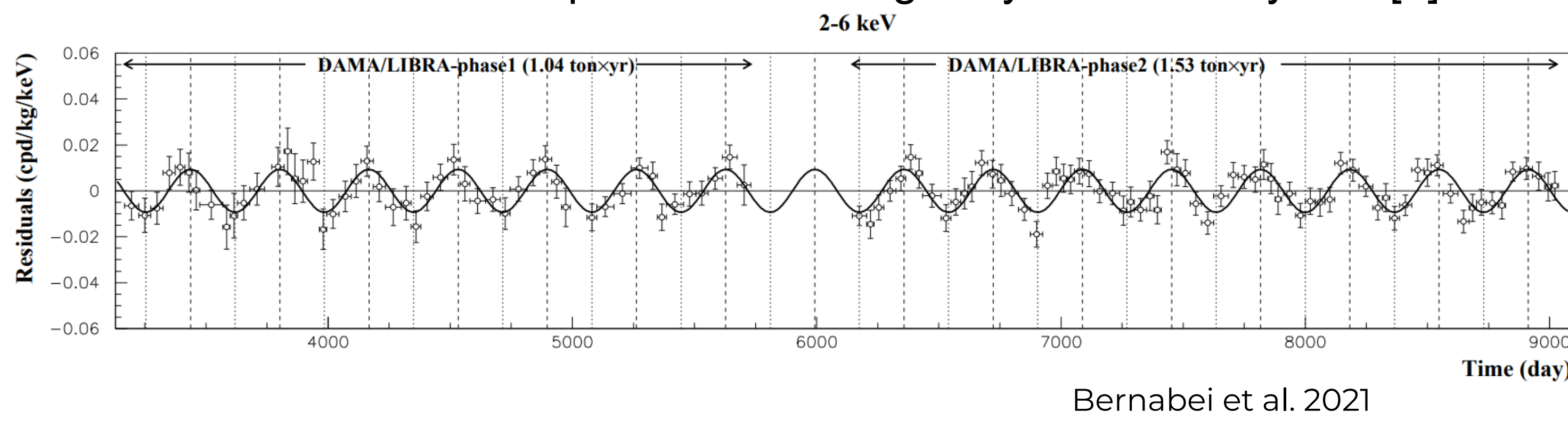




SABRE MOTIVATION

Astronomical observations provide evidence for the existence of dark matter
→ **85% of the mass density of the Universe.**

The **DAMA/LIBRA experiment** has been observing a modulating signal consistent with dark matter presence in our galaxy for over 20 years [1]:



However, no other dark matter experiment could confirm their observation.

Is it dark matter signal or seasonal modulation?

The **SABRE experiment** [2] is designed to **test the annual rate modulation** from dark matter interactions observed by DAMA/LIBRA using the same detector material, NaI(Tl) crystals.

This requires for SABRE to be a **double-site experiment with two separate detectors in two underground laboratories on both hemispheres.**

→ **SABRE South** at SUPL in Australia and **SABRE North** at LNGS in Italy.
SABRE South talk: Lachlan Milligan (Thu 15:06), SABRE North talk: Gabriella Cataldi (Thu 15:26)

SABRE detectors are an array of **ultra-high purity NaI(Tl) scintillating crystals**, aiming to **directly detect dark matter** through scattering off target nuclei.

SABRE South is located 1025 km underground in the **Stawell Underground Physics Laboratory** (cosmic rays are shielded down to a flux of $< 10^{-8} \text{ cm}^{-2} \text{ s}^{-1}$).

BACKGROUND PROJECTIONS

Monte Carlo simulations predict the background with these components [2]:

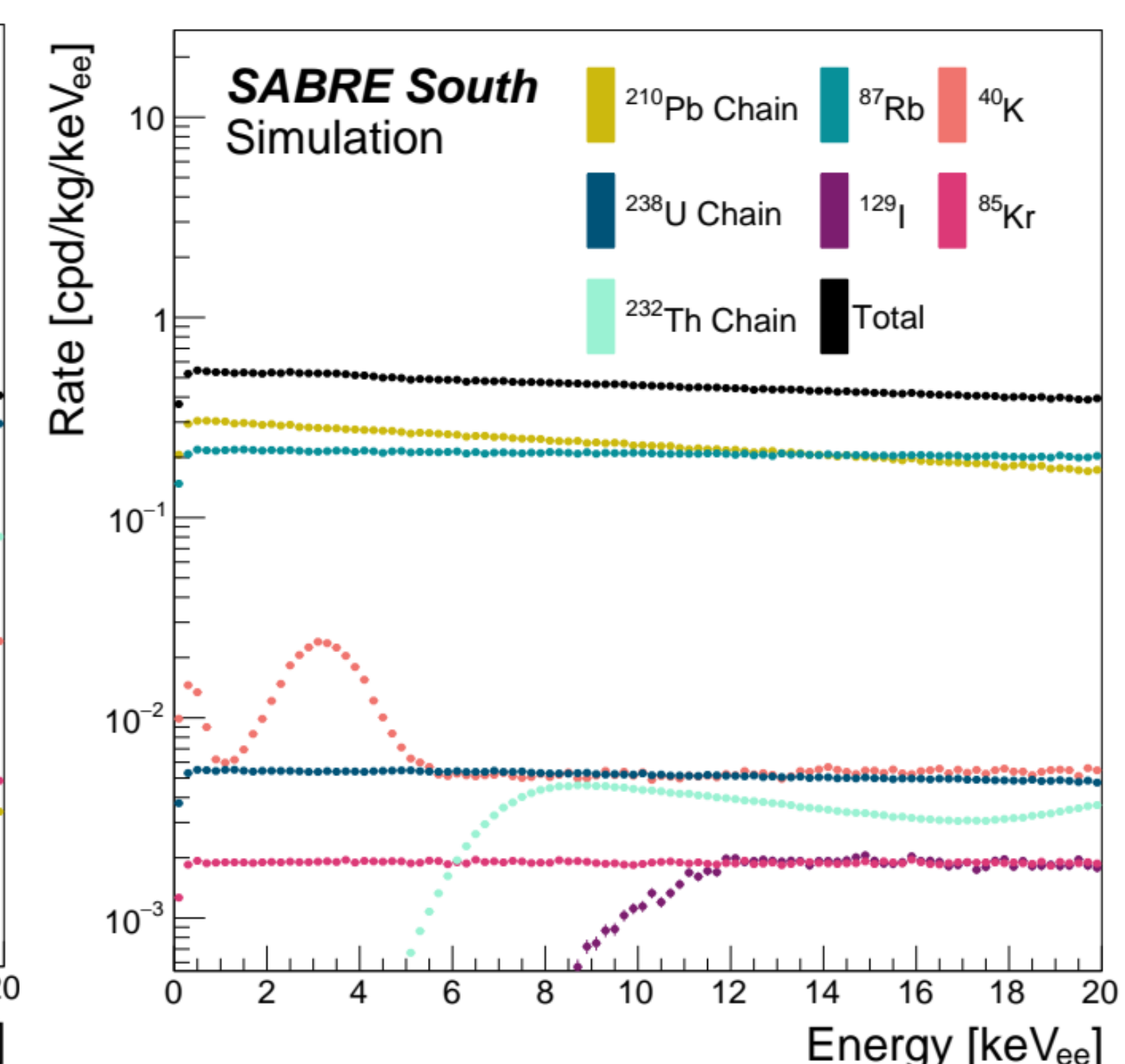
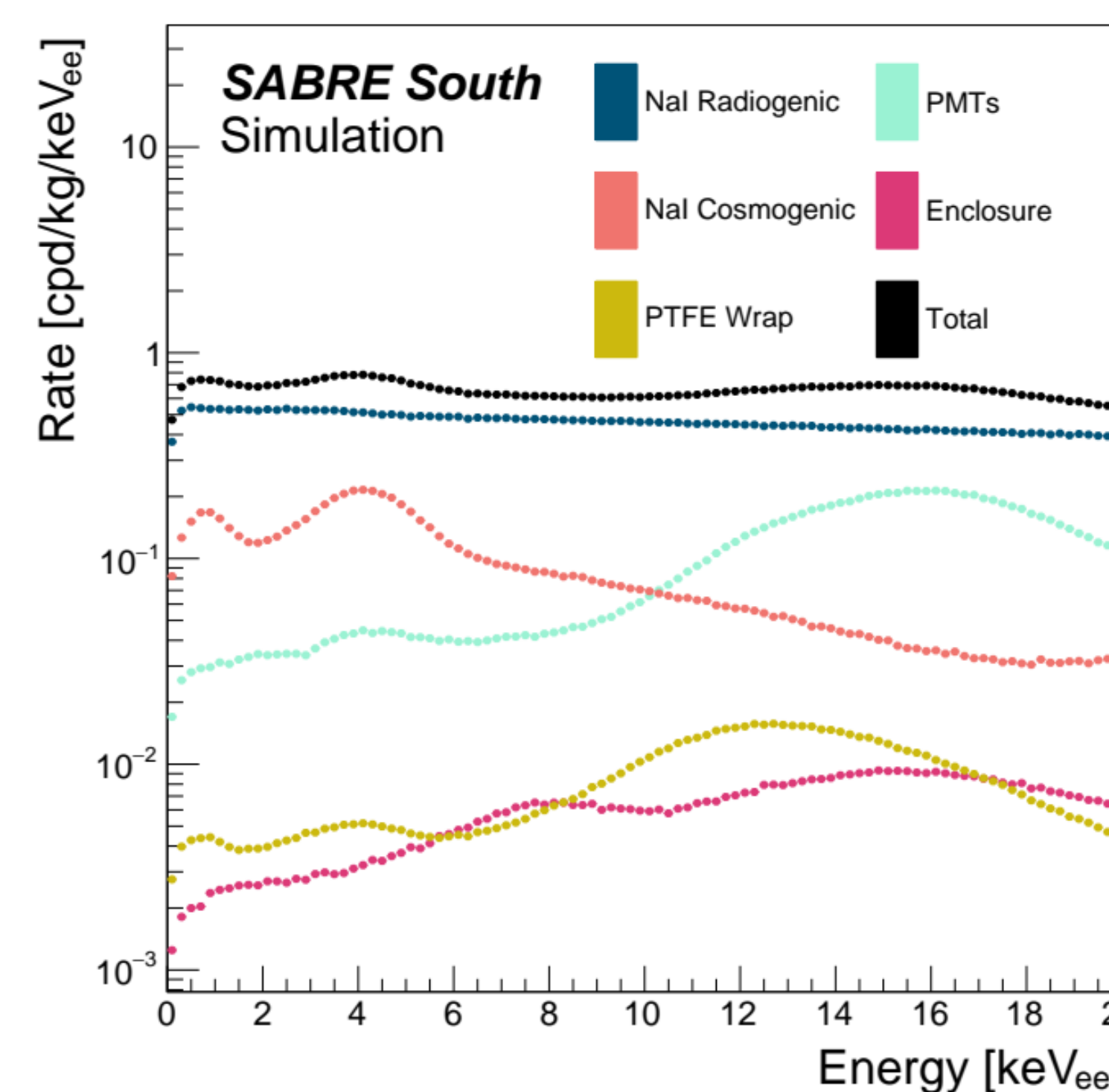
- **crystal intrinsic background** – isotopes naturally occurring in the NaI(Tl),
- **cosmogenic background** – through exposure to cosmic rays,
- **background from material radioactivity** - PMTs, crystal enclosure, PTFE.

The average background in the 1-6 keV ROI for SABRE is **0.72 counts /day /kg /keV_{ee}**.

The activity of **radiogenic isotopes** in SABRE crystals is shown in the table on the right.

²¹⁰Pb and ⁴⁰K are the key backgrounds.

Isotope	Activity [mBq/kg]
⁴⁰ K	$1.4 \cdot 10^{-1}$
²³⁸ U	$< 5.9 \cdot 10^{-3}$
²³² Th	$< 1.6 \cdot 10^{-3}$
⁸⁷ Rb	$< 3.1 \cdot 10^{-1}$
²¹⁰ Pb	$4.1 \cdot 10^{-1}$
⁸⁵ Kr	$< 1.0 \cdot 10^{-2}$
¹²⁹ I	1.3



RADIO-IMPURITY STUDIES

Low background sets a fundamental **limit to SABRE sensitivity**, radio-impurities need to be **identified, screened, quantified and reduced.**

⁴⁰K – half-life of 1.25 Ga, primordial origin

SABRE South crystals are embedded in 10 tonnes of liquid veto for active background rejection → **⁴⁰K is efficiently suppressed by the veto** down to $1.3 \cdot 10^{-2} \text{ cpd /kg /keV}_{ee}$. However, we aim to quantify the ⁴⁰K concentration.

⁴⁰K QUANTIFICATION:

Due to the low abundance of: ⁴⁰K (0.01%), we measure the highly abundant ³⁹K (93%) and use the well-known ⁴⁰K/³⁹K ratio to get the ⁴⁰K concentration.

The required **⁴⁰K concentration** in the SABRE South crystals is **< 10 ppb**.

MEASUREMENT TECHNIQUES to characterize ultra-low ⁴⁰K :

- 1) **Inductively-Coupled Plasma Mass Spectrometry** - in this analytical technique, plasma is used to ionize and detect elements in a dissolved diluted sample.
- 2) **Super-Secondary Ion Mass Spectrometry** [3] - using solid samples allows two orders of magnitude lower sensitivity requirement as no dilution needed. Campaign started at HZDR in Germany.

We can measure potassium concentrations down to a few ppb, reaching the required level of < 10 ppb.
We are developing and improving the measurement techniques.

²¹⁰Pb – half-life of 22.2 a, naturally in environment

²¹⁰Pb is the **highest radiogenic contaminator** in the crystals.

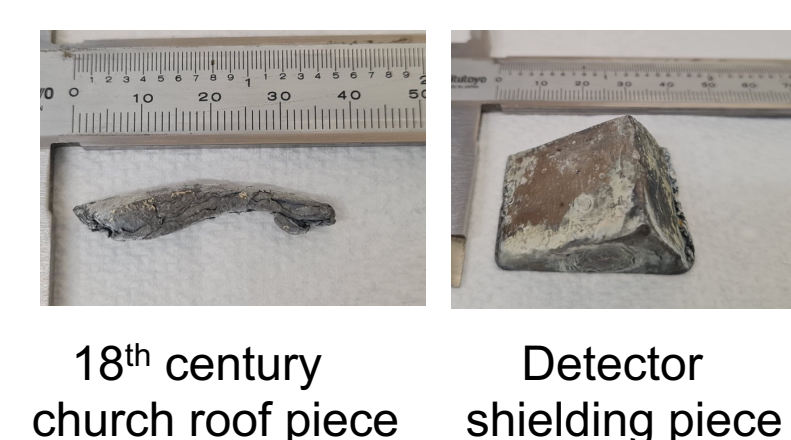
²¹⁰Pb atoms can be precisely counted via **accelerator mass spectrometry (AMS)** as an isotopic ratio of ²¹⁰Pb/²⁰⁸Pb.

The goal is a detection sensitivity in ultra-pure NaI(Tl) powder at a **²¹⁰Pb specific activity of 30 μBq /kg NaI**.

The AMS system has a high selectivity and background suppression capability as it suppresses molecular background by electron stripping in the high voltage terminal of a tandem accelerator.

CHALLENGE:

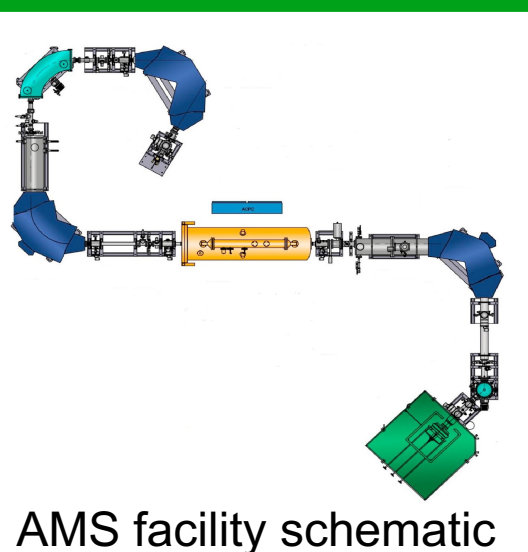
There is not enough lead to produce an AMS sample after extraction from NaI(Tl) → we search for an **optimal carrier with as low ²¹⁰Pb content as possible.**



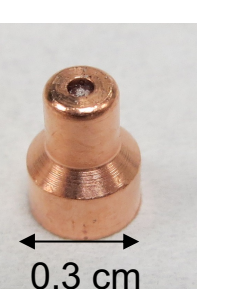
18th century church roof piece Detector shielding piece

Several aged materials (centuries old) were examined. We found a material with a **²¹⁰Pb/²⁰⁸Pb ratio in the order of 10⁻¹⁶ – We found a suitable carrier!**

We can measure ²¹⁰Pb/²⁰⁸Pb down to the required 10⁻¹⁶, two orders of magnitude lower than previously reported [4].



AMS facility schematic



AMS sample holder

TAKEAWAY MESSAGES

SABRE aims to test the annual rate modulation measured by DAMA/LIBRA.

Development of radio-impurity quantification methods allowed us to **improve the radio-impurity quantification by two orders of magnitude.**

We can measure **potassium concentrations down to a few ppb** and the **²¹⁰Pb/²⁰⁸Pb ratio down to 10⁻¹⁶.**

REFERENCES and ACRONYMS

- [1] R. Bernabei et al., Eur. Phys. J. C 67 (2010): 39-49
- [2] E. Barberio et al., Eur. Phys. J. C 83.9 (2023): 878
- [3] G. Rugel et al., Nucl. Instrum. Meth. B 532 (2022): 52-57
- [4] M. B. Froehlich et al., Nucl. Instrum. Meth. B 529 (2022): 18-23

SABRE = Sodium Iodide with Active Background REjection

DAMA/LIBRA = DArk MATter Large sodium Bulk RAre processes