

# Study of the criticality for QGP formation in AA and $pp$ collisions

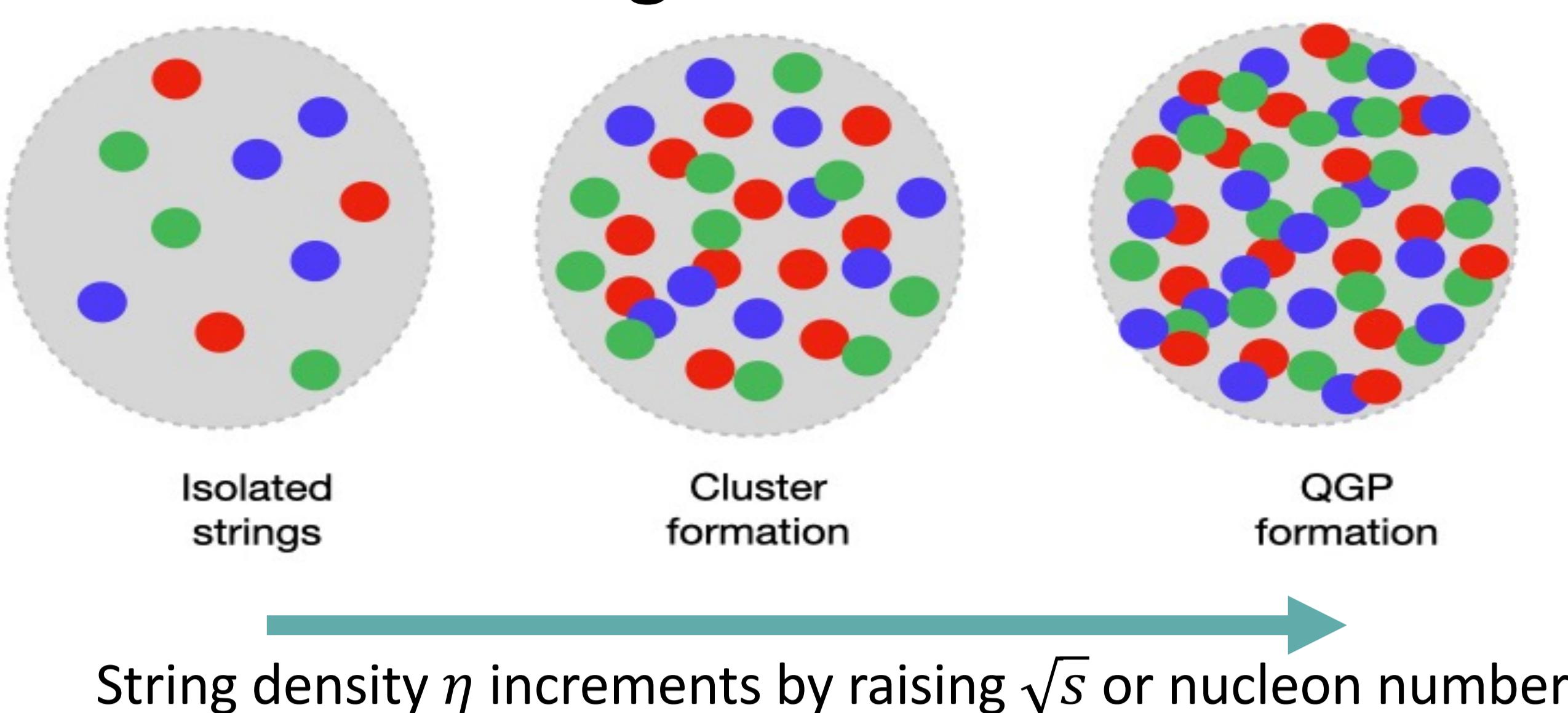
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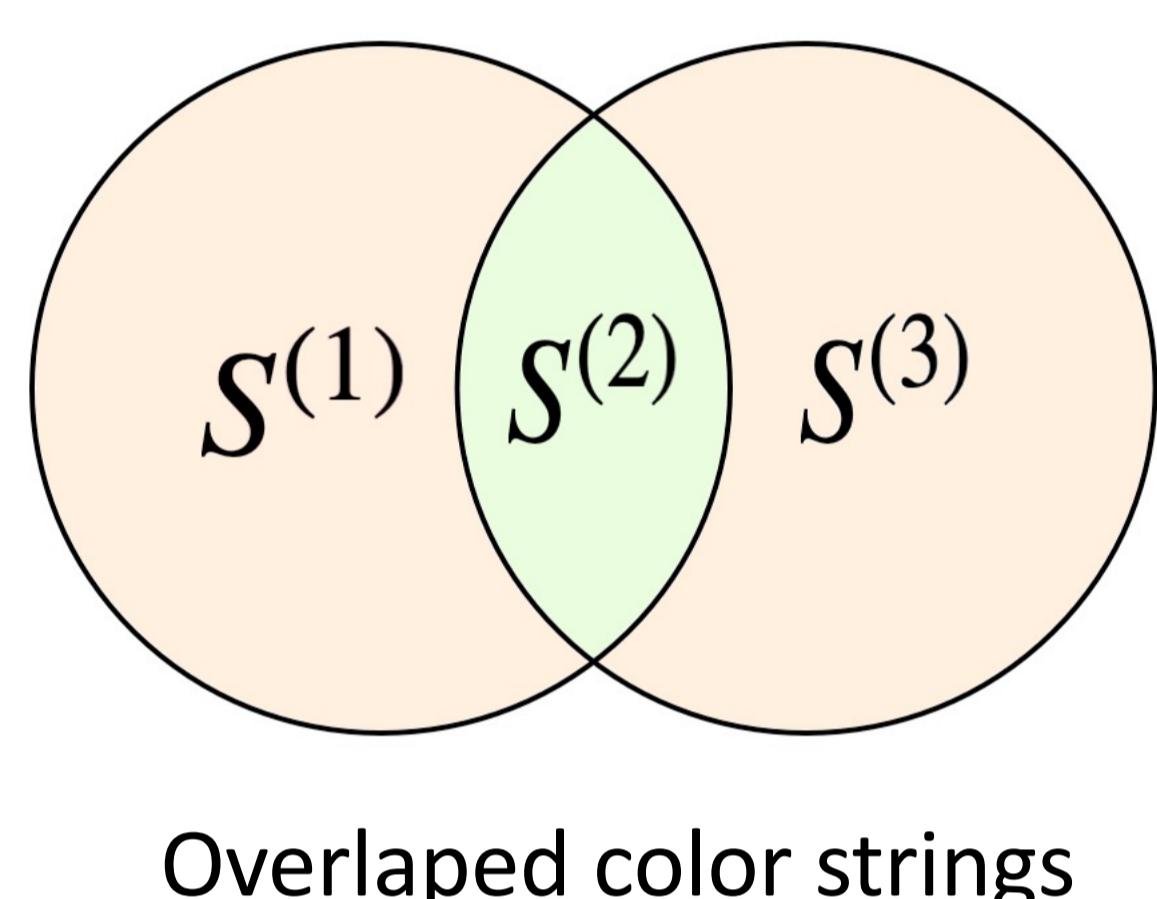
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## Color String Percolation Model

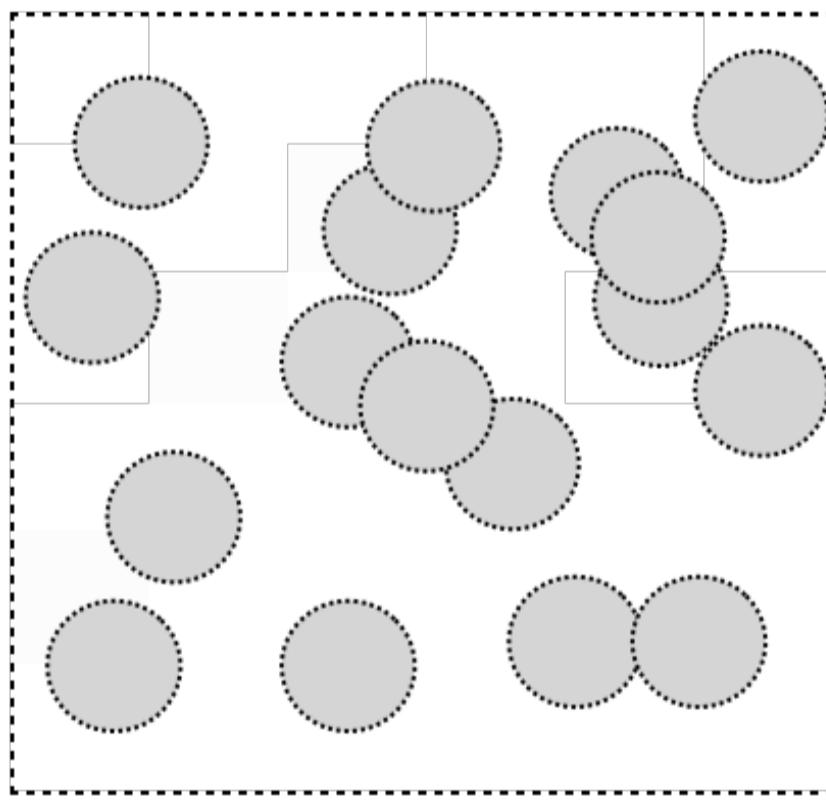


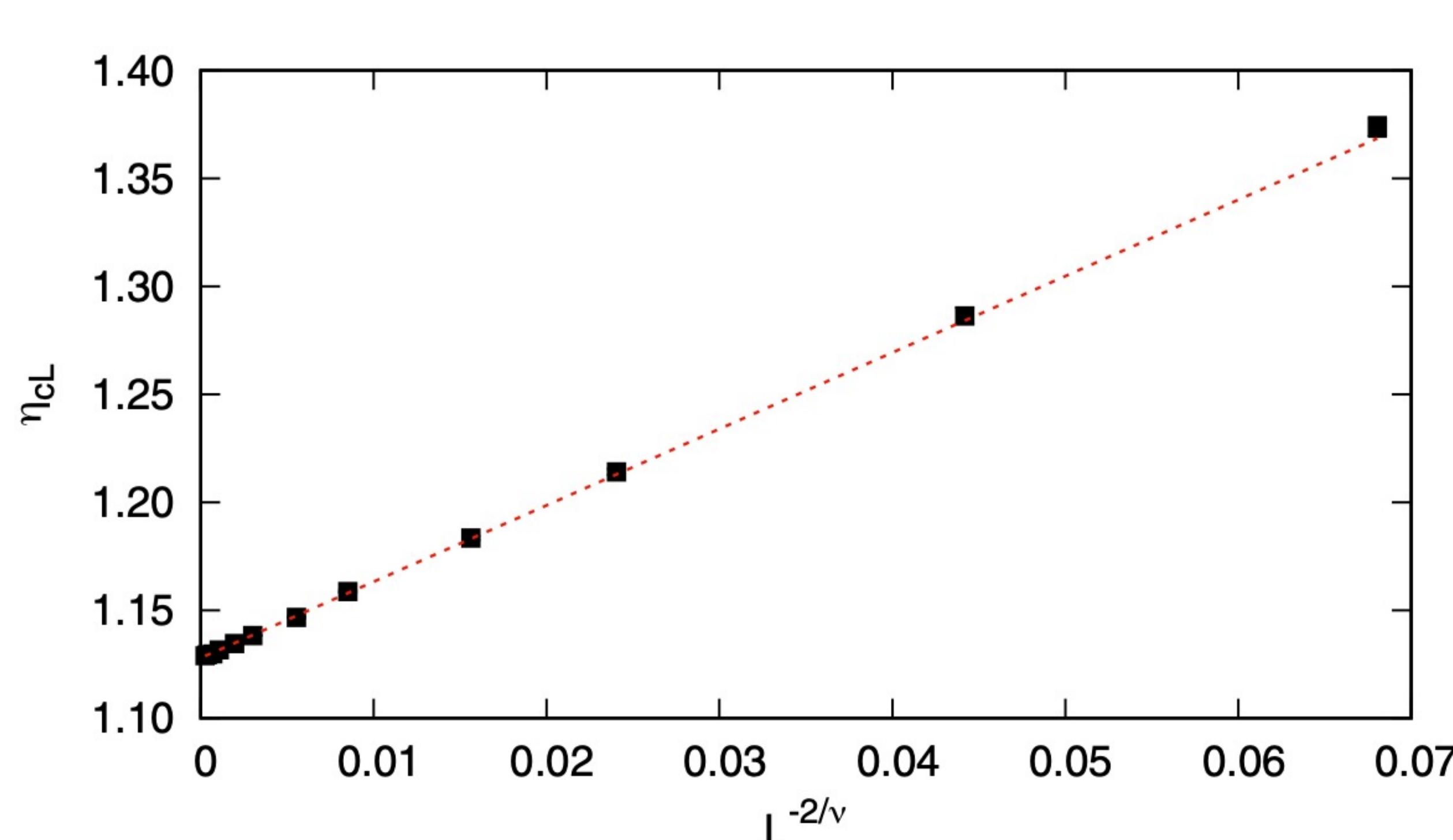
## Color suppression factor $F(\eta)$ and CSPM observables



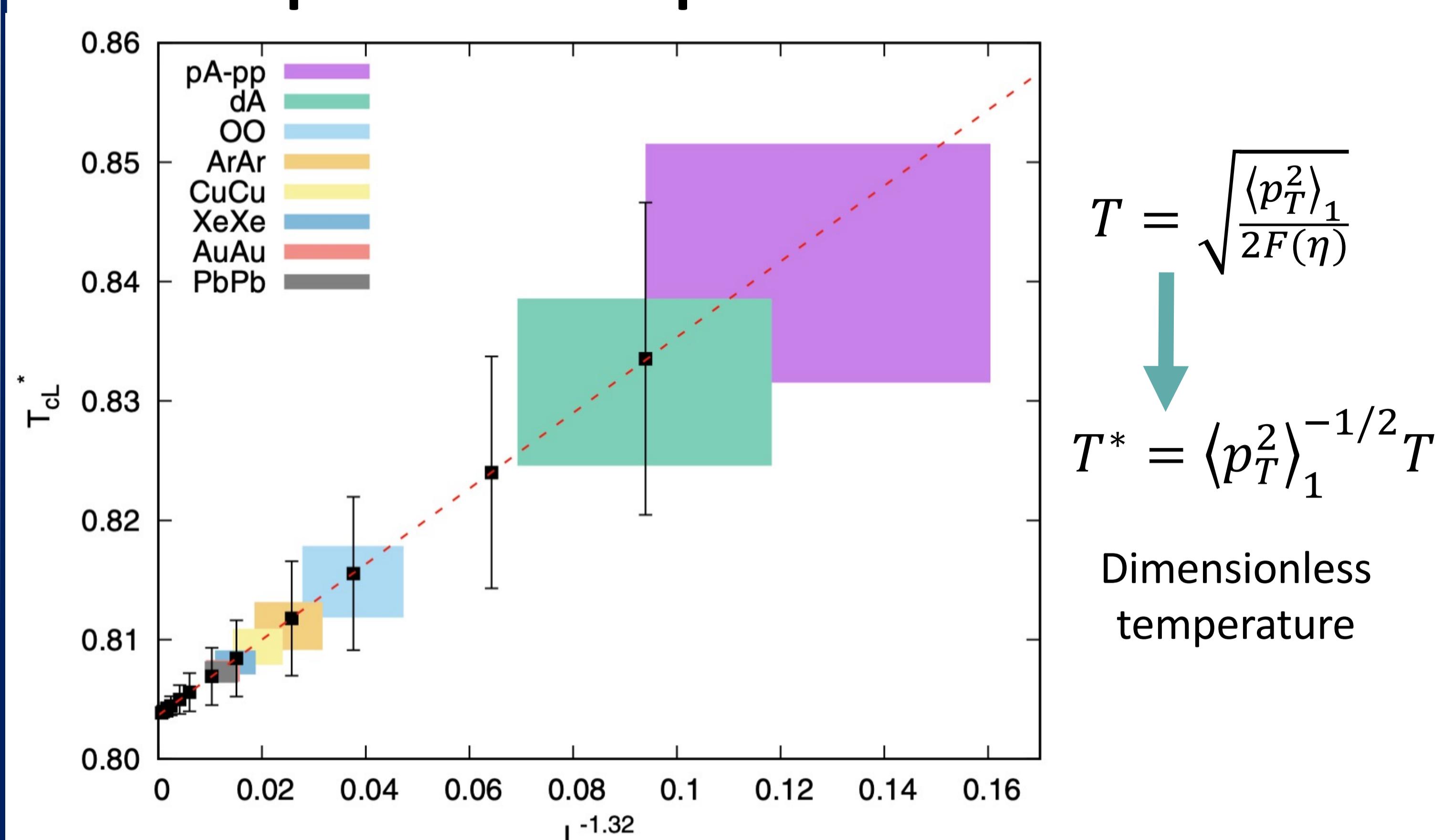
- $F(\eta)$  reflects the effects on the interaction between strings
- Since  $F(\eta)$  is a decreasing function  $\mu = NF(\eta)\mu_1$  decreases, and  $\langle p_T^2 \rangle = \langle p_T^2 \rangle_1 / F(\eta)$  is enhanced.

## Finite Size Effects

Percolation theory	Ultrarelativistic collisions
Thermodynamic limit $L \rightarrow \infty$	PbPb, AuAu collision
How can we include pp collisions?	Density is an intensive quantity
Mertens-Moore algorithm	 $L = \frac{R_A}{r_0} = \frac{r_o^* A_M^{1/3}}{r_0} \sim 5 A_M^{1/3}$
Scaling laws	$X_c - X_{cL} \propto L^z$ $X_c - X_{cL} \propto (5A_M)^{\frac{z}{3}}$ <p>The observable can be scaled by the nucleon number !!</p>



## Temperature-like parameter in the CSPM



Dimensionless transition temperature associated with the QGP formation is defined by  $T^*(\eta_c)$ .

For CSPM we estimate  $T_c^* = 0.80365$  in the thermodynamic limit.

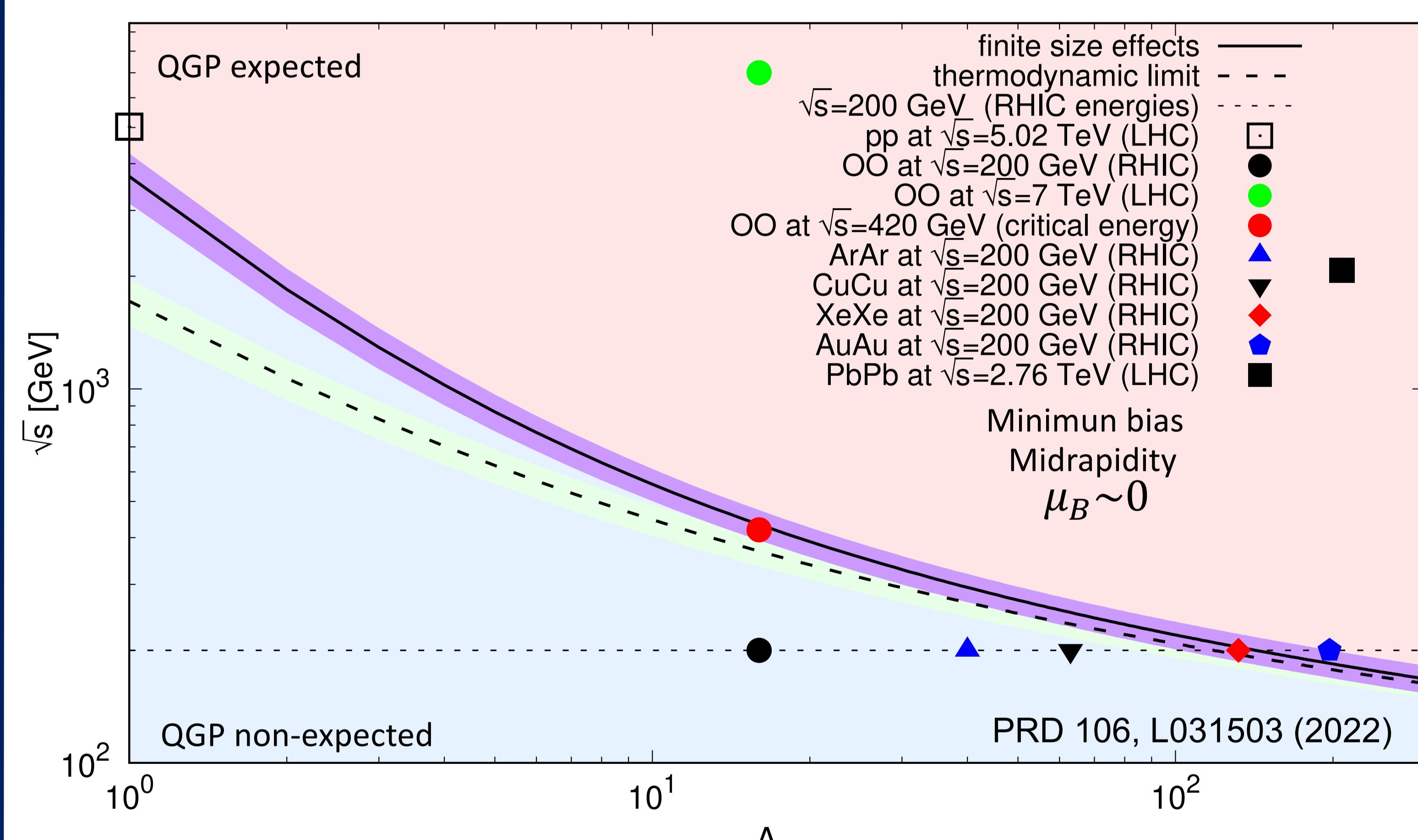
## Minimal center of mass energy needed for QGP formation

We express string density for pp and AA collisions as function of  $\sqrt{s}$  through

$$\eta^{pp}(\sqrt{s}) = \frac{\pi}{25} \left[ 2 + 4 \left( \frac{r_0}{R_p} \right)^2 \left( \frac{\sqrt{s}}{m_p} \right)^{2\lambda} \right] \quad \text{and} \quad \eta^{AA}(\sqrt{s}) = \eta^{pp}(\sqrt{s}) A_M^{\alpha(\sqrt{s})}.$$

We estimate the critical  $\sqrt{s}$  by solving  $\eta^{AA}(\sqrt{s}) = \eta_{cL}^{AA}$

Collision	$A_M$	$\sqrt{s}_{cL}$ GeV	$\eta_{cL}^{AA}$
pp	1	3700	1.444
OO	16	420	1.2057
ArAr	40	290	1.1768
CuCu	63	246	1.1668
XeXe	132	201	1.1547
AuAu	197	184	1.1497
PbPb	208	182	1.1491



## Conclusions

- CSPM presents finite size effects that can be expressed in terms of  $A_M$
- Transition temperature is higher for small systems than for large ones.
- Small systems require higher energy or higher multiplicity.
- Our estimation of  $\sqrt{s}$  for QGP formation is consistent with those reported in literature.

## References

- J.C. Texa et al, PRD 106, L031503 (2022)
- Braun and Pajares, EPJC 16, 349 (2000)
- J. E. Ramírez et al, PRC 103, 094029 (2021)
- J. Dias de Deus and C. Pajares, PLB 642, 455 (2006)