

Federal University of Santa Catarina

# Coherent photoproduction of light vector mesons off nuclear targets in the dipole picture.

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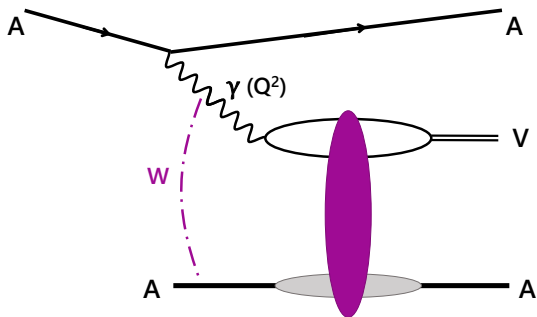
Preprint: 2310.06965



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# Coherent production

In the coherent vector meson photoproduction, the nucleus target remains intact.



The rapidity differential cross section for the  $AA \rightarrow AVA$  process can be factorized:

$$\frac{d\sigma^{AA \rightarrow AVA}}{dy} = \int d^2b \frac{\omega dN_\gamma(\omega, b)}{d\omega} \frac{d\sigma^{\gamma A \rightarrow VA}(\omega, b)}{d^2b} + (y \rightarrow -y).$$

# Photon flux and ultraperipheral collisions

Photon flux at  $b$  emitted by the projectile nucleus excluding strong interactions is [Glauber, Matthiae, Nucl. Phys. B 21, 135 (1970)]

$$\omega \frac{dN_\gamma(\omega, b)}{d\omega} = \int d^2 b_\gamma \frac{\omega N_\gamma(\omega, b_\gamma)}{d\omega d^2 b_\gamma} \exp\left(-\sigma_{\text{NN}}^{\text{tot}} \int d^2 b' T_A(b') T_A(|\vec{b}_{AA} - \vec{b}'|)\right). \quad (1)$$

The Weizsäcker-Williams method calculated for a punctual charge is used to calculate the  $b$ -dependent photon flux:

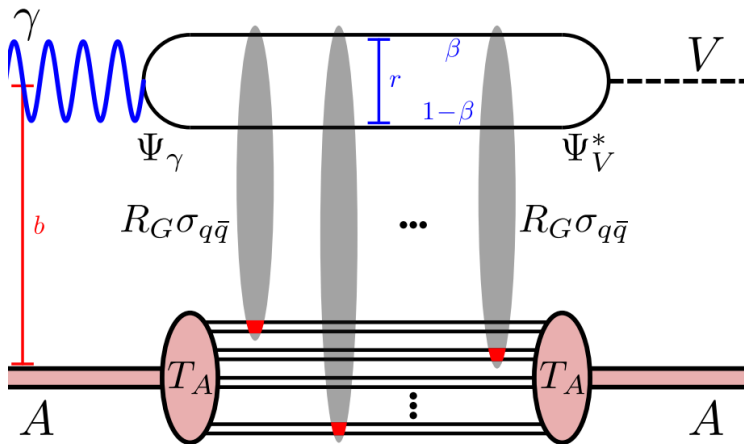
$$\frac{\omega d^3 N_\gamma(\omega, b_\gamma)}{d\omega d^2 b_\gamma} = \frac{Z^2 \alpha_{\text{em}} \omega^2}{\pi^2 \gamma^2} K_1^2\left(\frac{b_\gamma \omega}{\gamma}\right), \quad (2)$$

The Woods-Saxon distribution for the lead nuclear density was used in the thickness function

$$T_A(b) = \int_{-\infty}^{+\infty} dz \rho_A(b, z) \quad \text{with} \quad \rho_A(b, z) = \frac{N_A}{1 + \exp\left[\left(\sqrt{b^2 + z^2} - R_C\right) / \delta\right]}. \quad (3)$$

where  $N_A$  is the normalization and the parameters for lead (Pb) nucleus are  $R_C = 6.62$  fm and  $\delta = 0.546$  fm.

# Color dipole approach



# Wave functions

The wave function overlap is:

$$[\Psi_V^* \Psi_\gamma](\beta, r) = \frac{e_f e \sqrt{N_c}}{(2\pi)^{3/2} 2\beta(1-\beta)} \left[ m_f^2 K_0(\epsilon r) \psi_{n,L}(\zeta) - (\beta^2 + (1-\beta)^2) \epsilon K_1(\epsilon r) \partial_r \psi_{n,L}(\zeta) \right], \quad (4)$$

with  $\zeta^2 = \beta(1-\beta)r^2$ . The QED photon wavefunction is used. The vector meson wave function is calculated from AdS/QCD holographic model in the light cone. The scalar part is factorized to:

$$\psi(\beta, \zeta, \varphi) = e^{iL\varphi} \sqrt{\beta(1-\beta)} \frac{\phi(\zeta)}{\sqrt{2\pi\zeta}} e^{\frac{m_q^2}{\beta} + \frac{m_q^2}{1-\beta}},$$

The  $\phi(\zeta)$  wave function is found by solving the relativistic equation

$$\left( -\frac{d^2}{d\zeta^2} - \frac{1-4L^2}{4\zeta^2} + U(\zeta) \right) \phi(\zeta) = M^2 \phi(\zeta)$$

with effective confining potential:

$$U(\zeta, J) = \kappa^4 \zeta^2 + 2\kappa^2(J-1)$$

The solutions are HO wave functions, with eigenvalues

$$M^2 = 4\kappa^2 \left( n + \frac{J}{2} + \frac{L}{2} \right).$$

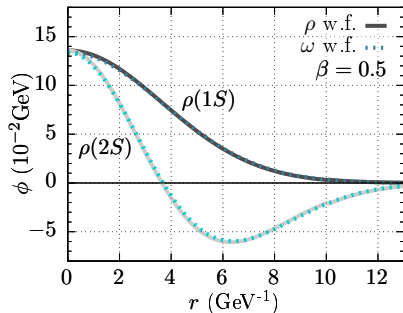
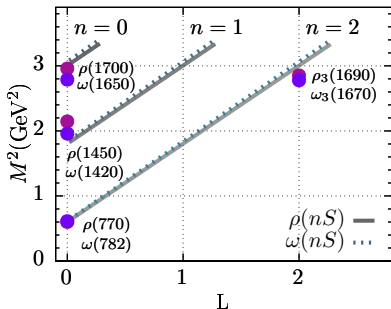
# Vector meson wave function

## AdS/QCD holographic model

Mass dependent  $\kappa$

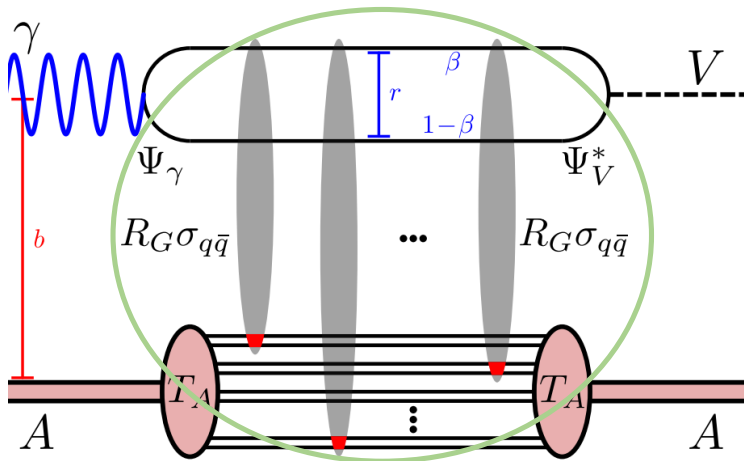
$$\kappa = M_{V(n=0)}/\sqrt{2}$$

The left panel shows the mass spectroscopy of  $\rho$  and  $\omega$  vector mesons.



The right panel presents the LF wave functions for the  $\rho$  vector meson at ground and excited states.

# Dipole-nucleus cross section



The total cross sections for the  $\gamma A \rightarrow VA$  process is

$$\frac{d\sigma^{\gamma A \rightarrow VA}(\omega, b)}{d^2b} = \left| \int d\beta d^2r \Psi_V^* \Psi_\gamma(\beta, r) \left( 1 - e^{-\sigma_{q\bar{q}}(x, r) T_A(b)/2} \right) \right|^2$$

The 2017 fit of the GBW model was used to describe the dipole cross section  $\sigma_{q\bar{q}}$ .

$$\sigma_{q\bar{q}}(x, r) = \sigma_0 \left(1 - e^{-r^2 Q_s^2(x)/4}\right), \quad (5)$$

where parameters were taken from the Golec-Biernat & Sapeta 2017 fit to HERA DIS data:  $\sigma_0 = 27.32$  mb,  $Q_s^2(x) = Q_0^2 (x_0/x)^\lambda$ , with  $Q_0^2 = 1$  GeV<sup>2</sup>,  $x_0 = 0.42 \times 10^{-4}$  and  $\lambda = 0.248$ .

## 1 Real part

$$\sigma_{q\bar{q}}(x, r) \Rightarrow \sigma_{q\bar{q}}(x, r) \left(1 - i \frac{\pi}{2} \lambda\right)$$

## 2 Skewness

$$R_g(\lambda) = \frac{2^{2\lambda+3}}{\sqrt{\pi}} \frac{\Gamma(\lambda + 5/2)}{\Gamma(\lambda + 4)}$$

$$\lambda = \frac{\partial \ln \sigma_{q\bar{q}}(x, r)}{\partial \ln (1/x)}$$

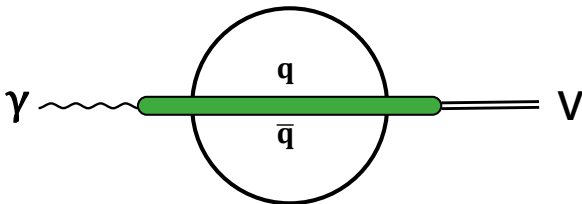


# Coherence length

At high energies, the  $q\bar{q}$  pair lifetime, called coherence length, is defined as

$$l_c = \frac{2\omega'}{M_V^2}. \quad (6)$$

The dipole model considers it to be much bigger than the nucleus radius.



At high energies, we need to consider that the photon can split into higher states with gluons  $|q\bar{q}g\rangle, |q\bar{q}gg\rangle, \dots$ , with **smaller** coherence lengths.

[Kopeliovich, Schafer and Tarasov, Phys. Rev. D 62 054022, 2000]

# Gluon shadowing

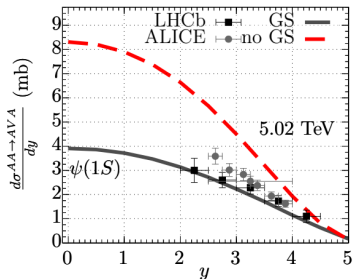
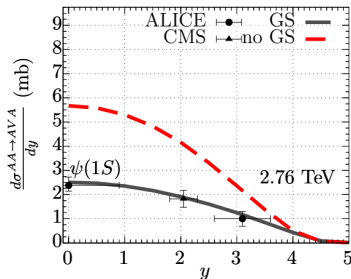
- Higher states fully considered when the proton is the target. Effectively, they are included in the GBW fit.
- When the nucleus is the target, higher states contribute less since they do not survive long enough.
- This reduction of the  $\gamma A$  cross section is called **gluon shadowing**.
- As a first approximation, it is just a rescaling of the dipole cross section by the factor  $R_G$

$$\sigma_{q\bar{q}}(r, x) \rightarrow \sigma_{q\bar{q}}(r, x)R_G(x, \mu^2).$$

# Gluon shadowing for heavy vector mesons

In the heavy vector meson case it is calculated with EPPS16 and a factorization scale  $\mu = M_V/2$ .

[Henkels, E.G.O., Pasechnik, Trebien, Phys. Rev. D 102, 014024 (2020)].

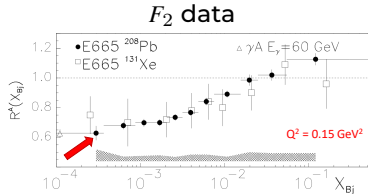
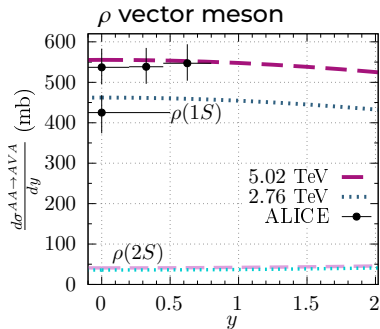


The evolution does not reach the light vector meson scale.

We do not have gluon shadowing for smaller scales

↪ This is a **PROBLEM**.

# Gluon shadowing extraction from data



[Data taken from the Fermilab E665 Collaboration,  
Z.Phys.C 67 (1995) 403-410]

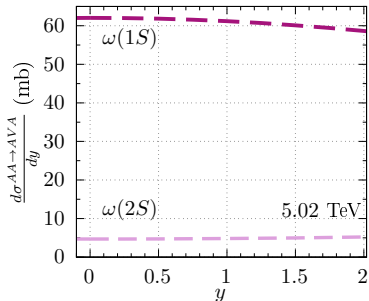
$$\frac{F_2^A}{AF_2^P} = 0.537$$

The available data led to an optimal value for the gluon shadowing

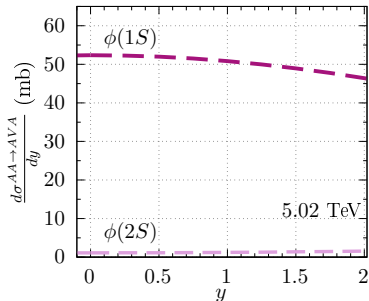
$$R_G = 0.85$$

# Coherent photoproduction predictions

$\omega$  vector meson



$\phi$  vector meson

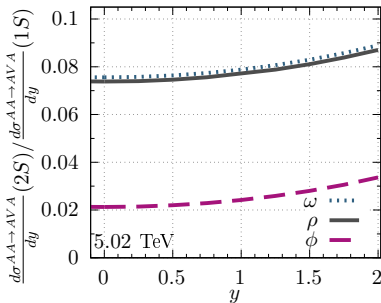


We used the same  $R_G = 0.85$  factor to make predictions for other light vector mesons.

Measurements can bring information about the  $R_G$  scale dependence.

# 2S to 1S ratio predictions

$\rho$ ,  $\omega$ , and  $\phi$  vector mesons



Predictions for the ratio between the excited state cross section and its corresponding ground state cross section.

This can be a tool for finding the best dipole cross section model.

# Conclusion

- **Light vector meson** photoproduction: **small hard scales  $Q^2$** .
- The **Glauber-Gribov** approach is used for nuclear targets.
- The inclusion of **gluon shadowing**  $R_g$  is necessary.
- Fitting to **ALICE  $\rho(1S)$  UPC** and **E665  $F_2$**  data gives  **$R_g = 0.85$** .
- We make **predictions**:  $\rho(2S)$ ,  $\omega(1S, 2S)$ , and  $\phi(1S, 2S)$ .
- The **ratio** between excited and ground state cross sections depend on **vector meson wavefunctions** and probe the **dipole cross section**.

**See more:**

Haimon Trebien & Bruna Stahlhöfer talks

# Thank you







EXTRA SLIDES

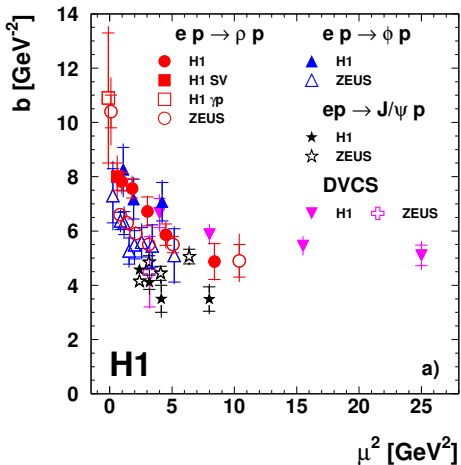
The slope parameter for heavy vector mesons takes the form:

$$B(W, Q^2) \approx B_0 + 4\alpha'(0) \ln\left(\frac{W}{W_0}\right) - B_1 \ln\left(\frac{Q^2 + M_V^2}{M_{J/\Psi}^2}\right), \quad (7)$$

while the one for light vector mesons is given by:

$$B = N \left[ 14.0 \left( \frac{1\text{GeV}^2}{Q^2 + M_V^2} \right)^{0.2} + 1 \right], \quad (8)$$

# Slope parameter



# Color dipole parametrizations

without impact parameter dependence

## 1 GBW

It has the form:

$$\sigma_{q\bar{q}}(x, r) = \sigma_0 \left( 1 - e^{-\frac{r^2 Q_s^2(x)}{4}} \right), \quad (9)$$

with saturation scale defined as  $Q_s^2(x) = Q_0^2 \left( \frac{x_0}{x} \right)^\lambda$ .

## 2 KST

It includes corrections for small  $\hat{s} \equiv W^2$ .

$$\sigma_{\bar{q}q}(r, \hat{s}) = \sigma_0(\hat{s}) \left[ 1 - e^{-r^2/R_0^2(\hat{s})} \right]. \quad (10)$$

In this case, the  $\hat{s}$ -dependence appears in  $\sigma_0(\hat{s})$  and  $R_0(\hat{s})$ , which are given by

$$R_0(\hat{s}) = 0.88 \text{fm} (s_0/\hat{s})^{0.14}, \quad \sigma_0(\hat{s}) = \sigma_{\text{tot}}^{\pi p}(\hat{s}) \left( 1 + \frac{3R_0^2(\hat{s})}{8 \langle r_{\text{ch}}^2 \rangle_\pi} \right). \quad (11)$$

# Color dipole parametrizations

with impact parameter dependence

### 3 bSat

The partial dipole amplitude is

$$N(x, \mathbf{r}, \mathbf{b}) = 1 - \exp\left(-\frac{\pi^2}{2N_c} r^2 \alpha_s(\mu^2) x g(x, \mu^2) T(b)\right) \quad (12)$$

with the following scale dependence  $\mu^2 = 4/r^2 + \mu_0^2$ .

### 4 BK

The BK evolution equation is given by

$$\frac{\partial N(r, b, Y)}{\partial Y} = \int d\mathbf{r}_1 K(r, r_1, r_2) (N(r_1, b_1, Y) + N(r_2, b_2, Y) - N(r, b, Y) - N(r_1, b_1, Y) N(r_2, b_2, Y)) . \quad (13)$$

Gluon recombination effects are taken into account in the non-linear term

# Color dipole parametrizations

## 5 bCGC

It is the interpolation of solutions for the BFKL and the BK :

$$N(x, \mathbf{r}, \mathbf{b}) = \begin{cases} N_0 \left( \frac{r Q_s}{2} \right)^{2[\gamma_s + (1/(\eta \Lambda Y)) \ln(2/r Q_s)]} & r Q_s \leq 2 \\ 1 - e^{-\mathcal{A} \ln^2(\mathcal{B} r Q_s)} & r Q_s > 2 \end{cases} . \quad (14)$$

The  $b$ -dependence is introduced in the saturation scale

$$Q_s \equiv Q_s(x, b) = \left( \frac{x_0}{x} \right)^{\Lambda/2} \left[ \exp \left( -\frac{b^2}{2B_{\text{CGC}}} \right) \right]^{1/(2\gamma_s)} . \quad (15)$$

