

AXION PHOTON COUPLINGS IN TYPE IIB STRING THEORY

based on [\[2308.xxxxx\]](#) with Naomi Gendler, Doddy Marsh,
and Liam McAllister

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AXIONS AND STRING THEORY

Axions are a hallmark of weak coupling limits in string theory:

$$\tau = \frac{\theta}{2\pi} + i\frac{4\pi}{g^2} \quad \text{and} \quad \delta V \sim e^{2\pi i\tau} + c.c.$$

Perturbative shift symmetry $\theta \rightarrow \theta + \text{const}$
forbids large non-derivative couplings:

- ▶ naturally light
- ▶ evade fifth force constraints

And: can solve strong CP problem via PQ mechanism!

[Peccei, Quinn'77]

Unique portal to UV physics via dimension five coupling

$$g_{a\gamma\gamma} \varphi \frac{1}{2} F \wedge F$$

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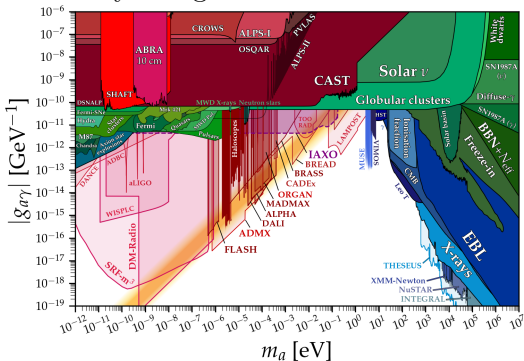
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$$g_{a\gamma\gamma} \varphi \frac{1}{2} F \wedge F$$

Axions are actively being searched for: [\[https://github.com/cajohare\]](https://github.com/cajohare)



In Calabi-Yau compactifications, usually get large number of axions, the famous **axiverse**.

[Arvanitaki, Dimopoulos, Dubovsky, Kaloper, March-Russell'09]

Important goal: Understand axiverse predictions!

[Demirtas, McAllister, Rios-Tascon'18][Demirtas, Gendler, Long, McAllister, JM'21]

[Cicoli, Hebecker, Jaeckel, Wittner'22][Cicoli, Guidetti, Righi, Westphal'21]...

+see Michele's & Sang Hui's talks!

KEY TAKE AWAYS

We develop machinery to compute axion-photon and axion-axion couplings in generic regime of hierarchical axion masses, at large number of axions.

Key lesson: The axiverse is hard to see. Most (non-)abelian gauge groups couple at most to a few axions.

Two effects conspire:

- ▶ Exponential mass hierarchies suppress axion couplings.

[cf. Arvanitaki, Dimopoulos, Dubovsky, Kaloper, March-Russell'09]

[cf. Agrawal, Nee, Reig'22]

- ▶ Sparse graph of divisor intersections in CY threefolds suppresses couplings (in type IIB axiverse of C_4 -axions).

[cf. Halverson, Long, Nelson, Salinas'19]

Also: we apply our tools to the type IIB axiverse of C_4 axions, to study astrophysical constraints and cosmic birefringence.

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PLAN

1. Hierarchies in axion interactions (in type IIB string theory)
2. Astrophysical bounds in an ensemble of semi-realistic models
3. Conclusions

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DIAGONAL TOYMODEL

Start with an oversimplified toymodel. Assume an EFT of

- ▶ N axions $\phi^a \simeq \phi^a + 2\pi$, with **diagonal kinetic terms**

$$\mathcal{L} \supset - \sum_{a=1}^N \frac{f_a^2}{2} (\partial\phi_a)^2 .$$

- ▶ N “basis” instantons generating a potential

$$V(\phi) = \sum_a \Lambda_a^4 (1 - \cos(\phi_a)) , \quad \Lambda_a \geq \Lambda_{a+1} > 0 .$$

- ▶ E&M coupling to a single **QED-axion** $\theta = \phi_{a^*}$:

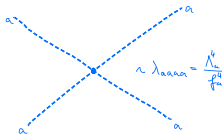
$$\mathcal{L} \supset \alpha_{EM} \cdot \frac{\theta}{2\pi} \cdot \frac{1}{4} F_{\mu\nu} \tilde{F}^{\mu\nu} .$$

DIAGONAL TOYMODEL (continued)

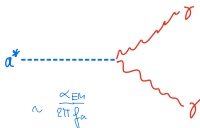
This theory is almost trivial. We get

- ▶ N mass eigenstates $\varphi^a = f_a \cdot \phi_a$ with masses $m_a^2 = \frac{\Lambda_a^4}{f_a^2}$

- ▶ only axion self interactions



- ▶ a single axion coupling to E&M



This model is of course too naïve!

STRUCTURELESS MODEL

Generically one would expect to find far more general EFTs in string theory, i.e.

- ▶ Arbitrary kinetic terms

$$\mathcal{L} \supset - \sum_a \frac{1}{2} K^{ab} \partial \phi_a \partial \phi_b .$$

- ▶ $N + k$ instantons with generic integer charges \vec{q}_I , generating a potential

$$V(\phi) = \sum_{I=1}^{N+k} \Lambda_I^4 \left(1 - \cos \left(2\pi \vec{q}_I \cdot \vec{\phi} \right) \right) , \quad \Lambda_I \geq \Lambda_{I+1} > 0 .$$

- ▶ E&M coupling to a **general linear combination**

$\theta_{EM} = 2\pi \vec{q}_{EM} \cdot \vec{\phi}$ with generic integer **charge vector** \vec{q}_{EM} :

$$\mathcal{L} \supset \alpha_{EM} \cdot \frac{\theta_{EM}}{2\pi} \cdot \frac{1}{4} F_{\mu\nu} \tilde{F}^{\mu\nu} .$$

STRUCTURELESS MODEL (continued)

Generally, understanding the large N axiverse would seem to be tedious. One has to

- ▶ go to canonical normalization, $K^{ab} = e^{ca}e^{cb}$,

$$\hat{\varphi}^a = e^{ab} \phi_b.$$

- ▶ Numerically find an axion vacuum $\langle \hat{\varphi}^a \rangle$,
- ▶ Numerically identify the eigenmodes $\varphi^a = \mathcal{O}^a_b \hat{\varphi}^b$ of the Hessian $\mathcal{H}_{ab} = \langle \partial_{\hat{\varphi}^a} \partial_{\hat{\varphi}^b} V \rangle$.
- ▶ Compute axion photon couplings by expanding

$$\frac{\theta_{EM}}{2\pi} = \vec{q}_{EM} \cdot \vec{\phi} \equiv \vec{\tilde{q}}_{EM} \cdot \vec{\hat{\phi}} \equiv \vec{\bar{q}}_{EM} \cdot \vec{\varphi}, \quad g_{a\gamma\gamma} = \alpha_{EM} \mathbf{q}_{EM}^a$$

No particular structure is evident in the $g_{a\gamma\gamma}$ or axion-axion interactions...

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SIMPLICITY AT LARGE N

Somewhat surprisingly, it turns out that type IIB orientifold models are almost like the naïve diagonal model!

The first “emergent” simplification occurs for hierarchical mass scales $\Lambda_a \gg \Lambda_{a+1}$ (**Numerically unstable!**).

One proceeds as follows:

- ▶ Canonically normalize kinetic terms.
- ▶ Select the leading N instantons \vec{q}_a , by deleting those with charges \vec{q}_I contained in the linear span of the $\vec{q}_{1,\dots,I-1}$.
- ▶ Apply the Gram-Schmidt process to the \vec{q}_a to find an axion basis, s.t. **instanton charge matrix is lower triangular:**

$$q_{ab} = \begin{pmatrix} q_{11} & 0 & 0 & \dots & 0 \\ q_{21} & q_{22} & 0 & \dots & 0 \\ \vdots & & \ddots & & \\ q_{N1} & & \dots & & q_{NN} \end{pmatrix}$$

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SIMPLICITY AT LARGE N (continued)

Now the instanton term with smallest action depends only on the first axion φ^1 :

$$\mathcal{L} = -\frac{1}{2}\partial\varphi^a\partial\varphi^a - \Lambda_1^4 (1 - \cos(2\pi\mathbf{q}_{11}\varphi^1)) + \dots$$

In limit $\frac{\Lambda_2}{\Lambda_1} \rightarrow 0$ the axion φ^1 is a mass eigenstate.

Integrating out φ^1 , get EFT of $N - 1$ axions with upper triangular instanton charge matrix, so by induction:

“Gram-Schmidt” basis = Mass eigenbasis + $\mathcal{O}\left(\frac{\Lambda_a^4}{\Lambda_{a+1}^4}\right)$.

SOME CONSEQUENCES

- ▶ Axion masses: $m_a^2 \approx \frac{\Lambda_a^4}{f_a^2}$ with $f_a := (2\pi\mathbf{q}_{aa})^{-1}$.

$$V(\varphi) = \sum_{a=1}^N \Lambda_a^4 \left(1 - \cos \left(\sum_b \theta_{ab} \frac{\varphi^b}{f_b} \right) \right)$$

with hierarchical **mixing angles** θ_{ab} :

- ▶ Mass suppressed mixing angles

$$\theta_{ab} \approx \frac{\mathbf{q}_{ab}}{\mathbf{q}_{bb}} \quad \text{for } b \leq a$$

$$\theta_{ab} \approx -\frac{\Lambda_b^4}{\Lambda_a^4} \frac{\mathbf{q}_{ba}}{\mathbf{q}_{aa}} \ll 1 \quad \text{for } b > a$$

Light axions do not mix into “heavy” theta angles.

SOME CONSEQUENCES (continued)

- ▶ If \vec{q}_{EM} also contributes an instanton term to potential,

$$V(\varphi) \supset \Lambda_{EM}^4 (1 - \cos(2\pi\vec{q}_{EM} \cdot \vec{\varphi})) ,$$

axions lifted by instantons with $\Lambda_a \ll \Lambda_{EM}$ do not couple to EM:

$$g_{a\gamma\gamma} f_a = \mathcal{O} \left(\frac{\Lambda_a^4}{\Lambda_{EM}^4} \right)$$

- ▶ Mixed quartics suppressed by lightest mass scale:

$$\lambda_{a \leq b \leq c < d} \sim \frac{\Lambda_d^4}{f_a f_b f_c f_d} \theta_{da} \theta_{db} \theta_{dc}$$

Axions preferentially decay to their mass neighbors.

KINETIC DECOUPLING IN TYPE IIB AXIVERSE

So far: discussed suppression of couplings in ratios of mass scales Λ_a .

\Rightarrow pure QFT effect

Next: Planck-suppression of axion interactions from kinetic decoupling at large $N \equiv h_+^{1,1}$

We will study Calabi-Yau hypersurfaces X in toric fourfolds V at large $h^{1,1}$, i.e. the KS-database.

Axions given by $h_+^{1,1}$ zero modes of C_4 , lifted by D3-brane instantons wrapped on four-cycles.

ASSUMPTIONS AND APPROXIMATIONS

- ▶ All BPS D3-brane instantons contribute to W .
- ▶ SUSY is broken at low enough scales s.t. BPS instantons dominate over non-BPS instantons, i.e.

$$m_{\frac{3}{2}} \ll M_P$$

- ▶ The Kähler and CS moduli of X are stabilized at high scales compared to the axion masses $m_a = \Lambda_a^2/|f_a|$.
- ▶ No two-form axions, i.e. $h_-^{1,1} = 0$.
- ▶ Standard model realized on (intersecting) 7-branes.
- ▶ Approximate the cone of effective divisors by the one inherited from toric ambient variety.

AXION FIELD SPACE METRIC

The smallest irreducible divisors in a CY X are the
prime toric divisors:

$$D_i := \{x_i = 0\} \cap X, \quad i = 1, \dots, h^{1,1}(X) + 4$$

Letting D_a be basis of $H^2(X)$ furnished by subset of smallest linearly independent $h^{1,1}$ prime toric divisors:

$$q_{ab} = \delta_{ab}$$

Their overlap, measured by inverse Kähler metric is

$$K_{ab} = \langle D_a, D_b \rangle = 2\mathcal{V} \left(c_{ab} + \frac{\text{Vol}(D_a)\text{Vol}(D_b)}{\mathcal{V}} \right)$$

with $c_{ab} = 0$ if $D_a \cap D_b = \emptyset$.

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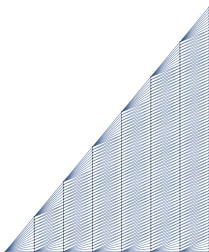
Off-diagonal terms K_{ab} from non-intersecting divisors are Planck-suppressed, i.e. $\propto 1/\mathcal{V}$.

AXION FIELD SPACE METRIC (continued)

Crucial: At large $h^{1,1}$ divisor intersections are sparse!

[cf. Halverson, Long, Nelson, Salinas'19]

Reason: Toric ambient variety arises from triangulating reflexive polytope Δ° . A pair of divisors (D_a, D_b) intersects iff points (p_a, p_b) in Δ° are in same two-face of Δ° and connected by triangulation:



(largest two-face of Δ° with $h^{1,1} = 491$, Delaunay triangulation)

LOCALIZATION OF AXION WAVEFUNCTION

In absence of off-diagonal terms we would be back to our naïve toy model with trivial mixing angles $\theta_{ab} = 0$ for $a \neq b$. Every mass eigenstate overlaps precisely with one basis divisor.

This gets corrected by

- ▶ Cross terms c_{ab} for intersecting divisors
⇒ **Local in intersection graph.**
- ▶ Planck suppressed off-diagonals $\frac{\text{Vol}(D_a)\text{Vol}(D_b)}{\mathcal{V}}$
⇒ **Non-local in intersection graph,**
negligible for small pairs of divisors.

Mixing angles θ_{ab} are

- ▶ localized around $a = b$ in intersection graph,
 - ▶ and mass suppressed for $a < b$.

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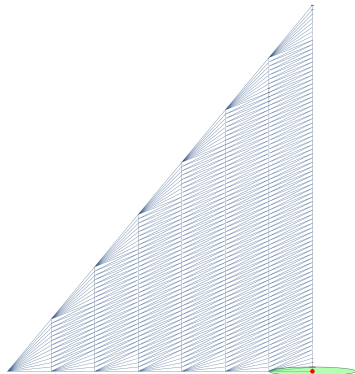
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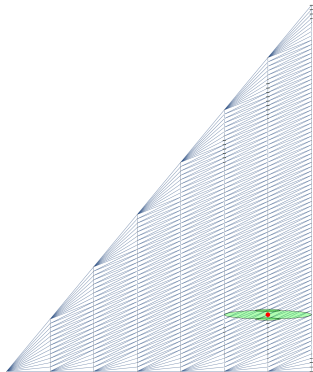
LOCALIZATION OF AXION WAVEFUNCTION

Example at $h^{1,1} = 491$,
at “tip of the stretched Kähler cone”: $\mathcal{V} \approx 1.7 \times 10^{12}$
QED on smallest divisor: $\text{Vol}(D) = 0.5$



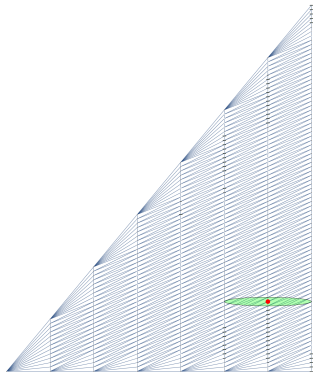
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QED on 20th divisor: $\text{Vol}(D) = 11$



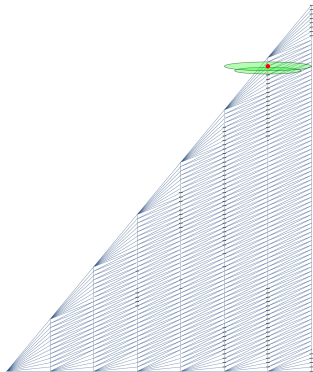
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QED on 50th divisor: $\text{Vol}(D) = 23$



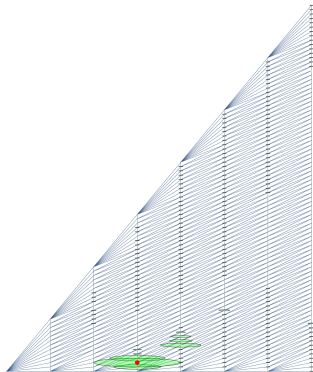
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Example at $h^{1,1} = 491$,
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QED on 100th divisor: $\text{Vol}(D) = 63$



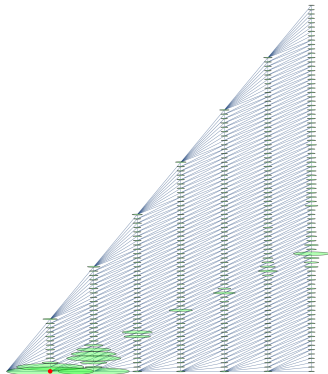
LOCALIZATION OF AXION WAVEFUNCTION

Example at $h^{1,1} = 491$,
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QED on 200th divisor: $\text{Vol}(D) = 655$



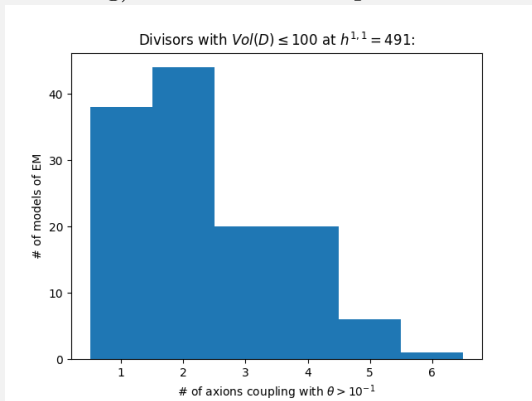
LOCALIZATION OF AXION WAVEFUNCTION

Example at $h^{1,1} = 491$,
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QED on 343th divisor: $\text{Vol}(D) = 6.8 \times 10^6$



IS AN AXIVERSE INVISIBLE ?

Even at very large $h^{1,1}$ most standard models realized on (intersecting) seven-branes couple to few axions:



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AXION HELIOSCOPIES

How relevant are the suppression effects for real world experiments?

One of the most stringent bounds on axion photon couplings:
axion helioscopes



Effective coupling probed:

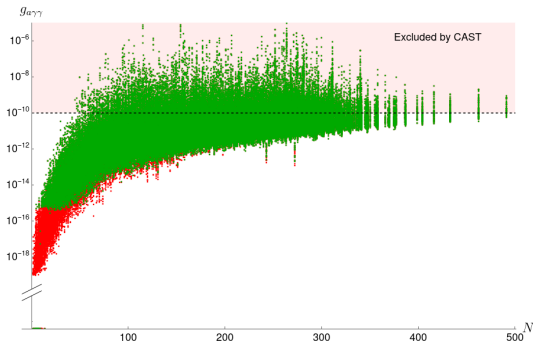
$$g_{\text{eff}} = \sqrt{\sum_{m_a < 1\text{eV}} (g_{a\gamma\gamma})^2} = \sqrt{\sum_{m_a < 1\text{eV}} \left(\frac{\alpha_{EM}}{2\pi f_a} \theta_{EM,a} \right)^2}$$

Bound from CAST: $g_{\text{eff}} \lesssim 10^{-10} (\text{GeV})^{-1}$.

Without mixing suppression, i.e. $\theta_{EM,a} = \mathcal{O}(1)$,
get dominated by smallest f_a .

TYPE IIB AXIVERSE EXCLUSIONS (naïve)

Under this assumption, significant part of the type IIB axiverse would seem to be excluded: [Demirtas, Gendler, Long, McAllister, JM'21]



Key fact: f_a vary over a few orders of magnitude

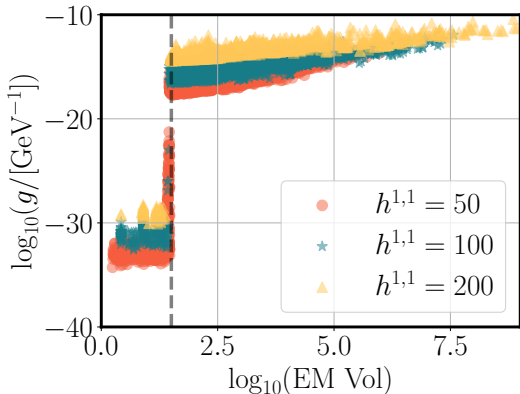
[Demirtas, McAllister, Rios-Tascon'20]

But: With significant mixing suppression, expect

$$g_{eff} \sim \frac{\alpha_{EM}}{2\pi \langle f_a \rangle} \ll \frac{\alpha_{EM}}{2\pi \min_a(f_a)}$$

MASSES AND COUPLINGS

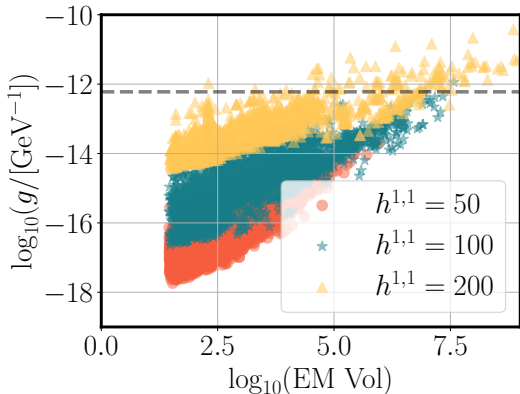
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- ▶ effective coupling for Chandra (similar for CAST) drops off as volume of EM divisor drops below ~ 32 : axions heavier than QED axion outside of Chandra mass range.

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COSMIC BIREFRINGENCE

Currently there is $\sim 3.6\sigma$ evidence for
cosmic birefringence in CMB := non-zero rotation of
polarization plane by angle β between CMB and today.

$$\beta = (6.0 \pm 1.6) \times 10^{-3}$$

[Minami, Komatsu'20][Eskilt, Komatsu'22] +Donghui's talk!

Possible explanation:

[Carroll, Field, Jackiw'90]

axion φ with $10^{-33} eV \sim H_{\text{now}} < m < H_{\text{CMB}} \sim 10^{-28} eV$

relaxes to minimum from generic initial field displacement φ_0

$$\beta \approx g_{a\gamma\gamma}\varphi_0 = \frac{\alpha_{\text{EM}}}{2\pi}\theta_{\text{EM},a}\frac{\varphi_0}{f_a} \approx 3 \times 10^{-3} \times \theta_{\text{EM},a} \quad \text{for } \varphi_0 \sim \pi \cdot f_a$$

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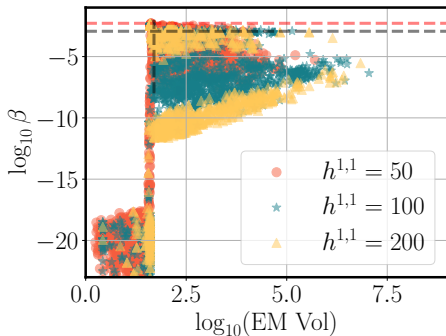
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COSMIC BIREFRINGENCE (continued)

To account for this, need axion(s) with $\mathcal{O}(1)$ mixing angle to standard model.

We find this to be possible only if QED-axion itself falls in correct mass window, $\text{Vol}(D_{QED}) \gtrsim 40$:

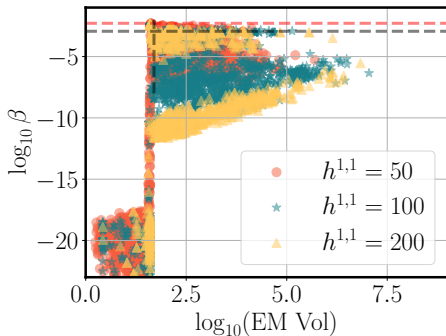


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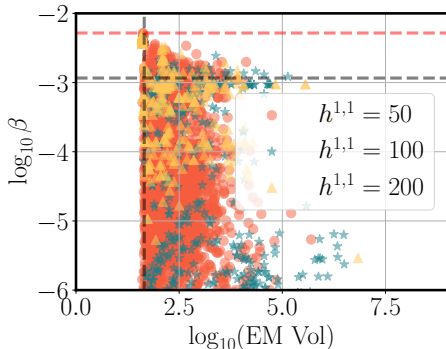


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COSMIC BIREFRINGENCE (continued)

To account for this, need axion(s) with $\mathcal{O}(1)$ mixing angle to standard model.

This appears to be possible only if QED-axion itself falls in correct mass window, $\text{Vol}(D_{QED}) \gtrsim 40$:



Incidentally, this coincides with GUT scale $U(1)_Y$ coupling in standard model $\frac{4\pi}{g^2(10^{16}\text{GeV})} \sim 40\dots$

CONCLUSIONS

We have developed tools to systematically compute axion photon couplings in large axiverses with large separation of scales.

We have learned that a true axiverse is almost invisible: QED realized on a small divisor couples to $\mathcal{O}(1)$ axions even in models with many axions. This significantly reduces constraints from astrophysical processes.

WIP not mentioned: dark matter abundances from misalignment, freeze-in/out, constraints from decaying dark matter, moduli problem, dark radiation...

```
3 : M = m/m^3
3 = subquotient (|- x y |, | x^3 x2y xy2 y^3 |)
3 : R-module, subquotient of
4 : C = resolution M
4 : R <- R <- R <- R <- R
4 : 0 1 2 3
4 : ChainComplex
5 : T = (**(R^1/m)
5 = cokernel (1) | x y 0 0 | <- cokernel (2) | x y 0 0 0 0 0 0 0 0 | <- cokernel (4) | x y 0 0 0 0 |
```

Macaulay2, Computational Algebraic Geometry and String Theory



To be held at the University of Utah

May 28th - June 1st, 2024 (School)
June 3rd - June 5th, 2024 (Conference)



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May 28-June 5, 2024

Organizers: D. Eisenbud, D. Grayson, A. Leykin, A. Lukas, L. McAllister, K. Schwede, M. Stillman

감사합니다