



Incoherent Electron Pair Production Background @ C³

Tim Barklow (SLAC), Lindsey Gray (FNAL),
Elias Mettner, Abdollah Mohammadi (UW-Madison),
Dimitris Ntounis, Caterina Vernieri (SLAC)

ECFA 2nd Topical Meeting on Generators

21 June 2023



WISCONSIN
UNIVERSITY OF WISCONSIN-MADISON



NATIONAL
ACCELERATOR
LABORATORY

bold - did most of the work :-)

Background Simulations at C³

- Linear collider machine and beam backgrounds play a significant role in:
 - Detector design (occupancies of innermost tracking layers)
 - Ultimate physics reach (fake rate / misreconstructions from spurious hits)
- Last workshop: demonstrated similarity of C³ to ILC
 - We have been assuming ILC-like physics performance of an SiD-like detector @ C³
 - Here we show refined estimates for the pair-production backgrounds without hadron photoproduction (effectively a 10% increase to what we will see)
 - Other machine backgrounds (tertiary muons, etc.) to come later, they are smaller effects
- We will see:
 - C³ is quantitatively equivalent to ILC from the perspective of backgrounds at the IR
 - While detector re-optimization is required (bunch structure) physics reach considerations are the same.

C3 Parameters

- Input values to simulation derived from C3 optics and dynamics simulations @ 250 GeV CoM
 - Started this project with some guesses due to incomplete information
 - Now have complete configuration of the machine from background simulation perspective
- Note that bunch/repetition structure at C3 different from ILC

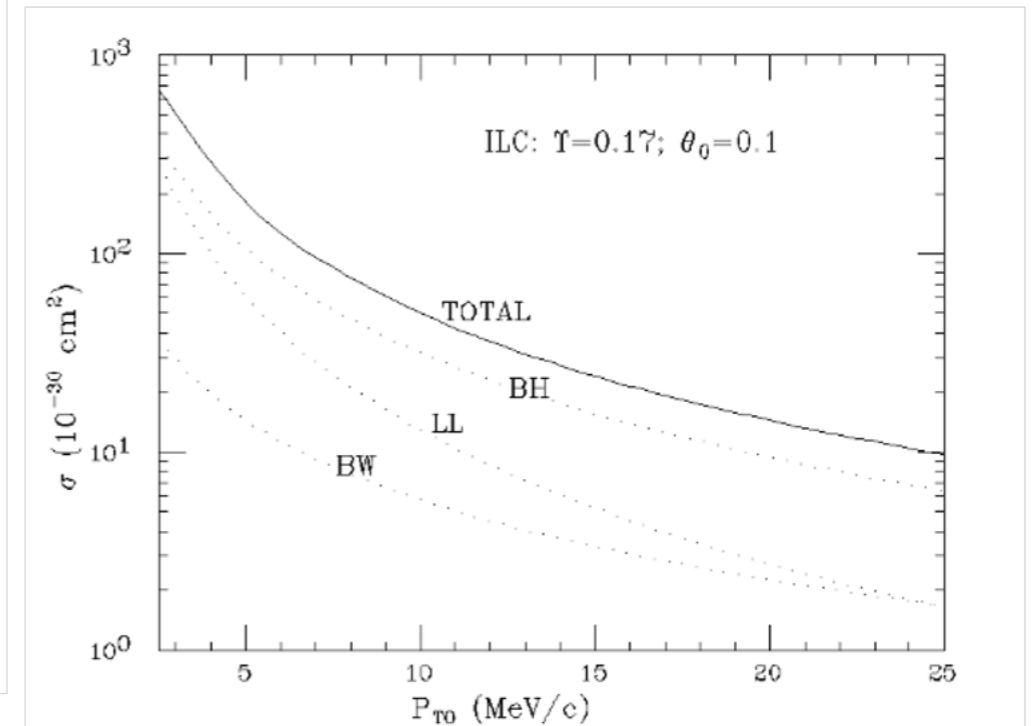
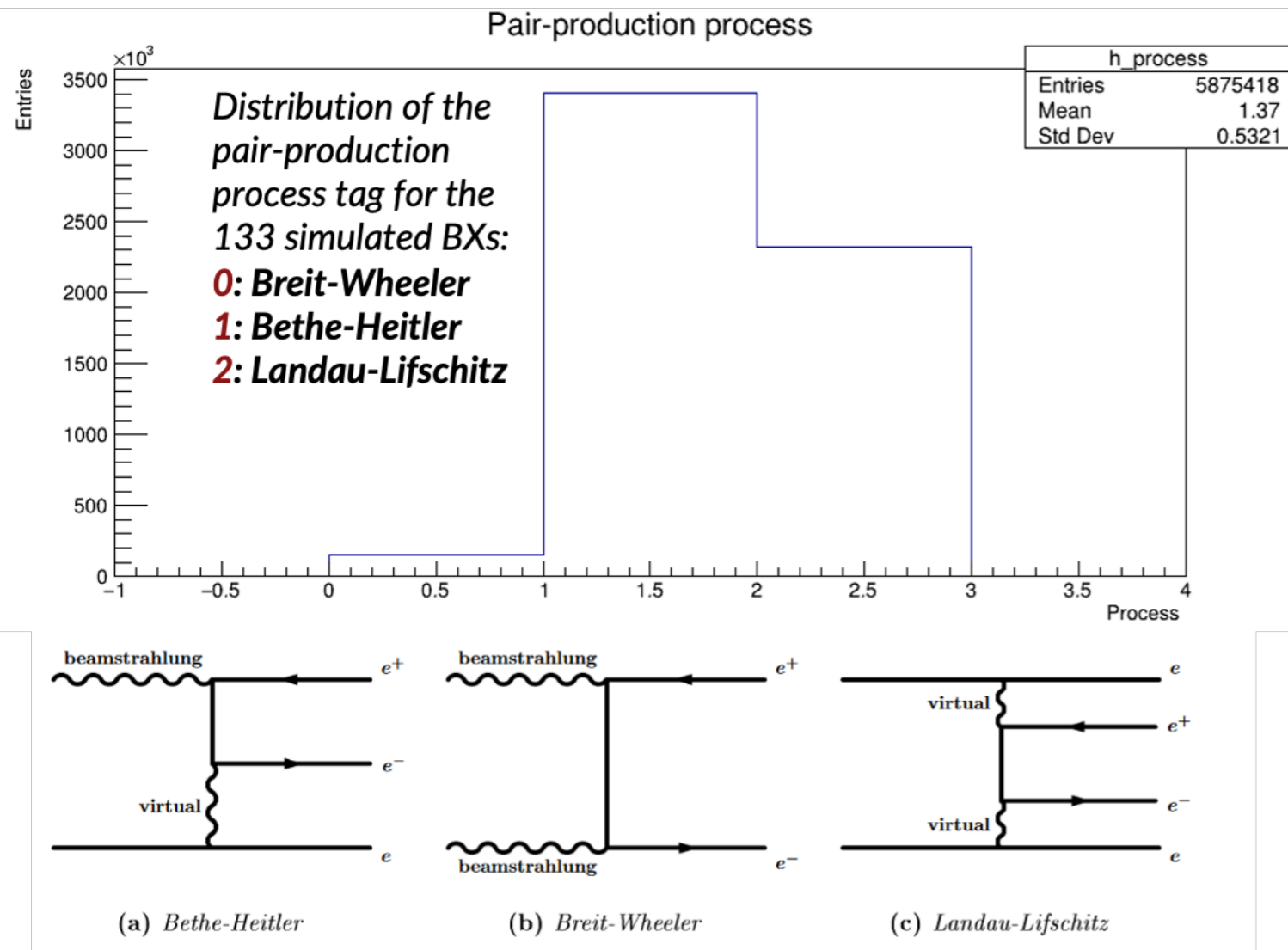
Parameter	Units	Value
β_x^*	mm	12
β_y^*	mm	0.12
$\epsilon_{N,x}^*$	nm	900
$\epsilon_{N,y}^*$	nm	20
σ_x^*	nm	210.12
σ_y^*	nm	3.13
σ_z^*	μm	100
n_b		133
f_{rep}	Hz	120
N		$6.25 \cdot 10^9$
θ_c	rad	0.014

- The emittances on the table are **normalized**. The transverse beam size is calculated as:

$$\sigma_{x,y}^* = \sqrt{\epsilon_{x,y}^* \beta_{x,y}^*} = \sqrt{\frac{\epsilon_{L,x,y}^* \beta_{x,y}^*}{\gamma}}, \quad \gamma = \frac{E}{m_e c^2} = \frac{\sqrt{s}}{2m_e c^2}$$

	Initial Tests	Emilio's Values
Energy spread	0.1%	0.3%
Energy spread distribution	Gaussian	Flat
Offset in x direction (nm)	0	5
Offset in y direction (nm)	0	0.2
Waist shift in x direction (μm)	0	0
Waist shift in y direction (μm)	0	Thanks Emilio! 0
Crossing angles (not compensated by crab scheme)	0	0

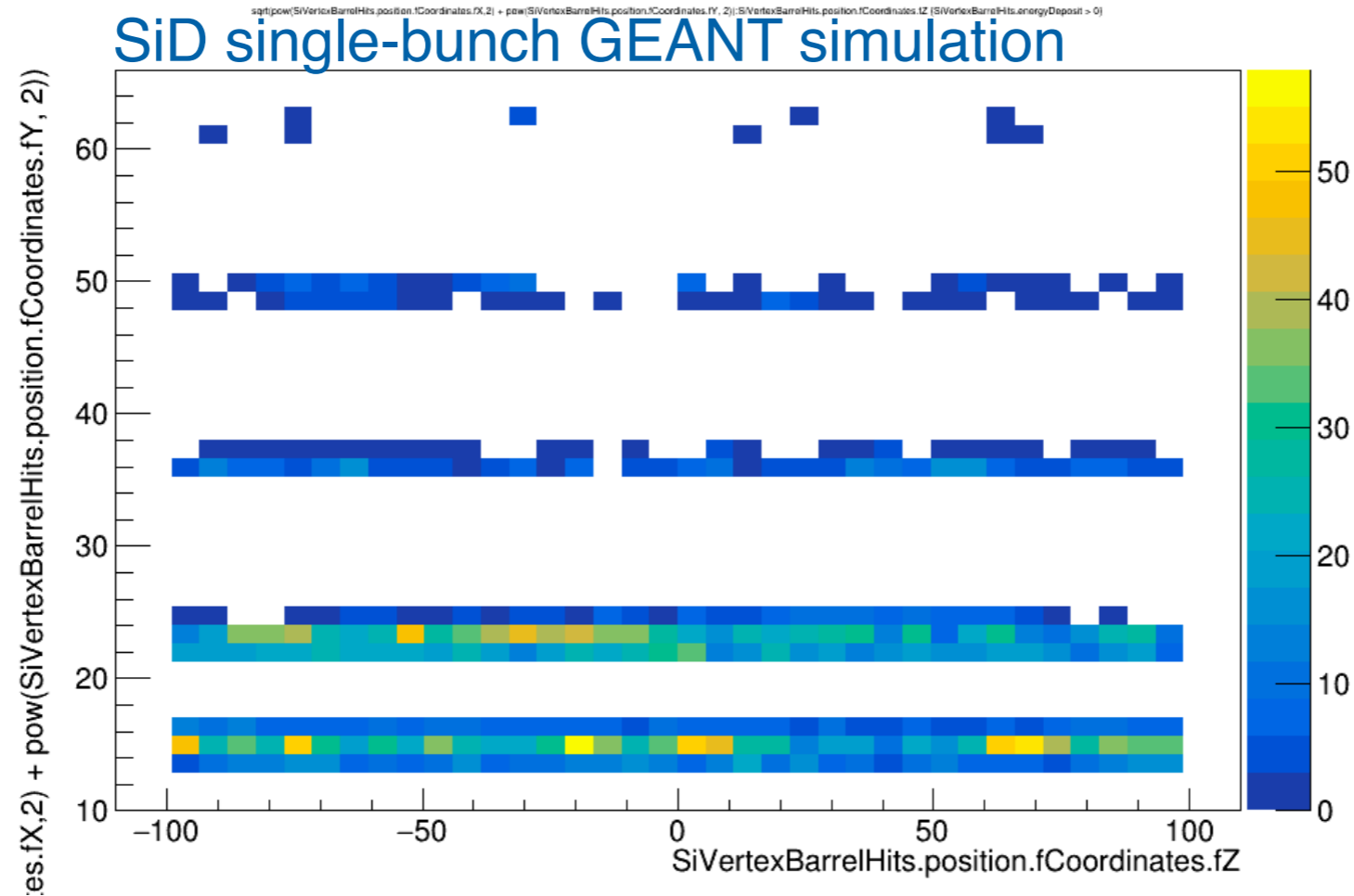
Guinea Pig and C³



Source: https://bib-pubdb1.desy.de/record/405633/files/PhDThesis_ASchuetz_Publication.pdf

- To simulate the pair background we use the Guinea-Pig (GP) program
 - As configured for this study, simulates the primary production modes production of e^+/e^- pairs from beam and beamstrahlung initiated backgrounds
 - There are additional handles for hadron photoproduction but GP's implementation is known to be inaccurate (work beginning on more accurate simulation)

Towards Full Simulation



- Using slightly modified geometry shipped with dd4hep with most recent SiD pixel barrel description
 - ~2000 hits in the first barrel layer in a single bunch (0.24% average occupancy per pixel per train 25x25um)
 - Simulation time is roughly **1 hour** per bunch (several routes for improvement)
- First results ~consistent with back-of-the-envelope expectation
 - Need to check endcaps, occupancy very angle-dependent

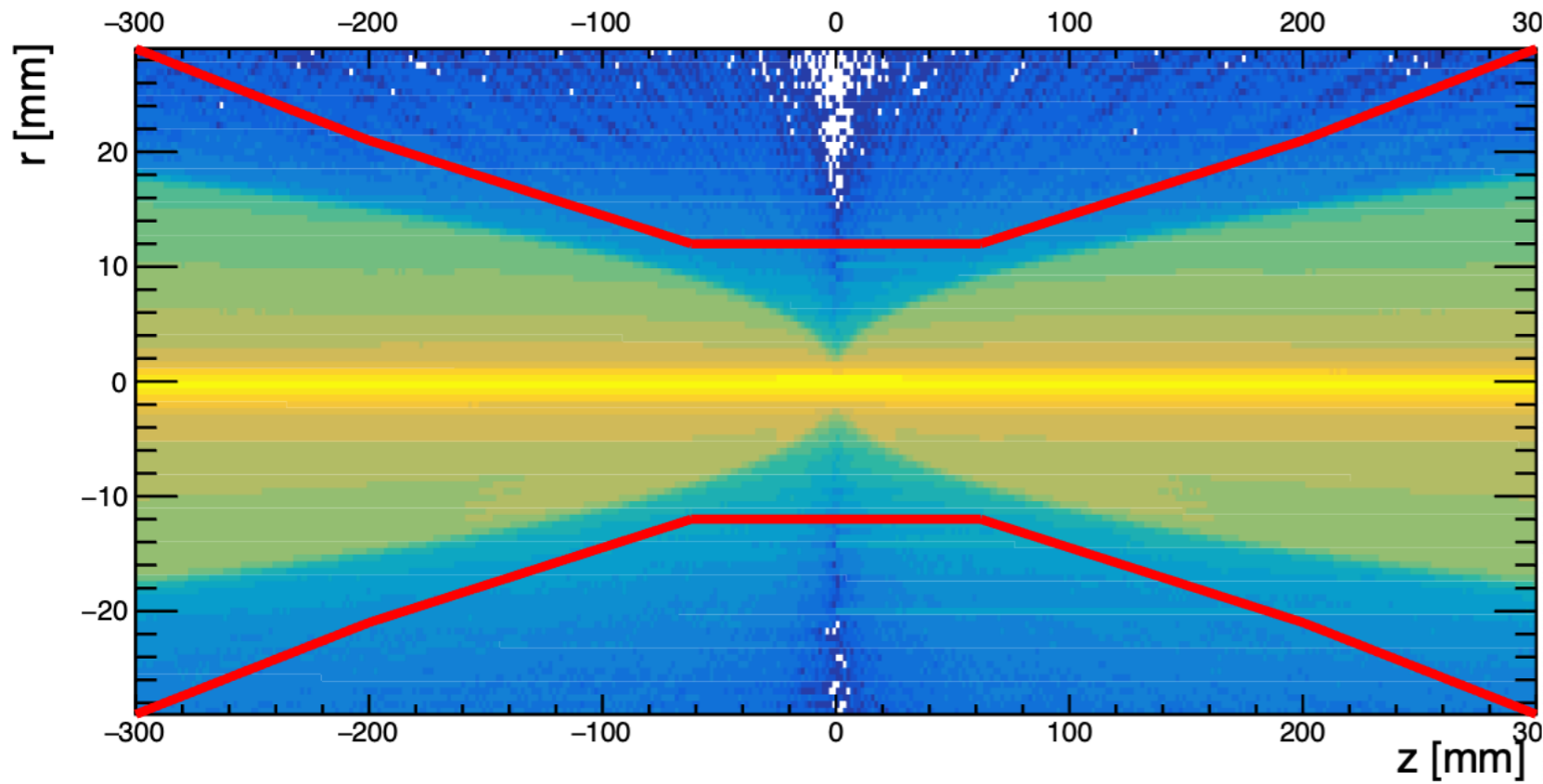
What's new?

Doing better than 1 hour per bunch...

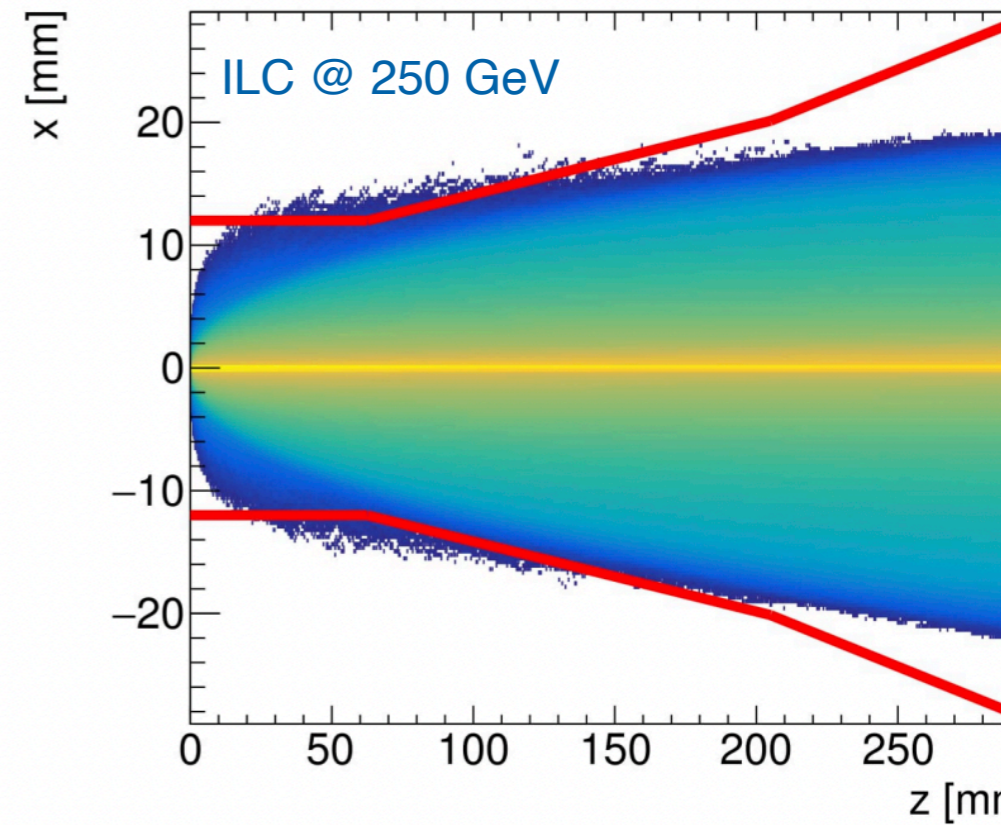
Vertex Barrel layer	Mean number of hits – 0 MeV cut	Mean number of hits – $E > 10$ MeV cut	Mean number of hits – $pT > 5$ MeV cut
1 st layer	341.5 ± 2.6	218.5 ± 2.0	239.0 ± 2.2
2 nd layer	113.9 ± 1.8	101.5 ± 1.7	97.6 ± 1.6
3 rd layer	70.9 ± 1.8	63.2 ± 1.8	51.4 ± 1.7
4 th layer	51.1 ± 1.6	43.4 ± 1.7	38.6 ± 1.5
5 th layer	34.3 ± 1.4	25.1 ± 1.2	20.7 ± 1.2
All 5 layers	614.8 ± 4.2	451.6 ± 4.1	447.3 ± 3.9

- Investigated possibility of cutting out particles that would not reach first layer
 - Simulated 400 random-seed variants of the same event with different kinematic cuts
 - Significant impact on mean number of hits in first layer, not viable for accurate simulations
- Main improvement came from choosing a more efficient random number generator
 - Mersenne Twister took us down to 15 minutes per background event which is at least not glacial

Envelope Plots a la ILC

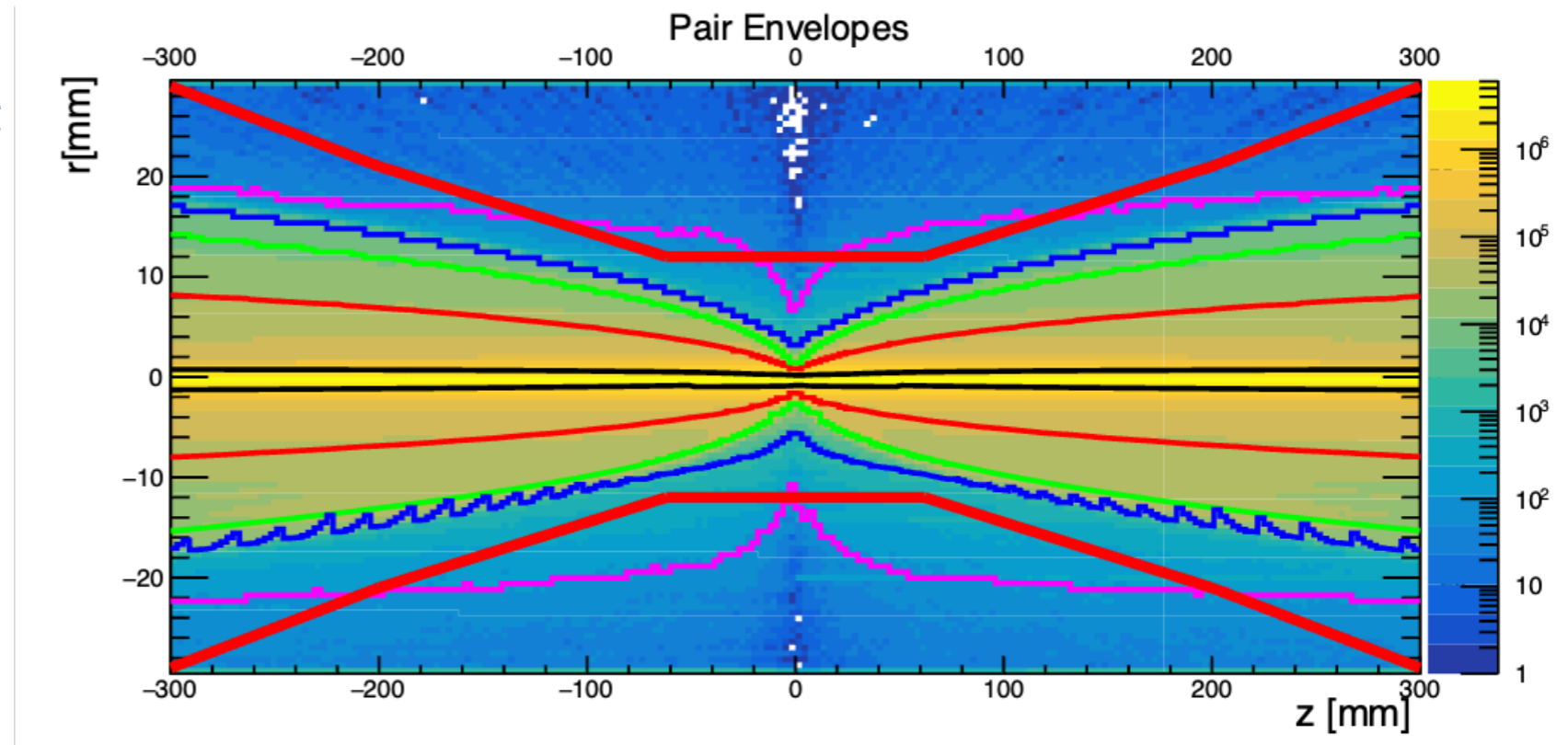


Pairs spiraling in the magnetic field

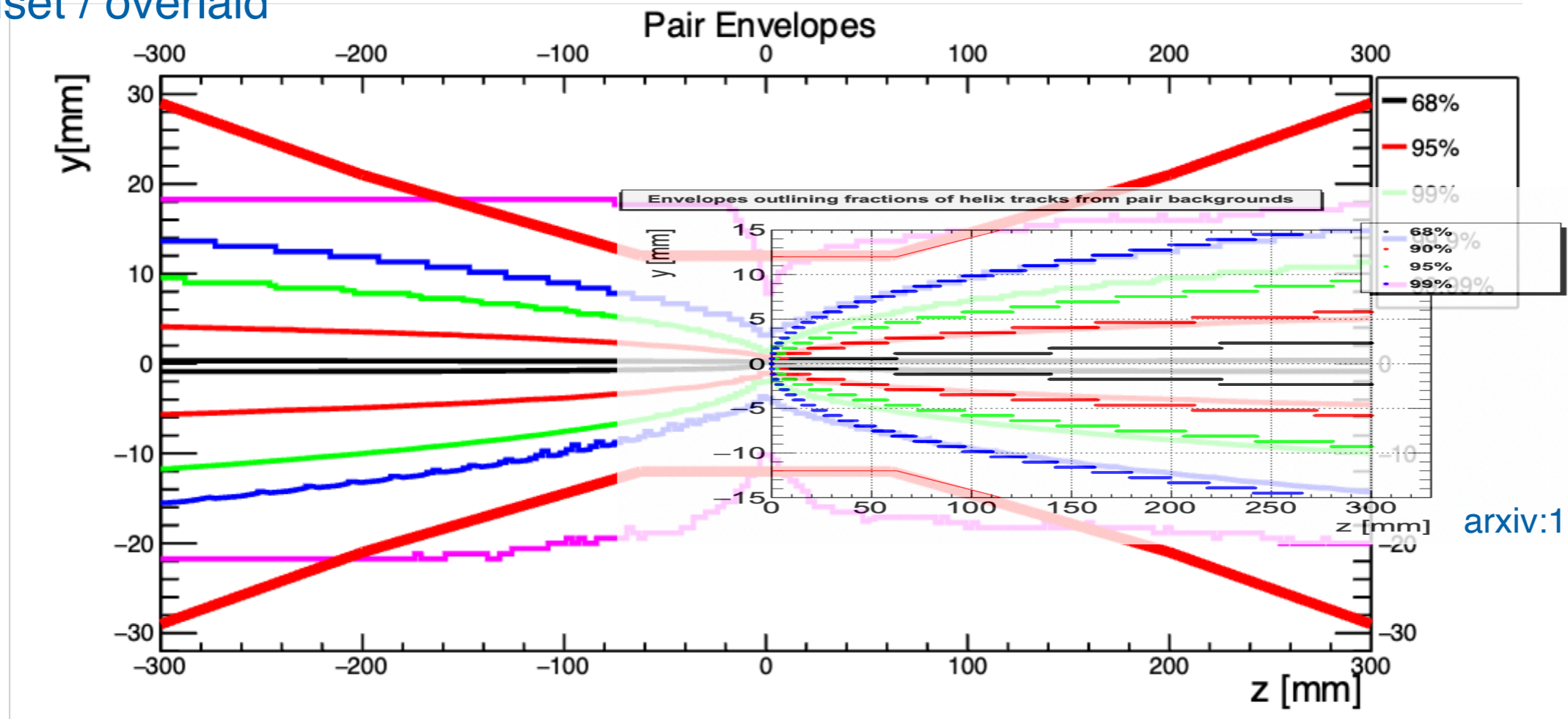


- Red line is latest placement of beam pipe at C³, most recent SiD geometry has first layer at 14mm away from IR
- Qualitatively similar results to ILC in this view...

Envelope Plots a la ILC

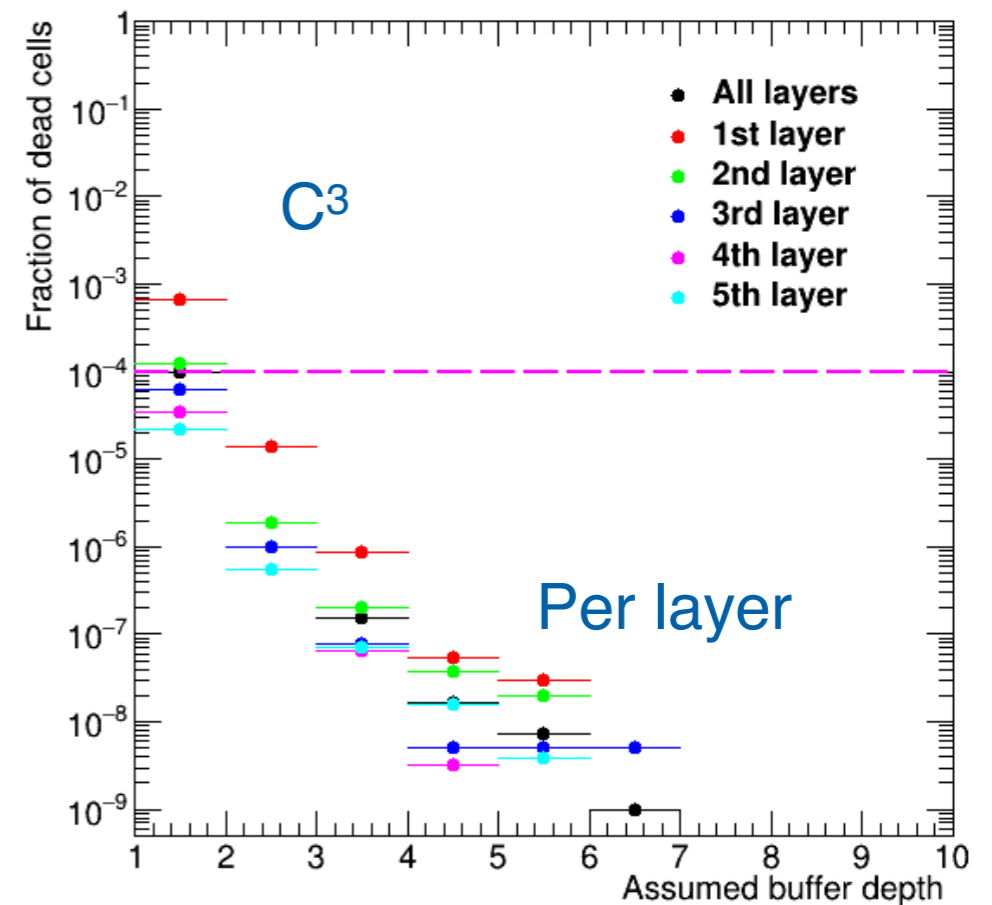
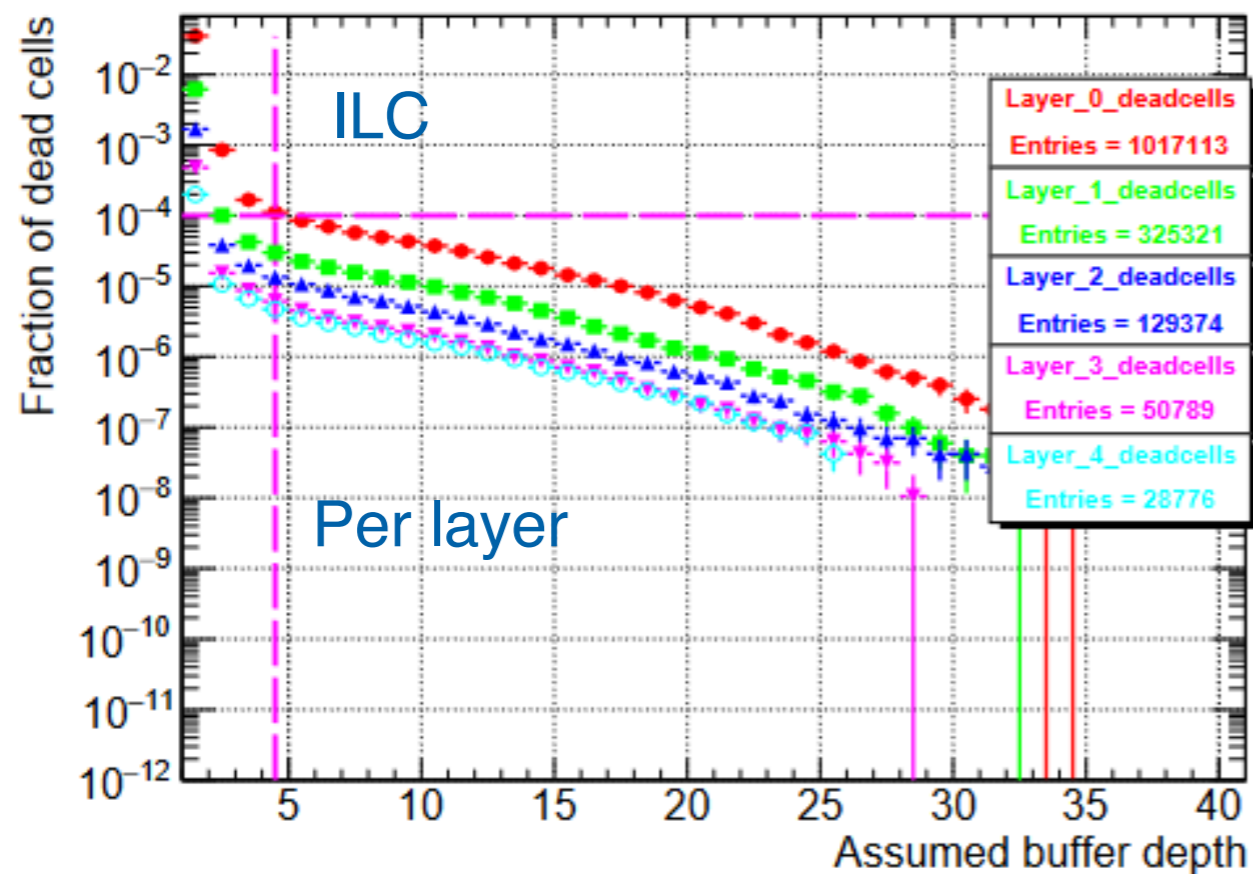


ILC inset / overlaid



arxiv:1703.05737

Latest SiD Geometry in Full Simulation



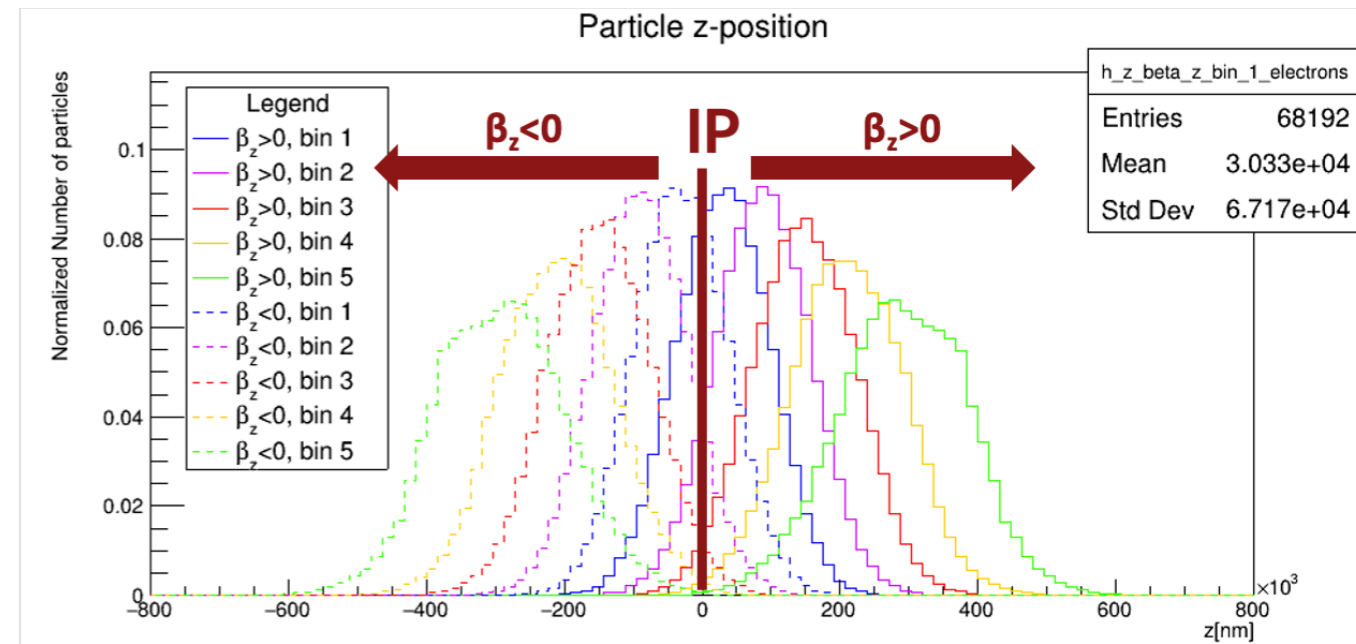
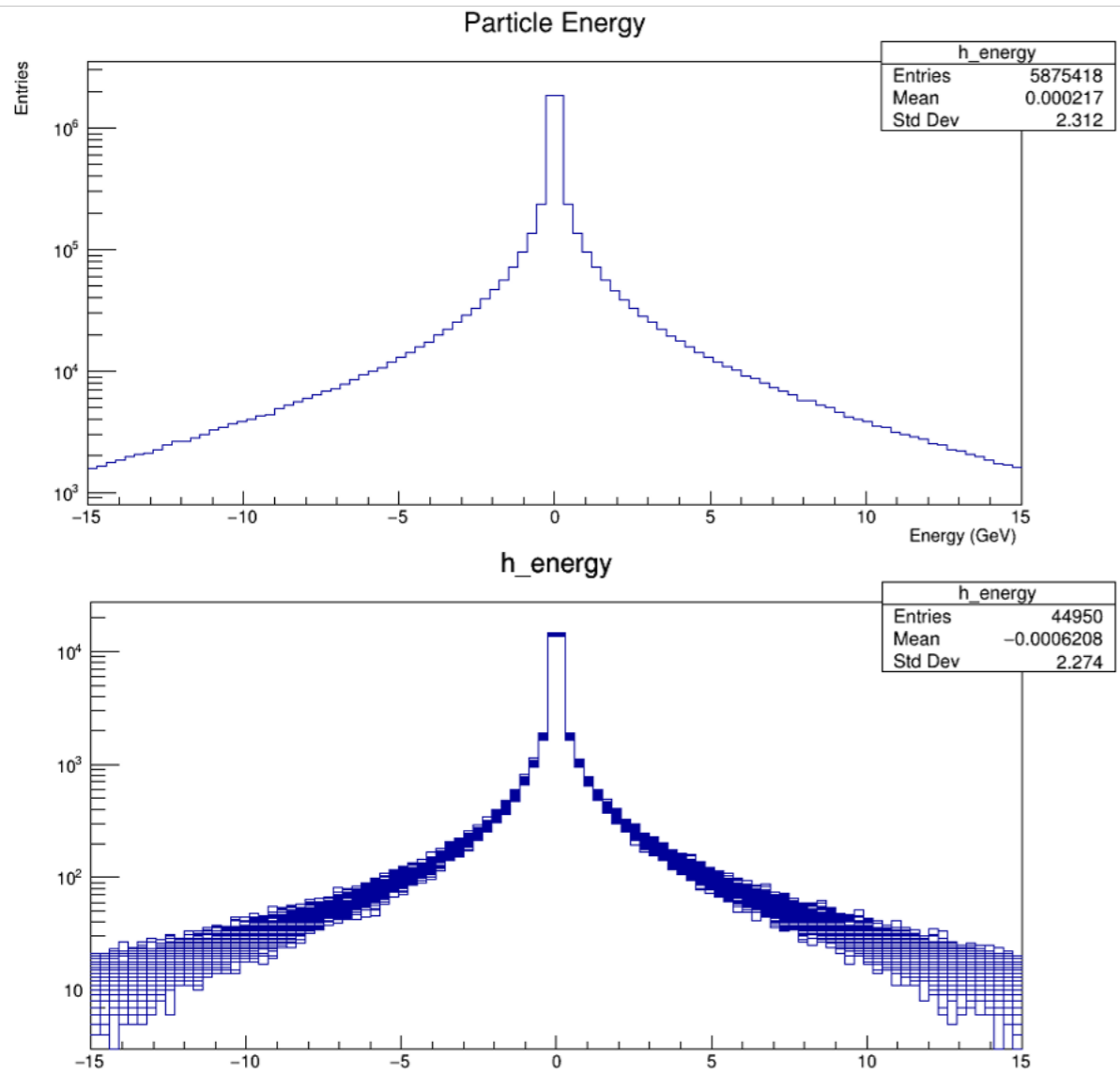
- Checked many times to ensure fidelity of simulation and outcome of results
 - Concerns about magnetic field, exact versions of geometry, etc.
- Together with envelope confirmation indications that we could move the inner pixel layer closer
 - Closer hit: improved sagitta determination, HF tagging, triggering, electron reco.

Conclusions / Plans

- Very close beam-profile match between C³ and ILC 250 (and 500)
 - For a full C³ bunch train, 5T field
- General intuition from ILC background & physics studies applicable
 - We will continue to collect the revitalize all previous backgrounds
 - Will store copies and recreation instructions for all relevant code
- I think we are in a place to confidently back up the C³ physics case with prior studies
- Your scrutiny, feedback, and participation is appreciated! Thanks in advance!

Extras

Raw GP Results



Distribution of the z-position of beam-induced e^+/e^- for the 133 simulated BXs for different bins of β_z :

- bin 1: $0.0 < |\beta_z| < 0.2$
- bin 2: $0.2 < |\beta_z| < 0.4$
- bin 3: $0.4 < |\beta_z| < 0.6$
- bin 4: $0.6 < |\beta_z| < 0.8$
- bin 5: $0.8 < |\beta_z| < 1.0$

- We generated 133 bunches configured with the C³ parameters ensuring unique random seeds to simulate a full bunch train
 - Simulation of e^+/e^- propagation through bunch charge is apparent and consistent with expectations
 - Sub-distributions per bunch consistent with each other
 - Average of 44176 particles per bunch, observed expected steeply falling energy spectrum