

# Exploring dark QCD dark matter models with heavy quarks

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## **General Setup**

Lagrangian of a dark SU(N<sub>d</sub>) gauge theory with heavy quarks ( $m_Q > \Lambda_{dQCD}$ ), which can be charged under the SM or which are SM singlets.

$$\mathcal{L}_{ ext{dark}} \supset -rac{1}{2} ext{tr} \left( G_{d,\mu
u} G_d^{\mu
u} 
ight) + \sum_{j=1}^{N_F} \overline{Q}_{d,j} \left( i \gamma^\mu D_\mu - m_{Q,j} 
ight) Q_{d,j}$$

Gauge groups SO(N) and Sp(2N) have also been considered in the literature Accidental Composite Dark Matter, O. Antipin et al. (2015) Low-enegy effective description of dark Sp(4) theories, S. Kulkarni et al. (2022)

Confinement scale  $\Lambda_{dQCD}$  is identified with the one-loop Landau pole

$$lpha_d(m_Q) = rac{2\pi}{eta_0 \ln \left( rac{m_Q}{\Lambda_{
m dQCD}} 
ight)}, \quad eta_0 = rac{11 N_d - 2 N_F}{3}$$

If no other portal to the SM is included, confinement scale and quark mass(es) fix all the parameters of the theory.

Dark baryon number and dark species number are *accidental global symmetries* that ensure stability of dark hadrons.



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**Baryons** made of  $N_d$  dark quarks  $\rightarrow$  **DM candidate**.

Stable up to dimension  $3/2^*(N_d + 1)$  operators for SU(N<sub>d</sub> odd) (fermionic DM). and stable up to dimension  $(3/2^*N_d + 2)$  operators for SU(N<sub>d</sub> even) (bosonic DM).



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**Mesons** made of quark-antiquark: Stable against a decay into the SM up to dimension 5. Without accidental symmetries (G-parity, flavour conservation) protecting them, they decay quickly into glueballs.

*Model for Thermal Relic Dark Matter of Strongly Interacting Massive Particles*, Y. Hochberg et al. (2015) *A Theory of Dark Pions*, H.C. Cheng et al. (2022)







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**Glueballs** (GBs) as the lightest hadrons in the heavy quark case. Stable up to dimension 6. Pure glue *thermal* dark matter is excluded by overclosure. Hidden SU(N) glueball dark matter, A. Soni and Y. Zhang (2016) Non-Abelian Dark Forces and the Relic Densities of Dark Glueballs, L. Forestell et al. (2017) Glueball dark matter, precisely, P. Carenza et al. (2023)







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**Dark nuclei**: If a light mediator is added to the model, dark nucleosynthesis is feasible.

Big Bang Darkleosynthesis, G. Krnjaic and K. Sigurdson (2014) Dark Nuclei I & II, W. Detmold et al. (2014) Big Bang Synthesis of Nuclear Dark Matter, E. Hardy et al. (2014)









## Squeezeout of Dark Matter



It has been found that in the case of a **first order** confining phase transition with **heavy quarks**, the phase transition drastically depletes the dark matter abundance via the squeezeout effect.

Accidentally Asymmetric Dark Matter, P. Asadi et al. (2021) Thermal Squeezeout of Dark Matter, P. Asadi et al. (2022) Glueballs in a Thermal Squeezeout Model, P. Asadi et al. (2022)

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Heavy fundamental quarks can not enter the confined-phase bubble since colour string breaking is exponentially suppressed. Only a statistical excess fraction of  $\sqrt{N_q^{\text{initial}}}$  quarks in the deconfined pockets survives the phase transition  $\rightarrow$  **dramatic decrease of the dark matter abundance!** 



## Thank you for your attention

