

#### **GWAK: Gravitational-Wave Anomalous Knowledge**

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## **Binary Mergers**



Produces: **time-series** [1-D strain + auxiliary channels]

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#### What are anomalies?

Anomalies are unmodelled waveforms

- GWs have been detected in Matched-Filter pipelines from:
  - Binary Black Holes (BBH)
  - Binary Neutron Stars (BNS)
  - Black Hole Neutron Star (BHNS)
- Anomalous signals:
  - Core Collapse Supernovas (CCSNe)
  - Neutron Star Glitches
  - Cosmic Strings
  - NSs collapsing to BHs
  - Gravitational Bremsstrahlung
  - Other stochastic processes

#### **GW Dataset - Nontrivial**

- Length measurements are ~  $10^{-22} m$
- Constantly changing detector noise usually clouds signal
- Detector glitches occur every O(10 sec) resembling GWs in excess power!



## **Unsupervised** Learning: Detection



## **Quasi-Anomalous Knowledge - QUAK**

[Park et al. (2021) (JHEP)]



# **Quasi-Anomalous Knowledge - QUAK**



#### QUAK -> GWAK





#### **Real-Time Animations**





#### **QUAK Spaces on Supernovae**



# **Quantitative Results**

- LSTM-AE currently best performing model
  - Model architecture input size favors shorter signals
- Dedicated searches perform well at low-SNR
- Able to consistently identify midrange SNR anomalies at incredibly low FAR
- Test on full LIGO O(3) incoming
- Ideas for real-time O(4) in the works



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