

UON Collider Collaboration





Fast timing detectors for the muon system of a muon collider experiment: requirements from simulation and prototype performance

16th Topical Seminar on Innovative Particle and Radiation Detectors 25-29 September 2023 Siena, Italy

Ilaria Vai

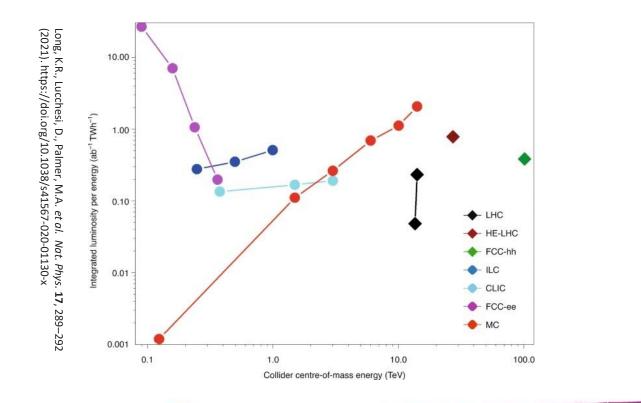
on behalf of the International Muon Collider Collaboration

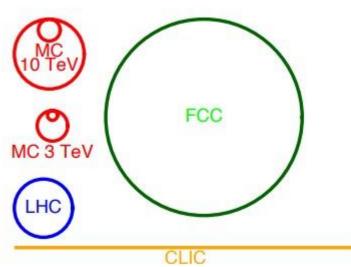


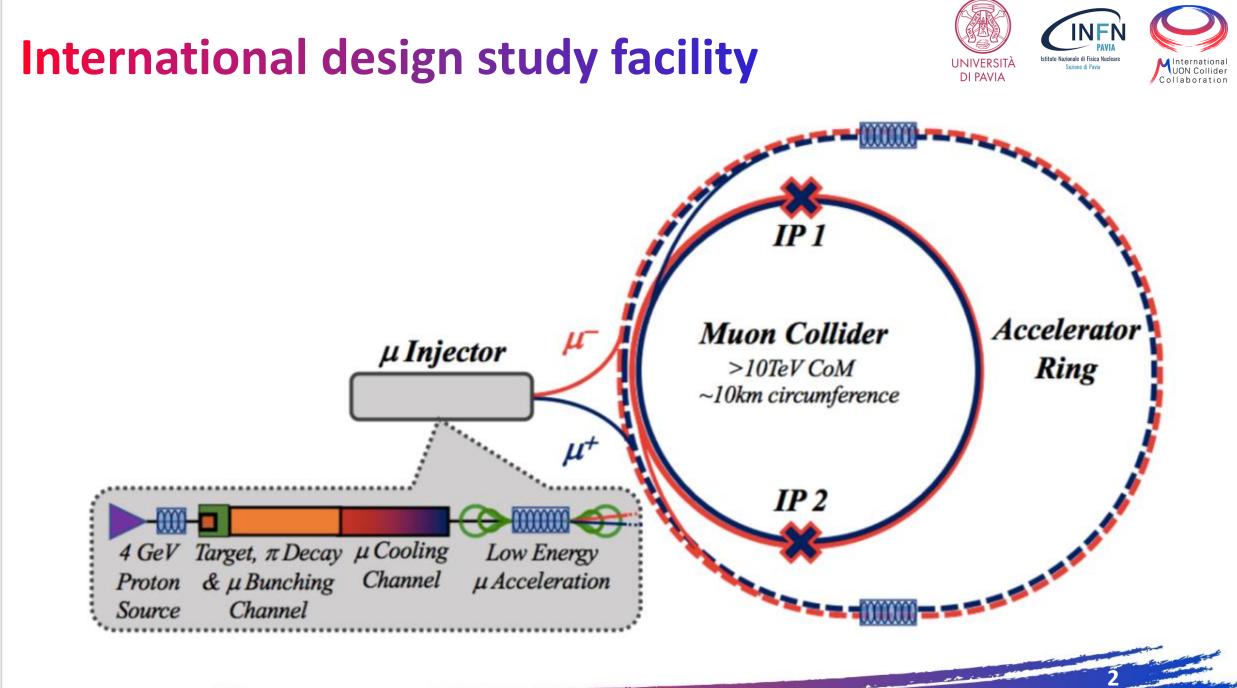


A Muon Collider is a unique possibility of combining high energy with very precise measurements

- pp machines \rightarrow no limitations from synchrotron radiation
- e+e- machines \rightarrow nominal E_{cm} entirely available





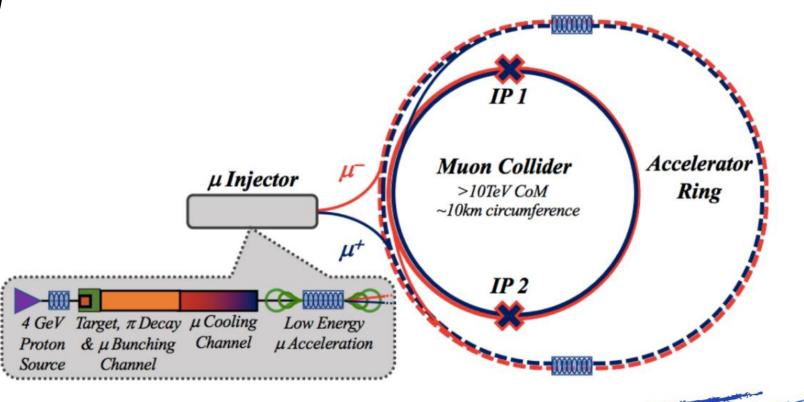


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Main challenges



- *Muon beams*: need to produce large numbers of muons in groups with small emittance
- Beam-Induced-Background



PRD2023

Beam-Induced-Background



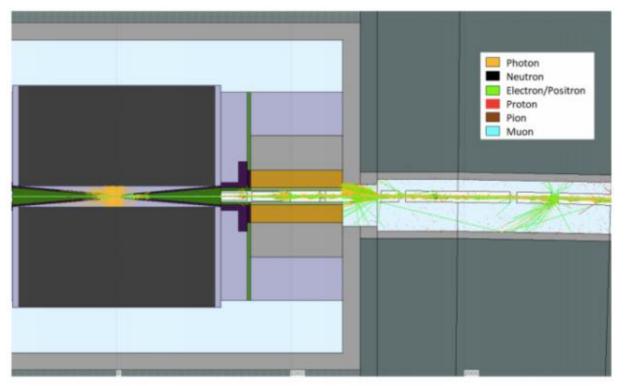
$\mu^- \to e^- \overline{\nu_e} \nu_\mu$

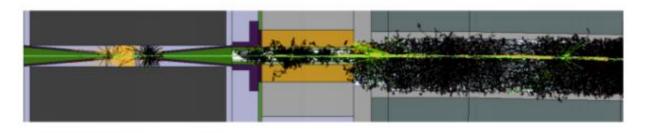
Secondary particles are produced from interaction with machine components:

- charged hadrons
- neutral hadrons
- Bethe-Heitler muons
- electrons
- photons

BIB may affect detector performance:

- shielding nozzles
- optimization of detector design

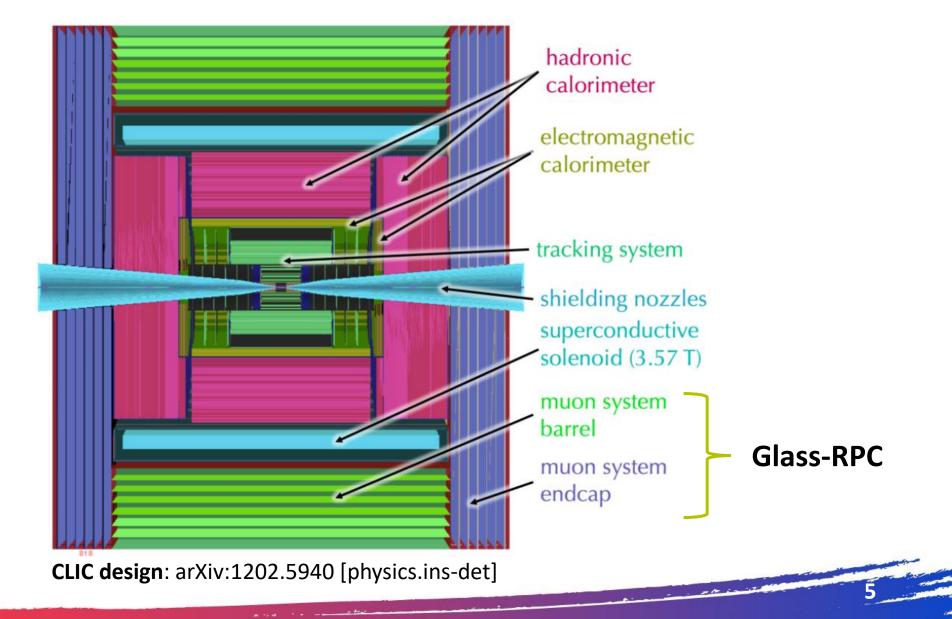




2021 JINST 15 P11009

Initial detector design

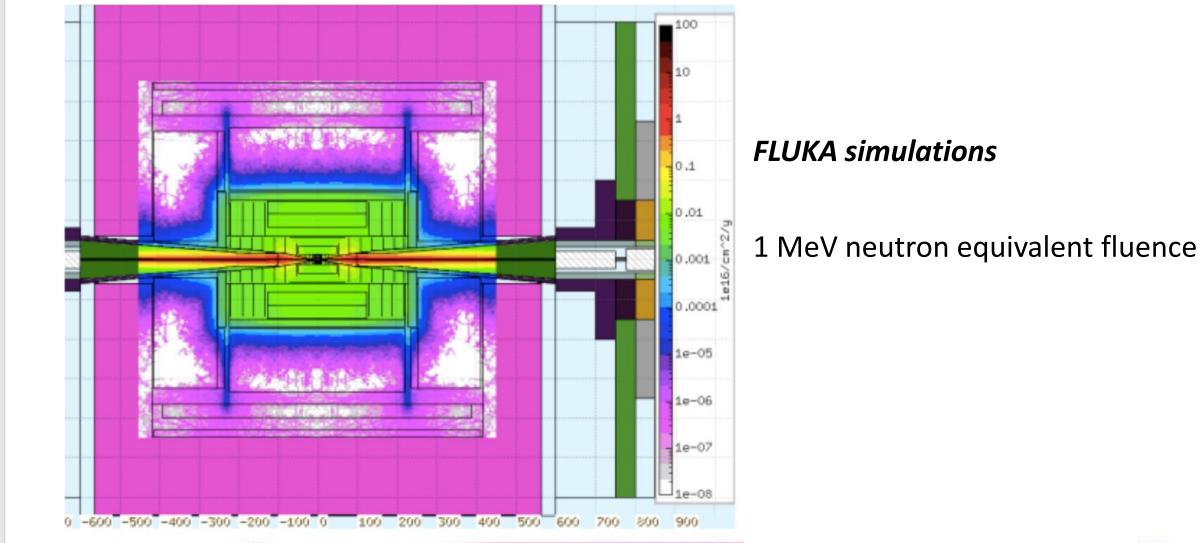




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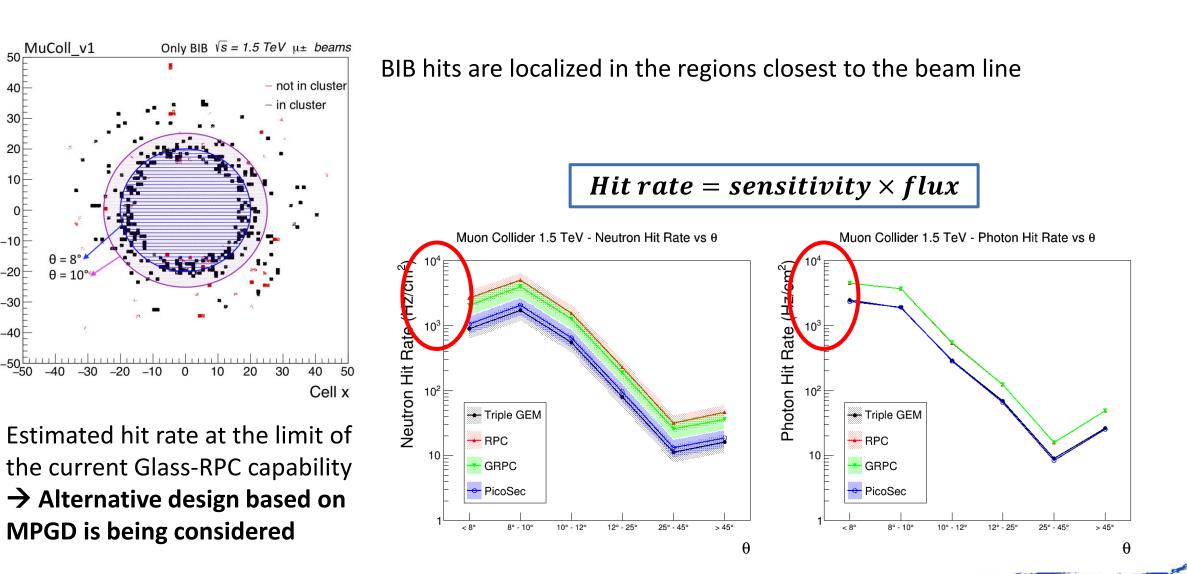
Beam-Induced-Background





BIB in the muon system





40

30

20

10F

-10

-20

-30

-40

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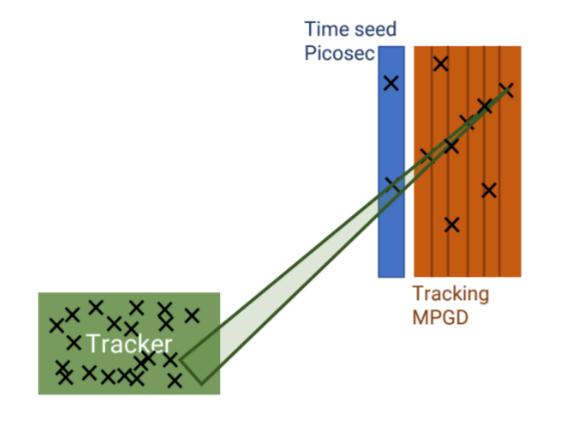
PRD202

Alternative muon system design



Out-to-In muon tracking approach is currently being considered: reconstruct muons in the muon system and then propagate back to the tracker \rightarrow reduce combinatorial background

We need a time resolution comparable to the tracker one \rightarrow ~100 ps



Alternative muon system design:

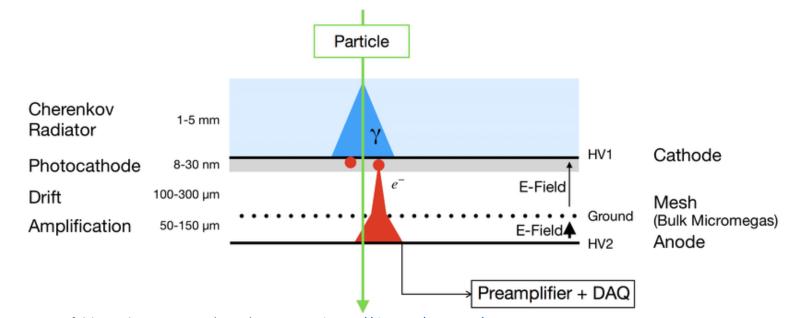
- Timing layer \rightarrow <u>Picosec</u> detector
- Tracking based on classical MPGDs (eg. Triple-GEM)

Picosec Micromegas detector





PICOSEC Micromegas



J. Bortfeldt et al, NIM A 903 (2018) 317–325, https://doi.org/10.1016/j.nima.2018.04.033

Working principle

- Highly energy particles cross the Cherenkov radiator and produce Cherenkov photons
- Photons are converted by the **photocatode**
- Electrons enter the micromegas and are amplified

Why this idea?

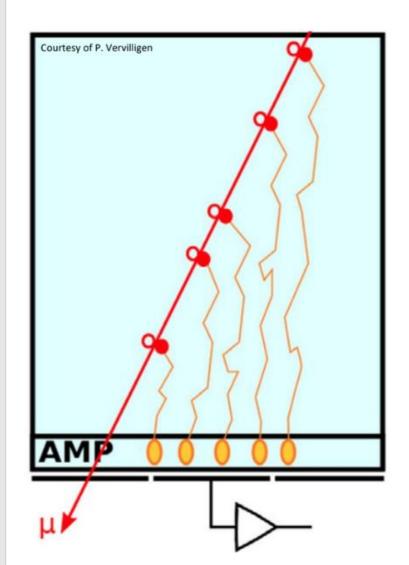
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Timing in MPGD



PICOSEC

Micromegas

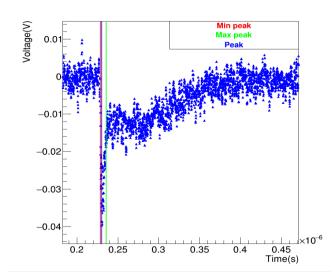


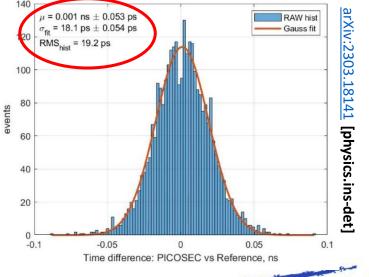
Time resolution is dominated by the position of the ionization cluster in the drift gap \rightarrow *large drift gap = large fluctuation = worse time resolution*

Classical MPGDs usually have $\sigma_t \sim few \ ns$

Picosec advantages:

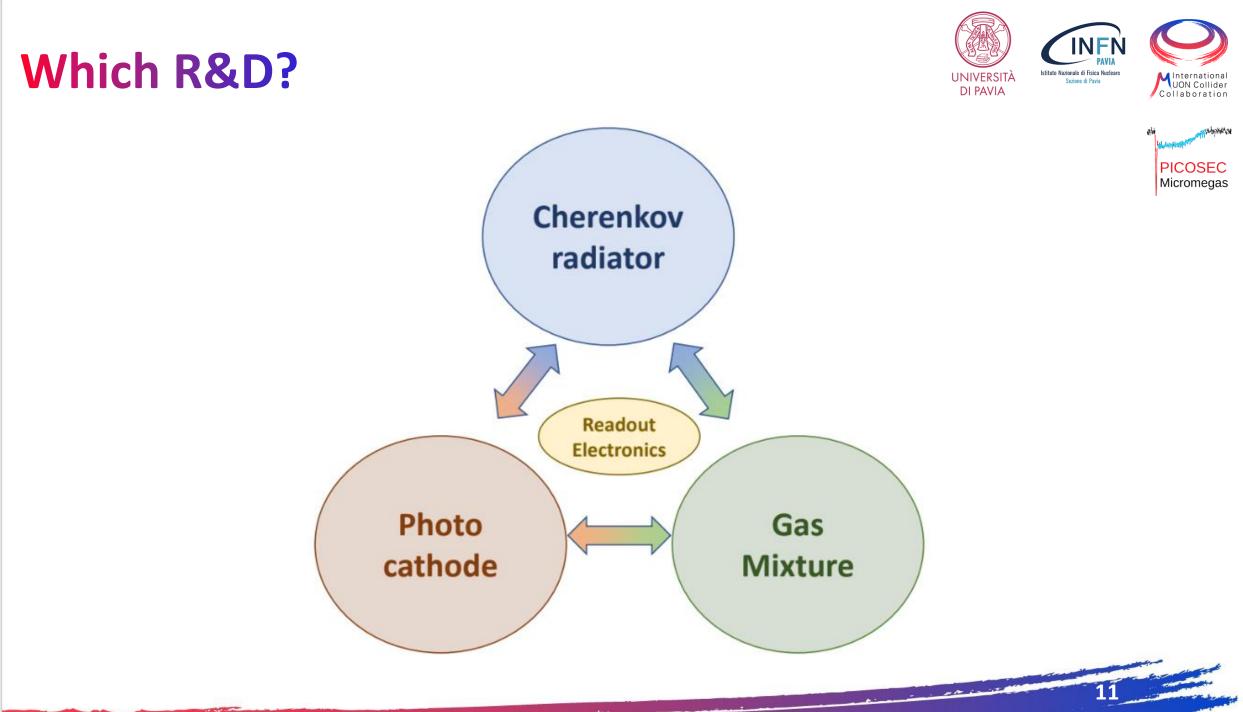
all electrons produced in the same position + very thin drift gap = *fluctuation greatly reduced*





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Cherenkov radiator

Baseline: *MgF*₂

Pro:

• High UV transparency

Cons:

- Fragile
- High cost
- Max 100 cm²





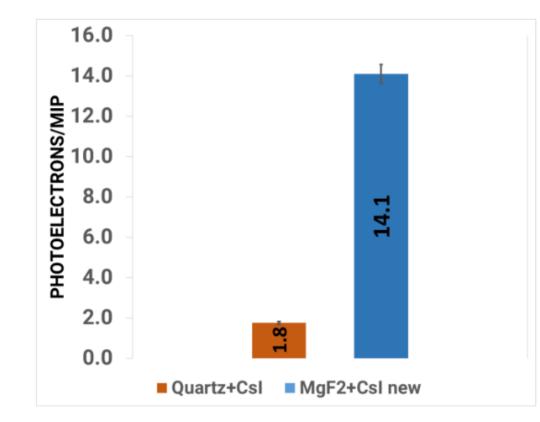


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Micromegas

Alternatives:

- Quartz
- Sapphire



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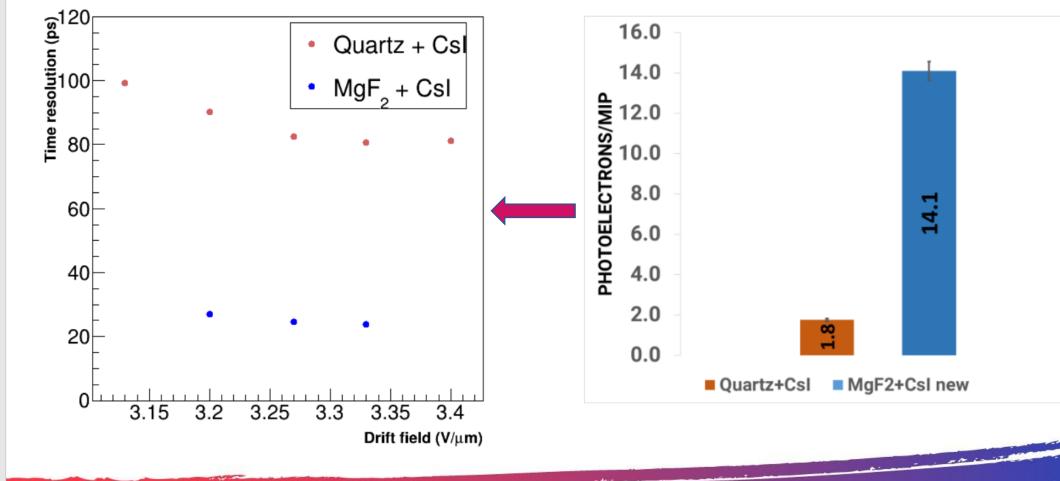




Quartz ٠



Sapphire •



Photocathode

Baseline: CsI

Pro:

- High quantum efficiency to UV (≈10p.e./MIP)
- Easy to coat by Chemical Vapour Deposition

Cons:

- Hygroscopic (sealed operation, dry gas)
- Damage by ion bombardment

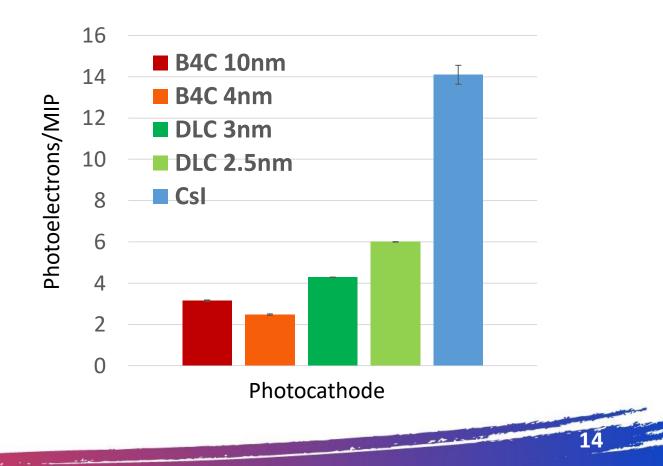


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Micromegas

Alternatives:

- Metallic (Al, Cr, Au...)
- Carbon-based (DLC, B4C)



Time Resolution (ps) 100 **IPRD2023**

6

20



Time resolution vs amplitude B4C_4

Time resolution vs amplitude B4C_7

Time resolution vs amplitude B4C_10

Time resolution vs amplitude DLC_2_5

Time resolution vs amplitude CsI

0.10 0.15 0.20 0.25 0.30 0.35 0.40 0.45 0.50

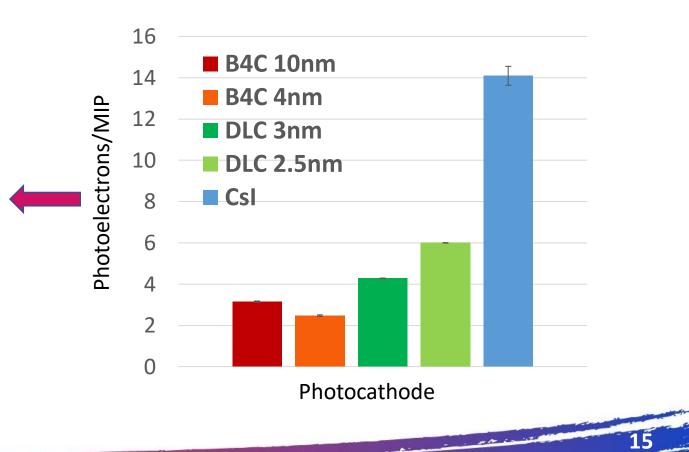
Mean Amplitude (V)





Alternatives:

- Metallic (Al, Cr, Au...)
- Carbon-based (DLC, B4C) •



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Gas mixture

Baseline: $Ne/C_2H_6/CF_4$

Pro:

- High gain and discharge quenching (up to 2-3x10⁵)
- High drift velocity ($10\div15 \text{ cm}/\mu s$)

Cons:

- Very expensive
- Very high GWP (≈740)







Alternatives:

Gas mixture used	Global Warming Potential 100-years (normalized to CO ₂)
Ne/C ₂ H ₆ /CF ₄ (80/10/10)	740
Ne/iC ₄ H ₁₀ (94/6)	0.2
Ar/CO ₂ (93/7)	0.07
Ar/CO ₂ /iC ₄ H ₁₀ (93/5/2)	0.11

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IPRD202

Results

Ar-based gas mixtures have too small operation range \rightarrow extremely difficult to find an operational configuration in which the detector is stable & the signal is visible

Ne/iC_4H_{10} gives interesting performance:

 χ^2 / NDF = 67.3 / 69

= 64.0 ps RMS_{tot} = 56.4 ps

9.6

9.65

9.7

9.75

9.8

 μ = 9.812 ns \pm 1.910 ps

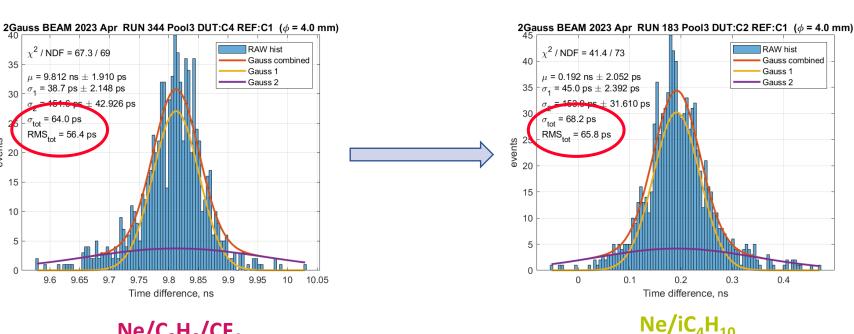
= 38.7 ps \pm 2.148 ps

35

events

10

 $Ne/C_2H_6/CF_4$ **Photocathode:** B4C 6nm (3 PE/MIP) \rightarrow Lower time resolution wrt CsI is expected



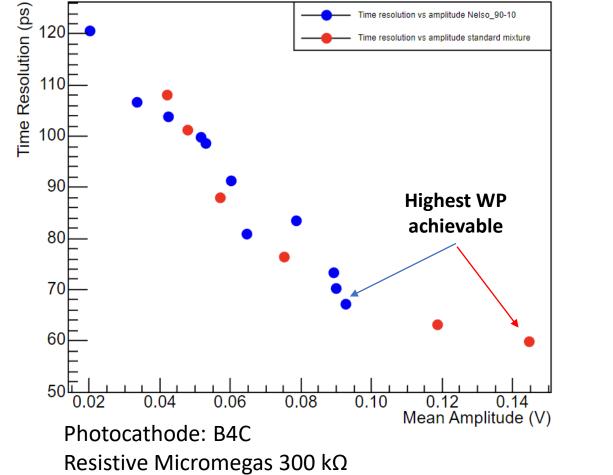




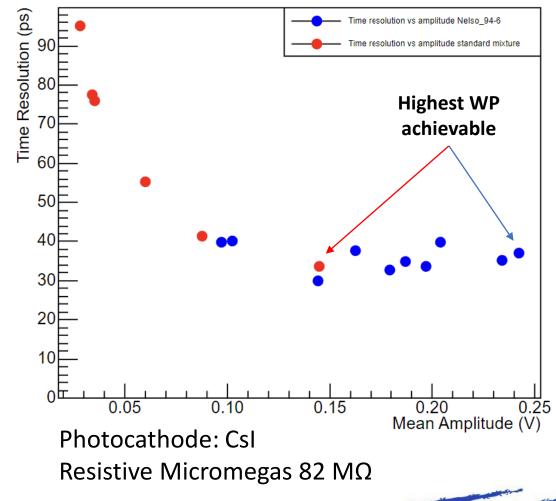
Results



Ne/iC₄H₁₀ 90/10 vs standard mixture



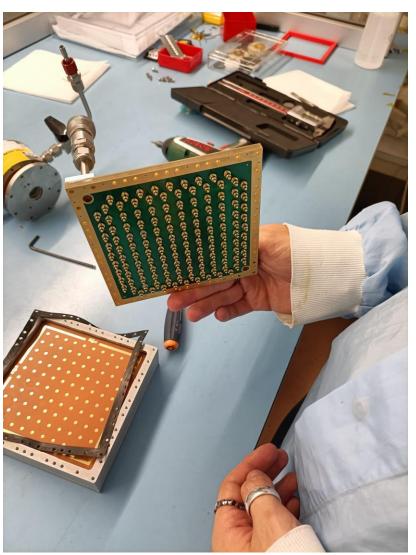
Ne/iC₄H₁₀ 94/6 vs standard mixture



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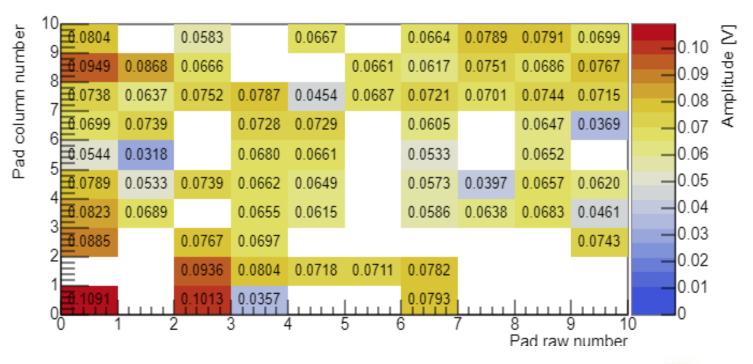
10x10 cm² multipad prototype





10x10 cm2 multipad prototype assembled in spring 2023:

- 100 readout channels
- tested in July test beam with standard gas mxture
- readout with ORTEC preamp (lab test) and SAMPIC digitizer (test beam)







The **muon collider** is an interesting opportunity for the future of high energy physics. It poses however challenging operational conditions, mainly related to the presence of BIB.

The design of a **proper experiment** is ongoing: MPGDs are being considered as candidate detectors for the muon system.

Picosec is a candidate technology for a timing layer in the muon system: a dedicated R&D is ongoing, maily focused on optimization of the Cherenkov radiator, photocathode and gas mixture.

Thanks to the <u>RD51 Picosec-Micromegas Collaboration</u> for the precious support and help!

Next year R&D activities will focus on on the mechanics needed for the instrumentation of large surfaces and on the assessment of a environmental-friendly gas mixture.

Thanks for your attention!