

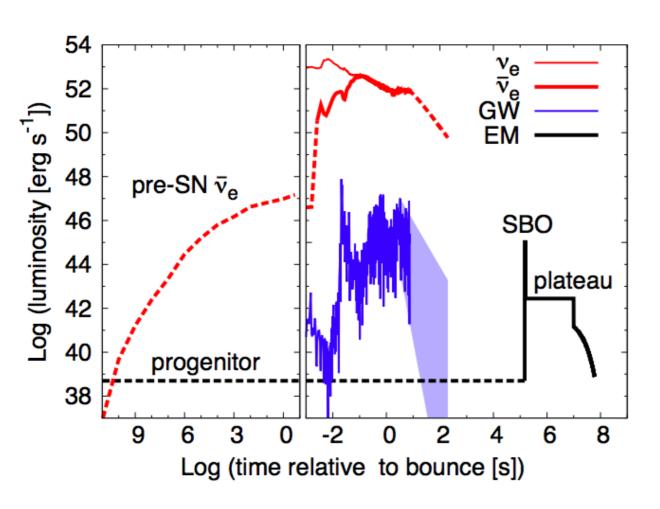
Low-latency pointing to a supernova with neutrinos

for Duke University A3D3 group by Kate Scholberg

Janina Hakenmüller (postdoc) Josh Queen (grad) Van Tha Bik Lian (postbacc) → new postbacc: Lucie Afko

July 10, 2023 - A3D3 High-Throughput AI Methods and Infrastructure Workshop

Multimessenger astronomy



K. Nakamura et al., MNRAS Vol461 Issue 3 2016

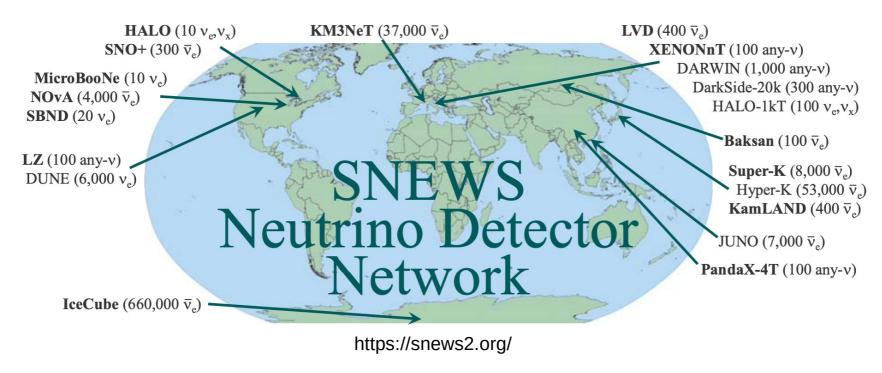
Supernova bursts:

- >99% of energy released in the form of neutrinos
- time scale:
 - neutrinos: 10⁻²s -10s
 - grav. Waves: 10⁻²s -10s
 - el.-mag.: >10⁴s
- neutrinos interact only weakly with a very small cross section
- => neutrino detection for early warning system of burst
- => detection has enormous physics potential for astronomy and particles physics

SNEWS

(SuperNova Early Warning System)

- SNEWS1.0: simple 10s coincidence, running in automatic mode since 2005
- Upgrade to SNEWS2.0: improved latency, neutrino based pointing, including triangulation

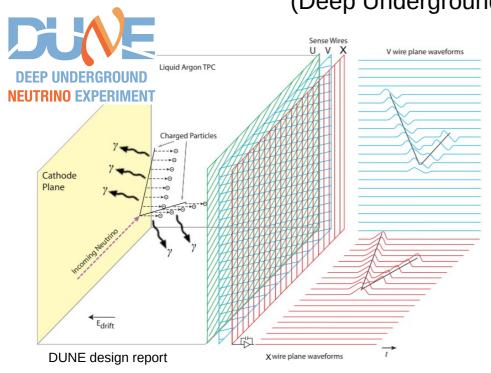


Detection materials: water-based, argon, scintillator, lead, noble liquid dark (dark matter exp.)



DUNE

(Deep Underground Neutrino Experiment)



- liquid argon time-projection chambers: ionization and scintillation light totally active calorimeter
- far detector in South Dakota: LAr mass 4 x 10kt
- underground: 1.5km rock/4300 m w.e.
- GeV scale: Neutrino oscillations in long baseline beam, study of atmospheric neutrinos, BSM physics, baryon number violation
- low energies: supenova neutrinos, solar neutrinos, diffuse supernova background

large mass

high statistics

underground

background suppression

excellent 3D imaging

pointing to supernova, multi messenger

complementarity to other experiments

(JUNO, Hyper-Kamiokande)

Supernova neutrino detection with DUNE

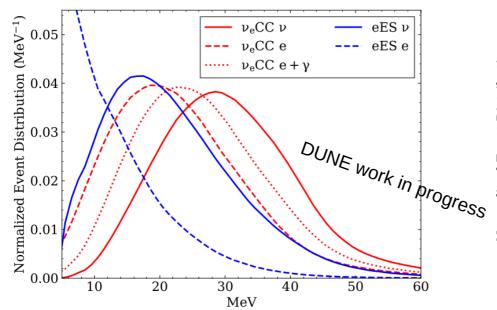
Dominant channels: charge-current interaction (v_eCC) and elastic electron scattering (eES)

Additional channels: ES of other flavors, NC scattering on Ar, ...

$$v_e + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$$

~3000 events @10kpc (GKVM model) flat angular distribution

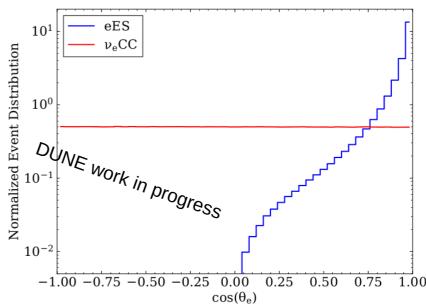
Energy spectra:



$$\nu_e + e^- \rightarrow \nu_e + e^-$$

~300 events @10kpc (GKVM model) primary e^{-} direction ~ v_{a} direction

Overall angular distributions:

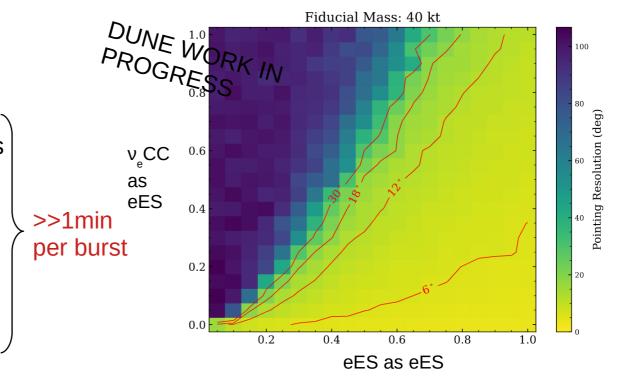


 \rightarrow channel tagging: distinguish v_e CC and eES

DUNE supernova pointing results

Procedure:

- simulation of supernova events, noise and background
- reconstruction of single events:
 - Identification of supernova events
 - energy
 - direction, daughter flipping for head-tail disambiguation
- combination of events to burst:
 - minimum energy: 10MeV
 - · maximum log likelihood
 - assumptions on channel tagging
- → publication in preparation



Likelihood output:

LAr volume	40kt	10kt (one module)
Perfect disambiguation	3.7 deg	7.4 deg
$4\% v_e$ CC as eES	5.0 deg	10.6 deg

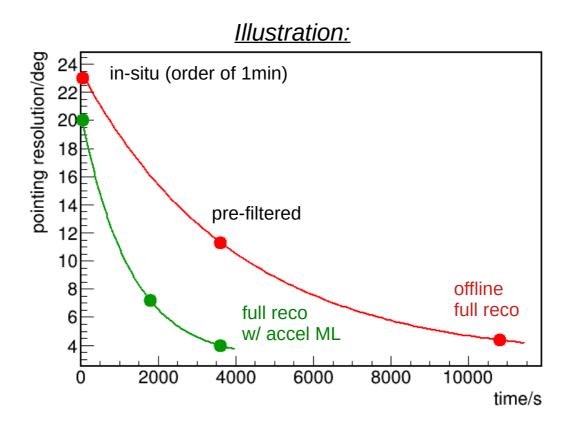
GKVM model, 10kpc distance

(68% confidence level)

Online Pointing and machine learning

<u>Strategy:</u> prompt, low-resolution followed by successive improvements ("kaizen" approach, coined for this by K. Scholberg at Accelerating physics with ML in Boston)

- → convert offline code to complete standalone pipeline, upgrade with ML
- → also test upgrades of offline code with ML



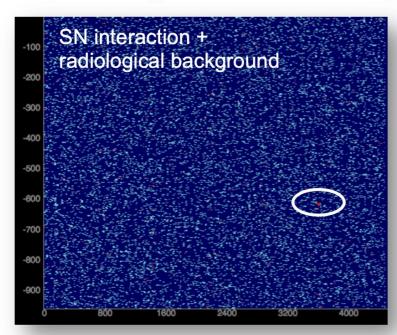
poster by JH

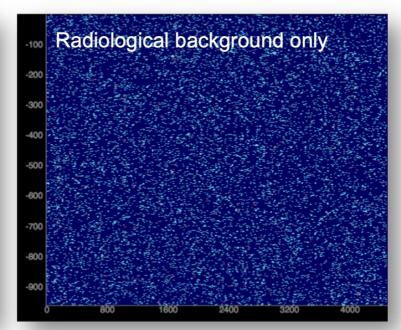
Identifying supernova events

High resolution "video stream": 11.5 Mega Pixel frames per 2.25ms, 12 bits resolution

Special challenge: neutrinos from supernova core collapse

Very low energy and small (in extent) topology, similar to radiological background activity in the detector





Need O(10⁴) overall background suppression, while maintaining high efficiency to a frame containing a supernova neutrino interaction

[simulation]

Real-Time Inference With 2D Convolutional Neural Networks on Field Programmable Gate Arrays for High-Rate Particle Imaging Detectors

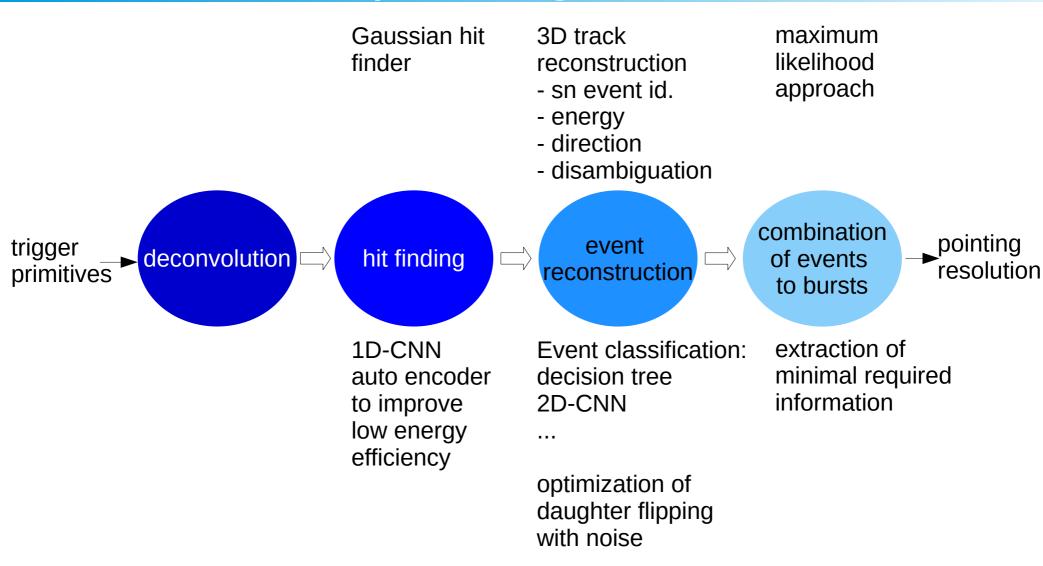
Yeon-jae Jwa (Columbia U.), Giuseppe Di Guglielmo (Columbia U.), Lukas Arnold (Columbia U.), Luca Carloni (Columbia U.), Georgia Karagiorgi (Columbia U.) (Jan 14, 2022)

Published in: Front. Artif. Intell. 5 (2022) 855184 • e-Print: 2201.05638 [physics.ins-det]

Georgia Karagiorgi, Columbia @ Fast Machine Learning - 2019

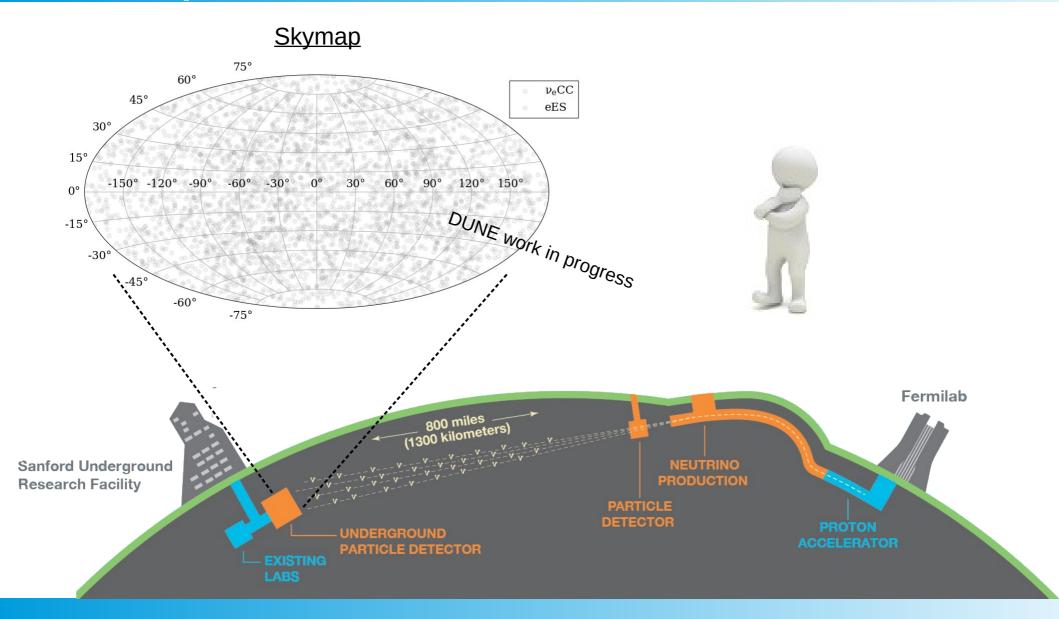
by G. Karagiorgi group, Columbia

Standalone pointing code for DUNE

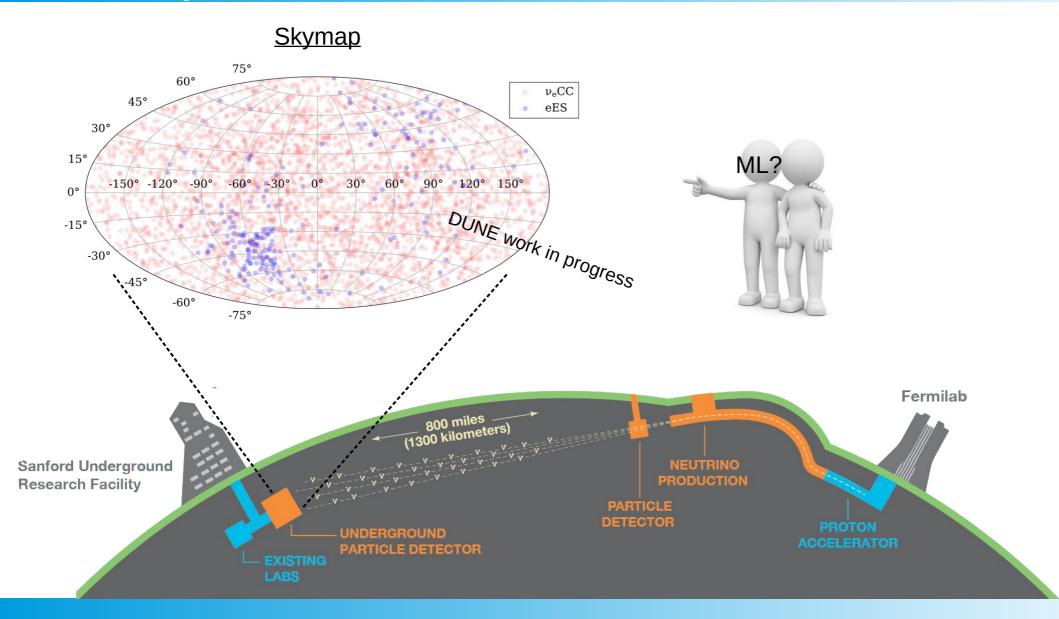


Remark: This list of ML efforts is for sure not complete.

Supernova detection with DUNE

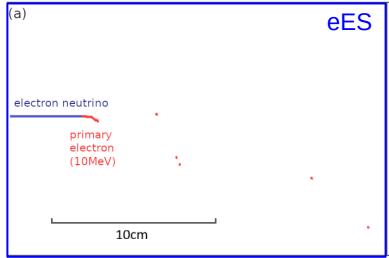


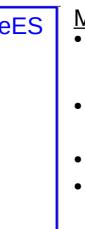
Supernova detection with DUNE



Event classification

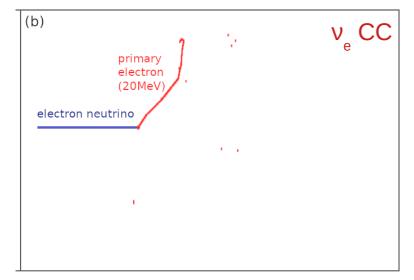
Geant4 event displays:





- ML approaches ongoing:
 E. Conley (Duke) https://zenodo.org/record/4122909#.Y9pfpy-B3pA

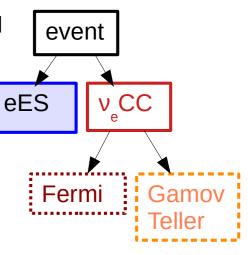
 → boosted decision tree (ave. track length, charge,...)
 → ~5% missidentification seems feasible
- G. Karagiorgi's group: YOLO (CNN for object detection)
- M. H.L.S. Wang et al.: 2D CNN



Possible to derive information for pointing? overall flat angular distribution, BUT:

the two types of deexcitation of ⁴⁰K* carry directional information

$$v_e + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$$

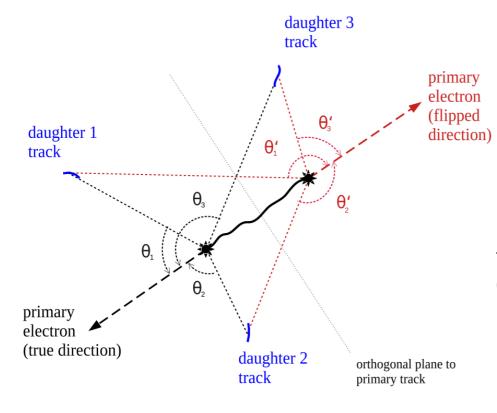


Directional head-tail disambiguation

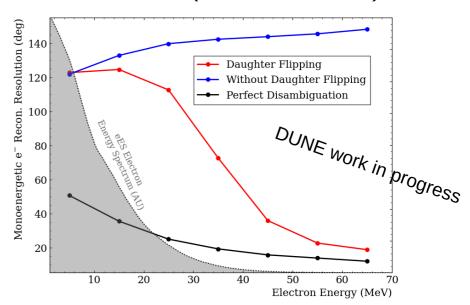
Daughter flipping:

mean $cos(\theta_i) > mean cos(\theta_i)$

→ resolving head-tail direction disambiguties



Performance (from simulations):



test daughter flipping performance with Michel electrons (from cosmic muons) in **ICEBERG** (~1t LArTPC with Dune electronics)

- → study daughter flipping with data in presence of noise
- → acquire test data set for ML applications
- → examine performance improvements with ML

poster by Joshua Queen

Low energy events

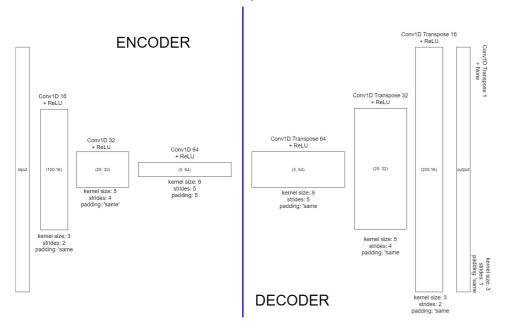
In cooperation with Michael H.L.S. Wang (Fermilab):

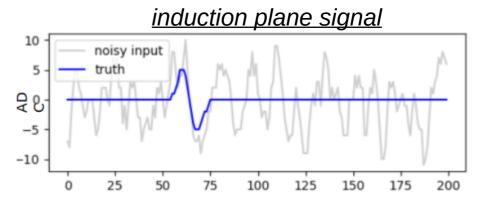
Extracting low energy signals from raw LArTPC waveforms using deep learning techniques — A proof of concept

Lorenzo Uboldi (CERN), David Ruth (Unlisted, US, IL), Michael Andrews (Carnegie Mellon U.), Michael H.L.S. Wang (Fermilab), Hans Joachim Wenzel (Fermilab) et al. (2021)

Published in: Nucl.Instrum.Meth.A 1028 (2022) 166371 • e-Print: 2106.09911 [physics.ins-det]

→ 1D-CNN => next step auto encoder





- improve LAr TPC detection efficiency at low energies (<1 MeV energy deposition)
- method to denoise input waveforms

poster by Van Tha Bik Lian

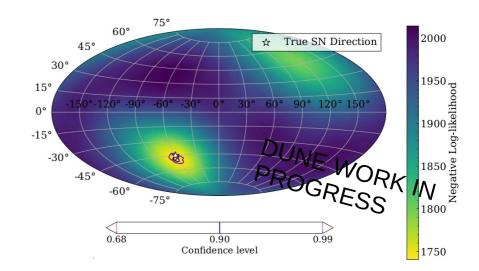
Summary and outlook

Multimessenger astronomy for supernova burst detection:

- Neutrino emission: high luminosity, fast, direct pointing back to source
 - → detect all the neutrinos to provide direction for el-mag. detection systems
- SNEWS early warning system

DUNE: 4 x 10kt LArTPC

- detection of electron neutrinos
- 3D imaging reconstruction
- Offline pointing resolution without ML:
 3.7 deg in 40kt (perfect channel tagging)
 5.0 deg in 40kt (4% ν₂CC as eES)



→ online pointing code and ML

- "kaizen": provide first result ~1min, improve precision over time
- set up complete pipeline, upgrade with ML: channel tagging, directional reconstruction and daughter flipping, low energy events, combination to burst,...