



cherenkov
telescope
array



Machine Learning in LST: advanced trigger and sensitivity

Tjark Miener (tjark.miener@unige.ch)



UNIVERSITÉ
DE GENÈVE

FACULTÉ DES SCIENCES

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CTLearn community

- **CTLearn** is a high-level Python package for using **Deep Learning** for **IACT event reconstruction** maintained by Nieto D. (Madrid) and Miener T. (UniGe)
- Perez A. (PhD student at Madrid), Burmistrov L. (UniGe) and Abellan Beteta C., Bezshyiko Ia., Hijano Mendizabal G., Serra N. (Uni Zurich) joined the team for the **AI Trigger system** project. Expertise on **porting** neural networks to **FPGAs**.
- Lacave B. (PhD student at UniGe with Heller M.) started to work on the **CTLearn** analysis pipeline for the **SST1M** and **LST-1** prototypes.



DOI [10.5281/zenodo.3345947](https://doi.org/10.5281/zenodo.3345947)

<https://github.com/ctlearn-project/ctlearn>

<https://ctlearn.readthedocs.io>



Come and join us

- Quick introduction to IACT low-level analysis with Deep Learning
- Ongoing performance studies of LSTs equipped with high-resolution advanced SiPM cameras using CNNs
- Developments on the AI Trigger system of the adv. SiPM camera
- CTLearn at CSCS

IACT low-level analysis

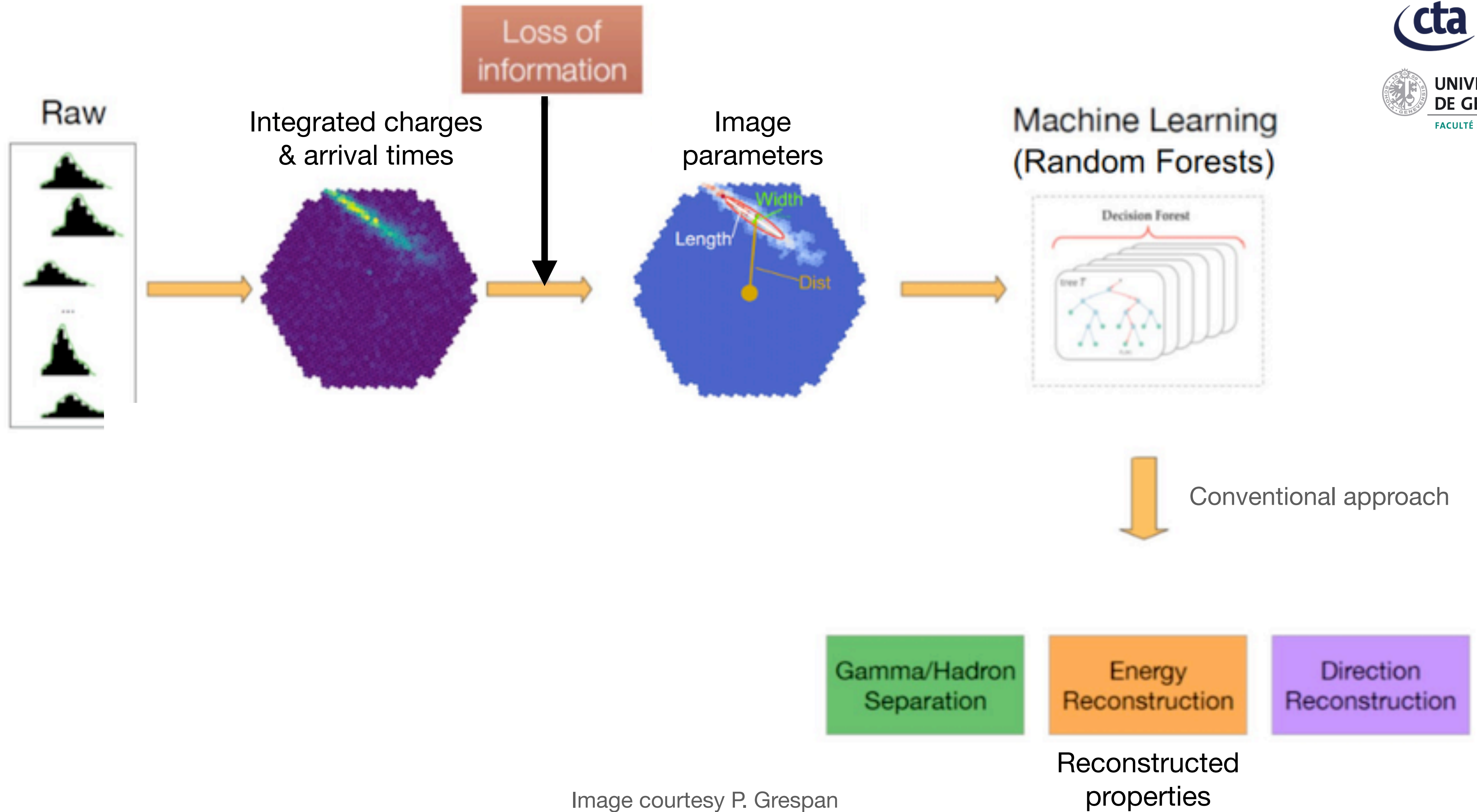
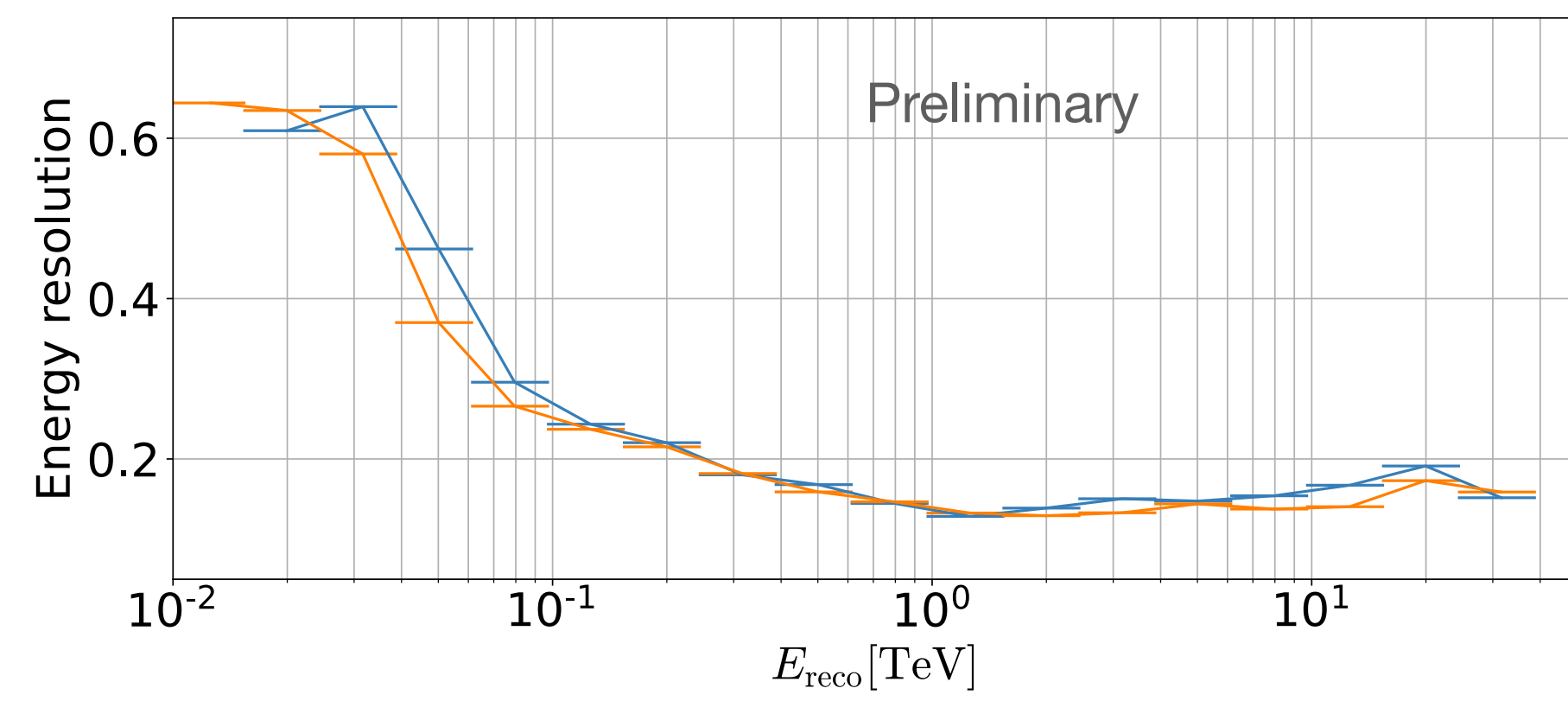
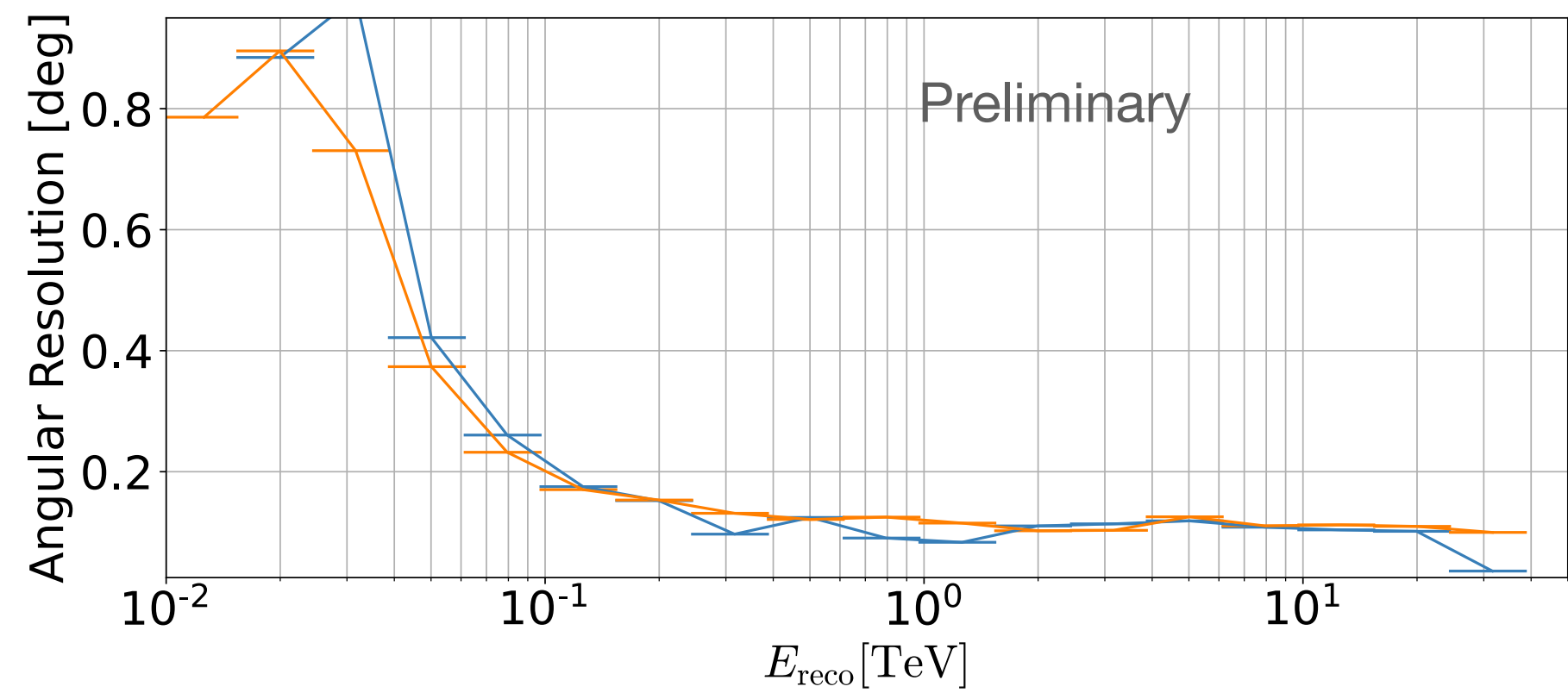
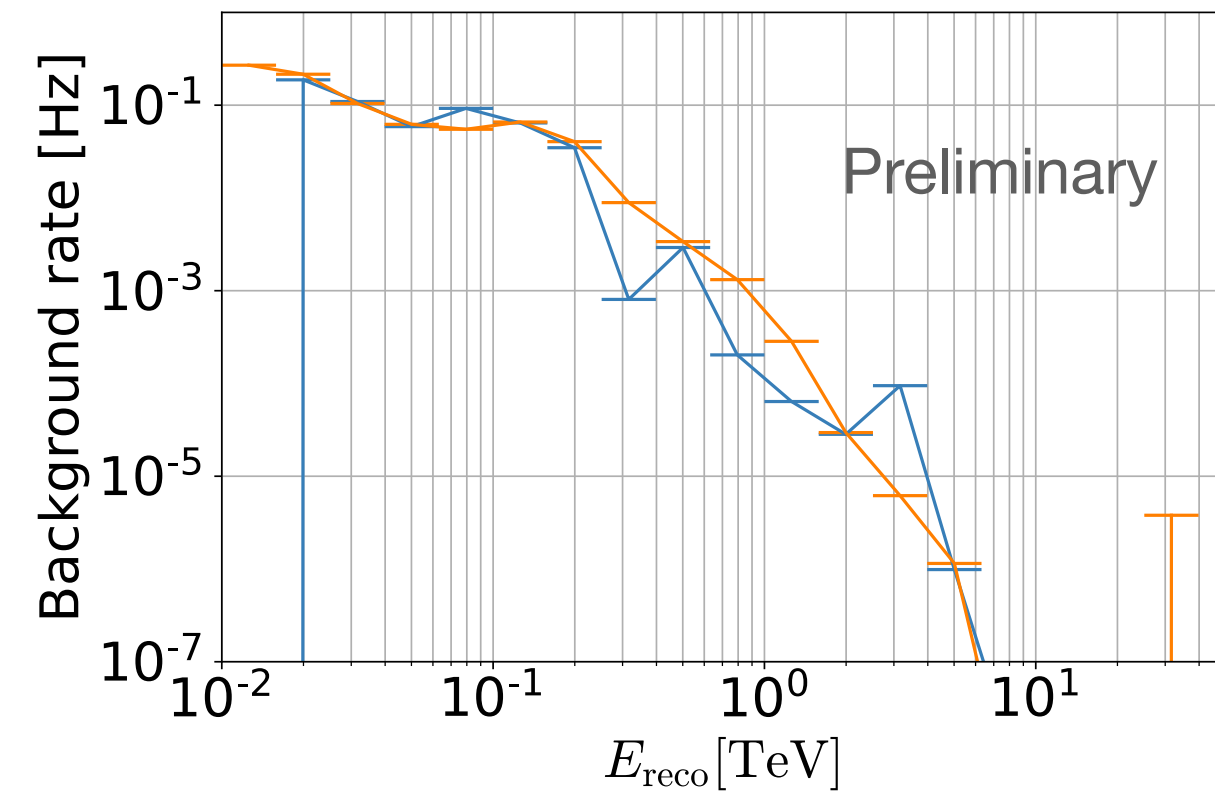
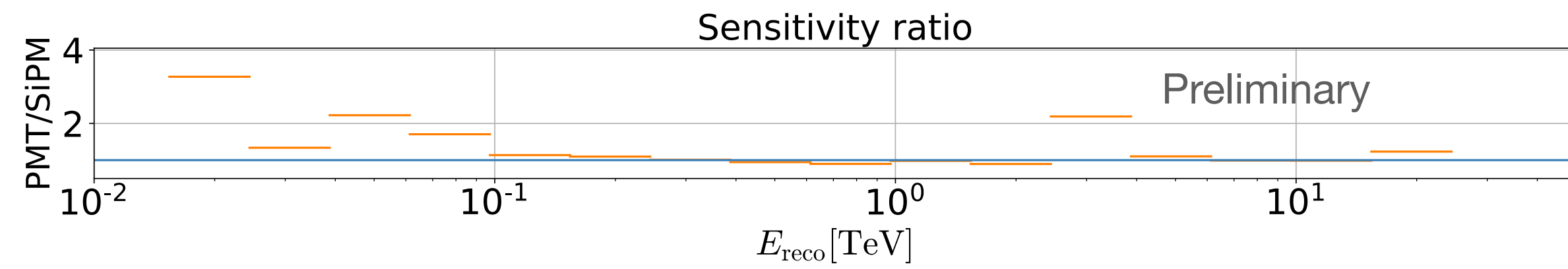
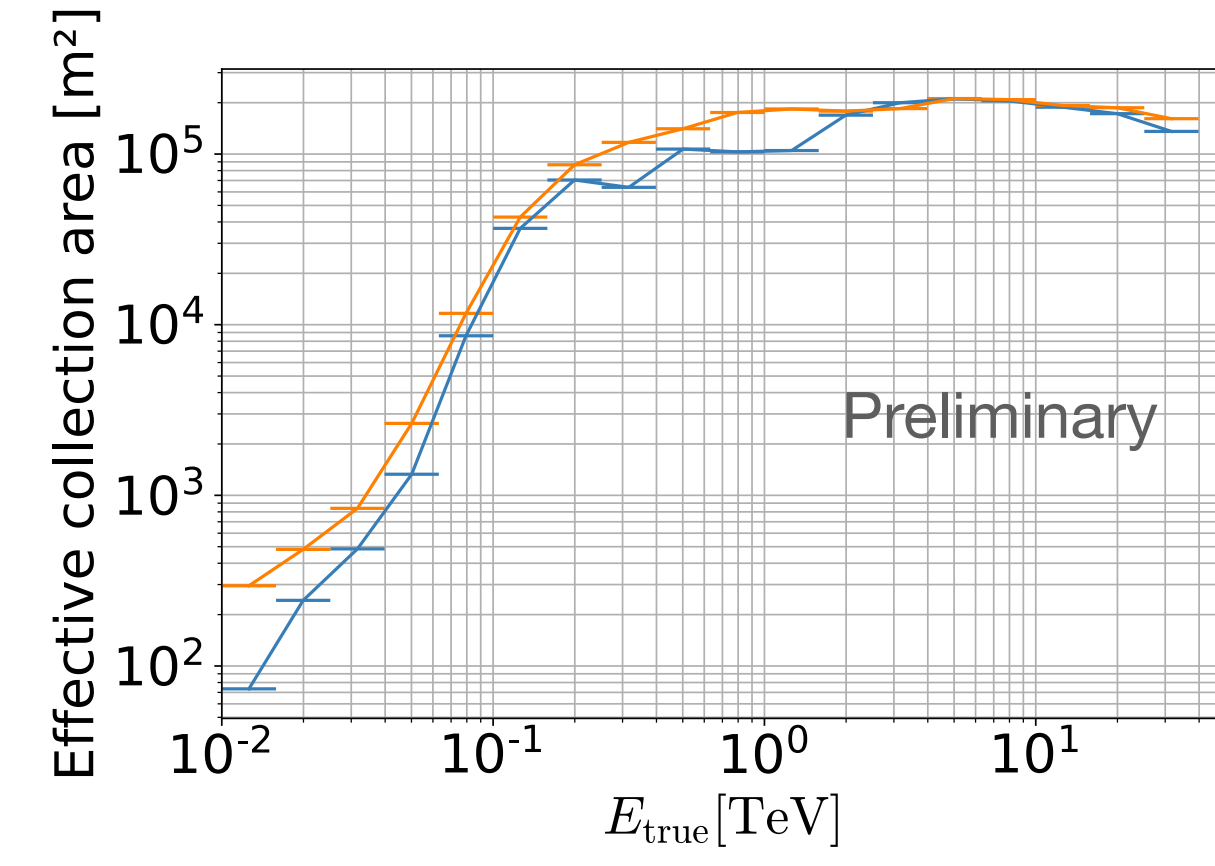
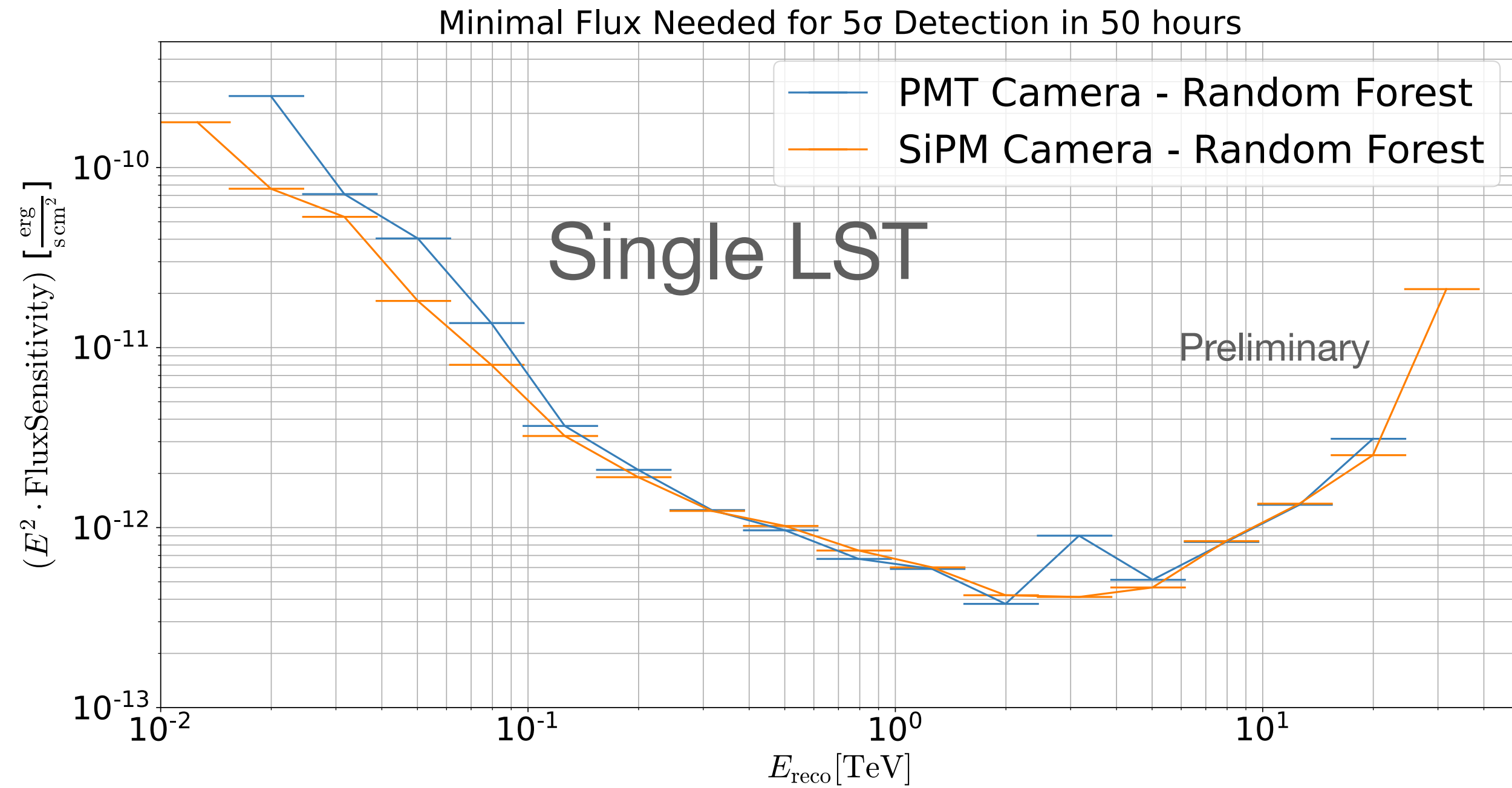
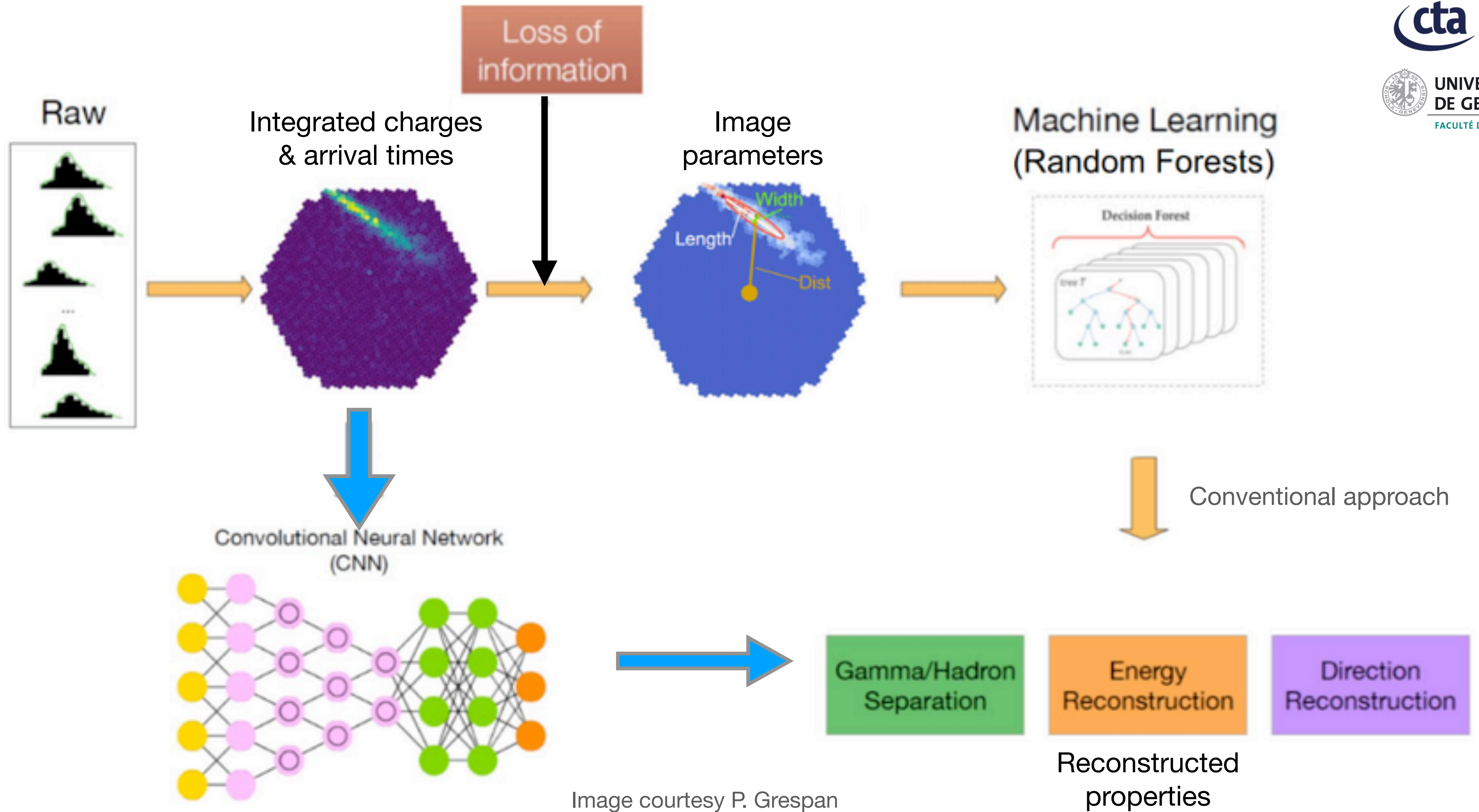


Image courtesy P. Grespan

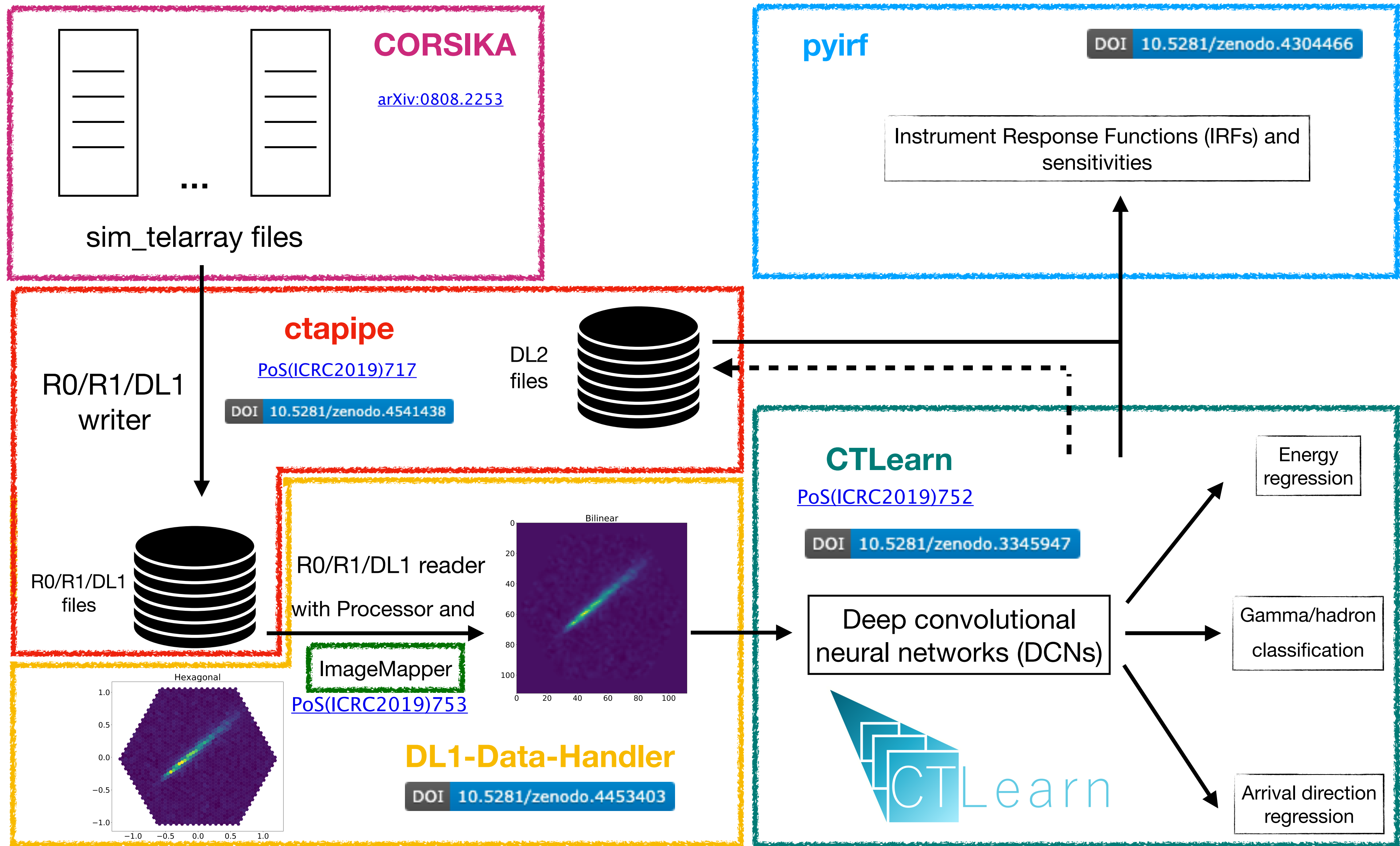
Random Forest: current PMT camera vs high-resolution SiPM camera



IACT low-level analysis with Deep Learning

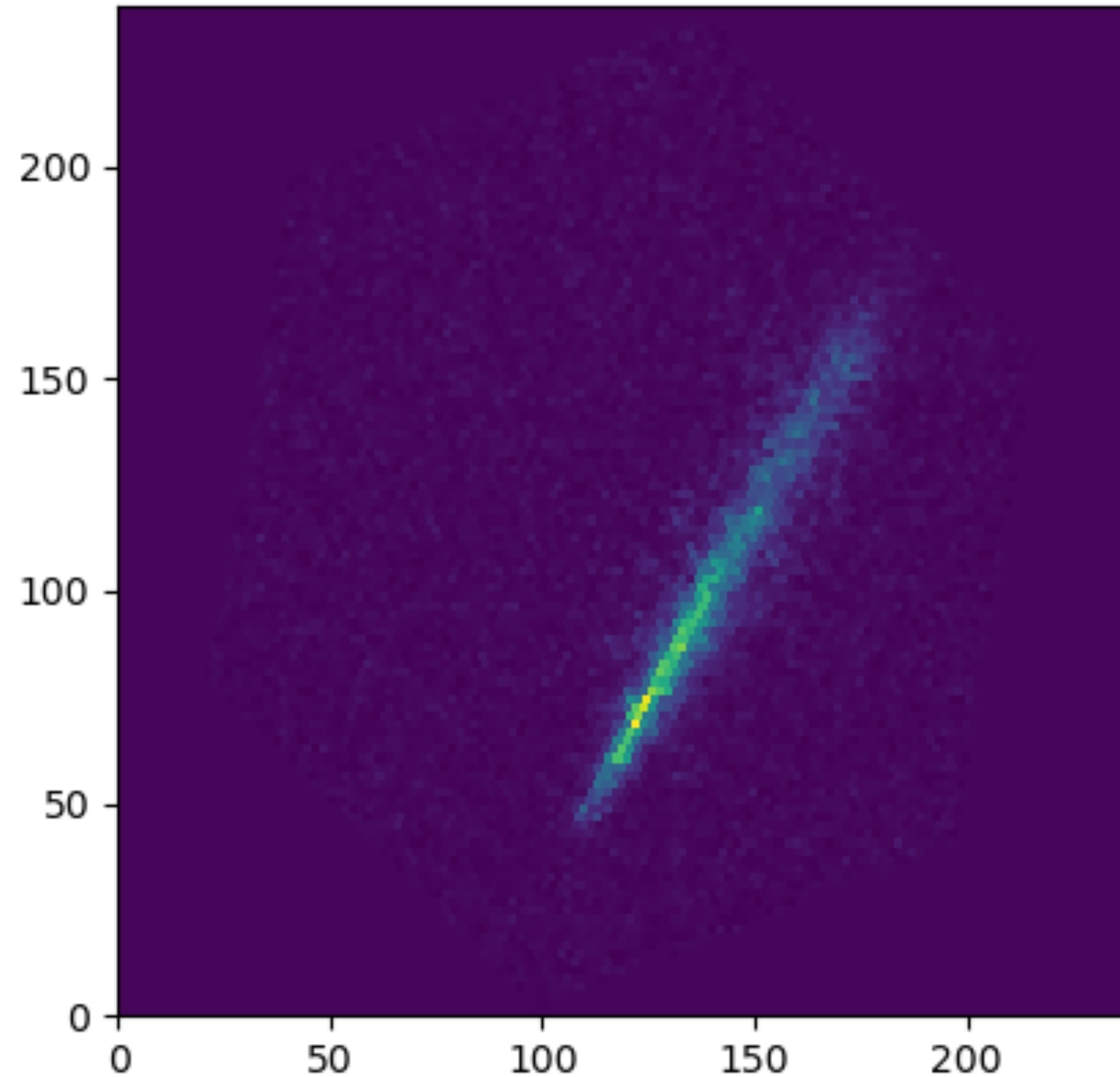


CTA analysis workflow with CTLearn

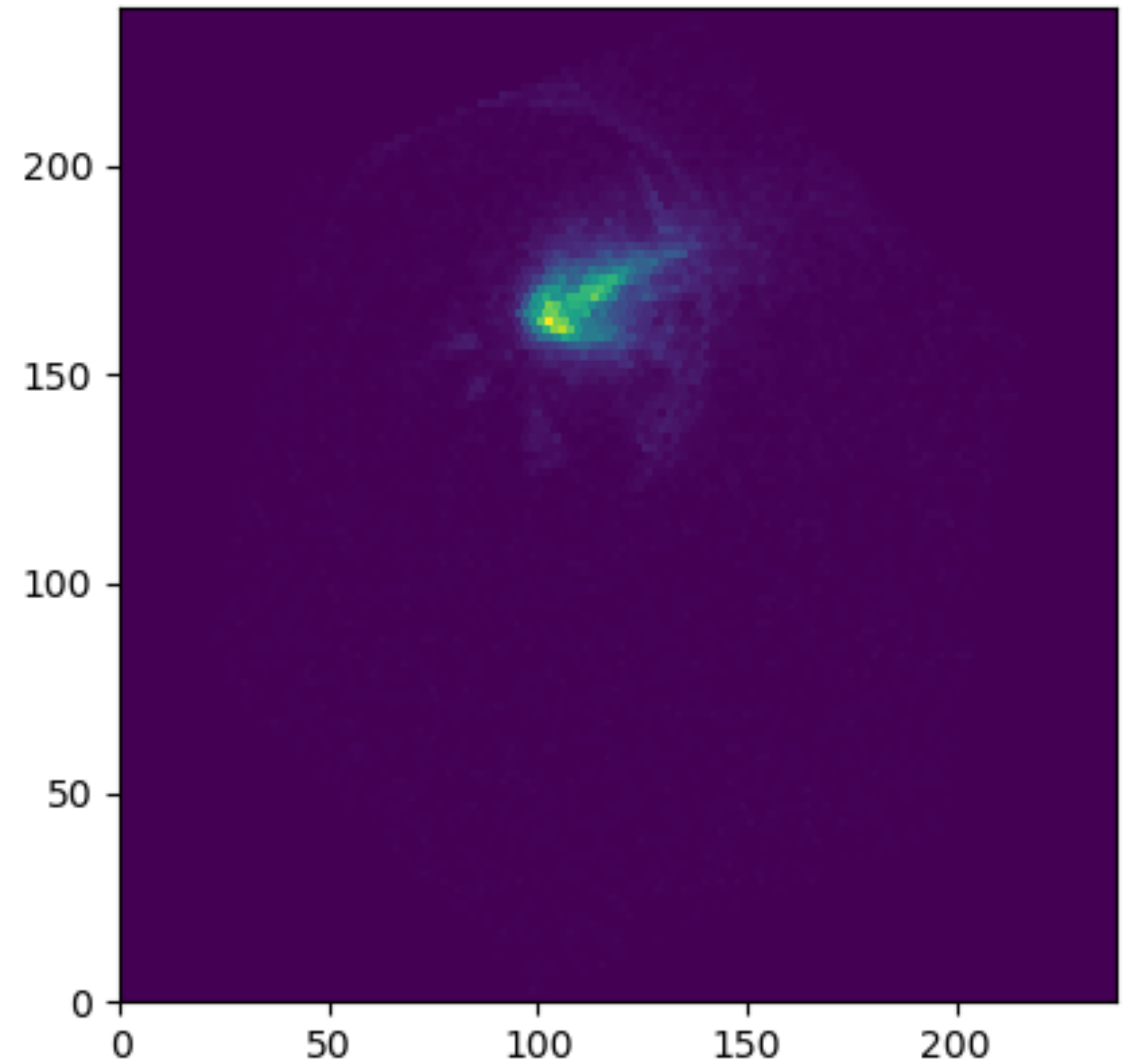


Shower images captured by the adv. LSTSiPM camera through the ImageMapper

Gamma

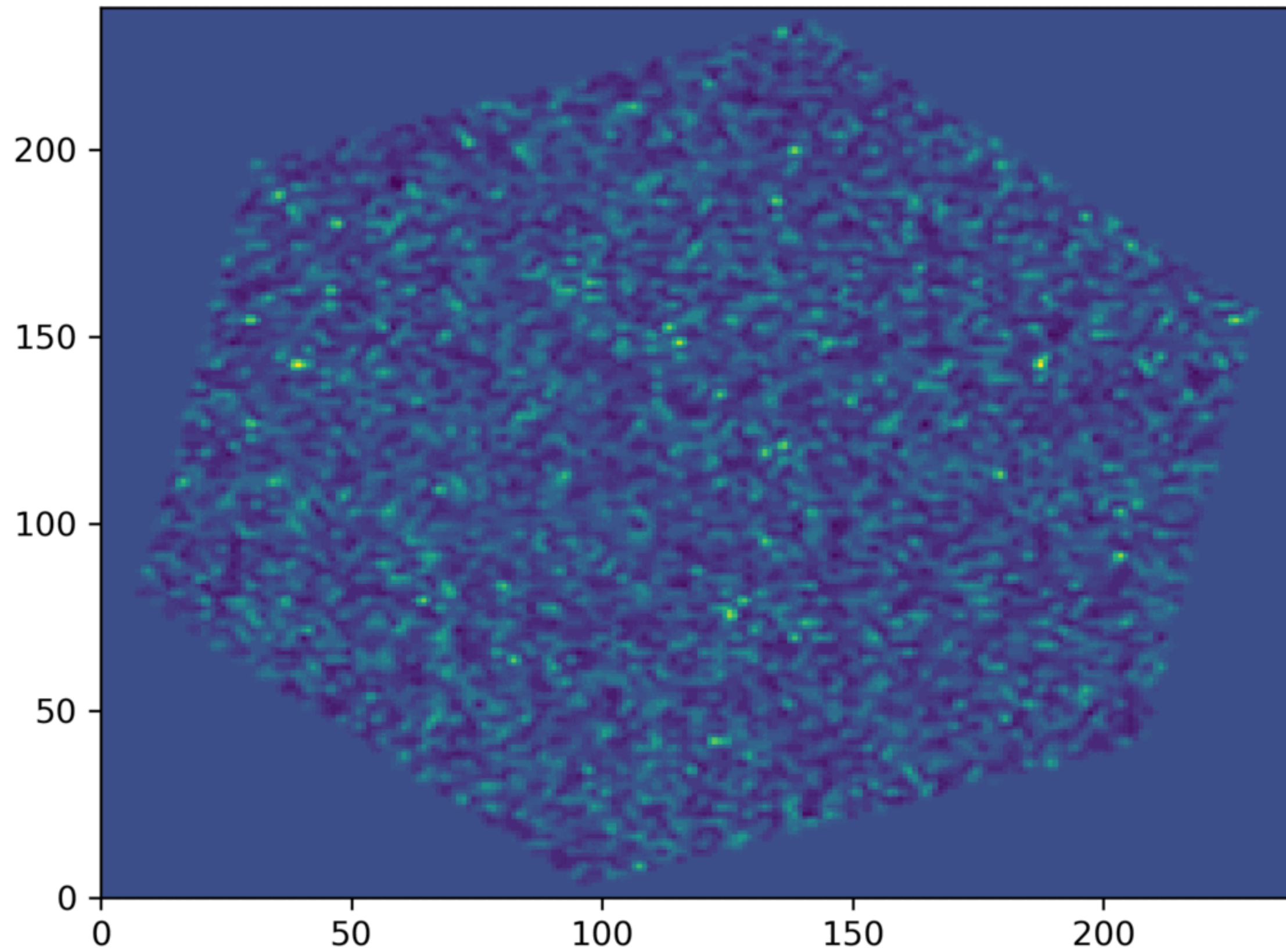


Proton

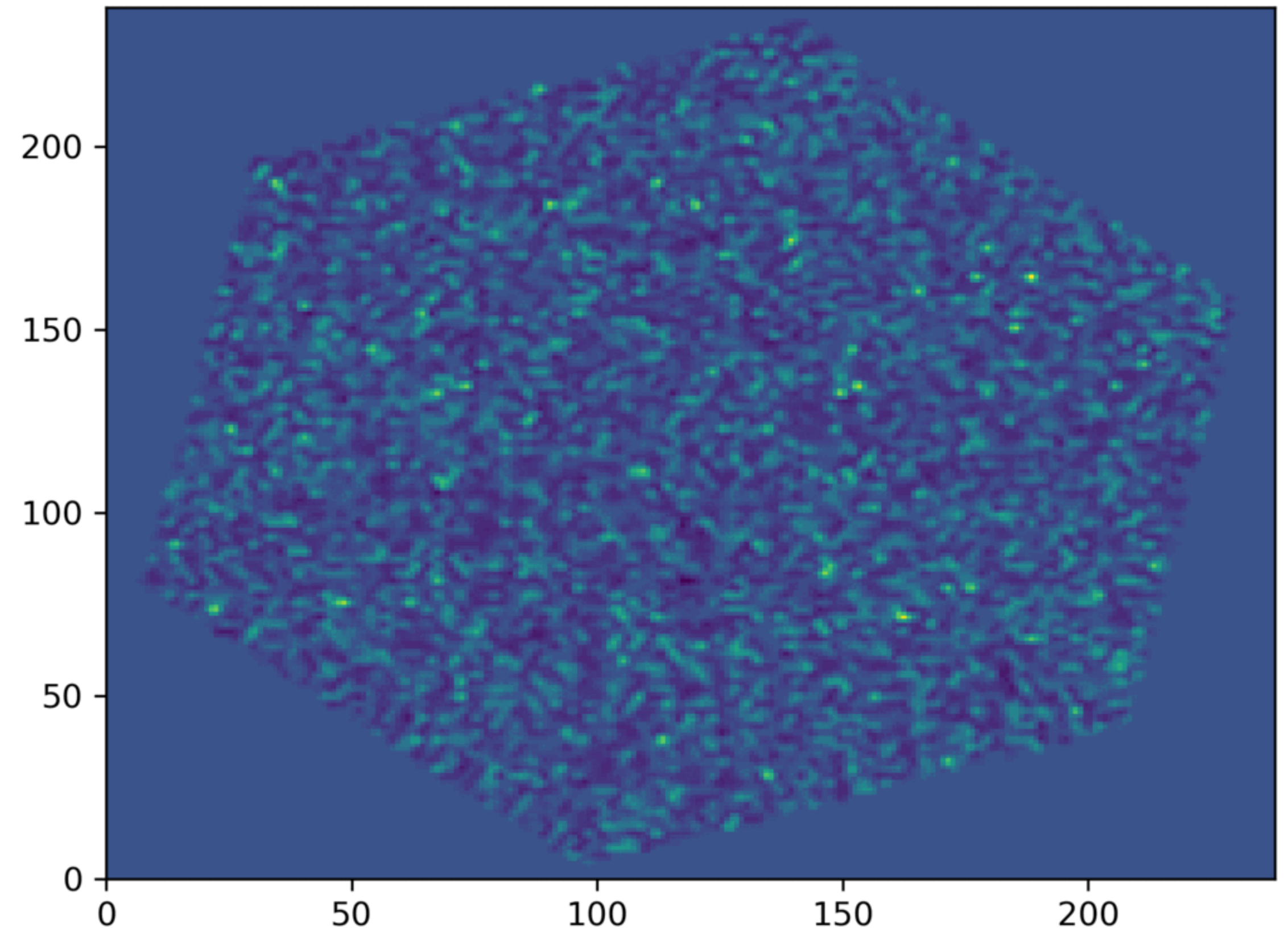


Calibrated waveforms captured by the adv. LSTSiPM camera through the ImageMapper

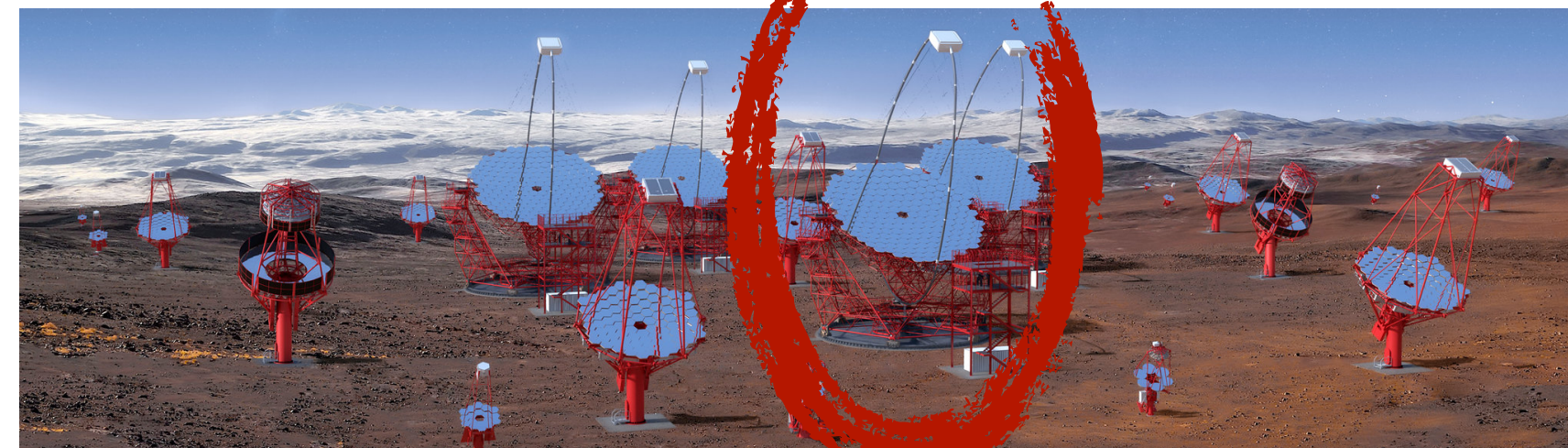
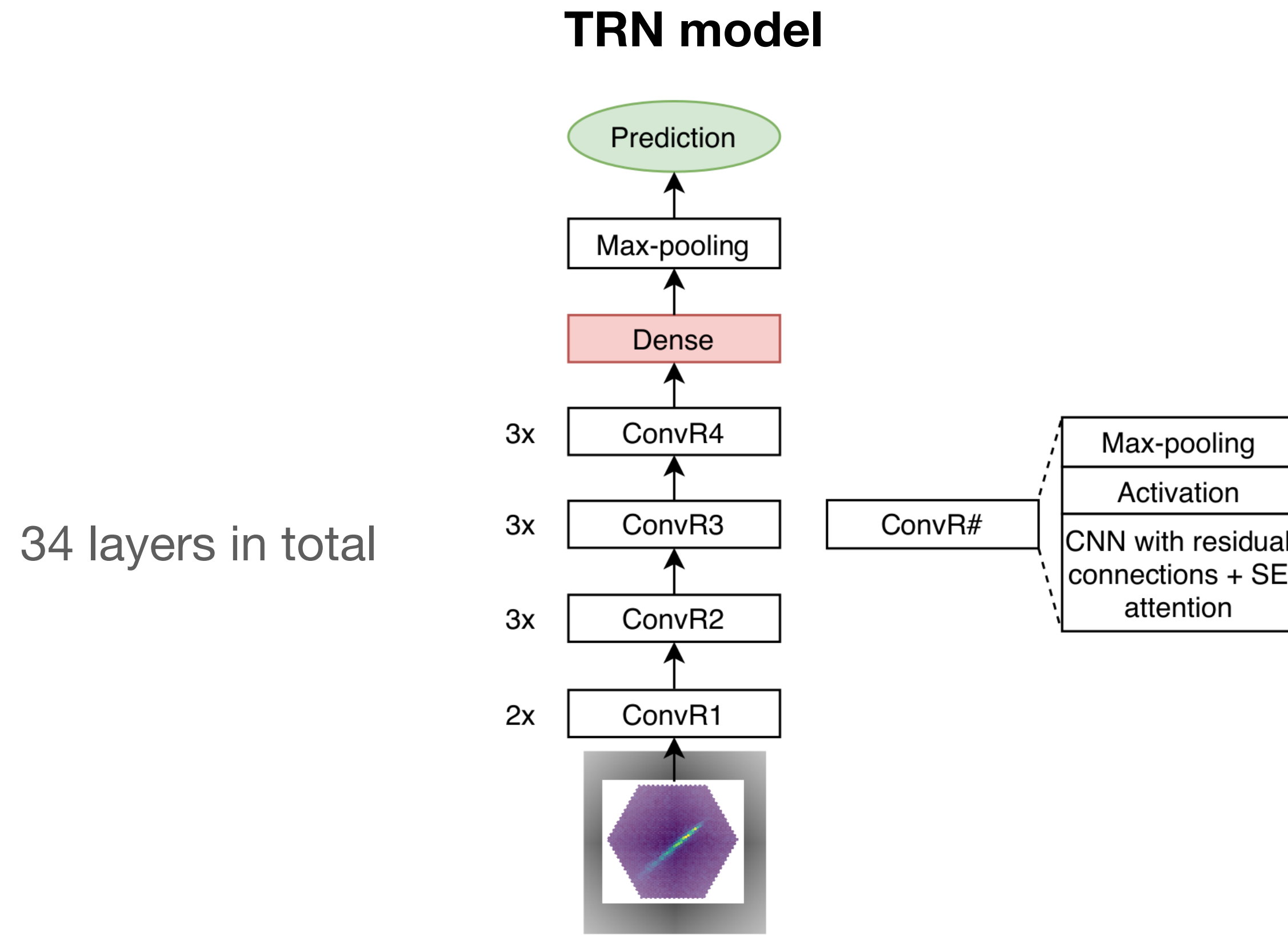
Gamma



Proton

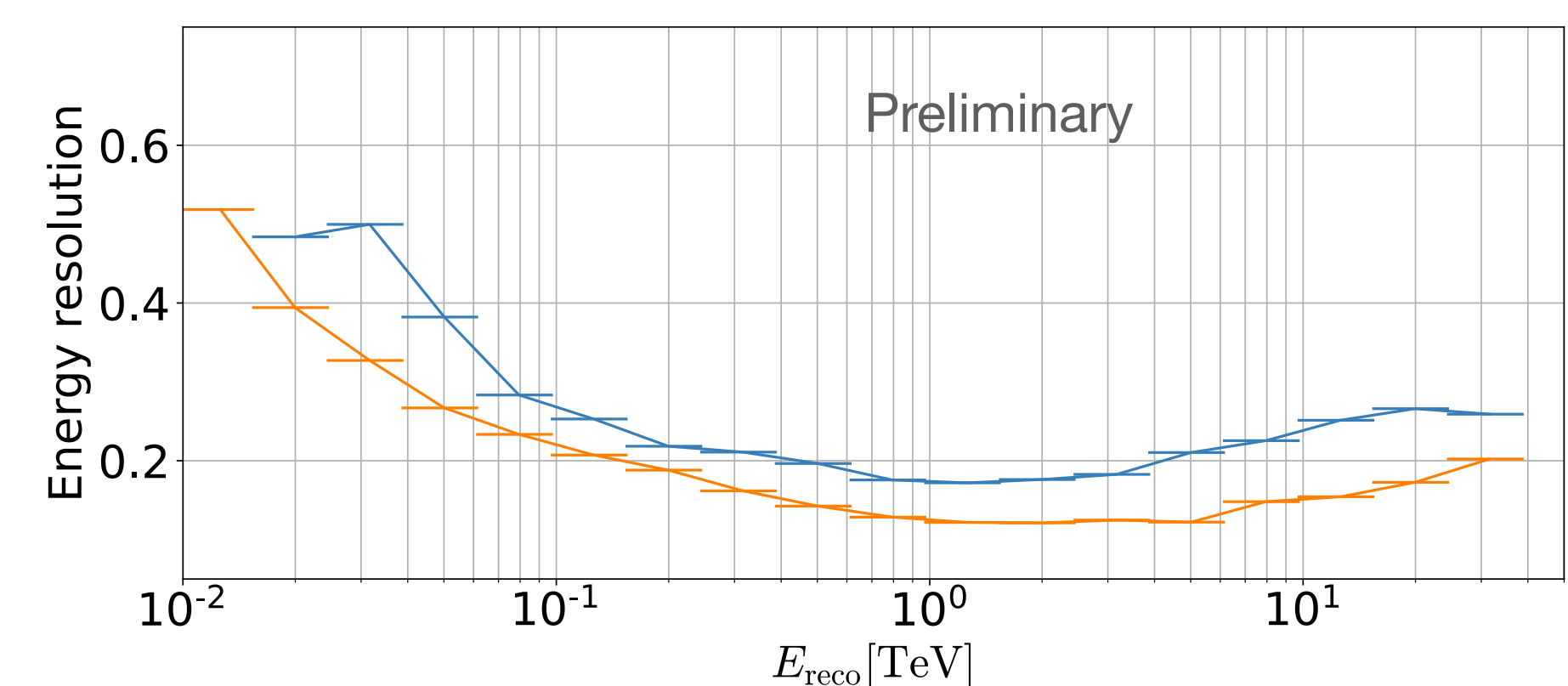
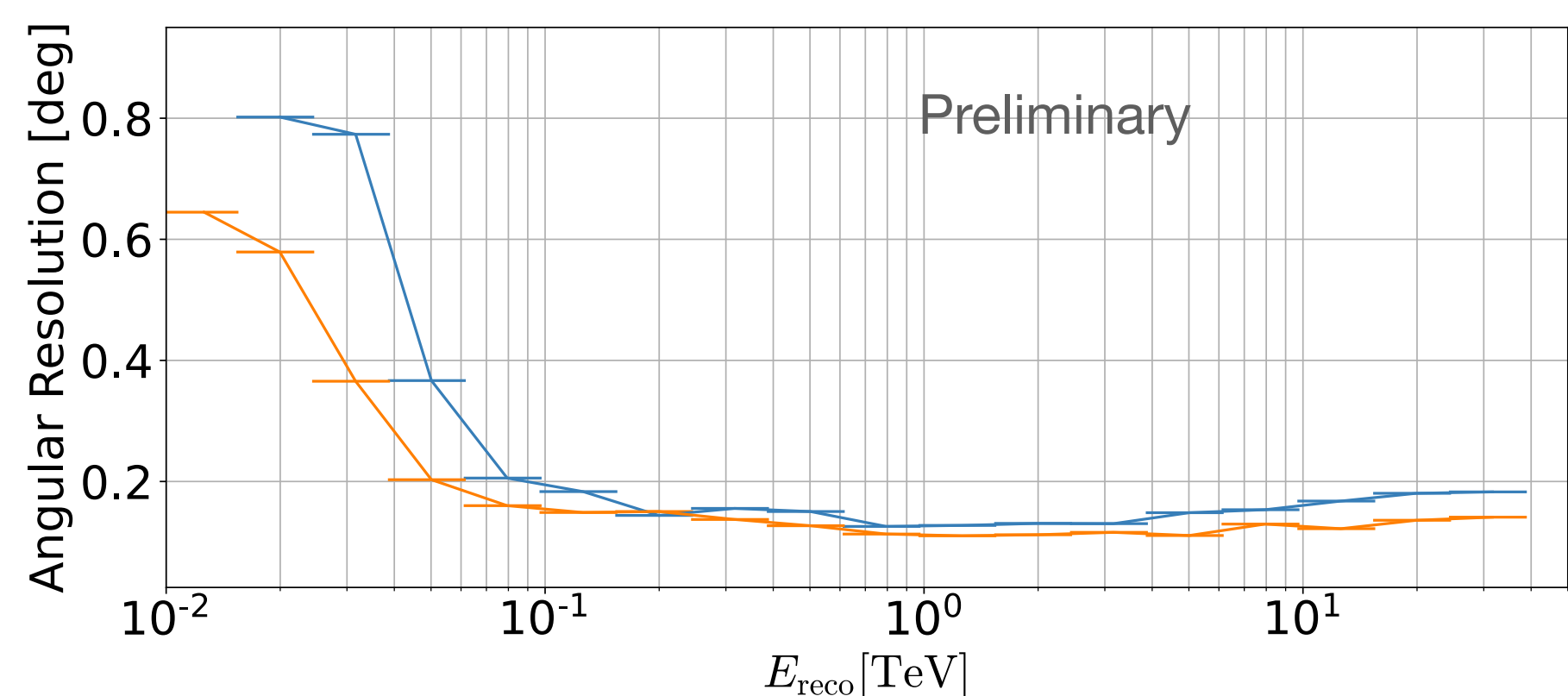
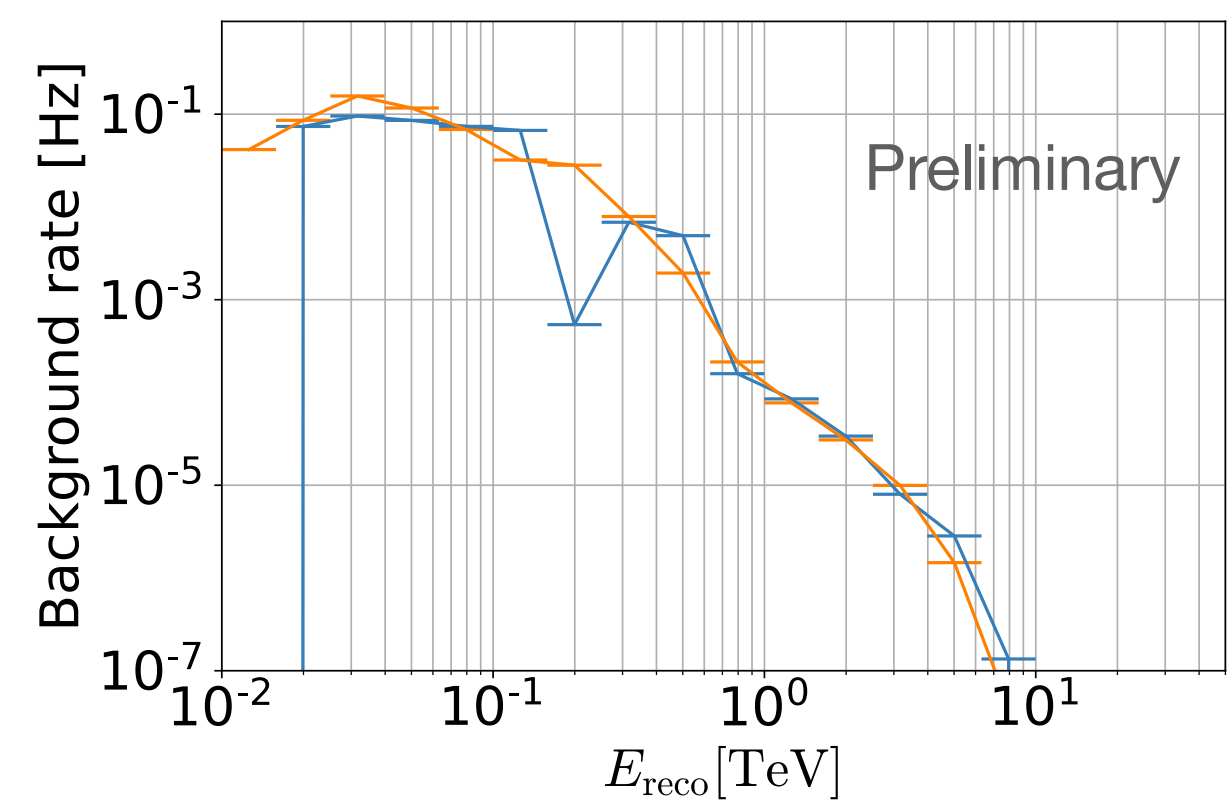
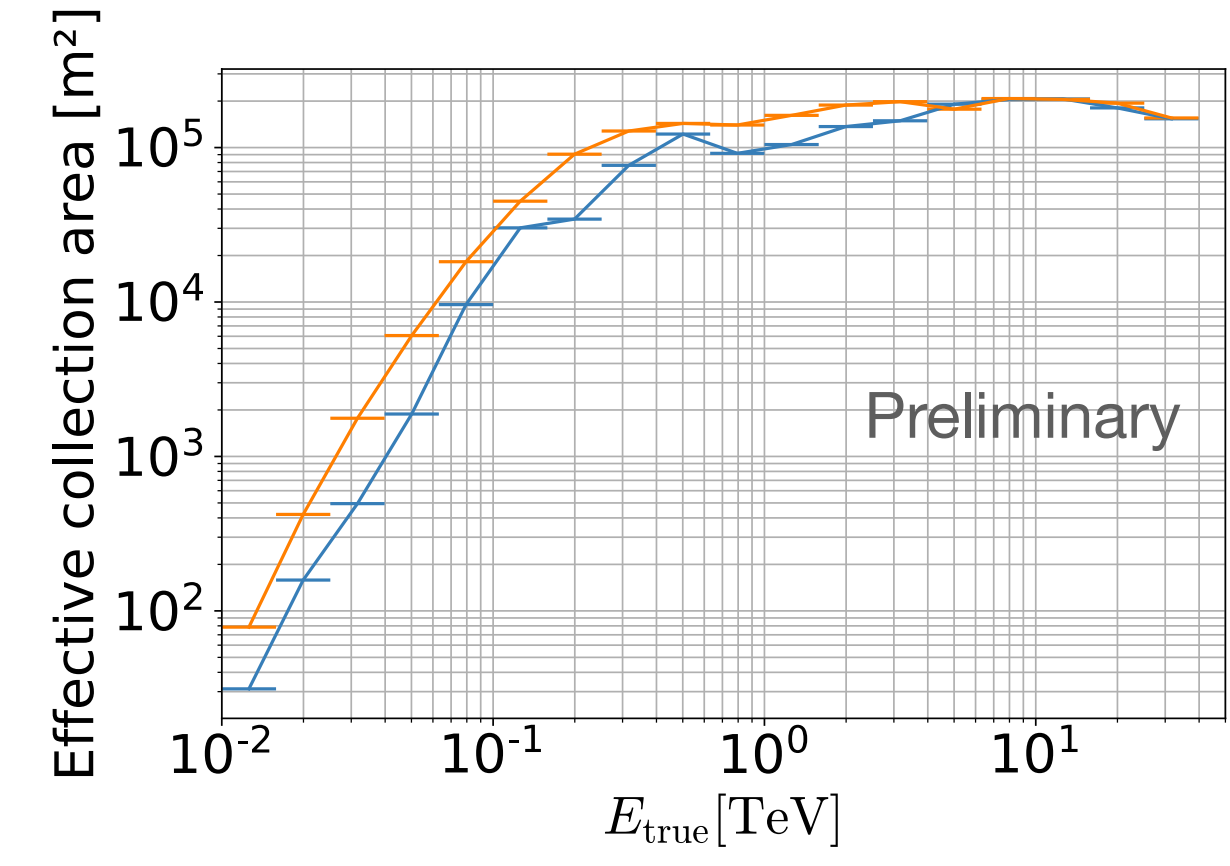
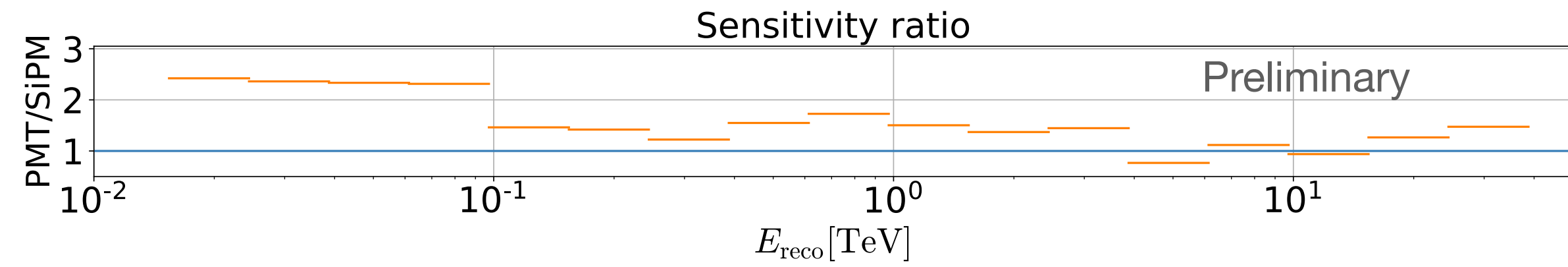
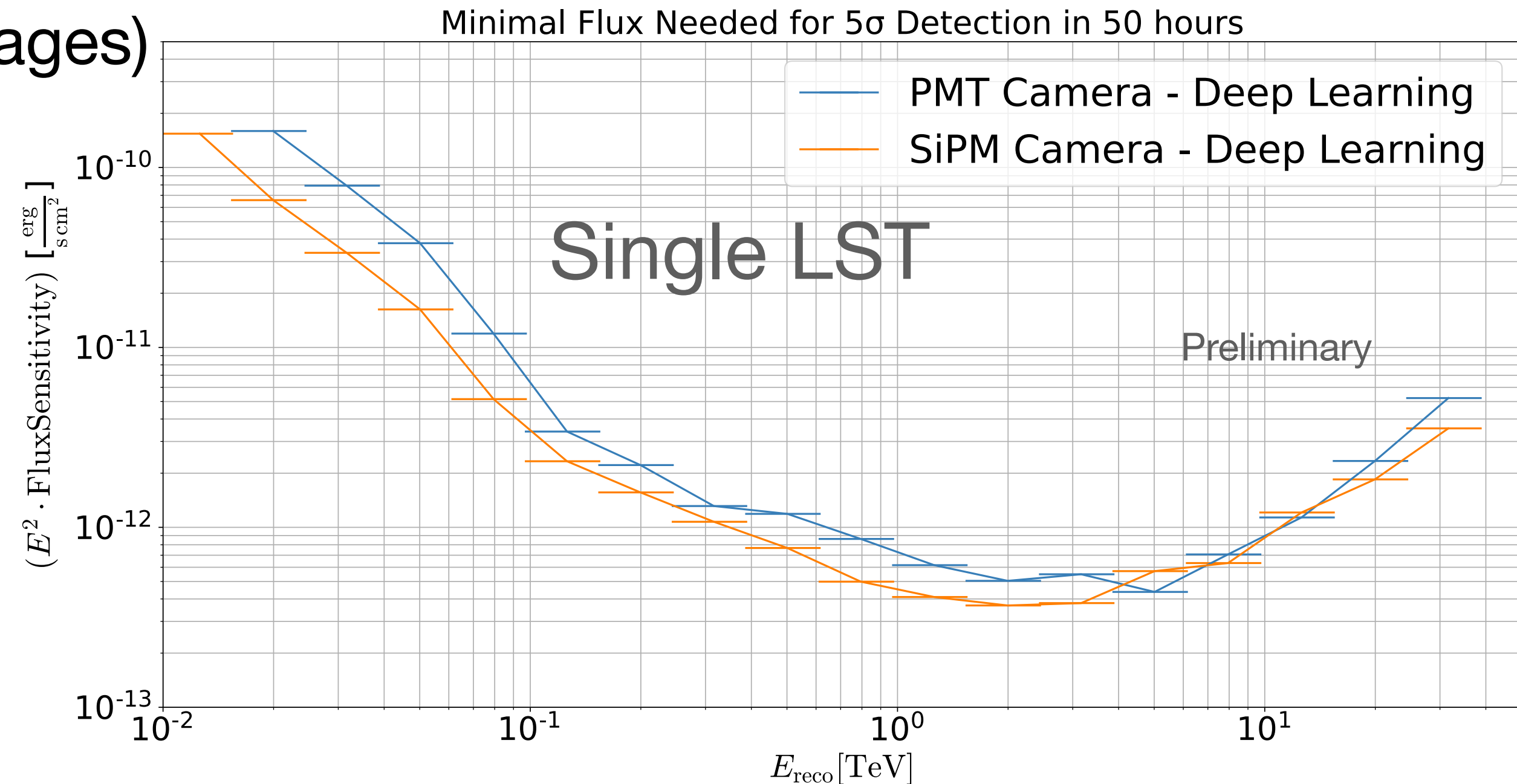


Thin-ResNet (default monoscopic model in CTLearn)

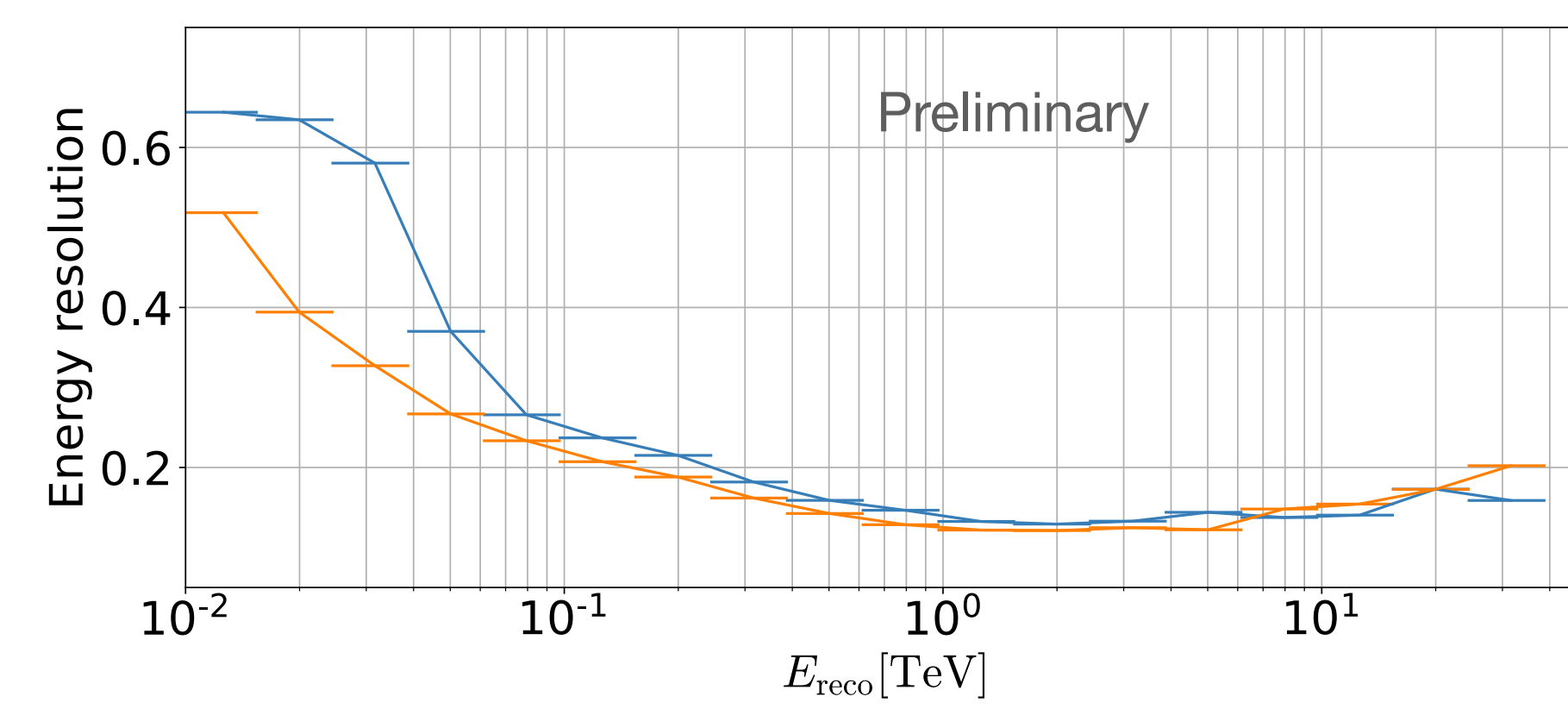
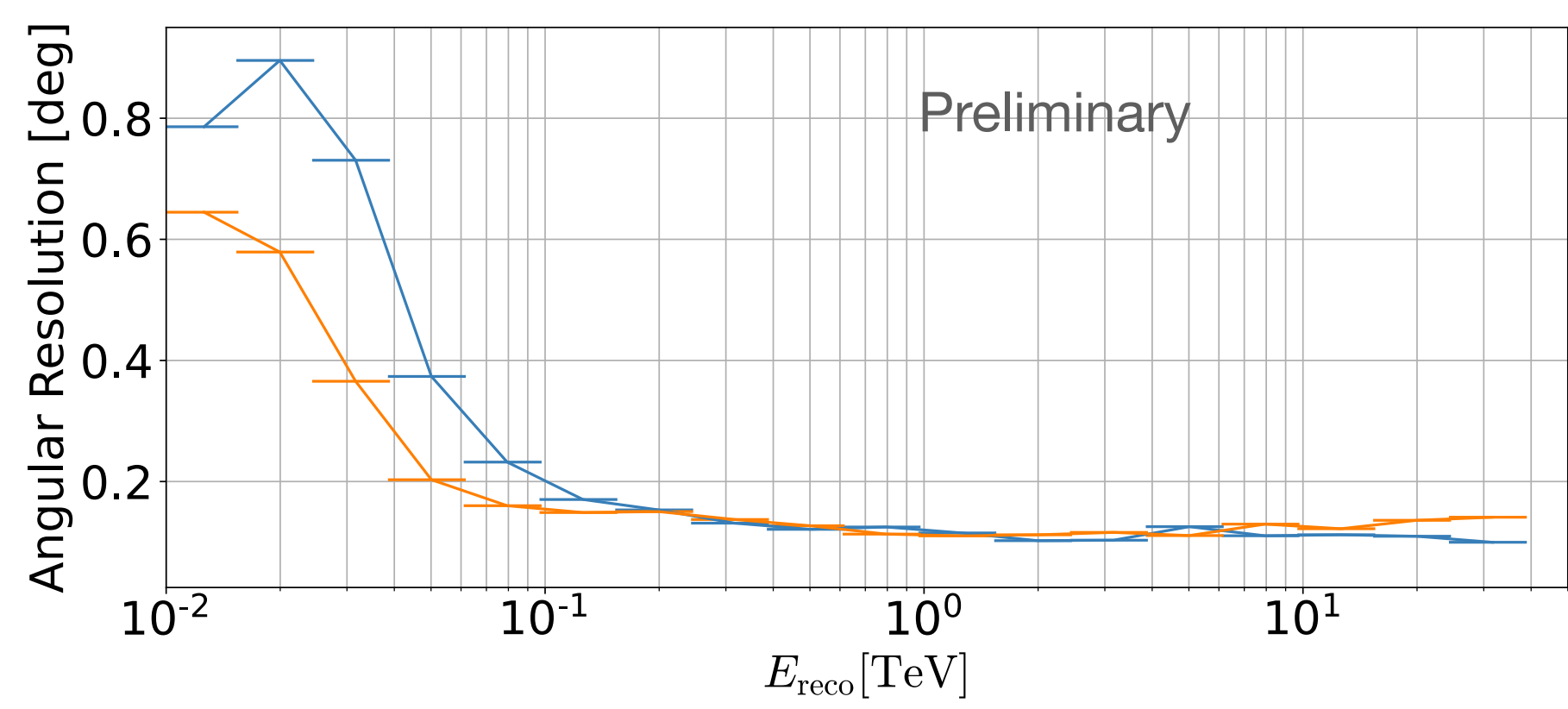
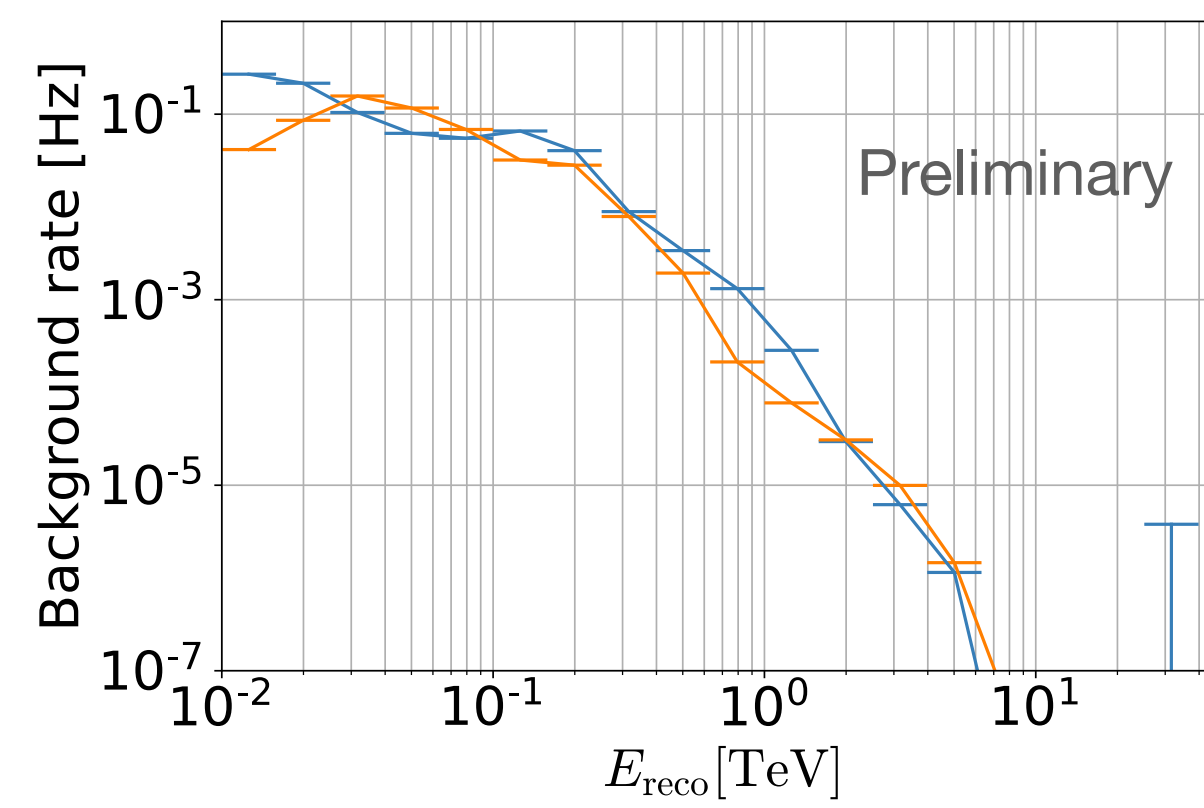
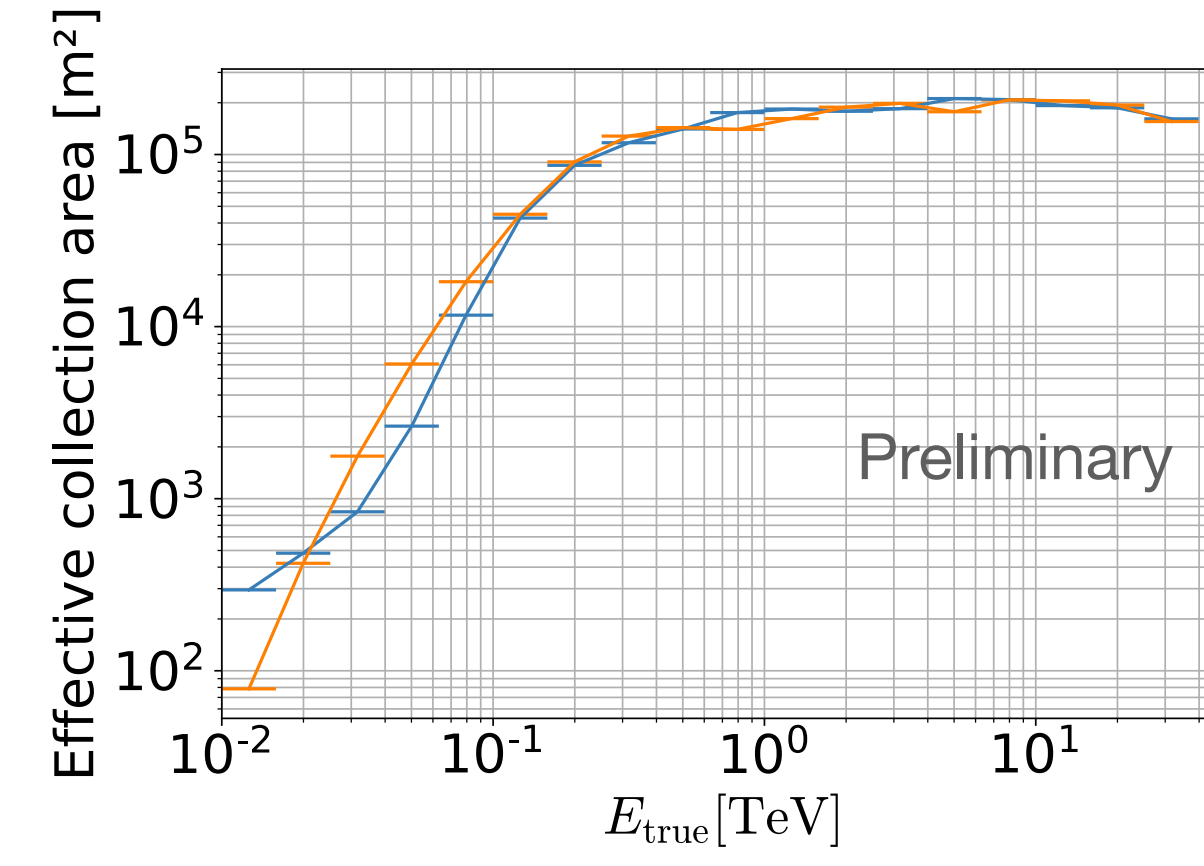
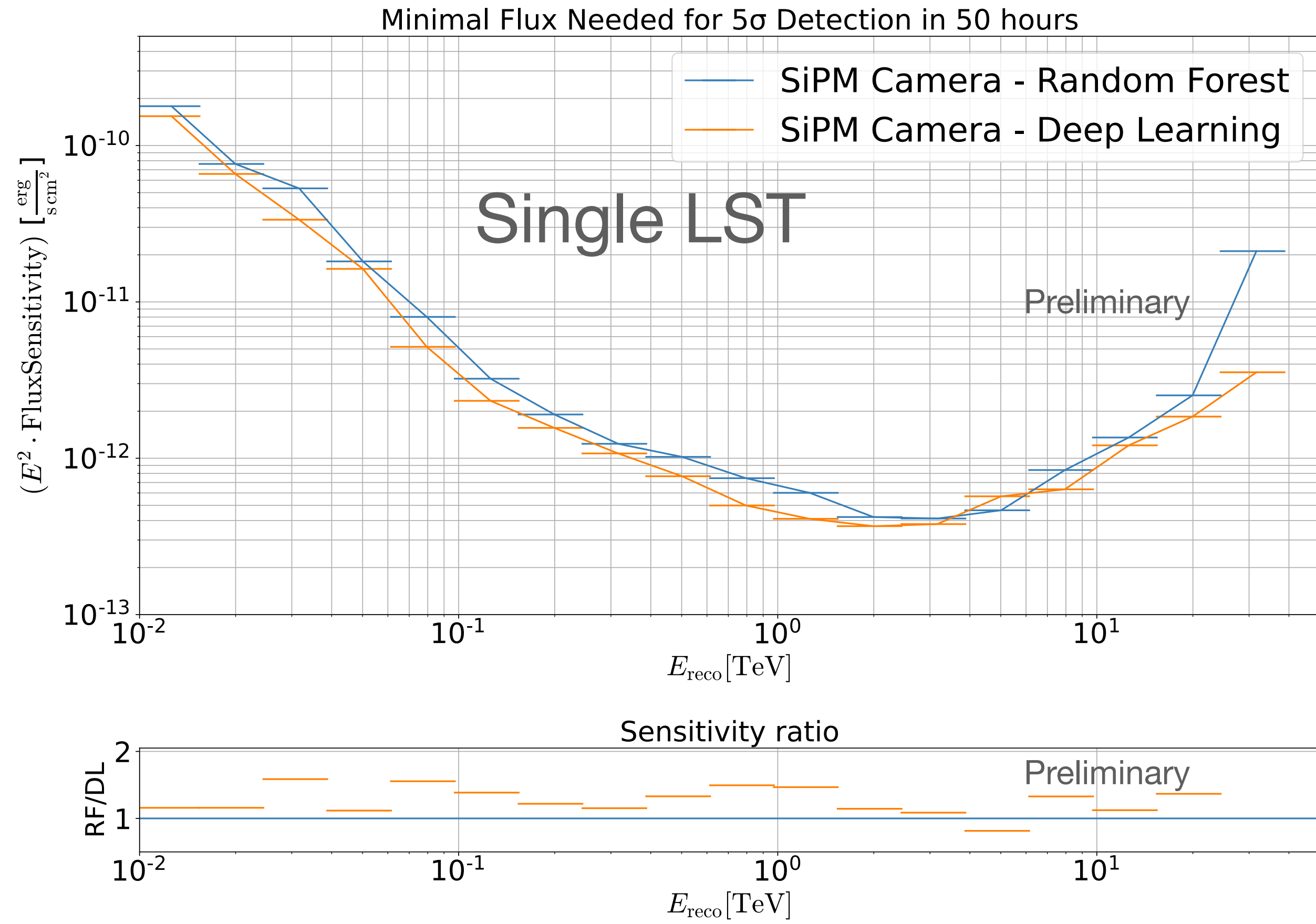


Deep Learning: current PMT camera vs high-resolution SiPM camera

(cleaned images)



High-resolution SiPM camera: Random Forest vs Deep Learning



Trigger architecture of the adv. LSTSiPM camera

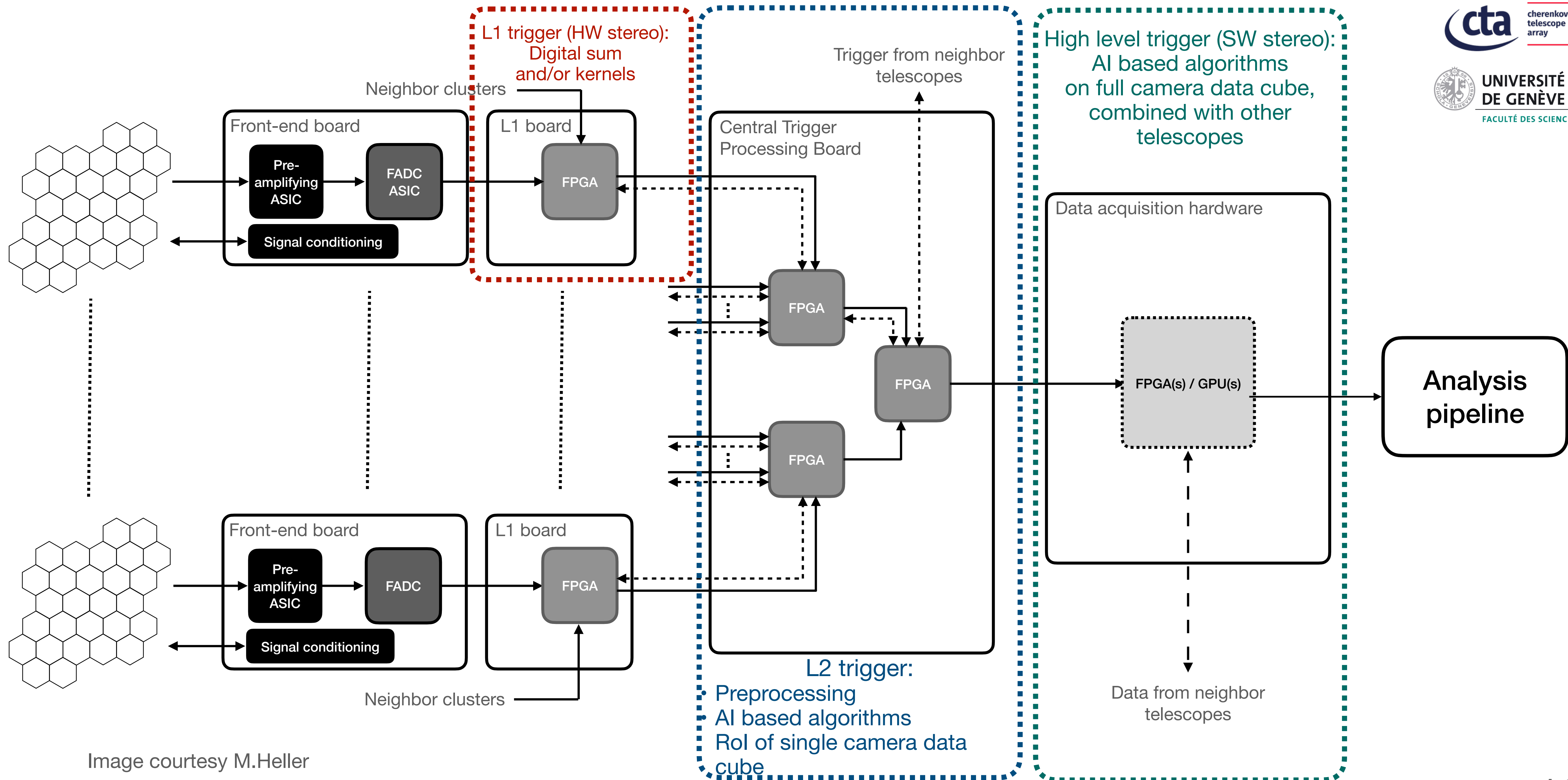
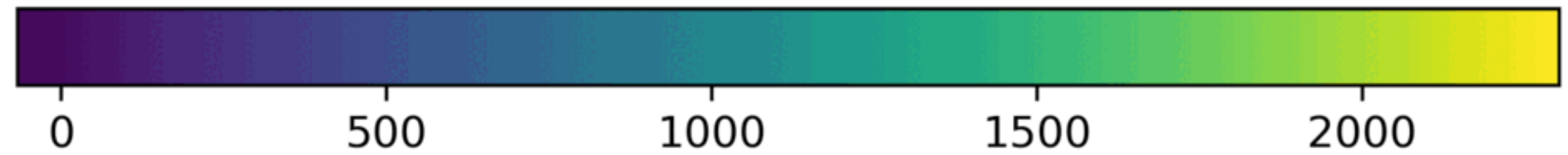
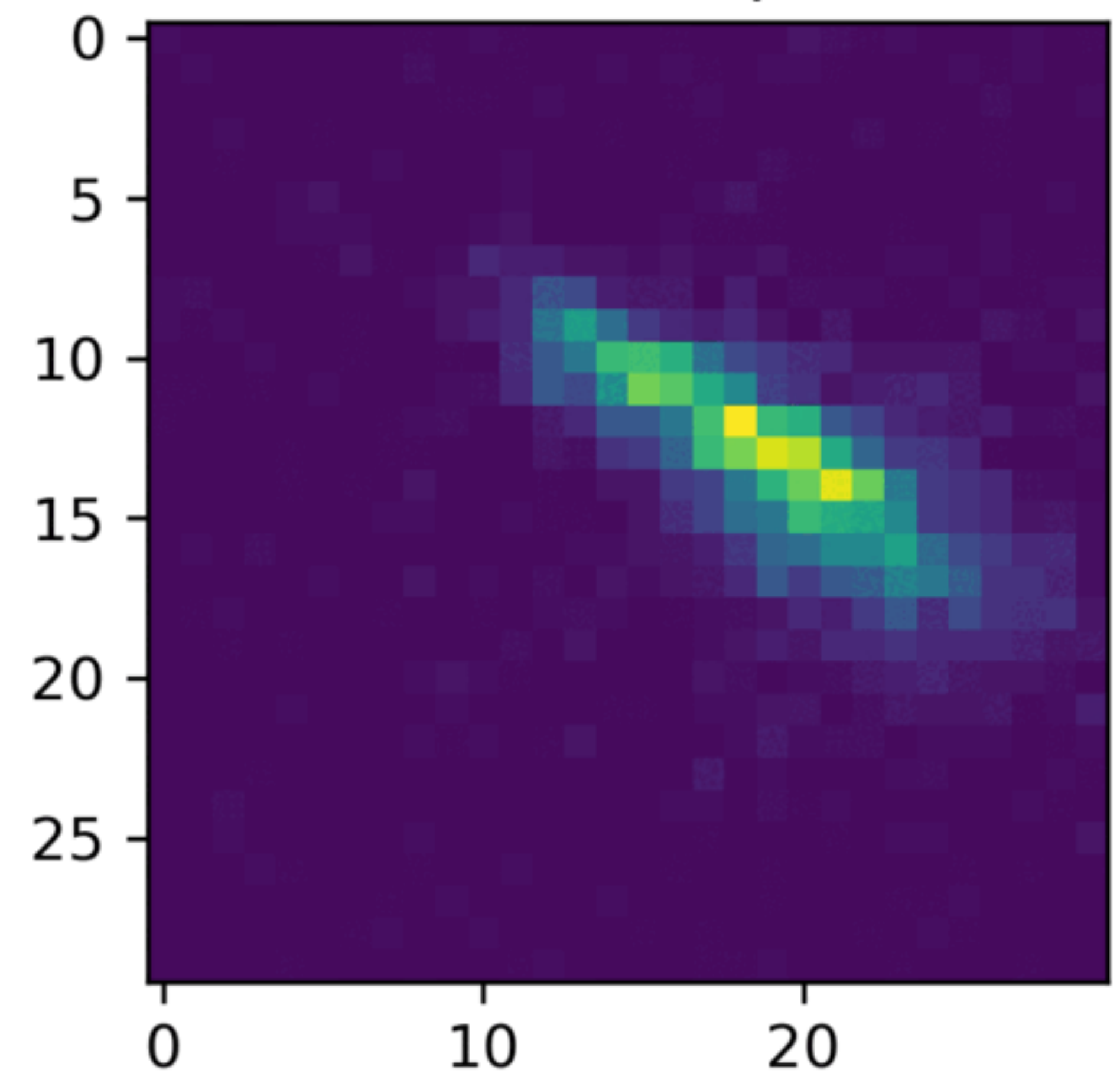
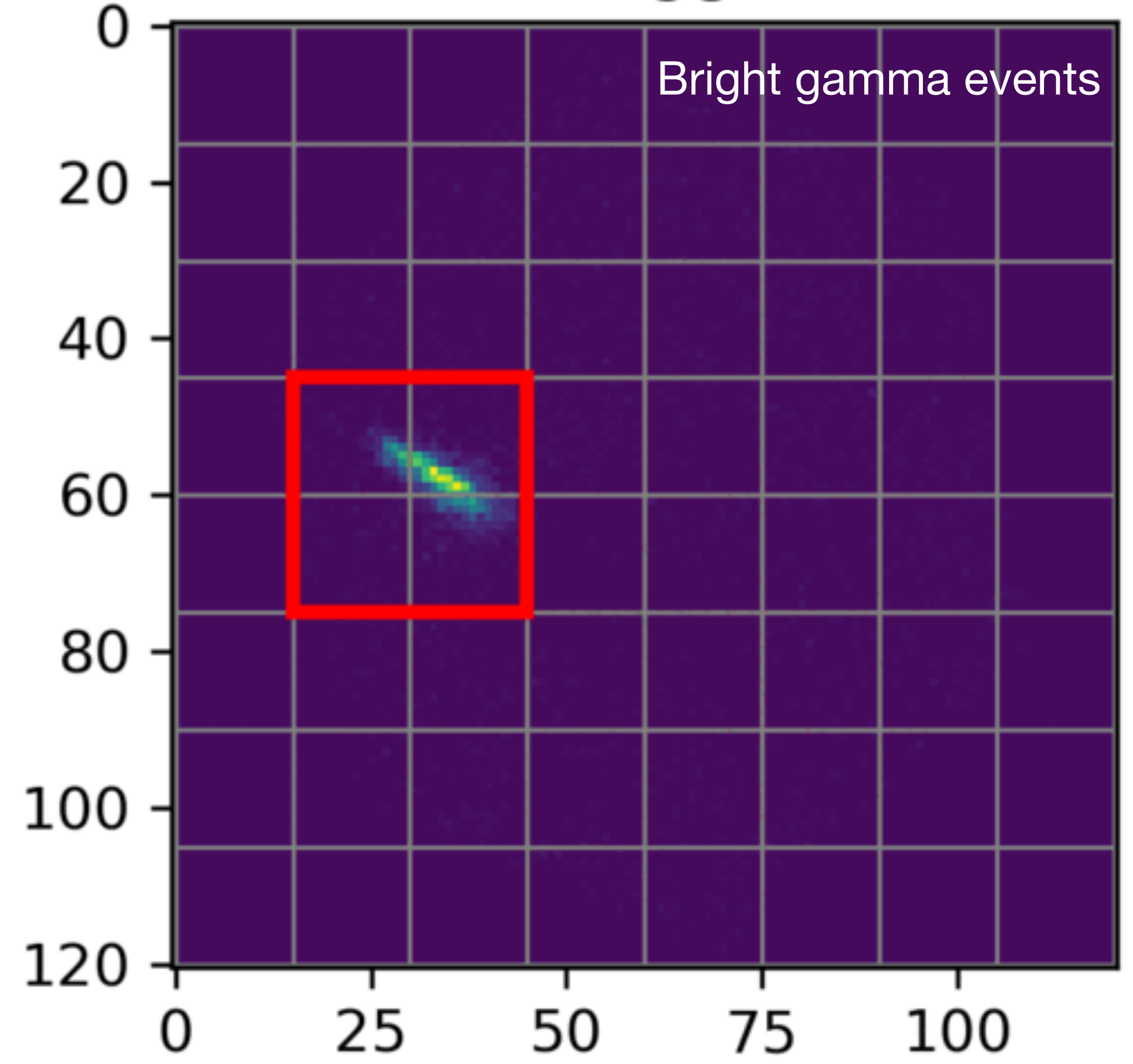


Image courtesy M.Heller

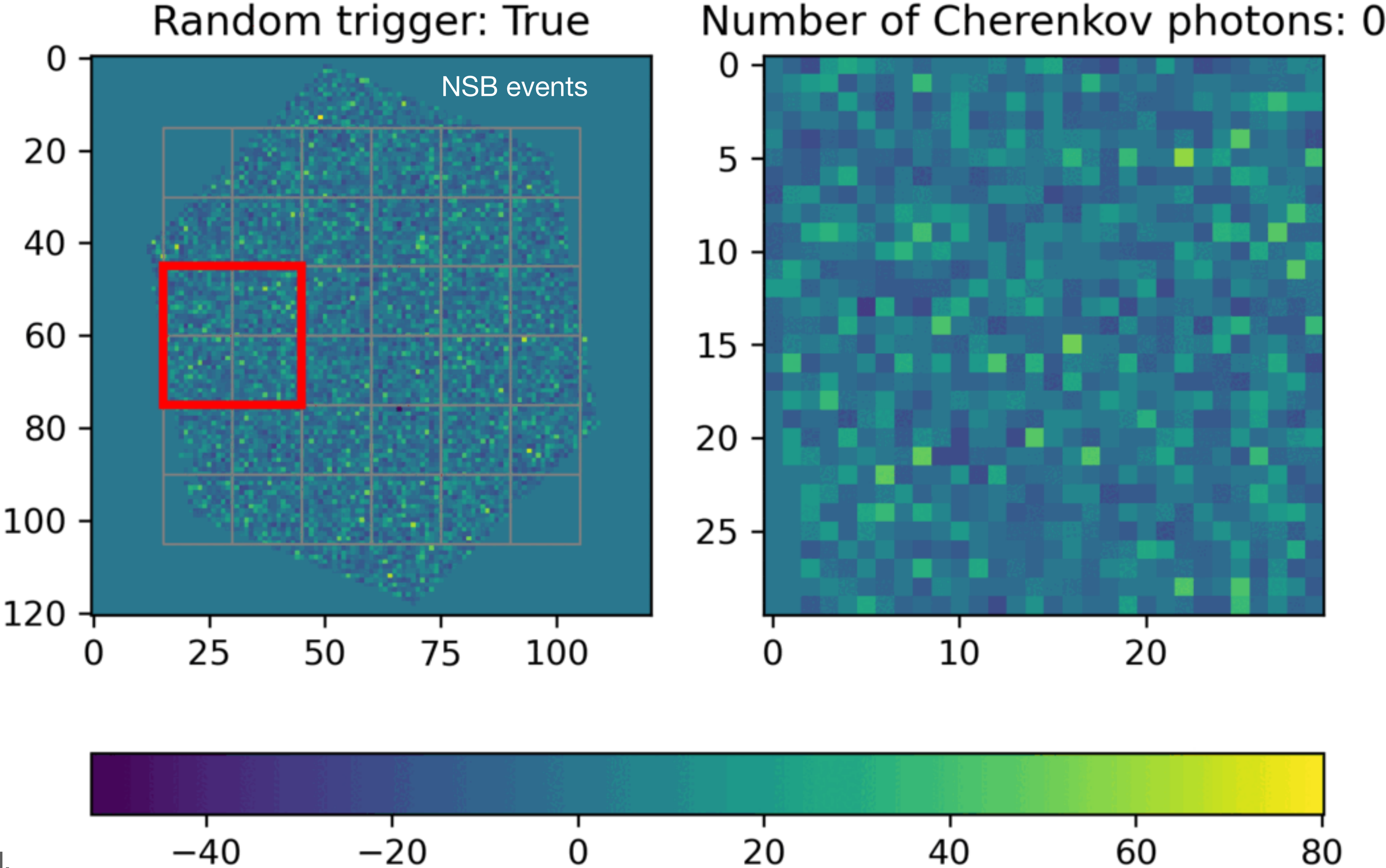
AI Trigger system of the adv. LSTSiPM camera

Random trigger: False

Number of Cherenkov photons: 10526

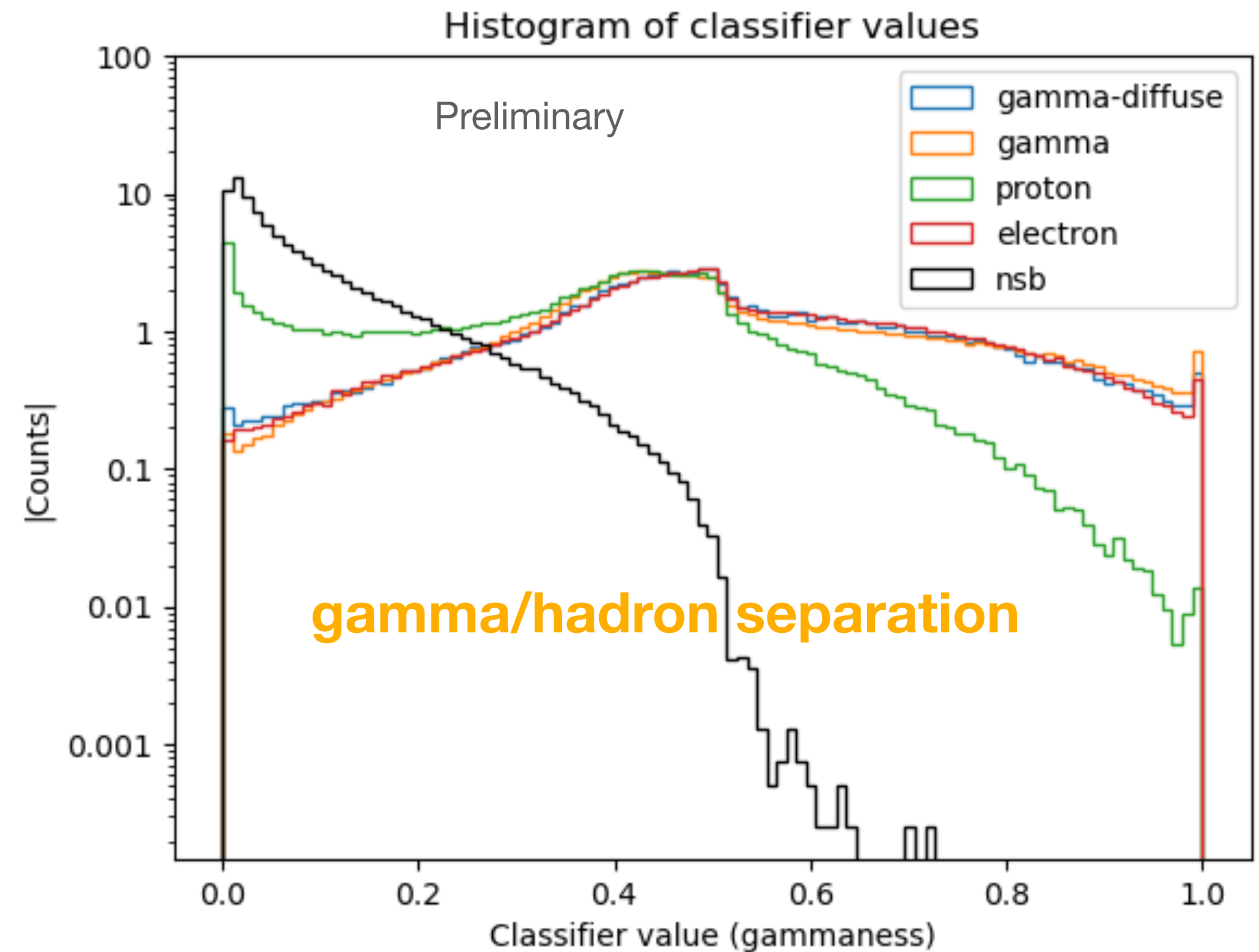
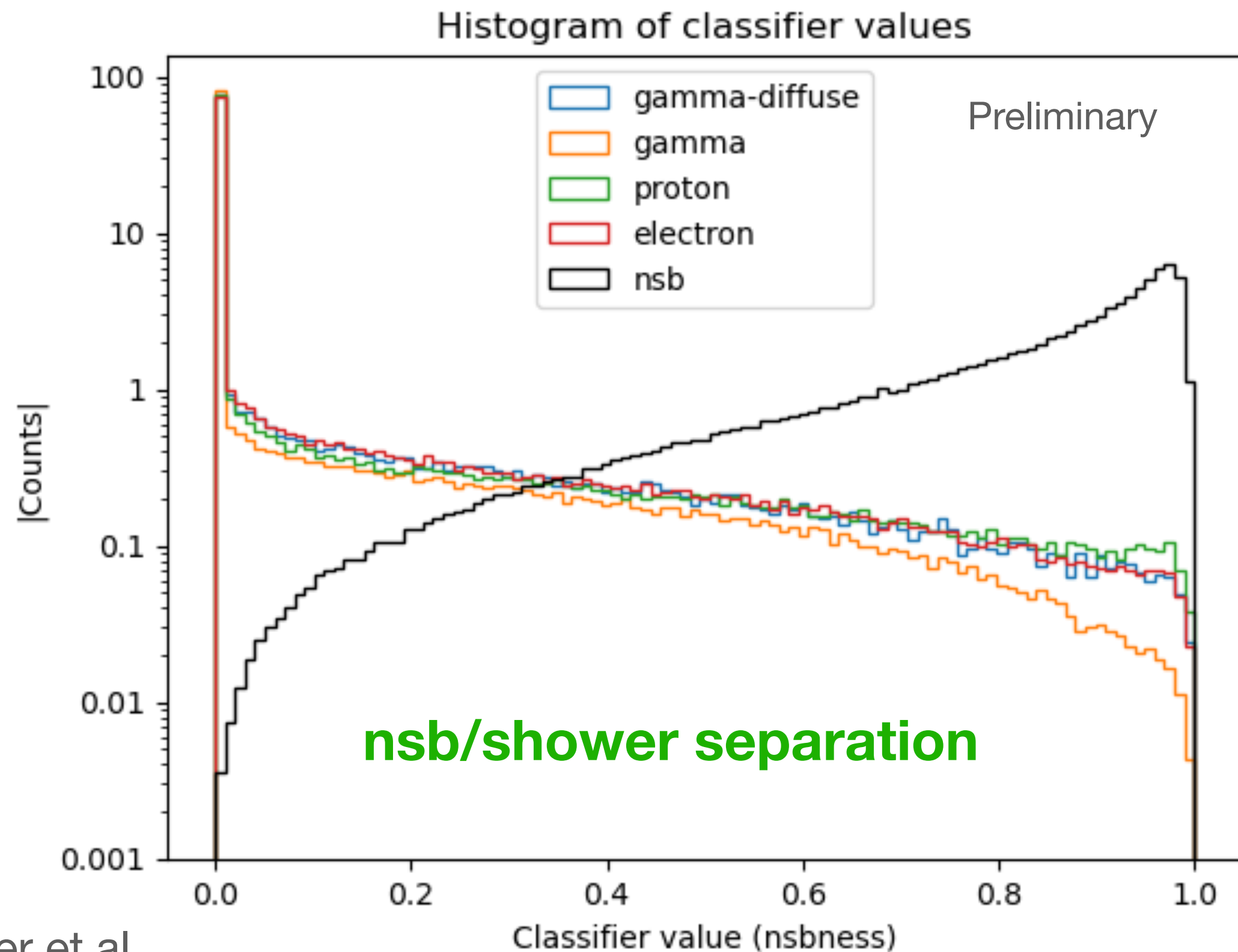


AI Trigger system of the adv. LSTSiPM camera



AI Trigger system of the adv. LSTSiPM camera

- **Categorical classification task** is performed including a ‘new class’ for **NSB patches**. Later on it is possible to add muons at training phase, which allows to do **muon tagging**.
- A **non-complex** plain CNN-based model is used that potential can run on the **FPGA** in the **CTP board**.





CSCS

Centro Svizzero di Calcolo Scientifico
Swiss National Supercomputing Centre



- All crucial MC simulation datasets including a trigger-less dataset are available at CSCS.
- CTA Learn software and dependencies successfully installed and deployed on the CSCS cluster.
- Different projects for the various applications: cta03 for ML-related studies with LST and cta04 for the AI Trigger system of the adv. SiPM camera.
- New project regarding the processing on real observational data from the LST-1 prototype using CTA Learn is submitted. Common interests with the DPPS for processing DVR-DL0 data offline.

Conclusion & outlook

- **CNN**-based methods show **superior** results for the full-event reconstruction with the high-resolution advanced SiPM cameras. **Performance studies in progress.**
- Preliminary studies with **waveforms** (shower developments) is promising and open the doors for various new applications such as the **AI Trigger system.**
- On going effort on **porting** those algorithms on **FPGAs** to run them **real-time** inside the cameras.
- Application on **real observational data** from the SST1M (stereo) and LST-1 (mono) prototypes is the next step. Our focus will lay on working with **cleaned** images and waveforms (or DVR-DL0 data) which is typically more **robust** when applying those algorithms to real observational data.
- Utilizing graph neural networks (GNNs) for **stereoscopic** reconstruction for CTA subarrays in the future.

Merci pour votre attention!



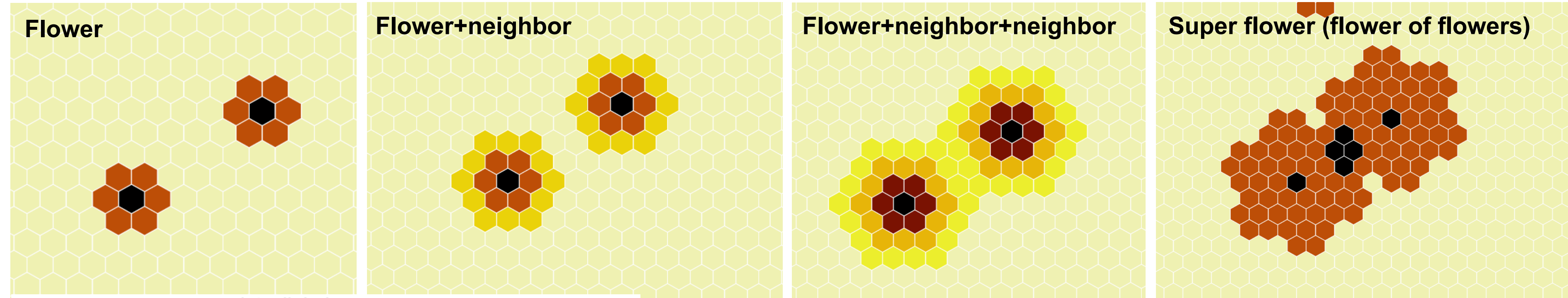
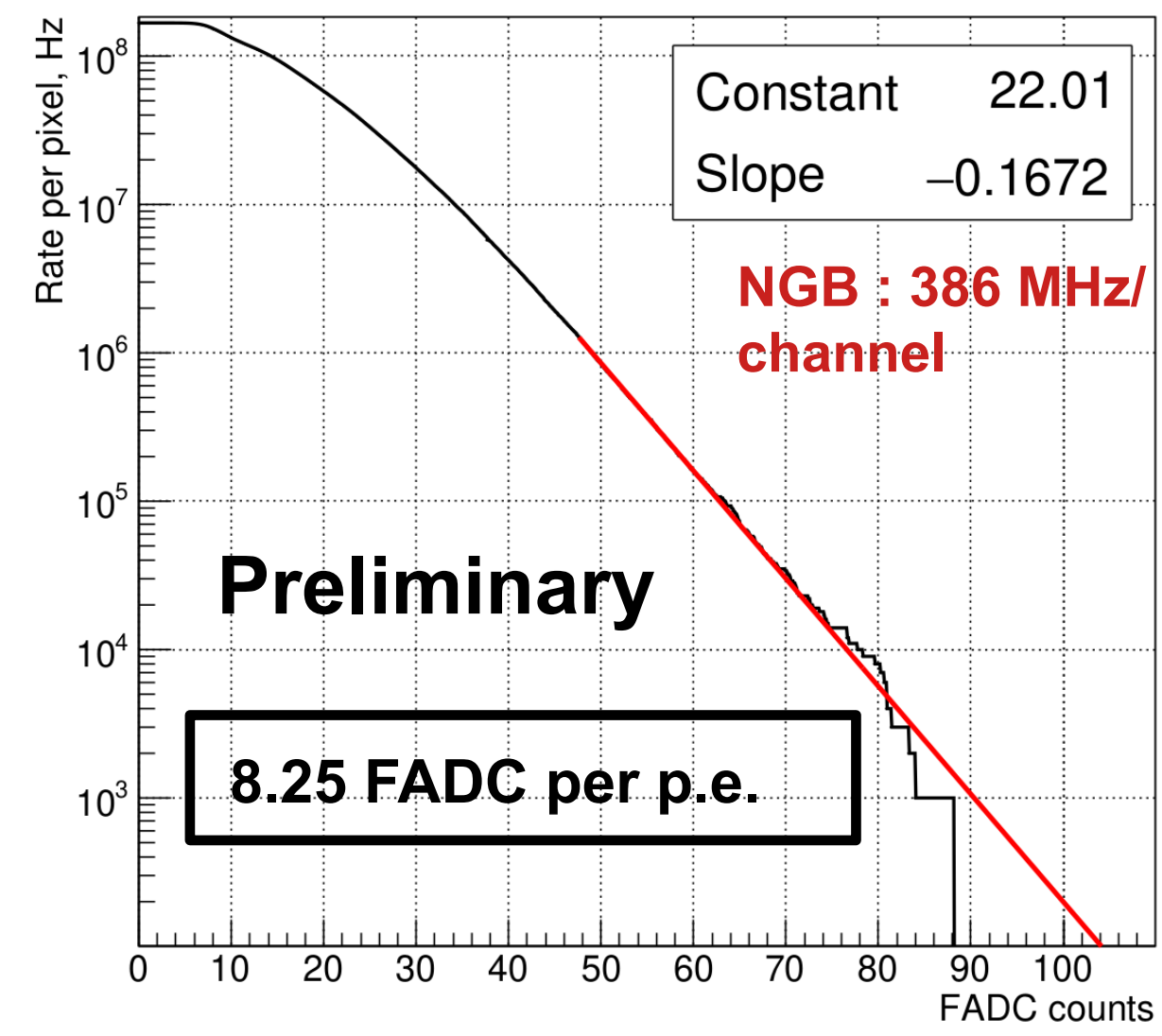
Backup

L1 Digital Sum Trigger

by Leonid Burmistrov (UniGe)

Motivation:

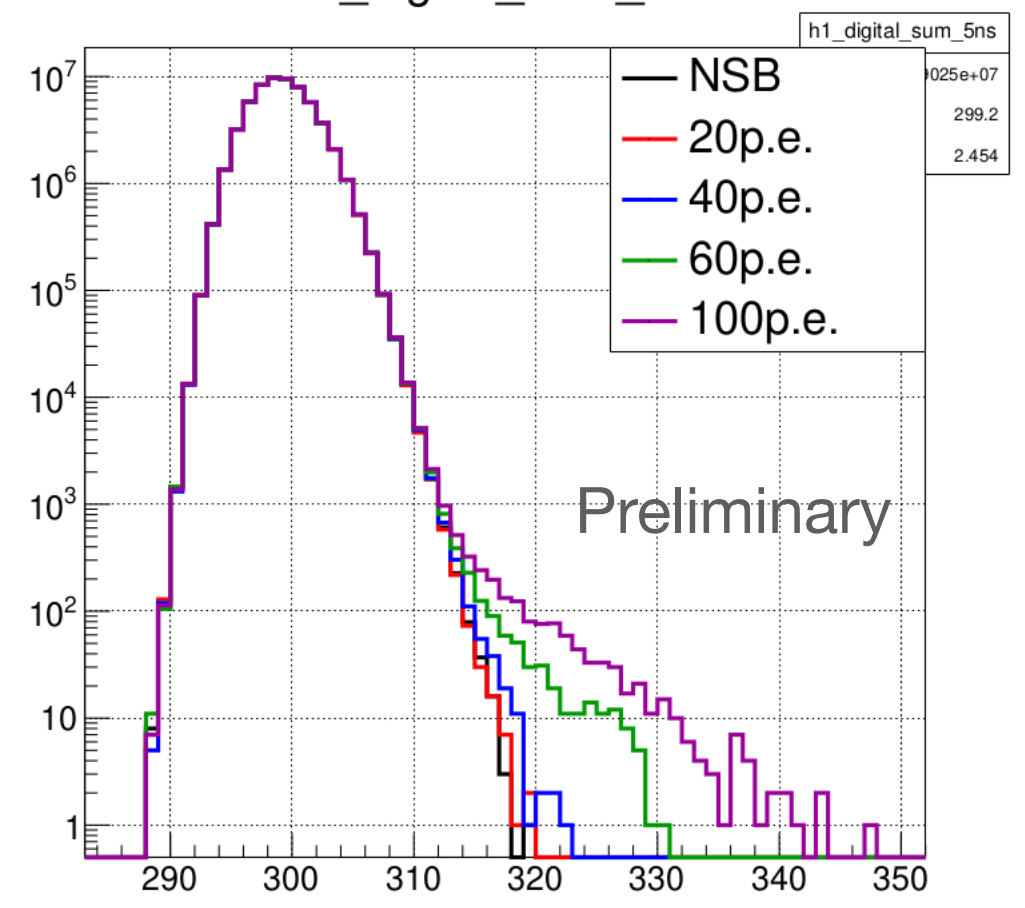
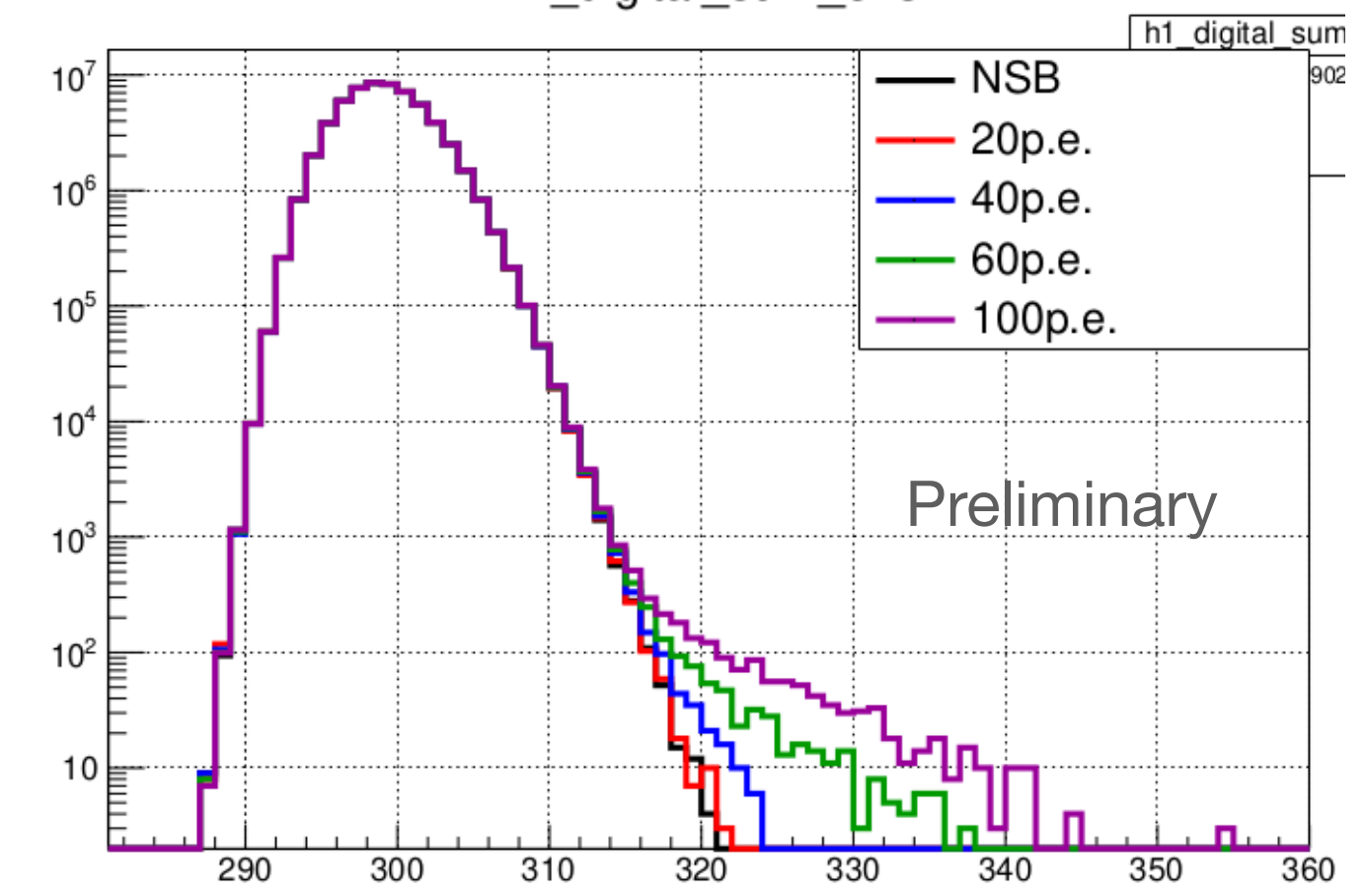
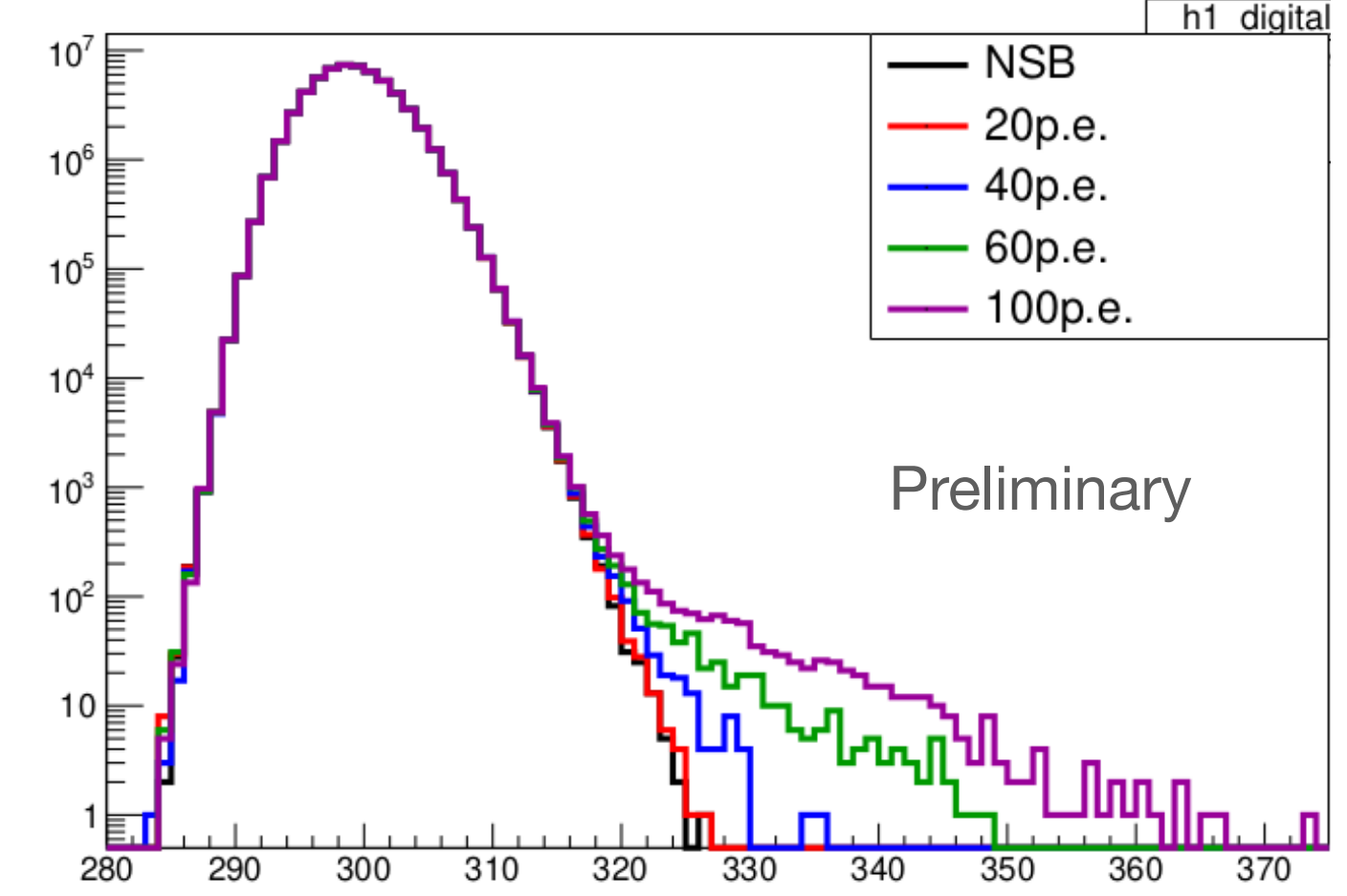
- Have better behaviour than threshold approach
- Build different clusters
- Integrate / differentiate over time samples
- Use different digital filters
- Use “coincidence” between clusters



h1_digital_sum

h1_digital_sum_3ns

h1_digital_sum_5ns



Waveform-CNN-processing available in CTLearn

- <https://github.com/cta-observatory/dl1-data-handler/pull/119> enables raw (R0) and calibrated (R1) waveform reading and preserved all previous DL1DH features (event-wise reading using generators, quality cuts, reading stereoscopic array-level data,...)
- <https://github.com/cta-observatory/dl1-data-handler/pull/124> enables pixel-wise pedestal subtraction for raw (R0) waveform.
- <https://github.com/ctlearn-project/ctlearn/pull/173> enables CNN-based models to process waveforms.
- PRs are all merged. New releases soon!

AI-based trigger system available soon

- <https://github.com/cta-observatory/dl1-data-handler/pull/126> enables data reading of the AI-based trigger system with the ability to crop trigger patches.
- <https://github.com/ctlearn-project/ctlearn/pull/180> enables the reconstruction of the true Cherenkov photons in a given trigger patch as a regression task. Not fully tested yet. Proof-of-concept needed.
- PRs are not 100% finalised. Some commits will follow.

Datasets - config settings (recommended by Istchain for PMT & Matthieu for SiPM)



```
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},
"ImageProcessor": {
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},
},
```

PMT

```
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},
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```

SiPM

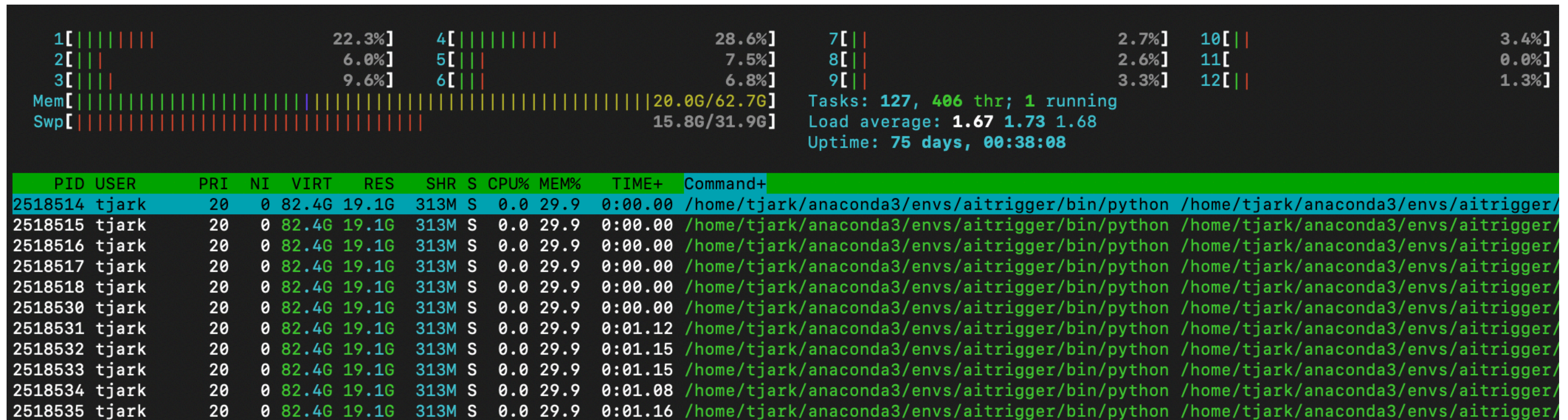
RF ctapipe config settings (inspired by lstchain standard config)

```
# =====  
# ctapipe-ml-train-particle-classifier config file.  
# version: VERSION  
#  
# Configuration for training machine-learning models  
# =====  
TrainParticleClassifier:  
  CrossValidator:  
    n_cross_validations: 5  
  
  ParticleClassifier:  
    model_cls: ExtraTreesClassifier  
    model_config:  
      max_depth: 30  
      min_samples_leaf: 10  
      n_jobs: -1  
      n_estimators: 100  
      bootstrap: true  
      criterion: "gini"  
      max_features: 1.0  
      max_leaf_nodes: null  
      min_impurity_decrease: 0.0  
      min_samples_split: 10  
      min_weight_fraction_leaf: 0.0  
      oob_score: false  
      random_state: 42  
      verbose: 0  
      warm_start: false  
      class_weight: null
```

```
features:  
  - log_hillas_intensity  
  - hillas_fov_lon  
  - hillas_fov_lat  
  - hillas_length  
  - hillas_width  
  - hillas_skewness  
  - hillas_kurtosis  
  - area  
  - timing_slope  
  - leakage_intensity_width_2  
  
#stereo_combiner_cls: "StereoMeanCombiner"  
#StereoMeanCombiner:  
# log_target: true  
  
QualityQuery:  
  quality_criteria:  
    #- ["HillasValid", "HillasReconstructor_is_valid"]  
    - ["enough intensity", "hillas_intensity > 25"]  
    - ["Positive width", "hillas_width > 0"]  
    - ["enough pixels", "morphology_n_pixels > 3"]  
    - ["not clipped", "leakage_intensity_width_1 < 0.2"]  
  
FeatureGenerator:  
  features:  
    - ["area", "hillas_width * hillas_length"]  
    - ["log_hillas_intensity", "np.log(hillas_intensity)"]
```

DL1-DH I/O performance during training - CNN-based full-event reconstruction on calibrated waveforms with CTLearn

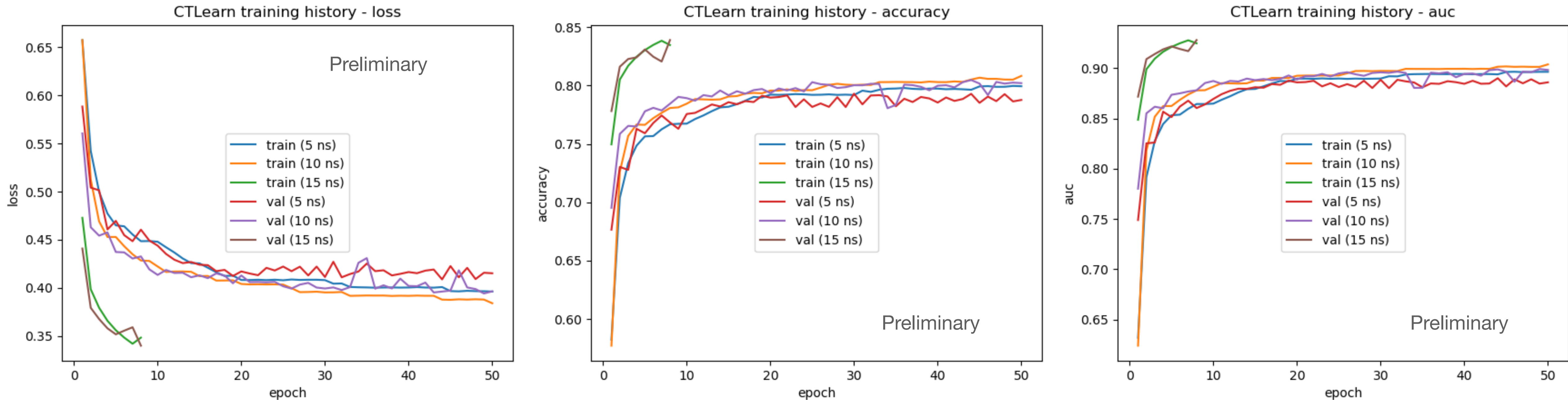
- Training with merged files is more efficient. We can now train with the full production (80 % reserved for the training set for gamma-diffuse and proton) without running into RAM issues.
- RAM usage for the training with a ~4TB dataset is less than 10GB per GPU.



Training process - CNN-based full-event reconstruction on calibrated waveforms with CTLearn

- **Gamma/hadron separation** is possible for CNN-based models on calibrated waveforms!
- Calibrated waveform (R1) training was performed with loose quality cuts ($\text{hillas_intensity} > 50$ & $\text{leakage_intensity_width_2} < 0.2$) extracting 5, 10, and 15 snapshots (nano seconds) around the shower maximum
- 5 and 10 snapshots were trained with a fraction of the training stats because I was using the run-wised files (less efficient). Training for 15 snapshots (15ns) was done with the whole training stats using the merged files.

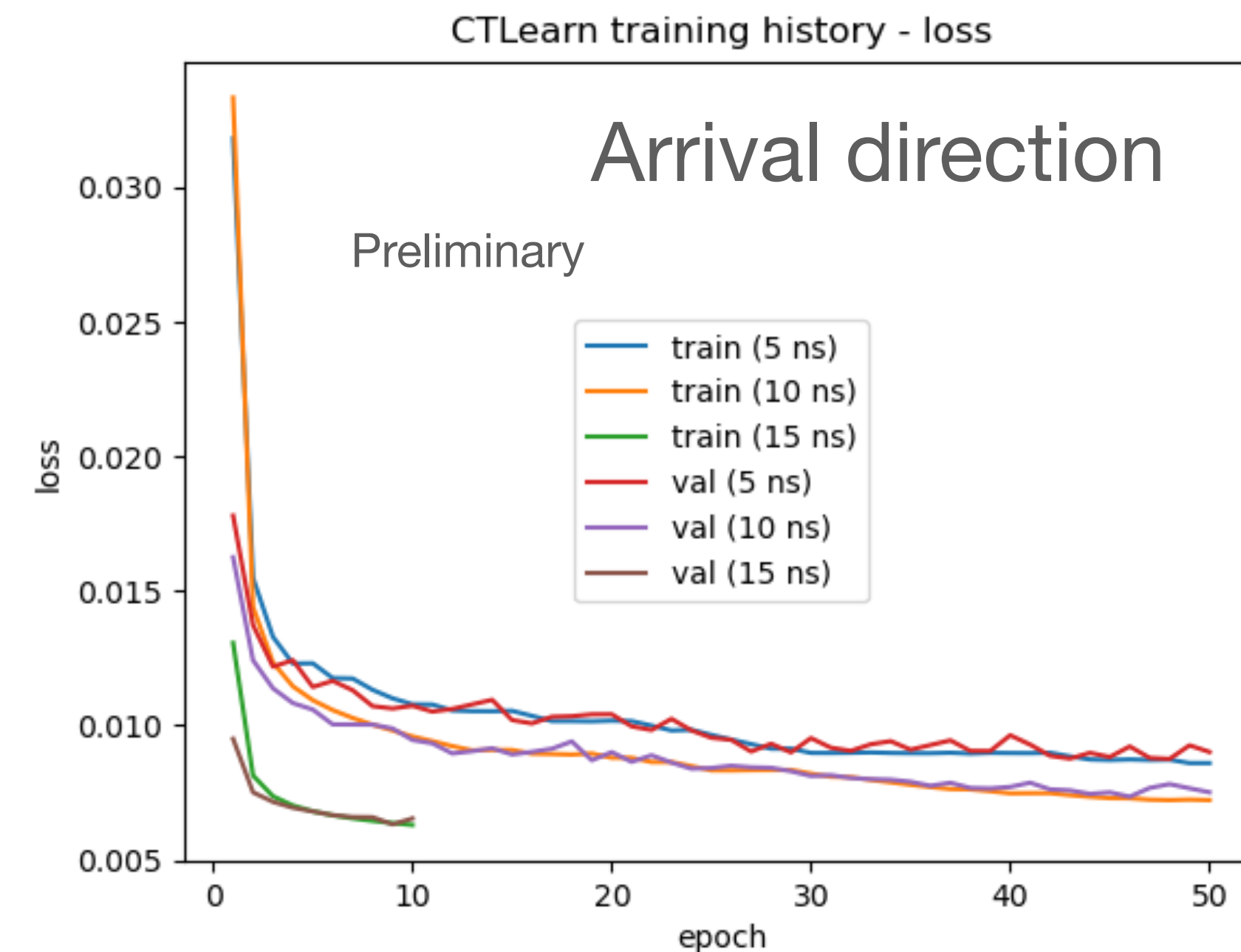
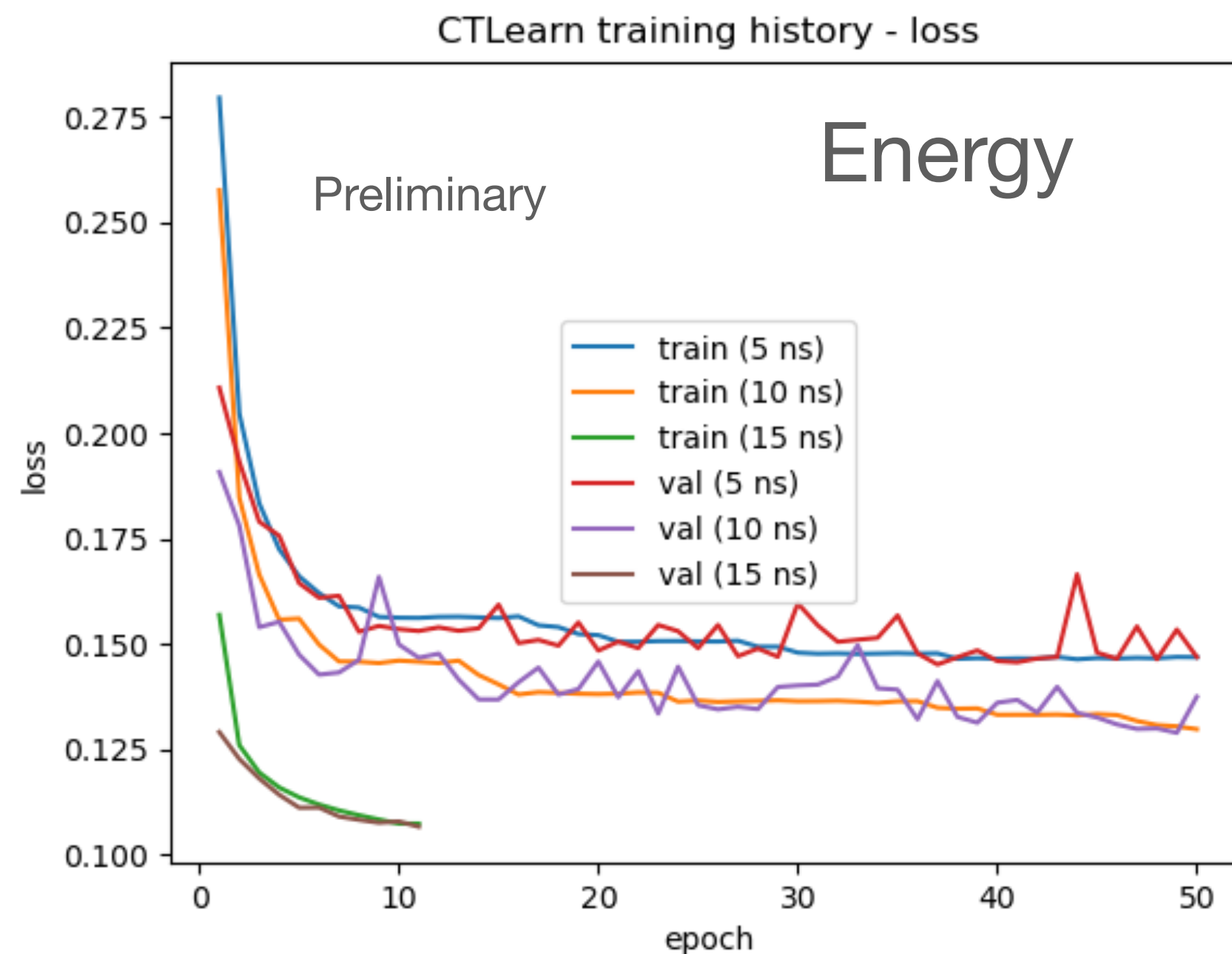
5 ns (low stats) vs 10 ns (low stats) vs 15 ns (full stats)



Training process - CNN-based full-event reconstruction on calibrated waveforms with CTLearn

- **Energy and arrival direction regression** is possible for CNN-based models on calibrated waveforms!
- Calibrated waveform (R1) training was performed with loose quality cuts ($\text{hillas_intensity} > 50$ & $\text{leakage_intensity_width_2} < 0.2$) extracting 5, 10, and 15 snapshots (nano seconds) around the shower maximum
- 5 and 10 snapshots were trained with a fraction of the training stats because I was using the run-wised files (less efficient). Training for 15 snapshots (15ns) was done with the whole training stats using the merged files.

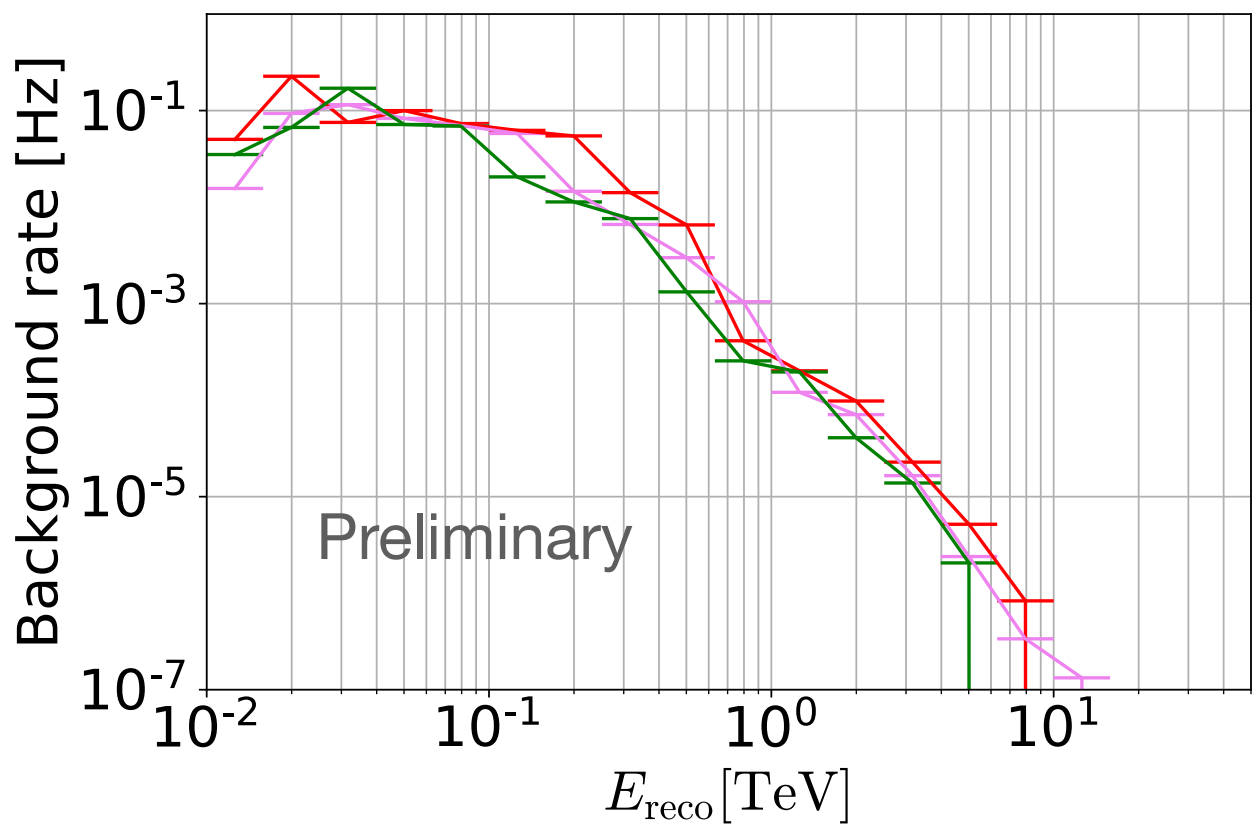
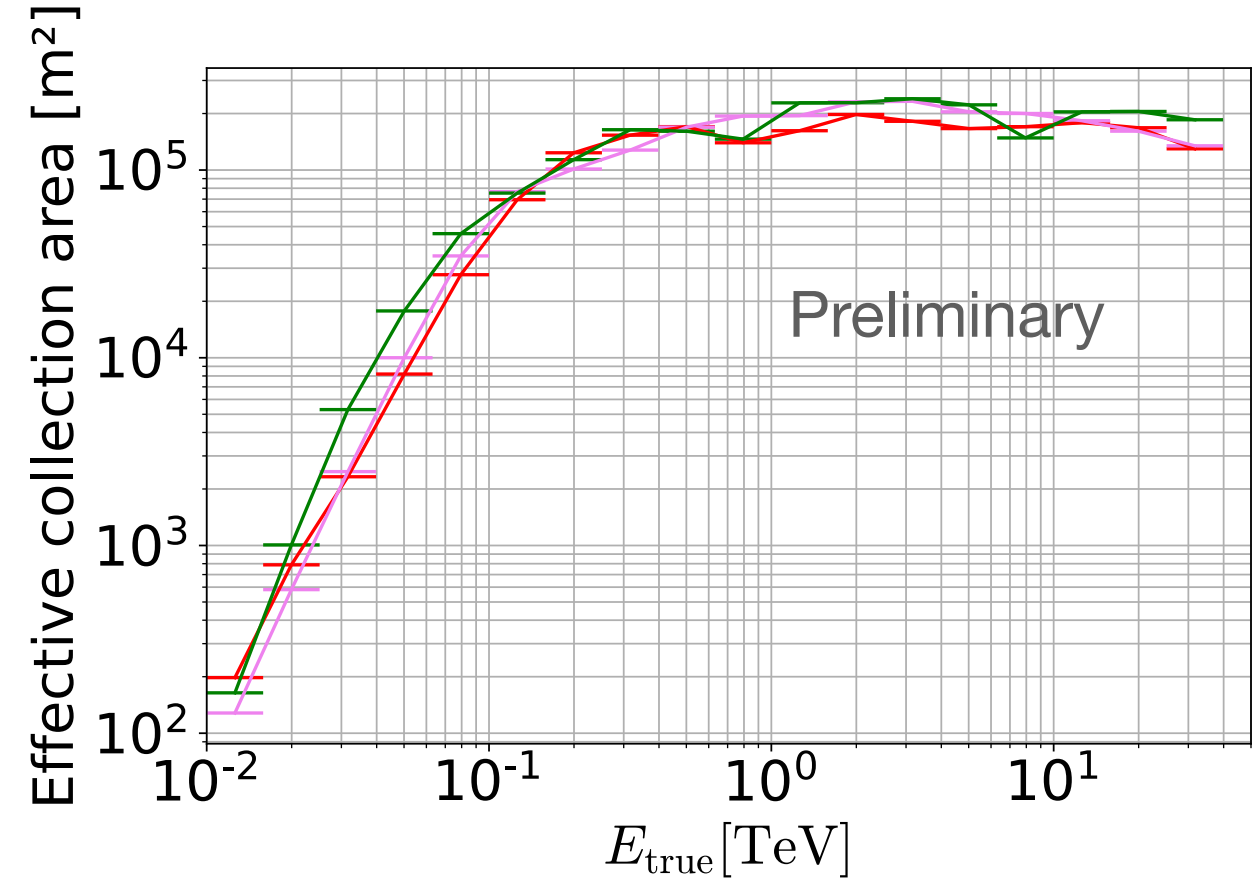
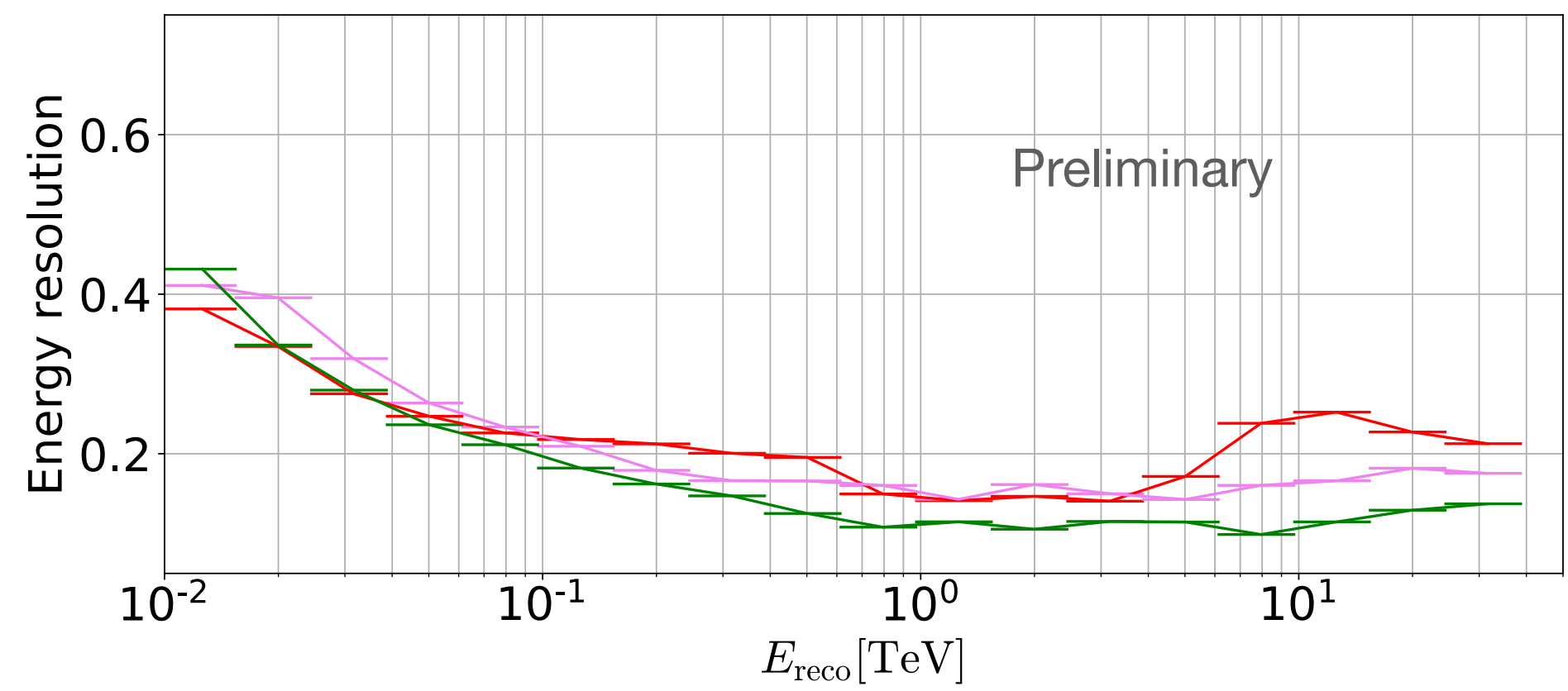
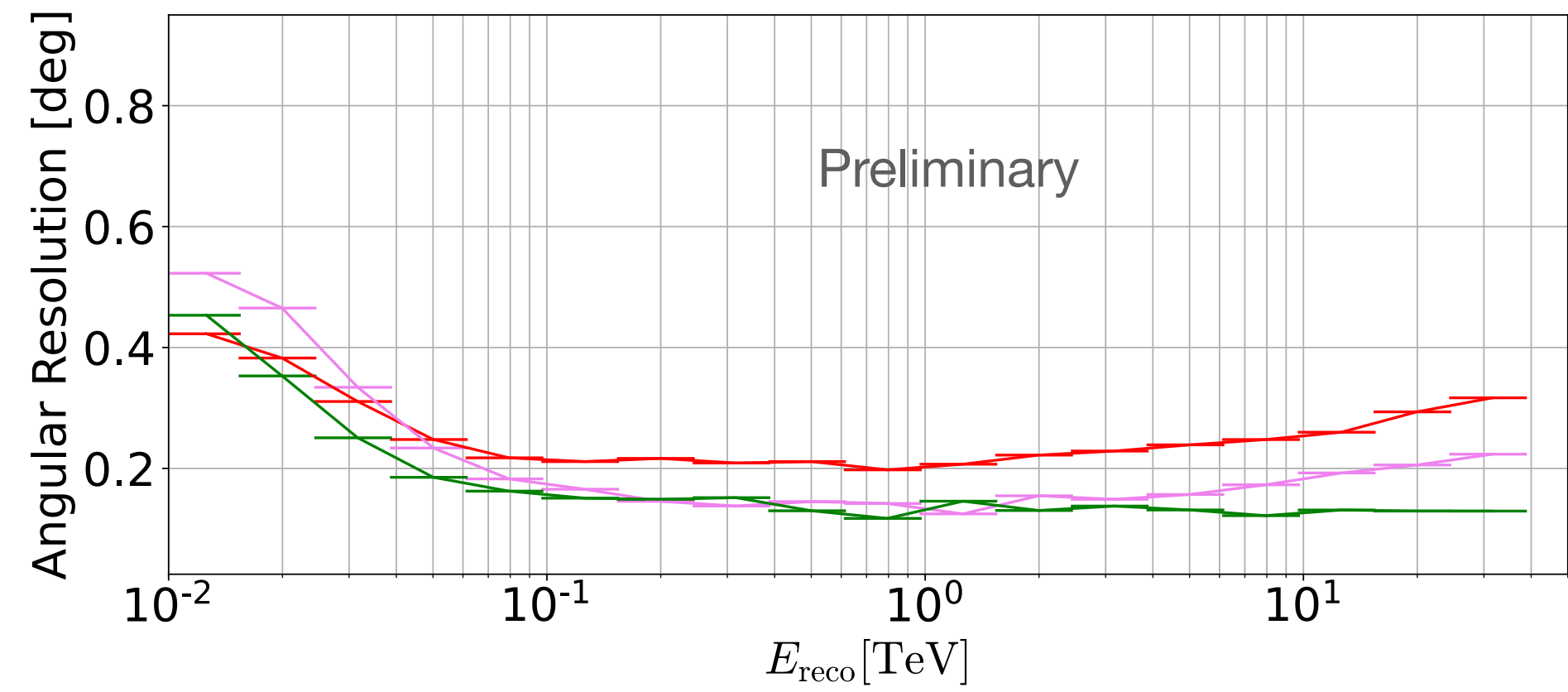
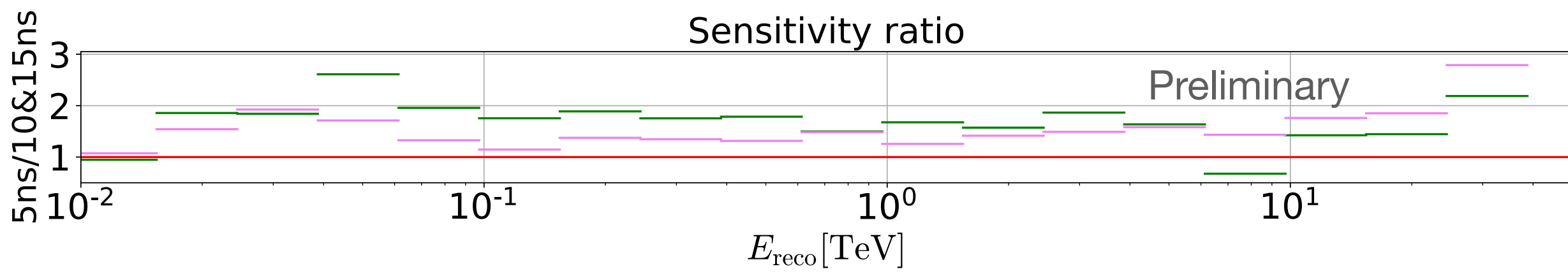
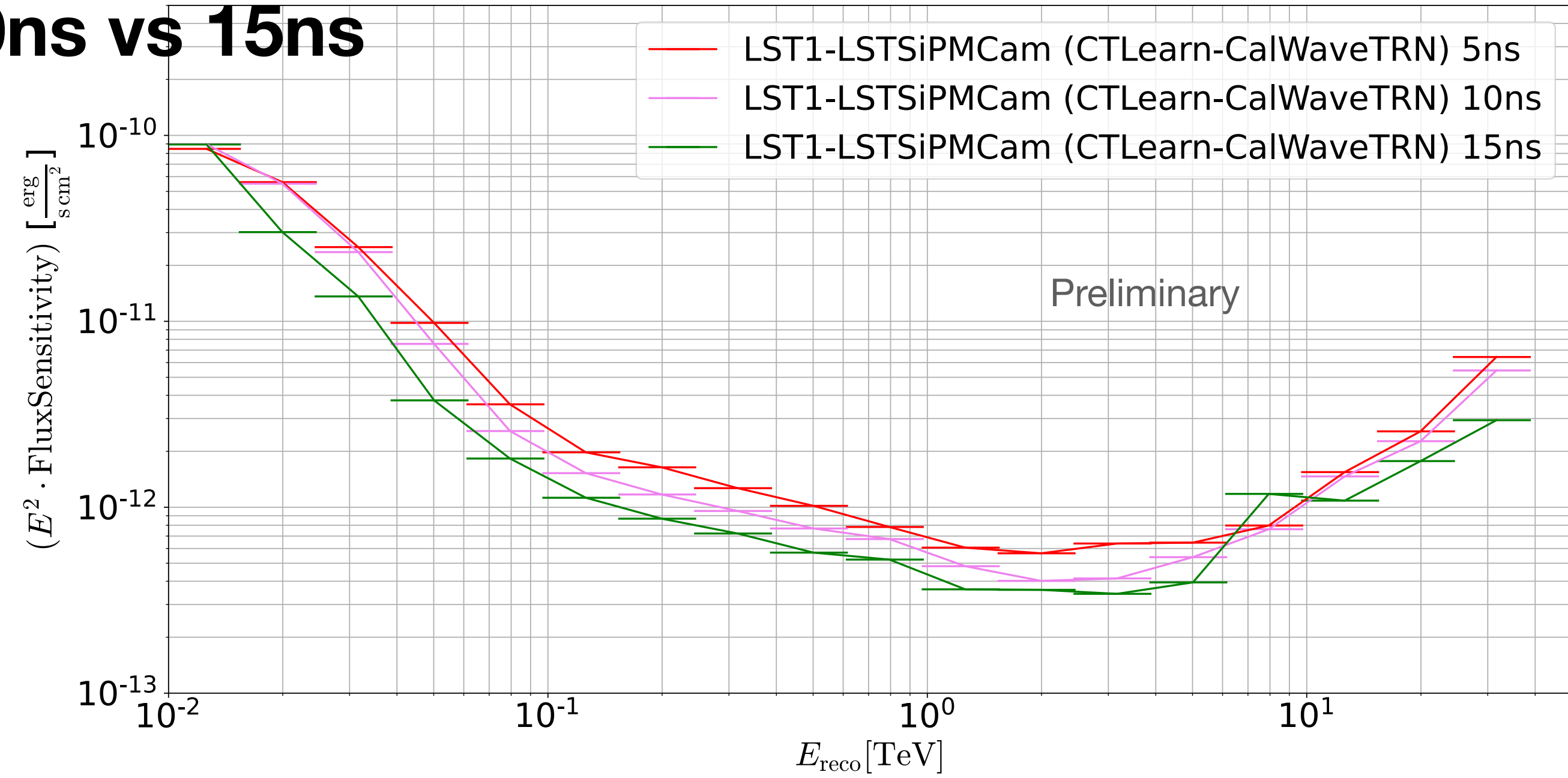
5 ns (low stats) vs 10 ns (low stats) vs 15 ns (full stats)



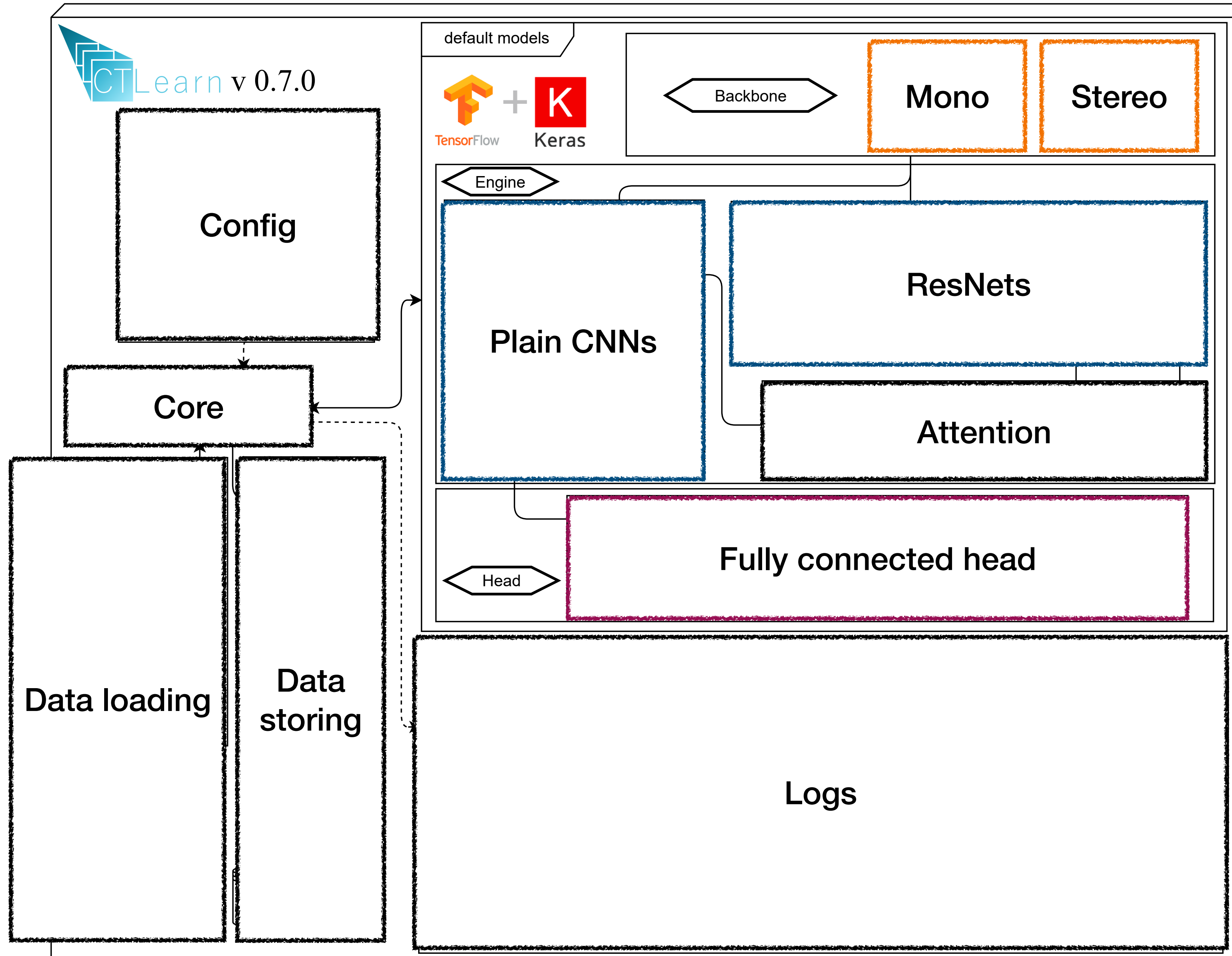
CalWaveTRN

5ns vs 10ns vs 15ns

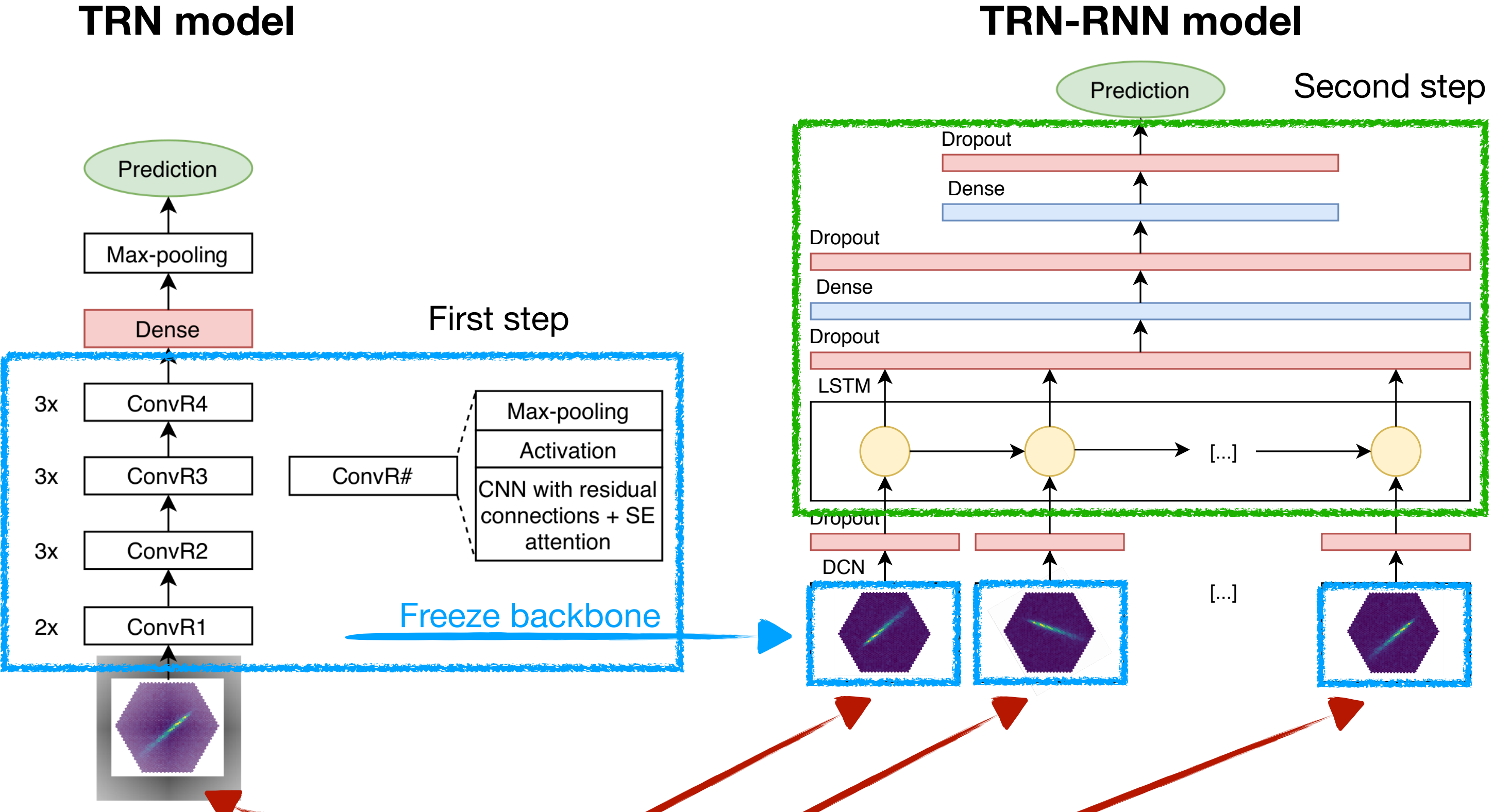
Minimal Flux Needed for 5 σ Detection in 50 hours



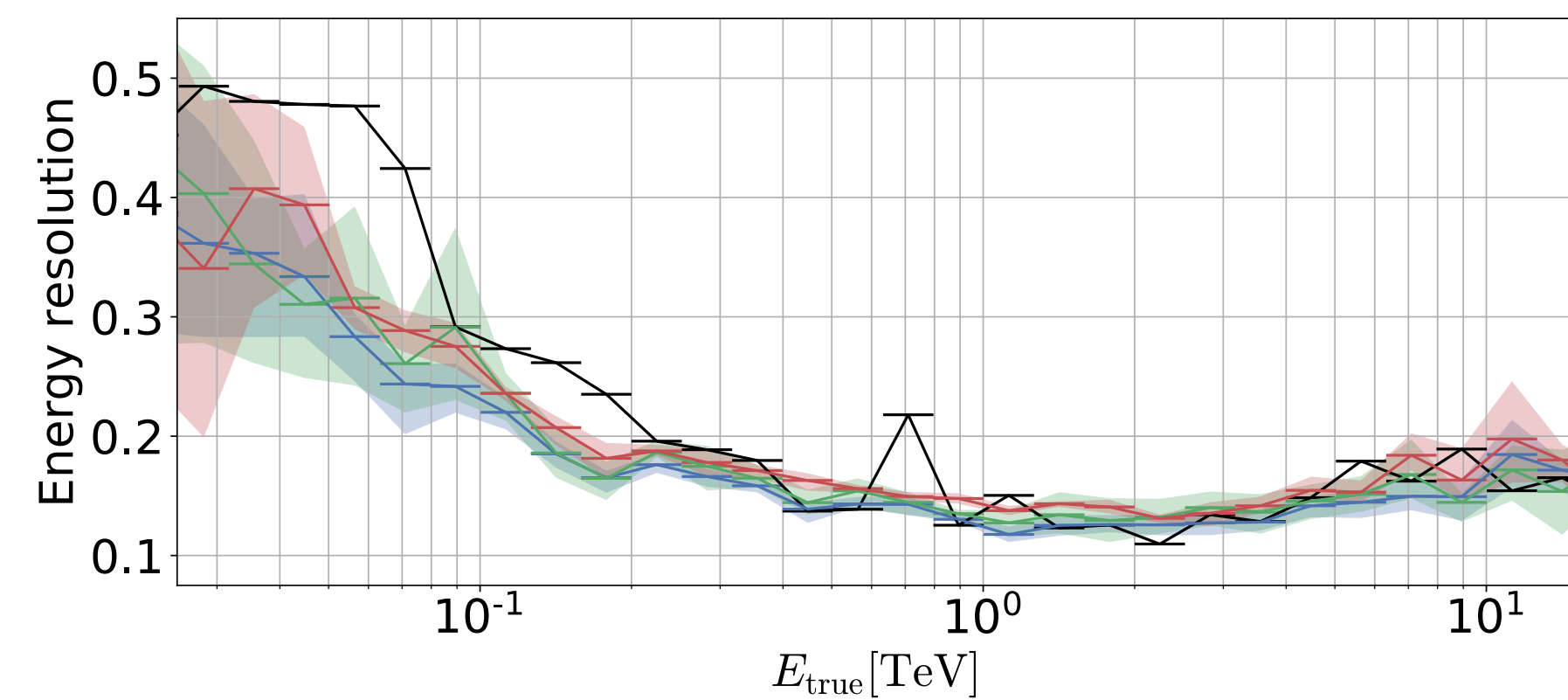
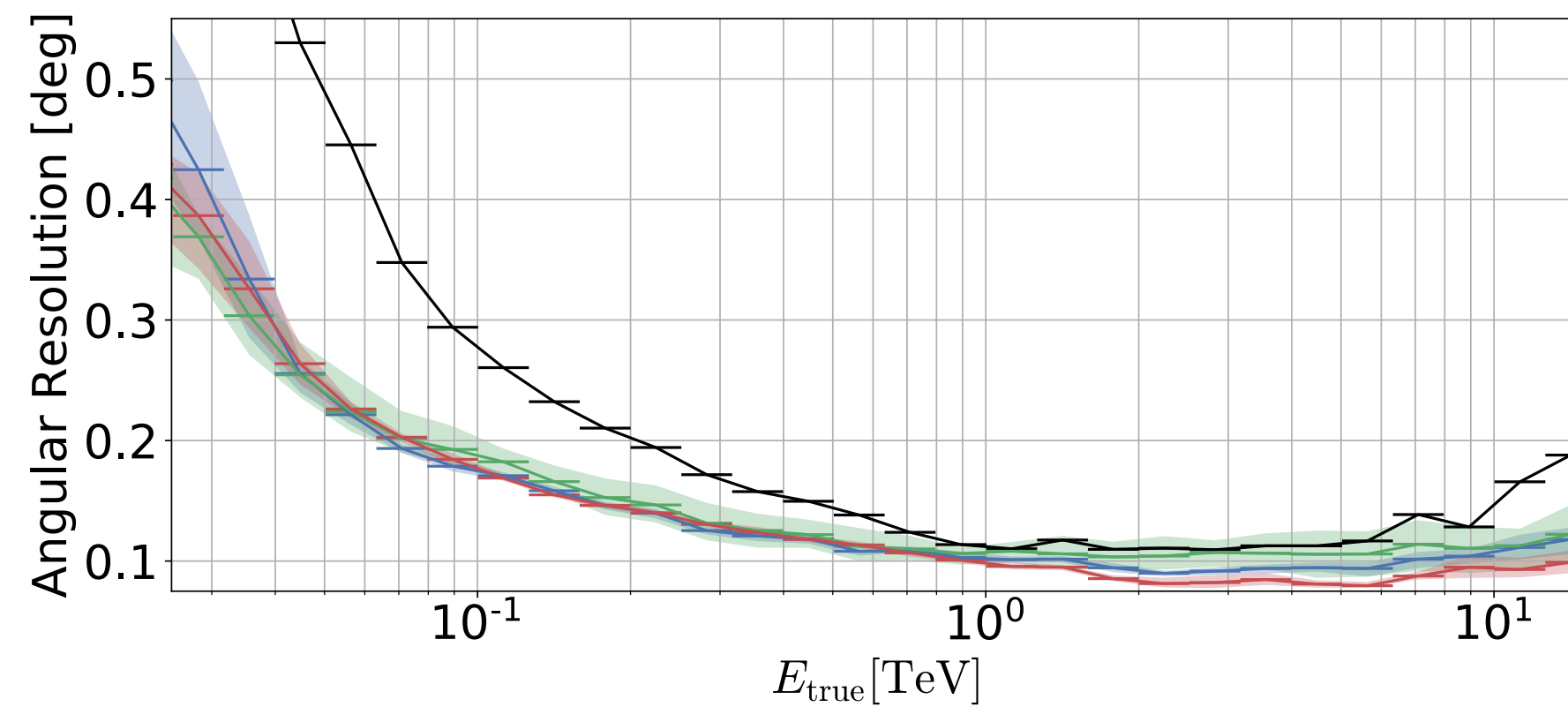
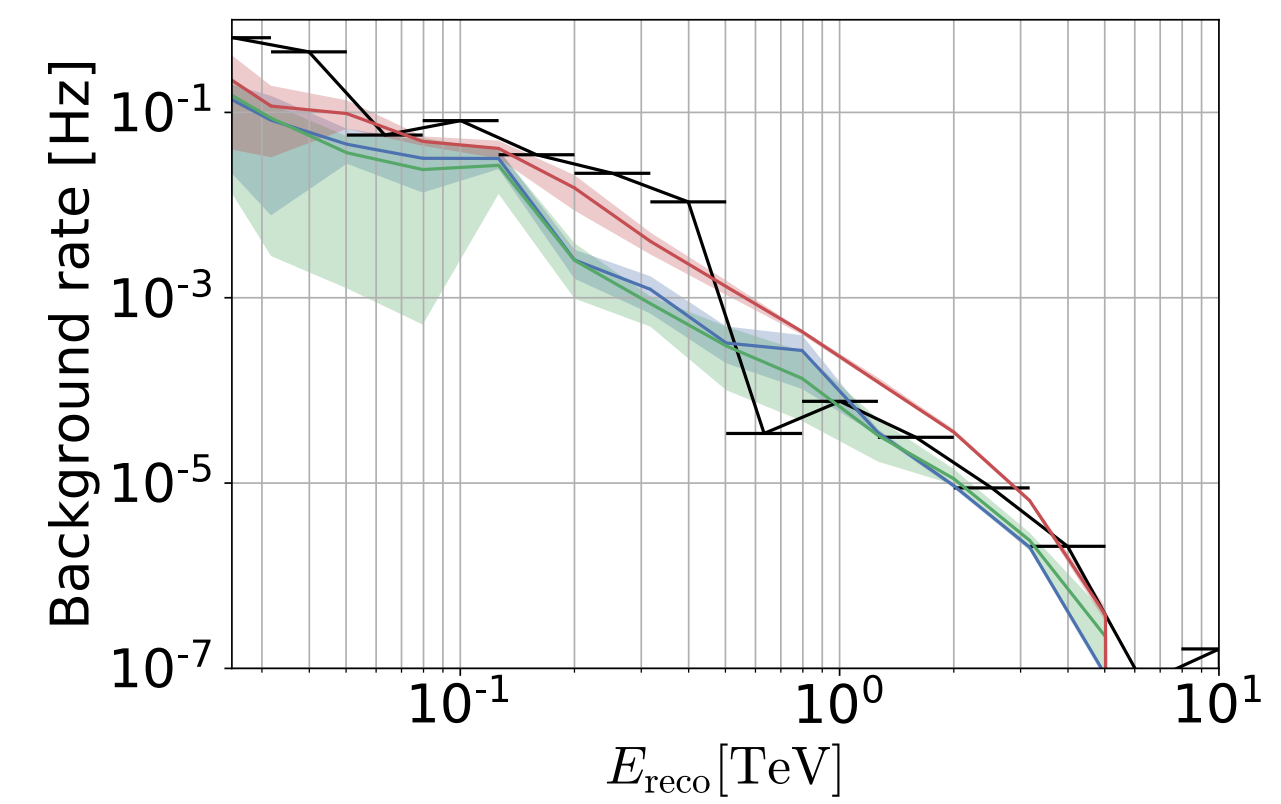
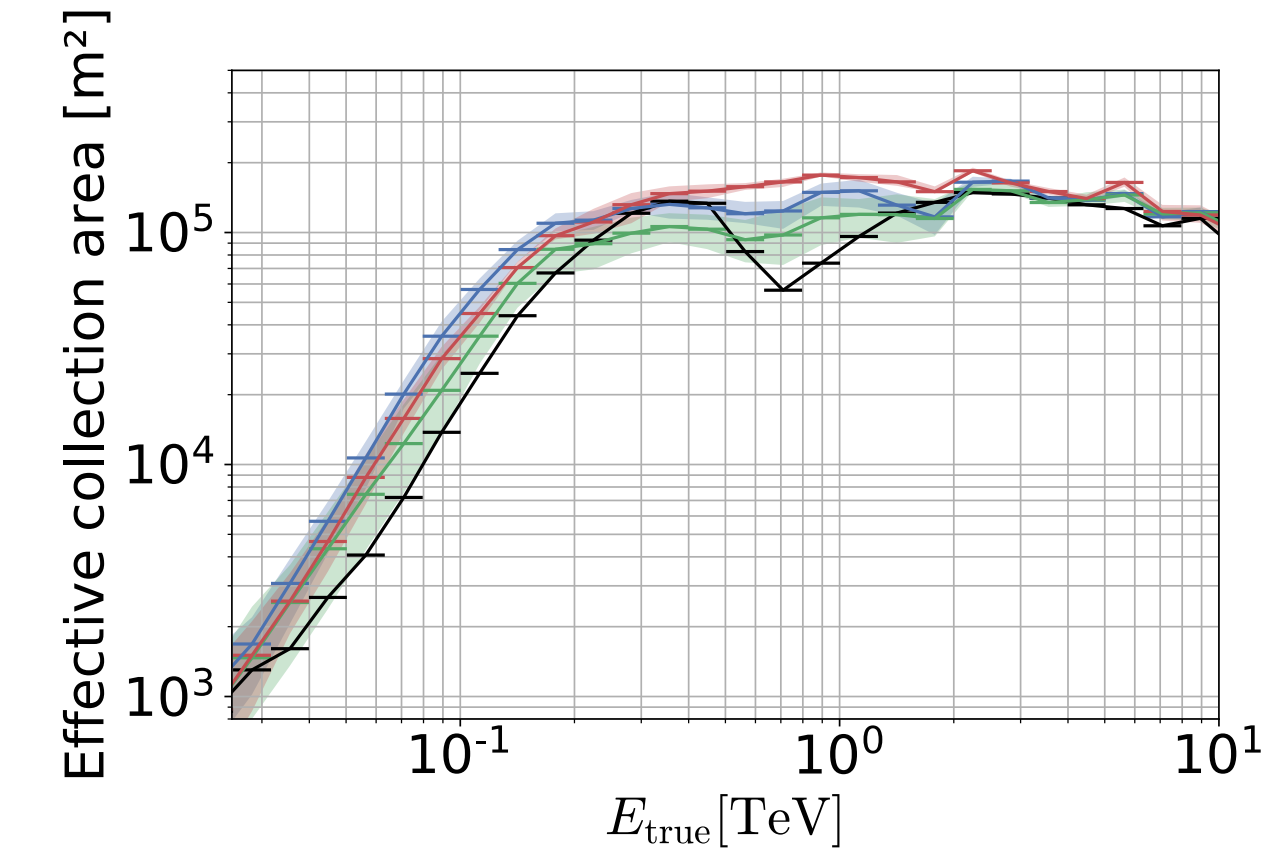
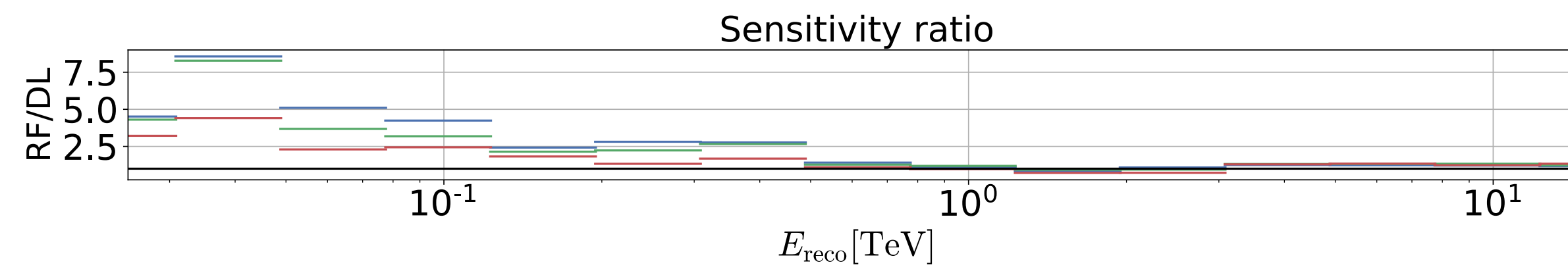
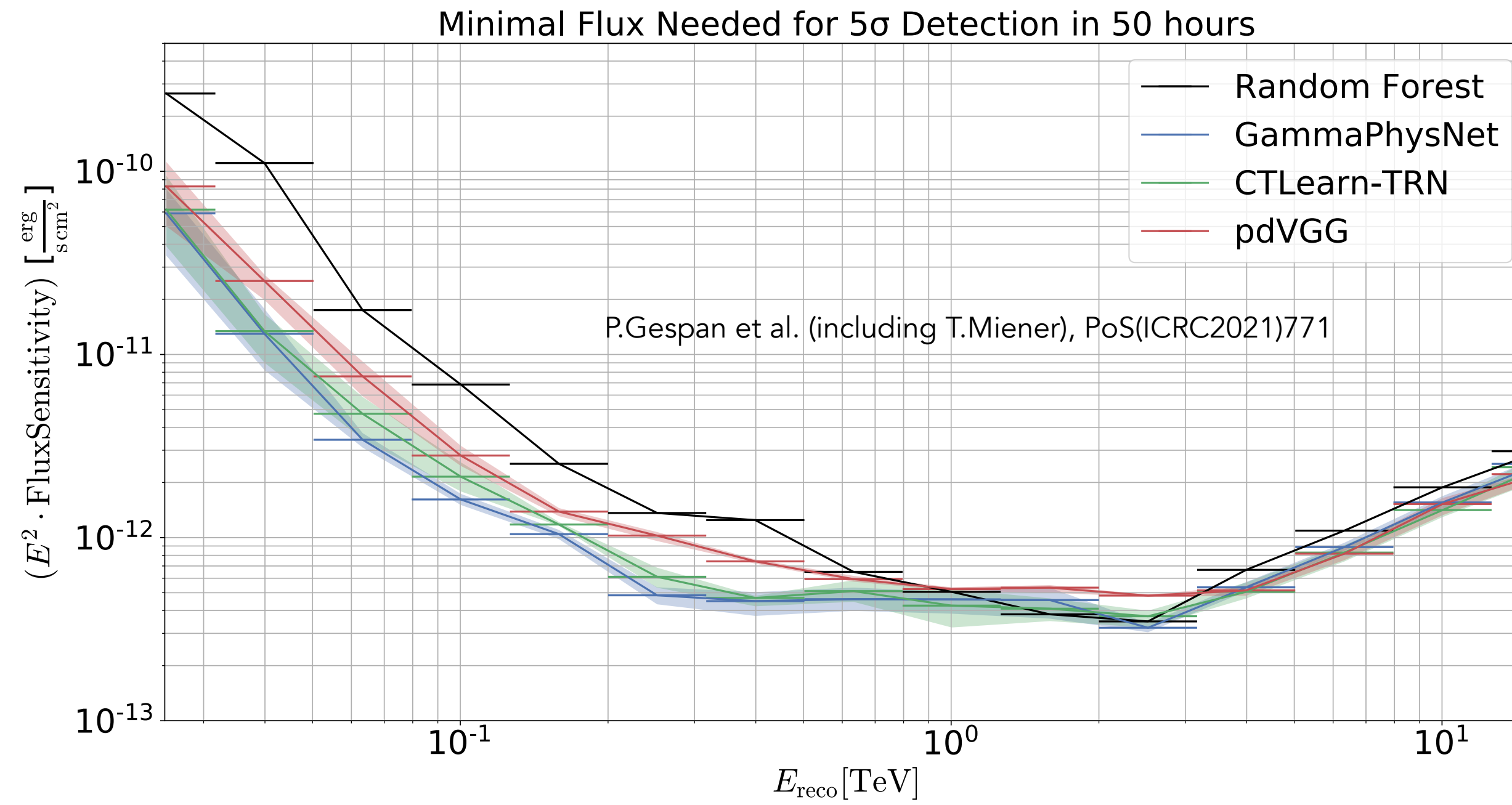
CTLearn workflow



Default CTLearn models (current best-performing)

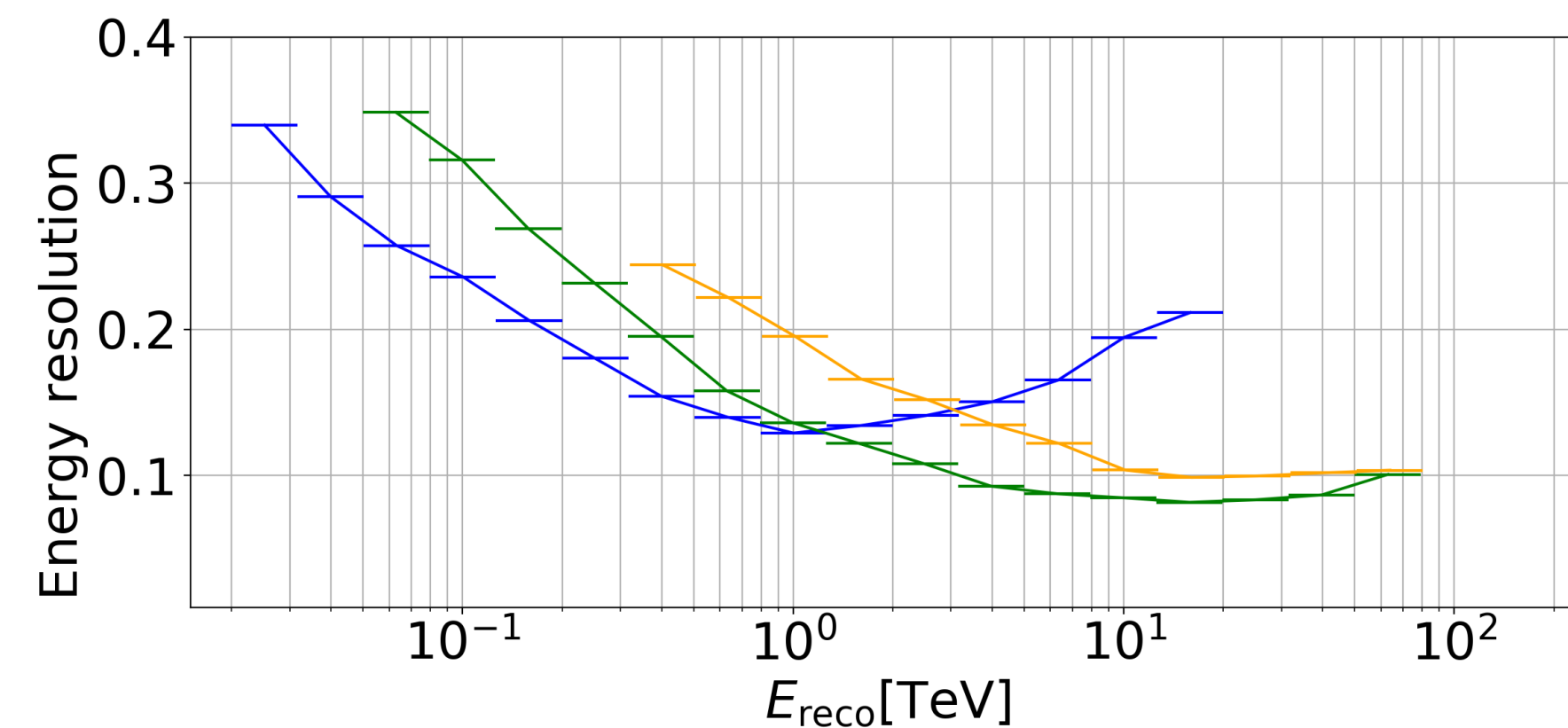
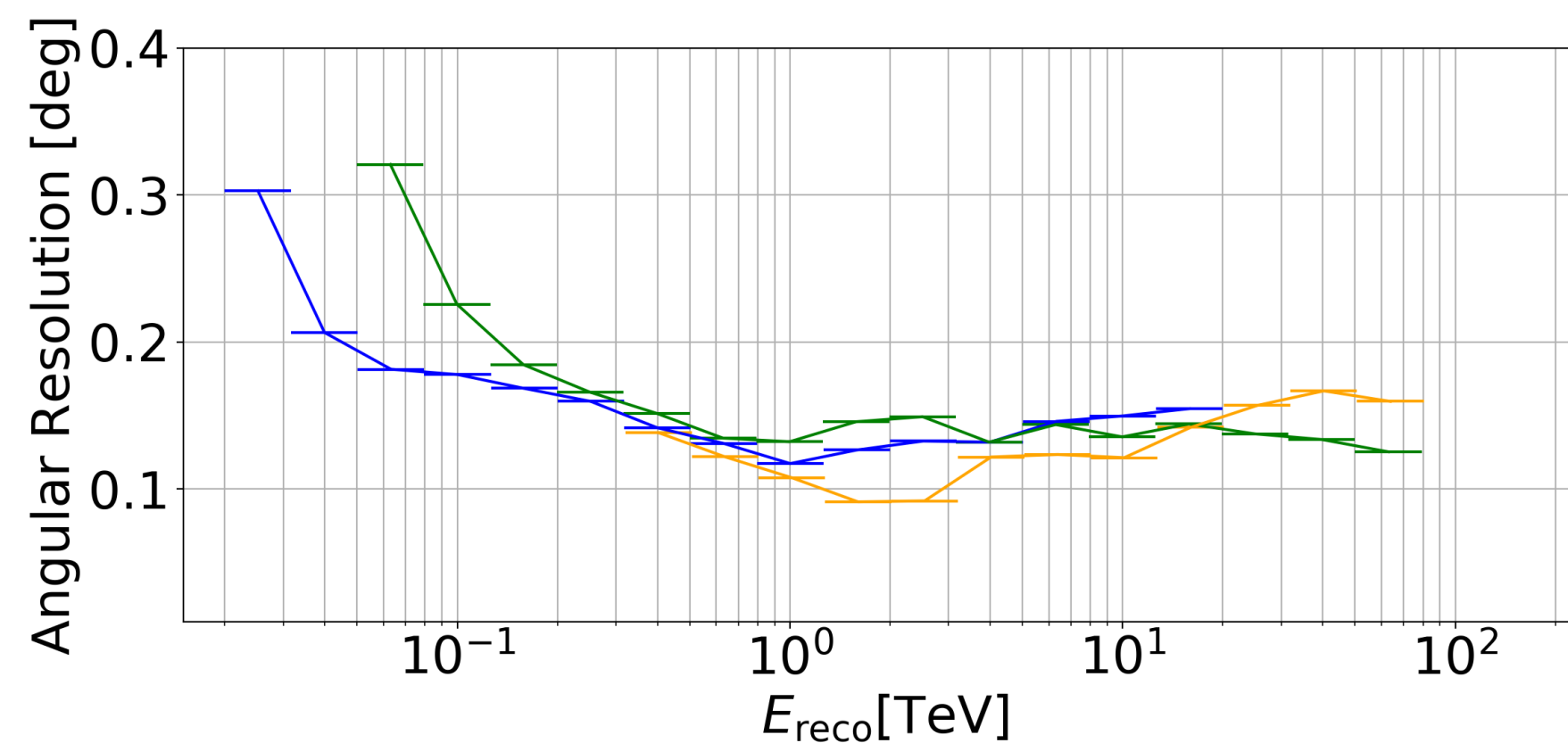
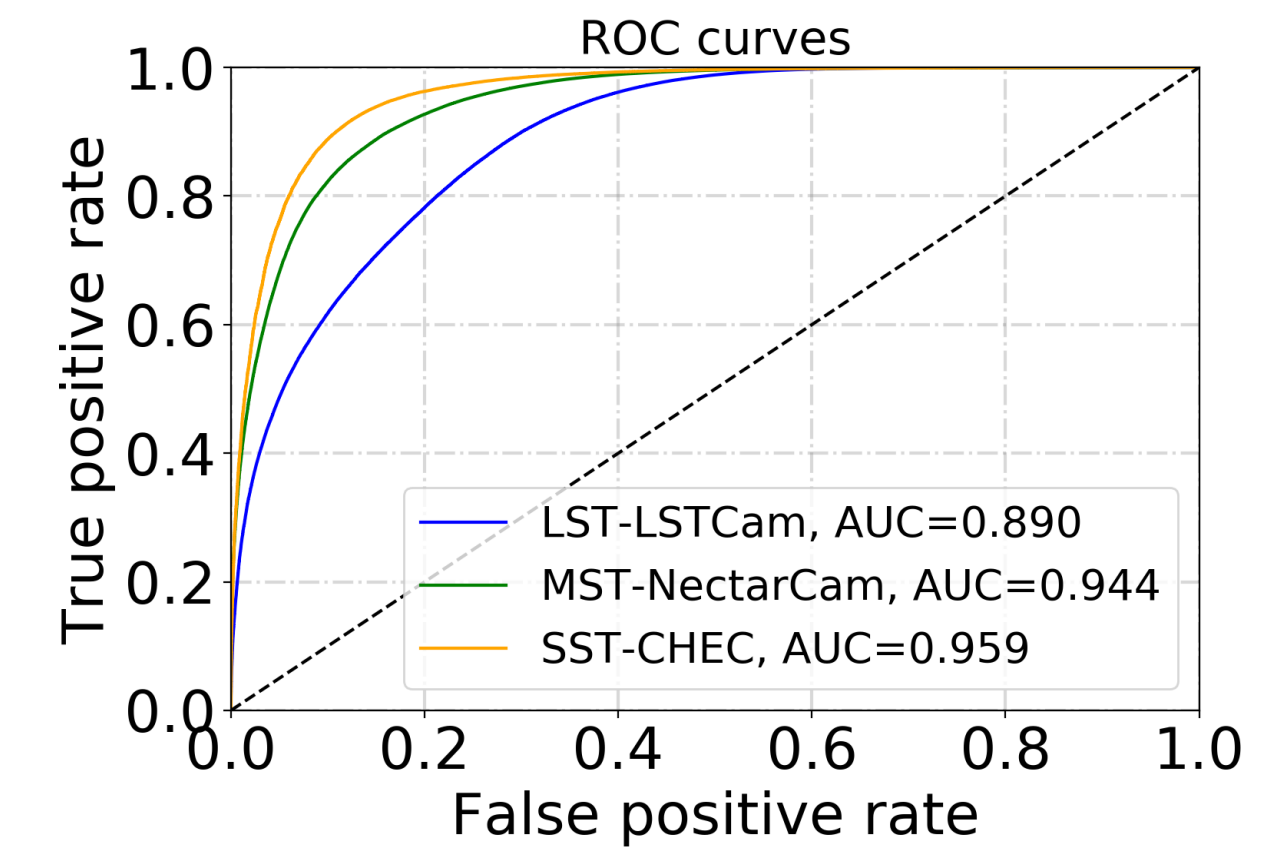
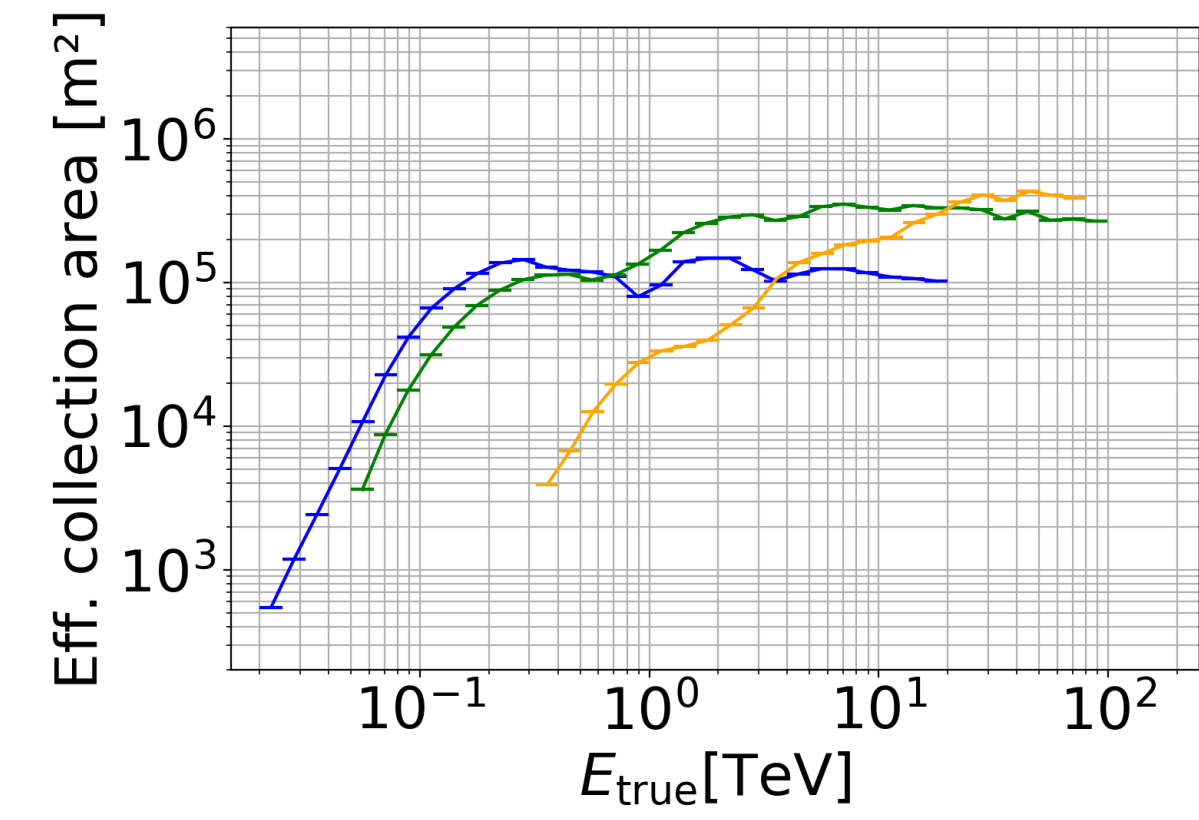
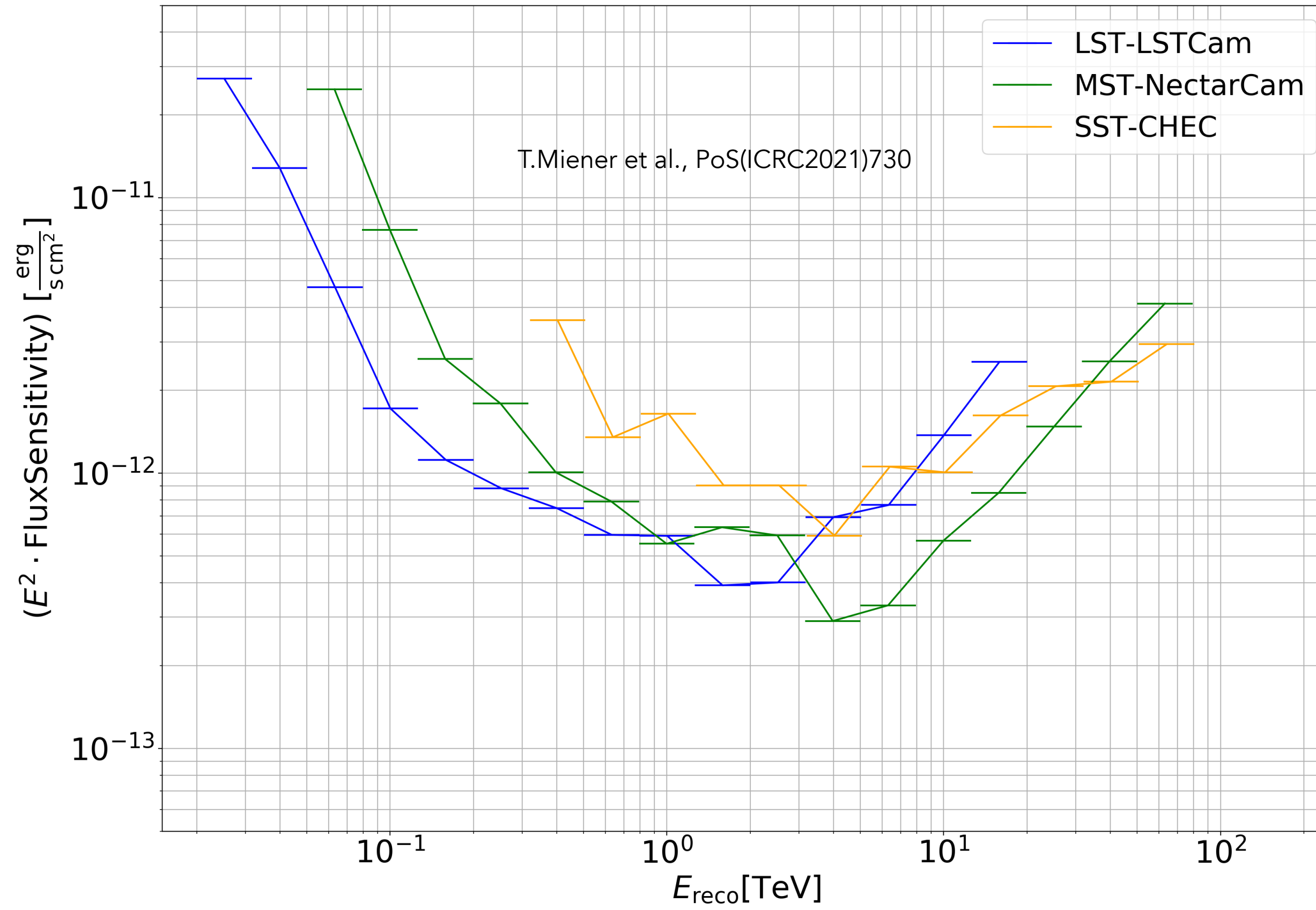


DL-driven event reconstruction applied to simulated data from a single LST of CTA

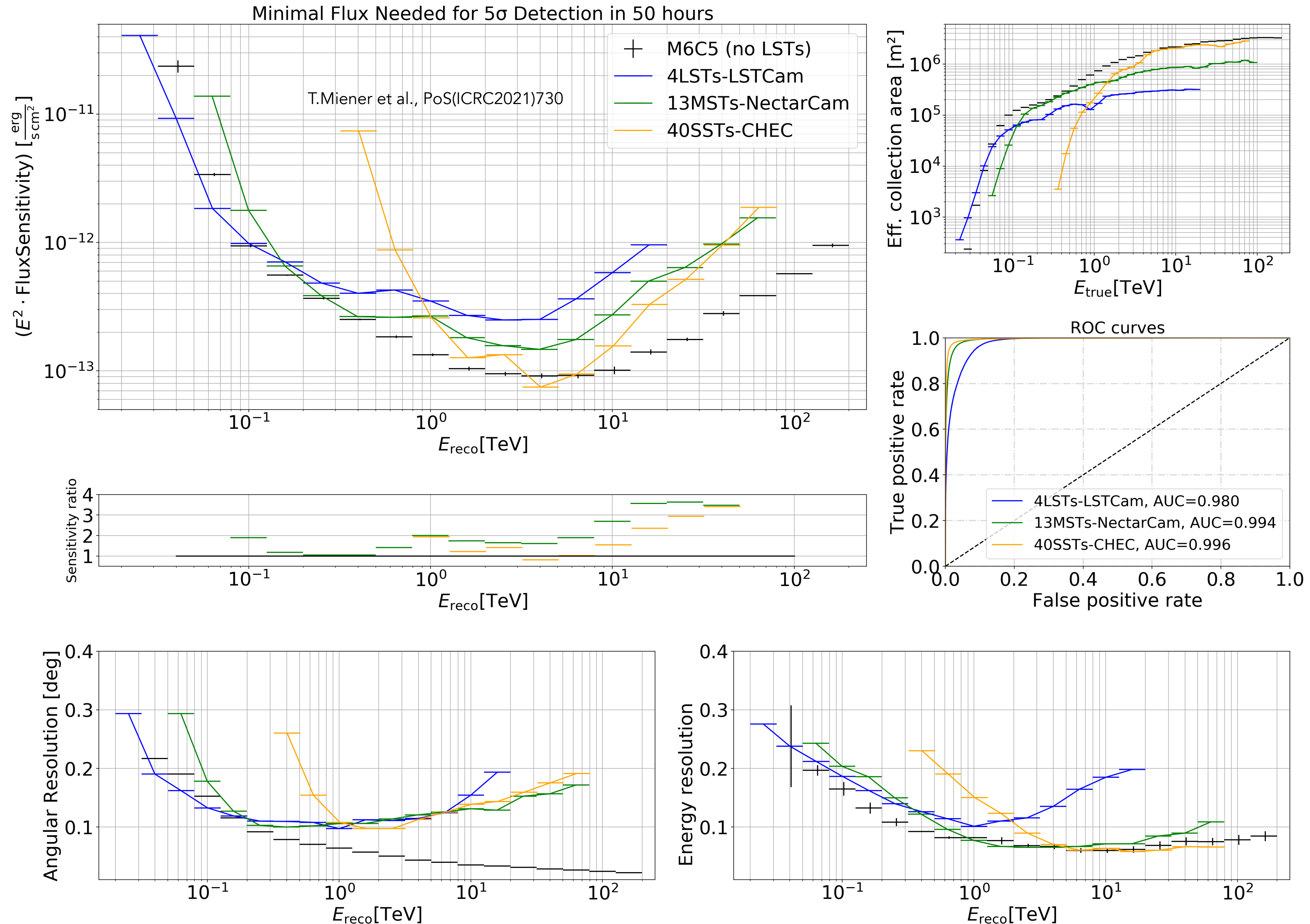


First DCN-based full-event reconstruction on all CTA telescope types (single-telescope)

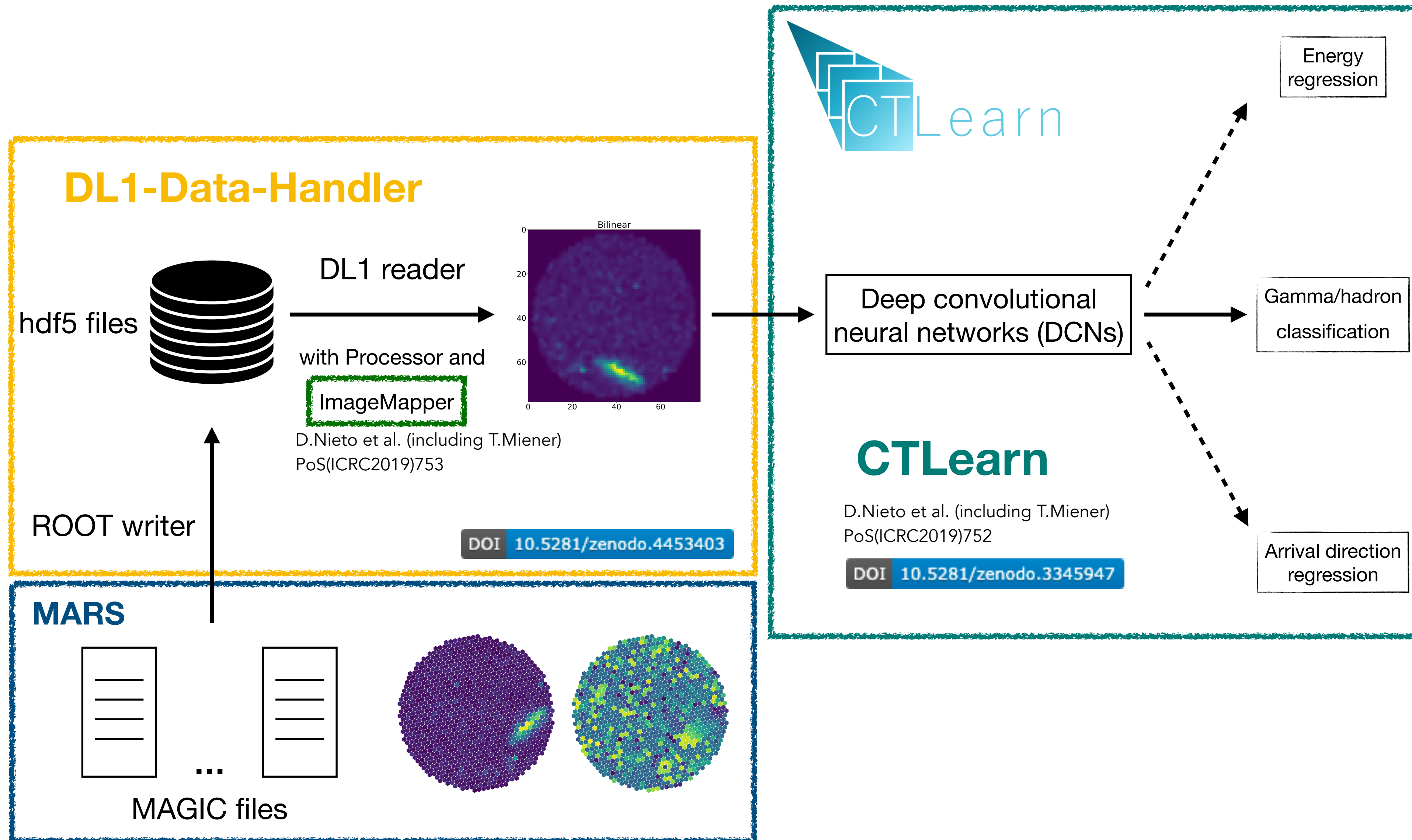
Minimal Flux Needed for 5σ Detection in 50 hours



First DCN-based full-event reconstruction on all CTA telescope types (multi-telescopes)

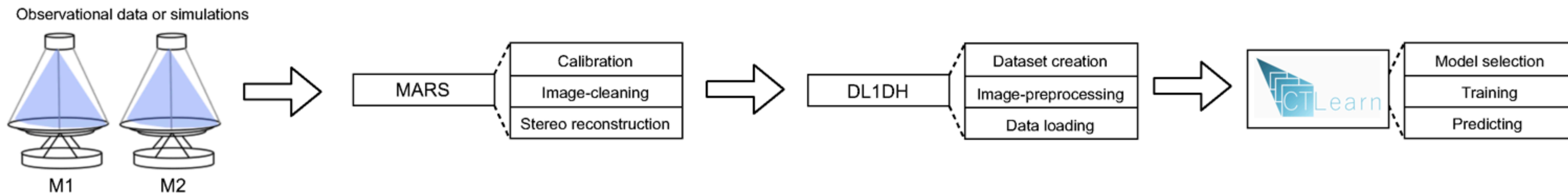
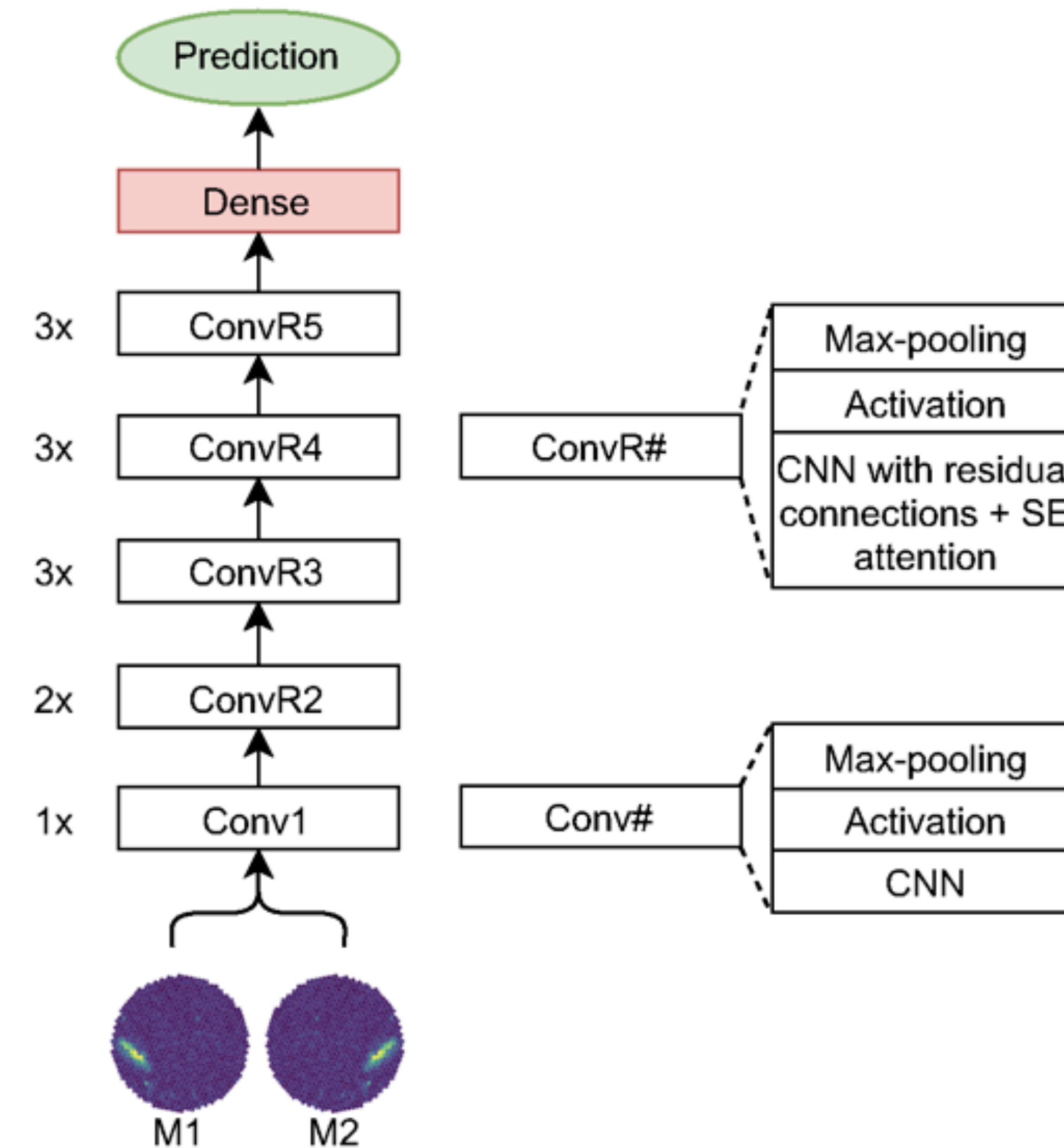


DL application on real data with the MAGIC telescopes

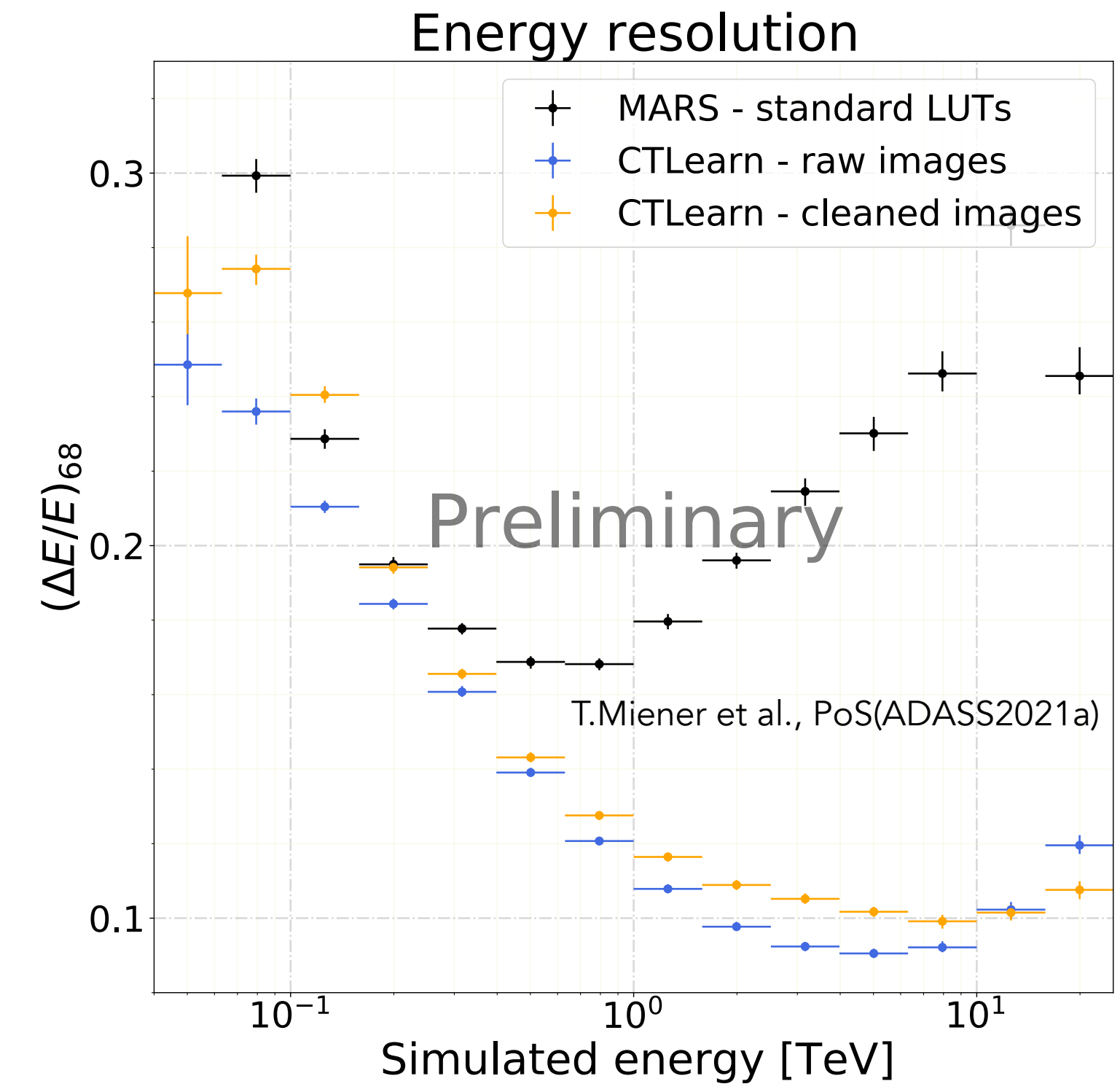
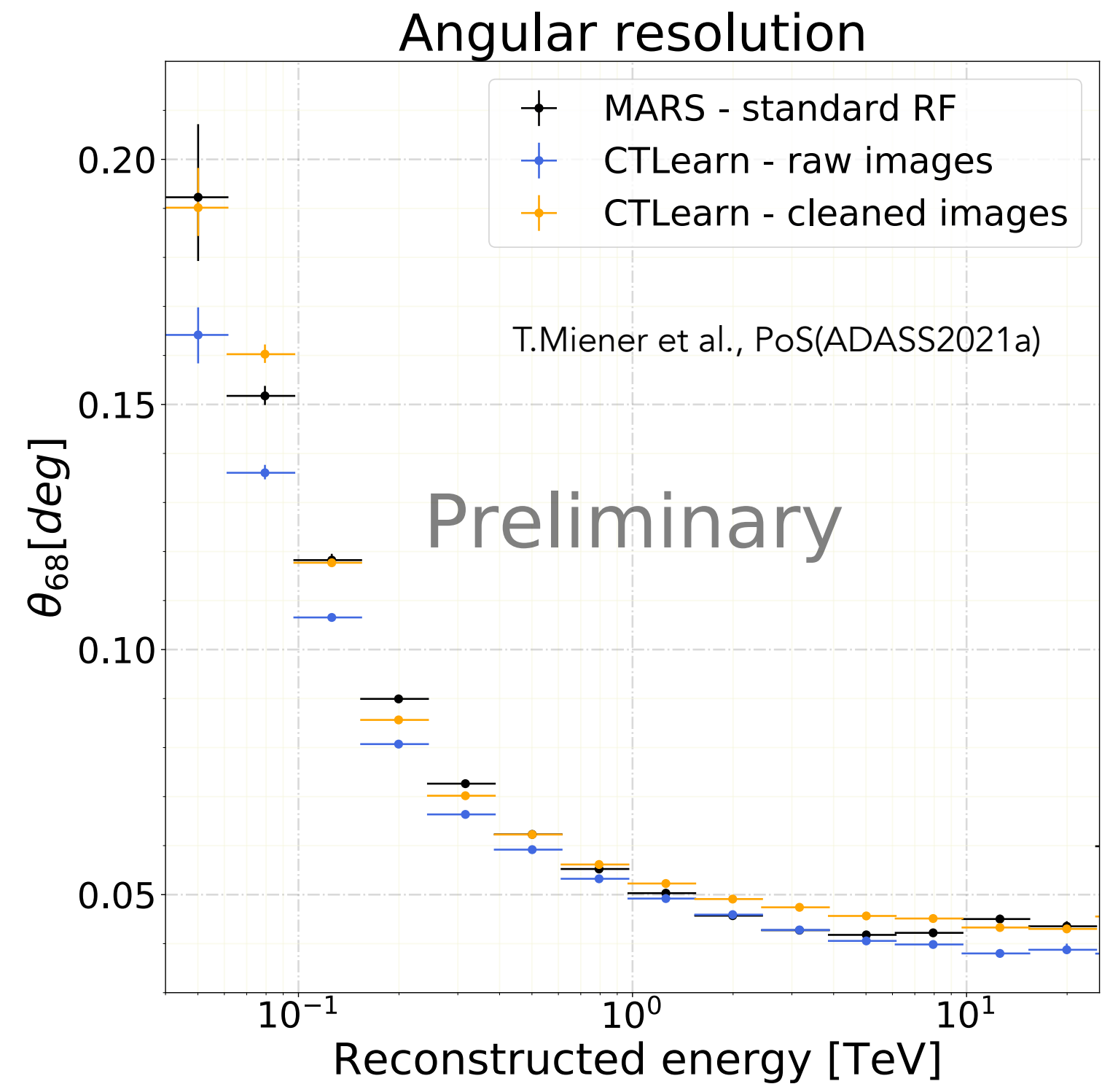
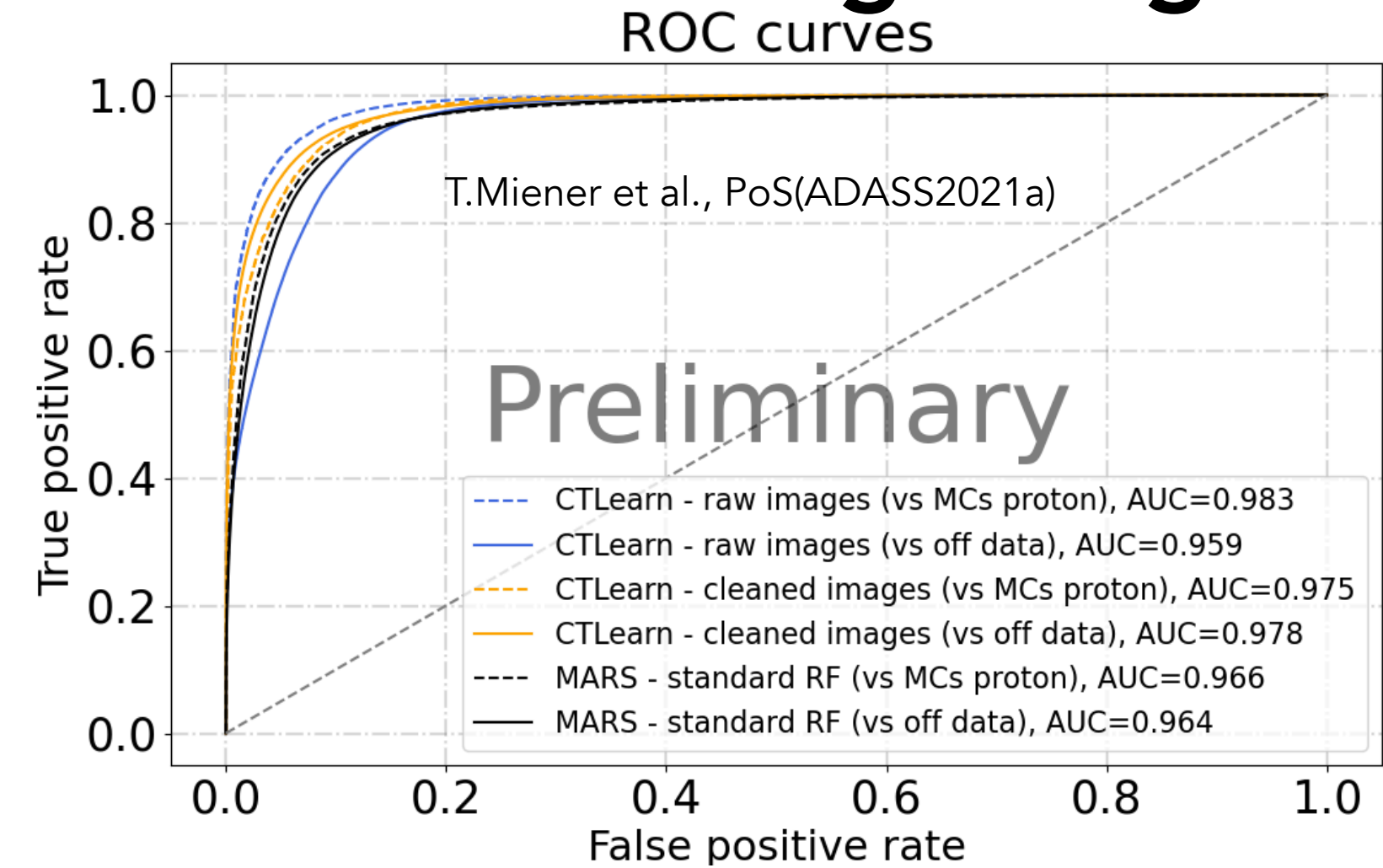


DL application on real data with the MAGIC telescopes

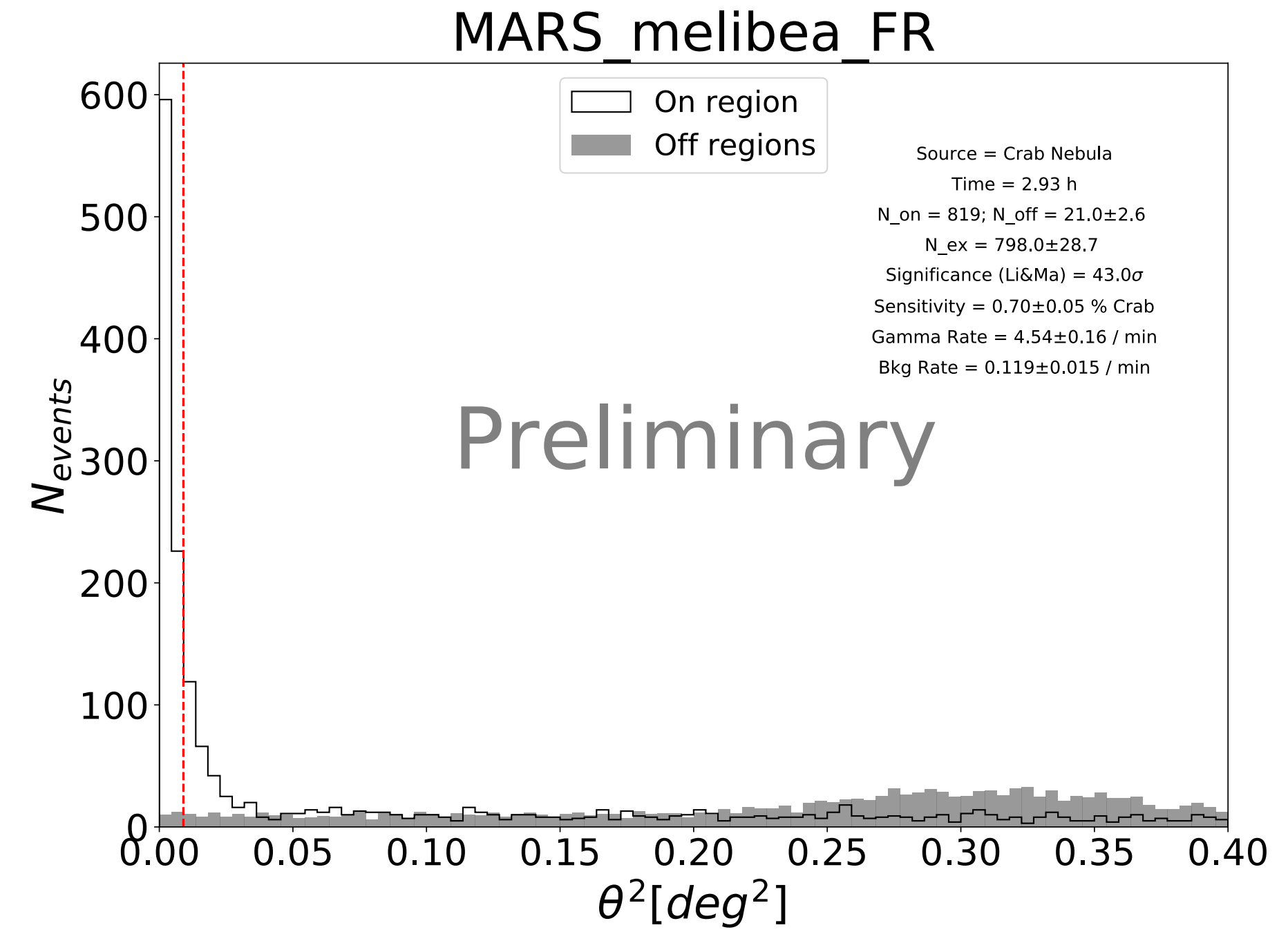
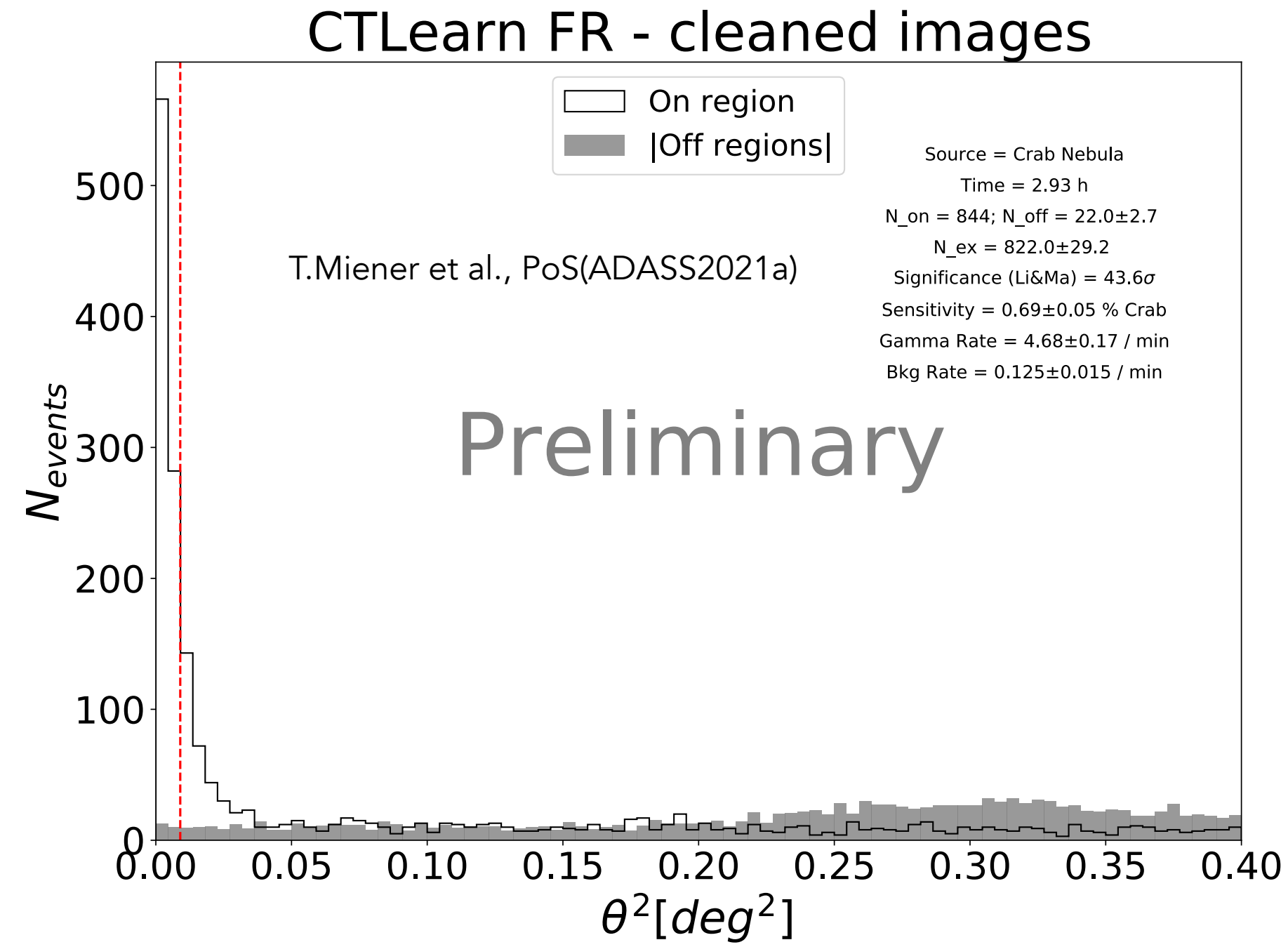
- Establishment of an alternative data processing pipeline for the MAGIC telescopes based on *MARS* and *DL1DH+CTLearn*
- First source detection (Crab Nebula) using DCNs from the *CTLearn* team
- We explored two different analysis schemes, where we trained the same TRN model with raw images and cleaned images
- A Crab Nebula sample of 2.93 h have been analyzed with *CTLearn* and the latest *MARS* software using the standard settings for the analysis focusing of the full range (FR; >250 GeV) - and low energy (LE; >100 GeV) range.



Reconstruction performance using MC gamma simulations



Crab Nebula analysis with CTLearn



Analysis	N _{on}	N _{off}	N _{ex}	γ rate [/min]	bkg rate [/min]	Sen. [% Crab]	Sig. (Li&Ma)
MARS – FR	819	21.0 ± 2.6	798.0 ± 28.7	4.54 ± 0.16	0.119 ± 0.015	0.70 ± 0.05	43.0σ
CTLearn – FR (raw)	629	23.3 ± 3.1	605.7 ± 25.3	3.45 ± 0.14	0.133 ± 0.018	0.97 ± 0.08	36.5σ
CTLearn – FR (cleaned)	844	22.0 ± 2.7	822.0 ± 29.2	4.68 ± 0.17	0.125 ± 0.015	0.69 ± 0.05	43.6σ
MARS – LE	3579	679.0 ± 15.0	2900.0 ± 61.7	16.49 ± 0.35	3.861 ± 0.086	1.09 ± 0.03	61.1σ
CTLearn – LE (raw)	2730	673.7 ± 20.0	2056.3 ± 56.0	11.70 ± 0.32	3.832 ± 0.114	1.53 ± 0.05	47.5σ
CTLearn – LE (cleaned)	3536	680.7 ± 15.1	2855.3 ± 61.3	16.24 ± 0.35	3.872 ± 0.086	1.11 ± 0.03	60.4σ

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