



Searches for dark matter with CMS

International Conference on High Energy Physics 2024

M. Waßmer on behalf of the CMS collaboration | July, 19th



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Motivation

measurements

gas

5

best-fit model

150

50

0

s-1) 100

V (km

Introduction

Astrophysical and cosmological observations motivate the existence of dark matter.

Spin-0 portal



10

stellar disk

Dark Matter halo

Largest part of matter in the universe is cold dark matter (DM). No suitable DM candidate available in the standard model (SM) of particle physics.



Extended dark sectors



References



Summarv



3/22 July, 19th M. Waßmer: Searches for dark matter with CMS



Search for new physics with a mono-top signature



CMS-PAS-SUS-23-004

5/22

- spin-1 portal realized in a simplified model [1, 2]
- vector or axial-vector coupling of the mediator V
- Dirac-type DM candidate χ (WIMP)
- E_{T} + top guark (mono-top) at tree level via flavor-changing neutral current (FCNC)
- up-top production mode favored by parton density functions at LHC



Search for new physics with a mono-top signature



- only consider hadronic top guark decay
- cluster top decay products into a fat-jet
- ML-based discriminator (particleNet [3]): top-jets vs. QCD-jets
- hadronic recoil as observable
- control regions (CRs) based on number of leptons and b-tags
- robust statistical model using transfer factors
- \Rightarrow estimate main backgrounds in signal region (SR) from data in CRs



CMS-PAS-SUS-23-004

Spin-0 portal

Extended dark sectors

Search for new physics with a mono-top signature





CMS-PAS-SUS-23-004

Most stringent exclusion limits for (axial) vector coupled dark matter production via an up-top FCNC to date.

Introduction	Spin-1 portal	Spin-0 portal	Extended dark sectors	Summary	References
7/22 July, 19th	M. Waßmer: Searches for dark matter with CMS				

Summary of spin-1 simplified models (vector)



References

CMS 2000 Dark matter mass m_{DM} [GeV] 95% CL exclusions 1800 Observed $m_{med} = 2 m_{DM}$ 1600 \swarrow Ω_c h² \ge 0.12 Expected 1400 Boosted dijet (77 fb⁻¹) Phys. Rev. D 100 (2019) 112007 1200 Dijet w/ btag (19.7 fb-1) Phys. Rev. Lett. 120 (2018) 201801 1000 Dijet w/ ISR i (18.3 fb-1) 800 Phys. Lett. B 805 (2020) 135448 Dijet (35.9-137 fb⁻¹) JHEP 08 (2018) 130 600 JHEP 05 (2020) 033 DM + i/V(aa) (137 fb⁻¹) Vector mediator 400 JHEP 11 (2021) 153 Dirac DM DM + γ (35.9 fb⁻¹) 200 g_{DM} = 1.0 JHEP 02 (2019) 074 g_a = 0.25 DM + Z(II) (137 fb⁻¹) 0 g, = 0 Eur. Phys. J. C 81 (2021) 13 10² 10^{3} Mediator mass m_{med} [GeV] March 2024, CMS-EXO Public Results Introduction Spin-1 portal Spin-0 portal Extended dark sectors Summarv

Summary of spin-1 simplified models (axial-vector)





9/22 July, 19th M. Waßmer: Searches for dark matter with CMS



DM produced with a single top quark or a top quark pair

References

CMS-PAS-EXO-22-014

- spin-0 portal realized in a simplified model [2, 4, 5, 6]
- scalar (ϕ) or pseudo scalar (*a*) mediator, Dirac-type DM candidate (χ)
- fermion (f) coupling of ϕ/a of Yukawa-type, i.e., Higgs-boson-like couplings $\propto m_f$
- SM couplings preferrably to heavy third generation quarks (top and bottom quark)
- no mixing with SM Higgs boson assumed (while mixing exists in dark Higgs models)
- sensitive final states $t\overline{t} + \chi\overline{\chi}$ and $t/\overline{t} + \chi\overline{\chi}$
- $t\bar{t} + \chi\bar{\chi}$ also explored in top squark search CMS-SUS-20-002



search in all lepton multiplicity channels (0l, 1l, 2l) with full run 2 data

- $t\bar{t} + \chi\bar{\chi}$ and $t/\bar{t} + \chi\bar{\chi}$ categories according to b-tag multiplicity
- further categorization using forward jet-multiplicity for t- and tW-channel modes of $t/\bar{t} + \chi \overline{\chi}$
- hadronic recoil as observable in 01 & 11 channels
- neutral network discriminant in 2l channel
- control regions to improve background process estimations in signal regions
- signal-like excess with $< 2\sigma$ significance

CMS-PAS-EXO-22-014

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Spin-1 portal

Introduction 12/22 J Extended dark sectors



Spin-0 portal





Summary of spin-0 simplified models (scalar)





March 2024, CMS-EXO Public Results

 Introduction
 Spin-1 portal
 Spin-0 portal

 13/22
 July, 19th
 M. Waßmer: Searches for dark matter with CMS

Extended dark sectors

References

Summary of spin-0 simplified models (pseudoscalar)





March 2024, CMS-EXO Public Results

 Introduction
 Spin-1 portal
 Spin-0 portal

 14/22
 July, 19th
 M. Waßmer: Searches for dark matter with CMS

Extended dark sectors

References



Extended dark sectors

Search for DM with b-quark and lepton pairs

CMS-PAS-SUS-23-018

Introduction

- pseudoscalar a links SM and DM \rightarrow avoids DM-nucleon constraints from direct detection experiments
- Fermi-LAT gamma-ray excess [7] hypothesized to originate from $\chi \overline{\chi} \rightarrow b \overline{b} [8, 9, 10]$
- 2HDM model + additional portal pseudoscalar (2HDM+a) naturally allows pseudoscalar link between SM and DM and can produce couplings favoring b quarks [11]
- possible production of bb, DM candidates in association with a Z boson

Spin-0 portal

search for significant E_{T} , a dilepton pair, and a pair of b-jets





Search for DM with b-quark and lepton pairs



- baseline selection using leptons, b-tag multiplicity, *E*_T and analytical tt
 (DL) reconstruction
- neural network (MLP) to further discriminate signal from remaining backgrounds
- neural network output used to create bins of low and high sensitivity
- overall normalization of major background processes freely-floating and constrained from dedicated CRs



CMS-PAS-SUS-23-018



Search for DM with b-quark and lepton pairs

CIVIS-PAS-505-23-01

M. Waßmer: Searches for dark matter with CMS



Introduction

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Search for DM in association with a Higgs boson ($\rightarrow \tau \tau$)

Spin-0 portal

CMS-PAS-SUS-23-012

- production of a SM Higgs boson decaying to two *τ* leptons in association with a DM pair
- another signature possible in 2HDM+a
- selection: significant *E*_T, a di-τ pair compatible with SM Higgs boson mass, no additonal leptons or b-tagged jets, and large total transverse mass



Spin-1 portal







- wide variety of dark matter searches at CMS ranging from
 - simplified models and simple dark sectors to extended dark sectors
 - simple final states to more complex ones and new signatures
- new data collection and reconstruction techniques e.g. scouting and parking (not shown in this talk)
- detailed/extensive review of CMS dark sector searches: CMS-EXO-23-005
- so far no significant dark matter signal found
- more data, higher center-of-mass energy, refined analysis techniques, and detector upgrades will push sensitivity further in the future

Thank you for your attention. Keep your eyes open for new results!

References I



References

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