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KEK / Accelerator Laboratory

Workshop on Dust Charging and Beam-Dust Interaction in Particle Accelerators @ CERN

2023-06-13

## Normal-Conducting Accelerating Cavities for the SuperKEKB Rings





## BD Signal A: Fast drop of the accelerating field



## **BD Signal B: Current flash**

(During the high-power test of the RF cavity for the DR)



Field emitted e<sup>-</sup> → Impact on the metal surface → X-ray radiation

**X-ray detector** (plastic scintillator + PMT)























The information on the breakdown trigger seed can be mostly

- destroyed by
- hidden in

the much larger phenomenon of a vacuum arc.

# **Method of This Study**





## **3 TV cameras** for **Multi-directional and wide-field** observation



## Setup of the Experiment



(No beam injected into the RF cavity during the high-power test)

Accelerating Structure used in This Study: normal-conducting standing-wave single-cell cavity operated with continuous 509-MHz input wave developed for <u>SuperKEKB</u> Positron Damping Ring (DR)





- Accel. mode: 509 MHz Continuous Wave (CW) TM<sub>010</sub>
- Made of Oxygen Free Copper (Class1)
- *Q*<sub>0</sub> = ~30000
- $\blacksquare R_{\rm sh}/Q_0 = 150 \ \Omega$
- Spec.  $V_c = 0.8 \text{ MV} (\rightarrow E_{acc} = 3.1 \text{ MV/m})$
- Wall-loss power: ~150 kW @ $V_c$  = 0.80 MV

For more details, T. Abe, "Completion of the First Production Version of the Accelerating Cavity for the SuperKEKB Positron Damping Ring", KEK Accl. Lab. Topics 2013/10/7 (web article): <u>http://www2.kek.jp/accl/eng/topics/topics131007.html</u>

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## Details on the RF cavity used in this experiment



### The results obtained in this study have generality.





## Statistics on all the 205 breakdown events

- 10% "Pyrotechnic" breakdowns
  - Observed only in the initial stage of RF conditioning



## Example (2/3)

#### Stable bright spots on the end plate

at  $V_{\rm c}$  = 0.95 MV ( $E_{\rm acc}$  = 3.7 MV/m)



(a) 1 frame (1/30 s) before this cavity breakdown.

## Example (2/3)

One of the stable bright spotexploded at the momentat  $V_c = 0.95$  MV ( $E_{acc} = 3.7$  MV/m)of breakdown,



(a) 1 frame (1/30 s) before this cavity breakdown.

(b) At the moment of this cavity breakdown.

## Example (2/3)



(a) 1 frame (1/30 s) before this cavity breakdown.

(b) At the moment of this cavity breakdown.

(c) Shortly after recovering from this cavity breakdown at  $V_c = 0.95$  MV.

Mini-Supernova!



Time [s]

## Statistics on all the 205 breakdown events

- 10% "Pyrotechnic" breakdowns
  - Observed only in the initial stage of RF conditioning
- 25% accompanied by a bright-spot (BS) explosion



## Example (3/3)

#### Spot-type explosion not originating from a stable bright spot

No bright spot in this area  $(V_c = 0.65 \text{ MV} (E_{acc} = 2.5 \text{ MV/m}))$ 



(a) 1 frame (1/30 s) before this cavity breakdown.

A spot-type explosion at the moment of breakdown



(b) At the moment of this cavity breakdown.

Down



(c) 1 frame (1/30 s) after this cavity breakdown.



## Statistics on all the 205 breakdown events

- 10% "Pyrotechnic" breakdowns
  - Observed only in the initial stage of RF conditioning
- 25% accompanied by a bright-spot (BS) explosion
- 40% accompanied by a spot-type explosion not originating from a stable bright spot

No observed breakdown events were accompanied by two or more explosions. → Such an explosion must be a breakdown trigger!

For more details:

- T. Abe, "Visual Imaging of Radio-Frequency Cavity Breakdown", KEK Accl. Lab. Topics 2016/10/5 (web article): <u>http://www2.kek.jp/accl/eng/topics/topics161005.html</u>
- ✓ T. Abe, et al., "Breakdown Study Based on Direct In-Situ Observation of Inner Surfaces of an RF Accelerating Cavity during a High-Gradient Test", Physical Review Accelerators and Beams 19, 102001 (2016).





## Further investigations



## We observed emission spectra for 10 stable bright spots (BSs).

#### **Upstream Endplate**

**Downstream Endplate** 



## **Observed Spectra**



- ✓ Pixel-sensitivity calibration
- ✓ Wavelength calibration
- $\checkmark\,$  Corrections for the mirror and view-port windows
- ✓ Background subtraction

 No line or band spectrum in the visible light or near-infrared regions

2 The spectra move toward lower wavelengths at higher cavity voltages



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applied

Experimental setup to take a spectrum and temperature simultaneously to calibrate the spectral sensitivity of the hyperspectral camera



## The thermocouple temperature: 930 degC

#### Raw data shown at 800 nm



Dust Workshop @ CERN / Fireball-triggered Vacuum Breakdown of RF Cavities (2023-06-13)

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## How to derive the temperature from the spectrum



## Measured temperatures in degC @ $V_c = 0.95$ MV

Upstream Endplate





## Phase Diagram of Copper



## Candidates of the Fireball Materials

Should have a Sublimation point > 1,000°C in ultrahigh vacuum.

Element	Sublimation point [°C] @ 2 x 10 <sup>-5</sup> Pa	Remarks
W	2258.6	Materials of the SuperKEKB collimator heads
Та	2123.4	
C (Graphite)	1769.9	Heater materials of vacuum furnaces for RF- cavity fabrication
Мо	1705.7	
Zr	1565.7	Material of NEG pump strips (e.g. St707)
Ti	1162.6	Material of the KEKB collimator heads
Au	894.7	
Cu	795.3	
Al	765.0	
Ве	764.2	No chance of leading to fireball breakdown in RF cavities made of Cu
Ag	635.3	
	541.0	

Data from <a href="https://www.iap.tuwien.ac.at/www/surface/vapor\_pressure">https://www.iap.tuwien.ac.at/www/surface/vapor\_pressure</a>

## Simple Simulation using CST MPHYSICS STUDIO

Simplified model of graphite



## **Three High-Speed Cameras**



Model: "HAS-D3M"

- ✓ Made by DITECT Co., Ltd. (<u>https://www.ditect.co.jp/en/</u>)
- ✓ Max. 100,000 fps (1,000 or 2,000 fps used in this study)
- ✓ Three cameras can be synchronized frame by frame.

# We detected 40 BD events with this setup.

# In 4 BD events out of the 40 BD events, we observed a flying bright object!

# Slow motion video (slowed 200 times) of one of the strange breakdown events

#### During high-power operation with $V_c = 0.88 \text{ MV}$

(No beam injected into the RF cavity during this high-power test)



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## Downstream $\rightarrow$ Upstream $\rightarrow$ Impact $\rightarrow$ Breakdown



## Other 3 events

(during the high-power test)



## From the observation using the high-speed cameras

■ More flying bright objects (→ breakdowns) observed using the high-speed cameras than using the low-spec cameras

Breakdowns with spot-type explosion should be caused by flying bright objects.

- The high-speed cameras are not sensitive to infrared.
  - $\rightarrow$  (Temperature of the emitters) > 1,000 degC if the object size is  $\lesssim$  0.1 mm
    - → Flying "fireballs"



## Observation of the size of the bright-spot emitters



#### <mark>≲ ~0.1mm</mark>

# 2017/04/07 17:47:12 @ mm





## Physical process of the fireball hypothesis, leading to SBLs (Sudden Beam Losses)



## **Essential Situation for the fireball breakdown**

• Coexistence of different materials with largely different sublimation points in the same place.



- In the case of the SuperKEKB collimators
  - Heads made of W or Ta with a high sublimation point
  - Vacuum chamber made of Cu with a low sublimation point

## Simulation using CST / PIC solver for the SuperKEKB collimator



## Summary

Fireballs, > 1,000degC electrically-charged micro-particles, have been observed in UHF CW cavities at KEK.

- From the era of TRISTAN
- Adhering on copper surfaces or flying
- So far, we can operate the RF cavities with fireballs.
  - > The breakdown rates are low enough for beam operation so far.
- Fireballs have been found to be a dominant trigger seed in RF cavity breakdowns for UHF CW cavities
  - Except for early stage of high-power RF conditioning.
- T. ABE proposed the fireball hypothesis to explain the trigger of the sudden beam losses at SuperKEKB.
  - Simulation and experimental studies are on-going.

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## For more details on the fireball breakdown

#### ■ KEK Accl. Lab. Topics (Web article)

- "Minuscule Gremlins Cause Vacuum Breakdown in Radio-Frequency Accelerating Cavities"
  - https://www2.kek.jp/accl/eng/topics/topics190122.html

#### Presentations at Workshops

- <u>T. Abe, "Direct Observation of Breakdown Phenomena in Normal-Conducting Accelerating Structures: 509-MHz</u> <u>Continuous-Wave Cavity and 11.4-GHz Pulsed-Wave Cavity"</u>, presented at <u>the 12th International Workshop on</u> <u>Breakdown Science and High Gradient Technology (HG2019)</u>, Chamonix, France, June 2019.
- <u>T. Abe, "Updated Results of Breakdown Study for 509-MHz Continuous-Wave Accelerating Cavities based on Direct</u> <u>In-situ Observation</u>", presented at the 7th International Workshop on Mechanisms of Vacuum Arcs (MeVArc 2018), Puerto Rico, May 20-24, 2018.

#### Original paper

• <u>T. Abe, et al., "Direct Observation of Breakdown Trigger Seeds in a Normal-Conducting RF Accelerating Cavity",</u> Physical Review Accelerators and Beams **21**, 122002, 2018.

Thank you for your attention!

# **Backup information**

A graphite particle in the constant and uniform magnetic field (509 MHz)

Ø0.01mm Mo ---- Ø0.10mm Mo ---- Ø1.00mm Mo



#### ABE, KAGEYAMA, SAKAI, TAKEUCHI, and YOSHINO

PHYS. REV. ACCEL. BEAMS 21, 122002 (2018)

FIG. 22. Simulations of temperatures of spherical microparticles made of graphite (black lines), molybdenum (green lines), and copper (magenta lines) with diameters of 1.0 mm (solid lines), 0.1 mm (dashed lines), and 0.01 mm (dotted lines) located in a vacuum. Radiation cooling was calculated according to the Stefan–Boltzmann law with an emissivity of 0.8 for graphite and 0.1 for copper and molybdenum. Heat capacities of 0.71, 0.39, and 0.28 kJ/K/kg were used for graphite, copper, and molybdenum, respectively. (a) and (b) Application of a 508.9-MHz magnetic field of 7000 A/m to graphite microparticles with diameters of 1.0 and 0.1 mm, respectively, assuming an electric conductivity of  $1.0 \times 10^5$  S/m for graphite. (c) Temperature variation of the microparticles from an initial temperature of  $40 \,^{\circ}\text{C}$  with heat generation by a magnetic field of 7000 A/m. (d) Temperature variation of the microparticles from an initial temperature of 1200 °C without heat generation.