

Muon Collider Synergies
(Orsay, France)

Ideas for High-Energy
Neutrino Experiments

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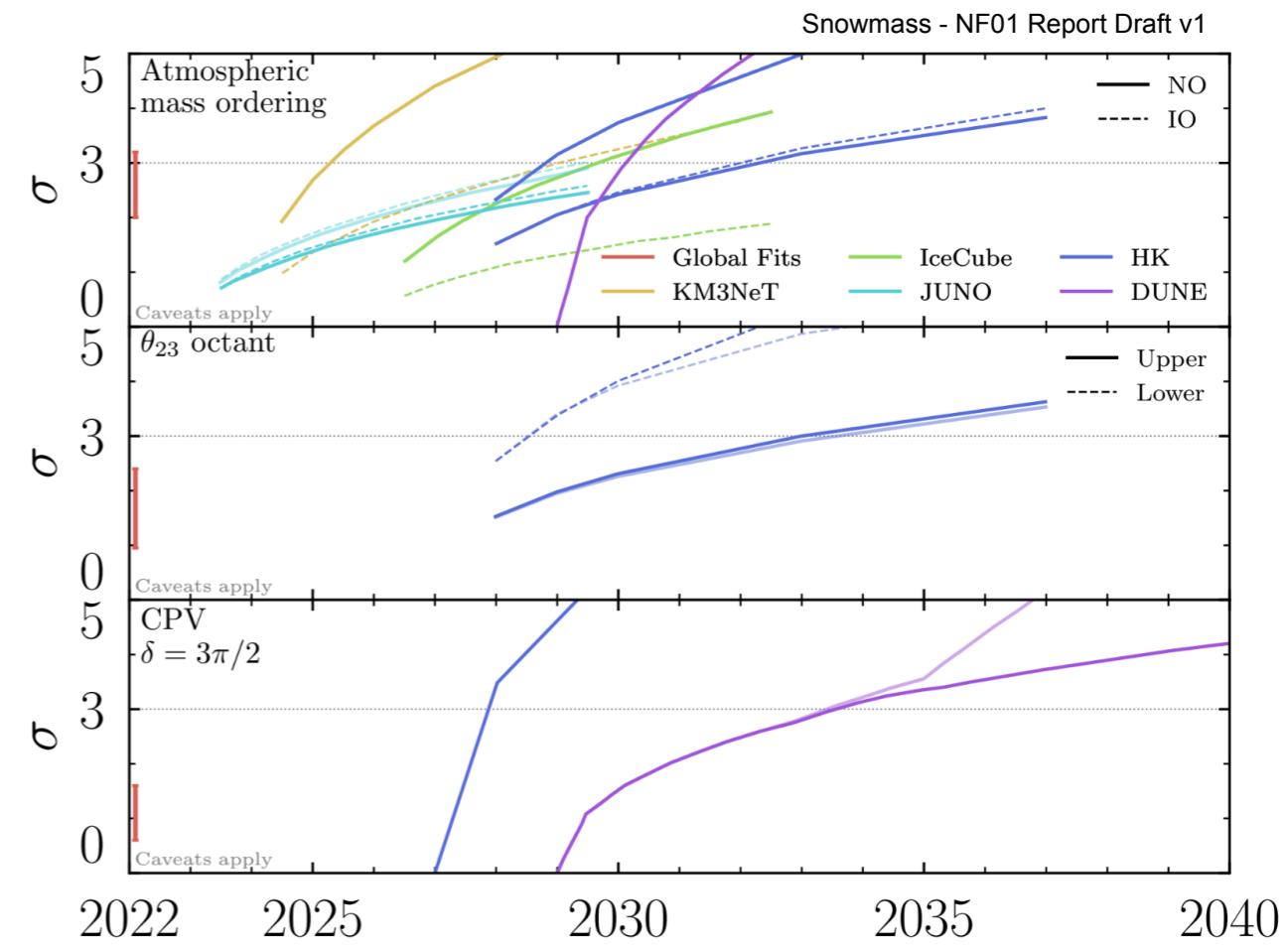
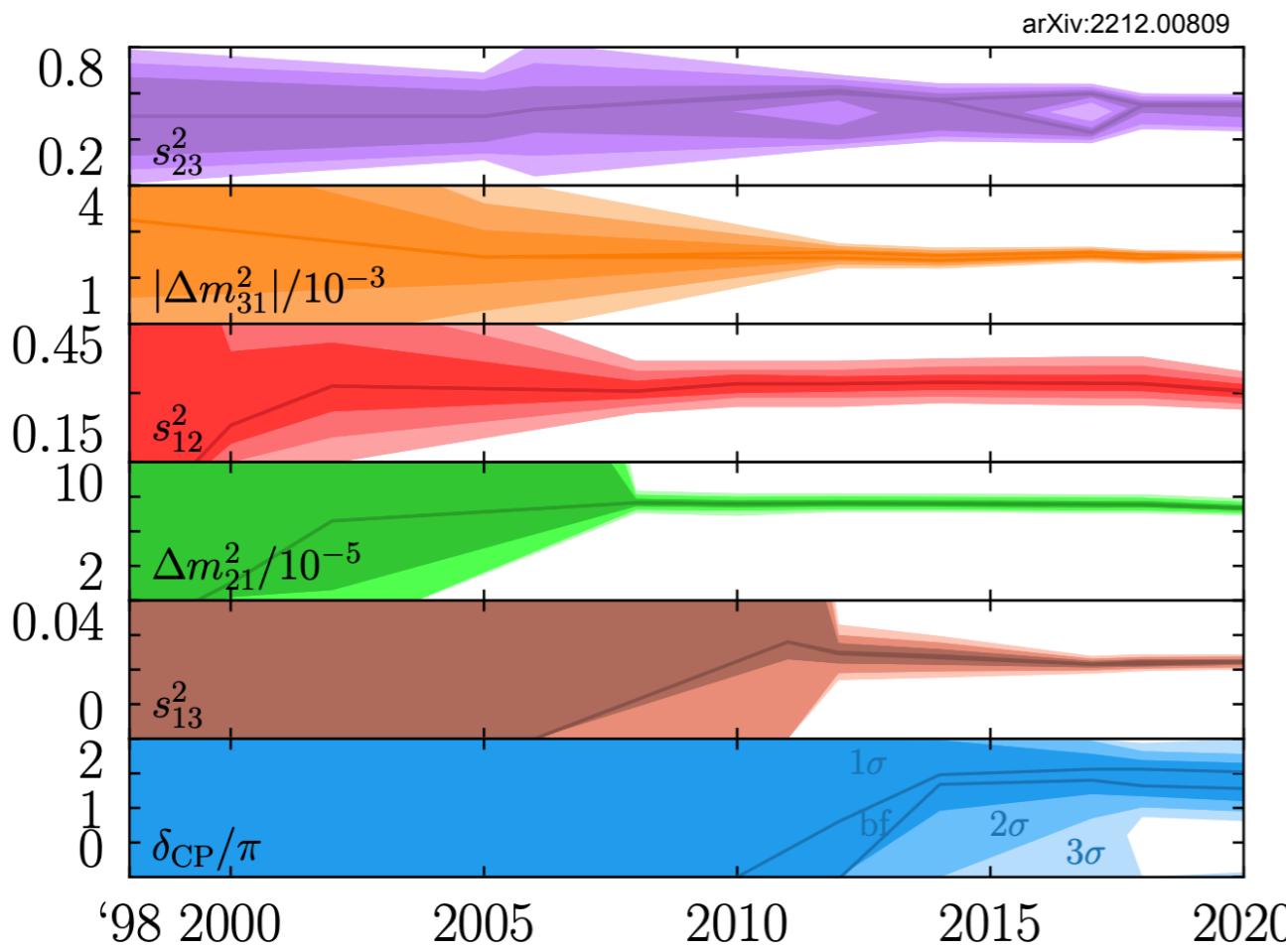


- What will we know about neutrinos in 2045?
- What else can we learn with a powerful neutrino beam -> specifications?



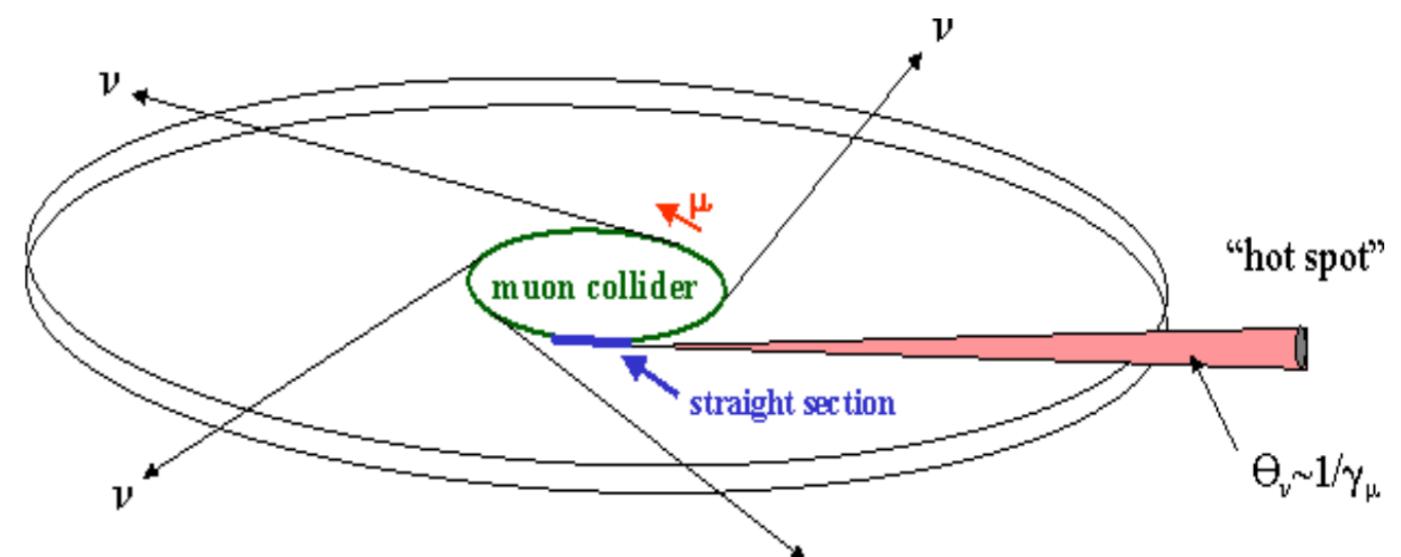
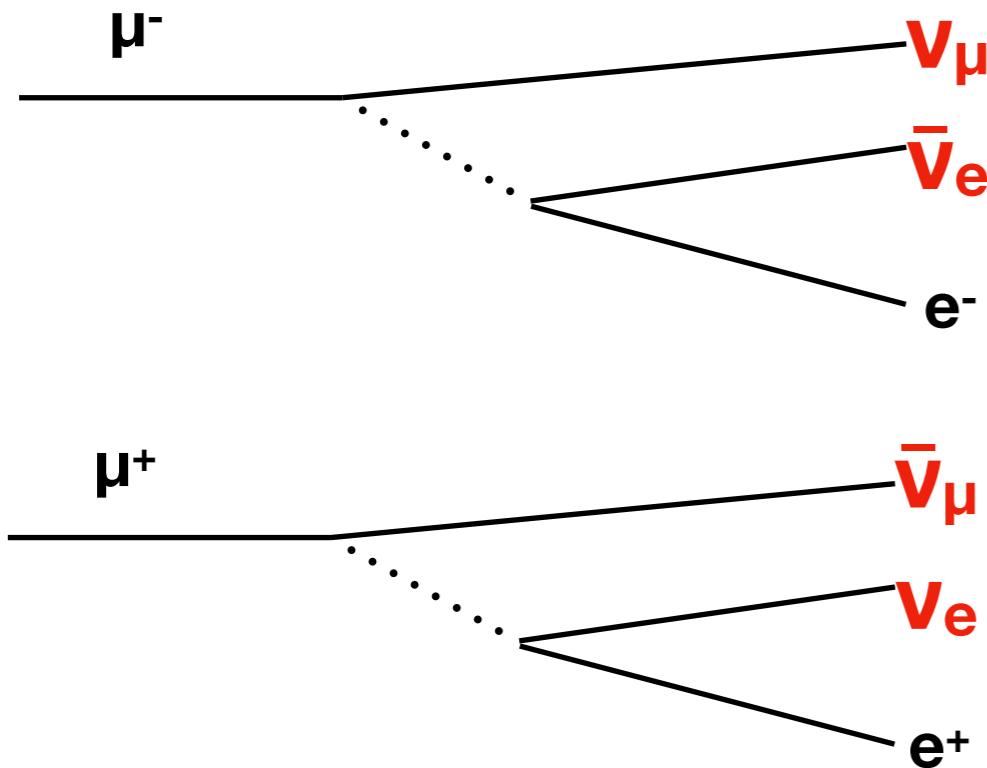
Neutrino oscillations

- Current challenges:
 - CP violation and mass ordering.
- Will they be an interesting topic in 2045?
 - DUNE, HyperK, JUNO, KM3NeT, IC-Upgrade should be there -> precession physics.



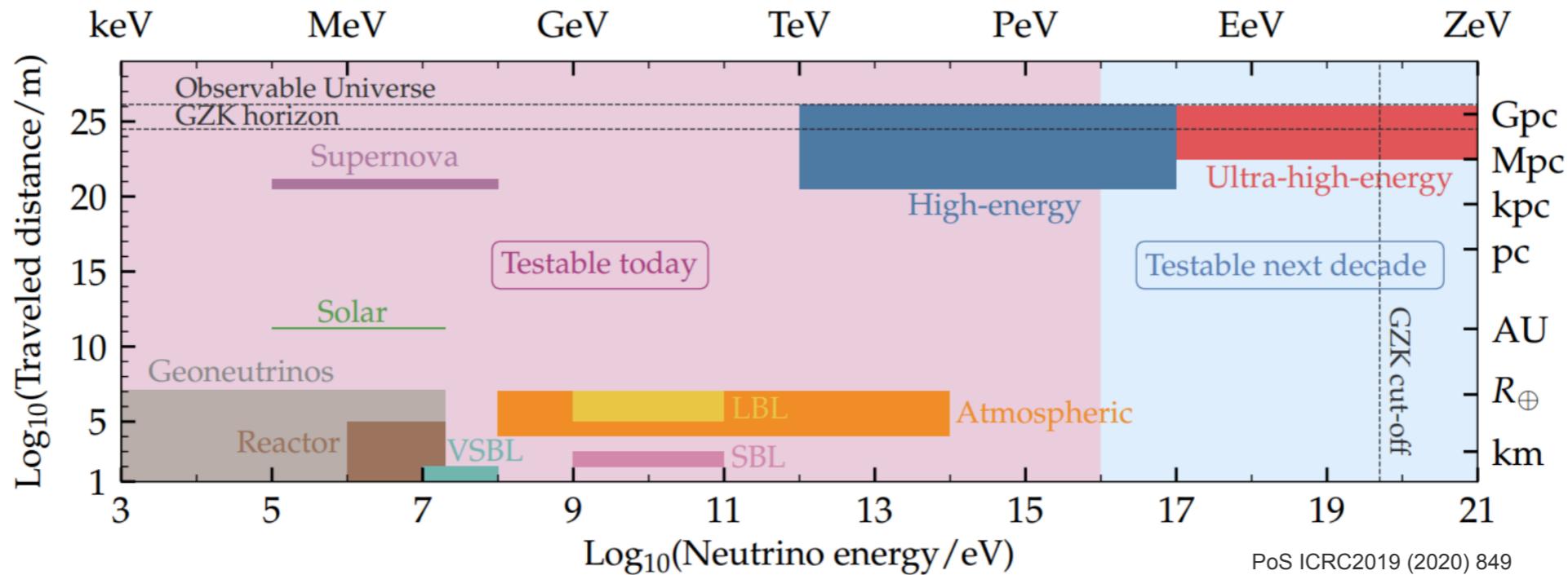
Muon collider

- Intense beam of neutrinos:
 - Known composition: ν_μ and $\bar{\nu}_e$.
 - Energies expanding from GeV to TeV.
 - Can we change polarity (i.e. direct μ^+ instead of μ^- decays)?



Where can Muon Collider make in impact?

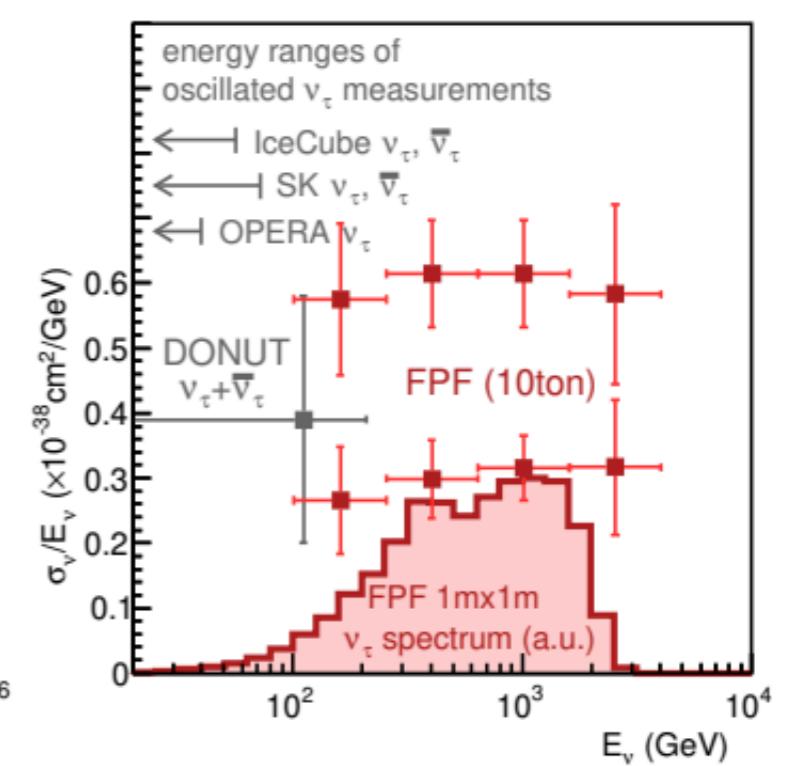
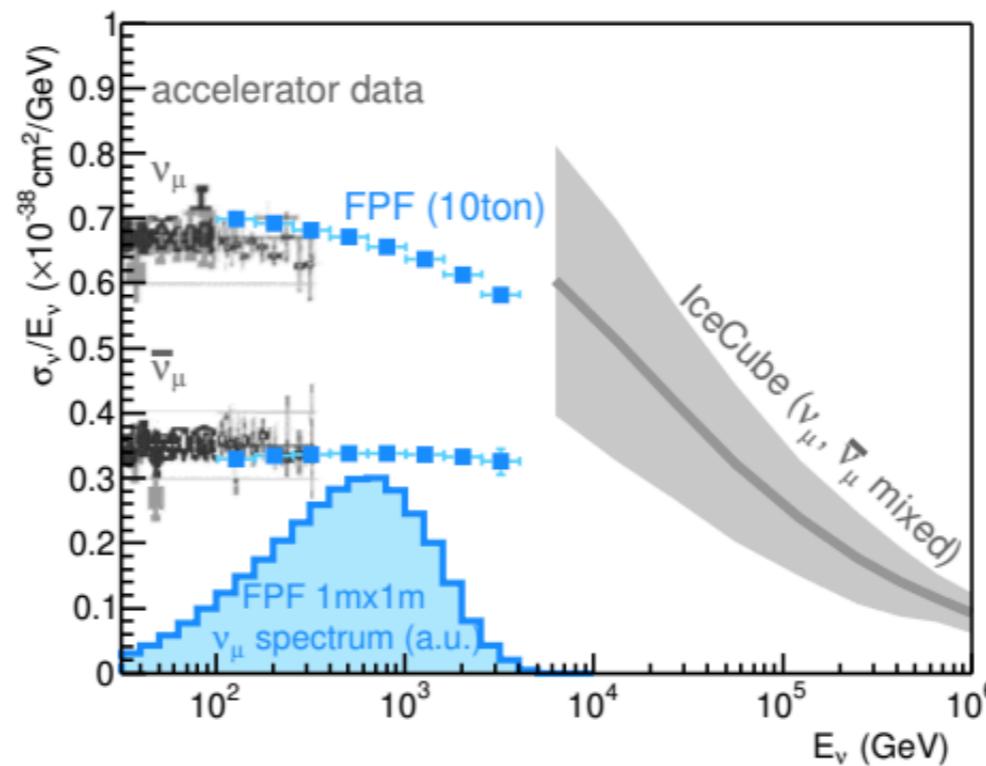
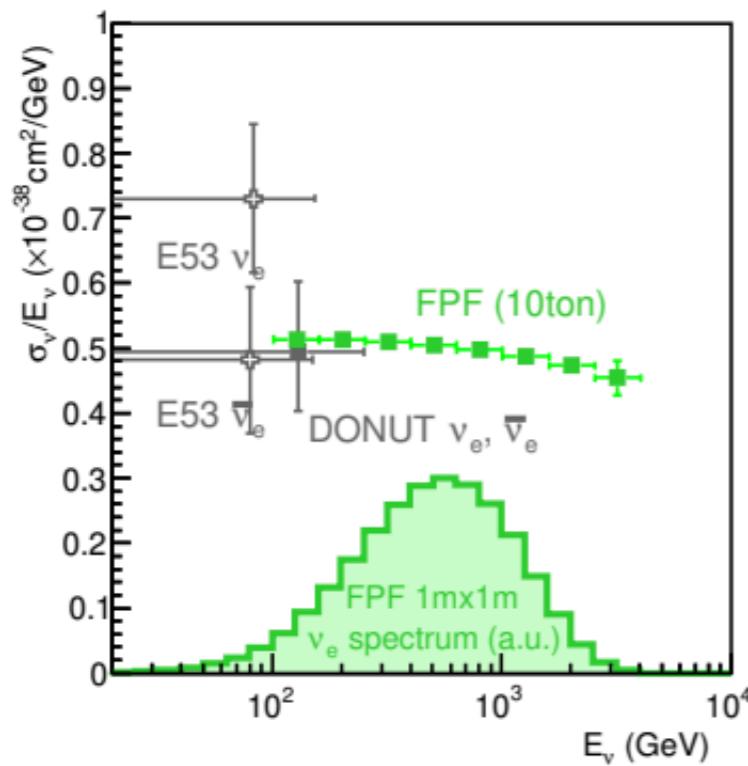
- $E < 10 \text{ GeV}$
 - Already explored with current and next generation of experiments.
- $10 \text{ GeV} < E < 100 \text{ GeV}$
 - Atmospheric neutrinos -> Large uncertainties in flux and poor resolutions!
 - NuTeV/NOMAD -> Focus on no-oscillation regime.
- $E > 100 \text{ GeV}$
 - Atmospheric neutrinos -> Large uncertainties in flux and poor resolutions!
 - FPF -> Uncertainties in flux!



Neutrino cross sections

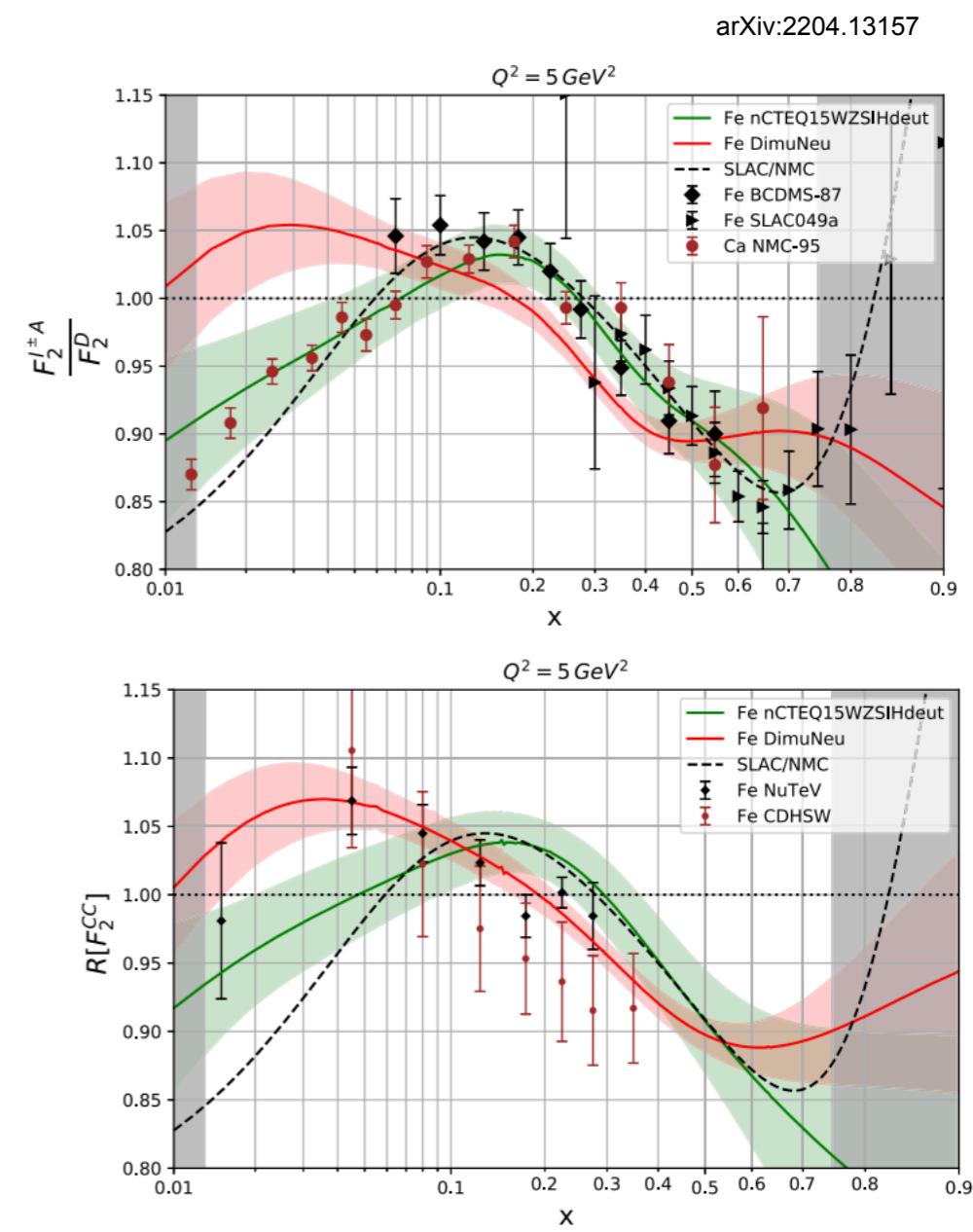
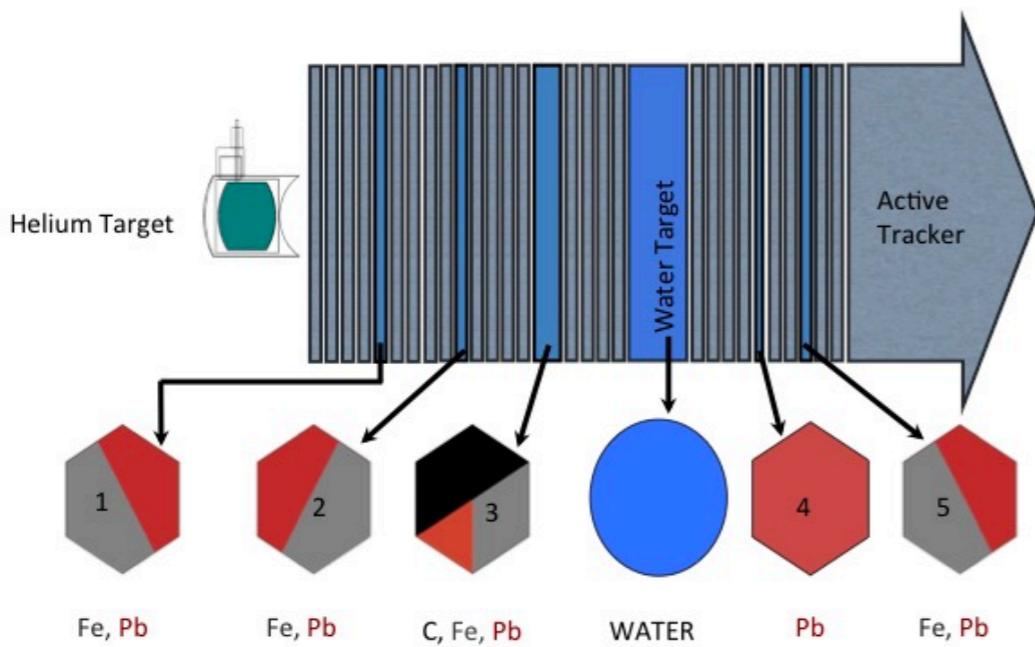
- Follow FPF strategy -> Detector close to the collider.
 - High statistics.
 - Control flux.
 - Precise measurement of neutrino cross section.

arXiv:2203.05090



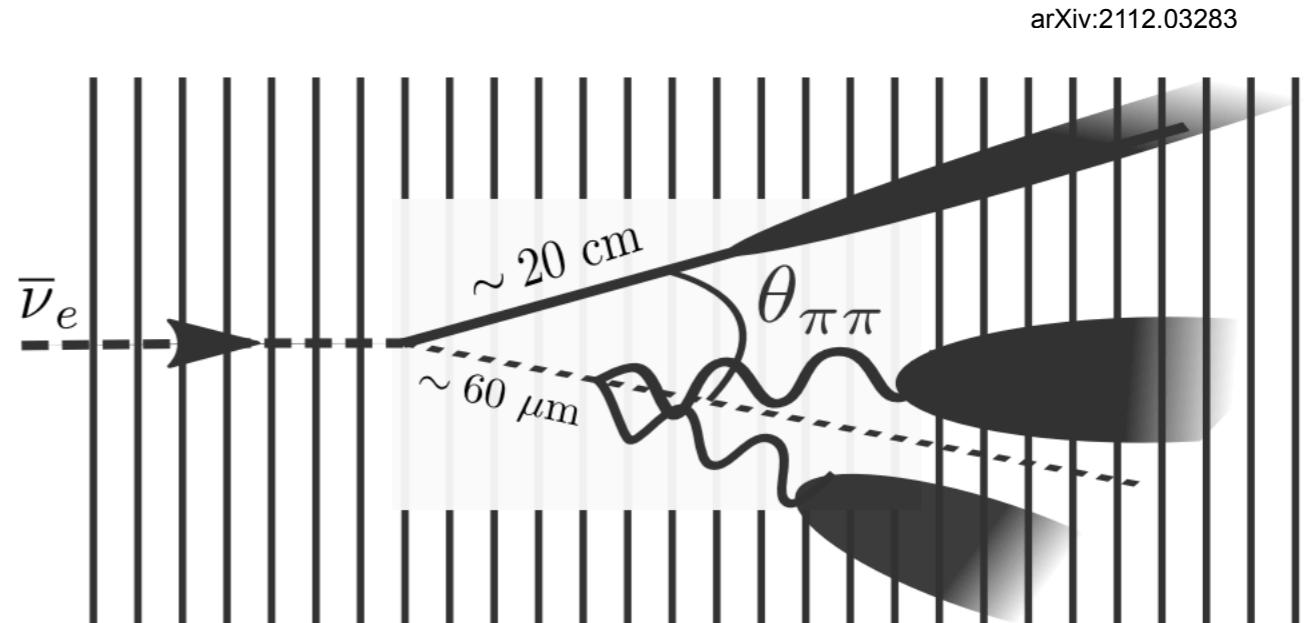
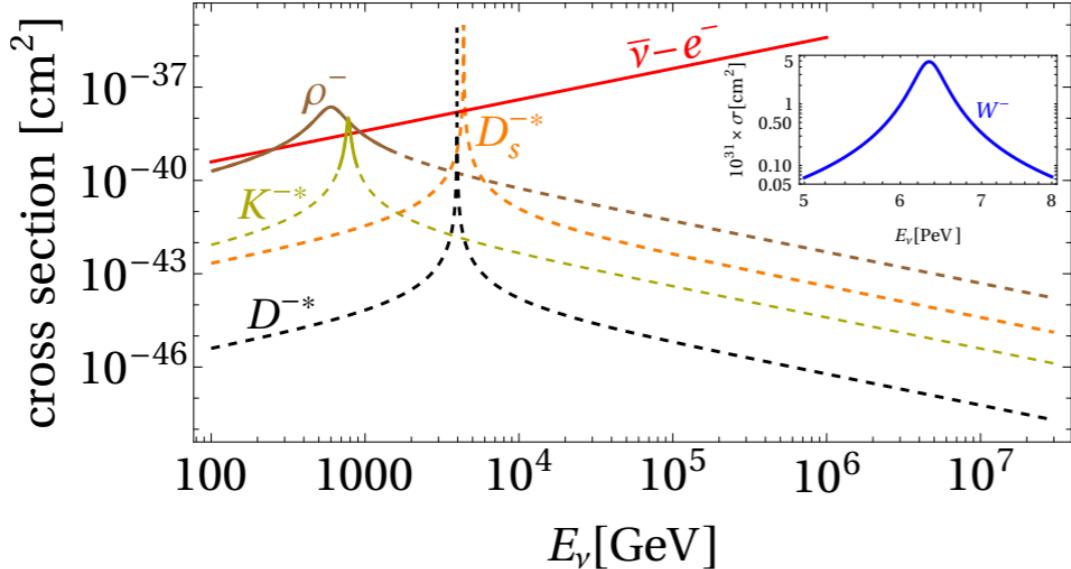
Neutrino cross sections

- Multi target experiment (a la MINERvA)?
 - Might allow us to study nuclear effects.
 - EMC, shadowing, etc.
 - We need H/He target!
- Dimuons -> strange composition of proton



Neutrino cross sections

- Meson resonance from antineutrino-electron scattering
 - $O(100)$ interactions in FPF experiments.
 - Background from DIS and through-going muons.
 - LArTPCs and emulsion detector are the most promising technologies.



arXiv:2112.03283

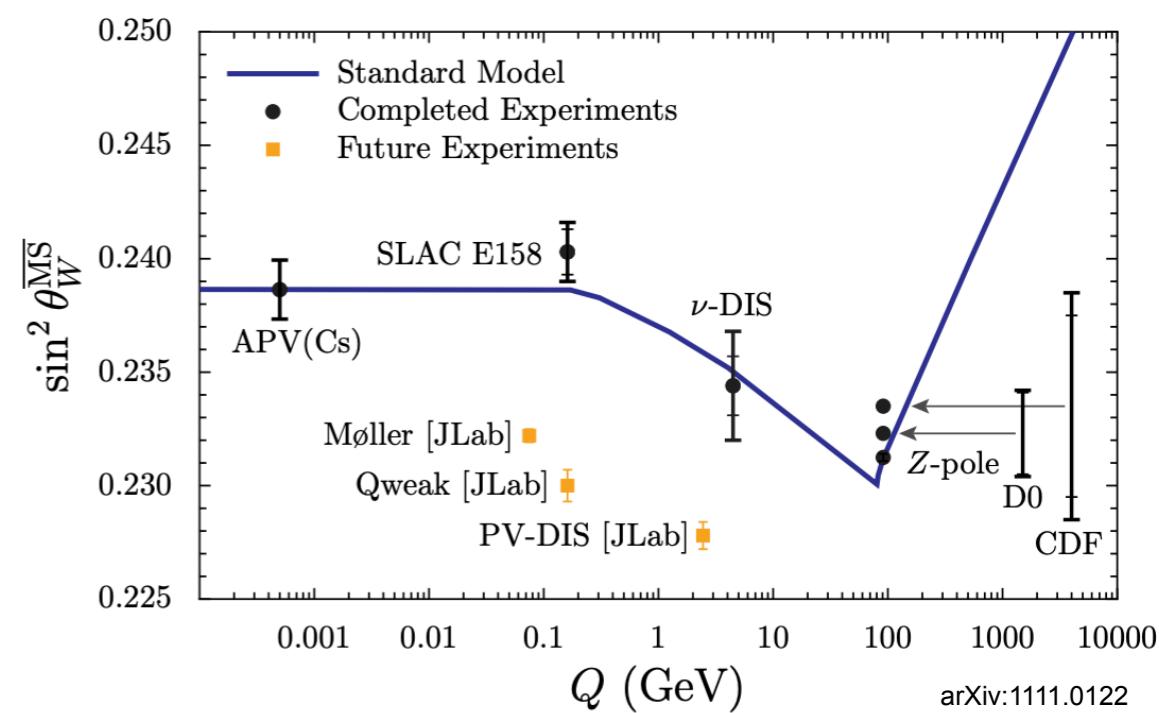
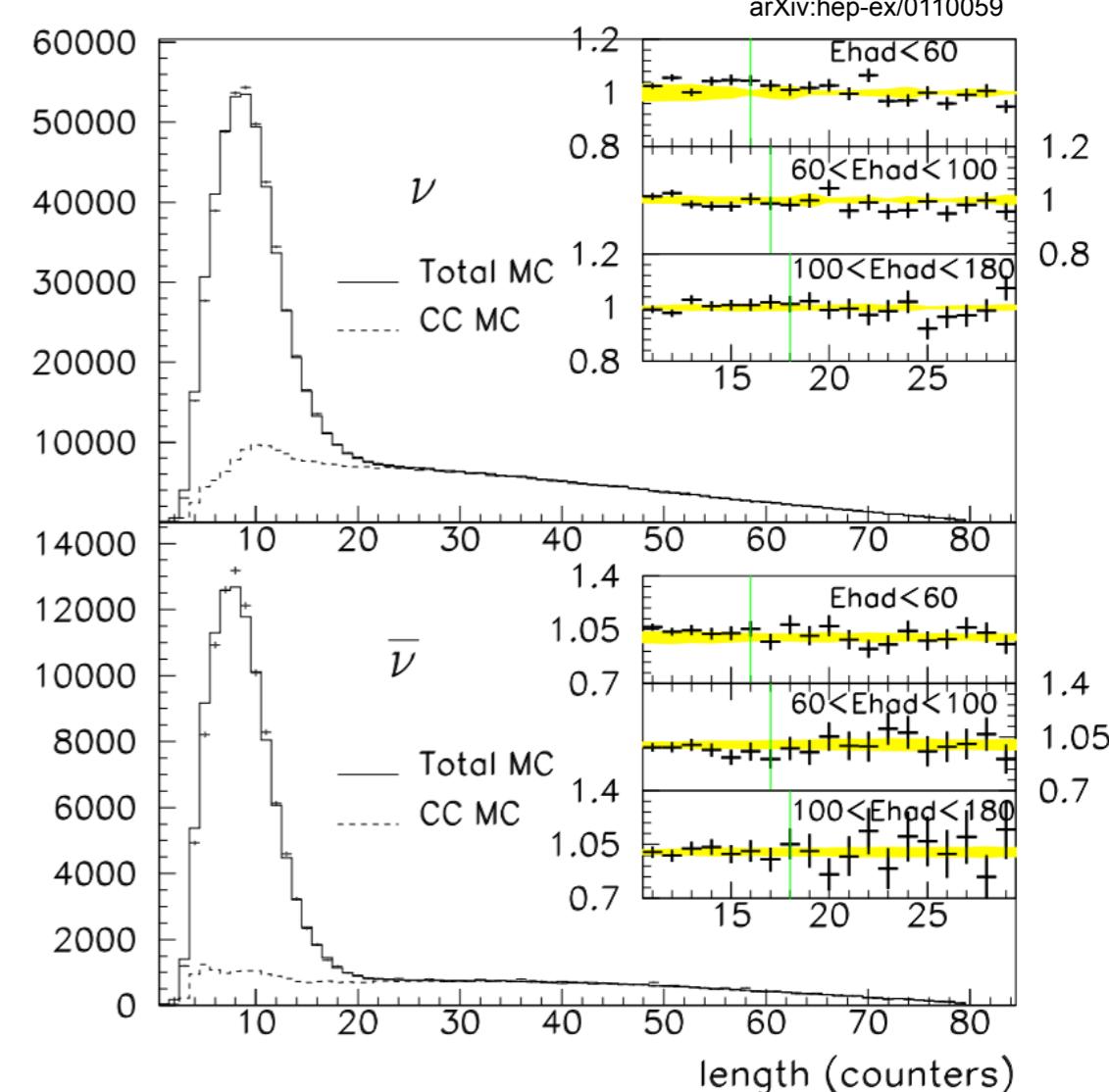
Test EW

- NuTeV "anomaly":

- Pure beam of muon (anti)neutrinos (<2% wrong sign+flavor).
- Paschos-Wolfenstein (PW) relation:

$$R = \frac{\sigma_{NC}^{\nu A} - \sigma_{NC}^{\bar{\nu} A}}{\sigma_{CC}^{\nu A} - \sigma_{CC}^{\bar{\nu} A}} \approx 1/2 - \sin^2 \theta_W$$

- 3 sigma tension with SM
- Overall consensus nucleon charge symmetry violating effects, strange sea quarks, and nuclear corrections.



Test EW

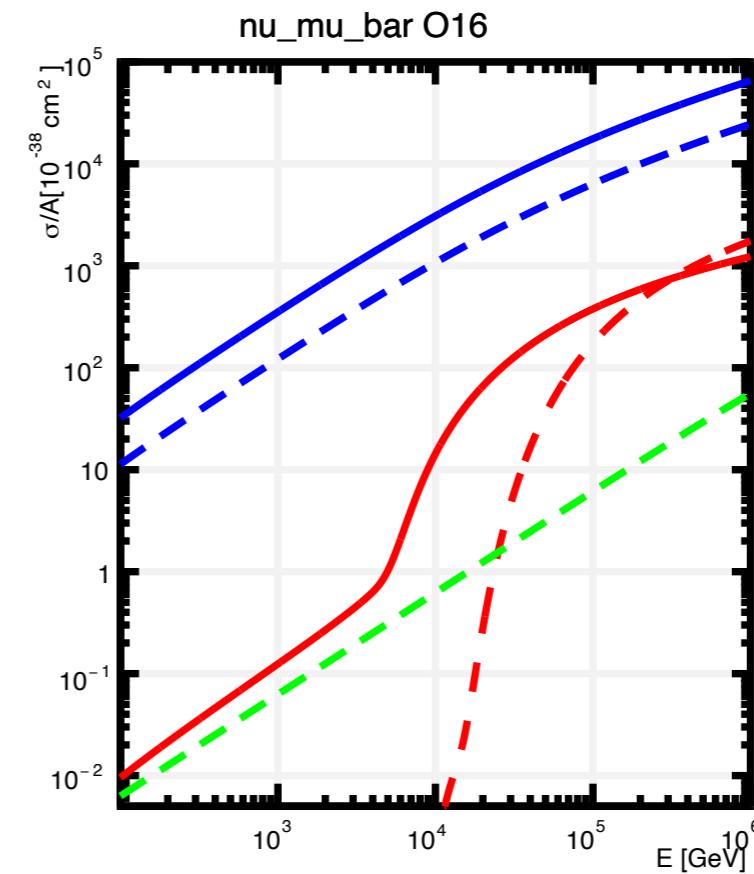
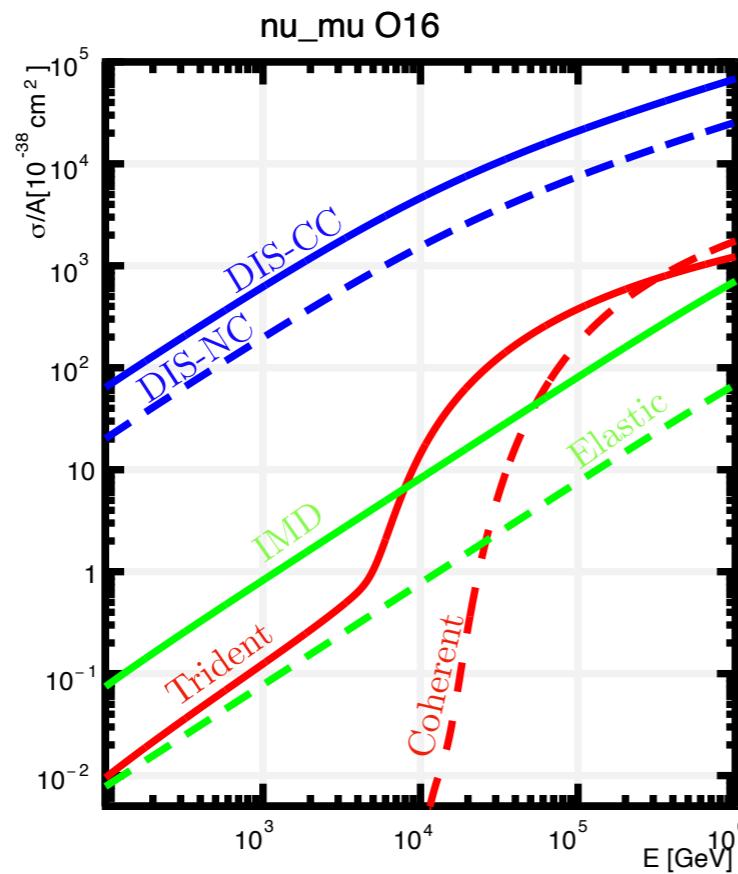
CAVEAT: Muon collider has large electron-flavor content unlike the NuTeV/NuSong style beams.
But it is a very well defined electron-flavor content

- NuSONG: arXiv:0907.4864

- Looking also at electron scattering from pure ν_μ and $\bar{\nu}_\mu$ beams.
- IMD and low-nu used to constrain the flux.

$$\frac{N(\nu_\mu e^- \rightarrow \nu_\mu e^-)}{N(\nu_\mu e^- \rightarrow \mu^- \nu_e)} = \frac{\sigma_{NC}^{\nu e} \times \Phi^\nu}{\sigma^{IMD} \times \Phi^\nu}.$$

$$\frac{N_{\nu DIS}^{low E_{had}}}{N_{\bar{\nu} DIS}^{low E_{had}}} \times \frac{N(\bar{\nu}_\mu e^- \rightarrow \bar{\nu}_\mu e^-)}{N(\nu_\mu e^- \rightarrow \mu^- \nu_e)} = \frac{\Phi^\nu}{\Phi^{\bar{\nu}}} \times \frac{\sigma_{NC}^{\bar{\nu} e} \times \Phi^{\bar{\nu}}}{\sigma^{IMD} \times \Phi^\nu}.$$



Test EW

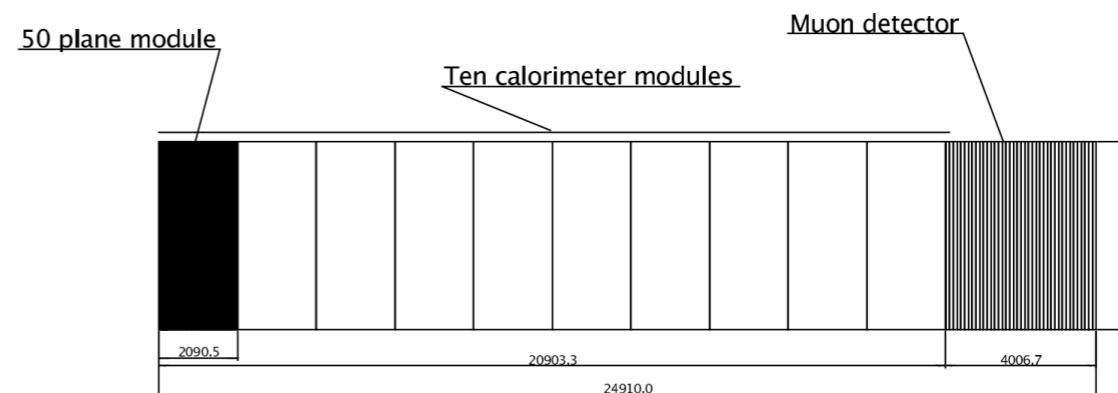


- **NuSOnG:** arXiv:0907.4864

- Four calorimeters+muon spectrometer separated by 15m.
- Each calorimeter has 500 layers of glass and active detectors.

- **Why glass?**

- Long radiation length -> PID
- Short detector -> calibration
- Isoscalar.



Neutrino mode

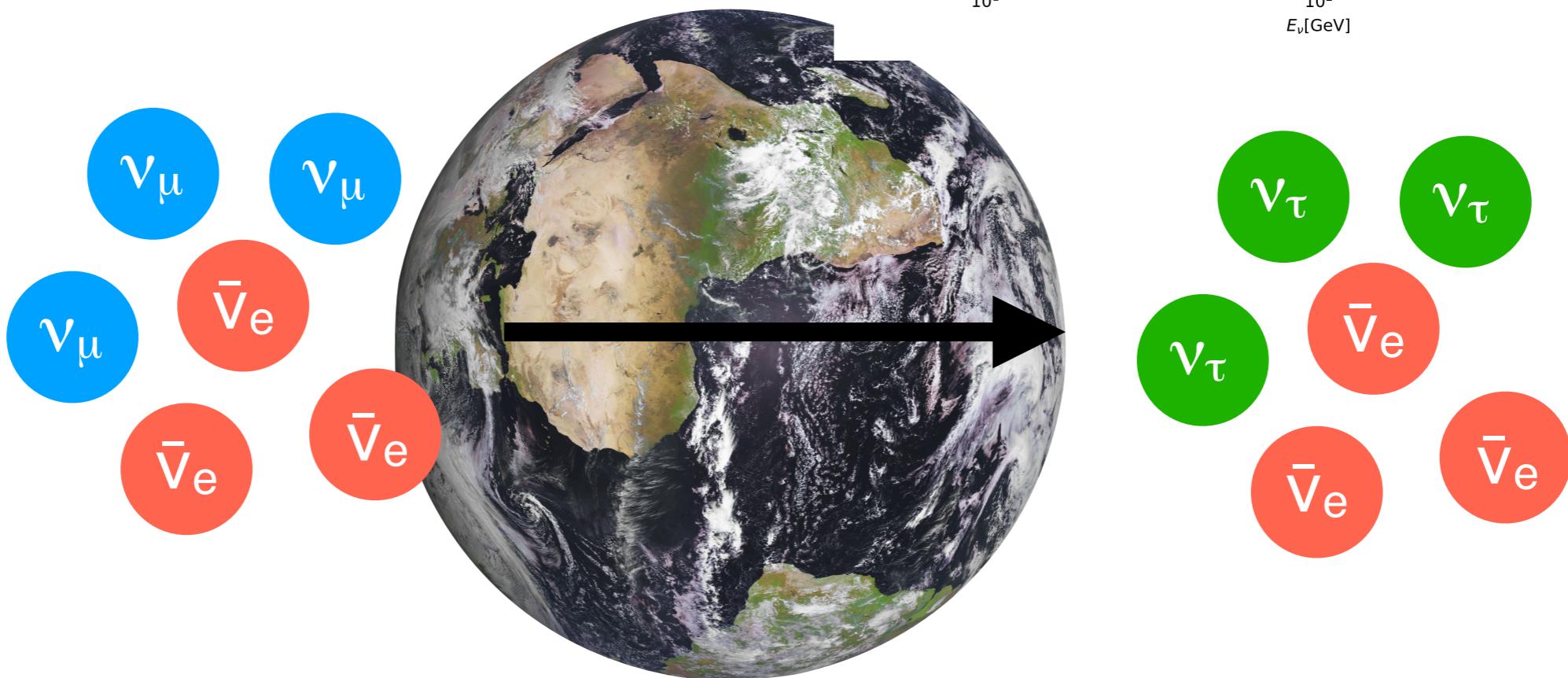
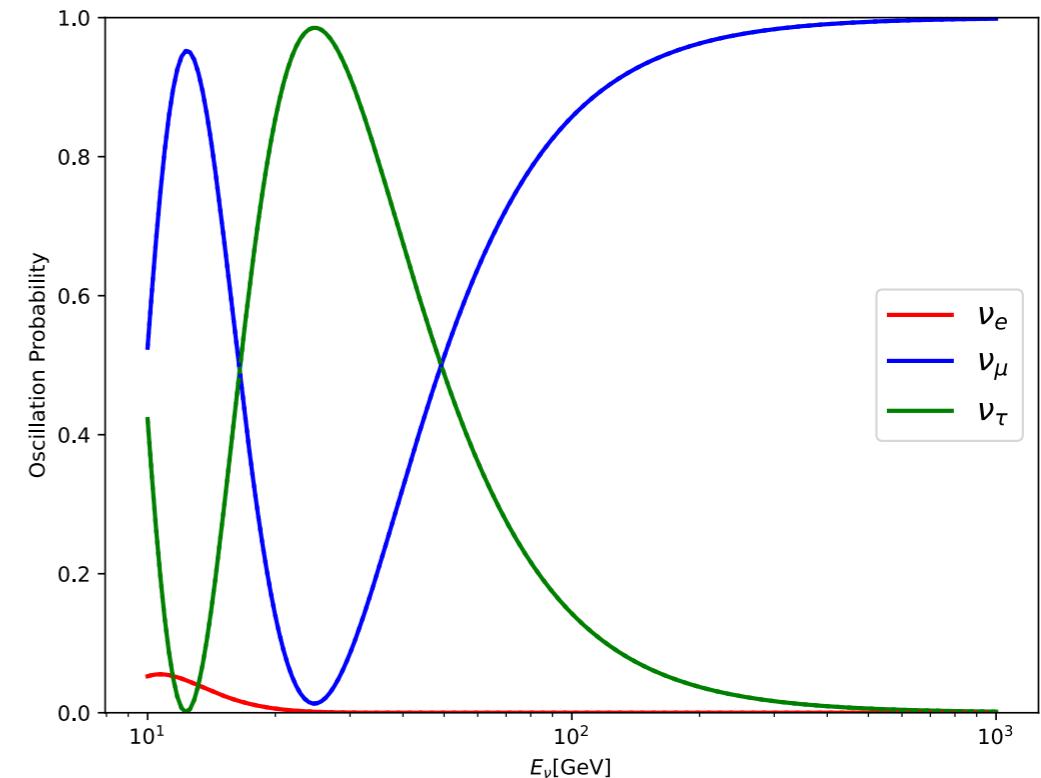
507k	ν_μ CC quasi – elastic scatters
178k	ν_μ NC – elastic scatters
1016k	ν_μ CC π^+
302k	ν_μ CC π^0
272k	ν_μ NC π^0
226k	ν_μ NC π^\pm
1379k	ν_μ CC and NC Resonance multi – pion
202M	ν_μ CC Deep Inelastic Scattering
63M	ν_μ NC Deep Inelastic Scattering
24k	ν_μ neutrino – electron NC elastic scatters
235k	ν_μ neutrino – electron CC quasielastic scatters (IMD)

Antineutrino mode

548k	$\bar{\nu}_\mu$ CC quasi – elastic scatters
195k	$\bar{\nu}_\mu$ NC – elastic scatters
1103k	$\bar{\nu}_\mu$ CC π^+
321k	$\bar{\nu}_\mu$ CC π^0
297k	$\bar{\nu}_\mu$ NC π^0
246k	$\bar{\nu}_\mu$ NC π^\pm
1516k	$\bar{\nu}_\mu$ CC and NC Resonance multi – pion
102M	$\bar{\nu}_\mu$ CC Deep Inelastic Scattering
36M	$\bar{\nu}_\mu$ NC Deep Inelastic Scattering
21k	$\bar{\nu}_\mu$ neutrino – electron NC elastic scatters
0k	$\bar{\nu}_\mu$ neutrino – electron CC quasielastic scatters (IMD)

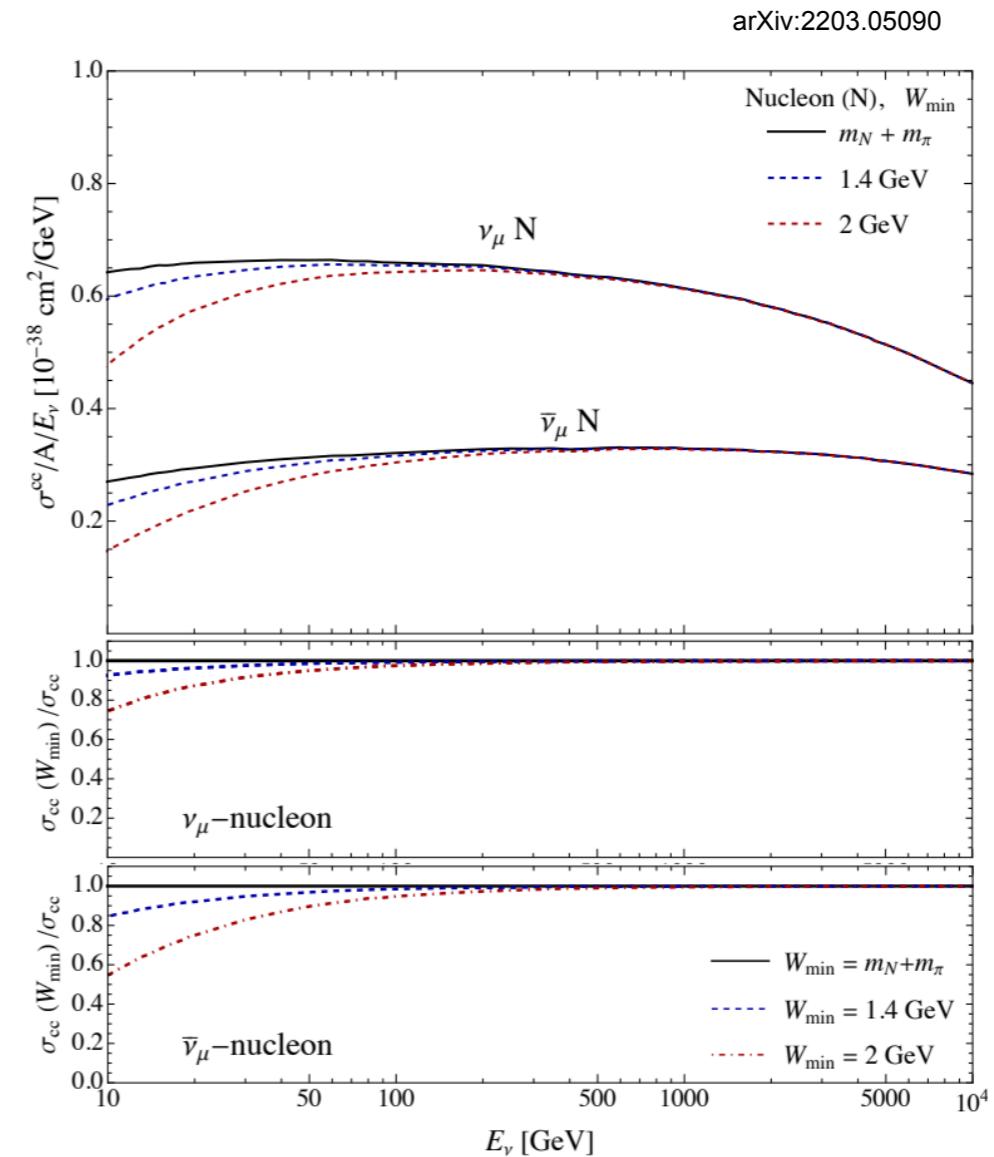
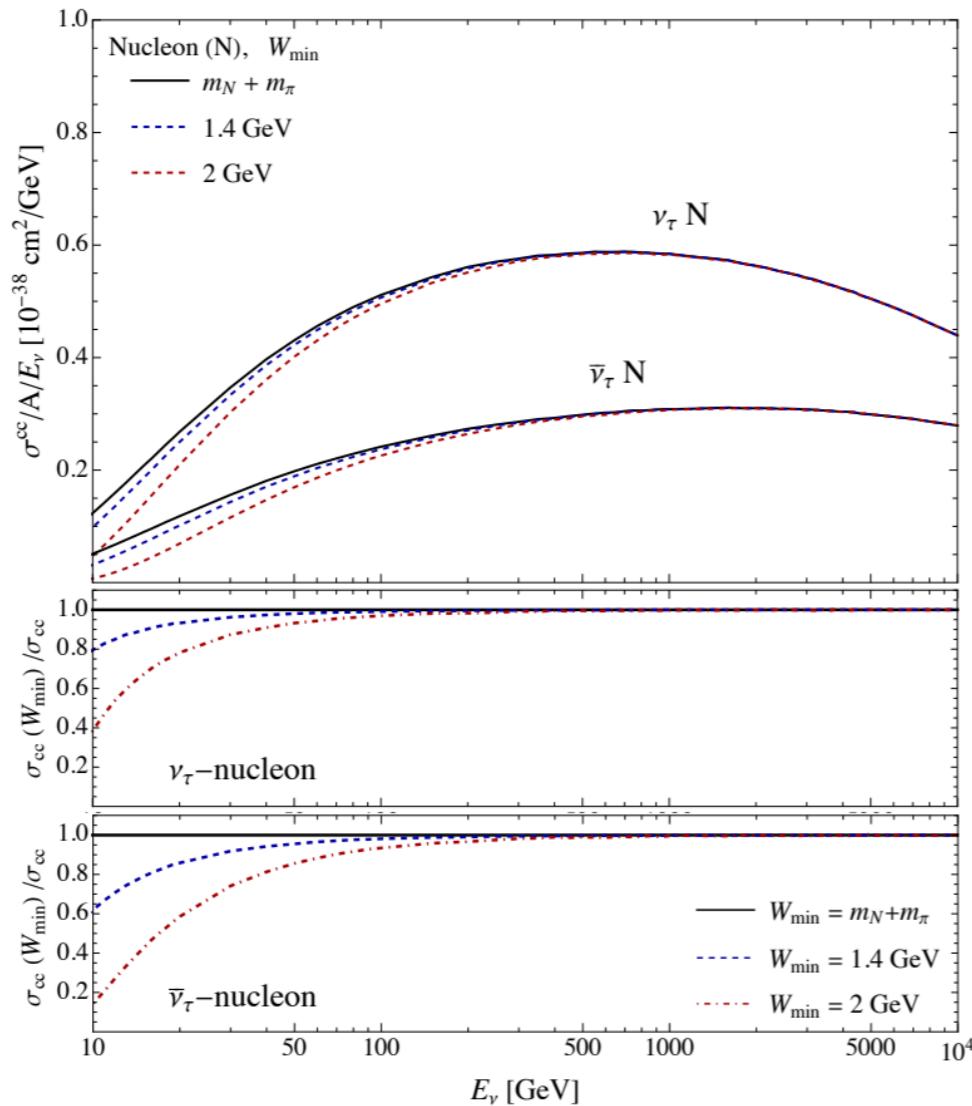
Tau appearance

- Earth as baseline ($L \sim 12000\text{km}$).
 - Large ν_τ appearance at 20-100 GeV.
 - $P_{ee} = 1$.



Tau appearance

- Enough energy to interact via charged current.
 - Factor $\sim 0.4(0.75)$ with respect to ν_μ at 20(100) GeV.

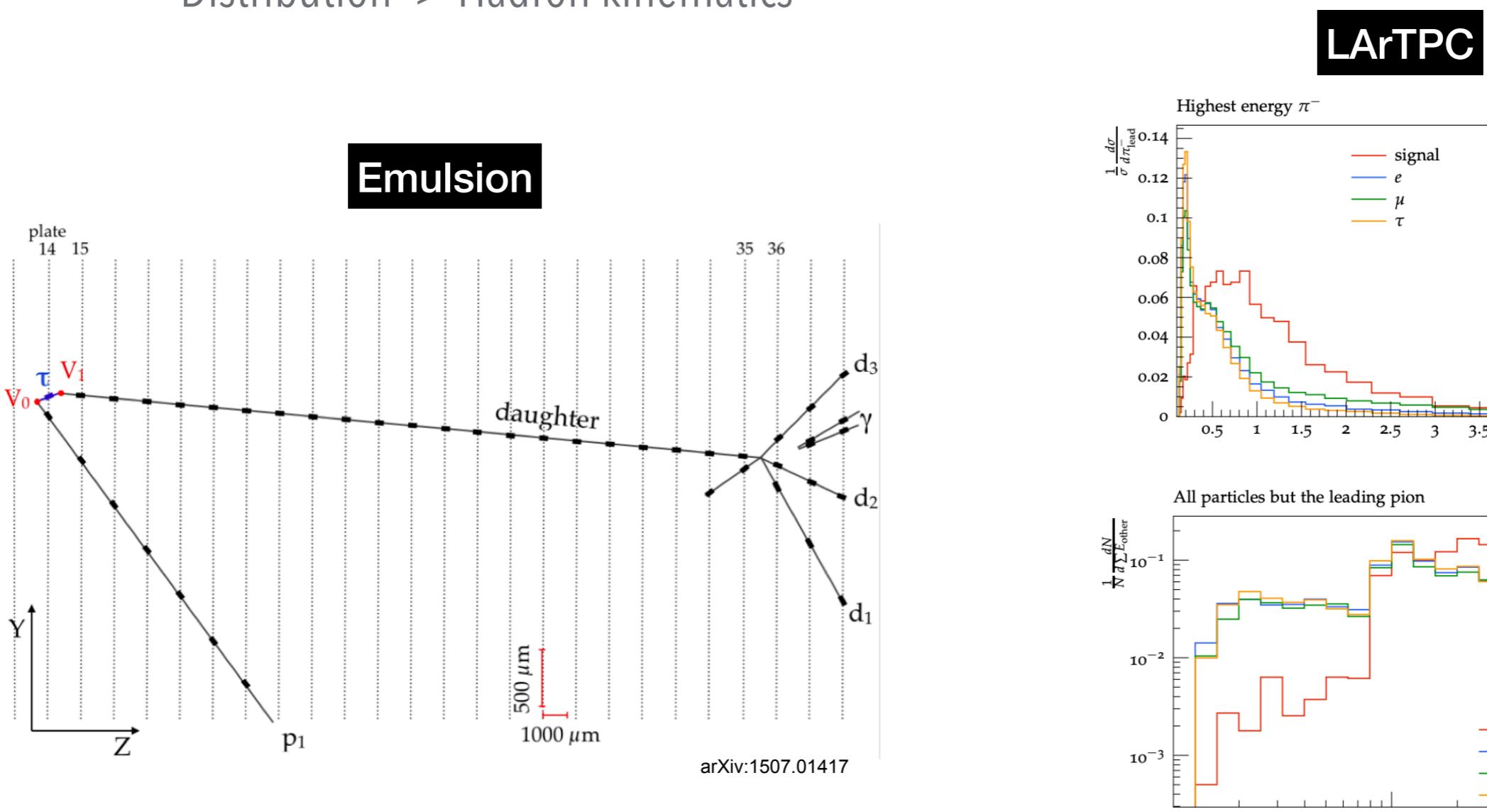


arXiv:2203.05090

Tau appearance

- Identify tau neutrino interactions

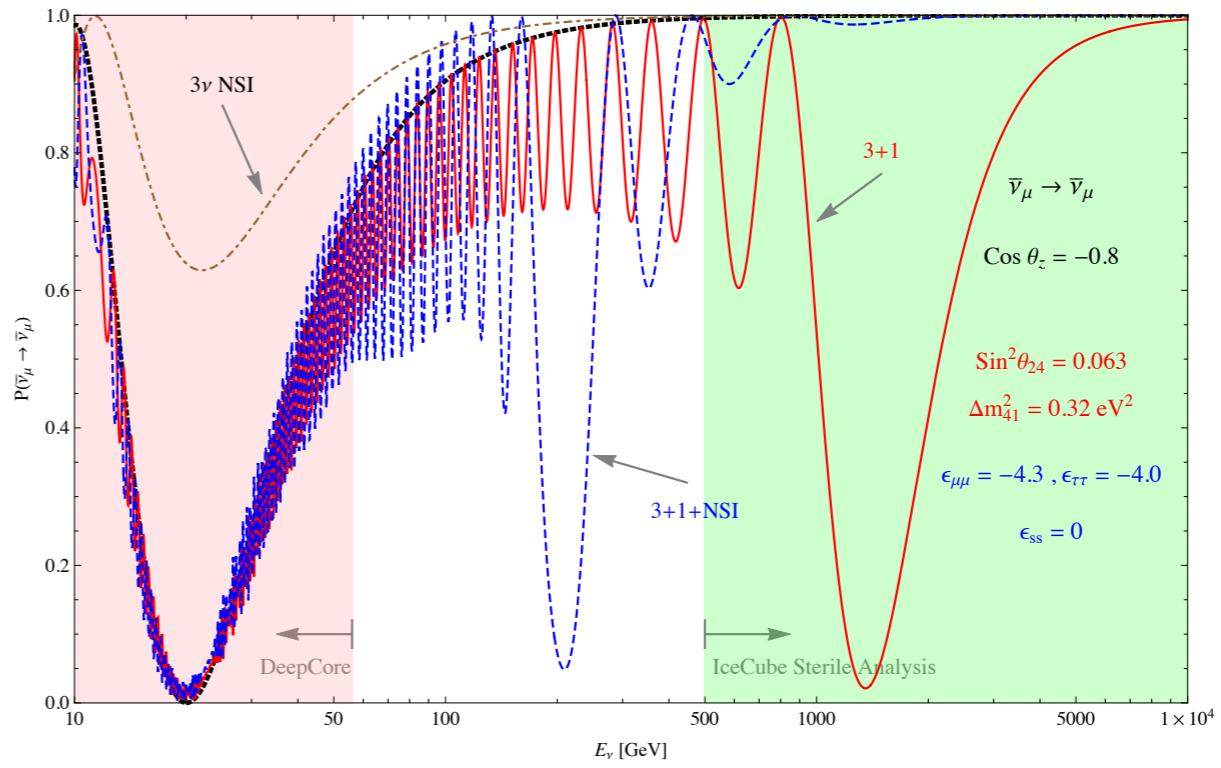
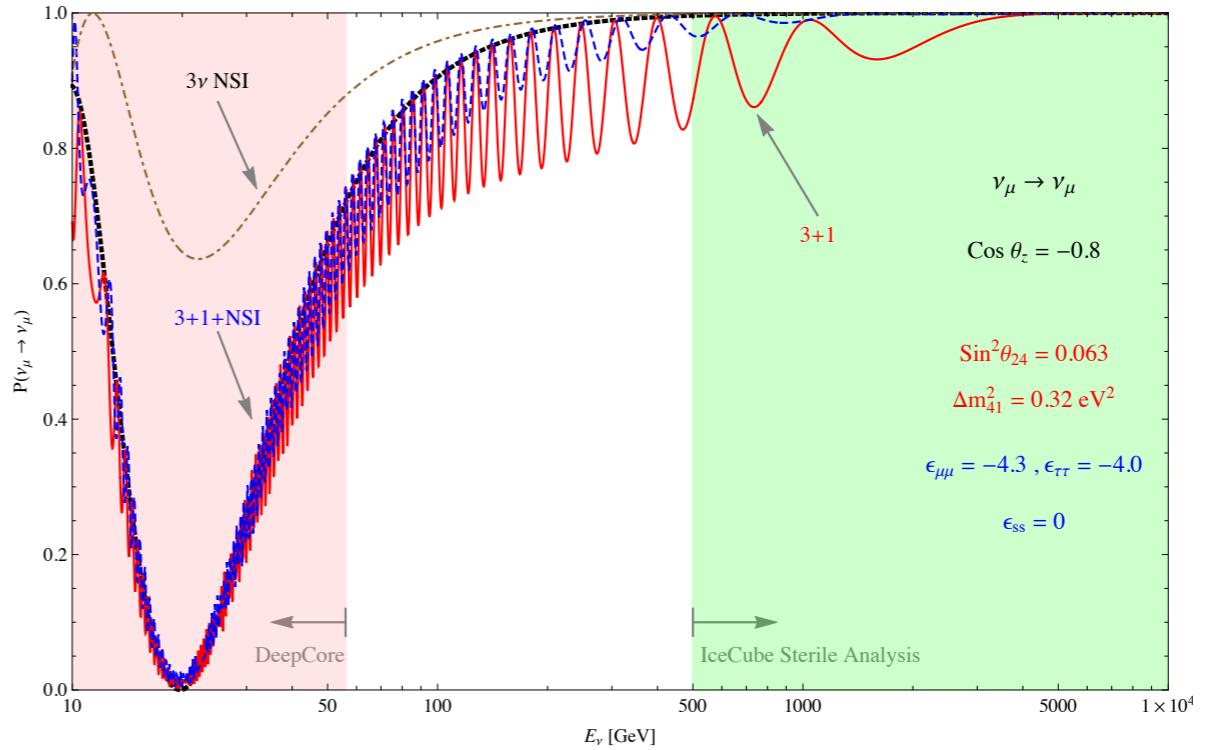
- Event by event -> Reconstruct secondary vertex with <1cm from the interaction point.
- Distribution -> Hadron kinematics



Tau appearance

• What can we study?

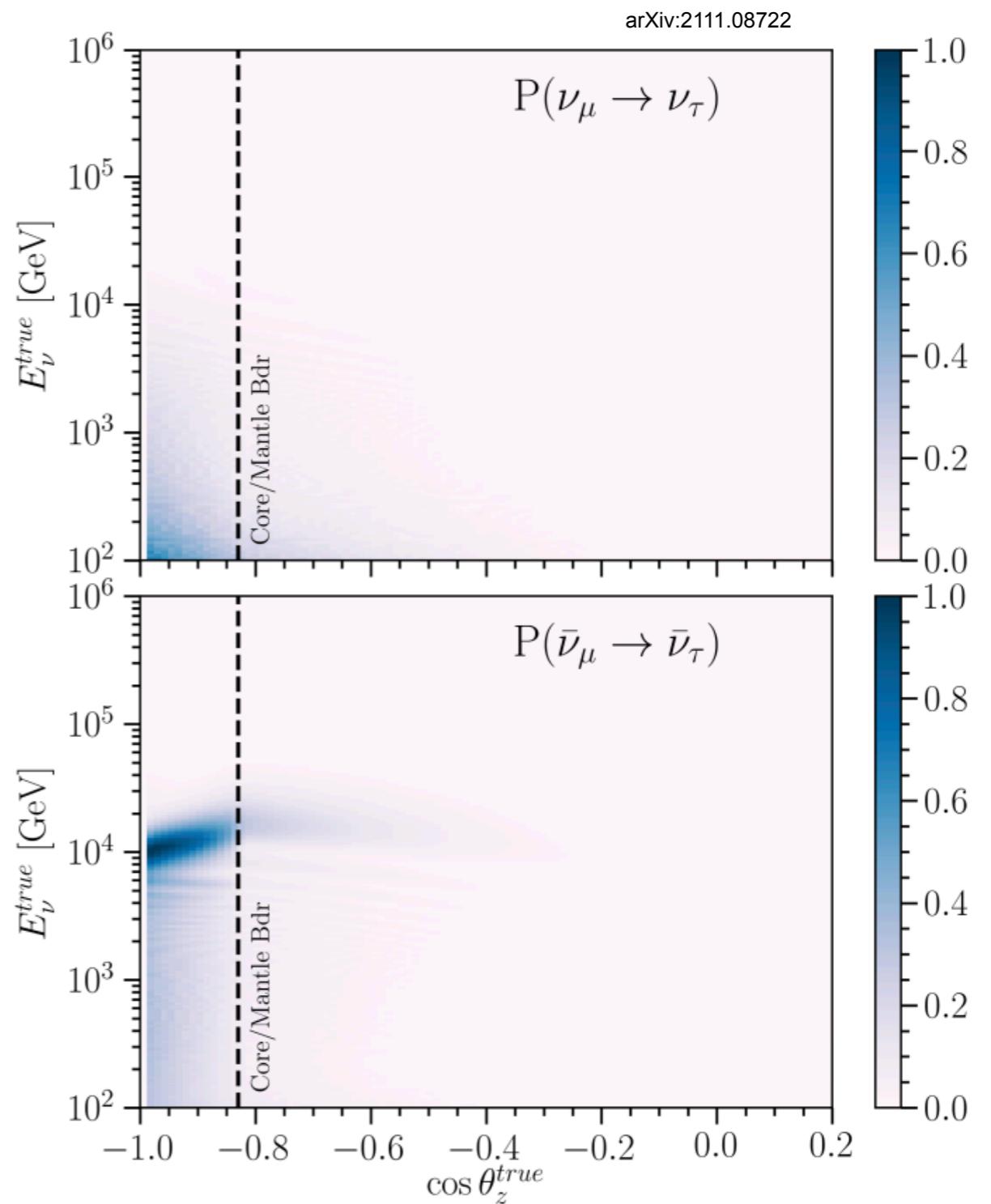
- ν_μ disappearance.
- Interesting portal to BSM.



arXiv:1810.11940

Tau appearance

- What can we study?
 - ν_τ vs $\bar{\nu}_\tau$ appearance.
 - Interesting portal to BSM.



Conclusion

- Neutrino beam from muon collider:
 - Expand along a wide energy range.
 - Very good understanding of flavor composition.
 - Intense!
- Which areas can benefit from it?
 - Neutrino cross sections -> nuclear effects, secondary interactions.
 - EW measurements.
 - Tau appearance.
- Concentrating on tomorrow's physics, not today's.
 - Technological challenges for engineers -> let them stew on it

Acknowledgements

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