

Interpolation and cosmological constant in non-SUSY heterotic strings

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- Y.K., NPB 990 (2023) 116160
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Introduction

■ Where is SUSY breaking scale?

- There is **no evidence** for SUSY in multi TeV scale according to the LHC.
- It is interesting to consider SUSY is already broken **at very high energies**.

■ We have more **non-SUSY** vacua than **SUSY** ones in 10D:

- Type IIA
 - Type IIB
 - Type I
 - Heterotic $SO(32)$
 - Heterotic $E_8 \times E_8$
 - Type 0A
 - Type 0B
 - Heterotic $SO(32)$
 - Heterotic $SO(16) \times E_8$
 - Heterotic $SO(16) \times SO(16)$
 - Heterotic $E_7^2 \times SU(2)^2$
 - Heterotic $SO(24) \times SO(8)$
 - ...
- difficulty: very large cosmological constant (vacuum energy density)

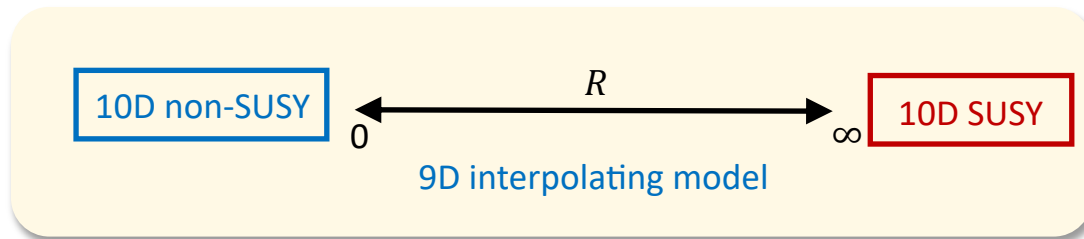
$$\Lambda^{(D)} \sim \mathcal{O}(M_S^D) \quad M_S : \text{string scale}$$

- We want to obtain small or vanishing cosmological const. **without SUSY**.

■ Our approach: “interpolating models”

- Low-dimensional models constructed by \mathbb{Z}_2 freely acting orbifolds
- The models are **Non-SUSY** where a radius R is finite.
- But **SUSY can be restored** in $R \rightarrow \infty$ (and/or 0) limits.

(Example)



Interpolate
between two 10D endpoints

■ In SUSY restored region ($R \approx \infty$),

$$\Lambda^{(9)} = \frac{\xi}{R^9} (n_F - n_B) + \mathcal{O}(e^{-R}) \quad \left[\begin{array}{l} \xi : \text{positive constant} \\ n_F, n_B : \# (\text{massless fermion, boson}) \end{array} \right]$$

Itoyama, Taylor (1986)

➤ $n_F = n_B \Rightarrow$ **exponentially suppressed cosmological constant**

■ The general heterotic models constructed by \mathbb{Z}_2 -twists:

• d -dim. compactified with $\#(\mathbb{Z}_2$ -twisted directions) **arbitrary**

• With **full set of moduli**: $C_{ij} = G_{ij} + B_{ij}$ & A_i^I turned on

$$(i = 10 - d, \dots, 9, I = 1, \dots, 16)$$

(\leftrightarrow with **all marginal deformations** considered)

➤ Show various interpolation patterns in $d = 1, 2$

➤ Find solutions of $n_F = n_B$ where cosmol. const. is exp. supp.

➤ Analyze Wilson-line stability of the effective potential

Outline

1. Introduction
2. Non-SUSY heterotic strings with general Z_2 twists
3. Interpolation in 9D and 8D
4. Cosmological constant
5. Summary

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Narain lattice of SUSY heterotic strings on T^d

- Modular inv. \Rightarrow internal momenta $P = Z\mathcal{E} \in \Gamma^{16+d,d}$
 - Narain lattice: even self-dual lattice w/ Lorentz. sign. $(16 + d, d)$
 - Labeled by an integer vector $Z = (q, \underline{m}, \underline{n}) \in \mathbf{Z}^{16} \times \mathbf{Z}^d \times \mathbf{Z}^d$
 - winding numbers
 - KK momenta

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winding numbers KK momenta
 - Turn on full moduli: $d(d + 16) = d^2 + 16d \Rightarrow (G_{ij} + B_{ij})$ & A_i^I
 - Consider a rectangular d -torus: $G_{ij} = R_i^2 \delta_{ij}$

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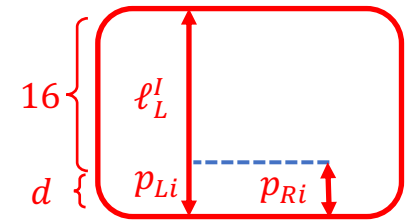
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➤ $P = (\ell_L, p_L; p_R)$

Narain, Sarmadi, Witten, (1986)

$$\begin{cases} \ell_L^I = \pi^I - m^i A_i^I, \\ p_{Li} = \frac{1}{\sqrt{2}R_i} \left(\pi \cdot A_i + n_i + m^j \left(G_{ij} + B_{ij} - \frac{1}{2} A_i \cdot A_j \right) \right) \\ p_{Ri} = \frac{1}{\sqrt{2}R_i} \left(\pi \cdot A_i + n_i - m^j \left(G_{ij} - B_{ij} + \frac{1}{2} A_i \cdot A_j \right) \right) \end{cases}$$



$\pi^I \equiv q^I \alpha_{16} \in \underline{\Gamma}^{16} \Leftarrow Spin(32)/\mathbf{Z}_2$ or $E_8 \times E_8$ lattice

Construction of non-SUSY heterotic strings

■ Z_2 freely acting orbifold (stringy Scherk-Schwarz comp.)

- Project out SUSY hetero on T^d by $\frac{1 + (-)^F \alpha}{2}$ (+ twisted sec. added)

$$Z_2 \text{ generator : } (-)^F \alpha \left\{ \begin{array}{l} F: \text{spacetime fermion \#} \\ \alpha: \text{shift of order 2 such as } \alpha |P\rangle = e^{2\pi i P \cdot \delta} |P\rangle \end{array} \right.$$

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↳ Non-SUSY strings depend on \hat{Z}

- Splitting the Narain lattice $\Gamma^{16+d,d}$ into $\Gamma_+^{16+d,d}$ and $\Gamma_-^{16+d,d}$:

$$\begin{aligned} \Gamma_+^{16+d,d} &= \{ P \in \Gamma^{16+d,d} \mid \delta \cdot P \in \mathbb{Z} \} \\ \Gamma_-^{16+d,d} &= \{ P \in \Gamma^{16+d,d} \mid \delta \cdot P \in \mathbb{Z} + 1/2 \} \end{aligned} \quad \Rightarrow \quad \alpha |P\rangle = \begin{cases} + |P\rangle & \text{for } P \in \Gamma_+^{16+d,d} \\ - |P\rangle & \text{for } P \in \Gamma_-^{16+d,d} \end{cases}$$

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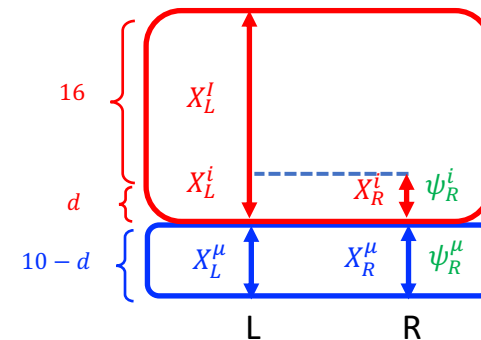
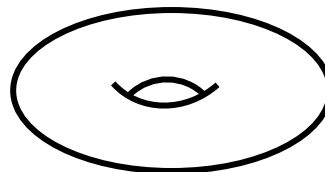
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Bosons/Fermions live in $\Gamma_+^{16+d,d} / \Gamma_-^{16+d,d}$ respectively ➡ SUSY breaking

1-loop partition function

- Heterotic strings on T^d (with maximal SUSY)

$$Z^{T^d} = \frac{Z_B^{(8-d)}}{X_L^\mu, X_R^\mu} \frac{(\bar{V}_8 - \bar{S}_8)}{\psi_R^\mu, \psi_R^i} \frac{Z_{\Gamma^{16+d,d}}}{X_L^i, X_L^i, X_R^i}$$



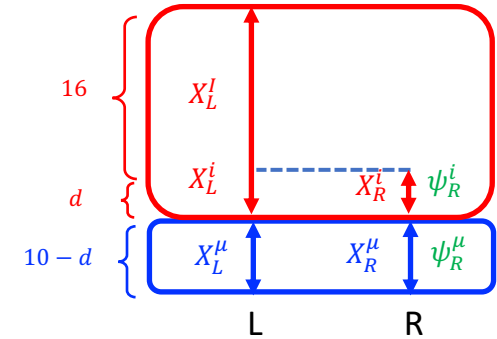
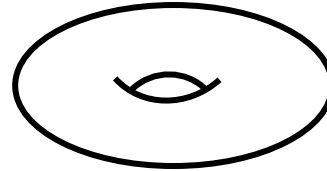
orbifolding by $(-)^F \alpha$

$$\left(\begin{array}{l} Z_B^{(8-d)} = \tau_2^{-\frac{8-d}{2}} (\eta\bar{\eta})^{-(8-d)} \\ Z_{\Gamma^{16+d,d}} = \eta^{-(16+d)} \bar{\eta}^{-d} \sum_{p \in \Gamma^{16+d,d}} q^{\frac{1}{2}p_L^2} \bar{q}^{\frac{1}{2}p_R^2} \\ q = e^{2\pi i \tau}, V_8, S_8, O_8, C_8: SO(8) \text{ characters} \end{array} \right)$$

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- Non-SUSY Heterotic strings

$$Z_{(\hat{Z})}^{SUSY} = Z_B^{(8-d)} \left\{ \frac{\bar{V}_8 Z_{\Gamma_+^{16+d,d}}}{\text{vector}} - \frac{\bar{S}_8 Z_{\Gamma_-^{16+d,d}}}{\text{spinor}} + \frac{\bar{O}_8 Z_{\Gamma_{\pm}^{16+d,d+\delta}}}{\text{scalar}} - \frac{\bar{C}_8 Z_{\Gamma_{\mp}^{16+d,d+\delta}}}{\text{co-spinor}} \right\}$$

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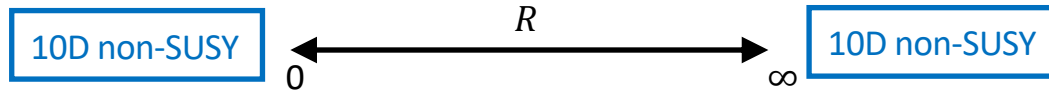
■ 9D Non-SUSY heterotic models ($d = 1$)

Itoyama, Koga, Nkajima (2021)

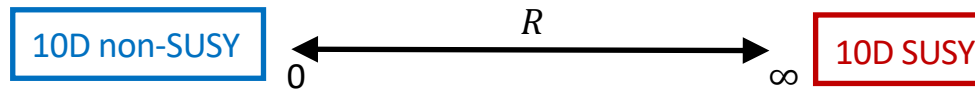
➤ $2^2 = 4$ classes exist

$$|\hat{\pi}|^2 + 2\hat{m}\hat{n}^t = 0 \pmod{4}$$

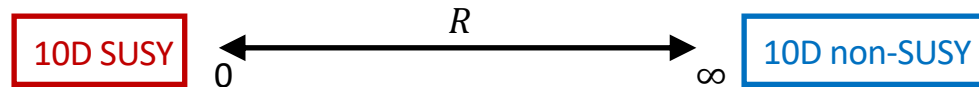
- class (1): $|\hat{\pi}|^2 = 0 \pmod{4}$, $(\hat{m}; \hat{n}) = (0; 0)$



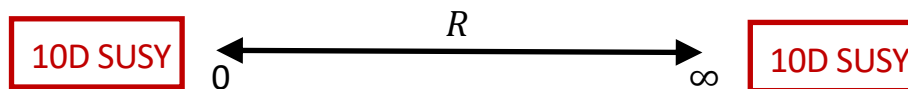
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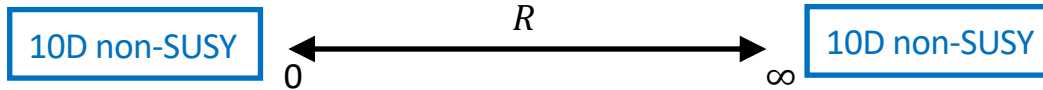
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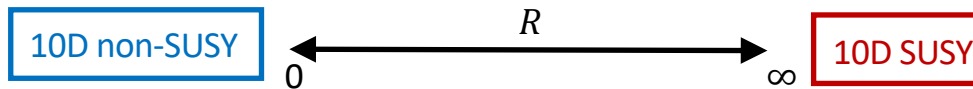
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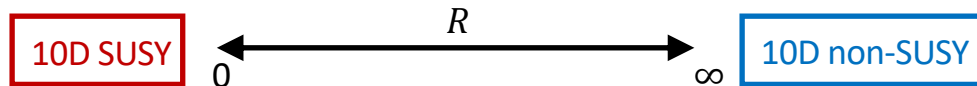


non-SUSY heterotic strings on a circle

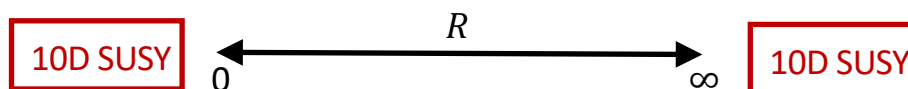
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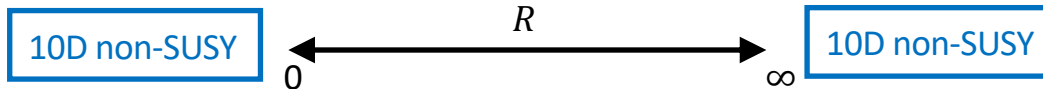
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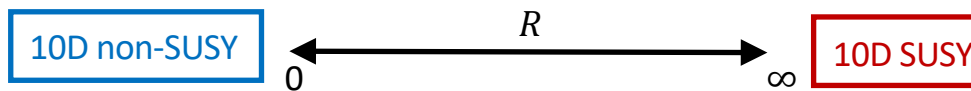
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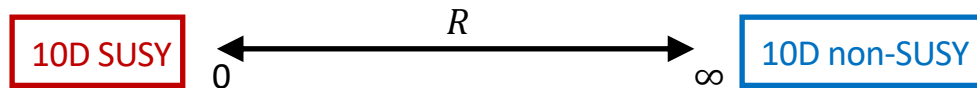


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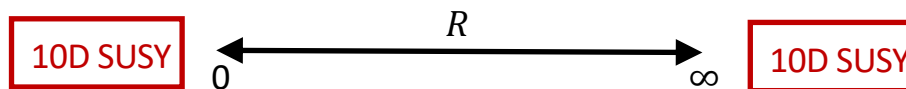


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interpolation between SUSY and non-SUSY vacua

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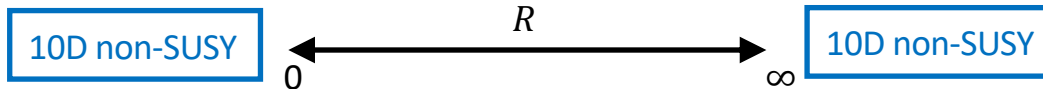
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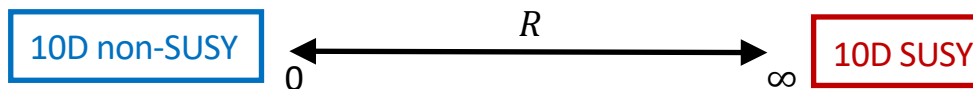
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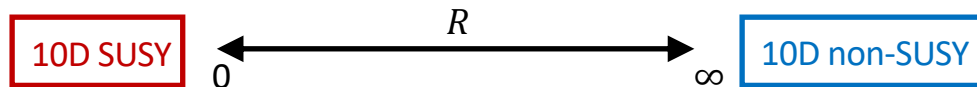


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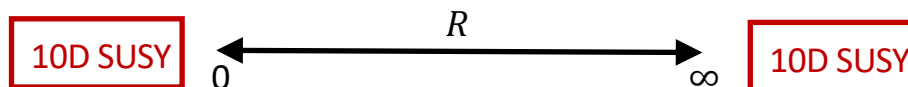


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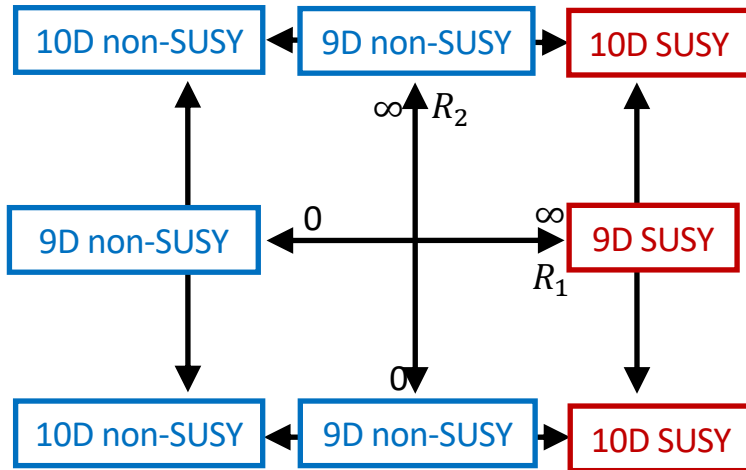
SUSY restored at both of the endpoints

■ 8D Non-SUSY heterotic models ($d = 2$)

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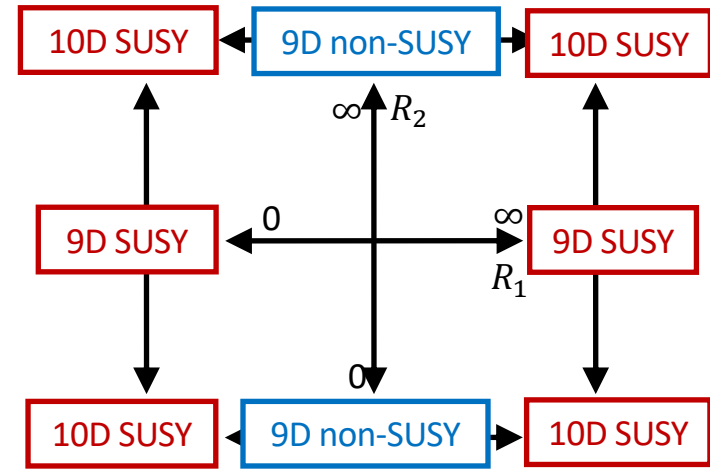
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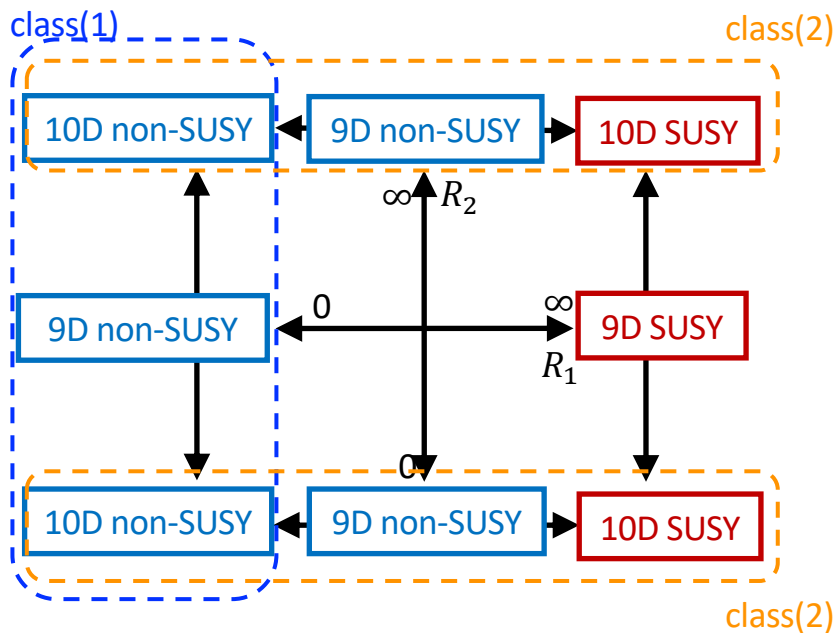


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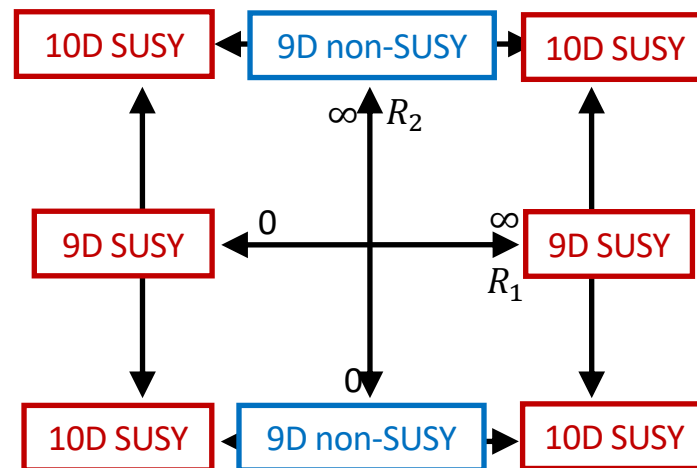
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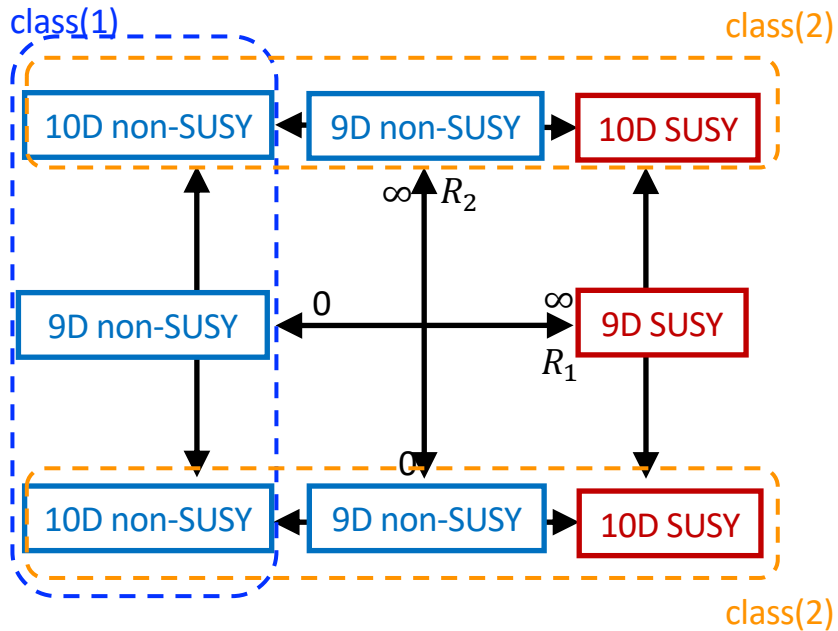


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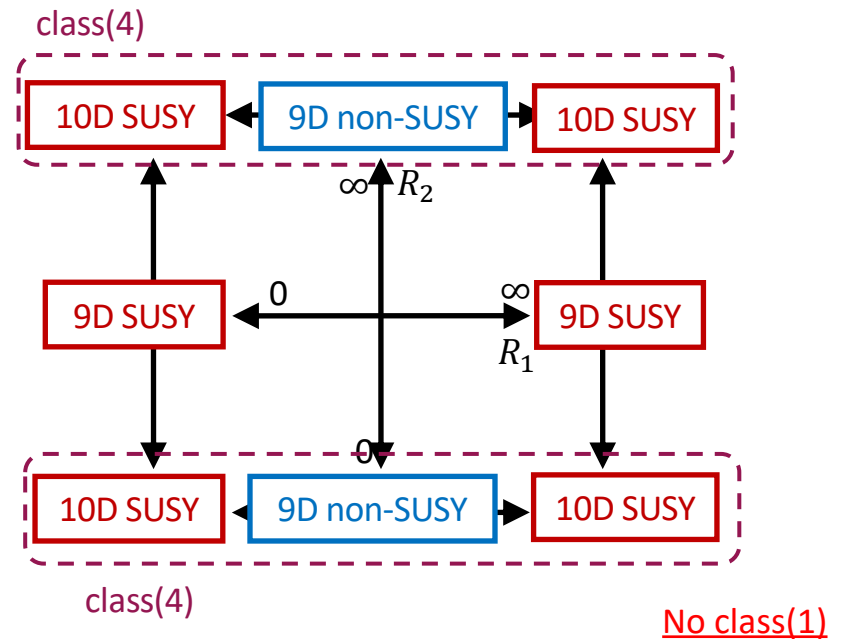
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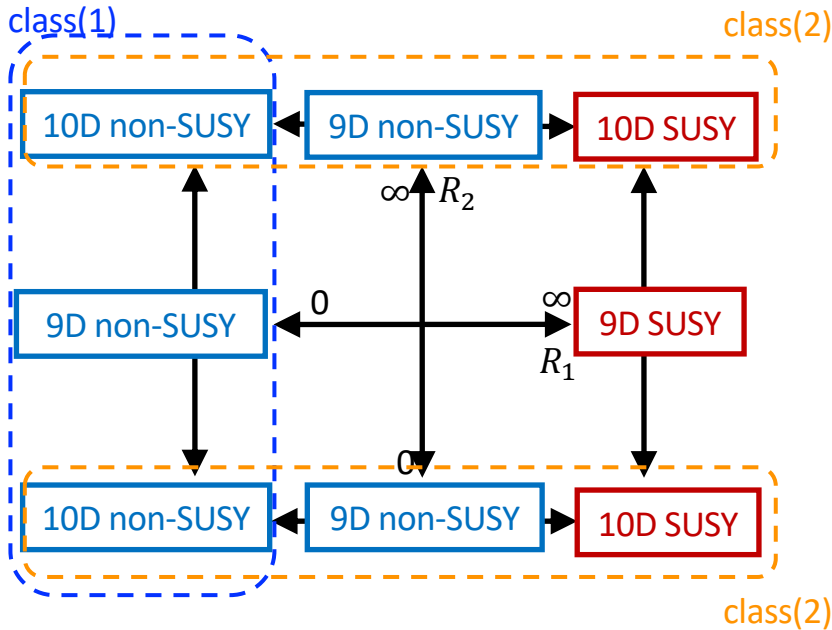
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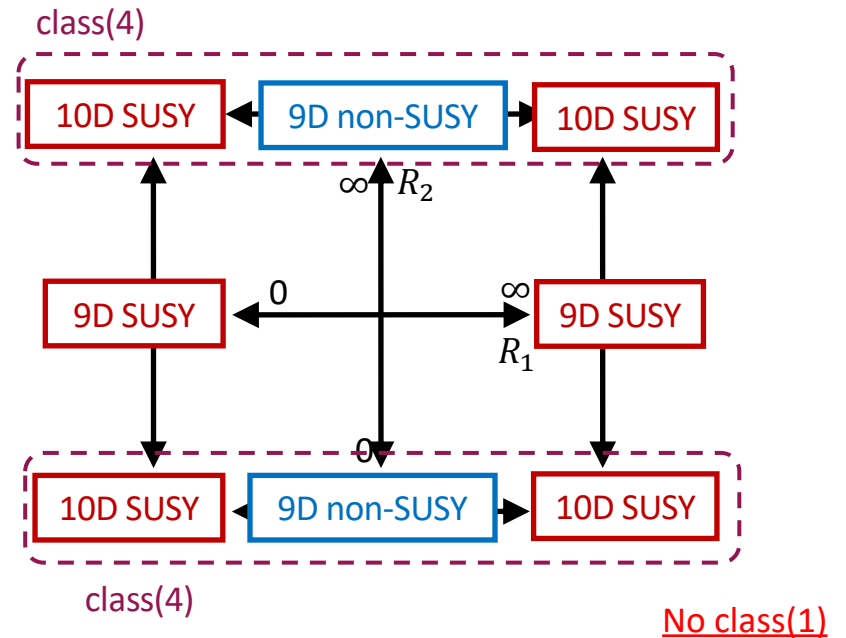
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10D (Non-)SUSY condition:

Limits of R_1, R_2	10D SUSY model	10D Non-SUSY model
$(R_1, R_2) \rightarrow (\infty, \infty)$	$\hat{m}^1 + \hat{m}^2 > 0$	$\hat{m}^1 + \hat{m}^2 = 0$
$(R_1, R_2) \rightarrow (\infty, 0)$	$\hat{m}^1 + \hat{n}_2 > 0$	$\hat{m}^1 + \hat{n}_2 = 0$
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Cosmological Constant

- 1-loop cosmological constant (effective potential) :

$$\Lambda^{(10-d)} = -\frac{1}{2} (2\pi\sqrt{\alpha'})^{-(10-d)} \int_{\mathcal{F}} \frac{d^2\tau}{\tau_2^2} Z_{(\hat{Z})}^{SUSY}$$

Fundamental Region :

$$\mathcal{F} = \left\{ \tau = \tau_1 + i\tau_2 \in \mathbb{C} \mid -\frac{1}{2} \leq \tau_1 \leq \frac{1}{2}, |\tau| \geq 1 \right\}$$

- Assignment of (\hat{m}, \hat{n}) :

$$\begin{array}{ll} (\hat{m}^{a(2)}, \hat{n}_{a(2)}) = (1, 0) & \text{for } a(2) = 1, \dots, D_2 \\ (\hat{m}^{a(4)}, \hat{n}_{a(4)}) = (1, 1) & \text{for } a(4) = D_2 + 1, \dots, D \\ (\hat{m}^{b(3)}, \hat{n}_{b(3)}) = (0, 1) & \text{for } b(3) = D + 1, \dots, D + D_3 \\ (\hat{m}^{b(1)}, \hat{n}_{b(1)}) = (0, 0) & \text{for } b(1) = D + D_3 + 1, \dots, d \end{array} \left. \begin{array}{l} \} \\ \} \end{array} \right\} \begin{array}{l} \text{SUSY at } R_a \rightarrow \infty \\ \text{Non-SUSY at } R_b \rightarrow \infty \end{array}$$

Formula for cosmological constant

■ Consider $D \geq 1$, all $R_i \approx \infty$: SUSY is restored

• Up to exponentially suppressed terms,

$$\Lambda^{(10-d)} \sim -\frac{4! \cdot 2^{d-1}}{\pi^{15-d} (\sqrt{\alpha'})^{10-d}} \left(\prod_{i=1}^d R_i \right) \sum_{\vec{n}} \left\{ \sum_{a=1}^D (2n_a - 1)^2 R_a^2 + \sum_{b=D+1}^d (2n_b)^2 R_b^2 \right\}^{-5} \\ \times 8 \left(24 + \sum_{\pi \in \Delta_g} \exp \left[2\pi i \left\{ \sum_{a=1}^D (2n_a - 1)(\pi \cdot A_a) + \sum_{b=D+1}^d n_b (\pi \cdot A_b) \right\} \right] \right)$$

Δ_g : nonzero roots of $SO(32)$ or $E_8 \times E_8$, not Δ_{\pm}

\Rightarrow CC does not depend on all the other endpoint models

massless
condition



$$2\pi \cdot A_a \in \mathbb{Z}$$

$$\pi \cdot A_b \in \mathbb{Z}$$

$$\Lambda^{(10-d)} \sim \frac{4! \cdot 2^{d-1}}{\pi^{15-d}} \left(\prod_{i=1}^d R_i \right) \sum_{\vec{n}} \left\{ \sum_a (2n_a - 1)^2 R_a^2 + \sum_b (2n_b)^2 R_b^2 \right\}^{-5} (n_F - n_B)$$

Solutions of $n_F = n_B$

■ SUSY $SO(32)$ endpoint model:

$$\Delta_{SO(32)} = \{(\pm, \pm, 0^{14})\}$$

■ Simplest configurations:

- A_a^I ($a = 1, \dots, D$) are the same configuration
- A_b^I ($b = D + 1, \dots, d$) are taken to be 0

$$A_a = \left(0^p, \left(\frac{1}{2}\right)^q\right) \quad (p + q = 16), \quad A_b = (0^{16})$$

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- $D \in 2\mathbf{Z} + 1$: $n_F - n_B = 4pq - \{2p(p - 1) + 2q(q - 1)\} - 24$

$$n_F = n_B \Rightarrow (p, q) = (9, 7), (7, 9)$$

Solutions of $n_F = n_B$

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Cosmological constant is exponentially suppressed
when the gauge group is $SO(18) \times SO(14)$

Outline

1. Introduction
2. Non-SUSY heterotic strings with general Z_2 twists
3. Interpolation in 9D and 8D
4. Cosmological constant
5. Summary

Summary

- $(10 - d)$ D Non-SUSY models are constructed by orbifolding by $(-)^F \alpha$
(α : shift of order 2 in Narain lattice)
- Various interpolations are shown in $d = 2$ case
- Cosmological constant of $(10 - d)$ D Non-SUSY models in $R_i \approx \infty$ is

$$\Lambda^{(10-d)} \sim \frac{4! \cdot 2^{d-1}}{\pi^{15-d}} \left(\prod_{i=1}^d R_i \right) \sum_{\vec{n}} \left\{ \sum_a (2n_a - 1)^2 R_a^2 + \sum_b (2n_b)^2 R_b^2 \right\}^{-5} (n_F - n_B) + \mathcal{O}(e^{-R/\sqrt{\alpha'}})$$

- Find the configurations of WLs which give exp. supp. cosmol. const.

Out look

Higher-loop correction, (meta)stable vacua, cosmology, ...