



Status of SABRE North at LNGS and radiopurity of SABRE NaI(Tl) crystals



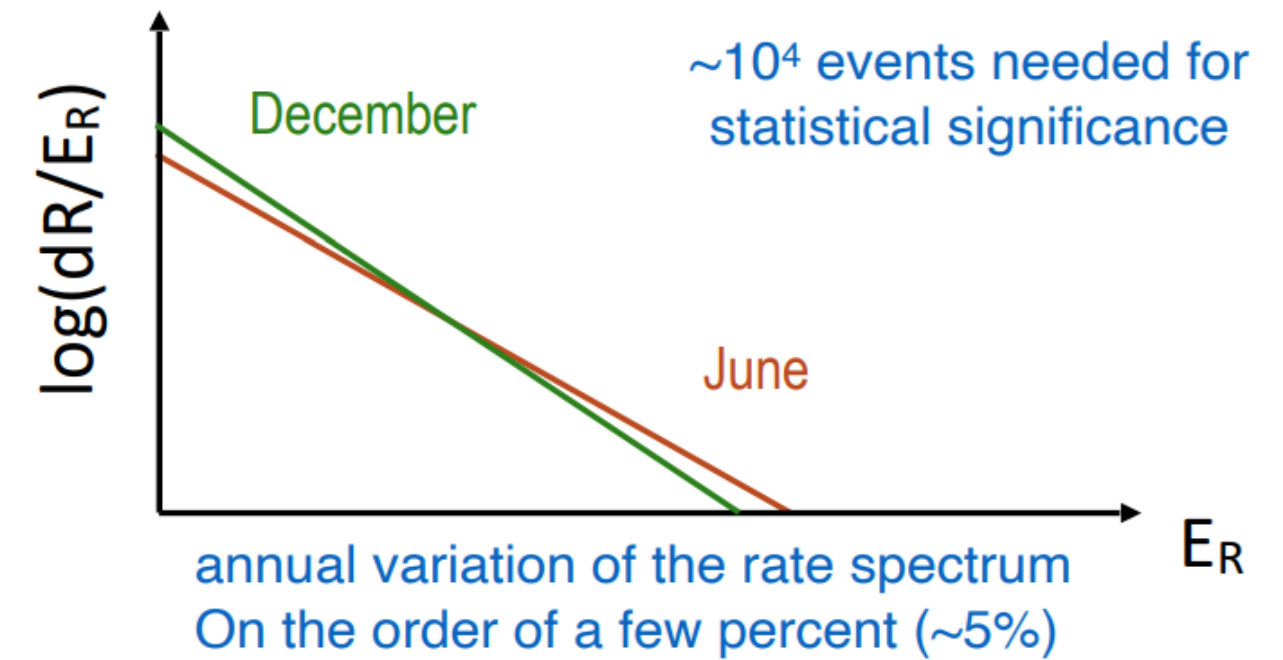
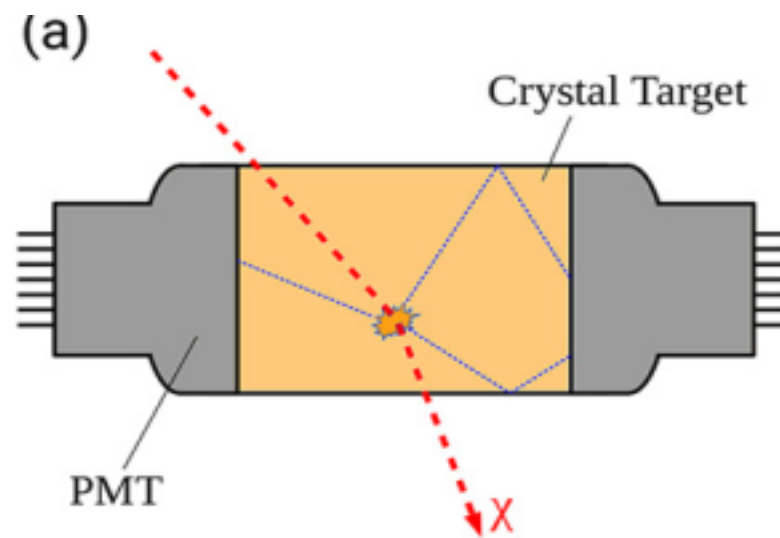
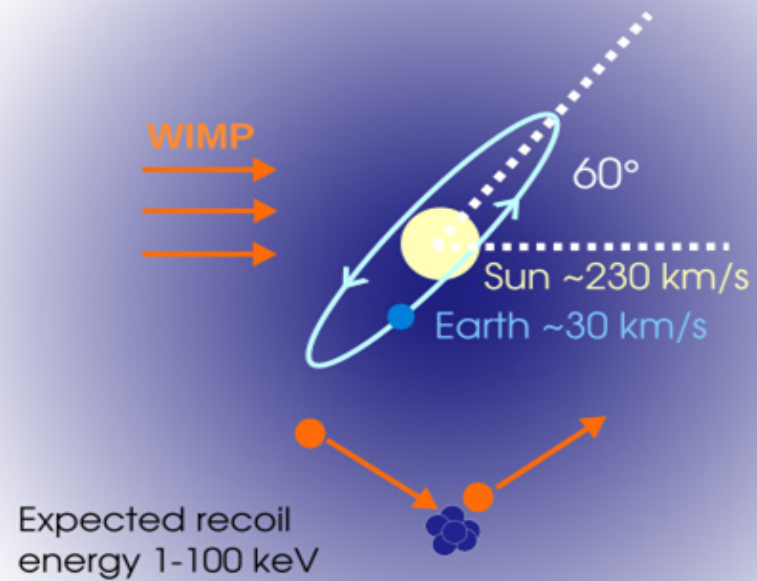
Gabriella Cataldi- INFN Lecce
on behalf of SABRE North Collaboration

ICHEP 2024 | PRAGUE

18 JUL 2024



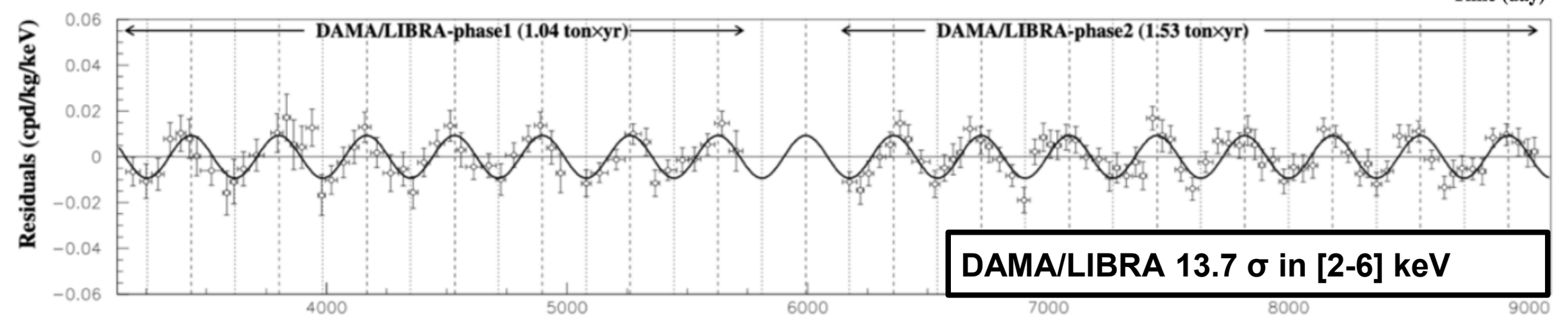
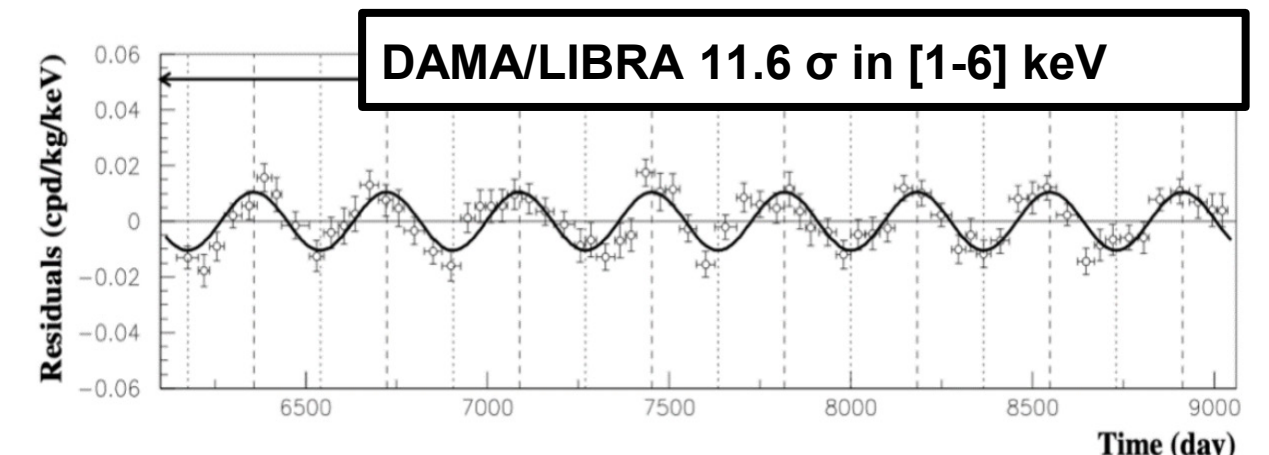
Dark Matter with annual modulation



- ✓ Expected rate in an Earth-based detector is modulated
- ✓ Small modulation fraction $S_m/S_0 = O(\sim \text{few } \%)$
- ✓ Region of interest [1-6] keV

Rate vs time

$$R = S_0 + S_m \cos\left(\frac{2\pi}{T}(t - t_0)\right)$$



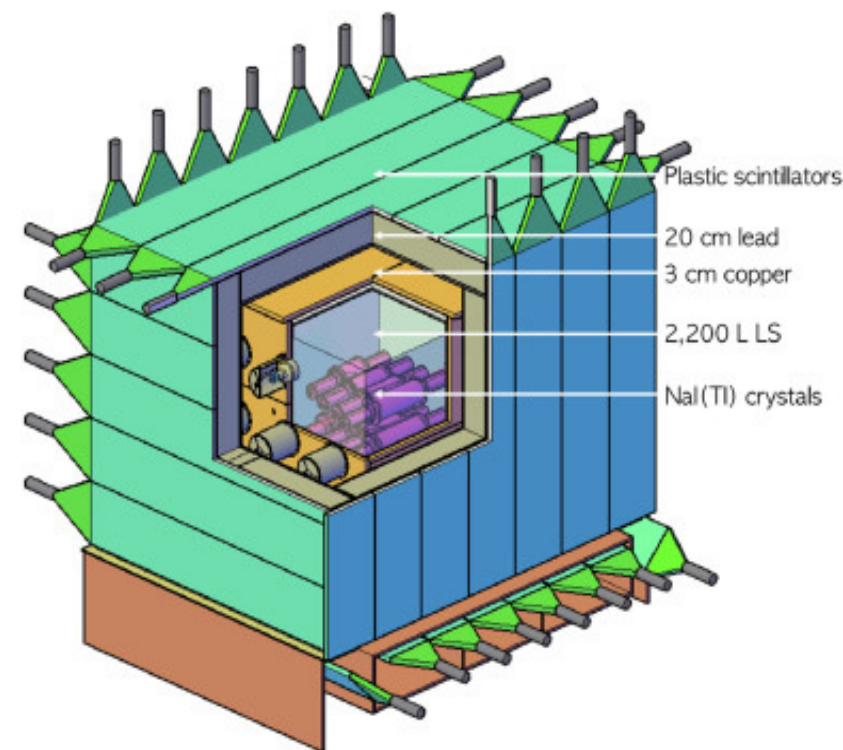
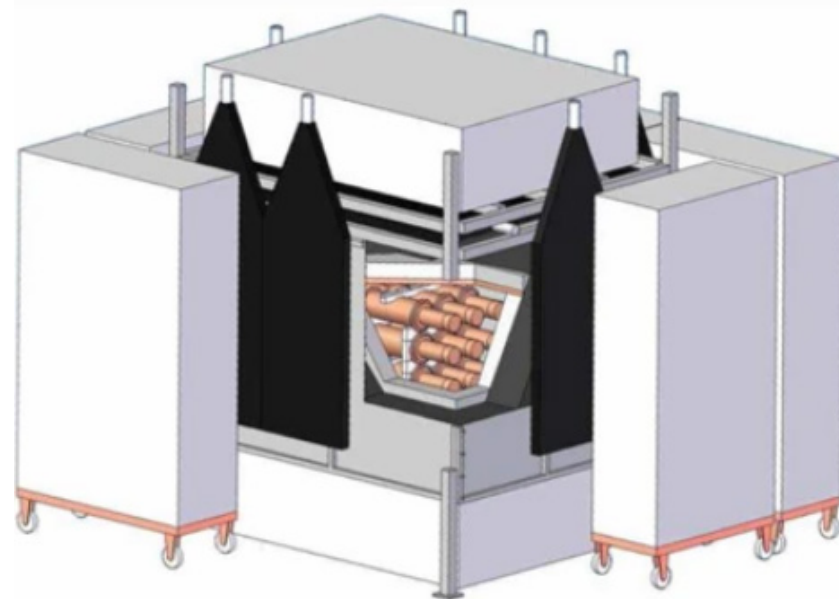
Nucl. Phys. At. Energy 22 (2021) 329-342

Nal(Tl) experimental landscape



ANAIS112
@Canfranc, Spain

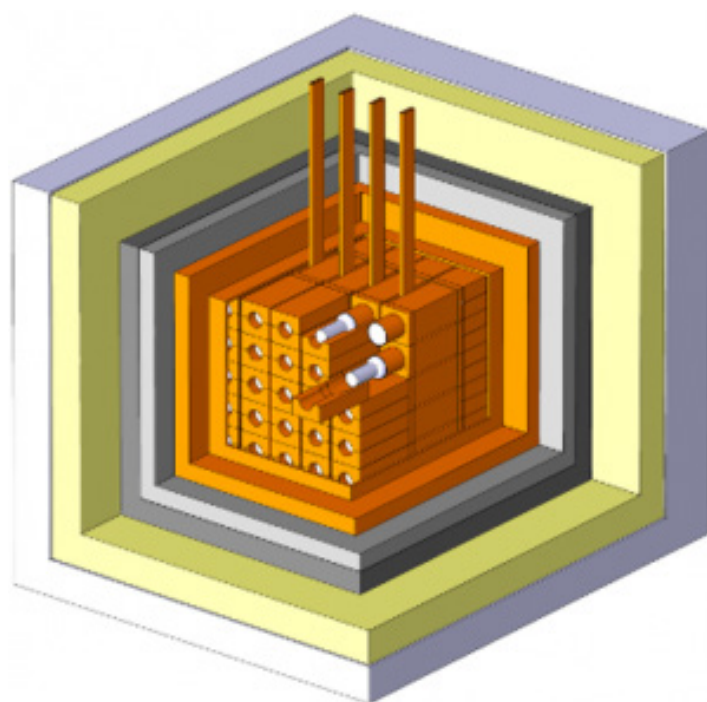
M. Martinez @this session.



COSINE-100
@Yang Yang, South Korea

InSoo Lee @this session.

DAMA/LIBRA @LNGS, Italy



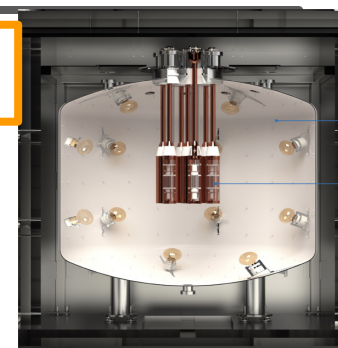
Ultra radio-pure crystals



SABRE South: L. Milligan@this session

@ SABRE South only

Active veto with liquid scintillator

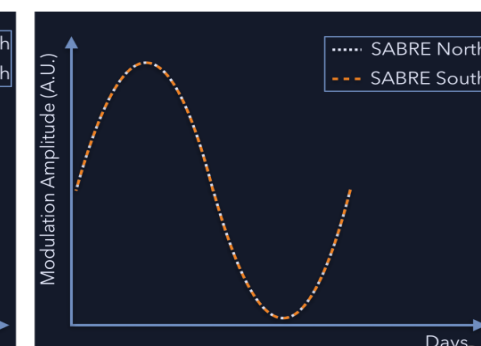
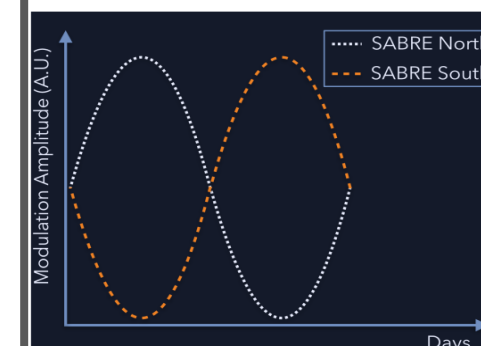


Double location:

Northern and Southern hemisphere

Seasonal effect

Dark Matter



SABRE North and South



- SABRE North at Laboratori Nazionali del Gran Sasso (LNGS) in Italy
- SABRE South at Stawell Underground Physics Laboratory (SUPL) in Australia

The map highlights the locations of the SABRE experiments. A blue dot in Italy is labeled "LNGS, Italy" with a box. A blue dot in Australia is labeled "Stawell Underground Laboratory, Australia" with a box. The Australian box also features the "STAWELL UNDERGROUND PHYSICS LAB" logo. Various partner logos are displayed around the map, including Princeton University, RMD (A Dynasil Company), Mellen (Coast to coast and around the globe), INFN (Istituto Nazionale di Fisica Nucleare), Università del Salento, Sapienza Università di Roma, Università degli Studi di Milano, Swinburne University of Technology, The University of Melbourne, The University of Sydney, Australian National University, The University of Adelaide, and KEK. A photograph of the Gran Sasso tunnel entrance is shown in the bottom left.

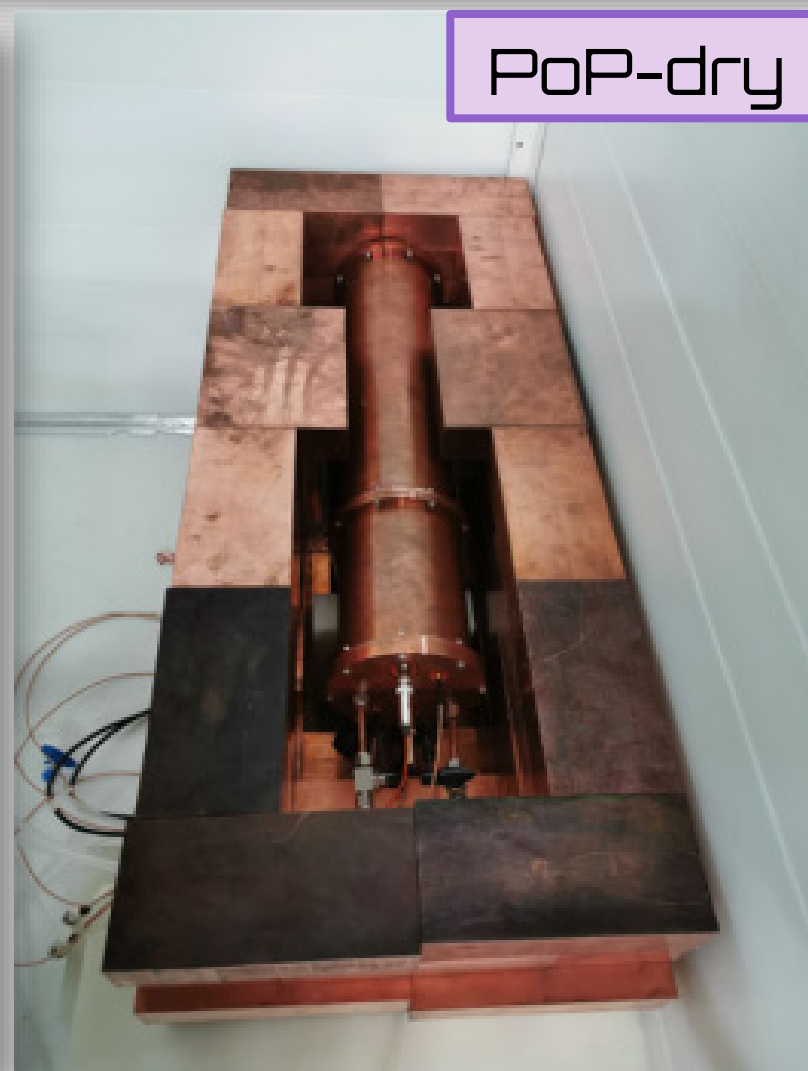
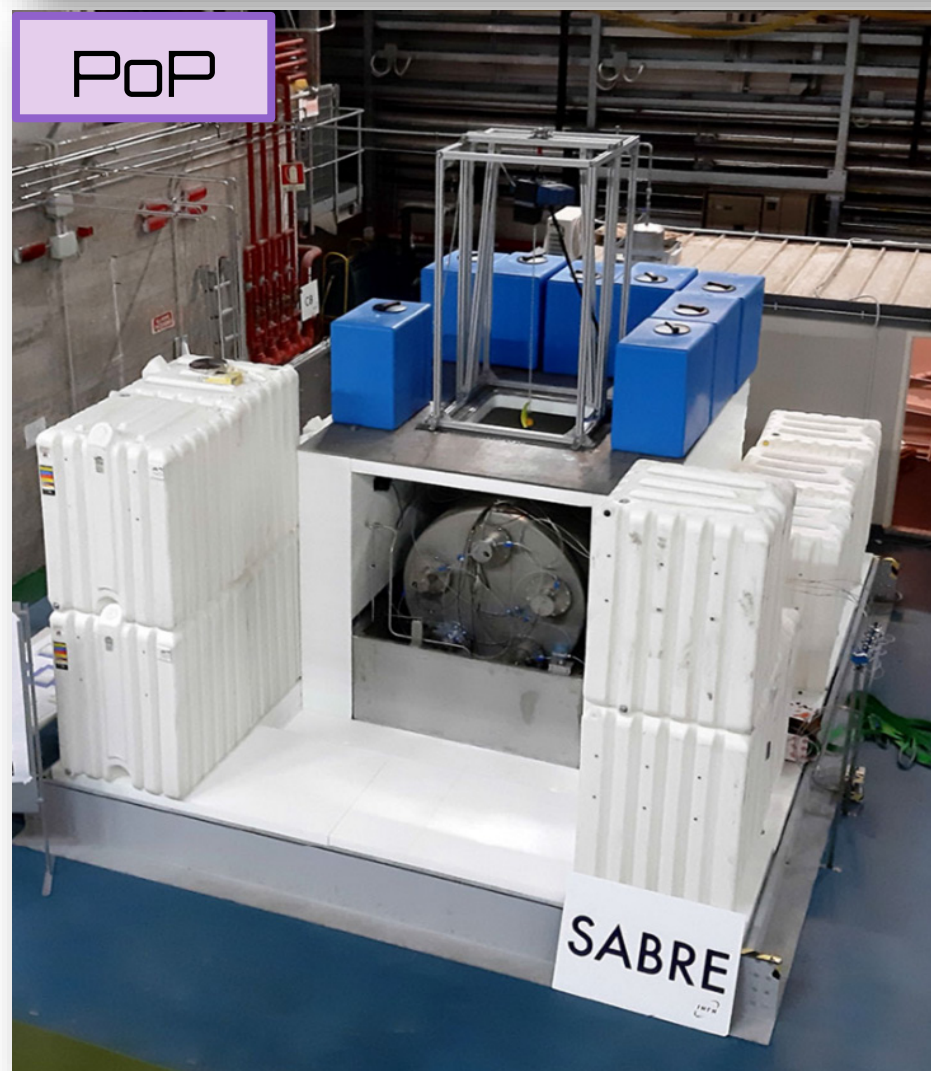
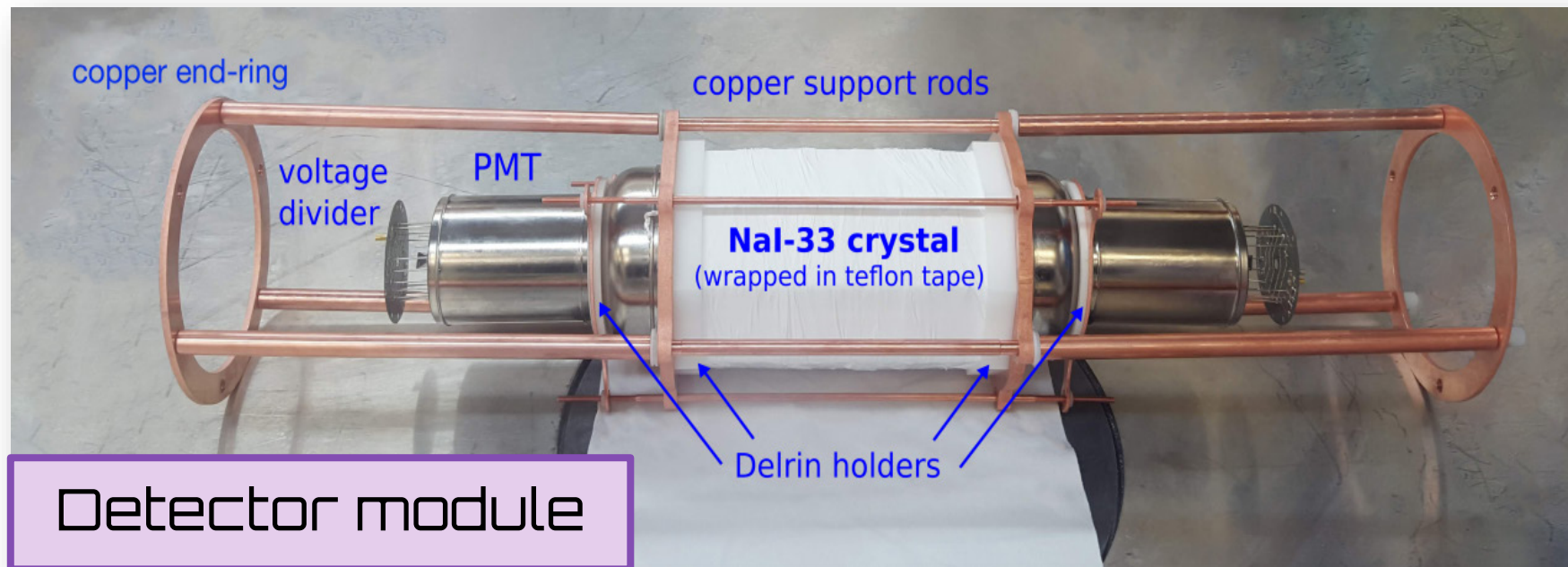


SABRE North and South detectors have **common core features**:

- Same crystal production and R&D.
- Same detector module concept (Ultra-pure crystals and HPK R11065 PMTs)
- Common simulation, DAQ and data processing frameworks
- Exchange of engineering know-how with official collaboration agreements between the ARC Centre of Excellence for Dark Matter and the INFN

SABRE North and South detectors have **different shielding designs**:

- ✓ SABRE North has opted for a fully passive shielding due to the phase out of organic scintillators at LNGS. Direct counting and simulations demonstrate that this is compliant with the background goal of SABRE North at LNGS.
- ✓ SABRE South will be the first experiment in SUPL, the liquid scintillator will be used for in-situ evaluation and validation of the background in addition to background rejection and particle identification.

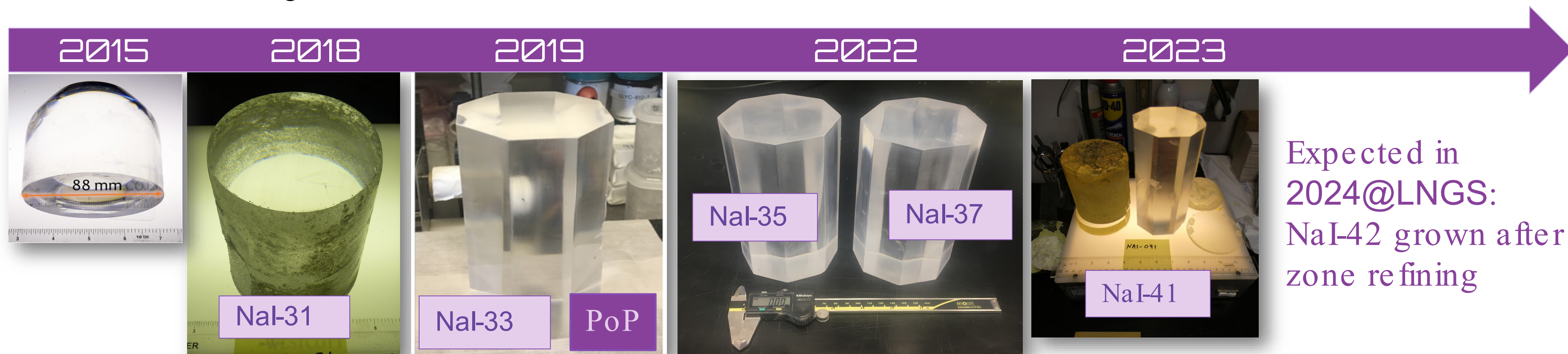


- ✓ Run in 2020 with Borexino liquid scintillator and NaI-33
 - ❖ 2 tons active veto with 10 8-inch PMTs + H₂O shielding
- ✓ Exploited successfully ⁴⁰K tagging with sensitivity at the level of 1 ppb
- ✓ Demonstration by direct counting of first crystal production after DAMA/LIBRA with **background in [1,6] keV of order 1 cpd/kg/keV**
- ✓ **PoP-dry run in 2021-2022:** passive shielding with additional layer of copper
 - ❖ confirmed background level

The SABRE crystals R&D



- ✓ R&D carried out by PU, INFN and ARC Centre of Excellence for DM
- ✓ Radioclean NaI powder *Astrograde* by Sigma Aldrich now Merck, Germany
- ✓ Crystals grown by RMD - Radiation Monitoring Devices, MA (USA)
 - ✓ Vertical Bridgman method in fused silica vessels

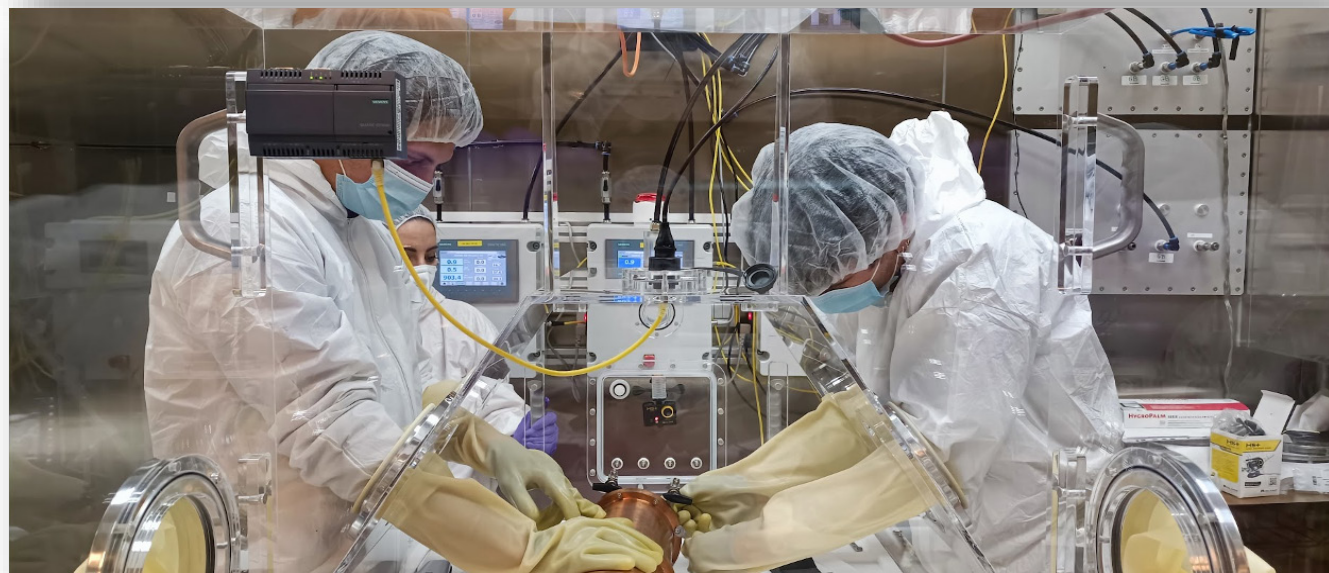
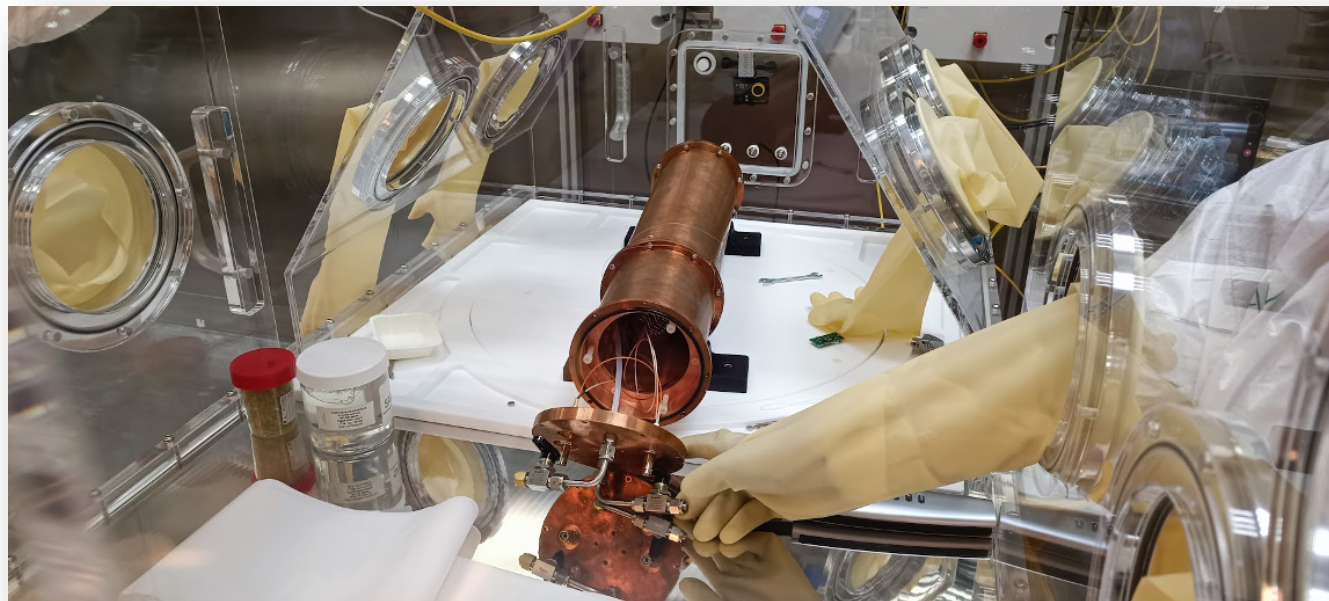


- NaI-33: background ~ 1 cpd/kg/keV \rightarrow close to DAMA/LIBRA Phase 1
- NaI-35, NaI-37: reproducibility within factor 2
- NaI-41: grown from chunks rather than powder \rightarrow demonstrated same optical quality

Crystal Operations in glovebox 2022-2023



- ❖ 27/09/2022 change of teflon reflector in Nal-33
- ❖ 29/11/2022 change of teflon reflector in Nal-33
- ❖ 7/12/2022 first assembly of Nal-37
- ❖ 24/01/2023 second assembly of Nal-37



All operations successful and moisture level in the glove-box kept always below 5% RH

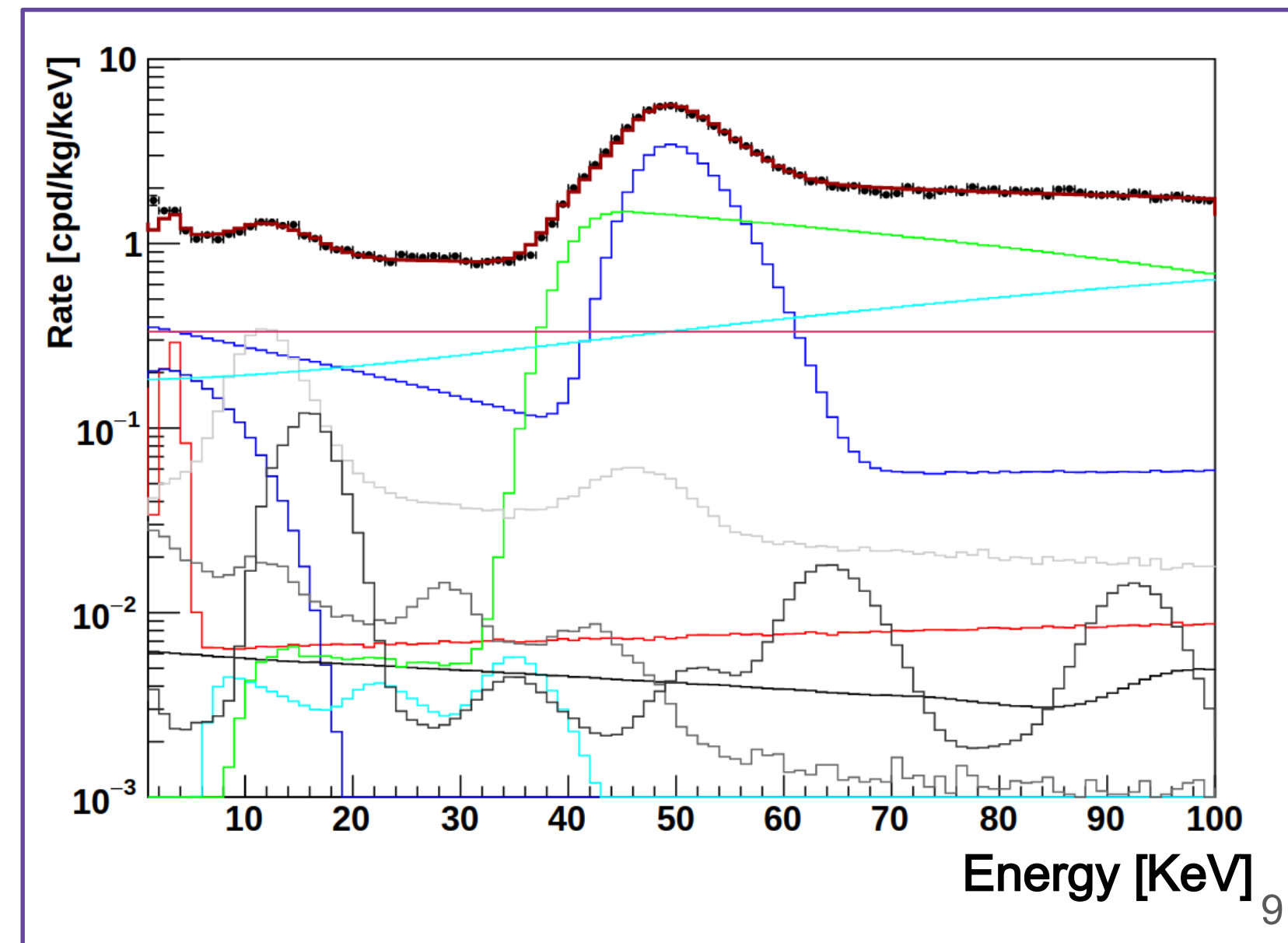


SABRE Background model (NaI-33)



- ❖ Background model updated since [Eur. Phys. J. C \(2022\) 82:1158](#)
- ❖ Background from reflector is not dominant (now constrained from direct measurements)
- ❖ Dominant backgrounds: ^{210}Pb in crystal bulk and external background

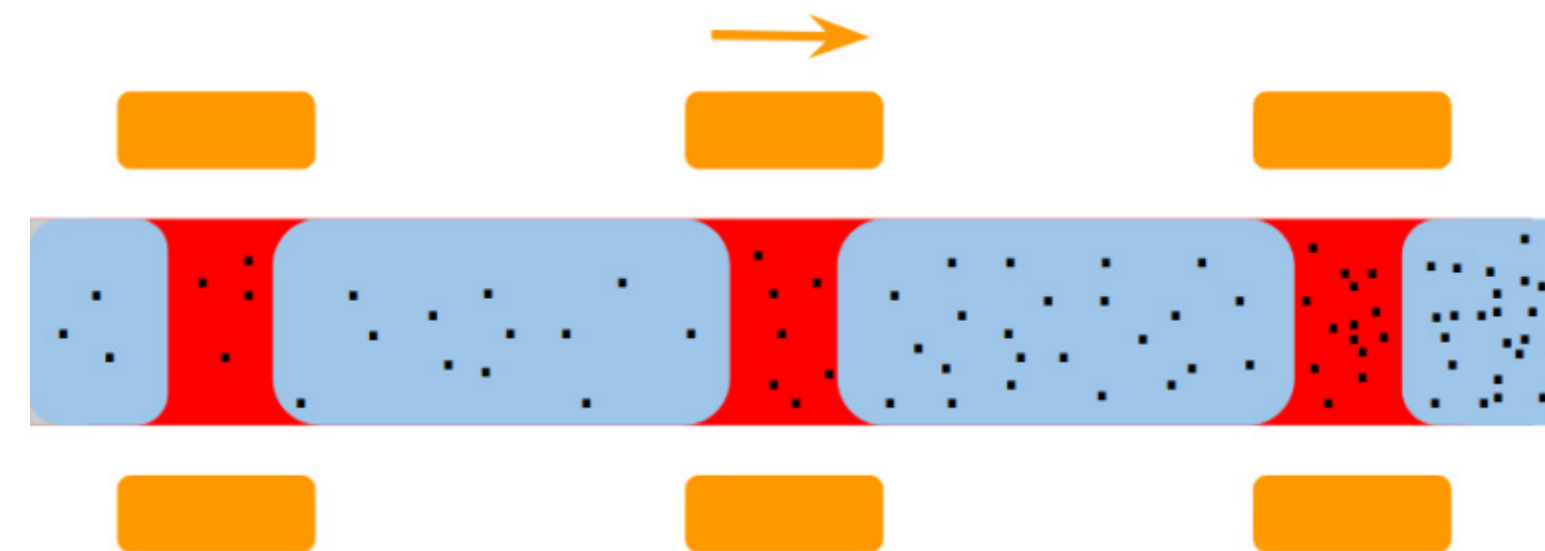
Source	Rate in ROI [1,6] keV [cpd/kg/keV]	Activity from fit
40K	0.125	0.16 ± 0.01 mBq/kg
^{210}Pb bulk	0.333	0.49 ± 0.05 mBq/kg
^{210}Pb reflector bulk	0.054	11 ± 1 mBq/kgPTFE
^{210}Pb reflector surface	0.023	< 0.6 mBq/m ²
3H	0.198	24 ± 2 mBq/kg
^{129}I	0.0003	1.03 ± 0.05 mBq/kg
^{238}U	0.006	5.9 ± 0.6 mBq/kg
^{232}Th	0.0003	1.6 ± 0.3 mBq/kg
PMT	0.003	1.9 ± 0.4 mBq/PMT
External	0.185	0.89 ± 0.05 relative unit to reference spectrum
Other b's	0.333	297 ± 15 counts
TOTAL	1.26 ± 0.27	



Zone refining technique



- ✓ Zone refining technique successfully used in semiconductor industry
- ✓ Impurities are segregated to one side of the ingot moving the ovens
- ✓ Tested on NaI Astro grade powder by Princeton group at Mellen company, Concord, NH (USA)



Zone refining could reduce to about 1/3 the Pb content, almost 1 order of magnitude K and possibly other internal contaminants like Rb.

Isotope	Impurity concentration (ppb)					
	Powder	S_1	S_2	S_3	S_4	S_5
^{39}K	7.5	< 0.8	< 0.8	1	16	460
^{85}Rb	< 0.2	< 0.2	< 0.2	< 0.2	< 0.2	0.7
^{208}Pb	1.0	0.4	0.4	< 0.4	0.5	0.5
^{24}Mg	14	10	8	6	7	140
^{133}Cs	44	0.3	0.2	0.5	3.3	760
^{138}Ba	9	0.1	0.2	1.4	19	330

[Phys. Rev. Applied 16, 014060 \(2021\)](#)

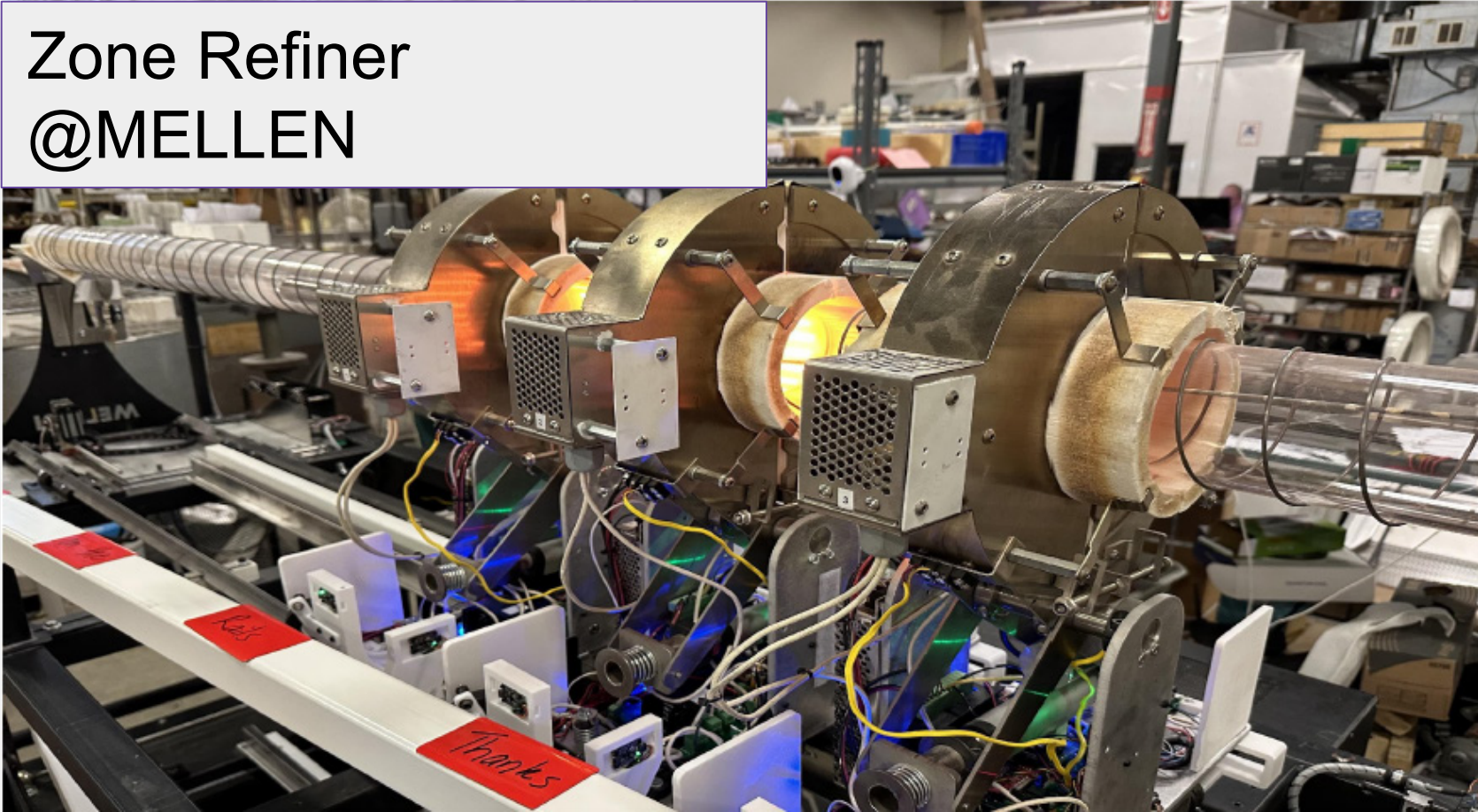
Zone refining activities 2023-2024



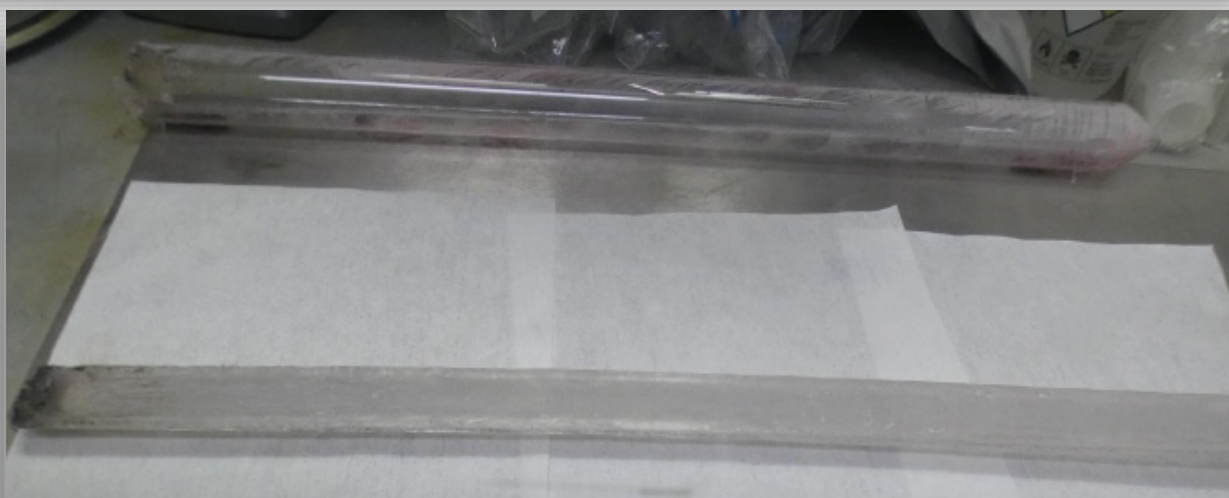
Four runs with 900 gr of AstroGrade NaI powder have been performed at MELLEN, NH, USA

- ❖ **RUN1:** Carbon coated ampoule
- ❖ **RUN2:** Carbon coated ampoule with increased number of passes
- ❖ **RUN3:** No coating + use of SiCl_4 to avoid sticking
- ❖ **RUN4:** No coating + use of SiCl_4
 - ❖ Ampoule sealed without gas inside
 - ❖ **Could be our preferred option**

For each run taken 5 samples from ingot of length equal to 60 cm taken and shipped to Canfranc Laboratory and Seastar for ICPMS measurements.

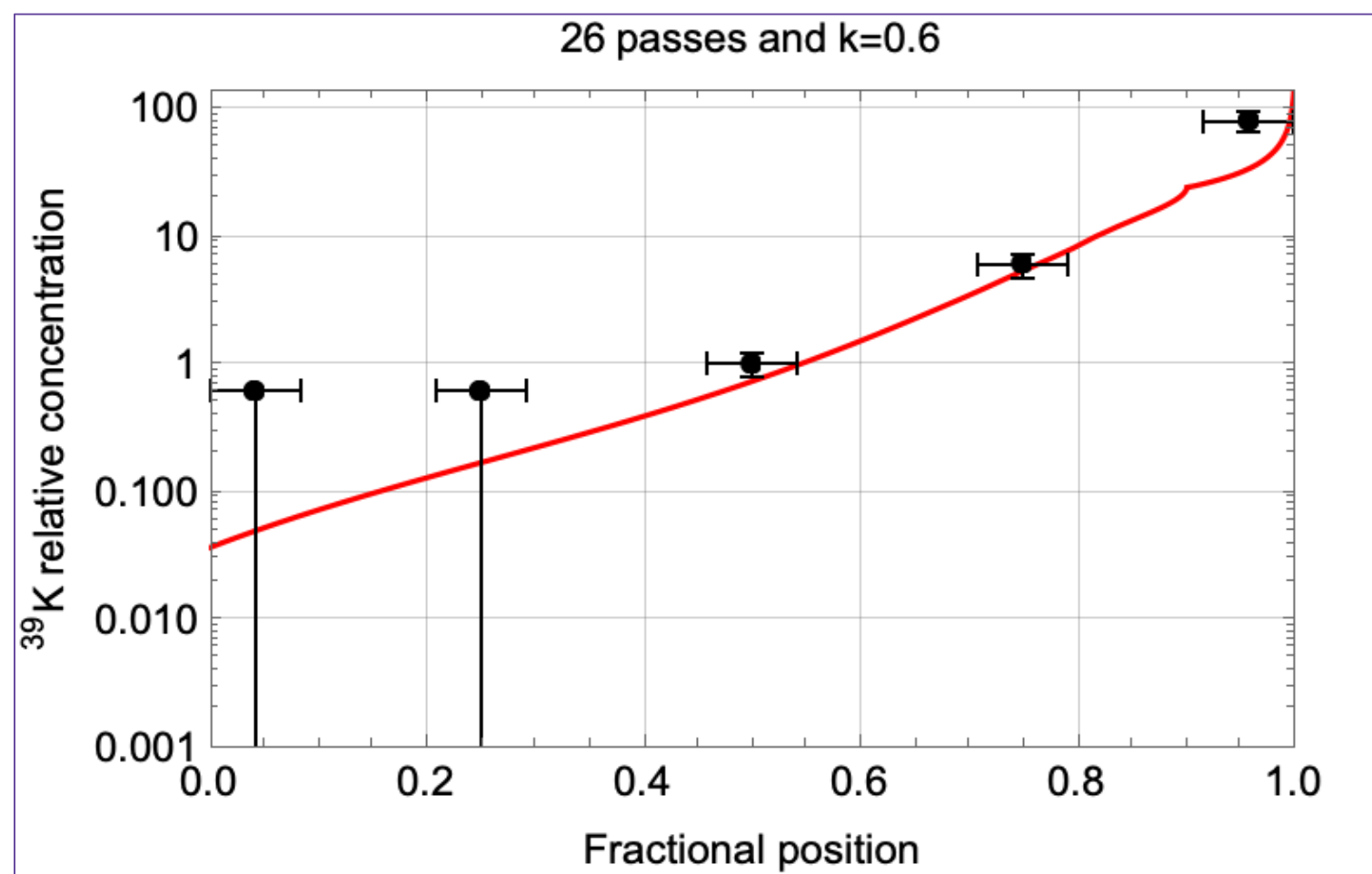


Carbon coated ampoule (RUN1)



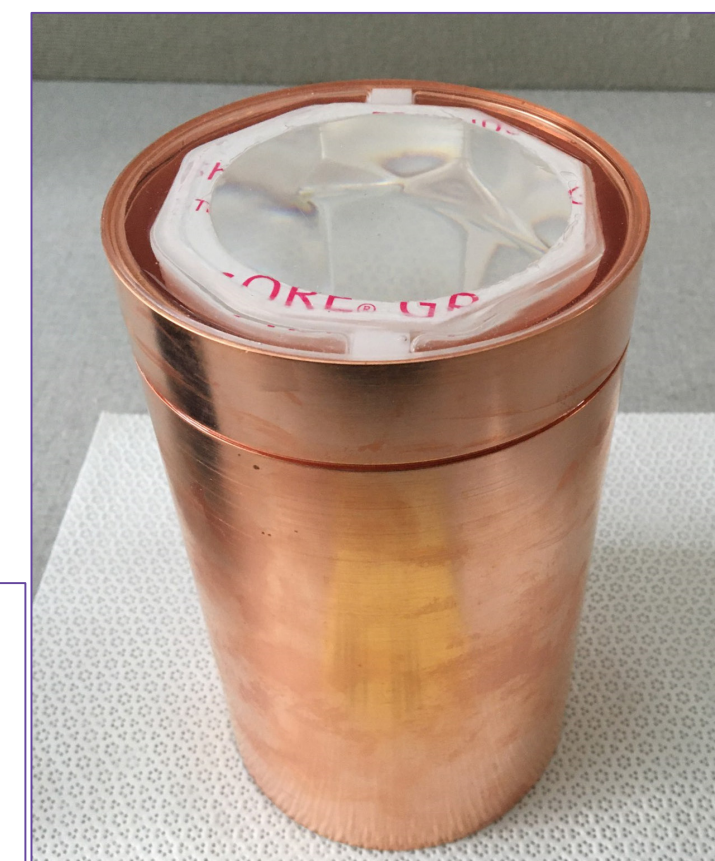
Ingot processed by Zone Refining (RUN3)

Zone refining preliminary results

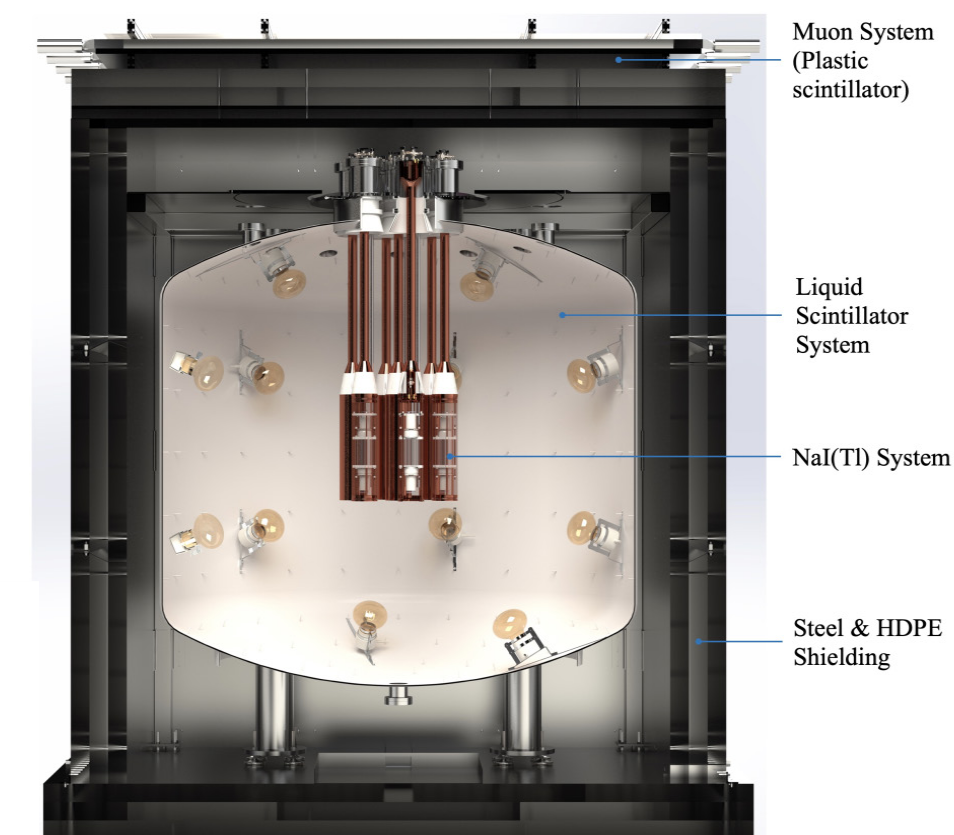
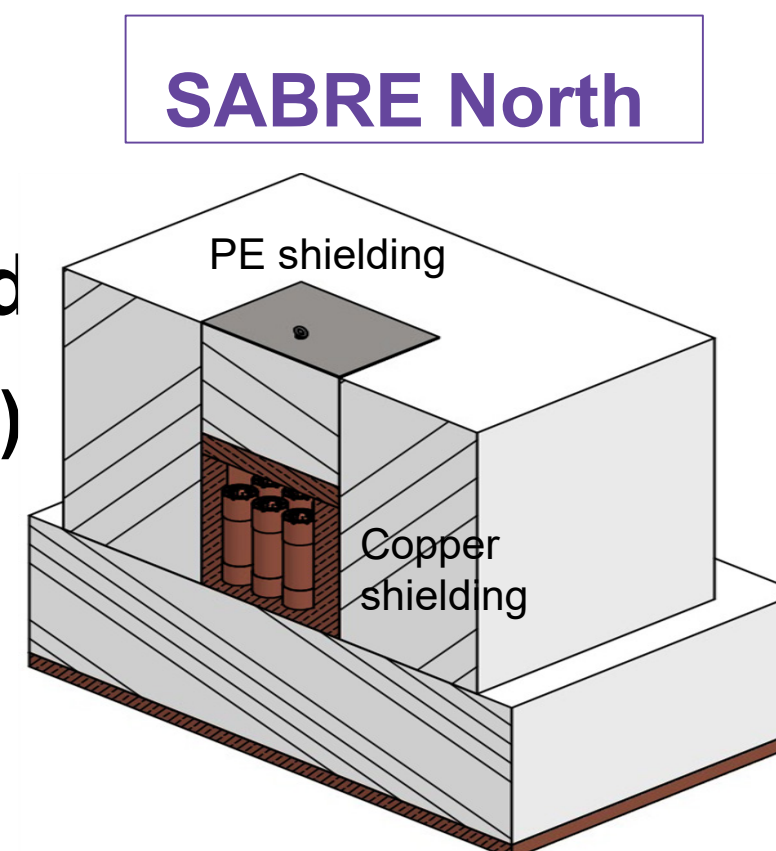


Comparison with the data from Run 1 with the model using a segregation factor of 0.6 for Potassium. The comparison is not based on a fit. Below the fractional position of 0.75, the average purification is approximately 87%, i.e. a contamination level of 10 ppb is reduced to about 1 ppb.

The successful growth of the NaI-41 crystal from chunks and its excellent optical properties represents an important step in our approach to producing high radiopurity crystals.

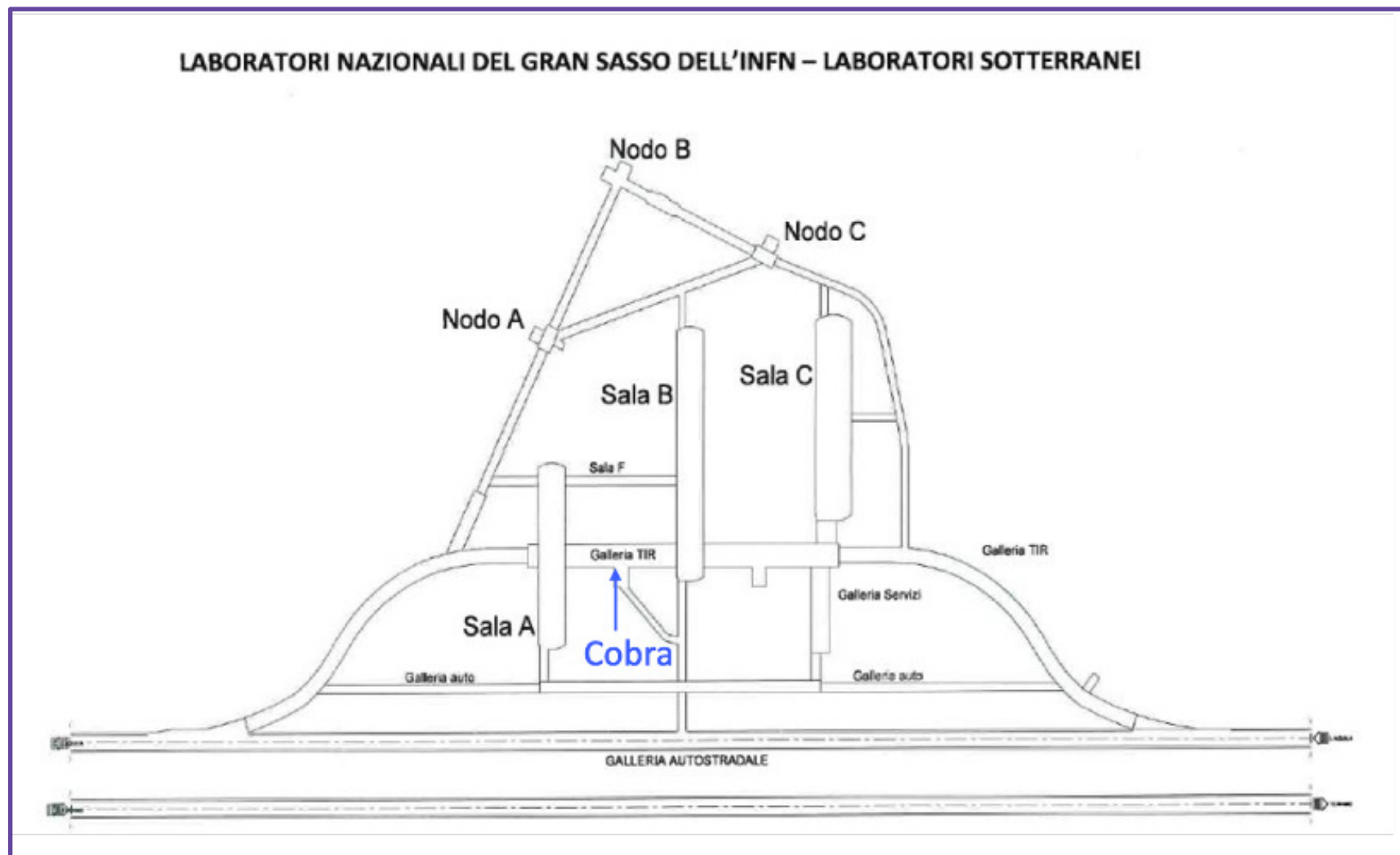


- **SABRE Proof-of-principle (PoP) and PoP-dry** achieved a background of **~ 1 cpd/kg/keV**
- Strategy to lower the **background**
 - **zone refining**
- For external background:
 - **→ SABRE North: improve passive shield**
 - **→ SABRE South: Liquid Scintillator (LAB) + Muon Veto**



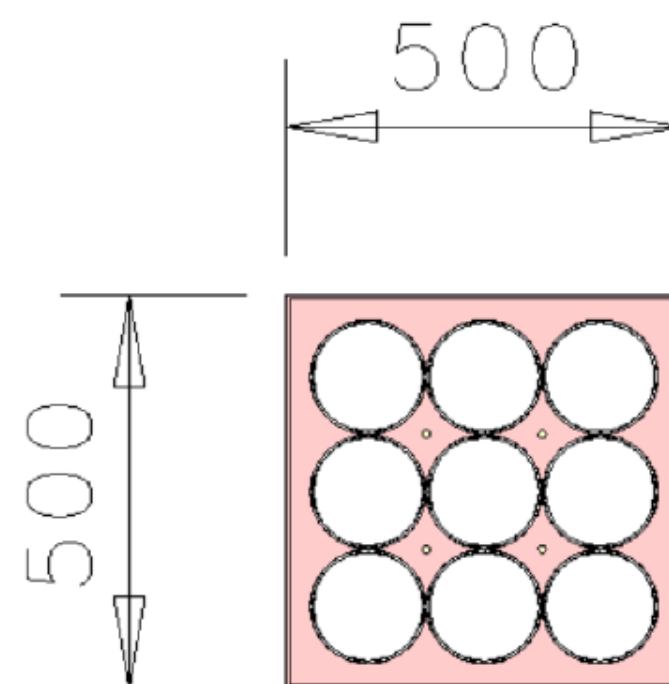
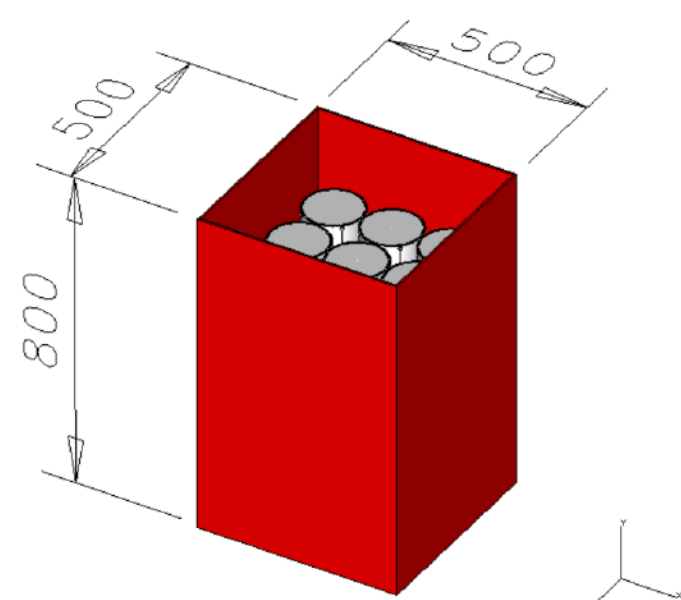
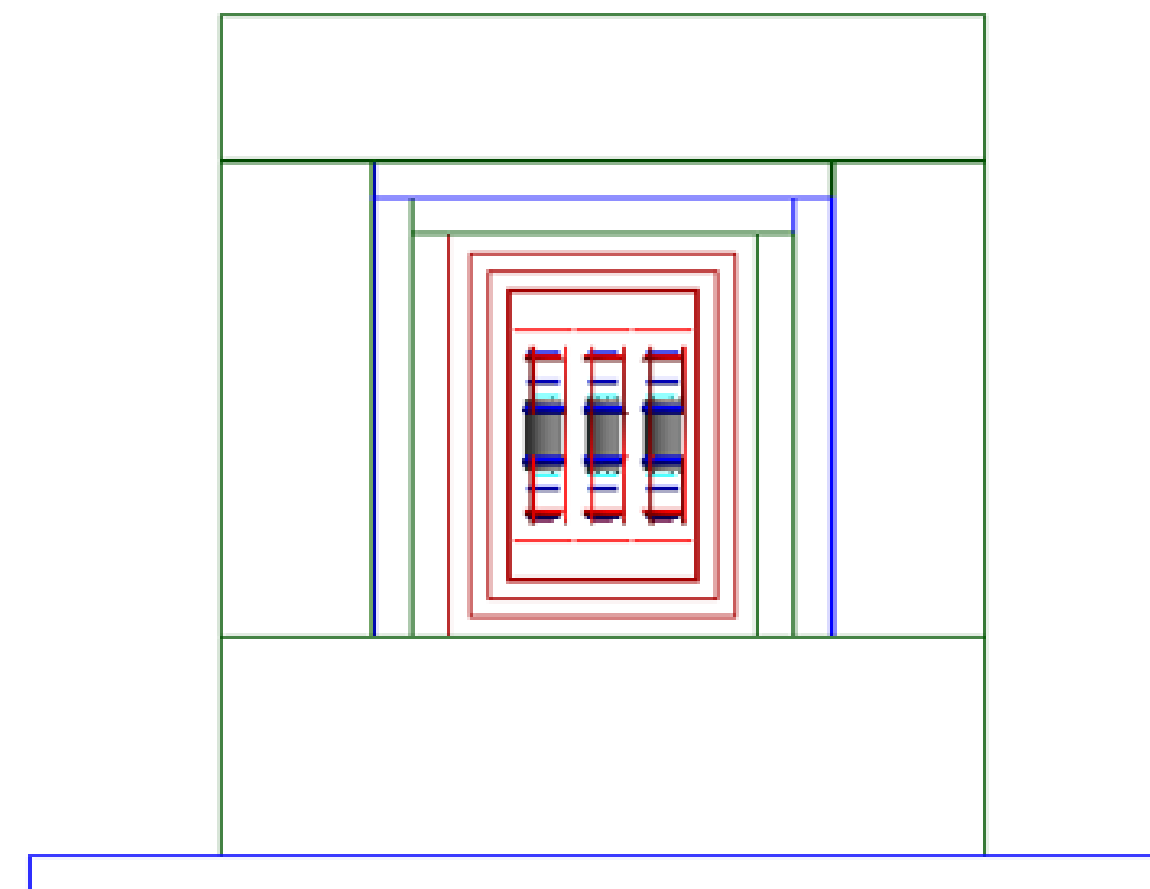
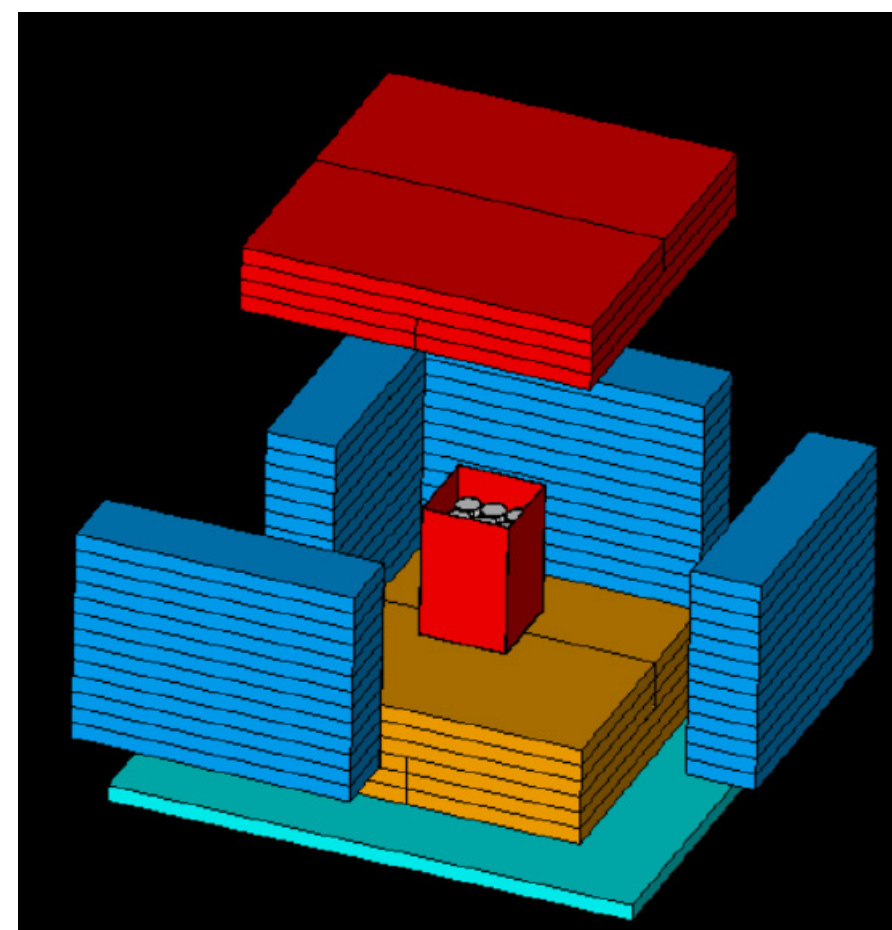
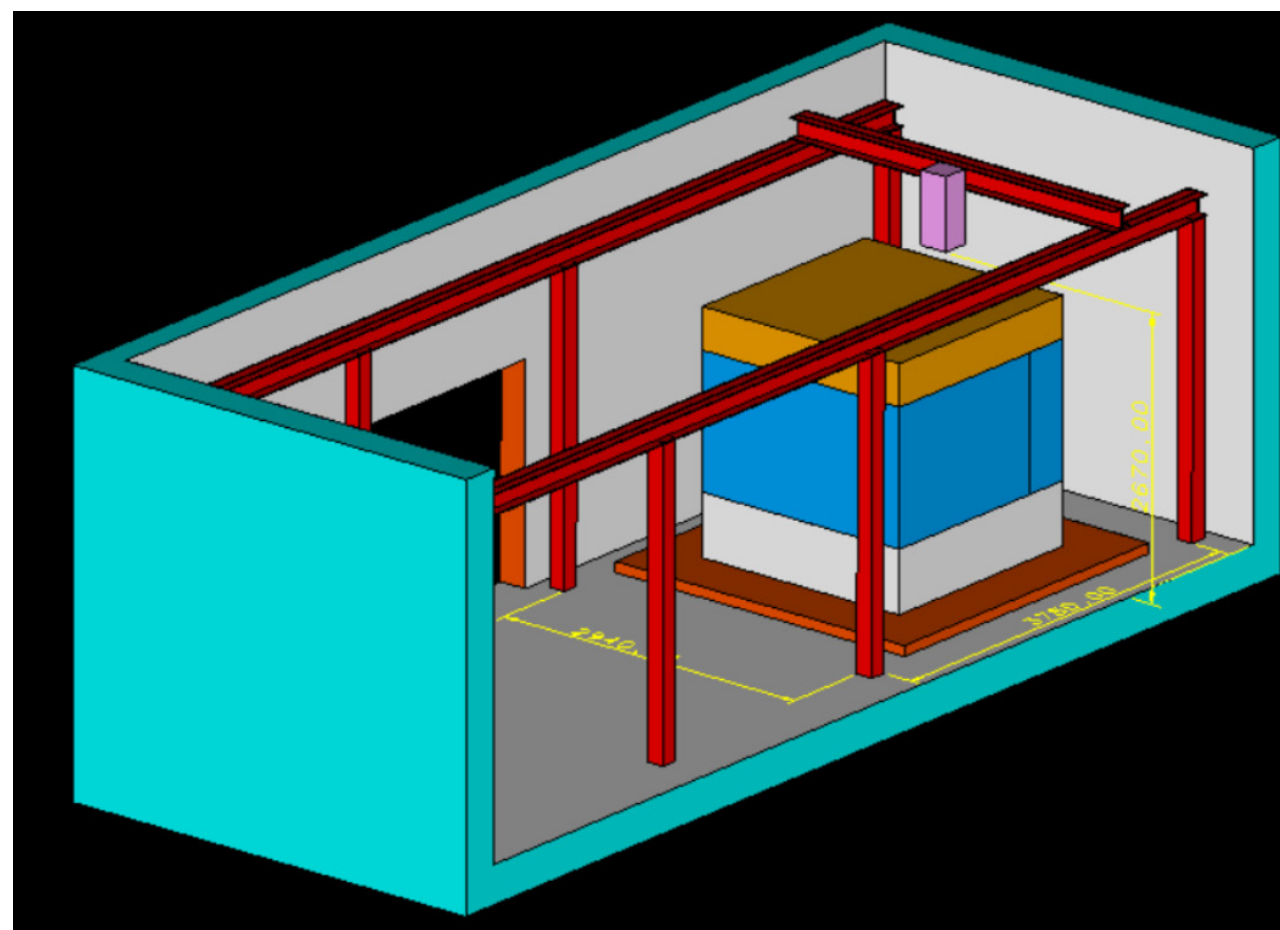
SABRE South

New SABRE experimental area is in the corridor between Hall B (Sala B) and Hall A (SALA A), in the so called «Cobra area».



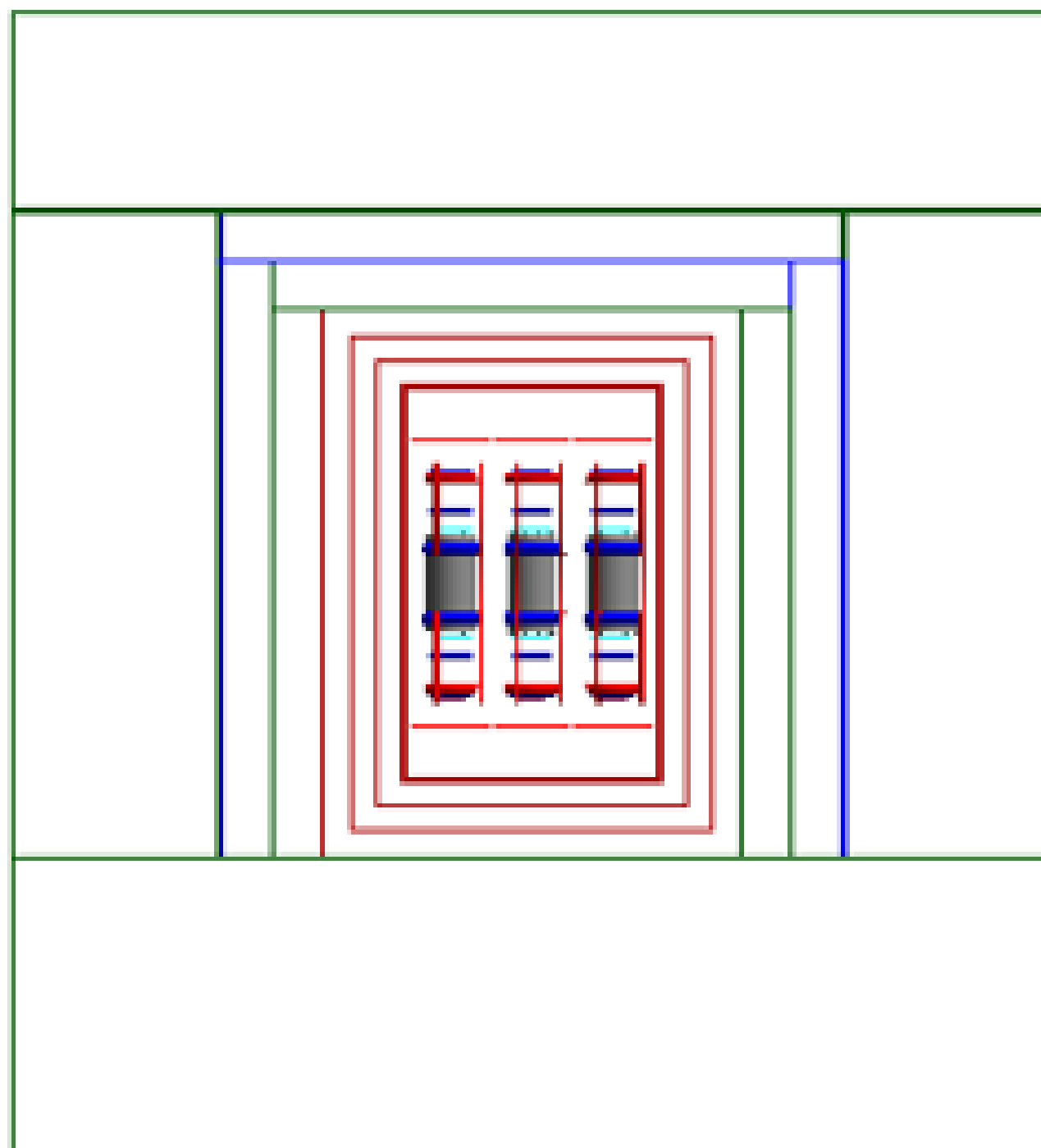
It consists of a two storeys building:
Ground floor (PT): set-up SABRE NORTH
First floor (P1). DAQ & counting room

SABRE North Experimental Area



3x3 NaI matrix with
25 cm copper and
50 cm polyethylene

Background contributions from Monte Carlo simulations



Source	Contribution in the ROI [1,6] keV units dru
Internal: NaI + PMTs + PTFE	0.5
Enclosure: Copper + Delrin parts	0.032
Shielding: Inner copper + Outer copper (negligible) + PE (negligible)	0.01
External gammas + neutrons (negligible)	0.001



- ❖ Conceptual design report presented in July 2021
- ❖ SABRE North TDR presented in June 2024 for the physic phase detector
- ❖ 3 x 3 matrix of crystals of ~5 kg mass each
- ❖ Fully passive shielding design: 25 cm copper + 50 cm PE
 - enough shielding power
 - negligible contribution to the total background
- ❖ Zone refining of Astro Grade has been tested across four runs with a reduction of some of the key background sources
- ❖ SABRE facilities are now installed in the final site at LNGS
- ❖ NaI-42 grown after zone refining will come to LNGS in 2024
- ❖ SABRE goal is to search for annual modulation with two similar NaI(Tl) detectors in the Northern and Southern Hemispheres
- ❖ SABRE expected to exclude/confirm annual modulation in 3-5 years of operation

THANKS FOR YOUR ATTENTION!



SABRE North



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Istituto Nazionale di Fisica Nucleare



PRINCETON
UNIVERSITY



UNIVERSITÀ
DEGLI STUDI
DI MILANO



SAPIENZA
UNIVERSITÀ DI ROMA

SABRE South



Australian
National
University



THE UNIVERSITY
of ADELAIDE



THE UNIVERSITY OF
SYDNEY



THE UNIVERSITY OF
MELBOURNE

BACKUP SLIDES



- ✓ Two passive shielding setups for crystal characterization
- ✓ A clean room with SABRE glovebox for crystal assembling



- 10 cm Cu + 15 cm of Pb
- Host 1 detector module
- Lexan box flushed with N₂



- 30 cm Cu shielding
- Host 1-3 detector modules
- Flushed with N₂



SABRE glovebox for the handling of hygroscopic NaI(Tl) crystals.