

L International Meeting
on Fundamental Physics
and XV CPAN Days
2 — 6 October 2023

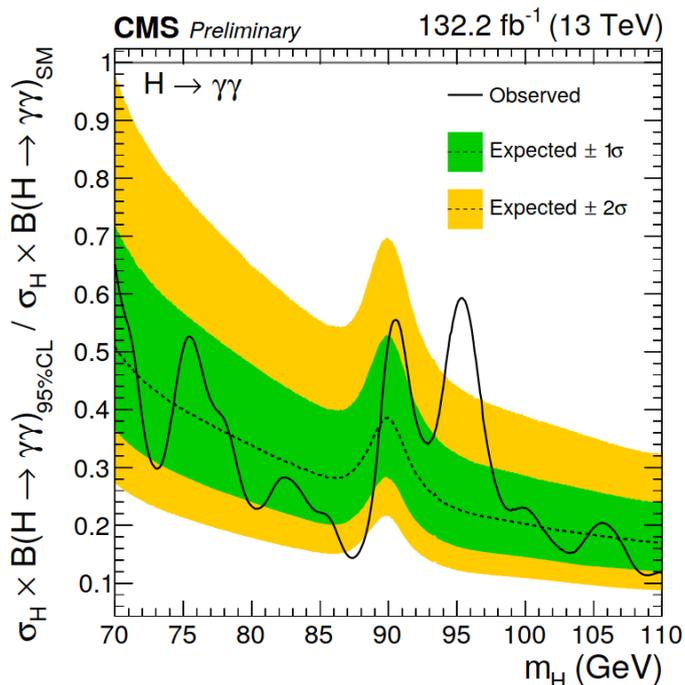
95 GeV excess within the scotogenic scenario

[based on 2306.03735 with A. Vicente and P. Escribano]

Víctor Martín Lozano

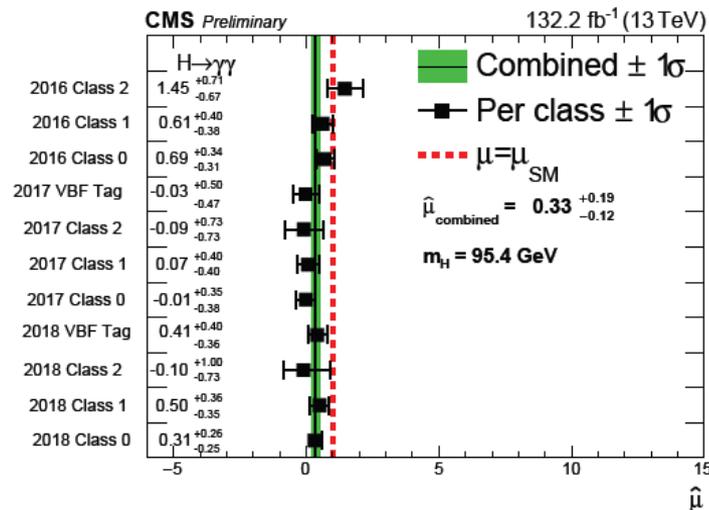
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The 95 GeV excess.

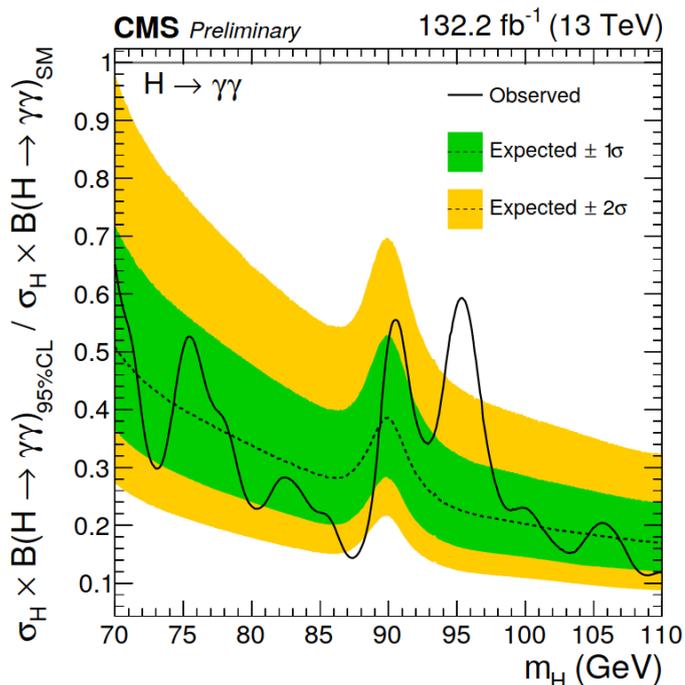


[CMS-PAS-HIG-20-002.]

- Run II data:
~ 2.9σ local significance
with signal strength:



The 95 GeV excess.

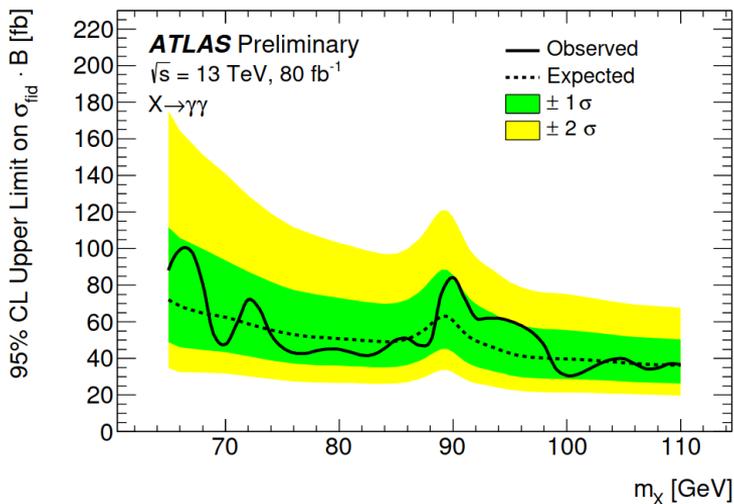


[CMS-PAS-HIG-20-002.]

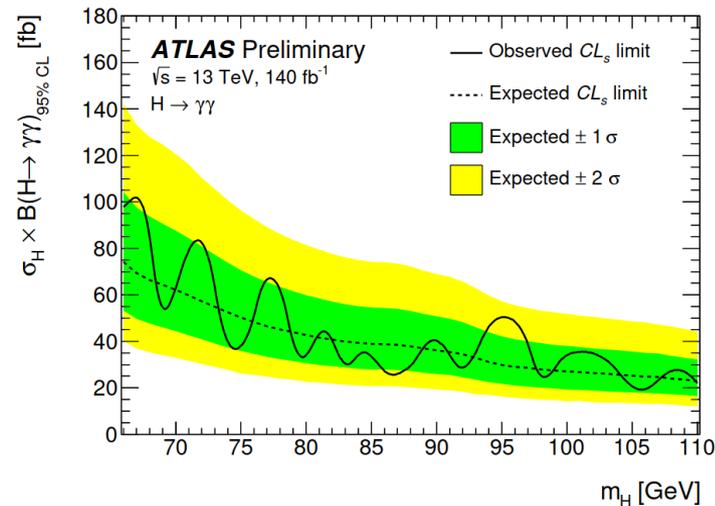
- Run II data:
~ 2.9σ local significance
with signal strength:

$$\mu_{\gamma\gamma}^{\text{CMS}} = \frac{\sigma^{\text{CMS}}(gg \rightarrow h_{95} \rightarrow \gamma\gamma)}{\sigma^{\text{SM}}(gg \rightarrow H \rightarrow \gamma\gamma)} = 0.33^{+0.19}_{-0.12}$$

The 95 GeV excess.



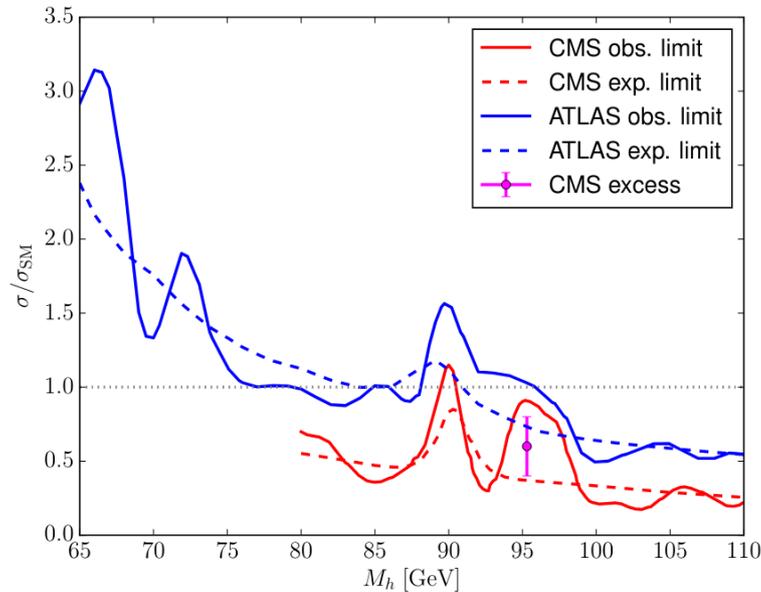
[ATLAS-CONF-2018-025]



[ATLAS-CONF-2023-035]

ATLAS shows a milder excess at the same mass range

The 95 GeV excess.

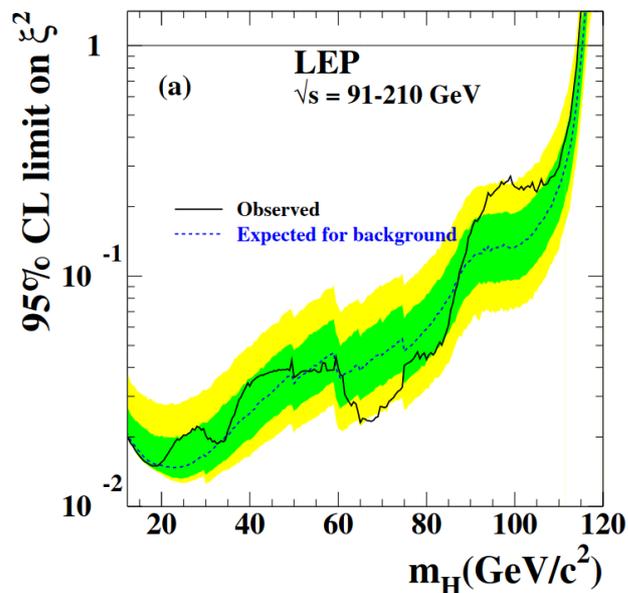


[S. Heinemeyer, T. Stefaniak 1812.05864]

However, the mild excess from ATLAS is in agreement with the stronger excess from CMS.

The 95 GeV excess.

(the origin)



[LEP collab. hep-ex/0306033]

All this comes from an excess of $\sim 2\sigma$ significance at LEP in the same range of masses.

$$\mu_{bb}^{\text{LEP}} = \frac{\sigma^{\text{LEP}}(e^+e^- \rightarrow Z h_{95} \rightarrow Z b\bar{b})}{\sigma^{\text{SM}}(e^+e^- \rightarrow Z H \rightarrow Z b\bar{b})} = 0.117 \pm 0.057$$

There are also hints in other channels:

$$\mu_{\tau\tau}^{\text{CMS}} = 1.2 \pm 0.5 \quad [\text{CMS 2208.02717}]$$

(No results from ATLAS so far)

The 95 GeV excess.

(the ambulance)

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The Scotogenic model.

σκότος = darkness

-γενής = that produces or generates

	gen	SU(2) _L	U(1) _Y	Z ₂
η	1	2	1/2	—
N	3	1	0	—

← One inert doublet

← Singlet fermion

[Ma, hep-ph/0601225]

(Dark Matter candidates due to the Z₂ symmetry)

$$\mathcal{L}_N = \overline{N}_i \not{\partial} N_i - \frac{M_{R_i}}{2} \overline{N}_i^c N_i + y_{i\alpha} \eta \overline{N}_i \ell_\alpha + \text{h.c.}$$

$$\mathcal{V} = m_H^2 H^\dagger H + m_\eta^2 \eta^\dagger \eta + \frac{\lambda_1}{2} (H^\dagger H)^2 + \frac{\lambda_2}{2} (\eta^\dagger \eta)^2 + \lambda_3 (H^\dagger H) (\eta^\dagger \eta) \\ + \lambda_4 (H^\dagger \eta) (\eta^\dagger H) + \frac{\lambda_5}{2} \left[(H^\dagger \eta)^2 + (\eta^\dagger H)^2 \right]$$

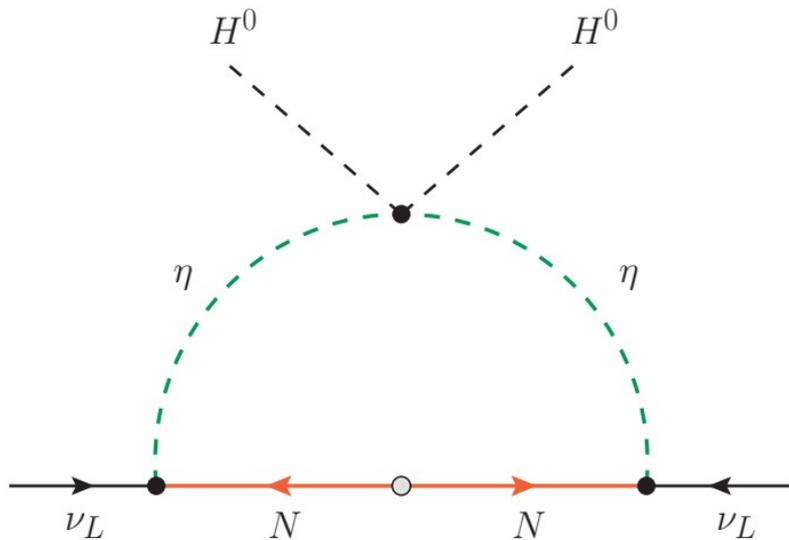
The Scotogenic model.

σκότος = darkness

-γενής = that produces or generates

(but also that is produced or generated by)

It generates neutrino masses at 1-loop level due to the dark particles in the loop.



$$m_\nu = \frac{\lambda_5 v^2}{32\pi^2} y^T M_R^{-1} f_{\text{loop}} y$$

The Scotogenic model.

$$\mathcal{L} \supset y_{na\alpha} \bar{N}_n \tilde{\eta}_a^\dagger \ell_L^\alpha - \kappa_{nm} S \bar{N}_n^c N_m - \frac{1}{2} (M_N)_{nn} \bar{N}_n^c N_n + \text{h.c.}$$

Field	Generations	SU(3) _c	SU(2) _L	U(1) _Y	\mathbb{Z}_2
ℓ_L	3	1	2	-1/2	+
e_R	3	1	1	-1	+
N	n_N	1	1	0	-
H	1	1	2	1/2	+
η	n_η	1	2	1/2	-
S	1	1	1	0	+

n_η generations of inert doublets

n_N generations of Singlet fermions

$$\mathcal{V}_H = m_H^2 H^\dagger H + \frac{1}{2} \lambda_1 (H^\dagger H)^2,$$

$$\mathcal{V}_\eta = (m_\eta^2)_{aa} \eta_a^\dagger \eta_a + \frac{1}{2} \lambda_2^{abcd} (\eta_a^\dagger \eta_b) (\eta_c^\dagger \eta_d), \quad \mathcal{V}_S = \frac{1}{2} m_S^2 S^2 + \frac{1}{3} \mu_S S^3 + \frac{1}{4} \lambda_S S^4,$$

$$\begin{aligned} \mathcal{V}_{\text{mix}} = & \lambda_3^{ab} (H^\dagger H) (\eta_a^\dagger \eta_b) + \lambda_4^{ab} (H^\dagger \eta_a) (\eta_b^\dagger H) + \frac{1}{2} \left[\lambda_5^{ab} (H^\dagger \eta_a) (H^\dagger \eta_b) + \text{h.c.} \right] \\ & + \mu_H H^\dagger H S + \frac{1}{2} \lambda_{HS} H^\dagger H S^2 + \mu_\eta^{ab} \eta_a^\dagger \eta_b S + \frac{1}{2} \lambda_{\eta S}^{ab} \eta_a^\dagger \eta_b S^2. \end{aligned}$$

Interpretation of the 95 GeV.

$$\langle H^0 \rangle = \frac{v}{\sqrt{2}}, \quad \langle \eta_a^0 \rangle = 0, \quad \langle S \rangle = v_S$$

$$\mathcal{M}_{\mathcal{H}}^2 = \begin{pmatrix} v_S (\mu_S + 2\lambda_S v_S) - \frac{\mu_H v^2}{2v_S} & \mu_H v + \lambda_{HS} v v_S \\ \mu_H v + \lambda_{HS} v v_S & \lambda_1 v^2 \end{pmatrix}$$

$$V_{\mathcal{H}} \mathcal{M}_{\mathcal{H}}^2 V_{\mathcal{H}}^T = \widehat{\mathcal{M}}_{\mathcal{H}}^2 = \text{diag} (m_{h_1}^2, m_{h_2}^2)$$

$$V_{\mathcal{H}} = \begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix}$$

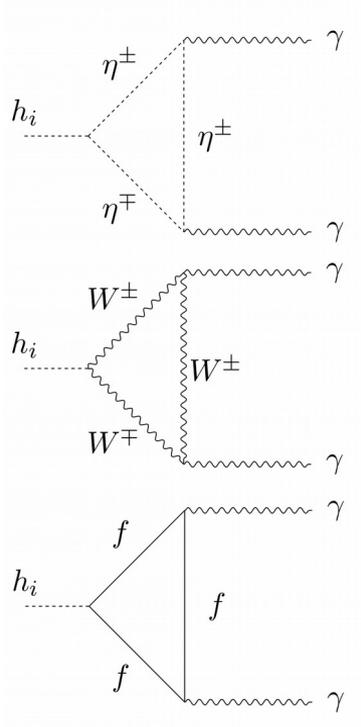
After EWSB takes place both H and S acquire a VEV and the physical states are an admixture of them.

After diagonalizing the mass matrix we can identify two states:

$$h_1 \equiv h_{95}, \quad h_2 \equiv h_{125}.$$

Interpretation of the 95 GeV.

$$h_1 \equiv h_{95}, \quad h_2 \equiv h_{125}.$$



Due to the presence of the inert doublets there is an additional channel for the $\gamma\gamma$ decay.

$$\Gamma(h_i \rightarrow \gamma\gamma) = \frac{G_F \alpha^2 m_{h_i}^3}{128 \sqrt{2} \pi^3} \left| \sum_f N_c Q_f^2 g_{h_i f f} A_{1/2}(\tau_f) + g_{h_i W W} A_1(\tau_W) + \sum_a \frac{v}{2m_{\eta_a}^2} g_{h_i \eta \eta}^{aa} A_0(\tau_\eta) \right|^2$$

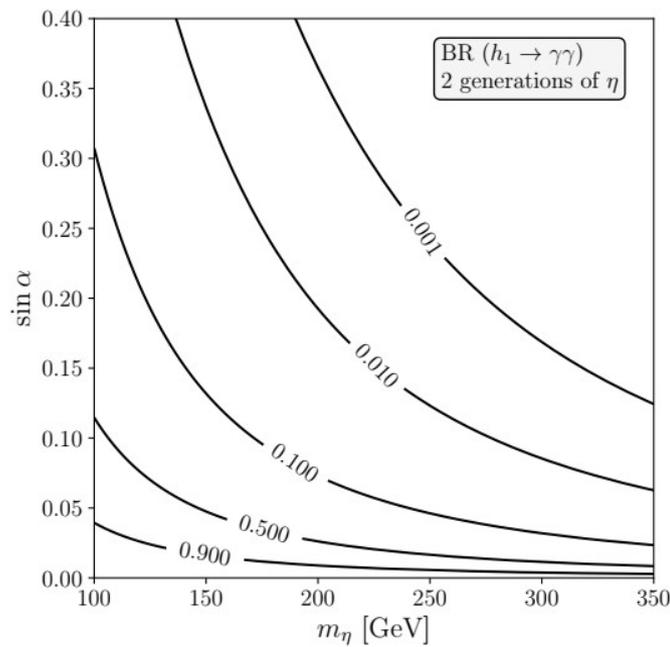
This decay will be proportional to the couplings to the doublets:

$$g_{h_1 \eta \eta} = \cos \alpha (\lambda_{\eta S} v_S + \mu_\eta) + \sin \alpha \lambda_3 v,$$

$$g_{h_2 \eta \eta} = -\sin \alpha (\lambda_{\eta S} v_S + \mu_\eta) + \cos \alpha \lambda_3 v.$$

Interpretation of the 95 GeV.

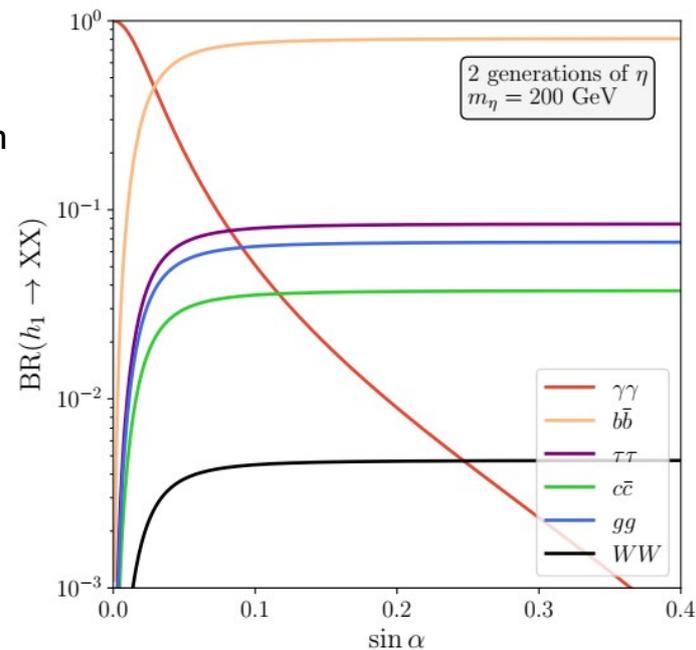
$$h_1 \equiv h_{95}, \quad h_2 \equiv h_{125}.$$



The lower the mixing the highest the decay rate into $\gamma\gamma$ decay.

$$g_{h_1\eta\eta} = \cos\alpha (\lambda_{\eta S} v_S + \mu_\eta) + \sin\alpha \lambda_3 v,$$

$$g_{h_2\eta\eta} = -\sin\alpha (\lambda_{\eta S} v_S + \mu_\eta) + \cos\alpha \lambda_3 v.$$



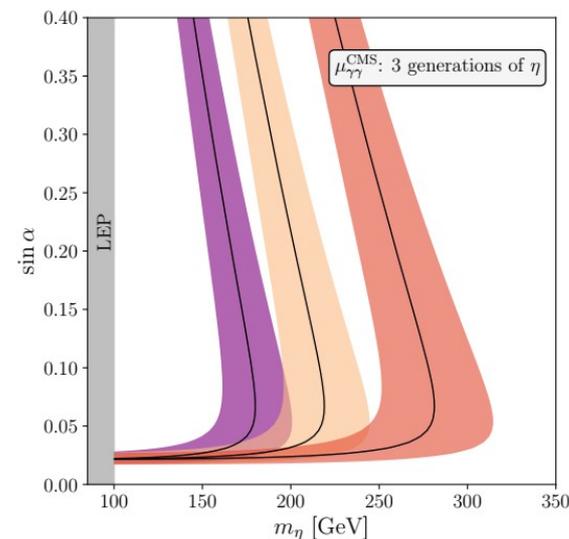
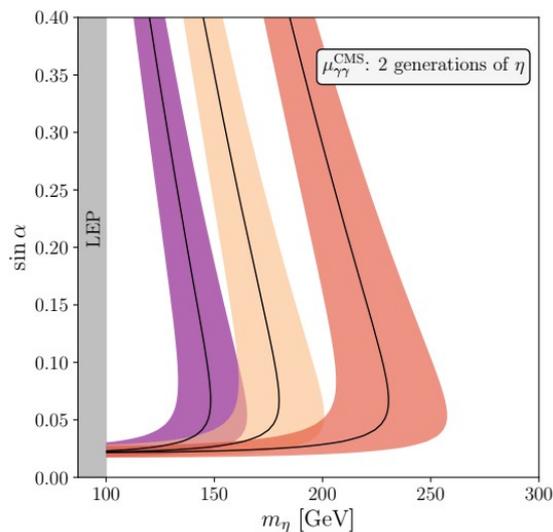
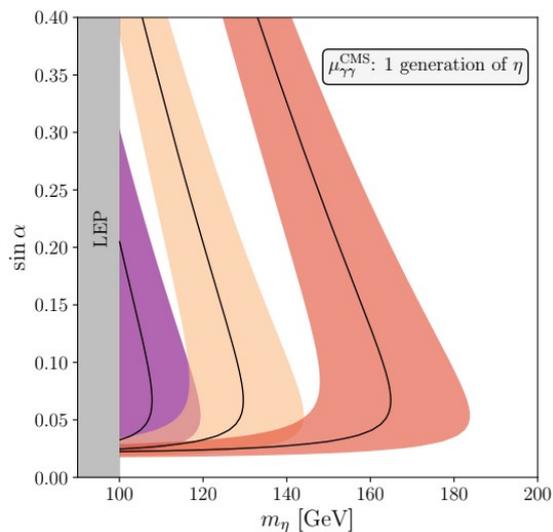
Numerical Results.

$$\bar{\lambda}_{\eta S} v_S = 500 \text{ GeV}, \bar{\lambda}_3 = 0.6,$$

$$\bar{\mu}_\eta = 500 \text{ GeV (purple)}, 1000 \text{ GeV (yellow)}, 2000 \text{ GeV (red)}$$

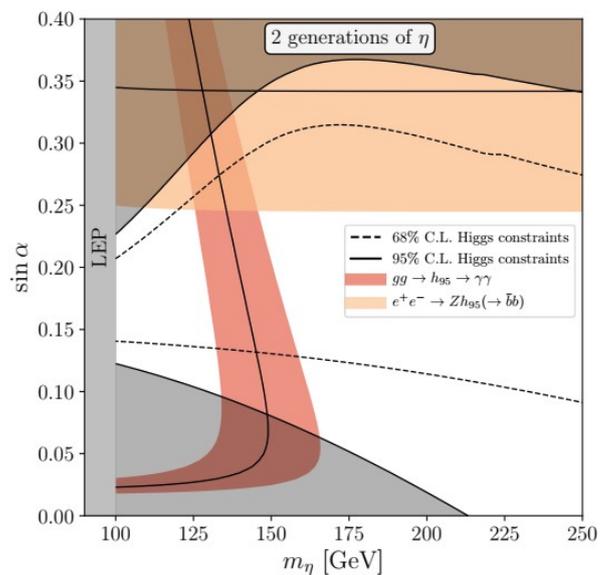
$$\mu_{\gamma\gamma} = \frac{\sigma(gg \rightarrow h_1)}{\sigma_{\text{SM}}(gg \rightarrow H)} \times \frac{\text{BR}(h_1 \rightarrow \gamma\gamma)}{\text{BR}_{\text{SM}}(H \rightarrow \gamma\gamma)} = \sin^2 \alpha \frac{\text{BR}(h_1 \rightarrow \gamma\gamma)}{\text{BR}_{\text{SM}}(H \rightarrow \gamma\gamma)}$$

$$\mu_{\gamma\gamma}^{\text{CMS}} = \frac{\sigma^{\text{CMS}}(gg \rightarrow h_{95} \rightarrow \gamma\gamma)}{\sigma^{\text{SM}}(gg \rightarrow H \rightarrow \gamma\gamma)} = 0.33^{+0.19}_{-0.12}$$

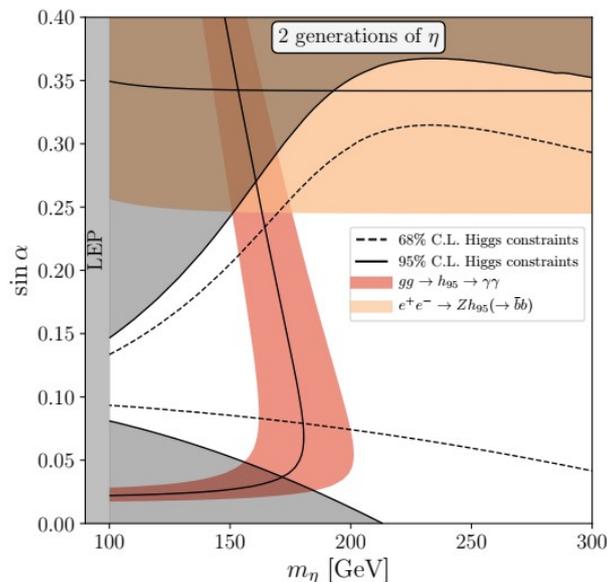


Numerical Results.

$$\bar{\lambda}_{\eta S} v_S = 500 \text{ GeV}, \bar{\lambda}_3 = 0.6,$$



$$\bar{\mu}_\eta = 500 \text{ GeV}$$



$$\bar{\mu}_\eta = 1000 \text{ GeV}$$

$$\mu_{\gamma\gamma}^{\text{CMS}} = \frac{\sigma^{\text{CMS}}(gg \rightarrow h_{95} \rightarrow \gamma\gamma)}{\sigma^{\text{SM}}(gg \rightarrow H \rightarrow \gamma\gamma)} = 0.33^{+0.19}_{-0.12}$$

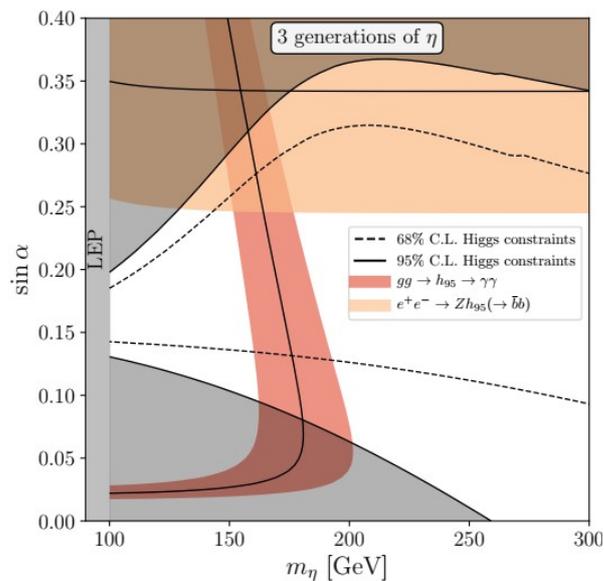
$$\mu_{bb}^{\text{LEP}} = \frac{\sigma^{\text{LEP}}(e^+e^- \rightarrow Zh_{95} \rightarrow Zb\bar{b})}{\sigma^{\text{SM}}(e^+e^- \rightarrow ZH \rightarrow Zb\bar{b})} = 0.117 \pm 0.057$$

HiggsTools (Higgs Constraints)

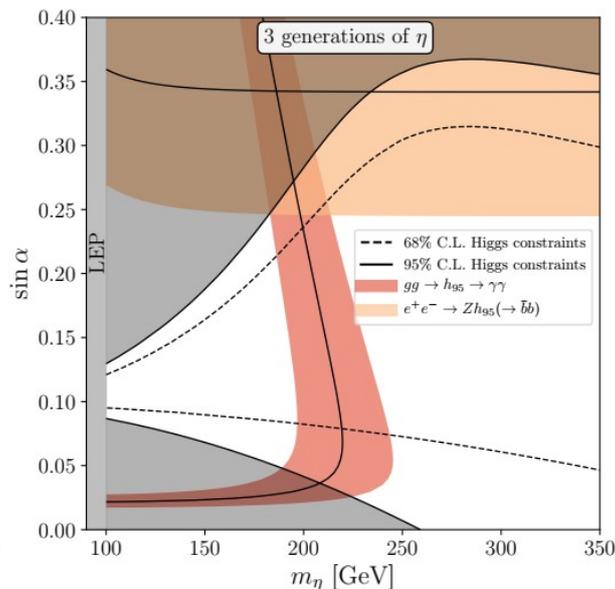
[Bahl, Biekötter, Heinemeyer, Li, Paasch,
Weiglein, Wittbrodt, 2210.09332]

Numerical Results.

$$\bar{\lambda}_{\eta S} v_S = 500 \text{ GeV}, \bar{\lambda}_3 = 0.6,$$



$$\bar{\mu}_\eta = 500 \text{ GeV}$$



$$\bar{\mu}_\eta = 1000 \text{ GeV}$$

$$\mu_{\gamma\gamma}^{\text{CMS}} = \frac{\sigma^{\text{CMS}}(gg \rightarrow h_{95} \rightarrow \gamma\gamma)}{\sigma^{\text{SM}}(gg \rightarrow H \rightarrow \gamma\gamma)} = 0.33^{+0.19}_{-0.12}$$

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HiggsTools (Higgs Constraints)

[Bahl, Biekötter, Heinemeyer, Li, Paasch,
Weiglein, Wittbrodt, 2210.09332]

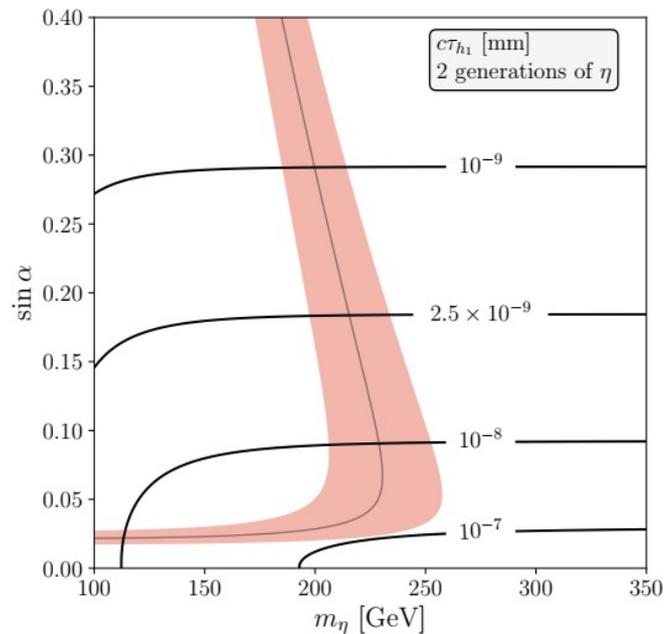
Conclusions.

- There are hints of a possible particle at a mass around ~ 95.4 GeV
- The scotogenic model is a phenomenologically viable model that predicts dark matter and the small neutrino masses.
- A generalization of the scotogenic model can explain the 95 GeV excess for different values of the parameter space.
- Different number of generations of inert doublets and singlet fermions can accommodate the excess.
- More data will be needed in the low mass region and different channels.

Thank you!

Decay width of the 95 GeV.

$$h_1 \equiv h_{95}, \quad h_2 \equiv h_{125}.$$



One may ask if the light scalar is long-lived for small values of the mixing angle.

However, for the region where the excess can be explained the lighter scalar decays promptly.