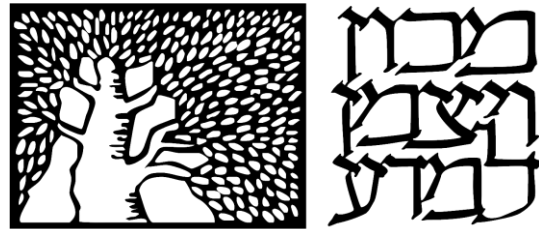


Novel noble-liquid detector concepts

Amos Breskin

Dept. of Astrophysics & Particle Physics



WEIZMANN INSTITUTE OF SCIENCE

Main NL-detector “users”: DM & ν physics

also $0\nu\beta\beta$ decay, HEP calorimetry, medical imaging, homeland security, ...

DM: mostly liquid XENON LXe (Darkside LAr)

Neutrino: liquid ARGON LAr

DM: WIMPS (GeV-TeV)

WIMP elastic scattering off atomic target nuclei

→ low deposited energy (keV) by nuclear recoils → **few – to few tens e-'s**

requires low detection thresholds!

→ e- multiplication & efficient scintillation-photon recording

→ background-radiation discrimination

Neutrino physics:

High MIP charge deposits ($\sim 10^4$ e-/mm)

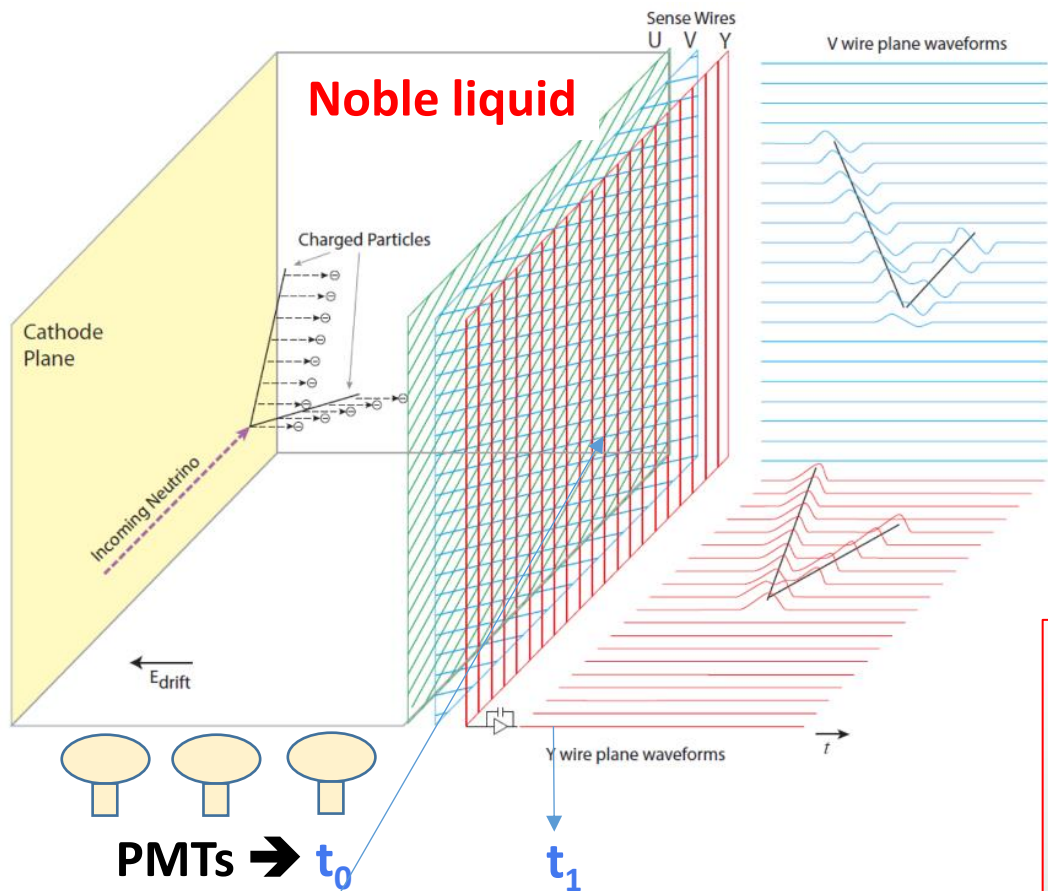
→ Can operate **without charge multiplication**


 multiplication would ease electronics; lower detection thresholds

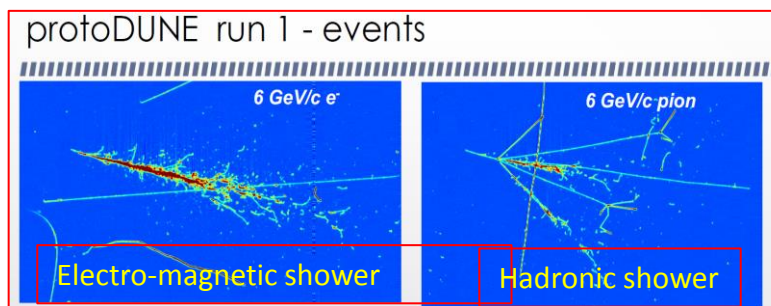
→ other physics goals? (e.g. *lower-energy phenomena*)

Single-phase TPC concept - Neutrinos

Single-phase Charge collection



- ν - induced charged particles ionize the noble liquid;
- \sim few 10^4 e^- /mm in LAr (MIPs)
- Electrons drift to, and induce charge on anode wires (2D);
- Primary-scintillation light provides the e^- drift time
 \rightarrow third coordinate; $(t_1 - t_0) v_e$
 \rightarrow 3D spatial reconstruction from wire planes;
- Charge collection with **no multiplication**
 \rightarrow good only for **large deposited charge**
 (e.g. neutrino physics)
- **Largest: ICARUS (760t), DUNE (68.000t)...** 

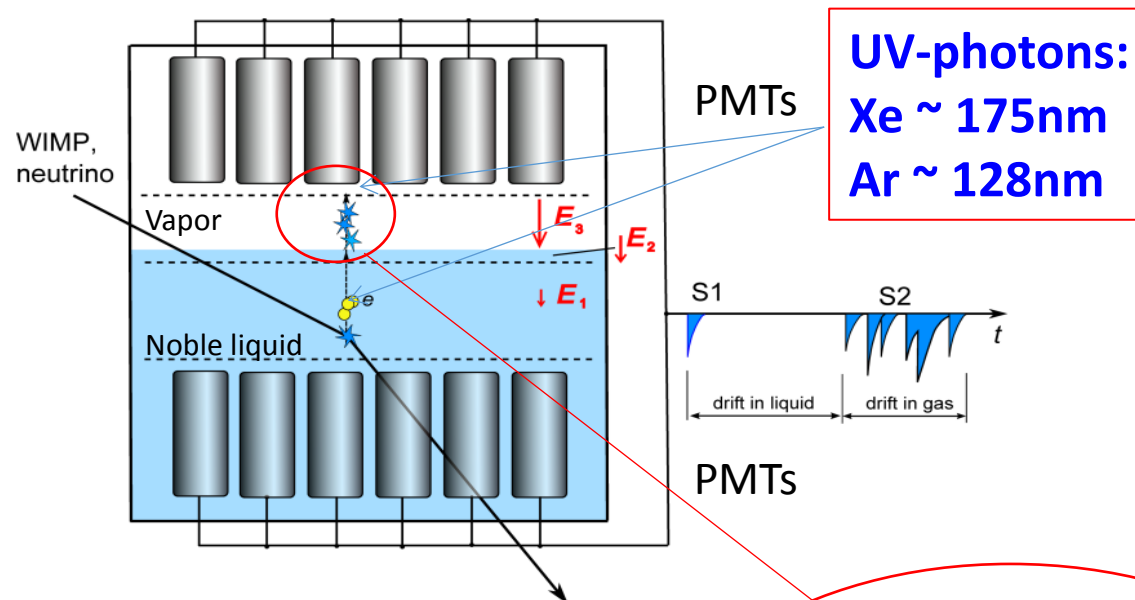
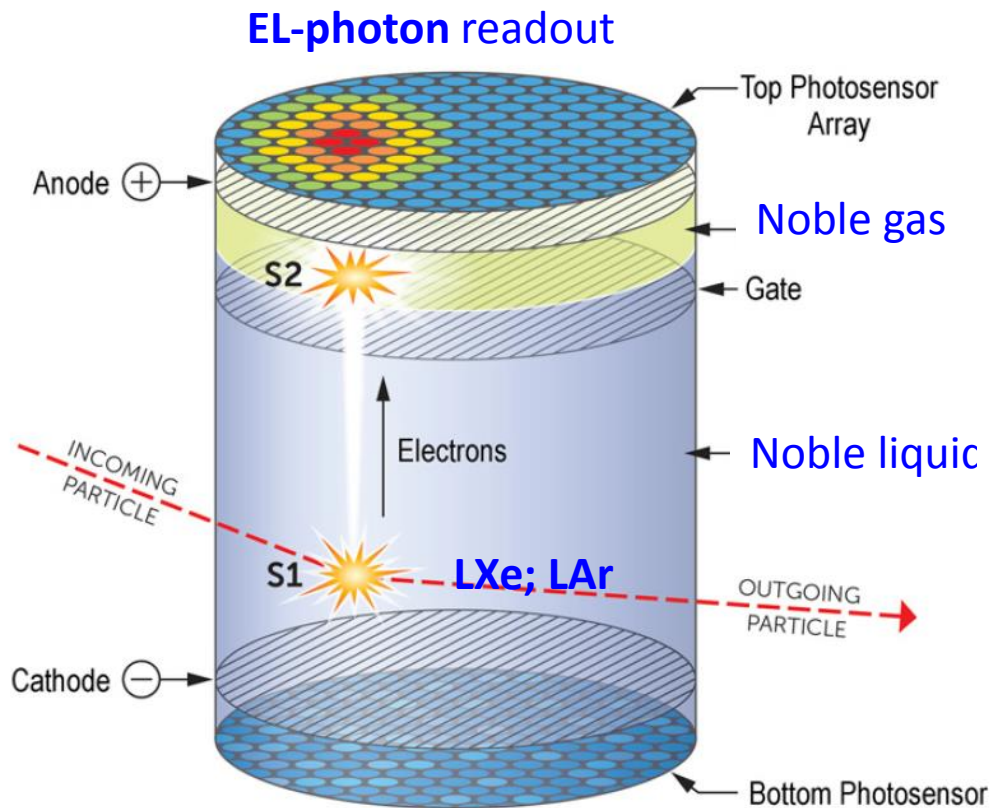


single proto-DUNE module...

- **3 wire planes or**
- **Multilayer perforated PCB** In 2 DUNE “vertical drift” modules

Dual-phase TPC concept - WIMPs

Direct detection of rare **WIMP** elastic scattering off atomic target nuclei



UV-photons:
Xe ~ 175nm
Ar ~ 128nm

S1 light signal:

- prompt scintillation photons

S2 charge signal:

- secondary scintillation photons from electroluminescence in GXe due to drifted electrons

3D event position reconstruction:

- X,Y: S2 hit pattern
- Z: drift time S2-S1

e.g. LXe 2-phase:

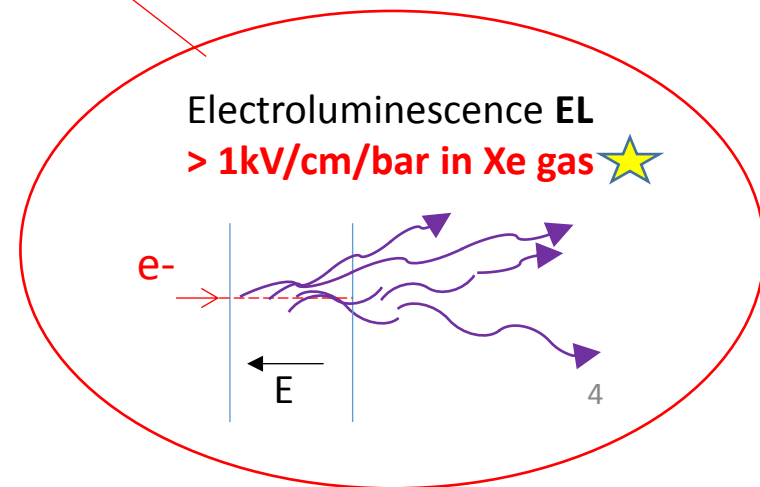
XENONnt, LUX-ZEPLIN (LZ), PNDA X, DARWIN/G3

8.5t

10t

4t

50t



Rare events: fighting background

WIMP signals:

Extremely small WIMP interaction cross sections $<10^{-48}\text{cm}^2$

→ VERY LARGE VOLUMES & HIGH SENSITIVITY

- Very rare:

1 event/ton x Y in LXe

- Buried in huge background:
> 10^6 -fold higher rates

- Very small: low-E recoils eV-KeV

→ **few-to-tens electrons in noble-liquid**

→ Underground

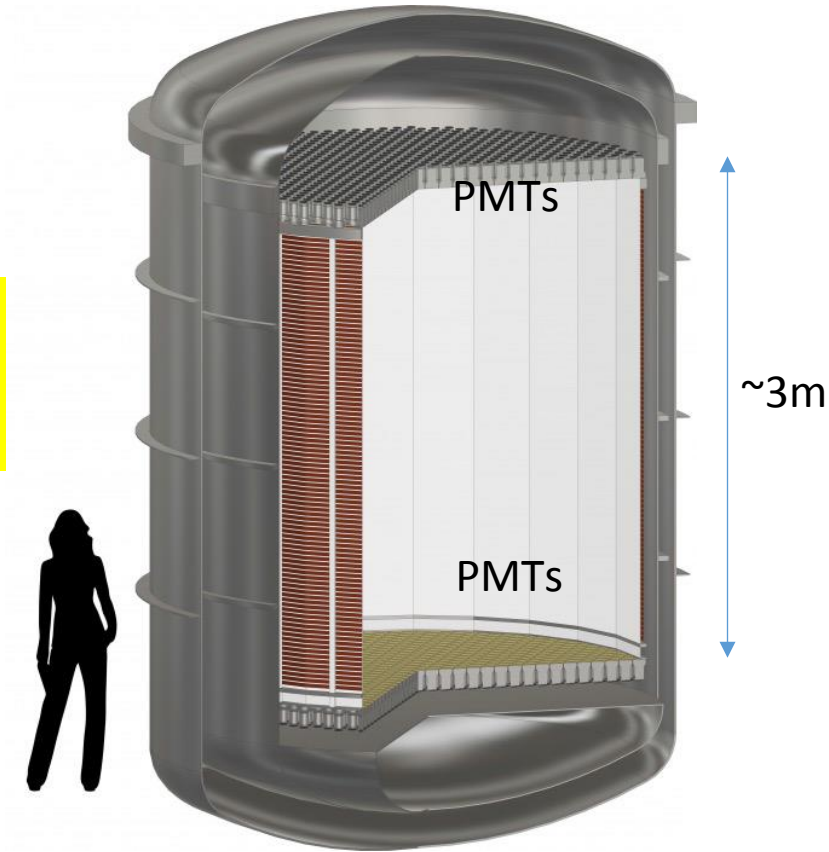
→ Low-radioactivity materials/gas

→ Cosmics VETO

→ ER background rejection

→ **Motivation for: novel detector concepts**

$10^{-49}\text{cm}^2 \rightarrow$
DARWIN LXe: 50t x 20Y
ARGO LAr: 300t x 10Y



Future 50-ton LXe

DARWIN/G3 Dark-matter observatory

XENON-LZ-DARWIN (XLZD consortium)

Aalbers et al 2023 J. Phys. G: Nucl. Part. Phys. 50 013001

DARWIN/G3 Baseline: Dual-phase/PMTs

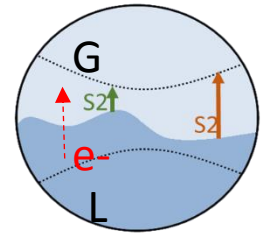
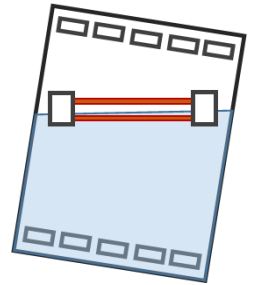
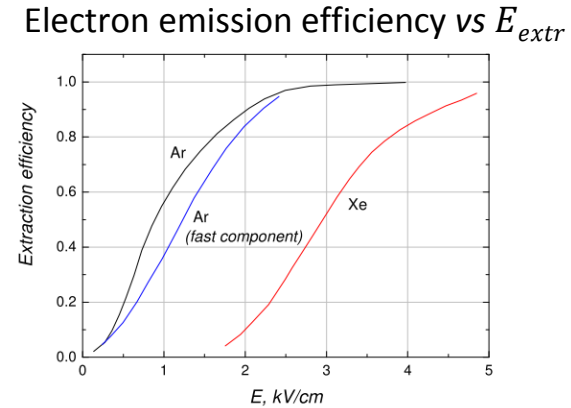
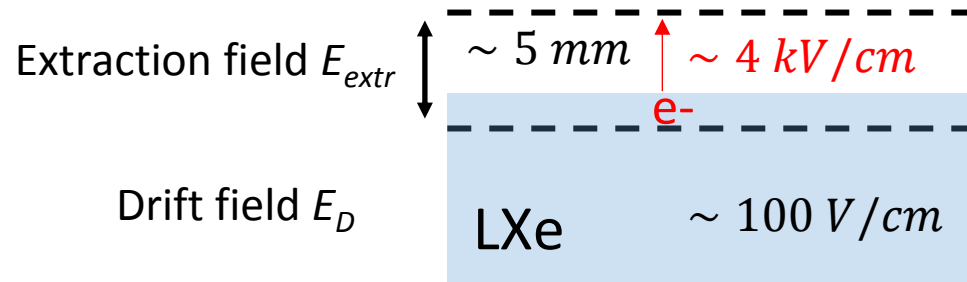
Ongoing extensive R&D on
detector & photo-sensors

Why new concepts?

- **Dual-phase detectors: Problems**

Current expected problems affecting resolution & efficiency:

→ liquid-gas interface instabilities - spontaneous electron emission, gas gap variations (tilt), electron extraction efficiency into liquid.



Low energy deposits → need multiplication & efficient photon detection

Limited avalanche gain in noble gases → electroluminescence multiplication

- **Single-phase detectors:**

Current limit: No multiplication in liquid.

→ OK only for **large energy deposits**.

Stable primary-charges multiplication & efficient detection in both configurations
→ lowering detection thresholds (& cost?)

Single-phase TPCs

Ultimate goal:

Combined efficient detection of both ionization electrons and VUV photons

- Advantages

- No liquid-gas interface
 - Reduced instabilities (interface ripples)
 - No delayed e^- emission or e^- transfer efficiency through interface
 - No gate-interface-anode alignment problems
 - Horizontal drift \rightarrow sporadic bubbling not a concern
 - Potential improvement to the ionization e^- 's energy resolution
- Different geometries possible
 - Radial TPC
 - Symmetric central cathode TPC \rightarrow lower voltages needed

- Challenges

- High EL and CM thresholds \rightarrow amplification requires high electric fields

Single-phase - from wires to microstructures

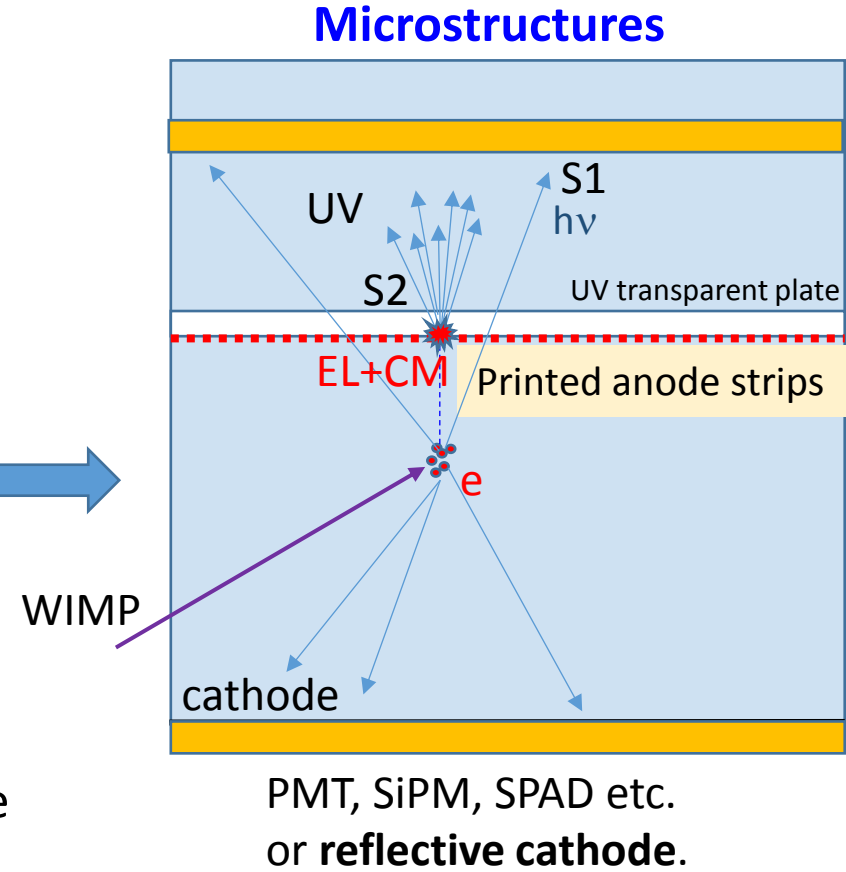
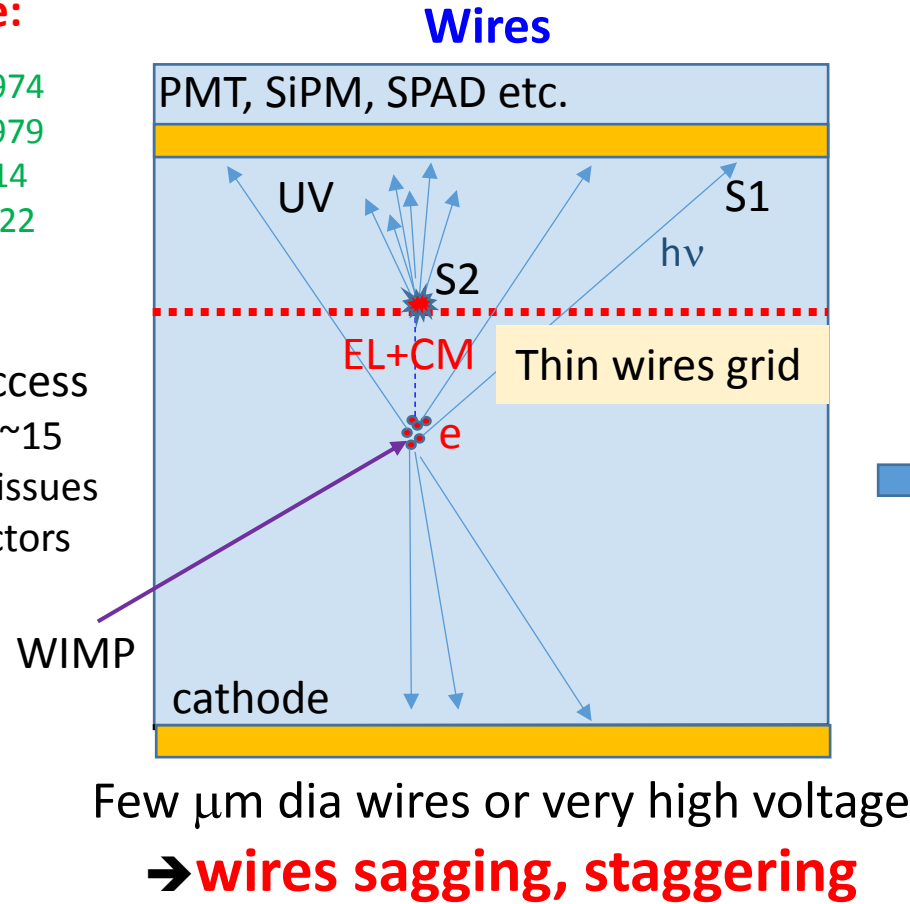
Examples: here multipliers sense **S2-e⁻ ONLY**
 Scintillation **S1-photons** recorded by photosensors

~50 years of Wires in LXe:

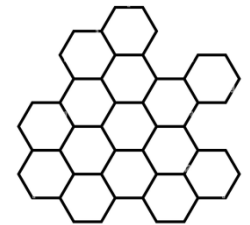
- Derenzo PRA 1974
- Masuda NIM 1979
- Aprile JINST 2014
- Brown JINST 2022

So far:

- No real success
- Charge gain ~15
- Mechanical issues @ large detectors



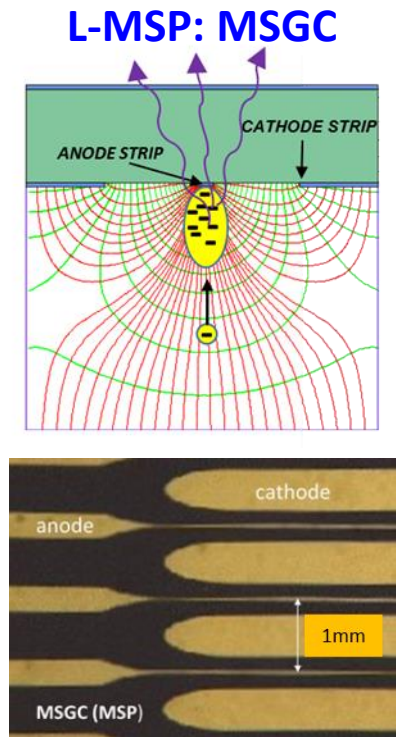
Micro Strip Plates & more!



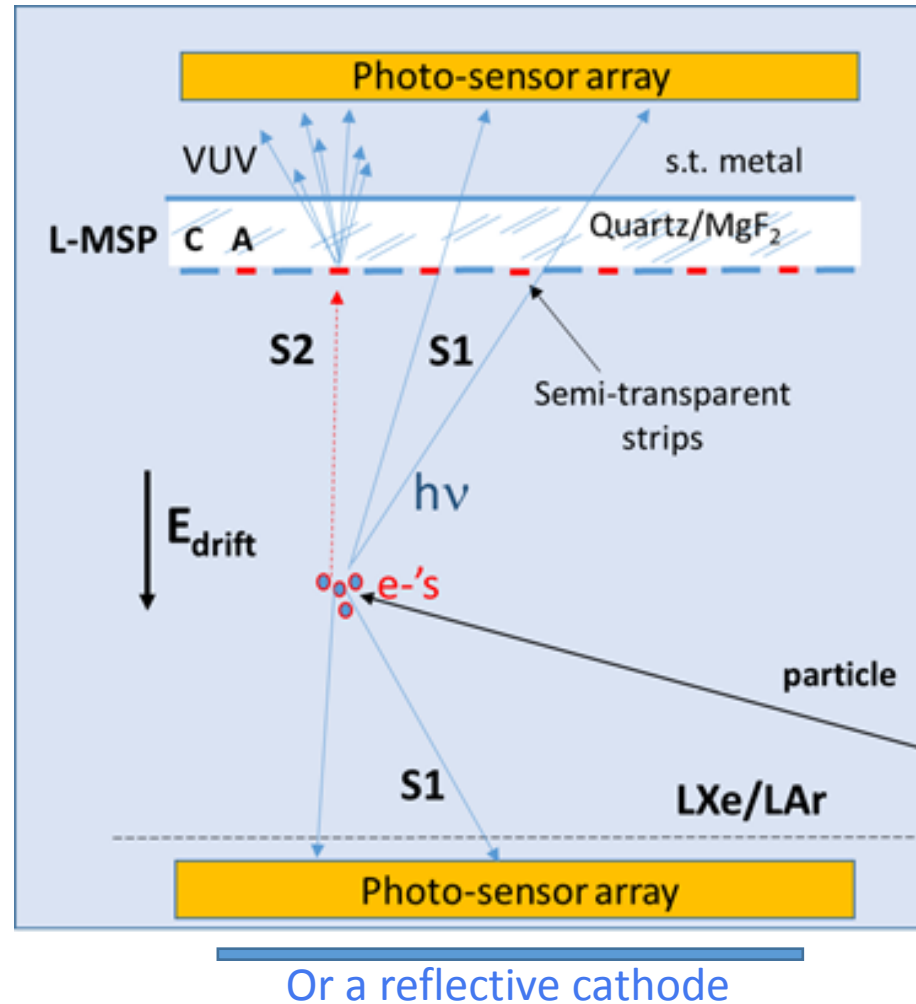
Robust Modular

Goals: Devise other robust high-gain solutions for detecting both **S2-e⁻ & S1-photons**

Single-phase with Micro Strip Plates (MSP): S2 e⁻ only



A. Oed 1988



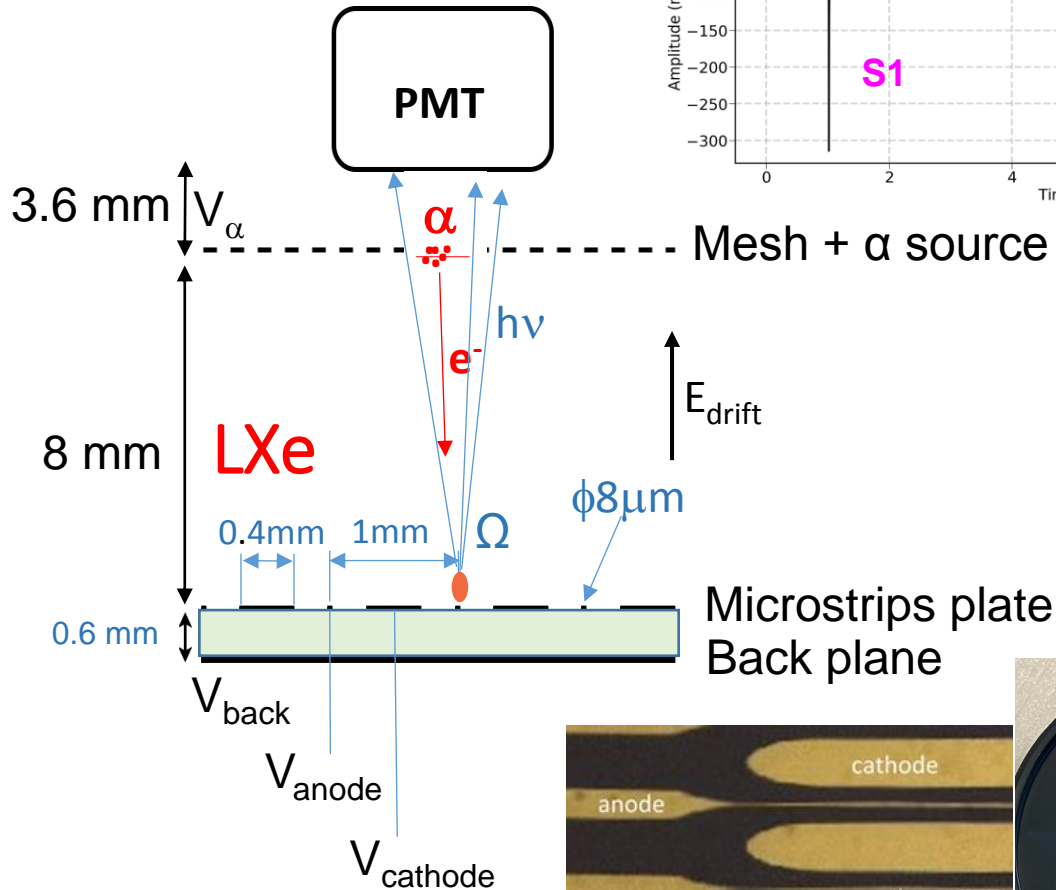
- MSP formed on VUV-transparent substrate, with semi-transparent Ni or Cr electrodes.
- Deposited charges drift in liquid; undergo **EL & small charge multiplication (CM)** near anode strips.
- The EL photo-yield depends on MSP type.
- EL Photons recorded above.
- S1 scintillation photons: with top & bottom photo-sensors (or reflective cathode)

Policarpo, Chepel 1995
x10 electron multiplication
 in LXe with MSGC;
 No EL photons recorded

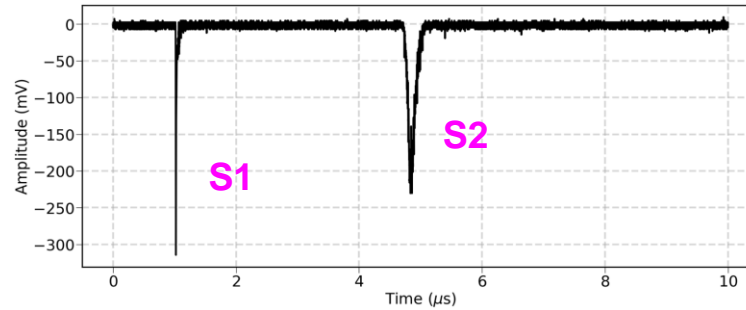
Microstrip Plate - Preliminary results in LXe

Paper underway

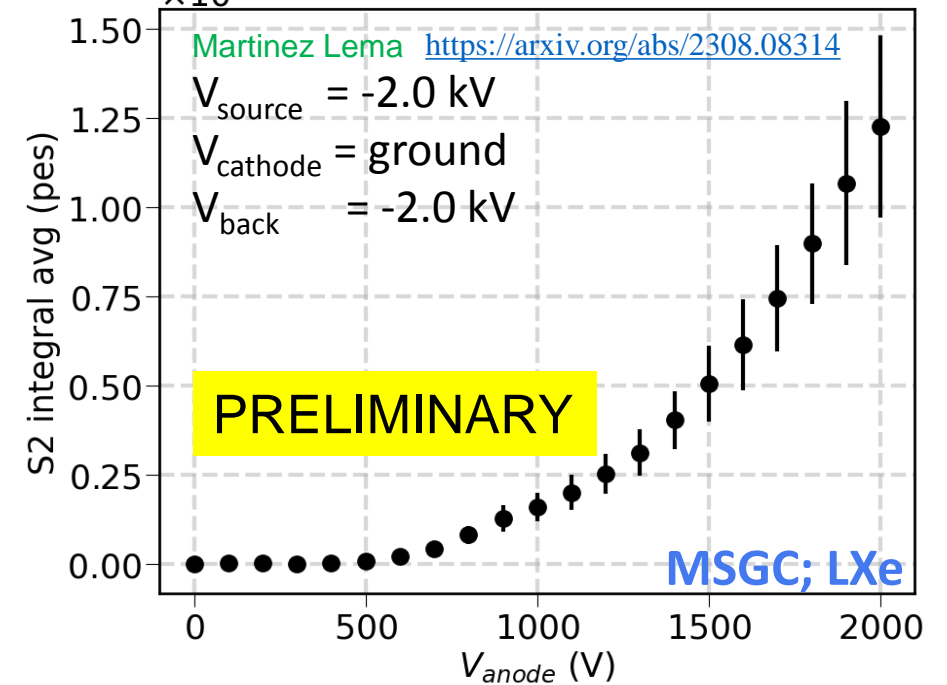
Setup



Single event

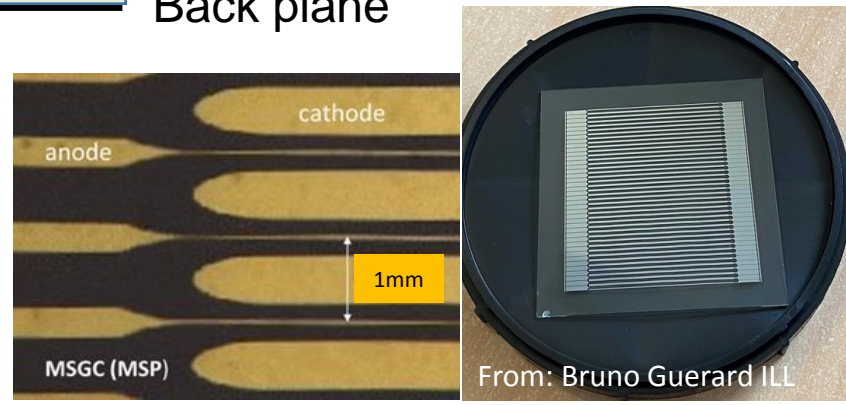


Electroluminescence vs strip voltage



At present discharge-limited **HV=2kV**:
 Charge gain **~3**
 Photo-yield **~ 33 photons/e-**

Much higher Q-gains/photoyields expected with other multipliers @ higher V_{anode}



Prospects: MSGC vs COCA-COLA vs VCC

For comparison:

wires in LXe:

Aprile 2014

10 μ m wires / 6.75 kV

~ charge gain x14

~ 290 photons/e-

MSGC has the best field configuration

- But operation is limited to ~2 kV a-to-c (~33 photons/ie)

COCA-COLA & VCC can operate at higher V without discharges

higher e⁻ multiplication and light yields

COCA-COLA COated CAthode - COnductive LAyer potential

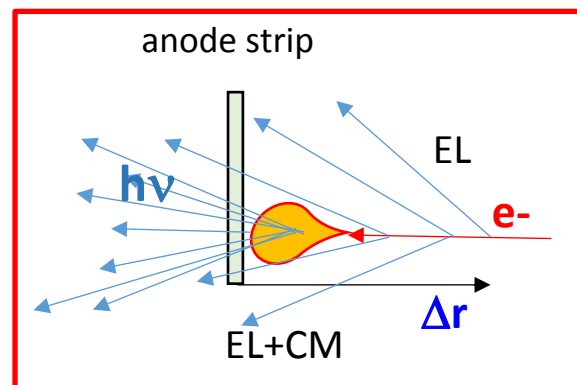
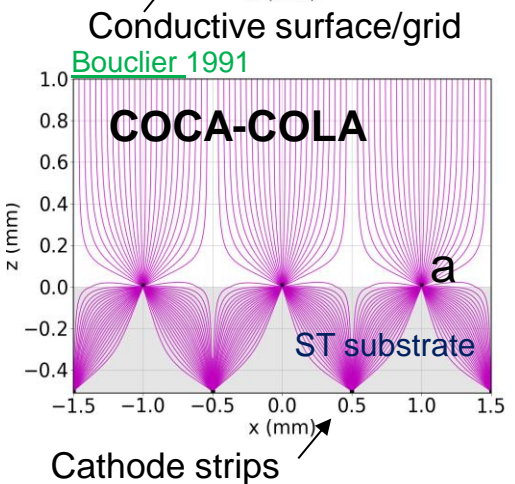
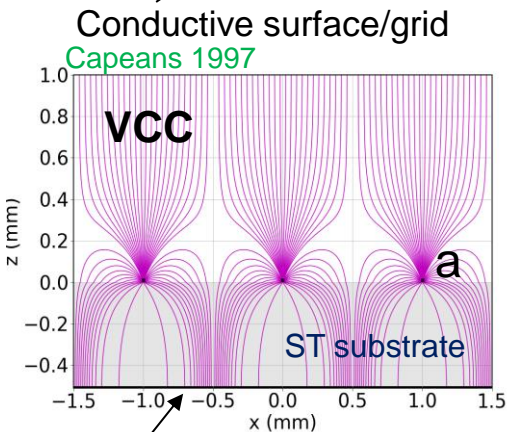
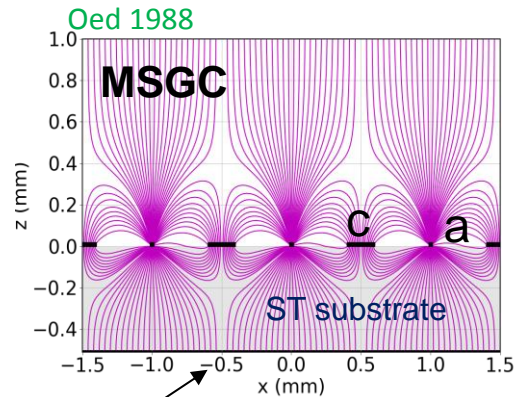
~x14 e⁻ multiplication @5kV

~525 photons / ie

VCC Virtual Cathode Chamber potential

~x40 e⁻ multiplication

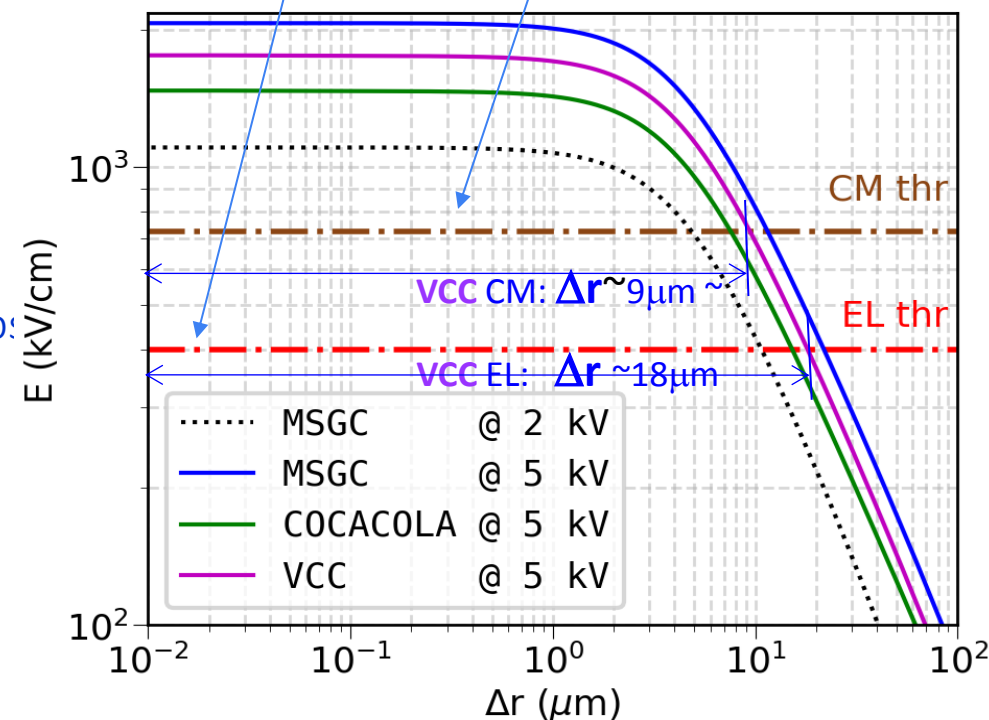
~1670 photons / ie @5kV (8 μ m strips)



Δr increases at lower strip widths

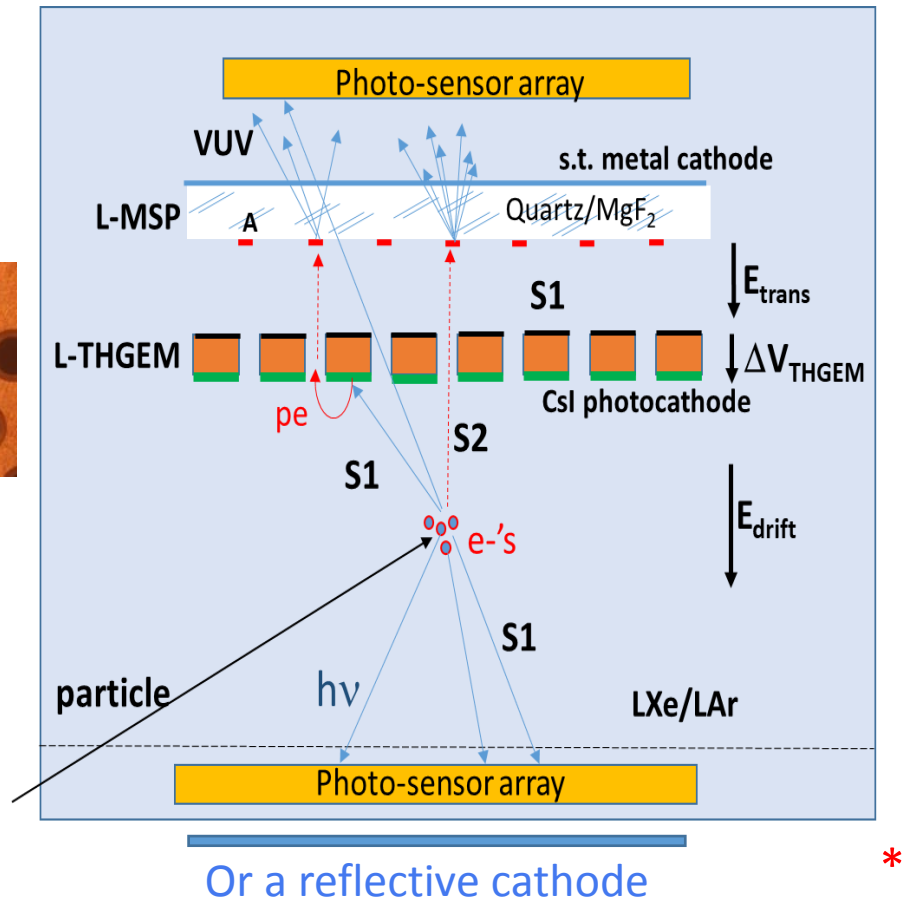
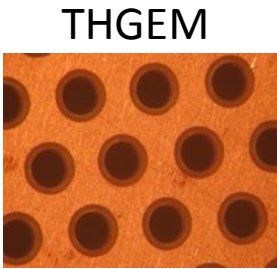
THRESHOLDS: Aprile JINST 2014

EL, ~400KV/cm; CM, ~700KV/cm



- 1-10 nm thick, 8 μ m wide anode strips
- COCACOLA cathode 200 μ m wide
- MSGC cathode 400 μ m wide

Single-phase with cascaded THGEM* + MSP → S1 hv & S2 e⁻



QE_{eff} of CsI in LXe @ E > 5kV/cm: 25%

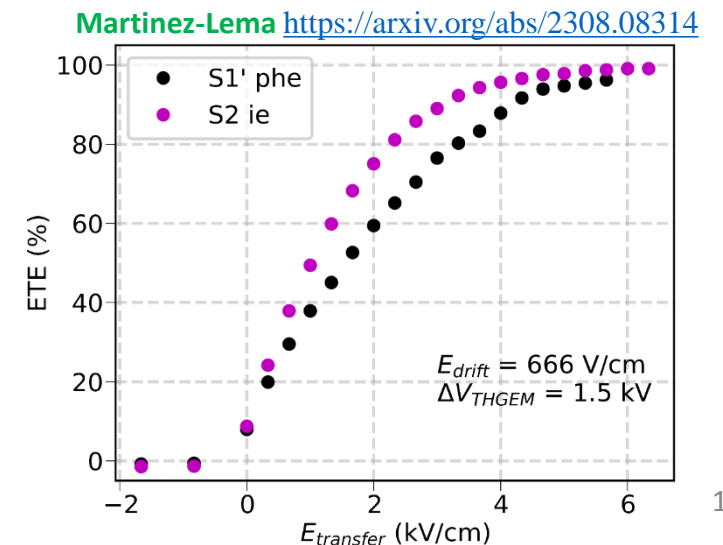
Erdal 2021

https://jinst.sissa.it/jinst/theses/2021_JINST_TH_002.jsp

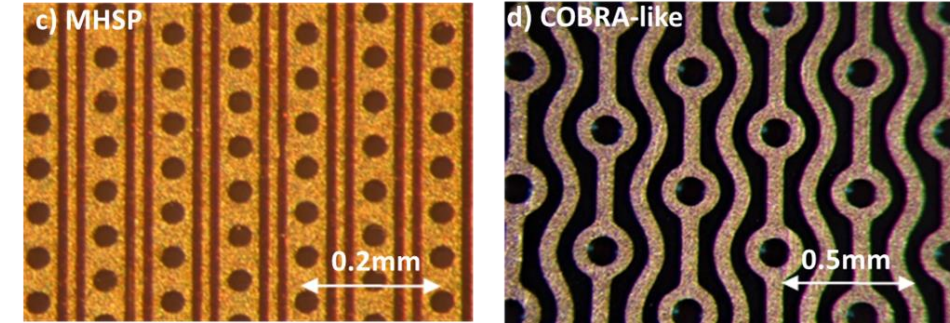
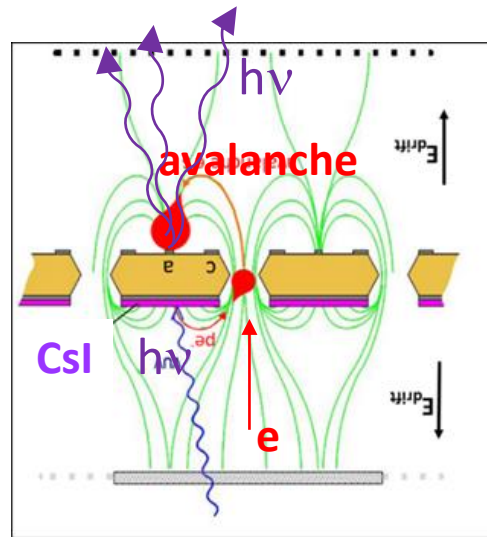
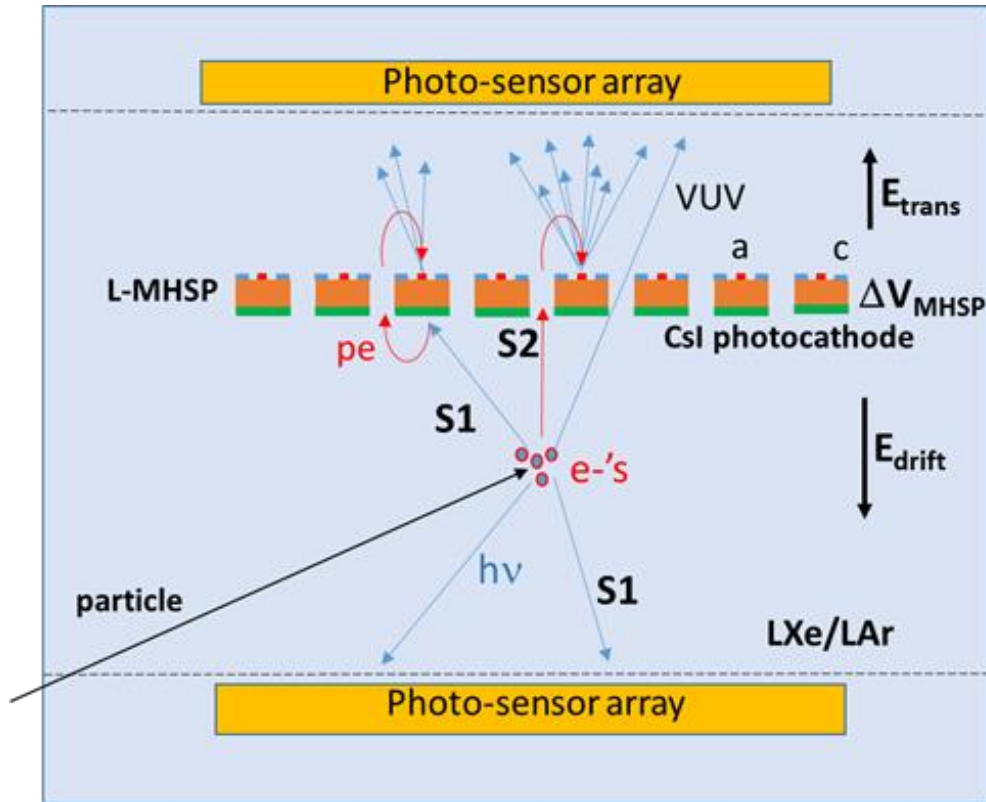
- 2-stage TPC with CsI-coated L-THEM + L-MSP; (here L-VCC with S.T. Cr\Ni strips on VUV- substrate)
- **S2 e⁻ & S1 UV-pe⁻** collected into L-THGEM holes & efficiently* transferred to the L-MSP.
- VUV photons emitted by **EL + small avalanche** near strips, detected through the substrate, by top photo-sensors.
- A fraction of S1 photons – detected by bottom photo-sensors or reflected by a reflective-cathode to the CsI.
- Option: top L-THGEM surface can be reflective or WLS-coated (→ visible-range photo-sensors, glass substrate).

*THGEM: Thick Gas Electron Multiplier

Experimental Transfer Efficiency of e⁻ and pe⁻ through THGEM holes in LXe.



Single-phase with single-element Micro Hole & Strip Plate (MHSP)

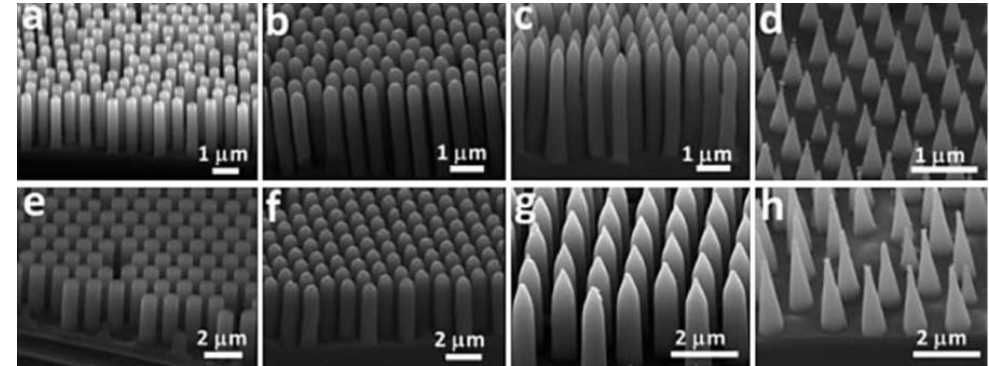
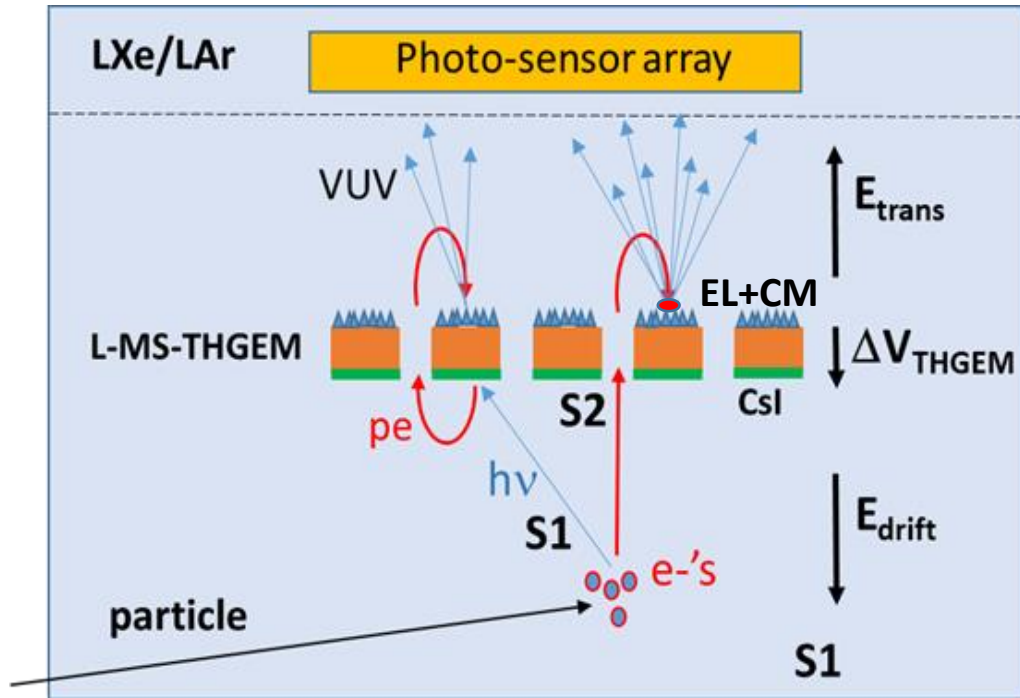


MHSP: Veloso, Rev Sci. Instr. 71, 2371 (2000)
<https://doi.org/10.1063/1.1150623>

Current technologies → few- μm strips possible

- A single-phase TPC with CsI-coated **L-MHSP**.
- Both S2 e- & VUV photoelectrons collected into the L-MHSP holes & collected by MHSP anode strips.
- **ALL** VUV photons (EL + CM): detected by the top photo-sensors.
- Other fraction of S1 photons - detected by bottom photo-sensors.

Single-phase with Micro-structured electrode



Hao Lin <https://doi.org/10.1039/C3TA11889D>

- Single-phase TPC with micro-structured THGEM top surface (**L-MS-THGEM**), under-coated with CsI.
- Both S2 e- & S1 VUV pe- are collected into the holes towards the micro-structured top surface
- VUV photons emitted by EL + CM at the “anode tips”, under high E, are detected by top photo-sensors.
- Other fraction of S1 photons are detected by bottom photo-sensors (not shown).

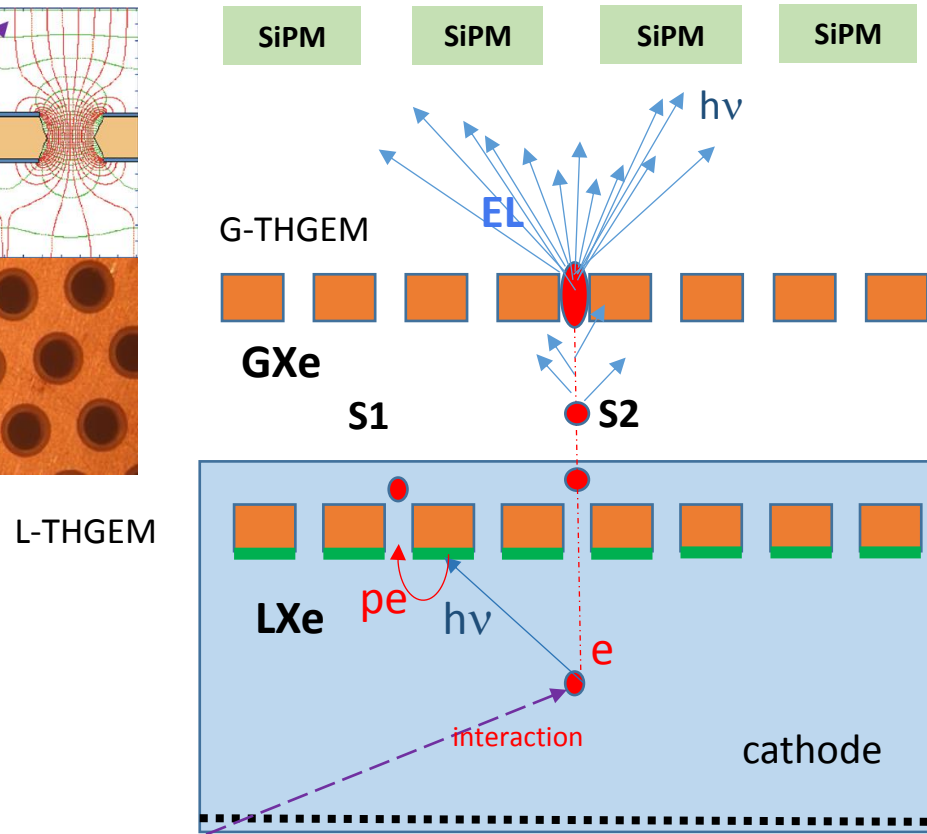
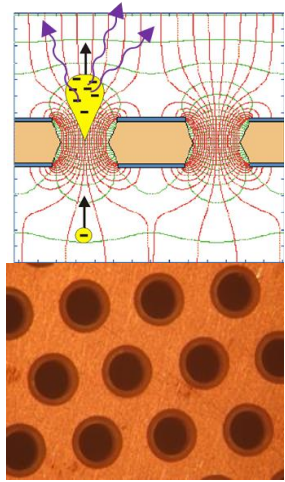
Novel Dual-Phase TPC Concepts with reduced interface instabilities

In case Single-Phase ones remain a dream...

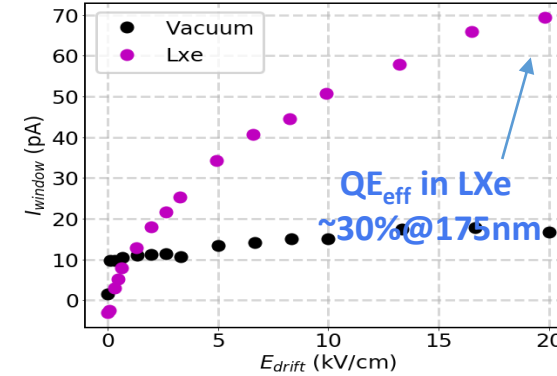
Dual-Phase: Cascaded Liquid Hole Multiplier (LHM) – S1 & S2

Martinez Lema 2022

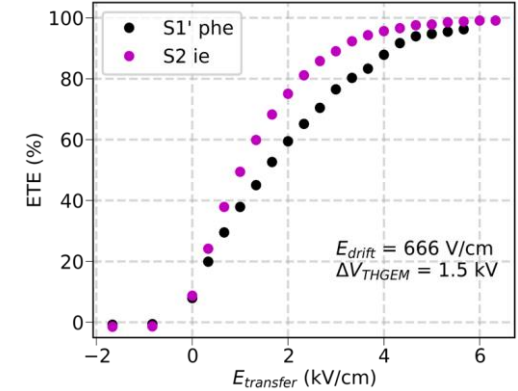
The multiplier: Perforated electrode with thin strips



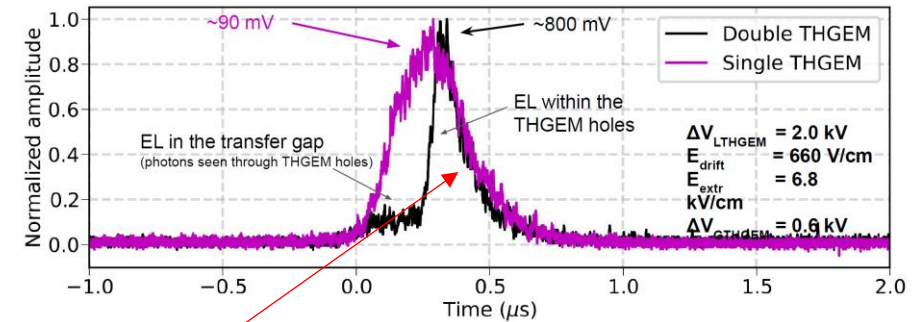
Martinez-Lema <https://arxiv.org/abs/2308.08314>



LXe: increased QE_{eff} by ~ 3.5 compared to vacuum! (Work function)



e- & pe- transfer efficiency through THGEM immersed in LXe



EL in holes is faster & does not depend on interface-gap variations

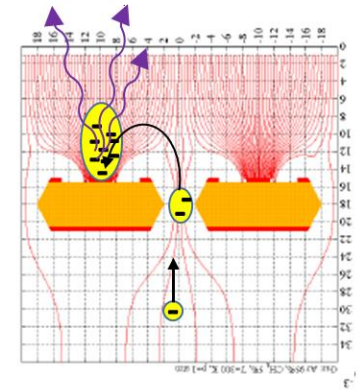
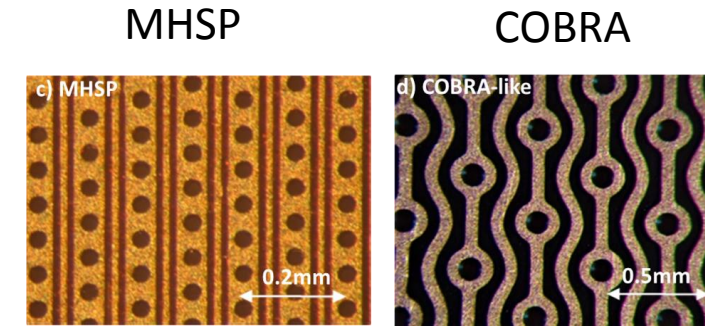
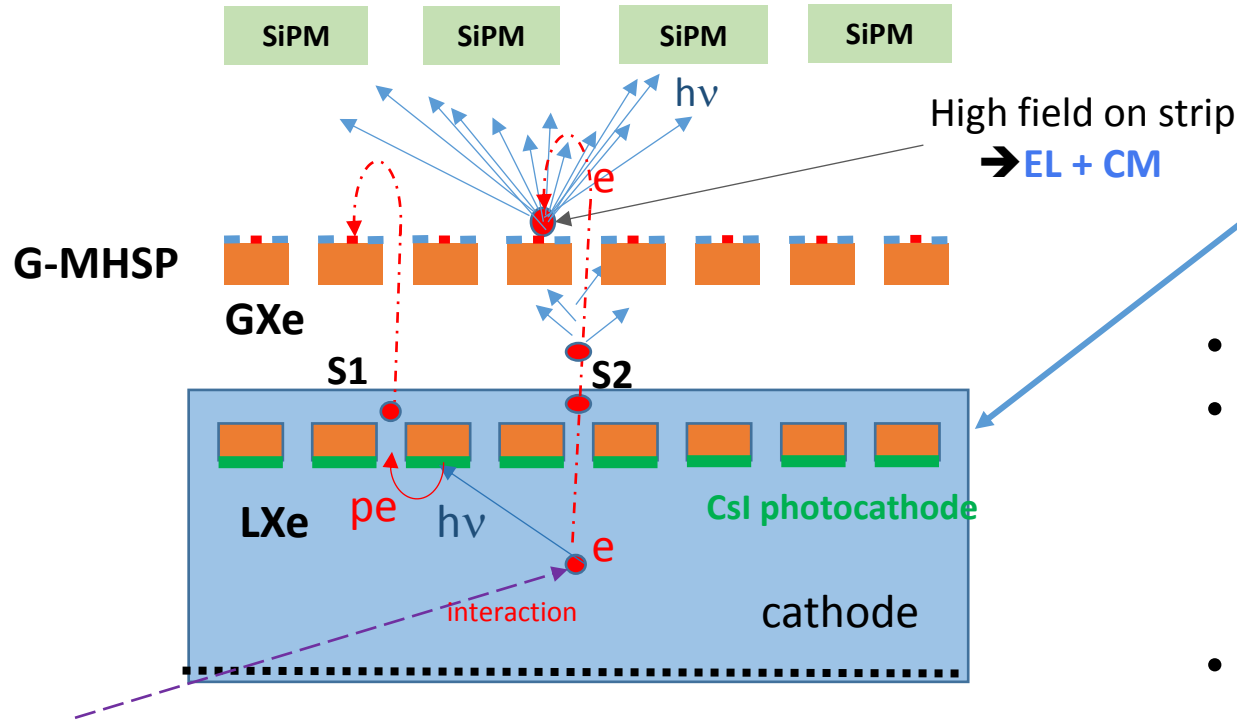
PRELIMINARY photoyield:
 $\sim 350 \text{ photons/e} \backslash 4\pi$

- **S2** e- and fraction of **S1**-induced pe- extracted efficiently **in liquid** through holes of a **CsI-coated** immersed L-THGEM electrode.
- They traverse the liquid-gas interface inducing EL **in gas phase**, mostly in G-THGEM holes.
- Readout by nearby photo-sensors. Bottom cathode: reflective, or photo-sensors underneath.

Dual-Phase: Cascaded Liquid Hole Multiplier (LHM) - S1 & S2

The multiplier: Perforated electrode with thin strips

Similar operation as with G-THGEM,
but EL photons emitted from strips



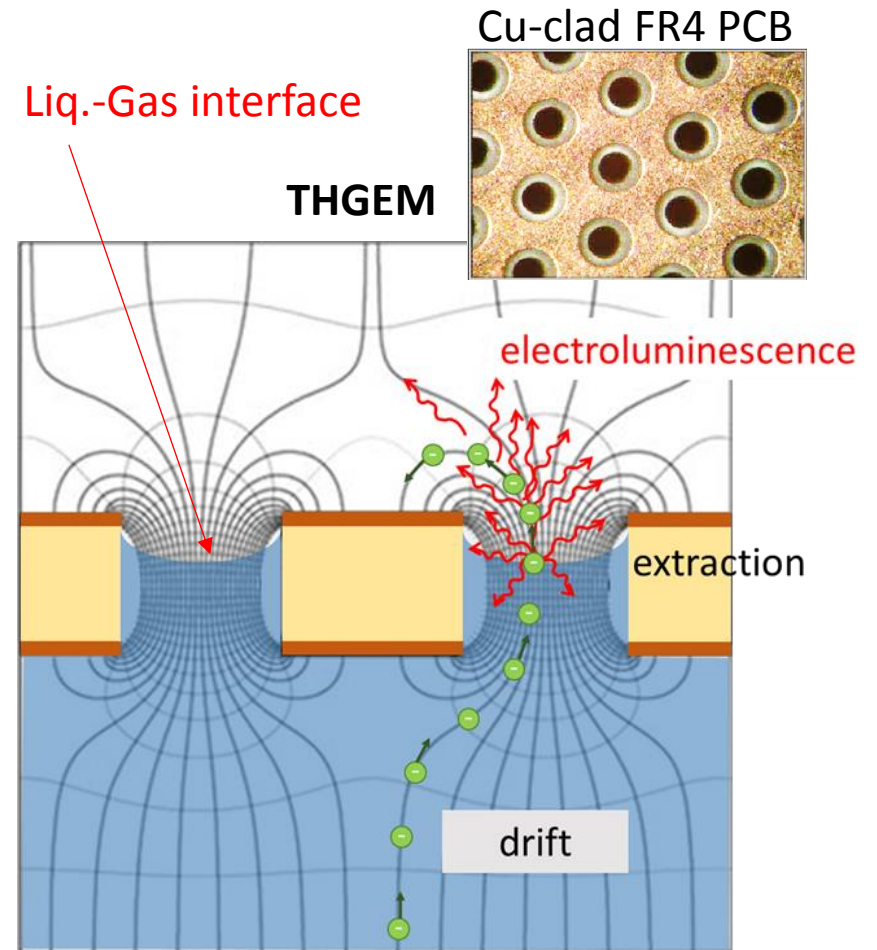
Current technologies → few- μm strips possible

- **High CsI QE** → expected high PDE
- **Detection of gas-amplified single photons above dark current**
→ Possible to use SiPM, CMOS-SPAD etc.
→ Lower detection threshold
- **Transmitted charges detected in G-THGEM holes:**
→ Interface fluctuations not critical
→ Faster EL signals

EXPECTED: higher field @strips → larger photon yield & faster signals

Dual-Phase: The Floating Hole Multiplier FHM

- Better surface-effects control by separation of the two phases;
- A “robust” interface: perforated electrode **FLOATING** on the liquid surface → e.g. **GEM, THGEM, COBRA**, etc
- Significantly smaller free liquid surface → **reduced surface instabilities**
- High E @ interface → High electron extraction probability
→ **Result – reduced single-electron noise**
- No need for fine detector levelling and liquid level control



LXe density **2.9** g/cm³

FR4 density **2.0±0.2** g/cm³

If copper cladding is not too heavy → THGEM
should float on the surface of LXe

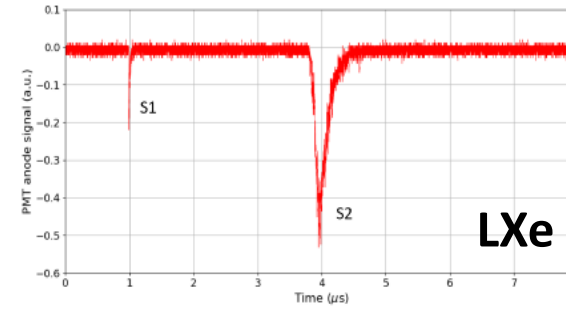
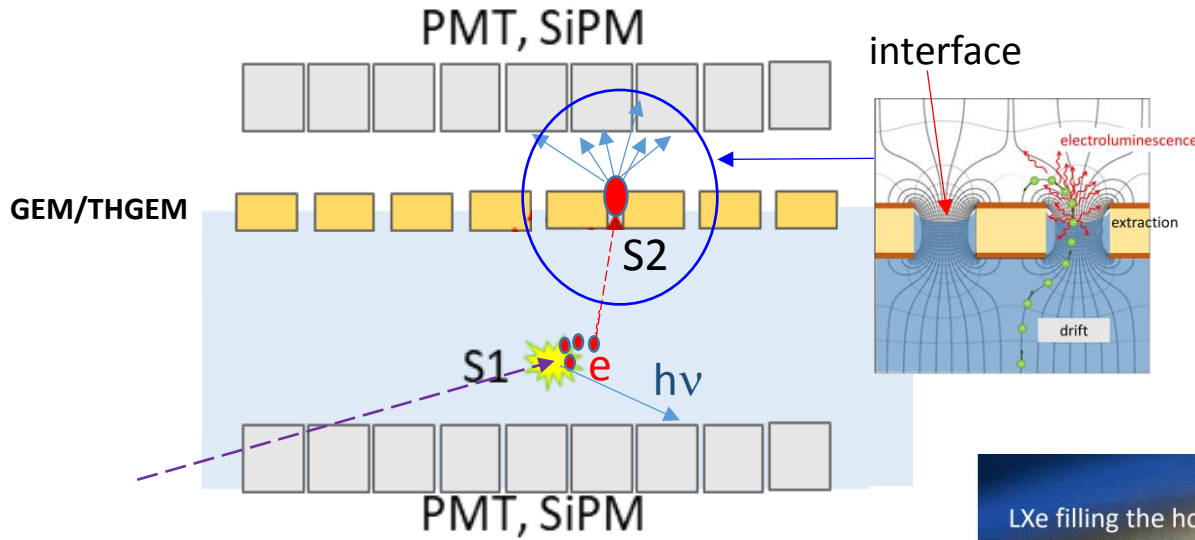


Idea V. Chepel

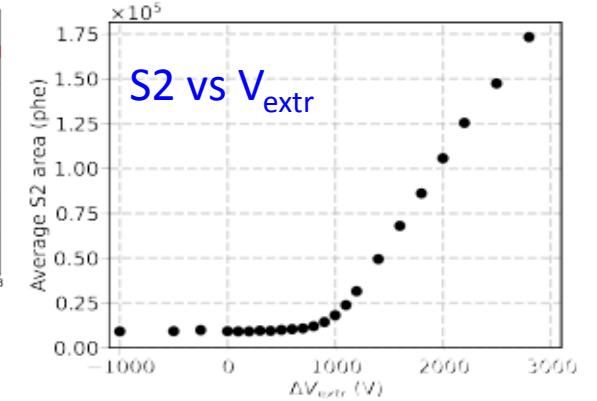
FHM – S2 e⁻ only

proof of FHM principle in LXe:

Chepel 2023 <http://arxiv.org/abs/2301.12990>



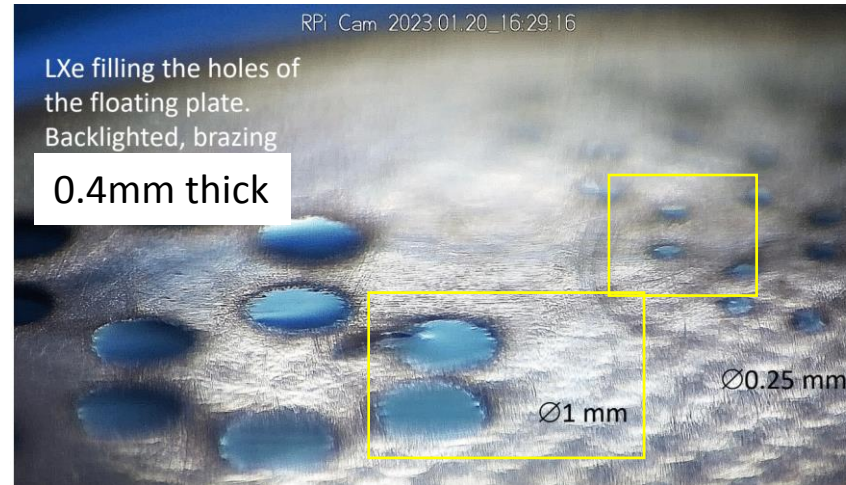
α-induced S1 & S2



PRELIMINARY:
 $Y \sim 500 \text{ photons/e} \backslash 4\pi$

- S2 electrons extracted efficiently through the interface within holes of a perforated electrode, **floating on the liquid**.
- They traverse the interface (under high fields) inducing EL in gas.
- Readout by nearby photo-sensors.

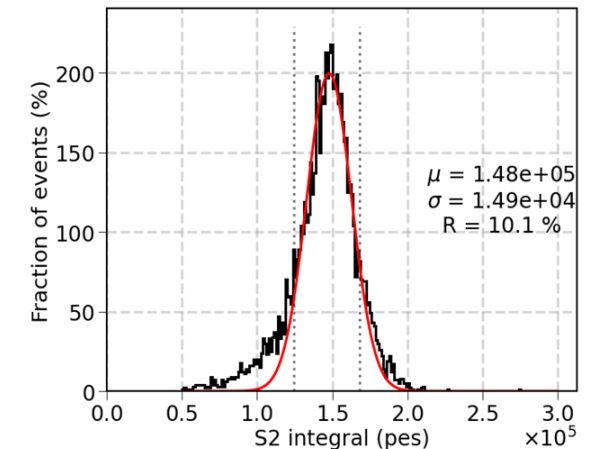
Reduced interface effects and instabilities



With 0.4mm thick electrode floating on LXe, the holes are rather filled with the liquid.

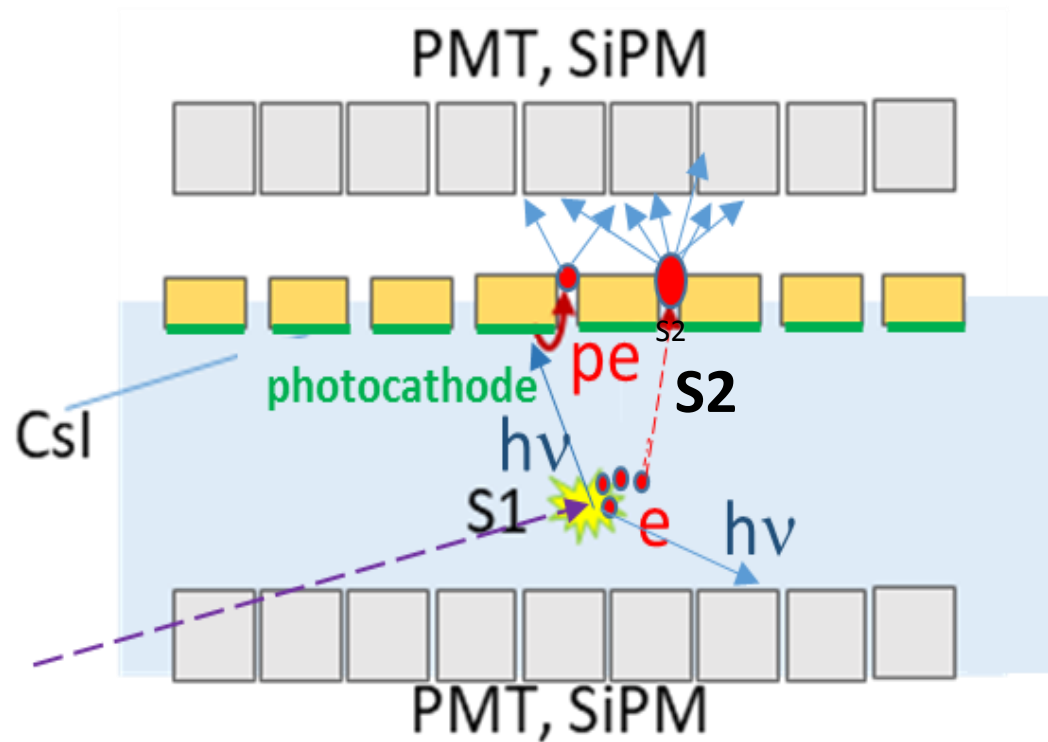
R&D: Thicker electrode:

➔ expected higher photon yields.



α-induced S2
 Energy resolution

FHM – S1 photons & S2 e⁻



An optional **CsI-coated** perforated floating electrode allows for detecting both, a fraction of **S1 photons** (photoelectrons from CsI) and **S2 electrons**. Both are collected into the holes and traverse the interface (under high fields) inducing EL in gas. Readout by nearby photo-sensors. Under R&D.

Conclusions

- **Generic R&D:** novel methods for future applications.
- Goals - novel sensing concepts for noble-liquid detectors.
- Focus - combined high-sensitivity light & charge detection methods.
- Readout - Optical readout of EL & CM photons
- Preference – **Single-phase TPCs**. (novel 2-phase in case...)
- Aims - solving current physical & technical issues; lowering detection thresholds
→ potential applications in large-volume experiments.
- Main (but not sole) “users” - future DM experiments and neutrino physics.

Some proofs of principle; others in progress



Many open questions requiring simulations & exp. R&D (LAr, LXe)

Open to collaborations – seeking for students/postdocs!

Thank you!

Backup

Noble liquids

Element	Z	A	Liquid density (g/cc)	Boiling point (K)	Photon yield (γ /keV)	Triplet decay time	Emission Wavelength (nm)	Radio activity
He	2	4.00	0.13	4.2	22	13(s)	80	None
Ne	10	20.18	1.2	27.1	32	15(μ s)	78	None
Ar 	18	39.95	1.4	87.3	40	1.5(μ s)	128	³⁹ Ar 1Bq/kg
Kr	36	83.80	2.4	119.9	49	85(ns)	148	⁸⁵ Kr 1MBq/kg
Xe 	54	131.30	3.1	165.0	64	22(ns)	175	¹³⁶ Xe < 10uBq/kg

Nikkel 2012

XENON: - higher density, operation T, photo-yield, emission WL + faster but **HIGH COST**
- smaller detector volumes.

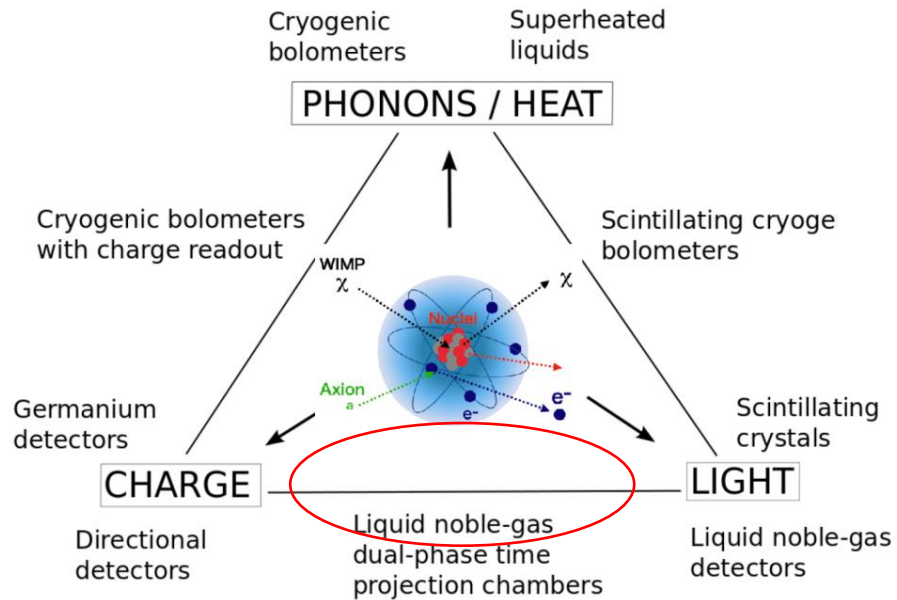
Neutrino: liquid ARGON LAr

DM: mostly liquid XENON LXe

Unsolved problems in physics 1 – Dark matter

- Galaxies do not appear to obey current laws of physics.
Stars' velocity in galaxies > expected by Newtonian mechanics.
- Is there **Dark Matter (DM)**? What is its nature? Is it a **particle**? or:
- Do the phenomena require a **modification of the laws of gravity**?
MOND? - MODified Newtonian Dynamics **Milgrom 1982**

Several DM detection techniques:



J. Phys. G43 (2016) 1, 013001& arXiv:1509.08767

Dark matter candidates:

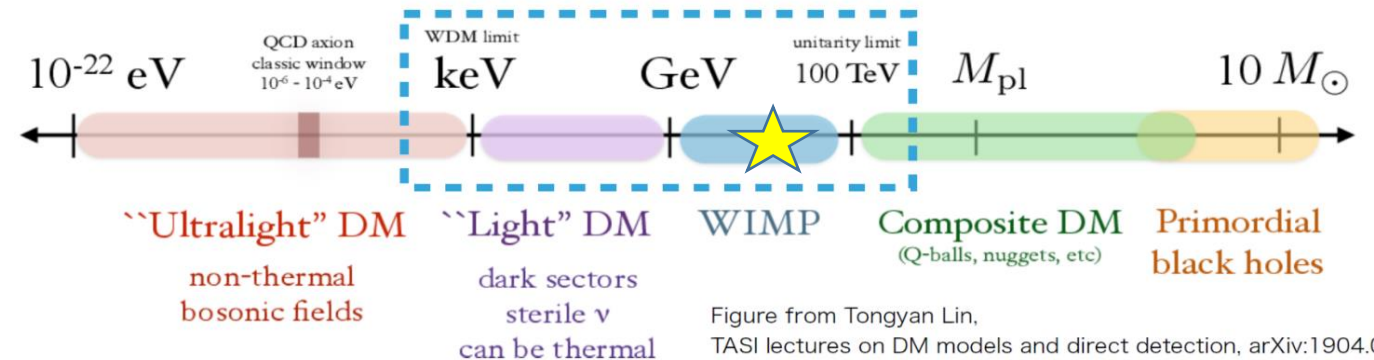


Figure from Tongyan Lin, TASI lectures on DM models and direct detection, arXiv:1904.07915

- ~100 orders of magnitude in mass.
- A variety of detectors related to mass range.
- Direct-detection of **WIMPs**: with **noble-liquid detectors**:
- Many experiments

Unsolved problems in physics 2 - Neutrino

Neutrinos: Fundamental particles detected >70Y ago.

Several open questions:

- Do neutrinos and anti-neutrinos oscillate differently (CP violation)?
- Their mass order? (mass hierarchy)



- Other neutrino types? Interactions?
- Are neutrinos their own anti-particles? (***Majorna vs Dirac***)
- Neutrino Masses
-

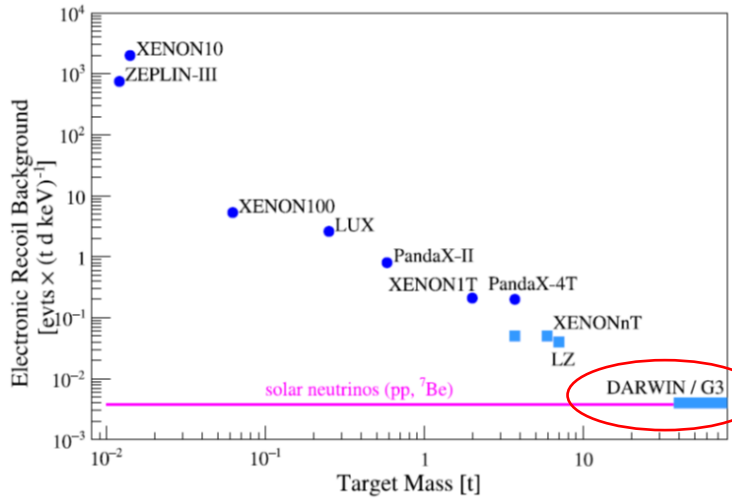
Long baseline exp.

Short baseline exp.

$0\nu\beta\beta$

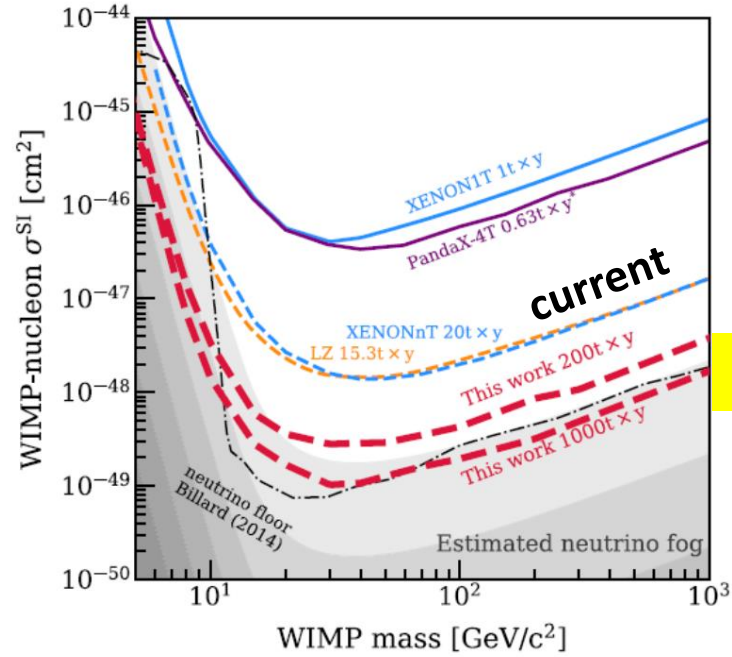
Many experiments!

DARWIN/G3 50t LXe observatory for DM & neutrino physics



DARWIN/G3

Low background \rightarrow



\leftarrow 50t x 20Y

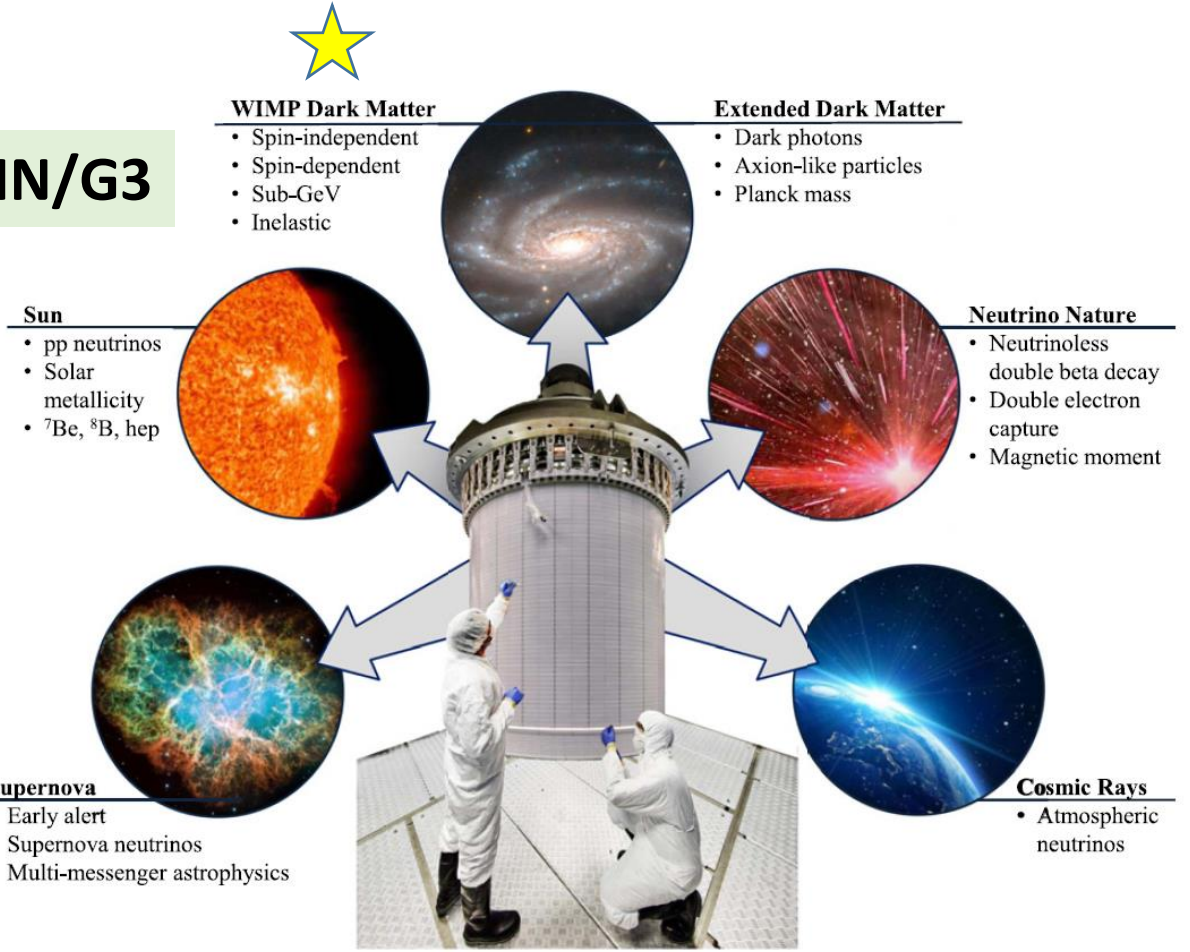


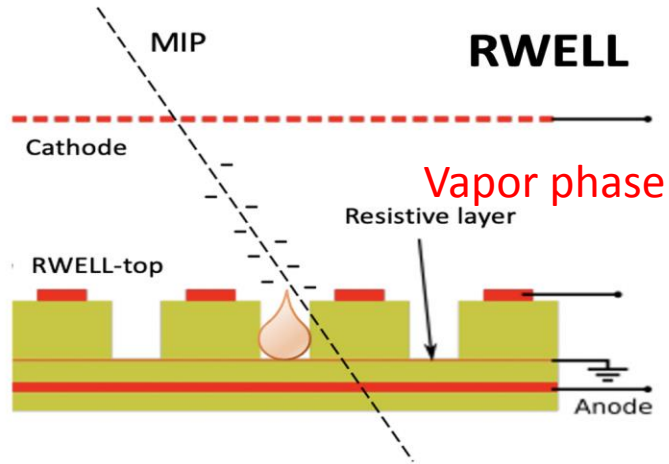
Figure 2. The science channels of a next-generation liquid xenon observatory for rare events spans many areas and is of interest to particle physics, nuclear physics, astrophysics, solar physics, and cosmology.

Aalbers et al 2023 J. Phys. G: Nucl. Part. Phys. 50 013001
DARWIN/G3

Cryogenic charge multipliers with resistive electrodes

Motivation: Charge multiplication in noble gases limited by discharges due to secondary effects to **<10 (DUNE)**

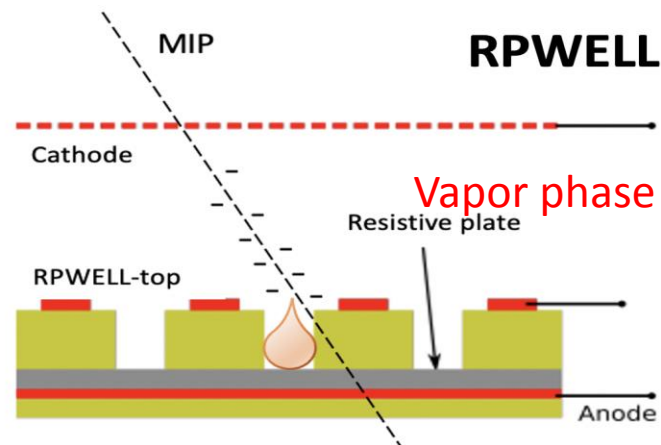
Goal: Quench harmful discharges by deploying **resistive materials** in the gas-avalanche multiplier



- THGEM electrode, coupled to readout anode via **resistive electrode**.
- Deposited charges collected & undergo avalanche multiplication in the holes.

Resistive WELL (RWELL)

- **Resistive film**, e.g. Diamond-Like-Carbon (DLC), on an insulator.
- Charges evacuated sideways to ground via resistive layer
- Signals are induced on readout anode by charges movement

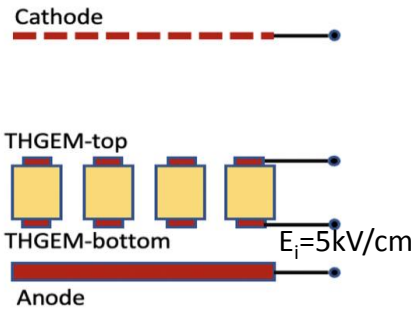


Resistive-Plate WELL (RPWELL)

- **resistive plate**, e.g. a ceramic material Yttria-Stabilized-Zirconia (YSZ) doped with ferrite oxide (Fe_2O_3), to readout anode.
- Amplified charge travel through the resistive plate to the anode.

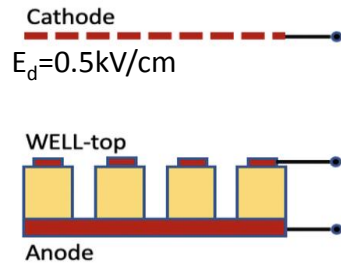
challenges: find/develop resistive materials of the “right” surface/bulk resistivity at noble-liquid temperature

Comparison: THGEM/WELL & RWELL/RPWELL @ 90K

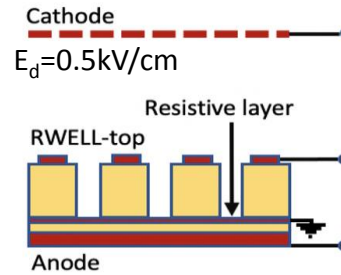


THGEM

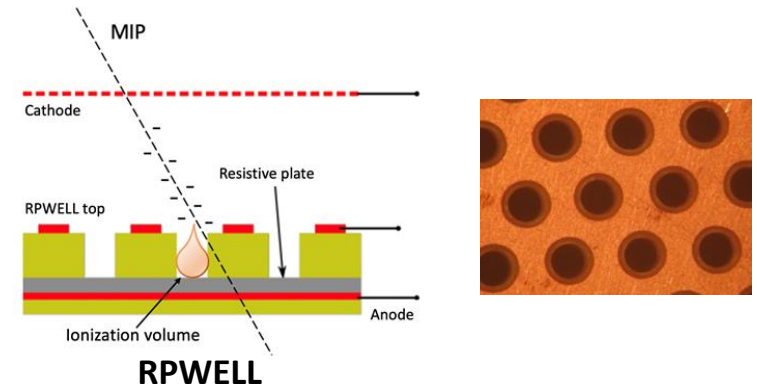
DUNE LEM-like



WELL

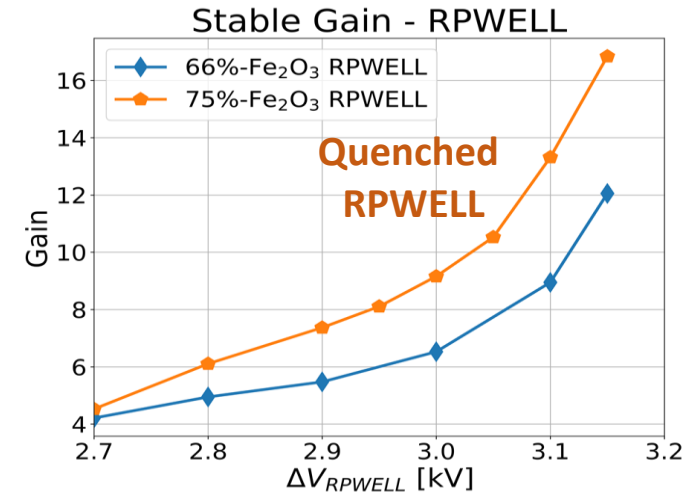
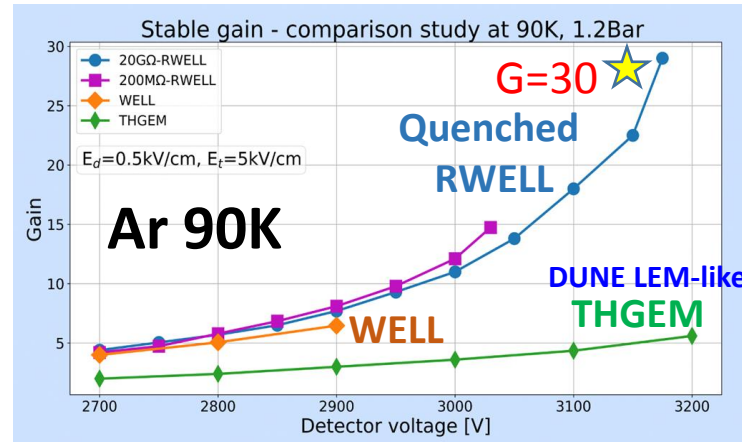


RWELL ★



RPWELL

Optimal resistivities:
 temperatures LAr ~87K
 Resistive FILM: ~ 20GΩ/□
 Resistive PLATE: ~ 10⁹-10¹²Ω·cm



Tesi

<https://doi.org/10.1088/1748-0221/18/06/C06017>

<https://arxiv.org/abs/2307.02343>

So far, with experiments performed under equal conditions:

RWELL/DLC gain > THGEM (LEM) gain;

RWELL gain > RPWELL gain (but near zero discharges only @ gain < 17)