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# Study of neutrino mass matrices with vanishing trace and one vanishing minor

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<span id="page-2-0"></span>Neutrino physics originated from  $\beta$ -decay when it was established that the average energy of the electrons produced in the  $\beta$ -decay is significantly smaller than the total energy released.





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- <span id="page-3-0"></span>**Pauli** suggested that some of the energy has been taken away by a new particle neutrino which is emitted in the decay process, which carries energy and have spin  $\frac{1}{2}$  , but which is **massless, neutral and weak interacting**.
- $\blacksquare$  From the standard model of particle physics, we find that the neutrinos  $\nu_e$ ,  $\nu_\mu$  and  $\nu_\tau$  are **massless**.
- In 1956, **Frederick Reines and Clyde L. Cowan, Jr.** first discovered electron neutrino in a nuclear fission reactor.

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## **Neutrino Oscillation**

However **Super-Kamiokande experiment** carried out in 1996 confirmed that neutrinos undergo oscillation giving the concept of neutrino mass.



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**Masses and mixing of three flavors of neutrino (** $\nu_e$ ,  $\nu_u$ ,  $\nu_{\tau}$ ) can be described by a  $(3 \times 3)$  mass matrix  $M_{\nu}$ .

$$
M_{\nu} = \begin{pmatrix} M_{ee} & M_{e\mu} & M_{e\tau} \\ M_{e\mu} & M_{\mu\mu} & M_{\mu\tau} \\ M_{\tau e} & M_{\tau\mu} & M_{\tau\tau} \end{pmatrix}
$$
 (1)

We can rewrite the neutrino mass matrix as

$$
M_{\nu} = V \begin{pmatrix} m_1 & 0 & 0 \\ 0 & m_2 & 0 \\ 0 & 0 & m_3 \end{pmatrix} V^T
$$
 (2)

 $\blacksquare$  V is the diagonalizing PMNS matrix, parametrized as  $V = UP_\nu$  and  $(m_1, m_2, m_3)$  are the neutrino mass eigenvalues.

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- $\blacksquare$   $M_{\nu}$  is parametrized by a total of nine parameters  $(m_1, m_2, m_3, \theta_{12}, \theta_{13}, \theta_{23}, \delta, \alpha, \beta).$
- Out of these nine parameters only five of them are measured by neutrino oscillation experiments. They are the three mixing angles  $(\theta_{12}, \theta_{13}, \theta_{23})$  and two mass squared differences ( $\Delta m^2_{21}, \Delta m^2_{32}$ ).

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- Inspite of the tremendous effort and progress, the neutrino mass matrix has not been fully determined, by any of the conceivable set of feasible experiments.
- **Although there is increasing information on the numerical** values of these parameters, the origin of leptonic flavor structure is still a mystery.
- $\blacksquare$  The smallness of the neutrino mass, mass hierarchy, origin of CP violation are some of the striking questions whose answers are still lacking in the standard model of particle physics.

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<span id="page-8-0"></span>At first we reconstructed the neutrino mass matrix  $M_{\nu}$  as

$$
M_{\nu} = V \begin{pmatrix} m_1 & 0 & 0 \\ 0 & m_2 & 0 \\ 0 & 0 & m_3 \end{pmatrix} V^T
$$
 (4)

Here  $(m_1, m_2, m_3)$  are the neutrino mass eigenvalues and V is the diagonalising PMNS matrix.

$$
M_{\nu} = U \begin{pmatrix} \lambda_1 & 0 & 0 \\ 0 & \lambda_2 & 0 \\ 0 & 0 & \lambda_3 \end{pmatrix} U^T,
$$
 (5)

$$
ext{ where } \lambda_1 = m_1, \lambda_2 = m_2 e^{2i\alpha}, \lambda_3 = m_3 e^{2i(\beta + \delta)}.
$$

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<span id="page-9-0"></span>■ We have six cases of one vanishing minor.  $\blacksquare$  The minors of the off-diagonal elements can be given by:

$$
C_{mn} = (-1)^{m+n}((M_{\nu})_{(m+1,n-1)}(M_{\nu})_{(m+2,n+1)}
$$
  
-(M<sub>ν</sub>)<sub>(m+1,n+1)</sub>(M<sub>ν</sub>)<sub>(m+2,n+2)</sub>) (6)

## and for diagonal elements:

$$
C_{mm} = (-1)^{2m} ((M_{\nu})_{(m+1,m+1)} (M_{\nu})_{(m+2,m+2)} - (M_{\nu})_{(m+1,m+2)} (M_{\nu})_{(m+2,m+1)})
$$
(7)

For  $m+l$ ,  $n+l > 3$ , we have to take the values  $(m + l) - 3$ ,  $(n + l) - 3$ . Here m, n can take values (1, 2, 3) and  $l = 1, 2$ .

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## $\blacksquare$  The condition for vanishing minor

$$
C_{mn} = 0, \quad C_{mm} = 0 \tag{8}
$$

## Solving  $Eq(8)$  we get

$$
m_1 m_2 e^{2i\alpha} A_3 + m_2 m_3 e^{2i(\alpha + \beta + \delta)} A_1 + m_3 m_1 e^{2i(\beta + \delta)} A_2 = 0
$$
<sup>(9)</sup>

$$
A_i = (U_{pj}U_{qj}U_{rk}U_{sk} - U_{tj}U_{uj}U_{vk}U_{wk}) + (j \longleftrightarrow k)
$$
 (10)

## here (i,j,k) is a cyclic permutation of (1,2,3). Therefore the two constraint equations becomes

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$$
\lambda_1 \lambda_2 A_3 + \lambda_2 \lambda_3 A_1 + \lambda_3 \lambda_1 A_2 = 0 \tag{11}
$$

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$$
\lambda_1 + \lambda_2 + \lambda_3 = 0 \tag{12}
$$

Considering 
$$
\lambda_1 > 0
$$
 and defining  $X = \frac{\lambda_2}{\lambda_1}$  and  $Y = \frac{\lambda_3}{\lambda_1}$ 

$$
XA_3 + XYA_1 + YA_2 = 0 \tag{13}
$$



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$$
\mathcal{L}^{\mathcal{L}}(\mathcal{L}^{\mathcal{L}}(\mathcal{L}^{\mathcal{L}}))
$$

<span id="page-11-1"></span>
$$
1 + X + Y = 0 \tag{14}
$$

## On solving Eqs.[13 a](#page-11-0)nd [14 w](#page-11-1)e have

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$$
X_{\pm} = \frac{(A_3 - A_1 - A_2) \pm \sqrt{(A_3 - A_1 - A_2)^2 - 4A_1A_2}}{2A_1}
$$
\n(15)

$$
Y_{\pm} = \frac{(A_2 - A_1 - A_3) \pm \sqrt{(A_3 - A_1 - A_2)^2 - 4A_1A_2}}{2A_1}e^{-2i\delta}
$$
\n(16)

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For the solution pairs  $(X_+, Y_-)$  and  $(X_-, Y_+)$  we get the Majorana phases as

$$
\alpha = \frac{1}{2} Arg \left[ \frac{(A_3 - A_1 - A_2) \pm \sqrt{(A_3 - A_1 - A_2)^2 - 4A_1 A_2}}{2A_1} \right] \tag{17}
$$

$$
\beta = \frac{1}{2} Arg \left[ \frac{(A_2 - A_1 - A_3) \pm \sqrt{(A_3 - A_2 - A_1)^2 - 4A_1 A_2}}{2A_1} e^{-2i\delta} \right]
$$
 (18)

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**Symmetry [Realization](#page-26-0)** ■ The solution pairs  $(X_+, Y_+)$  and  $(X_-, Y_-)$  satisfy the constraint equations under the condition

$$
(A_3 - A_1 - A_2)^2 - 4A_1A_2 = 0.
$$

$$
\alpha = \frac{1}{2} Arg \left[ \frac{(A_3 - A_1 - A_2)}{2A_1} \right],\tag{19}
$$

$$
\beta = \frac{1}{2} Arg \left[ \frac{(A_2 - A_1 - A_3)}{2A_1} e^{-2i\delta} \right].
$$
 (20)

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## $\blacksquare$  The ratios of the neutrino masses are

### $\rho = \left| \frac{m_2}{\cdots} \right|$  $\frac{m_2}{m_1}e^{2i\alpha}|$  and  $\sigma = |\frac{m_3}{m_1}|$  $\frac{m_3}{m_1}e^{2i\beta}$  (21)

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We have checked the viability of the model by calculating the ratio of the solar and atmospheric mass squared difference 
$$
(R_{\nu})
$$
.

$$
R_{\nu} = \frac{\delta m^2}{\Delta m^2} = \frac{2(\rho^2 - 1)}{|2\sigma^2 - \rho^2 - 1|}
$$
 (22)

where  $\delta m^2 = m_2^2 - m_1^2$  is the solar mass splitting and  $\Delta m^2 = |m_3^2 - \frac{1}{2}(m_2^2 + m_1^2)|$  the atmospheric mass splitting. For NH,  $R_{\nu} = \frac{2\epsilon}{\sigma^2}$ , if we consider  $\rho = 1 + \epsilon$  for  $m_1$ and  $m<sub>2</sub>$  being very close to each other with the values  $0.013 < \epsilon < 0.017$  on  $3\sigma$  [nufit2021]. For IH,  $R_{\nu} = \frac{2(\rho^2 - 1)}{\rho^2 + 1}$ .

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- The range of  $\delta$  is further checked by plotting the atmospheric mixing angle  $\theta_{23}$  against  $\delta$ .
- $\blacksquare$  Finally we have calculated the allowed range of the Majorana Phases for the allowed cases for the viable range of  $\delta$ .
- We have studied the Neutrinoless double beta decay and Jarlskog Invariant for all the viable cases.

$$
|m_{ee}| = |m_1 c_{12}^2 c_{13}^2 + m_2 s_{12}^2 c_{13}^2 e^{2i\alpha} + m_3 s_{13}^2 e^{2i\beta}|
$$
 (23)

$$
J_{cp} = \frac{1}{8}\sin 2\theta_{12}\sin 2\theta_{23}\sin 2\theta_{13}\cos \theta_{13}\sin \delta \qquad (24)
$$

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TABLE I. Neutrino oscillation parameters from global fits [nufit2021].  $\Delta m^2_{3l} = \Delta m^2_{31} > 0$  for normal hierarchy and  $\Delta m^2_{3l} = \Delta m^2_{32} < 0$  for inverted hierarchy.



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$$
A_1 = c_{12}^2 c_{13}^2, \ \ A_2 = s_{12}^2 c_{13}^2, \ \ A_3 = s_{13}^2 e^{2i\delta} \tag{25}
$$

Now we consider the pair  $(X_+, Y_-)$  and plot  $R_\nu$  for both NH and IH.



FIG. 1.  $R_{\nu}$  plots (a) for NH and (b) for IH.

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<span id="page-19-0"></span>Plot for  $\alpha$  and  $\beta$  for this solution pair for the allowed ranges of  $\delta$ .



FIG. 2.  $\alpha$  and  $\beta$  plots for NH for the pair  $(X_+, Y_-)$  with  $\delta = (50^\circ, 150^\circ) \oplus (220^\circ, 320^\circ).$ 

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- **Similar procedure is followed for the solution pairs**  $(X_-, Y_+), (X_+, Y_+)$  and  $(X_-, Y_-)$  of the texture but plots show that  $R_{\nu}$  acquires values far beyond the experimental range. Hence all these solutions of the texture are ruled out.
- Now to explore further phenomenology of the texture,  $|m_{ee}|$  -  $m_{liahtest}$  and  $|m_{ee}|$ - $\beta$  are plotted for neutrinoless double beta decay where the mass of the lightest neutrino is bound within 0.037 eV and 0.042 eV for NH and IH respectively at 95% confidence.

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FIG. 3.  $|m_{ee}|$  plots for  $(X_+, Y_-)$  for NH versus  $m_{lightest}$  and  $\beta$ .

### <span id="page-21-0"></span>From Fig[.3](#page-21-0) we observe that  $|m_{ee}|$  lies within the  $\mathcal{L}_{\mathcal{A}}$ experimental bounds.

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Now we plot  $J_{cp}$ - $\delta$  for CP violation.



FIG. 4.  $J_{cp}$  plot for NH for the solution pair  $(X_+, Y_-)$ .

<span id="page-22-0"></span>In Fig[.4,](#page-22-0) we find  $J_{cp}$  within the range  $(0.018 - 0.04)$ . Thus case  $C_{11} = 0$  is viable under the phenomenological study for normal mass ordering in case of the solution pair  $(X_+, Y_-)$  of the texture.

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All the remaining textures  $C_{12} = 0, C_{13} = 0, C_{22} = 0$ ,  $C_{23} = 0$  and  $C_{33} = 0$  have been examined following our procedure of analysis.

TABLE II. Viable cases under normal hierarchy, inverted hierarchy and neutrinoless double beta decay.



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**Symmetry [Realization](#page-26-0)** TABLE III. Allowed ranges of CP phases  $\delta$ ,  $\alpha$ ,  $\beta$ ,  $|m_{ee}|$  and  $J_{cp}$ for the viable cases.



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 $\blacksquare$  We implement  $Z_5$  Abelian flavor symmetry group to realize all the viable textures. A  $Z_5$  consists of the group elements

$$
(1, \omega, \omega^2, \omega^3, \omega^4)
$$

where  $\omega=e^{\frac{i2\pi}{5}}$  is the generator of the group.

The Lagrangian can be written as **Tall** 

$$
\mathcal{L} = \left(\frac{<\Phi>}{\Lambda}\right)^{Q_{D_{Li}} + Q_{l_{Rj}}} Y_{ij}^{(k)} \overline{D}_{Li} \phi_{k} l_{Rj} + \left(\frac{<\Phi>}{\Lambda}\right)^{Q_{D_{Li}} + Q_{\nu_{Rj}}} Y_{ij}^{(k)} \overline{D}_{Li} \tilde{\phi}_{k} \nu_{Rj}
$$

$$
+ \left(\frac{<\Phi>}{\Lambda}\right)^{Q_{\nu}} Y_{Ri}^{+ Q_{\nu}} Y_{ij}^{(k)} \chi_{k} \overline{\nu}_{Ri} \nu_{Rj} + h.c.
$$
(26)

The  $Q_{\alpha}(\alpha = D_{Li}, l_{Rj}, \nu_{Rj})$  are the FN charges for the SM fermion ingredients.

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## $\blacksquare$  For all the cases, we assign the FN charges for the Lepton sector as

$$
\overline{D}_{1,2,3}:(a+1,a,a),l_{R1,2,3}:(0,1,2),\nu_{R1,2,3}:(d,c,b). \text{ (27)}
$$

Here  $D_{Li}$ ,  $l_{Rj}$ ,  $\nu_{Ri}$ ,  $(i, j = 1, 2, 3)$  represents the  $SU(2)_L$ doublets, the RH  $SU(2)_L$  singlets and the RH neutrino singlets respectively.

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 $M_R$  and  $M_D$  for vanishing minor of the element  $M_{11}$ .

$$
M_R = \begin{pmatrix} 0 & \xi & \zeta \\ \xi & \eta & \upsilon \\ \zeta & \upsilon & \kappa \end{pmatrix}, \ M_D = \begin{pmatrix} x & 0 & 0 \\ 0 & y & 0 \\ 0 & 0 & z \end{pmatrix},
$$

$$
M_{\nu} = -M_D M_R^{-1} M_D^T = \frac{1}{\Gamma} \begin{pmatrix} (-v^2 + \eta \kappa)x^2 & (\zeta - \xi y)xy & (-\zeta \eta + \xi v)xz \\ (\zeta - \xi y)xy & -\zeta^2 y^2 & \xi \zeta yz \\ (-\zeta \eta + \xi v)xz & \xi \zeta yz & -\xi^2 z^2 \end{pmatrix},
$$
\n(28)

where 
$$
\Gamma = -\zeta \eta^2 + 2\xi \zeta \upsilon - \xi^2 \kappa
$$
.

- $\blacksquare$  On implementing  $Z_5$  symmetry, the fields of the relevant particles transform as:
	- $\nu_{R1} \rightarrow \omega^3$  $\nu_{R1}$ ,  $\overline{D}_{L1} \rightarrow \omega^2 \overline{D}_{L1}$ ,  $l_{R1} \rightarrow \omega^3$  $(29)$

$$
\nu_{R2} \to \omega^2 \nu_{R2}, \qquad \qquad \overline{D}_{L2} \to \omega^3 \overline{D}_{L2}, \qquad l_{R2} \to \omega^2 l_{R2} \tag{30}
$$

 $\nu_{R3} \rightarrow \nu_{R3}$ ,  $\overline{D}_{L3} \rightarrow \overline{D}_{L3}$ ,  $l_{R3} \rightarrow l_{R3}$  (31)

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## Forming the required bilinears dictated by  $Z_5$  symmetry we obtain

$$
\nu_{Ri}^T \nu_{Rj} = \begin{pmatrix} \omega & 1 & \omega^3 \\ 1 & \omega^4 & \omega^2 \\ \omega^3 & \omega^2 & 1 \end{pmatrix}, \quad \overline{D}_{Li} \nu_{Rj} = \begin{pmatrix} 1 & \omega^4 & \omega^2 \\ \omega & 1 & \omega^3 \\ \omega^3 & \omega^2 & 1 \end{pmatrix}, \quad (32)
$$

$$
\overline{D}_{Li} l_{Rj} = \begin{pmatrix} 1 & \omega^4 & \omega^2 \\ \omega & 1 & \omega^3 \\ \omega^3 & \omega^2 & 1 \end{pmatrix}.
$$

■ We consider the transformation of the singlet scalars  $\chi_k(k=1,2,3)$  which is responsible for the Majorana neutrino mass matrix  $M_R$  and SM-like doublet scalar  $\phi$ which is responsible for the Dirac neutrino mass matrix  $M_D$  and the lepton mass matrix  $M_l$  under  $Z_5$ transformation as

$$
\chi_1 \to \omega^2 \chi_1, \ \chi_2 \to \omega^3 \chi_2, \chi_3 \to \omega \chi_3 \tag{34}
$$

 $\phi \rightarrow \phi$ <br>Sangeeta Dey [Study of neutrino mass matrices with vanishing trace and one vanishing minor](#page-0-0) 30/38

# Symmetry Realization

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 $\mathcal{L}$ 

## Now the lagrangian dictated by  $Z_5$  is

$$
Z_{M}^{5} = \epsilon^{d+c} m_{12} \nu_{R1}^{T} c^{-1} \nu_{R2} + \epsilon^{d+b} Y_{\chi_{13}}^{(1)} \chi_{1} \nu_{R1}^{T} c^{-1} \nu_{R3} + \epsilon^{2c} Y_{\chi_{22}}^{(3)} \chi_{3} \nu_{R2}^{T} c^{-1} \nu_{R2} + \epsilon^{c+b} Y_{\chi_{23}}^{(2)} \chi_{2} \nu_{R2}^{T} c^{-1} \nu_{R3} + \epsilon^{2b} m_{33} \nu_{R3}^{T} c^{-1} \nu_{R3} + \epsilon^{a+d+1} Y_{D_{11}} \overline{D}_{L1} \tilde{\phi} \nu_{R1} + \epsilon^{a+c} Y_{D_{22}} \overline{D}_{L2} \tilde{\phi} \nu_{R2} + \epsilon^{a+b} Y_{D_{33}} \overline{D}_{L3} \tilde{\phi} \nu_{R3} + \epsilon^{a+1} Y_{l_{11}} \overline{D}_{L1} \phi l_{R1} + \epsilon^{a+1} Y_{l_{22}} \overline{D}_{L2} \phi l_{R2} + \epsilon^{a+2} Y_{l_{33}} \overline{D}_{L3} \phi l_{R3}.
$$
\n(36)

## Now we construct the mass matrix  $M_R$ ,  $M_D$  and  $M_l$  as

$$
M_R = \begin{pmatrix} 0 & e^{d+c}m_{12} & y_{\lambda13}^{(1)} \chi_1 e^{d+b} \\ e^{d+c}m_{12} & y_{\lambda23}^{(2)} \chi_3 e^{2c} & y_{\lambda23}^{(2)} \chi_2 e^{c+b} \\ y_{\lambda13}^{(1)} \chi_1 e^{d+b} & y_{\lambda23}^{(2)} \chi_2 e^{c+b} & e^{2b}m_{33} \end{pmatrix}, \qquad (37)
$$

$$
M_D = \begin{pmatrix} y_{D_{11}} \tilde{\phi}e^{a+d+1} & 0 & 0 \\ 0 & y_{D_{22}} \tilde{\phi}e^{a+c} & 0 \\ 0 & 0 & y_{D_{33}} \tilde{\phi}e^{a+b} \end{pmatrix}, M_I = \begin{pmatrix} y_{l_{11}} \phi e^{a+1} & 0 & 0 \\ 0 & y_{l_{22}} \phi e^{a+1} & 0 \\ 0 & 0 & y_{l_{33}} \phi e^{a+2} \end{pmatrix}.
$$

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 $(1)$ 

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$$
w_{\nu} = \Omega \begin{pmatrix} e^{2\beta y} y_{D_{11}}^{0} (y_{\nu 1}^{0.2} x_2^2 - m_{33} y_{\nu 2}^{0.1} x_3) & e^{2y} y_{D_{11} y_{D_{22}} (m_{11} m_{33} - y_{\nu 13}^{0.1} y_{\nu 2}^{0.1} x_1 x_2) & -e^{2y} y_{D_{11} y_{D_{22}} (m_{11} m_{\nu 1}^{0.2} x_2 - y_{\nu 13}^{0.1} y_{\nu 2}^{0.2} x_1 x_3) \\ e^{2y} y_{D_{11} y_{D_{22}} (m_{11} m_{33} - y_{\nu 13}^{0.1} y_{\nu 2}^{0.2} x_1 x_2) & -\beta y_{D_{21} y_{\nu 1}}^{0.1} y_{\nu 13}^{0.1} x_1^2 & -m_{11} \beta^2 y_{D_{22} y_{D_{23}} y_{\nu 13}^{0.1} x_1 x_3} \\ -e^{2y} y_{D_{11} y_{D_{3}} (m_{11} y_{\nu 2}^{0.2} x_2 - y_{\nu 13}^{0.1} y_{\nu 2}^{0.2} x_1 x_3) & -\beta^2 y_{D_{22} y_{D_{23}} y_{\nu 13}^{0.1} x_1 m_{11}} & \beta^2 y_{D_{23} m_{11}}^{0.1} x_1^2 & -m_{11} \beta^2 y_{D_{23} m_{11}}^{0.
$$



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**The textures**  $C_{11} = 0$ ,  $C_{13} = 0$ ,  $C_{22} = 0$  and  $C_{33} = 0$  have been found viable for normal hierarchy only and the case  $C_{12} = 0$  has been found viable for both normal and inverted hierarchies.

Again the case  $C_{23} = 0$  is completely ruled out.

- Interestingly the solution pair  $(X_+, Y_-)$  supports all the cases except the case  $C_{13} = 0$ .
- The solution pairs  $(X_+, Y_+)$  and  $(X_-, Y_-)$  support the cases  $C_{12} = 0$  and  $C_{13} = 0$  for normal hierarchy.
- **Further the solution pair**  $(X_-, Y_+)$  supports  $C_{12} = 0$  for inverted hierarchy and  $C_{22} = 0$  for normal hierarchy.

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**The Majorana phase**  $\alpha$  for the textures  $C_{12} = 0$  and  $C_{13} = 0$  is vanishingly small and the range is highly constrained.

- $\blacksquare$  For all the viable textures, the atmospheric mixing angle  $\theta_{23}$  lies in the range (40 $^{\circ}$ , 45 $^{\circ}$ ). Thus the phenomenology of these textures favors the first quadrant for atmospheric mixing.
- $\blacksquare$  For all the cases both the neutrinoless double beta decay rate,  $|m_{ee}|$  and the strength of the Dirac CP violation,  $J_{CP}$  remain within the experimental bounds.
- We have found that all the cases favour normal hierarchy of neutrino mass pattern using FN mechanism.

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