



# Status on COSINE-100 experiment

In-Soo Lee

Center for Underground Physics (IBS)

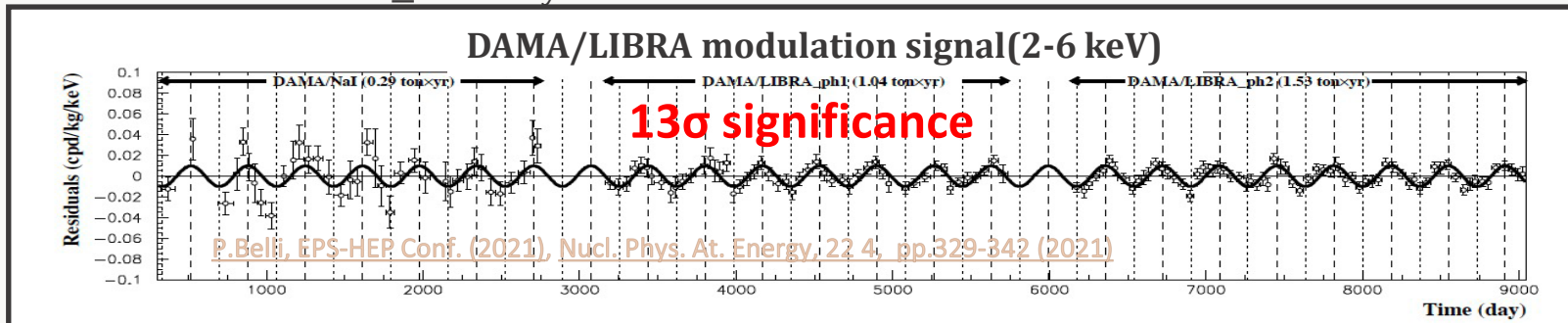
On behalf of the COSINE-100 collaboration

July 18th, 2024

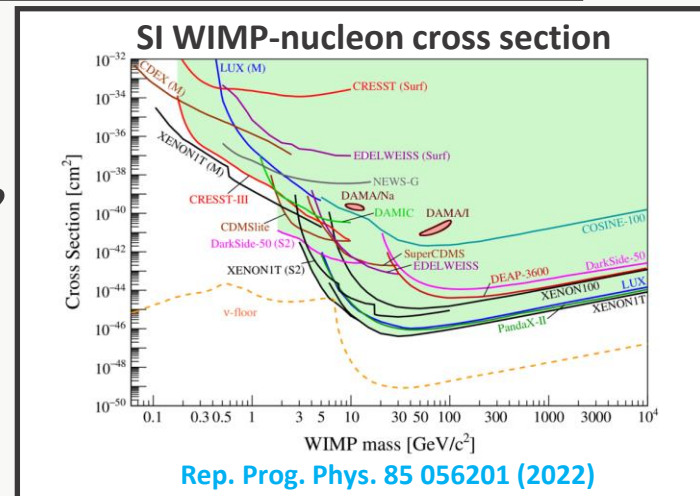


# Motivation-DAMA/LIBRA

- DAMA/LIBRA experiment
  - Search for **Dark matter (DM)** annual modulation signature
  - Using  $25 \times 9.7$  kg NaI(Tl) detectors
  - **Claim an observation** of the DM (WIMP modulation signal) at  **$13\sigma$  C.L.** (2-6 keV)
  - **Amplitude** :  $0.00996 \pm 0.0007$  counts/day/kg/keV
    - Phase =  $145 \pm 5$  days
    - Period =  $0.997 \pm 0.0008$  year



- **No other experiment** has succeeded direct detection of DM except for **DAMA/LIBRA**.
  - **Is NaI(Tl) special** for DM interaction?



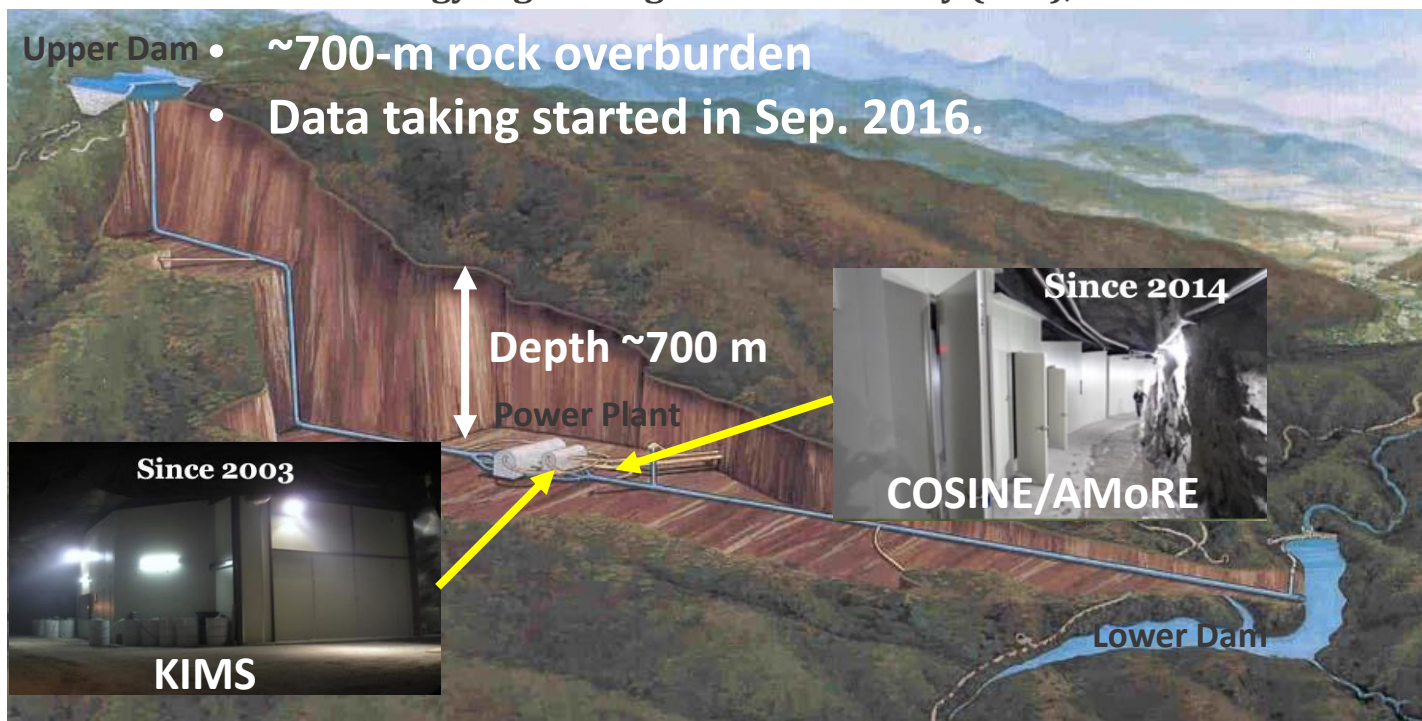
# COSINE-100 experiment



- Joint collaboration of DM-ICE & KIMS
- ~50 collaborators in 18 institutes
- To confirm DAMA/LIBRA's claim using same target material, NaI(Tl)

## Yangyang underground laboratory (Y2L), Korea

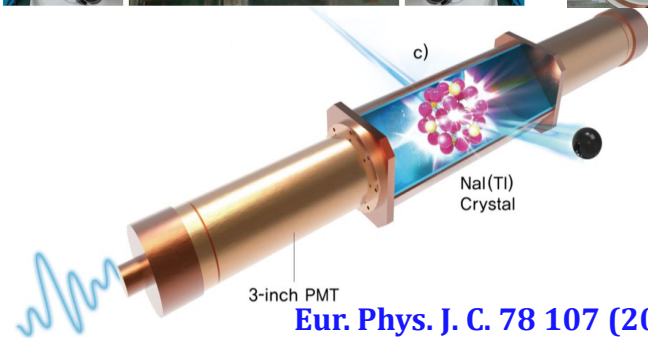
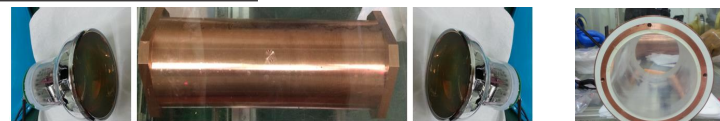
- Upper Dam • ~700-m rock overburden
- Data taking started in Sep. 2016.



# COSINE-100 experiment

- 8 NaI(Tl) Crystals (106 kg)
  - Higher light yield (15 PE/keV) than DAMA/LIBRA (5-10 PE/keV)
- Shielding structure
  - 2200-L LAB-based Liquid scintillator(LS)
  - $4\pi$  plastic scintillator
  - 3-cm thick copper box
  - 20-cm thick lead castle

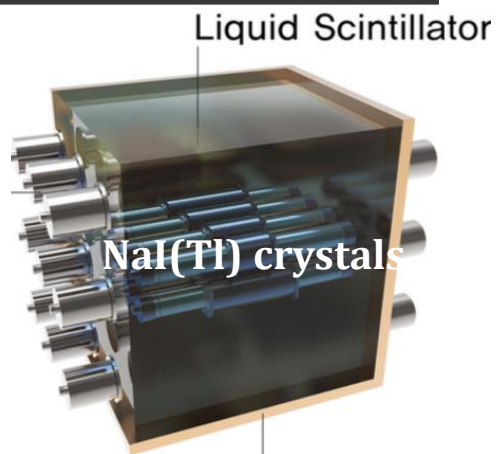
## NaI(Tl) crystal



[Eur. Phys. J. C. 78 107 \(2018\)](#)

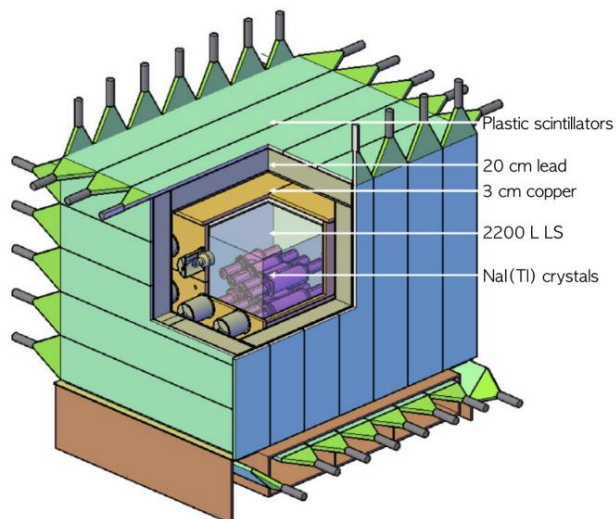
[Eur. Phys. J. C. 78 490 \(2018\)](#)

## Shielding structure



[Nucl. Instrum. Meth. A 106, 165431 \(2021\)](#)

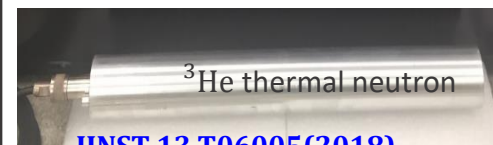
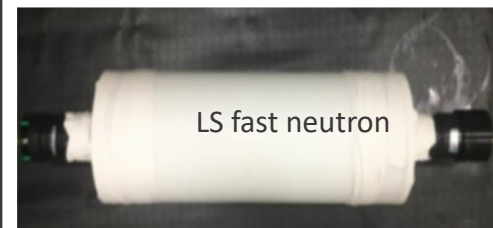
[Nucl. Instrum. Meth. A 851 103 \(2017\)](#)



[2018 JINST 13 T02007](#)

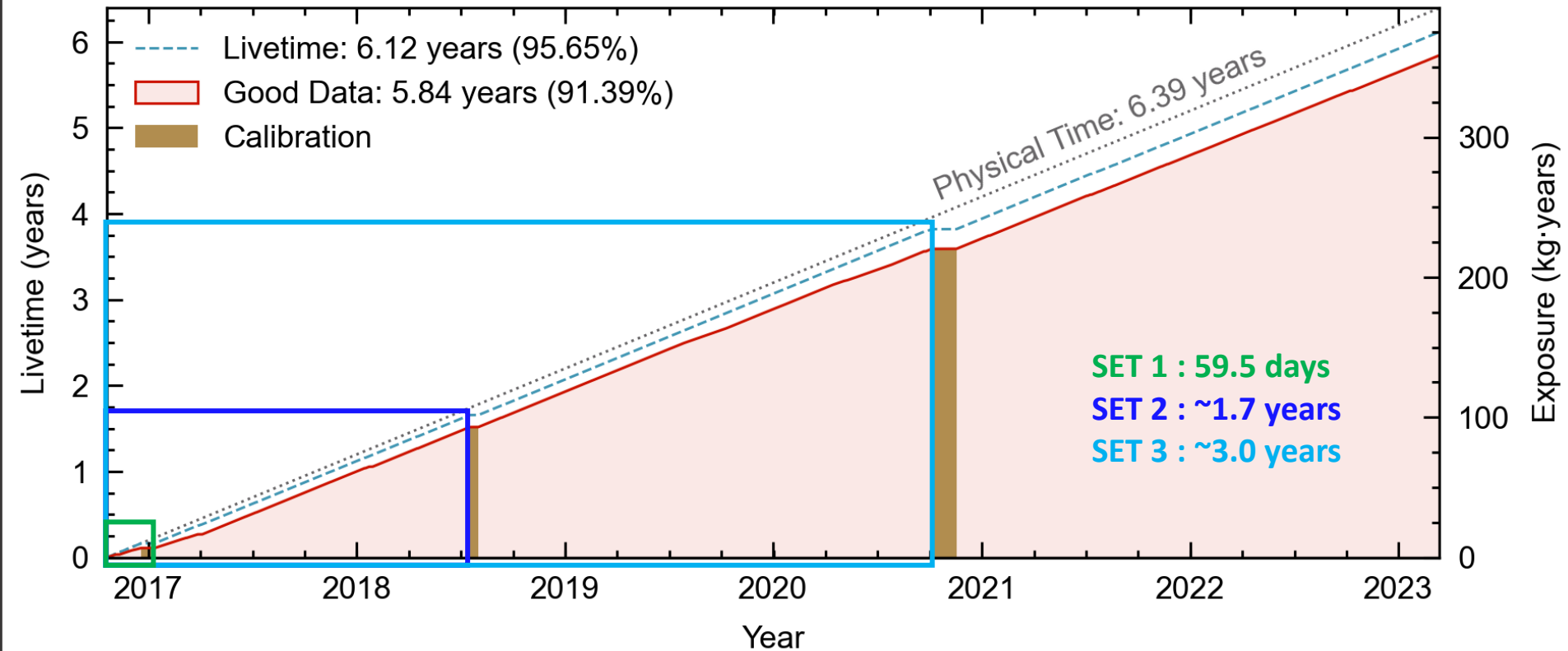
[Eur. Phys. J. C 78, 107 \(2018\)](#)

## Neutron monitoring



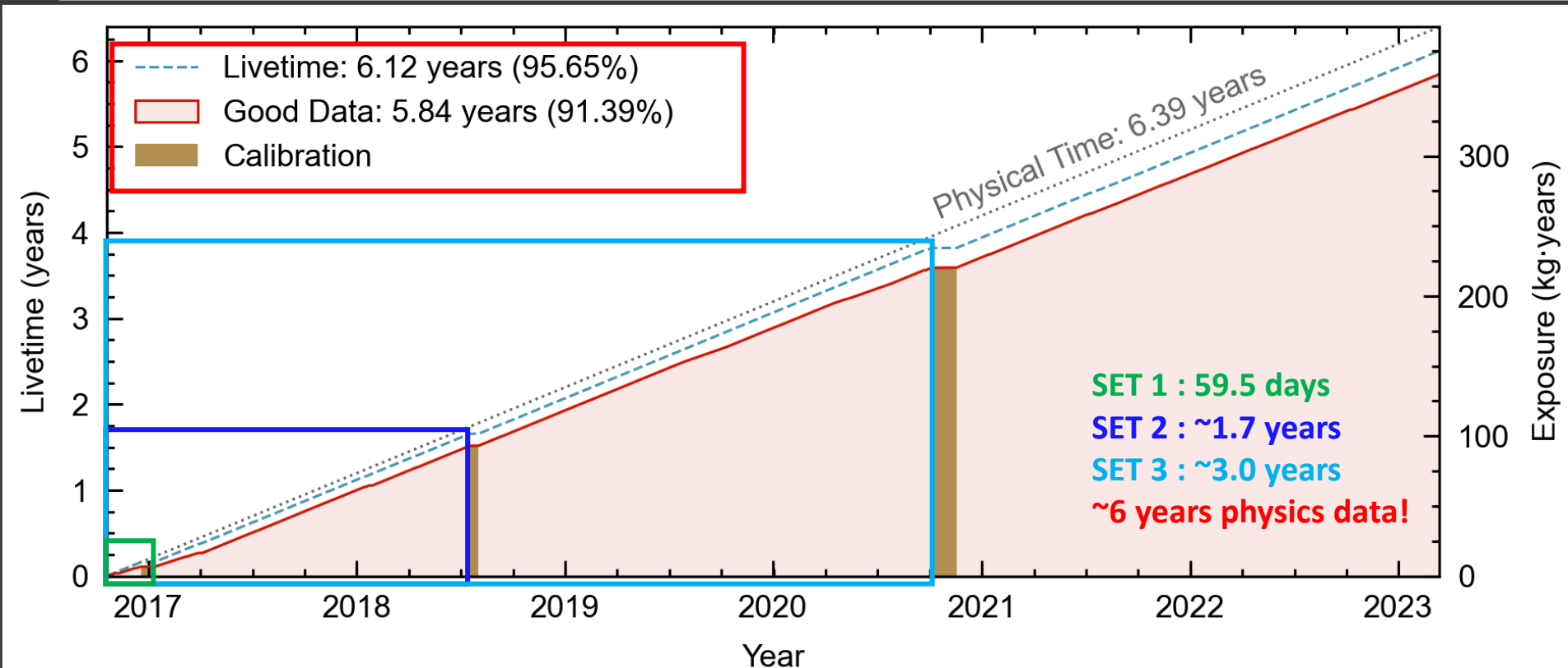
[JINST 13 T06005\(2018\)](#)

# Data exposure



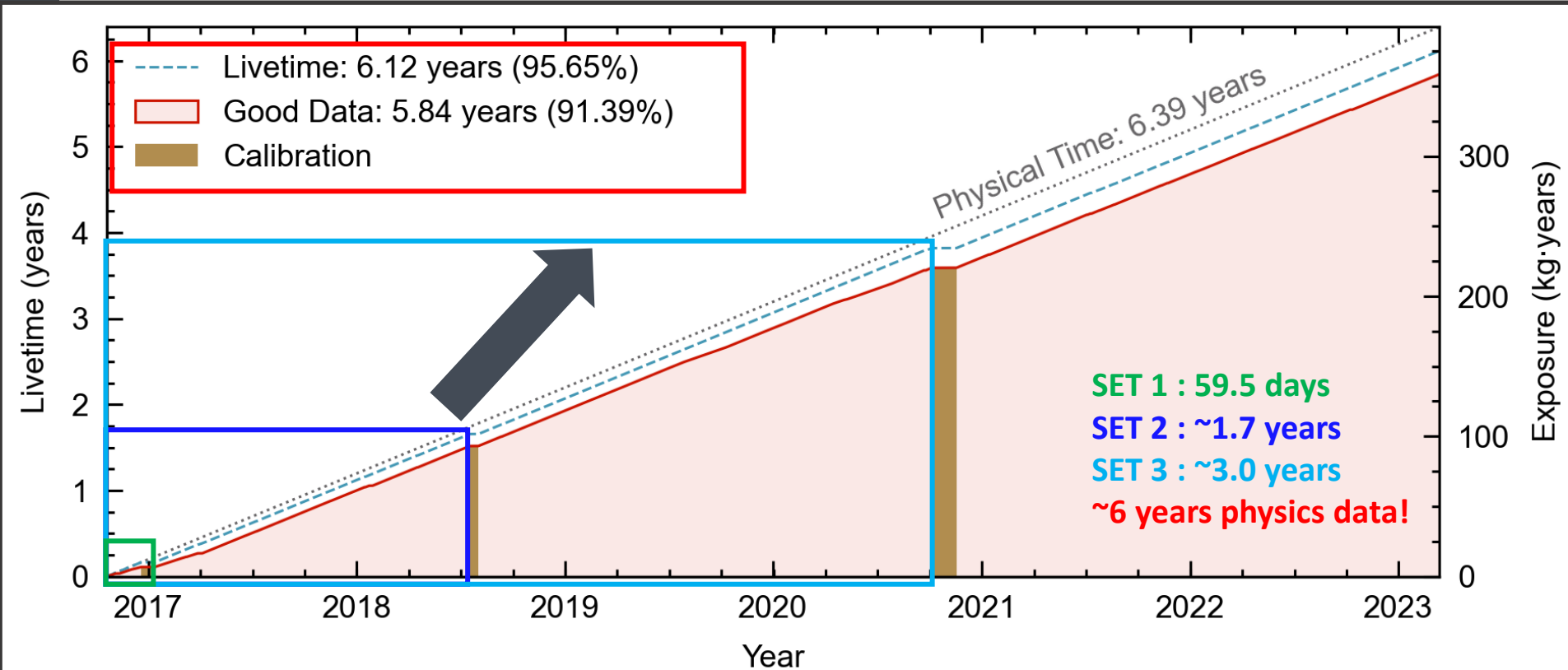
- **Operation from Sep. 2016 to Mar. 2023**
  - Physics run : 6.12 years
  - Good data : 5.84 years

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  - Good data : 5.84 years

# Data exposure



- Main analysis is being performed with **SET3**

# Detector understanding update with SET3 data

- **Event selection**

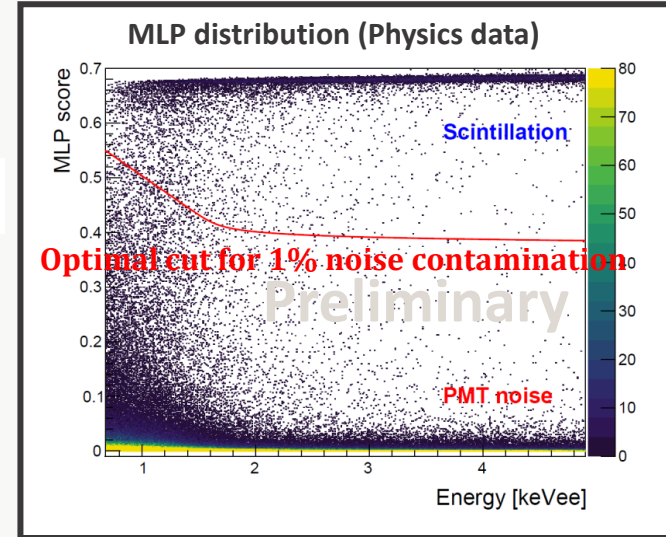
- **Multivariable analysis with pulse-shape parameters**

- Meantime, likelihood parameters (SET2)

$$F_{\mathcal{L}}^{N1} = \frac{\ln \mathcal{L}_{\text{FFT}}^{N1} - \ln \mathcal{L}_{\text{FFT}}^S}{\ln \mathcal{L}_{\text{FFT}}^{N1} + \ln \mathcal{L}_{\text{FFT}}^S}$$

- + **Fast Fourier Transformation** parameter(for SET3)

- 1 keV  $\rightarrow$  **0.7 keV**





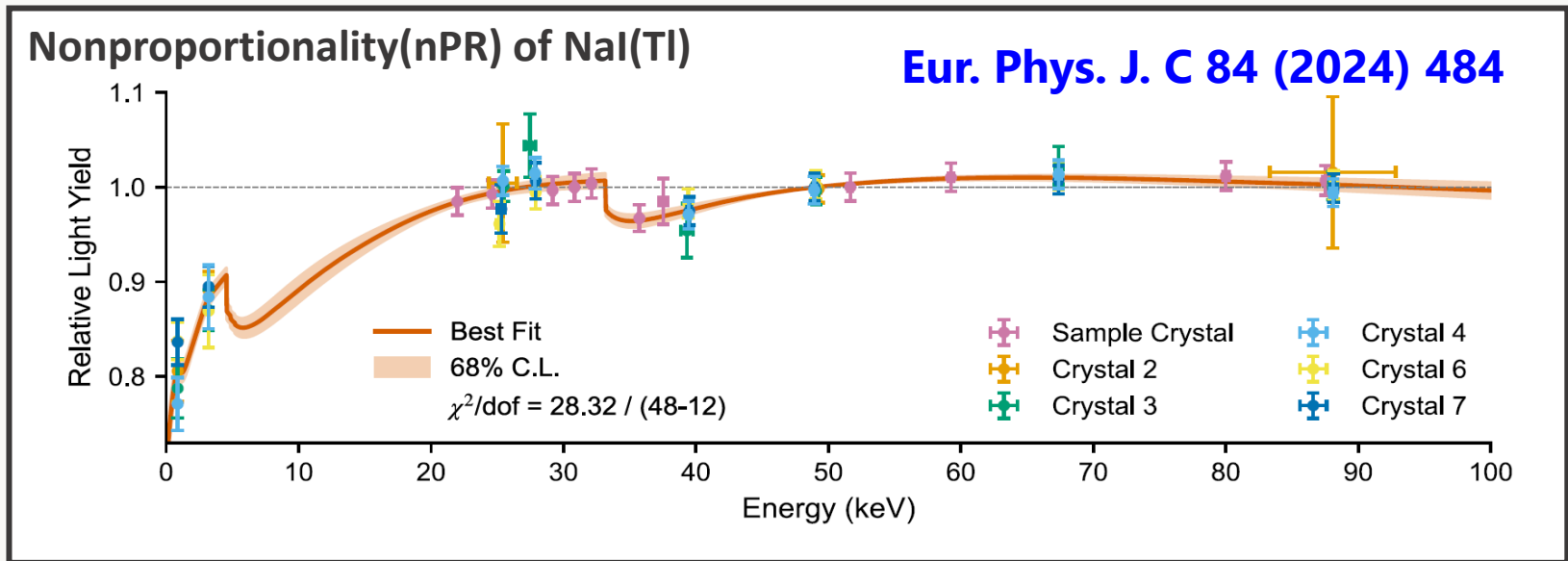
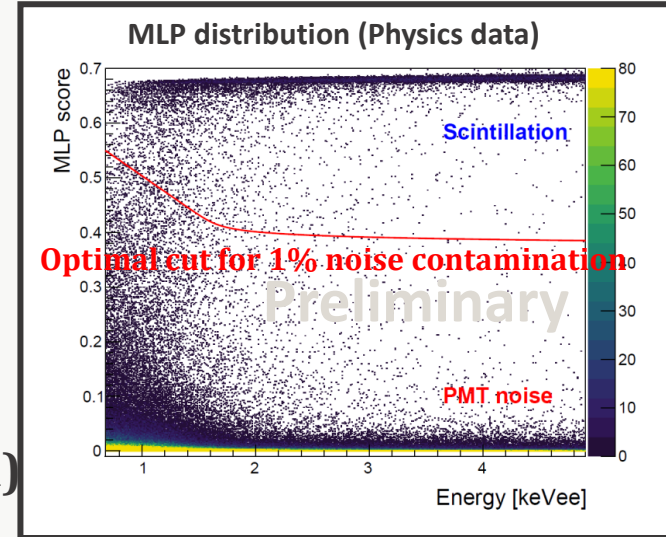
# Detector understanding update with SET3 data

- **Event selection**

- **Multivariable analysis with pulse-shape parameters**
  - Meantime, likelihood parameters (SET2)
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- 1 keV  $\rightarrow$  **0.7 keV**

- **Background modeling**

- Considers updated **nonproportionality of NaI(Tl)**



# Detector understanding update with SET3 data

## • Event selection

- Multivariable analysis with pulse-shape parameters

- Meantime, likelihood parameters (SET2)

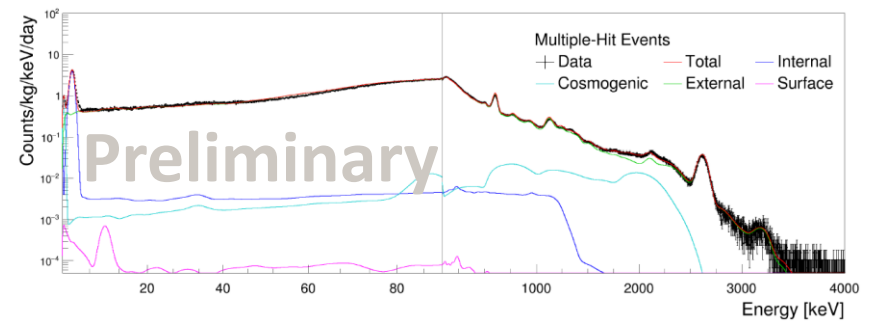
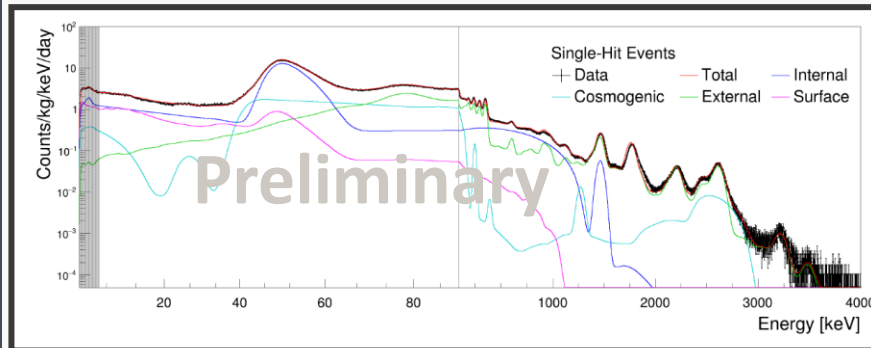
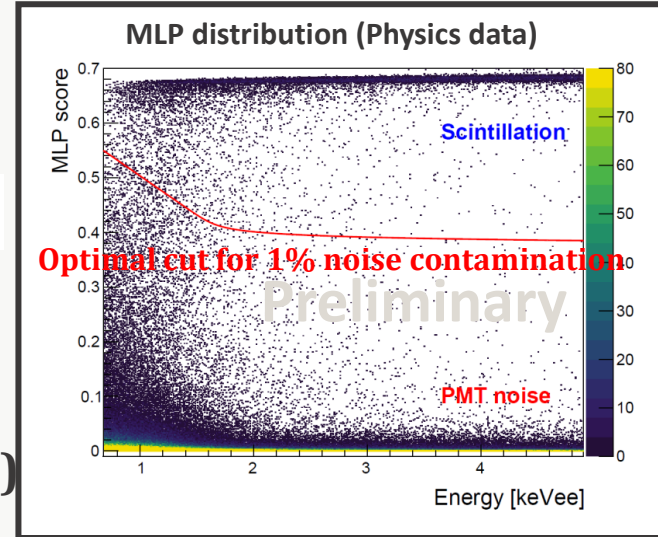
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## Event selection

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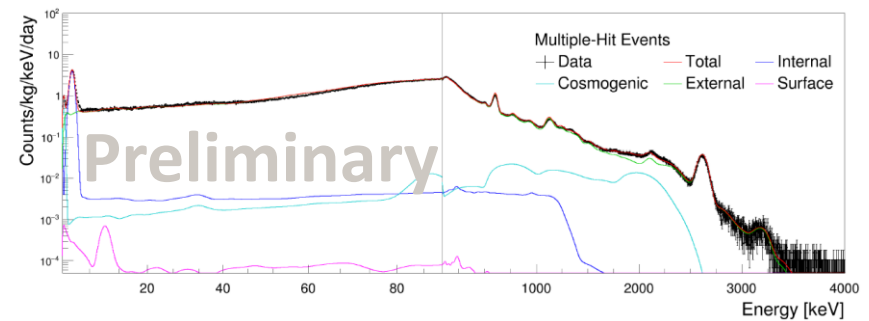
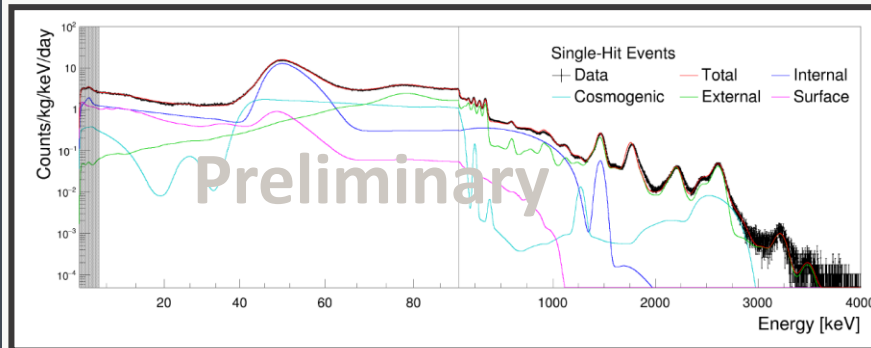
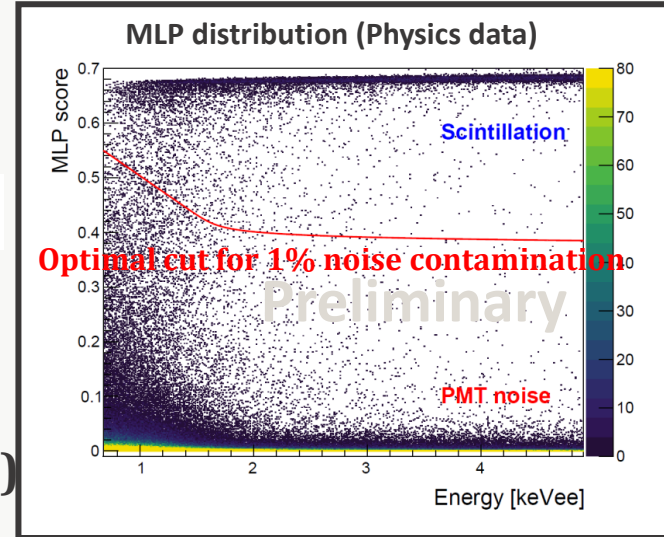
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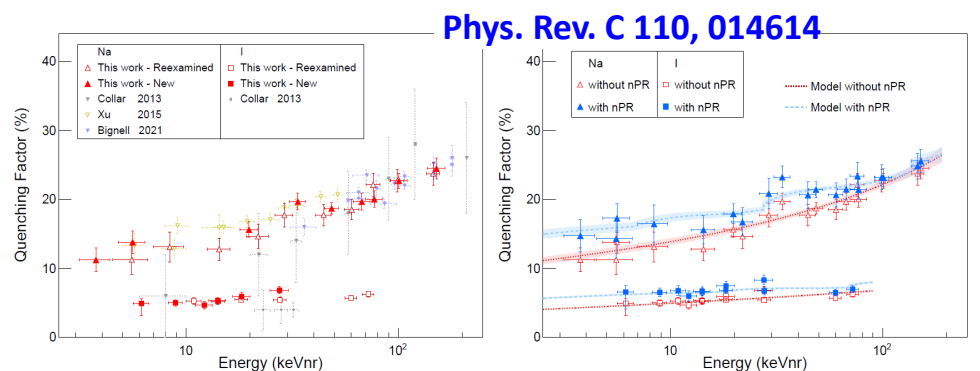
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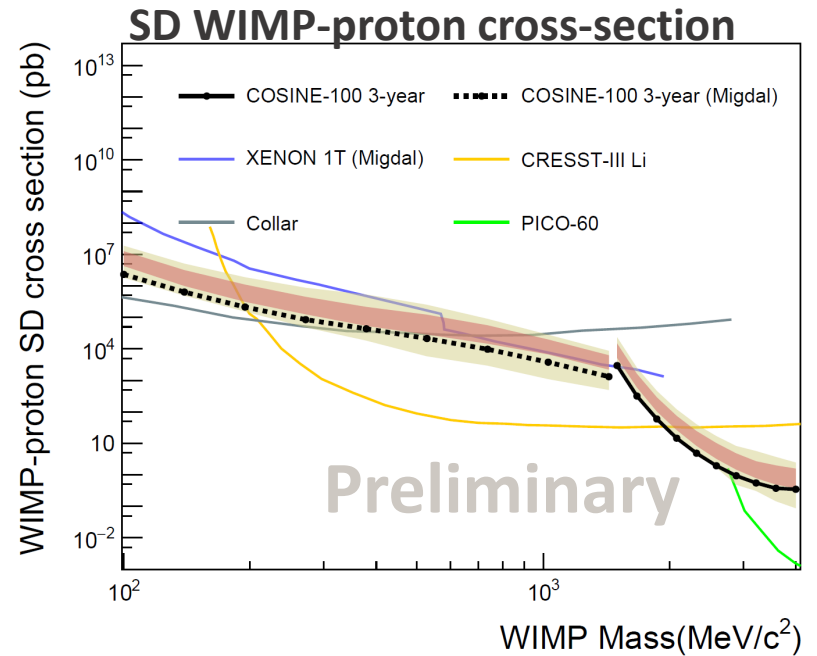
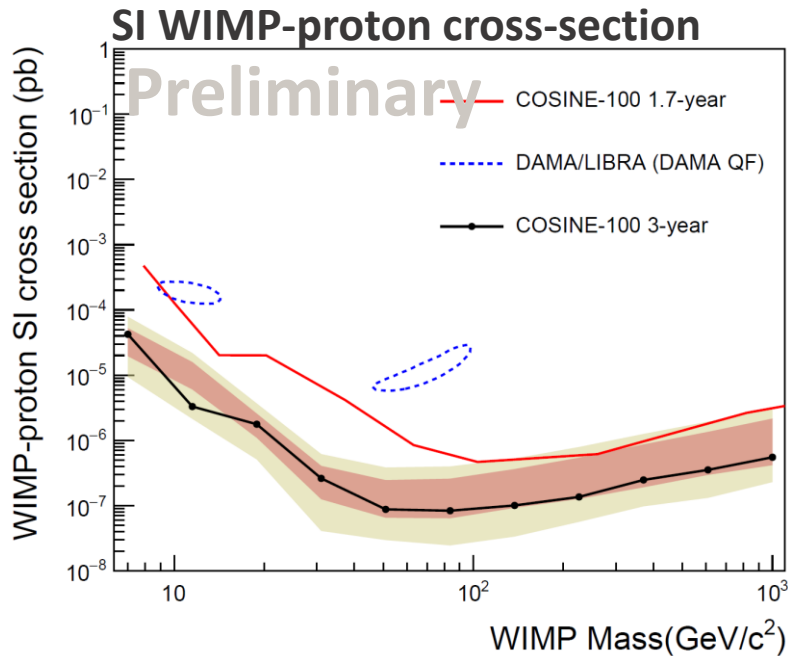
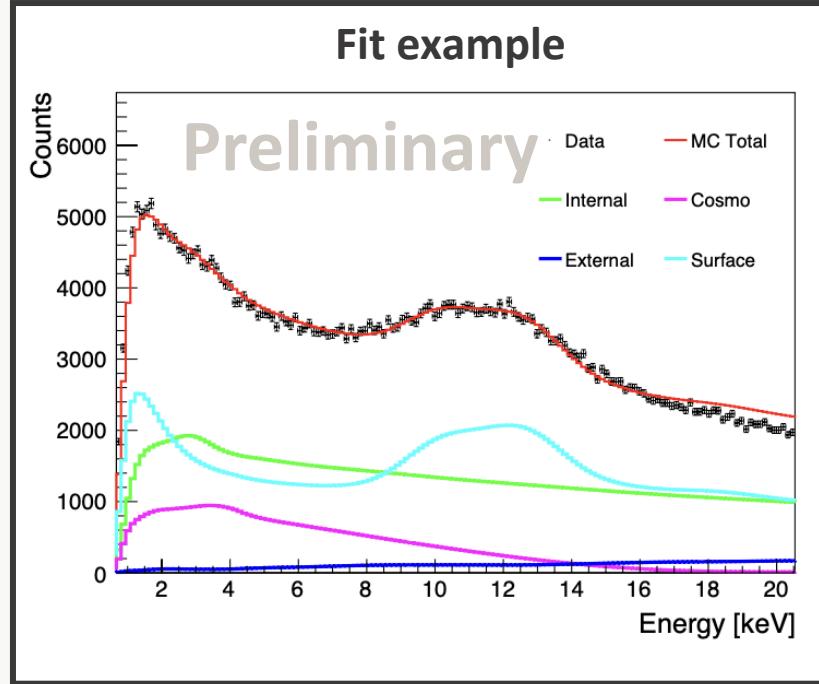
## Quenching factor of NaI(Tl)

- Update low energy region



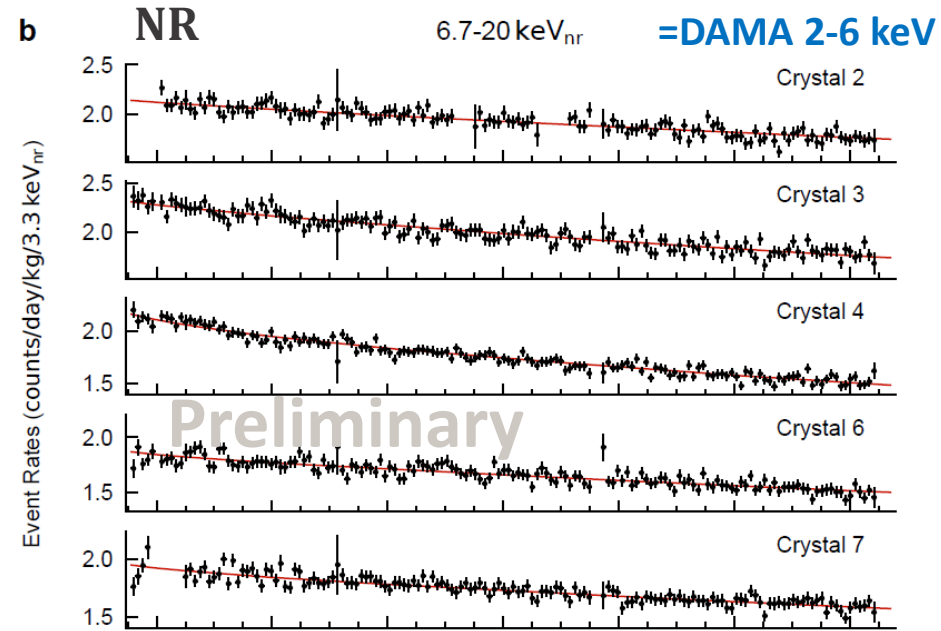
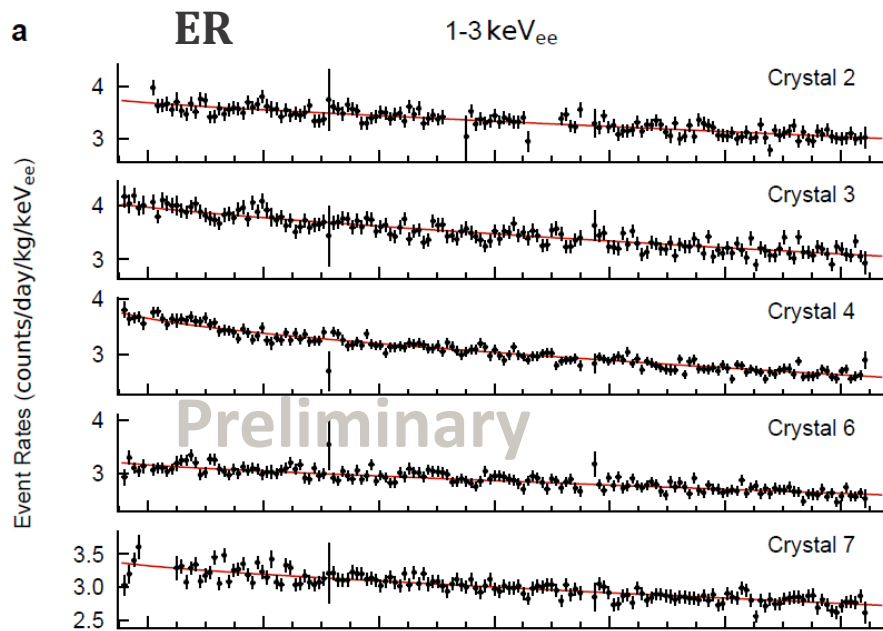
# WIMP Extraction

- Analyze the **SET3** data with **new event selection** and **new background modeling**
- **O(10)** improvement of SI WIMP-proton cross-section **limit** from **previous result**



# Annual modulation analysis

- Using **full COSINE-100 data (~6 years)**
  - Using SET 3 event selection and modeling
  - Calibration method to test DAMA's claim
    - Calibration for **Electron Recoil (ER)**
    - Calibration for **Nuclear Recoil with new quenching factor (NR)**



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  - Applying **full model of time dependent background component**

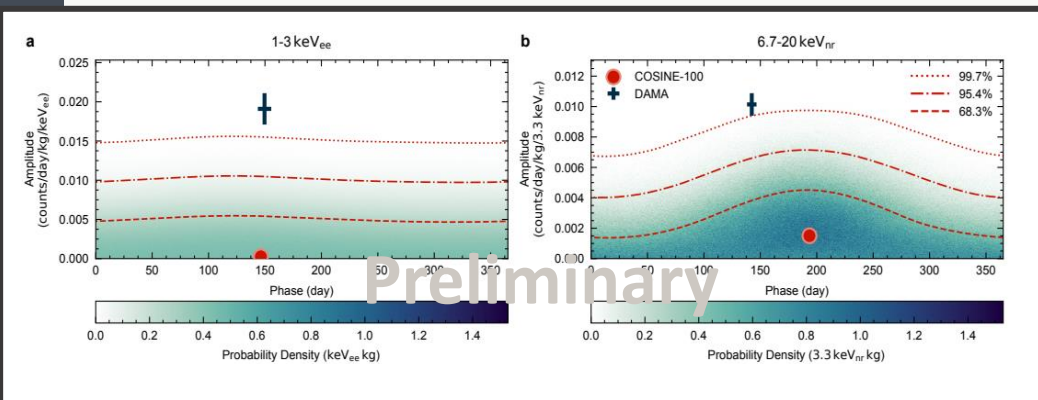
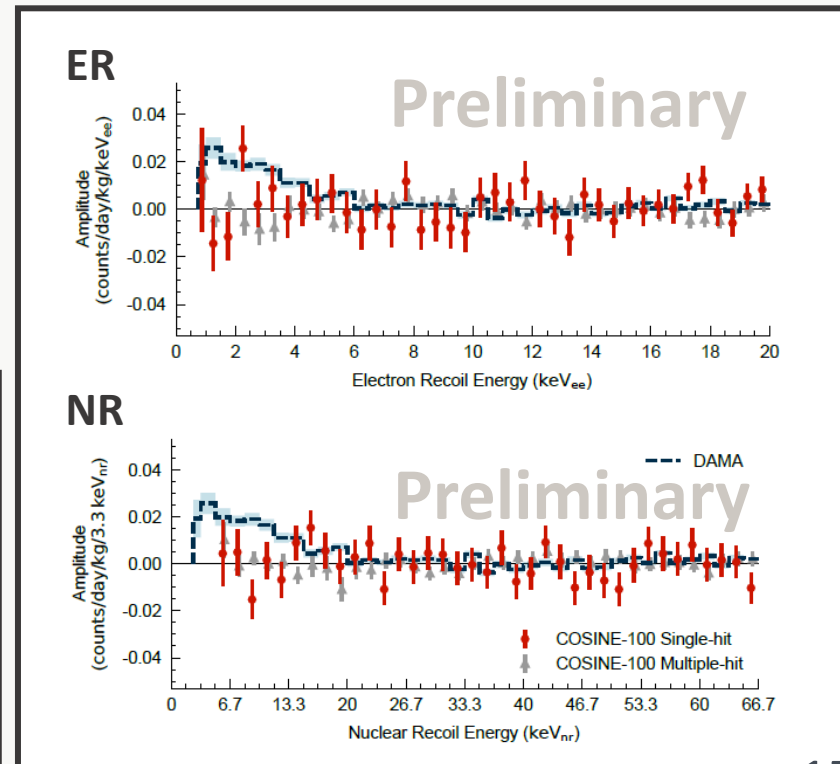
$$R^i(t|S_m, \alpha^i, \beta_k^i) = \alpha^i + \sum_{k=1}^{N_{bkgd}} \beta_k^i e^{-\lambda_k t} + S_m \cos(\omega(t - t_0))$$

Constant from long-lived backgrounds  
Exponential decays from short-lived  
Modulation signal

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Energy (keV)		$A_{\text{COSINE}}$ ( $10^{-3}$ DRU)	$A_{\text{DAMA}}$ ( $10^{-3}$ DRU)	Exclusion C.L. ( $\sigma$ )
ERC	1-3	$0.4 \pm 5.0$	$19.1 \pm 2.0$	3.57
NRC	6.7-20	$1.3 \pm 2.7$	$10.0 \pm 0.7$	3.23

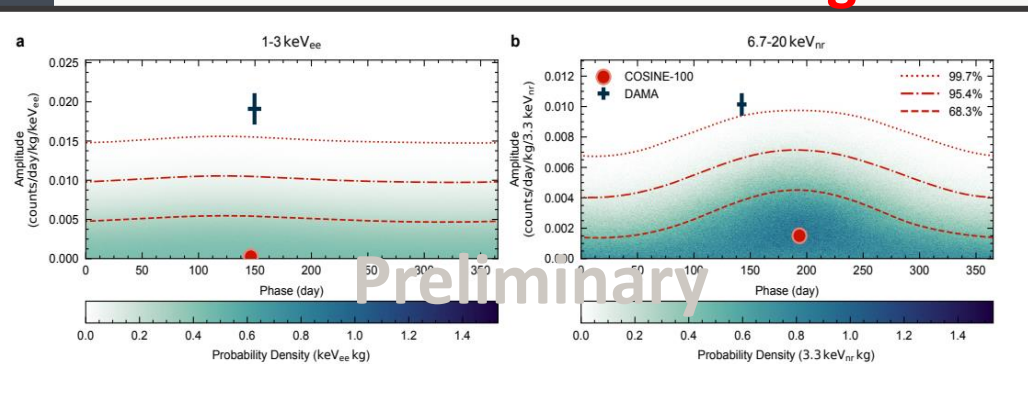
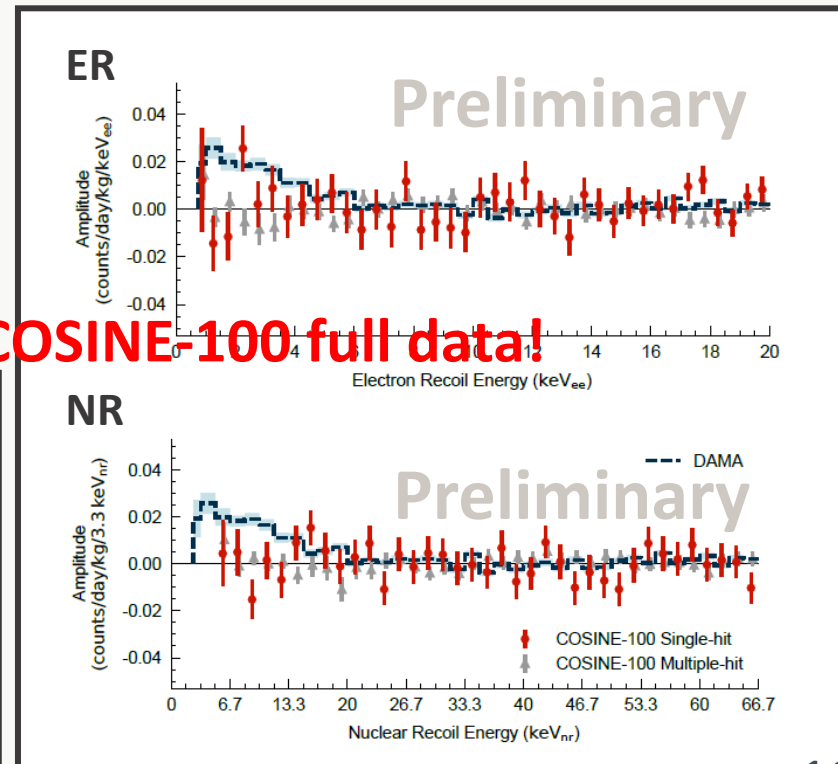


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**No DAMA like signal in COSINE-100 full data!**

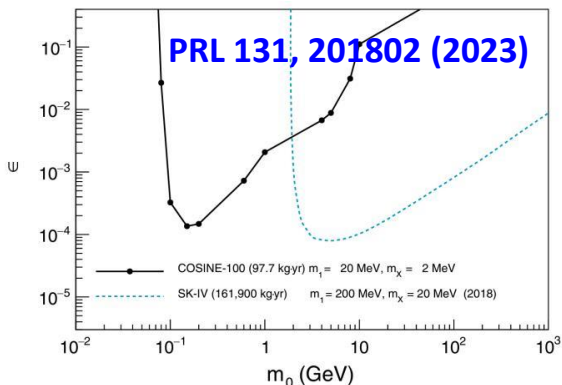
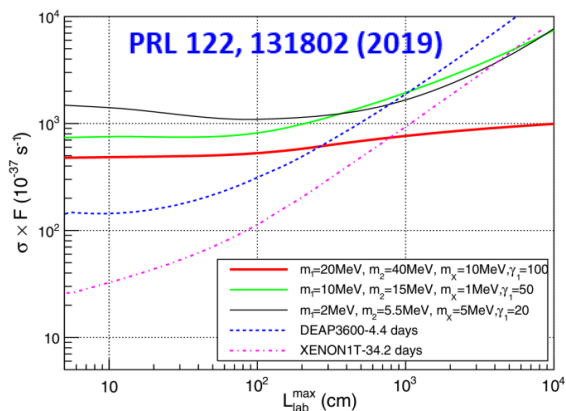




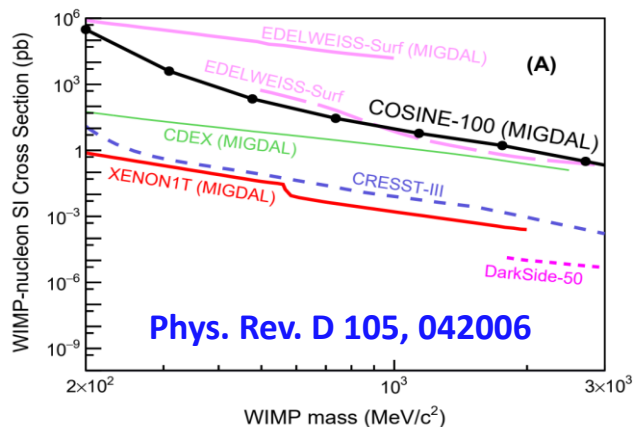
# Other DM scenarios

- Performed with SET2 data
- No dark matter signal

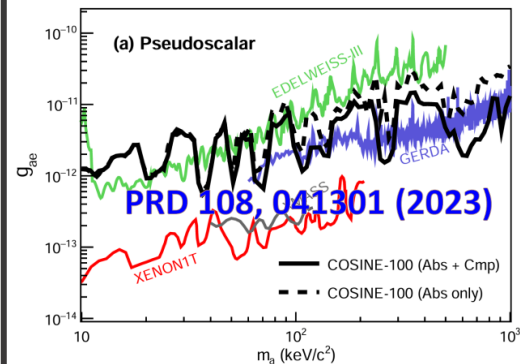
## Boosted DM



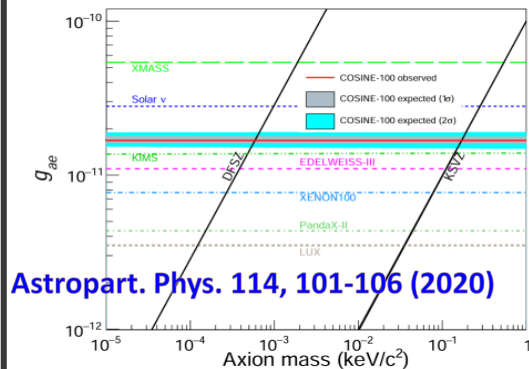
## Migdal effect



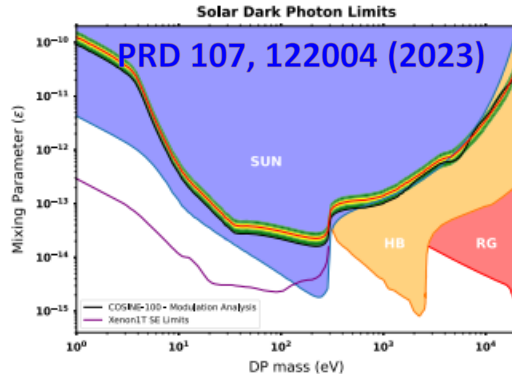
## Bosonic Super WIMP



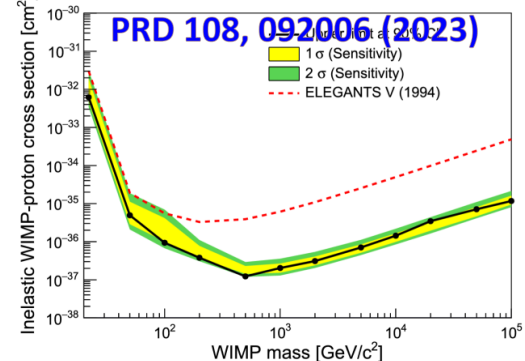
## Solar Axion



## Solar Bosonic Particles

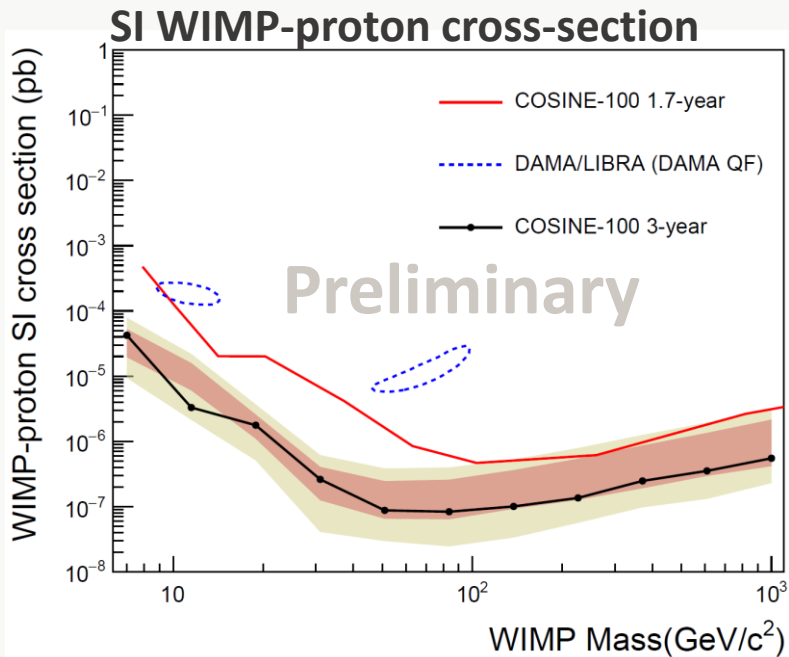


## WIMP-I inelastic scattering

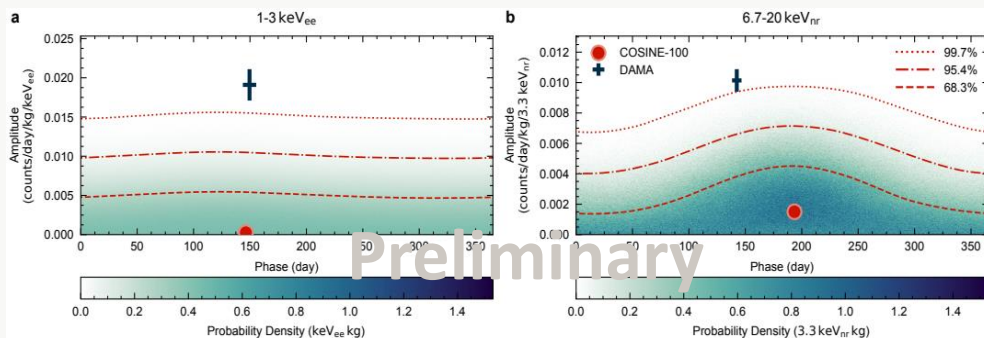


# What is next for COSINE-100?

- Spectral Analysis



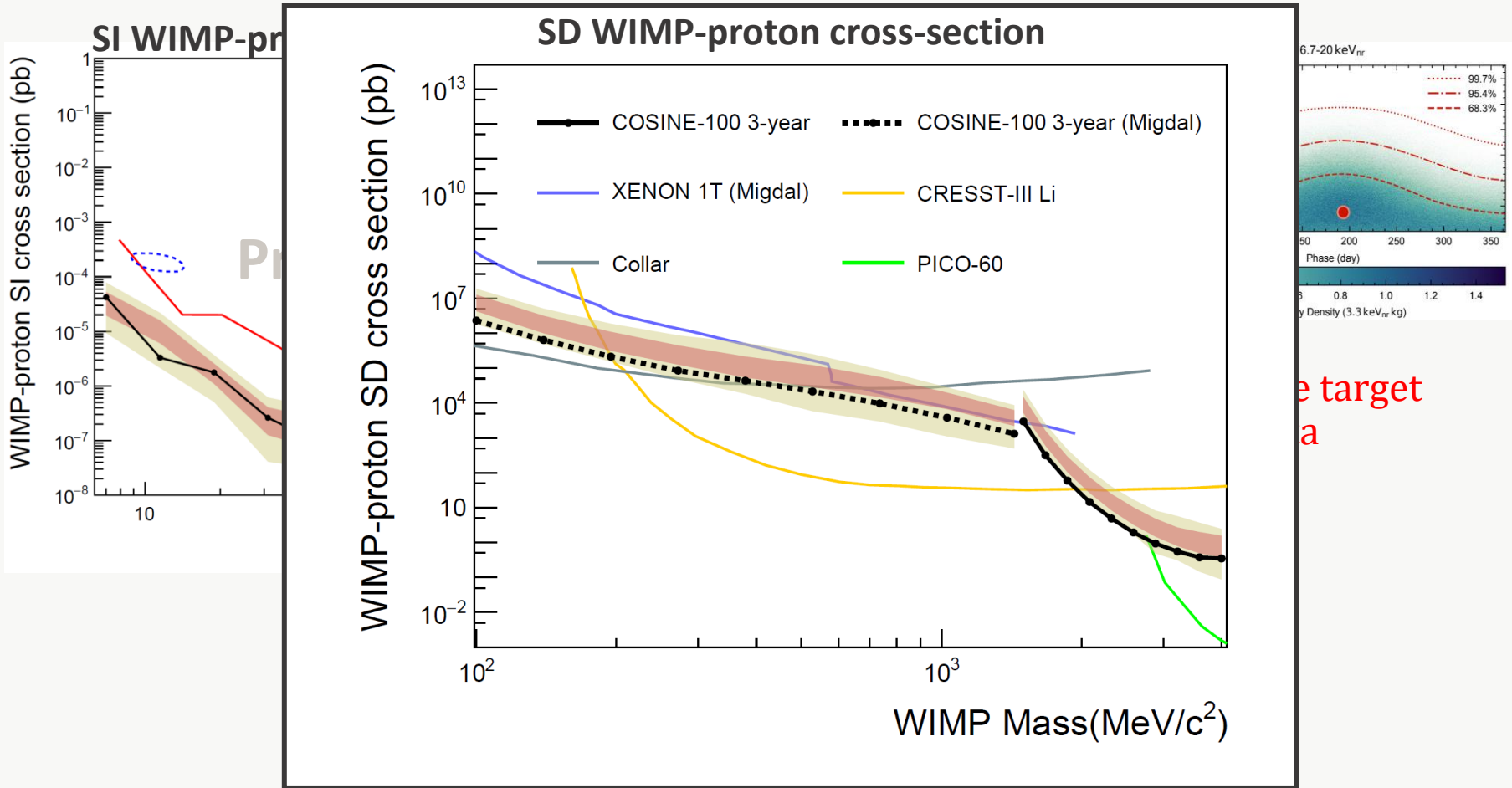
- Modulation analysis



- **No DAMA-like signal in same target NaI(Tl) with COSINE-100 data**

# What is next for COSINE-100?

- Spectral Analysis
- Modulation analysis

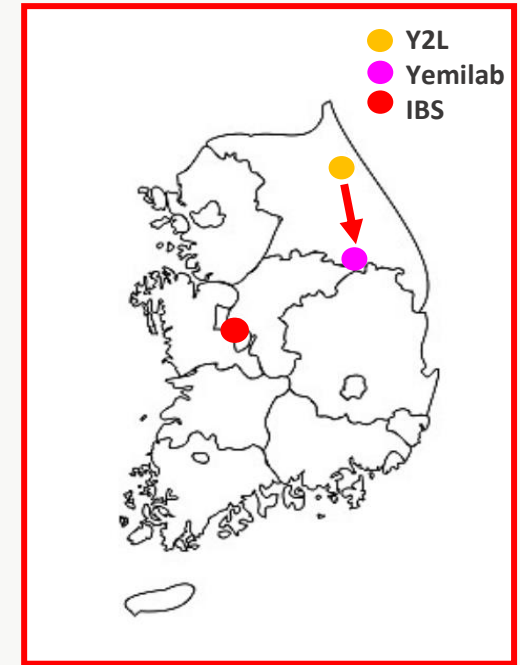


- Explore **low mass WIMP** with **lower threshold!**

→ **Need higher light yield !**

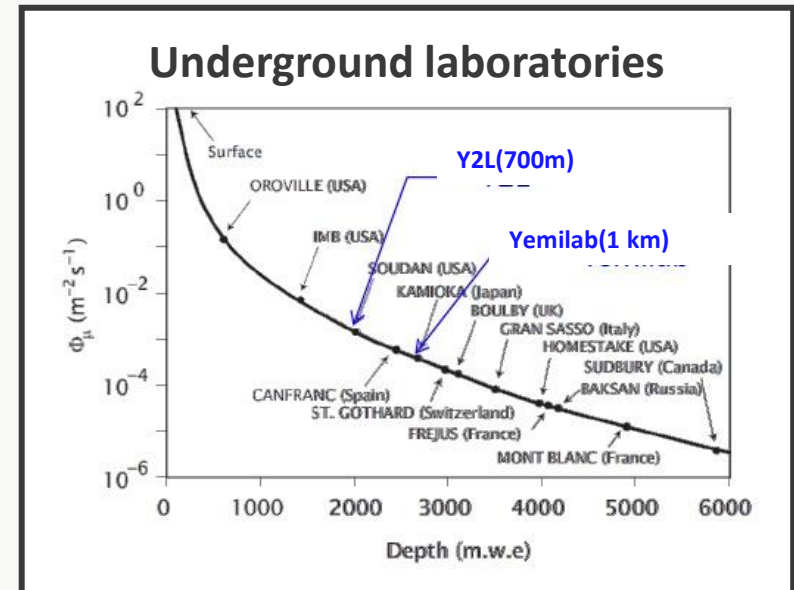
# COSINE-100 Upgrade

- **For lower threshold!**
- **Enhance light yield of COSINE-100 crystals**
  - Operate at  $-35^{\circ}\text{C}$  environment
  - Upgrade encapsulation design
- **Relocate COSINE-100 setup @ Yemilab**
  - Newly constructed underground laboratory in Korea
  - Y2L  $\rightarrow$  Yemilab from 2023 to 2024 [Front. Phys. 02 April 2024](#)
  - 700 m  $\rightarrow$  1000 m overburden
- **Same shielding structure**



COSINE-100 NaI(Tl) crystals

Crystal #	Size (diameter x length)	Light yield(PEs/keV)
1	5.0" x 7.0"	14.9 $\pm$ 1.5
2	4.2" x 11.0"	14.6 $\pm$ 1.5
3	4.2" x 11.0"	15.5 $\pm$ 1.5
4	5.0" x 15.3"	14.9 $\pm$ 1.5
5	5.0" x 15.5"	7.3 $\pm$ 0.7
6	4.8" x 11.8"	14.6 $\pm$ 1.5
7	4.8" x 11.8"	14.0 $\pm$ 1.4
8	5.0" x 15.5"	3.5 $\pm$ 0.3



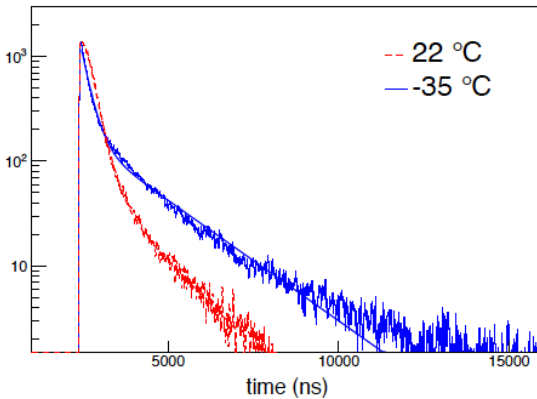
# NaI(Tl) at $-35^{\circ}\text{C}$

[Astroparticle Physics 141 \(2022\) 102709](#)

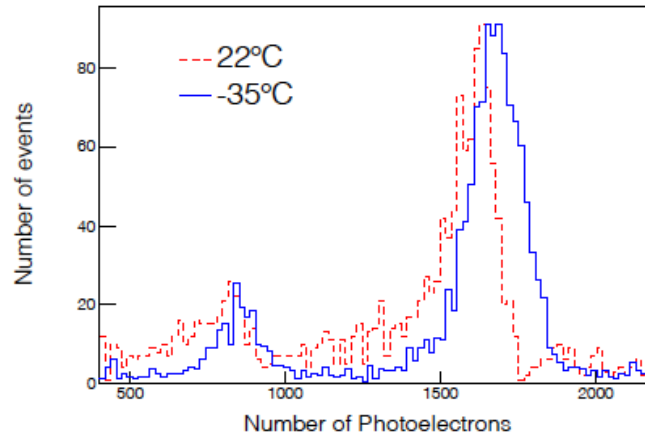
- **Low temperature operation (at  $-35^{\circ}\text{C}$ )**
  - $\sim 5\%$  increased **Light yields**
  - $\sim 9\%$  increased **alpha quenching**

## Accumulated waveform

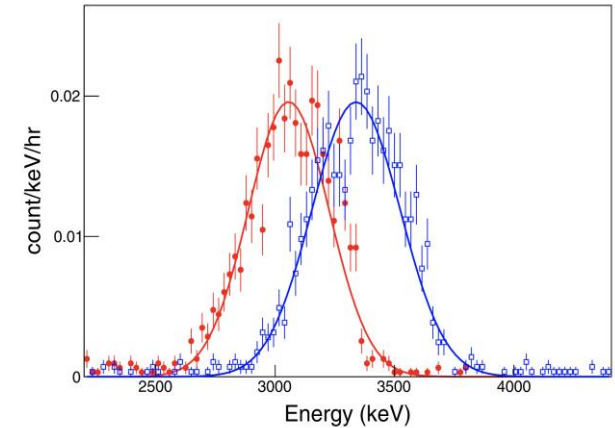
Accumulated waveform of  $^{241}\text{Am}$  Events



## NPE Measurement



## Alpha response using $^{210}\text{Po}$



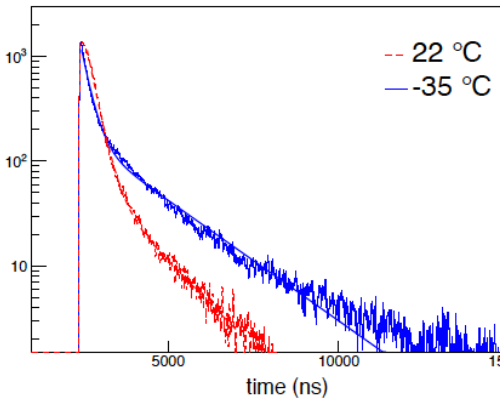
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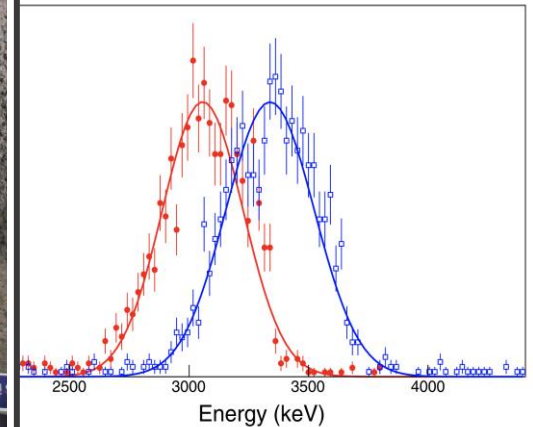
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## COSINE detector fridge room @Yemilab



## Alpha response using $^{210}\text{Po}$



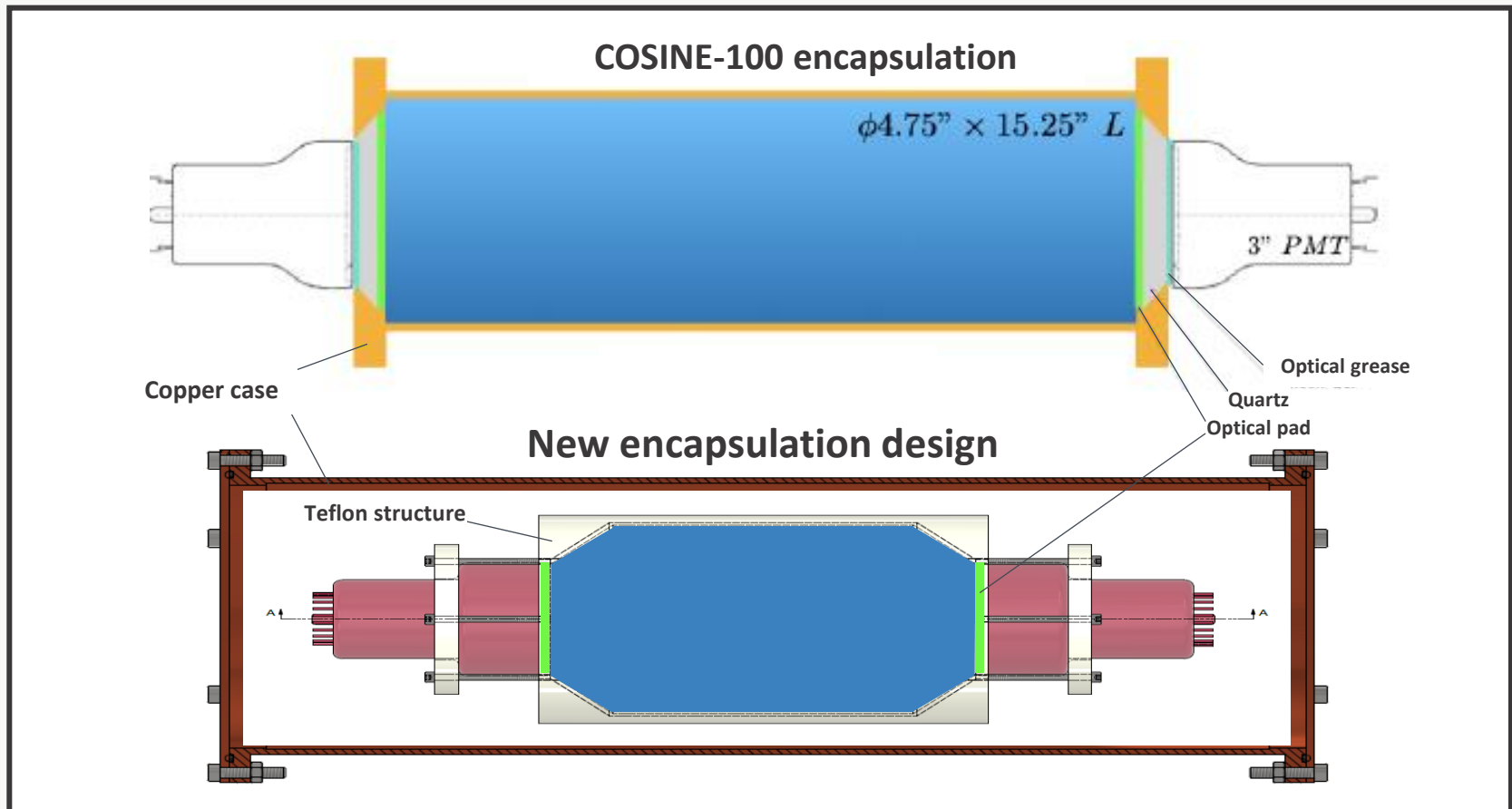
- **The detector fridge room was prepared for COSINE-100U!**

# New encapsulation design

- **Minimize optical coupling layer** between PMT and crystal
  - 3 layers  $\rightarrow$  1-layer optical pad
- **Teflon structure** for stable optical contact
- **Copper case** for protection from outside
- **Expect ~50% increase in light yield**
  - 15 P.E.  $\rightarrow$  >20 P.E.
- **Machining crystal edge** with **light guide-like shape**
  - 7 cm diameter = PMT photocathode size
  - Maximize light collecting

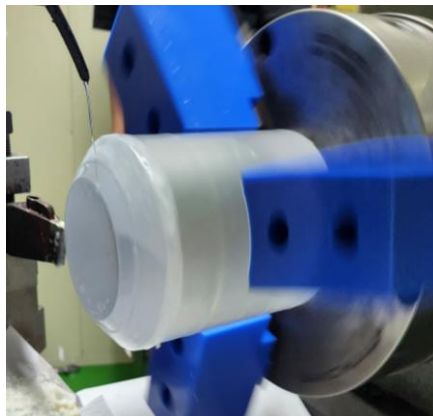
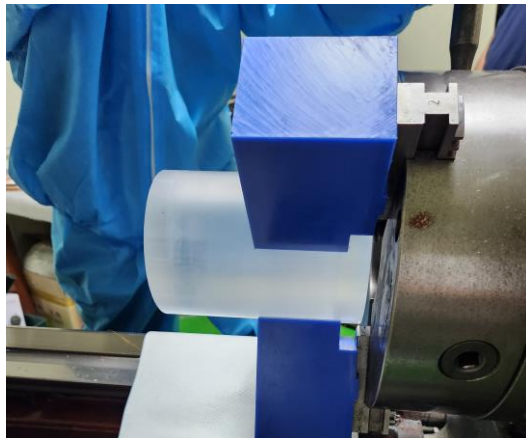
[Nucl. Instrum. Meth. A 981, 164556 \(2020\)](#)

[arXiv:2404.03691](#)

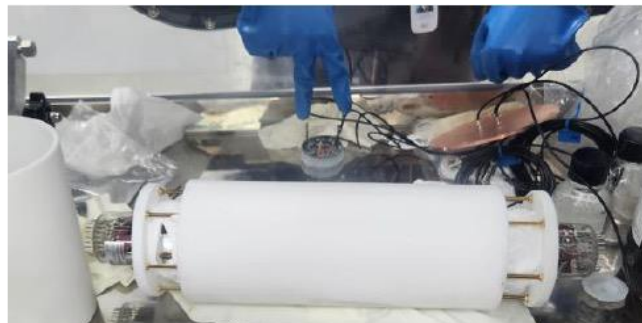
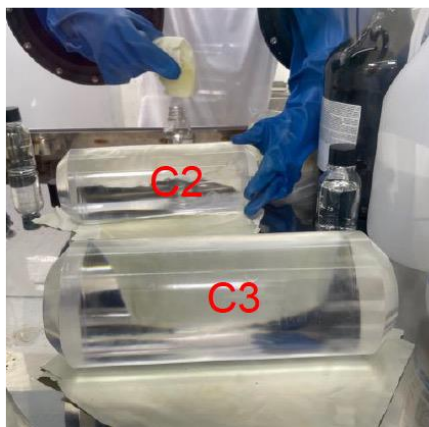


# New encapsulation assembling

## Machining crystal edge



## Assembling @glovebox ( $H_2O < 10$ ppm)

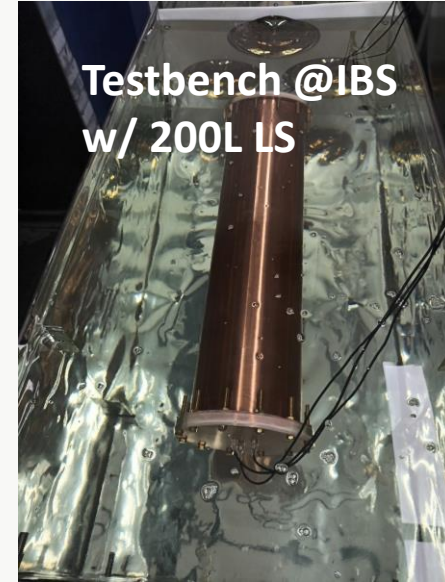




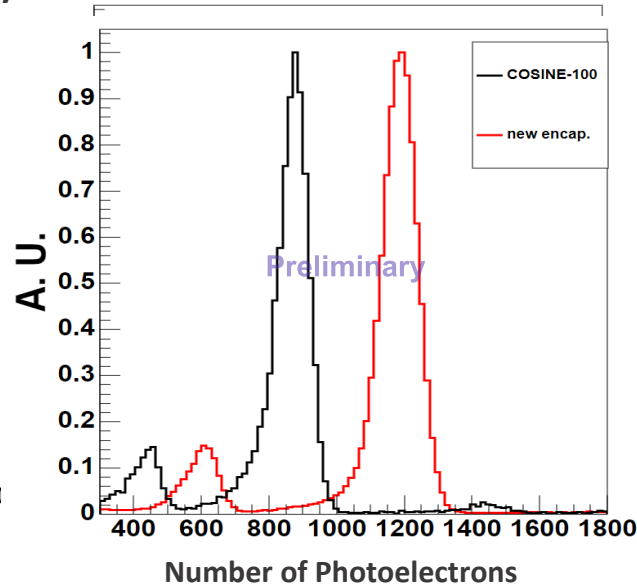
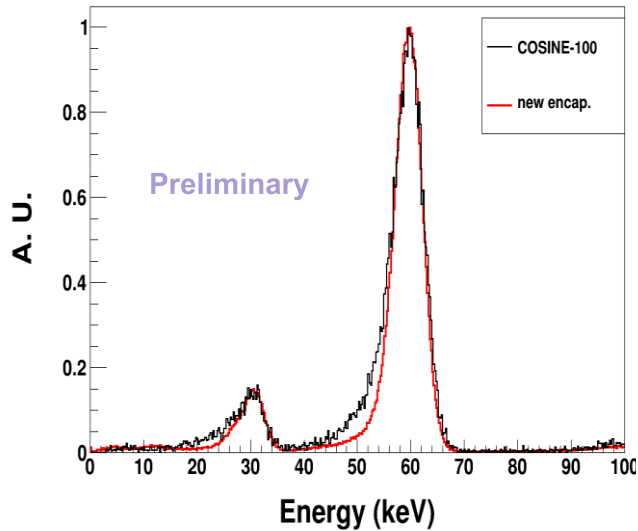
# New encapsulation performance

Light yields

- Test performance in HQ testbench  
(On ground w/200 L LS)



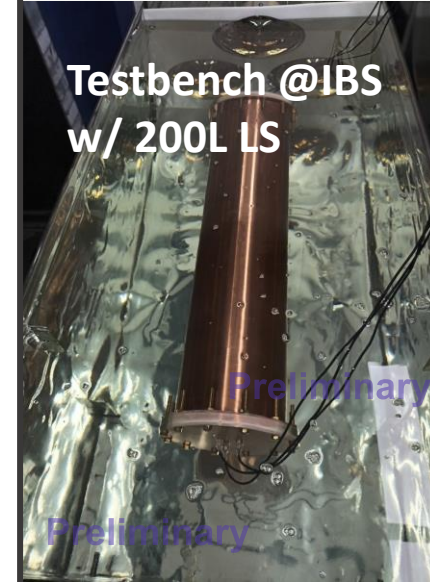
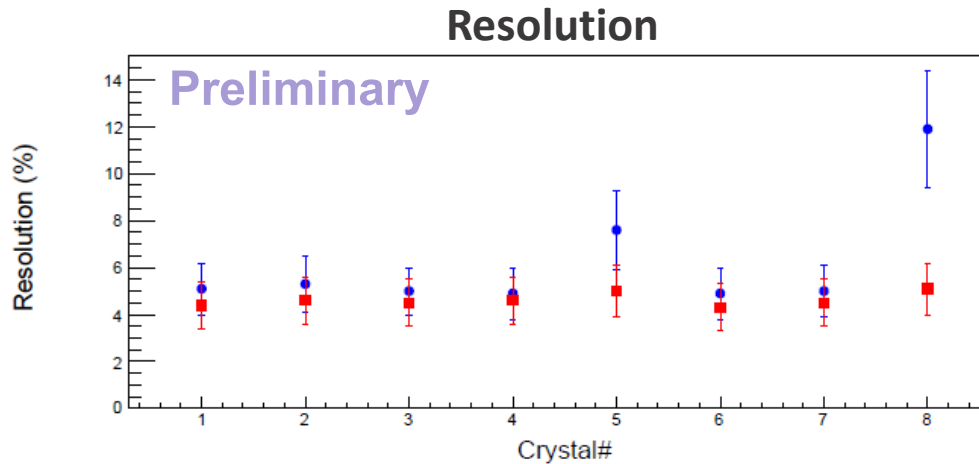
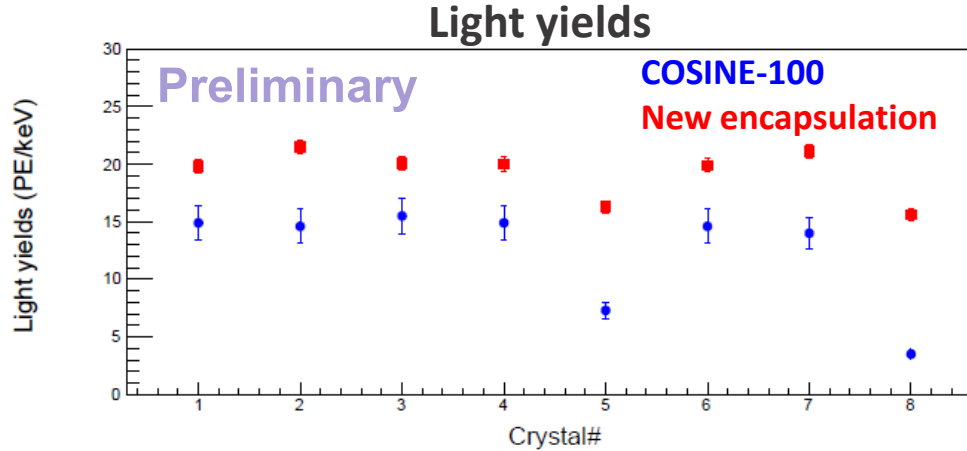
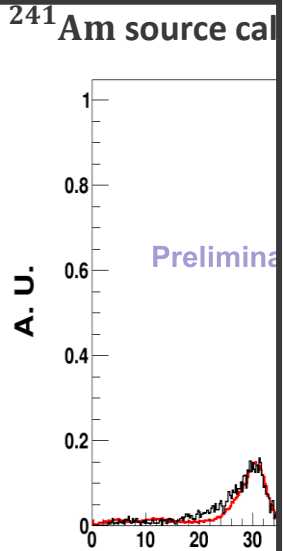
<sup>241</sup>Am source calibration (59.54 keV), Crystal# 2



- Resolution :  $5.3 \pm 1.2\% \rightarrow 4.6 \pm 1.0\%$   
Light yield :  $14.6 \pm 1.5 \rightarrow 21.0 \pm 0.6 \text{ PE/keV}$

# New encapsulation performance

- Test per  
(On grou



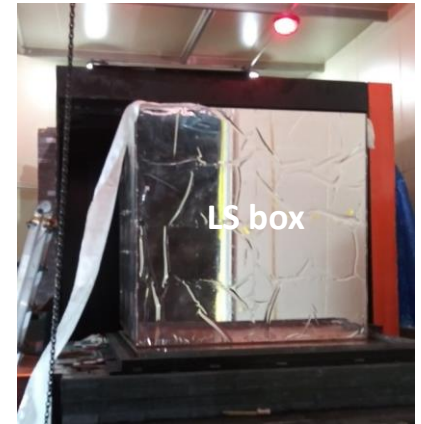
- Resolut
- Light yi

- **Assembling** for all crystals was **finished!**
  - Up to **45% light yield** improvement
  - **Resolution** improved by **~10% @ 59.57keV**

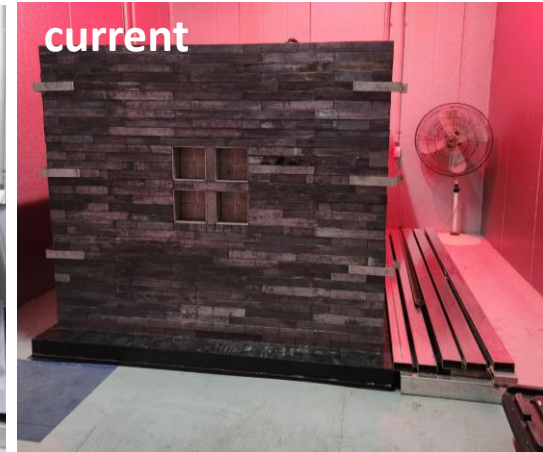
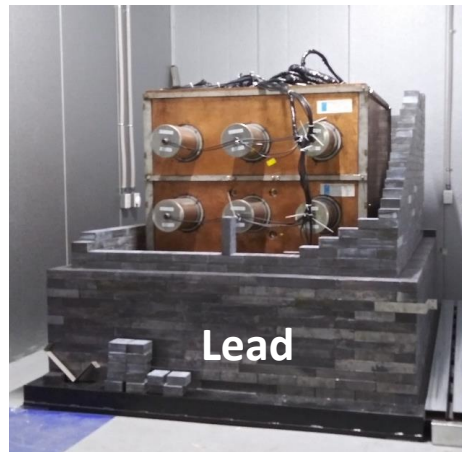
# Status on COSINE-100U

- Y2L → Yemilab (From Sep. 2023)

## COSINE-100 decommissioning @Y2L



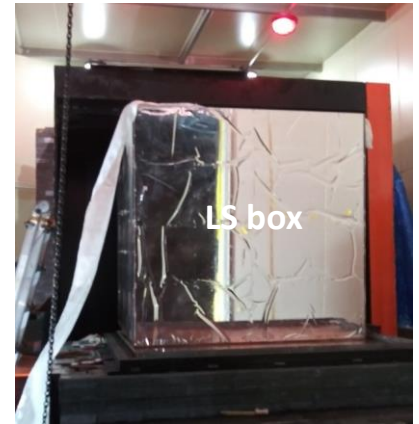
## Shield installation in fridge room @ Yemilab



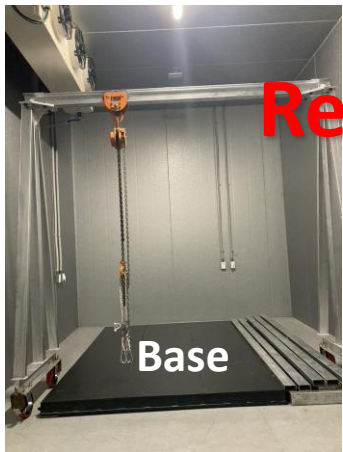
# Status on COSINE-100U

- Y2L → Yemilab (From Sep. 2023)

## COSINE-100 decommissioning @Y2L



## Shield installation in fridge room @ Yemilab



**Ready to install COSINE-100U detectors!**

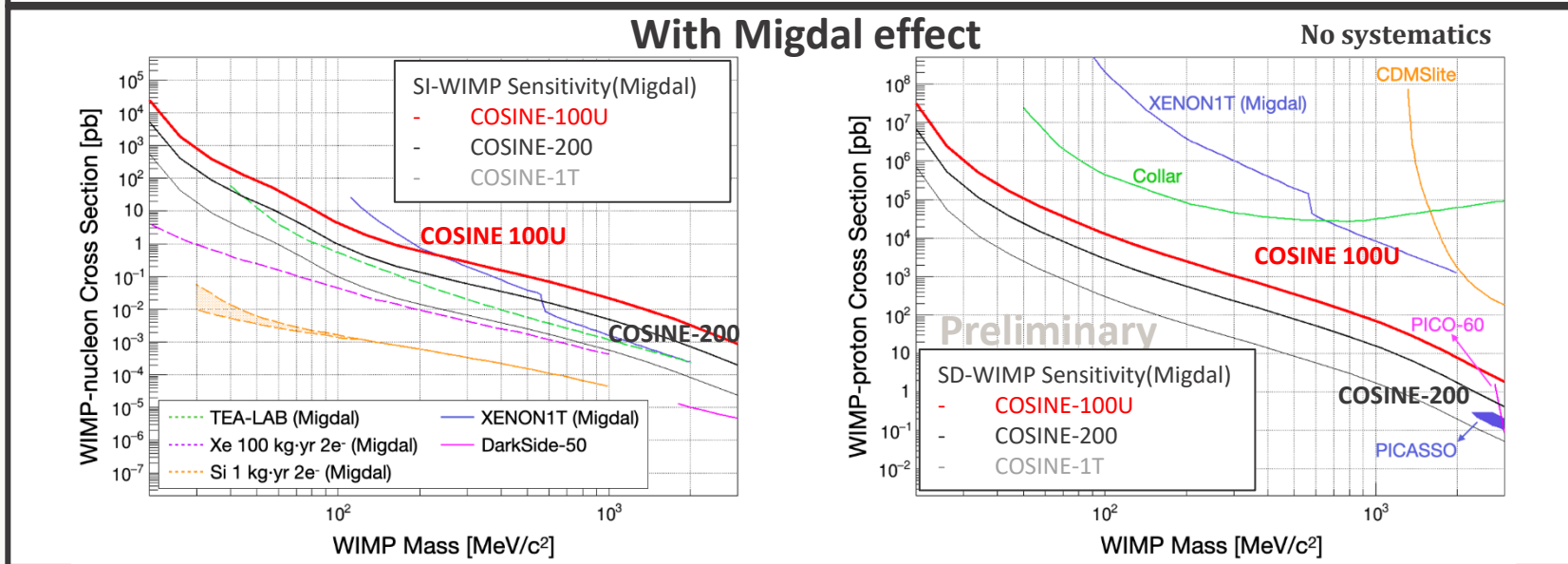
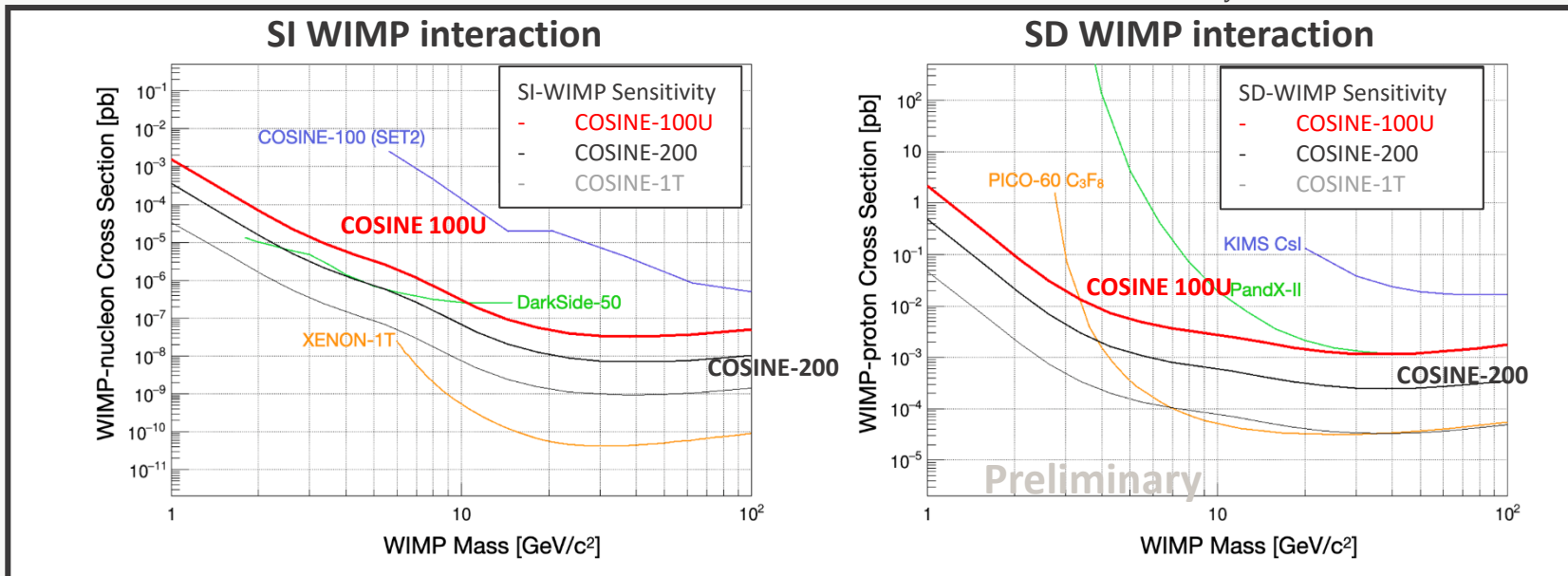
# Plan for COSINE-100U

	2024-01	2024-02	2024-03	2024-04	2024-05	2024-06	2024-07	2024-08
Crystals		Assembling&Test					Install	Physics run!!
Lead	Bottom	Side						
LS		PMTs ,LS production						
Muon	Fix							
Hardware					DAQ, PCs, HV, Cable, Monitoring			
Software		DAQ, production, monitoring						

- Under preparation to start **physics run in August!**

# Sensitivity estimation

- Condition for upgrade COSINE-100 (COSINE-200)
  - 1 year
  - Light yields 22 P.E./keV
  - COSINE-100 (R&D crystal) background
  - No systematics



# Summary

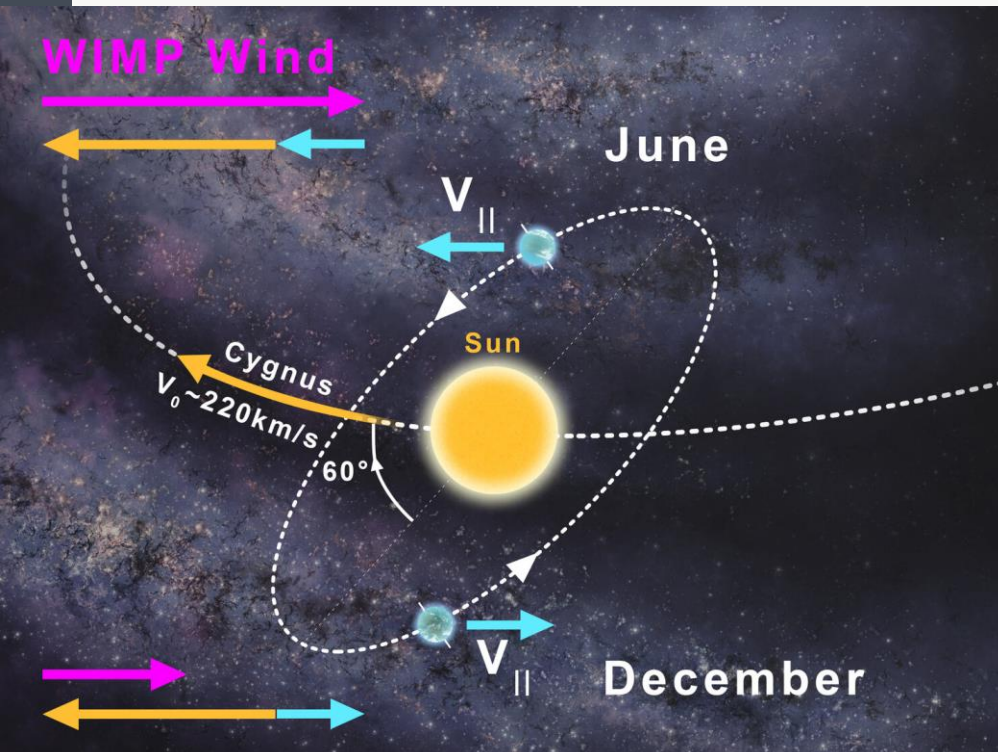
- COSINE-100 was designed to **test the DAMA/LIBRA signal** via the **same target material, NaI(Tl)** from Sep. 2016 to Mar. 2023
- Spectral Analysis
  - **Improved WIMP-nucleon cross section using 3-years data**
  - Investigation of various DM scenarios
- Modulation analysis
  - **No DAMA like modulation** with more than  **$3\sigma$  significance** using full COSINE-100 data
- **COISNE-100U** experiment is being prepared using detectors with **enhanced light yields** detectors in newly constructed underground laboratory.
  - **40% light yield improvement** with new encapsulation technique.
  - Detector fridge room for  **$-35^{\circ}\text{C}$  operation** is ready
- **COSINE-100U** operation will start in **August 2024** for searching low mass WIMP!

# Back up





# Annual modulation signal



- Spherical halo of dark matter in our galaxy is expected to make dark matter flux to Solar system (WIMP Wind)
- Speed of WIMP wind changes according to seasonal rotation of the Earth.
- Changing WIMP speed induces changing interaction rate.

## Rate of WIMP elastic scattering

$$R \propto N_T \cdot \sigma_{\chi N} \cdot \frac{\rho_{\chi}}{m_{\chi}} \int_{v_{min}} \frac{f(v)}{v} dv$$

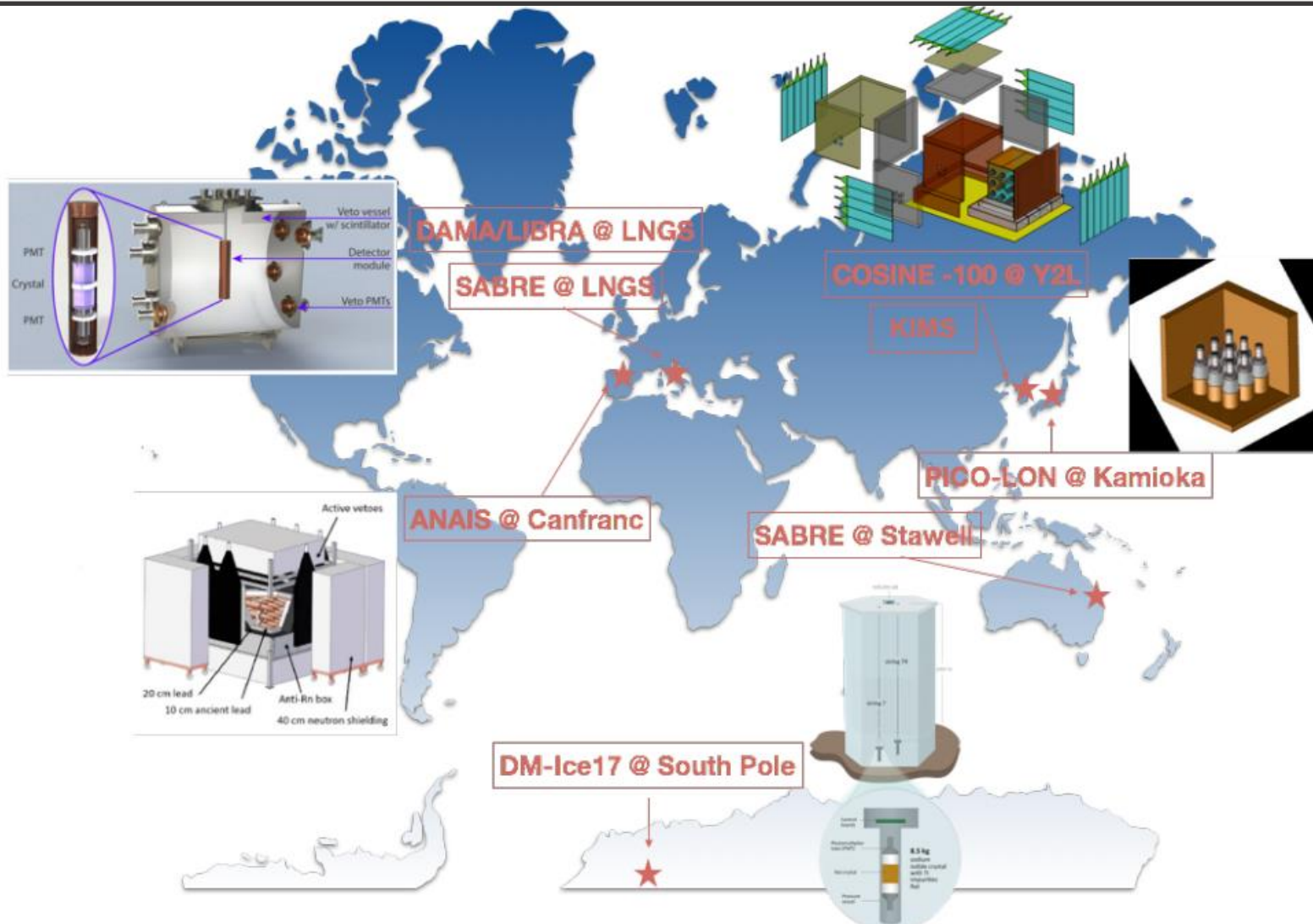
$f(v)$  : WIMP velocity distribution

$v_{min}$  : minimum velocity for interaction

$\rho_{\chi}$  : local WIMP density

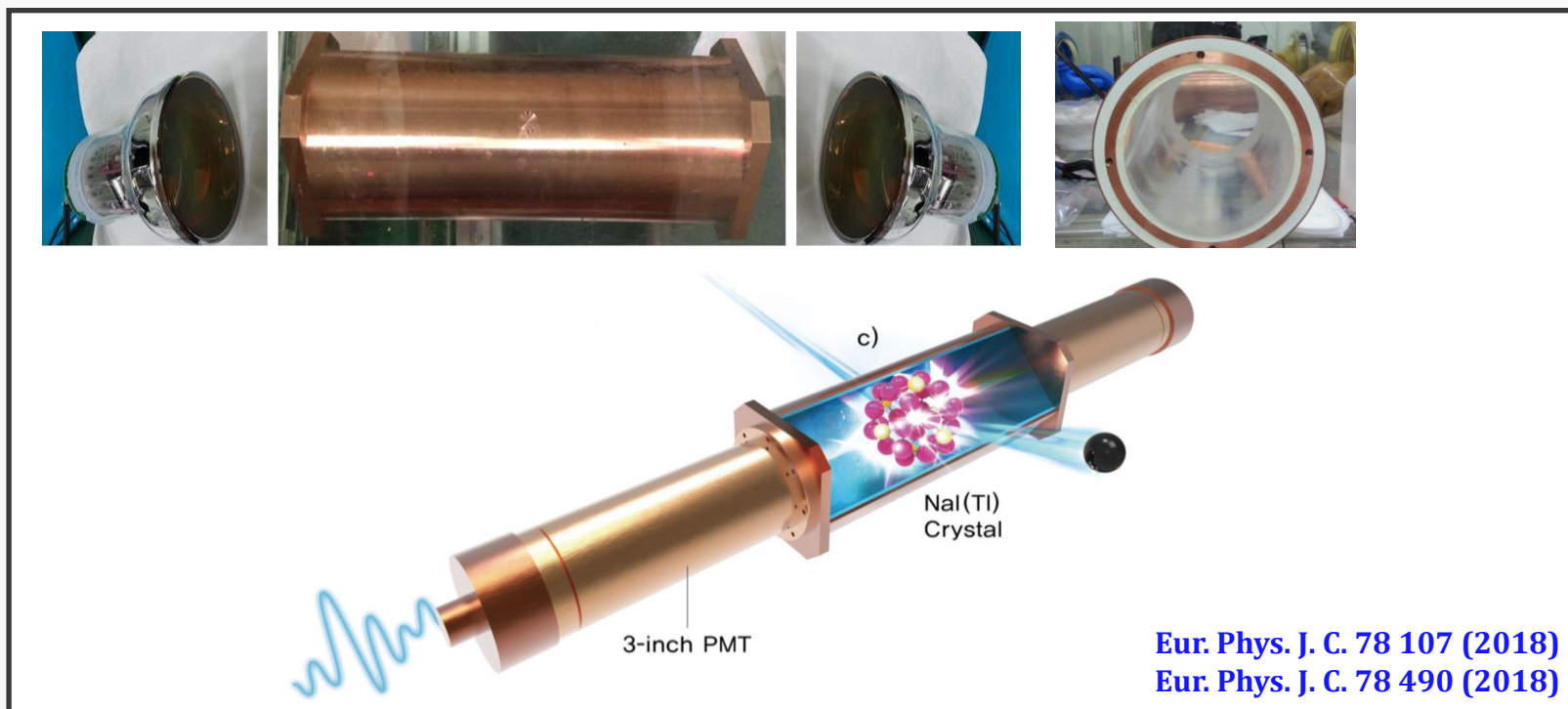
$m_{\chi}$  : WIMP mass

# NaI(Tl) Dark matter search



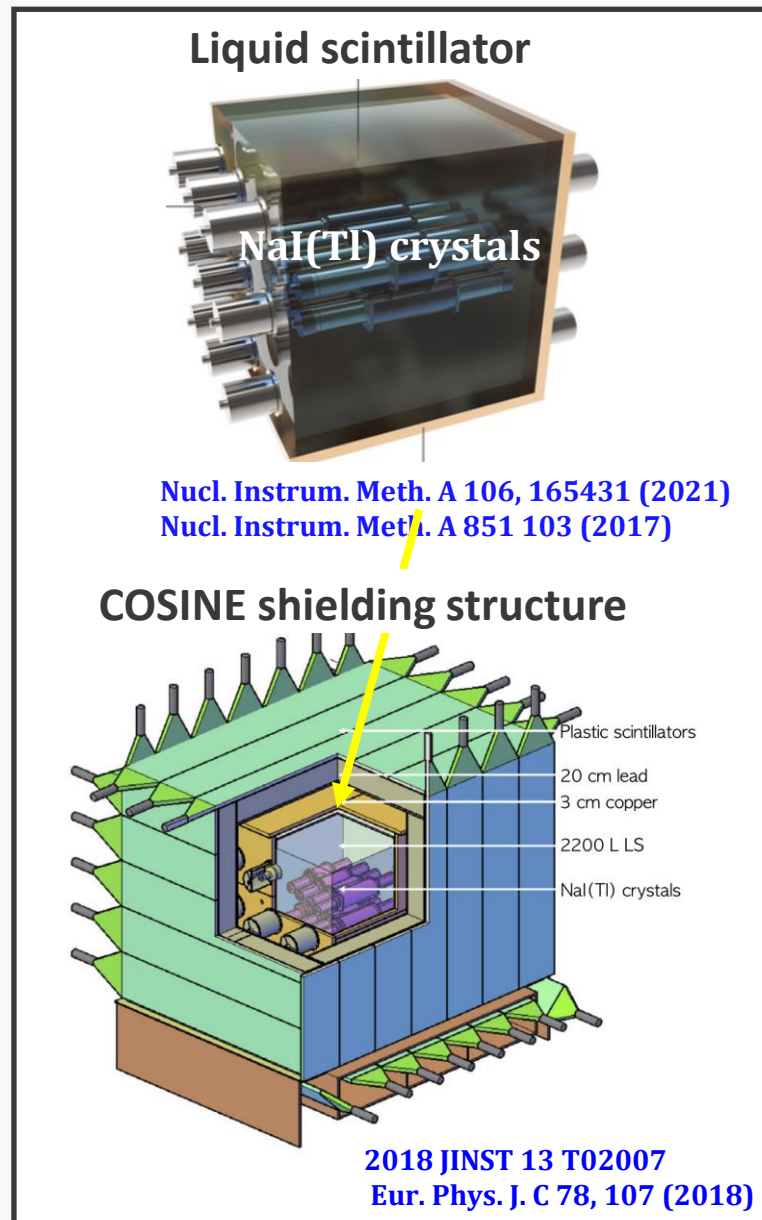
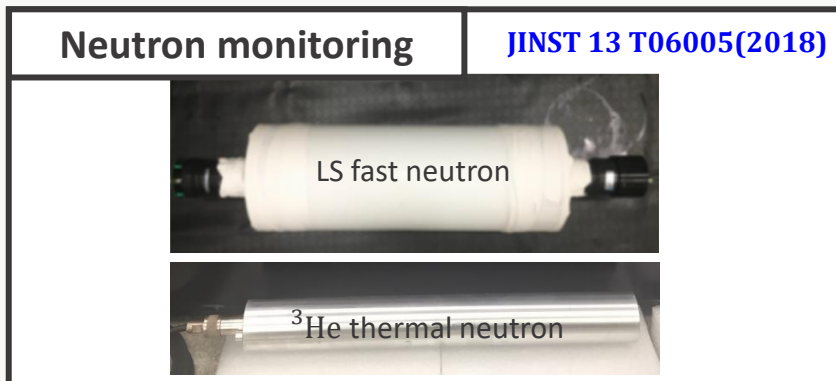
# COSINE-100 experiment-NaI(Tl)

- 8 low-background NaI(Tl) crystals with 106 kg in total
  - U/Th/K level is less than DAMA, but total alphas ( $^{210}\text{Pb}$ ) are higher than DAMA.
  - Total background level is 2-3 times that of DAMA/LIBRA.
  - Higher light yield (15 P.E./keV) than DAMA/LIBRA (5-10 P.E./keV)
    - Can make the threshold lower easily
  - Each crystal is encapsulated in copper and quartz windows.
  - Two 3-inch PMTs (R12669SEL) are attached to each crystal.
    - Quantum efficiency: 35% @ 420 nm

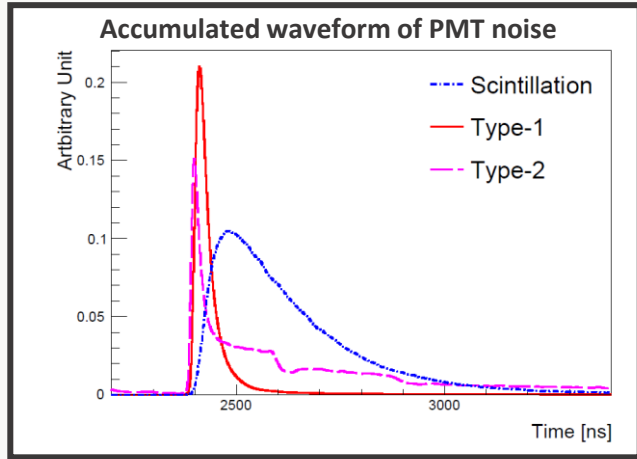


# COSINE-100 experiment-Shielding

- **Active veto**
  - **Liquid scintillator (LS)**
    - 2200-L LAB-based LS
    - 5-inch PMT(R877) for LS detector
  - **4 $\pi$  Muon counter**
    - 37 plastic scintillator panels
    - 2-inch PMT(H7195) for muon counter
- **Passive veto**
  - 3-cm thick copper box
  - 20-cm thick lead castle
- **Neutron monitoring**
  - Fast neutron detector (Liquid Scintillator)
  - Thermal neutron detector ( $^3\text{He}$  gas detector)



# Event Selection (SET2)

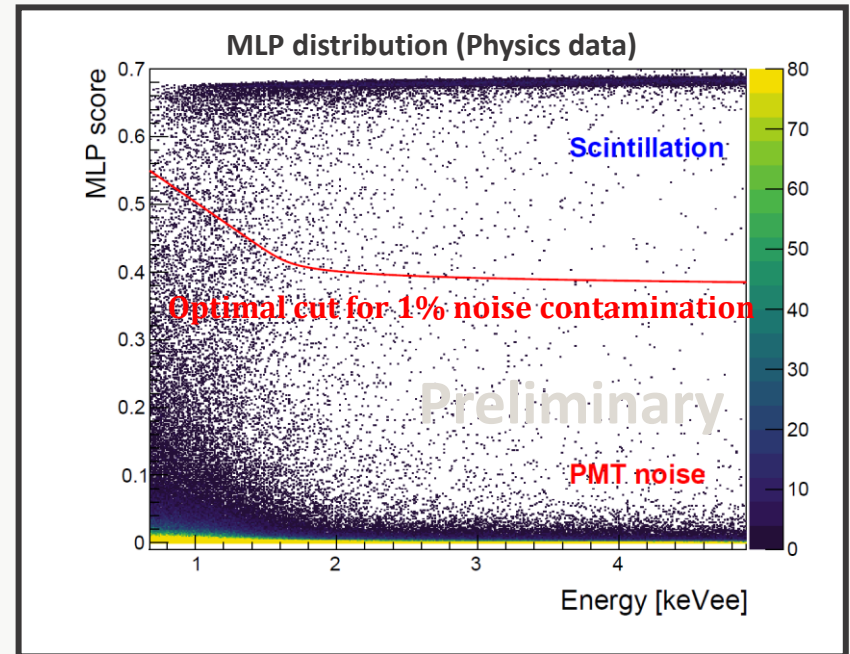
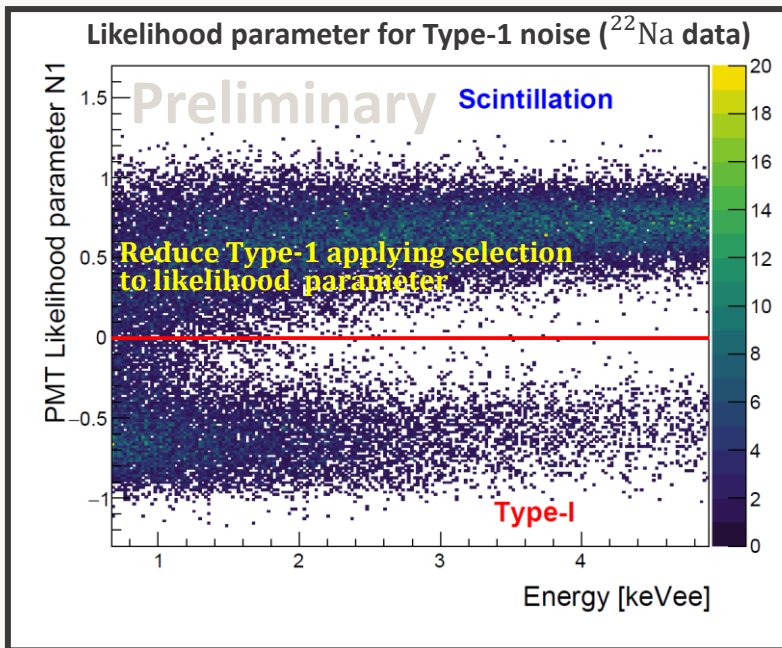


- Event selection updated since Astropart. Phys 130, 102581 (2021)
- To select **scintillation event** from **PMT-induced noise**
  - Use **pulse-shape parameters**
    - Meantime, likelihood parameters
  - + **Fast Fourier Transformation** parameter (for Set3)

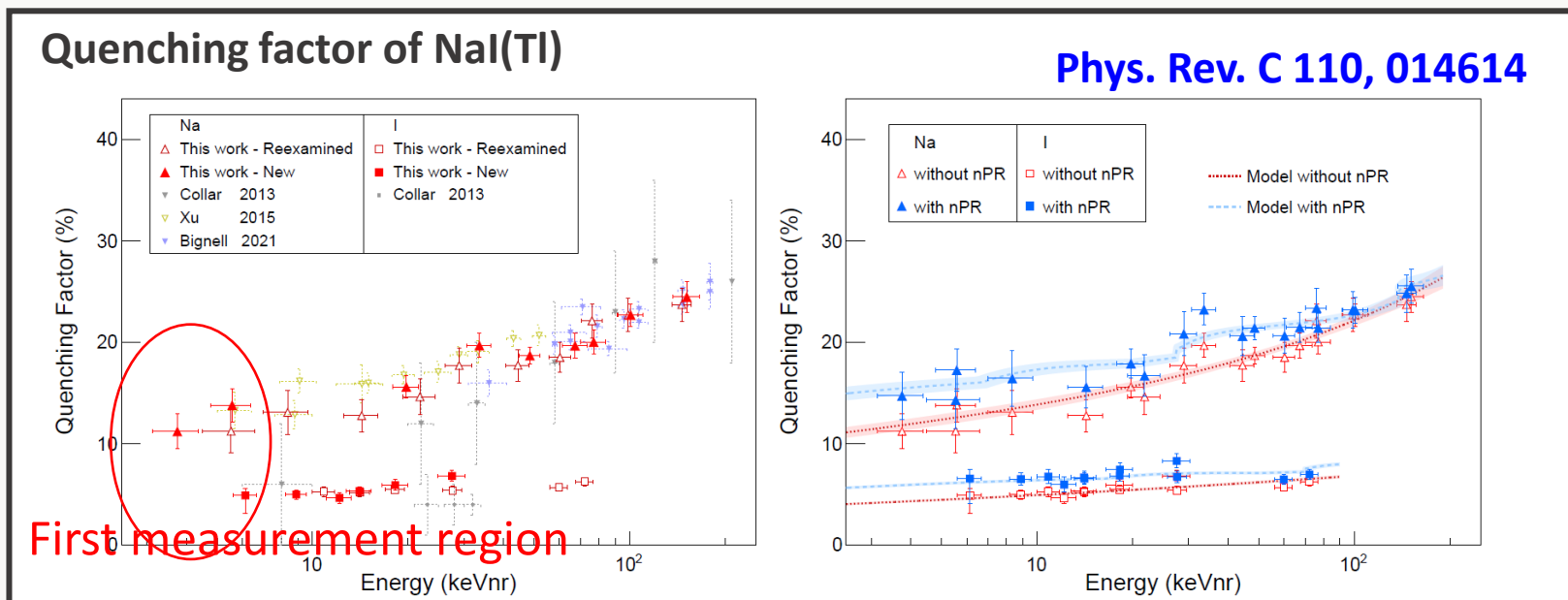
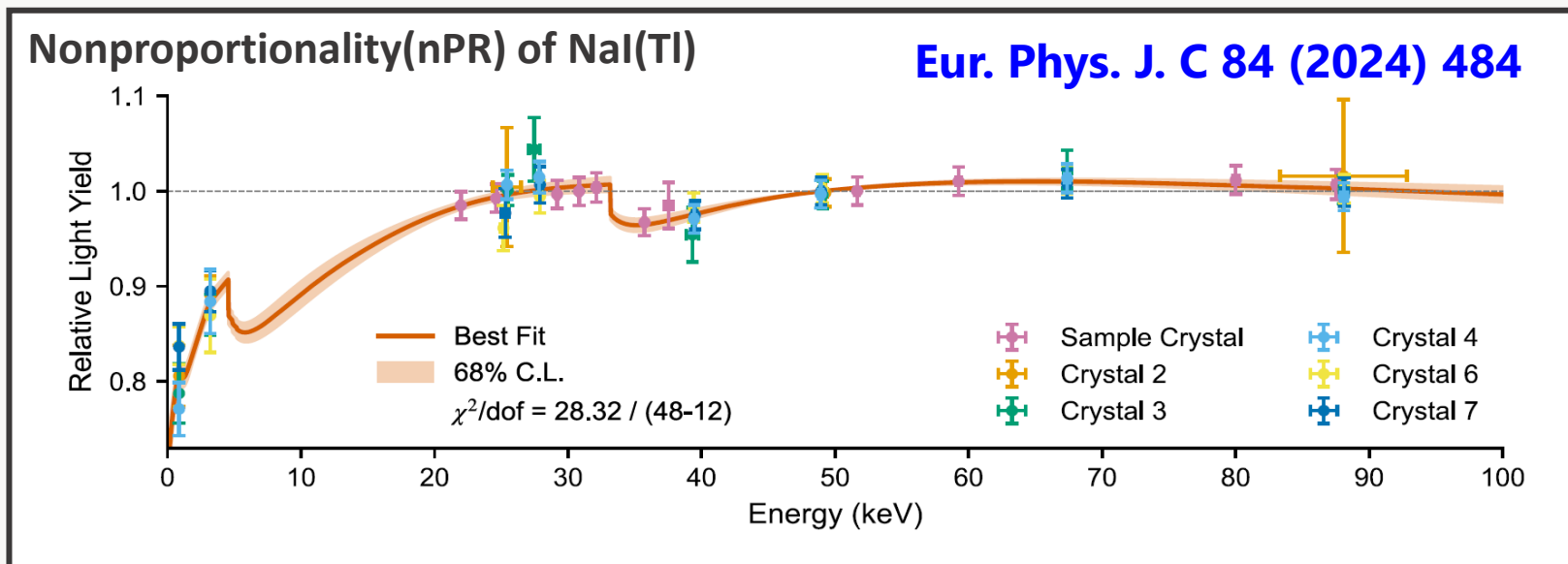
• Multivariable analysis

- **BDT(SET2)**  $\rightarrow$  **MLP(SET3)**  
Boosted Decision Tree      Multilayer perceptron
- 1 keV  $\rightarrow$  **0.75 keV**

$$F_{\mathcal{L}}^{N1} = \frac{\ln \mathcal{L}_{\text{FFT}}^{N1} - \ln \mathcal{L}_{\text{FFT}}^S}{\ln \mathcal{L}_{\text{FFT}}^{N1} + \ln \mathcal{L}_{\text{FFT}}^S}$$



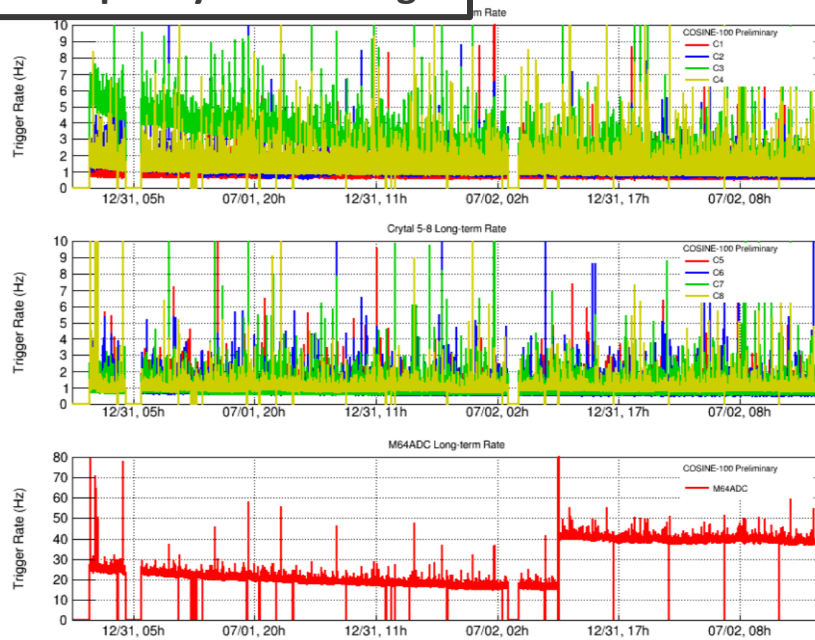
# Update Nonproportionality and Quenching factor of NaI(Tl)



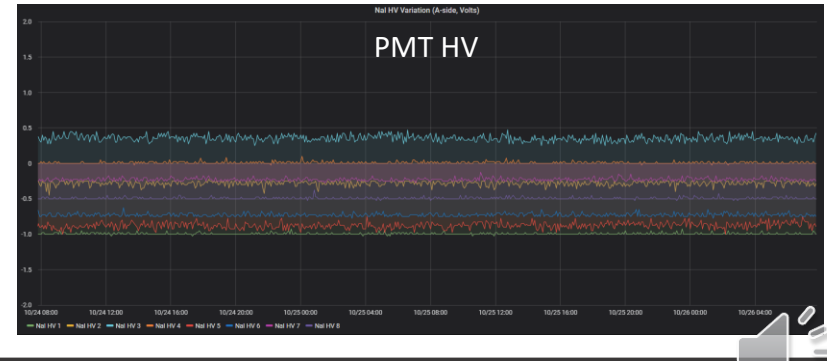
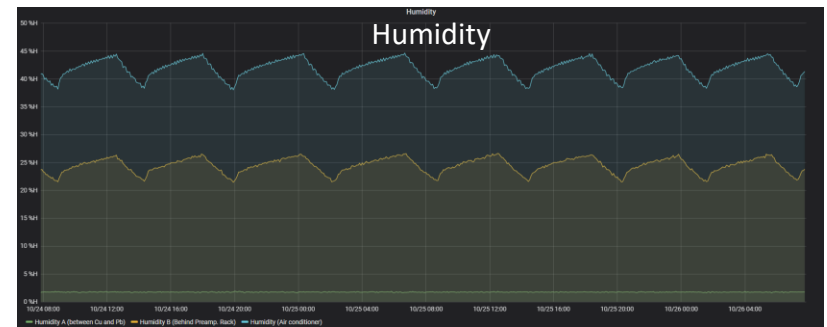
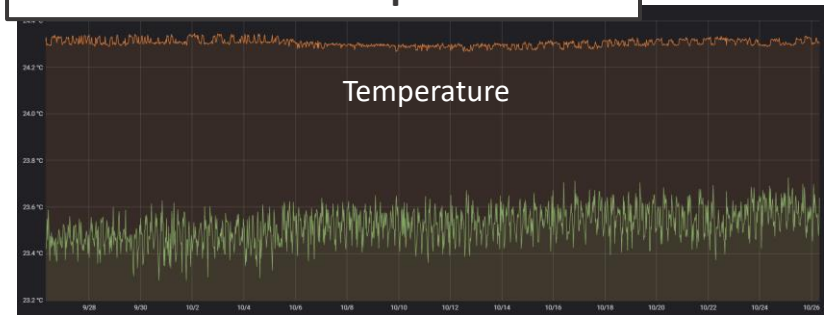
# COSINE-100 operation **cont'd**

- Slow monitoring system

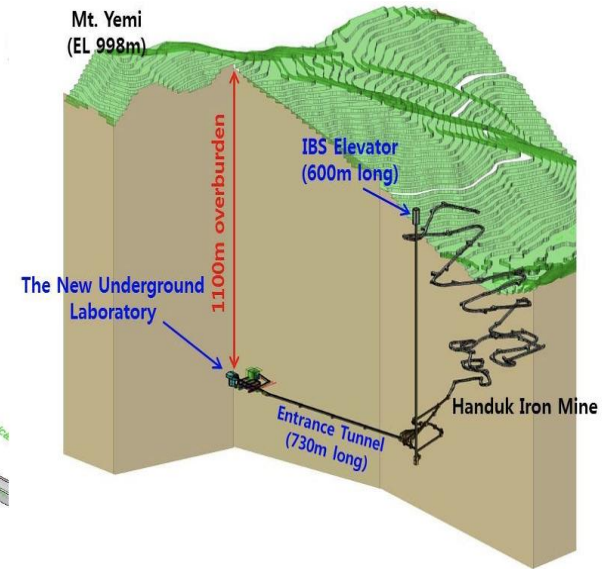
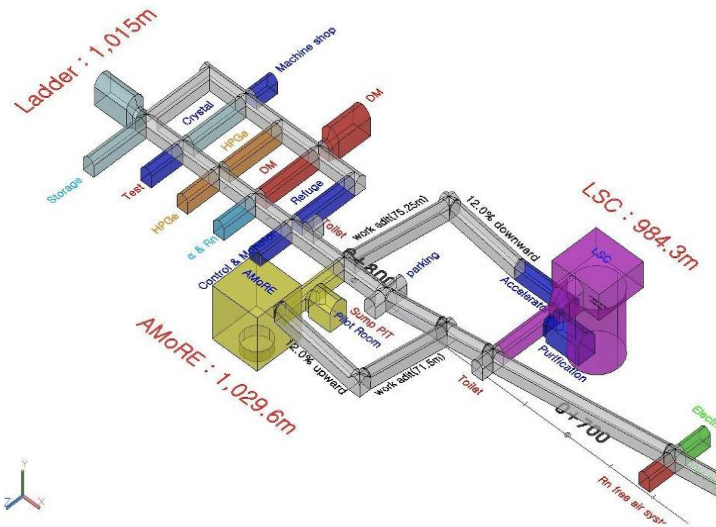
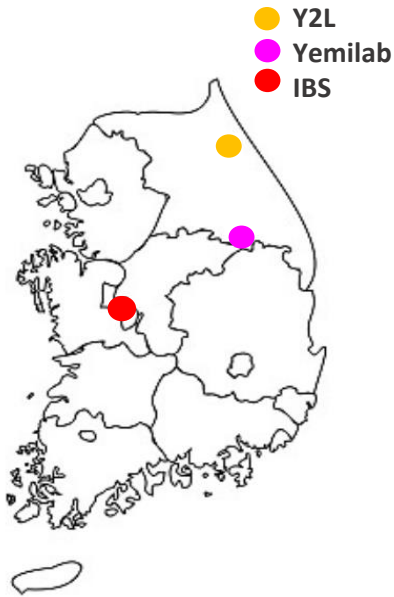
## Data quality monitoring



## >200 Environmental parameters



# Yemilab @ Jeongseon

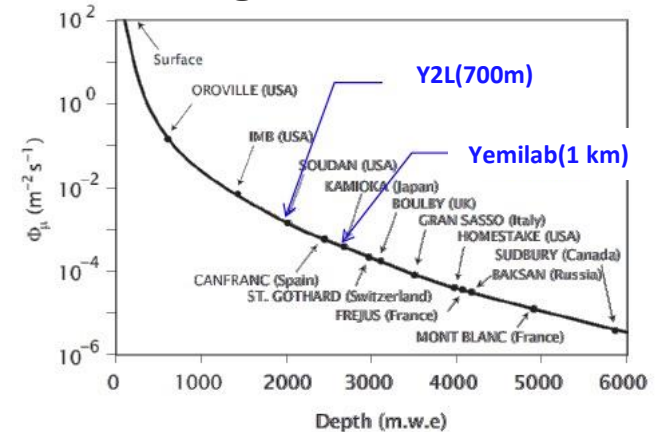


- **Newly constructed underground laboratory at Jeongseon in 2022**

- 1000 m rock overburden
- 3000 m<sup>3</sup> experimental area

Front. Phys., 02 April 2024

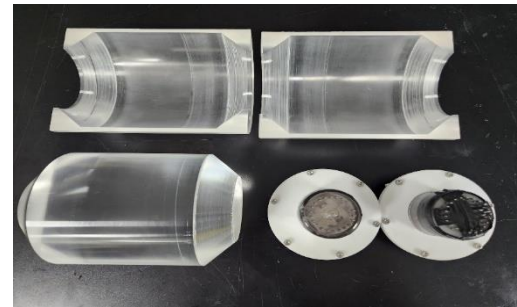
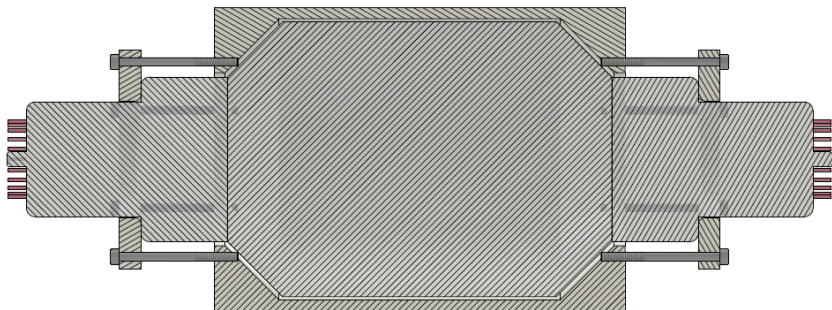
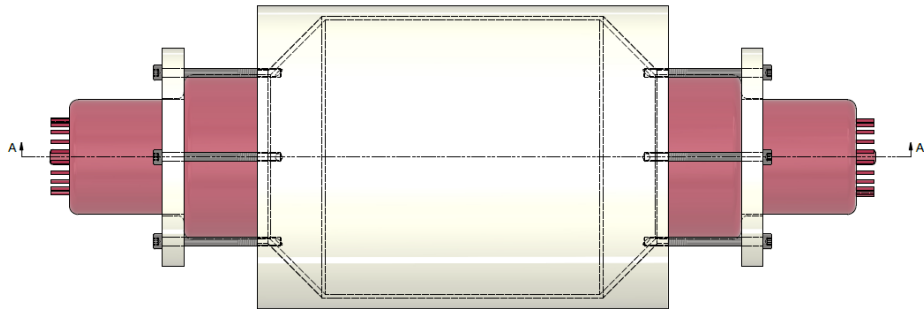
## Underground laboratories





# Encapsulation

- **Inner** structure for PMT-crystal connection
  - Direct contact PMT and crystal with optical pad
  - **PTFE** body, **Brass** bolt
  - PMT base shield (PTFE)

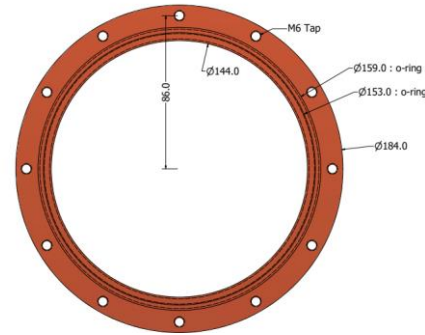
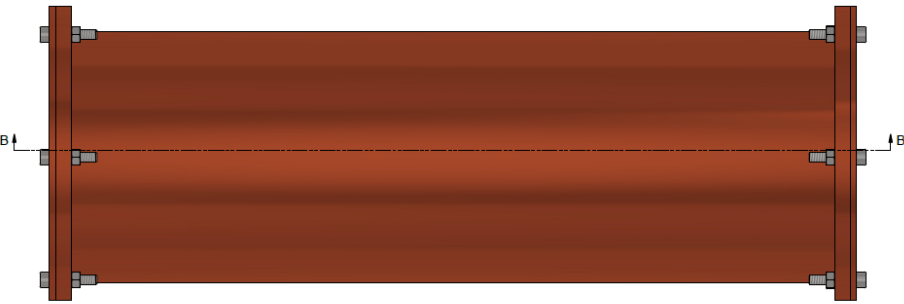


PMT base shield

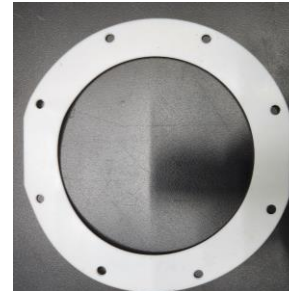


# Encapsulation cont'd

- **Copper case**
  - Same as **NEON** experiment
  - To prevent LS & air leak
    - **PTFE Gasket**
    - Cable gland



PTFE gasket



cable gland

