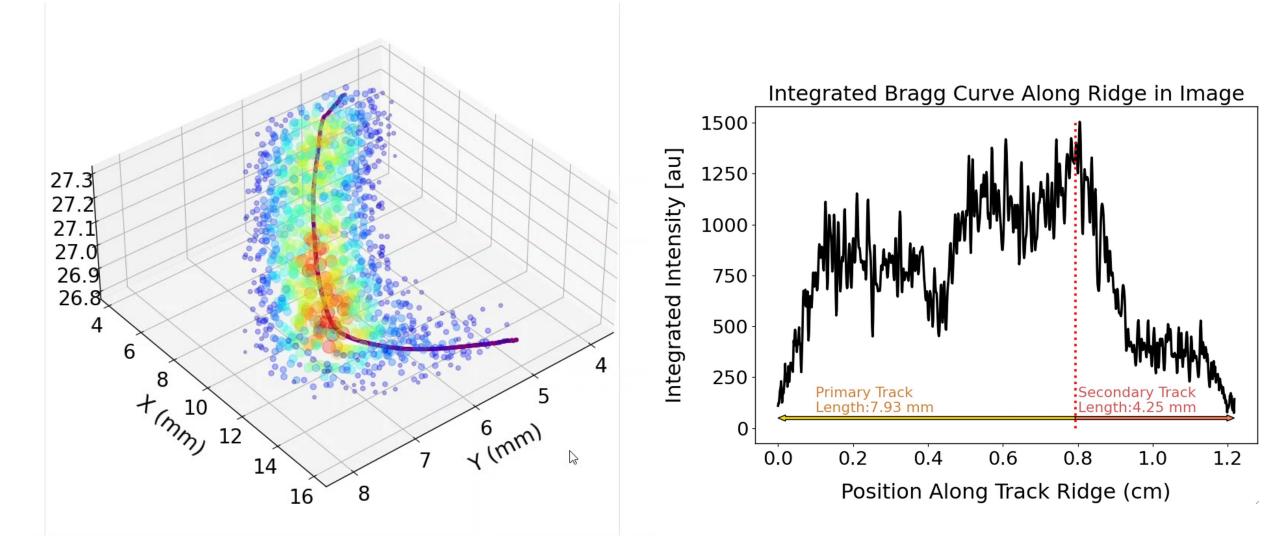
DOI 10.1088/1748-0221/18/07/C07013

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Ζ

3D track reconstruction



Preparations for the Second Science RUN

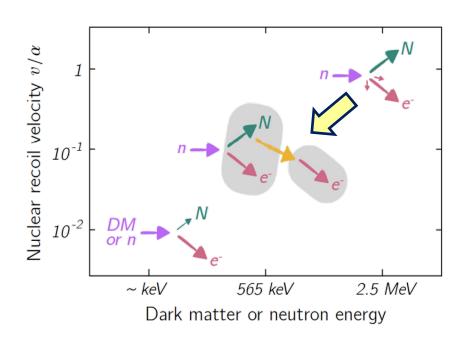
Improvements for the Second Science Run (SSR):

- Fixed communication between DAQ PC and camera use CoaXPress instead of USB decreasing exposure time from 20 ms to 8.3 ms - less chance for coincidence events.
- 2. Fixed missing bonded connection to central ITO strips.
- 3. Use FPGA to record timing information of every trigger also within 3 ms of DAQ dead time.
- 4. New GEMs w/o metallisation around mounting holes
- 5. Use of mask to provide better overlap of ITO and camera field-of-view
- 6. At the later stage of the SSR a new collimator increasing NR rate x3.

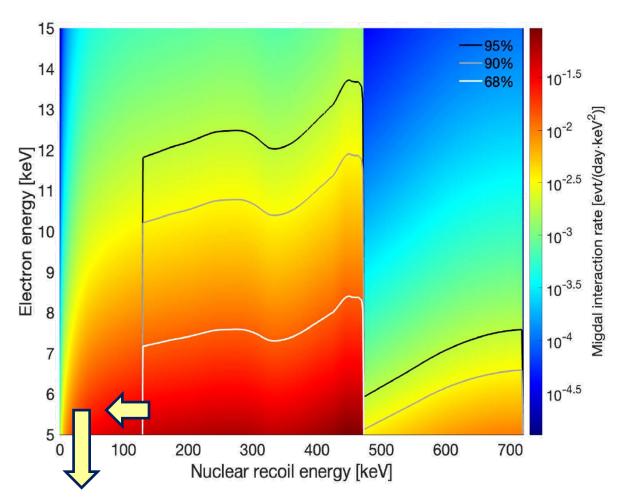
R&D for a MIGDAL Phase II

MIGDAL Phase II - Motivation

- Probe lower energies
- Attain higher rates

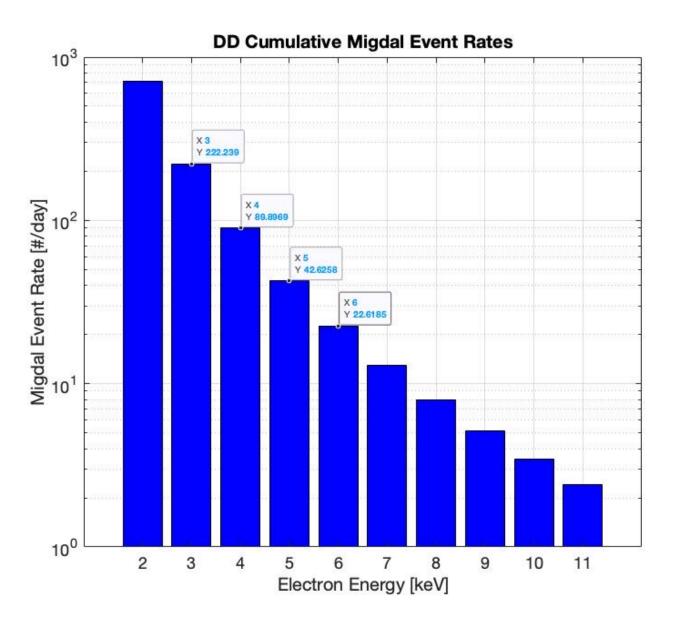


P. Cox *et al* 2023 *Phys. Rev. D* **107**, 035032 https://doi.org/10.1103/PhysRevD.107.035032



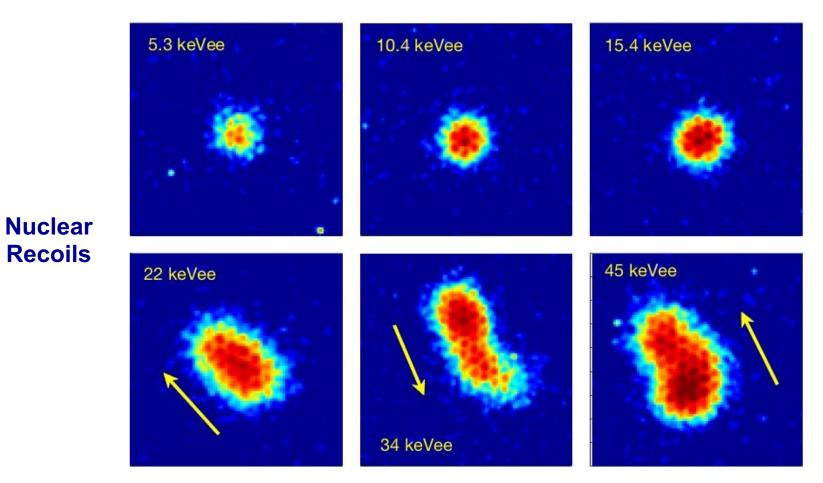
H.M. Araújo *et al* 2023 *Astropart. Phys.* **151** 102853 https://doi.org/10.1016/j.astropartphys.2023.102853

- Probing lower electron energies
 → higher rates
- Will require better spatial resolution → NID
- → NI-OTPC results motivate this R&D

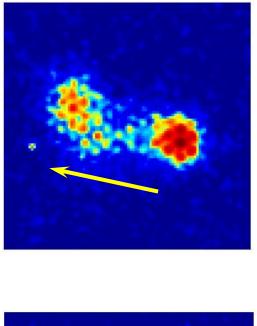


Pure CF4

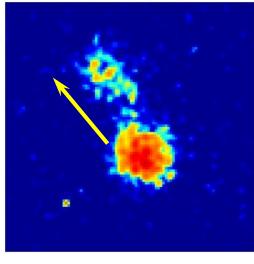
- E_{discrimination} ~ 4 keVee
- E_{directionality} ~ 20 keVee NR* *Disappointing,
- $E_{directionality} \sim a \text{ few keVee ER} \text{ due to } \sigma \sim 0.7 \text{ mm}$

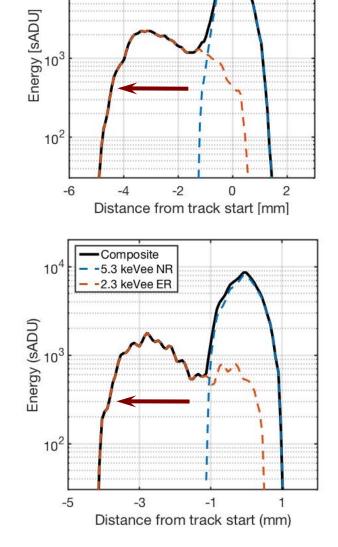


Low energy "Migdal Events"??



- 3D and lower diffusion will help
- Backgrounds?



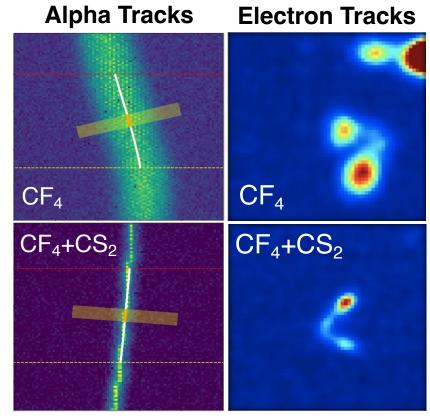


Composite
10.2 keVee NR
3.9 keV ER

10

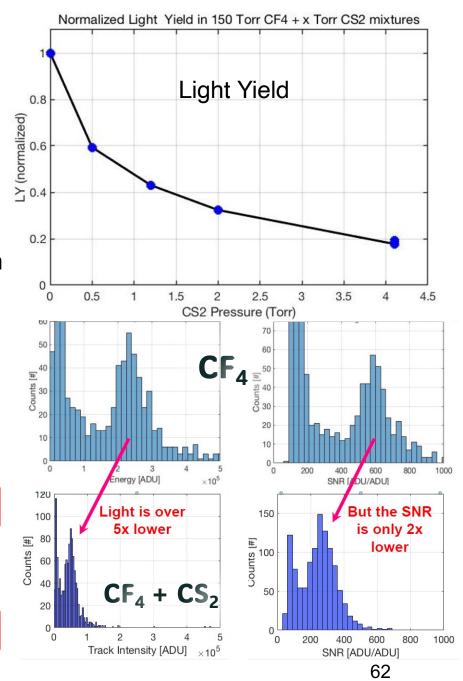
NI-OTPC Results

 Results show promise for NID but will be challenging in O-TPC



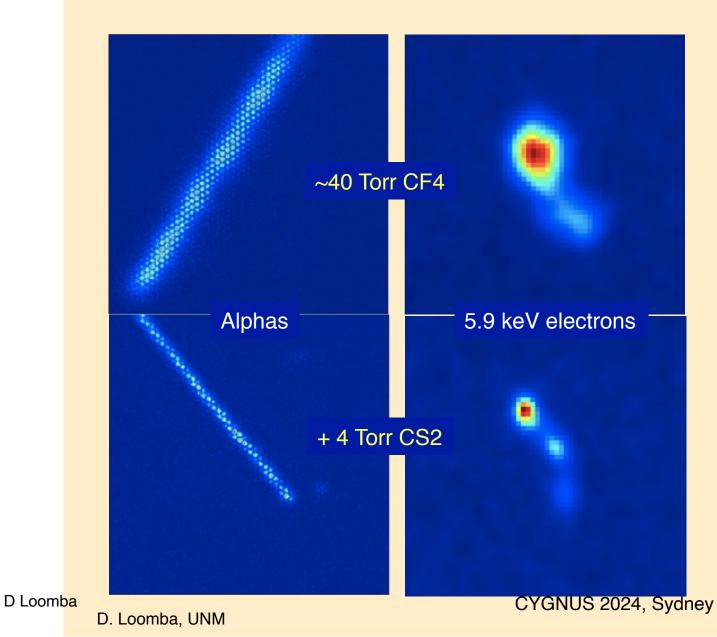
Electron Tracks Measured Transverse Diffusion

150 Torr CF₄ + X torr CS₂			
CS, (Torr)	σ(μm)		
0	~400		
2.9	133.53		
4.2	126.10		
5.4	125.09		
45 Torr CF ₄ + X torr CS ₂			
0	~550		
4	~150-200		



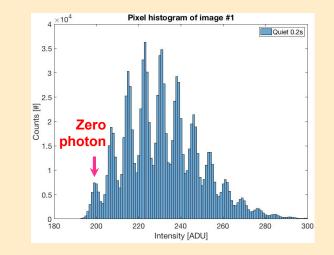
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Negative-ion OTPC



Hamamatsu ORCA-Quest

• Photon Resolving Power:

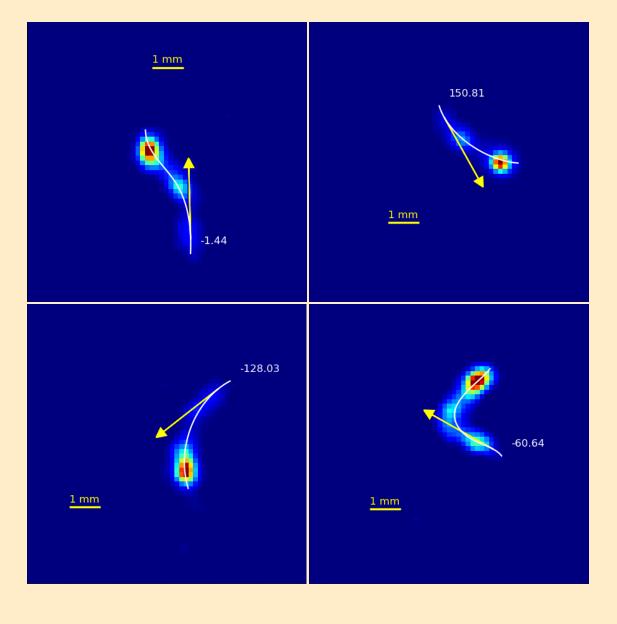


Radiment Glass-GEMs

• 270 micron pitch

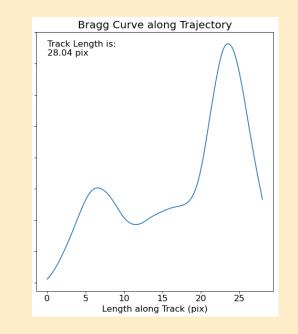
~45 Torr CF ₄ + x Torr CS ₂				
	CS ₂ (Torr)	σ(μm)		
	0	~500		
	4	~150-200		

<mark>6</mark>3



Low diffusion, high spatial resolution enables detailed reconstruction of particle's trajectory:

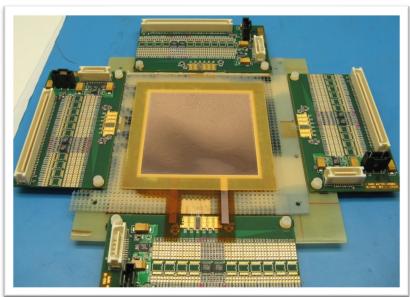
- Head/tail of track
- Initial direction
- Range
- **dE/dx** (Bragg curve):

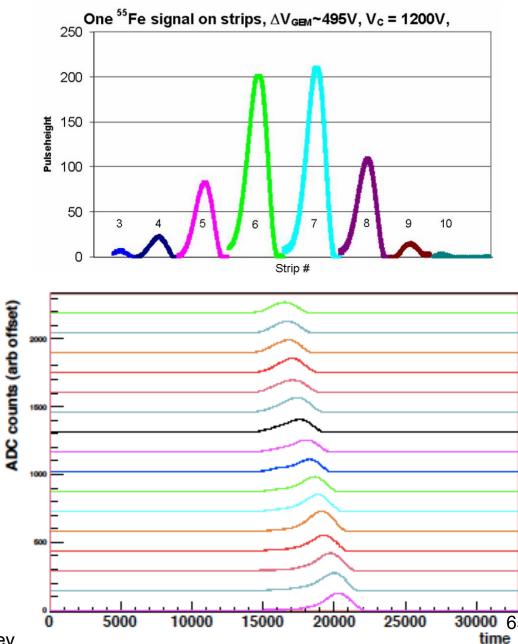


Charge readout(!) NI-TPC

Advantages of a NI *charge* readout TPC:

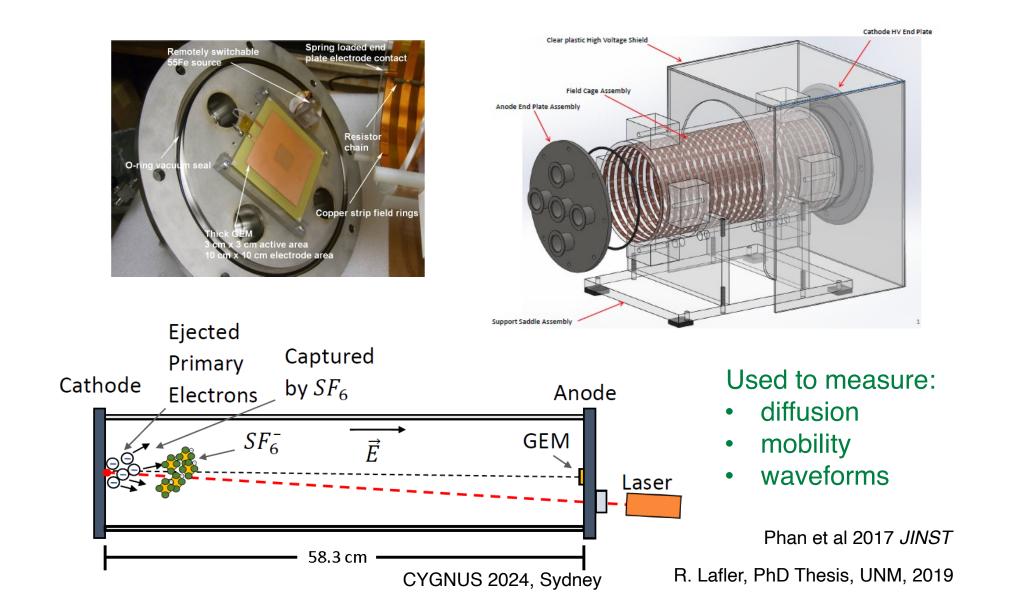
- Achieving high dynamic range does not require high gas gains
- Gas mixtures can be optimized for target atoms w/o worrying about light yield
- The detector will be built and characterized in the next year or two



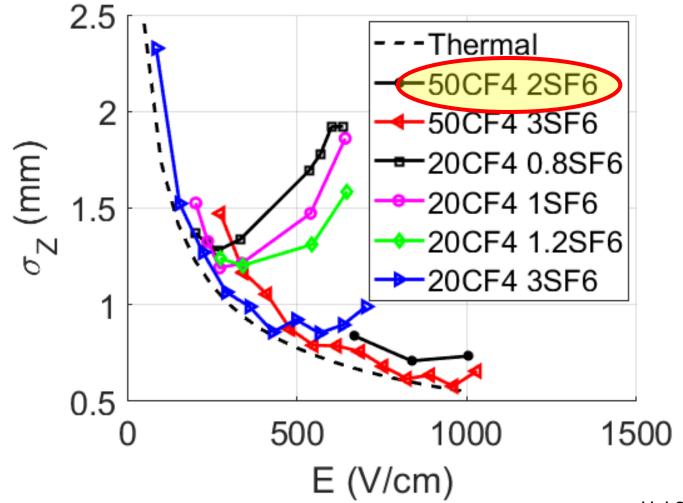


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R&D program to measure/optimize NI+noble gas properties



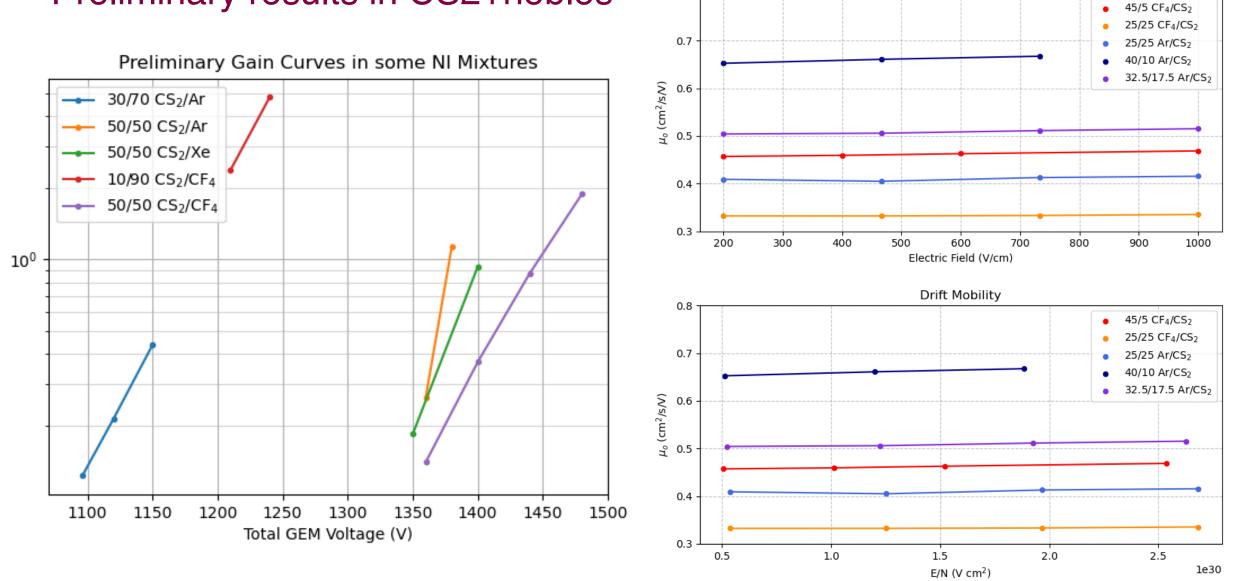
Results for σ_L in CF4/SF6 mixtures (60 cm drift):



н. Latler, PhD Thesis, UNM, 2019

Preliminary results in CS2+nobles

⁵⁵Fe Peak Mean Current [Au] (∝ gain)



0.8

Drift Mobility

Summary

- The MIGDAL experiment aims to perform an unambiguous observation of the Migdal effect.
- First science run took place with DD neutron source at the NILE facility at RAL.
- The detector performed well through the weeks of operation with highly ionizing NRs.
- Analysis of recorded data underway.
- 50% of recorded data are blinded.
- Second science run starting soon.
- Work is ongoing to create a next generation NI-TPC to probe lower energy Migdal events
- Stay tuned for results !

Thank You!



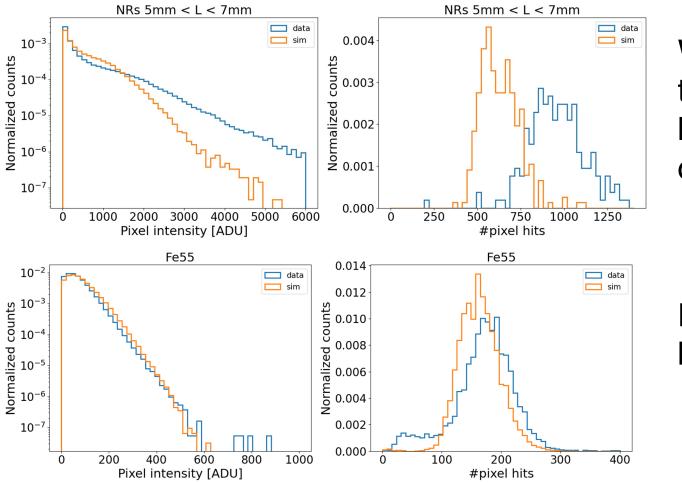




Expected Migdal backgrounds per 1 million DD-induced nuclear recoils with E > 100 keV

Component	Topology	D–D neutrons	
		>0.5	5–15 keV
Recoil-induced δ -rays	Delta electron from NR track origin	≈0	0
Particle-Induced X-ray Emission (PIXE)			
X-ray emission	Photoelectron near NR track origin	1.8	0
Auger electrons	Auger electron from NR track origin	19.6	0
Bremsstrahlung processes ^a			
Quasi-Free Electron Br. (QFEB)	Photoelectron near NR track origin	112	≈0
Secondary Electron Br. (SEB)	Photoelectron near NR track origin	115	≈0
Atomic Br. (AB)	Photoelectron near NR track origin	70	≈0
Nuclear Br. (NB)	Photoelectron near NR track origin	≈ 0	≈0
Neutron inelastic γ -rays	Compton electron near NR track origin	1.6	0.47
Random track coincidences			
External γ - and X-rays	Photo-/Compton electron near NR track	≈ 0	≈0
Trace radioisotopes (gas)	Electron from decay near NR track origin	0.2	0.01
Neutron activation (gas)	Electron from decay near NR track origin	0	0
Muon-induced δ -rays	Delta electron near NR track origin	≈ 0	≈0
Secondary nuclear recoil fork	NR track fork near track origin	-	≈ 1
Total background	Sum of the above components		1.5
Migdal signal	Migdal electron from NR track origin		32.6

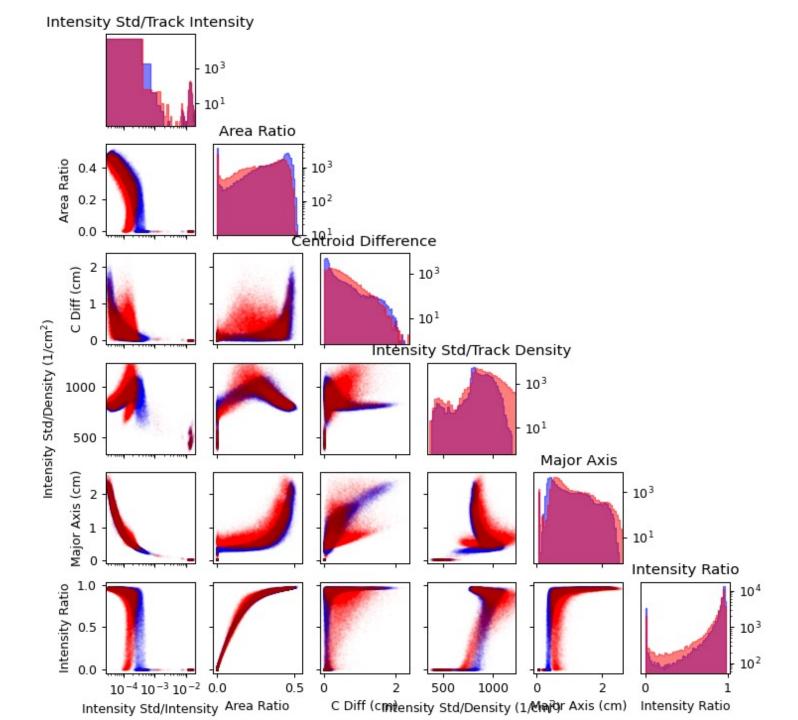
Simulation vs. Data

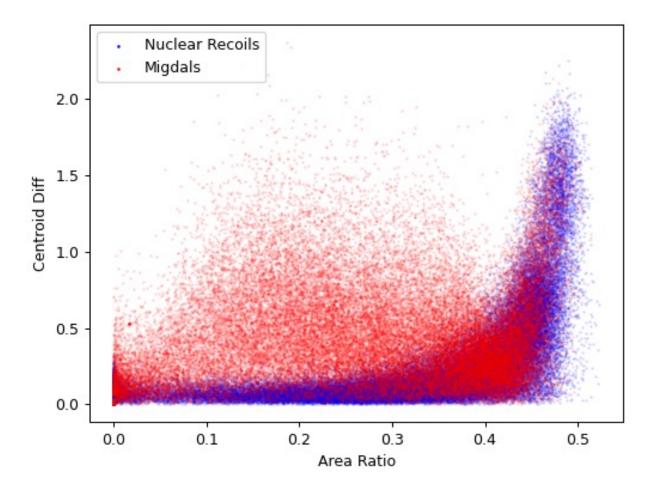


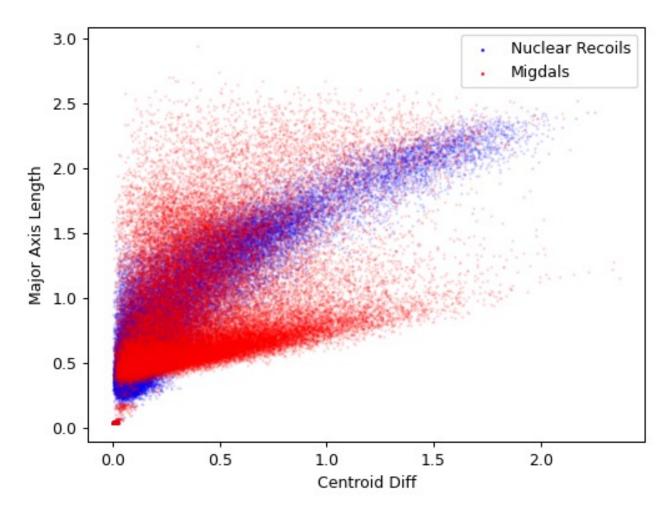
We believe that reflections are the origin of the differences between sim vs data NR distributions.

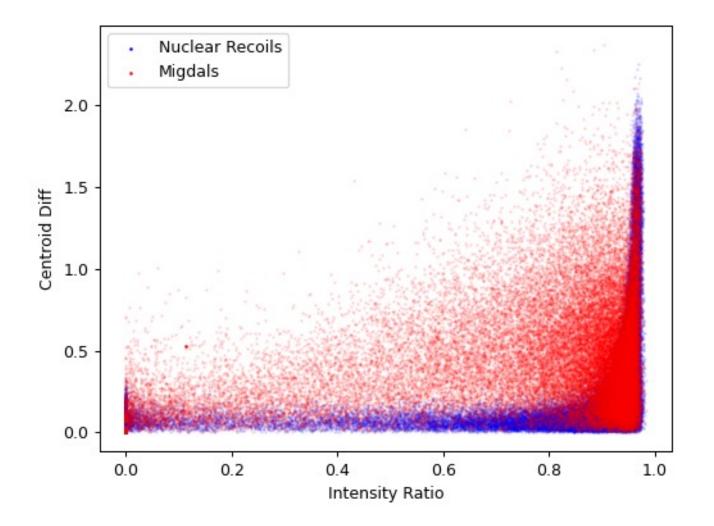
Data and simulation agree much better for simulated ERs

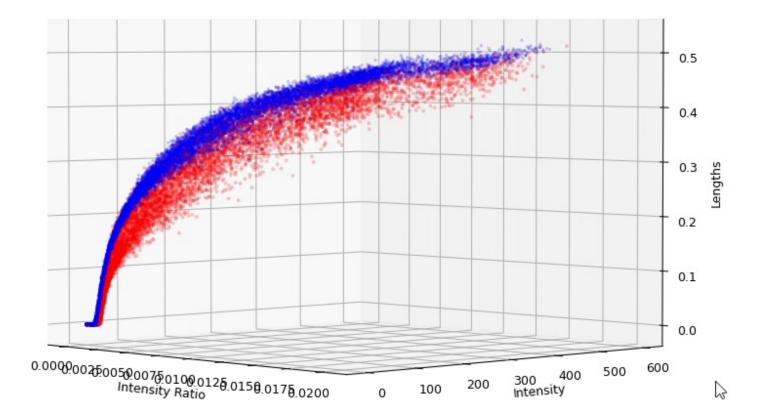
The main problem is that simulating NRs in our optical readout is very challenging. We plan to pursue an idea to stitch simulated ERs on data NRs to form Migdal signals for YOLO (or CNN!) to be trained and tested on.



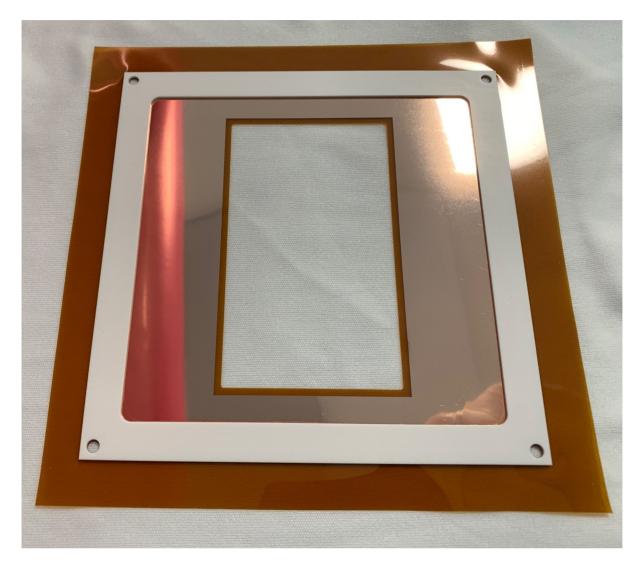


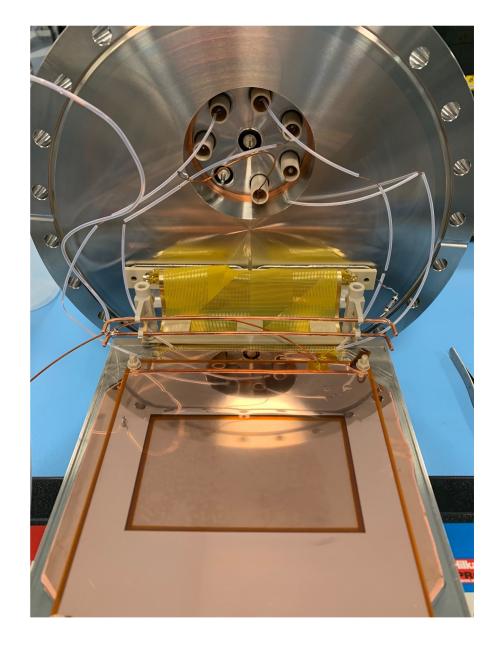




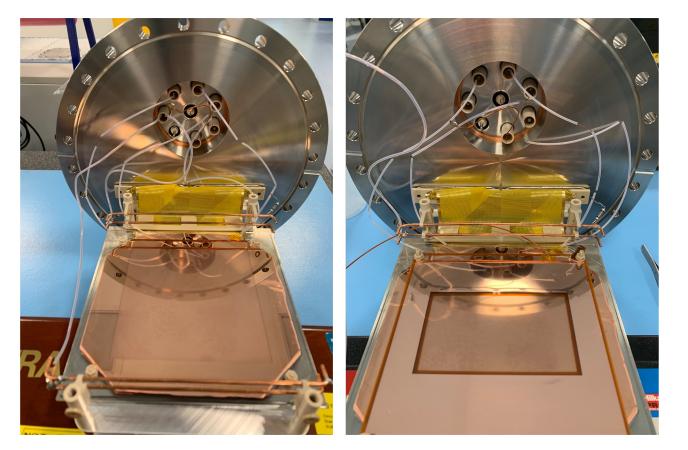


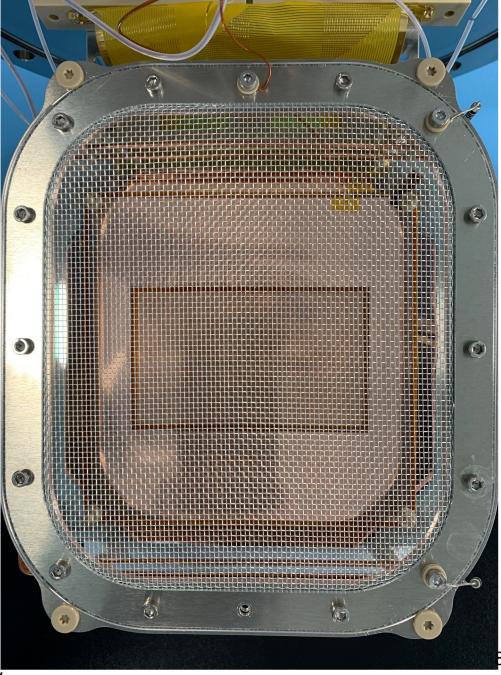
Mask assembly





Mask assembly





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The Migdal effect has also been studied for neutrons on helium....an intriguing prediction:

Neutron-impact ionization of He

M S Pindzola¹, T G Lee¹, Sh A Abdel-Naby¹, F Robicheaux², J Colgan³ and M F Ciappina⁴

J. Phys. B: At. Mol. Opt. Phys. 47 (2014) 195202

We present energy and angle differential cross sections for the neutron-impact single ionization of He at 100 keV in figure 4. The TDCC results are for single ionization leaving He⁺ in the ground state, where the outgoing electron momentum k = 2.0 (E = 54.4 eV) and $\phi = 0$. We find that the electrons prefer to leave in the opposite direction to the target nucleus.

We would also like to measure the angular distribution of the Migdal electron...

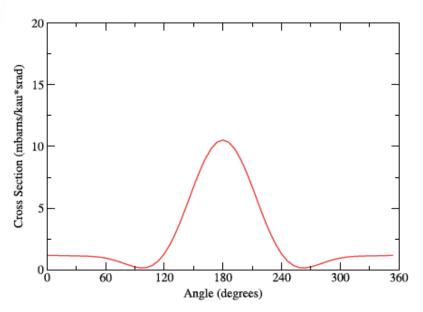
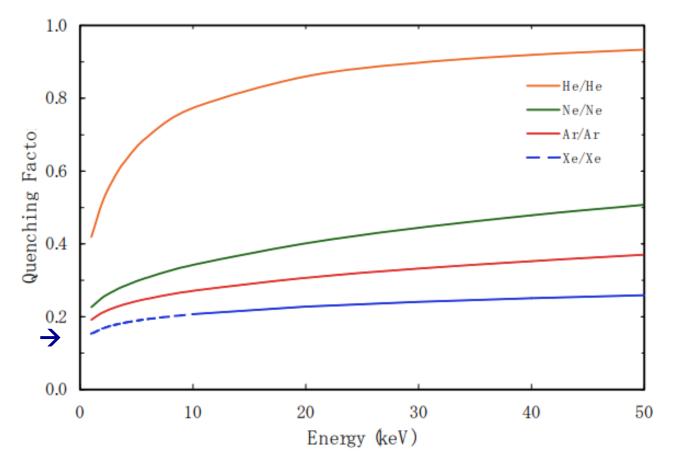


Figure 4. Neutron-impact single ionization of He at 100 keV. Solid line (red): TDCC method for the single ionization differential cross section with k = 2.0 and $\phi = 0$ (1.0 mbarn = 1.0×10^{-27} cm², kau = momentum in au, srad = solid angle in radians).

Punch line is that only 10-20% of the NR energy is measured by XENON in the region of interest.



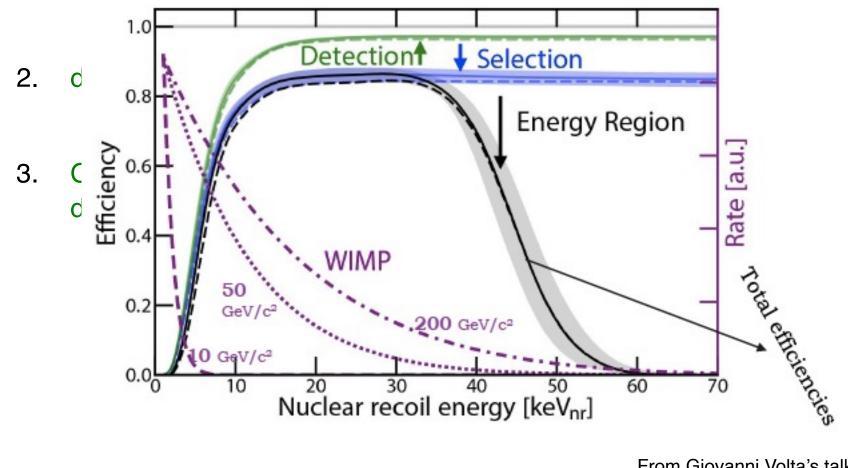
 \rightarrow This means ~0.01 – 0.02 keV for 1 GeV WIMP!

Figure 2. The nuclear quenching factor $q_{\rm nc}$ for recoil ions in, from top to bottom, He, Ne, Ar and Xe as a function of the recoil energy at 1-50 keV.

From Akira Hitachi 2007 J. Phys.: Conf. Ser. 65 012013

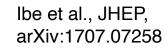
Summarizing the reasons for XENON's insensitivity to low mass WIMPs:

1. Maximum recoil energy for a 1 GeV/c² WIMP:

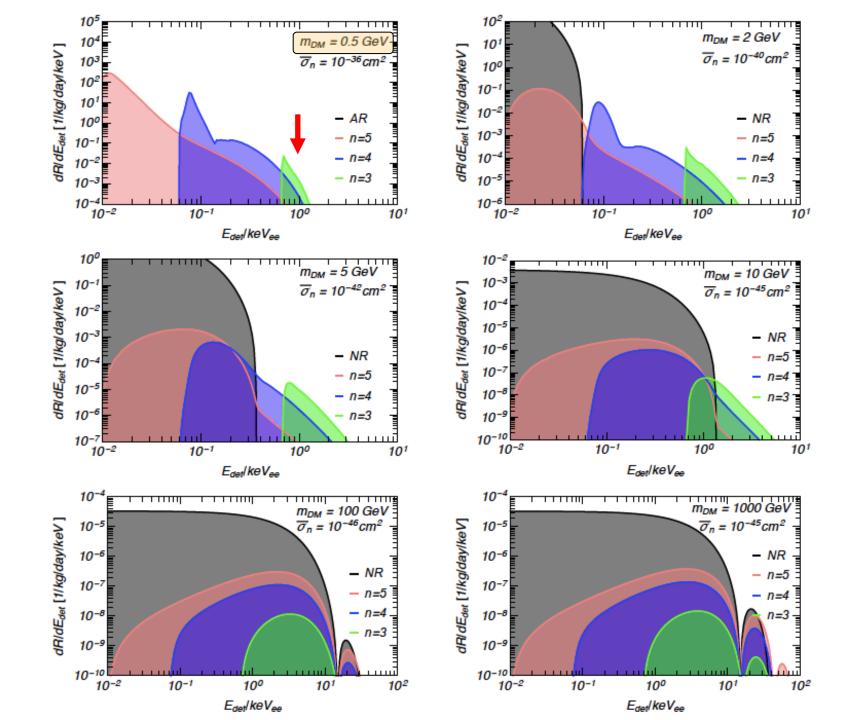


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From Giovanni Volta's talk at the Swiss Physics Society Annual Meeting, August 29, 2019



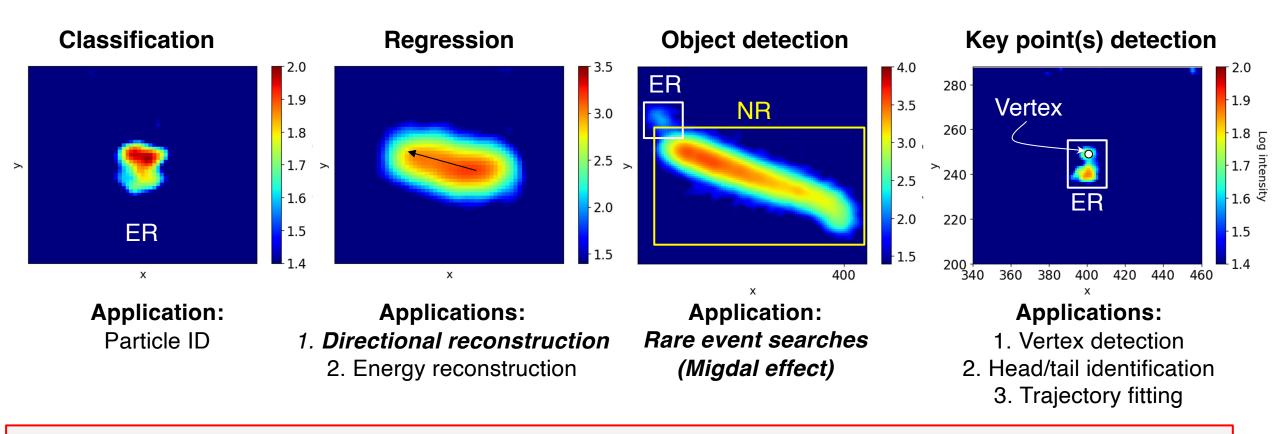
Differential event rate as a function of detected energy



D Loomba

84

YOLO Capabilities



Many implementations of algorithms for object detection and key point detection already exist (see <u>here</u> and <u>here</u>), and there are many tools for developing custom classification and regression models for 2D and 3D image data (<u>PyTorch tutorial</u>, <u>TensorFlow tutorial</u>, <u>spconv for sparse 3D convolutional neural networks</u>). *Today I'll highlight examples of (1) directional reconstruction for a CYGNUS prototype BEAST TPC and (2) object detection for the rare event Migdal search*

D Loomba