#### **R&D on GAGG single crystals for fast timing detectors** in high-rate and radiation environments

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 <sup>4</sup>Institute of Physics of the Czech Academy of Sciences, 16200, Prague, Czech Republic

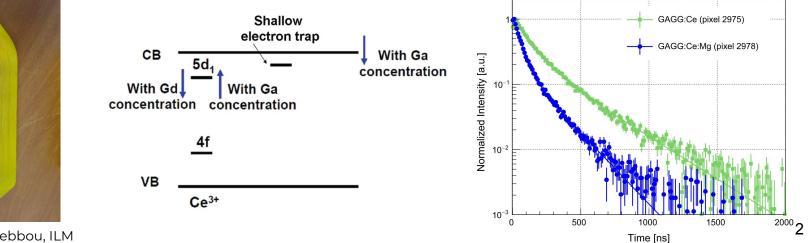
#### Gadolinium Aluminium Gallium Garnet

	light yield (photons/MeV)	first decay time (ns)	second decay time (ns)
$\mathrm{Gd}_3\mathrm{Al}_4\mathrm{Ga}_1\mathrm{O}_{12}$ $\mathrm{Gd}_3\mathrm{Al}_3\mathrm{Ga}_2\mathrm{O}_{12}$	15 895 <b>45 931</b>	316 (100%) 221 (100%)	
$\mathrm{Gd}_3\mathrm{Al}_2\mathrm{Ga}_3\mathrm{O}_{12}$	42 217	52.8 (73%)	282 (27%)
$Gd_{3}Al_{1}Ga_{4}O_{12} \\$	17912	42.2 (34%)	90.5 (66%)
$Gd_{3}Al_{0}Ga_{5}O_{12} \\$	0	*ND	*ND

K. Kamada et al., Cryst. Growth Des. 2011, 11, 10, 4484–4490 Ce-doped multi-component garnets discovered in 2011. Amongst them is Gadolinium Gallium Aluminium Garnet  $Gd_3Al_2Ga_3O_{12}$  (GAGG):

- High light yield and fast scintillation
- Tunable composition
  - K. Kamada et al., Optical Materials 36 (2014) 1942–1945
- Acceleration of scintillation via divalent ions codoping (e.g. Magnesium)

K. Kamada et al., Optical Materials 41 (2015) 63–66 M. Lucchini et al., NIM A816 (2016) 176-183

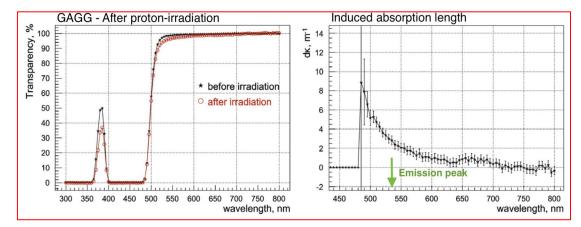


Courtesy of K. Lebbou, ILM

# **Radiation Hardness**

#### Garnets are radiation hard

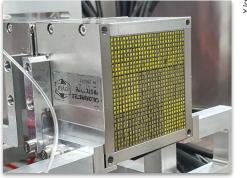
- GAGG irradiated w 3.1x10<sup>15</sup> protons/cm<sup>-2</sup> (24 GeV/c)
  - Dose: 910 kGy
    Induced absorption below 4m<sup>-1</sup> at
  - the emission peak See: V. Alenkov et al., NIM A 816 (2016) 176

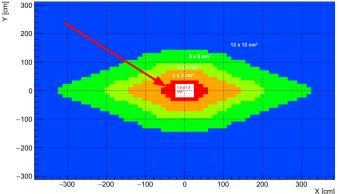


--> GAGG fibres are the baseline solution for the innermost region of the LHCb PicoCal

(see talk by E. Picatoste Tue 26)



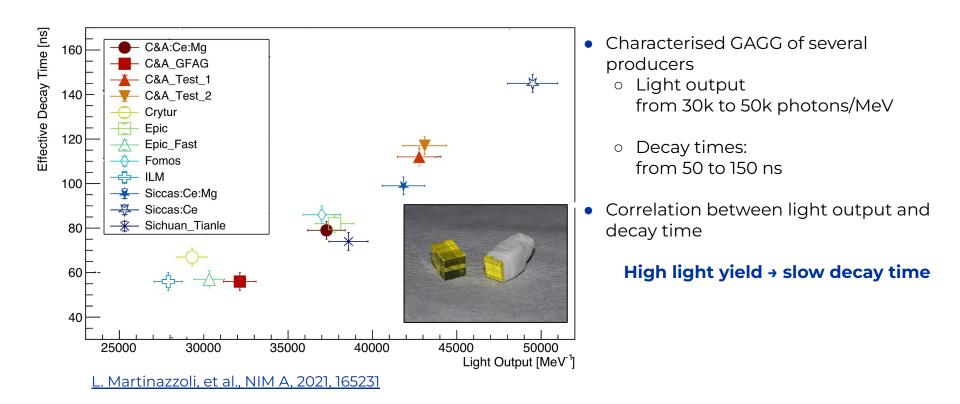




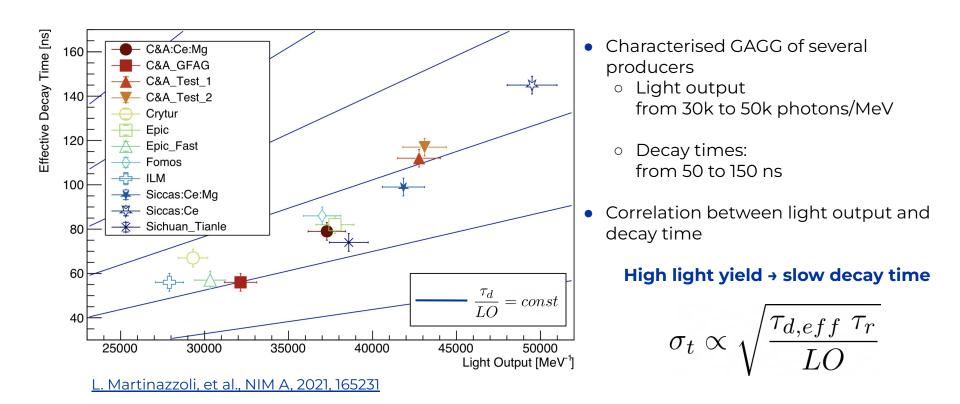
# I. Present GAGG

II. Accelerating Scintillation

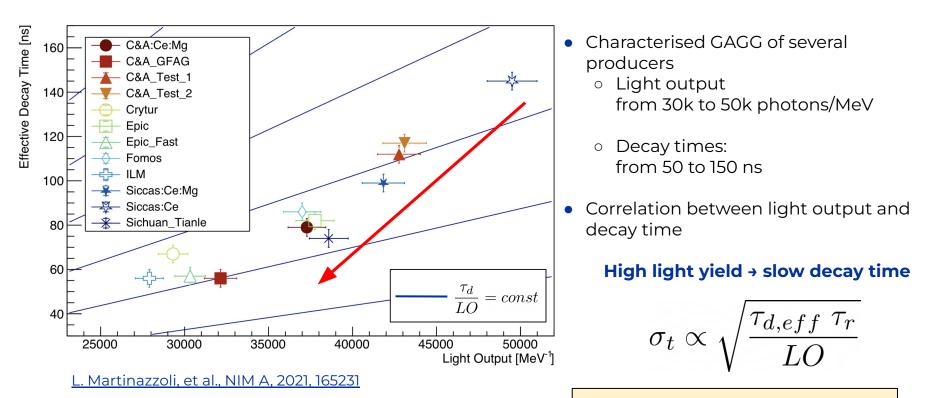
#### Light Output and Effective Decay Time



#### Light Output and Effective Decay Time



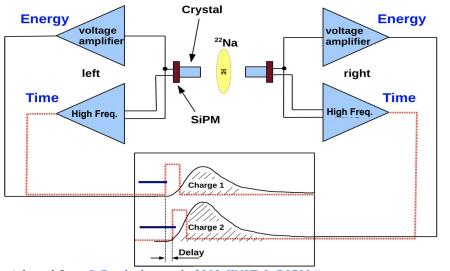
#### Light Output and Effective Decay Time



**Faster samples have better timing** 

7

#### **Coincidence Time Resolution**



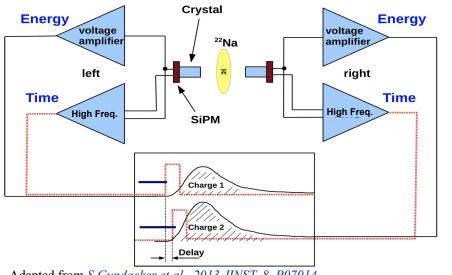
Adapted from S. Gundacker et al., 2013 JINST, 8, P07014

- Time resolution measured with two 511 keV photons in coincidence with <sup>22</sup>Na (CTR)
- High-Frequency amplifier (~1.5 GHz bandwidth) for timing

CTR of GAGG in line with LYSO:Ce

Crystal	$\begin{array}{c} \mathrm{CTR} \ \mathrm{(HF)} \ \mathrm{[ps]} \\ \mathrm{2x2x3} \ \mathrm{mm}^{3} \end{array}$
C&A GFAG	$87\pm2$
ILM GAGG	$90 \pm 2$
EPIC Fast GAGG	$90 \pm 2$
LYSO:Ce	$86\pm2$

#### **Coincidence Time Resolution**



Adapted from S. Gundacker et al., 2013 JINST, 8, P07014

However, decay time of ~ 50 ns (+ slow component)...

--> Pile-up at high rates!

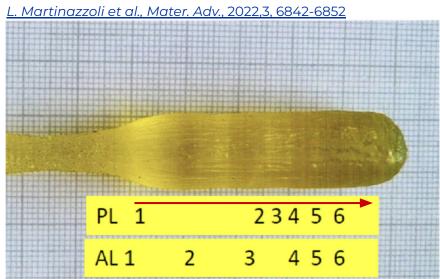
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# I. Present GAGGII. Accelerating Scintillation

## The Samples



	g	Gd	Ce	Mg	Ga	Al
SC		2.955	0.015	0.03	3	2
AL1	0.065	2.917	0.0065	0.0051	2.787	2.284
AL2	0.155	2.944	0.0063	0.0053	2.813	2.231
AL3	0.416	2.978	0.0096	0.0058	2.845	2.160
AL4	0.618	2.954	0.0131	0.0039	3.043	1.986
AL5	0.703	2.959	0.0164	0.0060	3.108	1.982
AL6	0.789	2.952	0.0279	0.0112	3.251	1.758
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Samples produced from plates cut along the ingot

- AL to study composition
- PL to study luminescence

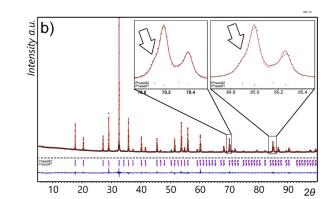
Along the ingot:

- Both Ce and Mg increasing

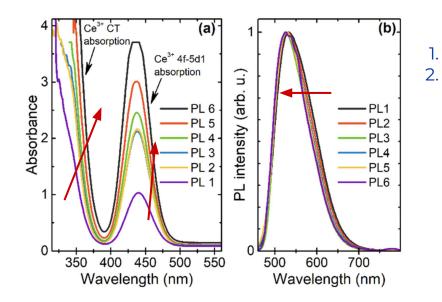
   Up to l at%!
- Ga increasing

Rietveld analysis found 2 garnet phases (80%-20%)

- Possibly due to Al/Ga subclusters See: Phys. Status Solidi A 2018, 215, 1701034
- No perovskite phase



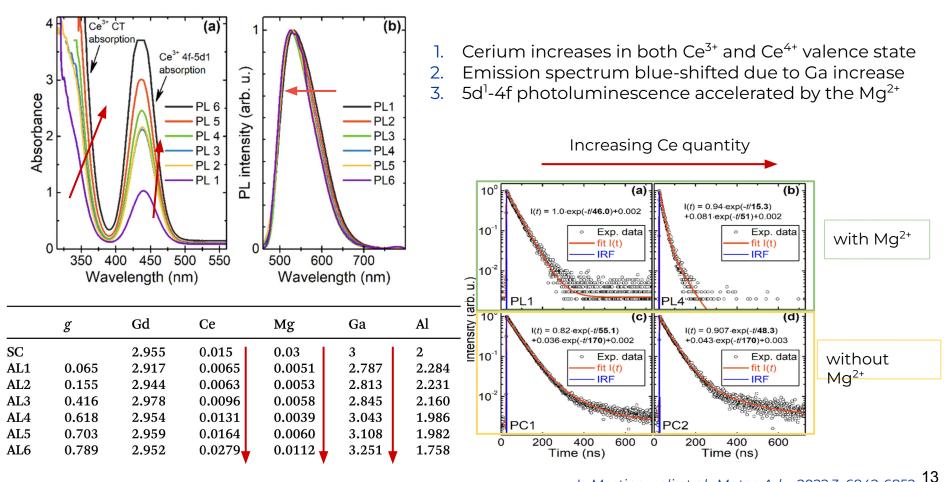
#### Photoluminescence



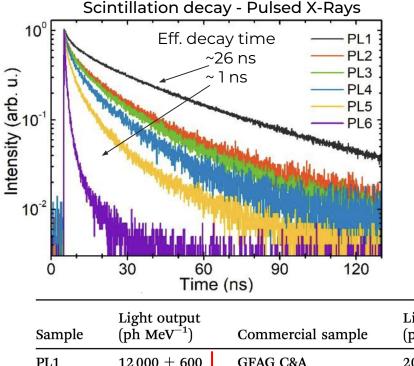
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Cerium increases in both Ce<sup>3+</sup> and Ce<sup>4+</sup> valence state Emission spectrum blue-shifted due to Ga increase

#### Photoluminescence



## Scintillation

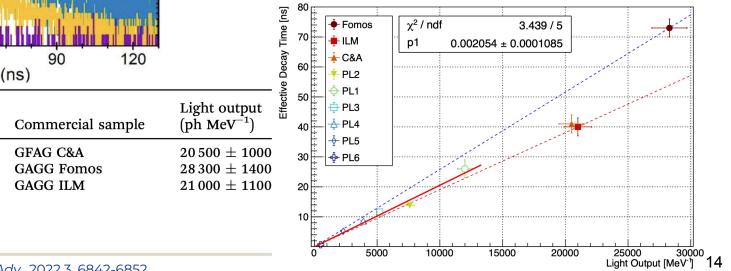


#### Samples of 1x1x5 mm<sup>3</sup> used

Light quenching observed moving along the ingot:

- Light output reduction by > 10x
- Decay time accelerated by > 10x

Acceleration compensating the Light Yield reduction to maintain the same initial photon time-density



L. Martinazzoli et al., Mater. Adv., 2022,3, 6842-6852

 $7600 \pm 380$ 

 $5200 \pm 260$ 

 $3900 \pm 190$ 

 $2300 \pm 230$ 

 $500 \pm 300$ 

PL2

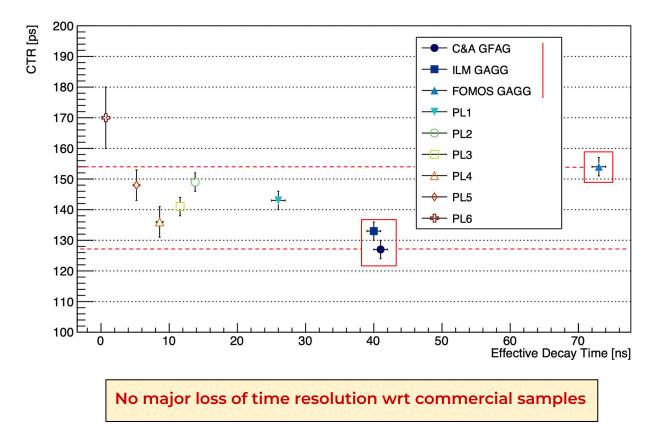
PL3

PL4

PL5

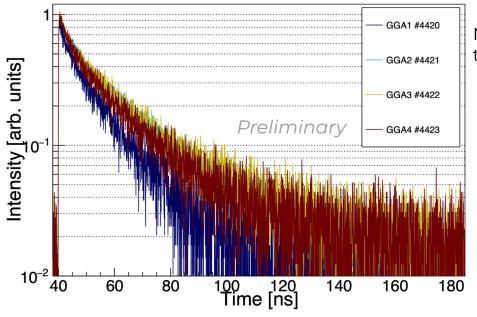
PL6

#### Timing



Note: different (worse) surface state for the new samples

#### New Samples

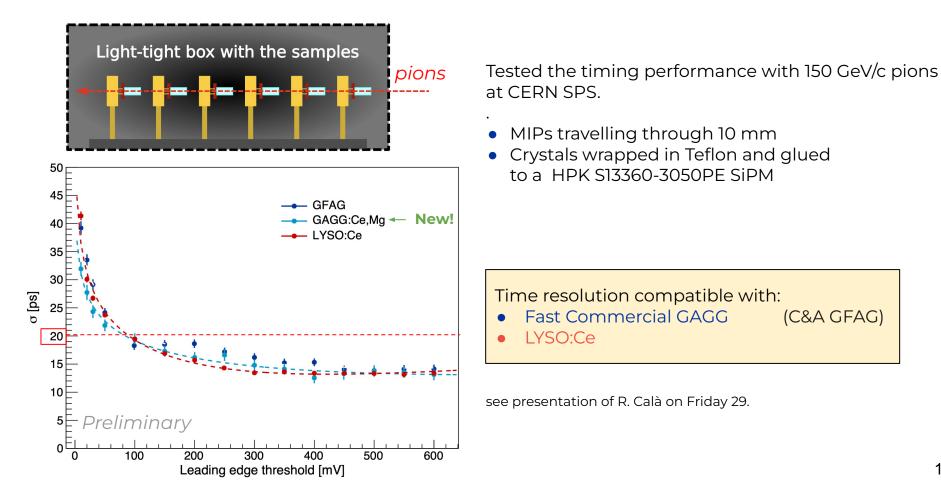


New and larger samples of **2x2x10 mm<sup>3</sup>** produced by the FZU institute, Prague.

- 4 samples with similar composition (and scintillation)
- Main decay time of 15-20 ns
- Slow component under study

Crystal	Photons Output $[MeV^{-1}]$	$ au_{d,1} \; [\mathrm{ns}]$	%	$\tau_{d,2} \; [\mathrm{ns}]$	%	$ au_{d,3} \; [\mathrm{ns}]$	%	$ au_{d,eff} \; [\mathrm{ns}]$
Sample 1	6'200	3	10	15	66	99	24	12
Sample 2	7'400	3	8	19	<b>67</b>	77	25	16
Sample 3	7'400	3	7	18	66	82	27	15
Sample 4	6'200	3	10	17	67	80	23	14

#### Timing with MIPs



Cerium-doped **GAGG** is:

- radiation hard
- fast and with high light output
- commercially available today

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Decay time can be accelerated with heavy (co)-doping:

- Heavy doping achieved up to ~1 at%
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  - No perovskite phase observed
- Acceleration of the scintillation by >10x paired by reduction in light output
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Ongoing R&D to develop large-size ingots with accelerated scintillation in:

• **Crystal Clear Collaboration** and the **TWISMA project** TWIN European project between ISMA, CERN, ILM Lyon

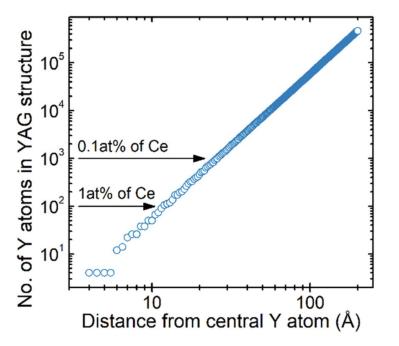
(Grant agreement 101078960)

• Institutes taking part in the LHCb PicoCal



# **Back-up**

# Ce-Mg pairs

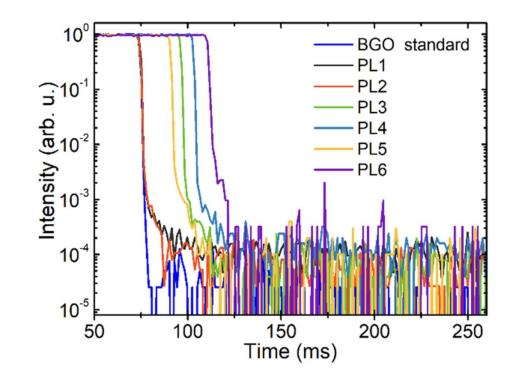


 Recent literature suggests that PL quenching in co-doped garnets is due to {Ce-Mg} pairs see:

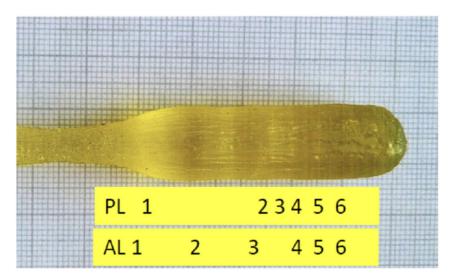
Babin et al., Optical Materials **83** (2018) 290 Babin et al., J. Lumin. **215** (2019) 116608

- Doping of 1 at% reduces the distance of dopants at the level of the lattice parameter
- Different components explained by discrete available distances between dopants.
- Similar behaviour observed in CeF<sub>3</sub> see: M. Nikl, et al., J.Phys.Cond.Mat. 7, 6355-6364 (1995)

#### Afterglow



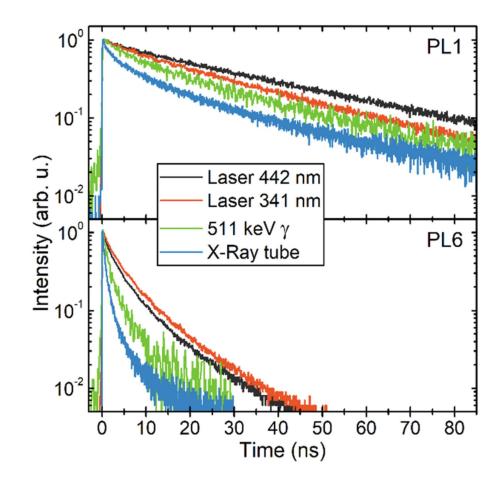
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- Diameter: 12.5 mm
- Czochralski Iridium crucible
- Atmosphere: N + 2% O
- Cyberstar Mini-oxypuller machine
- Elementary cell parameters, 12.28279(5) Å and 12.29436(13) Å, with their volume ratio of 0.8 : 0.2.

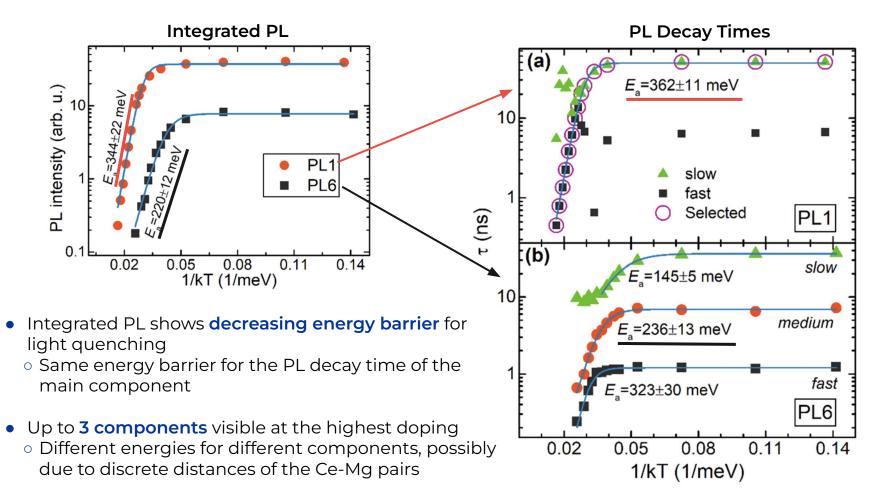
#### **Kinetics Comparison**



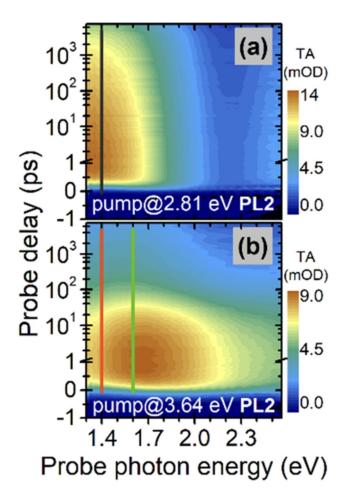
# **X-Rays Kinetics**

Sample	$\tau_{r}$ (ps)	τ <sub>d1</sub> (ns)	$R_1$ (%)	$\tau_{d2}$ (ns)	$R_2$ %	$\tau_{d3}$ (ns)	R <sub>3</sub> (%)	$\tau_{d,eff}$ (ns)
C&A GFAG	32	6.0	4.6	44.5	69.2	222	26.3	41
ILM GAGG	37	4.0	3.2	40.4	56.4	138	40.4	40
Fomos GAGG	30	2.2	0.5	53.1	41.7	166	57.8	73
PL1	13	2.5	3.3	25.4	48.0	79.2	48.8	26
PL2	8	2.1	7.2	16.6	54.6	66.2	38.2	13.8
PL3	5	1.6	6.2	12.6	47.5	46.0	46.3	11.6
PL4	5	1.5	9.2	11.3	53.9	45.4	36.9	8.6
PL5	5	1.0	11.0	7.1	51.8	40.8	37.2	5.2
PL6	5	0.2	19.5	1.5	53.0	14.9	27.5	0.7

#### Temperature dependence of Photoluminescence



# **Transient Absorption**



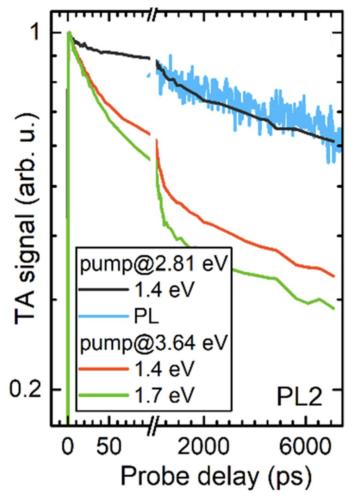
Studied non-equilibrium absorption bands:

• Pump 4f-5d<sup>1</sup> - 2.8 eV (a) :

absorption band in line with literature for GAGG
 see: G. Tamulaitis et al., J. Appl. Phys. 124, 215907 (2018)
 and Keynote presentation on Wednesday by G. Tamulaitis

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  - $\circ\,$  new band extending towards higher eV increasing dopants

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  - new absorption band peaking at 1.7 eV
  - $\circ\,$  new band extending towards higher eV increasing dopants
- Non-exponential decay of electrons in 5d<sup>1</sup> and 5d<sup>2</sup> confirms non-radiative recombination channels