

Axion Dark Radiation from the Primordial Thermal Bath



This project has received funding from the European Union's Horizon 2020 research and innovation programme under the Marie Skłodowska-Curie grant agreement No 860881

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D'Eramo**



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DI PADOVA

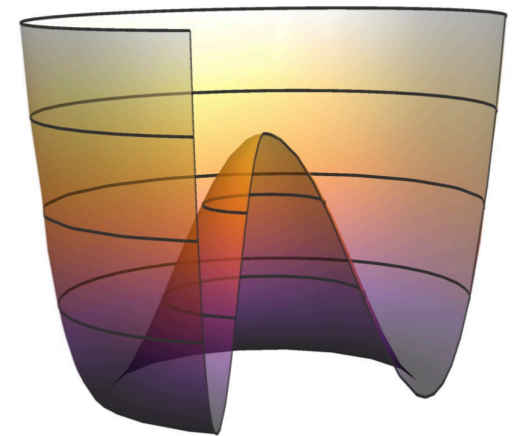


The QCD Axion

New global $U(1)_{PQ}$ symmetry

- spontaneously broken at the scale f_a (with $f_a \gg$ weak scale)
- anomalous under strong interactions

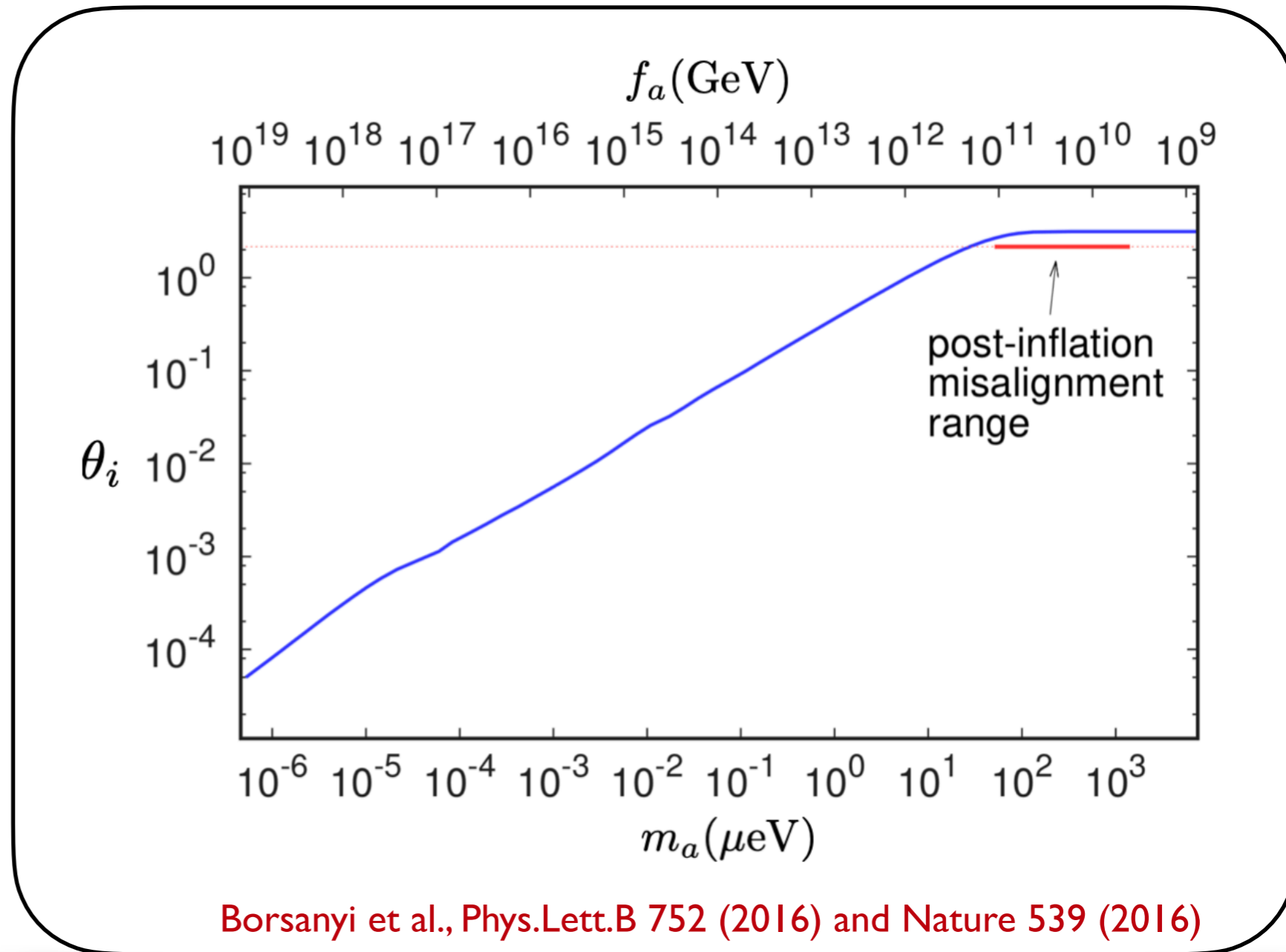
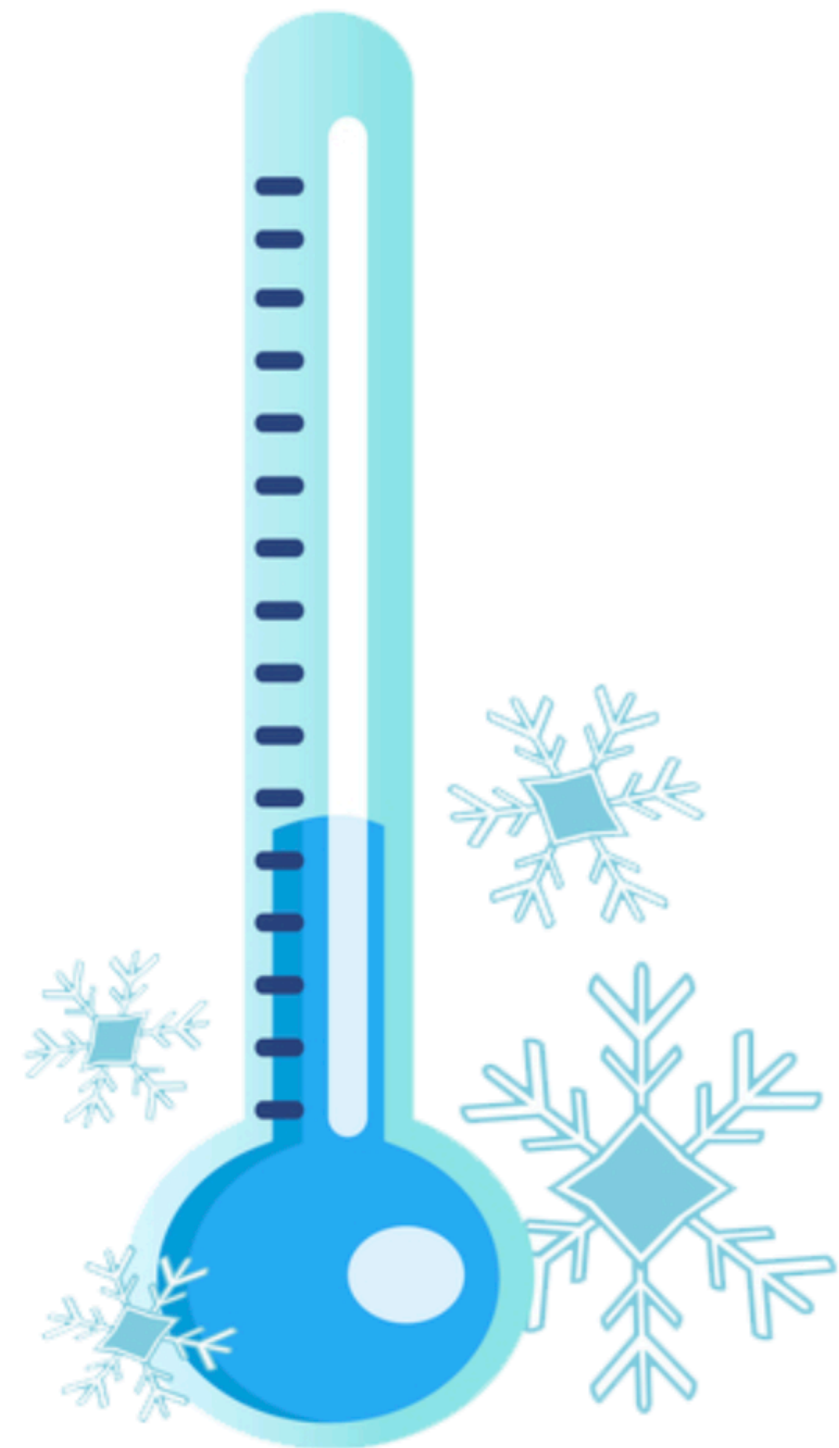
Pseudo Nambu-Goldstone boson in the low-energy spectrum (“**QCD axion**”) with (“anomalous”) coupling to gluons



$$\frac{a}{f_a} \frac{\alpha_s}{8\pi} G^{\mu\nu} \tilde{G}_{\mu\nu}$$

$$m_a \simeq 5.7 \left(\frac{10^{12} \text{ GeV}}{f_a} \right) \mu\text{eV}$$

Cold Axions – Dark Matter



Recent work on
post-inflationary
axion dark matter:

Gorghetto, Hardy, Villadoro
JHEP07 (2018) and SciPost Phys.10 (2021)
Buschmann, Foster, Safdi
PRL124 (2020) and Nature Commun. 13 (2022)

Hot Axions — Dark Radiation

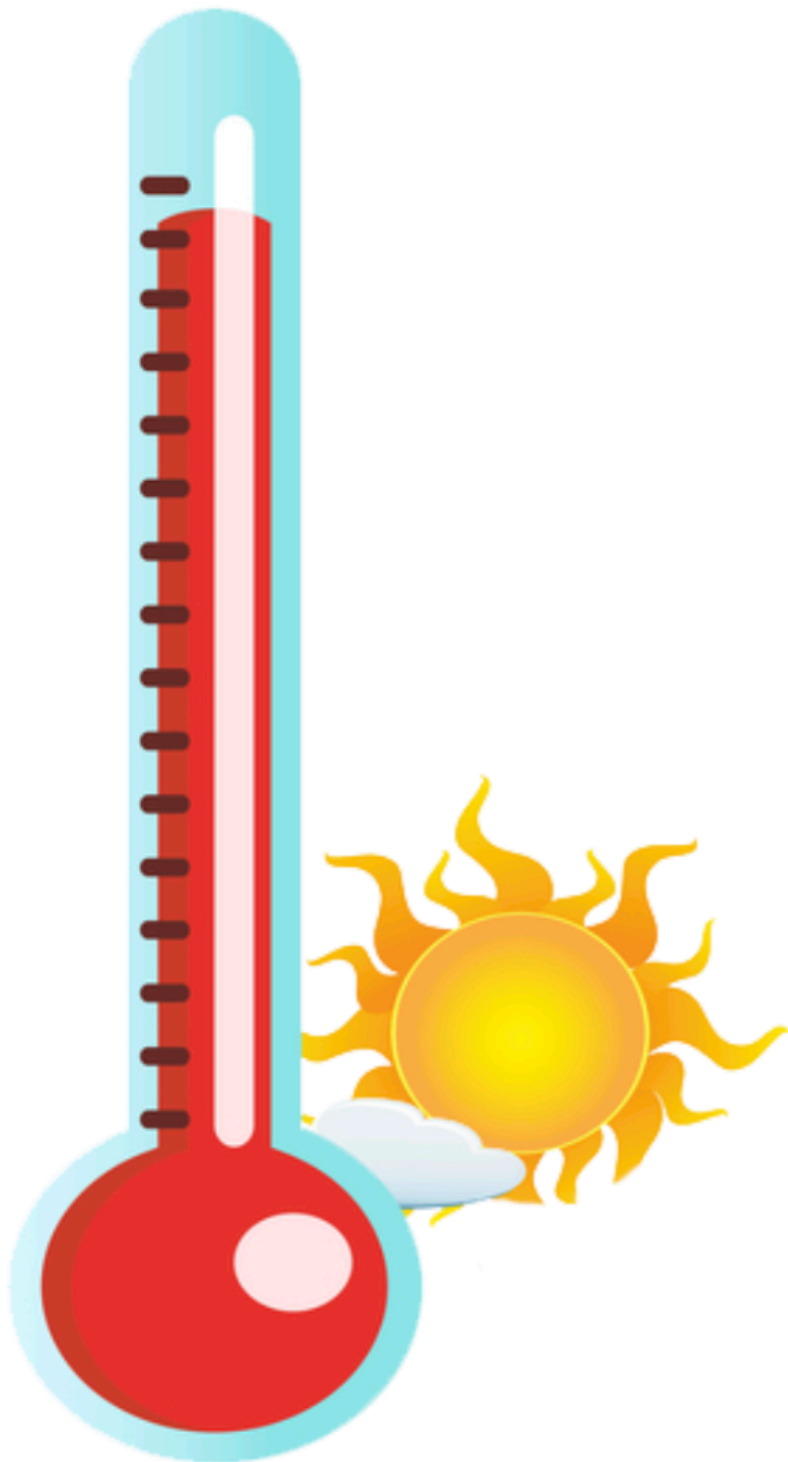
Axions produced with kinetic energy much larger than their mass (i.e. “hot”)

Additional radiation at:

- BBN ($m_a \approx \text{MeV}$)
- CMB formation ($m_a \approx 0.3 \text{ eV}$)

$$\rho_{\text{rad}} = \left[1 + \frac{7}{8} \left(\frac{T_\nu}{T_\gamma} \right)^4 N_{\text{eff}} \right] \rho_\gamma$$

$$\Delta N_{\text{eff}} = \frac{8}{7} \left(\frac{11}{4} \right)^{4/3} \frac{\rho_a}{\rho_\gamma}$$



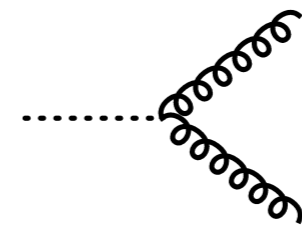
Thermal Axions

Scatterings and/or decays involving primordial thermal bath particles (axion energy $\gg m_a$, i.e. “hot”)

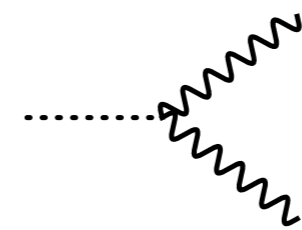
Unavoidable
Production Source!

GOALS:

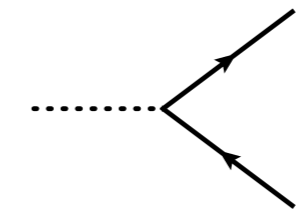
- Compute how many axions are produced in the early universe
- Quantify the resulting effect on cosmological observables



$$\frac{\alpha_s}{8\pi} \frac{a}{f_a} G^{\mu\nu} \tilde{G}_{\mu\nu}$$



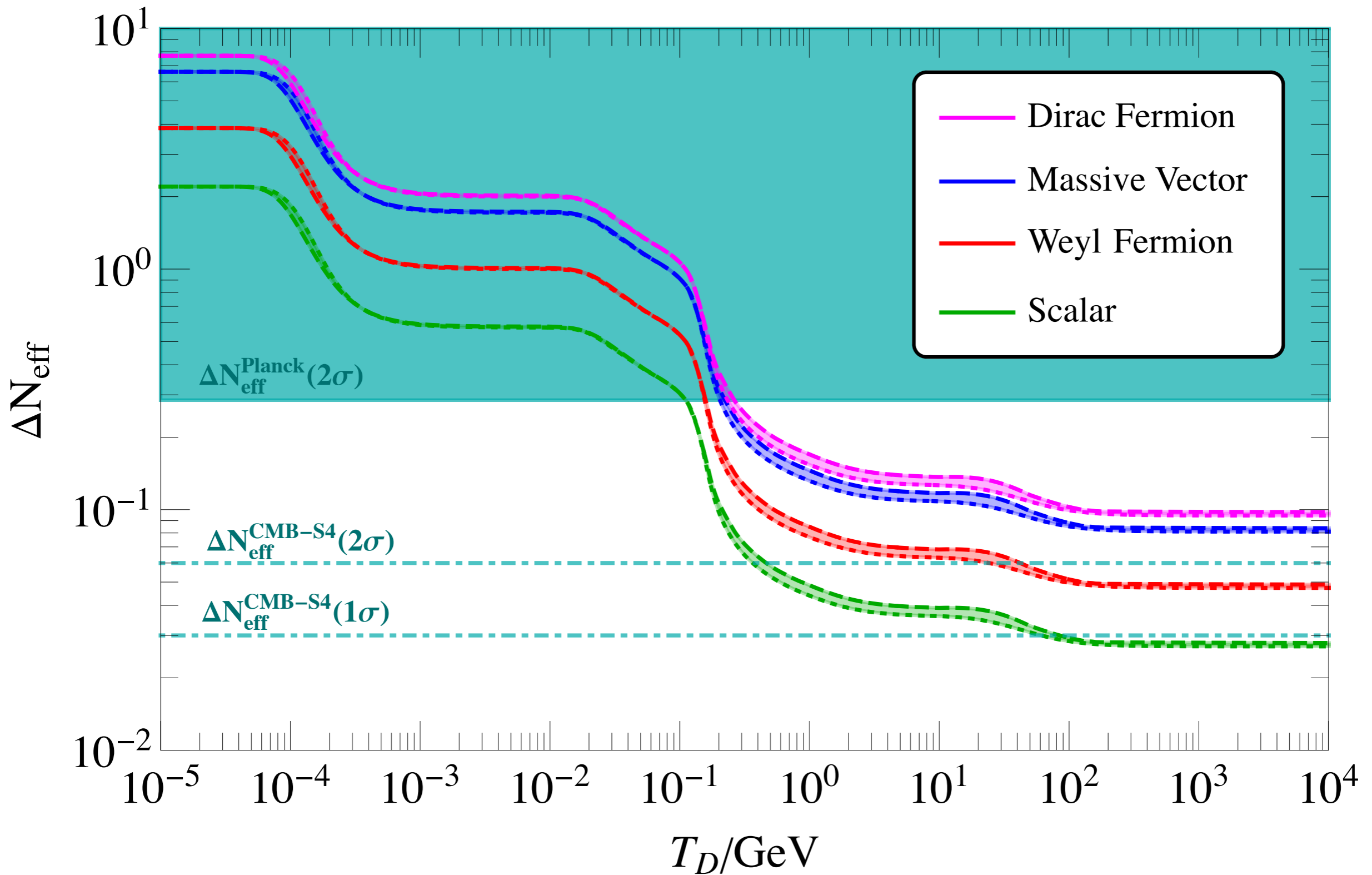
$$c_{\gamma\gamma} \frac{\alpha_{\text{em}}}{8\pi} \frac{a}{f_a} F^{\mu\nu} \tilde{F}_{\mu\nu}$$



$$c_\psi \frac{\partial_\mu a}{f_a} \bar{\psi} \gamma^\mu \gamma^5 \psi$$

$$\frac{dn_a}{dt} + 3Hn_a = \gamma_a$$

Bounds and Prospects



Single Coupling Switched On

Simplified Scenario

Axion coupled to a single Standard Model field

Ferreira, Notari, Phys.Rev.Lett. 120 (2018)

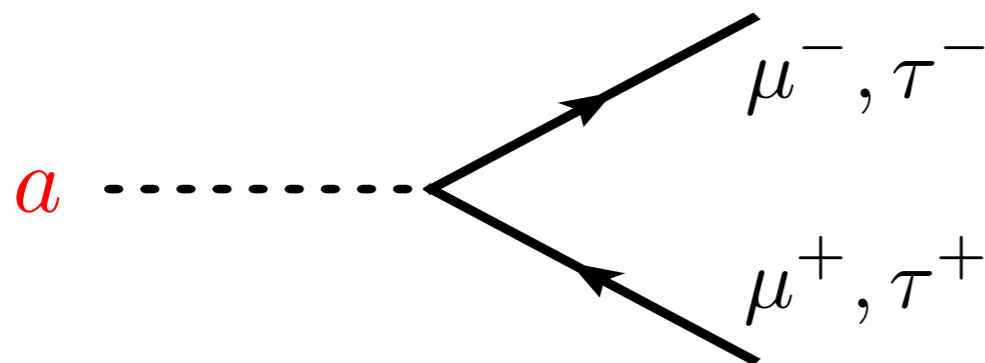
FD et al, JCAP 11 (2018)

Arias-Aragón et al., JCAP 11 (2020) and JCAP 03 (2021)

Green et al., JCAP 02 (2022)

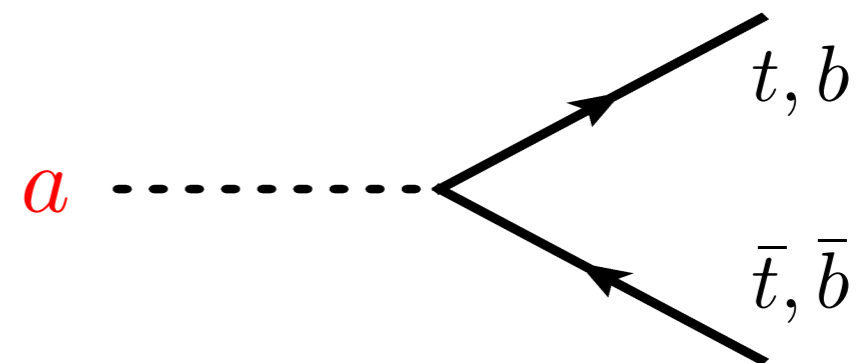
FD et al., Phys.Rev.Lett. 128 (2022)

Leptons



FD, Ferreira, Notari, Bernal, JCAP 1811 (2018)

3rd Gen. Quarks



Arias-Aragon, FD, Ferreira, Merlo, Notari, JCAP 03 (2021)

Irreducible Part for the QCD Axion

Strong CP Problem

$$\frac{a}{f_a} \frac{\alpha_s}{8\pi} G^{\mu\nu} \tilde{G}_{\mu\nu}$$

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Above

$$gg \rightarrow ga$$

$$\bar{q}q \rightarrow ga$$

$$q/\bar{q} g \rightarrow q/\bar{q} a$$

Long-range of gluon-mediated interactions
give rise to unpleasant IR behavior

Studied for
 $T > 10^4$ GeV

Masso, Rota, Zsembinski, PRD66 (2002)

Graf, Steffen, PRD 83 (2011)

Salvio, Strumia, Xue, JCAP 01 (2014)

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Below

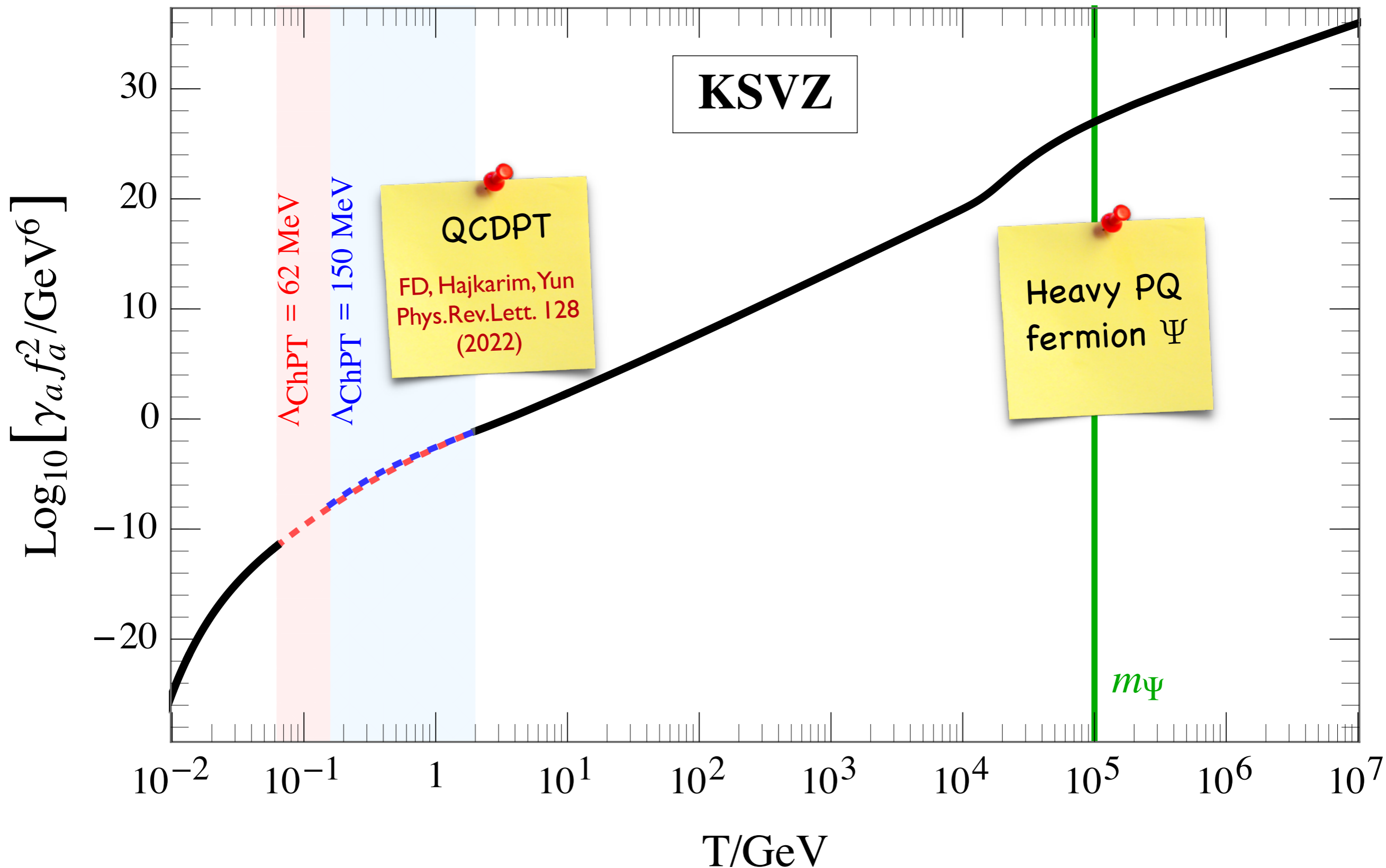
$$\pi\pi \rightarrow \pi a$$

Chang, Choi, Phys.Lett.B 316 (1993)

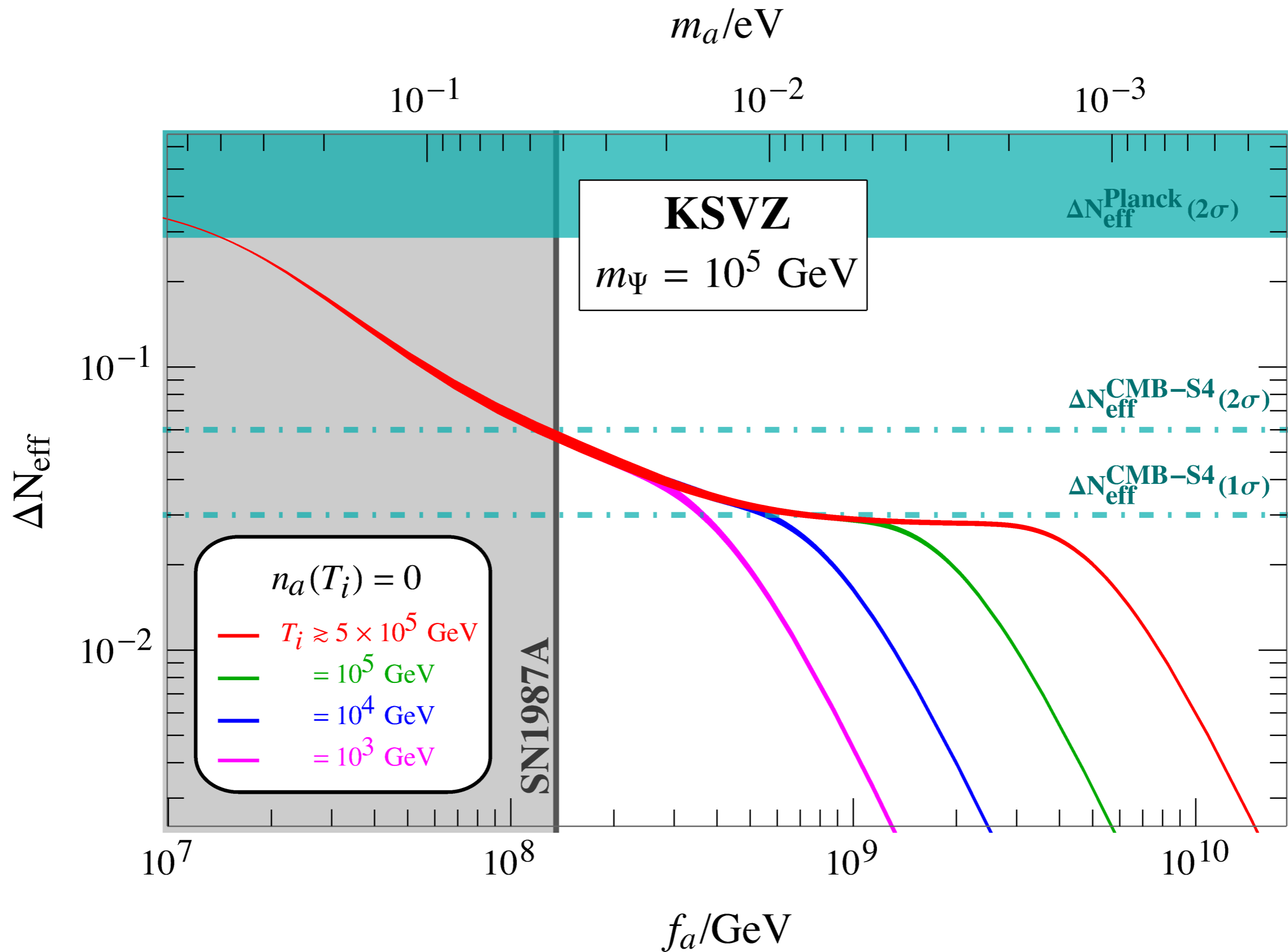
Pion scattering

Recent studies:
Di Luzio, Martinelli, Piazza, Phys.Rev.Lett. 126 (2021)
Notari, Rompineve, Villadoro, Phys.Rev.Lett. 131 (2023)
Di Luzio, Camalich, Martinelli, Phys. Rev. D 108 (2023)

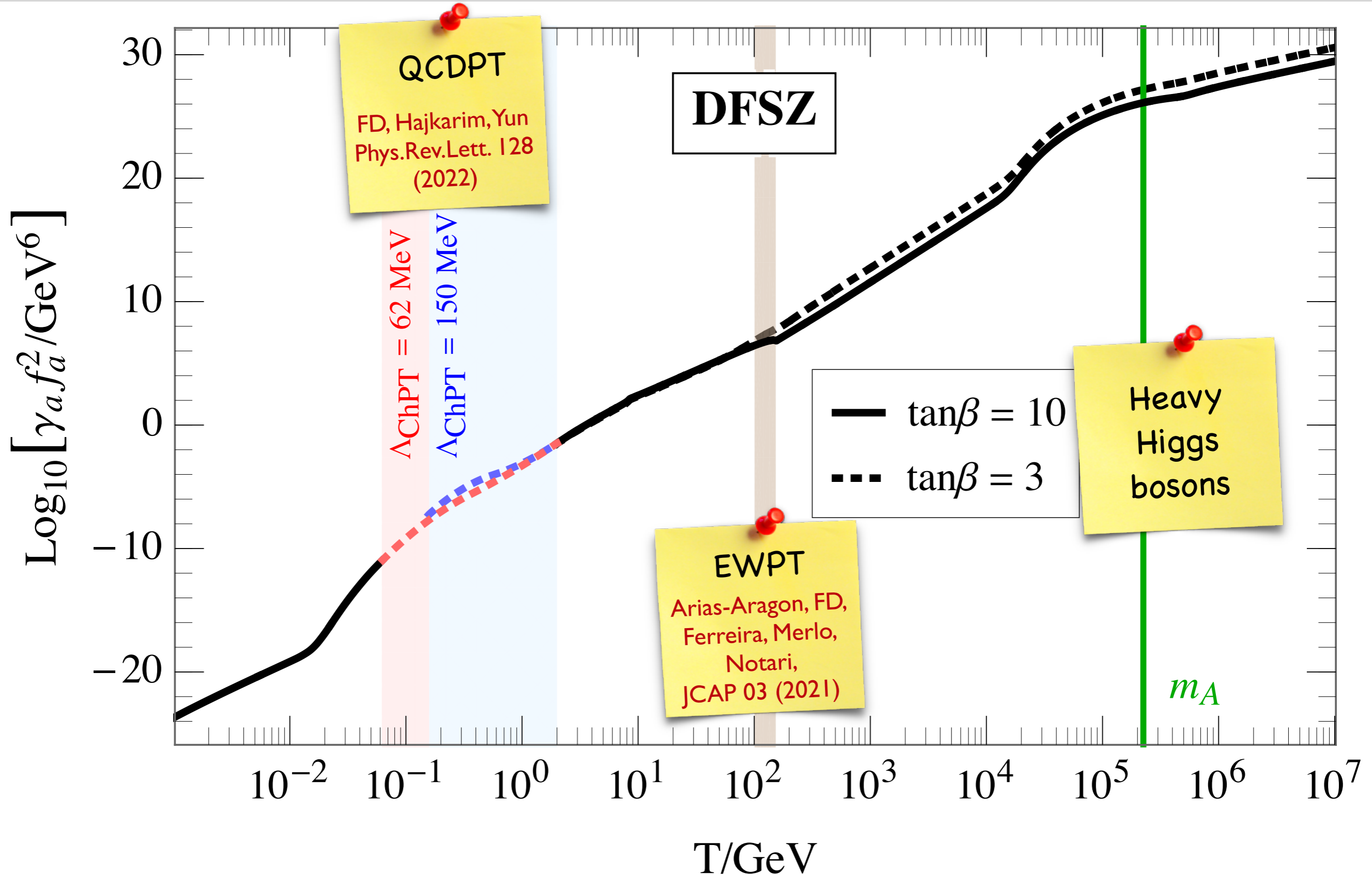
KSVZ Axion — Production Rate



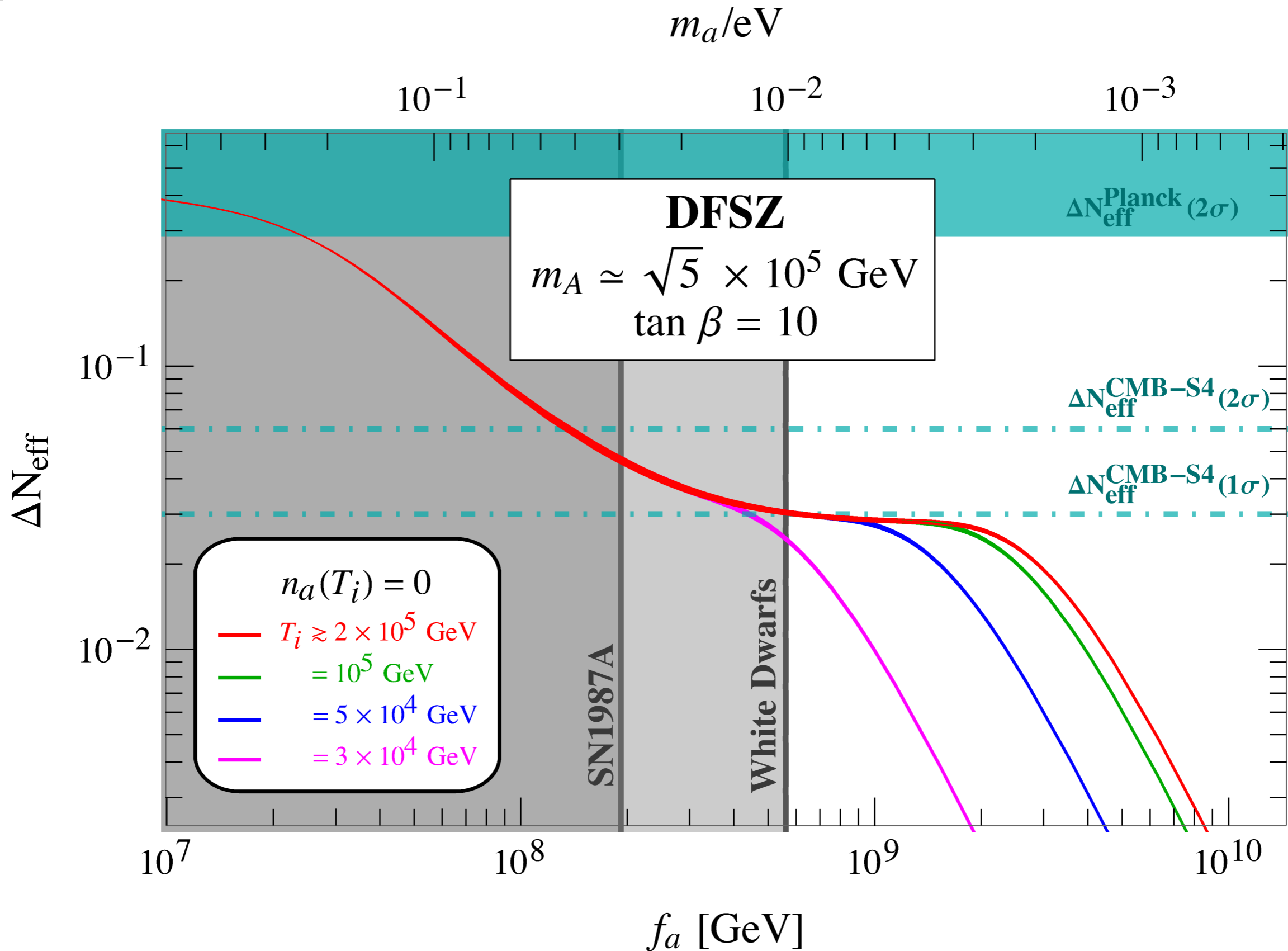
KSVZ Axion — ΔN_{eff}



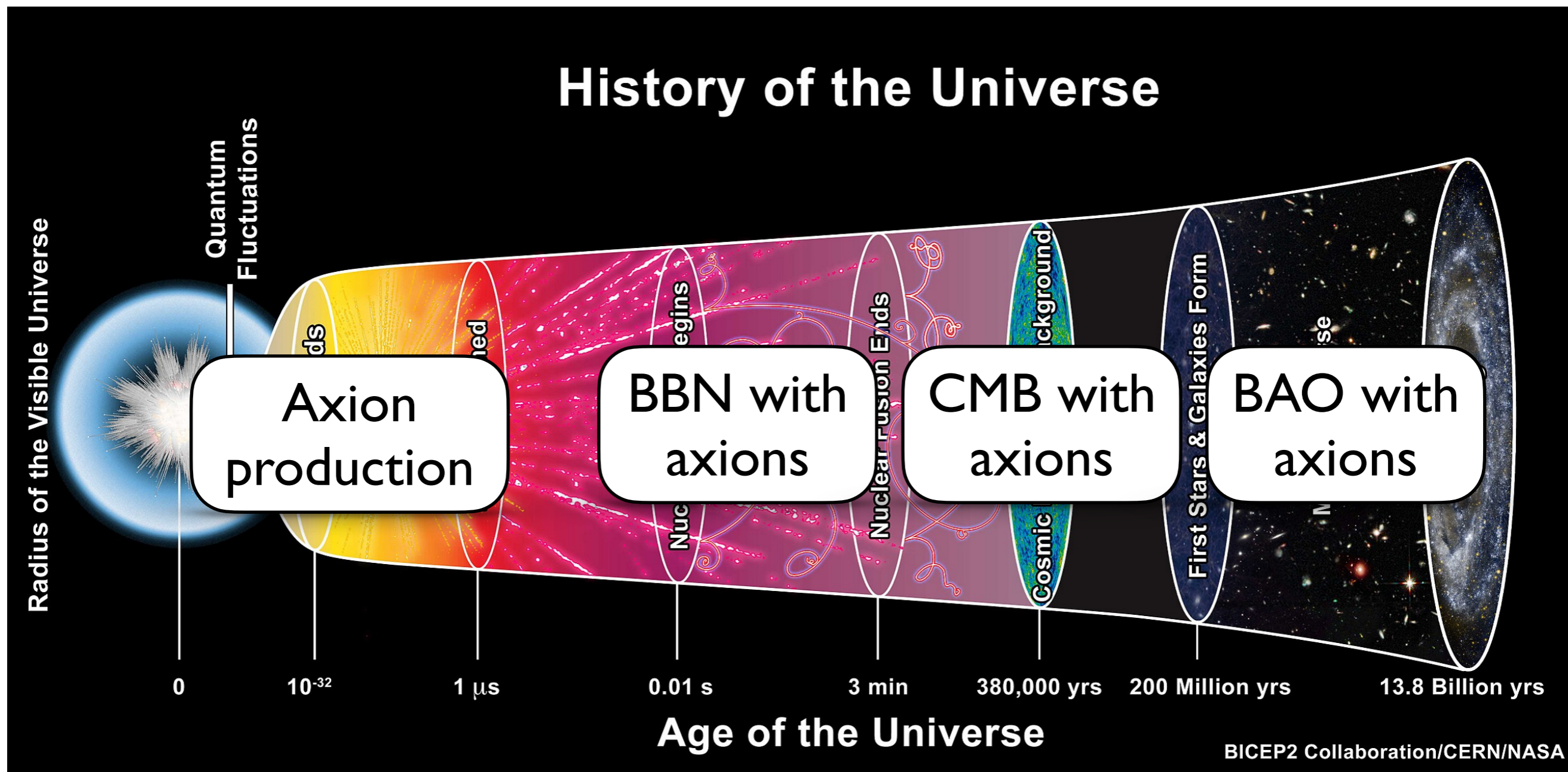
DFSZ Axion — Production Rate



DFSZ Axion — ΔN_{eff}



QCD Axion Mass Bound



KSVZ

$$m_a \leq 0.282(0.420) \text{ eV}$$

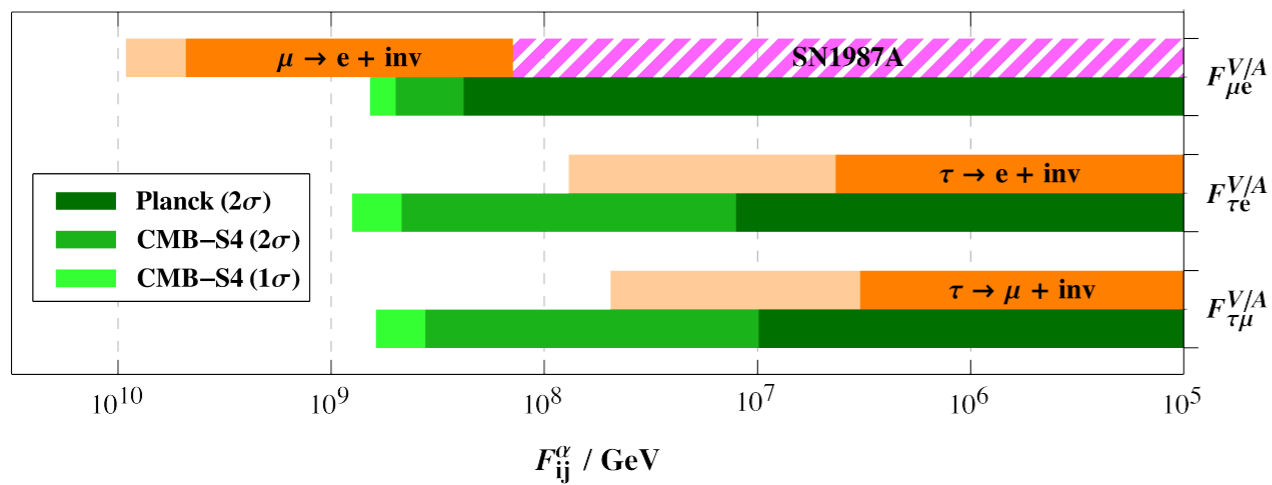
DFSZ

$$m_a \leq 0.209(0.293) \text{ eV}$$

A Minor Variation: FV Axions

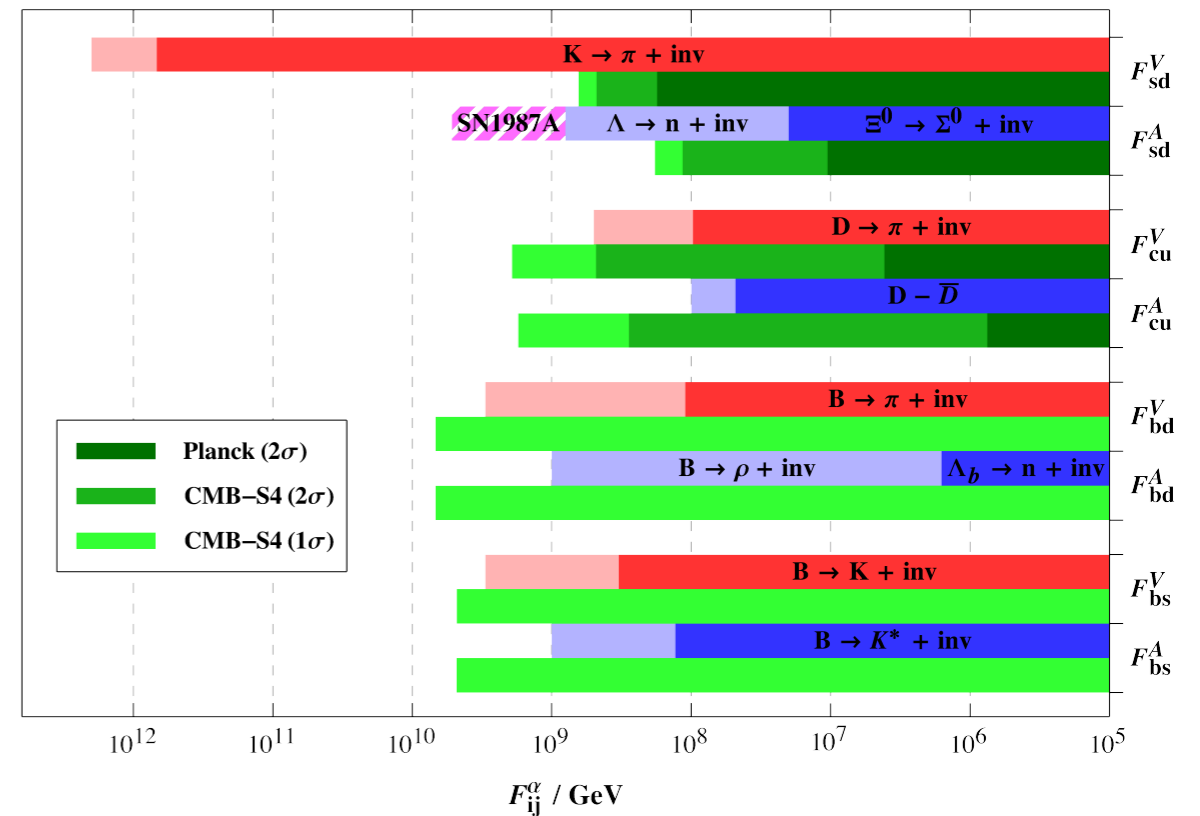
$$\mathcal{L}_{\text{FV}}^{(a)} = \frac{\partial_\mu a}{2f_a} \sum_{\psi_i \neq \psi_j} \bar{\psi}_i \gamma^\mu \left(c_{\psi_i \psi_j}^V + c_{\psi_i \psi_j}^A \gamma^5 \right) \psi_j$$

Leptonic FV

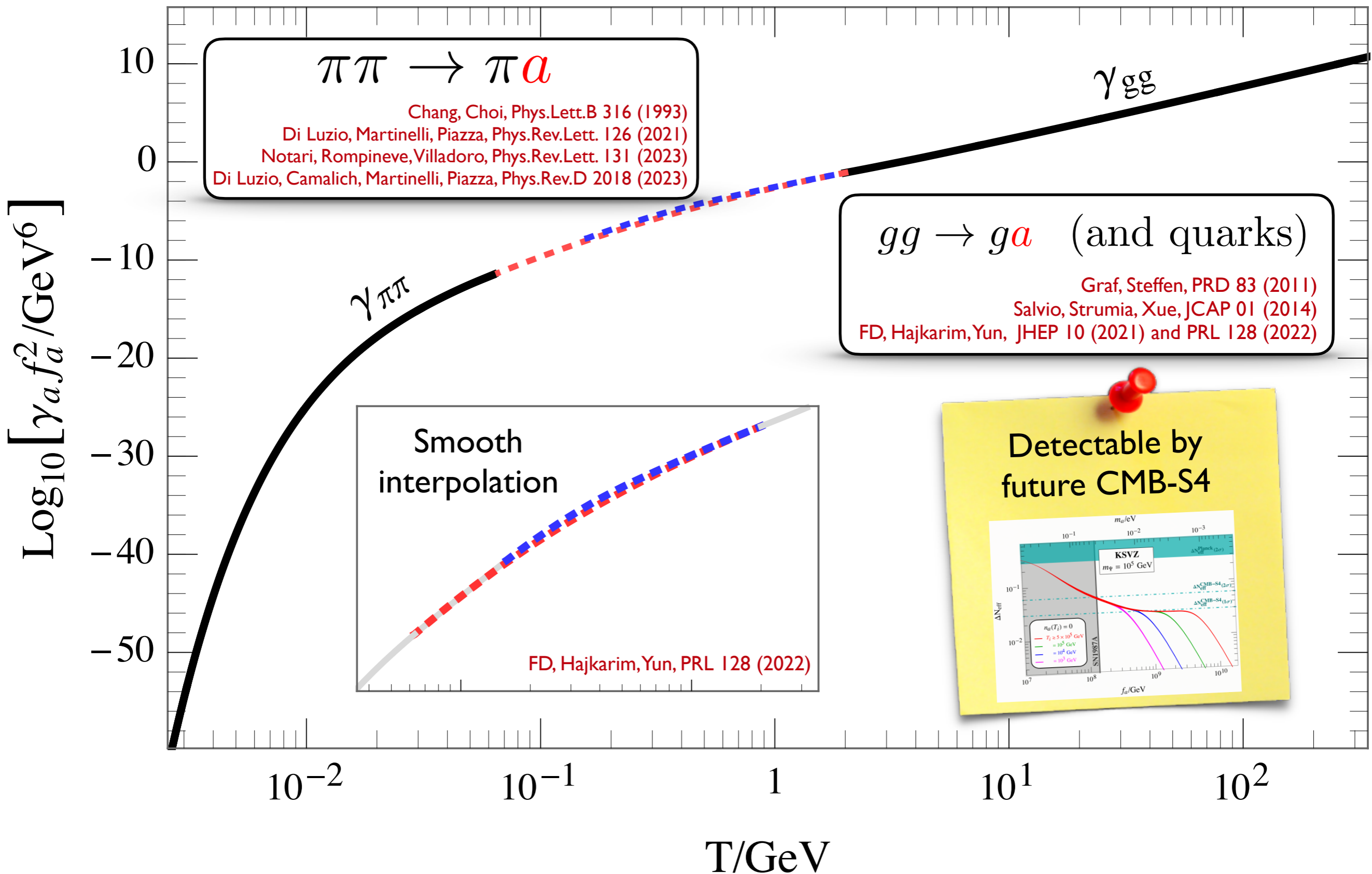


$$F_{\psi_i \psi_j}^\alpha \equiv \frac{2f_a}{c_{\psi_i \psi_j}^\alpha}$$

Hadronic FV



Where Do We Stand?



What's Next?

Axion production rate
across the confinement scale still unknown

$$\gamma_a = n_i n_j \times \langle \sigma_{ij \rightarrow ja} v_{\text{rel}} \rangle$$

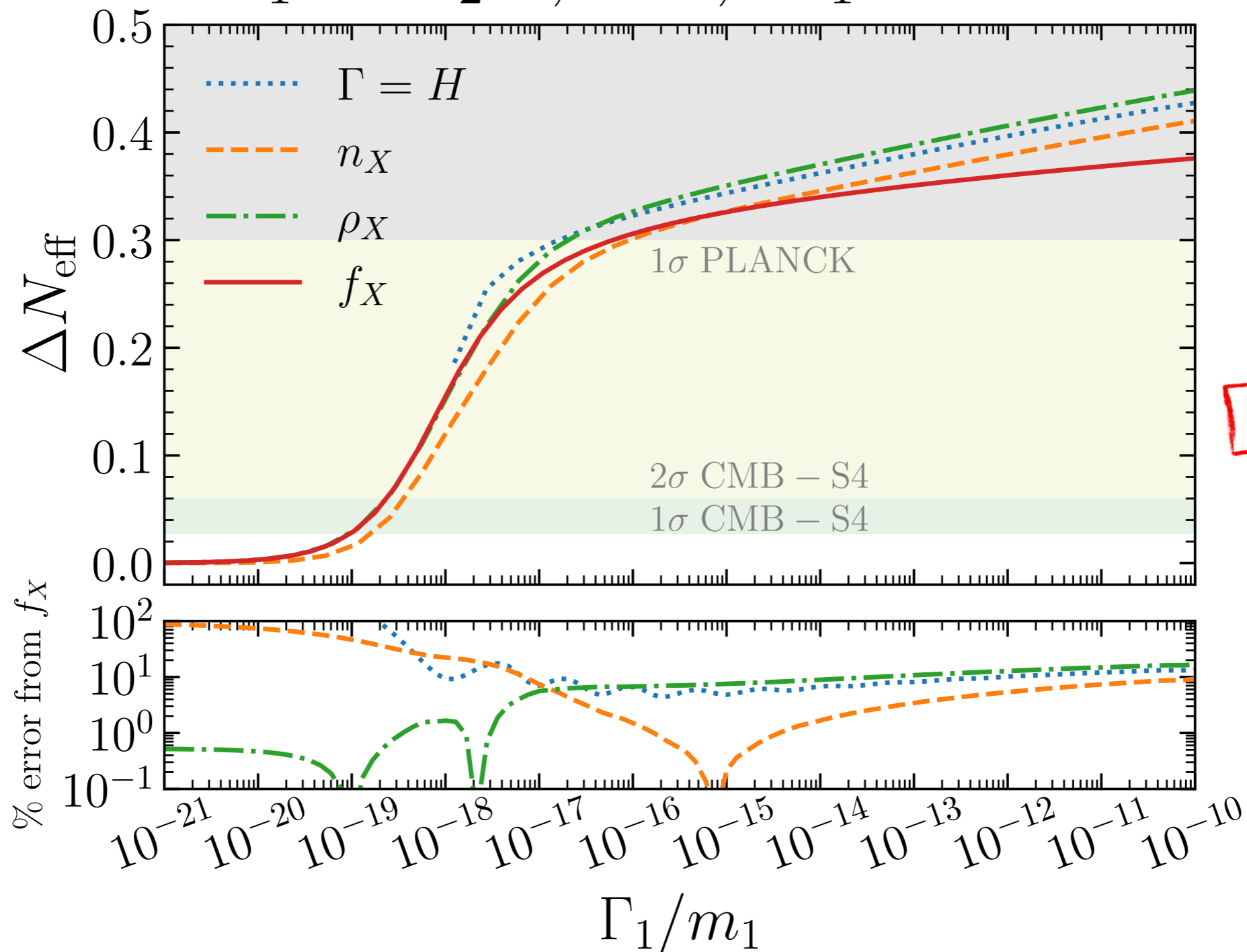
Thermal bath

Particle Physics

1. Cross sections with other hadrons?
2. Thermal bath description between 150 MeV and fews GeV?
3. Boltzmann equation evolution and cosmological observables?

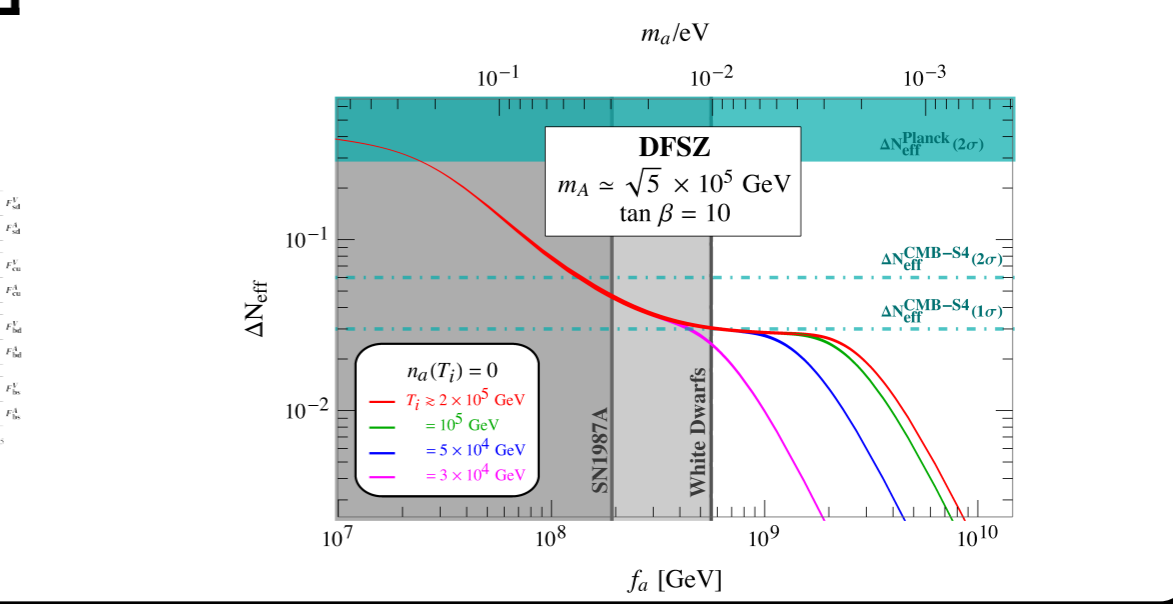
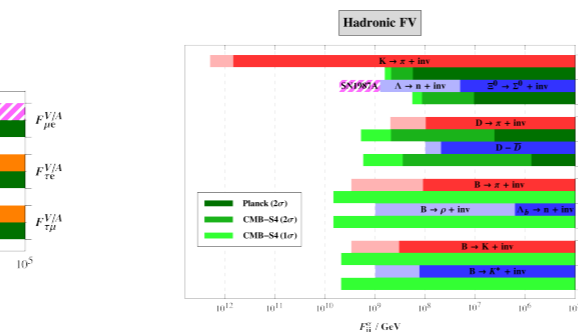
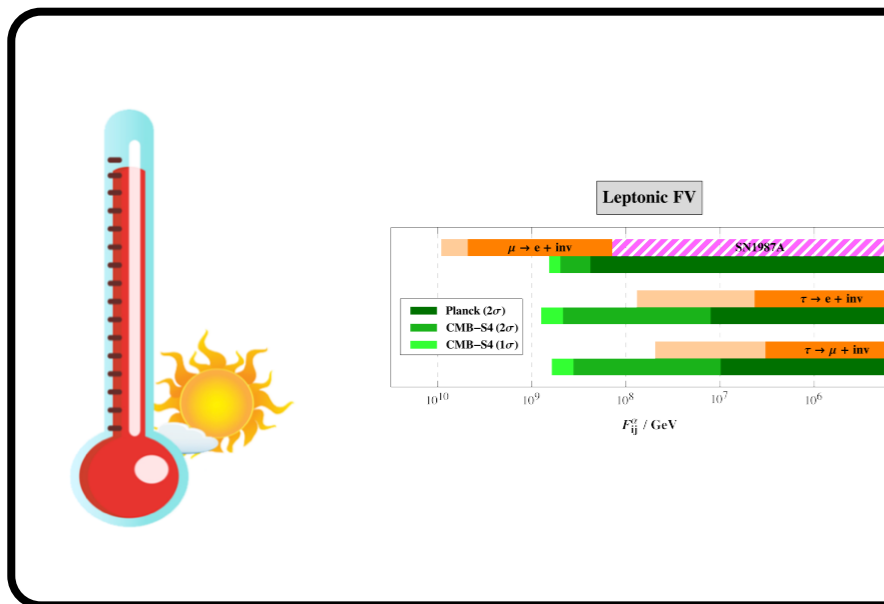
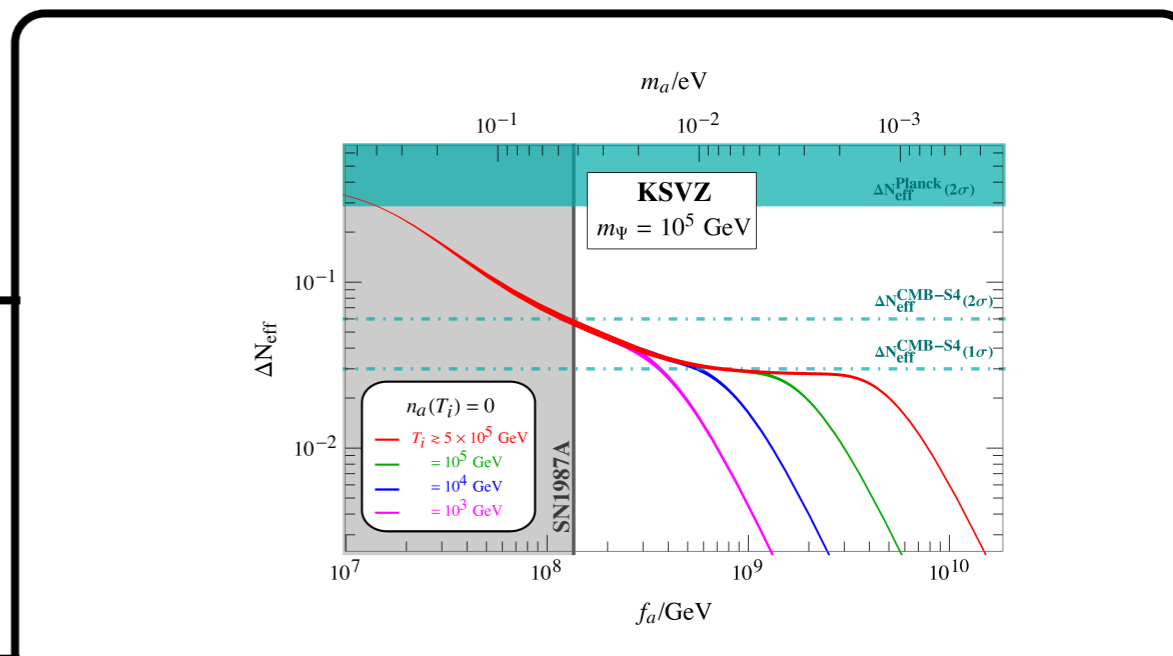
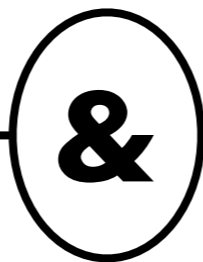
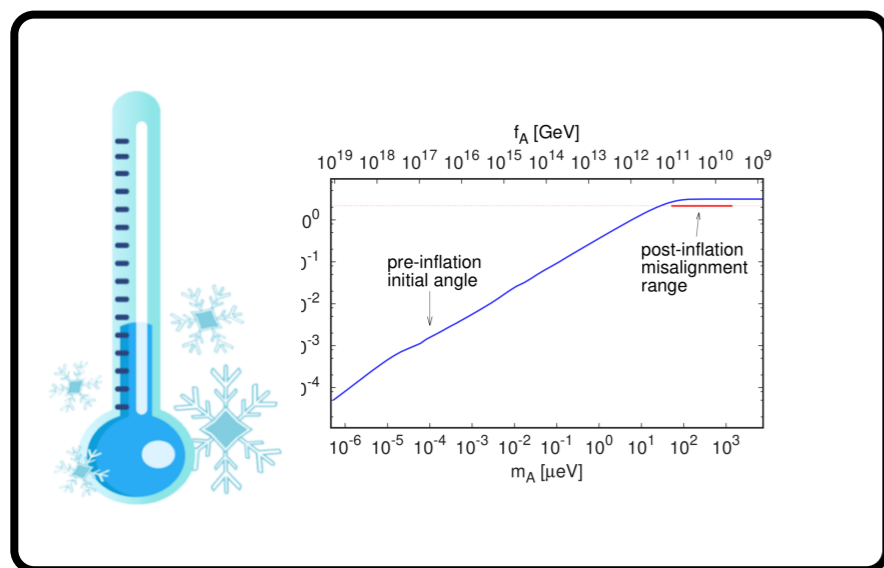
Phase-Space Analysis

$$\mathcal{B}_1 \rightarrow \mathcal{B}_2 X, \text{ MB}, m_1 = 1 \text{ GeV}$$



PRELIMINARY

Outlook



Thermal Axions

Complementary to other probes of the PQ mechanism

