# **Bounds on the Species Scale**

#### Max Wiesner Harvard University

Center for the Fundamental Laws of Nature

& Center of Mathematical Sciences and Applications



HARVARD UNIVERSITY CENTER OF MATHEMATICAL SCIENCES AND APPLICATIONS based on:

Damian van de Heisteeg, Cumrun Vafa, MW, David H. Wu 2212.06841, 2305.07701

Damian van de Heisteeg, Cumrun Vafa, MW 2303.13580



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Common feature of many theories of Quantum Gravity:

- 1. light scalar fields  $\phi^i$  determining couplings, masses, etc.
- 2. towers of massive states beyond EFT.

Distance Conjecture [Ooguri, Vafa'06] Along paths in scalar field space traversing distances  $d \gg l_p$ an infinite tower of states becomes light in Planck units as  $\frac{M(Q)}{M_{pl}} \sim e^{-Ad(P,Q)}$ 



Tested in string theory in great detail: Cf. reviews [Palti '19, van Beast, Calderon-Infante, Mirfendereski, Valenzuela '21; Agmon, Bedroya, Kang, Vafa '22]

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Problem: Would need to know the exact spectrum of the theory at any point in field space.

Way out: Instead of exact spectrum of states study the variation of the Quantum Gravity cut-off  $\Lambda$  over field space  $\mathcal{M}_{\phi} \rightarrow$  Species Scale!

[Dvali '07]

What is the species scale?

(See talks by David Andriot, Alvaro Herraez, Alberto Castellano, Niccolo Cribiori, Dieter Lüst, Marco Scalisi, Irene Valenzuela, ...)

- In the presence of *N* effective light degrees of freedom: entropy of any black hole needs to satisfy: *S* > *N*
- Compare to Schwarzschild black hole  $S_{\text{Schwarzschild}} = \left(r_H M_{\text{pl}}\right)^{d-2}$

 $\rightarrow$  radius of minimal black hole describable within EFT

$$r_{\rm H,min}M_{\rm pl} \sim N^{1/(d-2)}$$

• Interpretation: Higher derivative corrections break down for curvatures  $R \sim \mathcal{O}(r_{\rm H,min}^{-2})$ 

$$S = \int d^d x \sqrt{-g} \left[ \frac{M_{\rm pl}^{d-2}}{2} \left( R + \sum_{n} \underbrace{\underbrace{\mathcal{O}_n(R)}_{\Lambda_s^{n-2}} + \dots}_{n} \right) \right].$$

$$\int_{\Lambda_s} \frac{M_{\rm pl}}{N^{\frac{1}{d-2}}}$$

• Distance Conjecture: *N* varies over field space  $\Rightarrow \Lambda_s = \Lambda_s(\phi)$ 

### **Moduli Dependent Species Scale**

**Questions:** 1. From a pure EFT perspective: Can we give a bound on the variation of  $\Lambda_s(\phi)$ ?

**2.** Can we give a closed expression for  $\Lambda_s(\phi)$ ?

**3.** Can we use explicit form for  $\Lambda_s(\phi)$  to bound potentials?

### Moduli Dependent Species Scale

- **1.** From a pure EFT perspective: Can we give a bound on the variation of  $\Lambda_s(\phi)$ ? **Questions:** 
  - **2.** Can we give a closed expression for  $\Lambda_s(\phi)$ ?
  - **3.** Can we use explicit form for  $\Lambda_s(\phi)$  to bound potentials?
- **Answers:** 1. Consistency of the perturbative expansion of a theory of Einstein gravity + higher derivative corrections requires:  $\frac{|\nabla \Lambda_s|}{\Lambda_s} \lesssim \mathcal{O}(1)$ 
  - 2. In theories with Type II compactifications the moduli dependence of  $\Lambda_s$  is captured by the topological genus-one free energy  $F_1$ . Can be fixed in explicit examples

3. Slowly varying positive potentials bounded by  $V(\phi) < A \exp\left(\frac{2}{\sqrt{(d-1)(d-2)}}\Delta\phi\right)$ 

 $\rightarrow$  can bound the maximal field range — including interior parts of field space) cf. [Scalisi, Valenzuela '19] (See Marco's talk on Monday)

First: Bound on the slope of the species scale as a function of moduli.

• Consider general Einstein theory of gravity + scalar field:

$$S = \int d^d x \sqrt{-g} \left[ \frac{M_{\rm pl}^{d-2}}{2} R - \frac{1}{2} (\partial \phi)^2 + \dots \right] \,.$$

• Step 1: Integrate out heavy states  $\rightarrow$  generate higher-derivative terms suppressed by  $\Lambda_s$ .

$$S = \int d^d x \sqrt{-g} \left[ \frac{M_{\rm pl}^{d-2}}{2} \left( R + \sum_n \frac{\mathcal{O}_n(R)}{\Lambda_s^{n-2}(\phi)} + \dots \right) - \frac{1}{2} (\partial \phi)^2 + \dots \right].$$

- Step 2: Integrate out small-distance modes of  $\phi$  above  $\Lambda_s$ .
- Consistency requirement: Integrating out small-distance modes of a single field  $\phi$  should not considerably change the perturbative expansion!

A single field does not significantly change the entropy of the smallest black hole and therefore does not significantly affect the species scale!

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- Expand the species scale:  $\Lambda_{s}(\phi_{0} + \delta\phi) = \Lambda_{s}(\phi_{0}) + \delta\phi \Lambda_{s}'(\phi_{0}) + \mathcal{O}(\delta\phi^{2})$
- Get effective interaction through operator:

$$\frac{\mathcal{O}_n(R)}{\Lambda_s(\phi)^{n-2}} \to \frac{\mathcal{O}_n(R)}{\Lambda_s(\phi_0)^{n-2}} - \frac{(n-2)\Lambda'_s(\phi_0)}{\Lambda_s(\phi_0)^{n-1}}\,\delta\phi\,\mathcal{O}_n(R)$$

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$$\sim \int d^d x d^d y \ M_{\rm pl}^{2(d-2)} \mathcal{O}_n(x) \frac{\Lambda'_s(\phi_0)}{\Lambda_s(\phi_0)^{n-1}} G_\phi(x,y) \mathcal{O}_m(y) \frac{\Lambda'_s(\phi_0)}{\Lambda(\phi_0)_s^{m-1}}$$

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- Generate effective point-interaction:

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- Asymptotic regions in field space seem to saturate the bound.
- Since  $\Lambda_s \sim m_{\text{tower}}^{\gamma}$ : masses of towers predicted by distance conjecture should not decay *faster* than exponential!!

Similar results in [Calderon-Infante, Castellano, Herraez, Ibanez '23]

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• According to Emergent string conjecture can have: [Lee, Lerche, Weigand '19] Emergent string limits in d dimensions: Decompactification limits from  $d \rightarrow D$  dimensions:

$$\frac{|\Lambda'_s|}{\Lambda_s} \to \frac{1}{\sqrt{d-2}}$$

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cf. also [Etheredge, Heidenreich, Kaya, Qiu, Rudelius'22]

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Decompactification limits from  $d \rightarrow D$  dimensions:

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 $\rightarrow$  Universal bound?

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cf. also [Etheredge, Heidenreich, Kaya, Qiu, Rudelius'22]

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## **Explicit Expression for** $\Lambda_s$ ?

Question: Can we determine the constant c?

 $\rightarrow$  need to have an explicit expression for  $\Lambda_s$ 

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However ... in Type II string compactifications there exists an  $R^2$  correction that can be computed in perturbation theory and that does not get corrected!

## Moduli dependent species scale

Consider topological string with target space CY threefold:

- Effective action contains term:  $S \supset \int F_1(R^-)^2$ Topological genus-1 free energy
- Worldsheet theory described by CFT with  $\mathcal{N} = (2,2)$  supersymmetry in two dimensions.
- For such a CFT can define moduli-dependent index

[Cecotti, Vafa '92, Bershadsky, Cecotti, Ooguri, Vafa '93]

$$F_1 = \frac{1}{2} \int_{\mathscr{F}} \frac{d^2 \tau}{\tau_2} \operatorname{Tr} \left[ (-1)^F F_L F_R q^{H_0} \bar{q}^{\bar{H}_0} \right]$$

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- Satisfied holomorphic anomaly equation:

$$\partial_{\bar{j}}\partial_{i\bar{j}}F_{1} = \operatorname{Tr}(-1)^{F}C_{i\bar{C}_{\bar{j}}} - \frac{1}{12}G_{i\bar{j}}\operatorname{Tr}(-1)^{F}, \quad \text{[Cecotti, Vafa '92]}$$

Chiral ring structure constants

Zamolodchikov metric on field space

- Holomorphic anomaly equation can be integrated to

$$F_{1} = \frac{1}{2} \left( 3 + h^{1,1} - \frac{\chi}{12} \right) K - \frac{1}{2} \log \det G_{i\bar{j}} + \log |f|^{2},$$
  
Euler characteristic of mirror  $Y_{3}$  Kähler potential Holomorphic ambiguity

- Holomorphic ambiguity can be fixed by matching the behavior of  $F_1$  at boundary of moduli space.

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## $F_1$ and the number of species

**Question:** Why can  $F_1$  be used to estimate the number of light species?

#### <u>A first look:</u>

- $N_{\rm sp}$  measures the accumulation of light modes in the spectrum of the theory.
- Spectrum of Laplacian Δ gives a measure for number of light states.
   → growth of |log(det Δ)| would measure accumulation of light states.
- Problem: spectrum of  $\Delta$  difficult to compute!

## $F_1$ and the number of species

**Question:** Why can  $F_1$  be used to estimate the number of light species?

<u>A first look:</u>

- $N_{\rm sp}$  measures the accumulation of light modes in the spectrum of the theory.
- Spectrum of Laplacian  $\Delta$  gives a measure for number of light states.  $\rightarrow$  growth of  $|\log(\det \Delta)|$  would measure accumulation of light states.
- Problem: spectrum of  $\Delta$  difficult to compute!
- However: exists combination of Laplacian  $\Delta_{p,q}$  acting on (p,q)-forms that is computable since it is an index-like quantity.

$$\frac{1}{2}\sum_{p,q}(-1)^{p+q}\left(p-\frac{3}{2}\right)\left(q-\frac{3}{2}\right)\,\log(\det\Delta_{(p,q)})$$

[Bershadsky, Cecotti, Ooguri, Vafa '93]

- This combination can be identified with  $F_1$ !
- General expectation:

$$\frac{\Lambda_s}{M_{\rm pl}} \sim F_1^{-1/2} \quad \rightarrow reproduces \ right \ behavior \ in \ asymptotic \ regions$$

black hole perspective in [Cribiori, Lust, Staudt '22]

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How does  $F_1$  behave asymptotically?

<u>Three kinds of infinite distance limits:</u>

[Lee, Lerche, Weigand '19]

- 1. Large volume limit (decompactification to 5d M-theory)
- 2. Emergent String limit (requires K3-fibration for CY)
- 3. Decompactification to 6d (requires  $T^2$ -fibration for CY)



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Parametrize all limits by  $s \to \infty$ :  $F_1$  has asymptotic form  $F_1 = c_2 s - \beta \log s$ 

$$\implies \text{slope of } \Lambda_s \text{ given by } \quad \frac{|\nabla \Lambda_s|}{\Lambda_s} \to \alpha + \frac{\beta}{s} \log s$$



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Limit	α	$\beta$
Large volume	$\sqrt{\frac{1}{6}}$	$\frac{1}{4}(18 + h^{1,1} + h^{2,1})$
Emergent String	$\sqrt{\frac{1}{2}}$	$\frac{1}{12}(6+5h^{1,1}+h^{2,1}-6h^{1,1}_{\text{vert}})$
6d decompactifcation	$\frac{1}{2}$	$\frac{1}{12} \left( 42 + 4h^{1,1} + 2h^{2,1} - 6h^{1,1}(B_2) \right)$



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 slope of  $\Lambda_s$  given by  $\frac{|\nabla \Lambda_s|}{\Lambda_s} \rightarrow \alpha + \frac{\beta}{s} \log s$ 

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 $\frac{|\nabla \Lambda_s|}{\Lambda_s}$  typically approaches asymptotic value from **above!** (see also David's talk on Monday)

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[van de Heisteeg, Vafa, MW, Wu '23]

Use the **species scale** to bound slowly varying potentials with V > 0!

 $\rightarrow$  For slowly rolling field bound  $V(\phi) \leq \Lambda_s^2$  cf. [Hebecker, Wrase '18, Scalisi, Valenzuela '18]

(see Marco's talk)

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Motivation:

→ dS entropy  $S_{dS} \sim H^{-2}$  should *at least* account for light species! →  $H^2 < \Lambda_s!$ 

(alternatively: smallest black hole should 'fit' into dS space!

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Unlike TCC: species scale argument can be used to fix A!

(At least in cases where species scale is known everywhere and supersymmetry is broken mildly)

 $\rightarrow$  again in CY compactifications of Type II string theory!

Simple application: Type II CY compactifications with fluxes!

- [van de Heisteeg, Vafa, MW, Wu '23]
- Supersymmetry broken to  $\mathcal{N} = 1 \rightarrow$  species scale still given by  $F_1$ ??
- Yes! Since fluxes do not affect the topological string amplitudes!

[Vafa '00]

• Effective action contains two terms: [Ooguri, Vafa '03]

$$S_{\mathcal{N}=1} \supset \int d^4x \, d^2\theta \mathcal{F}^{2g} N_i \frac{\partial F_g}{\partial S_i}, \qquad S_{\mathcal{N}=1} \supset g \int d^4x \, d^2\theta \, \mathcal{W}^2 \mathcal{F}^{2(g-1)} F_g,$$
For g=1 is just the same term

as in N=2!

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For g=1 is just the same term as in N=2!

We can constrain the field range over which the potential is approximately flat  $V \simeq V_0$  as:

$$\Delta \phi \le -\sqrt{6} \log \frac{V_0}{M_{\rm pl}^4} + b$$

Can we get large field distances in the interior?

We can further constrain the field range over which the potential is approximately flat  $V \simeq V_0$  as:



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Can we get large field distances in the interior?

From asymptotic expression:  $\Lambda_s^2 = A e^{-2\lambda\Delta\phi}$ 

$$\rightarrow \Delta \phi_R(V_0) = -\frac{1}{2\lambda_R} \log[V_0] + \frac{1}{2\lambda_R} \log[A_R]$$

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![](_page_48_Figure_3.jpeg)

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![](_page_50_Figure_3.jpeg)

## Conclusions

In this talk:

- Discussed the moduli-dependence of the effective quantum gravity cut-off, i.e., the species scale beyond asymptotic regimes.
- Showed a universal bound  $\frac{|\nabla \Lambda_s|}{\Lambda_s} \leq \mathcal{O}(1)$  just from consistency of the effective action!
- Argued that in Type II CY compactifications the topological genus-1 free energy reliably computes the species scale everywhere in (vectormultiplet) moduli space.
- Allows to explicitly verify EFT bound in examples → naive asymptotic bound not correct!
- Species scale can be used to bound scalar potentials  $\rightarrow$  leads to same constraint as TCC!
- In Type II setup:  $F_1$  robust enough to still give species scale even in  $\mathcal{N} = 1$ 
  - → allows to fix  $\mathcal{O}(1)$  coefficients and to bound the range for approx. flat, positive potentials (including the interior)

# Thank you!!

Max Wiesner

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