



Parton energy loss in inhomogeneous matter

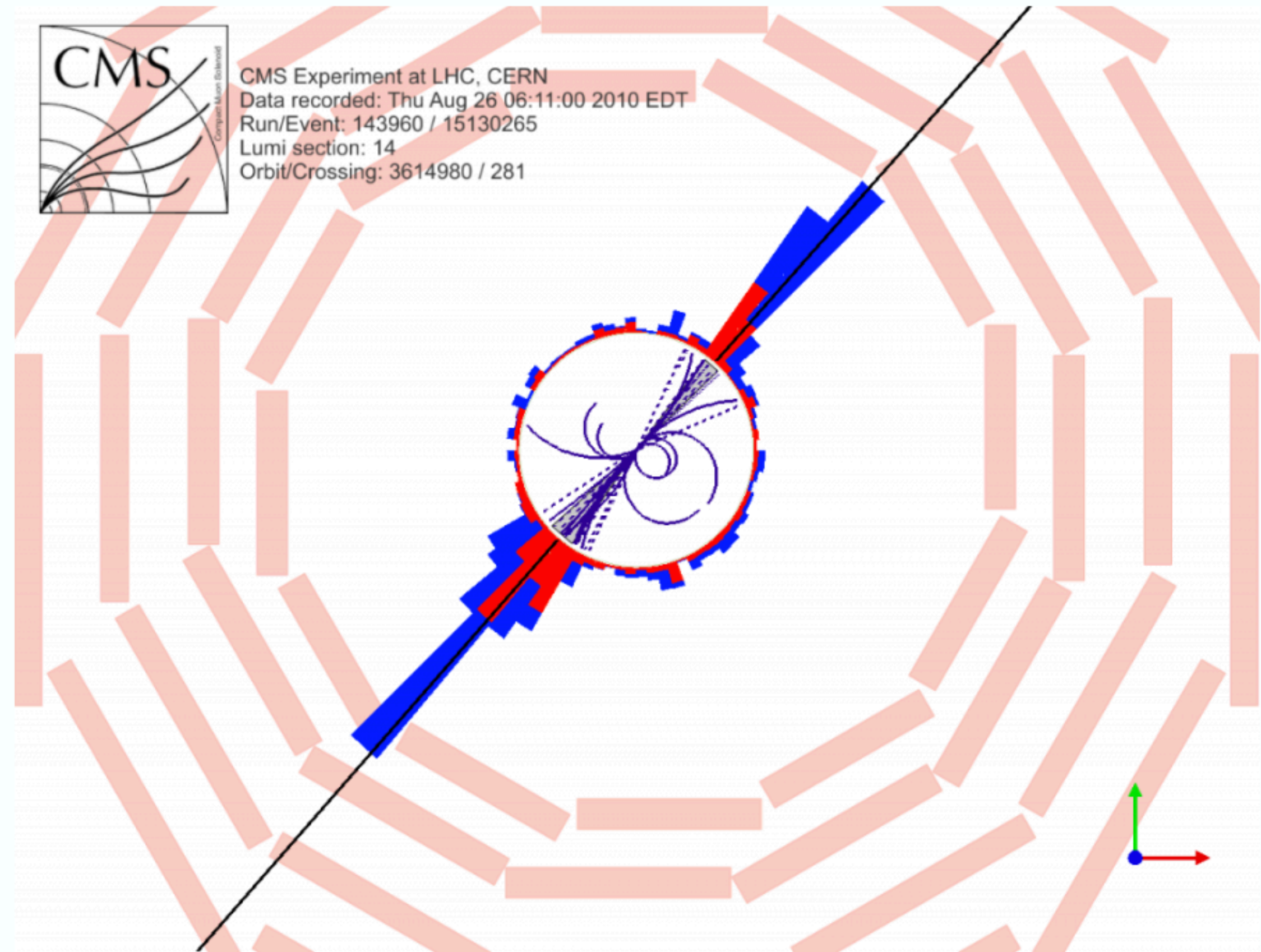
28th July 2023, JETSCAPE Summer School

João Barata, BNL

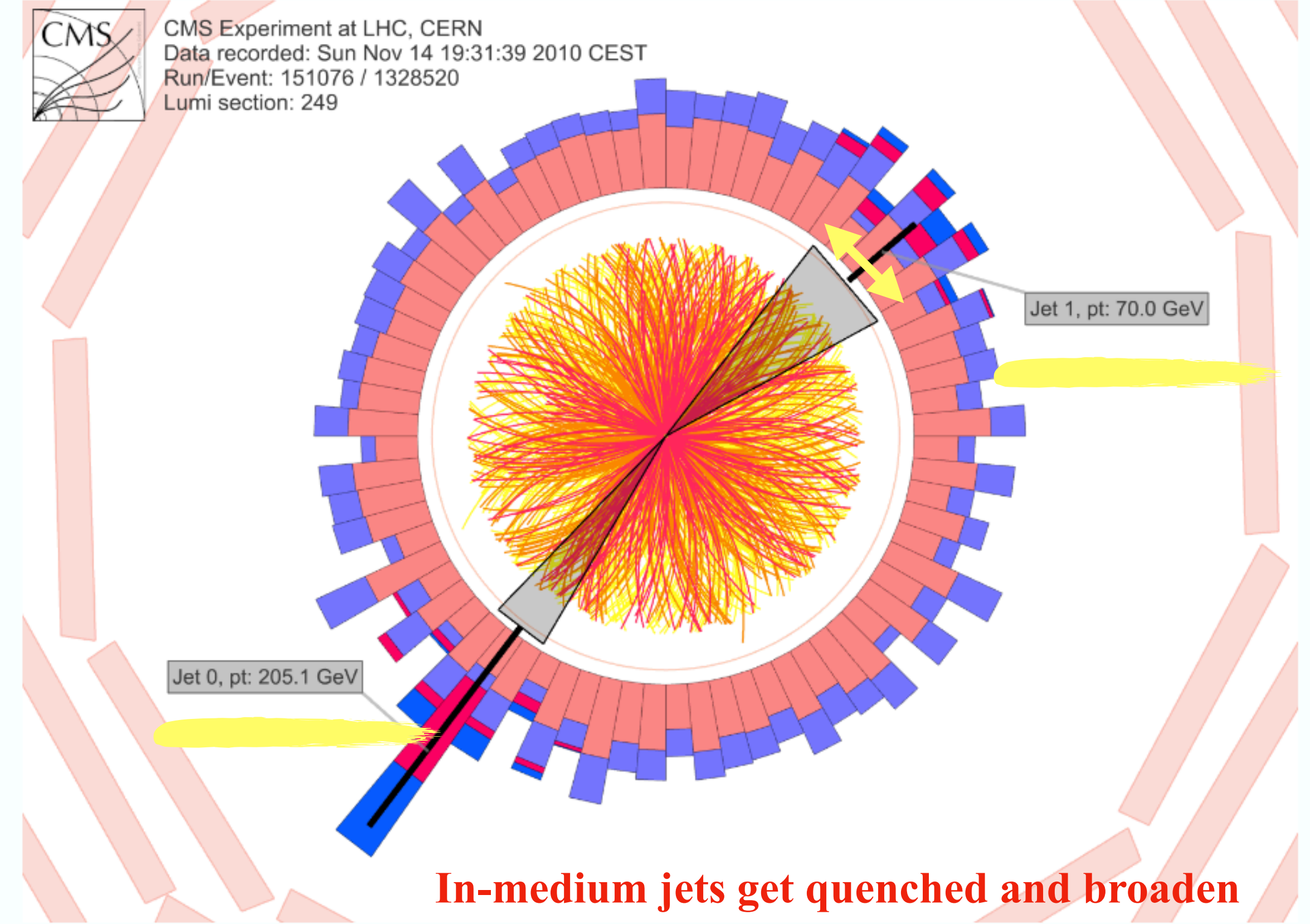
Based on work done with T. Luo, X. Mayo, G. Milhano, A. Sadofyev, C. Salgado, X.-N. Wang

2202.08847 2210.06519 2304.03712 2307(8).xxxx ongoing

Jets in hot plasmas

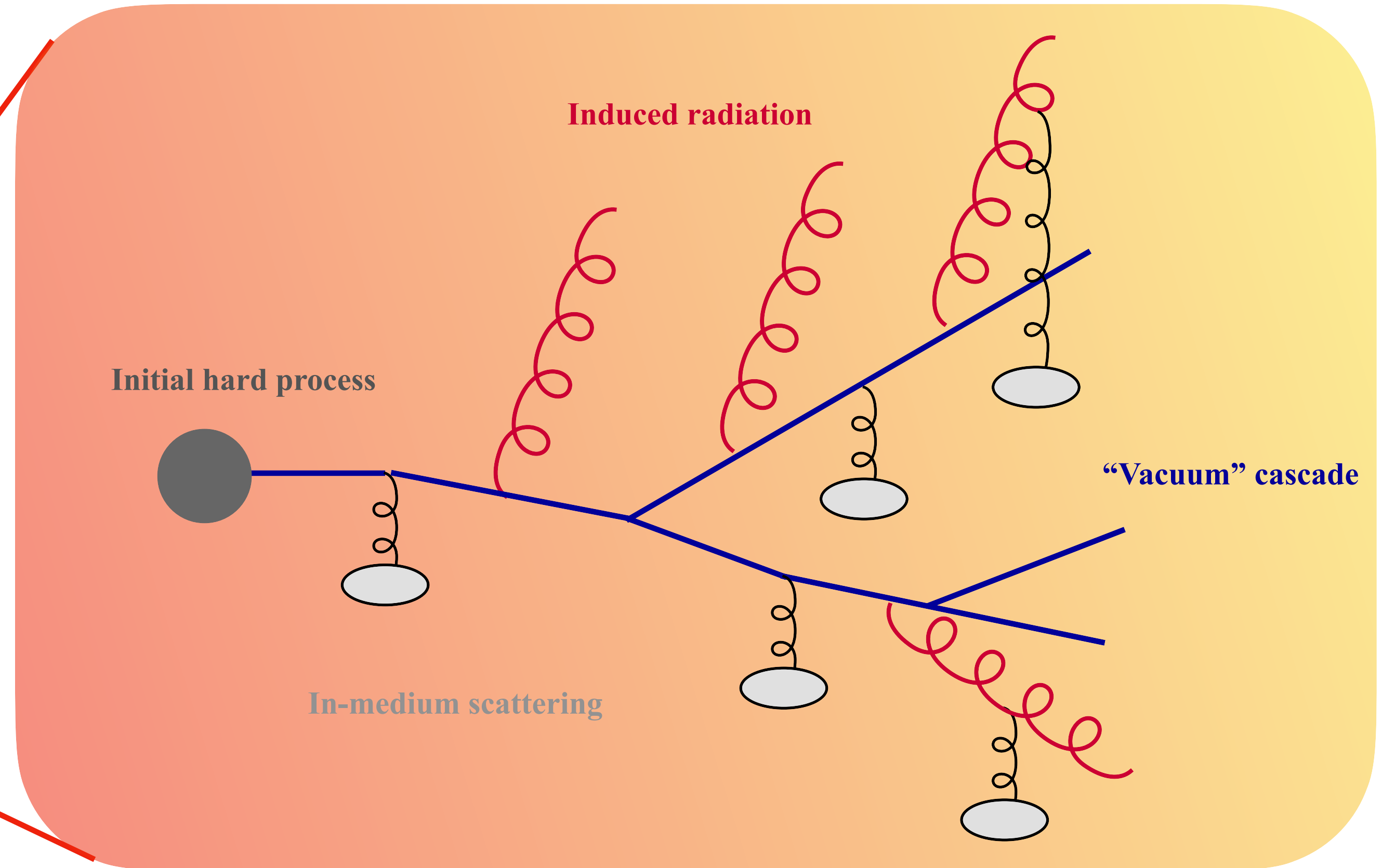
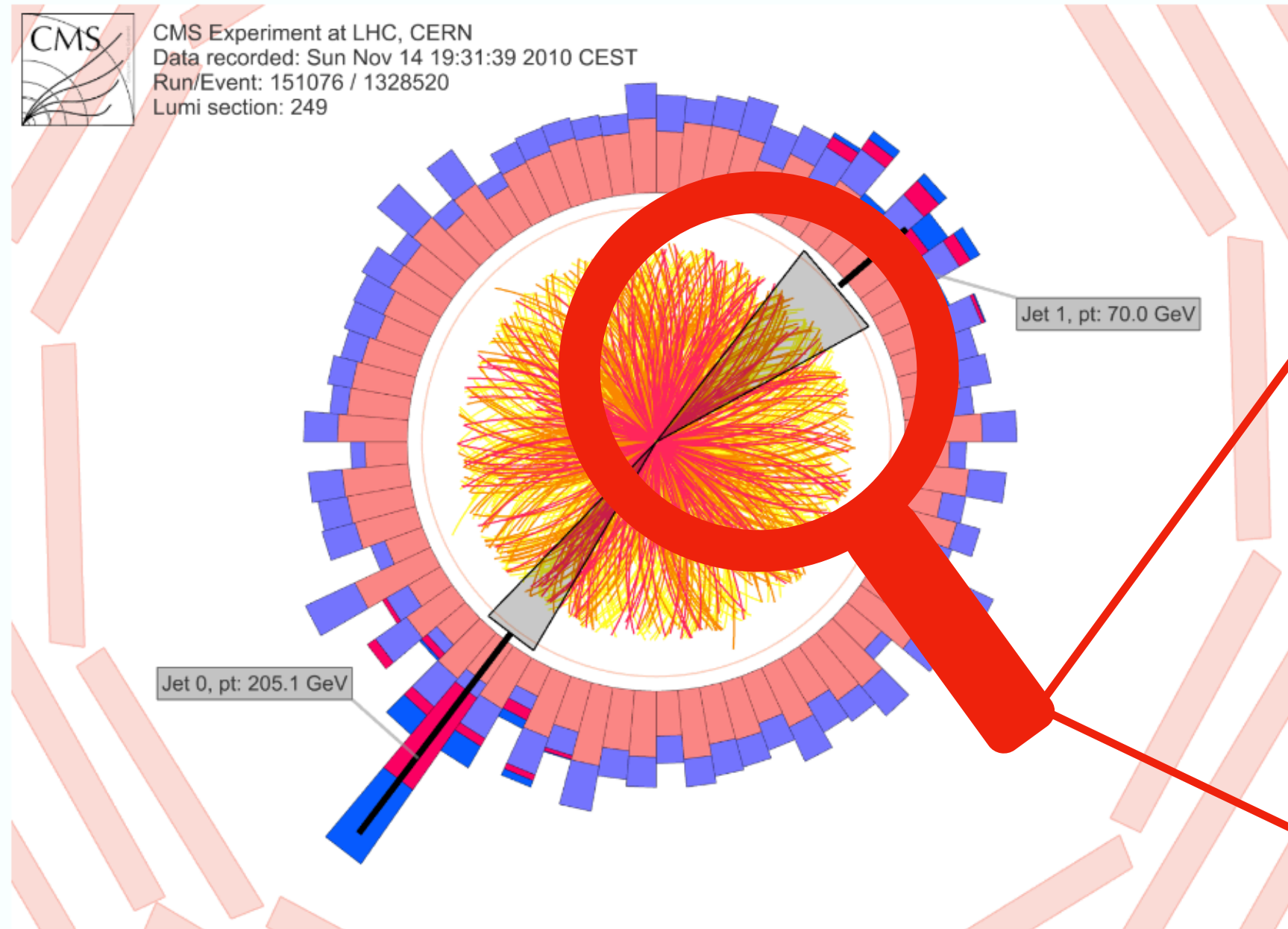


pp dijet event in CMS



PbPb dijet event in CMS

Jets in hot plasmas



- Theoretical task:**
- 1) describe medium induced modifications to jet structure
 - 2) extract information about the medium from jet observables

How do we treat jet evolution in theory?

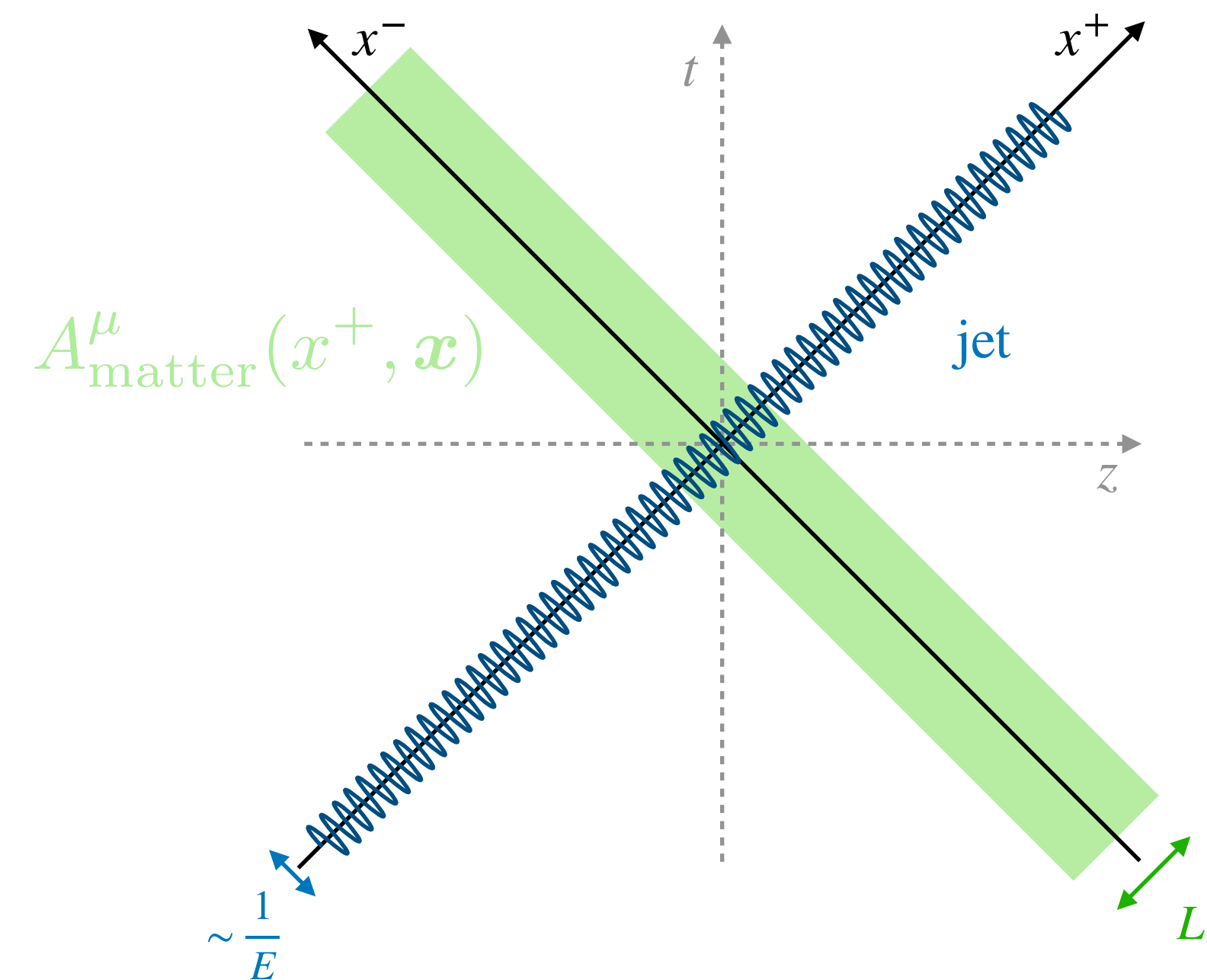
1) Use eikonal expansion, i.e. expansion in inverse powers of jet energy; keep kinetic phases

$$k^- L \sim \frac{k^2}{2E_{\text{jet}}} L$$

2) Matter enters through classical background field; usually assumed: **homogeneous**, infinitely long, static, ...

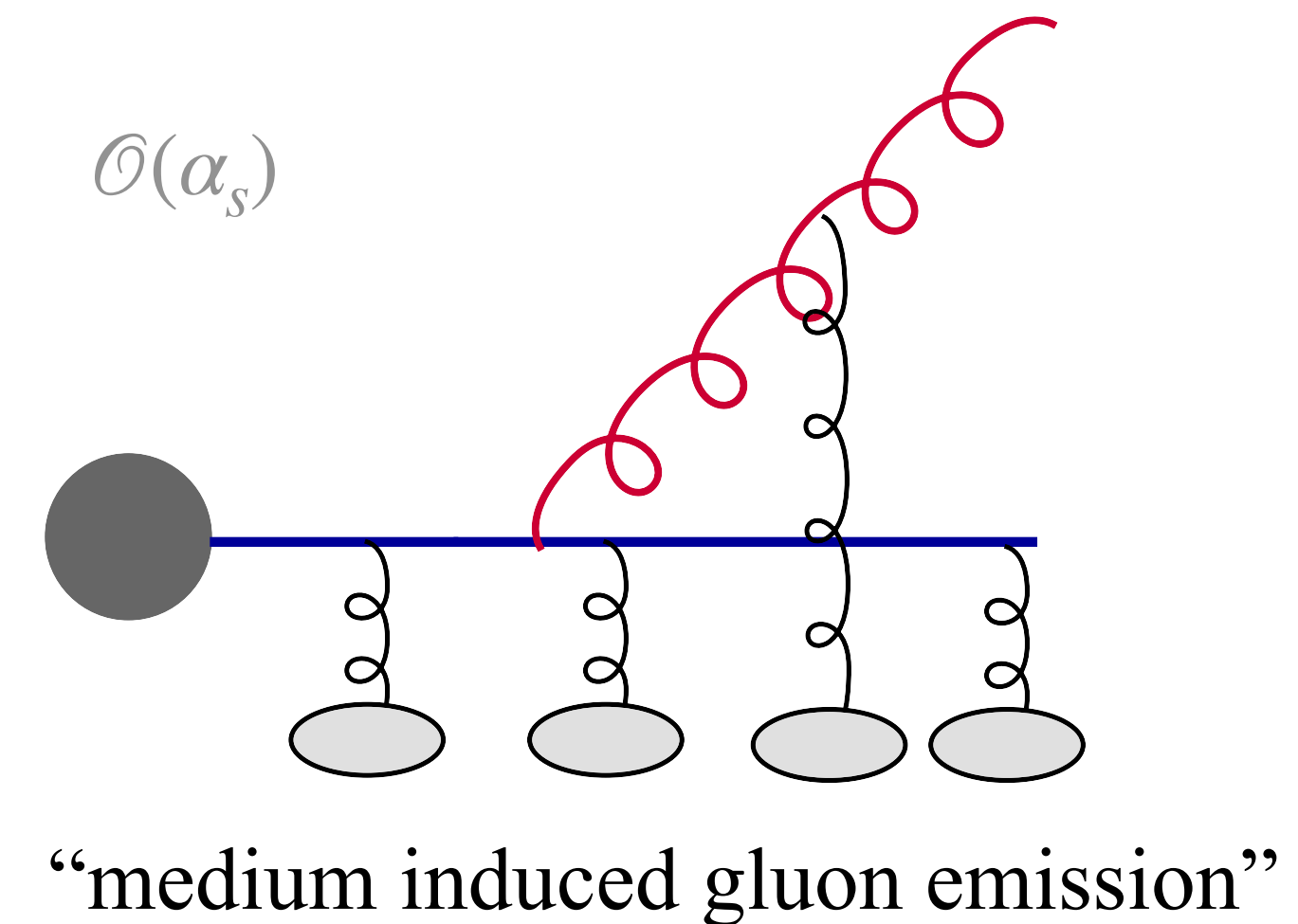
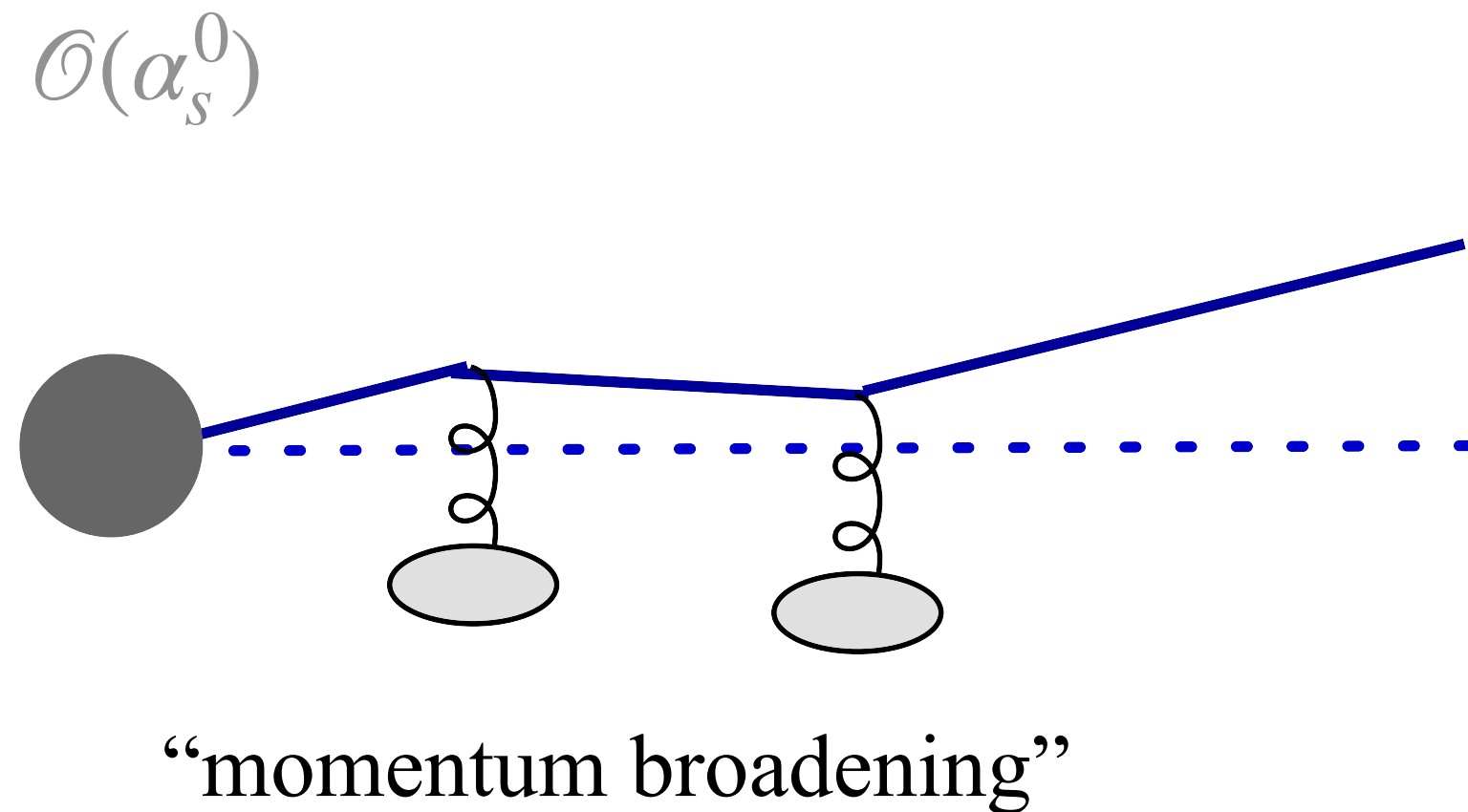
$$A^\mu(x^+, x^-, \mathbf{x}) \approx A_{\text{matter}}^\mu(x^+, \mathbf{x}) + \delta A^\mu(x^+, x^-, \mathbf{x})$$

$$\langle A_{\text{matter}}(x) A_{\text{matter}}(y) \rangle \sim \delta(x - y)$$



How do we treat jet evolution in theory?

Even with these approximations this is a **challenging problem!** Focus on lowest order processes



One can gain analytical insight into the problem: BDMPS-Z, GLV, AMY, Higher-Twist, ...

Talks by L. Apolinário, I. Soudi



Jets decouple more from plasma evolution: less sensitivity to medium properties

Why should we want to go beyond ?

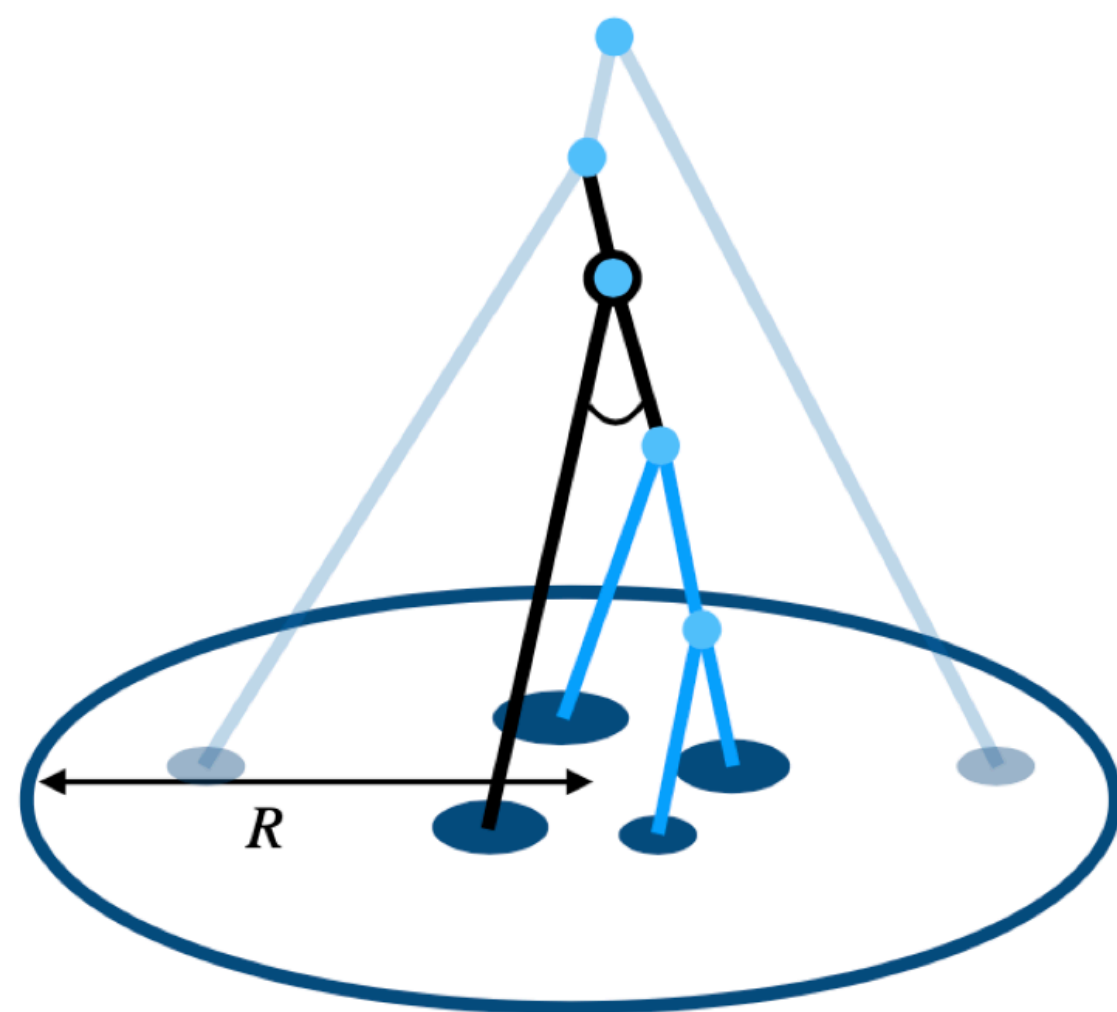
Corrections to this picture will be NLO effect, power correction, subeikonal, ...

logs

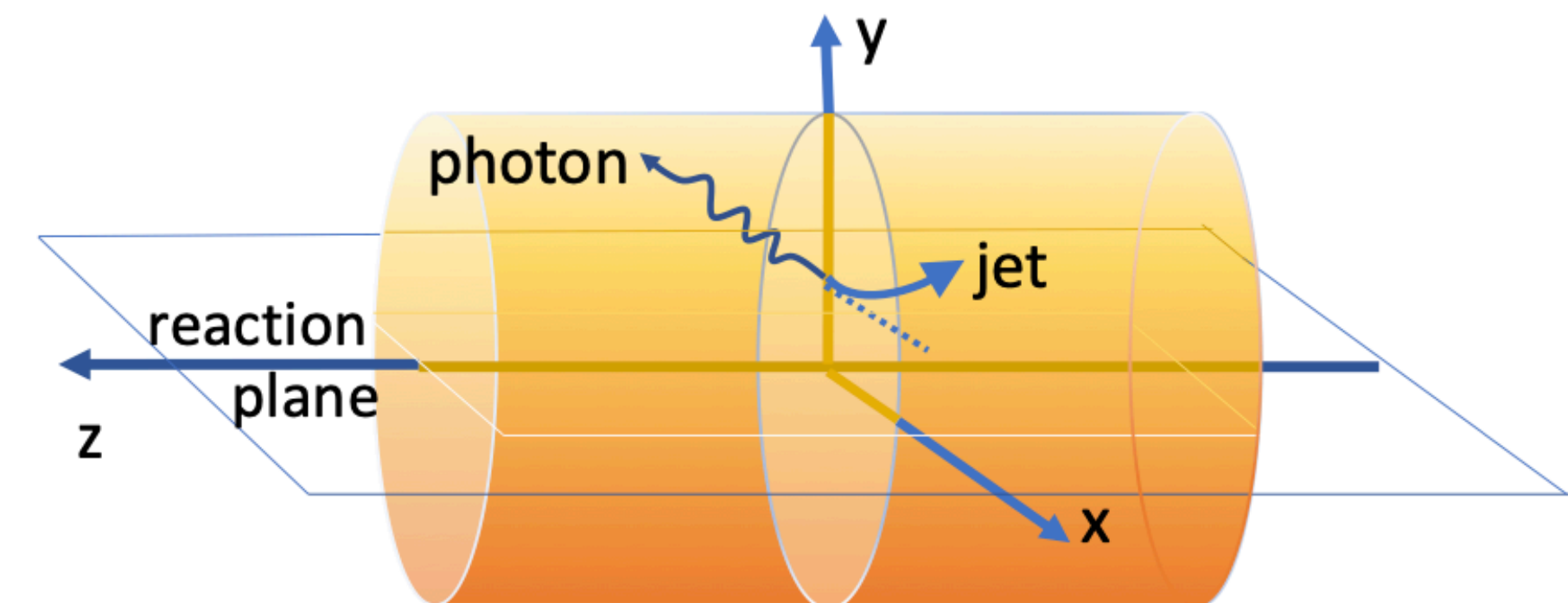
Enhanced by medium parameter

Leading order effect

General move towards more differential access to medium information from jets



Jet sub-structure



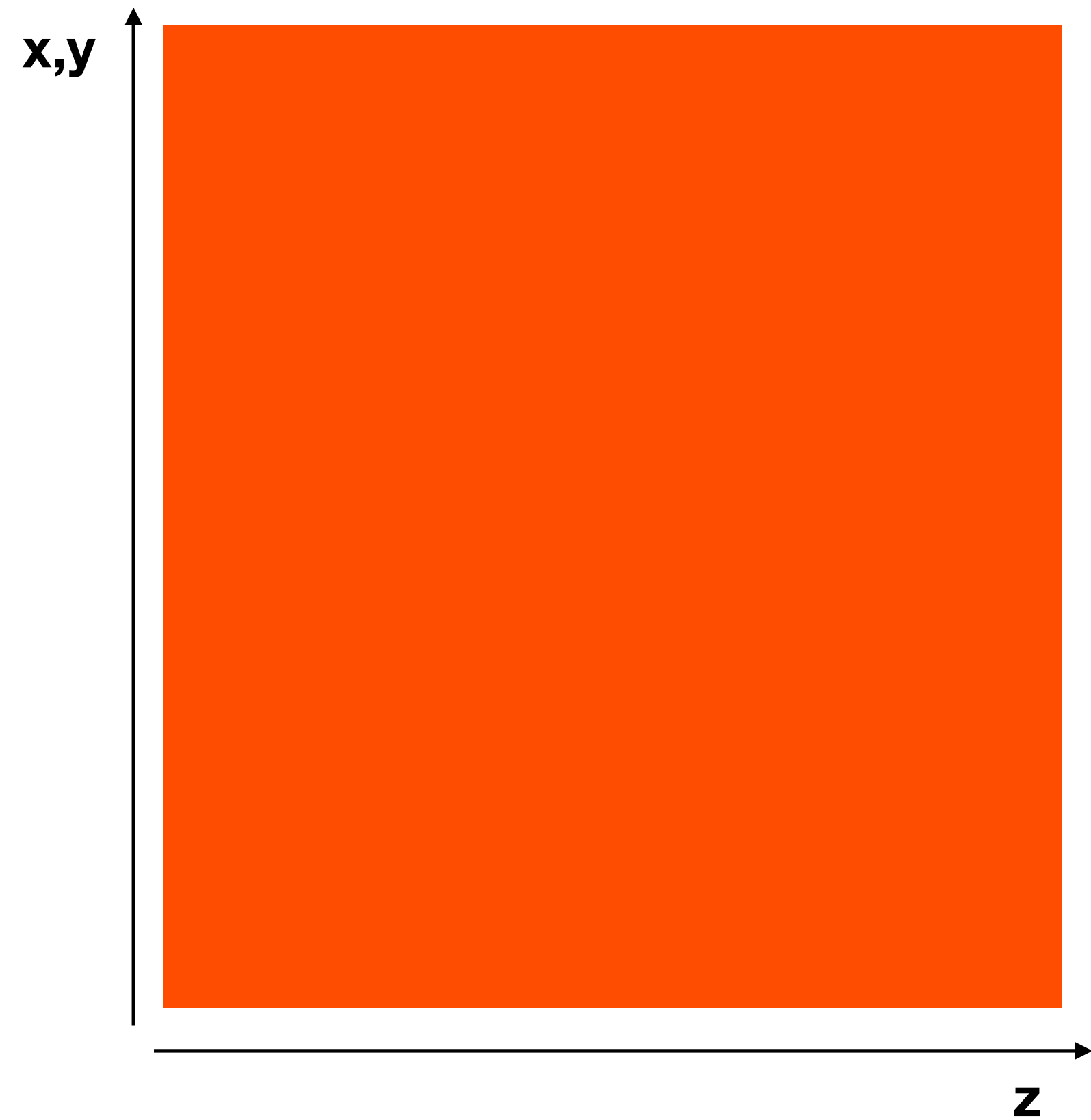
Jet tomography

Going beyond: anisotropic matter

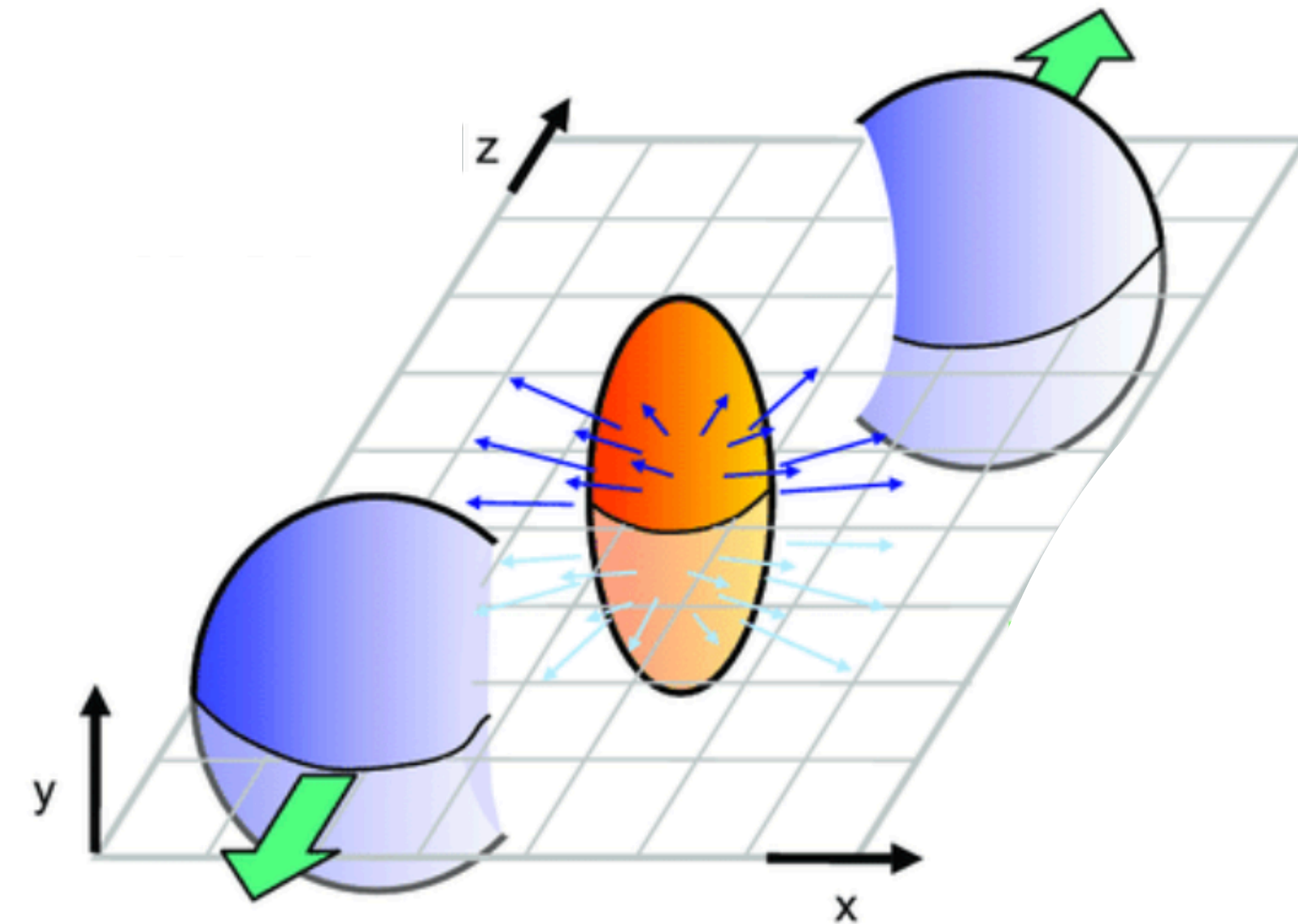
Why?

1

An infinitely long static medium ...

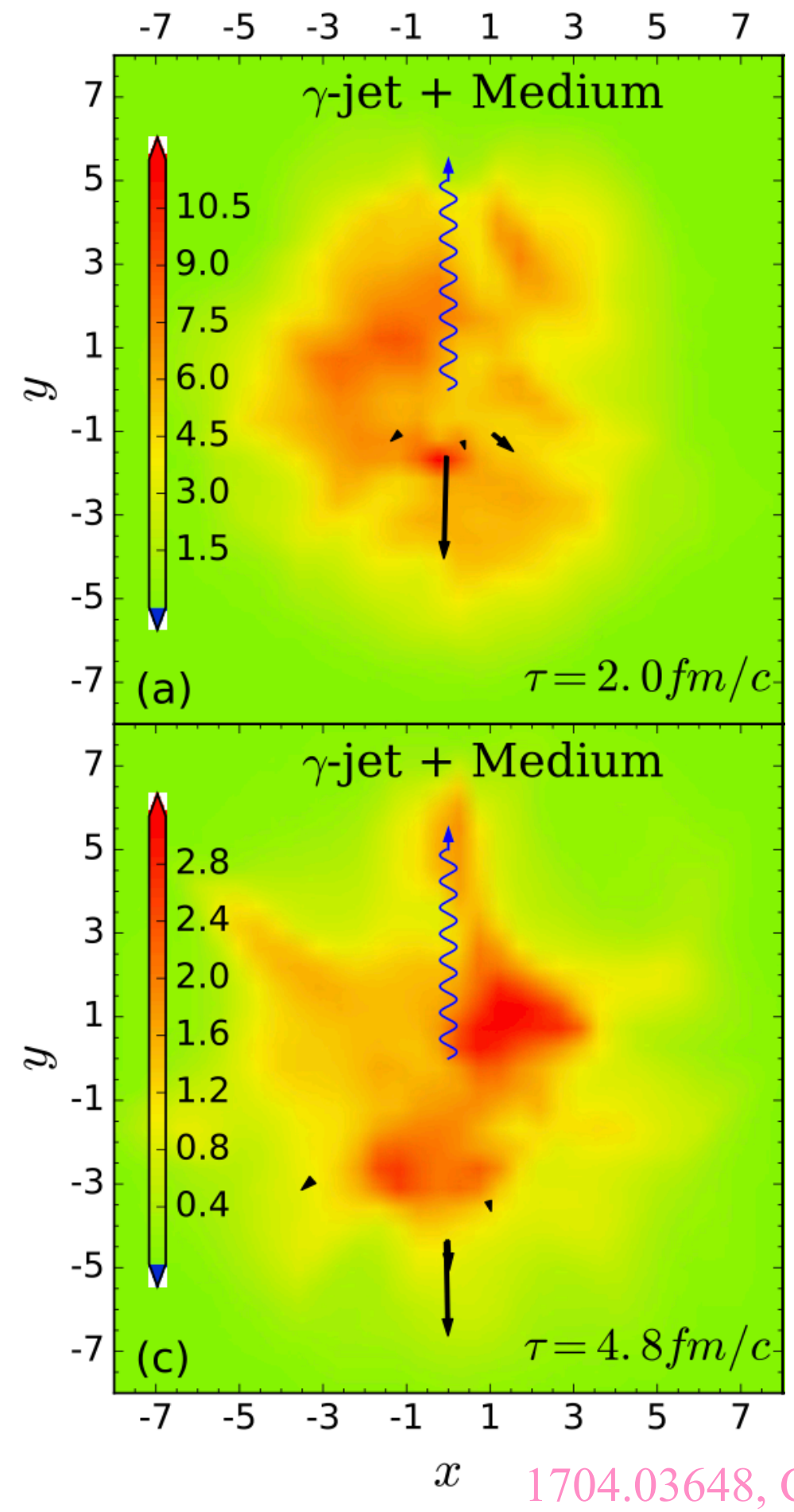


... might not be a good approximation to a real droplet

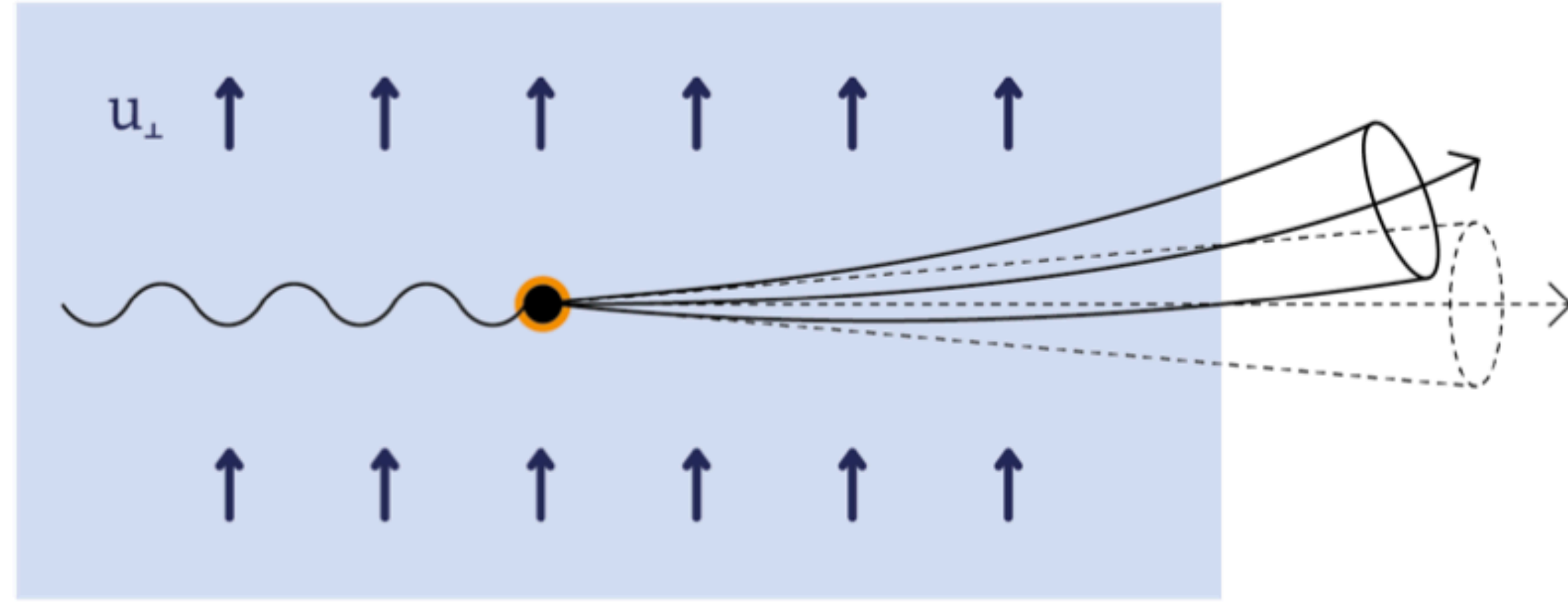


Going beyond: anisotropic matter

2



1704.03648, Chen et al



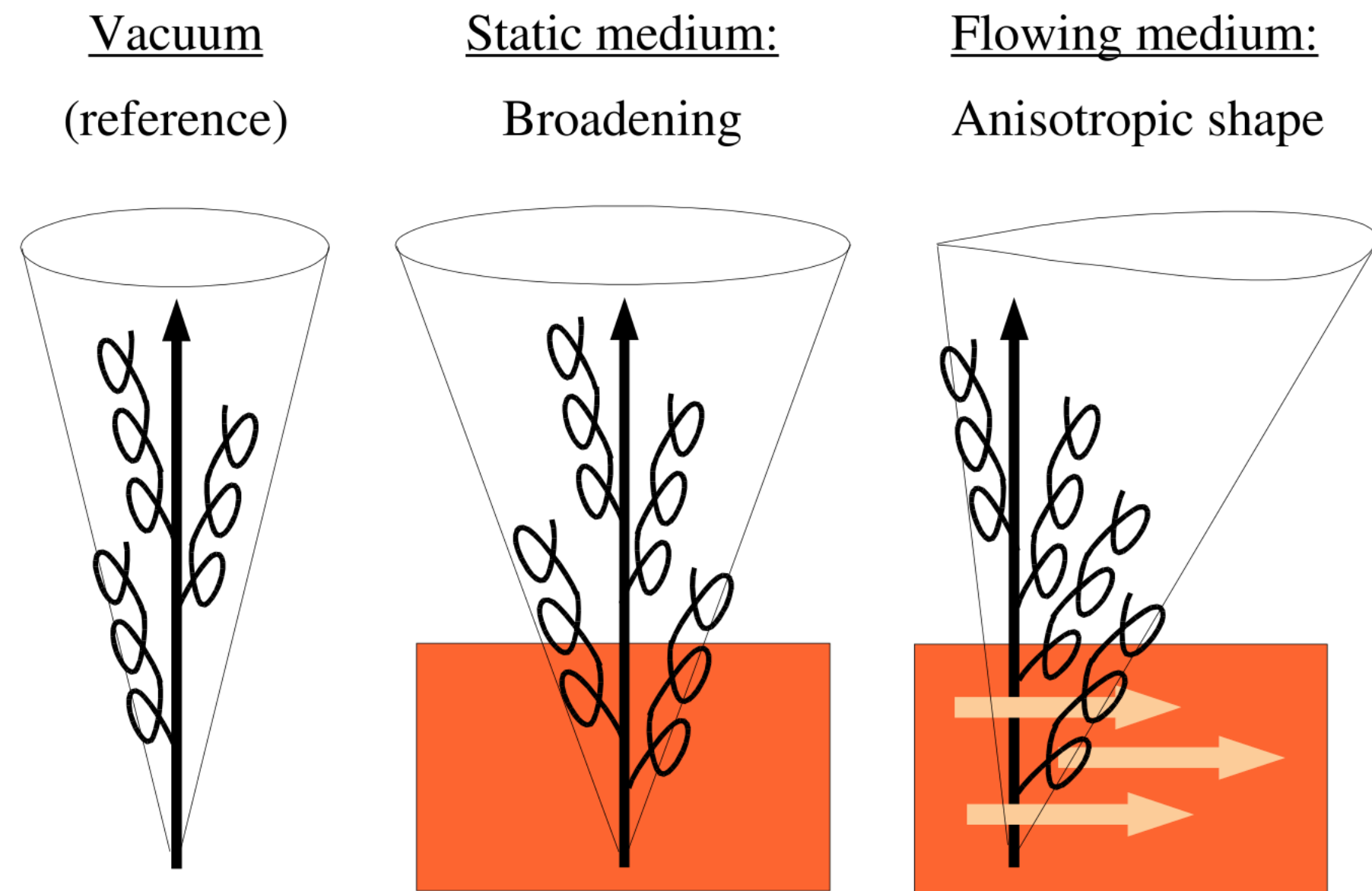
2104.09513, A. Sadofyev, M. Sievert, I. Vitev

Medium response to jet

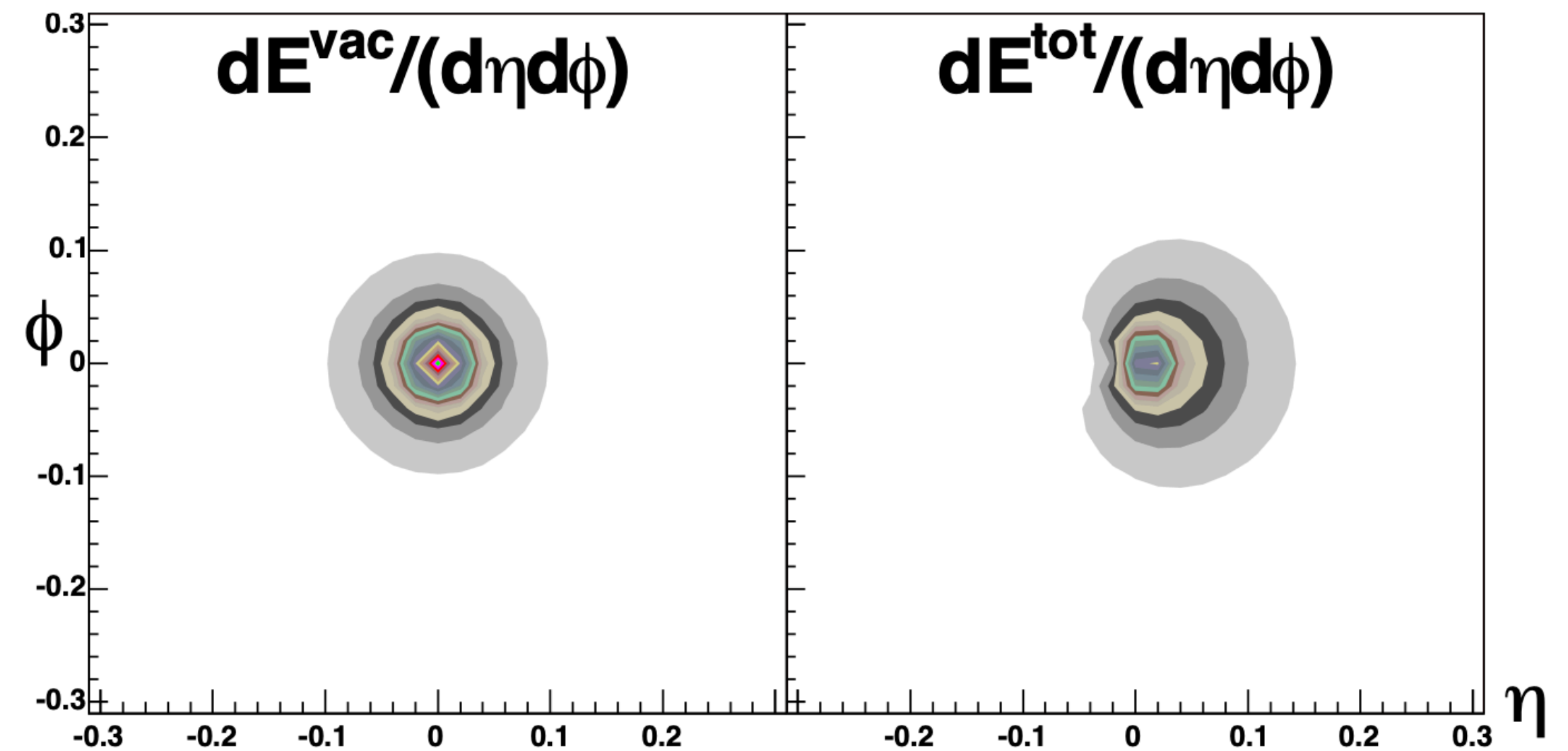
Jet response to medium

Going beyond: anisotropic matter

3 Previous studies considered similar effects, but at pheno level



2004, Armesto, Salgado, Wiedemann

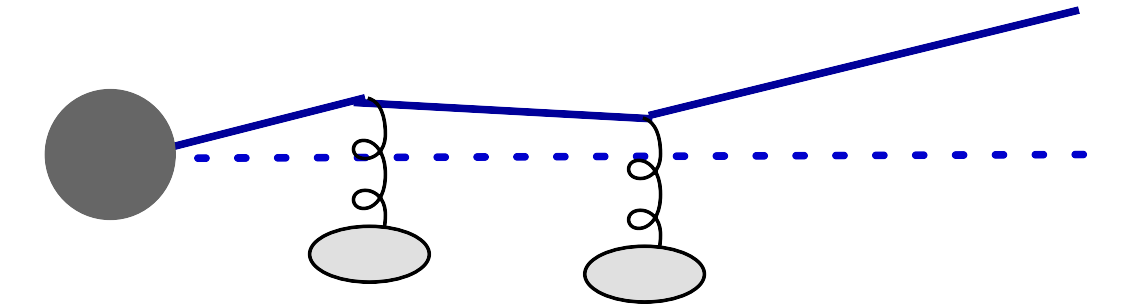


$$|a(\mathbf{q})|^2 = \frac{\mu^2}{\pi [(\mathbf{q} - \mathbf{q}_0)^2 + \mu^2]^2}$$

← Debye mass

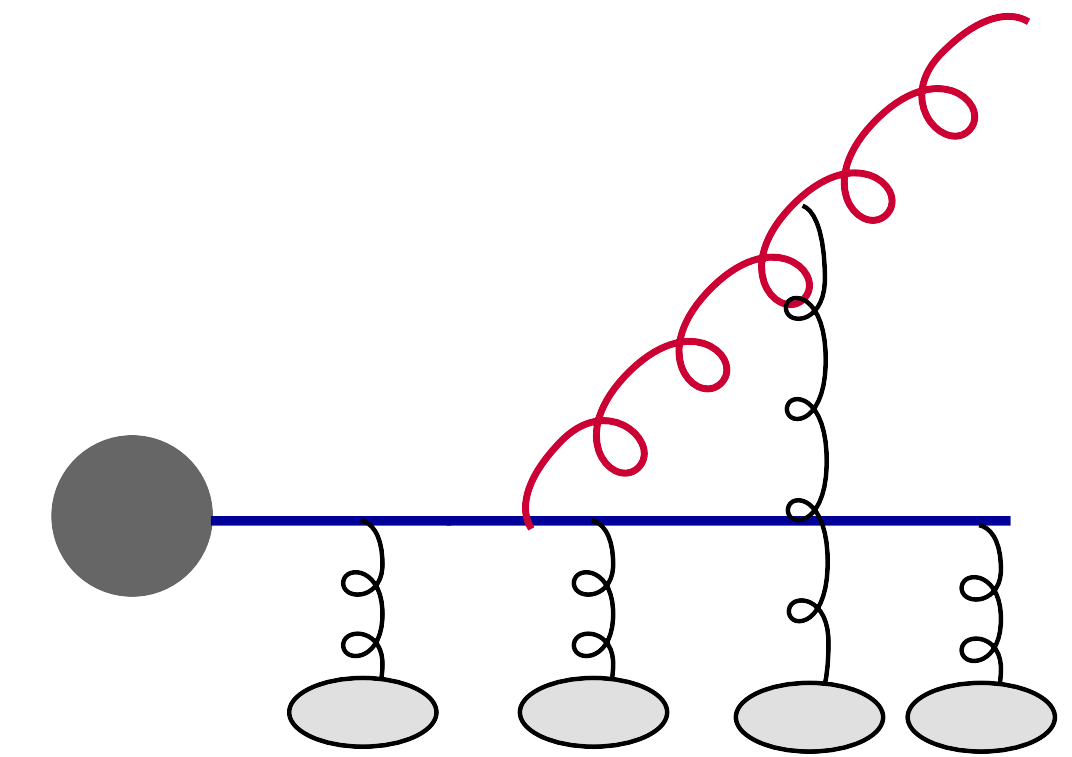
① Momentum broadening in dense anisotropic media

2202.08847 2210.06519



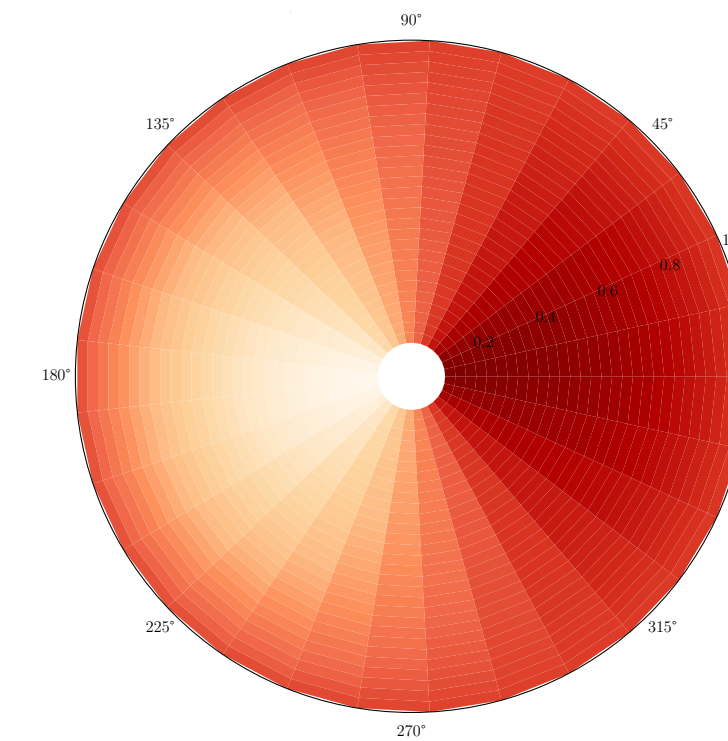
② Radiative energy loss in dense anisotropic media

2304.03712

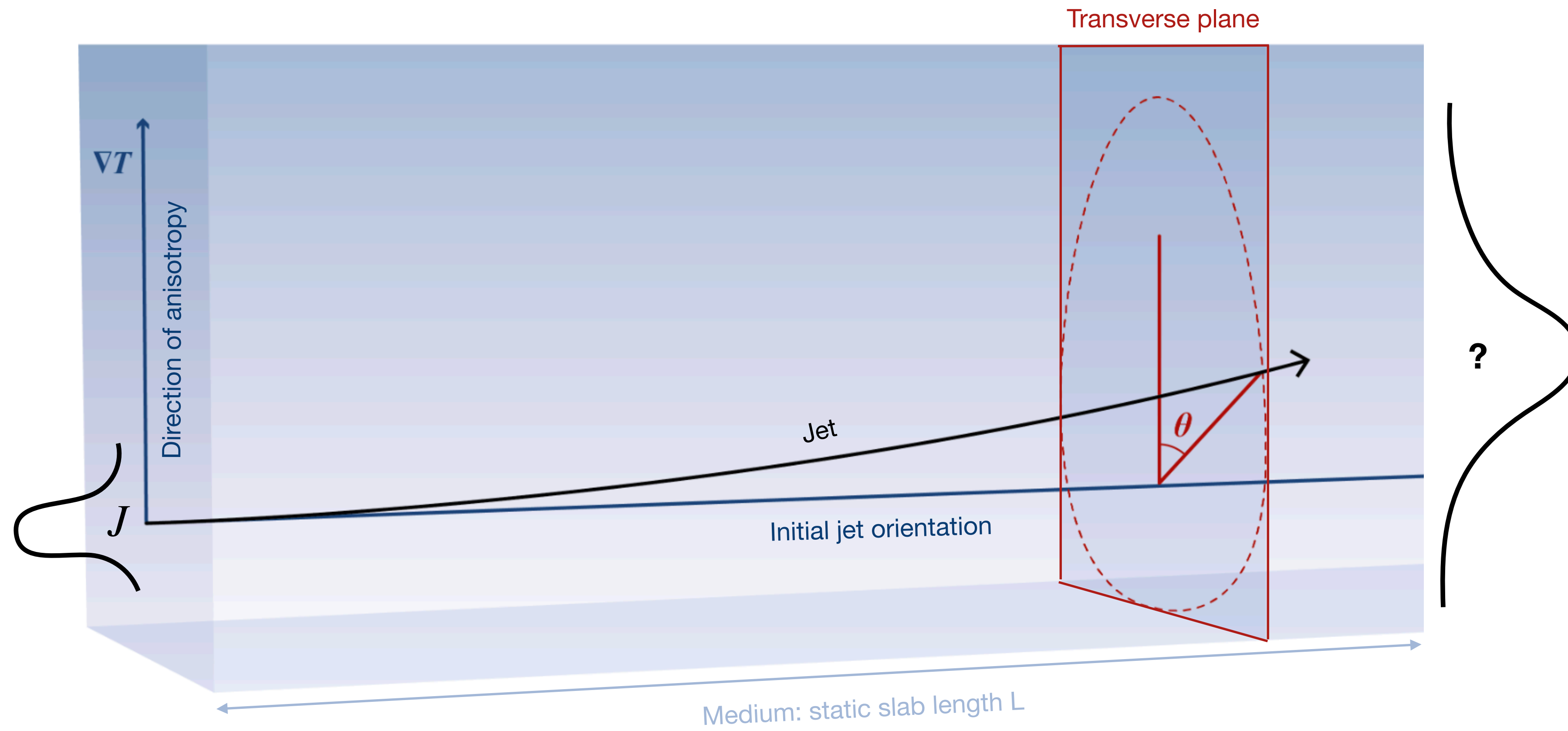


③ Jet observables in inhomogeneous matter

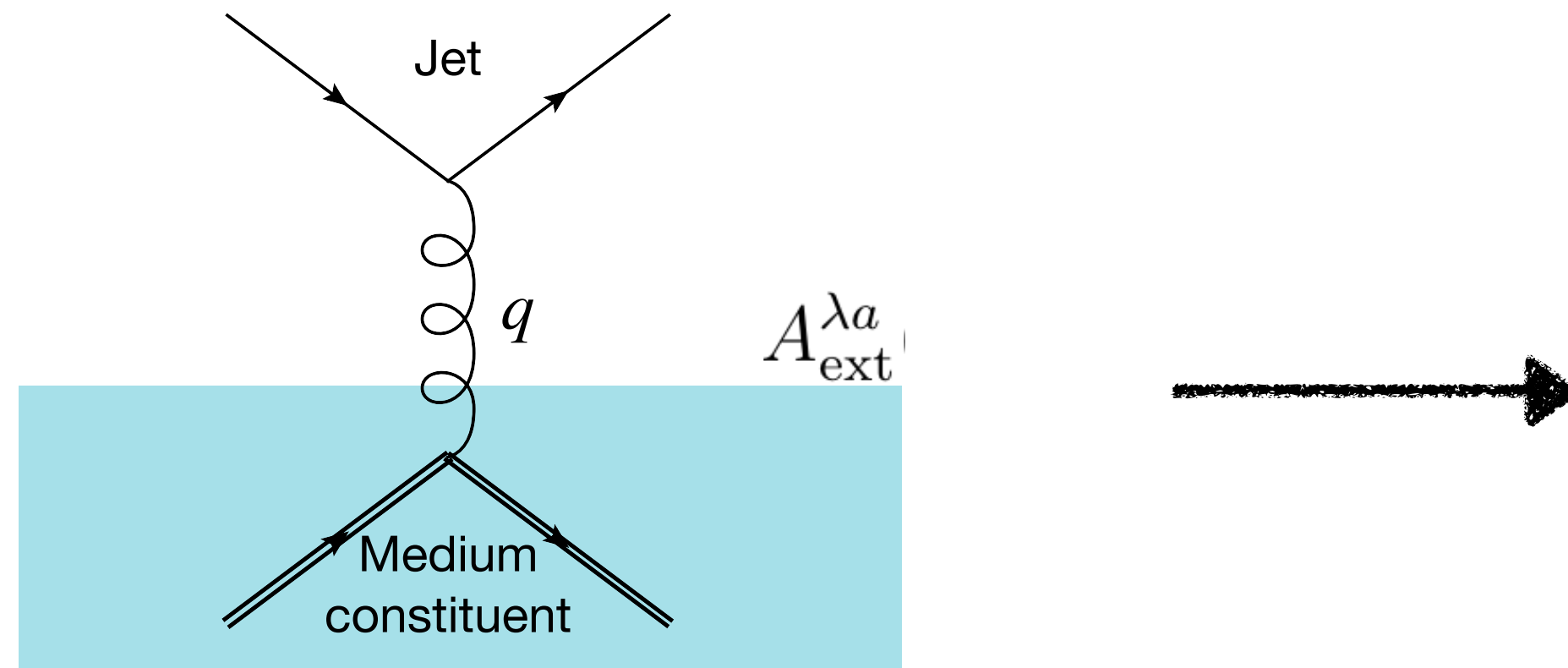
2308.xxxx



Momentum broadening in anisotropic media



The medium is described by a classical field



Model dependent elastic scattering potential for source j

No energy transfer in each scattering: transverse t-channel gluon exchanges only

$$gA_{\text{ext}}^{\lambda a}(q) = -(2\pi) g^{\lambda 0} \sum_i e^{-i(\mathbf{q} \cdot \mathbf{x}_j + q_z z_j)} \underline{t_j^a} \underline{v_j(q)} \delta(q^0)$$

where we use the GW model

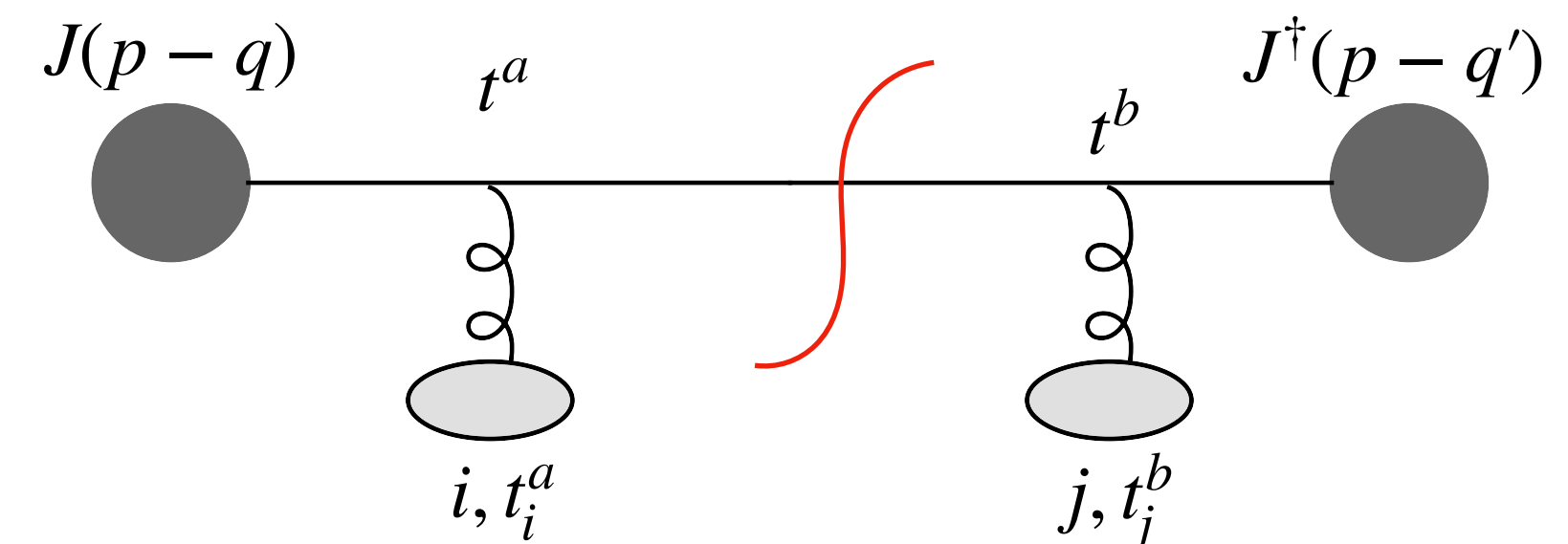
$$v_i(q) \equiv \frac{-g^2}{-q_0^2 + \mathbf{q}^2 + q_z^2 + \mu_i^2 - i\epsilon}$$

Medium statistics follow from 2-gluon approximation

$$\langle \underline{t_i^a t_j^b} \rangle = \frac{1}{d_{\text{tgt}}} \text{tr} (t_i^a t_j^b) = \frac{1}{2C_{\bar{R}}} \underline{\delta_{ij} \delta^{ab}}$$

Only non-trivial correlator

Probe interacts with the same scattering center in amplitude and conjugate amplitude



Momentum broadening in anisotropic media

1 Compute all diagrams up to $2N$ field insertions



2 For each N , square and average the respective diagrams

$$\langle |M|^2 \rangle = \underbrace{\langle |M_0|^2 \rangle}_{N=0} + \underbrace{\langle |M_1|^2 \rangle + \langle M_2 M_0^* \rangle + \langle M_0 M_2^* \rangle}_{N=1} + \underbrace{\langle |M_2|^2 \rangle + \langle M_3 M_1^* \rangle + \langle M_1 M_3^* \rangle + \langle M_4 M_0^* \rangle + \langle M_0 M_4^* \rangle}_{N=2} + \dots$$

The averaging is performed by taking the limit of continuous distribution in the medium

$$\sum_i f_i = \int d^2 \mathbf{x} dz \underbrace{\rho(\mathbf{x}, z)}_{\text{Density of scattering centers}} \underbrace{f(\mathbf{x}, z)}_{\substack{\rho(\mathbf{x}, z) \\ \mu^2(\mathbf{x}, z)}} \xrightarrow{\nabla T = 0} \int d^2 \mathbf{x}_n e^{-i(\mathbf{q}_n \pm \bar{\mathbf{q}}_n) \cdot \mathbf{x}_n} = (2\pi)^2 \delta^{(2)}(\mathbf{q}_n \pm \bar{\mathbf{q}}_n)$$

3 Resum the Opacity Series

A detailed derivation shows that the square amplitude for $2N$ insertions has the form

$$\langle |M|^2 \rangle^{(N)} = \prod_{n=1}^N \left[(-1) \int_0^{z_{n+1}} dz_n \int \frac{d^2 \mathbf{q}_n}{(2\pi)^2} \mathcal{V}(\mathbf{q}_n, z_n) \right] |J(E, \mathbf{p}_{in})|^2$$

where we identify the effective scattering potential

$$\mathcal{V}(\mathbf{q}, z) \equiv -\mathcal{C} \rho(z) \left(\underbrace{|v(\mathbf{q}^2)|^2}_{\text{diagram}} - \delta^{(2)}(\mathbf{q}) \underbrace{\int d^2 \mathbf{l} |v(\mathbf{l}^2)|^2}_{\text{diagram}} \right)$$

The resummation in this case gives:

$$\frac{d\mathcal{N}}{d^2 \mathbf{x} dE} = \sum_{N=0}^{\infty} \int \frac{d^2 \mathbf{p} d^2 \mathbf{r}}{(2\pi)^2} e^{i\mathbf{p} \cdot (\mathbf{x} - \mathbf{r})} \frac{(-1)^N [\mathcal{V}(\mathbf{r})L]^N}{N!} \frac{d\mathcal{N}^{(0)}}{d^2 \mathbf{r} dE} = e^{-\mathcal{V}(\mathbf{x})L} \frac{d\mathcal{N}^{(0)}}{d^2 \mathbf{x} dE}$$

When averaging in 2 we used

$$\sum_i f_i = \int d^2\mathbf{x} dz \rho(\mathbf{x}, z) f(\mathbf{x}, z) \xrightarrow{\nabla T = 0} \int \underline{d^2\mathbf{x}_n e^{-i(\mathbf{q}_n \pm \bar{\mathbf{q}}_n) \cdot \mathbf{x}_n}} = (2\pi)^2 \delta^{(2)}(\mathbf{q}_n \pm \bar{\mathbf{q}}_n)$$

For **anisotropic media** this no longer holds; **hard to tackle** in general

We perform a gradient expansion for the 2 relevant parameters: ρ and μ 2104.09513, A. Sadofyev, M. Sievert, I. Vitev

$$\rho(\mathbf{x}, z) \approx \rho(z) + \nabla \rho(z) \cdot \mathbf{x} \quad \mu^2(\mathbf{x}, z) \approx \mu^2(z) + \nabla \mu^2(z) \cdot \mathbf{x}$$

So that when averaging instead of a momentum space Dirac delta one obtains

$$\int \underline{d^2\mathbf{x}_n x_n^\alpha e^{-i(\mathbf{q}_n \pm \bar{\mathbf{q}}_n) \cdot \mathbf{x}_n}} = i (2\pi)^2 \frac{\partial}{\partial (q_n \pm \bar{q}_n)_\alpha} \delta^{(2)}(\mathbf{q}_n \pm \bar{\mathbf{q}}_n)$$

Proceeding as in 3 we find that

$$\mathcal{V}'(\mathbf{x}) \equiv \frac{\partial}{\partial \mu^2} \mathcal{V}(\mathbf{x})$$

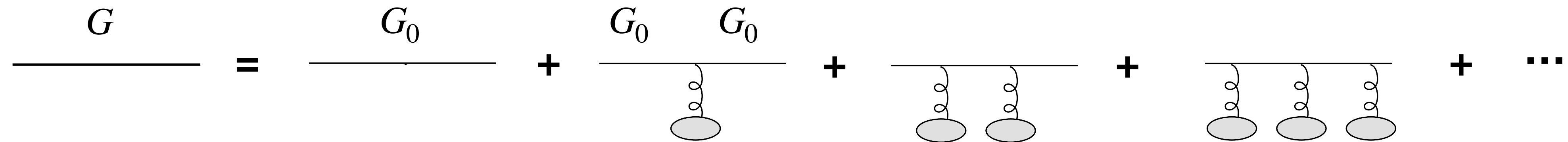
$$\begin{aligned} \frac{d\mathcal{N}^{(N)}}{d^2\mathbf{x}dE} = \int \frac{d^2\mathbf{p}d^2\mathbf{r}}{(2\pi)^2} e^{i\mathbf{p}\cdot(\mathbf{x}-\mathbf{r})} (-1)^N [\mathcal{V}(\mathbf{r})L]^N \left\{ \frac{1}{N!} + \frac{L}{E(N+1)!} \times \sum_{m=1}^N \left[(N+1-m)\mathbf{p} \cdot \left(\frac{\mathcal{V}'(\mathbf{r})}{\mathcal{V}(\mathbf{r})} \nabla \mu^2 + \frac{1}{\rho} \nabla \rho \right) + i(N+1-m)^2 \frac{\nabla \mathcal{V}(\mathbf{r})}{\rho \mathcal{V}(\mathbf{r})} \cdot \nabla \rho \right. \right. \\ \left. \left. + i(N+1-m) \left(\frac{\nabla \mathcal{V}'(\mathbf{r})}{\mathcal{V}(\mathbf{r})} + (N-m) \frac{\mathcal{V}'(\mathbf{r})}{\mathcal{V}(\mathbf{r})} \frac{\nabla \mathcal{V}(\mathbf{r})}{\mathcal{V}(\mathbf{r})} \right) \cdot \nabla \mu^2 \right] \right\} \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{r}dE} \end{aligned}$$

Resumming the opacity series then leads to the compact expression

$$\frac{d\mathcal{N}}{d^2\mathbf{x}dE} = e^{-\mathcal{V}(\mathbf{x})L} \left\{ \left[1 - i \frac{\mathcal{V}(\mathbf{x})L^3}{6E} \left(\frac{\mathcal{V}'(\mathbf{x})}{\mathcal{V}(\mathbf{x})} \nabla \mu^2 + \frac{1}{\rho} \nabla \rho \right) \cdot \nabla \mathcal{V}(\mathbf{x}) \right] \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x}dE} + i \frac{\mathcal{V}(\mathbf{x})L^2}{2E} \left(\frac{\mathcal{V}'(\mathbf{x})}{\mathcal{V}(\mathbf{x})} \nabla \mu^2 + \frac{1}{\rho} \nabla \rho \right) \cdot \nabla \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x}dE} \right\}$$

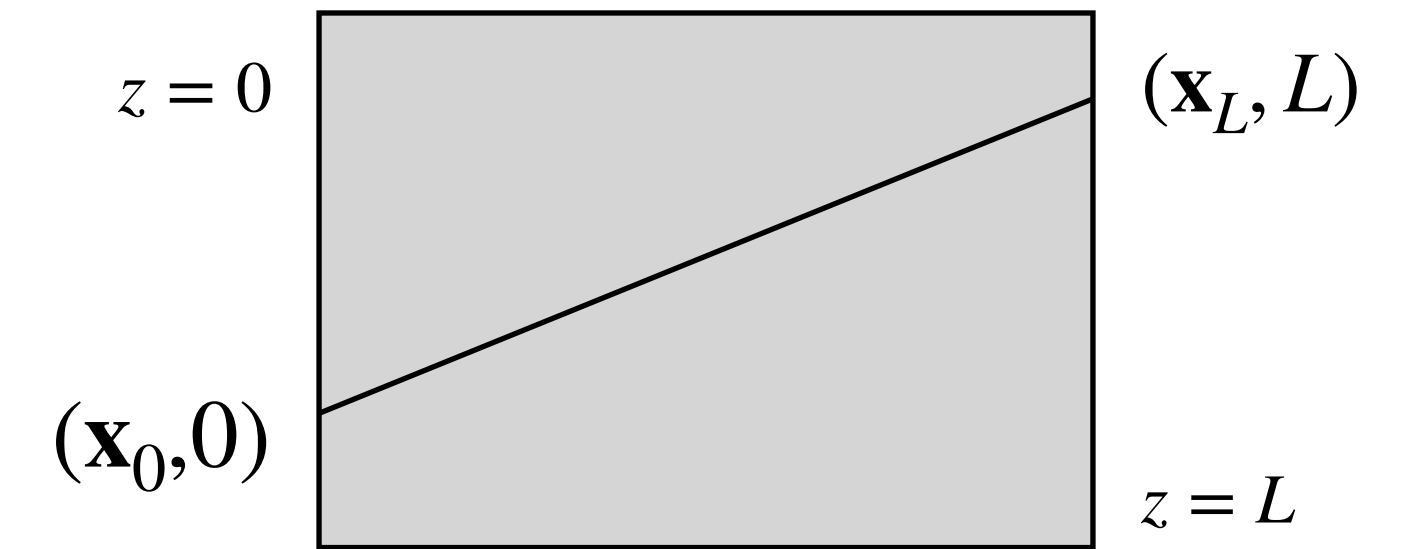
Momentum broadening in anisotropic media

1 Compute an effective in-medium propagator

$$\frac{G}{\text{---}} = \frac{G_0}{\text{---}} + \frac{G_0}{\text{---}} \frac{G_0}{\text{---}} + \frac{G_0}{\text{---}} \frac{G_0}{\text{---}} \frac{G_0}{\text{---}} + \dots$$


This results in an effective propagator G

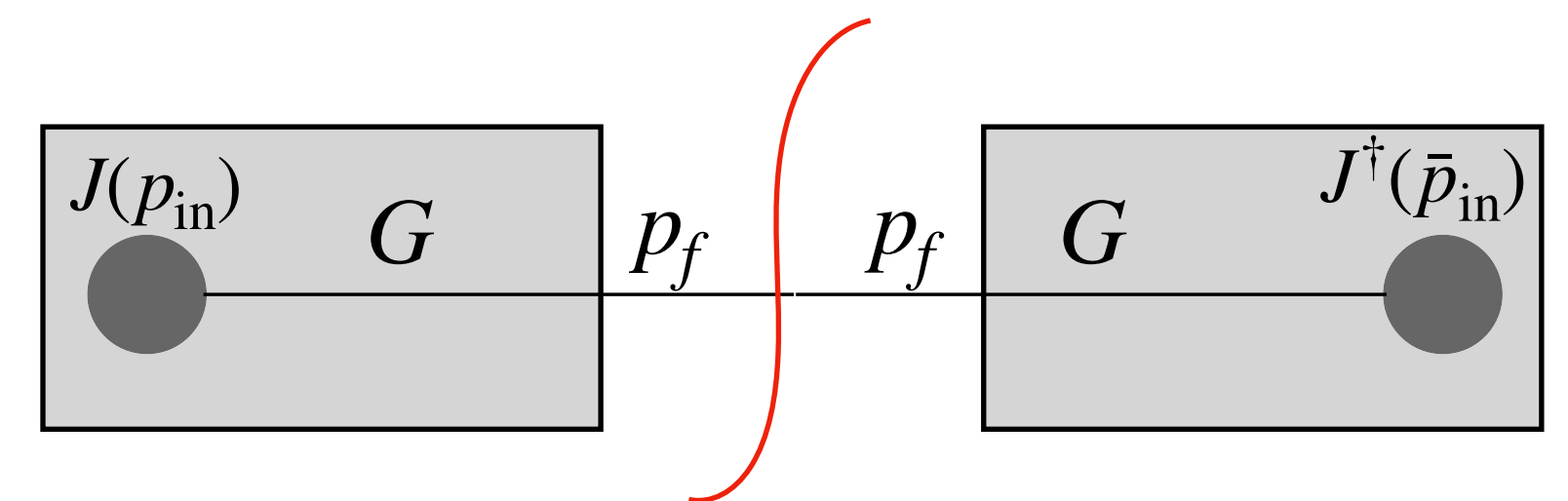
$$G(\mathbf{x}_L, L; \mathbf{x}_0, 0) = \int_{\mathbf{x}_0}^{\mathbf{x}_L} \mathcal{D}\mathbf{r} \exp\left(\frac{iE}{2} \int_0^L d\tau \dot{\mathbf{r}}^2\right) \mathcal{P} \exp\left(-i \int_0^L d\tau t_{\text{proj}}^a v^a(\mathbf{r}(\tau), \tau)\right)$$



2 Compute the relevant Feynman diagrams

$$\langle |M|^2 \rangle = \int \frac{d^2\mathbf{p}_{in} d^2\bar{\mathbf{p}}_{in}}{(2\pi)^4} \langle G(\mathbf{p}_f, L; \mathbf{p}_{in}, 0) G^\dagger(\mathbf{p}_f, L; \bar{\mathbf{p}}_{in}, 0) \rangle J(\mathbf{p}_{in}) J^*(\bar{\mathbf{p}}_{in})$$

$$gA_{\text{ext}}^{\mu a}(q) = -(2\pi) g^{\mu 0} v^a(q) \delta(q^0) \quad v^a(q) = \sum_i e^{-i\vec{q} \cdot \vec{x}_i} t_i^a v_i(q)$$



3 Solve the remaining average of dressed propagators

Option 1) Solve first the path integrals and then average Equivalent to Opacity Series approach

Option 2) Perform the average before integration

In practice, by solving the remaining integrals one performs the resummation of averaged quantities directly

The **key step** is to use the fact that the color average of the potential at different positions

$$\langle t_{\text{proj}}^a v^a(\mathbf{r}, \tau) t_{\text{proj}}^b v^{\dagger b}(\bar{\mathbf{r}}, \bar{\tau}) \rangle = C g^4 \int dz d^2\mathbf{x} \rho(\mathbf{x}, z) \int \frac{d^2\mathbf{q} dq_z d^2\bar{\mathbf{q}} d\bar{q}_z}{(2\pi)^6} \frac{e^{i\mathbf{q}\cdot(\mathbf{r}-\mathbf{x})} e^{-i\bar{\mathbf{q}}\cdot(\bar{\mathbf{r}}-\mathbf{x})} e^{iq_z(\tau-z)} e^{-i\bar{q}_z(\bar{\tau}-z)}}{(\mathbf{q}^2 + q_z^2 + \mu^2(\mathbf{x}, z))(\bar{\mathbf{q}}^2 + \bar{q}_z^2 + \mu^2(\mathbf{x}, z))}.$$

implies for $\nabla T = 0$

$$\left\langle \mathcal{P} \exp \left(-i \int_0^L d\tau t_{\text{proj}}^a v^a(\mathbf{r}(\tau), \tau) \right) \mathcal{P} \exp \left(i \int_0^L d\bar{\tau} t_{\text{proj}}^b v^b(\bar{\mathbf{r}}(\bar{\tau}), \bar{\tau}) \right) \right\rangle = \exp \left\{ - \int_0^L d\tau \mathcal{V}(\mathbf{r}(\tau) - \bar{\mathbf{r}}(\tau)) \right\}$$

Momentum broadening in anisotropic media

To linear order in gradients from 3 we find now

$$\langle t_{\text{proj}}^a v^a(\mathbf{r}, \tau) t_{\text{proj}}^b v^{\dagger b}(\bar{\mathbf{r}}, \bar{\tau}) \rangle \simeq \left(1 + \frac{\mathbf{r}(\tau) + \bar{\mathbf{r}}(\tau)}{2} \cdot \left(\nabla \rho \frac{\delta}{\delta \rho} + \nabla \mu^2 \frac{\delta}{\delta \mu^2} \right) \right) \mathcal{C} \delta(\tau - \bar{\tau}) \rho g^4 \int \frac{d^2 \mathbf{q}}{(2\pi)^2} \frac{e^{i\mathbf{q} \cdot (\mathbf{r} - \bar{\mathbf{r}})}}{(\mathbf{q}^2 + \mu^2)^2}$$

One can still show that the 2-point correlator exponentiates

before

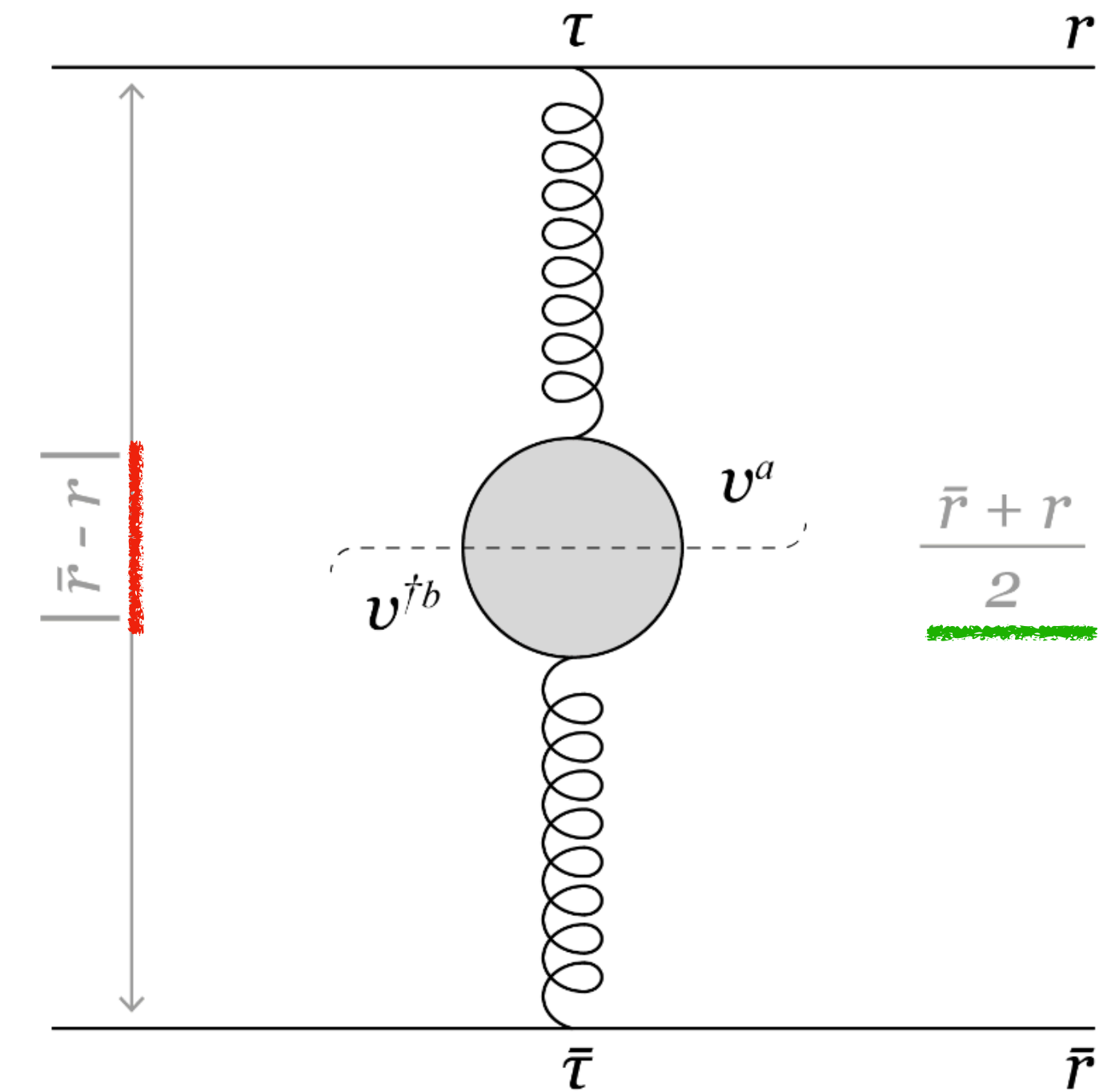
$$\left\langle \mathcal{P} \exp \left(-i \int_0^L d\tau t_{\text{proj}}^a v^a(\mathbf{r}(\tau), \tau) \right) \mathcal{P} \exp \left(i \int_0^L d\bar{\tau} t_{\text{proj}}^b v^b(\bar{\mathbf{r}}(\bar{\tau}), \bar{\tau}) \right) \right\rangle = \exp \left\{ - \int_0^L d\tau \mathcal{V}(\mathbf{r}(\tau) - \bar{\mathbf{r}}(\tau)) \right\}$$



$$\left\langle \mathcal{P} \exp \left(-i \int_0^L d\tau t_{\text{proj}}^a v^a(\mathbf{r}(\tau), \tau) \right) \mathcal{P} \exp \left(i \int_0^L d\bar{\tau} t_{\text{proj}}^b v^b(\bar{\mathbf{r}}(\bar{\tau}), \bar{\tau}) \right) \right\rangle = \exp \left\{ - \int_0^L d\tau \left[1 + \frac{\mathbf{r}(\tau) + \bar{\mathbf{r}}(\tau)}{2} \cdot \left(\nabla \rho \frac{\delta}{\delta \rho} + \nabla \mu^2 \frac{\delta}{\delta \mu^2} \right) \right] \mathcal{V}(\mathbf{r}(\tau) - \bar{\mathbf{r}}(\tau)) \right\}$$

Center of mass of dipole

Dipole size



Momentum broadening in anisotropic media

The final distribution has the form

$$\frac{d\mathcal{N}}{d^2\mathbf{x}dE} \simeq \exp\{-\mathcal{V}(\mathbf{x})L\} \left\{ \left[1 - \frac{iL^3}{6E} \nabla\mathcal{V}(\mathbf{x}) \cdot \left(\nabla\rho \frac{\delta}{\delta\rho} + \nabla\mu^2 \frac{\delta}{\delta\mu^2} \right) \mathcal{V}(\mathbf{x}) \right] \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x}dE} + \frac{iL^2}{2E} \left(\nabla\rho \frac{\delta}{\delta\rho} + \nabla\mu^2 \frac{\delta}{\delta\mu^2} \right) \mathcal{V}(\mathbf{x}) \cdot \nabla \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x}dE} \right\}$$

$$\frac{d\mathcal{N}}{d^2\mathbf{x}dE} = \mathcal{P}(\mathbf{x}) \hat{S}(\mathbf{x}) \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x}dE}$$

In the literature referred to as single particle broadening distribution (when Fourier transformed)

Usually a unit operator, but now it acts with ∇ on initial distribution

Effective factorization no longer holds in general due to operator nature

Still

$$\int d^2\mathbf{p} \frac{d\mathcal{N}}{d^2\mathbf{p}dE} = \frac{d\mathcal{N}}{d^2\mathbf{x}dE} \Big|_{\mathbf{x}=0} = \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x}dE} \Big|_{\mathbf{x}=0} = \int d^2\mathbf{p} \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}dE}$$

Consider the case of a source with finite width

$$E \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}dE} = \frac{f(E)}{2\pi w^2} e^{-\frac{\mathbf{p}^2}{2w^2}}$$

It is possible to show that even though $\langle \mathbf{p} \rangle = 0$

higher odd moments can be generated, for example

$$\langle p^\alpha \mathbf{p}^2 \rangle = \frac{w^2 L^2 \mu^2}{E \lambda} \frac{\nabla^\alpha \rho}{\rho} \ln \frac{E}{\mu} + \frac{L^3 \mu^4}{6E \lambda^2} \frac{\nabla^\alpha \rho}{\rho} \left(\ln \frac{E}{\mu} \right)^2$$

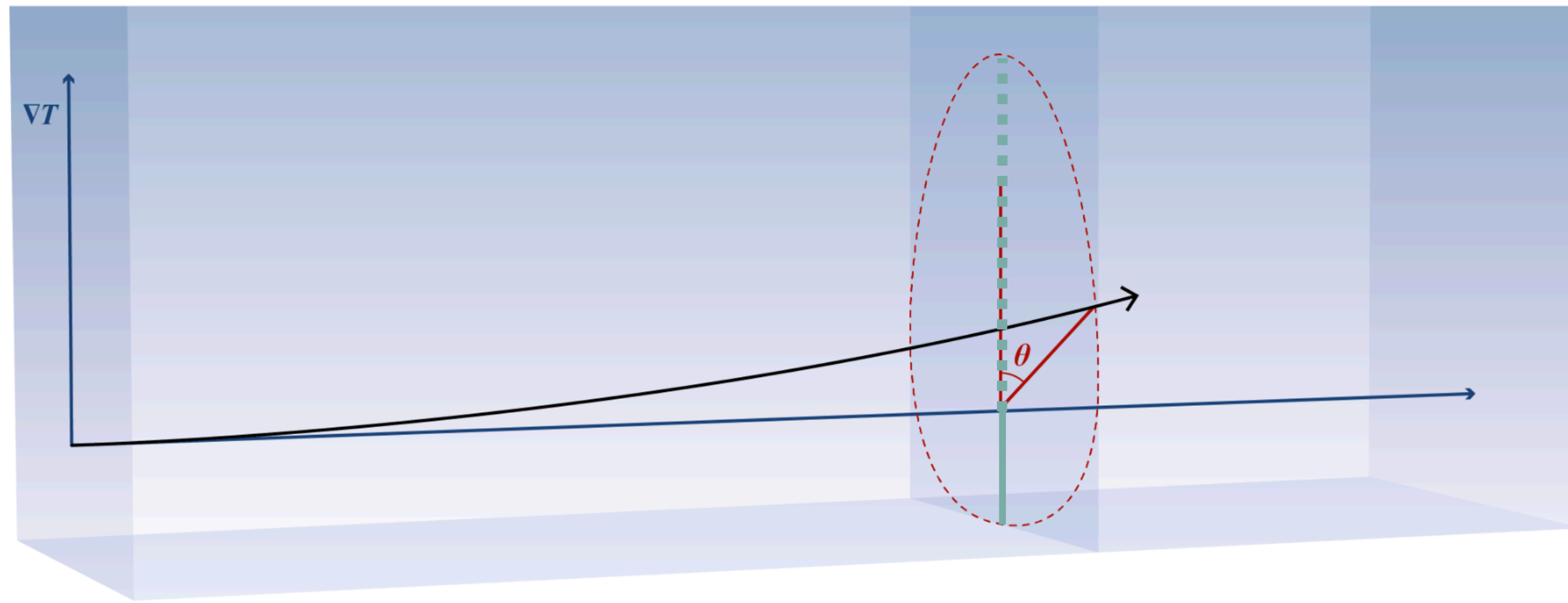
$N = 1$ $N = 2$

Higher N terms dominate due to diverging potential at large momenta

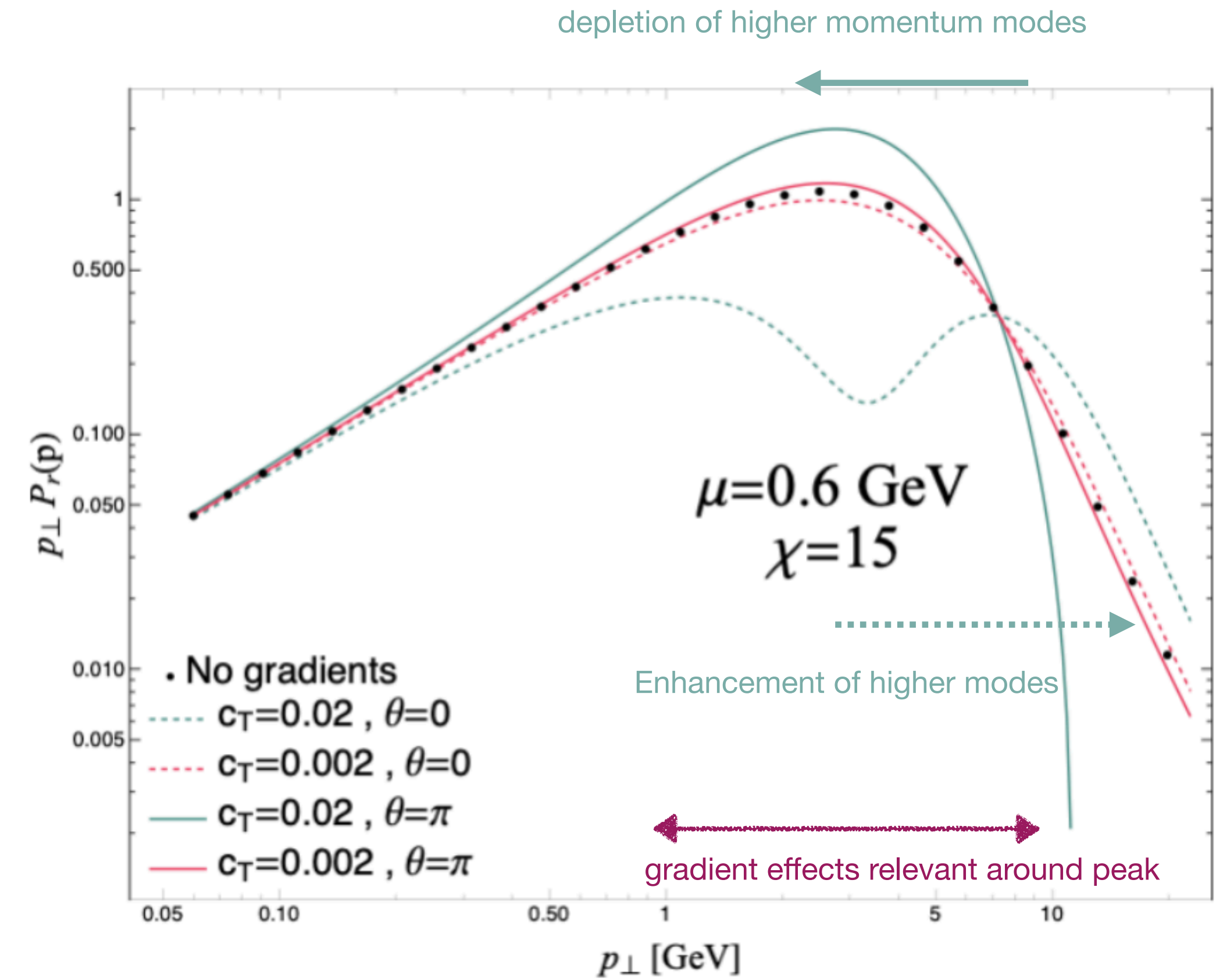
Coulomb logarithm

Momentum broadening in anisotropic media

For the full GW model we have

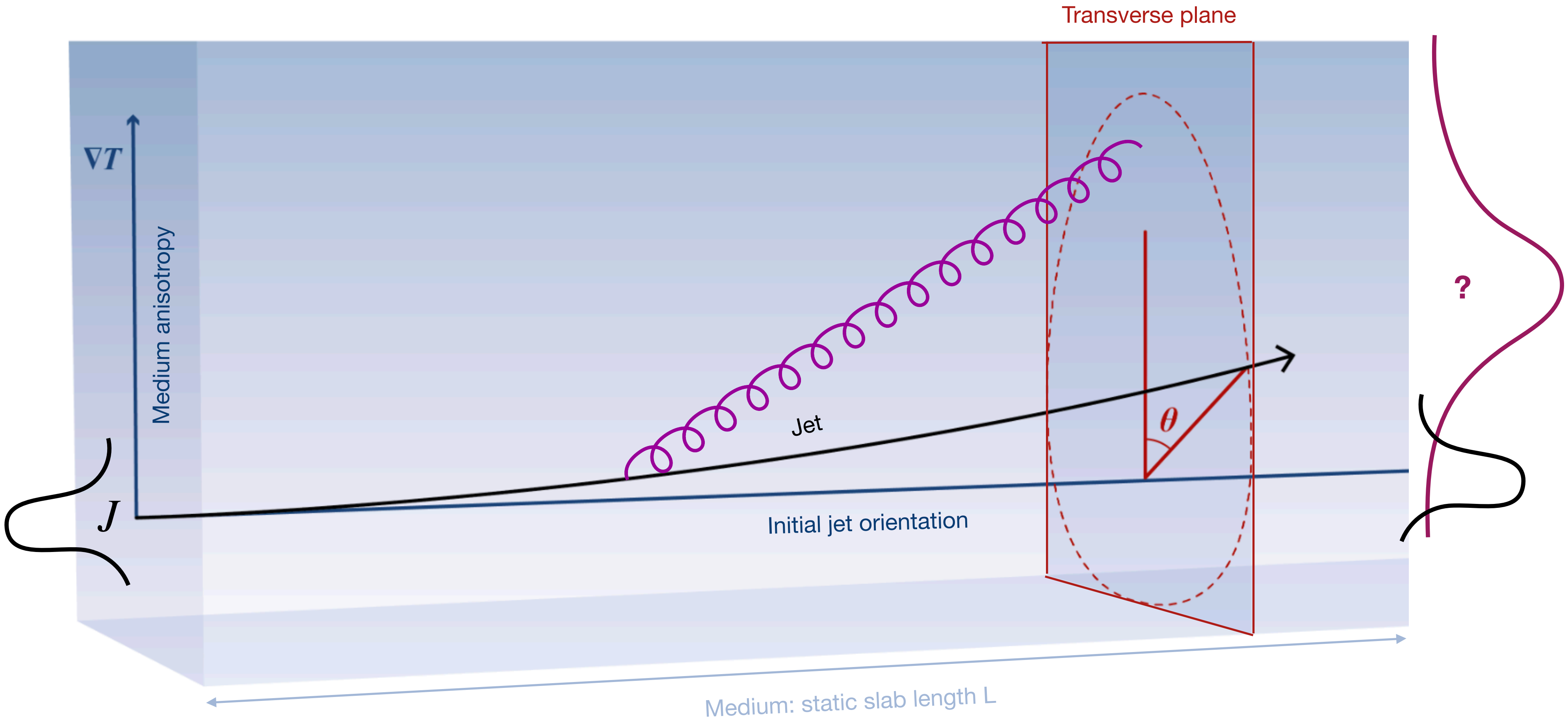


The full distribution is written in terms of the angle θ and parameter $c_T \equiv \left| \frac{\nabla T}{ET} \right|$.



$$\frac{\nabla \rho}{\rho} \sim 3 \frac{\nabla T}{T}, \quad \frac{\nabla \mu^2}{\mu^2} \sim 2 \frac{\nabla T}{T}$$

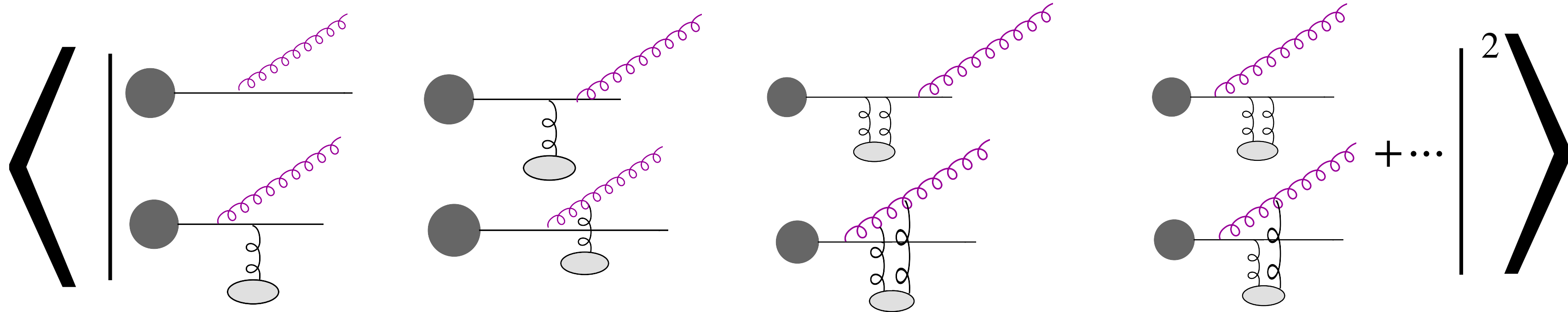
Radiative energy loss in dense anisotropic media



Radiative energy loss in dense anisotropic media

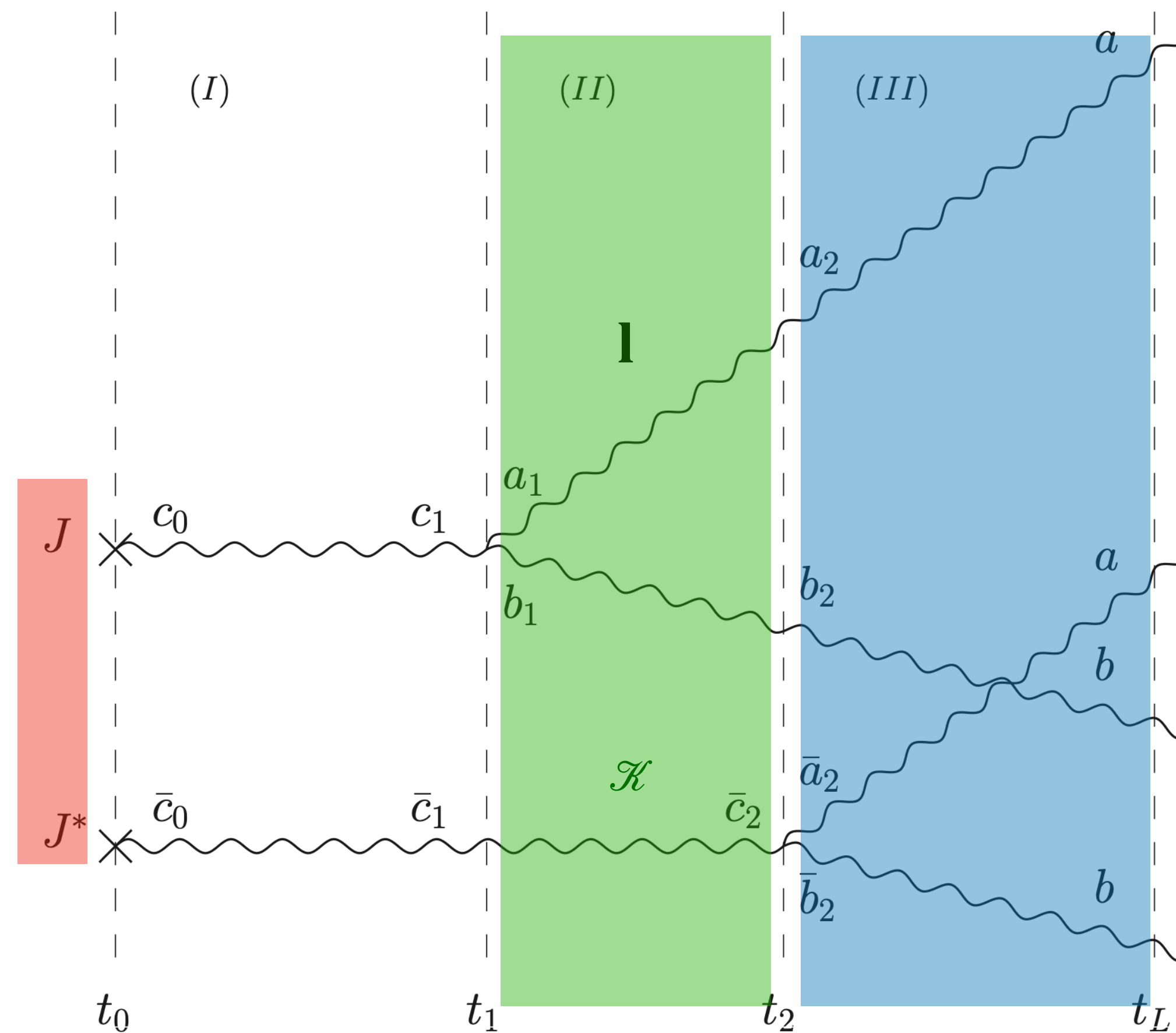
In **Opacity Expansion style** calculation one needs to compute

Beyond $N=1$, this
seems to be quite
hard to do



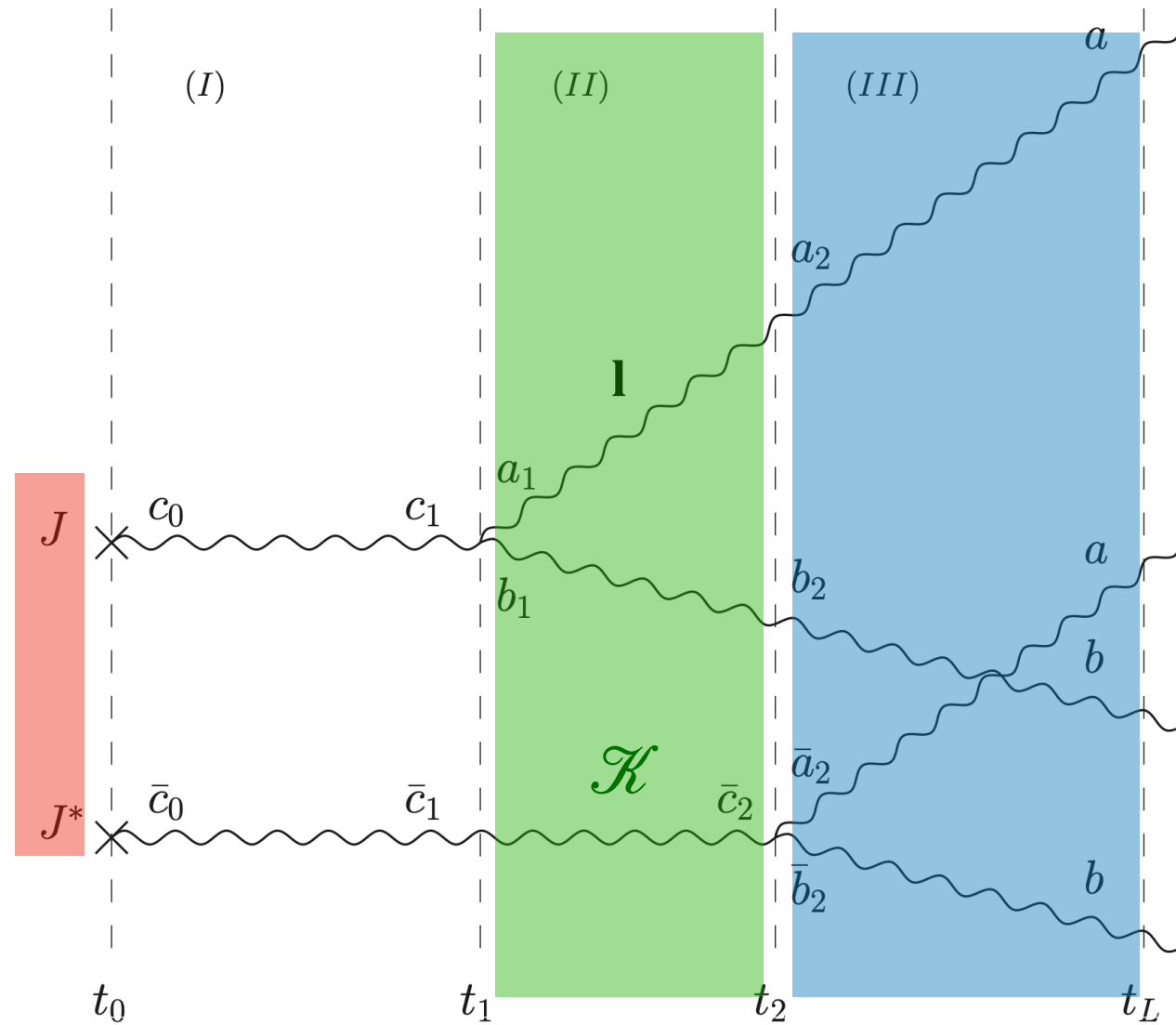
Radiative energy loss in dense anisotropic media

Using **resummed propagators** calculation becomes simpler



see e.g. 1209.4585, J.-P. Blaizot, F. Dominguez, E. Iancu, Y. Mehtar-Tani

Radiative energy loss in dense anisotropic media



In this case, we can write the squared amplitude as

$$2(2\pi)^3 \omega E \frac{d\mathcal{N}}{d\omega dE d^2\mathbf{k}} = \frac{2\alpha_s C_F}{\omega^2} \text{Re} \int_0^\infty d\bar{z} \int_0^{\bar{z}} dz \int_{\mathbf{x}_{in}, \mathbf{y}} |J(\mathbf{x}_{in})|^2 \left[\nabla_{\mathbf{x}} \cdot \nabla_{\bar{\mathbf{x}}} \left[\underbrace{S_2(\mathbf{k}, \mathbf{k}, \infty; \mathbf{y}, \bar{\mathbf{x}}, \bar{z})}_{\text{Solved!}} \mathcal{K}(\mathbf{y}, \mathbf{x}_{in}, \bar{z}; \mathbf{x}, \mathbf{x}_{in}, z) \right] \right] \Big|_{\mathbf{x}=\bar{\mathbf{x}}=\mathbf{x}_{in}}$$

Gluon production is determined by a correlator of the form

$$\mathcal{K}(\mathbf{y}, \mathbf{x}_{in}, \bar{z}; \mathbf{x}, \mathbf{x}_{in}, z) = \int_{\mathbf{x}}^{\mathbf{y}} D\mathbf{r} e^{\frac{i\omega}{2} \int d\tau \dot{\mathbf{r}}^2 - \int_z^{\bar{z}} d\tau \left(1 + \frac{\mathbf{r}(\tau) + \mathbf{x}_{in}}{2} \cdot \hat{\mathbf{g}} \right) \mathcal{V}(\mathbf{r}(\tau) - \mathbf{x}_{in})}$$

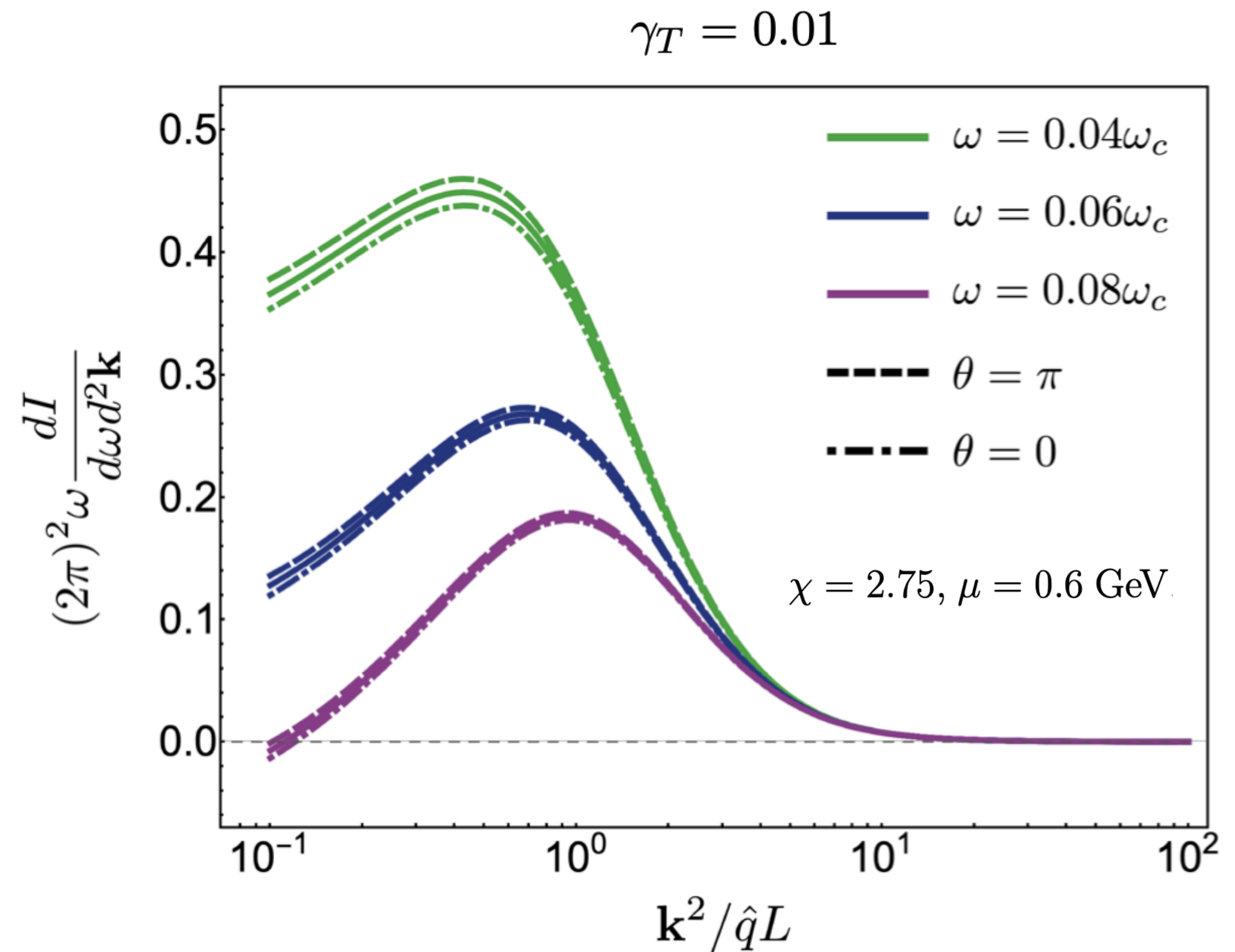
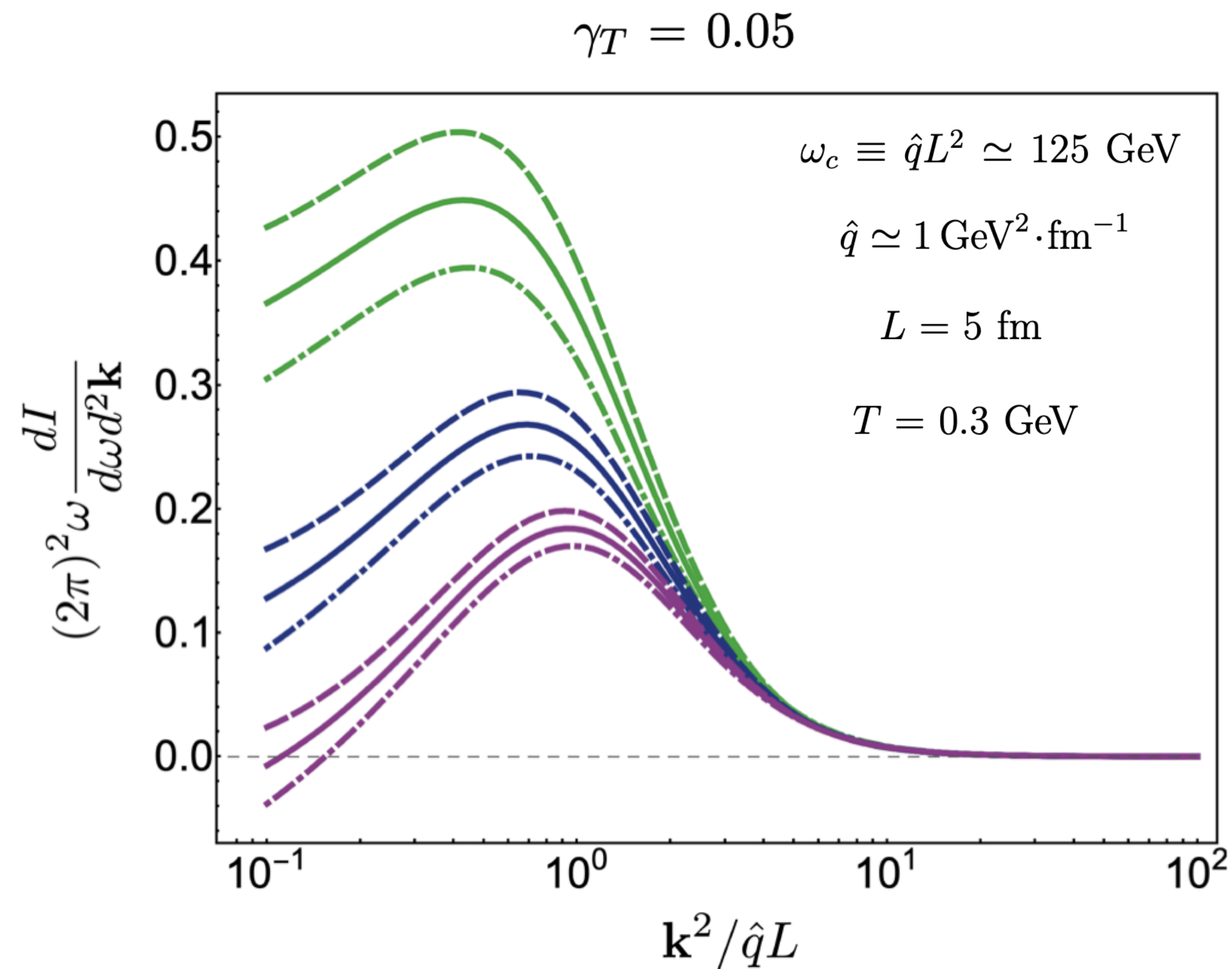
This object satisfies a simple recursive equation; expanding to first order in gradients it allows to perturbatively compute the spectrum in the form

$$\omega \frac{dI}{d\omega d^2\mathbf{k}} = \omega \frac{dI_0}{d\omega d^2\mathbf{k}} + (\hat{\mathbf{g}} \cdot \mathbf{k}) \omega \frac{dI_1}{d\omega d^2\mathbf{k}} + \mathcal{O}(\hat{\mathbf{g}}^2)$$

$$\omega \frac{dI}{d\omega d^2\mathbf{k}} = \omega \frac{dI_0}{d\omega d^2\mathbf{k}} + \omega \frac{dI_{\mathcal{P}}}{d\omega d^2\mathbf{k}} + \omega \frac{dI_{\mathcal{K}}}{d\omega d^2\mathbf{k}} + \omega \frac{dI_{\hat{\mathbf{g}}}}{d\omega d^2\mathbf{k}}$$

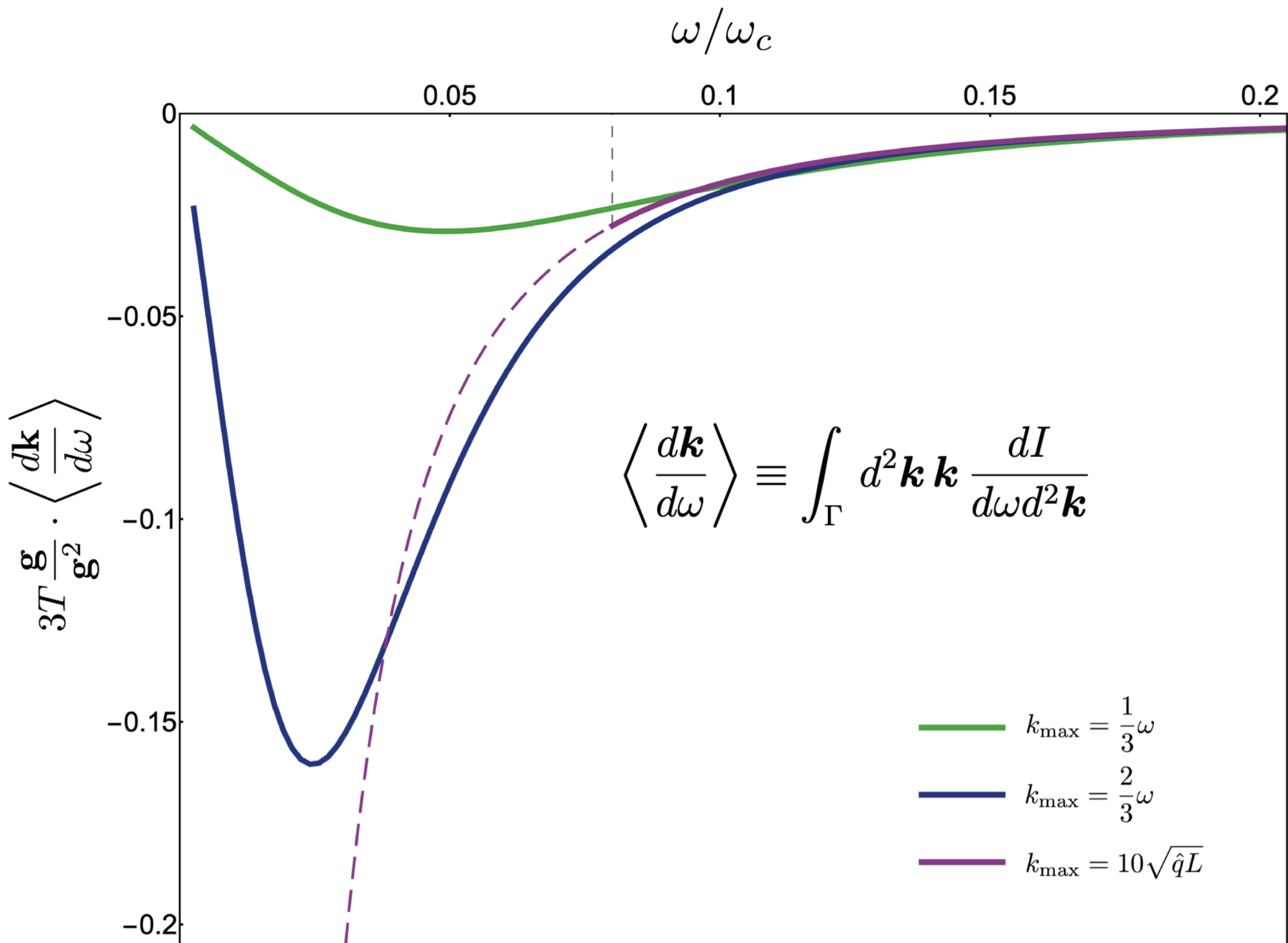
Radiative energy loss in dense anisotropic media

Numerical results, in the harmonic approximation for the in-medium scattering cross-section

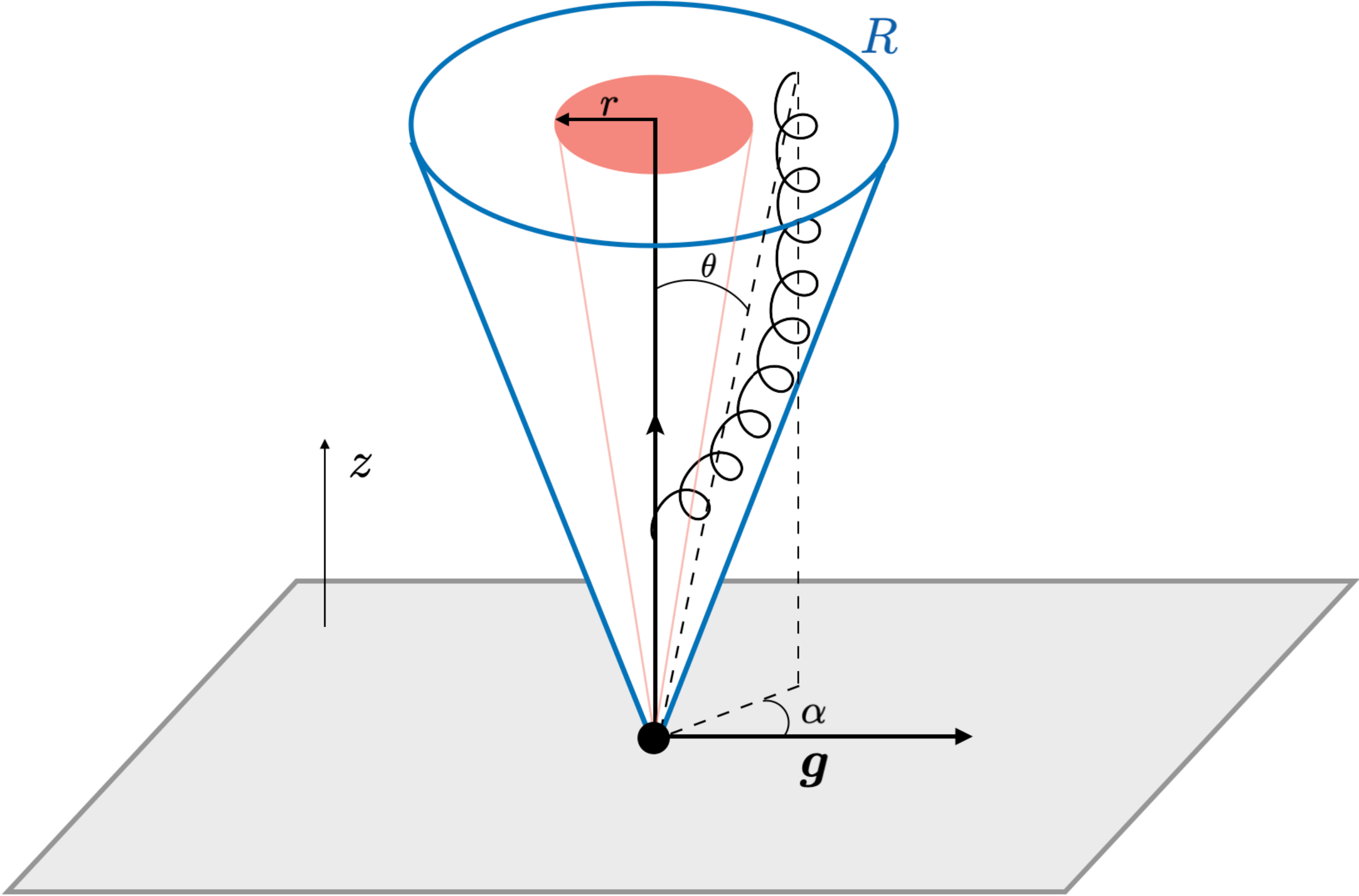


$$\gamma_T = |\nabla T / T^2|$$

Radiative energy loss in dense anisotropic media



Jet observables in inhomogeneous matter



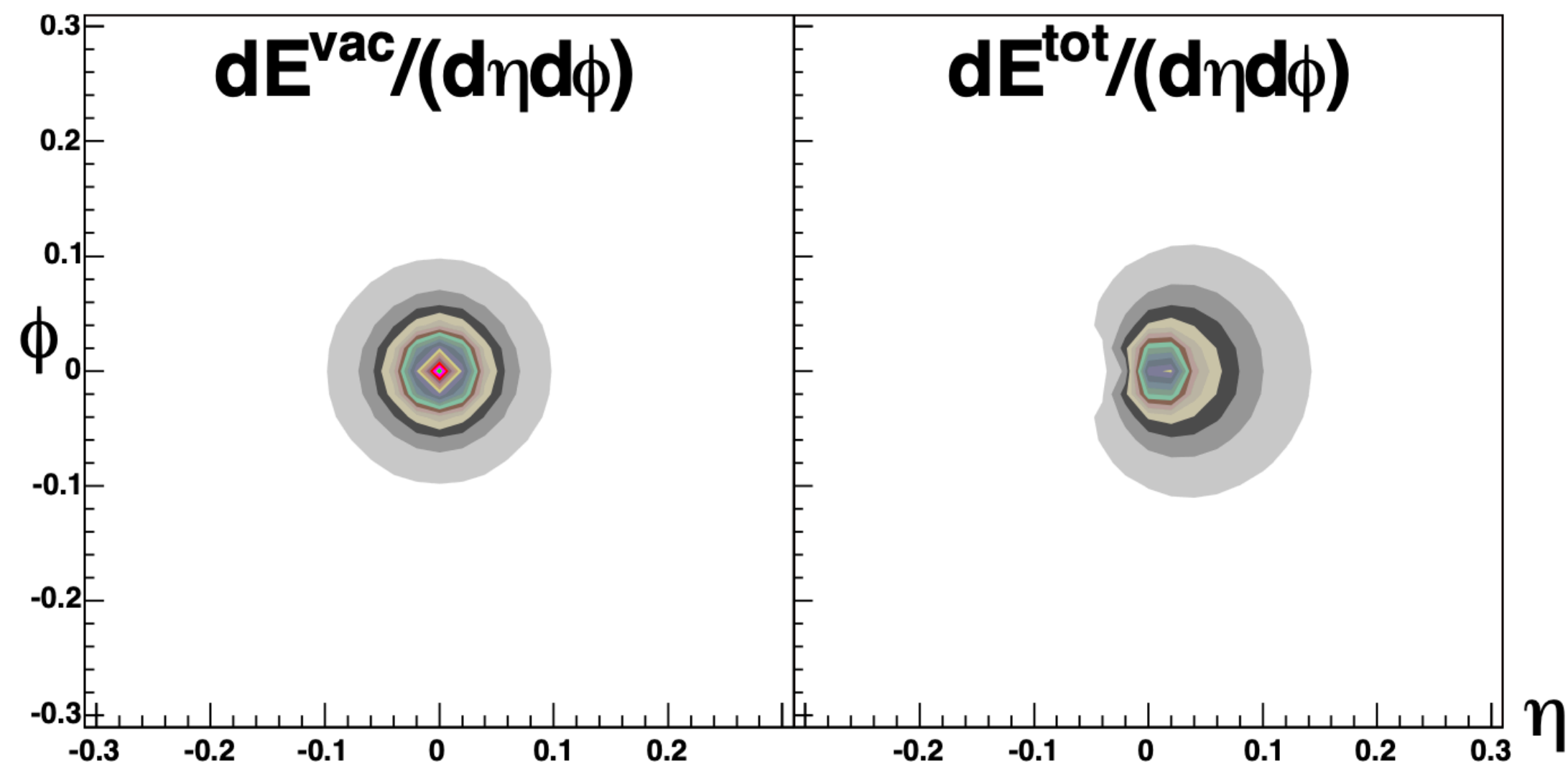
Jet observables in inhomogeneous matter

We have now the tools to compute jet observables (at least at leading order in strong coupling)

Observable 1: jet shape

$$(2\pi)p_t^{\text{jet}} \frac{d\rho(r)}{d\omega d\alpha} = 1 - 2\pi \int_{\omega r}^{\omega} dk k \omega \frac{dI}{d\omega d^2k}$$

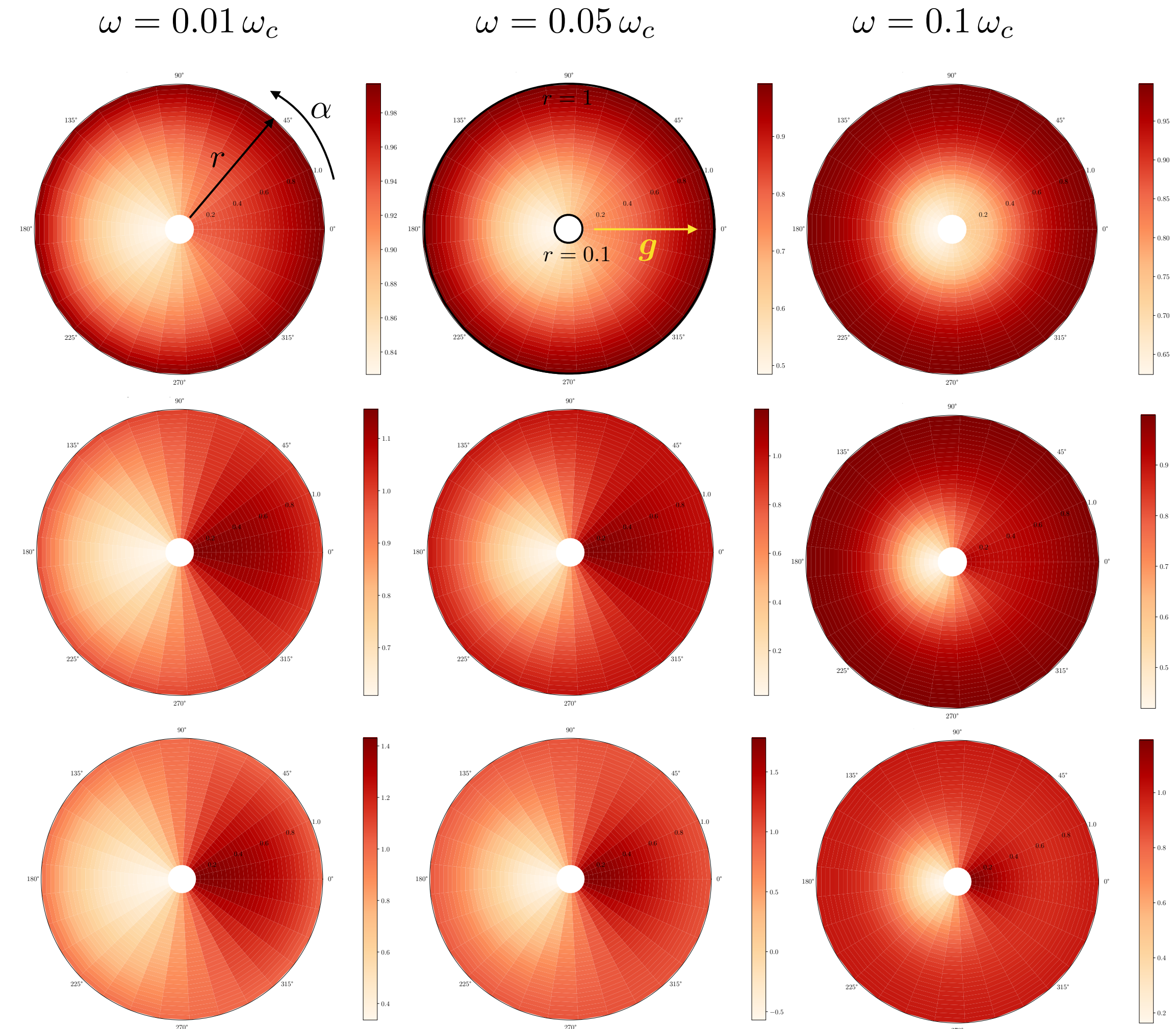
2004, Armesto, Salgado, Wiedemann



$\gamma_T = 0.1$

$\gamma_T = 0.5$

$\gamma_T = 1$



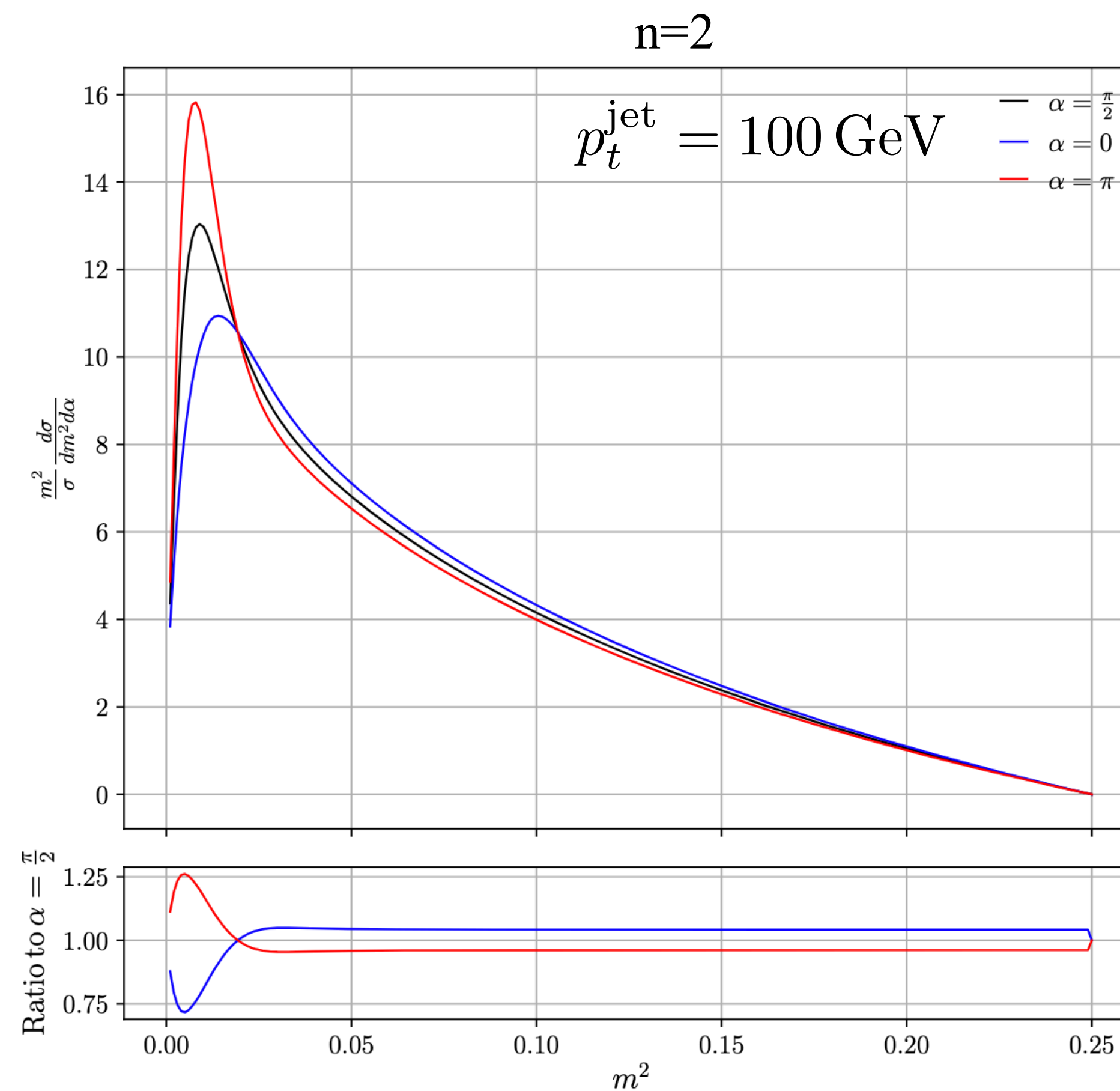
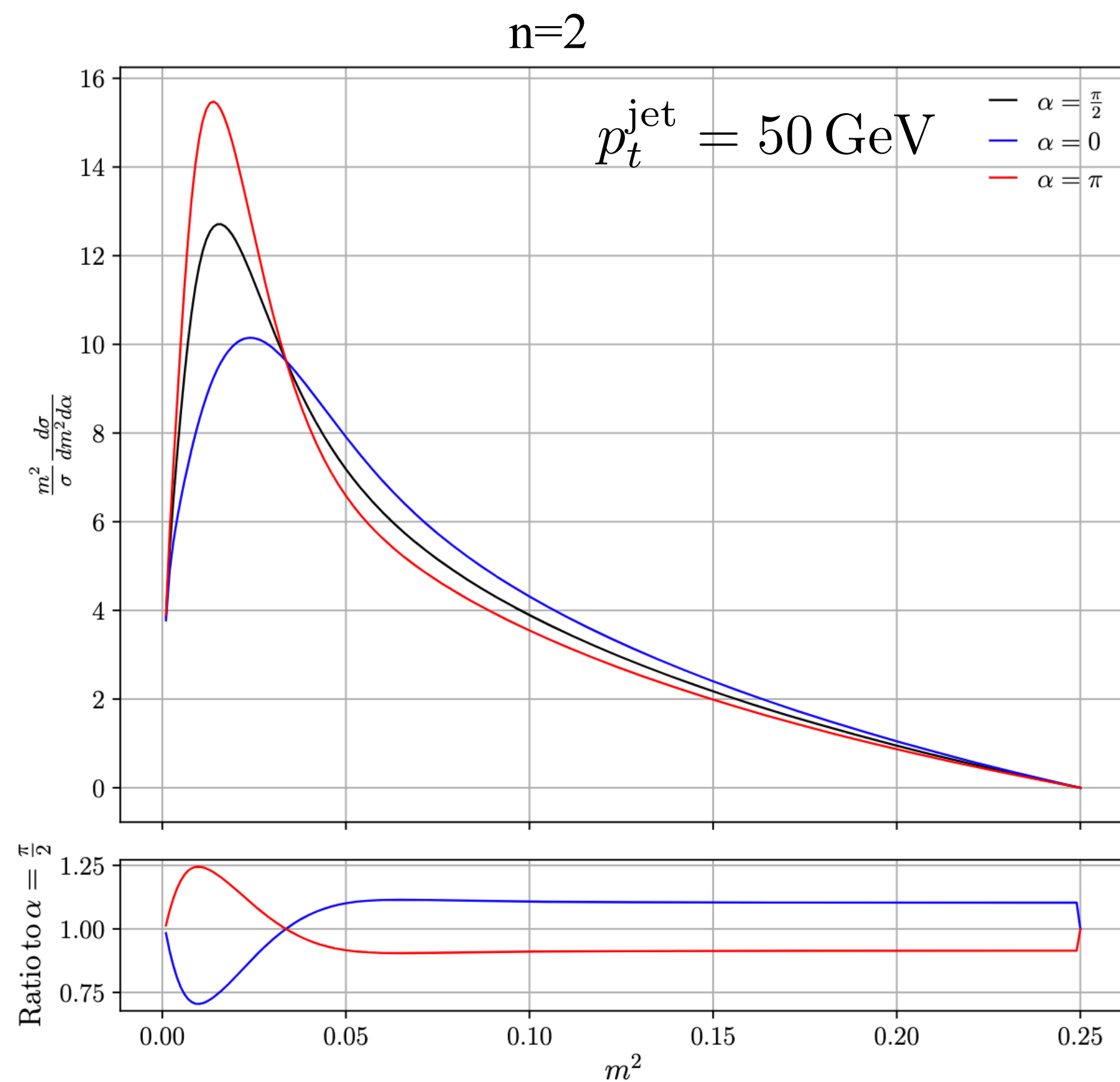
Jet observables in inhomogeneous matter

Observable 2: jet angularities

$$G_n = \sum_{i \in \text{jet}} \frac{p_t^i}{p_t^{\text{jet}}} g^{(n)}(r_i)$$

At leading logarithmic order we have

$$\frac{g_n}{\sigma} \frac{d\sigma}{dg_n d\alpha} = \left(\int_{\frac{g_n}{R^n}}^1 dx \left(\frac{\omega dI}{d\omega d^2\mathbf{k}} \frac{(p_t^{\text{jet}})^2 x^{1-\frac{2}{n}} g_n^{\frac{2}{n}}}{n} \right)_{\theta^n = \frac{g_n}{x}} + \frac{\alpha_s C_F}{\pi^2 n} \log \frac{R^n}{g_n} \right) e^{-\frac{\alpha_s C_F}{n\pi} \log^2 \frac{R^n}{g_n}}$$



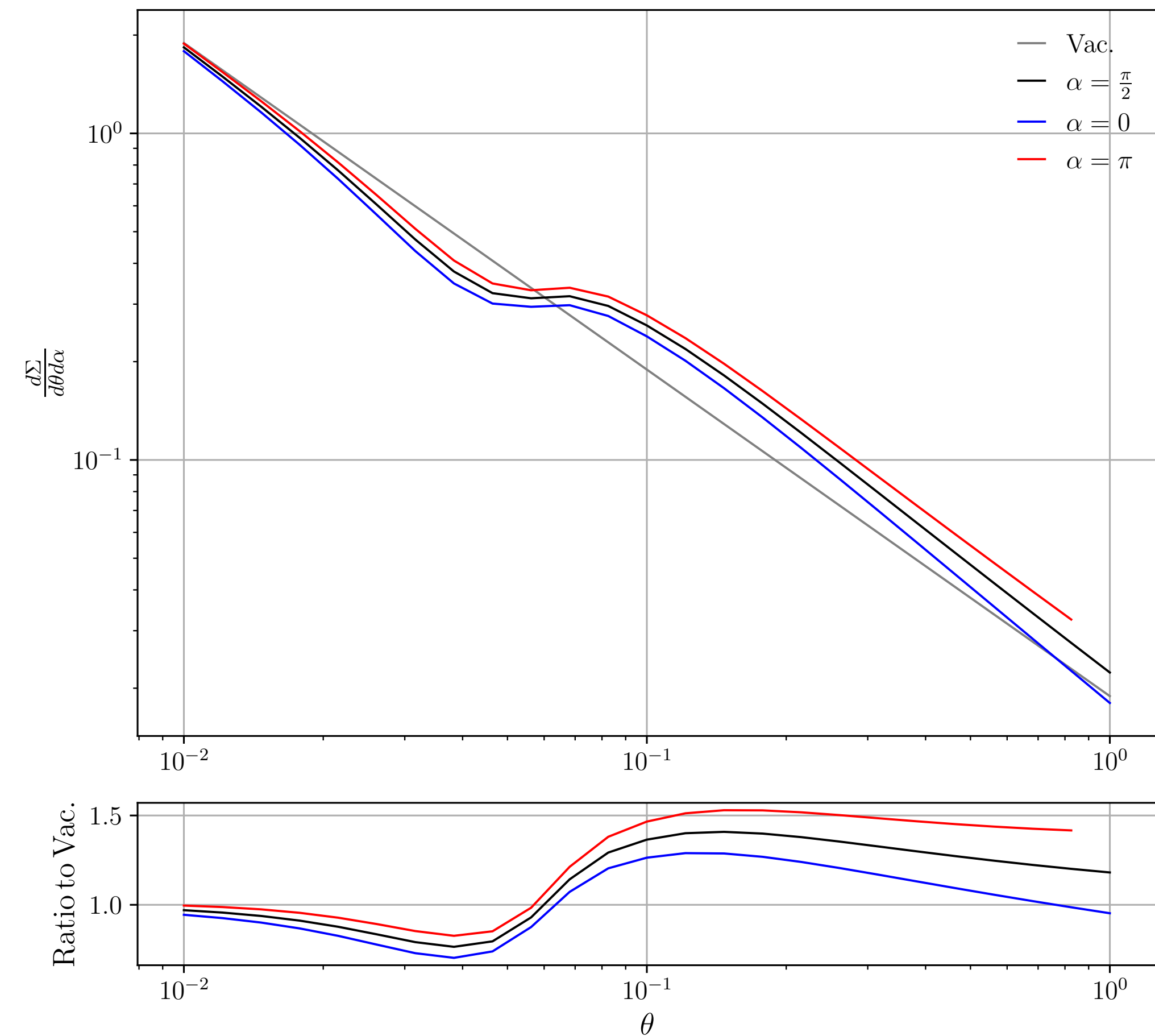
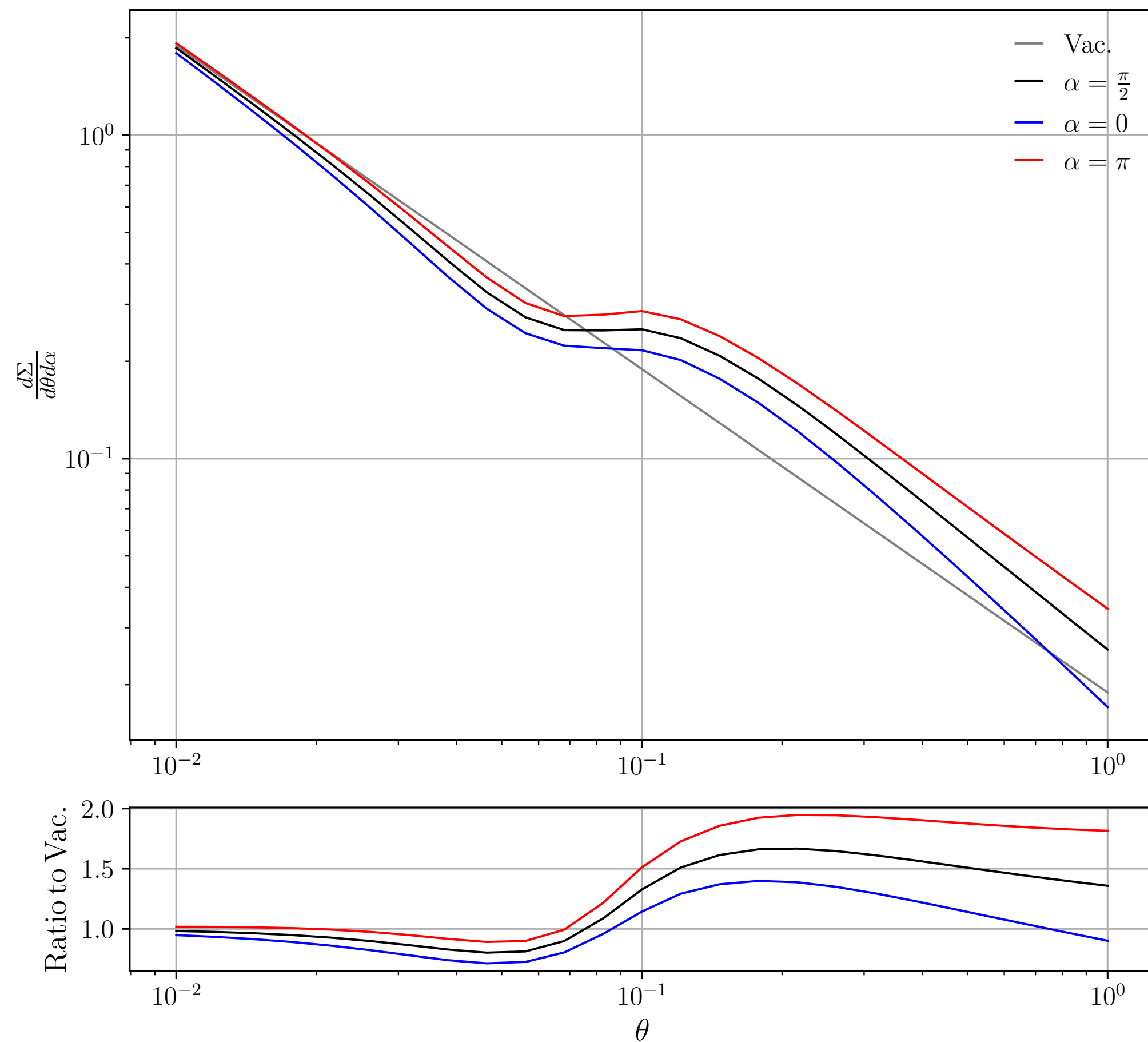
Jet observables in inhomogeneous matter

Observable 3: ENCs

Ideally we would need E3C, but already with EEC we have

$$\frac{d\Sigma}{d\theta d\alpha} = \int d\vec{n}_1 d\vec{n}_2 \frac{\langle \mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \rangle}{(p_t^{\text{jet}})^2} \delta(\cos(\theta_2 - \theta_1) - \cos(\theta)) \delta(\alpha - (\alpha_1 - \alpha_2))$$

for HICs and refs see Andres et al (2307.06226)



① I only discussed modifications due to anisotropies

Similar effects can be studied for medium flow

Jet Broadening in Flowing Matter – Resummation

Carlota Andres,^{1,*} Fabio Dominguez,^{2,†} Andrey V. Sadofyev,^{2,‡} and Carlos A. Salgado^{2,§}

② To go beyond we need MC with realistic geometries:

LBT model seems to agree with our approach !

