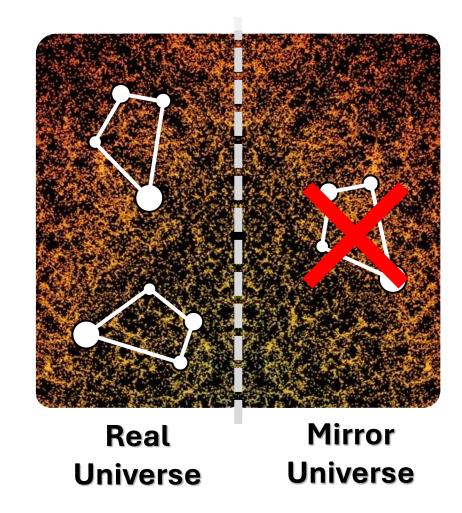


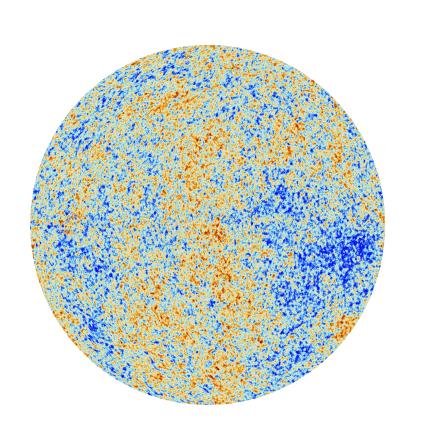
What is Parity Violation?

Different physics after a parity inversion: $(x, y, z) \rightarrow (-x, -y, -z)$

We look for different **patterns**



Why Search for Parity Violation?





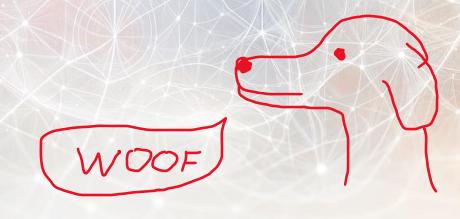


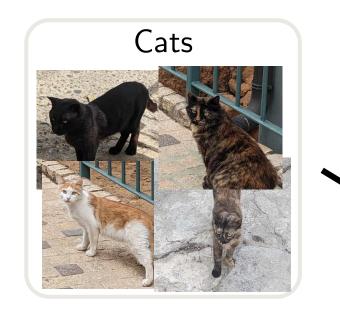


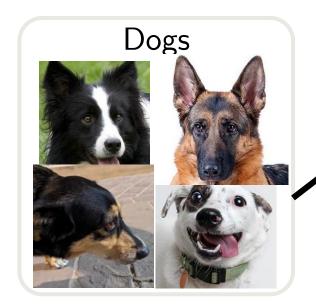
ATS ADOGS

Detecting Parity Violation with Unsupervised Learning

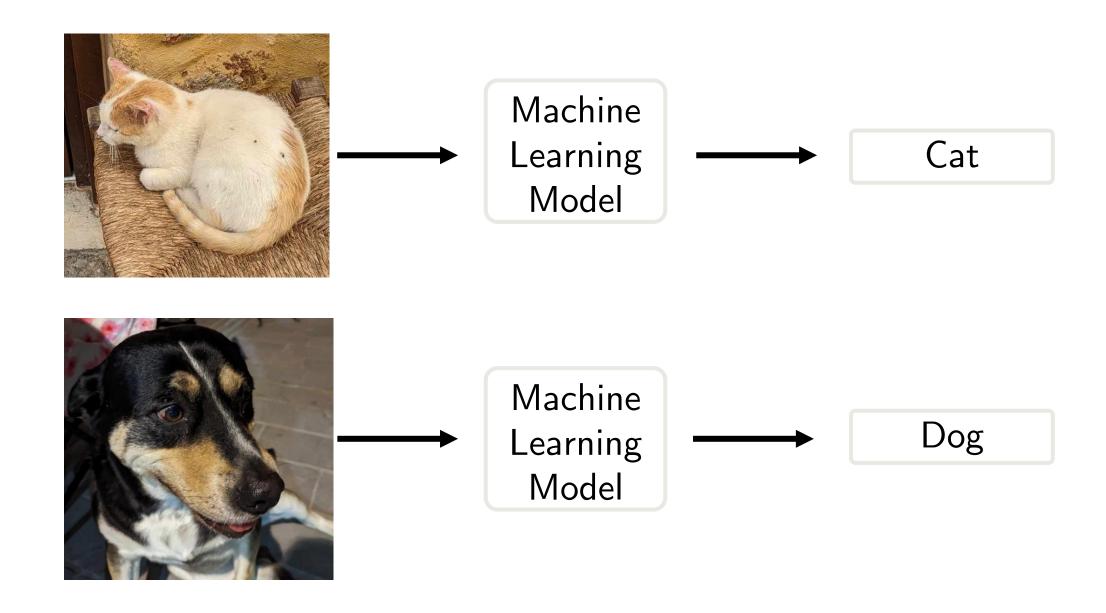




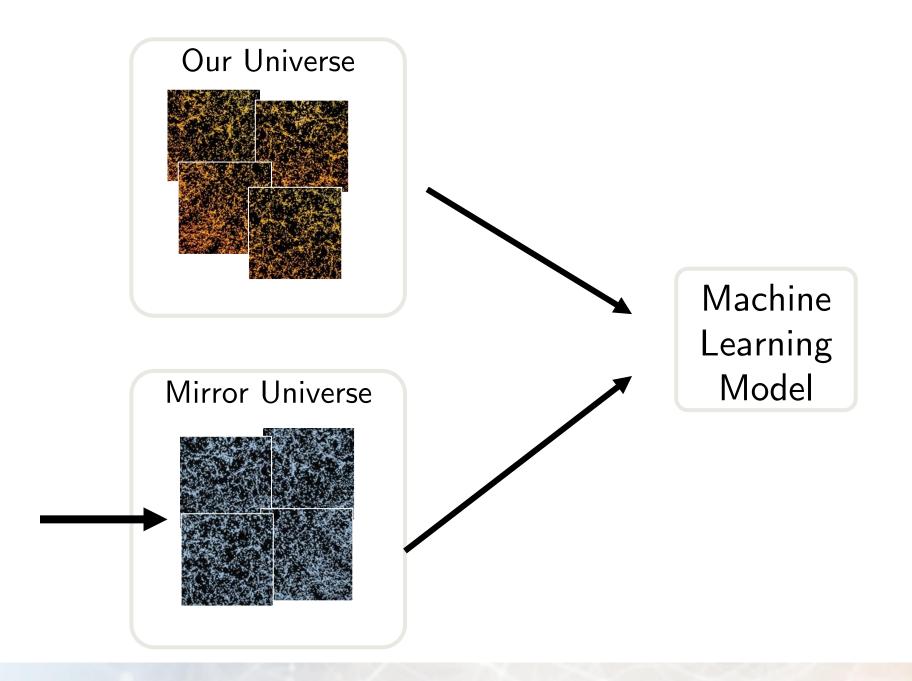


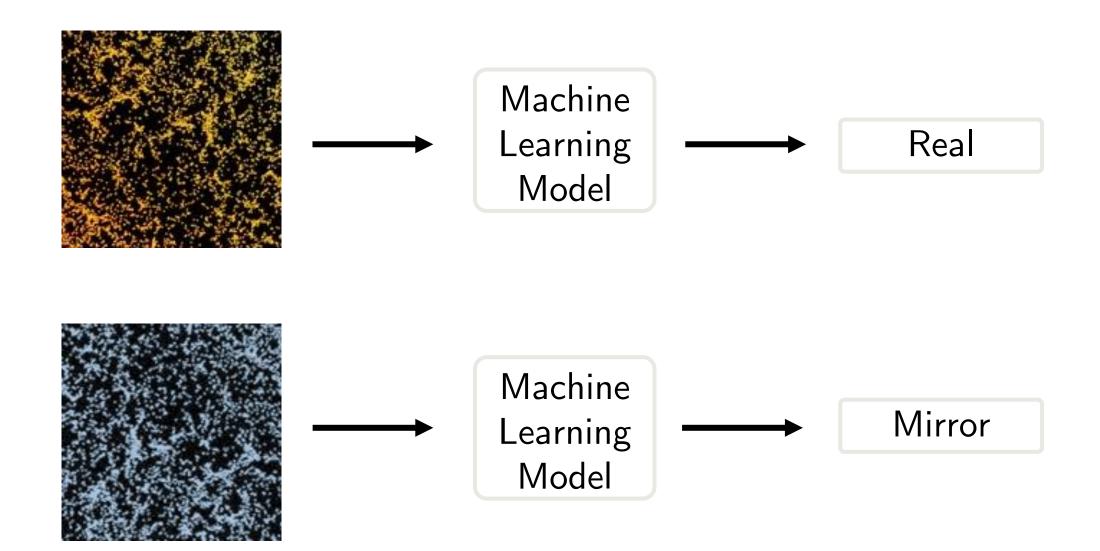


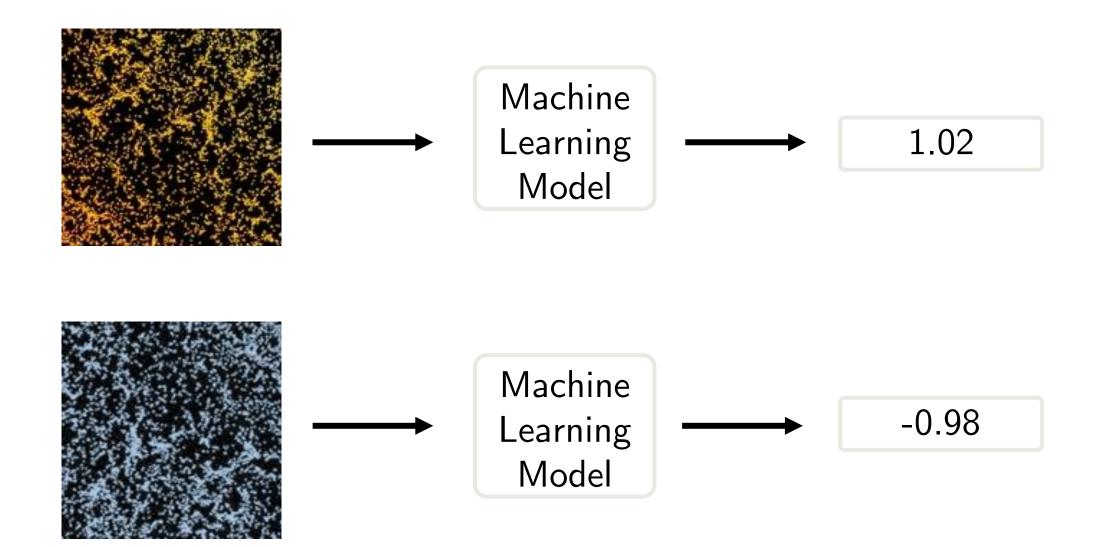
Machine Learning Model



Our Universe Machine Learning Model Mirror Universe



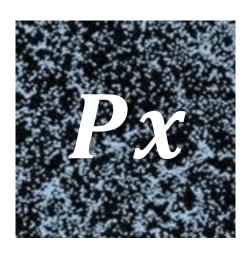






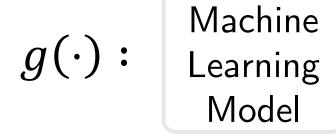






 $g(\cdot)$: Machine Learning Model





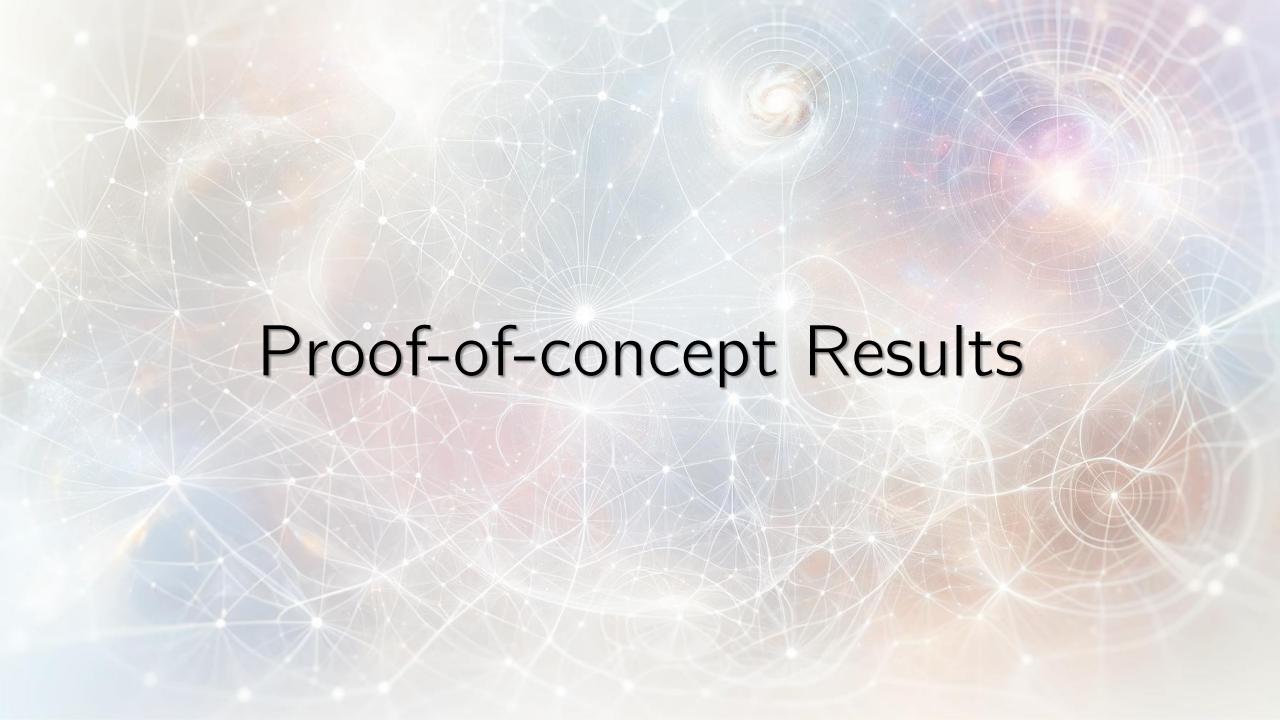


The difference: g(x) - g(Px)

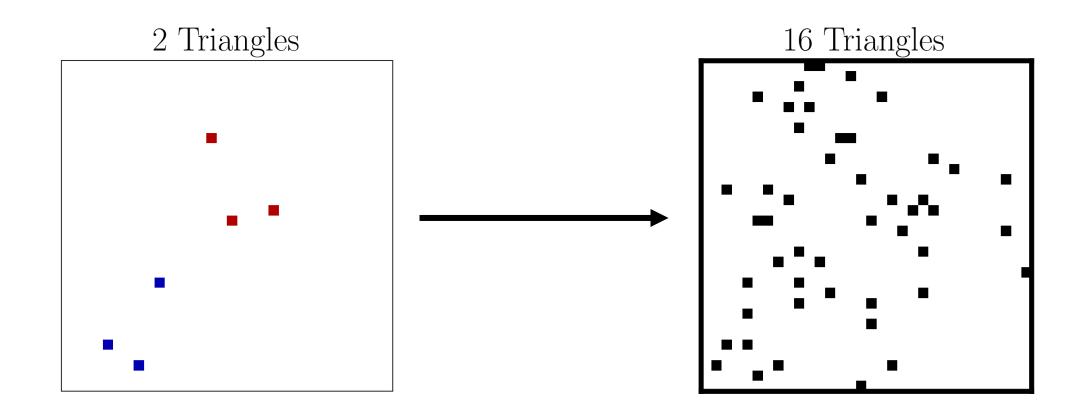
How do we make a detection?

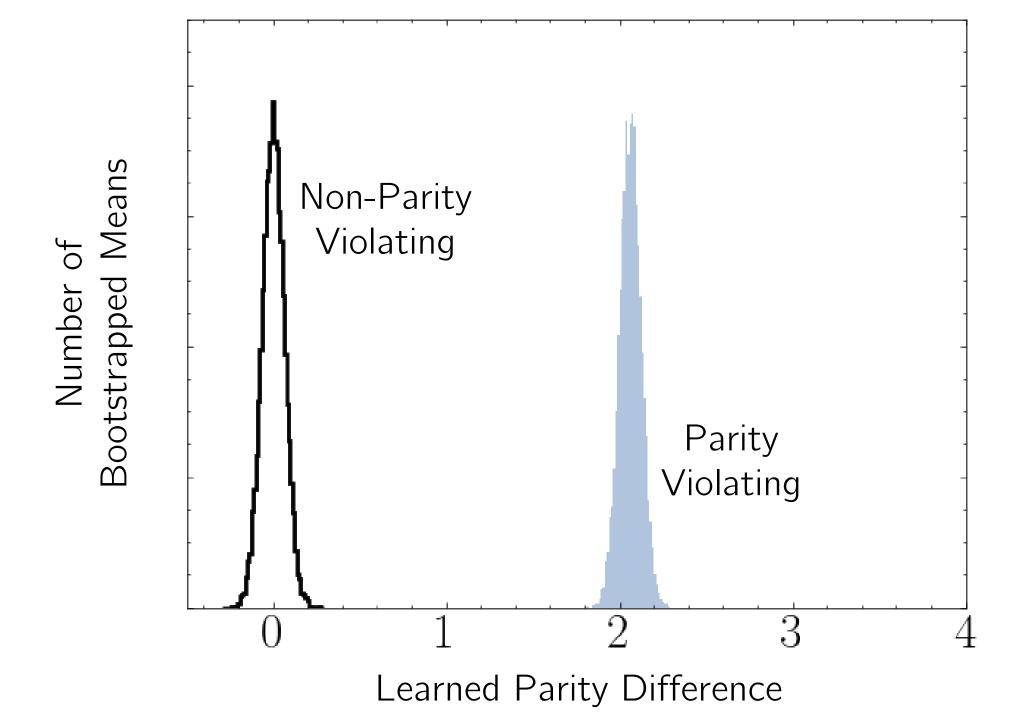
Look for a difference in *unseen* test data, $g(x) - g(Px) \neq 0$

- 1. Calculate the mean g(x) g(Px) over test data
- 2. Bootstrap resample to understand natural variability
- 3. Calculate a σ -level of detection



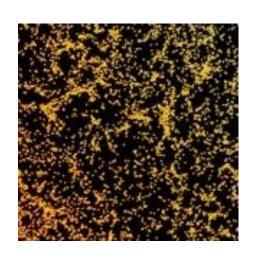
Simplified Data



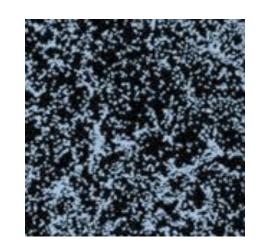


The perks of Unsupervised Learning

1. It's simple and familiar



VS

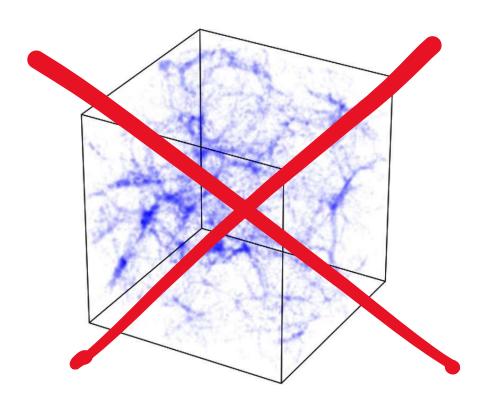






The perks of Unsupervised Learning

2. We only need observational data



The perks of Unsupervised Learning

3. No false detections

$$g(x) - g(Px) \neq 0$$



Future Directions

- 1. Extend to 3D
- 2. Search in simulations with known parity violation
- 3. Search in the real universe

- Parity violation might be present in the galaxy distribution.
- Unsupervised learning is a good way to find it, with many perks!
- Look out for the 3D application soon!

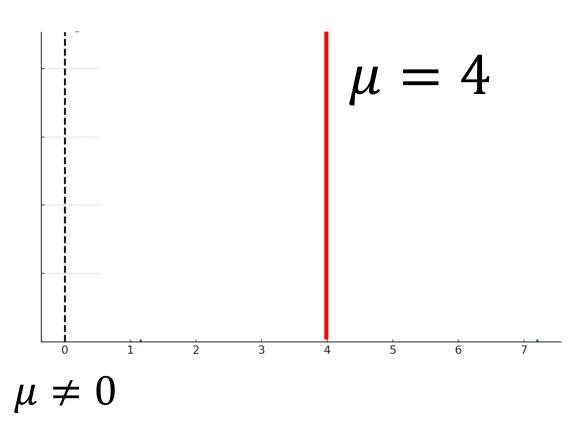
arxiv: 2405.13083

arxiv: 2312.09287



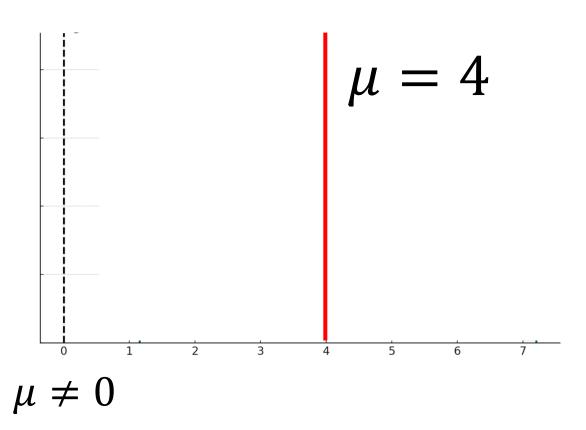
EXTRAS

 $\mu = \text{Mean}[g(x) - g(Px)]$ over test set



$$\mu = \text{Mean}[g(x) - g(Px)]$$
 over test set

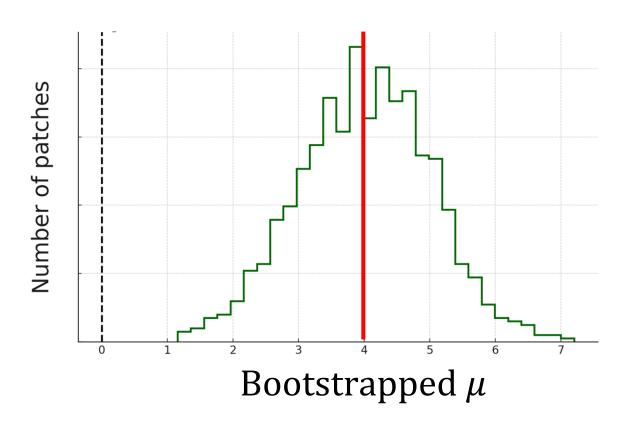
2. Is this nonzero because of signal or just noise?



$$\mu = \text{Mean}[g(x) - g(Px)]$$
 over test set

2. Is this nonzero because of signal or just noise?

 $\mu^* = \text{Bootstrapped } \mu \text{ over test set}$

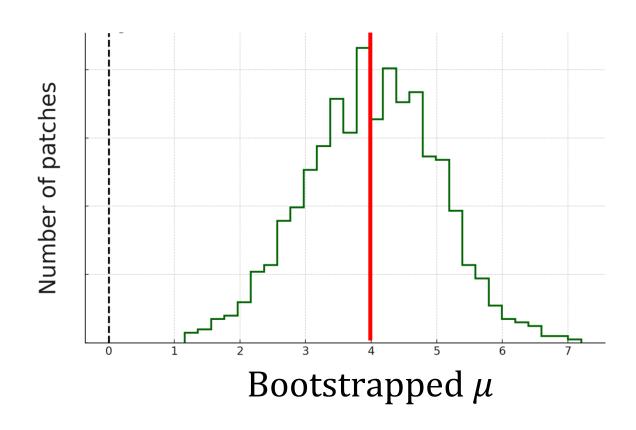


$$\mu = \text{Mean}[g(x) - g(Px)]$$
 over test set

2. Is this nonzero because of signal or just noise?

 $\mu^* = \text{Bootstrapped } \mu \text{ over test set}$

3. How confident are we in this detection?



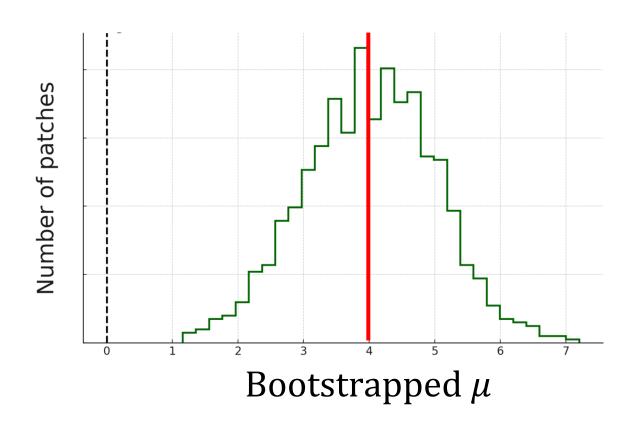
$$\mu = \text{Mean}[g(x) - g(Px)]$$
 over test set

2. Is this nonzero because of signal or just noise?

$$\mu^* = \text{Bootstrapped } \mu \text{ over test set}$$

3. How confident are we in this detection?

$$\eta = \frac{\text{mean}[\mu^*]}{\text{std}[\mu^*]}$$



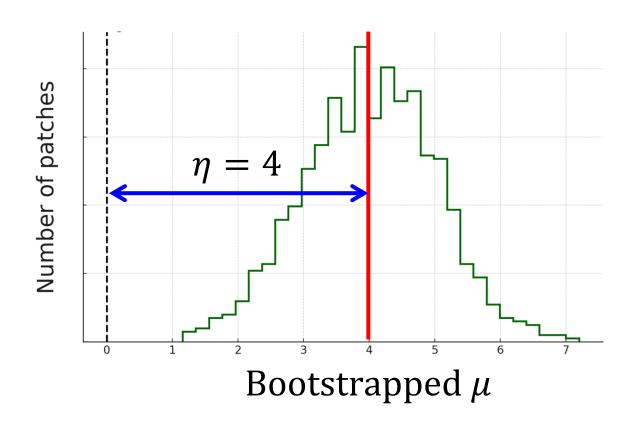
$$\mu = \text{Mean}[g(x) - g(Px)]$$
 over test set

2. Is this nonzero because of signal or just noise?

$$\mu^* = \text{Bootstrapped } \mu \text{ over test set}$$

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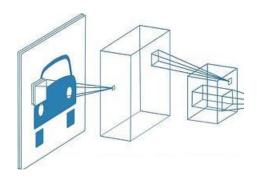
$$\eta = \frac{\text{mean}[\mu^*]}{\text{std}[\mu^*]}$$



g(x)Fully Connected Network Data Compression Field Parity Summary Statistic Statistics

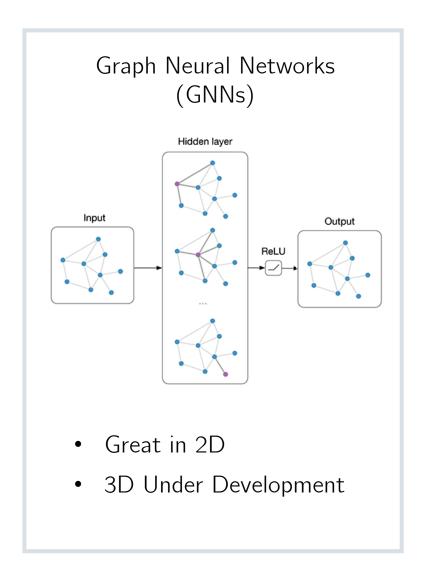
Types of Compression Model

Convolutional Neural Networks (CNNs)

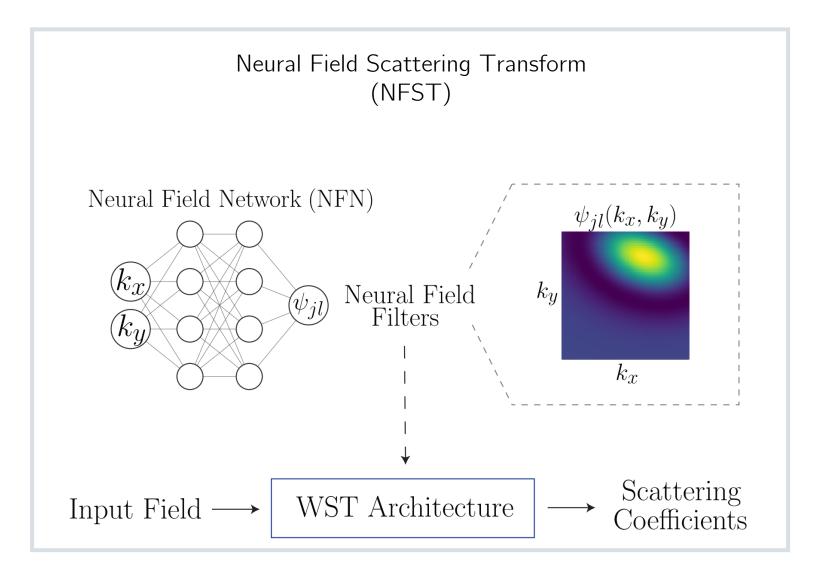


- Very flexible
- Fails in 3D (overtraining)

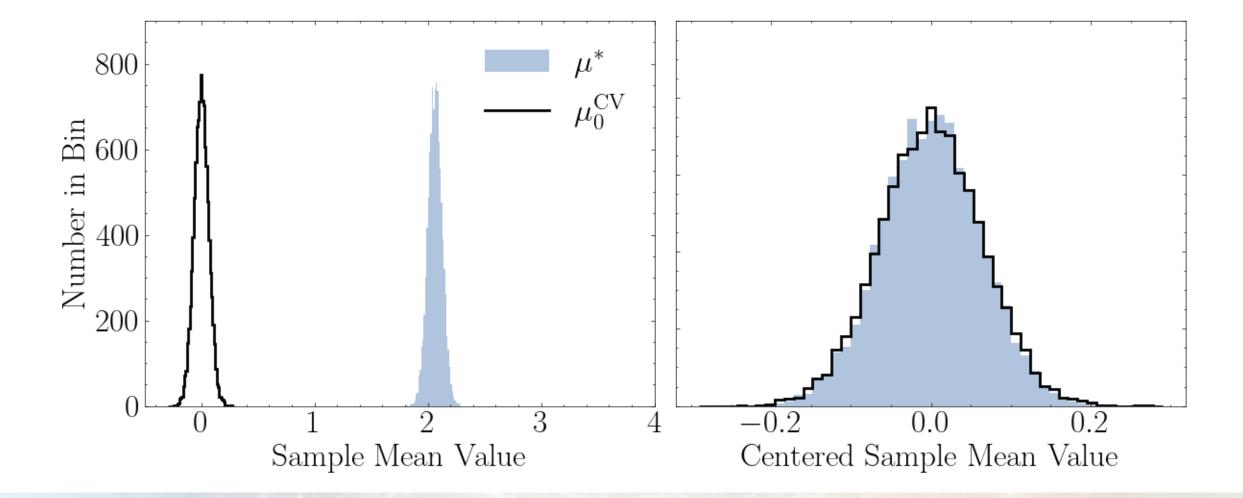
Wavelet Scattering Transforms (WSTs) Real Imag **Fixed Convolution Filter** Very robust - Symmetries Not flexible enough

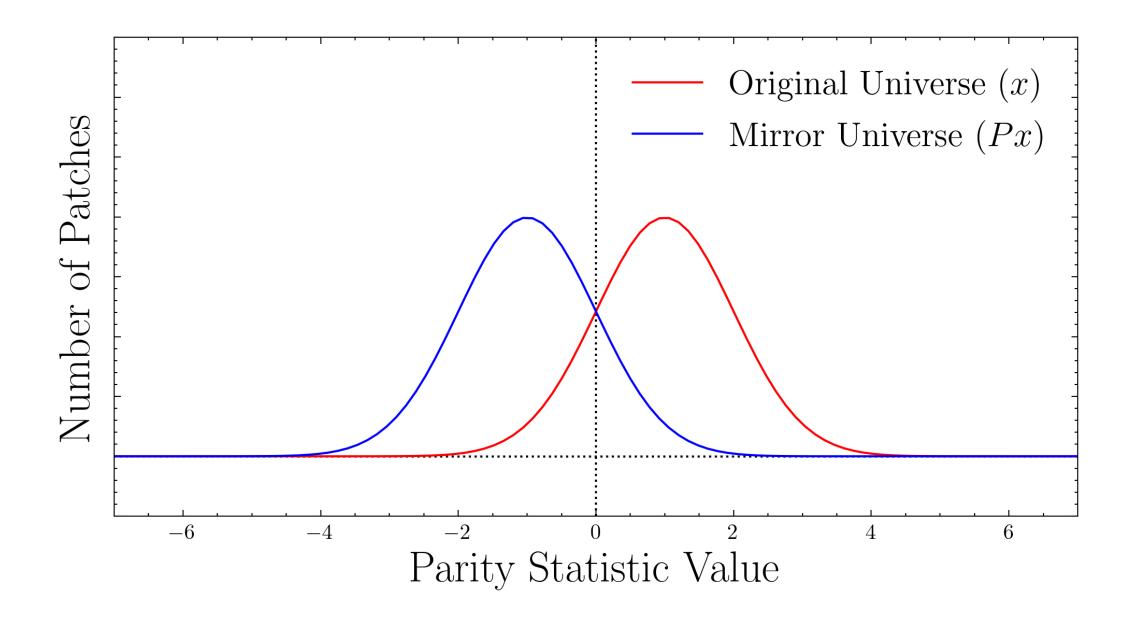


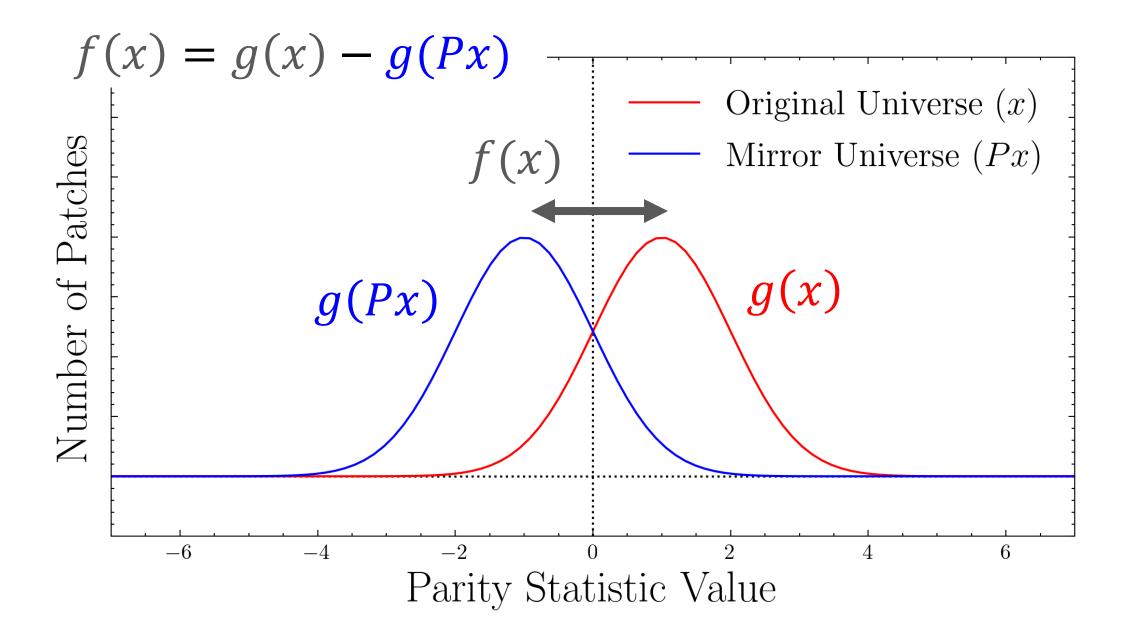
Types of Compression Model

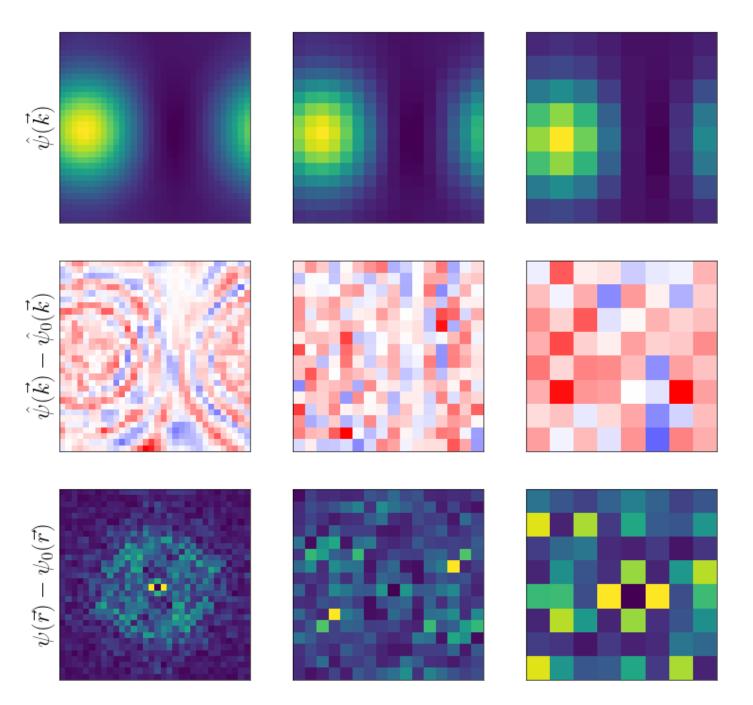


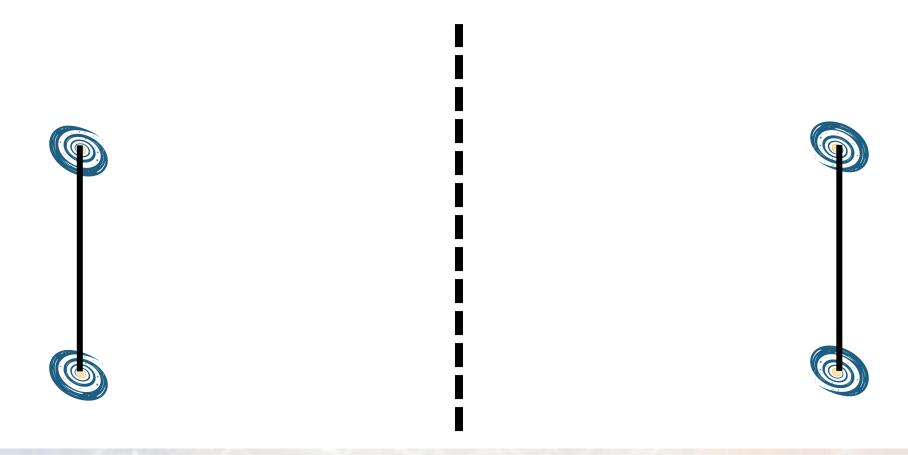
- Flexible
- Robust
- Works in 3D!

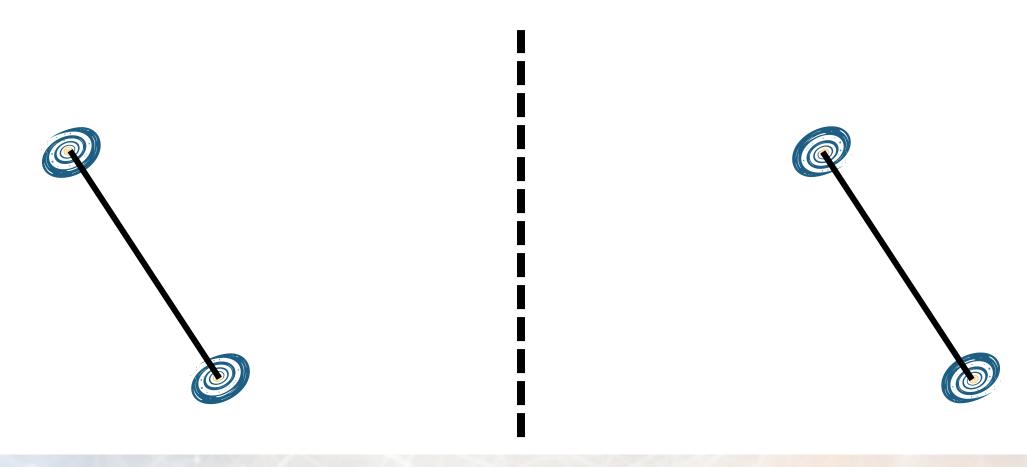


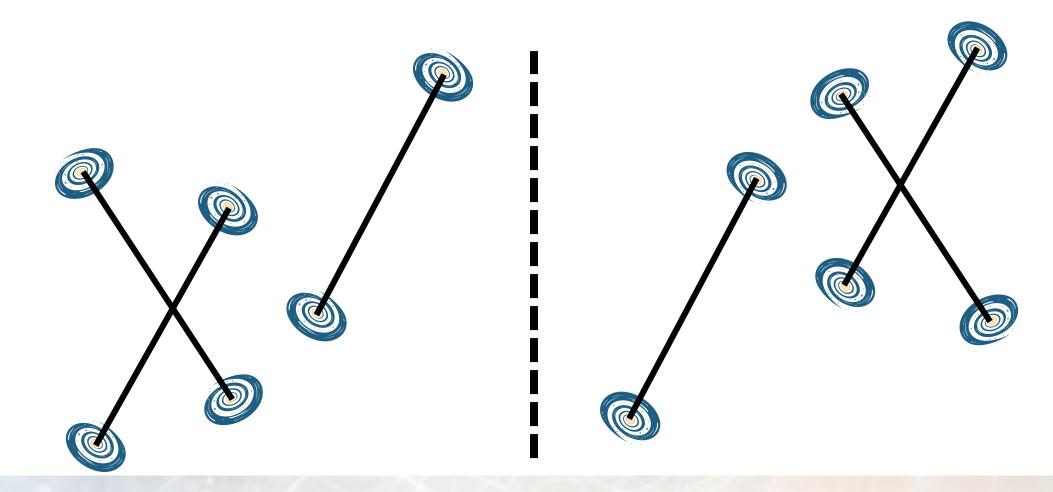




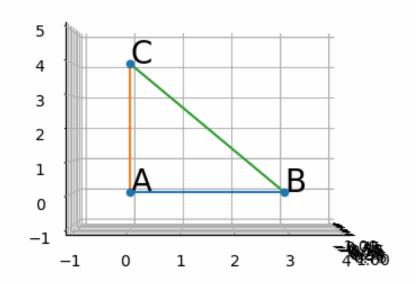


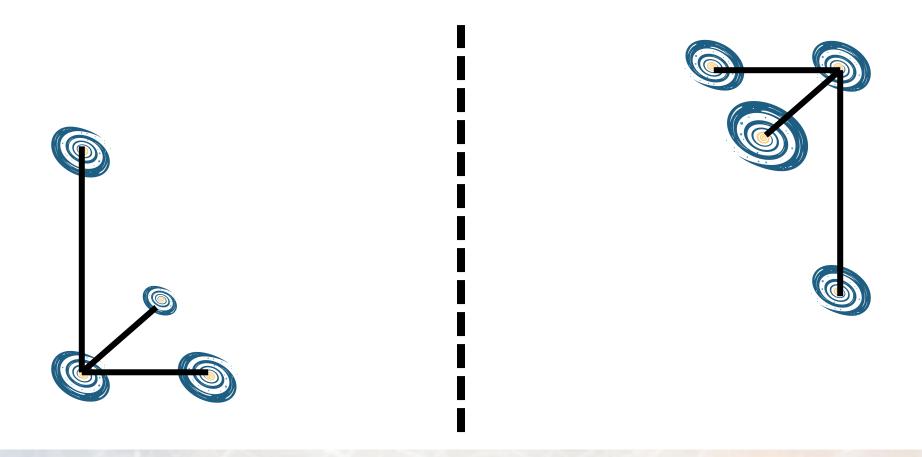


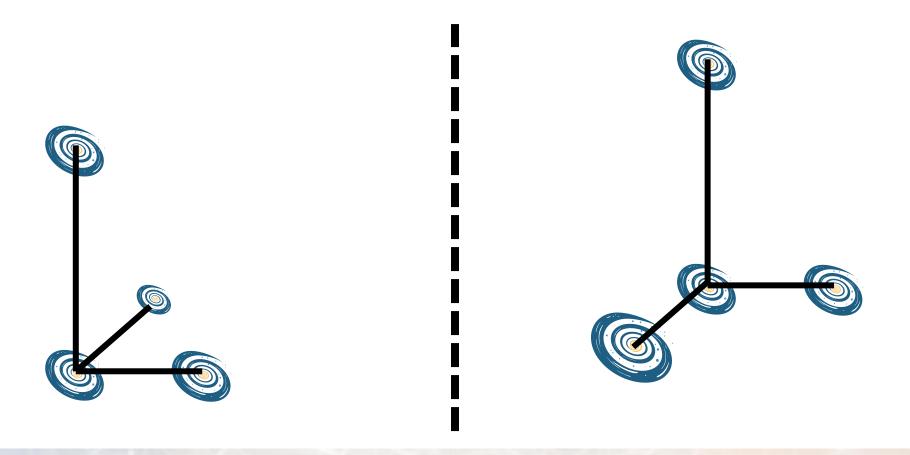




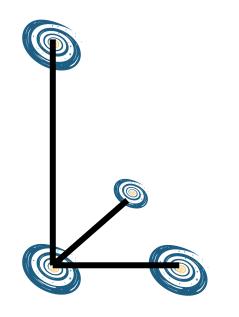




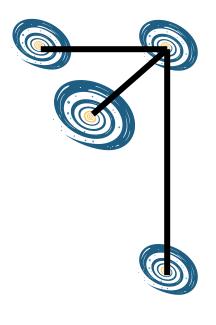




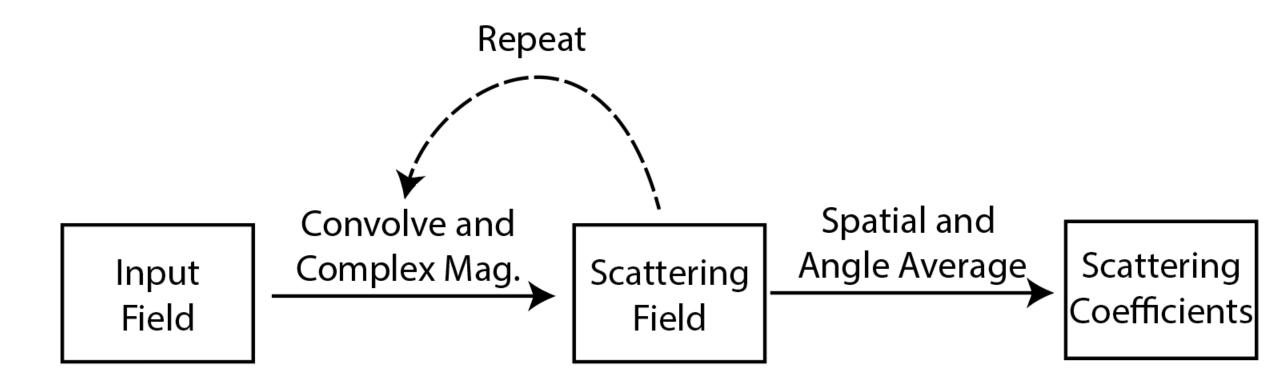
In the case of parity violation:



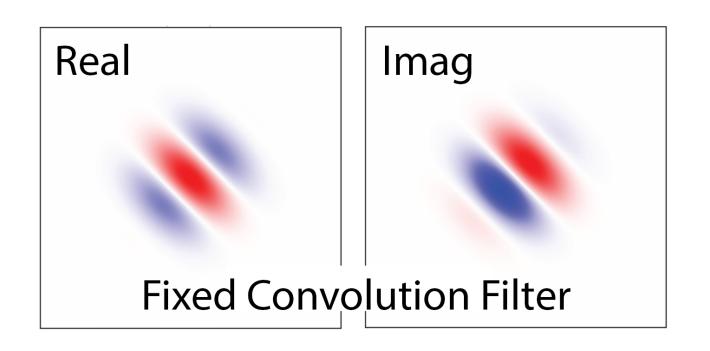


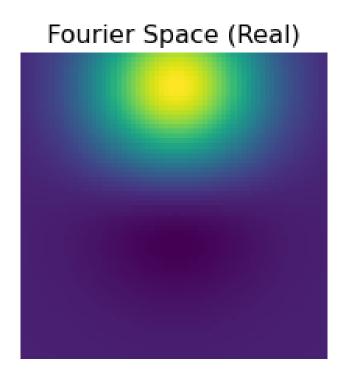


The Wavelet Scattering Transform



The Wavelet Scattering Transform

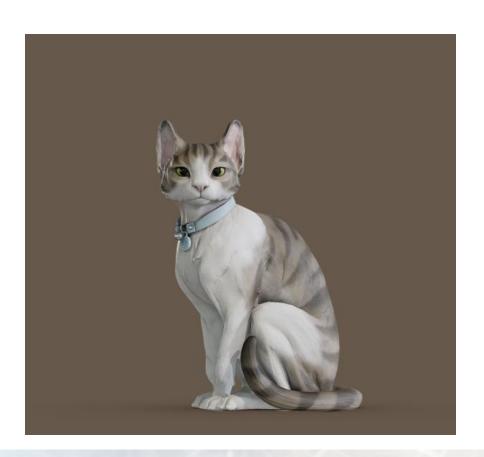




Parity Violation in the Galaxy Distribution

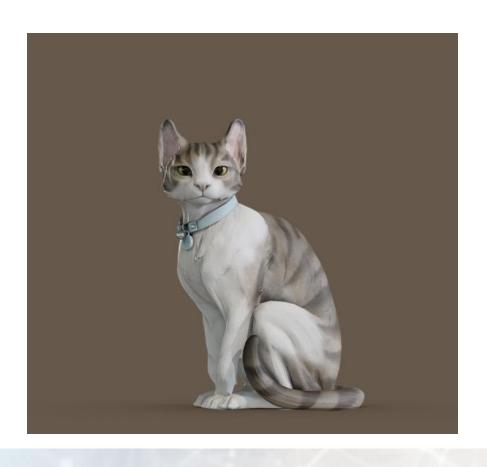
The Parity Inversion

$$(x, y, z) \rightarrow (-x, -y, -z)$$



The Parity Inversion

$$(x, y, z) \rightarrow (-x, -y, -z)$$





The Parity Inversion

$$(x, y, z) \rightarrow (-x, -y, -z)$$



