

Cosmic Inflation at the Crossroads

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Outline

Qualitative Inflation

Quantitative inflation

Bayesian inference

Qualitative Inflation

- What is inflation?
- Triggering and stopping acceleration
- Cosmological perturbation

Quantitative inflation

- Encyclopædia Inflationaris: Opiparous Edition
- Primordial power spectra
- Reheating consistent model predictions

Bayesian inference

- Machine-learning an effective likelihood
- Marginalized posteriors
- Computing bayesian evidences
- Bayes factors for all models
- Information gain on the reheating
- Vanilla inflation predicts. . .
- Conclusion

J. Martin, CR and V. Vennin: [arXiv:2404.10647](#), [arXiv:2404.15089](#),
[arXiv:1303.3787v4](#)

Qualitative Inflation

- ❖ What is inflation?
- ❖ Triggering and stopping acceleration
- ❖ Cosmological perturbation

Quantitative inflation

Bayesian inference

Qualitative Inflation

What is inflation?

Qualitative Inflation

❖ What is inflation?

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Bayesian inference

- Predictive, testable and tested early universe paradigm
 - ◆ Hypothetical accelerated expansion of the universe at $E_{\text{inf}} > \text{MeV}$ (BBN)
 - ◆ Addresses some unexplainable features of the Friedmann-Lemaître model
- For the simplest incarnation of inflation...
 - ◆ Historically introduced to dilute **monopoles** formed at GUT
 - ◆ **Flatness** of the spatial sections ($\Omega_K = 0.0009 \pm 0.0018$)
 - ◆ **Statistical isotropy** of the observable universe (horizon problem)
 - ◆ **Origin** of CMB and LSS (quantum fluctuations)
 - ◆ **Gaussianities** of the cosmological perturbations ($f_{\text{NL}} < -0.9 \pm 5$)
 - ◆ **Adiabaticity** of the cosmological perturbations (isocurv. $< 1\%$)
 - ◆ **Almost scale invariance** ($n_s = 0.9649 \pm 0.004$)

Triggering and stopping acceleration

Qualitative Inflation

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- The simplest way: single-field inflation

$$S = \int dx^4 \sqrt{-g} \left[\frac{1}{2\kappa^2} R + \mathcal{L}(\phi) \right] \quad \text{with} \quad \mathcal{L}(\phi) = -\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V(\phi)$$

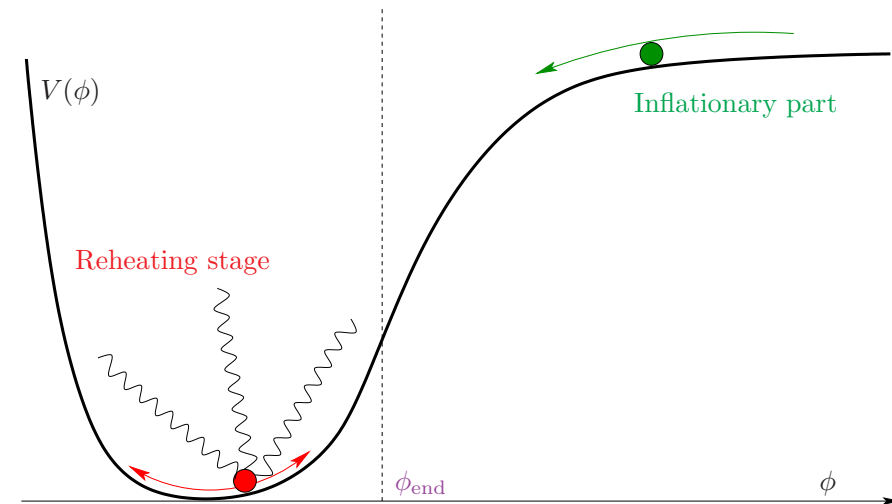
- Inflation occurs in the plateau and is followed by a **reheating** era

◆ Friedmann-Lemaître

$$H^2 = \frac{1}{3} \left(\frac{1}{2} \dot{\phi}^2 + V \right)$$

$$\frac{\ddot{a}}{a} = -\frac{1}{3} \left(\dot{\phi}^2 - V \right)$$

◆ $H \simeq \text{Constant} \rightarrow a \propto e^{Ht}$



- The reheating stage: everything after ϕ_{end} till radiation domination

Cosmological perturbations of quantum origin

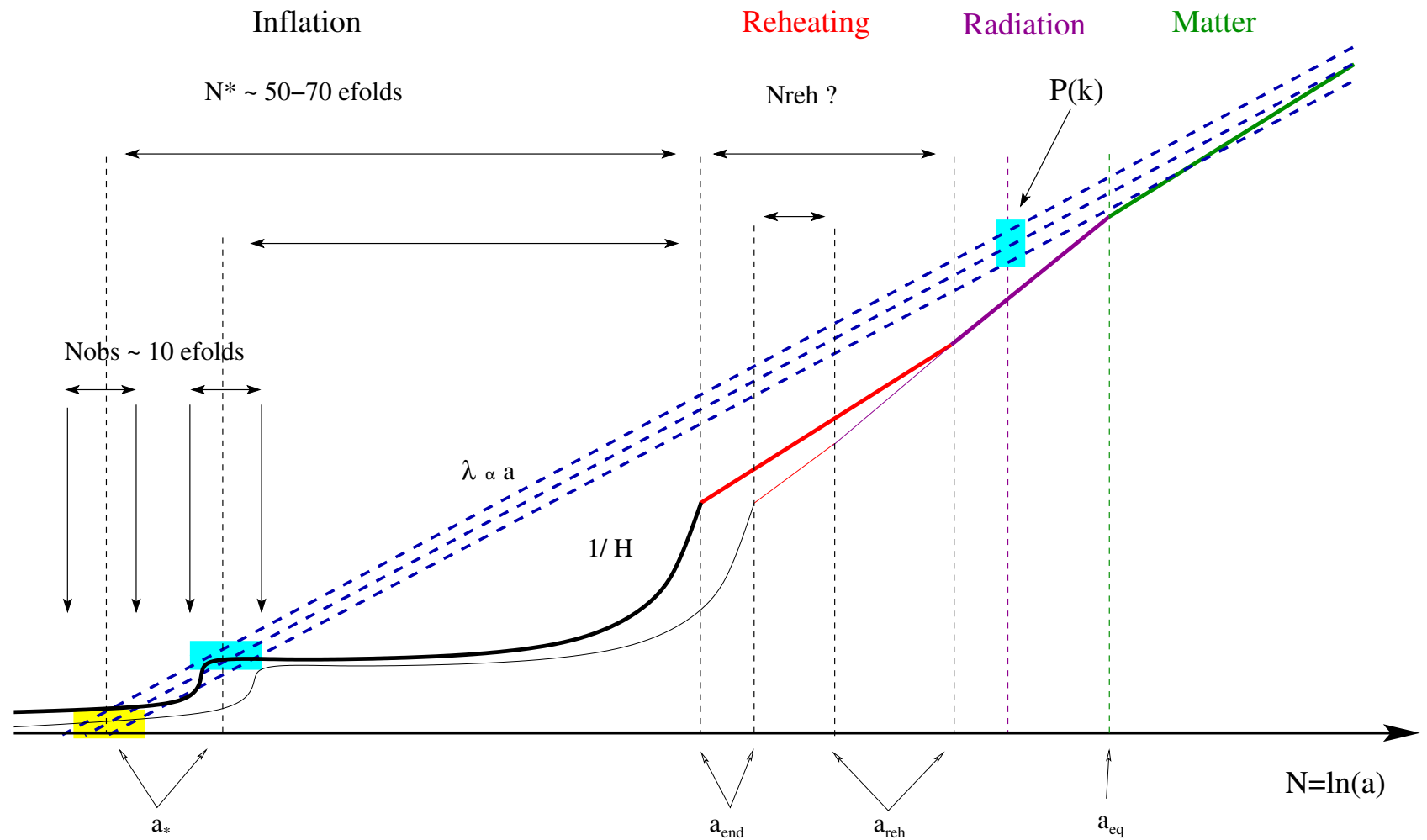
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- **Model testing:** reheating effects must be included!

Redshift at which inflation ends

- Depends on the redshift of reheating

$$1 + z_{\text{end}} = \frac{a_0}{a_{\text{end}}} = \frac{a_{\text{reh}}}{a_{\text{end}}} (1 + z_{\text{reh}}) = \frac{a_{\text{reh}}}{a_{\text{end}}} \left(\frac{\rho_{\text{reh}}}{\tilde{\rho}_\gamma} \right)^{1/4} = \frac{1}{R_{\text{rad}}} \left(\frac{\rho_{\text{end}}}{\tilde{\rho}_\gamma} \right)^{1/4}$$

- ◆ The reheating parameter $R_{\text{rad}} \equiv \frac{a_{\text{end}}}{a_{\text{reh}}} \left(\frac{\rho_{\text{end}}}{\rho_{\text{reh}}} \right)^{1/4}$
- ◆ Encodes **any observable deviations** from a radiation-like or instantaneous reheating $R_{\text{rad}} = 1$

- R_{rad} can be expressed in terms of $(\rho_{\text{reh}}, \bar{w}_{\text{reh}})$ or $(\Delta N_{\text{reh}}, \bar{w}_{\text{reh}})$

$$\ln R_{\text{rad}} = \frac{\Delta N_{\text{reh}}}{4} (3\bar{w}_{\text{reh}} - 1) = \frac{1 - 3\bar{w}_{\text{reh}}}{12(1 + \bar{w}_{\text{reh}})} \ln \left(\frac{\rho_{\text{reh}}}{\rho_{\text{end}}} \right)$$

- Scary Astrophysics (early universe)

$$10^{10} < z_{\text{end}} < 10^{28}, \quad -46 < \ln R_{\text{rad}} < 15$$

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Quantitative inflation

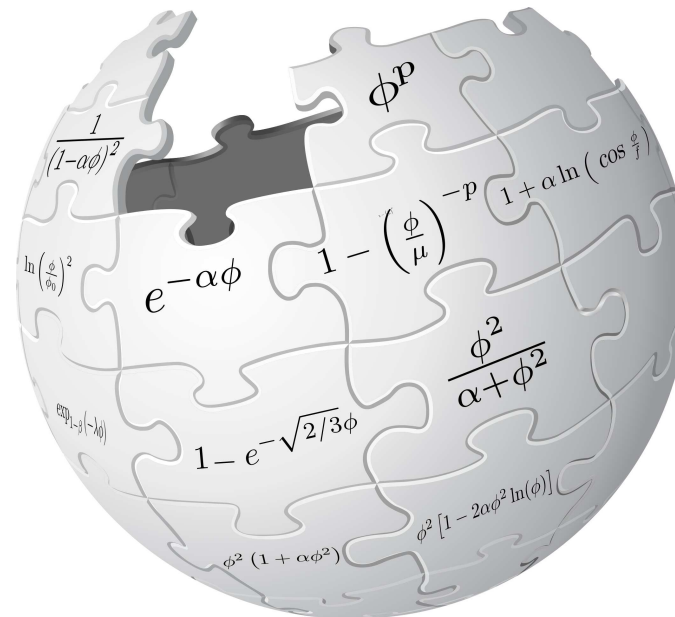
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- ❖ Primordial power spectra
- ❖ Reheating consistent model predictions

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Encyclopædia Inflationaris: Opiparous Edition

- New version published in May 2023 ([arXiv:1303.3787v4](https://arxiv.org/abs/1303.3787v4))
 - ◆ Deals with accurate slow-roll predictions for **287** models
 - ◆ Comes with a public runtime library **ASPIC**



- Computes the **Hubble-flow functions** from the model parameters θ_{inf}

$$(\theta_{\text{inf}}, R_{\text{rad}}) \longrightarrow \text{ASPIC} \longrightarrow \epsilon_i(\theta_{\text{inf}}, R_{\text{rad}}) \equiv \frac{d \ln |\epsilon_{i-1}|}{d \ln a}, \quad \epsilon_0 \propto \frac{1}{H}$$

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spectra

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Primordial power spectra

- Uniquely determined from the Hubble-flow functions

$$\mathcal{P}_\zeta = \frac{H_*^2}{8\pi^2 M_{\text{Pl}}^2 \epsilon_{1*}} \left\{ 1 - 2(1+C)\epsilon_{1*} - C\epsilon_{2*} + \left(\frac{\pi^2}{2} - 3 + 2C + 2C^2 \right) \epsilon_{1*}^2 \right.$$

$$+ \left(\frac{7\pi^2}{12} - 6 - C + C^2 \right) \epsilon_{1*}\epsilon_{2*} + \left(\frac{\pi^2}{8} - 1 + \frac{C^2}{2} \right) \epsilon_{2*}^2 + \left(\frac{\pi^2}{24} - \frac{C^2}{2} \right) \epsilon_{2*}\epsilon_{3*}$$

$$+ \left[-2\epsilon_{1*} - \epsilon_{2*} + (2+4C)\epsilon_{1*}^2 + (-1+2C)\epsilon_{1*}\epsilon_{2*} + C\epsilon_{2*}^2 - C\epsilon_{2*}\epsilon_{3*} \right] \ln \left(\frac{k}{k_*} \right)$$

$$+ \left[2\epsilon_{1*}^2 + \epsilon_{1*}\epsilon_{2*} + \frac{1}{2}\epsilon_{2*}^2 - \frac{1}{2}\epsilon_{2*}\epsilon_{3*} \right] \ln^2 \left(\frac{k}{k_*} \right) + \dots \left. \right\},$$

$$\mathcal{P}_h = \frac{2H_*^2}{\pi^2 M_{\text{Pl}}^2} \left\{ 1 - 2(1+C)\epsilon_{1*} + \left[-3 + \frac{\pi^2}{2} + 2C + 2C^2 \right] \epsilon_{1*}^2 + \left[-2 + \frac{\pi^2}{12} - 2C - C^2 \right] \epsilon_{1*}\epsilon_{2*} \right.$$

$$+ \left[-2\epsilon_{1*} + (2+4C)\epsilon_{1*}^2 + (-2-2C)\epsilon_{1*}\epsilon_{2*} \right] \ln \left(\frac{k}{k_*} \right) + (2\epsilon_{1*}^2 - \epsilon_{1*}\epsilon_{2*}) \ln^2 \left(\frac{k}{k_*} \right) + \dots \left. \right\}$$

- Currently known at third order: [arXiv:2205.12608](https://arxiv.org/abs/2205.12608) (involves ϵ_{4*})

Reheating consistent model predictions

- Quick check for which reheating history a model is compatible with the data

Qualitative Inflation

Quantitative inflation

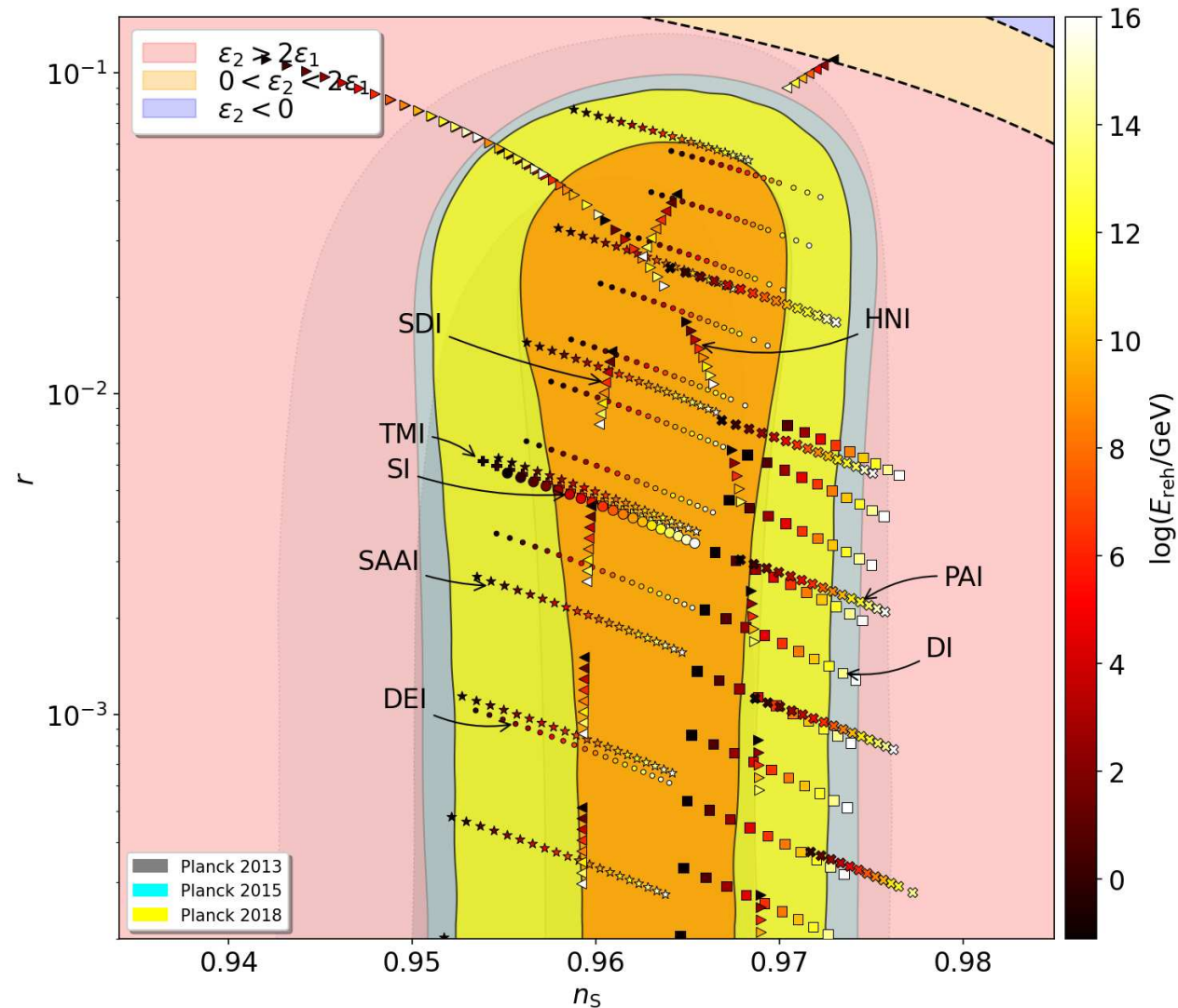
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Bayesian inference

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Bayesian inference

Machine-learning an effective likelihood

- To speed-up data analysis for 287 models

$$\mathcal{L}_{\text{eff}}(\mathbf{D}|P_*, \varepsilon_1, \varepsilon_2, \varepsilon_3) \propto \int P(\mathbf{D}|\boldsymbol{\theta}_s, P_*, \varepsilon_1, \varepsilon_2, \varepsilon_3, \varepsilon_4) \pi(\varepsilon_4) \pi(\boldsymbol{\theta}_s) d\varepsilon_4 d\boldsymbol{\theta}_s.$$

- The full likelihood has 55 parameters and is built upon
 - ◆ Planck 2020 post-legacy TT , TE , EE data (PR4/NPIPE maps)
 - ◆ Large scale EE polarization (lowE)
 - ◆ BICEP/Keck B -mode 2018 ([arXiv:2110.00483](https://arxiv.org/abs/2110.00483))
 - ◆ Small scale TE and EE from SPT-3G ([arXiv:2103.13618](https://arxiv.org/abs/2103.13618))
 - ◆ Baryon Acoustic Oscillations (SDSS collaboration)
- MCMC exploration of the 55 dimensions
 - ◆ COSMOMC up to 25 million samples ($R - 1 < 10^{-3}$)
 - ◆ GetDist marginalization in 4D: $P_*, \varepsilon_1, \varepsilon_2, \varepsilon_3$
- Basic machine learning: 1 hidden layer with 300 nodes

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Marginalized posteriors

● Cosmo and Hubble-flow parameters

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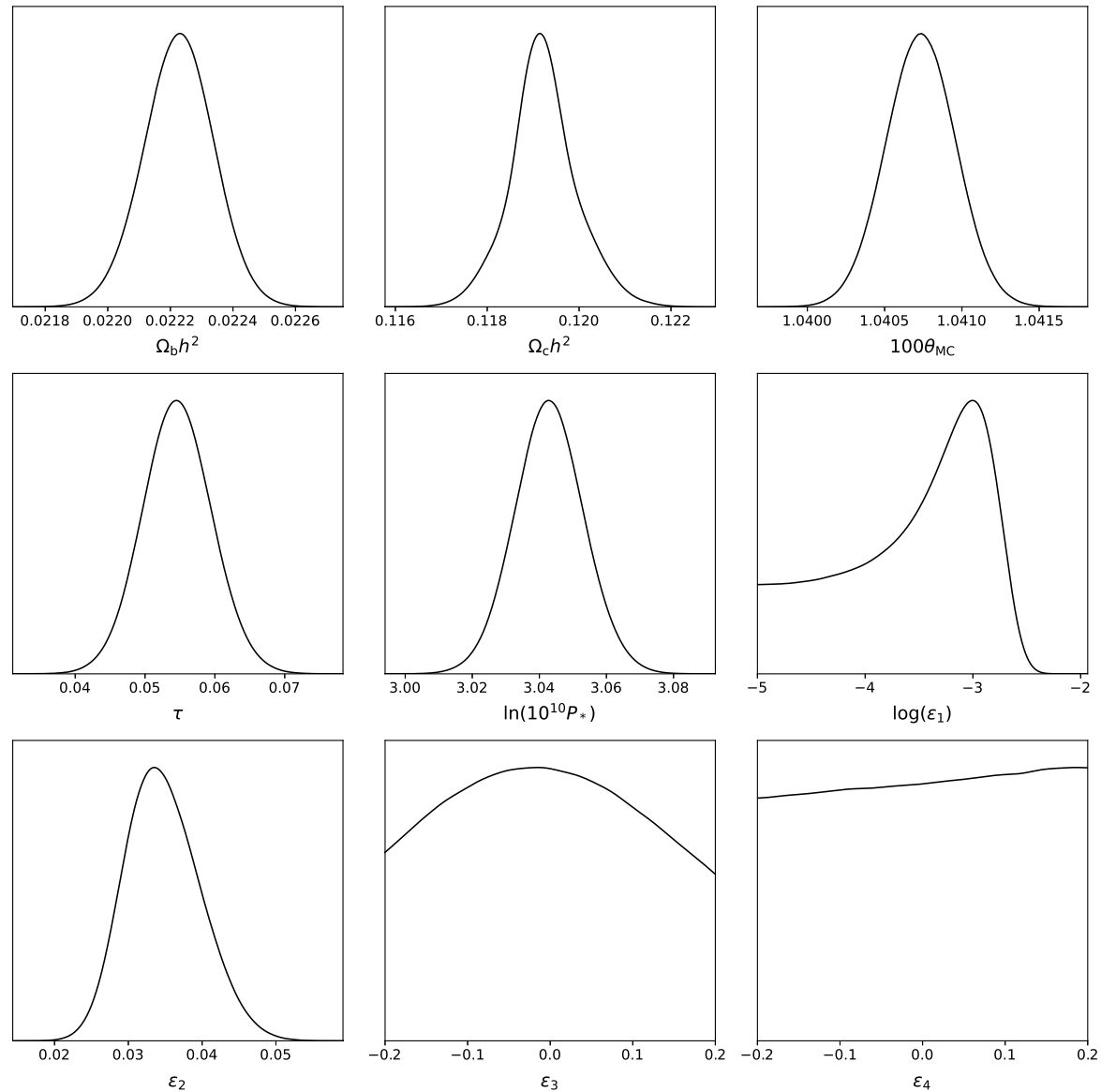
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Marginalized posteriors

● Effective likelihood vs exact: 2D

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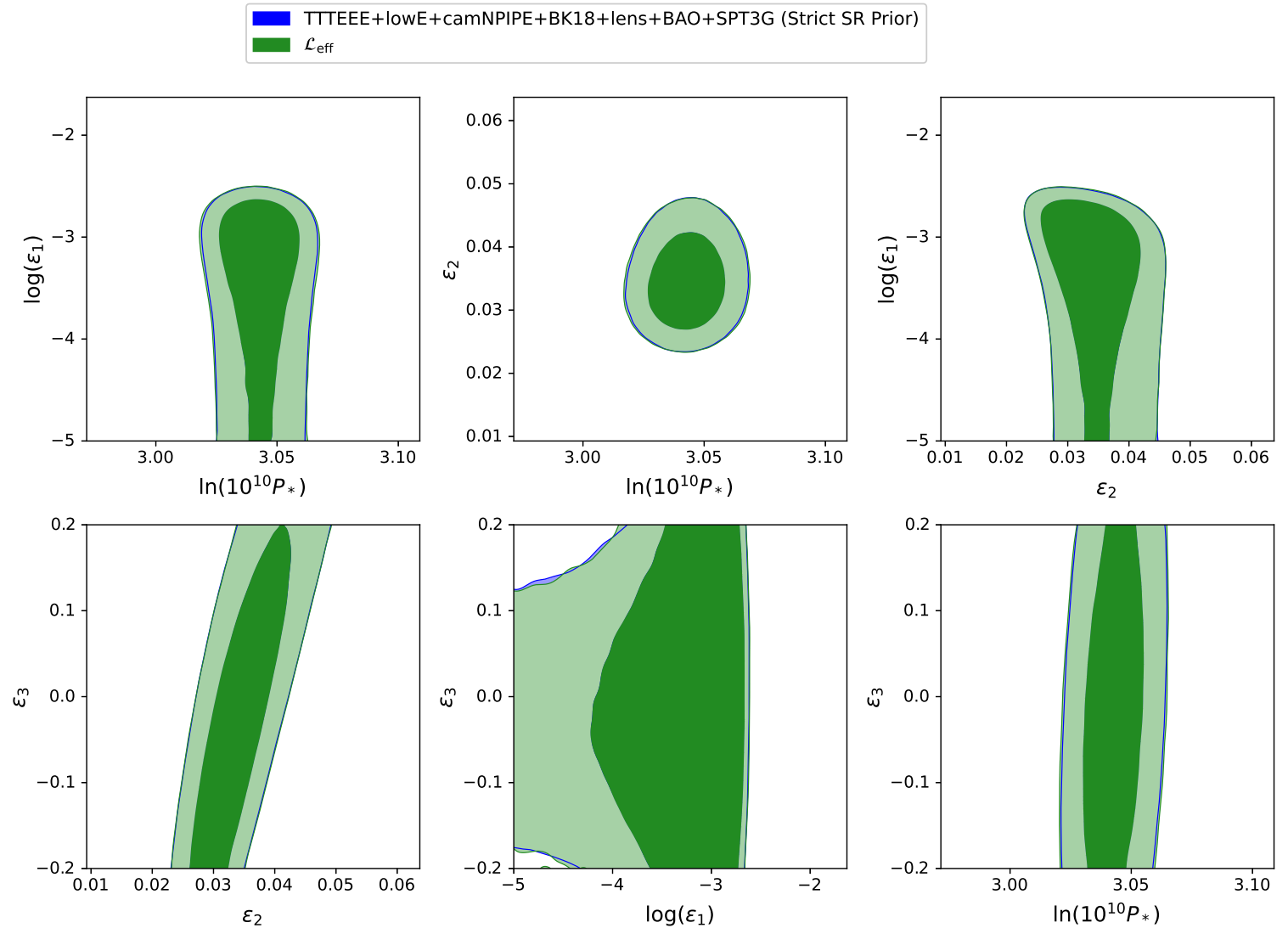
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Marginalized posteriors

- Effective likelihood vs exact: 1D

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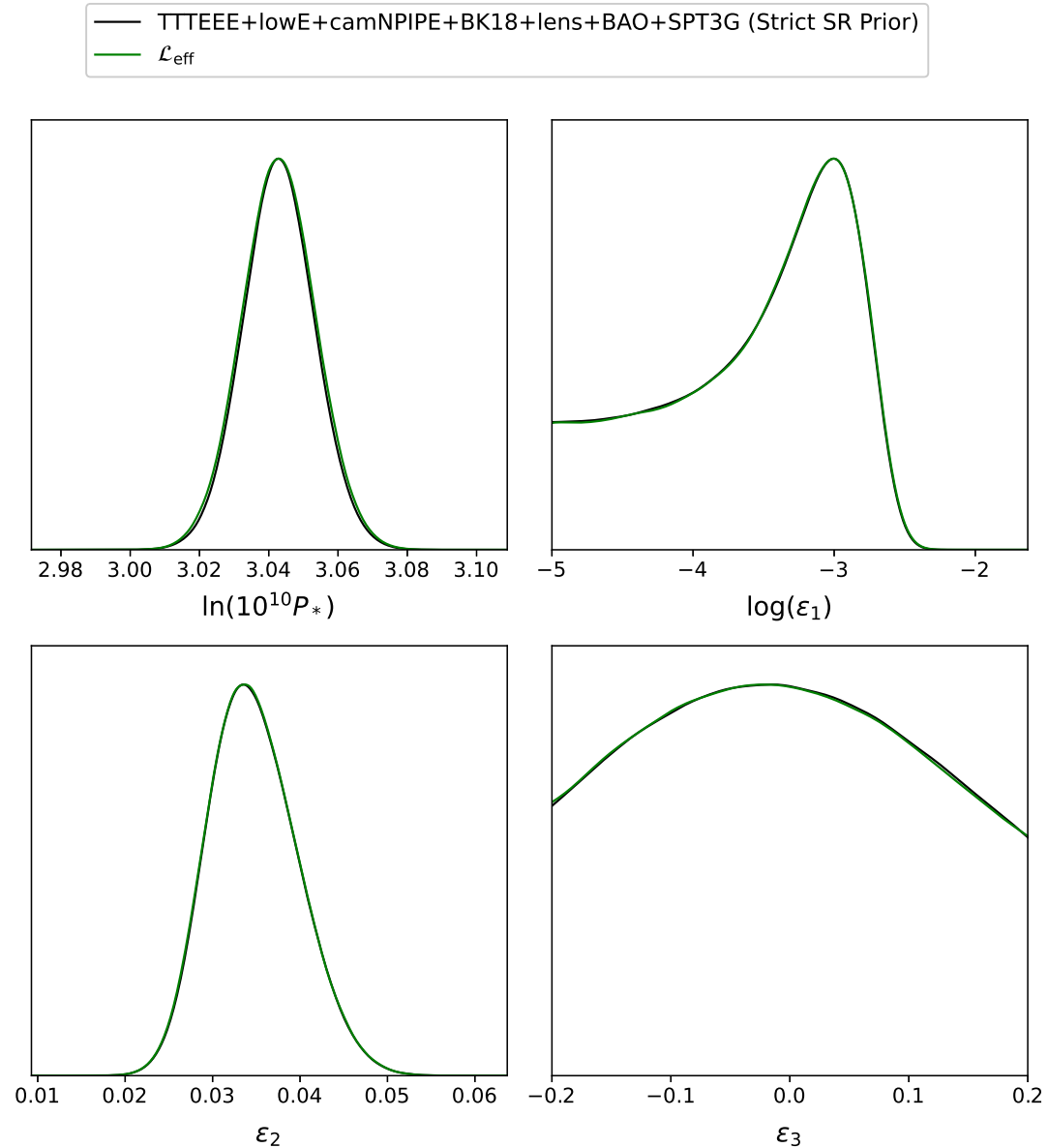
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Computing bayesian evidences

- Probability of a model \mathcal{M} to explain the data D

$$P(\mathcal{M}|D) = \frac{\mathcal{E}(D|\mathcal{M}) P(\mathcal{M})}{P(D)}$$

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- Bayesian evidence

$$\mathcal{E}(D|\mathcal{M}) \propto \int \mathcal{L}_{\text{eff}}(D|P_*, \varepsilon_1, \varepsilon_2, \varepsilon_3) \pi(\theta_{\text{inf}}, R_{\text{rad}}) d\theta_{\text{inf}} dR_{\text{rad}}$$

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- Computed with BAYASPIC running over 287 models
 - ◆ BAYASPIC \equiv ASPIC + PolyChord + \mathcal{L}_{eff}
 - ◆ A few cpu-hours per model \mathcal{M} (but up to 2 days for some)
- 1 TB of data output (nested chains, posteriors, plots,...)

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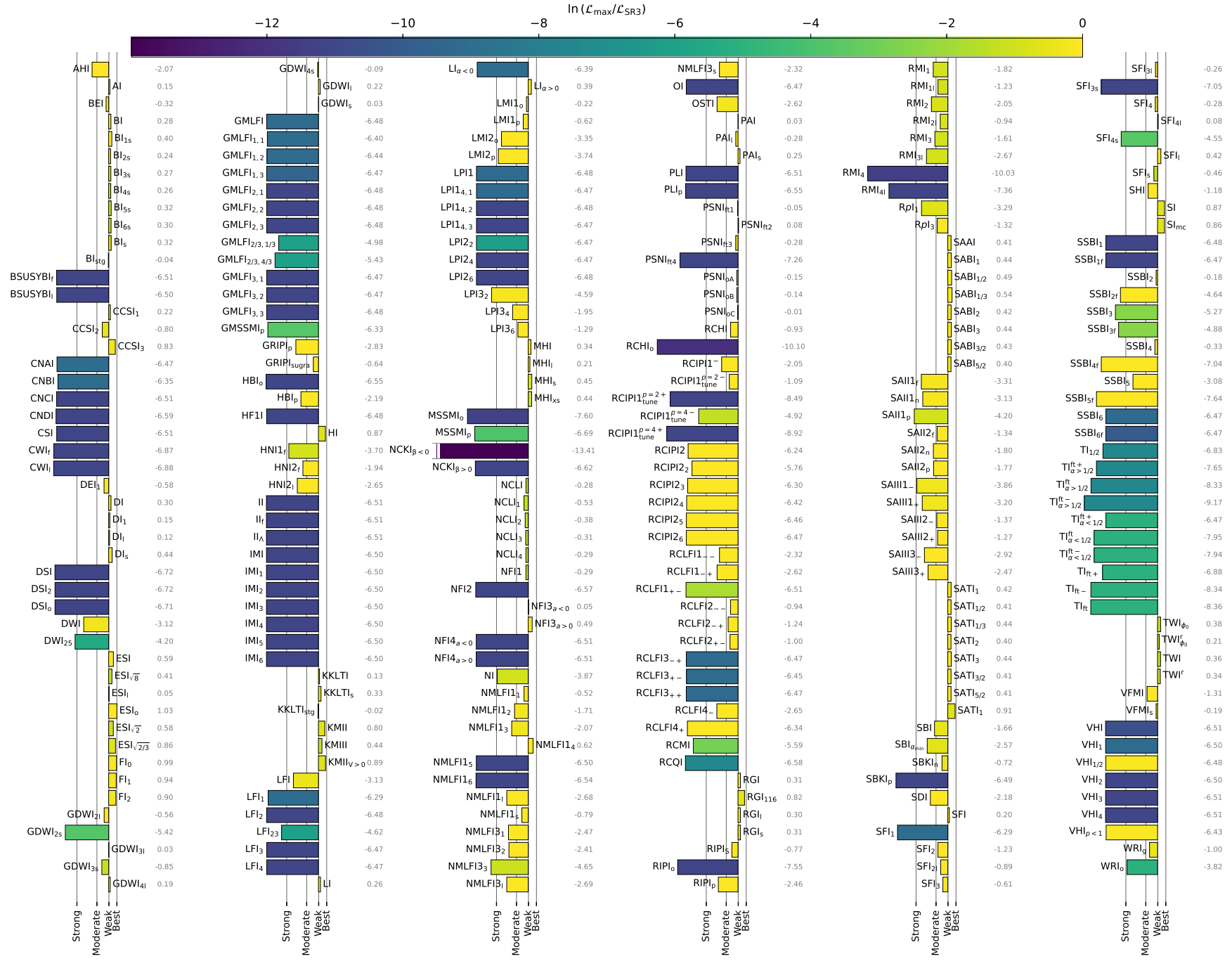
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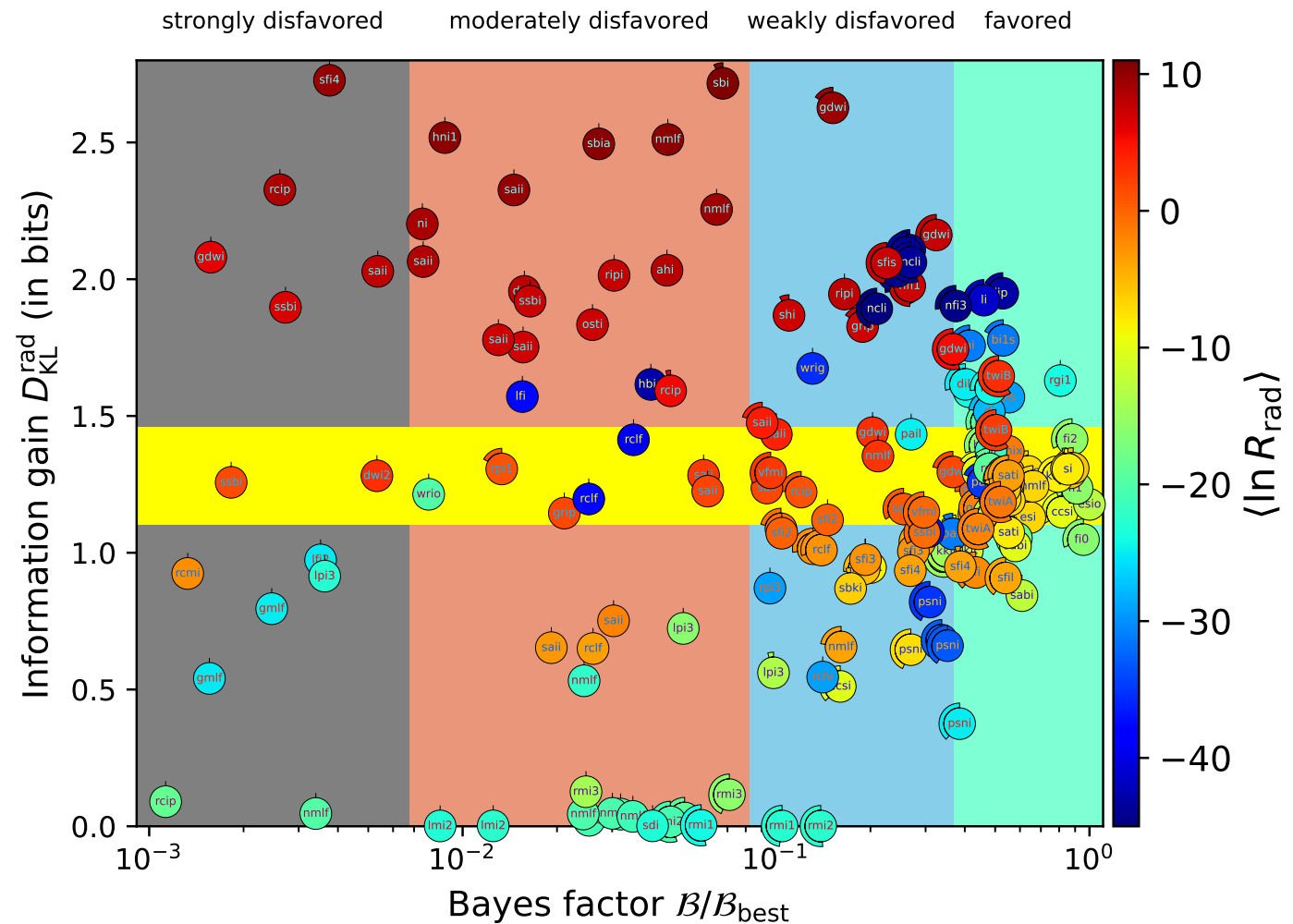
Bayes factors for all models



Information gain on the reheating

- Kullback-Leibler divergence between the prior and posterior

$$D_{\text{KL}}^{\text{rad}} = \int P(\ln R_{\text{rad}} | \mathbf{D}) \ln \left[\frac{P(\ln R_{\text{rad}} | \mathbf{D})}{\pi(\ln R_{\text{rad}})} \right] d \ln R_{\text{rad}},$$



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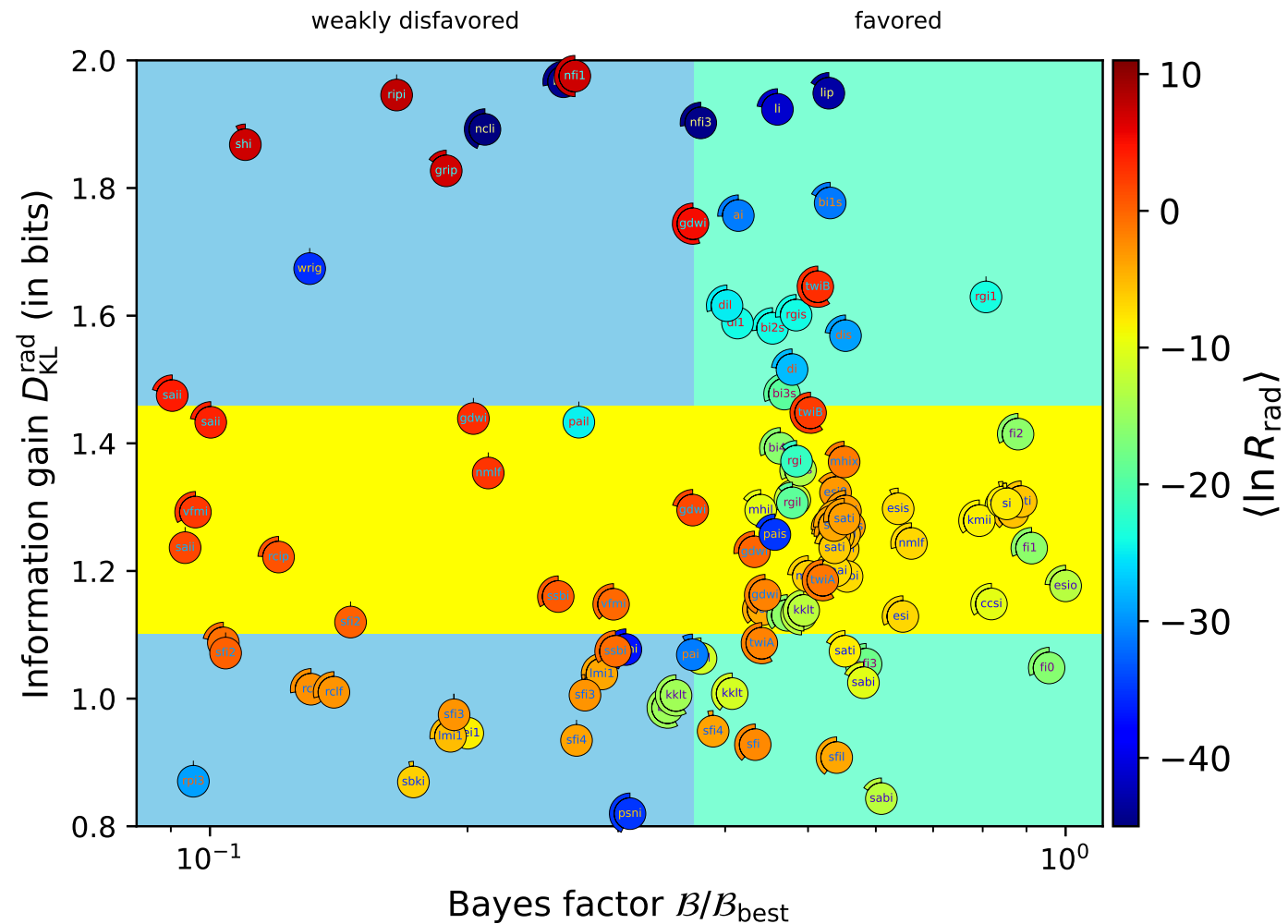
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Vanilla inflation predicts...

- Within the space of all slow-roll single-field models $\mathcal{I}_{\text{mod}} \equiv \{\mathcal{M}_i\}$
- Posterior probability of the running $\alpha_S \simeq -\epsilon_{2*} (2\epsilon_{1*} + \epsilon_{3*})$

$$P(\alpha_S | \mathbf{D}, \mathcal{I}_{\text{mod}}) = \sum_i P(\alpha_S | \mathbf{D}, \mathcal{M}_i) P(\mathcal{M}_i | \mathbf{D})$$

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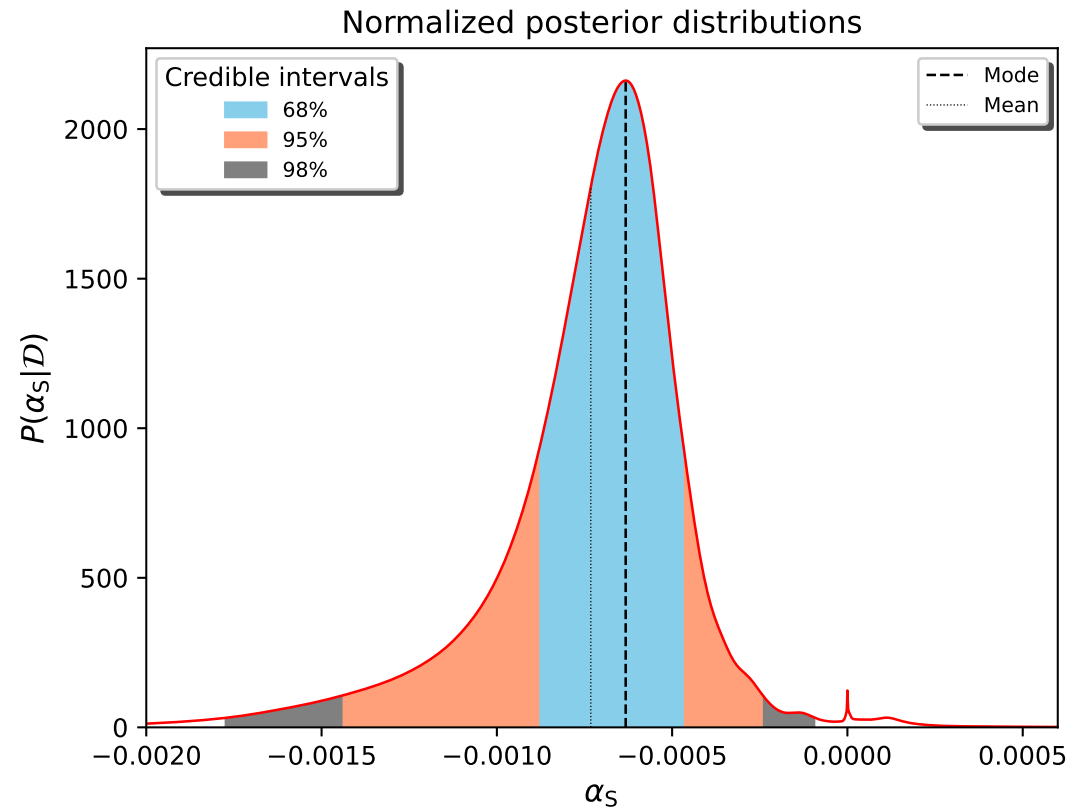
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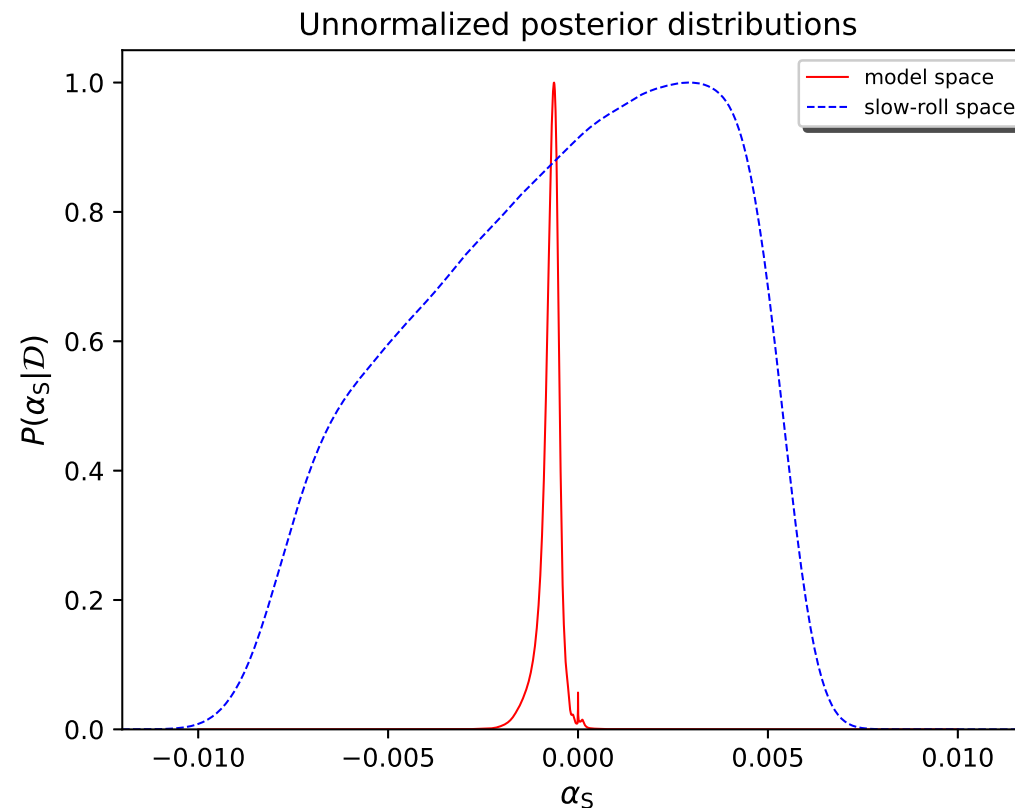
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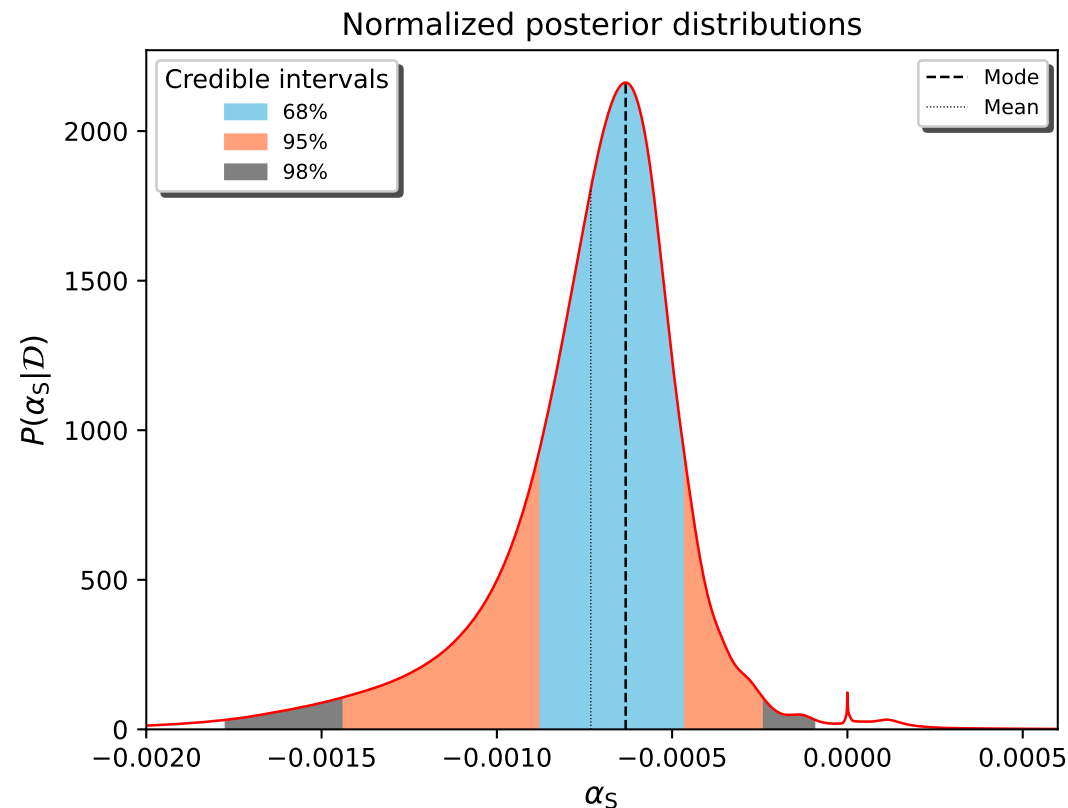
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- Credible interval: $-1.8 \times 10^{-3} < \alpha_S < -9.1 \times 10^{-5}$ (98%)

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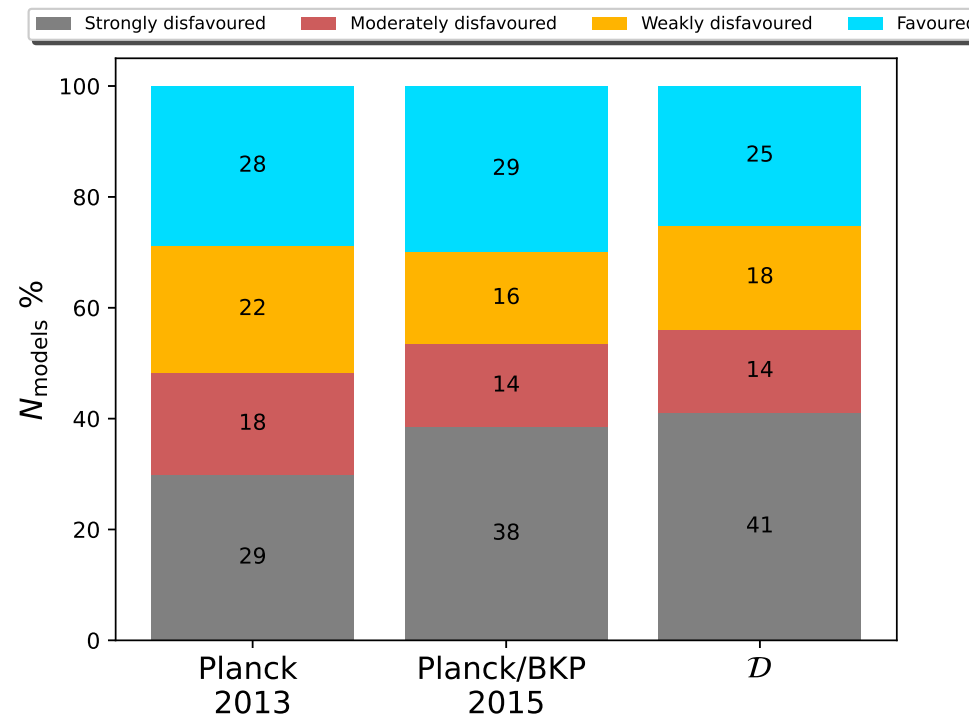
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Conclusion

- Bayesian data analysis in model space \mathcal{I}_{mod}
 - ◆ Enforces model consistency + new insights on the reheating era
 - ◆ Predicts: $\langle \alpha_s \rangle = -7.3 \times 10^{-4}$
- Data constraining power is winning against theoretical proposals



- Looking forward to the Euclid, LSS & CMB-S4 data!

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