

VZERO Status Report

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On behalf of
VZERO Team

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Alice Offline Week

Where is VZERO used?

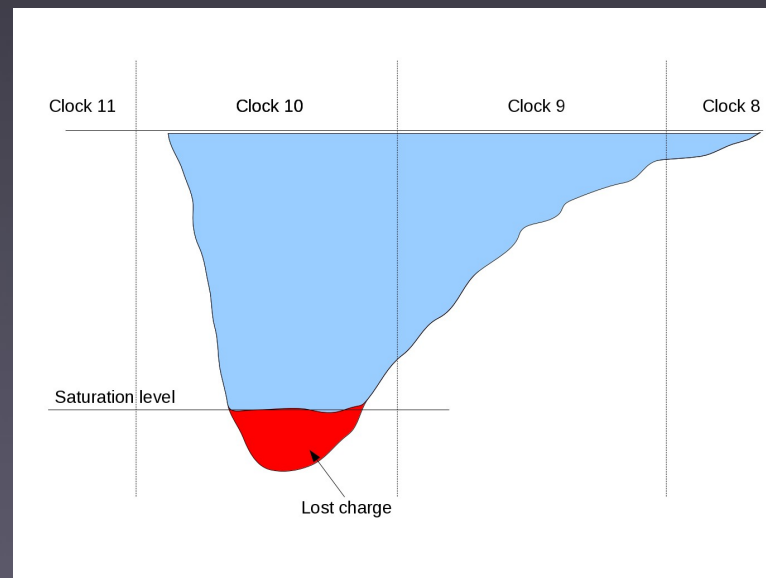
- Main L0/MB Trigger – pp & PbPb
- Offline trigger
 - Beam-Gas removal – pp & PbPb
 - Event classes (INEL, NSD, SD) - pp
- Luminosity & cross-section measurements – V0AND
 - Cross-section normalization
- Default centrality determination in PbPb
- Event-plane determination in PbPb
- New: multiplicity measurement in PbPb

Calibration

- Pedestals – online via DA
- Time offsets – no calibration, estimated directly from various sources (clock phase from BPTX, CTP & LTU configs, FEE config etc)
- Time delays (channel-by-channel) – done once in 2009, has to be redone only in case of hardware intervention (change of PM)
- Time-slewing – no calibration, parametrized was a function of all relevant data (signal amplitude and TDC thresholds), no need to redo it
- Light collection efficiencies – obtained once with 2010 pp data
- PM gain curves – so far using data from the lab
- TDC thresholds – effective vs FEE values, done once on HV-scan runs, intrinsic property of the detector
- Signal saturation in 2010 PbPb data – strictly speaking not a calibration, but rather a recovery

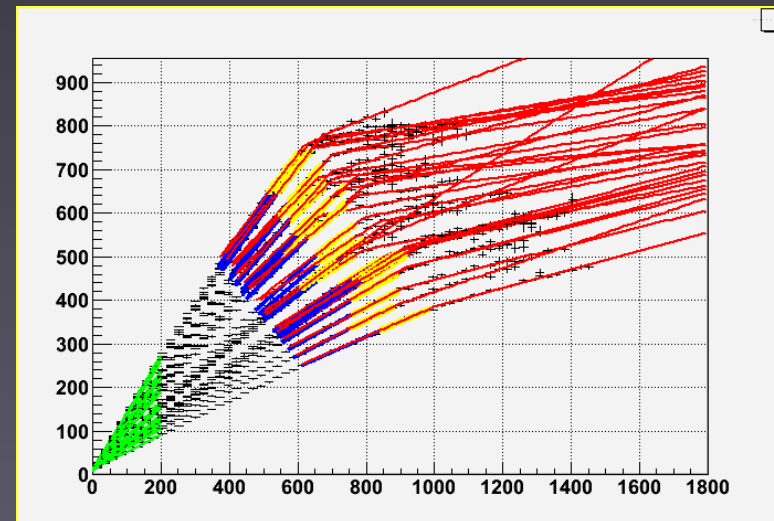
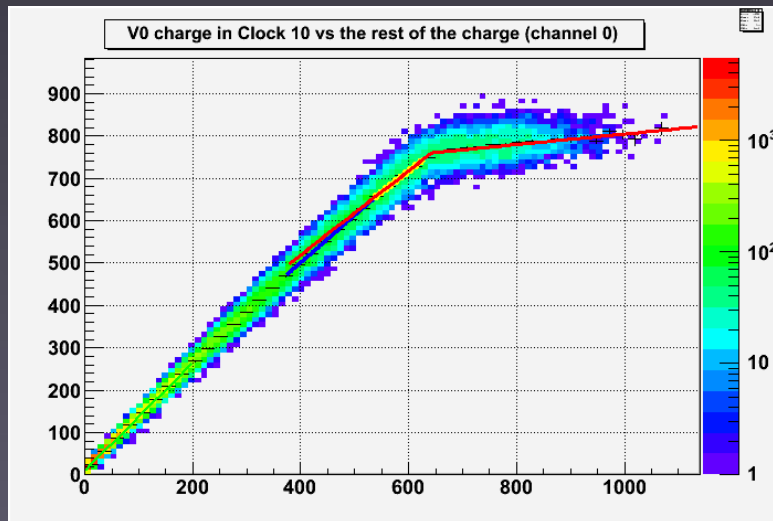
Saturation (1)

- From studies of charge distribution among ADC samples:
 - Preamplifier signal clipping
 - @ ~ 1300 ADC which corresponds $\sim 2.5V$
 - Clipping observed only in 'central' sample (clock 10)



Saturation (2)

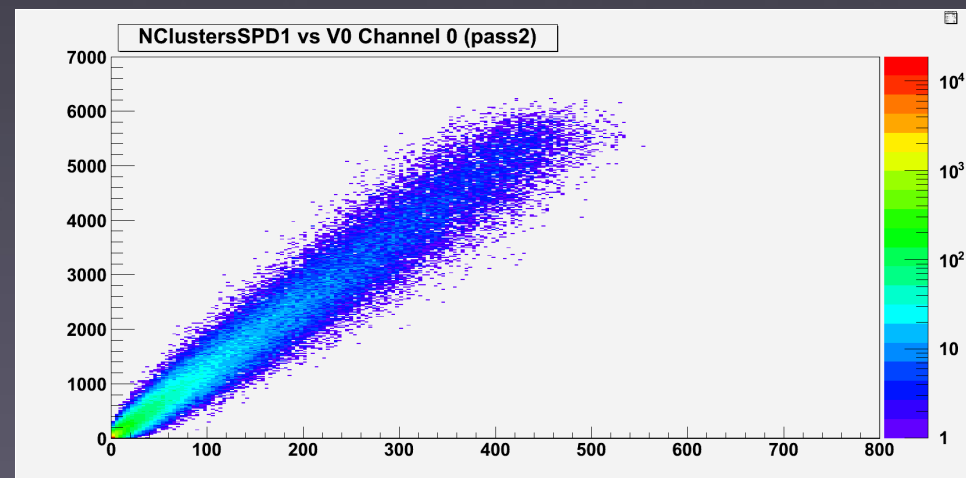
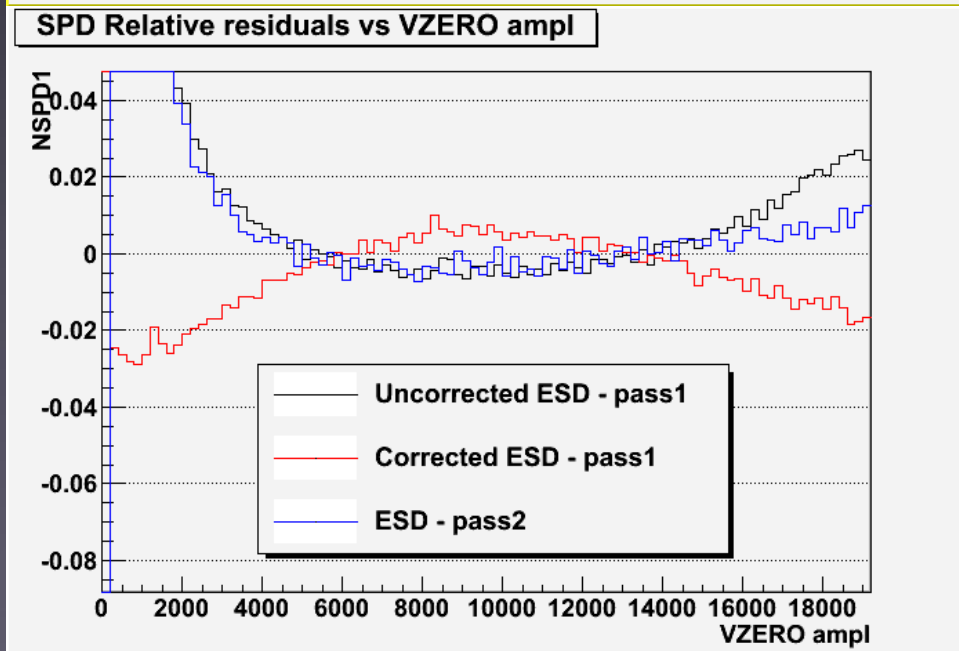
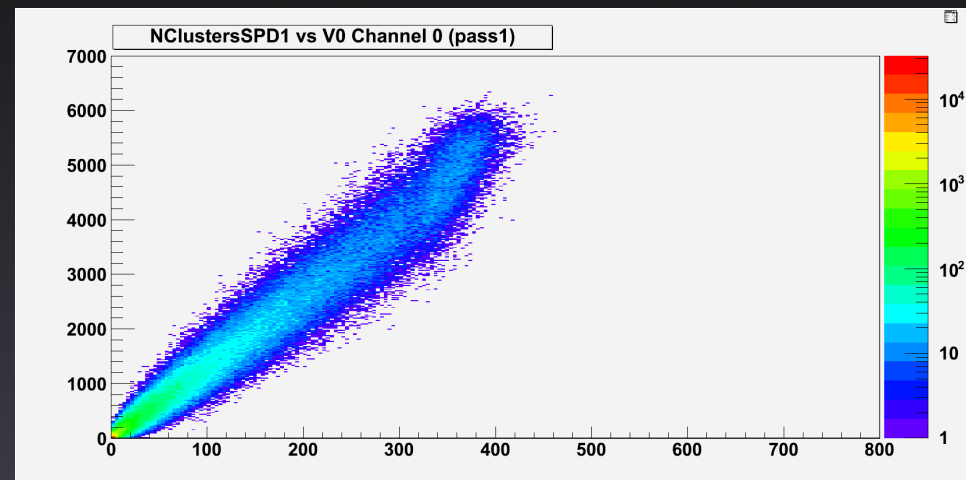
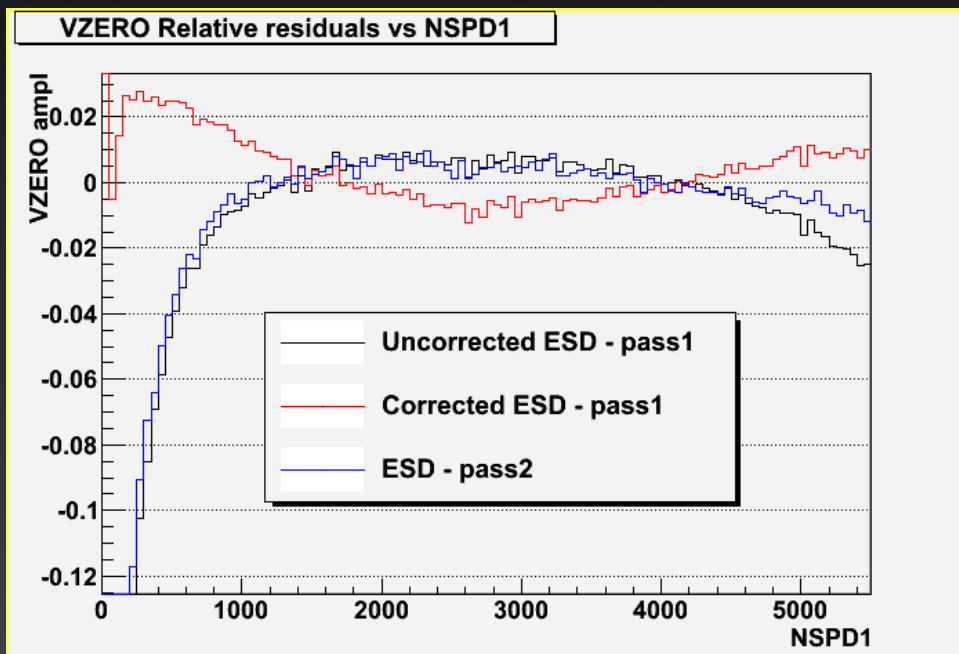
- Estimated by correlating signal in 'clock 10' to rest of the collected charge
- Fitted with two linear functions (below/above saturation threshold)
- Done for each channel individually



Saturation (3)

- Correction applied on total collected charge (including clock 10)
 - Assuming that the reason is signal clipping
 - More robust against clock phase drift
 - No need to calibrate the charge distribution among samples (run-by-run)
- Correction:
 - $A \rightarrow A + \alpha (A - \beta)$, if $A > \beta$

Saturation: Results

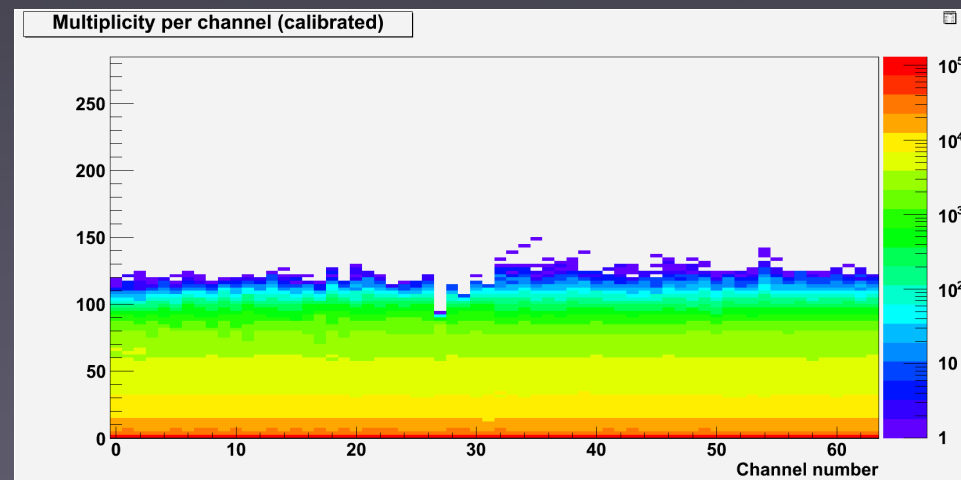
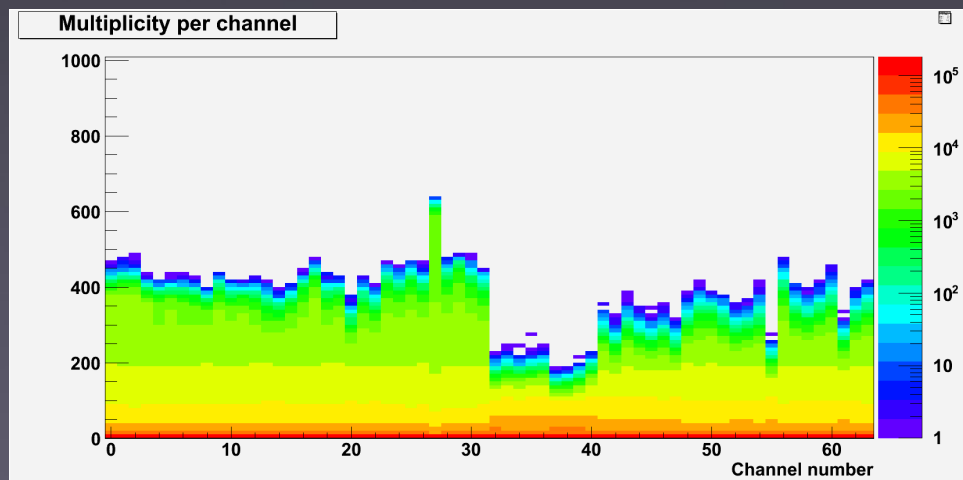


Light coll. eff & PM gains

- VZERO signal =
(# of prim + # of sec + geom effects) x
light yield x light collection eff x
PM gain x FEE response
- Therefore collection eff and PM gains are couple and can't be disentangled using the data
 - MIP signal shape (RMS) is rather broad due to sec + geom effects + low light levels
- So far we use:
 - PM gains from the lab (no field)
 - Fit MIP peak in the data & MC
 - Extract collection eff
- Even if the method works to first order it has caveats especially in channels which behave quite differently from the lab and in channels coupled to two scintillator units

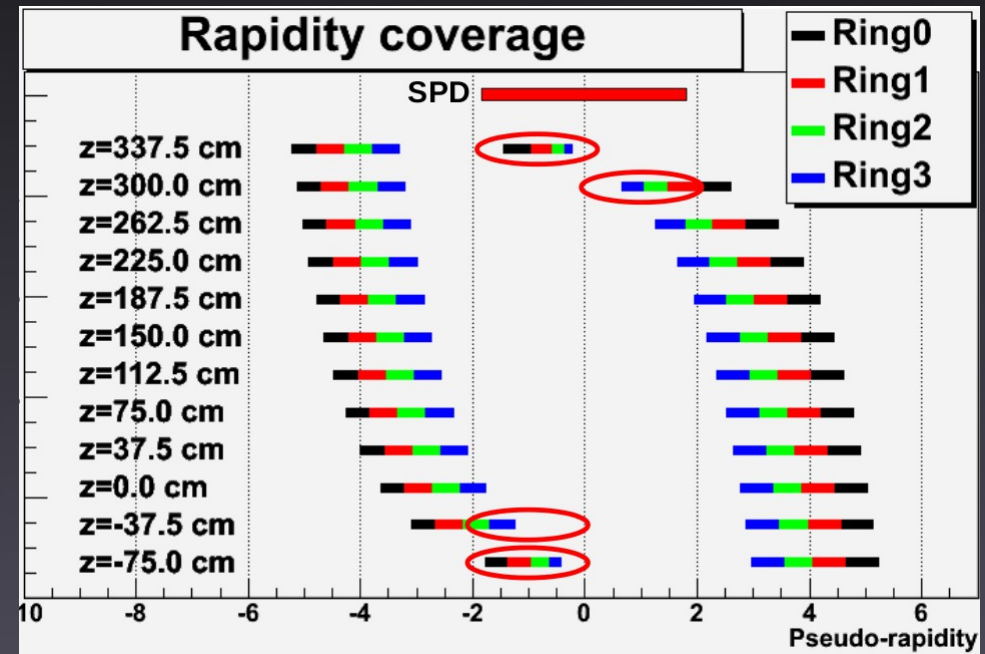
Event-plane measurement

- Needs 'calibrated'/equalized response among VZERO sectors
 - Equalization could be done with very simple 'analysis', no need to calibrate coll.eff/PM gains
 - Presently applied on ESD (should go to OADB)
 - In principle a simple DA would do the job – could be developed for this year PbPb run



Multiplicity measurement in PbPb

- Besides its physics value can provide also 'calibration' of the detector response – signal(# of prim)
- Idea is to use satellite collisions (z @-37.5,-75,..) and results from SPD in the central region
 - Selection of satellite events using ZDC timing cut a la phys sel
 - Secondaries + geom effects corrected using MC
 - ZDC stand-alone centrality selection
- $$dN_{deta} = A(z \sim 0) * \frac{dN_{deta}(SPD, z \sim 0)}{A(z \sim -75)} * \frac{A/dN(z \sim -75)}{A/dN(z \sim 0)}$$



The analysis is underway, more news next week on PWG2 meeting

Summary

- All is good with VZERO ;-)
- The problem with light coll.eff and PM gains
 - Is not a question of calibration framework, as the calibration needs to be done only once per year (assuming const HVs)
 - Is not a question of stat, in principle $\sim 100k$ events would be enough for the purpose
 - Is not a question of which data to use, we need only ESDs + small fraction of friends for pulse shape studies
 - Is a question of (new) ideas how to measure them