Freeze-In in different phases of the N2HDM

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Next to Two Higgs Doublet Model (N2HDM)

Freeze-Out vs. Freeze-In Numerical analysis Results Conclusions

Next to Two Higgs Doublet Model (N2HDM)

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Next to Two Higgs Doublet Model $_{[M\"ubileitner et al., 2017]}$

- Extending the scalar sector of the Standard Model with one complex scalar doublet Φ_2 and one real scalar singlet Φ_S
- \bullet Imposing two \mathbb{Z}_2 symmetries to the Lagrangian

$$\begin{split} \mathbb{Z}_2^1 &: \Phi_1 \to \Phi_1, \quad \Phi_2 \to -\Phi_2, \quad \Phi_S \to \Phi_S \\ \mathbb{Z}_2^2 &: \Phi_1 \to \Phi_1, \quad \Phi_2 \to \Phi_2, \quad \Phi_S \to -\Phi_S \end{split}$$

• The scalar potential is given by

$$\begin{split} V_{\text{N2HDM}} &= m_{11}^2 \Phi_1^{\dagger} \Phi_1 + m_{22}^2 \Phi_2^{\dagger} \Phi_2 + \frac{\lambda_1}{2} (\Phi_1^{\dagger} \Phi_1)^2 + \frac{\lambda_2}{2} (\Phi_2^{\dagger} \Phi_2)^2 + \lambda_3 \Phi_1^{\dagger} \Phi_1 \Phi_2^{\dagger} \Phi_2 \\ &+ \lambda_4 \Phi_1^{\dagger} \Phi_2 \Phi_2^{\dagger} \Phi_1 + \frac{\lambda_5}{2} \left[(\Phi_1^{\dagger} \Phi_2)^2 + h.c \right] + \frac{1}{2} m_s^2 \Phi_5^2 + \frac{\lambda_6}{8} \Phi_5^4 + \frac{\lambda_7}{2} \Phi_1^{\dagger} \Phi_1 \Phi_5^2 \\ &+ \frac{\lambda_8}{2} \Phi_2^{\dagger} \Phi_2 \Phi_5^2 \end{split}$$

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Phases of the N2HDM [Engeln et al., 2020]

Table 1: Different phases of the N2HDM

Phase	\mathbb{Z}_2^1	\mathbb{Z}_2^2	DM Candidate(s)	Suitable for Freeze-In
Broken Phase	broken	broken	0	no
Dark Doublet Phase	unbroken	broken	1	no
Dark Singlet Phase	broken	unbroken	1	yes
Full Dark Phase	unbroken	unbroken	2	yes

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Table 2: Scalar particle content of both phases

Phase	visible sector	dark sector
Dark Singlet Phase	1 • 1 • •	$H_{\rm DS}$
Full Dark Phase	$H_{\rm SM}$	$H_{\mathrm{DD}}, A_{\mathrm{D}}, H_{\mathrm{D}}^{\pm}, H_{\mathrm{DS}}$

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Freeze-Out vs. Freeze-In

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Freeze-Out [Gondolo and Gelmini, 1991]

Assumption: All particles, including the DM particle, are in thermal equilibrium.

Due to the expansion of the universe two phenomena happen:

 As the temperature of the universe is decreasing, the production of heavy particles decreases

 \Rightarrow The annihilation process becomes dominant and the number density of the heavy particles decreases

• The universe is expanding, hence interactions of particles become more unlikely \Rightarrow If $\Gamma_{Ann} = H$, the annihilation becomes irrelevant. This point in time is called Freeze-Out point.



Assumption: The DM particle is very weakly coupled to the thermal bath. Therefore it is decoupled from the thermal bath.

- The DM particle is produced through interactions between the thermal bath particles.
- Due to the expansion of the universe the production rate decreases and the number density increases only slowly.

Numerical analysis

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Constraints on parameter scans

- A parameter scan is performed with ScannerS [Coimbra, Sampaio, and Santos, 2013; Mühlleitner et al., 2020]
- Various experimental and theoretical constraints are applied.
- The dark matter properties are checked with MicrOMEGAs[Bélanger et al., 2018].

Relic density calculation

Full Dark phase

In the Full Dark phase one of the two DM candidates undergoes the Freeze-Out process, while the other one undergoes the Freeze-In process.

$$\Rightarrow \quad \Omega_c h^2 = \Omega_{\rm FI} h^2 + \Omega_{\rm FO} h^2$$

Dark Singlet phase

In the Dark Singlet phase only one DM candidate exists, which undergoes the Freeze-In process.

$$\Rightarrow \quad \Omega_c h^2 = \Omega_{\mathsf{FI}} h^2$$

Results

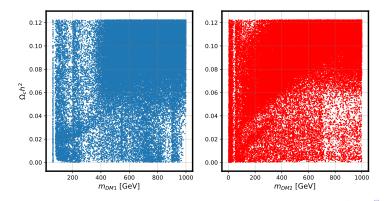
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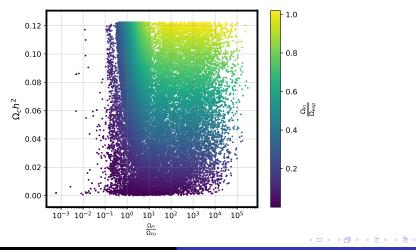
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Relic density in the FDP

Define $DM_1 = H_{DD}/A_D$ as the Freeze-Out particle and $DM_2 = H_{DS}$ as the Freeze-In particle

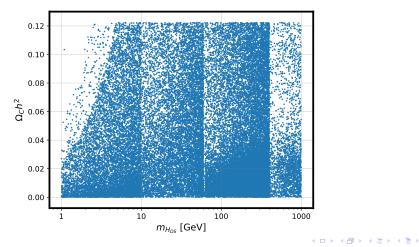


Relic density in the FDP



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Relic density in the DSP



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Conclusions

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- In two phases, namely the Full Dark phase and the Dark Singlet phase, it is possible to apply the Freeze-In mechanism.
- In both phases it has been shown, that in the scanned parameter regions, it is possible to comply with the observed relic density.
- In the FDP, it is possible to have both production mechanism, Freeze-Out and Freeze-In, for the generation of the observed relic density.