

# Current Status of Axion Searches and Future Prospects

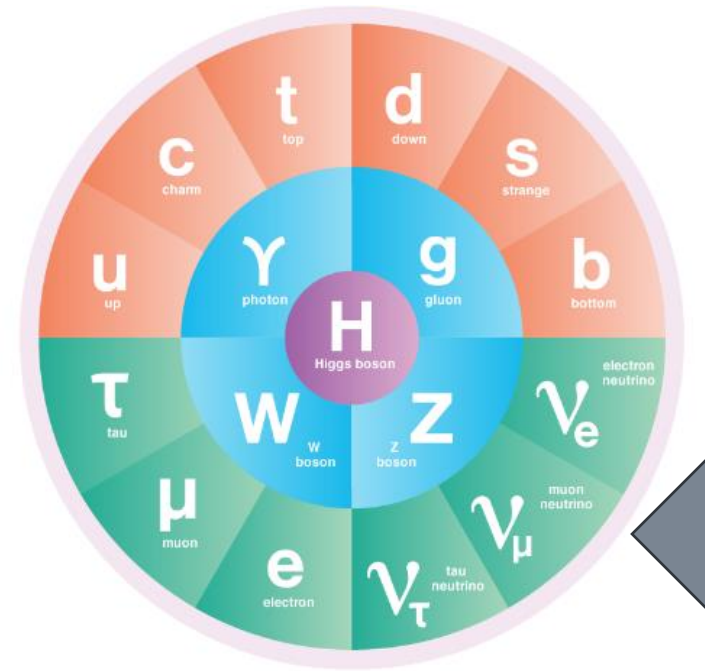
# AXION EXPERIMENTS

- Introduction
- Current efforts
- Prospects



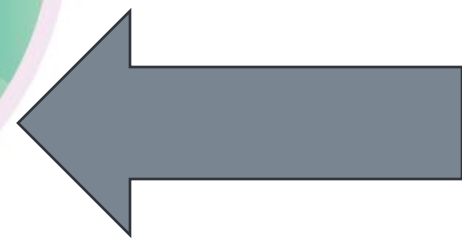
# What is it all about?

## - Standard model



● QUARKS  
 ● LEPTONS  
 ● BOSONS  
 ● HIGGS BOSON

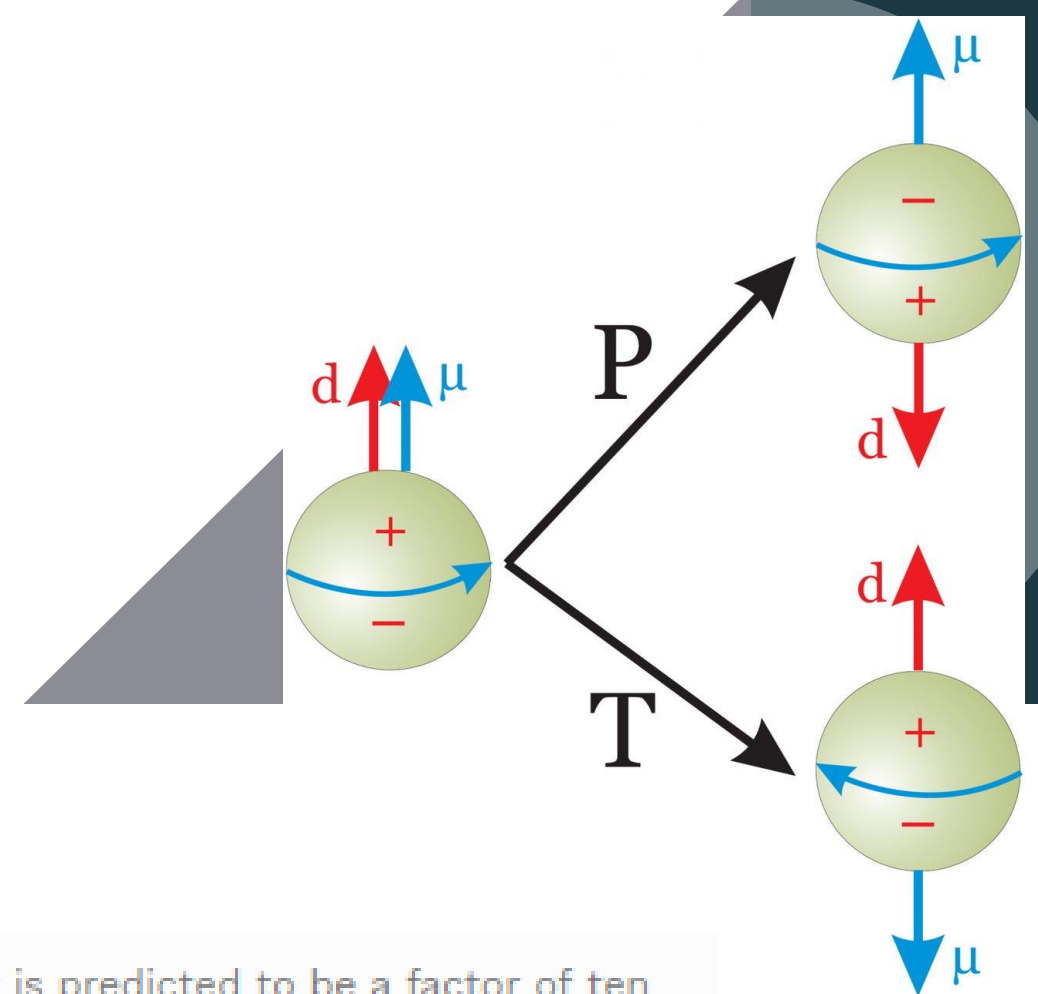
[https://www.energy.gov/sites/default/files/styles/full\\_article\\_width/public/2020/06/f75/gp-avg-standard-model-particle-physics.png?itok=dZHTSTIc](https://www.energy.gov/sites/default/files/styles/full_article_width/public/2020/06/f75/gp-avg-standard-model-particle-physics.png?itok=dZHTSTIc)



Yesterday  
--> 2040

# What is it all about?

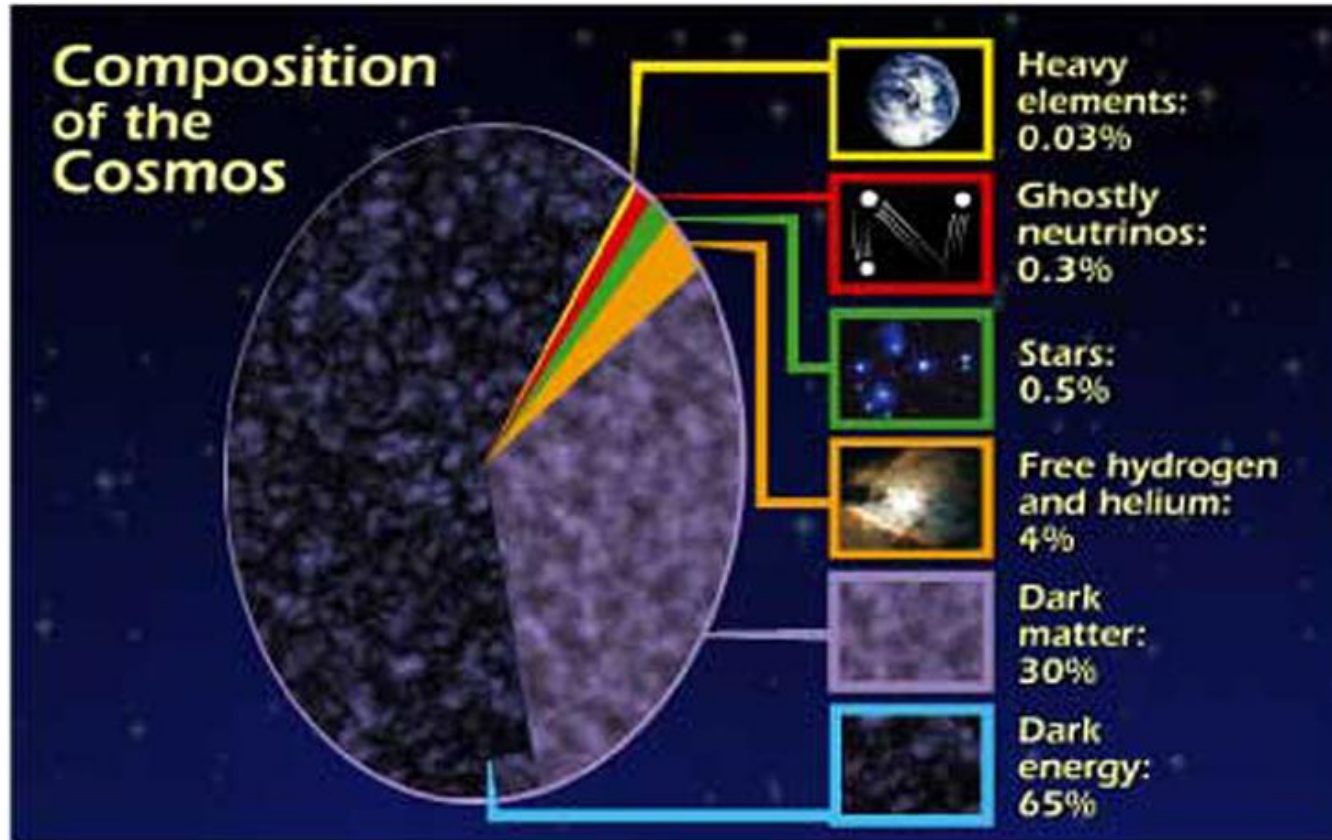
## - Strong CP problem



In the Standard Model, the neutron's electric dipole moment is predicted to be a factor of ten billion larger than our observational limits show. The only explanation is that somehow, something beyond the Standard Model is protecting this CP symmetry in the strong interactions. We can demonstrate a lot of things in science, but proving that CP is conserved in the strong interactions can never be done. However, solving the strong CP problem may be closer on the horizon than almost anyone realizes. [-] PUBLIC DOMAIN WORK FROM ANDREAS KNECHT

# What is it all about?

## - Composition of the Universe



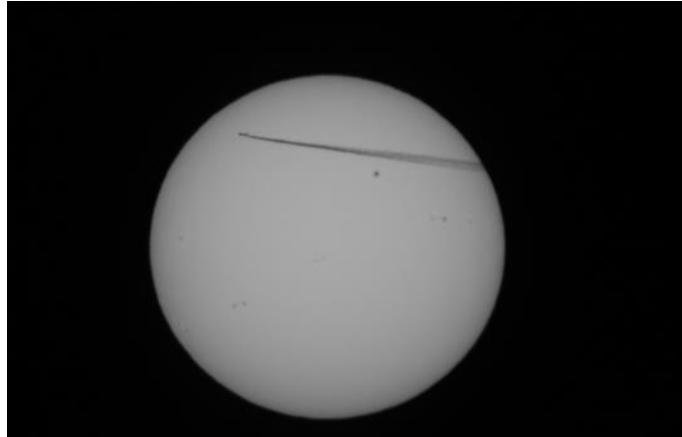
Ann Field (STScI)

# The answer? New particle. Axion!

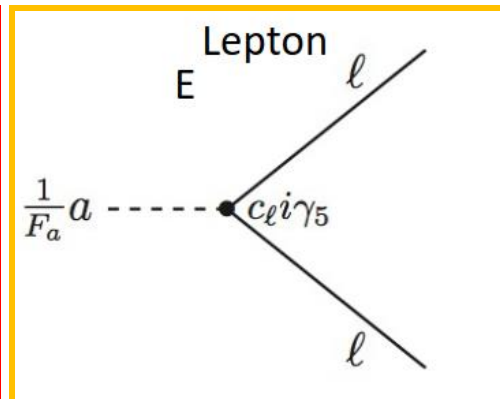
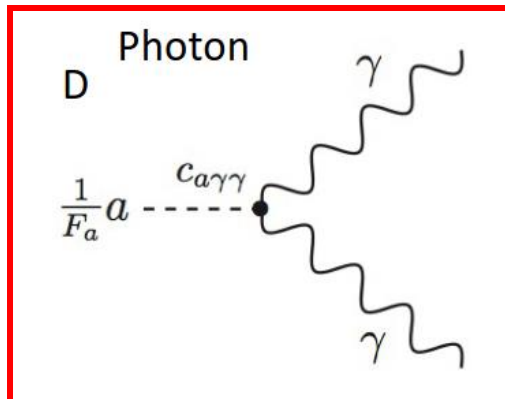
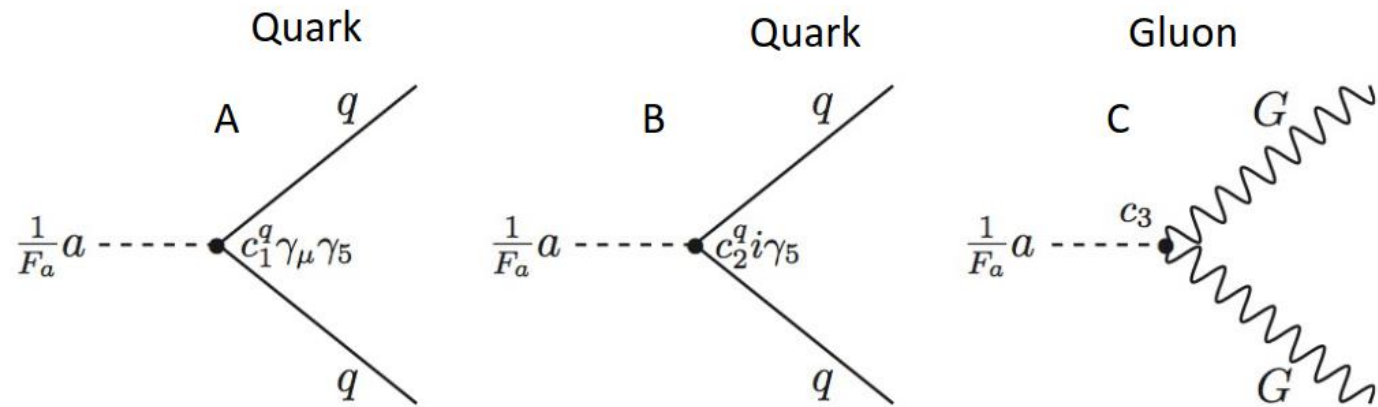
Strong CP problem  
1977 Peccei and Quinn --> new global symmetry  
--> spontaneously broken  
Wilczek and Weinberg --> new particle  
pseudo-Nambu-Goldstone boson



# Where find it?



# Strategy?



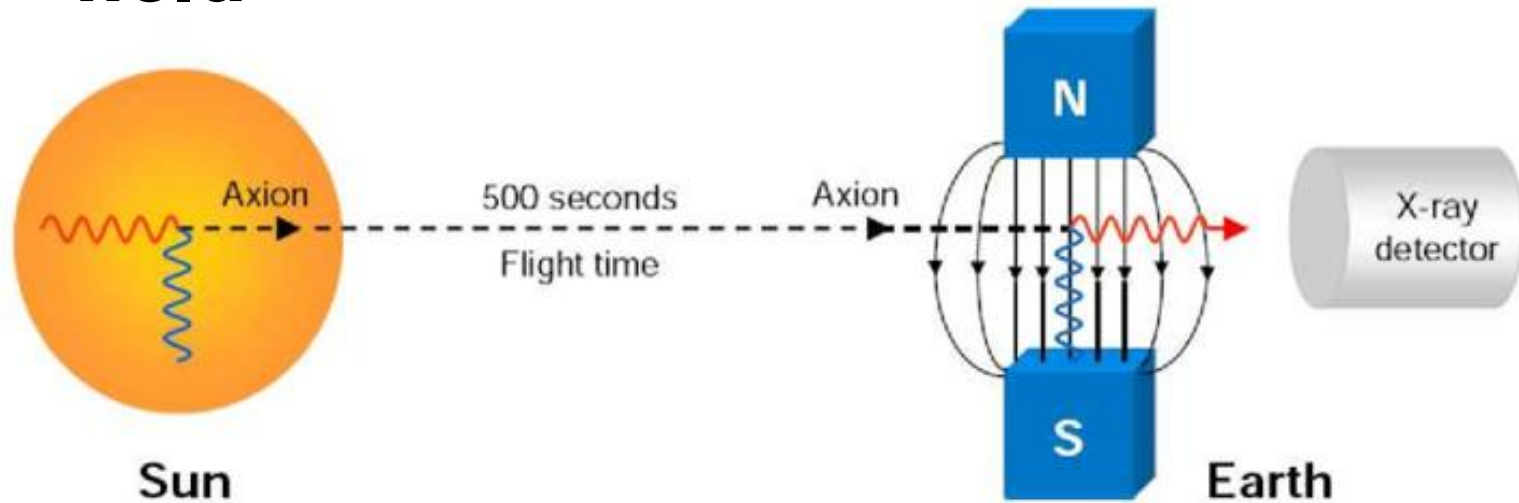
adapted from G. Ruoso



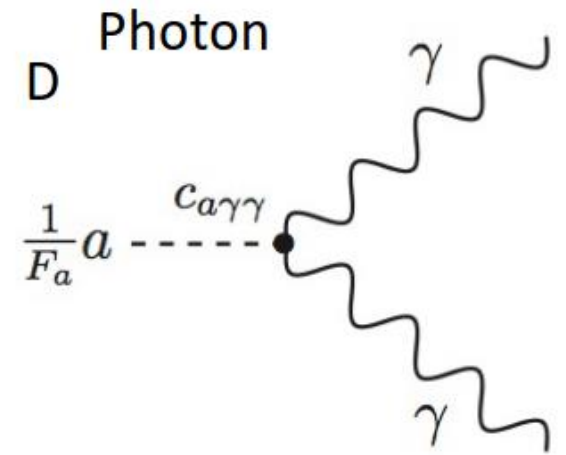


# Detection strategy

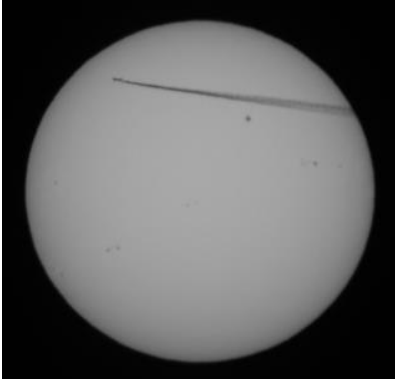
- Sun --> plenty of photons
- and strong magnetic field



*P. Sikivie, PRL 51, 1415-1417 (1983)*



# Most promising



## Helioscopes "a la Sikivie"

- CAST
- SUMICO
- BabyIAXO
- IAXO

See talk by J. Vogel



## Haloscopes

- CAST
- GRAHAL
- ADMX
- CAPP
- MADMAX
- QUAX
- DM Radio
- ABRACADABRA
- ORPHEUS
- WISPDMX
- HAYSTAC
- FLASH



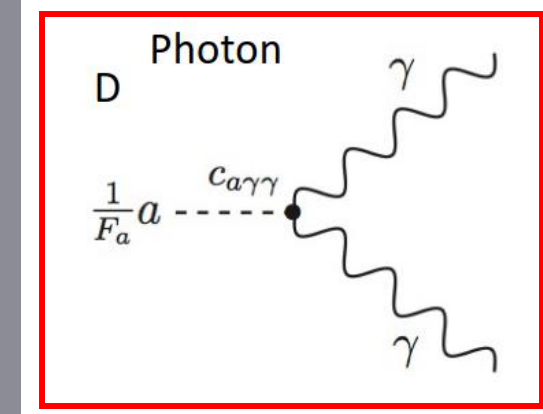
## Laboratory

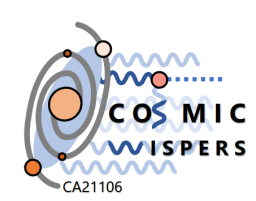
### LSW

- ALPS
- OSQAR

### Polarization

- PVLAS
- BFRT
- BMV
- OSQAR
- Q@A





# CAST

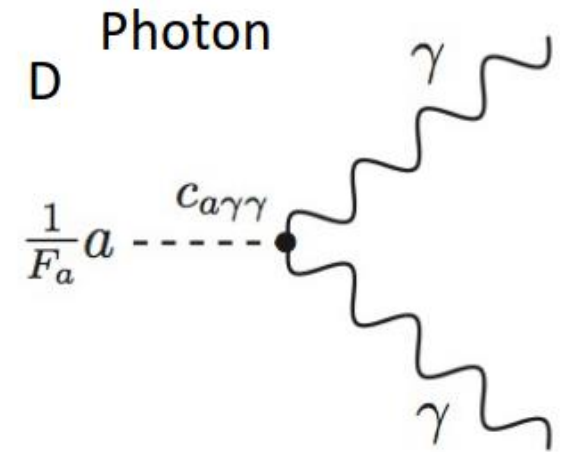
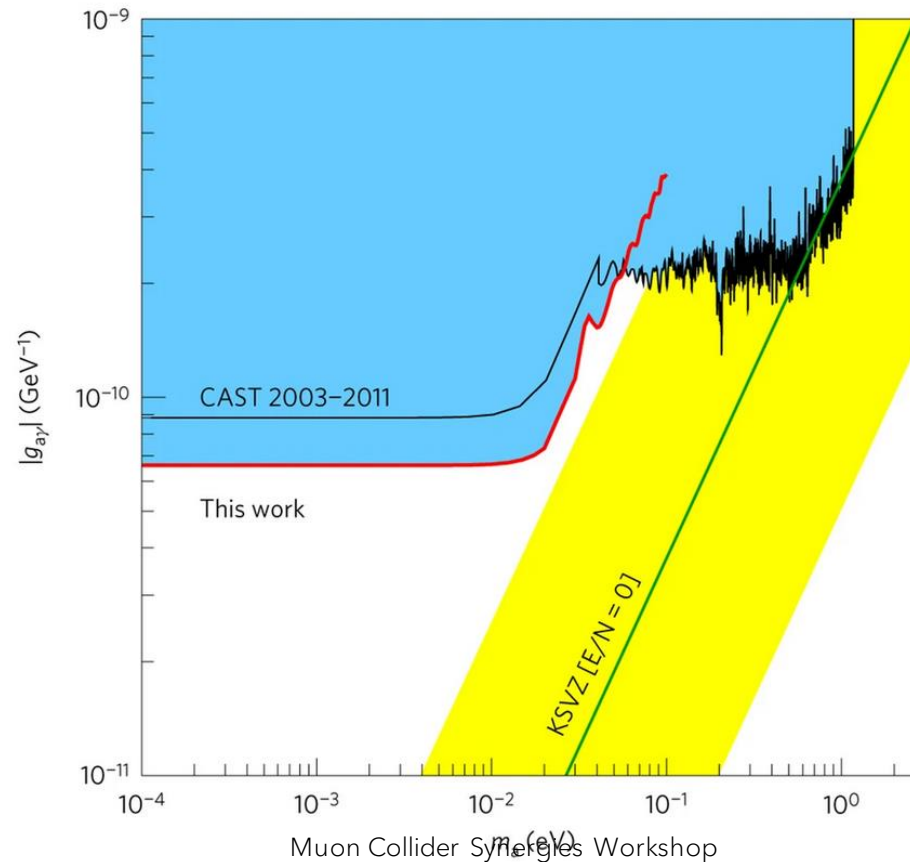
## Results - helioscope

Open Access | Published: 01 May 2017

### New CAST limit on the axion-photon interaction

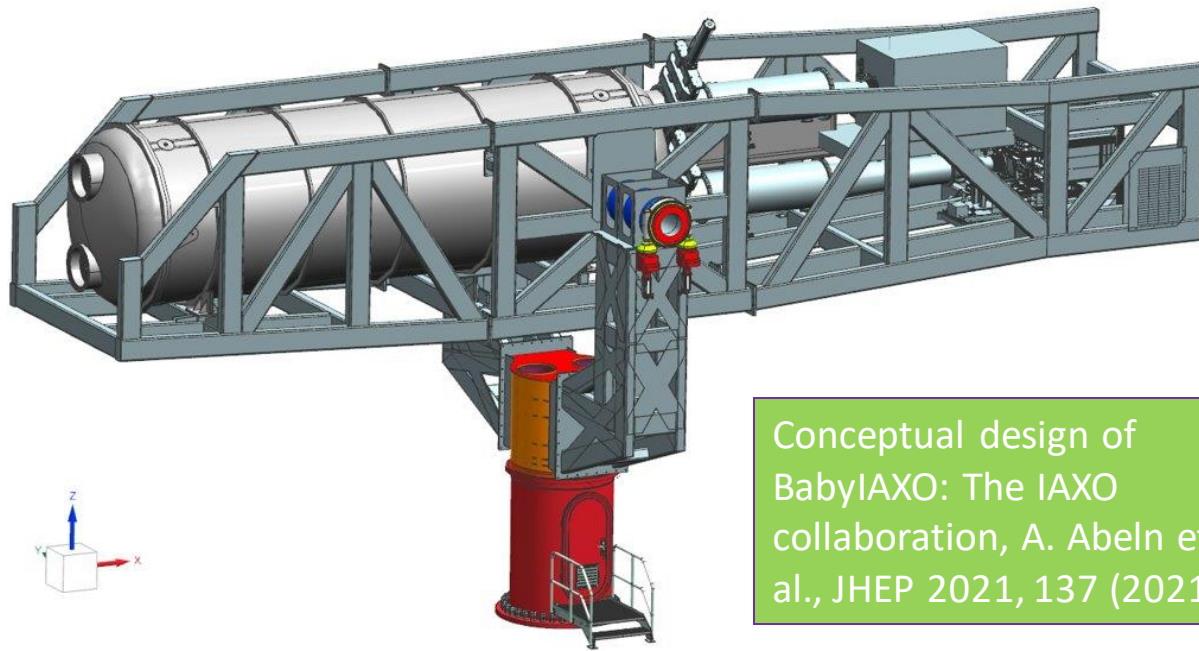
CAST Collaboration

Nature Physics 13, 584-590 (2017) | Cite this article



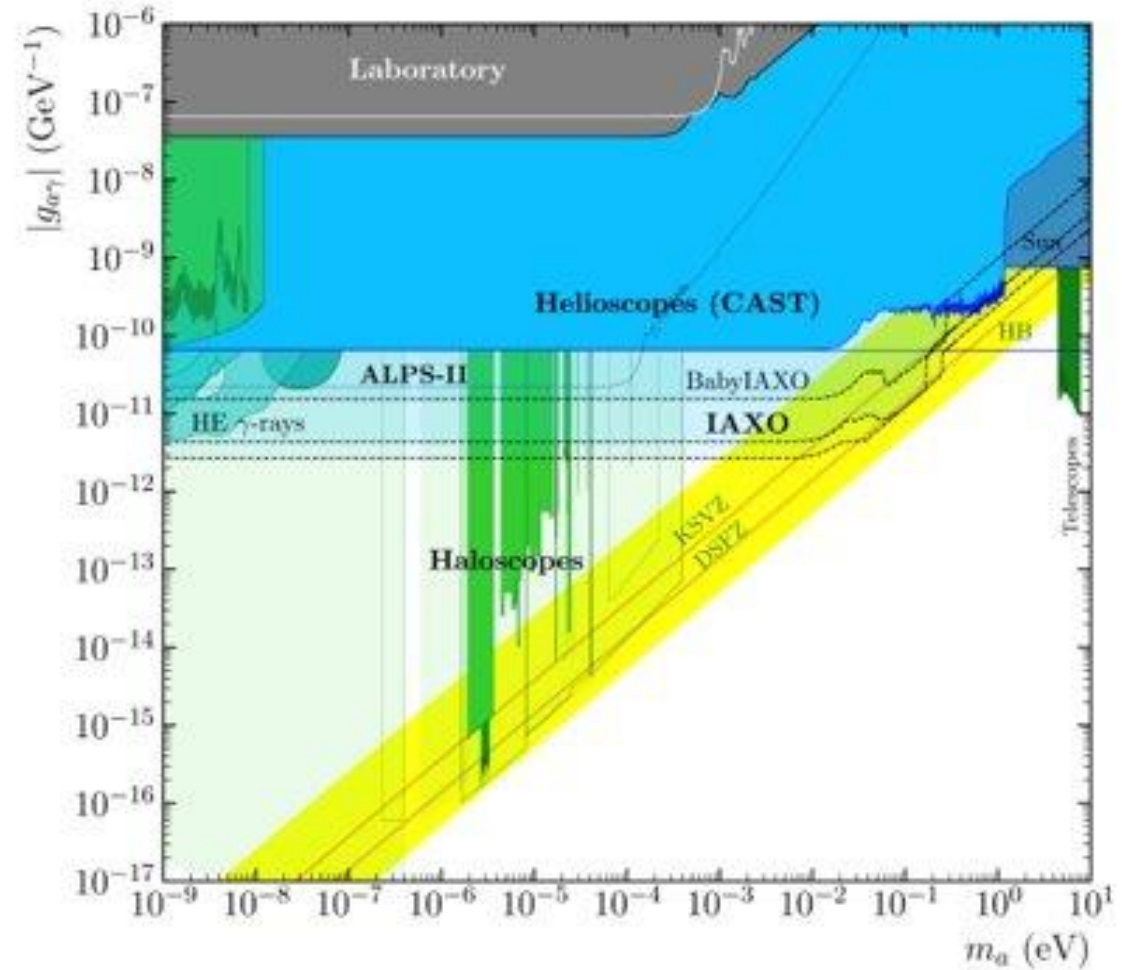


# The International AXion Observatory



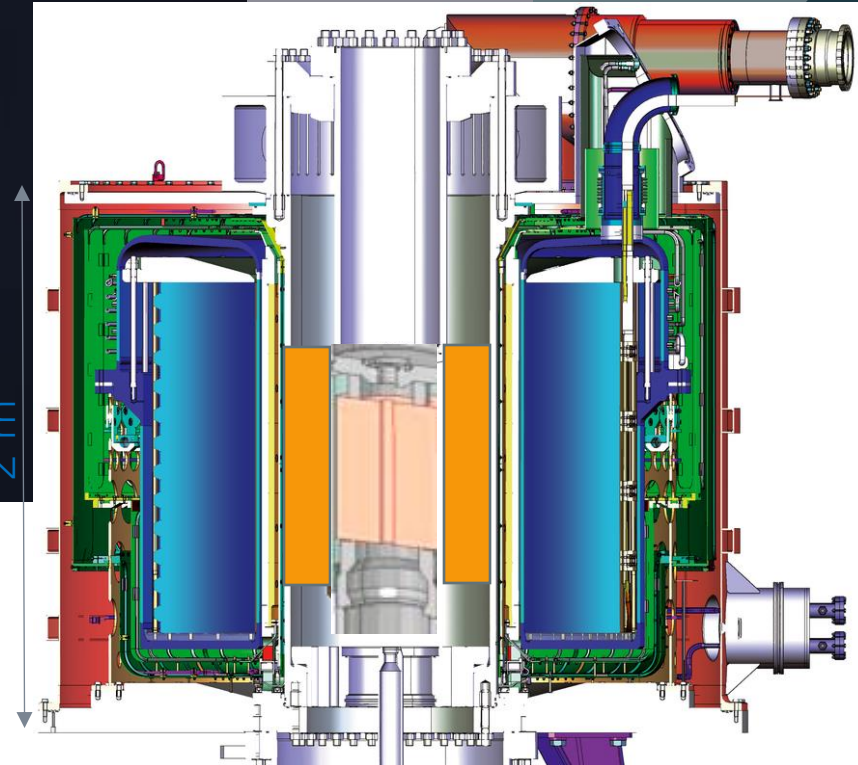
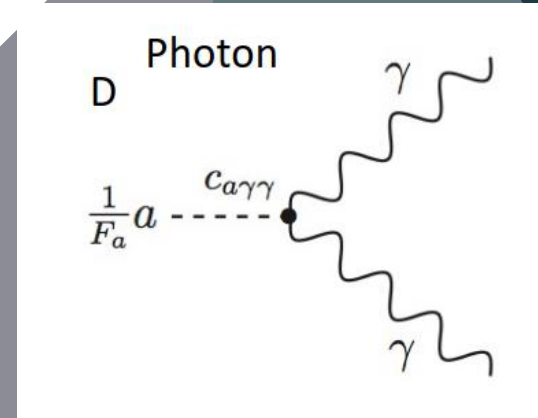
Conceptual design of BabyIAXO: The IAXO collaboration, A. Abeln et al., JHEP 2021, 137 (2021)

- Technological prototype with only two magnet bores (10 m,  $\varnothing$  70 cm)
- 2 detection lines with one x-ray telescope and one ultra-low background detector each
- Moving platform to track the Sun



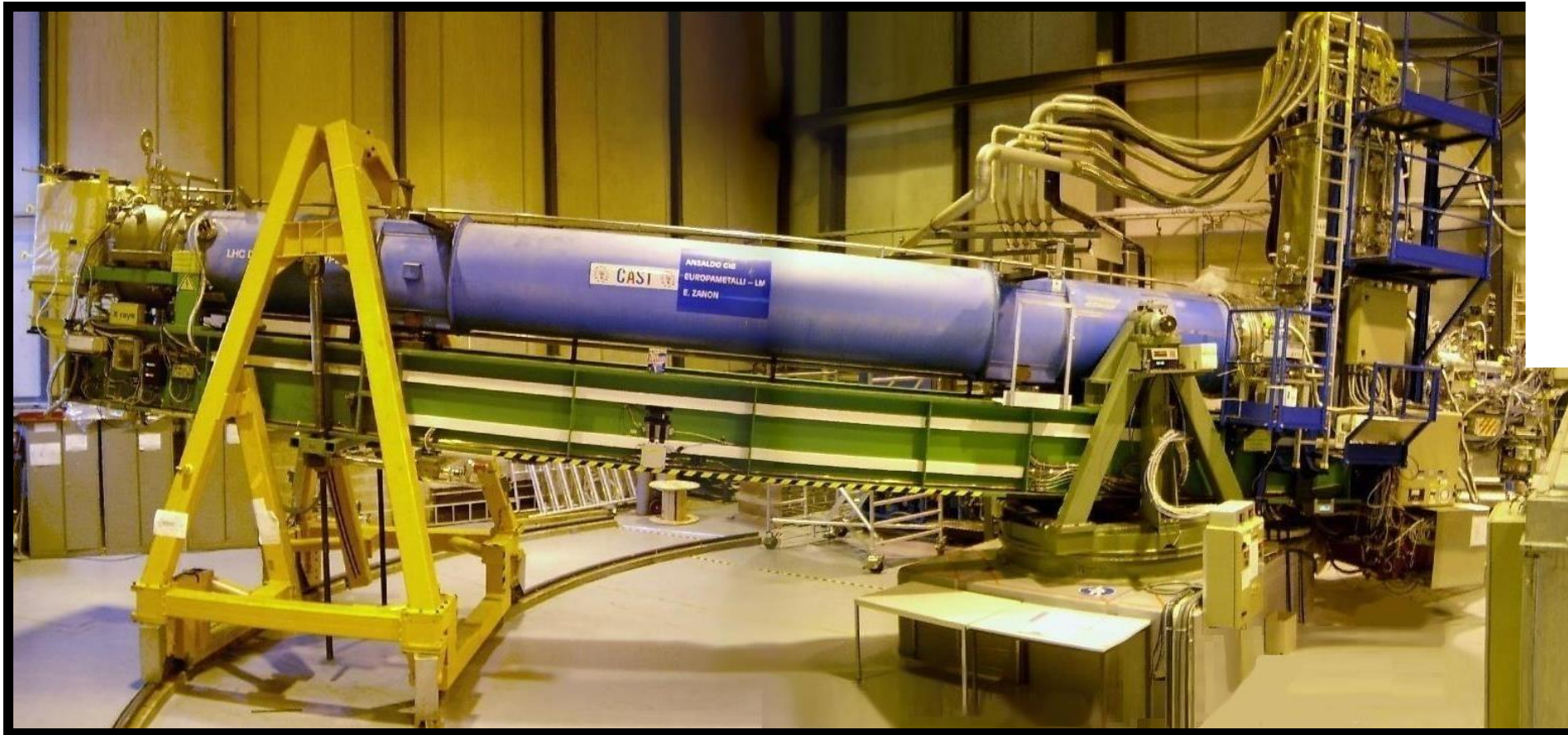
# Haloscopes

- CAST
- GRAHAL
- ADMX
- CAPP
- MADMAX
- QUAX
- DM Radio
- ABRACADABRA
- ORPHEUS
- WISPDMX
- HAYSTAC
- FLASH



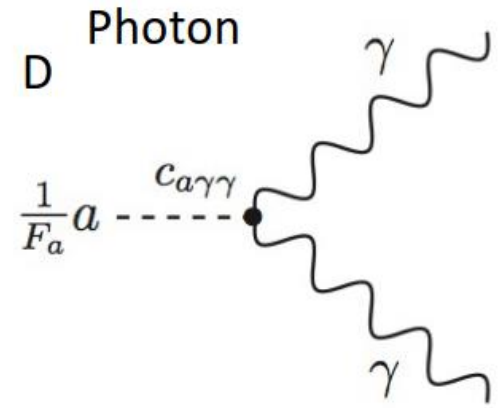
# CAST - CERN Axion Solar Telescope

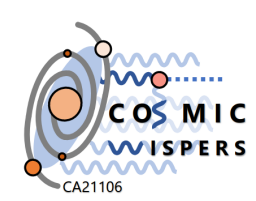
## - galactic halo



**CAPP**

**RADES**





# CAST Results - haloscope

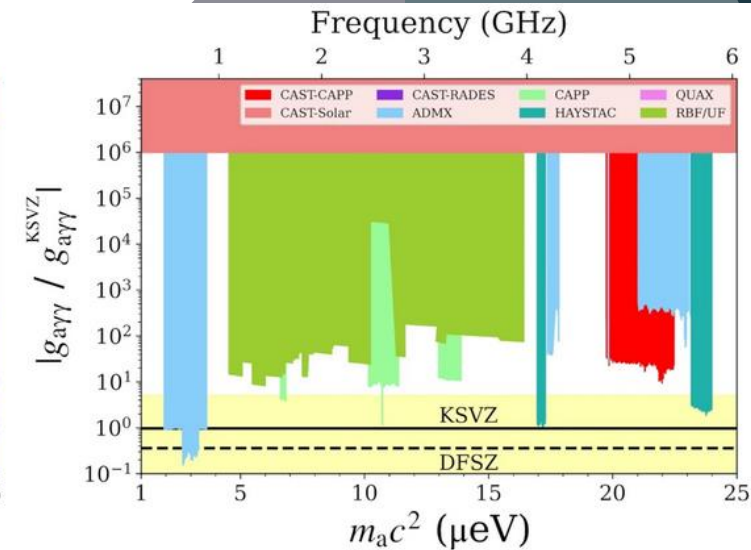
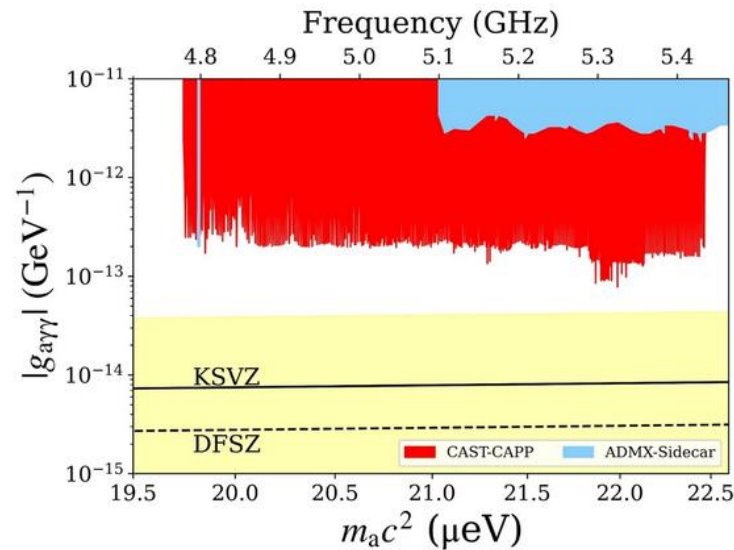
Article | [Open Access](#) | [Published: 19 October 2022](#)

## Search for Dark Matter Axions with CAST-CAPP

[C. M. Adair](#), [K. Altenmüller](#), [V. Anastassopoulos](#), [S. Arguedas Cuendis](#), [J. Baier](#), [K. Barth](#), [A. Belov](#), [D. Bozicevic](#), [H. Bräuninger](#), [G. Cantatore](#), [F. Caspers](#), [J. F. Castel](#), [S. A. Çetin](#), [W. Chung](#), [H. Choi](#), [J. Choi](#), [T. Dafni](#), [M. Davenport](#), [A. Dermenev](#), [K. Desch](#), [B. Döbrich](#), [H. Fischer](#), [W. Funk](#), [J. Galan](#), ... [K. Zioutas](#)

+ Show authors

[Nature Communications](#) **13**, Article number: 6180 (2022) | [Cite this article](#)



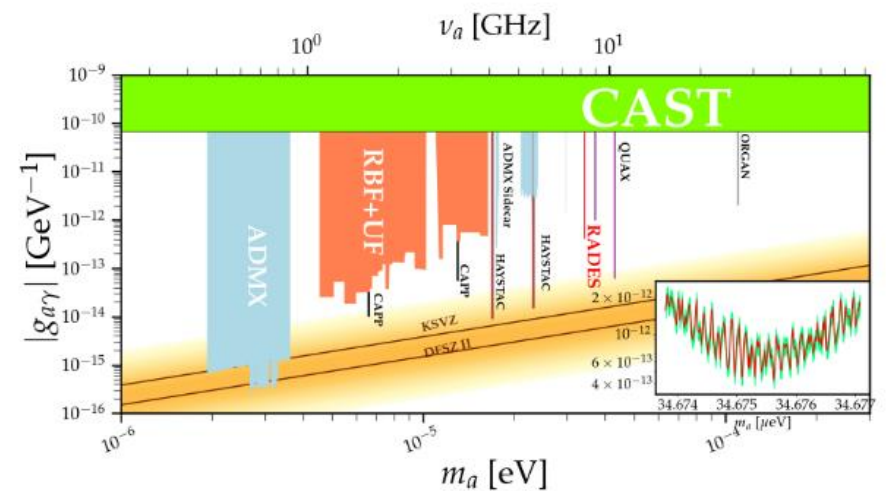
Regular Article - Experimental Physics | [Open Access](#) | [Published: 08 October 2021](#)

## First results of the CAST-RADES haloscope search for axions at $34.67 \mu\text{eV}$

[A. Álvarez Melcón](#), [S. Arguedas Cuendis](#) ✉, [J. Baier](#), [K. Barth](#), [H. Bräuninger](#), [S. Calatroni](#), [G. Cantatore](#), [F. Caspers](#), [J. F. Castel](#), [S. A. Cetin](#), [C. Cogollos](#), [T. Dafni](#), [M. Davenport](#), [A. Dermenev](#), [K. Desch](#), [A. Díaz-Morcillo](#), [B. Döbrich](#), [H. Fischer](#), [W. Funk](#), [J. D. Gallego](#), [J. M. García Barceló](#), [A. Gardikiotis](#), [J. G. Garza](#), [B. Gimeno](#), ... [K. Zioutas](#)

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[Journal of High Energy Physics](#) **2021**, Article number: 75 (2021) | [Cite this article](#)





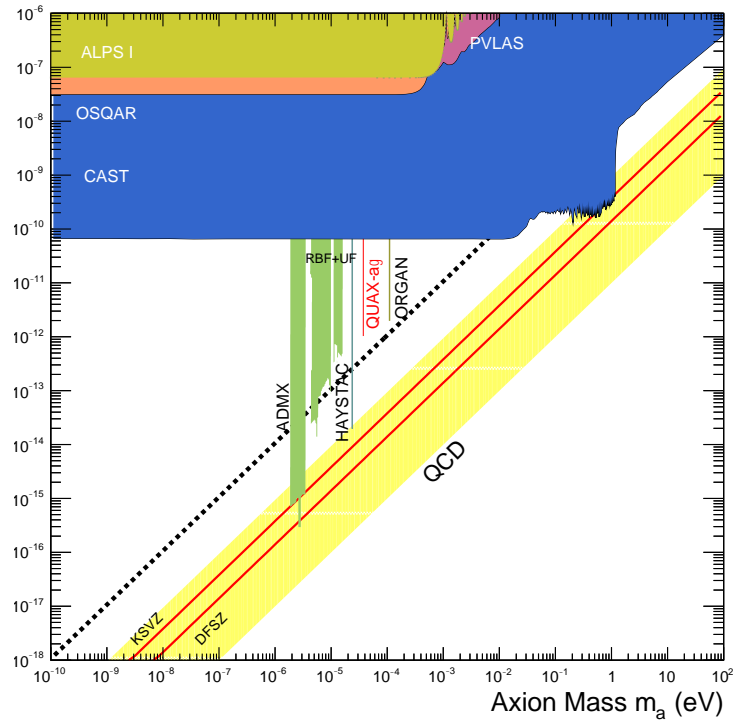
# Haloscopes QUAX

QUAX- $\alpha\gamma$  Result with Superconductive Resonant Cavity at  $m_a = 37.5$  meV, Phys. Rev. D **99**, 101101(R) (2019).

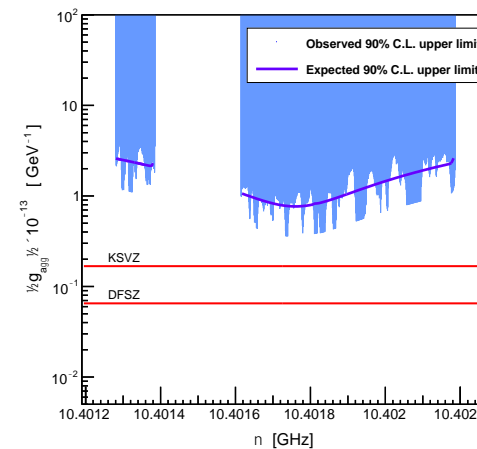
Search for Invisible Axion Dark Matter of mass  $m_a = 43$  meV with the QUAX- $\alpha\gamma$  Experiment, Phys. Rev. D **103**, 102004 (2021).

Search for Galactic Axions with high-Q Dielectric Cavity, Phys. Rev. D **106**, 052007 (2022).

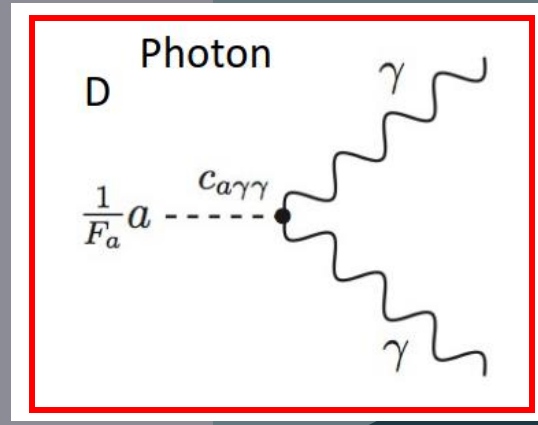
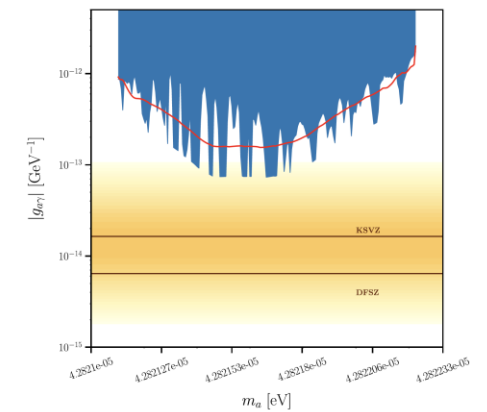
Quax- $\alpha\gamma$  2019



Quax- $\alpha\gamma$  2021



Quax- $\alpha\gamma$  2022



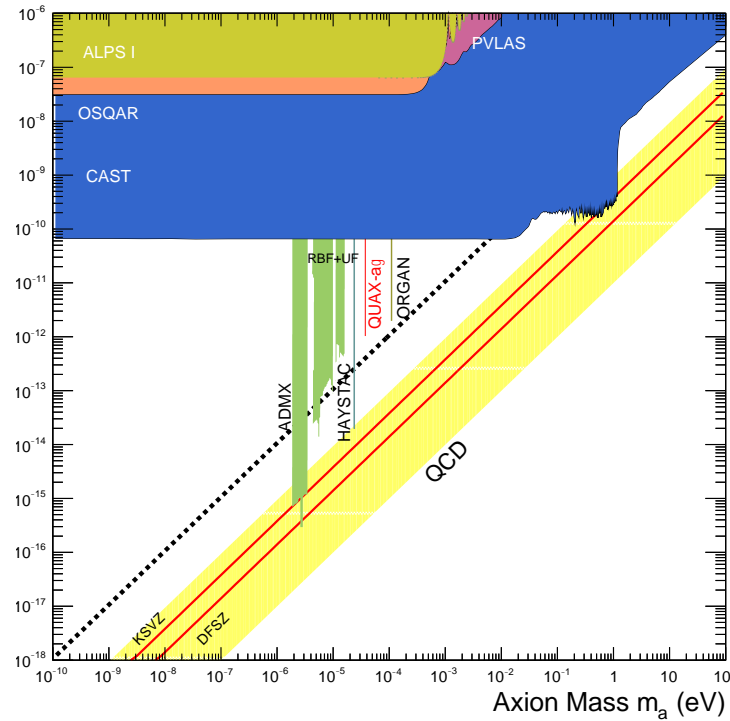
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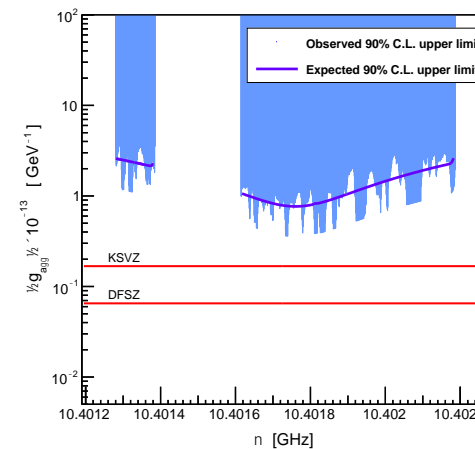
Search for Invisible Axion Dark Matter of mass  $m_a = 43$  meV with the QUAX- $\alpha\gamma$  Experiment, Phys. Rev. D **103**, 102004 (2021).

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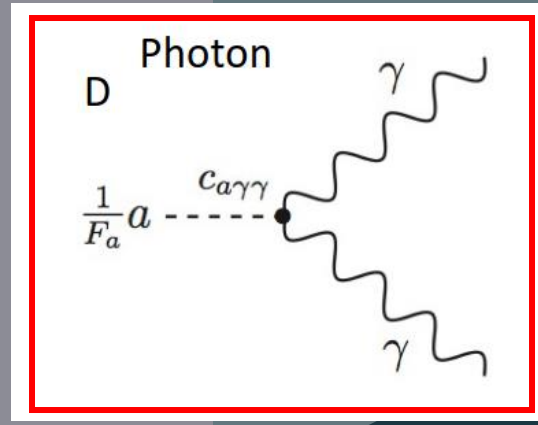
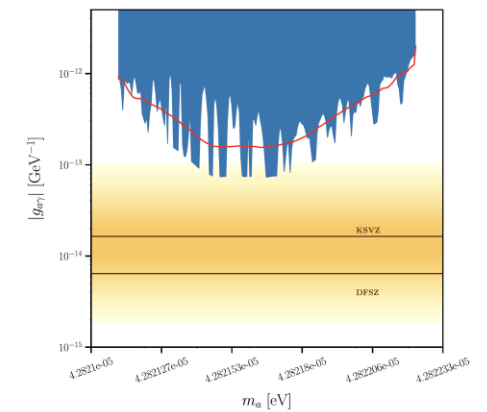
Quax- $\alpha\gamma$  2019



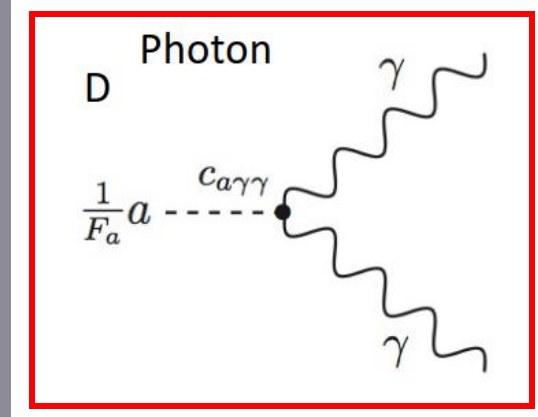
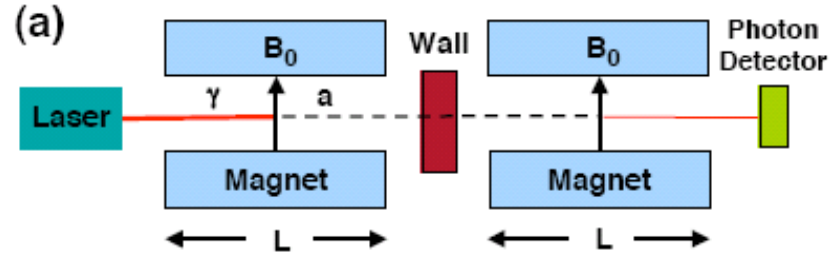
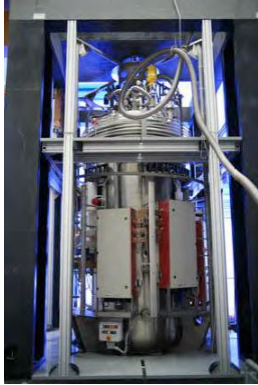
Quax- $\alpha\gamma$  2021



Quax- $\alpha\gamma$  2022

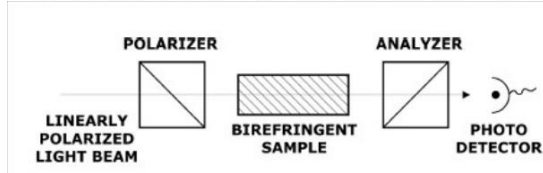
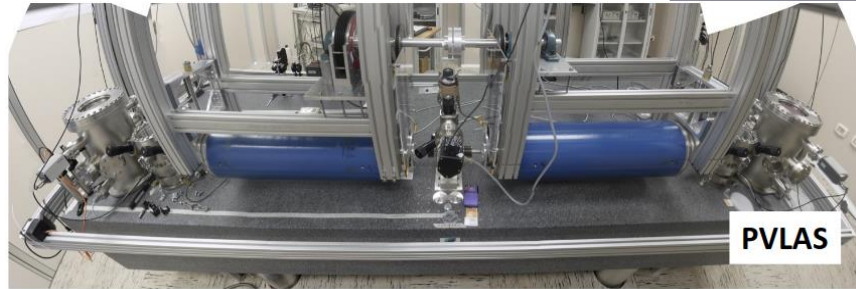


# Laboratory

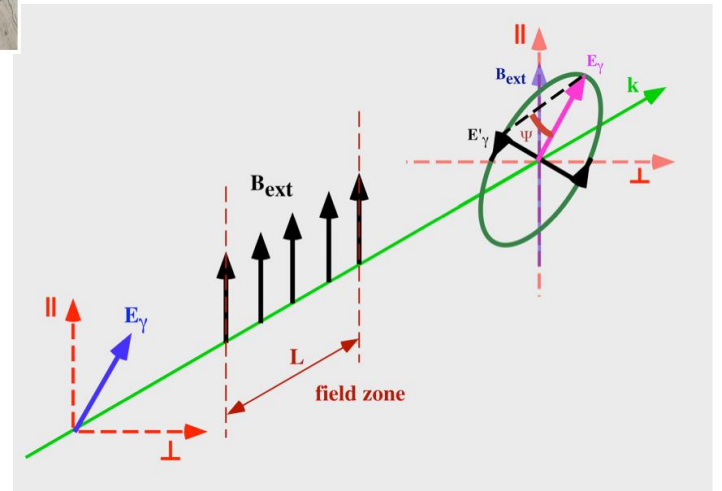


- LSW
- ALPS
  - OSQAR

- Polarization
- PVLAS
  - BFRT
  - BMV
  - OSQAR
  - Q@A

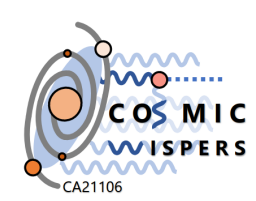


L.Maiani, R. Petronzio, E. Zavattini, Phys. Lett B, Vol. 173, no.3 1986



It's between these last two magnets where light is supposed to shine through the wall when ALPS II goes into operation. Credit: DESY / Heiner Müller-Elsner

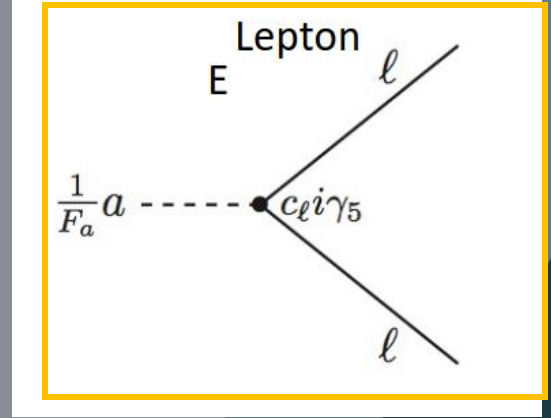




# QUAX

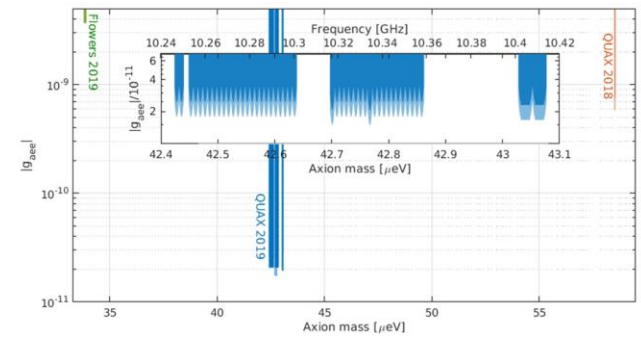
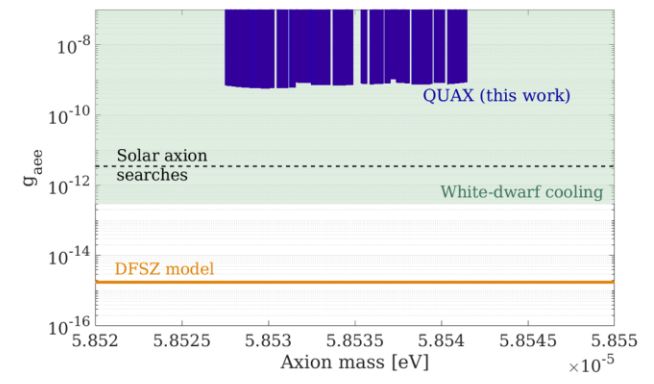
QUAX-ae result with Ferromagnetic Axion Haloscope at  $m_a = 58 \text{ meV}$ , EPJC (2018) 78:703.

QUAX-ae with Quantum-Limited Ferromagnetic Haloscope, Phys. Rev. Lett. 124, 171801 (2020).



Quax-ae 2018

Quax-ae 2020



QUAX



Trento Institute for Fundamental Physics and Applications



And new collaborations with



# NOT ONLY

## First axion results from the XENON100 experiment

E. Aprile *et al.* (XENON100 Collaboration)

Phys. Rev. D **90**, 062009 – Published 9 September 2014; Erratum [Phys. Rev. D \*\*95\*\*, 029904 \(2017\)](#)

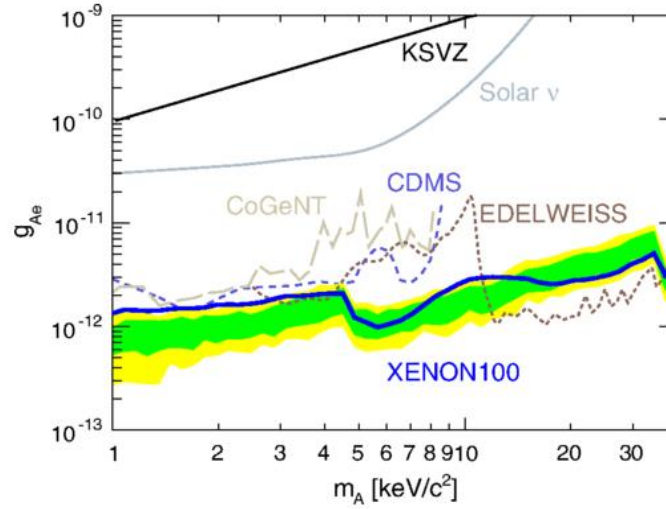
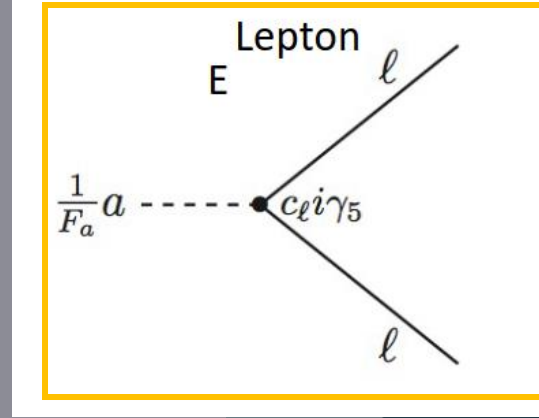


Figure 7

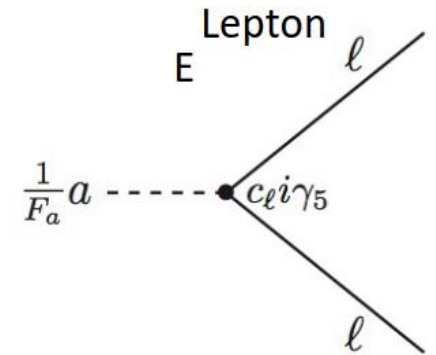
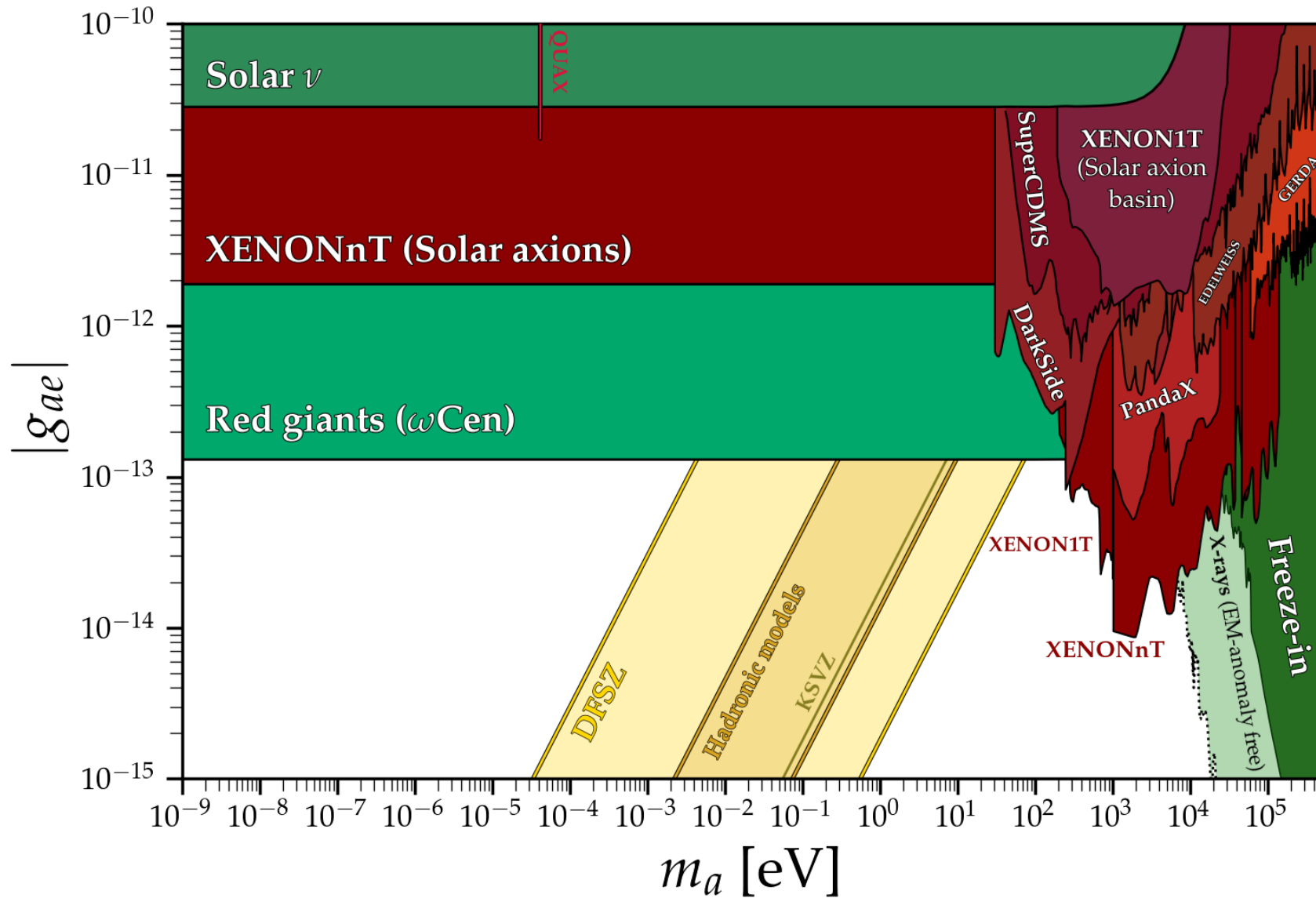
The XENON100 limits (90% C.L.) on ALP coupling to electrons as a function of the mass, under the assumption that ALPs constitute all of the dark matter in our galaxy (blue line). The expected sensitivity is shown by the green/yellow bands  $1\sigma/2\sigma$ . The other curves are constraints set by CoGeNT [39] (light brown dashed line), CDMS [40] (blue dashed line, more dotted), and EDELWEISS-II [31] (ochre dashed line, extending up to 40 keV/c<sup>2</sup>). Indirect astrophysical bound from solar neutrinos [34] is represented as a continuous light grey line. The benchmark KSVZ model is represented by a dark grey line [6, 7].

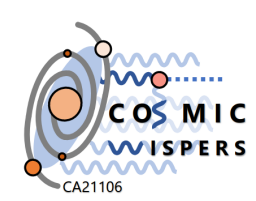
Reuse & Permissions



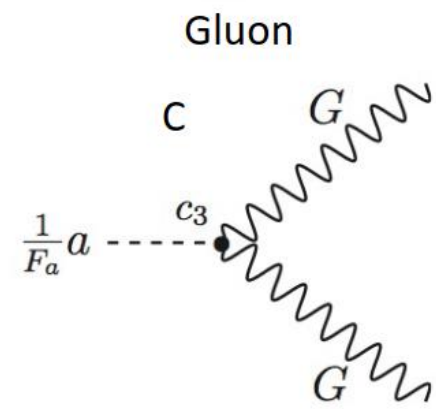
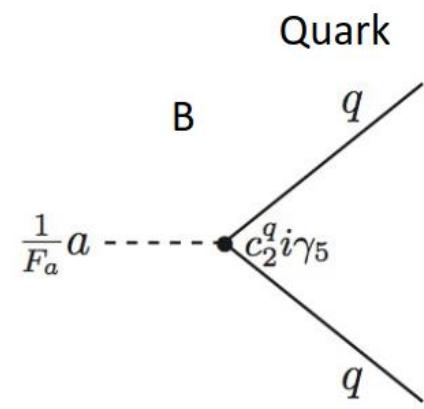
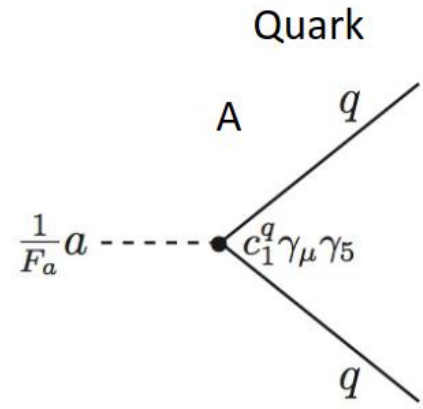
# Results - ae

<https://github.com/cajohare/AxionLimits>

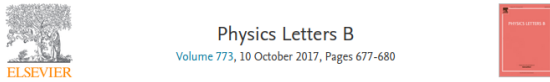




# "Exotic"

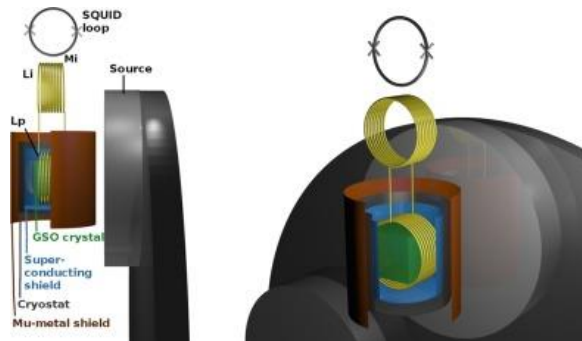


## QUAX



### Improved constraints on monopole-dipole interaction mediated by pseudo-scalar bosons

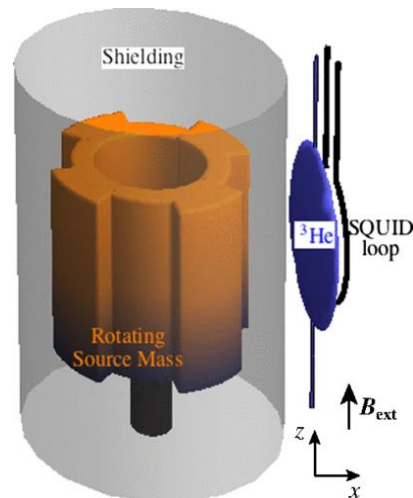
N. Crescini,<sup>a,b</sup> C. Braggio,<sup>c</sup> G. Carugno,<sup>c</sup> P. Falferi,<sup>d</sup> A. Ortolan,<sup>b</sup> G. Ruoso,<sup>b</sup>



## ARIADNE

### Resonantly Detecting Axion-Mediated Forces with Nuclear Magnetic Resonance

Asimina Arvanitaki and Andrew A. Geraci  
Phys. Rev. Lett. **113**, 161801 – Published 14 October 2014



Regular Article - Theoretical Physics | [Open Access](#) | [Published: 10 June 2020](#)

### Dark matter induced Brownian motion

[Ting Cheng](#), [Reinard Primulando](#) & [Martin Spinrath](#) ✉

[The European Physical Journal C](#) **80**, Article number: 519 (2020) | [Cite this article](#)

Letter | [Published: 25 August 2022](#)

### Experiments with levitated force sensor challenge theories of dark energy

[Peiran Yin](#), [Rui Li](#), [Chengjiang Yin](#), [Xiangyu Xu](#), [Xiang Bian](#), [Han Xie](#), [Chang-Kui Duan](#), [Pu Huang](#) ✉, [Jianhua He](#) ✉ & [Jiangfeng Du](#) ✉

[Nature Physics](#) **18**, 1181–1185 (2022) | [Cite this article](#)



### First results on the search for chameleons with the KWISP detector at CAST

[S. Argüedas Cuendis](#)<sup>g</sup>, [J. Baier](#)<sup>g</sup>, [K. Barth](#)<sup>g</sup>, [S. Baum](#)<sup>g</sup>, [A. Baviri](#)<sup>1,1</sup>, [A. Belov](#)<sup>1</sup>, [H. Bräuninger](#)<sup>f</sup>, [G. Cantatore](#)<sup>1,5</sup> ✉, [J.M. Carmona](#)<sup>g</sup>, [J.E. Castel](#)<sup>g</sup>, [S.A. Cetin](#)<sup>1</sup>, [I. Dafni](#)<sup>g</sup>, [M. Davenport](#)<sup>g</sup>, [A. Dermenev](#)<sup>1</sup>, [K. Desch](#)<sup>g</sup>, [B. Döbrich](#)<sup>g</sup>, [H. Fischer](#)<sup>g</sup> ✉, [W. Funk](#)<sup>g</sup>, [J.A. García](#)<sup>g,2</sup>, [A. Gardikiotis](#)<sup>m</sup>, ... [K. Zioutas](#)<sup>m</sup> ✉



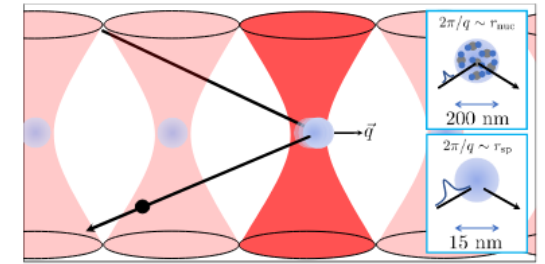
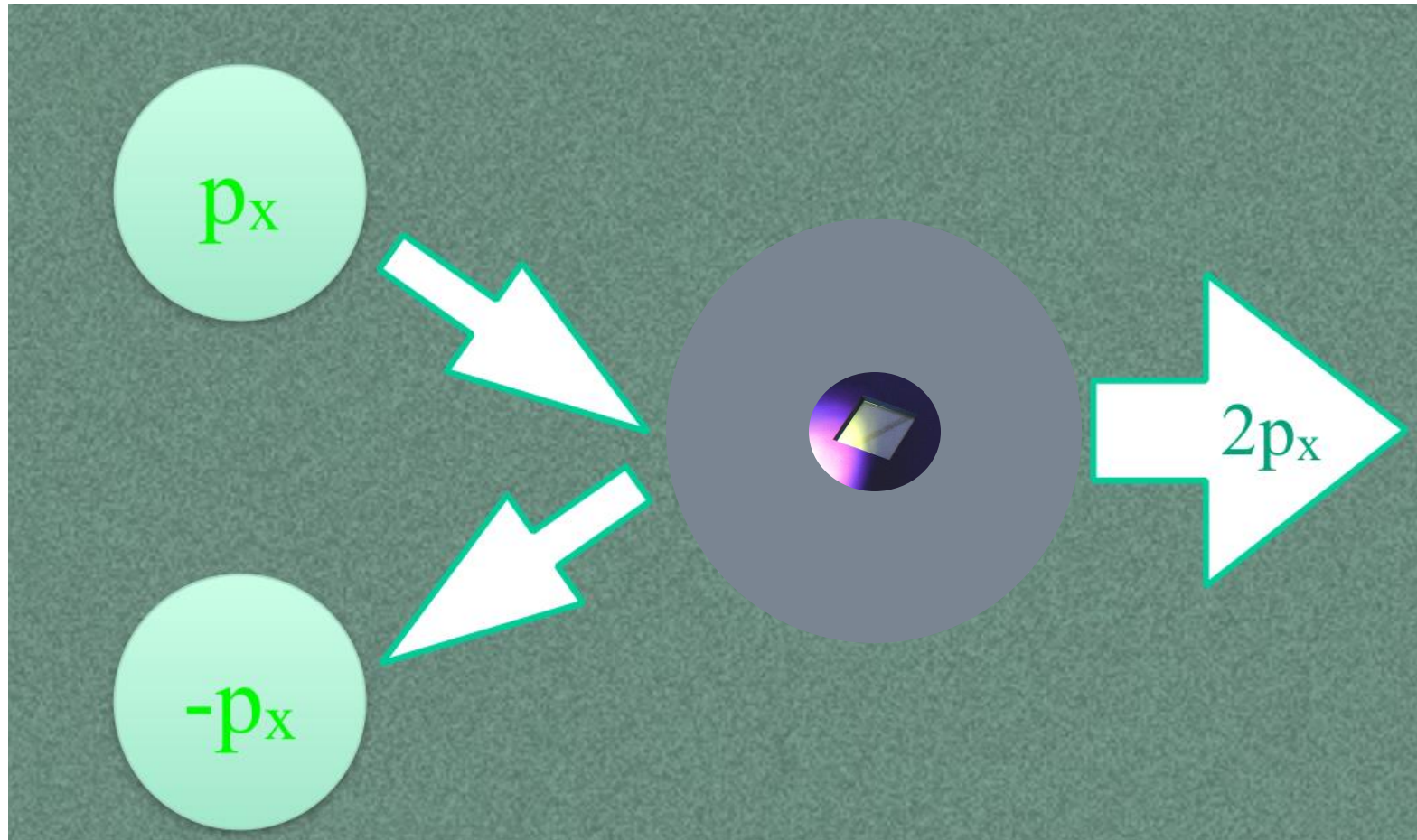
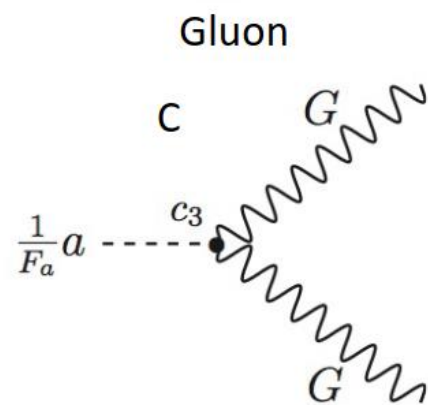
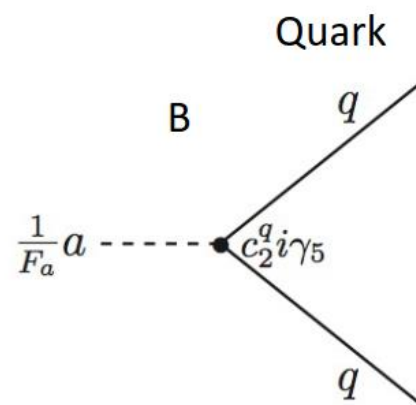
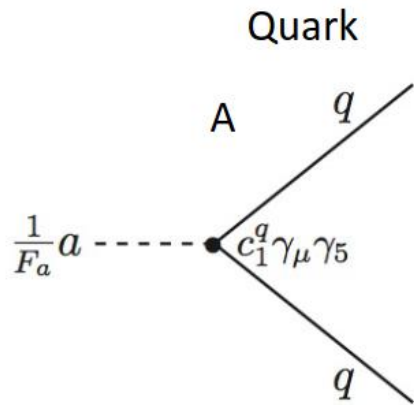


FIG. 1. As a dark matter particle scatters from a levitated optomechanical sensor (possibly part of a large array), it transfers to it momentum  $\vec{q}$ . For “large” sensors (upper inset) the interaction is coherent over a single nucleus. For “small” enough sensors, such that the inverse transferred momentum  $2\pi/q$  of the dark matter particle is comparable to the size of the sensor, leading to a large increase in scattering cross-section.

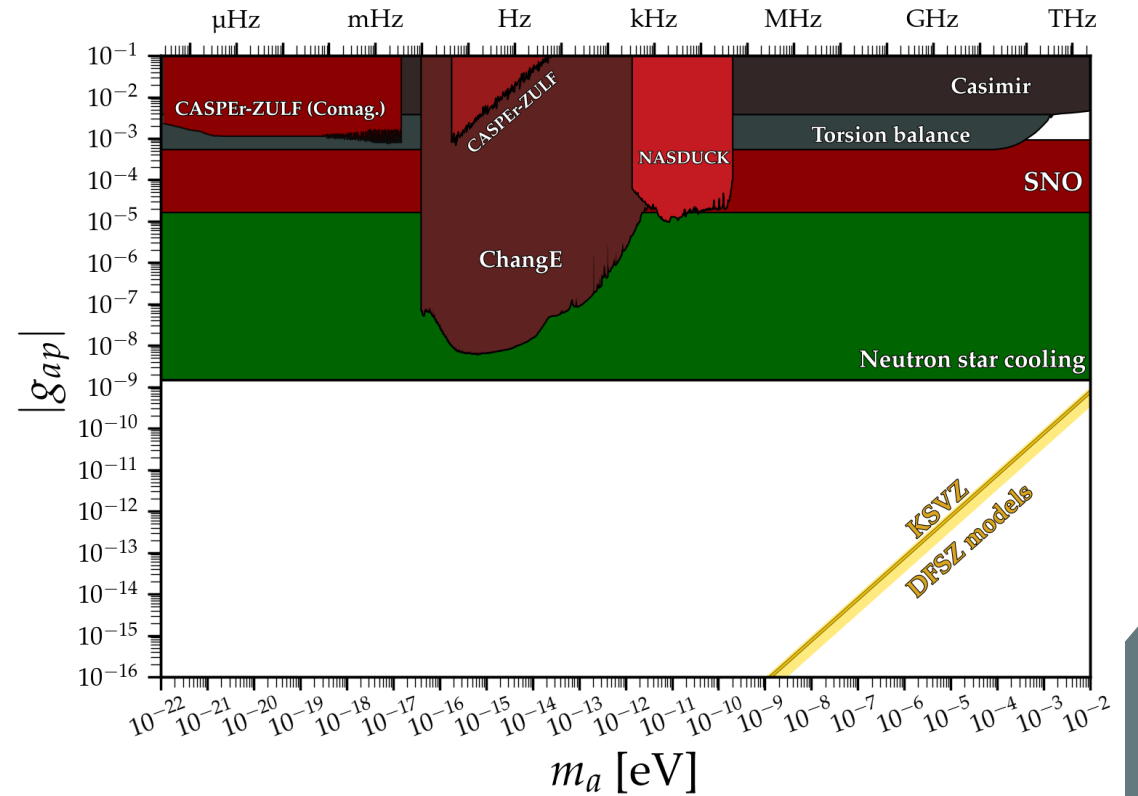
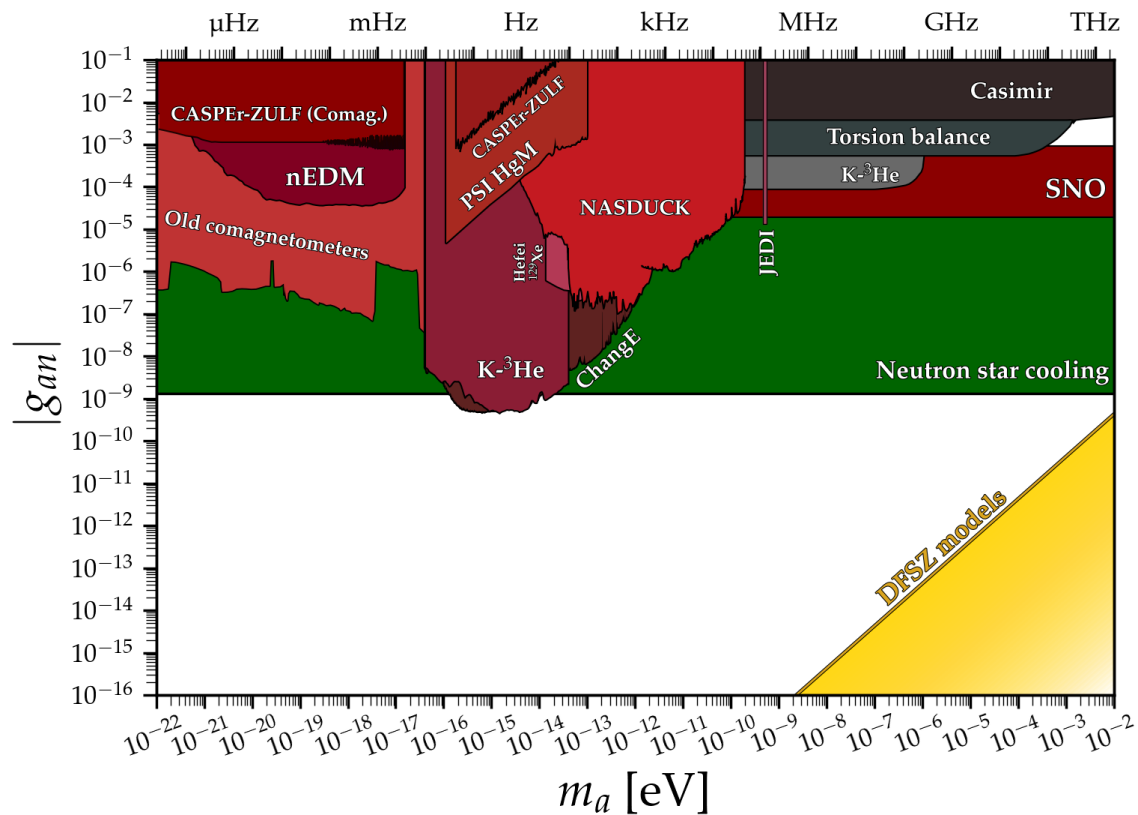


## Develop a radiation pressure detector!

- measure momentum transfer
- high sensitivity



<https://github.com/cajohare/AxionLimits>



# CONCLUSION

- The field is developing rapidly
- Different directions
  - bigger
  - more sensitive, SQL and beyond
- Interesting times ahead



# 18<sup>TH</sup>

# PATRAS WORKSHOP ON AXIONS, WIMPS AND WISPS

03–07 JULY 2023  
UNIVERSITY OF  
RIJEKA, FACULTY  
OF PHYSICS  
RIJEKA, CROATIA

## SCIENTIFIC PROGRAMME

The Physics Case for WIMPs,  
Axions, WISPs /// Direct and  
Indirect Searches for Dark  
Matter and Dark Energy ///  
Direct and Indirect Searches  
for Axions and WISPs ///  
Signals from Astrophysical  
Sources /// Review of  
Collider Experiments ///  
New Theoretical Developments

# THANK YOU

