# Theory of quenching in small systems

How do we know what to expect if we're not calculating energy-loss in an honestly small system?

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Faraday et.al. 2305.13182







## Many side-effects

#### L-dependence is not obviously L<sup>2</sup>

DGLV

All-pa \_corre

$$\begin{split} \Delta E_{ind}^{(1)} &= \frac{C_R \alpha_s L E}{\pi \lambda_g} \int dx \int \frac{d^2 \mathbf{q}_1}{\pi} \frac{\mu^2}{(\mu^2 + \mathbf{q}_1^2)^2} \frac{d^2 \mathbf{k}}{\pi} \\ &\times \int d\Delta z \bar{\rho}(\Delta z) \bigg[ - \frac{2(1 - \cos\{(\omega_1 + \tilde{\omega}_m)\Delta z\})}{(\mathbf{k} - \mathbf{q}_1)^2 + M^2 x^2 + m_g^2} \\ &\times \bigg( \frac{(\mathbf{k} - \mathbf{q}_1) \cdot \mathbf{k}}{\mathbf{k}^2 + m_g^2 + x^2 M^2} - \frac{(\mathbf{k} - \mathbf{q}_1)^2}{(\mathbf{k} - \mathbf{q}_1)^2 + M^2 x^2 + m_g^2} \bigg) \\ &+ \frac{1}{2} e^{-\mu_1 \Delta z} \bigg\{ \bigg( \frac{\mathbf{k}}{\mathbf{k}^2 + m_g^2 + x^2 M^2} \bigg)^2 \\ &\times \bigg( 1 - \frac{2C_R}{C_A} \bigg) \bigg( 1 - \cos\{(\omega_0 - \tilde{\omega}_m)\Delta z\} \bigg) \\ &+ \frac{\mathbf{k} \cdot (\mathbf{k} - \mathbf{q}_1)}{(\mathbf{k}^2 + m_g^2 + x^2 M^2)((\mathbf{k} - \mathbf{q}_1)^2 + M^2 x^2 + m_g^2)} \\ &\times \big( \cos\{(\omega_0 - \tilde{\omega}_m)\Delta z\} - \cos\{(\omega_0 - \omega_1)\Delta z\} \big) \bigg\} \bigg] \end{split}$$

 $\Delta E_{LO}^{(1)} = \frac{C_R \alpha_s}{4} \frac{L^2 \mu^2}{\lambda_g} \log \frac{E}{\mu}$ 

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ction  
$$\Delta E_{NLO}^{(1)} = \frac{EC_R \alpha}{\pi \lambda_g} \left( -\frac{2C_R}{C_A} \right) \frac{L}{2 + L\mu} \times \left( \log \left\{ \frac{2EL}{(2 + L\mu)} \right\} - 1 \right)$$



Formation time of gluon becomes critically important

This is GLV, are multiple interactions important?

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What about interference with other diagrams?

## Temperature is also important

 $\lambda_{mfp} \sim \frac{1}{\rho\sigma} \sim \frac{1}{g^2T}$ 

 $\mu_D \sim gT$ 

### Also important for $v_2$ and $R_{AA}$



Sensitivity to the nature of radiation in the early stages



Or sensitivity to late-time temperature profile?



How can we be faithful to rapidly evolving temperature profile in a pQCD calculation?

What does the temperature profile have to be to get R\_pPb~1 but v\_2 >0?

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## Speculation on a system-size scan

- Can we get to the space-time structure of medium-induced modification with a system-size scan?
- What would we have to compute?
  - Eg. Korinna's suggestion of color-coherence and formation time. Maybe these provide a scale?
  - Are there observables that predict an experimentally measurable modification in OO but predict a too-small modification in pPb?
- Can one isolate the effect of the pre-hydrodynamic phase in small systems and see it become sub-dominant in larger systems?

## Backups

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## Description of $R_{AA}$ and $v_2$ is non-trivial in AA



Sensitivity to the nature of radiation in the early stages

Or sensitivity to late-time temperature profile?

40-50%

v<sub>2</sub>(SP)

p\_(GeV/c)

ALICE

2

ο π

ΔŘ

D P

10-20%

6

#### Or do you need comprehensive hadronization picture?

## RAA and v2 are not useful at high-pT in small systems

- Reliance on a reference system
- Steeply falling production spectrum
  - Sensitive only to large  $\Delta E$
  - Sensitive to PDFs and nPDFs
  - Species-dependent
- Sensitive to initial condition
  - Geometry
  - Momentum anisotropy
- Sensitive to jet fragmentation
- Supposed to quantify  $\Delta E$ , but
  - $\Delta E \leftarrow L \leftarrow N_{coll}$ : uncontrolled  $\Delta E = \Delta E(T)$ : T is uncontrolled





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## All-pathlength energy loss - Details



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