

# *Background simulations for the BULLKID-DM experiment*



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# Outline

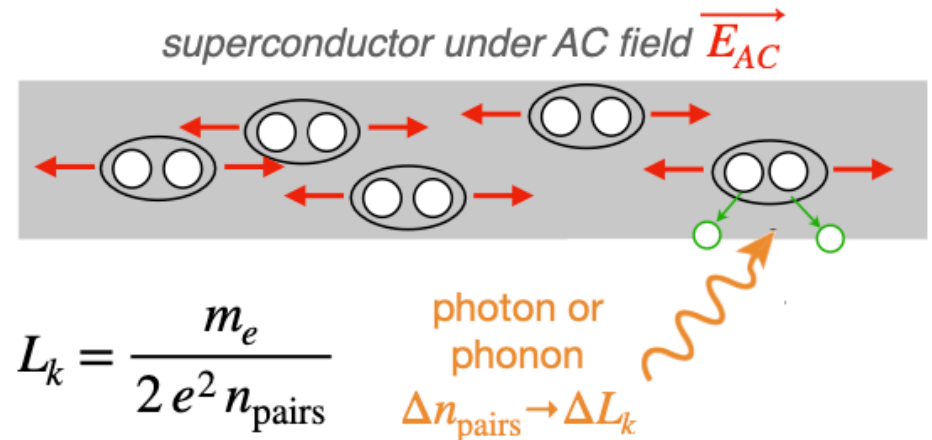
- Experiment overview
- DANAE setup at Sapienza
- Monte Carlo simulations for BULLKID-DM with GEANT4
- BULLKID-DM experiment underground
- Final remarks

# BULLKID: BULky and Low-threshold Kinetic Inductance Detectors

# Physics with kinetic inductance detectors

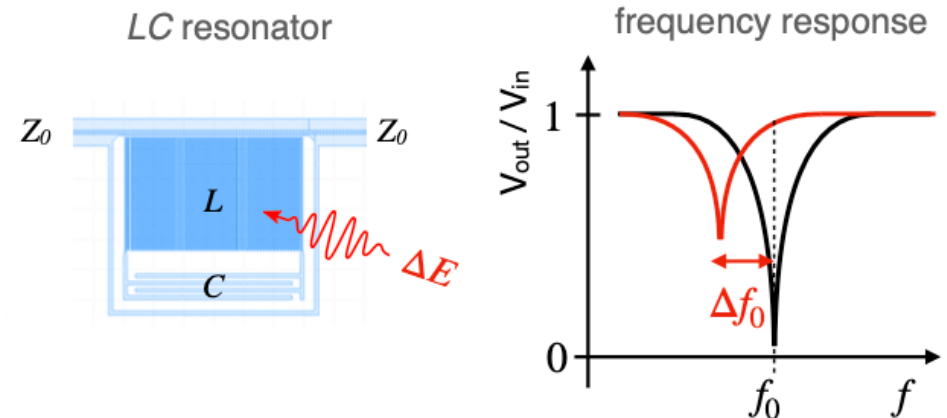
## AC superconductivity

- Electrons bound into Cooper pairs
- Kinetic inductance from physical inertia of mass pairs dependent on Cooper pair density
- High quality factors ( $Q \sim 10^4 - 10^6$ )



## Kinetic Inductance Detectors

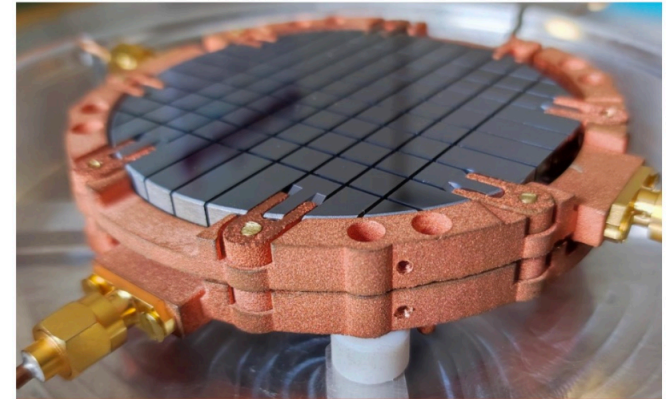
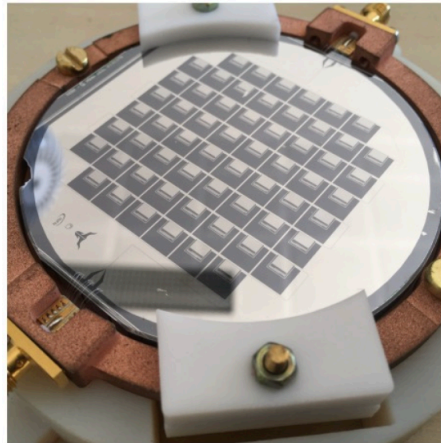
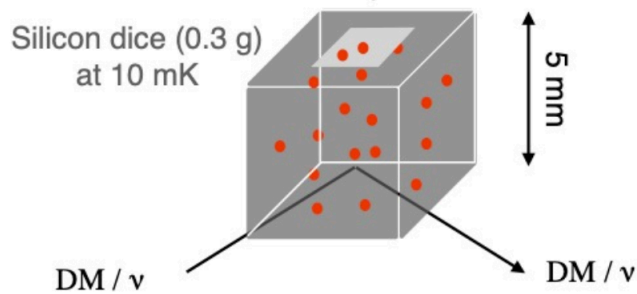
- Superconductor at  $T < 200$  mK (Aluminium)
- Resonant circuit
- Energy deposition breaks Cooper pairs





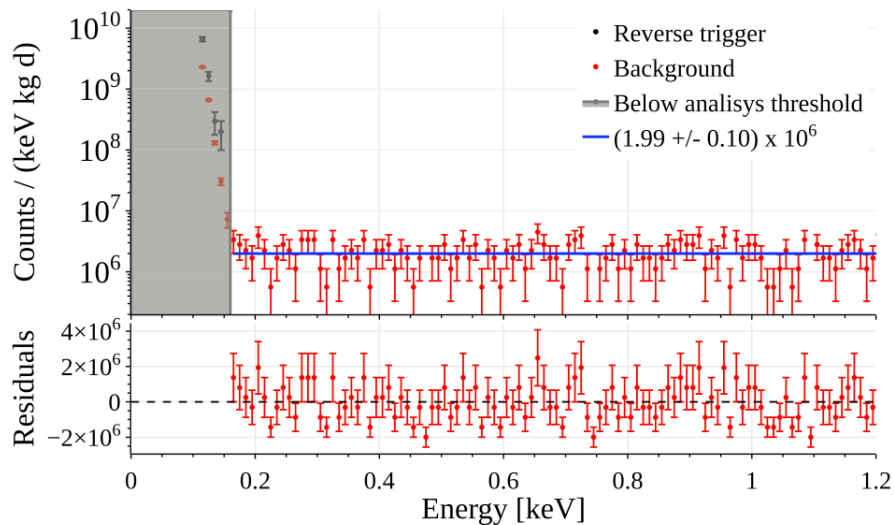
## A scalable detector for rare event searches

- Detection of phonons created by nuclear recoils in a silicon dice (0.3 g)
- Multiplexed readout: several KIDs coupled to the same feedline at different frequencies
- KID:  $\sim 2 \times 2 \text{ mm}^2 \times 50 \text{ nm}$ ,  $0.5 \mu\text{g}$
- Dices carved in a thick silicon wafer: 60 detectors in 1
- Calibration using optical photons of known energy



# First demonstration of BULLKID-DM

Eur. Phys. J. C (2024) 84:353



- Exposure of 39 hrs.  
(environmental backgrounds)
- Flat spectrum observed:  
( $2.0 \pm 0.1$  stat.  $\pm 0.2$  syst.)  $\times 10^6$   
counts/keV kg days
- Energy threshold of  $160 \pm 13$  eV
- Energy resolution:  $27 \pm 2$  eV

Experiment at Sapienza:

- Array of 60 cubic silicon particle absorbers (0.3 g each)
- Analysis on one of the central elements of the array using surrounding elements as veto



# BULLKID Collaboration



*25 people  
6 institutions  
4 countries*



**BULKID: Monte Carlo simulations for BULKID-DM with GEANT4**



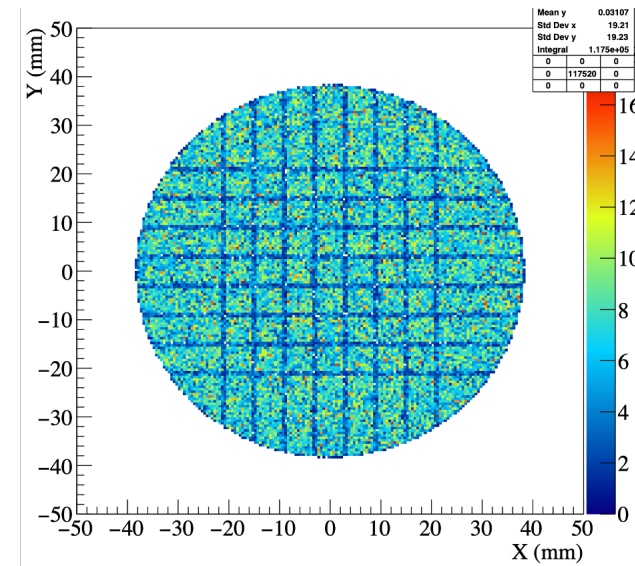
# Experiment at Sapienza

## GEANT4 model:

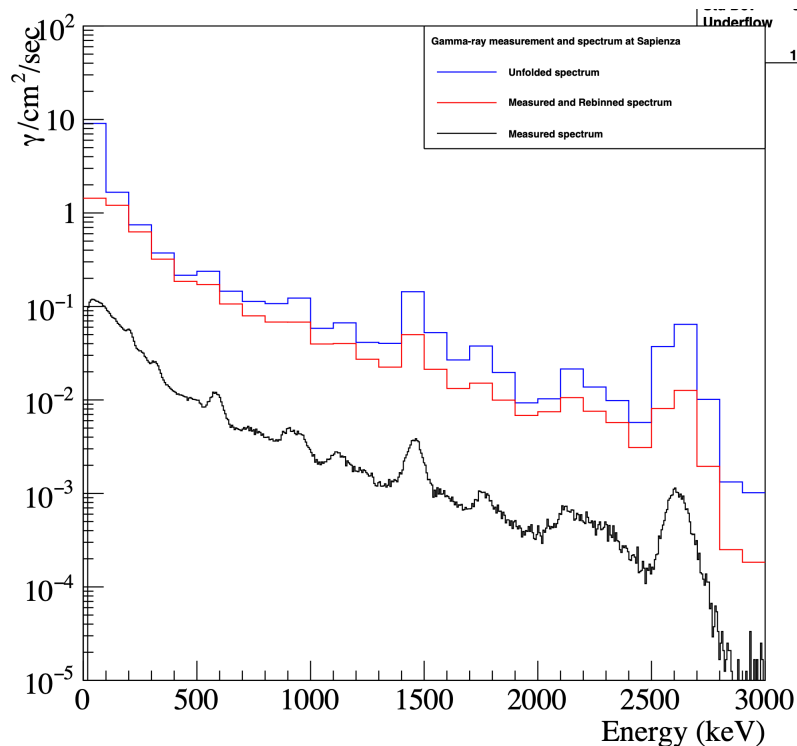
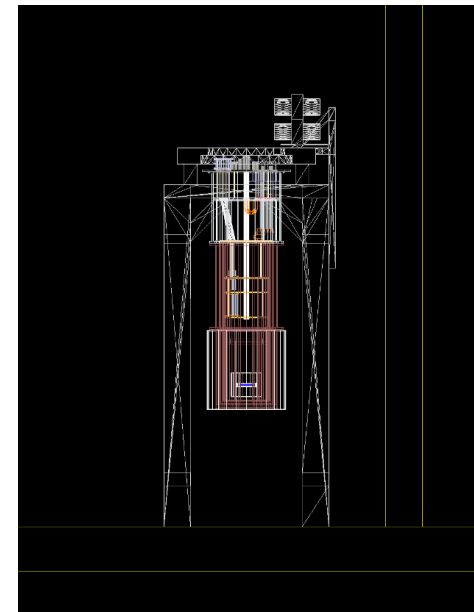
- GEANT4 v11.1.3 using Shielding Physics List

Simulation of backgrounds on surface:

- Gamma-rays and neutrons measured in laboratory
- CRY generator for cosmic rays

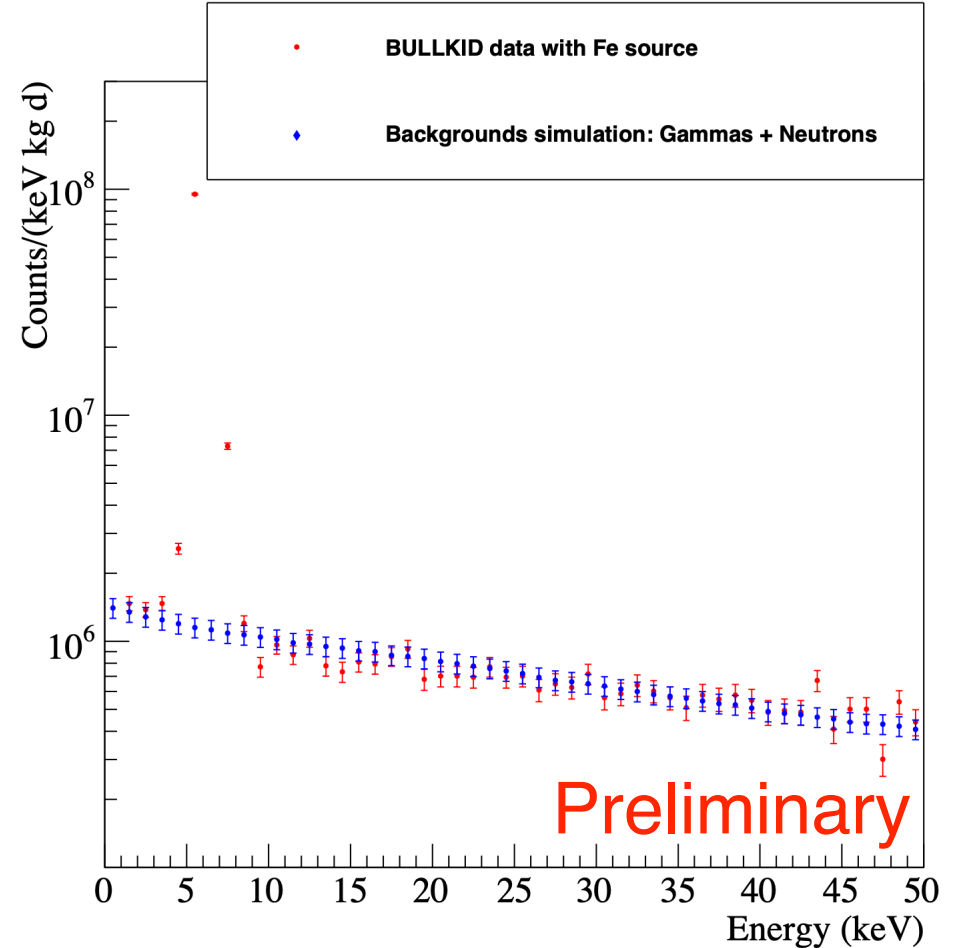
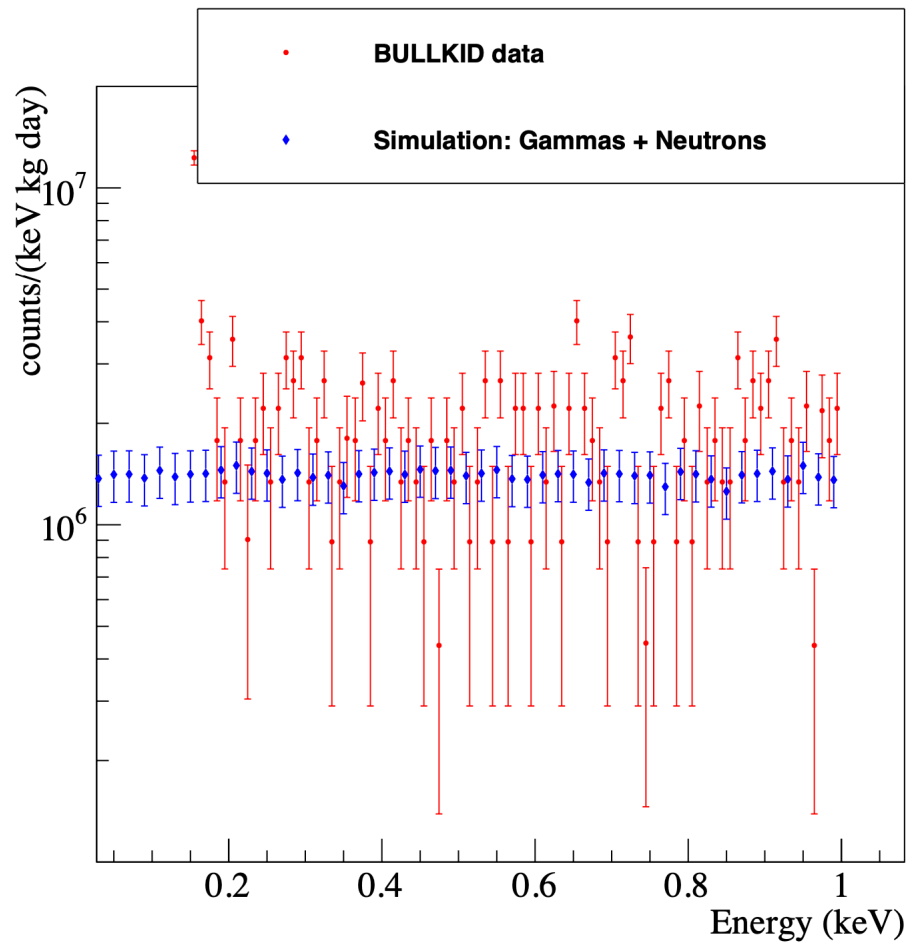


Detailed implementation of cryostat and inner detector



# Data vs Simulations

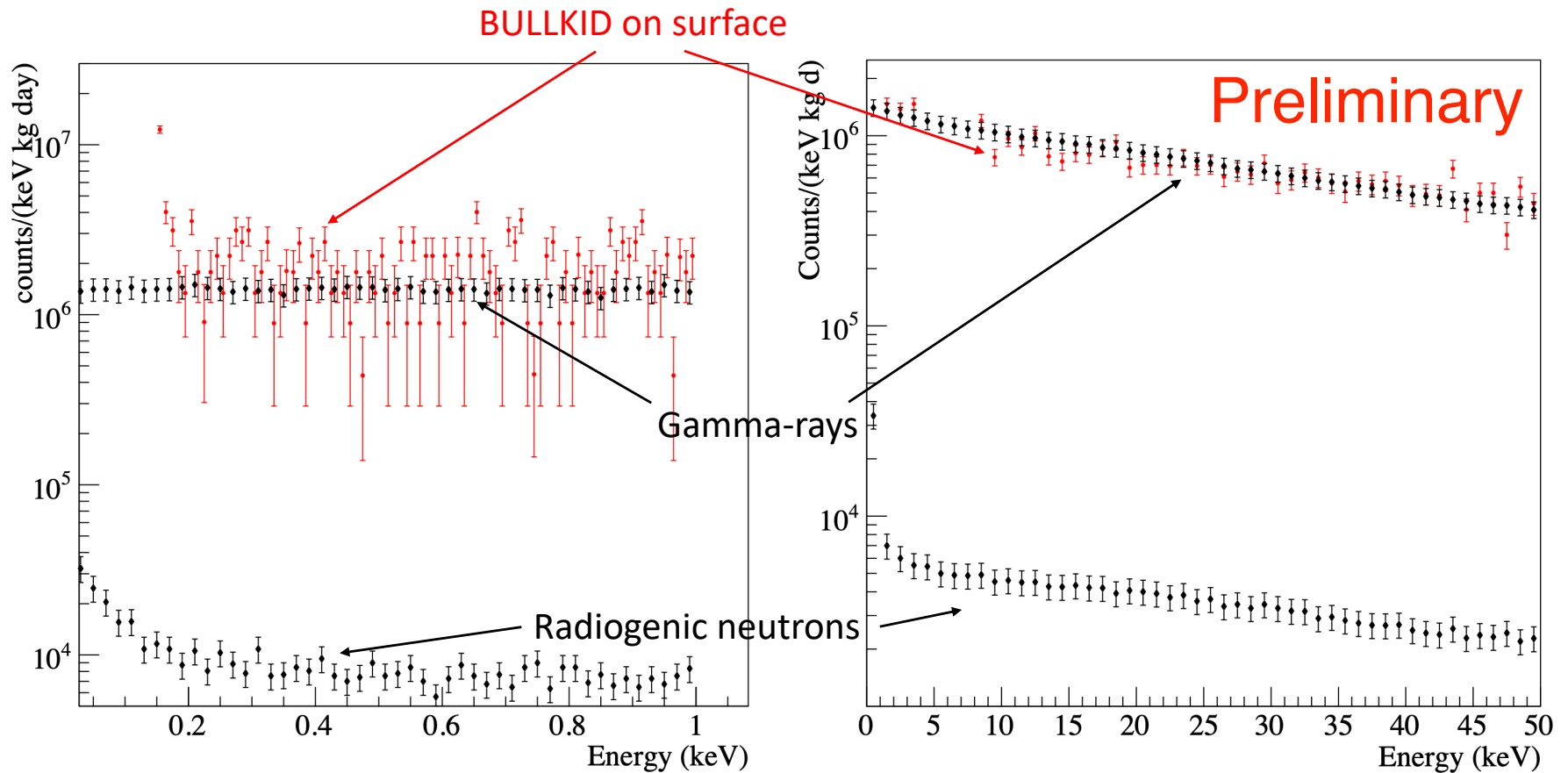
- Backgrounds dominated by environmental  $\gamma$ -rays
- Agreement between data and simulations



- Simulations:  $1.41 \times 10^6$  d.r.u. between 0 and 1 keV
- Very good agreement up to 50 keV

## Subleading backgrounds on surface

- $13.47 \text{ } \gamma/\text{cm}^2/\text{sec}$  from the environment are the main background source
- Neutrons have a contribution two orders of magnitude lower, while cosmogenic backgrounds (muons and neutrons) are below  $10^4$  d.r.u.



## Shielding on surface

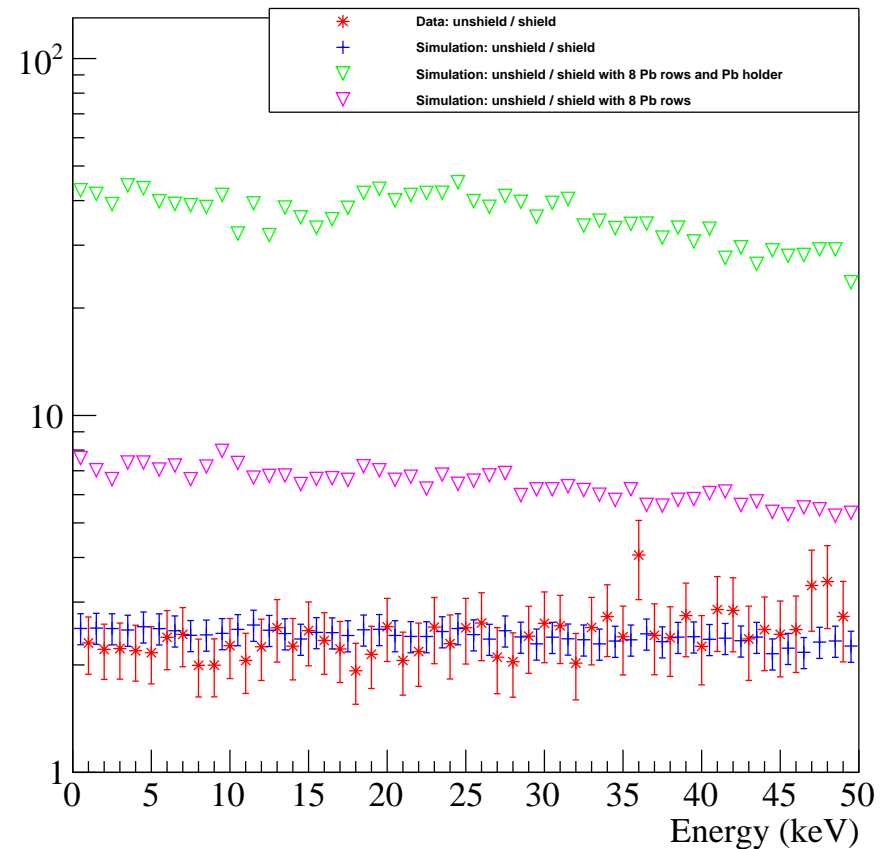
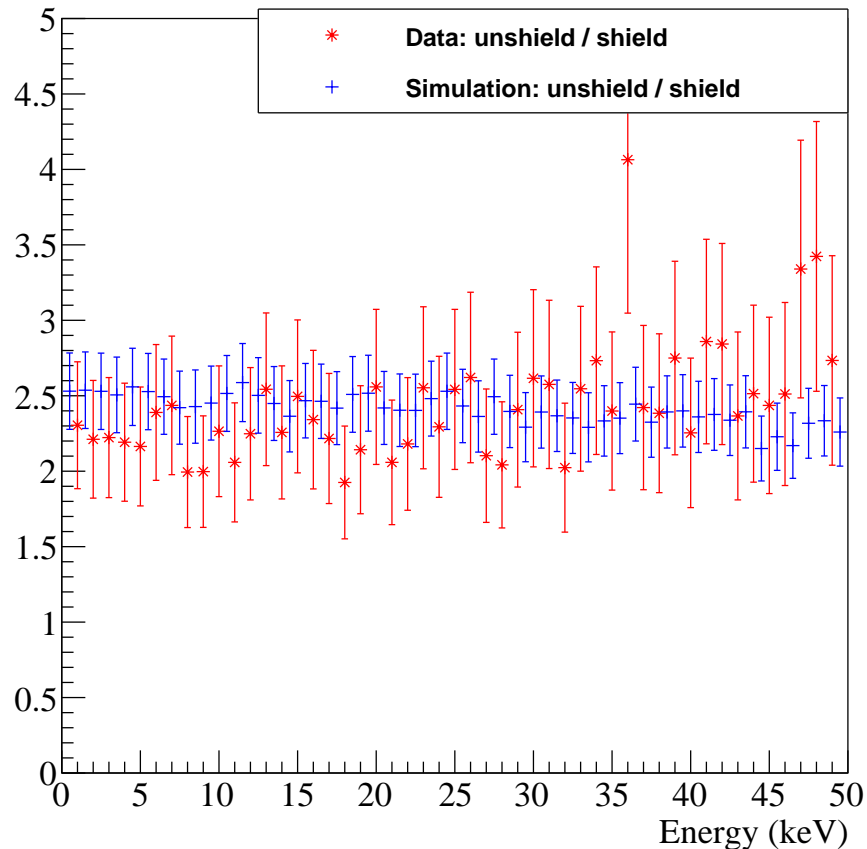
- Moderate  $\gamma$ -rays with Pb to reduce backgrounds by at least a factor of two
- Continue validating the GEANT4 model on surface





# Data vs Simulations with shielding

- Unshielded/shielded ratio in very good agreement
- Further reduction by adding additional rows of Pb



- Replacement of the wafer holder and 8 rows of Pb would reduce gamma contribution by nearly two orders of magnitude

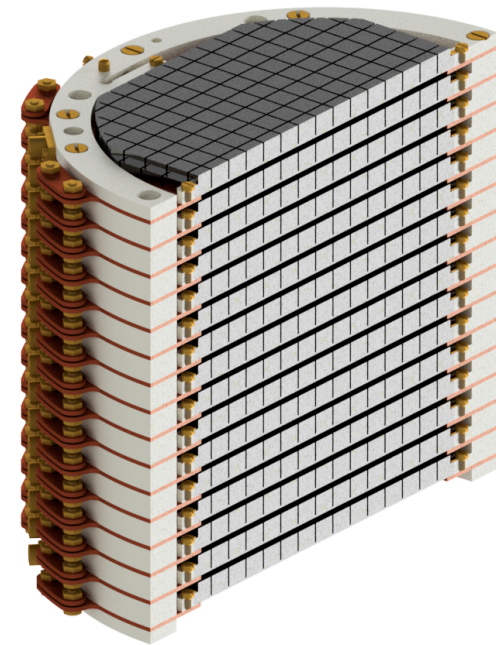
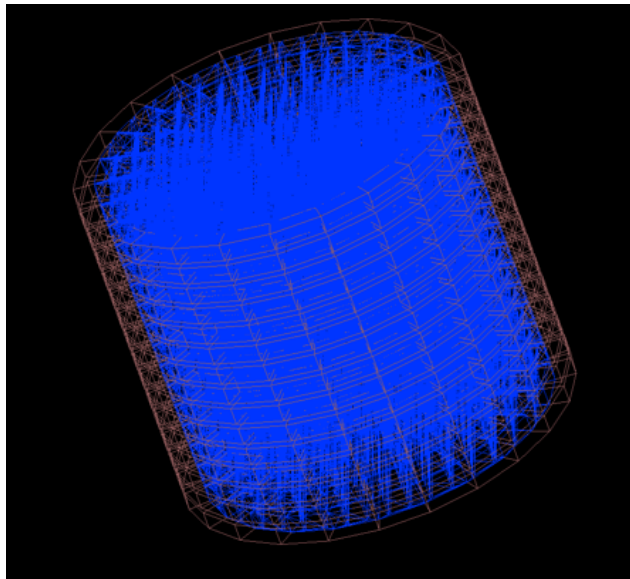
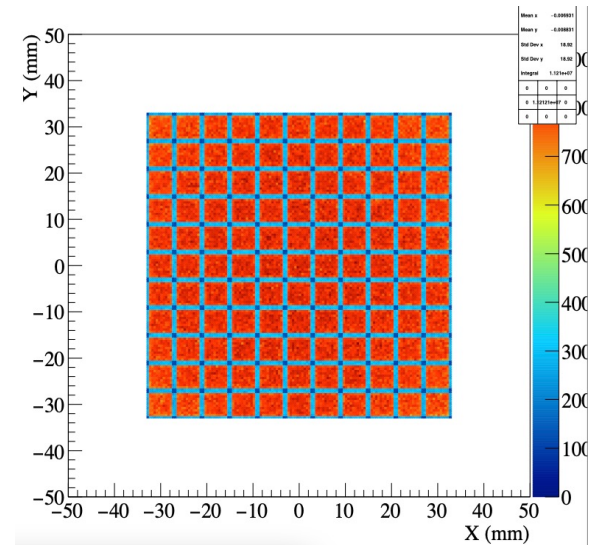
Monte Carlo simulations for the underground experiment at Gran Sasso

## Underground experiment at Gran Sasso

- Active silicon target:  $\sim 600$  gr.
- 16 waffers each 5 cm radius and 5 mm thick

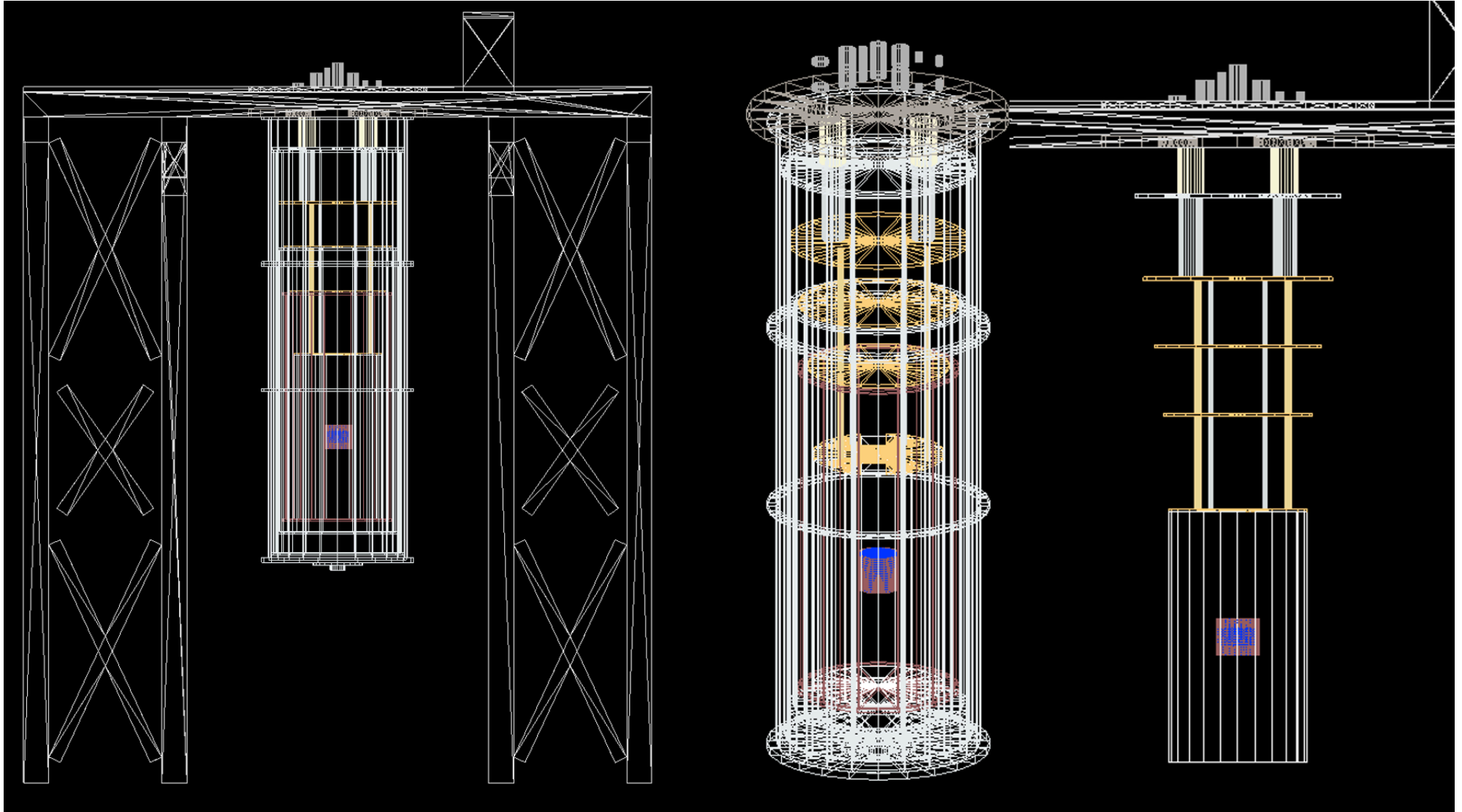
Initial simulation of most relevant backgrounds underground to define shielding configuration:

- Gamma-rays:  $0.729 \text{ } \gamma/\text{cm}^2/\text{sec}$
- Neutrons in several energy ranges: thermal, radiogenic, cosmogenic
- Muon flux:  $3.2 \times 10^{-8} \text{ } \mu/\text{cm}^2/\text{sec}$



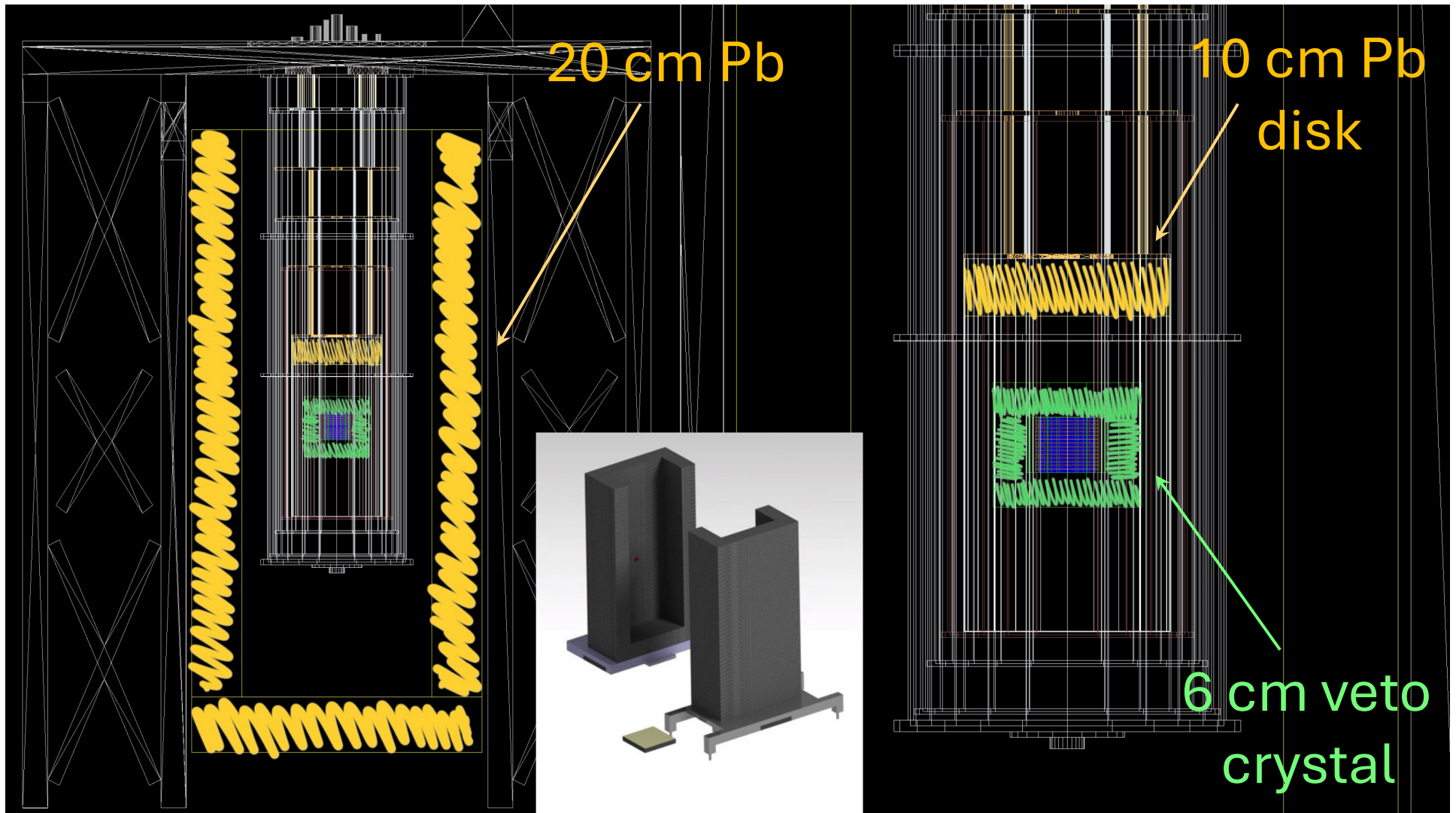
## Model for underground experiment

- Installation planned in the cryo facility at Gran Sasso (Hall B)
- Large volume available inside cryostat for additional shielding and veto



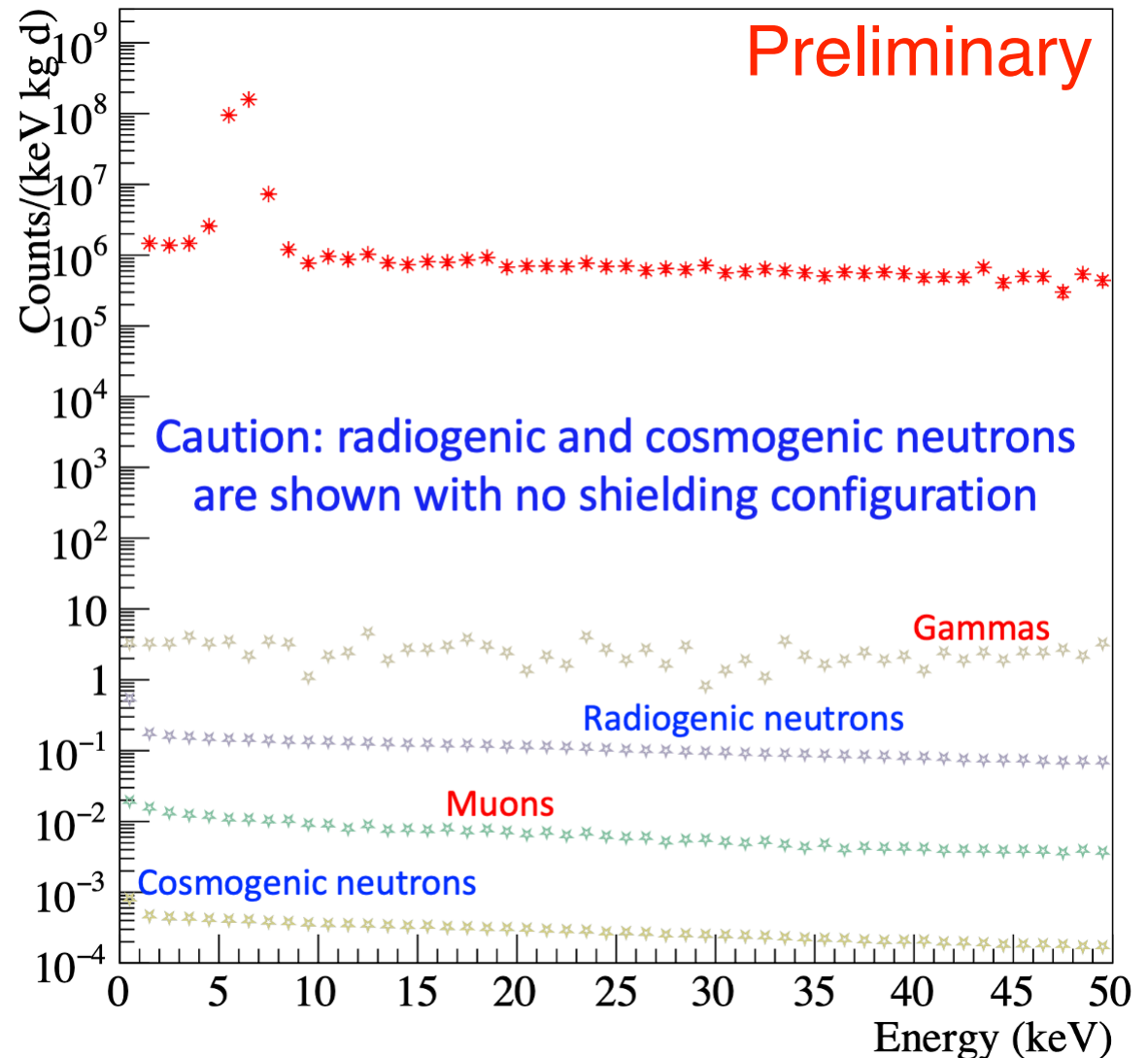
## Shielding for underground experiment: preliminary

- Baseline shielding to reduce external backgrounds: optimization in progress
- Addition of inner veto with high Z material (BGO, GSO)



## Baseline shielding: preliminary

- External gammas reduced at the level of a few d.r.u.
- Radiopurity level required for internal components in progress, including requirement for veto crystal purity
- Neutron moderator to reduce contribution from internals: ( $\alpha, n$ ) reactions, spontaneous fission, n captures





## Final remarks

- BULLKID-DM proved as a promising technology for low-mass dark matter searches
- Possible to reduce backgrounds on surface by two orders of magnitude to explore background excess in phonon experiments
- Underground detector has the potential to be a leading experiment
- Monte Carlo simulations are in good shape with excellent agreement and validation

